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#### NASA/TM-2003-212236, Vol. 7



**Topography Experiment (TOPEX)** Software Document Series

# Volume 7

TOPEX Mission Radar Altimeter Engineering Assessment Report

## February 1994

D.W. Hancock III G.S. Hayne C.L. Purdy R.L. Brooks

TOPEX Contact: D.W. Hancock III

September 2003

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# Volume 7

# TOPEX Mission Radar Altimeter Engineering Assessment Report

### February 1994

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September 2003

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The TOPEX Radar Altimeter Technical Memorandum Series is a collection of performance assessment documents produced by the NASA Goddard Space Flight Center Wallops Flight Facility over a period starting before the TOPEX launch in 1992 and continuing over greater than the 10 year TOPEX lifetime. Because of the mission's success over this long period and because the data are being used internationally to redefine many aspects of ocean knowledge, it is important to make a permanent record of the TOPEX radar altimeter performance assessments which were originally provided to the TOPEX project in a series of internal reports over the life of the mission. The original reports are being printed in this series without change in order to make the information more publicly available as the original investigators become less available to explain the altimeter operation and details of the various data anomalies that have been resolved.

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#### ACKNOWLEDGMENTS

The authors gratefully acknowledge the contributions of the members of the Wallops TOPEX Software Development Team and TOPEX Hardware Development Team:

Ronald Forsythe (NASA) Hayden Gordon (CSC) Jeff Lee (CSC) Dennis Lockwood (CSC) Carol Purdy (CSC) Rob Ryan (CSC) Bill Shoemaker (SMSRC)

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#### 1.0 INTRODUCTION

The primary objective of the NASA Ocean Topography Experiment (TOPEX) project is to develop, launch, and operate a satellite-borne dual-frequency radar altimeter; the altimeter is to provide ocean topographic range measurements with a nominal precision of ±2.4cm and an overall accuracy of ±14cm. The Johns Hopkins University Applied Physics Laboratory (JHU/APL) developed the altimeter hardware for NASA's Goddard Space Flight Center/Wallops Flight Facility (GSFC/WFF). The TOPEX altimeter was integrated into the TOPEX/POSEIDON spacecraft by Fairchild Space (FS) Company.

POSEIDON is a single-frequency solid-state altimeter developed by the Centre National d'Etudes Spatiales (CNES) of France. TOPEX and POSEIDON share a NASA-provided antenna on the spacecraft. The TOPEX/POSEIDON launch date was August 10, 1992.

This document describes the GSFC/WFF analysis of the on-orbit engineering data from the TOPEX radar altimeter, to establish altimeter performance. In accordance with Project guidelines, neither surface truth nor precision orbital data are used for the engineering assessment of the altimeter. The use of such data would imply not only a more intensive and complete performance evaluation, but also a calibration. Such evaluations and calibrations are outside the scope of this document and will be presented in a separate Verification Report.

TOPEX Mission Radar Altimeter Engineering Assessment Report

#### 2.0 RELATIONSHIPS TO OTHER DOCUMENTS

The TOPEX Project Plan (Reference 1) is the controlling document for Project and Mission goals, and specifies the methods to be used to achieve those goals. TOPEX Project Requirements and Constraints (Reference 2) and TOPEX Project Mission and Systems Requirements (Reference 3) are the controlling documents for requirements and constraints placed on the Project. TOPEX/POSEIDON Project Joint Verification Plan (Reference 4) is in consonance with these plans and is the controlling document for specifying verification activities and plans. References 5 through 28 are supporting documents for this verification effort.

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- (3) TOPEX Project Mission and Systems Requirements, April 1989, JPL D-5901, 633-103.
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#### 3.0 RADAR ALTIMETER DESCRIPTION

TOPEX is a dual-frequency radar altimeter with complete redundancy of all active circuitry. It is a nadir-looking radar which transmits RF energy towards the earth's surface, then receives and processes the reflected energy. It measures height above the earth's surface (pulse transmit time), ocean significant waveheight (via return pulse shape characteristics), and surface radar backscatter coefficient (via received energy).

The TOPEX radar altimeter design has evolved from a number of previous designs (for example, Seasat and Geosat), but also includes a second frequency to yield information on propagation delay due to the ionosphere. The two frequencies are Ku-Band (13.6 GHz) and C-Band (5.3 GHz); the height of the altimeter above the ocean surface will be measured at each frequency. The desired bandwidth for Ku-Band and for C-Band is 320 MHz. However, because of discrete geographic areas where 320 MHz may produce interference with ground-based C-Band systems, the contingency capability exists to change the C-Band bandwidth to 100 MHz for those geographic areas. The technical characteristics of the TOPEX radar altimeter are summarized in Table 3.0a, and a block diagram of the altimeter appears in Figure 3.0.

The TOPEX spacecraft has an onboard recording capability. Altimeter data are telemetered in two streams, the engineering stream and the science stream. The engineering stream includes engineering measurements, and is transmitted approximately every eight seconds. The science stream, which is transmitted at approximately one-second intervals, includes: Ku heights, C-minus-Ku height differences, significant waveheight voltage, waveform samples, and Automatic Gain Control values (signal strength measurement). The altimeter waveforms are processed to calculate significant wave height and provide corrected off-nadir angles. The Automatic Gain Control (AGC) is converted to sigma naught for windspeed estimation. Table 3.0b lists the availability of altimeter data streams for each altimeter mode.

Initiation of altimeter modes occurs by ground command, either real-time or delayed. The altimeter modes are defined in **Table 3.0c**.

Parameter	Ku/C
Mean Altitude (KM)	1336
Frequency (GHz)	13.6/5.3
Antenna Beamwidth (deg)	1.0/2.7
Peak RF Power (W)	20
Average RF Power (W)	9.2/2.3
Pulsewidth (uncompressed)	102.4 msec*
Pulsewidth (nsec) (compressed)	3.125*
Repetition Frequency (Hz)	4500/1100 **
Mission Duration (yrs)	3-5
Inclination (deg)	66.0
Repeat Cycle (days)	10
Beam Limited Footprint Diameter (KM)	24.3/63 ***
Pulse Limited Footprint Diameter (KM)	2.2 ***

\* Nominal pulsewidth

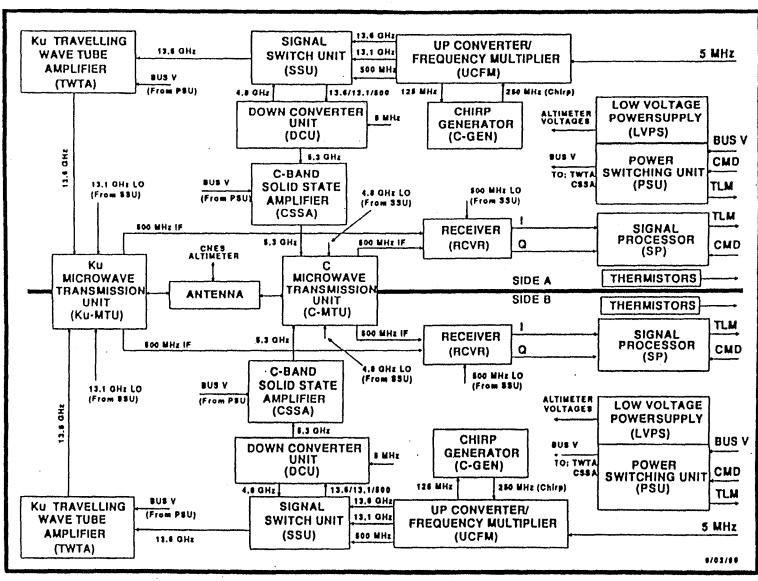
\*\* Varies with altitude.

\*\*\* 320 MHz, varies with pulsewidth and seastate.

#### Table 3.0aTOPEX Technical Characteristics

TOPEX Mission Radar Altimeter Engineering Assessment Report

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# TOPEX RADAR ALTIMETER - BLOCK DIAGRAM

Figure 3.0 Altimeter Block Diagram (from APL Document 7301-9028)

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Altimeter Mode	Telemetry Produced				
	Engineering	Science			
OFF	No	No			
IDLE	Yes	No			
STANDBY	Yes	Yes			
CALIBRATION	Yes	Yes			
TRACK	Yes	Yes			

Table 3.0b - Telemetry Produced in Each Altimeter Mode

8

Altimeter Mode	Definition
OFF 1	The spacecraft power system is <b>not</b> applying 28 volts to the radar altimeter. The spacecraft power bus relay to the altimeter is open.
OFF 2	The spacecraft power system is applying 28 volts to the radar altimeter, but all systems within the altimeter are in the OFF state. The internal altimeter relays are open. Power is supplied to the charging capacitors.
IDLE	The altimeter does not transmit. Primarily used when the CNES altimeter is on.
STANDBY	The altimeter does not transmit.
CALIBRATION - CAL-I	CAL-I is the first of two internal calibration modes. The transmitted pulse is fed back to the altimeter through a series of attenuators in 17 discrete steps. Provides monitoring of height bias, total loop gain characteristics, and waveform sample operations.
- CAL-II	This is the second of two internal calibration modes. CAL-II is a single-step process wherein the AGC operates on noise only. Provides receiver and waveform characteristics.
TRACK - COARSE ACQUISITION	The flight software searches for a return signal from the surface in low resolution (50 ns pulsewidth).
- COARSE TRACK	Surface return waveforms are tracked in coarse resolution.
- FINE ACQUISITION	After a signal is detected in coarse resolution, fine resolution acquisition (3.125 ns) begins. If the tracking performance in the Coarse Track mode is of sufficient quality, this step is omitted and the altimeter goes directly to Fine Track mode.
- FINE TRACK	Surface return waveforms are tracked in fine resolution.

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Table 3.0cDefinition of Altimeter Modes

#### 4.0 ON-ORBIT INSTRUMENT PERFORMANCE REQUIREMENTS

The source for the requirements that follow is the *TOPEX Project Radar Altimeter Development Requirements and Specification* document, August 25, 1988, WFF-672-85-004.

#### 4.1 Height Measurements

The term "height measurement", as used in this document, is synonymous with the term "range measurement." It refers to the fundamental process of the altimeter, the precise measurement of the distance between the sensor and the Earth's surface at nadir.

At an averaged output rate of height measurement over a minimum duration of one minute, the combined and single channel noise level of the Ku- and C-Band data shall be such that 68% of the height data will be within the specifications listed in **Tables 4.1a and 4.1b**.

H 1/3	Combined Ku/C Bands	Ku - Band 320 MHz BW	C-Band 320 MHz BW	C-Band 100 MHz BW
2m	2.4cm	2.0cm	3.1cm	6.3cm
4m	2.7cm	2.2cm	3.5cm	8.0cm
8m	3.2cm	2.6cm	4.3cm	8.4cm

# Table 4.1aHeight Measurement Noise Specifications for Three-Second<br/>Averaging Interval

H 1/3	Combined Ku/C Bands	Ku - Band 320 MHz BW	C-Band 320 MHz BW	C-Band 100 MHz BW
2m	4.1cm	3.5cm	5.4cm	10.9cm
4m	4.6cm	3.8cm	6.1cm	13.9cm
8m	5.5cm	4.5cm	7.4cm	14.5cm

# Table 4.1bHeight Measurement Noise Specifications for One-Second<br/>Averaging Interval

On-orbit measurement specifications assume a wave skewness value of 0.1, and a rainfall rate of less than 2 mm/hour.

#### 4.2 Height Measurement Drift Rate

After appropriate internal calibration, the maximum residual height drift rate shall be less than 2 cm per 10 days, during combined or single channel operation.

#### 4.3 Sea State (H1/3)

At an averaged output rate of one per second, the accuracy of H1/3 shall be within 10% or 50 cm, whichever value is larger, of the true sea state. This accuracy is to be attainable over a wave height range of 1 to 20 m, and pertains to both Ku- and C-Band measurements.

#### 4.4 Altitude Velocity and Acceleration

The altimeter is to provide a measurement of the height velocity for combined or single channel operation to  $\pm 1$  cm/s over the range of 0 to  $\pm 50$  m/s. The derived acceleration lag values, after ground-processing, are to have an accuracy of  $\leq 0.2$  cm under height acceleration conditions of  $\leq 1$  m/s/s for combined or single channel operation.

#### 4.5 Acquisition and Data Quality Settling Time

The altimeter shall provide data of the specified quality within 5 seconds of the initiation of acquisition over the ocean. This is applicable to combined or single channel operation for height accelerations of -1 to +1 m/s/s.

#### 4.6 Absolute Internal Calibration

An absolute internal calibration capability will be provided for both the Ku- and C-Band channels. This calibration will be performed twice per day, and will monitor:

- Changes in internal height bias to ±1.5 cm or less.
- Gate bias and gain to ±1% of maximum average value at nadir and zero H1/3.
- Transmit power to ±1 dB.
- Loop Gain Changes in Automatic Gain Control values related to Radio Frequency (RF) path loss, RF output power, receiver gain, or processor performance with precision and accuracy of ±0.25 dB.

#### 4.7 Radar Reflectivity (Backscatter Coefficient)

The ocean backscatter coefficient, sigma-zero, is to be determined within a measurement precision of  $\pm 0.25$  dB, with a measurement accuracy of  $\pm 1.0$  dB in both the Ku- and C-Band channels.

#### 4.8 Ancillary Altimeter System Specifications

#### 4.8.1 Time-Tag Precision and Accuracy

The Frequency Reference Unit (FRU) specifications are to provide a frequency stability of 1 part in  $7 \times 10^{11}$  in a 24-hour period, and an absolute frequency accuracy of 1 part in  $7 \times 10^8$  over a 3-year period, for the 5 MHz altimeter input.

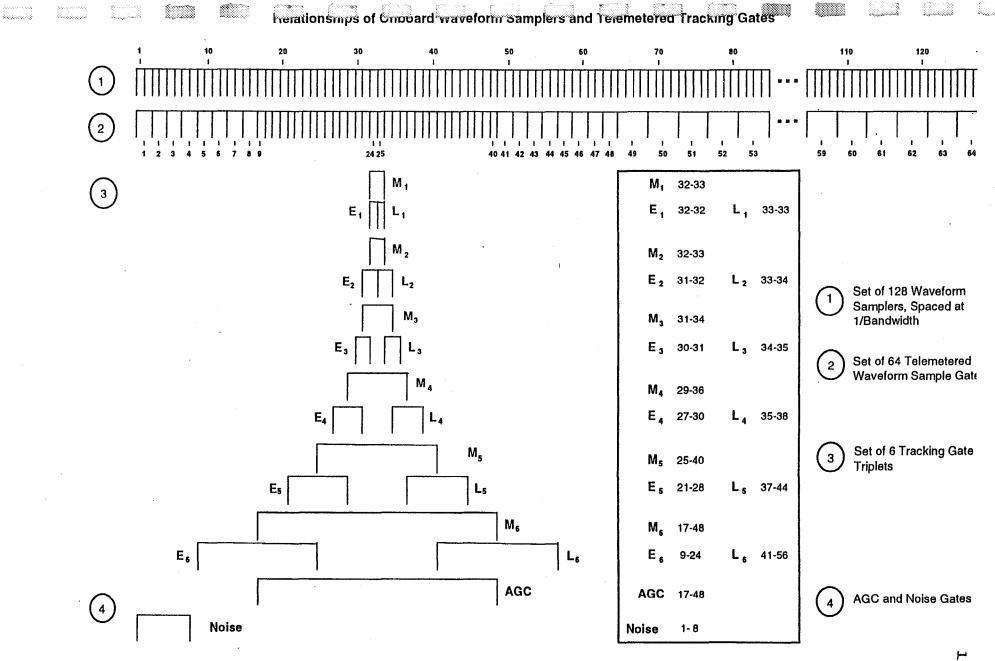
The maximum time-tag bias error within the altimeter will be  $\leq 10 \mu$ sec. The timetag bias is defined as the difference between the spacecraft time when a particular altitude measurement is performed by the altimeter and the time inserted in the telemetry data within the altimeter. Time-tag bias error in this assessment context does **not** include clock errors which may exist in the timing system of the spacecraft. The overall time-tag error, including spacecraft timing errors, is to be  $\leq 100 \mu$ sec.

#### 4.8.2 Waveform Samples

The altimeter will acquire 128 waveform samples of the ocean surface return in each of the two channels. Sixty-four amalgamations of the waveform samples, as shown in **Figure 4.8.2** (Item 2), will be telemetered in the data stream. The 64 samples are telemetered at a 10-per-frame rate (primary channel) and a 5-per-frame rate (secondary channel) with a selection available to designate which channel is to be high or low rate. Figure 4.8.2 pertains to Ku-Band waveforms; the C-Band waveform relationships are identical except that C-Band waveforms are shifted three gates to the right (later in time).

#### 4.8.3 Command Capability and Programmability

The altimeter will be commandable to various modes through the spacecraft interface via ground uplink. The ground uplink may also be used to reprogram the onboard altimeter operational programs.





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#### 5.0 ON-ORBIT INSTRUMENT PERFORMANCE ASSESSMENT

The TOPEX Altimeter (ALT) was initially turned on to IDLE mode, on August 21, 1992, eleven days after launch. It was switched to TRACK mode for the first time on-orbit on August 26, 1992. From initial turn-on to February 1, 1994, the ALT track status has been:

Mode	<u>Total Hours</u>
OFF	72
IDLE	1490
TRACK	11,130

The ALT is switched to IDLE whenever the CNES Altimeter is turned ON.

The performance assessment which follows is based solely on Side A. Side B of the altimeter has not been turned on since pre-launch testing. The Project plan is that Side B will not be turned on unless Side A fails.

The C bandwidth used throughout the performance assessment is 320 MHz. The 100 MHz bandwidth was very briefly tested after launch to affirm its operability.

The performance assessment shown below is keyed, by section numbering, to the performance requirements of Section 4.0.

5.1 Height Measurements

#### 5.1.1 Measurement Noise

The measurement noise of the ALT heights has been estimated by calculating the rms of one-per-frame Ku-minus-C height differences, after a linear fit over a 60-second interval (Hayne, 1993). The height rms for each of Ku, C, and combined Ku-and-C was then determined by scaling the rms-of-fit by the ratio of the Ku/C noise.

This method takes the effects of geoid noise out of the calculation that would be present on a straight fit to either frequency, since both frequencies follow the same geoid tracking.

The height measurement noise estimates resulting from this technique are listed in **Table 5.1.1a**, where the rms specification is for one-second averaging intervals as listed in Table 4.1b.

#### SIGNIFICANT WAVEHEIGHT

	2M		4M		8M	
	<u>SPEC</u>	<u>ON-ORBIT</u>	<u>SPEC</u>	<u>ON-ORBIT</u>	<u>SPEC</u>	<u>ON-ORBIT</u>
		Min. Max.		Min. Max.		Min.Max.
Combined Ku C320	4.1 3.5 5.4	1.4 - 2.9 1.2 - 2.4 1.8 - 3.7	4.6 3.8 6.1	2.2 - 3.3 1.8 - 2.7 2.9 - 4.4	5.5 4.5 7.4	3.7 - 4.5 3.1 - 3.8 5.1 - 6.3

#### Table 5.1.1a One-Second Range Measurement Noise Performance (in cm)

The pooled height measurement noise estimates from this technique are shown in **Table 5.1.1b**. Although this is a new noise-estimation technique, the results are in good agreement with those presented by Purdy and Marth at the February 1993 Verification Workshop; for that presentation, noise estimates were based on the use of repeating groundtracks to subtract the geoid signatures. These estimates are also similar to noise spectra computed by Richard Sailor, which were part of a presentation by Hayne, et al, at the same Workshop, and are generally within  $\pm 10\%$  of the estimates emanating from pre-launch thermal-vacuum testing. C100 range measurement noise estimates for 2 m and 4 m SWH are included in Table 5.1.1b, based on an analysis of the small amount of C100 on-orbit tracking data; no C100 data are available for 8 m SWH.

#### SIGNIFICANT WAVEHEIGHT

		2M		4M		8M
	<u>SPEC</u>	<u>ON-ORBIT</u>	<u>SPEC</u>	<u>ON-ORBIT</u>	<u>SPEC</u>	<u>ON-ORBIT</u>
Combined Ku C320 C100	4.1 3.5 5.4 10.9	2.1 1.7 2.6 4.1	4.6 3.8 6.1 13.9	2.9 2.4 3.8 6.5	5.5 4.5 7.4 14.5	4.0 3.4 5.6 n/a

Table 5.1.1b

#### One-Second Pooled Range Measurement Noise Performance (in cm)

#### 5.1.2 Ku/C Height Differences

The ionosphere introduces an error in height because of increased propagation delay; the propagation delay is proportional to the electron constant of the ionosphere and is also related to the frequency of the radar pulse. The C-Band is affected more by the ionosphere than Ku-Band and therefore C heights are generally greater than Ku heights. The height differences are much larger in daylight hours when the total electron content (TEC) along the radar propagation path is larger. Ionospheric effects are also latitude-dependent.

The Ku/C height differences provide a means of determining the TEC, and thus provide a technique for calculating the corrections for height errors attributed to the ionosphere. The Ku height corrections are equal to the Ku-minus-C heights multiplied by 0.179065468, per TOPEX Algorithm G1043: Combined Height/Ionospheric Correction.

An analysis of calculated ionospheric height corrections, performed shortly after launch, revealed that the altimeter-derived corrections were generally +1.8 cm during nighttime passes when the corrections should have been near-zero, and should have been about 1.8 cm more negative during daytime passes. A +100 mm offset was added to the C-Band heights to rectify this observed effect. The effective date of this C-Band offset was day 238 of 1992, at 23:12:00 UTC.

Two typical plots of height corrections based on ionospheric effects, subsequent to the implementation of the +100 mm C-Band height offset, are illustrated in Figure 5.1.2. The top plot in Figure 5.1.2 is for a daytime pass; the shape of the corrections appropriately contains the classic ionospheric maxima at about  $\pm 15$  degrees latitude, and the minima are at the high latitudes. The bottom plot in Figure 5.1.2 is for a nighttime pass, with appropriately smaller ionospheric effects.

#### 5.2 Height Measurement Drift Rate

The ALT has an internal Calibration Mode, which is routinely initiated twice per day, at approximately 12-hour intervals. A review of the internal height calibration results, discussed later in Section 5.6, indicates that there has not been any discernible height measurement drift with respect to the specified interval of 10 days.

#### 5.3 Sea State (H1/3)

The ALT determination of H 1/3, significant waveheight, appears to be quite good; it is in good agreement with H 1/3 derived from waveform fitting. A representative histogram of Ku significant wave height (SWH) for a full 10-day cycle is shown in **Figure 5.3a**; the distribution of measured wave heights is comparable with those of the earlier Geosat mission. The C SWH is typically 5 cm higher than the Ku SWH.

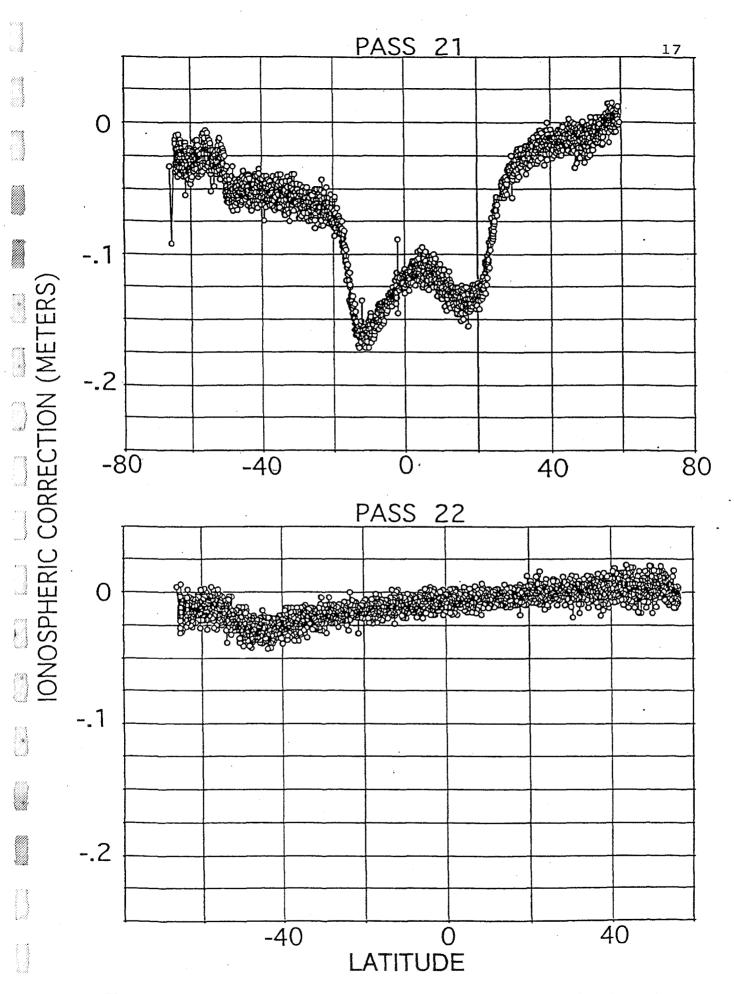
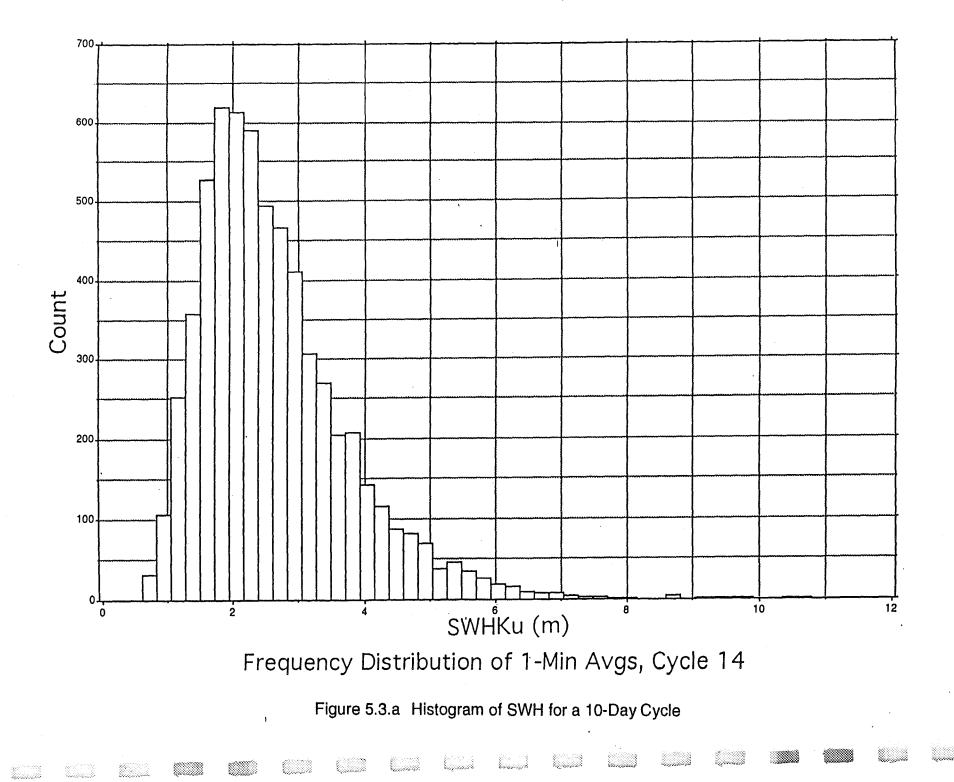


Figure 5.1.2 Calculated ionospheric Corrections for a Daytime Pass (top) and a Nighttime Pass (bottom)



Morris and Callahan of JPL have compared ALT-derived significant wave heights with wave heights independently measured by a NOAA buoy and by the Colorado University tide gauge, all in the vicinity of the Harvest platform. Their results, based on analysis of Cycles 1-10, were presented at the Verification Workshop in February 1993, and are shown here as Figure 5.3b. The agreement is excellent, with the ALT measurement generally about 10 cm high, but well within the specification of 50 cm.

#### 5.4 Altitude Velocity and Acceleration

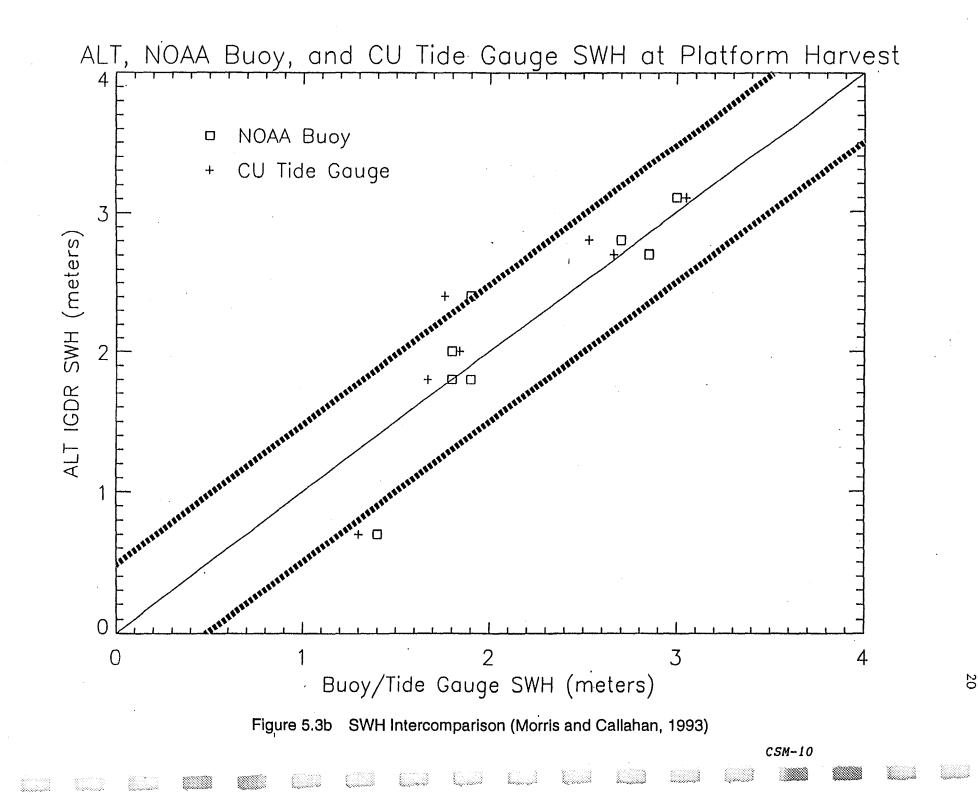
The achieved orbit for the TOPEX/POSEIDON spacecraft is very circular (mean eccentricity <0.0001); therefore, the orbital parameters contribute very little to altitude velocity or altitude acceleration. The altitude velocities as measured by the altimeter are, for the most part, due to the oblateness of the earth. The altitude accelerations, computed on the ground, are primarily due to localized changes in geoid slopes.

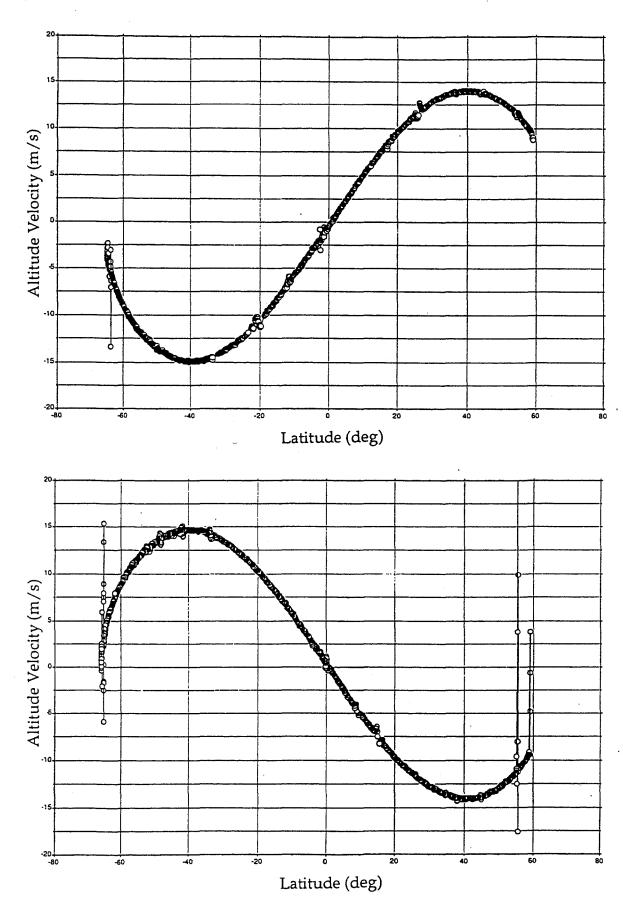
Typical measured altitude velocities, in m/s, for an ascending groundtrack (pass 21) and for a descending groundtrack (pass 22) are plotted vs. latitude in Figure 5.4a. The velocity is zero at the equator where the earth flattening effect changes sign, and approaches zero at the extreme orbital latitude of +66 degrees where the orbit levels out with respect to latitude. The maximum rates are at +40 degrees latitude where the earth flattening rate-of-change is the greatest. The absolute altitude velocity is slightly higher at -40 degrees than at +40 degrees latitude due to the different flattening characteristics between the southern and northern hemispheres.

Typical altitude accelerations, in m/s/s, are shown in Figure 5.4b for the same two passes as the previous Figure. The accelerations were computed by first-differencing the velocities of Figure 5.4a. The accelerations are observed to have a mean of zero, are generally banded within +0.3 m/s/s, and tend to be more variable near latitude -50 degrees. Studies by Hancock <u>et al</u>. (1990), using Geosat measurements, concluded that the effects of tracker lag on the altimeter range (dR), in meters, due to altitude acceleration (a), in m/s/s, are correctable by dR=at<sup>2</sup>/B, where t is the tracker update interval in seconds, and B is the tracker parameter Beta (typically B = 1/64). Algorithm S1039: Acceleration. Based on simulations, the calculated acceleration corrections have an accuracy of -0.05 <u>+</u> 0.04 cm.

#### 5.5 Acquisition and Data Quality Settling Time

During the first week after the on-orbit ALT was initially commanded to TRACK mode, acquisition tests were performed. The altimeter was commanded into 15 acquisition test sequences for each of seven different parameter sets. These parameter sets varied acquisition bandwidth, sweep range and threshold levels.





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Section 3

Figure 5.4a Altitude Velocities for Typical Ascending Groundtrack (Pass 21, top) and Typical Descending Groundtrack (Pass 22, bottom). In the Bottom Plots, the Groundtrack Moves from Right-to-Left.

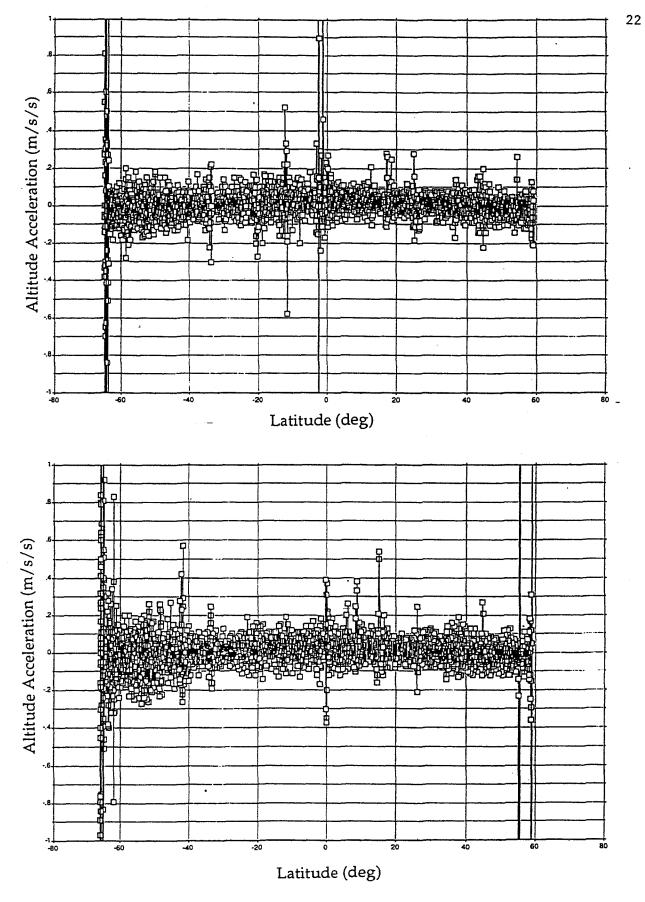


Figure 5.4b - Altitude Accelerations for Typical Ascending Groundtrack (Pass 21, top) and Typical Descending Groundtrack (Pass 22, bottom).

The results of the acquisition analyses were as follows:

- 1. For a cold start (no lock-on), the acquisition and data quality settling time varies from 3-7 seconds, depending on the particular parameter file configuration.
- 2. For a cold start, for the parameter file which is operationally in use, the acquisition and data quality settling time varies from 4-5 seconds.
- 3. Generally, the 20 MHz (50 ns equivalent pulse width) acquisition bandwidth performed better than the 5 MHz (200 ns equivalent pulse width) acquisition.
- 4. For the situation where the altimeter groundtrack goes from land to water, the acquisition and data quality setting time is usually 0-1.5 seconds depending on the land characteristics.
- 5. For many of the analyzed land-to-water transitions, the altimeter was already in FINE TRACK over the land areas.

The TOPEX altimeter uses an acquisition scheme which allows it to degrade from high resolution tracking to coarser tracking over land or ice without completely losing lock. This allows the altimeter to bypass much of the acquisition time when coming off land or ice.

#### 5.6 Absolute Internal Calibration

The altimeter has an internal calibration which provides for in-flight self-calibration. Currently it is scheduled two times per day at approximately 12-hour intervals (over land).

The calibrate command causes the altimeter to exercise two distinct calibration modes. In the first, CAL1, the transmitted pulse is fed back through digital attenuators with 17 sequential steps; the duration of each step is approximately ten seconds. The pulse is tracked during each of the steps. When all CAL1 attenuator steps are completed, the flight processor automatically changes to CAL2. CAL2 operates for approximately 60 seconds. Of the seventeen steps of CAL1, it has been observed that the initial seven steps have a larger signal level so are more reliable for calibration than the later steps.

The CAL1 mode allows for monitoring changes in height bias, changes in system loop gain and waveform sample operation.

In CAL2, the transmit pulse is not linked to the receiver, so it has a noise input only. The AGC is active and adjusts to the noise level. This allows for monitoring the receiver health and waveform sampler operation. The CAL1, step 5, Ku height and C height calibration results since launch are plotted in Figure 5.6a, along with the Ku MTU Cal Attenuator temperature. In all the plots of calibration results, the reference is the average of nine consecutive calibrations during cycle 001 (days 270-275 of 1992). The Ku height is observed to have decreased about 1.0 mm during these 530 days. The 5 mm height spikes in the Figure are assumed to be the effects of bit toggling when the altimeter was switched from IDLE Mode to TRACK Mode after an interval during which the French altimeter was turned on or when the spacecraft had a problem or did a maneuver. Normal stable operation shows none of these spikes. The C height has decreased about 5.0 mm since launch; the apparent recent increase in measurement noise is an artifact of bit toggling.

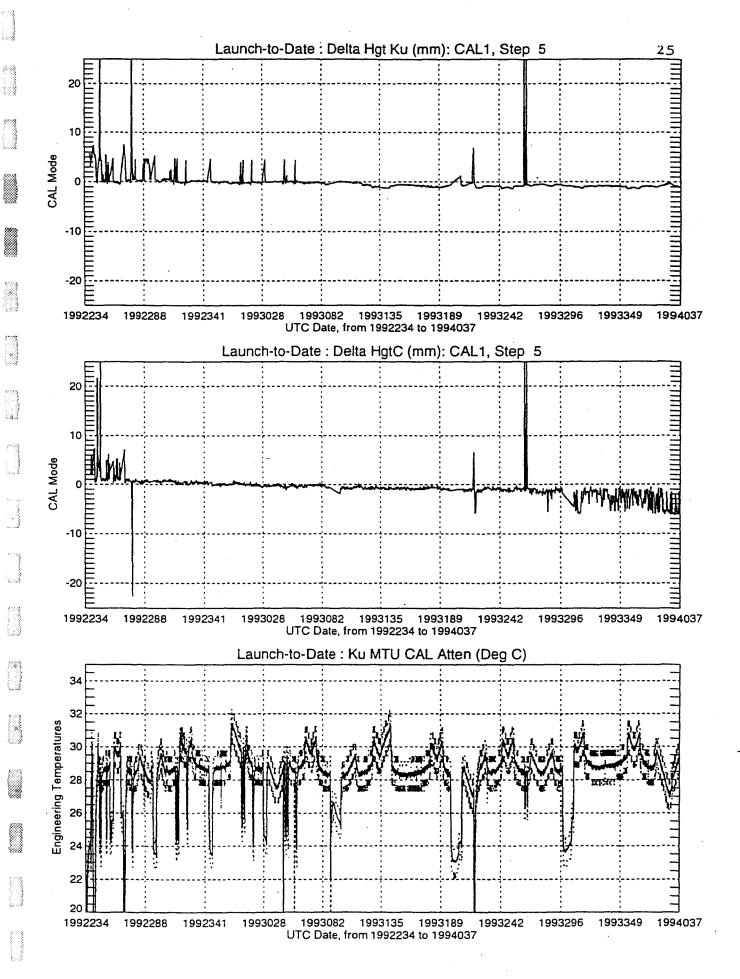
The CAL1, step 5, Ku AGC and C AGC calibration results since launch are plotted in **Figure 5.6b**, along with the AGC Receiver Section temperature. The Ku AGC is observed to have linearly decreased about 0.25 dB, from launch to day 242 of 1993; since that time, it has leveled off. The C AGC trend shows a linear decrease of about 0.10 dB since launch. At WFF's recommendation, and with the TOPEX Measurement Systems Engineer's concurrence, calibration-based AGC corrections have been entered into the production of GDR data at JPL. The AGC corrections and their start times are listed in Table 5.6a, where the values are to be added to the calculated sigma-naughts. The AGC corrections were not applied to the GDR data until the beginning of Cycle 48 on January 1, 1994.

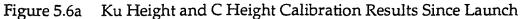
<u>Side A</u>	<u>Frequency</u>	<u>Year</u>	<u>Day-of-Year</u>	<u>Time</u>	<u>Correction</u>
А	Ku ,	1993	39	23:46	+0.10 dB
Α	Ku	1993	99	11:37	+0.15 dB
Α	Ku	1993	178	19:25	+0.25 dB
Α	С	1993	149	01:30	+0.10 dB

#### Table 5.6aAGC Corrections for GDR Production

The CAL2 Ku AGC and C AGC calibration plots are shown in **Figure 5.6**c. Early in the mission, Ku AGC increased about 0.1 dB; it is presently about 0.2 dB below its initial value. The C AGC also increased about 0.1 dB early in the mission; it has returned to its original level.

The AGC calibrations for Ku and C-Band, as shown in Figures 5.6b and 5.6c are not in agreement. Table 5.6b summarizes these two results and adds a third result wherein Ku and C-Band transmit power values, during periods of calibration, have been converted from watts to dB. The watts-to-dB calibration results for Ku- and C-Bands are shown in Figure 5.6d.





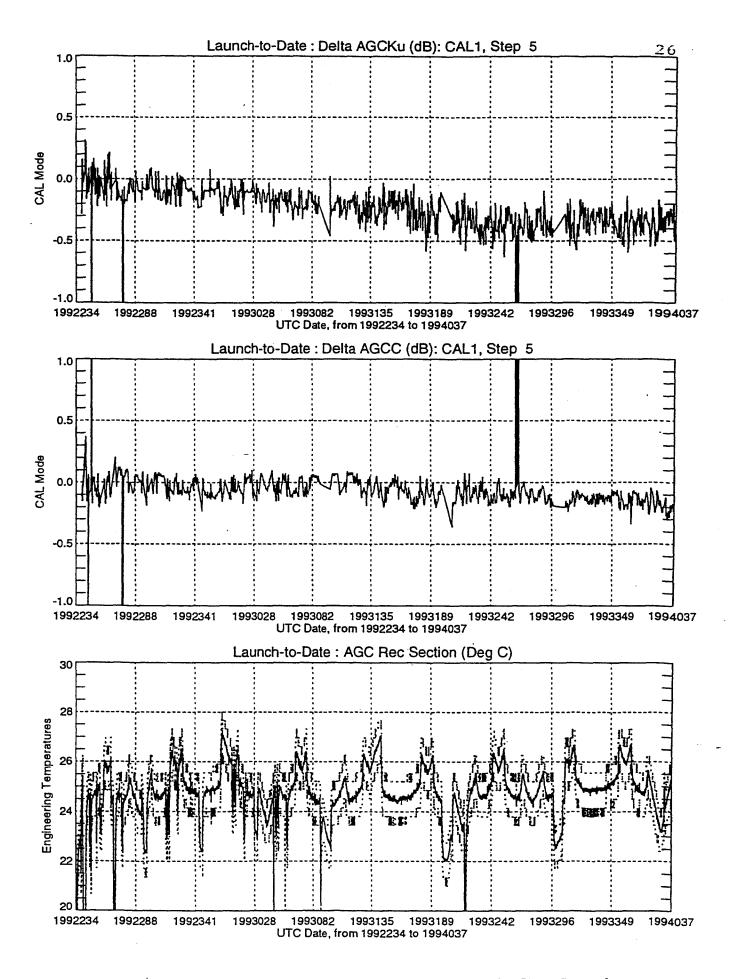


Figure 5.6b Ku AGC and C AGC Calibration Results Since Launch

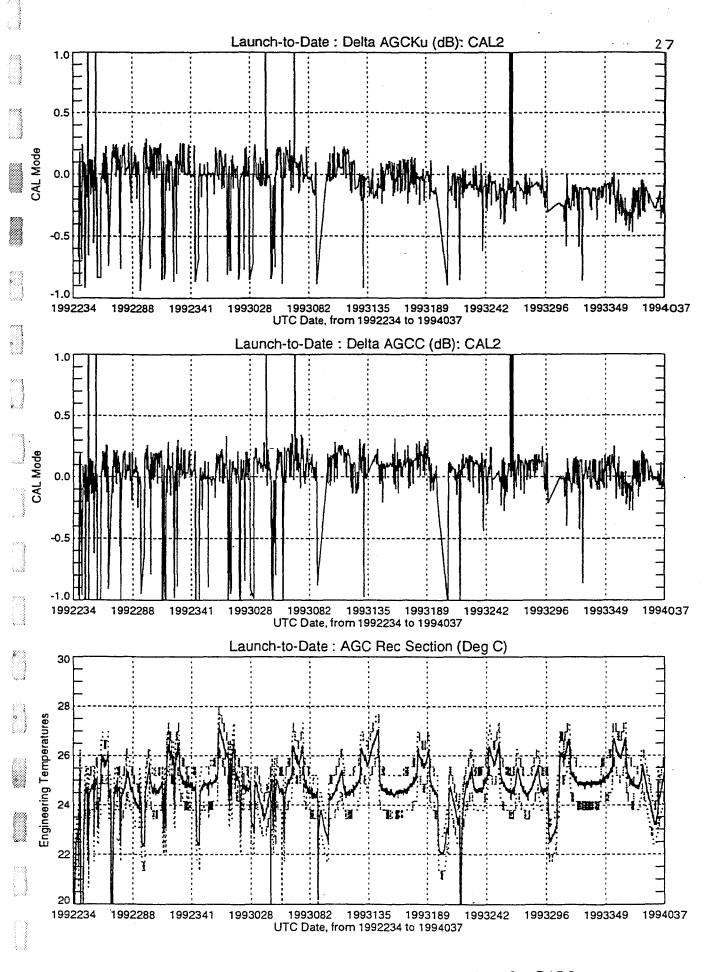


Figure 5.6c Ku and CAGC Calibration Plots for CAL2

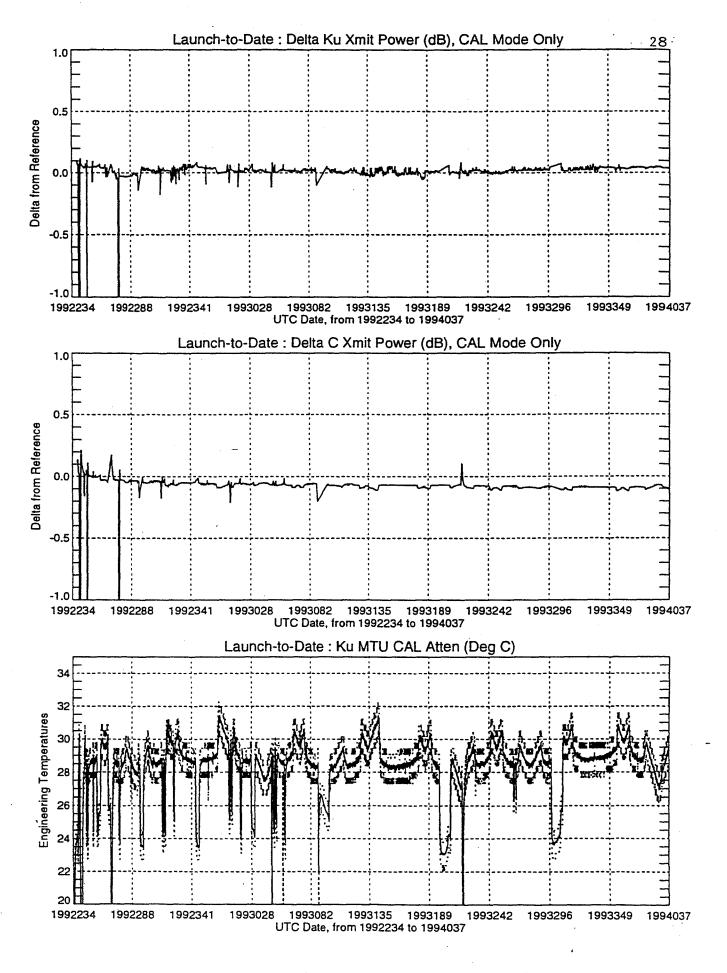


Figure 5.6d Ku and C Transmit Power During Calibration, Converted from Watts to dB

<u>Method</u>	<u>Ku-Band</u>	<u>C-Band</u>
CAL1	-0.25 dB	-0.10 dB
CAL2	-0.20 dB	0.00 dB
Watts-to-dB	-0.05 dB	-0.10 dB

Table 5.6bAGC Calibration Comparisons: Launch to Date

## 5.7 Radar Reflectivity (Backscatter Coefficient)

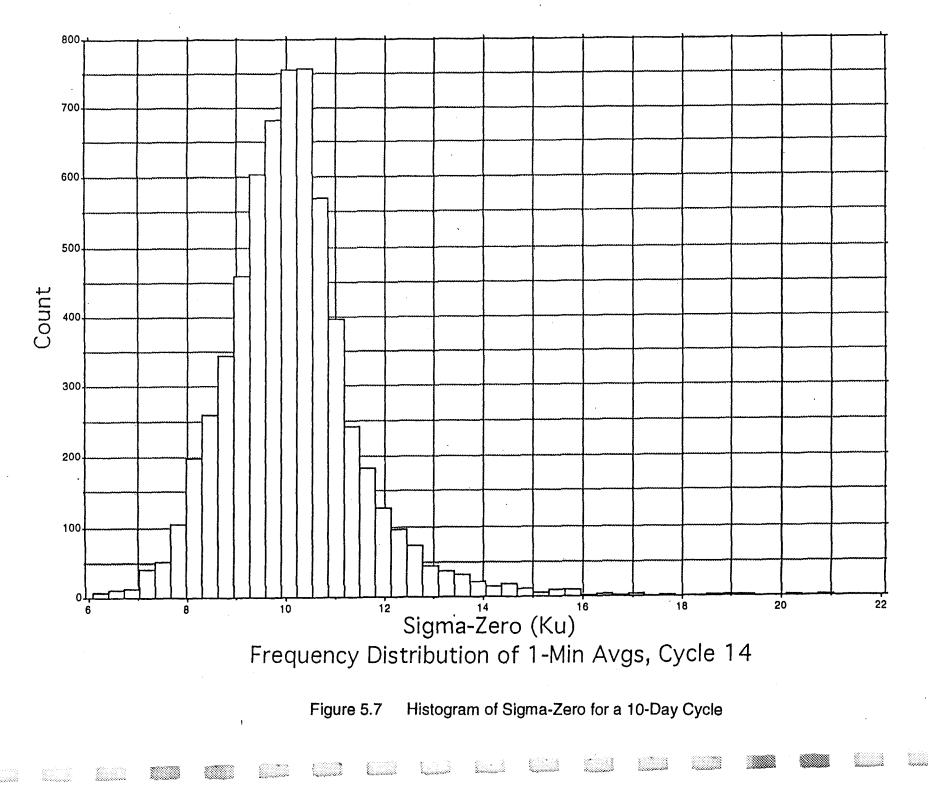
A representative distribution of Ku sigma zero, for a full cycle, is shown in **Figure 5.7**. While the ALT-derived histograms are similar in shape to those of Geosat, the ALT sigma zero values are shifted about +0.7 dB higher (Callahan, 1993, personal communication). In a June 1993 WFF informal memorandum, Hayne references Shoemaker who attributes 0.5 dB of this difference to the Geosat sigma-naught computations not taking into account the Earth's finite radius of curvature.

The histograms of C sigma zero have a narrower power distribution than the Ku, and the C sigma zero values are typically 3.4 dB higher.

Frequent occurrences of AGC "blooms" have been observed. Over areas of very calm seas, the waveforms become more specular, and the AGC increases by 5-8 dB. The duration of the blooms is typically on the order of a minute, which corresponds to about 350 km alongtrack. The effect of the specular waveforms on the ALT range tracker is to generally increase the noise of the range measurements during those intervals. A study of AGC blooms by Hayne et al (1993) concludes that:

- 1. The blooms are ocean surface effects, appearing in both Ku and C altimeter data.
- 2. Most of the affected altimeter measurements in the vicinity of the blooms are being appropriately flagged in the GDR processing.
- 3. Perhaps 5% of the over-ocean, normal-track data are in the vicinity of AGC blooms.

Hayne (1994) estimated the precision of one-per-frame GDR Ku sigma-naught measurements by fitting five short (70-160 seconds duration) segments to a quadratic. The rms of the residuals from the five fits varied from 0.065 to 0.091 dB. Hayne concludes that 0.08 dB is a conservative estimate of Ku sigma-naught measurement precision.



## 5.8 Ancillary Altimeter Systems Performance

#### 5.8.1 Time-Tag Accuracy

A plot of the Frequency Reference Unit (FRU) frequency stability is shown in **Figure 5.8.1**. This plot, from Cleven (1993), depicts the difference in milliHertz between the nominal frequency (5 MHz) and the measured frequency. The negative sign indicates that the measured frequency is less than the nominal. A line representing the per-day specification of 1 part in 7 x  $10^{11}$  is shown in the Figure. The absolute frequency accuracy specification of 1 part in 7 x  $10^{8}$  is being met since the measured frequency offset is well within ±350 milliHertz. Algorithm S1034: Oscillator Drift corrects the height measurements for the effects of FRU frequency drifts.

Shum <u>et al</u>, of the University of Texas, have made estimates of ALT time-tag accuracy determined from crossover analyses. The crossover solution computes the time shift which minimizes the crossover residuals. Their results, based on IGDR data from Cycles 1-12, were presented at the Verification Workshop in February 1993. Their preliminary conclusion, based on all 12 cycles, is that the overall TOPEX time-tag error is  $-0.1 \pm 0.2$  msec. The assessments of time-tag accuracy will become more refined as the crossover solutions use additional cycles of altimeter data, and as GDR data are utilized rather than IGDR data.

5.8.2 Waveform Samples

Waveform samples are discussed in Section 6.0.

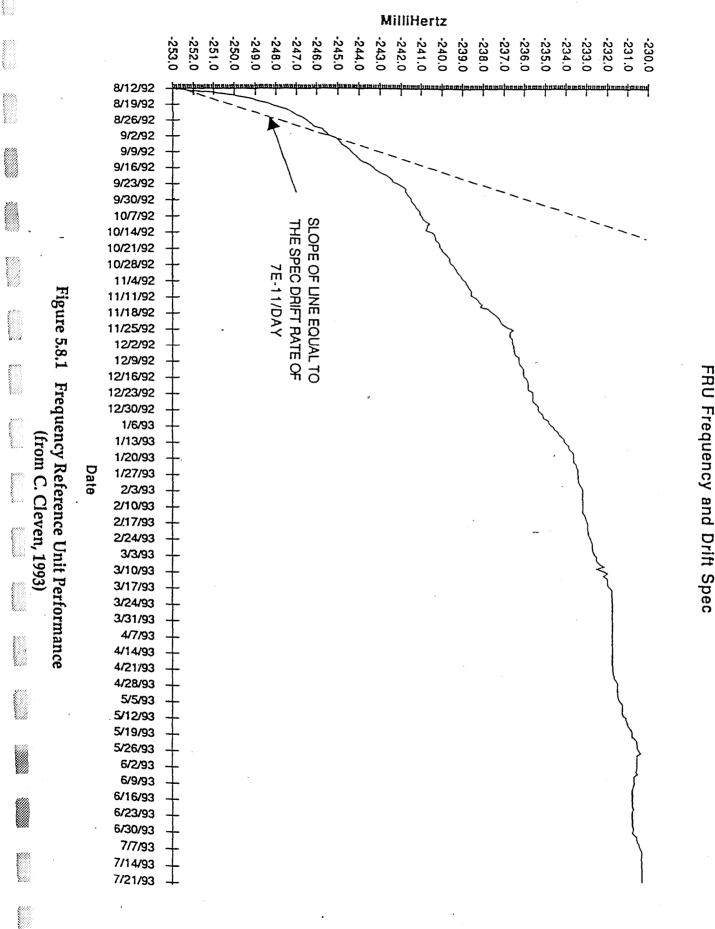
5.8.3 Single Event Upsets

As of February 1, 1994, there had been 59 Single-Event Upsets (SEUs). The altimeter automatically recovered from 50 of the SEUs, with the loss of a few seconds of data during each of those occurrences. The nine anomalous SEU occurrences were as follows:

• Day 247 of 1992 - The improper SEU recovery on this date was due to the

- corruption of the Pulse Count Variable. The altimeter was commanded to IDLE mode and then commanded to return to TRACK mode. Approximately seven hours of data were lost. A software patch to refresh the Pulse Counts was made on day 328 of 1992.
- Day 354 of 1992 On this date, the altimeter recovery from an SEU took 16-1/2 hours. At the end of that interval, another SEU apparently occurred and reset the processor.

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- Day 12 of 1993 The altimeter self-recovered from an SEU on this date, but 12 minutes of data were lost. The hypothesis for this recovery is that the SEU itself was not detected by the processor, but that the range eventually swept to the lower limit causing an internal reset.
- Day 230 of 1993 This upset was similar to that of day 12, but the altimeter self-recovery took 1-1/4 hours.
- Day 264 of 1993 Approximately 14.5 hours of data were lost when an interface lockup occurred, apparently due to an SEU. To reset the processor, the altimeter was commanded from TRACK to IDLE mode, a ground reset command was executed, and then the altimeter was commanded back to TRACK mode.
- Day 266 of 1993 Approximately 7.5 hours of data were lost when an interface lockup occurred, apparently due to an SEU. During this anomalous period, the altimeter generally remained in acquisition mode with some half-frame transfers to coarse track. To reset the processor, the altimeter was commanded from TRACK to IDLE mode, a ground reset command was executed, and then the altimeter was commanded back to TRACK mode.
- Day 307 of 1993 A science telemetry interface lockup was detected when the ALT was commanded to TRACK after the French Altimeter SSALT had been ON. The science word value remained at zero, even though the altimeter was tracking. The sequence to perform the ground reset was the same as for day 266. A total of 2-1/3 hours of data were lost.
- Day 330 of 1993 Approximately 8.2 hours of data were lost when an interface lockup occurred, apparently due to an SEU. During this period, the altimeter was tracking but did not transmit science data. To reset the processor, the altimeter was commanded from TRACK to SAFE IDLE mode, a ground reset command was executed, and then the altimeter was commanded back to TRACK mode.
- Day 001 of 1994 Approximately 3.7 hours of data were lost when an interface lockup occurred, apparently due to an SEU. During this anomalous period, the altimeter generally remained in TRACK mode, but science data were not being telemetered. An error reset command was executed to reset the processor.

The latitude/longitude locations of the 59 SEUs are shown in Figure 5.8.3. Almost all the SEUs have occurred within or near the outlined South Atlantic Anomaly, an area noted for its high solar activity. The circles in Figure 5.8.3 appear at the latitudes and longitudes where SEUs had automatic resets. The diamonds represent the locations of SEUs which required manual resets. The two diamonds at -90 degrees latitude are placed there because the locations of the SEUs could not be discerned.

The ALT is responding appropriately to all commands which it has received. Prelaunch, WFF developed 25 command blocks; three additional command blocks have been developed since launch. Most of the blocks have been transmitted to the altimeter at one time or another. The mode switching aspects of the 28 command blocks are depicted in **Figure 5.8.4**.

The programmability of the altimeter was demonstrated on November 23, 1993, when a software patch was successfully uploaded. The purpose of the patch is to refresh a pulse counter variable after a Single Event Upset.

The TOPEX onboard flight processor uses an uploaded programmable 274-byte parameter file. Revised parameter files have been sent to the altimeter for special testing and for performance improvement. The initial on-orbit parameter filename was PARMC320; the present parameter file, in operational use since early November 1992, is C35028SL. The contents of these two files are listed side-by-side in Table 5.8.4.

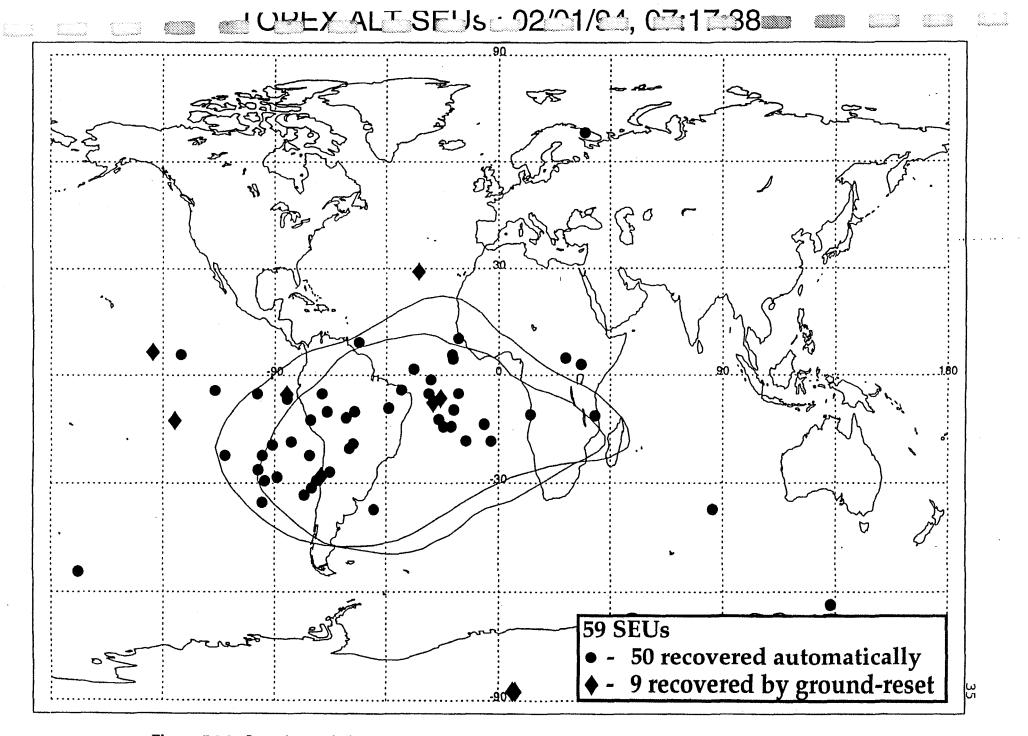
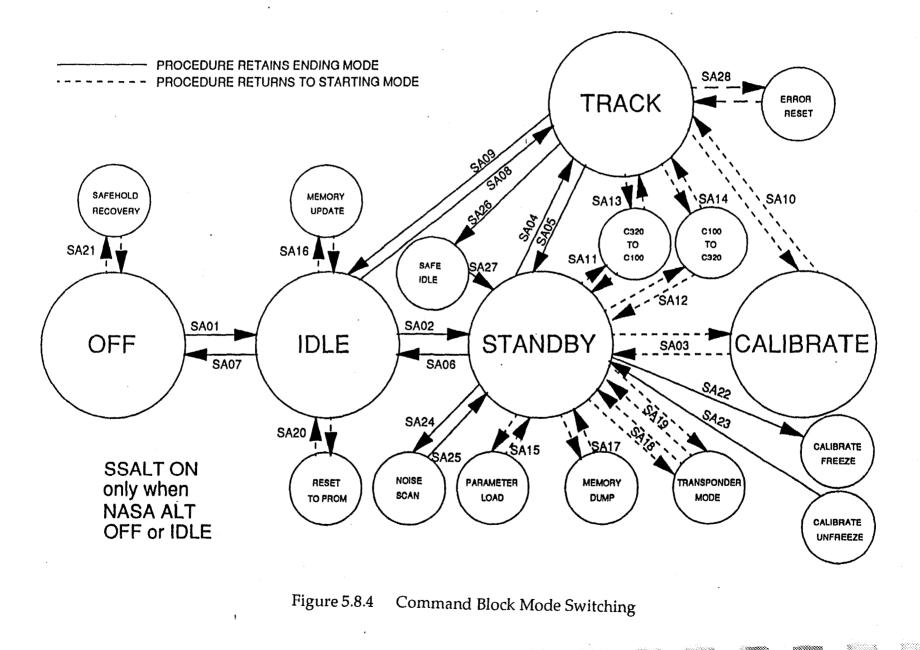


Figure 5.8.3 Locations of Single Event Upsets, in Relation to the South Atlantic Anomaly Area

# NASA ALTIMETER COMMAND PROFILE



	Filename = PARMC	Filename = C35028SL				
Byte	Parameter	Value	Hex	Value	Hex	Comment
	Iscan_Min_Hgt LSB/LSW	8.4992 mSec	00	8.4992 mSec	00	
	MSB/LSW		00		00	
2	LSB/MID		00		00	
3.	MSB/MID		00		00	· ·
-2	LSB/MSW		00		00	
-3-1	MSB/MSW		A6		A6	
	Iscan_Max_Hgt LSB/LSW	9.318 mSec	00	9.318 mSec	00	
	MSB/LSW		00		00	
-8-1	LSB/MID		00		00	
9	MSB/MID		00		00	
	LSB/MSW		00		00	
	MSW/MSW		<b>B6</b>		B6	
12	Iscan_Hgt_Inc LSB/LSW	200.0 nSec	00	200.0 nSec	00	
と渡	MSB/LSW		00		00	
-14-	LSB/MID		00		00	
15	MSB/MID		00		00	
	LSB/MSW	· ·	01		01	
	MSB/MSW		00		00	
18	Cal-I Index 1	77	4D	77	4D	
10,000	Cal-I Index 2	78	4E	78	4E	
	Cal-I Ku Min AGC Gate LSB	1024	00	1024	00	
21	MSB		04		04	
-?)	Cal-I C Min AGC Gate LSB	1024	00	1024	00	
3.	MSB		04		04	
<u>-</u> 24	CI_AGC_Threshold LSB	16384	00	16384	00	
25	MSB		40		40	
3	CI-AGC_Alpha	2	02	2	02	
27	CI_Track_Alpha	2	02	2	02	
28	CI_Ku_Hgt_Error_Scale	24	18	24	18	
3	CI_C_Hgt_Error_Scale	24	18	24	18	
30	CI_AGC_Error_Scale LSB	35	23	35	23	
	MSB		00		00	
3	AGC_Threshold . LSB	4096	00	4096	00	· · · · · · · · · · · · · · · · · · ·
33	MSB		10		10	
31   33   33    34	Low_Vres	6	06	4	04	Decrease Vres
3	Hgt_Adjustment	0	00	0	00	
30	AGC_Adjustment LSB/LSW	18 dB	00	10.5 dB	00	Decrease AGC Adjustment
3	MSB/LSW		00		00	
7 3 3 3 9	LSB/MSW		00		00	
39	MSB/MSW		48		2A	
4	AGC_Error_Scale LSB	139	8B	139	8B	
	MSB		00		00	
42	Ku_AGC_Gate_Scale	1	01	1	01	
4	C-100_AGC_Gate_Scale	32	20	32	20	
4	C_320_AGC_Gate-Scale	*4	04	4	04	
45	LRA_Min_Height LSB/LSW	1275 Km	00	1328 Km	00	Trim Range
4	MSB/LSW		00		00	L

Table 5.8.4Comparison of Initial (PARMC320) and Present (C35028SL)<br/>Programmable Parameter Sets

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		= <b>PARMC</b>			Filename =		والمراجع والمحاص والمتعالي والمتعالية والمتعار فالمتعار والمتعار والمتعالي والمتعالي والمتعار والمتعار والمتعار
Byte	Paramete		Value	Hex	Value	Hex	Comment
47		LSB/MID		14		00	
48		MSB/MID		1E		50	
49	· · · · · · · · · · · · · · · · · · ·	LSB/MSW		04		09	
50		MSB/MSW		A6		AD	l
51	LRA_Max_Height	LSB/LSW	1398 Km	00	1358 Km	00	Trim Range
52		MSB/LSW		00		00	
53		LSB/MID		FB		00	
54	}	MSB/MID		20		00	
55		LSB/MSW		08		F2	
56		MSB/MSW		<u>B6</u>		<u>B0</u>	
57	LRA_Height_Inc	LSB/LSW	3.24 Km	00	0.81 Km	00	Decrease Scan Step Size
58		MSB/LSW		00		00	· ·
59		LSB/MID		00		00	
60		MSB/MID		00		00	
61		LSB/MSW		6C		<u>1B</u>	
62		MSB/MSW		00		00	
63	LRA_AGC_Dec	LSB/LSW	4.25 dB	00	4.25 dB	00	
64		MSB/LSW		00		00	
65		LSB/MSW		00		00	
66		MSB/MSW		11		11	
67	AGC_Minimum	LSB/LSW	13 dB	00	13 dB	00	
68		MSB/LSW		00		00	
69		LSB/MSW		00		00	
70		MSB/MSW		34		34	
71	Delta_AGC	LSB/LSW	39 dB	00	30 dB	00	Decrease AGC Adjustment
72		MSB/LSW		00		00	
73		LSB/MSW		00		00	
74		MSB/MSW		<u>9</u> C		78	
75	HRA_Scan_Window		800.0 nSec	08	800.0 nSec	08	
76	HRA_Scan_Hgt_Inc		200.0 nSec	00	200.0 nSec	00	
77	······································	MSB/LSW		00		00	
78		LSB/MID	-	00		00	
79		MSB/MID		00		00	
80		LSB/MSW		01		01	
81		MSB/MSW		00		00	
82	HRA_AGC_NoiseI1		4	_04	4	04	
83	HRA_AGC_NoiseI2		7	07	7	07	
84	HRA_Min_Sig_Thr	LSB	1024	00	1024	00	
85		MSB		04		04	
86 -	LR_Ku_NoiseI1		4	04	4	04	
87	LR_Ku_NoiseI2		7	07	7	07	
88	LR_Ku_Thr_Hgt_Sca	le	31	1F	29	1D	Decrease Ku Threshold Ht. Scale
89	HR_Ku_AGC_I1		16	10	16	10	······································
90	HR_Ku_AGC_I2		47	2F	47	2F	
91	HR_Ku_NoiseI1		4	04	4	04	
92	HR_Ku_NoiseI2		7	07	7	07	

,

	Filename = PAR		1	Filename =		
Byte	Parameter	Value	Hex	Value	Hex	Comment
	Ku_HR_Thr_Hgt	25	19	25	19	
	Fine_Trk_Ku_Alpha	2	02	2	02	
95	Coarse_Trk_Ku_Alpha	1	01	1	01	
96	Ku_AGC_Scale1 LSB	33458	B2	33458	B2	
	MSB		82		82	
98	Ku_AGC_Scale2 LSB	33458	B2	33458	B2	
99 10 101	MSB		82		82	
0	Ku_AGC_Scale3 LSB	33511	E7	33511	E7	
101	MSB		82		82	
102 1 3	Ku-AGC_Scale4 LSB	33479	C7	33479	C7	
3	MSB		82		82	
104	Ku_AGC_Scale5 LSB	33309	1D	33309	1D	
105	MSB		82		82	
16	Ku_AGC_Scale6 LSB	32768	00	32768	00	
107	MSB		80		80	
108		.SB 77	4D	77	4D	
19	MSB		00		00	
1.0		.SB 105	69	105	69	
111	MSB		00		00	-
$\frac{1}{1}\frac{2}{3}$		SB 262	06	262	06	
	MSB		01		01	
114	Ku_EML_Hgt_Scale4 L	SB 696	<b>B</b> 8	696	<b>B</b> 8	
1.5	MSB		02		02	
15		SB 1882	5A	1882	5A	
117	MSB	· ·	07		07	
118	LR_C_NoiseI1	4	04	4	04	
1	LR_C_NoiseI2	7	07	7	07	
120	LR_C_Thr_Hgt_Scale	31	lF	29	1D	Tune Scale
121	HR_C_AGCI1	16	10	19	13	Shift C Gates +3
12	HR_C_AGC_I2	47	2F	50	32	Shift C Gates +3
123	HR_C_NoiseI1	4	04	7	07	Shift C Gates +3
124	HR_C_NoiseI2	7	07	10	0A	Shift C Gates +3
1 5	C_HR_Thr_Hgt	25	19	25	19	
126	Fine_Trk_C_Alpha	2	02	2	02	
127	Coarse_Trk_C_Alpha	1	01	1	01	
13	C_AGC_Scale1 LSB	32768	00	32768	00	
129	MSB		80		80	
130	C_AGC_Scale2 LSB	32768	00	32768	00	
$\frac{1}{132}$	MSB		80		80	
12	C_AGC_Scale3 LSB	32809	29	32809	29	
133	MSB		80	22007	80	
1	C-AGC_Scale4 LSB	32839	47	32839	47	
	MSB		80	52057	80	
136	C_AGC_Scale5 LSB	32829	3D	32829	3D	
130	MSB	52029	80	52027	80	
	C_AGC_Scale6 LSB	32768	00	32768	00	······································
139	C_AOC_Scale0 LSB MSB	-52700	80	52100	80	
14	C_EML_Hgt_Scale1 LSB	312	38	312	38	
			01		$\frac{38}{01}$	
14	MSB		01	I	10	

Table 5.8.4 (Continued)

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[	Filename = PARMC	320		Filename =	C3502	28SL
Byte	Byte Parameter Value Hex			Value	Hex	Comment
142	C_EML_Hgt_Scale2 LSB	440	B8	440	B8	
143	MSB		01		01	
144	C_EML_Hgt_Scale3_LSB	1040	10	1040	10	
145	MSB		04		04	
146	C_EML_Hgt_Scale4 LSB	2784	EO	2784	EO	
147	MSB		0A		0A	
148	C_EML_Hgt_Scale5 LSB	7112	C8	7112	C8	
149	MSB		1B		1B	
150	LR_Track_Point	63.5	7F	63.5	7F	
151	T1 LSB	24	18	96	60	Increase Signal Width
152	MSB		00		00	B
153	T2 LSB	160	A0	2560	00	Increase Signal
	202			2000	00	Variability
154	MSB		00		0A	
155	T3 LSB	16	10	64	40	Tune Signal Width
156	MSB		00		00	
157	T4 LSB	40	28	640	80	Tune Signal Variability
158	MSB		00		02	
159	T8 LSB	1024	00	1024	00	
160	MSB		04		04	
161	Ku_Early_Index1 1	31	1F	31	1F	
162	Ku_Early_Index2_1	31	1F	31	1 <b>F</b>	
163	Ku_Early_Index1_2	30	1E	30	1E	
164	Ku_Early_Index2_2	31	1F	31	1F	
165	Ku_Early_Index1_3	29	lD	29	1D	
166	Ku_Early_Index2_3	30	1E	30	lE	
167	Ku_Early_Index1_4	26	1A	26	1A	
168	Ku_Early_Index2_4	29	1D	29	1D	
169	Ku_Early_Index1_5	20	14	20	14	
170	Ku_Early_Index2_5	27	1B	27	1 <b>B</b>	
171	Ku_Early_Index1_6	8	08	8	08	
172	Ku_Early_Index2_6	23	17	23	17	
173	Ku_Middle_Index1_1	31	1F	31	1F	
174	Ku_Middle_Index2_1	32	20	32	20	
175	Ku_Middle_Index1_2	31	1F	31	1F	
176	Ku_Middle_Index2_2	32	20	32	20	
177	Ku_Middle_Index1_3	30	1E	30	1E	
178	Ku_Middle_Index2_3	33	21	33	21	
179	Ku_Middle_Index1_4	28	1C	28	1C	
180	Ku_Middle_Index2_4	35	23	35	23	
181	Ku_Middle_Index1_5	24	18	24	18	
182	Ku_Middle_Index2_5	39	27	39	27	
183	Ku_Middle_Index1_6	24	18	24	18	
184	Ku_Middle_Index2_6	39	27	39	27	
185	Ku_Late_Index1_1	32	20	32	20	
186	Ku_Late_Index2_1	32	20	32	20	
187	Ku_Late_Index1_2	32	20	32	20	
188	Ku_Late_Index2_2	33	21	33	21	
189	Ku Late Index1 3	33	21	33	21	
		11 504 //		<u></u>		

Г. <u></u>	Filename - PARMC320			Filename = C35028SL		
Byte	Parameter	Value	Hex	Value	Hex	Comment
190	Ku_Late_Index2_3	34	22	34	22	
91	Ku_Late_Index1_4	34	22	34	22	
192	Ku_Late_Index2_4	37	25	37	25	
193	Ku_Late_Index1_5	36	24	36	24	· ·
94	Ku_Late_Index2_5	43	2B	43	2B	
195	Ku_Late_Index1_6	40	28	40	28	
196	Ku_Late_Index2_6	55	37	55	37	
97	C Early Index 1 1	31	1F	34	22	Shift C Gates +3
198	C_Early_Index2_1	31	1F	34	22	Shift C Gates +3
199	C_Early_Index1_2	30	lE	33	21	Shift C Gates +3
00	C_Early_Index2_2	31	 1F	34	22	Shift C Gates +3
101	C_Early_Index1_3	29	1D	32	20	Shift C Gates +3
202	C_Early_Index2_3	30	1E	33	21	Shift C Gates +3
03	C_Early_Index1_4	26	1A	29	1D	Shift C Gates +3
$-\frac{00}{04}$	C_Early_Index2_4	29	ID	32	20	Shift C Gates +3
205	C_Early_Index1_5	20	14	23	17	Shift C Gates +3
06	C_Early_Index2_5	27	<u>1</u> B	30	1E	Shift C Gates +3
- 07	C_Early_Index1_6	8	08	11	OB	Shift C Gates +3
208	C_Early_Index1_0	23	17	26	1A	Shift C Gates +3
200	C_Middle_Index1_1	31	$\frac{17}{1F}$	34	22	Shift C Gates +3
$-\frac{0}{10}$	C_Middle_Index2_1	32	20	35	22	Shift C Gates +3
$-\frac{10}{211}$	C_Middle_Index2_1	31	1F	33	23	Shift C Gates +3
$\frac{211}{212}$	C_Middle_Index1_2	32	$\frac{11}{20}$	35	22	Shift C Gates +3
$-\frac{12}{13}$	C_Middle_Index1_3	30	- <u>20</u> 1E	33	23	Shift C Gates +3
-214		33	21	36	21	
	C_Middle_Index2_3		$\frac{21}{1C}$		$\frac{24}{1F}$	Shift C Gates +3
$\frac{215}{16}$	C_Middle_Index1_4	28 35	23	<u>31</u> 38		Shift C Gates +3
$-\frac{10}{217}$	C_Middle_Index2_4			have been a second s	26	Shift C Gates +3
	C_Middle_Index1_5	24	18	27	<u>1B</u>	Shift C Gates +3
218	C_Middle_Index2_5	39	27	42	2A	Shift C Gates +3
	C_Middle_Index1_6	24	18	27	1B	Shift C Gates +3
220	C_Middle_Index2_6	39	27	42	2A	Shift C Gates +3
221	C_Late_Index1_1	32	20	35	23	Shift C Gates +3
_ 22	C_Late_Index2_1	32	20	35		Shift C Gates +3
223	C_Late_Index1_2	32	20	35		Shift C Gates +3
224	C_Late_Index2_2	33	21	36	24	Shift C Gates +3
25	C_Late_Index1_3	33	21	36	24	Shift C Gates +3
_26	C_Late_Index2_3	34	22	37	25	Shift C Gates +3
227	C_Late_Index1_4	34	22	37	25	Shift C Gates +3
28	C_Late_Index2_4	37	25	40	28	Shift C Gates +3
29	C_Late_Index1_5	36		39	27	Shift C Gates +3
230	C_Late_Index2_5	43	2B	46	2E	Shift C Gates +3
<b>\$</b> 1	C_Late_Index1_6	40	28	43	2B	Shift C Gates +3
32	C_Late_Index2_6	55	37	58	3A	Shift C Gates +3
233	RAVE Time Constant LSB	8	08	8	08	
34	MSB		00		00	
and the second second	Ku_GI_Scale	82	52	82	52	
	C_320_GI_Scale	78	4E	78	4E	
237	C_100_GI_Scale	70	46	70	46	
	HR_Track_Point	63	3F	63	3F	
		0.5	51	0.00	<u></u>	·

Table 5.8.4 (Continued)

Filename - PARMC320				Filename =	C3502	8SL
Byte	Parameter	Value	Hex	Value	Hex	Comment
239	Thr_Hgt_Err_Win1	8	08	8	08	
240	Thr_Hgt_Err_Win2	8	08	8	08	
241	Thr_Hgt_Err_Win3	12	0C	12	0C	<u></u>
242	Thr_Hgt_Err_Win4	16	10	16	10	
243	Thr_Hgt_Err_Win5	32	20	32	20	
244	Fine_Trk_AGC_Alpha	3	03	3	03	
245	Fine_Track_Beta	6	06	6	06	
246	Coarse_Trk_AGC_Alpha	2.	02	2	02	
247	Coarse_Track_Beta	4	04	4	04	
248	T5 LSB	2560	00	2560	00	
249	MSB		0A		0A	
250	T6 LSB	10240	00	10240	00	
251	MSB		28		28	
252	T7 LSB	256	00	256	00	
253	MSB		01		01	
254	Ku_Pulse_Count	25	19	25	19	
255	C_Pulse_Count	26	1A	26	1A	
256	Acq_Pulse_Count	26	1A	26	1A	
257	AGC_Rate LSB	13	<b>0</b> D	13	0D	•
258	MSB _		00		00	
259	Xmit_Test_Height LSB/LSW	8609.6µSec	00	8609.6µSec	00	
260	MSB/LSW		00		00	
261	LSB/MID		00		00	
262	MSB/MID		00		00	
263	LSB/MSW		28		28	
264	MŚB/MSW		A8		A8	
265	Xmit_Test_Hgt_Rate LSB/LSW	0	00	0	00	
266	MSB/LSW		00		00	
267	LSB/MID		00		00	
268	MSB/MID		00		00	
269	LSB/MSW		00		00	
270	MSB/MSW		00		00	
271	Xmit_Test_AGC	60 dB	F0	60 dB	F0	
272	BC_Init	0	00	0	00	
273	WD_Init	25	19	25	19	

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#### 6.0 WAVEFORM SAMPLES ASSESSMENT

## 6.1 Introduction

The various waveform effects discussed in this section are produced by the TOPEX digital filter bank (DFB), and are present in both the Ku and the C waveforms. The DFB will be described briefly. Several departures of the signal from the ideal will be discussed. Some of the departures depend on the value of the fine-height word, so the relationship between height rate and fine-height word is also described. Sets of multiplicative and additive waveform corrections have been developed which, on average, compensate for the waveform departures from ideal. We acknowledge the efforts of P. C. Marth, Jr. of JHU/APL and W. B. Shoemaker of SM Systems and Research Corporation in trying to understand the details of waveform shape effects arising from the digital filter bank. J. R. Jensen, JHU/APL, contributed important simulations to clarify relationships of FFT hardware details to observed waveform features.

#### 6.2 Digital Filter Bank

#### 6.2.1 Brief Description

The digital filter bank (DFB) lies at the heart of the altimeter and its range tracking loop as shown in Figure 6.2.1. The transmitted pulse (320 MHz chirp bandwidth, 102.4 µsec pulsewidth, from the digital chirp generator) scatters off the ocean surface; the received return signal is dechirped and then mixed down to in-phase (I) and quadrature (Q) signals, filtered with a 625 kHz lowpass filter (the "anti-aliasing filter"), and digitized at a 1.25 MHz sample rate. At 1.25 MHz, the 102.4 µsec pulse results in 128 complex samples upon which the DFB performs a 128-point fast Fourier transform (FFT). The FFT output is 128 individual waveform samples, spaced effectively 9.765 kHz apart in frequency (equivalent to 3.125 n spacing in the time domain), which are tracked by the adaptive tracker unit (AMU). The tracking error in the tracking loop is nulled by frequency shifting. The frequency shift, controlled by the track loop, is accomplished by (complex) multiplication of the digitized I and Q samples by a rotating phaser just before the FFT is performed. The phaser rotation rate is controlled by the tracking loop's fine-height word. The 128 individual waveform samples are also compressed into 64 samples and placed into the telemetry stream. The compression depends upon mode, as summarized in **Table 6.2.1**. The DFB performs the same operations for both the Ku and the C altimeters in TOPEX.

In radar altimeters prior to TOPEX, the DFB's used the discrete Fourier transform with a maximum of 64 samples. For TOPEX, the FFT was implemented in digital electronics to lessen the time required to perform the full transform for the 128 samples. In a perfect DFB, the output would be a perfect Fourier transform of the

#### RE INTERFACE

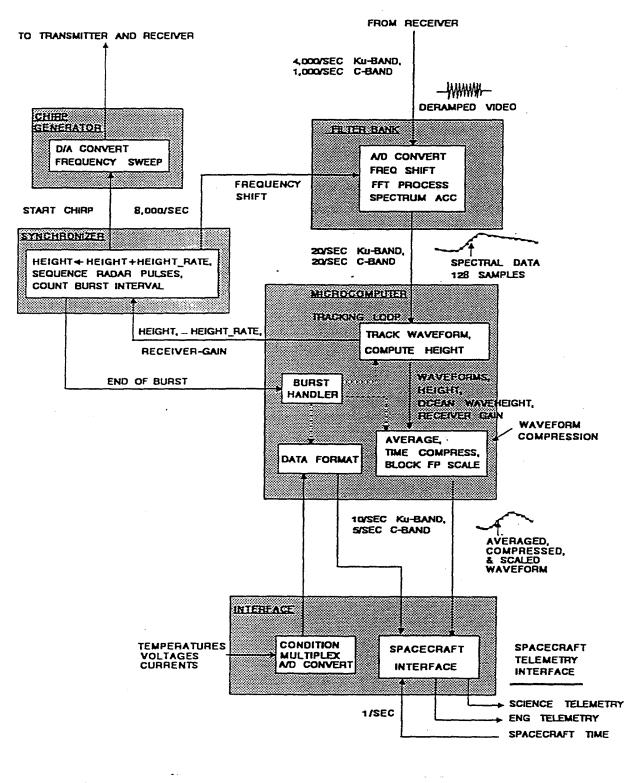


Figure 6.2.1 TOPEX Altimeter Data Processing (from Perschy, JHU/APL)

Table 6.2.1.	TOPEX	Waveform to	Telemetry	Sample	Compression in
		Differen	t Modes		

TELEMETRY Sample #	CAL MODE 2 or TRACK Waveform Sample #	CAL MODE 1, STANDBY, or TRANSMITTER TEST Waveform Sample #	TELEMETRY Sample #	CAL MODE 2 or TRACK Waveform Sample #	CAL MODE 1, STANDBY, or TRANSMITTER TEST Waveform Sample #
1	1 - 2	33	33	41	65
2	3 - 4	34	34	42	66
3	5-6	35	35	43	67
4	7 - 8	36	36	44	68
5	9 - 10	37	37	45	69
6	11 - 12	38	38	46	70
7	13 - 14	39	39	47	71
8	15 - 16	40	40	48	72
9	17	41	41	49 - 50	73
10	18	42	42	51 - 52	74
11	19	43	43	53 - 54	. 75
12	20	_ 44	44	55 - 56	76
13	21	45	45	57 - 58	77
14	22	46	46	59 - 60	78
15	23	47	47	61 - 62	79
16	.24	<sup>,</sup> 48	48	63 - 64	80
17	25	· 49	49	65 - 68	81
18	26	50	50	69 - 72	82
19	27	51	51	73 - 76	83
20	28	52	52	77 - 80	84
21	29	53	53	81 - 84	85
22	30	54	54	85 - 88	86
23	31	55	55	89 - 92	87
24	32	56	56	93 - 96	88
25	33	57	57	97 - 100	89
26	34	<b>58</b>	58	101 - 104	90
27	35	59	59	105 - 108	91
28	36	60	60	109 - 112	92
29	37	61	61	113 - 116	93 <sub>.</sub>
30	38	62	62	·117 - 120	94
31	39	63	63	120 - 124	95
32	40	64	64	125 - 128	96

input with no extraneous effects being introduced by the DFB itself. The shape of the transformed pulse in amplitude versus frequency would look exactly like the time domain average return from a 3.125 n transmitted pulse. But the TOPEX DFB does not provide a completely perfect transform, and there are several small effects of the DFB hardware/software on the shape of the output. Some of these effects are independent of value of the fine-height word, while others of the effects vary with fine-height word. In this discussion the waveform samples will always be numbered 1 through 128, and the telemetry samples will always be numbered 1 through 64.

#### 6.2.2 DFB Effects Not Depending on Fine-Height Word

#### 6.2.2.1 Zero Leakage

At the center of the filter bank (waveform sample 65 out of the full 128) there appears to be an excess signal, spread across several samples. Waveform sample 65 is the zero-frequency sample. This zero-frequency leakage appears to be present in all the TOPEX filter banks tested (*i.e.*, engineering model, breadboard unit, and both Side A and Side B in the flight unit), and is seen in all modes (standby, calibration, and tracking). The zero-frequency leakage is not affected by AGC or receiver gain, and appears to be an additive effect only. The probable explanation for this is provided by J. R. Jensen of JHU/APL who has done tracker simulations which include details of the mathematics performed in the digital FFT. Jensen has found a right shift-caused truncation in the DFB where round-off should have been used, and when the right shift is corrected in Jensen's simulation the zero leakage disappears. There is a subtle fine-height dependence in that round-off error will go to zero when the fine height is exactly zero with respect to any filter. Since in general the fine height is continually changing, the zero-leakage from round-off error will disappear only a small fraction of the time. The time average over even as short a time as one tracker update interval will exhibit this zero-leakage.

There may also be another DC or zero-frequency leakage from charges retained on the biasing capacitors in the analog to digital conversion; residual charge will appear as a DC bias and therefore appear at the zero-frequency filter and its immediate neighbors. As with wraparound, the way to handle the zero-leakage is to ignore the waveform samples in which it is present.

#### 6.2.2.2 FFT Finite Word Length Effects

The TOPEX FFT word length was 8 bits (a limit set by device availability at the time of the altimeter design), and the DFB output exhibits "sawteeth" and a "plateau excess". Jensen's simulation of DFB details included the effects of the finite word length (8 bits) of the DFB FFT devices. The simulation was able to demonstrate both sawteeth and plateau excess matching that seen in TOPEX data. While these effects

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can generally be removed (or compensated for) in ground processing, they are present in the waveforms presented to the TOPEX on-board tracker.

#### Sawteeth

The (uncompressed) waveform samples exhibit an alternating up-and-down neighbor-to-neighbor variation. This effect is also cyclic in groups of eight samples, so that within each group of eight neighboring sampler there will be one which is noticeably lower than the rest. The eighth sample away will also be lower than its neighbors. The sawteeth are smoothed by the compression from waveform to telemetry samples, so that in fine-tracking the sawteeth are seen only in samples number 17-48. The sawteeth are harder to see in actual fine-track data than in Calibration Mode 2 waveforms because in the latter the signal on which the sawteeth ride is an approximately straight horizontal line. **Figure 6.2.2.2a** shows the sawteeth in a typical Ku fine-track waveform averaged for 5 seconds, and **Figure 6.2.2.2b** is a magnified portion of part of the uncompressed waveform region of Figure 6.2.2.2a. **Figure 6.2.2.2c** shows the sawteeth in Calibration Mode 2. The sawteeth seem to be multiplicative, do not exhibit much variation with fine height, and in ground processing can be removed from Waveform samples by using multiplicative compensation-factors derived from Calibration Mode 2 data.

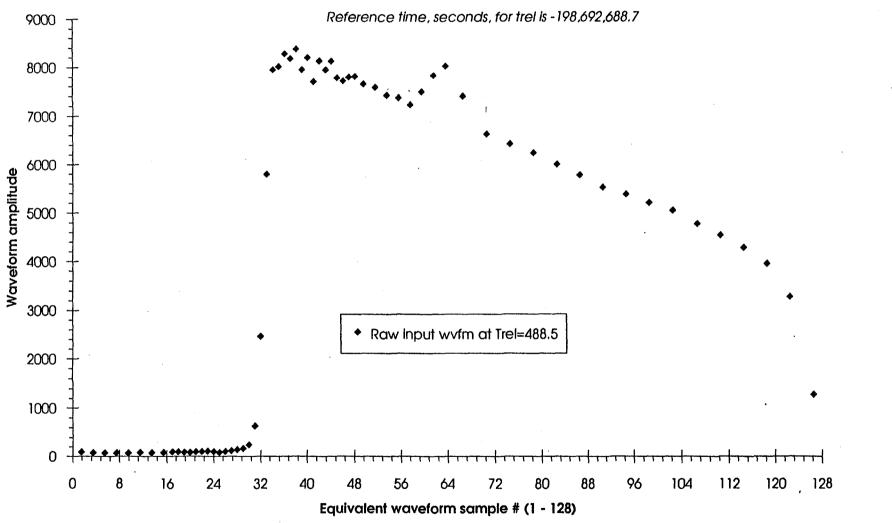
#### Plateau excess

Model TOPEX waveforms were generated at WFF for 2 meter SWH and attitude values of 0.0, 0.2 and 0.4 degrees: Jensen was supplied values for (internal) waveform samples 1 - 128, uniformly spaced at 3.125 n, with their track-point at sample 32.5, and having a (noise) baseline of zero. These model waveforms were used as input to the Jensen TOPEX simulation, and the simulation output waveform samples (1-128) were compared to the input. Jensen also produced a simulated set of Calibration Mode 2 output waveform samples.

The Jensen model can only do Cal-II at fixed values of the fine-height word, but the actual, on-board TOPEX hardware performs Cal-II using a continuous sweep of the height. Jensen supplied three simulated different Cal-II waveforms, for fine-height values of 0 and +/- three samples (*i.e.*, +/- 9.375 n). These simulated Cal-II waveforms are shown in Figure 6.2.2.2d, and with a magnified vertical scale in Figure 6.2.2.2e. These three different results were averaged, and that average used in subsequent waveform figures. The Cal-II average is shown by the solid line on Figures 6.2.2.2d and 6.2.2.2e.

Figure 6.2.2.2f summarizes the waveforms in and out of the Jensen simulation for the TOPEX Ku altimeter at 0.0 degrees attitude angle. The "+" symbols show the model waveform which was sent to Jensen, and the diamond symbols show the output from the Jensen simulation. The Figure 6.2.2.2d Cal-II result was converted into a set of gains used to "correct" the Jensen output, with the result shown by the solid line in Figure 6.2.2.2f. The Figure 6.2.2.2f results were scaled so that the sums





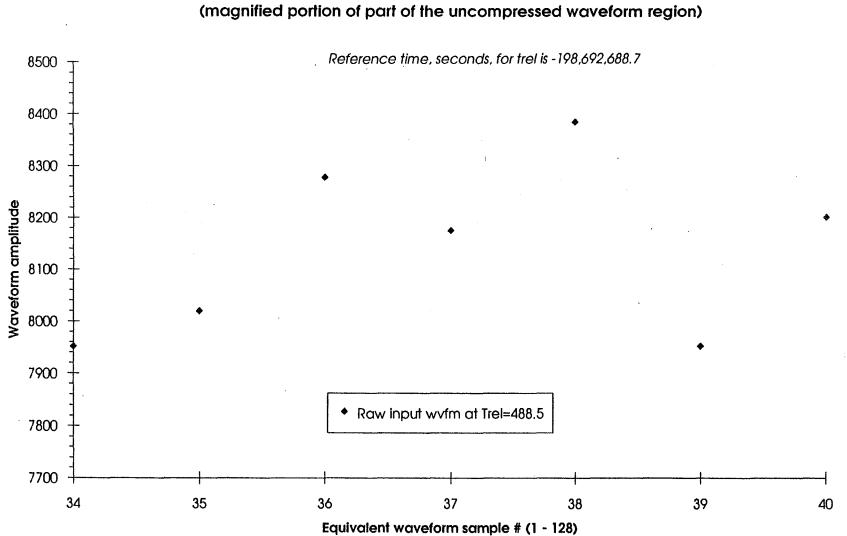
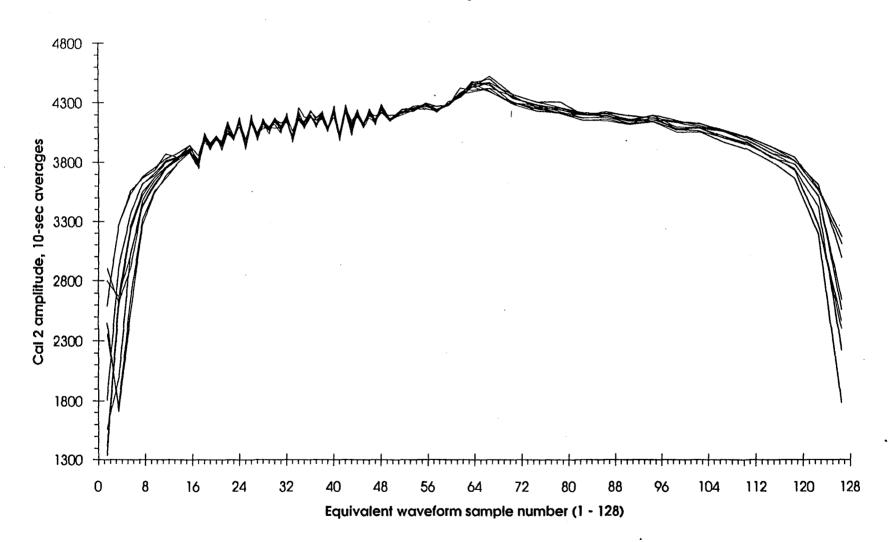
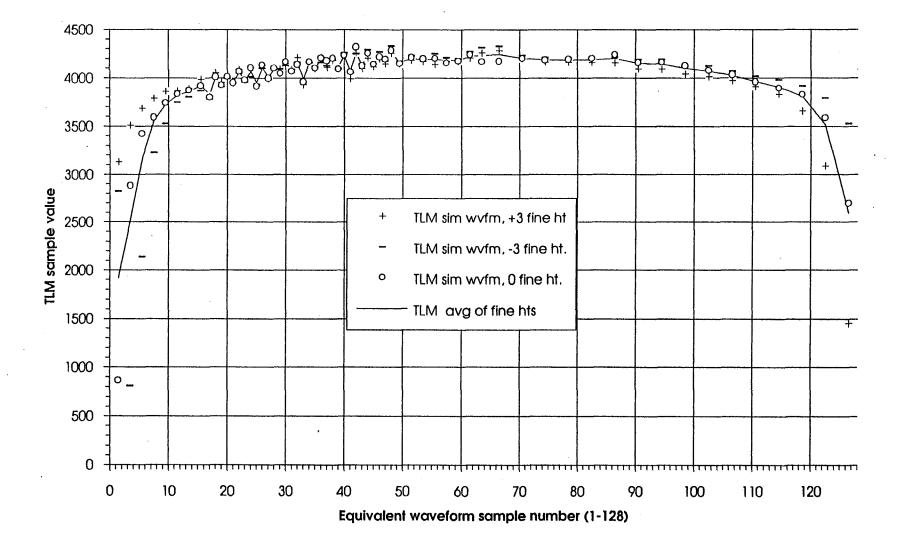


Figure 6.2.2.2b. TOPEX Ku Waveform for 1993 Day 257 Data magnified portion of part of the uncompressed waveform region)







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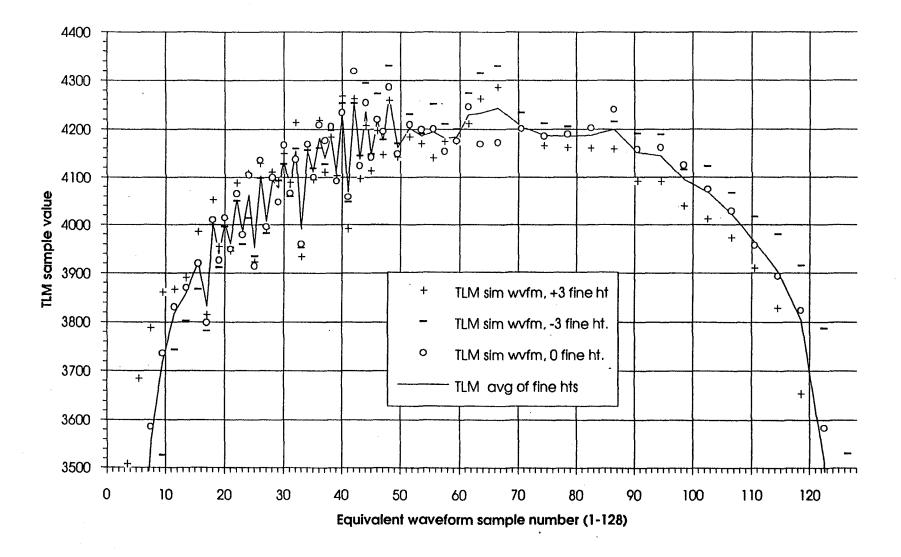
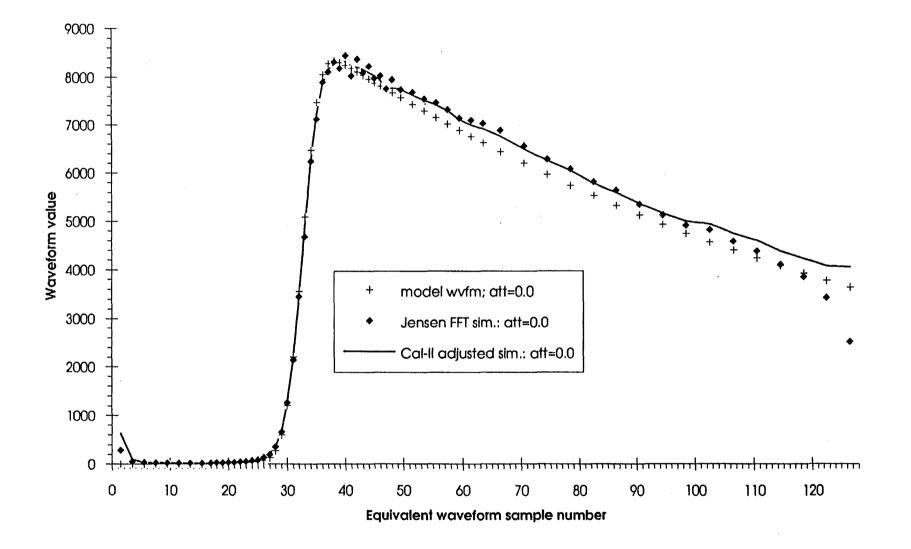




Figure 6.2.2.2f. Modeling Results for TOPEX Ku Altimeter



ა თ of telemetry samples 9-40 (corresponding to waveform samples 17-48) are the same for the three curves presented. Figure 6.2.2.2g presents the corresponding information for attitude of 0.2 degrees and Figure 6.2.2.2h is for an attitude of 0.4 degrees. Notice that these figures show the 64 telemetry sample results, but plotted on the equivalent waveform sample range of 1-128.

Comparing input ("+" symbols) to Cal-II-corrected output (solid lines) in Figures 6.2.2.2f - 6.2.2.2h there are two different effects: i) extra signal in the waveform plateau, and ii) non-zero baseline in the noise region of the waveform. Both of these effects may have some correlation with some of the observations of properties of actual TOPEX data. In these comparisons, one should ignore the first and last four TLM samples (equivalent waveform samples 1-8 and 113-128) since they may be affected by the bandpass filter and its wraparound.

The plateau excess can be compensated by a set of multiplicative corrections derived from the Jensen simulations. Fine-height dependence has been ignored to date in these corrections. Further simulation study is needed.

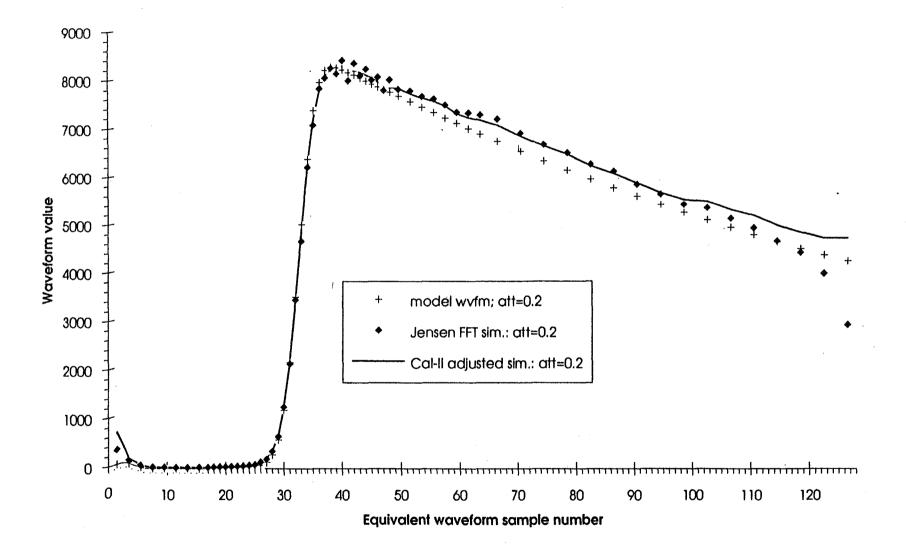
6.2.3 DFB Effects Which Do Depend Upon the Fine-Height Word

6.2.3.1 Relationship between Range Rate and Fine-Height Word

Because some waveform effects to be described depend on the value of the fineheight word, a brief description of the relationship between height rate and fineheight word is provided here. Figure 6.2.3.1 is a sketch of the fine-height word in the tracker. The least significant bit of the coarse-height word is 185.7 cm. When the height rate is negative (the height is decreasing), the height is kept in the upper half of the fine-height word. If the tracked height moves to the lower part of the fineheight word, the coarse-height is reduced by 187.5 cm and the fine-height is increased by 187.5 cm; the sum of the coarse-height and fine-height is the same but the tracked height will have been moved back to the upper half of the fine-height word. Similarly for increasing height the tracked height is kept in the lower half of the fine-height word.

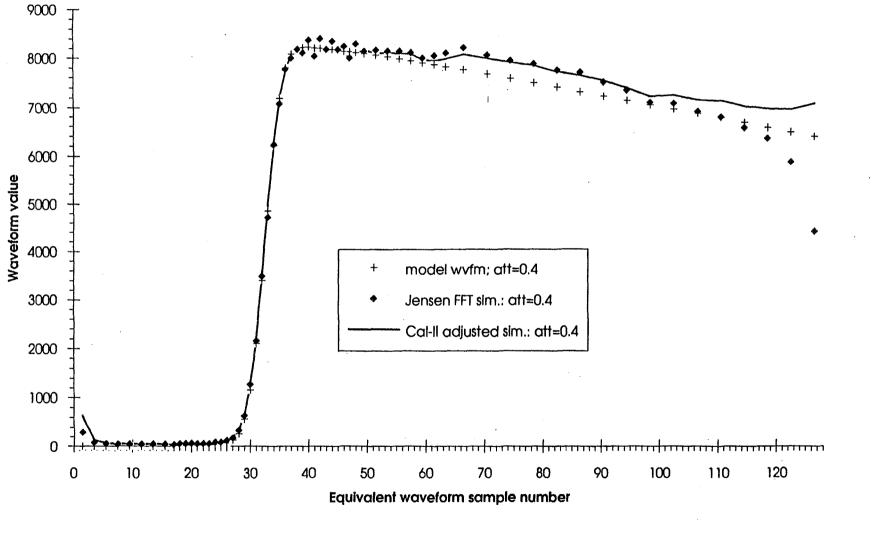
As TOPEX moves from a latitude extreme (either plus or minus 66 degrees) toward the equator, the height rate is negative as depicted in Figure 5.4a. After the equator crossing, the height rate is positive, resulting in four height rate sign reversals per orbit. Table 6.2.3.1 summarizes which half of the fine-height word the height will be in.



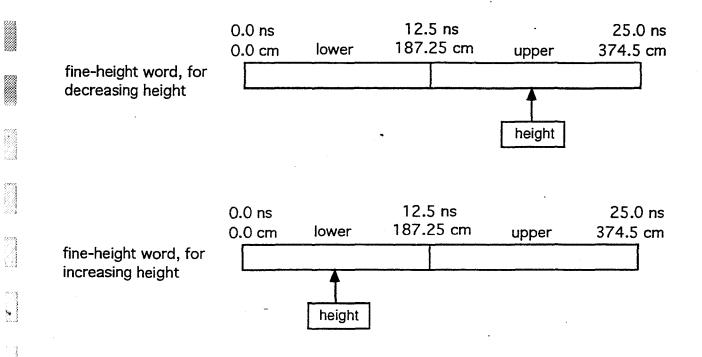


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	Positive Latitude (above equator)	Negative Latitude (below equator)
North-to-South	Upper Half	Lower Half
passes	(negative height rate)	(positive height rate)
South-to-North	Lower Half	Upper Half
passes	(positive height rate)	(negative height rate)

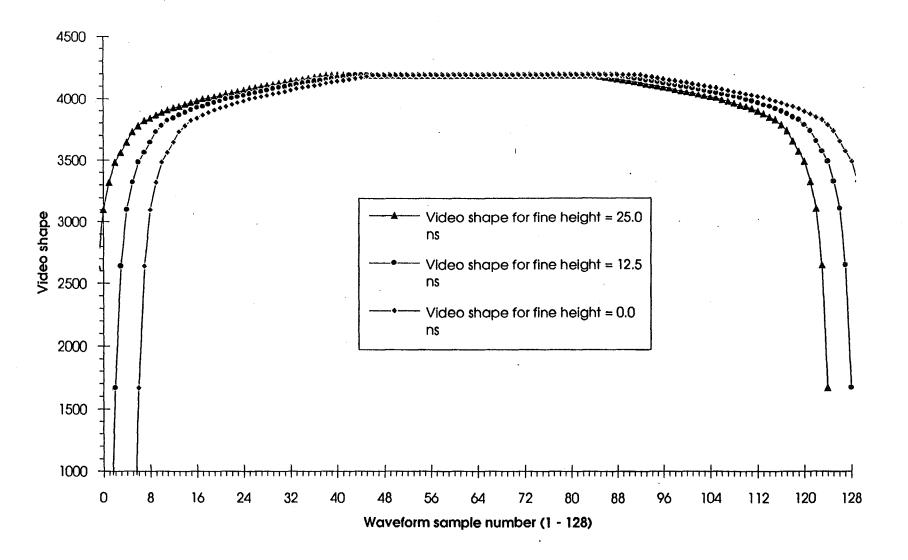
## Table 6.2.3.1 Fine Height Word Value Summary

## 6.2.3.2 Filter Bank End Fall-off

For different fine-height values, the effects of the receiver's video bandwidth shape (the shape resulting from the anti-aliasing filter) is moved horizontally relative to the individual waveform samples. **Figure 6.2.3.2** shows an idealized effect for three different fine-height values, and Figure 6.2.2.2d shows the effect as it appears in Calibration Mode 2 data. In this mode, the altimeter looks at noise only, and the AGC adjusts the receiver gain to attain a fixed level of 4096 for AGC minus noise. If the DFB created a replica of receiver input noise, there should be a fixed level seen across the entire band. The anti-aliasing filter modifies this, and the effects are seen at the low and high ends of the bandwidth. In Calibration Mode 2 the fine-height is constantly swept in one direction, incrementing the constantly the low-order bit of the fine-height word (so the fine-height will increase to its maximum, then start increasing again from zero) so that successive averages will show the anti-aliasing filter shape moving relative to the waveform samples. (Recall from Table 6.2.1 that the Cal Mode 2 telemetry compression is the same as for the fine-track mode.)

#### 6.2.3.3 Wraparound effect

At small values of fine height the filter bandpass is shifted upward and some of the energy coming through the filter bandpass will fall past the extent of the filter bank. This extra energy which should be at the end of the pulseshape will be wrapped around to the early waveform samples as sketched in Figure 6.2.2.2f. Figure 6.2.2.2c shows actual data exhibiting the wraparound in low samples. The wraparound affects only waveform samples at the low and high extremes.





#### 6.2.3.4 Leakage effects

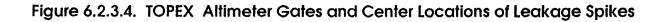
In the investigations of other anomalies in TOPEX waveforms, various low level "spikes" were found in some waveform samples. These were often referred to as leakage. The power of the leakage spikes is independent of receiver gain (AGC), and the location of these spikes in the waveform sample set moves with fine height value. These effects probably enter the altimeter just at or just after the A/D conversion of the I and Q video output of the receiver. Figure 6.2.3.4 sketches the location of these spikes and their relative magnitudes for the fine-height word at the middle of its range (i.e., at the highest height word value for positive height rate, or the lowest height word value for negative height rate). All these leakage effects are small, about 200 counts or less, compared to the AGC value of 4096 counts and the waveform peak of the order of 8000 counts. Figure 6.2.3.4 also indicates the locations of various gates (where gate designates the average value of a specified range of waveform samples) including the early (E), middle (M), and late (L) gates used in tracking. E5, M5, L5 and E4, M4, L4 are the track gates for gate index values 5 and 4, and are indicated by labeled horizontal lines in the figure. The E, M, L gates for gate index values 3, 2, and 1 are shown but labels are not attached in order avoid an excessively cluttered figure.

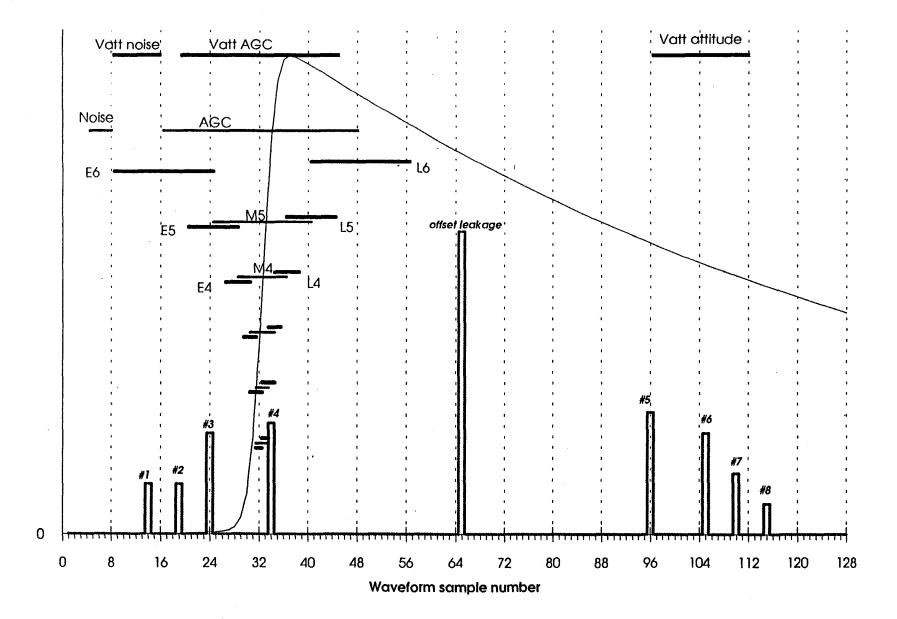
One way to see the leakages more clearly in the altimeter was to examine Standby Mode waveforms in a transmitter test mode. In Standby, the 64 telemetry samples contain the center 64 (of the total 128) uncompressed waveforms. In the transmitter test mode the height is continuously incremented, either up or down, so the fine height word will sweep over either the upper or the lower half of its range depending on the sign of the height rate. Only the leakages affecting track-mode telemetry samples 25-56 can be seen in this test mode, and the test must be run in two halves, each with a different sign of height rate, to see the full range over which a leakage spike will move.

Another way to see the leakage spikes has been to run a special Calibration Mode 2 test with the AGC attenuator set at a high level (approximately 25 dB) to attenuate the normal noise signal far enough that the small leakage spikes can be seen more clearly. The difficulty with this test mode is that some of the non-central leakages are made more difficult to see by the factor-of-2 or factor-of-4 compression in the telemetry samples.

Because the leakage spikes are so small and these test modes so relatively noisy, the interpretation of the test data is highly subjective. The characterizations of leakage spike amplitude could be in error by as much as 50%, and the waveform sample locations could be incorrect by perhaps one sample. **Table 6.2.3.4** lists the leakage spikes by designation, amplitude, and sample number location.







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Spike Designation	Amplitude,	Waveform Sample	TLM Sample #
	Counts	# Range	Range (track mode)
Offset Leakage	300	61-69	47-49
1	50	10-18	6-10
2	50	15-23	8-15
3	100	20-28	12-20
4	110	30-38	22-30
5	120	92-100	55-57
6	100	102-109	58-60
7	60	106-114	59-61
8	30	111-119	60-62

### Table 6.2.3.4 Characterization of TOPEX Waveform Leakage Spikes

#### Offset leakage spike

There is leakage in the center of the waveform sample set, at sample  $65 \pm 4$  depending on fine height word, which is designated as offset leakage in Table 6.2.3.4. This offset leakage is difficult to separate from the zero-frequency leakage (from truncation in digital FFT), but the offset leakage does move with fine-height word (unlike the zero-frequency leakage which does not). This offset leakage is believed to result from a slight offset in biasing of the A/D converter, based on some tests with existing engineering model and breadboard hardware. The engineering model DFB had the offset leakage whose behavior was approximately the same as the flight unit. The breadboard DFB did not originally exhibit the offset leakage but the offset leakage could be made to appear by changing the value of the biasing resistor on the A/D converter in the I channel. The resistor was adjusted until the breadboard DFB offset leakage matched that of the engineering model; this work indicates that the flight unit offset leakage is about 3/8 of the least significant bit in the I-channel A/D.

#### Other (numbered) leakage spikes

In the flight unit Side A there are several other spikes which are of the order of 100 counts or less, and these are designated by numbers 1 to 8 in Table 6.2.3.4. These spikes are probably caused by low level signal leakage into the receiver chain after the AGC adjustments but before the actual FFT process. There is no AGC effect on their amplitude above noise, but there does seem to be some temperature variation of their amplitudes. This temperature variation was found by reexamining some of the test data. The positions of most of these leakage spikes can be correlated with harmonics of the low voltage power supply switching frequencies, specifically 300, 400, 450, and 500 kHz; the power supply is known to generate these frequencies, based on unit EMF test. Notice that spikes # 4 and #5 are approximately equidistant

from the zero-frequency sample at the center of the 128 waveform sample set. Similarly spikes #3 and #6, #2 and #7, and #1 and #8 are also pairwise equidistant about the center. These particular (numbered) leakage spikes are present only in the flight unit Side A, but were only discovered and investigated after TOPEX launch. Side A had been chosen as the prime altimeter because of minor concerns on reliability of some components (*i.e.*, switches and UCFM).

The most worrisome of the numbered leakage spikes in Table 6.2.3.4 is #4 because of its position relative to the TOPEX tracking gates; this can been seen in Figure 6.2.3.4. All the spikes in Figure 6.2.3.4 will move by  $\pm$ 4 samples relative to their plotted center positions, depending on the value of the fine height word, and this moves spike # 4 in and out of the track gates. Some simulation studies at both JHU/APL and WFF indicate that the tracking errors arising from leakage spike #4 should generally be 1 centimeter or less for most operating conditions of general interest, but further study is needed.

## 6.3 Correction/Compensation for Waveform Effects

6.3.1 Waveform Samples to Avoid

Because of the zero-leakage and the offset leakage effects, any ground-based waveform processing should avoid using waveform sample 65 and at least 4 samples to either side of 65. Waveform fitting at WFF has skipped (by assigning zero weights to) telemetry samples 45 - 50, corresponding to waveform samples 57 - 72. Likewise, to avoid effects of fall-off and wraparound, telemetry samples 1 - 4 and 61 - 64 (corresponding to waveform samples 1 - 8 and 113 - 128) have been avoided in all WFF waveform fitting work.

#### 6.3.2 Multiplicative and Additive Factors

The several waveform effects can be partially compensated or corrected by sets of multiplicative and additive waveform factors. For each telemetry sample  $T_{s,i}$  there is an additive adjustment  $A_i$  and a multiplicative adjustment  $G_i$  to produce a corrected (compressed) waveform  $W_{c,i}$ ,

$$W_{c,i} = G_i x (T_{s,i} + A_i)$$
,

for i from 1 to 64. The multiplicative adjustments, which are obtained from a combination of the Cal Mode 2 data and additional modeling, compensate for the sawteeth, zero-frequency leakage, plateau excess, and end fall-off.

The additive adjustments A<sub>i</sub> compensate for the leakage effects but only in a way that averages over all possible fine height values. The spikes in Figure 6.2.3.4 are for the fine height at the middle of its entire range (i.e., at the highest value in the lower

half or the lowest value of the upper half). Since the fine height value is not available for ground post-processing (the fine height value was added to the coarse height in the Engineering Units conversion early in the TOPEX ground processing flow, and is not available as a separate entity after that conversion step) so that the spikes have to be treated as a probabilistic smear over all their possible positions. Figure 6.3.2 shows the spikes of Figure 6.2.3.4 as smeared over  $\pm 4$  samples. It would be possible to halve the width of the smear by treating data differently for increasing than for decreasing height rate, but that has not been done to date.

**Table 6.3.2** gives the TOPEX Ku and C-320 multiplicative and additive correction factors which are currently in the routine TOPEX ground processing at JPL. The additive factors are based on somewhat earlier leakage analyses than Figure 6.2.3.4 and 6.3.2 and Table 6.2.3.4. Specifically, spikes #7 and #8 were not used, and the two offset leakage spikes were used. These differences should be unimportant, because the offset leakage is not in any of the gates, and spike #7 contributes only a very small amount to the Vatt attitude gate used in attitude corrections in the ground-based processing. The Table 6.3.2 factors are currently used in various waveform fitting analyses at WFF. Although these factors are given for all the telemetry samples, the samples # 1 - 4, 45-50, and 61 - 64 have been given zero weighting in all WFF waveform fitting or analyses.

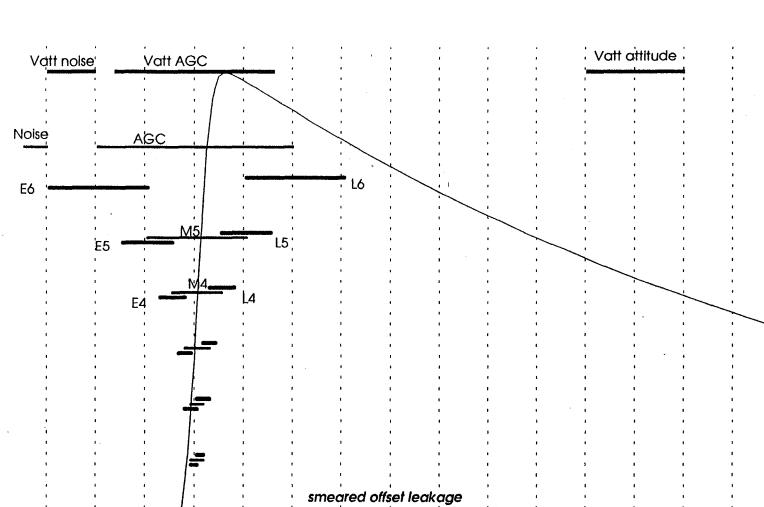
### 6.3.3 Example

**Figure 6.3.3** shows a typical input waveform (from scaled telemetry samples), the corrected waveform after application of the multiplicative and additive factors, and the waveform which is least-squares fitted to the corrected waveform. In this figure there is some fall-off and wraparound still visible, and there is still some extra energy in the vicinity of equivalent waveform sample 64. The waveform in this figure is the same 5-second average already shown earlier in Figure 6.2.2.2a. From waveform fitting to the corrected waveform data in Figure 6.3.3, the SWH estimate is 1.41 meters, and the attitude estimate is 0.12 degrees.

### 6.4 Waveform Effects Status

During the first half year of TOPEX operation (the evaluation period) approximately a half dozen new sets of range (and AGC and SWH) correction constants for the TOPEX data ground-processing system were generated through simulation procedures by the Wallops TOPEX team; each simulation set incorporated thencurrent knowledge of the various waveform features. Our waveform fitting results, after the final set of constants was produced in May 1993, do seem to confirm the adequacy of that set of processing constants for the range corrections for effects of significant waveheight and attitude. Some small improvements could conceivably be made, based on more extensive analyses, but we see no compelling argument for changes in the ground processing constants.





smeared spikes 1-4

TITT

Waveform sample number

Figure 6.3.2. TOPEX Altimeter Gates and Smeared Locations of Leakage



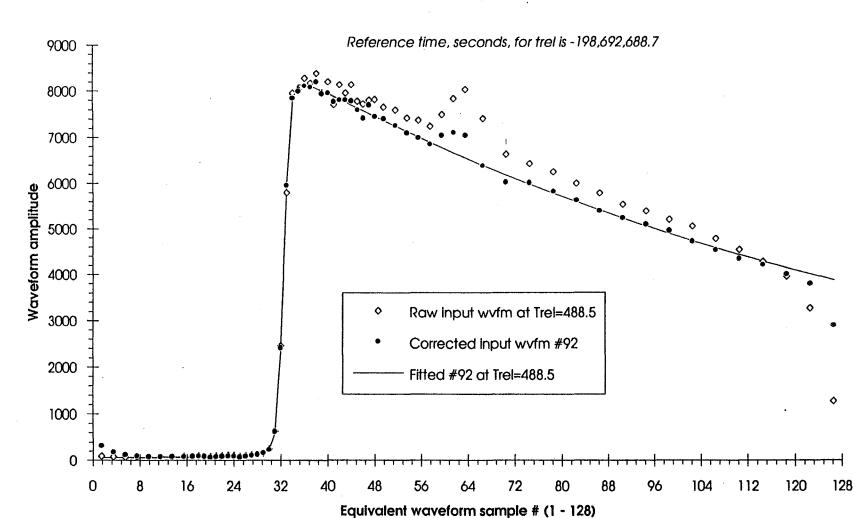
smeared spikes 5-8

80°

TTTTTT

Telemetry	Multiplicative	Ku	С	Telemetry	Multipicative	Ku	C
sample #	factors	additive	additive	sample #	factors	additive	additive
		factors	factors			factors	factors
1	3.355	0	0	33	1.007	0	0
2	2.327	0	0.	34	0.958	0	0
3	1.638	0	0	35	0.981	0	0
4	1.178	0	0	36	0.956	0	0
5	1.120	-1.39	-0.35	37	0.976	0	0
6	1.083	-2.78	-0.69	38	0.960	0	0
7	1.065	-2.78	-0.69	39	0.986	0	0
8	1.047	-9.44	-2.36	40	0.953	0	0
9	1.070	-9.44	-2.36	41	0.966	. 0	0
10	1.025	-9.44	-2.36	42	0.955	0	0
11	1.041	-6.67	-1.67	43	0.955	0	0
12	1.025	-17.78	-4.44	44	0.947	0	0
13	1.036	-17.78	-4.44	45	0.945	0	0
14	1.012	-17.78	-4.44	46	0.941	-15.56	-3.89
15	1.029	17.78	-4.44	47	0.913	-57.78	-14.44
16	1.009	-11.11	-2.78	48	0.882	-57.78	-14.44
17	1.037	-11.11	-2.78	49	0.868	-57.78	-14.44
18	0.999	-11.11	-2.78	50	0.909	-6.67	-1.67
19	1.023	-11.11	-2.78	51	0.934	0	0
20	1.002	-11.11	-2.78	52	0.932	0	0
21	1.006	0	0	53	0.937	0	0
22	0.992	-11.11	-2.78	54	0.932	0	0
23	1.005	-11.11	-2.78	55	0.945	-2.78	-0.69
24	0.987	-11.11	-2.78	56	0.948	-11.11	-2.78
25	1.027	-11.11	-2.78	57	0.954	-11.11	-2.78
26	0.987	-11.11	-2.78	58	0.936	-11.11	-2.78
27	0.997	-11.11	-2.78	59	0.949	-11.11	-2.78
28	0.981	-11.11	-2.78	60	0.957	-2.78	-0.69
29	0.990	-11.11	-2.78	61	0.983	0	0
30	0.979	-11.11	-2.78	62	1.013	0	0
31	0.997	0	0	63	1.158	0	0
32	0.969	Ó	0	64	2.273	0	0

Table 6.3.2. TOPEX Waveform Multiplicative and Additive Correction Factors



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Figure 6.3.3. Comparison of Input, Corrected, and Fitted Ku Waveforms for 1993 Day 257 Data

Although it would have been desirable for the waveform features to be smaller, we do find by comparing CAL2 telemetry sample data for more than a year of TOPEX operation that the individual samples are stable to better than 1% of their maximum magnitude and therefore have met the original hardware specification for waveform sampling. Considerable clarification has been gained from various post-launch laboratory testing of engineering model components, but we stress the importance of further testing and understanding of waveform features for any future radar altimeters.

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### 7.0 ENGINEERING MONITORS

### 7.1 Temperatures

The altimeter has 26 internal thermistors to monitor temperatures near key components. The specifications for the altimeter included the provision that the altimeter operate within temperature ranges of -10 degrees to +45 degrees Centigrade. The spacecraft was to maintain the baseplate temperatures between 0 and +35 degrees C, which it has done through February 1994.

Figure 7.1 comprises the temperature history plots for the 26 thermistors. The lowest temperatures occur after the altimeter has been in IDLE or OFF mode. The highest temperatures are observed to occur during periods of No Yaw Steering.

The minimum/maximum values for each of the thermistors during TRACK mode are summarized in Table 7.1. The temperature values are all within the specified temperature range and are within ranges predicted by the prelaunch spacecraft Hot and Cold balance tests. The thermal control systems of the altimeter and spacecraft appear to be performing very well.

### 7.2 Voltages, Powers and Currents

The altimeter has 17 monitors for voltages, powers and currents; the time history of each of these monitors is shown in Figure 7.2. Table 7.2 lists the minimum/maximum on-orbit output for each of the 17 monitors during TRACK mode, and compares them with measured TRACK mode values just prior to launch. All are performing nominally and within specifications.

7.2.1 Low Voltage Power Supplies (LVPS)

The time history of the LVPS [+12 V, +28 V, +15 V, -15 V, +5 V(5%), +5 V(1%), -5.2 V and -6 V], as depicted in Figure 7.2, illustrates that these power supplies are steady with little deviation.

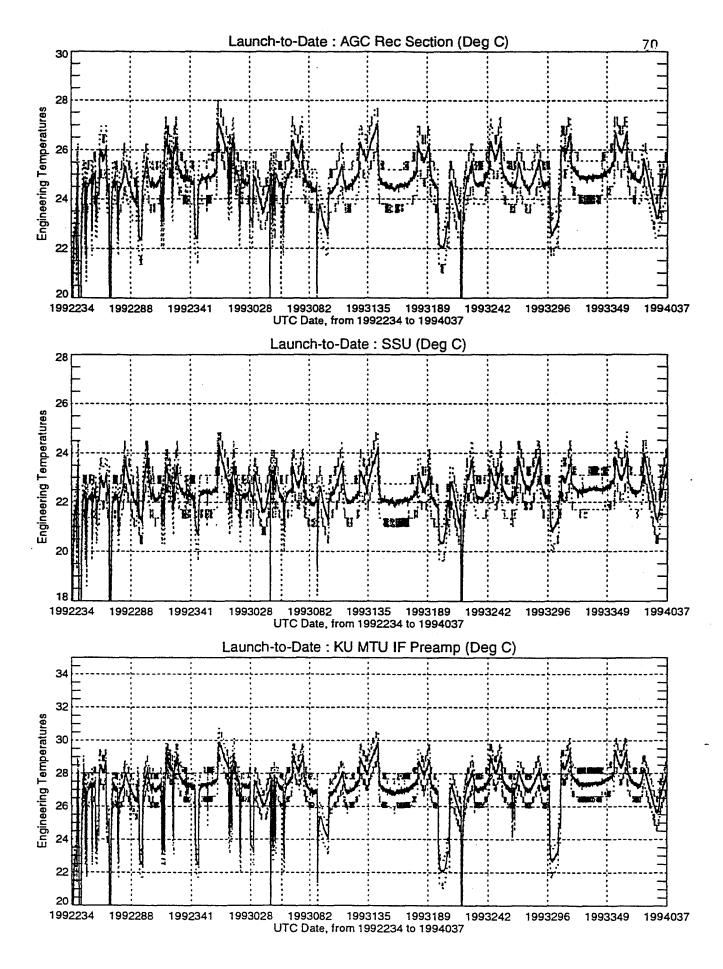
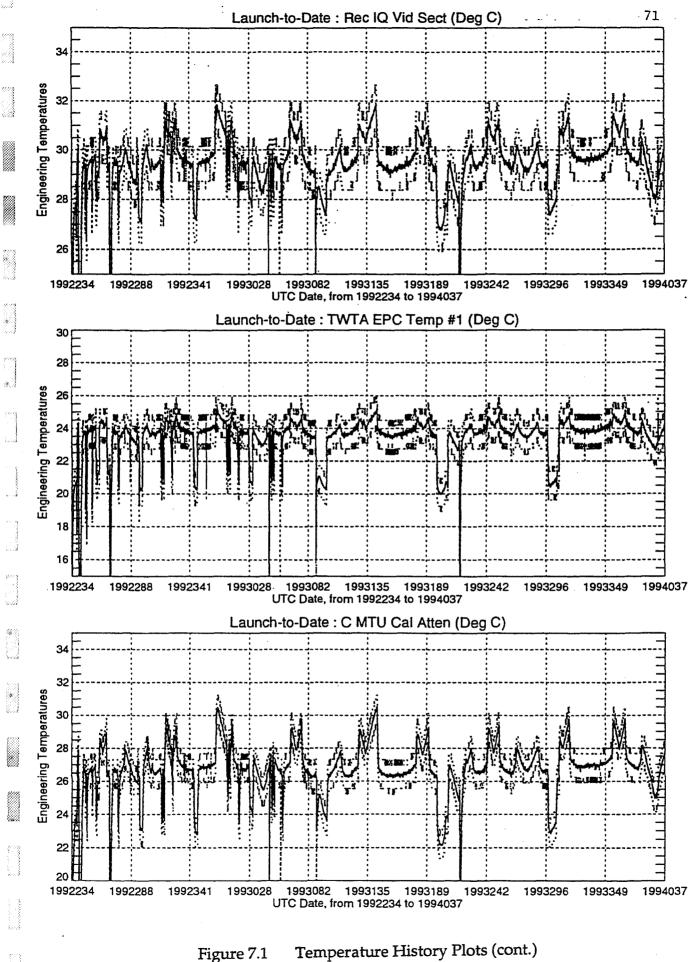
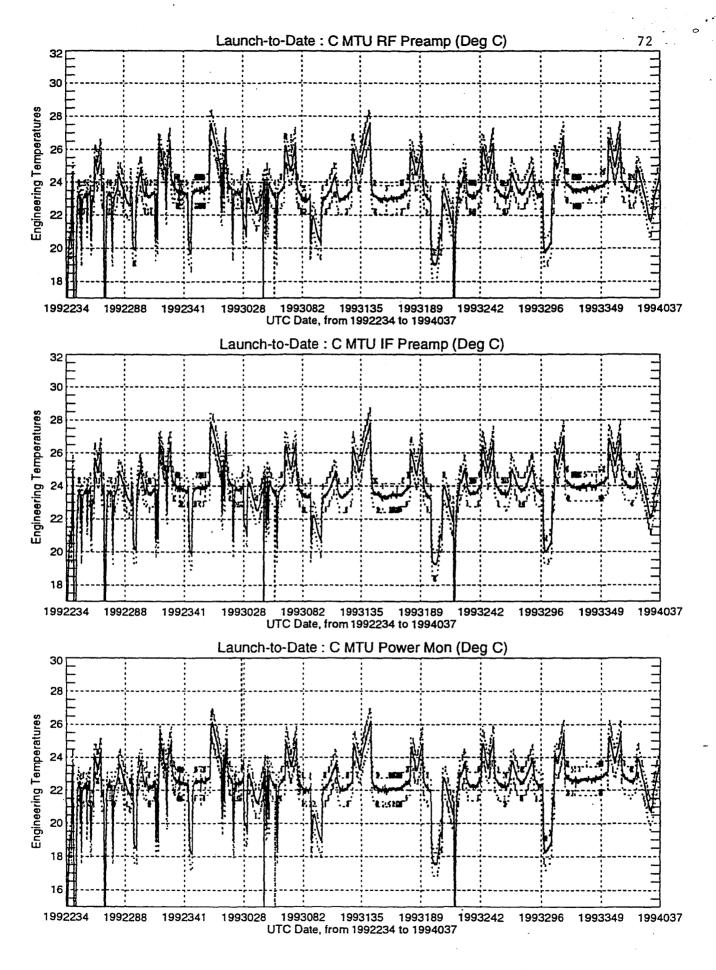
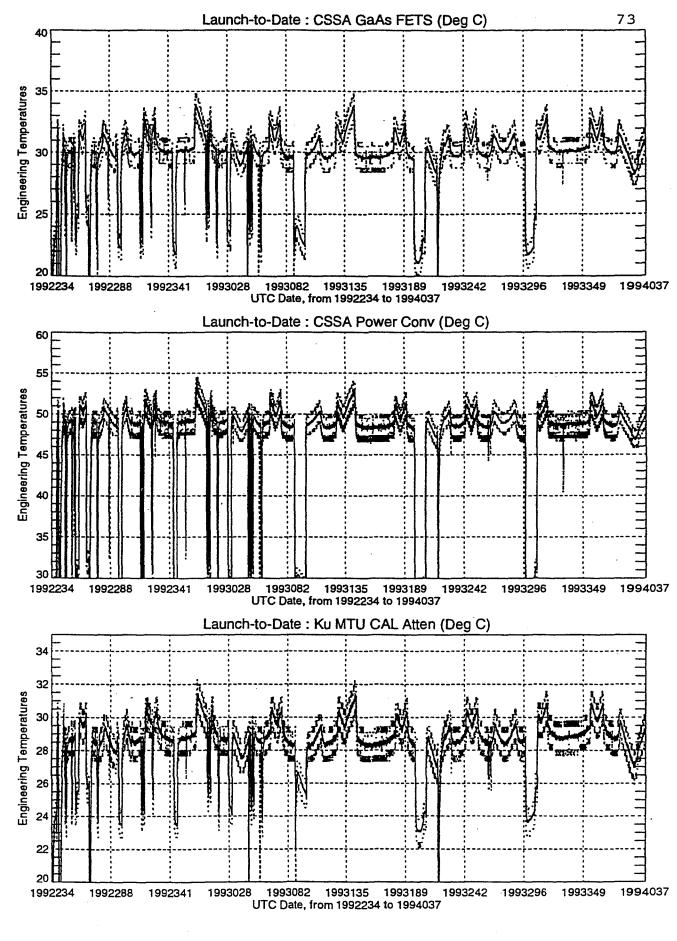


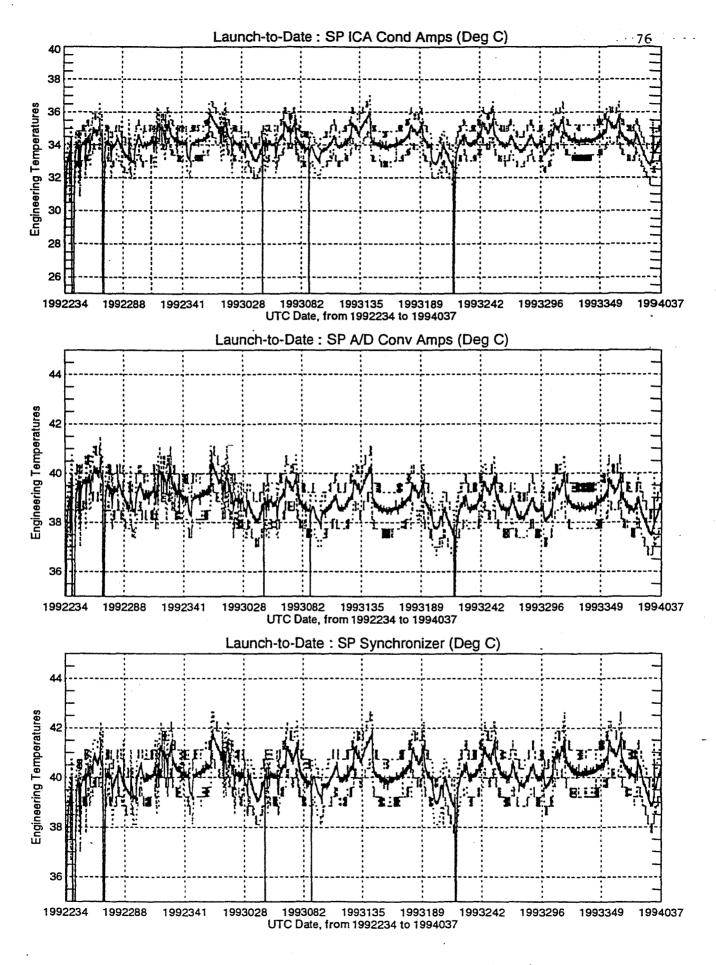
Figure 7.1 Temperature History Plots







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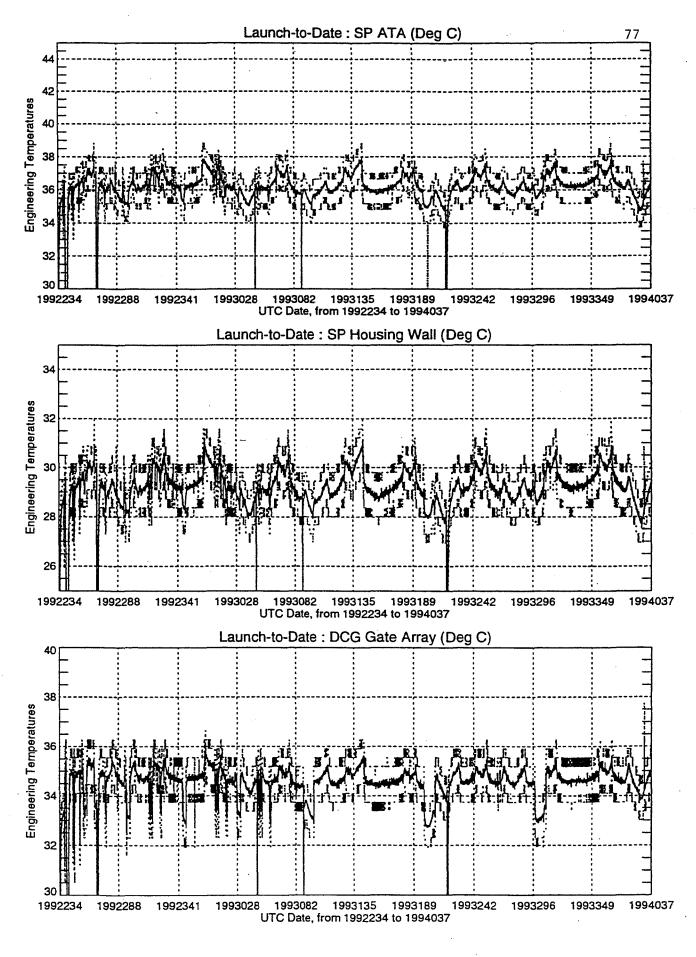
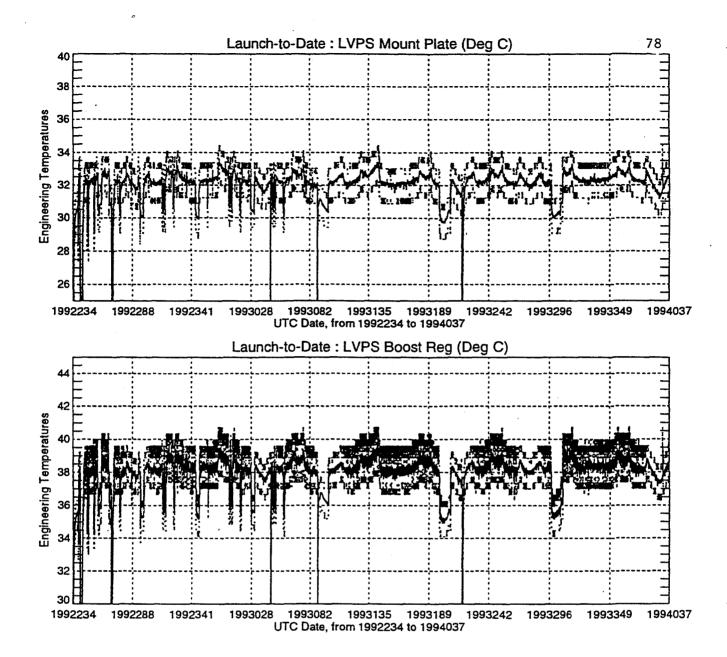


Figure 7.1 Temperature History Plots (cont.)

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Thermistor Location	Minimum (ºC)	Maximum (ºC)
Receiver AGC Selection	23.5	28.0
SSU	20.4	24.8
Ku MTU IF Preamp	25.0	30.6
Receiver IQ Video Select	28.4	32.9
Ku TWTA EPC	22.0	26.0
C MTU Cal Attenuator	24.5	31.2
C MTU RF Preamp	22.0	28.4
C MTU IF Preamp	22.3	28.7
C MTU Transmit Power Monitor	21.0	27.0
C SSA GaAs FETS	28.3	34.9
C SSA Power Converter	45.9	54.6
Ku MTU Cal Attenuator	27.2	32.2
Ku MTU Transmit Power Monitor	24.0	29.6
UCFM	24.2	28.0
Ku MTU RF Preamp	26.5	31.3
Downconverter	29.5	34.9
SP DFB Butterfly Board	42.8	46.6
SP DFB Memory	38.8	43.0
SP ICA Condition Amps	32.5	37.2
SP ICA A/D Converter	37.0	41.2
SP Synchronizer	38.8	42.7
SP ATA	34.8	38.9
SP Housing	27.5	32.0
DCG Gate Array	33.2	37.8
LVPS Transformer Mounting Plate	31.0	34.4
LVPS Boost Regulator Assembly	36.2	40.7

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# Table 7.1Thermistor Minimum/Maximum TemperaturesDuring TRACK Mode

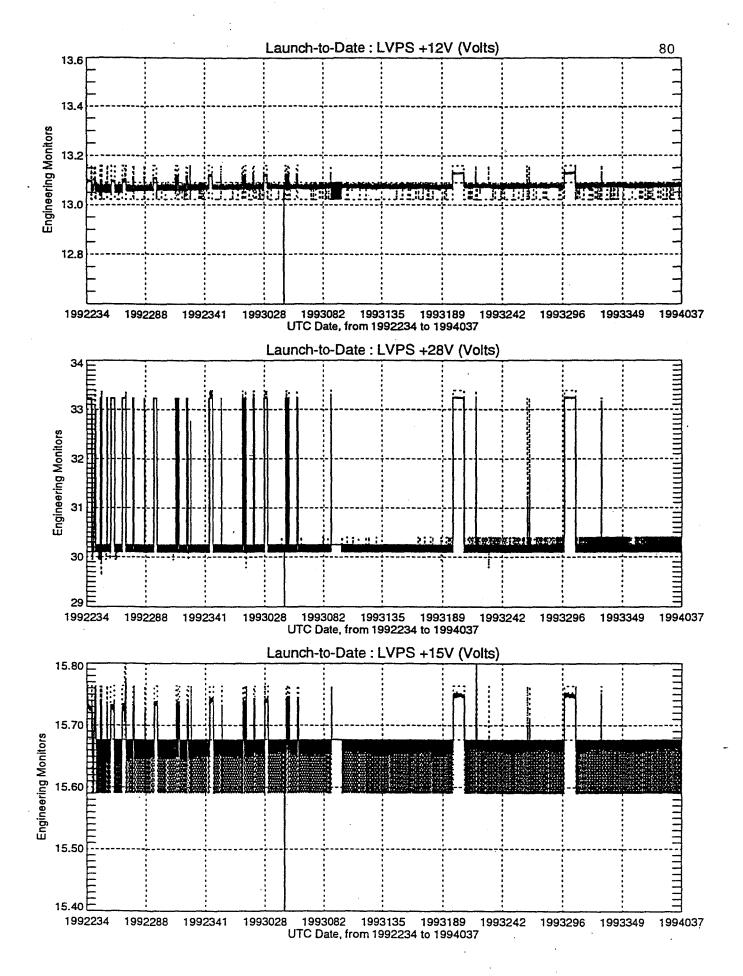
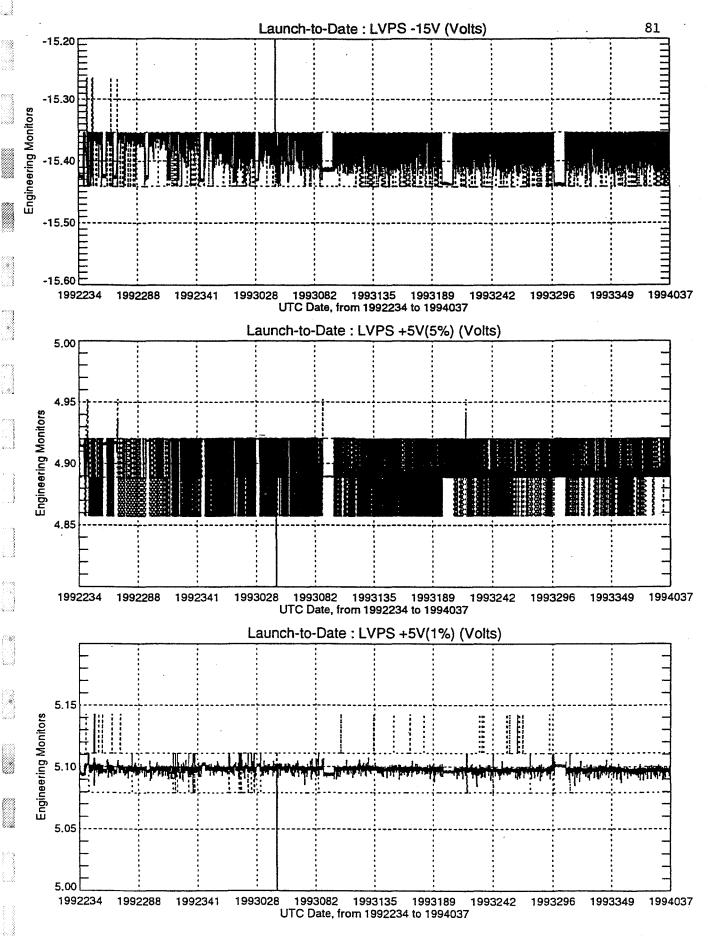
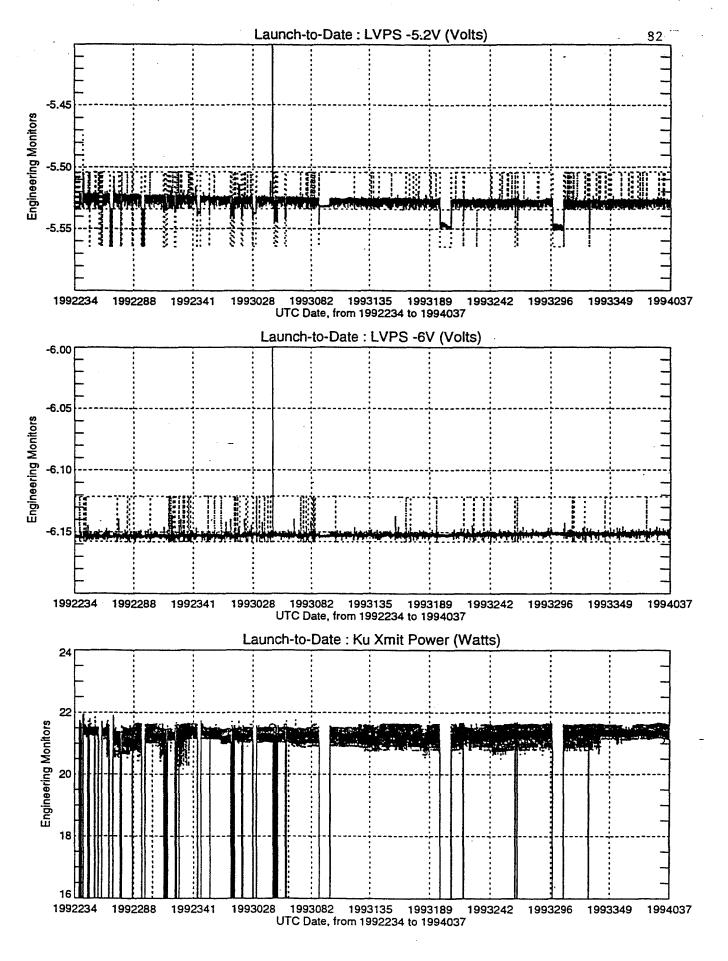
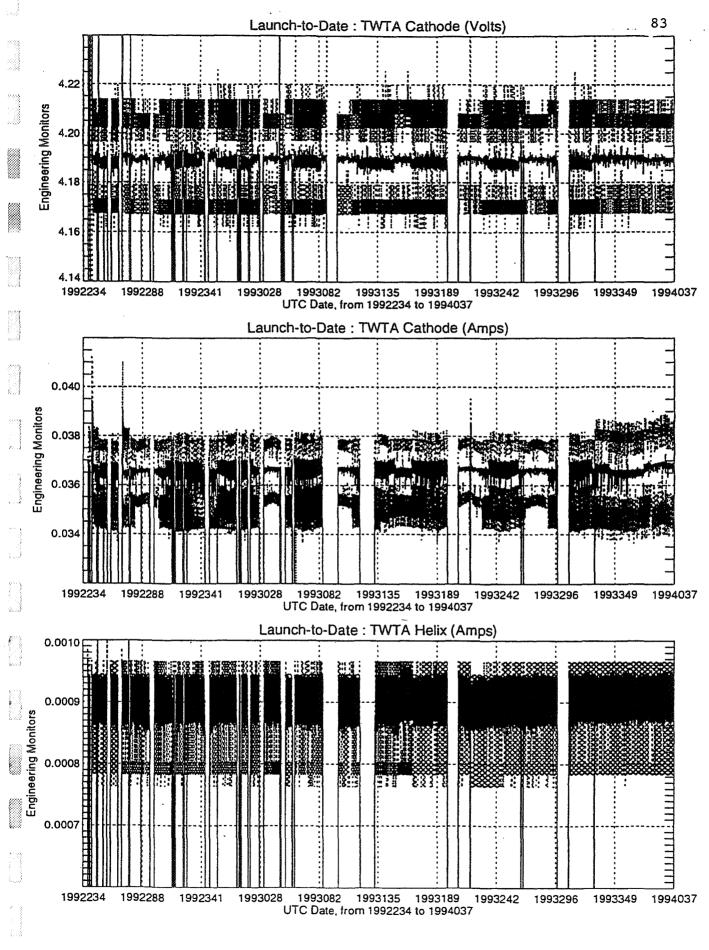


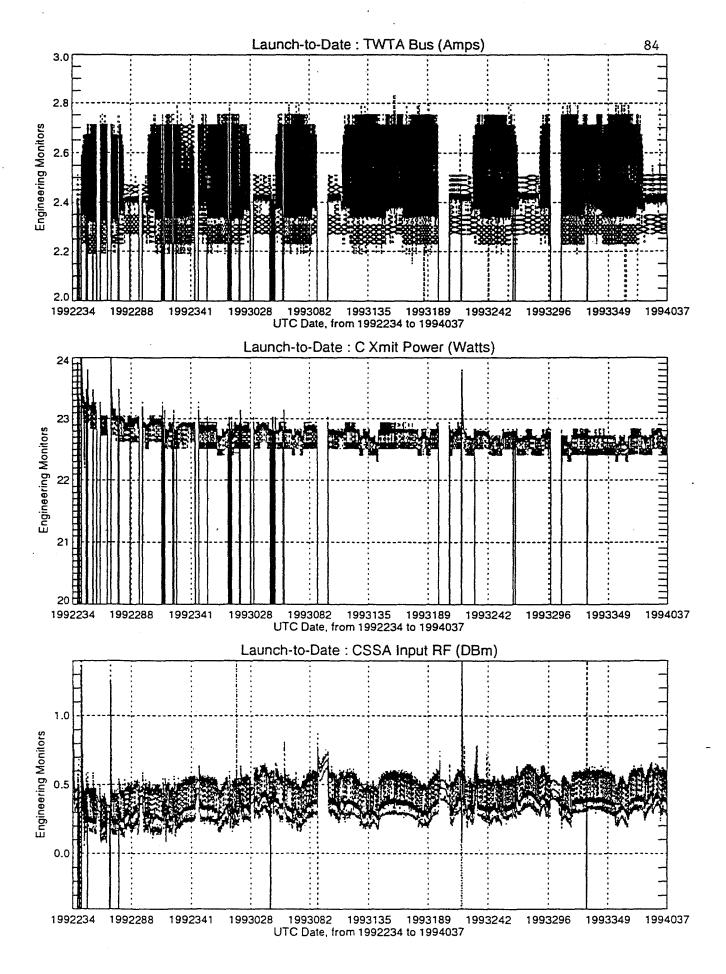
Figure 7.2 Monitor History Plots

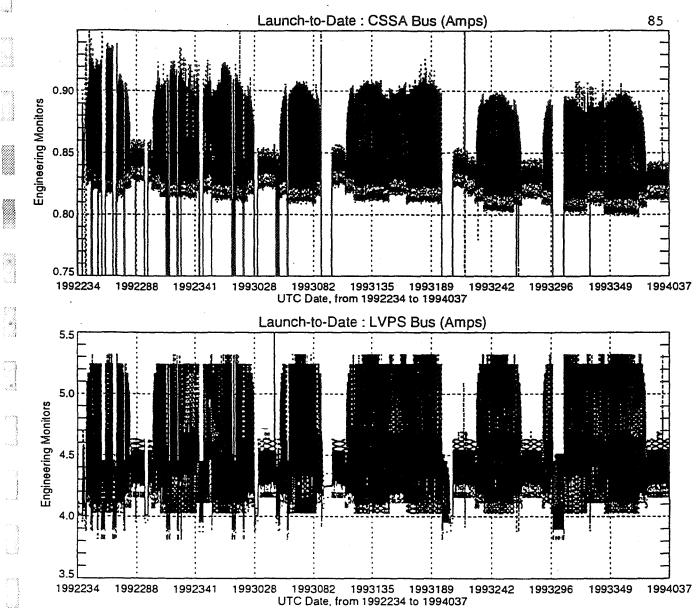


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Monitor History Plots (cont.) Figure 7.2

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Parameter	Ambient Limited Performance Tests at Launch Site (May 10, 1992)		On-Orbit	
	Minimum	Maximum	Minimum	Maximum
Volt. Mon LVPS +12 VDC	13.02	13.09	13.05	13.14
(Volts)	· .			
Volt. Mon LVPS +28 VDC (Volts)	29.93	30.37	30.10	30.40
Volt. Mon LVPS +15 VDC (Volts)	15.59	15.66	15.59	15.68
Volt. Mon LVPS -15 VDC (Volts)	-15.40	-15.35	-15.44	-15.35
Volt. Mon LVPS +5 VDC 5% (Volts)	4.87	4.89	4.86	4.92
Volt. Mon LVPS +5 VDC 1% (Volts)	5.08	5.11	5.10	5.10
Volt. Mon LVPS -5.2 VDC (Volts)	-5.53	-5.52	-5.53	-5.52
Volt. Mon LVPS -6 VDC (Volts)	-6.16	-6.14	-6.15	-6.15
Analog - Ku MTU Transmit Pwr. Monitor (Watts)	21.11	21.36	18.00	22.00
Analog - Ku TWTA Cathode Voltage (Volts)	4185.	4205.	4187.	4193.
Analog - Ku TWTA Cathode Current (Amps)	0.0350	0.0369	0.0363	0.0371
Analog - Ku TWTA Helix Current (Amps)	0.00083	0.00095	0.00087	0.00095
Analog - Ku TWTA Bus Current (Amps)	2.33	2.59	2.33	2.50
Analog - CMTU Transmit Power Monitor (Watts)	21.88	22.94	22.45	23.00
Analog - C SSA RF Power (dBm)	0.18	0.53	0.22	0.44
Analog - C SSA Bus Current (Amps)	0.835	0.951	0.815	0.860
Analog - LVPS Altimeter Bus Current (Amps)	4.05	4.79	3.90	4.70

# Table 7.2Comparison of Pre-Launch and On-Orbit Engineering ParametersDuring TRACK Mode

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### 7.2.2 Ku Transmit Power

The Ku band transmit peak power is monitored by a power detector monitor located in the Ku band Microwave Transmission Unit. It generates a voltage proportional to power which is inserted in the engineering telemetry data stream.

The time history of the Ku Transmit Power shows that the power levels have generally remained between 20.7 and 21.6 watts. Minor variations in Ku Transmit Power can be attributed to tracking versus acquisition cycles for which the PRF is different. Variations in PRF cause variations in Power monitor output.

In late November 1992 (days 326-339 in Figure 7.2), larger Ku transmit power variations were indicated. In general, these variations were less than 1 watt (21.25 to 20.70 W) but, occasionally, the monitor dipped all the way to 18 watts for a very short period. After about 2 weeks of these occasional "glitches", they have not reappeared.

When the glitches occurred they were not correlatable with any other altimeter measurement. The Helix current would be the best monitor for verifying power changes, but due to its noisy characteristics, it could not be correlated with the power drop. Also, some of the small changes occurred during calibration modes but, due to their relatively small magnitude (less than 0.1 dB), they could not be positively correlated with AGC changes.

Three possible causes for the glitches are: 1) the power monitor was erroneously indicating these changes possibly due to noise; 2) the telemetry or telemetry interface was erroneously reporting the power; 3) the output power of the TWTA was indeed varying. Although we cannot positively prove either of the above, it is believed that: 1) it probably was not the monitor since we had not seen the glitches before, they have cleared up, and the circuitry does not seem to lend itself to this phenomena; 2) the telemetry is an unlikely source since no other values appear to have been affected; 3) most likely there was real variation in the TWTA power. The TWTA could have become slightly gassy creating minor arcs, corona could have been occurring due to out-gassing of some other part of the spacecraft, or several other typical TWTA phenomena could have been occurring. This condition has not recurred, but it still remains a concern. The cause is unknown and close monitoring continues.

The minor variations in Ku Transmit Power can be attributed to tracking modes versus acquisition mode where the PRF is different. Variations in PRF cause variations in Power monitor output.

### 7.2.3 TWTA

The time history of the TWTA Cathode Voltage, TWTA Cathode Current, TWTA Helix Current and TWTA Bus Currents, as depicted in Figure 7.2, illustrates that these values have remained steady throughout the mission. The times of larger noise, e.g., days 75-105 and days 125-155, correspond to periods of occultation.

### 7.2.4 C Transmit Power

The time history of the C Transmit Power shows a linear decrease during the first 90 days, from 23.10 watts to 22.80 watts. This decrease may be due to the characteristics of the GaAs FETS used, and may be a long-term trend towards stabilization. Since that time, there has been a small linear decrease of an additional 0.20 watts. A long term stabilization effect was expected before launch based on test data and known GASFET characteristics.

### 7.2.5 CSSA

The CSSA Input RF Power time history displays a trend similar to the C Transmit power, but in the opposite direction. There has been an increase of about 0.2 dBm since the beginning of the on-orbit mission. Most of the observed short-term variations in CSSA Input RF are probably due to temperature variations in the power monitor located inside the CSSA; it is not a very precise monitor.

The CSSA Bus Current has remained steady throughout the mission, with the noisier amperage values occurring during periods of occultation when the spacecraft bus voltage varies more. Bus current inversely follows the spacecraft bus voltage.

### 7.2.6 LVPS Bus Current

The LVPS Bus Current has remained steady, becoming noisier during periods of occultation.

### 8.0 SYNOPSIS OF ENGINEERING ASSESSMENT

Data have been presented showing that the range noise is within the instrument specifications. The two-frequency altimeter ionospheric corrections do behave as expected, temporally and geographically. Significant waveheight estimation is also consistent with expectations, and the track acquisition time is adequate.

The NASA altimeter's Ku radar backscatter cross-section  $\sigma^{0}(Ku)$  is apparently shifted by about 0.7 dB relative to values observed by Geosat, but this may be partly due to the Geosat data processing's not having taken into account Earth curvature. The NASA altimeter's C-band cross-section  $\sigma^{0}(C)$  is about 3.4 dB higher than  $\sigma^{0}(Ku)$ .

The internal calibration mode has shown very small drifts in AGC, and several calibration-based changes in AGC have been entered into the routine ground data processing; these AGC drifts are to be expected, and are well within a reasonable drift magnitude.

A number of SEUs have been observed and the altimeter has automatically recovered from most of these. In nine of these, manual intervention was required; the ground monitoring has been improved to reduce the amount of data lost from anomalous SEU occurrences.

The NASA altimeter is responding correctly to all commands which it has received. The programmability has been successfully demonstrated by an uploaded software patch to improve SEU recovery. The altimeter's routine operation now uses an uploaded parameter file, and the current parameter set's values are included in this report.

There were a variety of small waveform effects produced by the altimeter's digital filter bank, and these have been described and values estimated for the size of the effects. The waveform samplers are performing within the specification of stability to within 1% of the waveform maximum, and no significant changes in the waveform effects have been detected over the past year and a half of operation of the NASA altimeter.

The engineering monitors (temperatures, voltages, currents, and powers) appear to indicate generally nominal performance, within specifications, although one 14-day period (1992, days 326-339) of Ku transmit power variations was described, in Section 7.2.2.

In brief, this report concludes that Side A of the NASA Radar Altimeter is fully operational and within prelaunch specifications. There are no indications of developing problems, and the performance requirements are being met. The Wallops Flight Facility TOPEX personnel are continuing to monitor and analyze the altimeter's performance.

# Other Documents in this Series

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Volume 2	WFF Topex Software Documentation Overview, May 1999 (Published May 2003)
Volume 3	WFF TOPEX Software Documentation Altimeter Instrument File (AIF) Processing, October 1998 (Published July 2003)
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REPOR	T DOCUMENTATIO	N PAGE		Form Approved OMB No. 0704-0188		
Public reporting burden for this collection of in gathering and maintaining the data needed, a	nformation is estimated to average 1 hour per	r response, including the time for	reviewing ins	tructions, searching existing data sources.		
collection of information, including suggestion Davis Highway, Suite 1204, Arlington, VA 22	ns for reducing this burden, to Washington He	adquarters Services, Directorate	for informatio	n Operations and Reports, 1215 Jefferson 1		
1. AGENCY USE ONLY (Leave blar		3. REPORT TYPE A				
	February 1994	Technical Me				
4. TITLE AND SUBTITLE	(TOPEX) Software Docu	ument Sories		DING NUMBERS		
	ltimeter Engineering A		t	n bir allan an a		
6. AUTHOR(S)			- data d	Code 972		
D.W. Hancock III, G.S	. Hayne, C.L. Purdy, 1	R.L. Brooks				
7. PERFORMING ORGANIZATION	NAME(S) AND ADDRESS (ES)		8. PEF	ORMING ORGANIZATION		
NASA Radar Altimeter				ORT NUMBER		
Observational Science						
Laboratory for Hydros			2003	-03169-0		
GSFC/Wallops Flight F Wallops Island, VA 2			1			
9. SPONSORING / MONITORING A		(ES)		ONSORING / MONITORING ENCY REPORT NUMBER		
National Aeronautics and S	pace Administration		NAGA	TM-2003-212236,		
Washington, DC 20546-000	-		Vol.			
11. SUPPLEMENTARY NOTES			<u> </u>			
	W. Hancock III, NASA ps Island, VA 23337	GSFC/Wallops Fl	ight Fa	acility		
12a. DISTRIBUTION / AVAILABILITY	STATEMENT		12b. DI	STRIBUTION CODE		
Unclassified - Unlimited			ſ			
Subject Category:42		r				
	ASA Center for AeroSpace 1 , Linthicum Heights, MD 21					
		.090; (301) 021-0390.				
13. ABSTRACT (Maximum 200 word	s)					
This document describes the GSFC/WFF analysis of the on-orbit engineering data from the TOPEX radar altimeter, to establish altimeter performance. In accordance with Project guidelines, neither surface truth nor precision orbital data are used for the engineering assessment of the altimeter. The use of such data would imply not only a more intensive and complete performance evaluation, but also a calibration. Such evaluations and calibrations are outside the scope of this document and will be presented in a separate Verification Report.						
14. SUBJECT TERMS				15. NUMBER OF PAGES		
IT, COULOI ILANIG				95		
				16. PRICE CODE		
17. SECURITY CLASSIFICATION OF REPORT	18. SECURITY CLASSIFICATION OF THIS PAGE	19. SECURITY CLASSIN OF ABSTRACT	FICATION	20. LIMITATION OF ABSTRACT		
Unclassified	of the time of Abolinat			UL		

Standard Form 298 (Rev. 2-89)

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