

1 Shocked Quartz in Polymict Impact Breccia from the
2 Upper Cretaceous Yallalie Impact Structure in Western
3 Australia

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14 **ABSTRACT**

15 Yallalie is a ~12 km diameter circular structure located ~200 km north of Perth,
16 Australia. Previous studies have proposed that the buried structure is a complex impact
17 crater based on geophysical data. Allochthonous breccia exposed near the structure has
18 previously been interpreted as proximal impact ejecta, however no diagnostic indicators
19 of shock metamorphism have been found. Here we report multiple (27) shocked quartz

20 grains containing planar fractures (PFs) and planar deformation features (PDFs) in the
21 breccia. The PFs occur in up to 5 sets per grain, while the PDFs occur in up to 4 sets per
22 grain. Universal stage measurements of all 27 shocked quartz grains confirms that the
23 planar microstructures occur in known crystallographic orientations in quartz
24 corresponding to shock compression from 5 to 20 GPa. Proximity to the buried structure
25 (~4 km) and occurrence of shocked quartz indicates that the breccia represents either
26 primary or reworked ejecta. Ejecta distribution simulated using iSALE hydrocode
27 predicts the same distribution of shock levels at the site as those found in the breccia,
28 which supports a primary ejecta interpretation, although local reworking cannot be
29 excluded. The Yallalie impact event is stratigraphically constrained to have occurred in
30 the interval from 89.8-83.6 Ma based on the occurrence of Coniacian clasts in the breccia
31 and undisturbed overlying Santonian to Campanian sedimentary rocks. Yallalie is thus
32 the first confirmed Upper Cretaceous impact structure in Australia.

33 INTRODUCTION

34 Impact cratering is a widespread geological process throughout the solar system.
35 The presence of shatter cones, shocked minerals, high-pressure phases, and geochemical
36 evidence for the presence of meteoritic material (e.g., French 1998; French and Koeberl
37 2010) has thus far led to the confirmation of about 190 terrestrial impact craters to date
38 (*Earth Impact Database*, 2018). Shock deformation microstructures in minerals form
39 when the shock wave produced by hyper-velocity impact travels through target rocks
40 (Stöffler and Langenhorst 1994 and references therein). Quartz is perhaps the most widely
41 reported shocked mineral in terrestrial impactites. Shock microstructures in quartz such as
42 planar fractures (PFs) and planar deformation features (PDFs) record shock pressures

43 from 5-30 GPa and are readily identified using a petrographic microscope (e.g., Stöffler
44 and Langenhorst 1994; Huffman and Reimold 1996; Ferrière and Osinski 2013). Here we
45 report the first diagnostic evidence of shock deformation in quartz grains in allochthonous
46 breccia from the Yallalie structure in Western Australia.

47 **Regional Geology**

48 The geology of the study area consists of Mesozoic sedimentary rocks, including
49 sandstone, siltstone, shale, and limestone (chalk), of the Dandaragan Trough within the
50 greater Perth Basin (Fig. 1). The Darling Fault lies ~10 km east of the previously
51 proposed rim of the Yallalie structure (Dentith et al. 1999), with the Archean Yilgarn
52 Craton located on the eastern side of the fault (Fig. 2). The Dandaragan Trough is an
53 asymmetrical graben that contains sedimentary successions up to 15 km thick (Harris
54 1994; Mory and Iasky 1996; Timms et al. 2015; Olierook et al. 2015). The trough is the
55 deepest part of the Perth Basin; sedimentary rocks that lie within it were deposited in
56 intracontinental rifts during the breakup of Gondwana and development of a passive
57 margin (Harris 1994; Sircombe and Freeman 1999; Song and Cawood 2000; Veevers et
58 al. 2005).

59 **The Yallalie Structure**

60 The Yallalie structure was first described in an impact context in a 1992
61 Meteoritical Society abstract (Dentith et al. 1992). Seismic surveys conducted across the
62 structure by Ampol Exploration Ltd. in 1988 and 1990 showed a zone of disruption that
63 extends to ~1500 m below the surface, with an abrupt contact between the structure and
64 overlying sedimentary units (Dentith et al. 1999; Hawke 2004; Hawke et al. 2006). A 3 to
65 4 km wide central uplift was identified by Dentith et al. (1999) from seismic data. A small

66 positive gravity anomaly of 30 gu (3 mGal) was described by Dentith et al. (1999) based
67 on a north-south transect across the center of the Yallalie structure, and attributed to
68 uplifted bedrock. Airborne magnetic surveys reveal a 12 km diameter circular feature,
69 consisting of concentric positive magnetic anomalies, further suggesting the presence of a
70 complex structure (Hawke et al. 2006). Two petroleum exploration wells, Cypress Hill-1
71 (Higgins 1988) and Yalallie-1 (Economo 1991), were drilled within the structure (Fig. 2).
72 Cypress Hill-1 (30°27'51" S, 115°48'42" E) was drilled in 1988 and penetrated 990 m
73 along the southeast rim, intersecting a rotated fault block. Yallalie-1 (30°20'40" S,
74 115°46'16" E) was drilled in 1990 and penetrated 3322 m into the center of the structure.
75 In both cases, the wells were not continuously cored, therefore geological contacts were
76 interpreted from wireline log data (Bevan 2012).

77 The Yallalie structure is thus buried below ~100-300 m of undisturbed sediment,
78 which includes thin layers of lacustrine sediment, laterite, and recent aeolian sand
79 (Dentith et al. 1999; Hawke et al. 2006). The target rocks are all sedimentary, and, in
80 general, consist of poorly sorted, interbedded, marine and fluvial sandstone, siltstone,
81 shale, chalk, and glauconitic sandstone (greensand) (Belford 1959; Playford et al. 1976;
82 Moray and Iasky 1996; Timms et al. 2015). Rocks affected by the proposed impact event
83 include the Jurassic Yarragadee Formation (below 320 m current depth), which is overlain
84 by the Cretaceous Parmelia Formation (Warnbro Group), and Leederville Formation
85 (Coolyena Group), respectively (Dentith et al. 1999; Olierook et al. 2015).

86 Allochthonous breccia, informally named the "Mungedar Breccia" (Bevan 2012),
87 outcrops 4 km west of the rim of the buried structure (Fig. 2), and is the only exposed
88 geological unit known that has been interpreted to have originated from the proposed

89 Yallalie impact event (e.g., Dentith et al. 1999). The breccia sits unconformably on Upper
90 Cretaceous siliciclastic rocks, and no other comparable anomalous breccia occurrences are
91 known from elsewhere in the Perth Basin (Dentith et al. 1999; Bevan et al. 2004; Bevan
92 2012). The breccia contains up to meter-scale blocks of sedimentary rocks, and has a
93 maximum thickness estimated at ~30 m. The breccia was previously interpreted as an
94 allochthonous polymictic breccia that formed from material ejected during the proposed
95 Yallalie impact event (Dentith et al. 1999).

96 Fractured quartz grains from cuttings in the Yallalie-1 borehole (Dentith et al.
97 1999) and in exposed Mungedar Breccia (Bevan, 2012) have been described previously,
98 however, none of the planar microstructures in those grains were indexed by universal
99 stage. Diagnostic evidence of shock microstructures has thus far not been documented in
100 any materials associated with the Yallalie structure. Despite the convincing geophysical
101 anomalies and reports of fractured quartz grains, until now no evidence of shock
102 deformation has been reported that could substantiate an impact origin of the Yallalie
103 structure to be confirmed.

104 The relative age of the structure is constrained primarily by stratigraphic relations
105 interpreted from cuttings in the Yallalie-1 core, along with tentative correlation of clasts
106 found within the allochthonous breccia exposed in the area (Dentith et al. 1999). The
107 depositional age of the regional Gingin Chalk (Santonian, 86.3-83.6 Ma, ages from Cohen
108 et al. 2013) has been previously cited as a minimum age constraint on the formation of the
109 Yallalie structure, as clasts of Gingin Chalk have not been identified in the Mungedar
110 breccia or from the two cores within the structure (Dentith et al. 1999). The presence of
111 clasts of the Molecap Greensand Formation of the Upper Cretaceous Coolyena Group

112 (Coniacian, 89.8-86.3, ages from Cohen et al. 2013) in the breccia currently provides the
113 best estimate for the maximum age constraint of the structure (Dentith et al. 1999).

114 **Shock Deformation Features in Quartz**

115 Quartz is widely used to document shock deformation, as it is abundant in crustal
116 rocks and is rather resistant to alteration. Planar fractures (PFs) form at relatively low
117 pressures (~5-10 GPa) and consist of open fractures with spacings from 10-15 μm that
118 form in specific crystallographic orientations (e.g. Stöffler and Langenhorst 1994; French
119 and Koeberl 2010; Poelchau and Kenkmann 2011). Quartz grains with multiple sets of
120 planar fractures have been experimentally produced during shock recovery experiments,
121 and form when pressures exceed the hugoniot elastic limit at ca. 5-10 GPa (Poelchau and
122 Kenkmann, 2011). Shocked quartz grains containing well-developed planar fractures have
123 been used to help confirm many impact events, including at the Rock Elm (French et al.
124 2004), Keurusselkä (Ferrière et al. 2010b), Hummeln (Alwmark et al. 2015), Saqqar
125 (Kenkmann et al. 2015) and Decorah (French 2018) impact structures and elsewhere (see
126 Poelchau and Kenkmann, 2011). We note that within the impact cratering community
127 some workers question whether or not multiple sets of PFs in quartz represent diagnostic
128 evidence of shock-deformation, as quartz grains with PFs (i.e., generally one single set,
129 rarely more) have been reported from non-impact settings (see discussion in French and
130 Koeberl, 2010). To avoid potential ambiguity in interpreting the significance of fractures
131 in putative shocked quartz grains, in this study we adhere to a strict criteria for
132 identification of shock-produced planar fractures in quartz. Our criteria required that: a)
133 grains must contain one or more sets of evenly spaced, parallel planar fractures that are
134 pervasive across the grain, and b) the orientation of each set of planar fractures must

135 correspond to crystallographic planes that have been demonstrated to form shock features
136 previously.

137 Planar deformation features (PDFs) form at higher pressures (10-30 GPa) and
138 consist of narrow, individual lamellae of amorphous SiO₂ in parallel (multiple) sets
139 spaced 2-10 µm apart. Similar to PFs, the orientations of PDFs are crystallographically
140 controlled. The high-pressure polymorphs coesite and stishovite, as well as diaplectic
141 quartz glass and melted quartz (i.e., lechatelierite) may also form at elevated pressures and
142 temperatures, depending on the properties of the target rocks, and local impact conditions
143 (Stöffler and Langenhorst 1994). During reversion from high temperature conditions,
144 silica can also form a distinctive texture called “ballen silica” which is commonly
145 observed in impactites (Ferrière et al. 2009a, 2010a).

146 **SAMPLES AND METHODS**

147 For this study, samples of Mungedar Breccia were collected from two different
148 outcrops located approximately 4 km west of the rim of the Yallalie structure (Fig. 2).
149 The breccia samples were selected to search for microstructural evidence of shock
150 deformation in minerals, in both rock and mineral clasts, as well as grains in the clastic
151 matrix. The samples include both clast-poor and clast-rich polymictic breccia varieties.
152 The breccia consists of green, tan and brown coloured rocks that contain a variety of
153 different clasts in a quartz-rich, sand-dominated matrix. Crude layering is locally visible
154 in outcrop (Fig. 3) and in thin section.

155 Samples were cut into slabs and impregnated with epoxy prior to preparing
156 polished thin sections (Fig. 4). A search for shocked minerals was conducted using a
157 petrographic microscope and scanning electron microscopy (SEM) at Curtin University

158 on clast-rich breccia sample 16YA09 (nine thin sections) and clast-poor breccia sample
159 16YL01 (two thin sections) (Fig. 5). In addition to searching for shocked minerals, a
160 systematic survey of clasts in the clast-rich breccia sample was conducted. Optical
161 microscopy was used to survey the sections for shocked minerals, and energy dispersive
162 spectroscopy with an SEM was used to determine the elemental composition of both
163 lithic clasts and grains with possible shock features. Two grains of quartz (Q4 and Q6)
164 were further mapped by electron backscatter diffraction (EBSD) using a Tescan Mira3
165 field emission SEM at Curtin University to evaluate the microstructure of grains with
166 PDFs and PFs. The EBSD data were collected using the Oxford program Aztec, and
167 maps were prepared with the Tango module in the Oxford software HKL Channel5. The
168 EBSD patterns were indexed using a match unit for quartz from the HKL database
169 (Sands 1969). Quartz grain Q4 was mapped with a 1.75 μm step size, and Q6 was
170 mapped with a 0.5 μm step size. EBSD data are shown in texture component maps,
171 where colors indicate either variations in lattice orientation or the presence of Dauphiné
172 twins (Wenk et al. 2011).

173 Crystallographic orientations of PFs and PDFs in grains of shocked quartz were
174 determined using a four-axis universal stage (u-stage) at the University of Vienna. A
175 Horiba LabRAM HR Evolution Raman microscope, also at the University of Vienna, was
176 used to identify silica polymorphs present in the samples. In addition to searching for
177 shocked quartz, we also conducted a preliminary survey to search for detrital shocked
178 zircon in Yallalie samples (e.g., Cavosie et al 2010), but have thus far not identified any
179 shocked accessory phases.

180 The iSALE shock physics hydrocode (Amsden et al. 1980; Collins et al. 2004;
181 Wünnemann et al. 2006) was used to model the formation of the Yallalie structure in
182 order to determine the shock-level of material within proximal ejecta. The cell resolution
183 in a 2D numerical mesh was 25 by 25 m. The projectile used in the simulation was 1.2
184 km in diameter, impacting Earth at 12 kms⁻¹ vertical speed; this speed also represents
185 faster speeds at moderately oblique impact angles (e.g., Pierazzo and Melosh 2000). The
186 impactor was modelled using a dunite analytical equation of state (ANEOS) (Benz et al.
187 1989) representative of a stony asteroid. The most voluminous target rocks at Yallalie are
188 the Jurassic Yarragadee Formation, and potentially the underlying Cattamarra Coal
189 Measures (Higgins 1988; Economo 1991; Olierook et al. 2015; Timms et al. 2015).
190 Quantitative analysis of both of these formations in nearby Gingin-1 and Gingin-2 wells
191 illustrate that they are dominated by quartz arenite to arkose with minor heterolithic
192 siltstone and shale, and variable proportions of water-filled porosity, typically up to
193 ~30% at 500 m dropping systematically to <10% at ~3.5 km depth (Delle Piane et al.
194 2013; Timms et al. 2015). Therefore, the target was simulated using the ANEOS equation
195 of state for granite (Pierazzo et al. 1997). Although granite is not present among the
196 target rocks where the Yallalie structure formed, it may more accurately represent the
197 target rock mineralogy than monomineralic quartzite. Granite is also a validated material
198 for modelling crater formation in hydrocodes (Pierazzo et al. 1997). We acknowledge
199 some level of uncertainty of the material properties of Perth Basin rocks at the time of the
200 Yallalie impact.

201 **RESULTS**202 **Petrology of the Mungedar Breccia**

203 The matrices of both clast-rich and clast-poor breccia samples are similar, and are
204 mainly composed of angular to sub-rounded quartz grains with minor alkali feldspar and
205 micas. Most matrix quartz grains range in size from 50 to 150 μm , although larger grains
206 are present. Lithic clasts within the clast-rich breccia are angular to sub-rounded, with at
207 least four different lithologies present, including mudstone, glauconitic sandstone,
208 siltstone, and palaeosol (Table 1). Mudstone is the most abundant lithology among lithic
209 clasts surveyed, comprising 36 of 66 clasts. Mudstone clasts are brown, very fine-grained
210 and range in size from 0.1 to 1.2 cm (Table 1). Some mudstone clasts contain deformed
211 biotite grains (Fig. 6). A total of four biotite grains with kink-bands were identified in two
212 different mudstone clasts. Glauconitic sandstone clasts (greensand) represent the second
213 most abundant lithic clast type, comprising 18 of 66 clasts surveyed (Table 1). Greensand
214 clasts range in size from 0.2 to 2 cm across, and have a distinctive green colour in hand
215 sample (Figs. 3A, 4A). Greensand clasts consist of aggregates of rounded $\sim 1 \mu\text{m}$
216 glauconite grains with minor amounts of detrital quartz and feldspar. The mudstone and
217 greensand clasts are both texturally heterogeneous, and locally contain large quartz and
218 feldspar grains that are easily distinguished from grains forming the matrix. Two other
219 lithologies were encountered. One is represented by a single 3 mm-long quartz-rich
220 siltstone clast (Fig. 5B), whereas the other is represented by a single 40 mm-long,
221 elongate, dark orange to brown palaeosol clast (Table 1). Minor veins are also present.

222 Shocked Quartz Grains - optical imaging

223 A total of twenty-one grains with planar fractures (PFs) were identified. Eighteen
224 sub-rounded to angular quartz grains with PFs were identified in the nine thin sections of
225 the clast-rich breccia (sample 16YA09), and three in the clast-poor breccia (sample
226 16YL01). The PF bearing grains range in diameter from 150 μm to 1.2 mm. Fifteen occur
227 as single grains in the matrix of the breccia, and the other six grains were found within
228 mudstone clasts. The PFs are well-developed and cut across the entire grain in most cases,
229 but are also localised within parts of the grains in others (e.g., Figs. 7,8). The PFs within
230 each grain are typically spaced between 10 to 20 μm apart. PFs in all grains were indexed
231 using the u-stage and found to be oriented along known crystallographic orientations for
232 PFs in quartz (Table 2). Each grain contains two to five sets of PFs that are oriented along
233 four main crystallographic planes (Table 2). Some PFs appear slightly curved (i.e., not
234 perfectly planar) in crystal-plastically strained quartz grains. However, these PFs are
235 locally planar and appear to be strictly crystallographically controlled; their deflections
236 appear to follow crystallographic orientation variations in the host grain.

237 A total of six quartz grains with decorated PDFs were identified; all from the
238 matrix of breccia sample 16YA09 (Fig. 9). The grains are sub-rounded to angular and are
239 100 μm to 4 mm in size, respectively. PDFs in all grains (9 sets) were also indexed using
240 the u-stage and found to be oriented along known crystallographic orientations for PDFs
241 in quartz (Table 2). One grain contains four orientations of PDFs, however only two are
242 visible on the untilted surface under the optical microscope (Fig. 9F). During close
243 inspection of the planar microstructures at high magnification, two additional PDF

244 orientations were identified using the u-stage; the brownish appearance of the grain is the
245 result of the dense network of PDFs.

246 **Shocked Quartz Grains - orientation mapping**

247 Orientation mapping using EBSD was conducted on part of a large quartz grain
248 (Q4) that contains four orientations of planar fractures (Figs. 7A, 10A). The orientation
249 map reveals two types of orientation domain; domains consisting of the host orientation,
250 and domains with irregular boundaries in a Dauphiné twin orientation that are misoriented
251 from the host grain by $60^\circ/\langle 0001 \rangle$ (Fig. 10B). The host and twin domains preserve up to
252 10° of distributed lattice misorientation across the area analyzed. The Dauphiné twin
253 domains are pervasively distributed throughout the mapped area, and in many cases their
254 boundaries terminate against the dominant set of $\{11\bar{2}2\}$ planar fractures.

255 A second orientation map was made on a quartz grain with decorated (0001) PDFs
256 that were measured by u-stage (Fig. 10C). In contrast to the PF-bearing grain (Fig. 10B),
257 the PDF-bearing grain contains relatively few Dauphiné twins (Fig. 10D). The orientation
258 map reveals $<10^\circ$ of misorientation throughout the grain, much of which appears to be
259 accommodated along low-angle boundaries (LABs), which are the dominant top-left to
260 bottom-right oriented features visible in the map (Fig. 10D). Some of the LABs are
261 planar, whereas others are irregular; the latter LABs locally form the boundaries of sub-
262 grains (top left area of Fig. 10D). The dominant set of top-left to bottom-right planar
263 LABs are conspicuously parallel to the measured (0001) PDF orientation (Fig. 10D), and
264 thus may have formed in association with the PDFs during impact.

265 Ballen Silica

266 One elongate, ~1 mm-long silica grain in clast-poor breccia sample 16YL01 was
267 found to exhibit a well-developed ballen texture (Fig. 11). The ballen texture is developed
268 throughout the grain, but is most visible near the grain margin where the presence of
269 secondary mineral films makes the texture more prominent. The ballen grain exhibits a
270 uniform extinction in cross-polarized light (Fig. 11B). MicroRaman spectra collected
271 from several points on the grain yield similar results that all show a dominant peak at 464
272 cm^{-1} that is associated with an O-Si-O bending mode (Ling et al. 2011), as well as other
273 minor peaks at 127, 205, 264, 355, and 805 cm^{-1} . These peaks are all characteristic of
274 unshocked α -quartz (McMillan et al. 1992), and no evidence of other silica polymorphs
275 was detected. The uniform extinction of the ballen in cross-polarized light (Fig. 11B) and
276 the fact that the grain is α -quartz further identifies the grain as type II ballen according to
277 the classification scheme of Ferrière et al. (2009a, 2010a). Type II ballen grains have only
278 been reported in rocks associated with meteorite impacts (Ferrière et al. 2009a).

279 iSALE Numerical Modelling of Yallalie Crater Formation

280 Impact simulation of the Yallalie crater formation produced a good match with
281 sub-surface structural maps from published seismic surveys and borehole stratigraphy
282 (Dentith et al. 1999; Hawke et al. 2006) (Fig. 12). The simulation shows a central uplift
283 that is approximately 4 km in diameter and vertically displaces stratigraphy by up to 2
284 km. The final crater depth is ≤ 1 km. Fault traces previously interpreted from seismic data
285 (Hawke et al. 2006) match well with fault traces suggested by the simulation, within a 5
286 km radius (Fig. 12). The iSALE simulation further suggests that the crater rim diameter

287 may be larger than previously estimated, approximately 14 km (if measured at pre-impact
288 surface level), or a 16 km diameter if measured rim-to-rim.

289 The spatial distribution and shock-level provenance of material forming the ejecta
290 deposit are calculated from the iSALE simulation (Fig. 12). The Mungedar Breccia
291 sample site (Fig. 2) corresponds to an approximate 10 km distance from the centre of the
292 crater. In the iSALE model, the ejecta blanket geometry at the breccia sample site was
293 taken to be 200 m wide and 100 m deep (as denoted by the grey box in Fig. 12). Each
294 numerical cell was tracked by a mass-less tracer particle, which provides information
295 about the material change with the flow. Analysis of each tracer particle inside the grey
296 box shows that the ejecta at this site is composed of 75% mass that is shocked to peak
297 pressures lower than 5 GPa, 17% of mass shocked to 5-10 GPa, and 8% mass shocked to
298 peak pressures above 10 GPa. Due to computational limitations, some of the mass-less
299 tracer particles are lost during the contact and compression cratering phase from
300 numerical cells that experience the largest tension or extension. Therefore, the 8% of
301 mass shocked at the highest pressures (reported here) may be higher in reality. Further
302 analyses suggested that the proportions of shocked ejecta were moderately insensitive to
303 variations in the size of the sampling area (grey box, Fig. 12); for example, a 100%
304 increase in the width of the sampled area results in a <10% variation in the shock
305 provenance statistics at this locality.

306 **DISCUSSION**

307 **Evidence of Shock Deformation at Yallalie**

308 Documentation of diagnostic shock features in quartz remains one of the most
309 reliable methods available to confirm an impact event (e.g., French 1998; French and

310 Koeberl 2010; Alwmark et al. 2014, Kenkmann et al. 2015; Holm-Alwmark et al. 2018).
311 Within the eleven thin-sections of Mungedar Breccia surveyed, a total of twenty-seven
312 shocked quartz grains were identified, including twenty-one with up to 5 sets of PFs and
313 six with up to 4 sets of PDFs. All of the planar microstructures in the grains were indexed
314 by u-stage, and are in orientations that are diagnostic of shocked quartz. The orientations
315 of indexed PFs (Table 2) record pressures from 5 to 10 GPa, while the orientations of
316 indexed PDFs (Table 2) record pressures from 7.5 to 20 GPa. Of the twenty-seven
317 shocked quartz grains identified, six PF-bearing grains occur in mudstone clasts (e.g., Fig.
318 7C,D), which definitively establishes mudstone as a shocked target rock. We interpret
319 these findings to represent the first diagnostic evidence of hypervelocity processes
320 documented within the Mungedar Breccia, which allows an impact origin for these
321 components to be confirmed. Based on the presence of PFs and PDFs in matrix quartz, we
322 estimate that these grains experienced pressures from 5-20 GPa, whereas quartz grains
323 with PFs in the mudstone clasts experienced pressures of <10 GPa (Stöffler and
324 Langenhorst 1994; Poelchau and Kenkmann 2011; Holm-Alwmark et al. 2018).

325 Shocked quartz grains in Mungedar Breccia are interpreted to have been excavated
326 from the Yallalie structure during the impact event. However, we also consider the
327 possibility that the provenance of shocked quartz and other grains identified in this study
328 (ballen silica, deformed biotite) in Mungedar breccia did not form during the Yallalie
329 impact event, and instead originated as detrital shocked grains (e.g., Cavosie et al. 2010)
330 in siliciclastic rocks from the pre-impact stratigraphy at Yallalie. This scenario is
331 considered unlikely, as no other impact structures are known regionally in the well-
332 characterized stratigraphy of the Perth Basin. The close proximity (~4 km) of breccia

333 outcrops in relation to the rim of the Yallalie structure, the presence of multiple grains of
334 shocked quartz containing PFs and PDFs, a grain of ballen α -quartz, the presence of
335 shocked lithic clasts with both PF-bearing quartz grains and deformed biotite grains, and
336 the absence of comparable polymict breccias regionally, all support the interpretation that
337 the Mungedar Breccia represents impact ejecta from the Yallalie structure. The closest
338 known impact structures with ages older than the basement stratigraphy at Yallalie (i.e.,
339 pre-Jurassic) and that could have contributed detrital shocked grains are all located >500
340 km from Yallalie (Earth Impact Database 2018). Moreover, no large-scale regional fluvial
341 or glaciogenic systems, the latter unlikely during the Cretaceous, are known that could
342 have delivered detrital shocked grains from the older impact structures over the required
343 distances, as has been demonstrated elsewhere (e.g., Erickson et al. 2013; Thomson et al.,
344 2014; Montalvo et al. 2017). It is also highly unlikely that the cm-scale shocked mudstone
345 clasts would survive multiple sedimentary cycles.

346 Deformation microstructures in the PF-bearing quartz grains described here from
347 the Mungedar Breccia are similar to shocked quartz grains with PFs from the Rock Elm
348 impact structure (USA), a 6.5-km-diameter structure formed in Cambrian sandstone
349 (French et al. 2004). PFs in shocked quartz were the principal evidence used to confirm a
350 hypervelocity origin for Rock Elm, as PDFs in quartz have not been reported from that
351 site. We note that an absence or low abundance of quartz grains with PDFs does not
352 preclude discovery of other higher-pressure shock indicators, given that reidite, a high-
353 pressure $ZrSiO_4$ polymorph that forms at >30 GPa, was later reported in polymict
354 sandstone breccia from the central uplift at Rock Elm (Cavosie et al. 2015). Shocked
355 quartz grains with PFs have been reported at an increasing number of impact structures; a

356 review by Poelchau and Kenkmann (2011) lists 26 impact structures where PFs (and
357 feather features, another shock-related feature) have been reported in shocked quartz; PFs
358 in quartz were most recently documented at the newly discovered ~5.5 km diameter
359 Decorah structure in Iowa, USA (French et al. 2018).

360 **Other Evidence of Impact Processes at Yallalie**

361 Additional lines of evidence that are commonly associated with, but not diagnostic
362 of, impact processes were found in this study. These include a quartz grain with type-II
363 ballen texture, kink-bands in biotite, and Dauphiné twins in shocked quartz grains.

364 Ballen silica is indicative of impact, although it is not considered as diagnostic
365 evidence of shock metamorphism. Ballen silica (shown here to be Type II, α -quartz)
366 forms upon cooling from β -quartz formed either by a solid-solid transformation involving
367 diaplectic quartz glass, or a solid-liquid transformation involving lechatelierite (Ferrière et
368 al. 2009a); either could have originated in melt formed during the Yallalie impact event,
369 although no occurrences of impact melt have been reported that could better establish its
370 provenance.

371 Biotite grains with kink-bands (Fig. 6) provide evidence for the deformation of
372 biotite, and so-called shocked mica grains have long been recognized in bedrock from
373 established impact structures (e.g., Schneider 1972). However, kink-bands in biotite alone
374 do not provide diagnostic evidence of shock deformation, as they also form during
375 endogenic deformation (e.g., Misra and Burg 2012). We note that in our survey of 36
376 mudstone clasts (Table 1), deformed biotite grains were only found in two clasts, both of
377 which also contain shocked quartz grains with PFs. These two clasts, tentatively identified

378 as Jurassic Yarragadee Formation, currently represent the only documented shocked target
379 rock fragments from the Yallalie structure (Fig. 7C).

380 Orientation maps show that both PF- and PDF-bearing shocked quartz grains from
381 Yallalie contain Dauphiné twins (Fig. 10). Dauphiné twins in quartz can form under low
382 stresses and stress rates that are associated with a wide range of endogenic processes (e.g.,
383 Hartley and Wilshaw 1973; Rahl et al. 2018). Therefore, Dauphiné twins do not provide
384 diagnostic evidence of shock deformation per se, but they have been reported in PDF-
385 bearing shocked quartz grains from the Vredefort Dome (Wenk et al. 2011), Charlevoix
386 (Trepmann and Spray 2007), and Rochechouart (Hamers et al. 2017) impact structures.
387 Termination of Dauphiné twin boundaries against PFs in the quartz grains in this study
388 suggests that they formed late during the impact deformation sequence, with their
389 propagation impeded against pre-existing PFs.

390 **The Mungedar Breccia: primary or secondary ejecta?**

391 The Mungedar Breccia is here shown to be a product of impact due to the presence
392 of shocked quartz grains, as well as lithic clasts (mudstone) that contain shocked quartz.
393 However, the question of whether the breccia represents a primary ejecta deposit, or one
394 that has been subsequently re-worked, remains unresolved; below we discuss various lines
395 of evidence to further explore the nature of the breccia.

396 The crude layering observed in breccia outcrops (Fig. 3) may have resulted from
397 primary deposition of the ejecta into a shallow water column. However, we cannot
398 exclude the possibility that the layering may also have resulted from subsequent
399 reworking. Shock levels experienced by clasts in the breccia, as recorded by deformation
400 microstructures in quartz (i.e., 5-20 GPa), are consistent with results of model simulations

401 of shock level in the primary ejecta blanket at the locations where breccia was sampled,
402 which is in turn consistent with a primary origin for the ejecta. However, if localized
403 reworking of the breccia occurred after deposition, the distribution of shocked material
404 may not have changed significantly. Likewise, the overall low abundance of material
405 shocked above 10 GPa in the modelled ejecta blanket (<10%) appears to agree with the
406 overall low abundance of shocked quartz grains identified (twenty-seven grains in 11 thin
407 sections). However, local reworking could also have diluted an originally higher
408 abundance of shocked minerals in the analyzed material.

409 The presence of a grain of ballen silica (Fig. 11) as a mineral clast in a sandstone
410 breccia that clearly did not melt (Fig. 5A) could be interpreted as evidence for reworking
411 of the original host rock (or melt) in which the ballen silica formed. However, flow paths
412 during crater excavation characteristically produce mixtures of both unshocked and highly
413 shocked material in ejecta (Melosh 1989), and so there is no conclusive evidence that the
414 clast-poor breccia sample with ballen silica was reworked.

415 We tentatively interpret the Mungedar Breccia to be a primary ejecta deposit based
416 on the overall good agreement between our petrographic observations and modelling
417 results. However, as noted above, the data do not allow a secondary origin by reworking
418 to be excluded. Regardless of the nature of the breccia (primary vs. secondary), given that
419 the Yallalie structure is buried below 100-300 m of post-impact sedimentary rocks
420 (Hawke et al. 2006), the exposures of the Mungedar Breccia at the present day surface
421 indicate that it may have been deposited on relatively high paleotopography (Dentith et al.
422 1999).

423 **Comparison of Mungedar Breccia with Other Impact Breccia**

424 Breccia comprised primarily of lithic clasts in a quartz-sand matrix has been
425 reported from other comparable-sized impact structures formed in sedimentary target
426 rock. Some of these breccias have similarities with the Mungedar Breccia. Of these, the
427 most similar may be that found at the Wetumpka impact structure in Alabama, USA, a 5-
428 km-diameter structure formed at ~84 Ma in a shallow marine environment (King et al.
429 2015). At the Wetumpka structure, there are deposits of ‘polymict impact breccia’ that
430 outcrop near the center of the structure (King et al. 2015). The Wetumpka surface
431 polymict impact breccia is described as a heterogeneous lithic clast-bearing breccia,
432 including local layering/laminations, with the dominant planar microstructures in
433 shocked quartz in that unit being PFs and feature features, rather than PDFs (King et al.
434 2015). While many aspects of the Wetumpka surface polymict impact breccia appear
435 similar to the Mungedar Breccia, the emplacement mechanisms cited at Wetumpka
436 include gravity-driven slumping and resurge (King et al. 2015). The location of the
437 Mungedar Breccia beyond the rim of the Yallalie impact structure effectively rules out
438 resurge, however, it is possible that slumping may have played a role if the Mungedar
439 Breccia was originally deposited on paleotopographic high surfaces or slopes.

440 **An Upper Cretaceous Impact Structure in Western Australia**

441 The discovery and characterization of PFs and PDFs in quartz grains from
442 allochthonous breccia at the Yallalie structure allows the recognition of Yallalie as an
443 established complex impact structure in Western Australia. The age of the Yallalie
444 impact is currently only constrained by stratigraphic relations to be Upper Cretaceous
445 (probably Santonian), as no materials appropriate for the use of radiometric dating
446 techniques have been discovered thus far. The buried structure appears to be well-

447 preserved overall, as interpreted from geophysical surveys and the overlying post-impact
448 sedimentary rocks (Hawke et al. 2006). This observation has been cited to argue that the
449 Yallalie structure may have formed in a quiescent shallow marine environment and was
450 buried quickly (Dentith et al. 1999; Hawke et al. 2006). Such an environment would be
451 similar to, for example, the Lockne (Sturkell 1998), Rock Elm (French et al. 2004), and
452 Flynn Creek (Schieber and Over 2005) impact structures. Studies have shown that
453 shallow marine impacts can result in rapid crater burial shortly after the impact, limiting
454 erosion and thus enhancing crater preservation (Dypvik et al. 2004). However, conclusive
455 evidence for a terrestrial vs. marine setting for the Yallalie impact remains elusive, and
456 requires further study.

457 **CONCLUSION**

458 Our survey revealed evidence of shock metamorphism within the Mungedar
459 Breccia at the Yallalie structure, resulting in the recognition of Yallalie as a newly
460 established impact structure in Western Australia. The PFs and PDFs found in twenty-
461 seven quartz grains that were characterized using u-stage provide diagnostic evidence of
462 shock deformation, and confirm that these quartz grains and the lithic clasts that contain
463 them experienced shock pressures from 5-20 GPa. Previous geophysical studies reported
464 structural disturbances created by the 12 km diameter Yallalie structure (Hawke 2004).
465 Complementary to existing observations, our iSALE modelling of ejecta geometry and
466 shock characteristics suggests the Yallalie impact structure may be up to 14-16 km in rim
467 diameter. The Earth Impact Database currently lists 27 impact structures within Australia
468 (*Earth Impact Database*, 2018); confirmation of a hyper-velocity impact origin for the
469 Yallalie structure increases the number of confirmed impact structures in Australia to 28.

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482 **REFERENCES**

- 483 Alwmark C., Alwmark-Holm S., Ormö J., and Sturkell E. 2014. Shocked Quartz Grains
484 from the Målingen Structure, Sweden-Evidence for a Twin Crater of the Lockne
485 Impact Structure. *Meteoritics & Planetary Science* 49(6):1076–1082.
- 486 Alwmark C., Ferrière L., Holm-Alwmark S., Ormö J., Leroux H., and Sturkell E. 2015.
487 Impact Origin for the Hummeln Structure (Sweden) and Its Link to the
488 Ordovician Disruption of the L Chondrite Parent Body. *Geology* 43(4):279–282.
- 489 Amsden A. A., Ruppel H. M., and Hirt C. W. 1980. SALE: A Simplified ALE Computer
490 Program for Fluid Flow at all Speeds. Report LA-8095. Los Alamos National
491 Laboratories, N. Mex. 105 pp.

- 492 Benz W., Cameron A. G. W., and Melosh H. J. 1989. The Origin of the Moon and the
493 single-impact hypothesis III. *Icarus* 81:113-131.
- 494 Bevan A. 2012. Yallalie Impact Structure, In: Western Australian Impact Craters, Field
495 Excursion. Excursion Guidebook for the 75th Annual Meeting of the Meteoritical
496 Society, Cairns, Australia.
- 497 Bevan A., Hough R., and Hawke P. 2004. Morphology and Origin of an Allocthonous
498 Breccia Near the Yallalie Structure, Western Australia: Evidence for Subaqueous
499 Impact? *AGC abstract* 227.
- 500 Carter J. D., Lippie S. L and Geological Survey of Western Australia. 1982. Moora,
501 Western Australia: sheet SH/50-10 international index. Geological Survey of
502 Western Australia, Perth, W.A
- 503 Cavosie A. J., Erickson T. M., and Timms N. E. 2015. Nanoscale Records of Ancient
504 Shock Deformation: Reidite (ZrSiO₄) in Sandstone at the Ordovician Rock Elm
505 Impact Crater. *Geology* 43(4):315–318.
- 506 Cavosie A. J., Quintero R. R., Radovan H. A. and Moser D. E. 2010. A Record of
507 Ancient Cataclysm in Modern Sand: Shock Microstructures in Detrital Minerals
508 from the Vaal River, Vredefort Dome, South Africa. *Geological Society of
509 America Bulletin* 122:1968–1980.
- 510 Cockbain A. E. 1990. Perth basin. Geology and Mineral Resources of Western Australia.
511 Western Australia Geological Survey, Memoir, 3:495–524.
- 512 Cohen K. M., Finney S. C., Gibbard P. L., and Fan J.-X. 2013. The ICS International
513 Chronostratigraphic Chart. *Episodes* 36(3):199-204.

- 514 Collins G. S., Melosh H. J., and Ivanov B. A. 2004. Damage and Deformation in
515 Numerical Impact Simulations. *Meteoritics & Planetary Science* 39:217–231.
- 516 Delle Piane C., Esteban L., Timms N. E., and Ramesh Israni, S. 2013. Physical Properties
517 of Mesozoic Sedimentary Rocks from the Perth Basin, Western Australia.
518 *Australian Journal of Earth Sciences* 60(6–7):735–745.
- 519 Dentith M. C., Bevan A. W. R., and Mcinerney K. B. 1992. A Preliminary Investigation
520 of the Yallalie Basin: a Buried 15 km Diameter Structure of Possible Impact
521 Origin in the Perth Basin, Western Australia. *Meteoritical society Meeting*
522 *Abstract*.
- 523 Dentith M. C., Bevan A. W. R., Backhouse J., Featherstone W. E., and Koeberl C. 1999.
524 Yallalie: A Buried Structure of Possible Impact Origin in the Perth Basin,
525 Western Australia. *Geological Magazine* 136(6):619–632.
- 526 Dypvik H., Burchell M. J., and Claeys P. 2004. Impacts into Marine and Icy
527 Environments — A Short Review. In *Cratering in Marine Environments and on*
528 *Ice*, 1–20. Impact Studies. Heidelberg: Springer Berlin.
- 529 Earth Impact Database. 2018. <http://www.unb.ca/passc/ImpactDatabase/>.
530 Accessed September 7, 2018.
- 531 Erickson T. M., Cavosie A. J., Moser D. E., Barker I. R., Radovan H. A., and Wooden J.,
532 2013. Identification and provenance determination of distally transported,
533 Vredefort-derived shocked minerals in the Vaal River, South Africa using SEM
534 and SHRIMP-RG techniques. *Geochimica et Cosmochimica Acta*, 107:170–188.
- 535 Economo M. 1991. Perth Basin EP 321, Western Australia, Yallalie No. 1 Well
536 Completion Report. Ampol Exploration Ltd. (unpublished).

- 537 Engelhardt W. v. and Bertsch W. 1969. Shock-induced Planar Deformation Structures in
538 Quartz from the Ries Crater, Germany. *Contributions to Mineralogy and*
539 *Petrology* 20:203–234.
- 540 Ferrière L. and Osinski R. 2013. Shock Metamorphism. In *Impact Cratering: Processes*
541 *and Products*, edited by Osinski G. R. and Pierrazo E. Oxford: Wiley-Blackwell.
542 pp. 106–124.
- 543 Ferrière L., Koeberl C., and Reimold W. U. 2009a. Characterisation of Ballen Quartz and
544 Cristobalite in Impact Breccias: New Observations and Constraints on Ballen
545 formation. *European Journal of Mineralogy* 21(1):203–217.
- 546 Ferrière L., Morrow J. R., Amgaa T., and Koeberl C. 2009b. Systematic Study of
547 Universal-Stage Measurements of Planar Deformation Features in Shocked
548 Quartz: Implications for Statistical Significance and Representation of Results.
549 *Meteoritics and Planetary Science* 44(6):925–940.
- 550 Ferrière L., Koeberl C., Libowitzky E., Reimold W. U., Greshake A., and Brandstätter F.
551 2010a. Ballen Quartz and Cristobalite in Impactites: New Investigations.
552 *Geological Society of America Special Paper* 465:609–618.
- 553 Ferrière L., Raiskila S., Osinski G. R., Pesonen L. J., and Lehtinen M. 2010b. The
554 Keurusselkä Impact Structure, Finland—Impact Origin Confirmed by
555 Characterization of Planar Deformation Features in Quartz Grains. *Meteoritics &*
556 *Planetary Science* 45(3):434–446.
- 557 French B. M. 1998. *Traces of Catastrophe: A Handbook of Shock-Metamorphic Effects*
558 *in Terrestrial Meteorite Impact Structures*. Houston: Lunar and Planetary Science
559 Institute.

- 560 French B. M., Cordua W. S., and Plescia J. B. 2004. The Rock Elm Meteorite Impact
561 Structure, Wisconsin: Geology and Shock-Metamorphic Effects in Quartz.
562 *Geological Society of America Bulletin* 116(1–2):200–218.
- 563 French B. M., and Koeberl C. 2010. The Convincing Identification of Terrestrial
564 Meteorite Impact Structures: What Works, What Doesn't, and Why. *Earth-*
565 *Science Reviews* 98(1–2):123–170.
- 566 French B. M., McKay R. M., Liu H. P., Briggs D. E.G., Witzke B. J. 2018. The Decorah
567 Structure, Northeastern Iowa: Geology and Evidence for Formation by Meteorite
568 Impact. *Geological Society of America Bulletin* doi:<https://doi.org/10.1130/B31925.1>
- 569 Hamers M. F., Pennock G. M., and Drury M. E. 2017. Scanning Electron Microscope
570 Cathodoluminescence Imaging of Subgrain Boundaries, Twins and Planar
571 Deformation Features in Quartz. *Physics and Chemistry of Minerals* 44:263–275.
- 572 Harris L. B. 1994. Structural and Tectonic Synthesis for the Perth Basin, Western
573 Australia. *Journal of Petroleum Geology* 17(2):129–156.
- 574 Hartley N. E. W., and Wilshaw T. R. 1973. Deformation and Fracture of Synthetic α -
575 quartz. *Journal of Materials Science* 8(2):265-278.
- 576 Hawke P. J. 2004. The Geophysical Signatures and Exploration Potential of Australia's
577 Meteorite Impact Structures. Ph.D. thesis, University of Western Australia, Perth,
578 Western Australia, Australia.
- 579 Hawke P. J., Buckingham A. J., and Dentith M. C. 2006. Modelling source depth and
580 possible origin of magnetic anomalies associated with the Yallalie impact
581 structure, Perth Basin, Western Australia. *Exploration Geophysics* 37:191–196.

- 582 Higgins R. 1988. Cypress Hill 1 - Well completion report for Ampol Exploration Ltd
583 (unpublished).
- 584 Holm-Alwmark S., Ferrière L., Alwmark C. and Poelchau M. H. 2018. Estimating
585 Average Shock Pressures Recorded by Impactite Samples Based on Universal
586 Stage Investigations of Planar Deformation Features in Quartz—Sources of Error
587 and Recommendations. *Meteoritics & Planetary Science* 53:110–130.
- 588 Huffman A. R., and Reimold W. U. 1996. Experimental Constraints on Shock-Induced
589 Microstructures in Naturally Deformed Silicates. *Tectonophysics* 256(1):165–217.
- 590 Kenkmann T., Afifi A. M., Stewart S. A., Poelchau M. H., Cook D. J., and Neville A. S.
591 2015. Saqqar: A 34 km Diameter Impact Structure in Saudi Arabia. *Meteoritics &*
592 *Planetary Science* 50(11):1925–1940.
- 593 King D. T., Morrow J. R., Petruny L. W., and Ormö J. 2015. Surficial Polymict Impact
594 Breccia Unit, Wetumpka Impact Structure, Alabama: Shock Levels and
595 Emplacement Mechanism. *Geological Society of America Special Paper*
596 518:149–164.
- 597 McMillan P. F., Wolf G. H., and Lambert P. 1992. A Raman Spectroscopic Study of
598 Shocked Single Crystalline Quartz. *Physics and chemistry of minerals* 19(2):71–
599 79.
- 600 Melosh H. J. 1989. Impact Cratering: A Geologic Process. New York: Oxford University
601 Press.
- 602 Misra S., and Burg J. P., 2012. Mechanics of Kink-Bands During Torsion Deformation of
603 Muscovite Aggregate. *Tectonophysics* 548: 22–33.

- 604 Montalvo S. D., Cavosie A. J., Erickson T. M., and Talavera C., 2017. Fluvial Transport
605 of Impact Evidence from Cratonic Interior to Passive Margin: Vredefort-derived
606 Shocked Zircon on the Atlantic Coast of South Africa. *American Mineralogist*
607 102(4):813–823.
- 608 Mory A. J., and Iasky R. P. 1996. Stratigraphy and Structure of the Onshore Northern
609 Perth Basin, Western Australia. Geological Survey of Western Australia. Report
610 46.
- 611 Olierook H. K. H., Timms N. E., Wellmann J. F., Corbel S., and Wilkes P. G. 2015. 3D
612 Structural and Stratigraphic Model of the Perth Basin, Western Australia:
613 Implications for Sub-Basin Evolution. *Australian Journal of Earth Sciences*
614 62(4):447–467.
- 615 Pierazzo E., Vickery A. M., and Melosh H. J. 1997. A reevaluation of impact melt
616 production. *Icarus* 127:408–423.
- 617 Pierazzo E. and Melosh H. J. 2000. Understanding Oblique Impacts from Experiments,
618 Observations, and Modelling. *Annual Reviews in Earth and Planetary Science*
619 28:141–167.
- 620 Playford P. E., Low G. H., and Cockbain A. E. 1976. Geology of the Perth Basin,
621 Western Australia. Geological Survey of Western Australia, 124 pp.
- 622 Poelchau, M. H., and Kenkmann T. 2011. Feather Features: A Low-Shock-Pressure
623 Indicator in Quartz. *Journal of Geophysical Research* 116(B2).
- 624 Rahl J. M., McGrew A. J., Fox J. A., Latham J. R., & Gabrielson T. 2018. Rhomb-
625 dominated crystallographic preferred orientations in incipiently deformed quartz

- 626 sandstones: A potential paleostress indicator for quartz-rich rocks. *Geology*
627 46(3):195–198.
- 628 Sands D. E. 1969. Introduction to Crystallography. New York: WA Benjamin.
- 629 Schieber J., and Over D. J. 2005. Chapter 4 Sedimentary Fill of the Late Devonian Flynn
630 Creek Crater: A Hard Target Marine Impact. Understanding Late Devonian And
631 Permian-Triassic Biotic and Climatic Events. 51–69.
- 632 Schneider H. 1972. Shock-Induced Mechanical Deformations in Biotites from Crystalline
633 Rocks of the Ries Crater (Southern Germany). Contributions to Mineralogy and
634 Petrology 37:75–85.
- 635 Sircombe K. N., and Freeman M. J. 1999. Provenance of Detrital Zircons on the Western
636 Australia coastline—Implications for the Geologic History of the Perth Basin and
637 Denudation of the Yilgarn Craton. *Geology* 27(10):879–882.
- 638 Song T., and Cawood P. A. 2000. Structural Styles in the Perth Basin Associated with
639 the Mesozoic Break-up of Greater India and Australia. *Tectonophysics* 317:55e72.
- 640 Stöffler D., and Langenhorst F. 1994. Shock Metamorphism of Quartz in Nature and
641 Experiment: I. Basic Observation and Theory. *Meteoritics & Planetary Science*
642 29(2):155–181.
- 643 Sturkell E. 1998. The Marine Lockne Impact Structure, Jämtland, Sweden: A Review.
644 Geologische Rundschau 87(3):253–267.
- 645 Thomson, O.A., Cavosie, A.J., Moser, D.E., Barker, I., Radovan, H.A., and French,
646 B.M., 2014, Preservation of detrital shocked minerals derived from the 1.85 Ga
647 Sudbury impact structure in modern alluvium and Holocene glacial deposits:

- 648 Geological Society of America Bulletin, v. 126, p. 720–737,
649 <https://doi.org/10.1130/B30958.1>.
- 650 Timms N. E., Olierook H. K. H., Wilson M. E. J., Delle Piane C., Hamilton P. J., Cope
651 P., and Stütenbecker L. 2015. Sedimentary Facies Analysis, Mineralogy and
652 Diagenesis of the Mesozoic Aquifers of the Central Perth Basin, Western
653 Australia. *Marine and Petroleum Geology* 60: 54–78.
- 654 Trepmann C. A., and Spray J. G. 2007. Shock-Induced Crystal-Plastic Deformation and
655 Post-Shock Annealing of Quartz. *European Journal of Mineralogy* 18(2):161.
- 656 Veevers J. J., Saeed A., Belousova E. A., and Griffin W. L. 2005. U–Pb Ages and Source
657 Composition by Hf-Isotope and Trace-Element Analysis of Detrital Zircons in
658 Permian Sandstone and Modern Sand from Southwestern Australia and a Review
659 of the Paleogeographical and Denudational History of the Yilgarn Craton. *Earth-
660 Science Reviews* 68(3):245–279.
- 661 Wenk H., Janssen C., Kenkmann T., Dresen G. 2011. Mechanical Twinning in Quartz:
662 Shock Experiments, Impact, Pseudotachylites and Fault Breccias. *Tectonophysics*
663 510:69–79.
- 664 Wünnemann K., Collins G. S., and Melosh H. J. 2006. A Strain-Based Porosity Model
665 for Use in Hydrocode Simulations of Impacts and Implications for Transient
666 Crater Growth in Porous Targets. *Icarus* 180:514–527.
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672 **TABLE**

Table 1. Characteristics of objects in clast-rich breccia sample 16YA09.

clast type	length (mm)	colour (PPL)	grain size	grain shape	minerals
<i>Lithic clasts</i>					
mudstone n=36	1-12	brn	silt + clay + sand	sr to a	Qz, Afs, Bt, clay, Ilm
glaucopitic sandstone n=18	2-20	grn/brn	sand + silt + clay	sr to sa	Qz, Afs, Zrn, Glt
siltstone n=1	3.5	grey	silt	sr to sa	Qz, Zrn, TiO ₂ ,
palaeosol n=1	40	drk or		r to sa	Qz, Glt
<i>Other objects</i>					
vein n=5	0.5-4 width	brn	clay	ind	clay, Qz, Afs, Ilm
opaque objects (Fe-rich clasts) n=5	1-3	opaque	sand/silt/clay	ind.	Qz, Glt
Notes: ind- indeterminate; grain shape: r = rounded; sr = sub-rounded, a = angular; sa = sub-angular; PPL= plane polarized light; Colours: lt.=light, brn=brown, drk=dark, grn=green, blk=black, or=orange; Minerals: Qz=quartz, Afs=alkali feldspar, Glt=glauconite, Bt=biotite, Ilm=ilmenite, Zrn=zircon					

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Table 2. Summary of PF and PDF set abundances and indexed PF and PDF crystallographic orientations in quartz grains from six thin sections from the Yallalie structure, as determined using the universal-stage.

	PF		PDF	
No. of investigated grains	21		6	
No. of measured sets	65		9	
No. of sets/grain (N)	3.1		1.5	
No. of sets/grain (N*) ^a	2.9		1.5	
No. of sets; % relative to total no. of quartz grains examined; number of grains				
	PF		PDF	
1 set	4.8	1	83	5
2 sets	28.6	6	n.d.	n.d.
3 sets	28.6	6	n.d.	n.d.
4 sets	28.6	6	16.7	1
5 sets	9.5	2	n.d.	n.d.
Total	100	21	100	6
Indexed PF and PDF crystallographic orientations; absolute frequency (%) ^b ; number of sets				
	PF		PDF	
c {0001}	13.8	9	11.1	1
a {1014}	n.d.	n.d.	n.d.	n.d.
ω {1013}	1.5	1	11.1	1
π {1012}	3.1	2	22.2	2
r, z {1011}	44.6	29	44.4	4
m {1010}	n.d.	n.d.	n.d.	n.d.
ξ {1122}	12.3	8	11.1	1
s {1121}	3.1	2	n.d.	n.d.
ρ {2131}	1.5	1	n.d.	n.d.
x {5161}	1.5	1	n.d.	n.d.
a {1120}	n.d.	n.d.	n.d.	n.d.
{2241}	3.1	2	n.d.	n.d.
{3141}	4.6	3	n.d.	n.d.
t {4041}	1.5	1	n.d.	n.d.
k {5160}	1.5	1	n.d.	n.d.
Unindexed	7.7	5	0.0	n.d.
Total	100	65	100	9

^aCalculated only on indexed sets (i.e., unindexed sets excluded).

^bMethod described in, e.g., Engelhardt and Bertsch (1969) and Stöffler and Langenhorst (1994).

n.d. = none detected.

681 **FIGURES**

682

683 **Figure 1.** Location map of the Yallalie Structure in Australia. (A) Map of the Perth
684 Basin, onshore sub-basins in the central-northern Perth Basin, and Yilgarn Craton (after
685 Cockbain 1990; Mory and Iasky 1994). Filled squares indicate cities and towns. (B) Map
686 of inset shown in A. The Yallalie structure is situated within the Dandaragan Trough, ~10
687 km west of the Darling Fault.

688

689 **Figure 2.** Location map and generalized geology of the area around the Yallalie structure
690 (after Carter et al. 1982 and Dentith et al. 1999). The breccia study area is located in the
691 lower left. The thick dashed line indicates the 12 km diameter size estimate previously
692 suggested by Dentith et al. (1999).

693

694 **Figure 3.** Photographs of Mungedar Breccia outcrops. (A) A large, elongate clast of
695 greensand (Molecap Greensand Fm.?) within the breccia. (B) Horizontal layering within
696 the breccia. (C) Centimetre-sized clasts found within the breccia. The shaft of the chisel
697 in all images is 2 cm wide.

698

699 **Figure 4.** Slabs of Mungedar Breccia from which thin-sections were made for this study.
700 (A) Green-to-tan colored clast-rich breccia, sample 16YA09. (B) Example of brown-to-
701 orange colored clast-poor breccia.

702

703 **Figure 5.** Thin-section scans of Mungedar Breccia samples. (A) Clast-poor sample
704 16YL01. The section was cut from a hand sample similar to that shown in Figure 4B, and
705 was collected proximal to the outcrops in Figure 3. This section contains a grain of ballen
706 quartz and three grains with PFs (see insets). (B) Clast-rich sample 16YA09. This section
707 contains six shocked quartz grains; two with planar fractures and four with decorated
708 planar deformation features (see insets). The section was cut from the hand sample shown
709 in Figure 4A, and was collected proximal to the outcrops shown in Figure 3.

710

711 **Figure 6.** Backscattered electron images of deformed biotite grains with kink-bands in
712 mudstone clasts from clast-rich Mungedar Breccia sample 16YA09. (A) Full grain image
713 of a biotite grain with one orientation of kink bands in a sub-vertical orientation (see
714 arrow in inset). (B) Full grain image of biotite grain with two orientations of kink-bands.
715 (C) Detailed view of two orientations of kink-bands from (B). The arrows indicate
716 orientations of the kink-bands. Note the small grain size.

717

718 **Figure 7.** Transmitted light images of shocked quartz grains with planar fractures from
719 clast-rich Mungedar Breccia sample 16YA09. (A) Image of a siltstone clast (outlined
720 with dashed line) containing shocked quartz grain Q7 and deformed biotite grains. (B)
721 Detailed view of inset shown in A, showing a full grain image of grain Q7 with three
722 orientations of indexed planar fractures. (C) Full grain image of shocked quartz grain Q1.
723 (D) Detailed view of inset from C, showing two orientations of indexed planar fractures.

724

725

726 **Figure 8.** Transmitted light images of shocked quartz grains with planar fractures from
727 clast-rich Mungedar Breccia sample 16YA09. (A) Full grain image of shocked quartz
728 grain Q4. (B) Detailed view of inset shown in A, showing four indexed orientations of
729 planar fractures. (C) Full grain image of shocked quartz grain Q3. (D) Detailed view of
730 inset from C, showing four indexed orientations of planar fractures.

731

732 **Figure 9.** Transmitted light images of shocked quartz grains with decorated planar
733 deformation features (PDFs) from clast-rich Mungedar Breccia sample 16YA09. (A)
734 Full grain image of shocked quartz grain Q6, with the indexed PDF orientation. (B) Full
735 grain image of shocked quartz grain Q2, with the indexed PDF orientation. (C) Full grain
736 image of shocked quartz grain Q16 which contains 3 PF sets and PDF set. (D) Detailed
737 view of inset shown in C, with the indexed PDF orientation. (E) Detailed view of inset
738 shown in C, with the one orientation of indexed PDFs (labelled) and three orientations of
739 indexed PFs. (F) Full grain image of shocked quartz grain Q24. Two orientations of PDFs
740 are indicated on the figure; two additional sets are visible with the u-stage along $\{1011\}$
741 and $\{1122\}$ orientations. A close inspection of the planar microstructures under the u-
742 stage at high magnification shows that the brownish appearance of the grain results from
743 the dense network of microstructures.

744

745 **Figure 10.** Orientation maps of shocked quartz grains. (A) Plain polarized light image of
746 a part of PF-bearing grain Q4 (see Figure 7A) that was mapped by EBSD. (B) Orientation
747 map showing strain in the host grain using a texture component with a maximum
748 misorientation of 10° (rainbow color scale). Dauphiné twins are shown in pink. (C) Plain

749 polarized light image of PDF-bearing grain Q6 that was mapped by EBSD. (D)
750 Orientation map showing strain in the host grain using a texture component with a
751 maximum misorientation of 10° (rainbow color scale). Dauphiné twins are shown in pink.
752 LAB= Low angle boundary. Red cross in B and D indicates the reference orientation.

753

754 **Figure 11.** Type-II ballen quartz grain from clast-poor breccia sample 16YL01. (A) Plain
755 polarized light image of the grain. (B) MicroRaman spectra confirming that the grain is
756 α -quartz. The inset shows a cross-polarized light image exhibiting homogeneous
757 extinction.

758

759 **Figure 12.** Numerical simulation of the Yallalie impact crater. Peak pressure is shown on
760 the left and partial displacement is shown on the right (r = radius, z = height relative to
761 the paleo surface). The grey box at the surface denotes the sampled ejecta site. This
762 simulation was made in iSALE-2D hydrocode using a 1.2 km impactor striking a granite
763 target at 12 km/s, forming a crater with a rim-to-rim diameter of 14-16 km. The grey
764 image below the model is a previously published interpreted seismic line across the
765 Yallalie structure (Bevan 2012). Interpreted fault traces on the seismic line (from Hawke
766 et al. 2006) were added to the iSALE model at the same location and scale. TWT= two
767 way time.

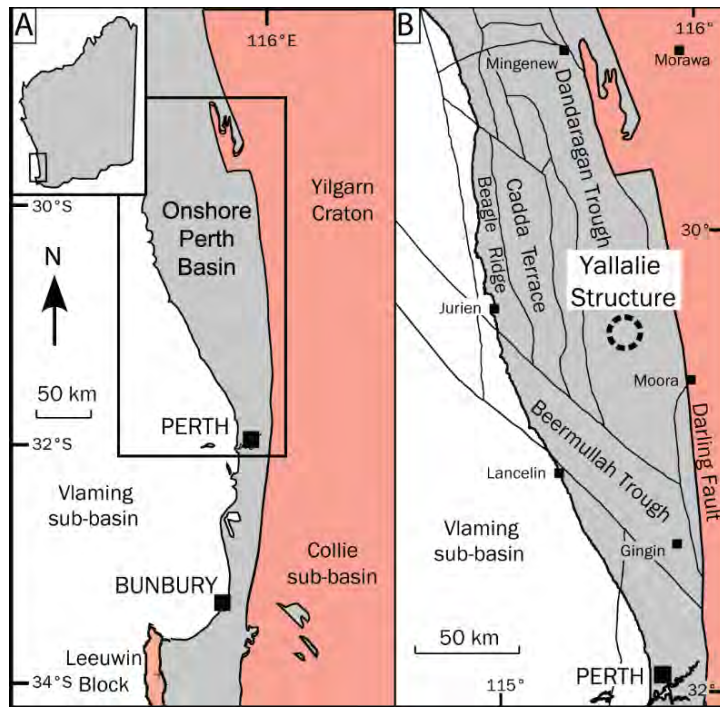


Figure 1

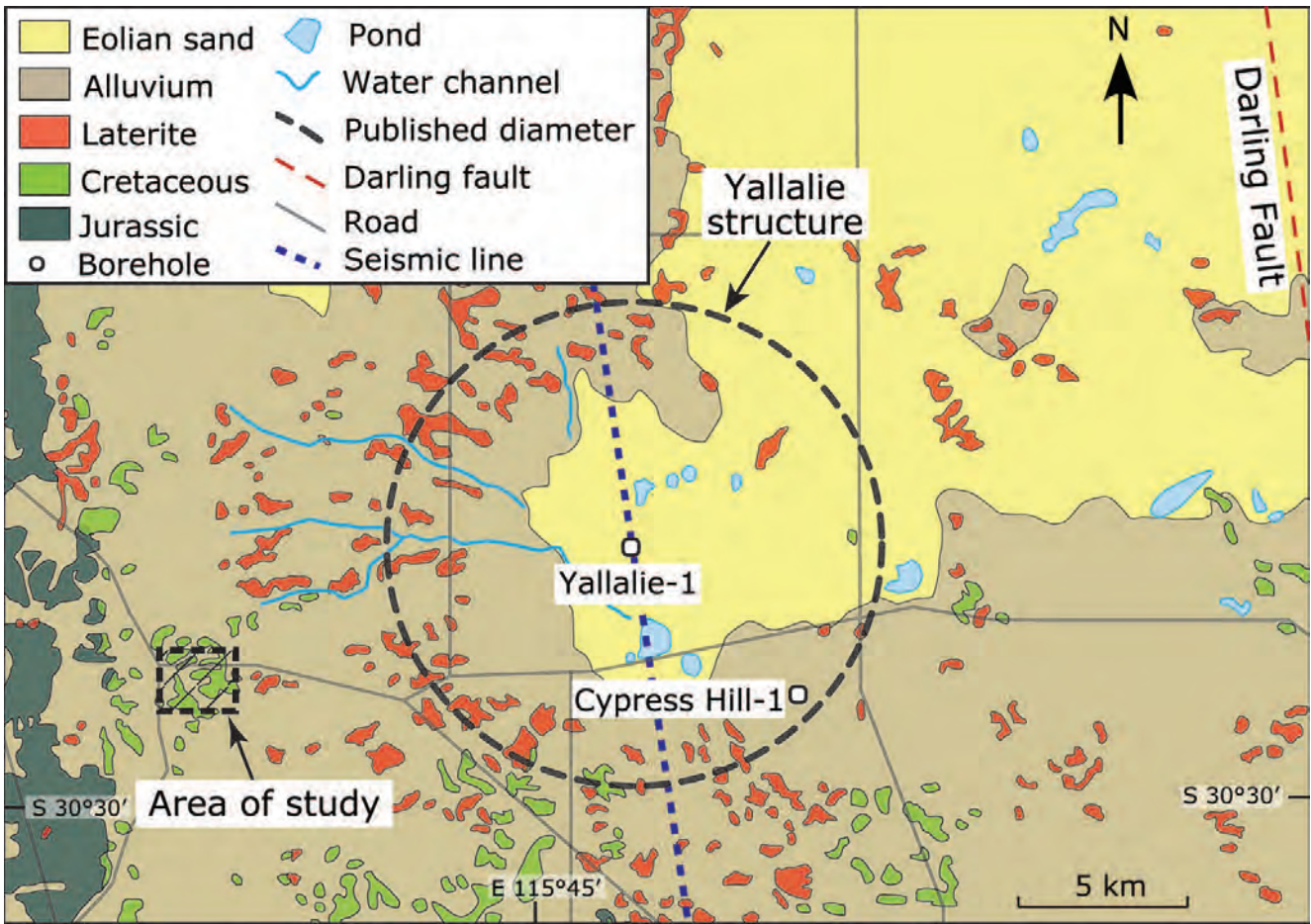


Figure 2

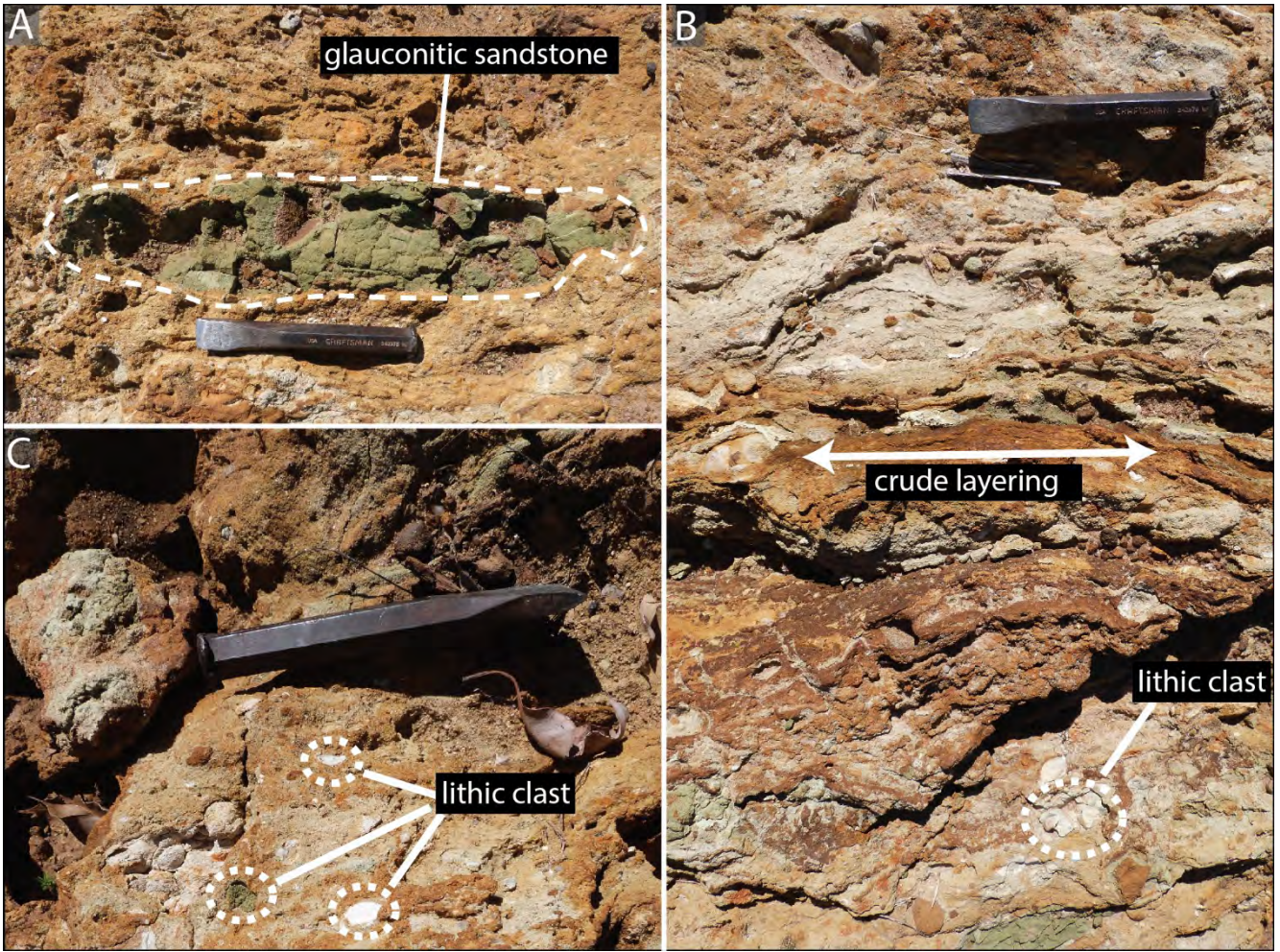


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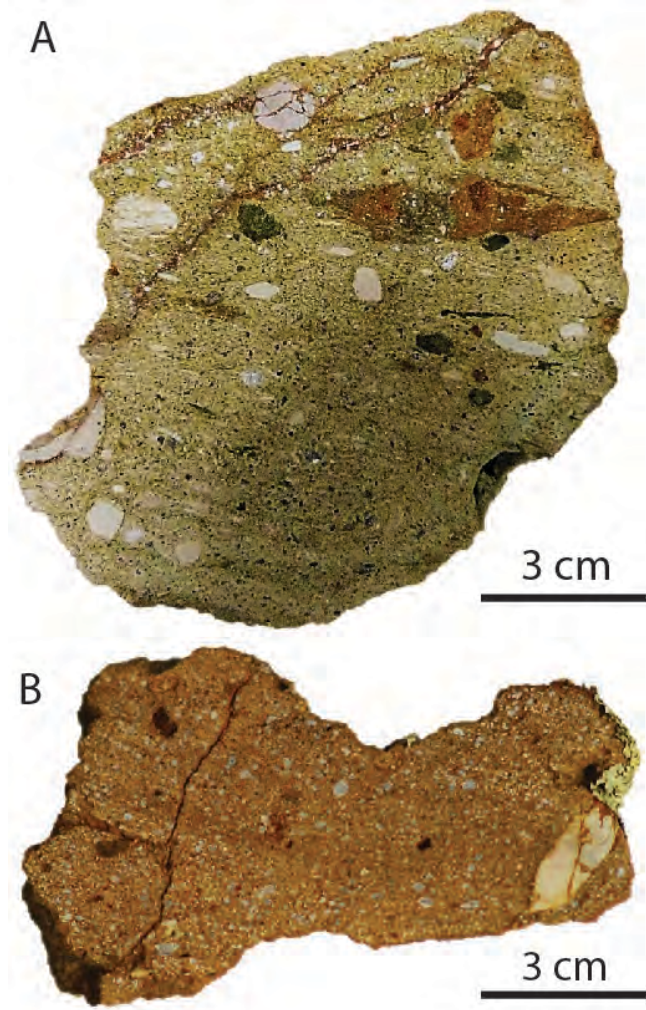


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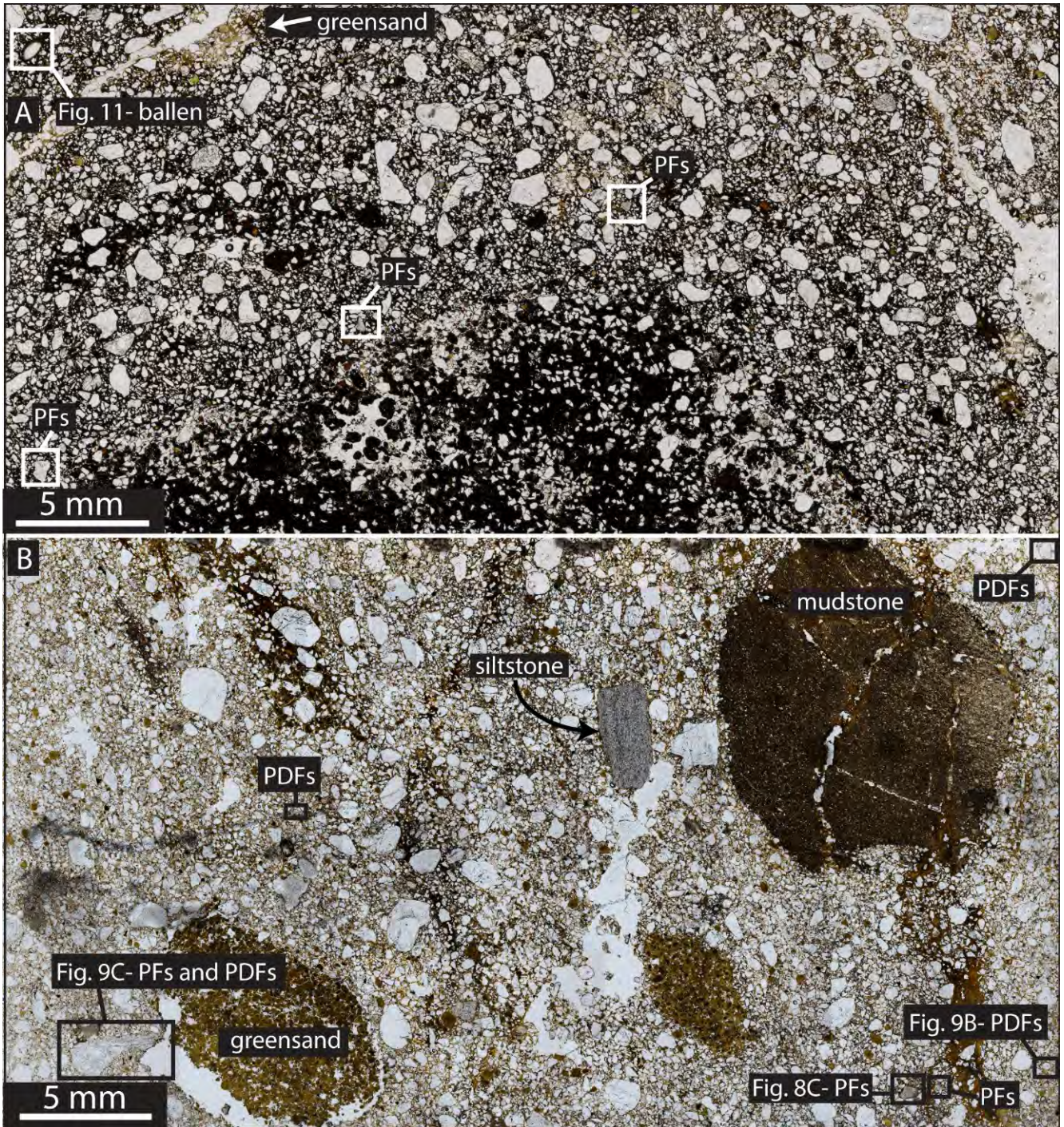


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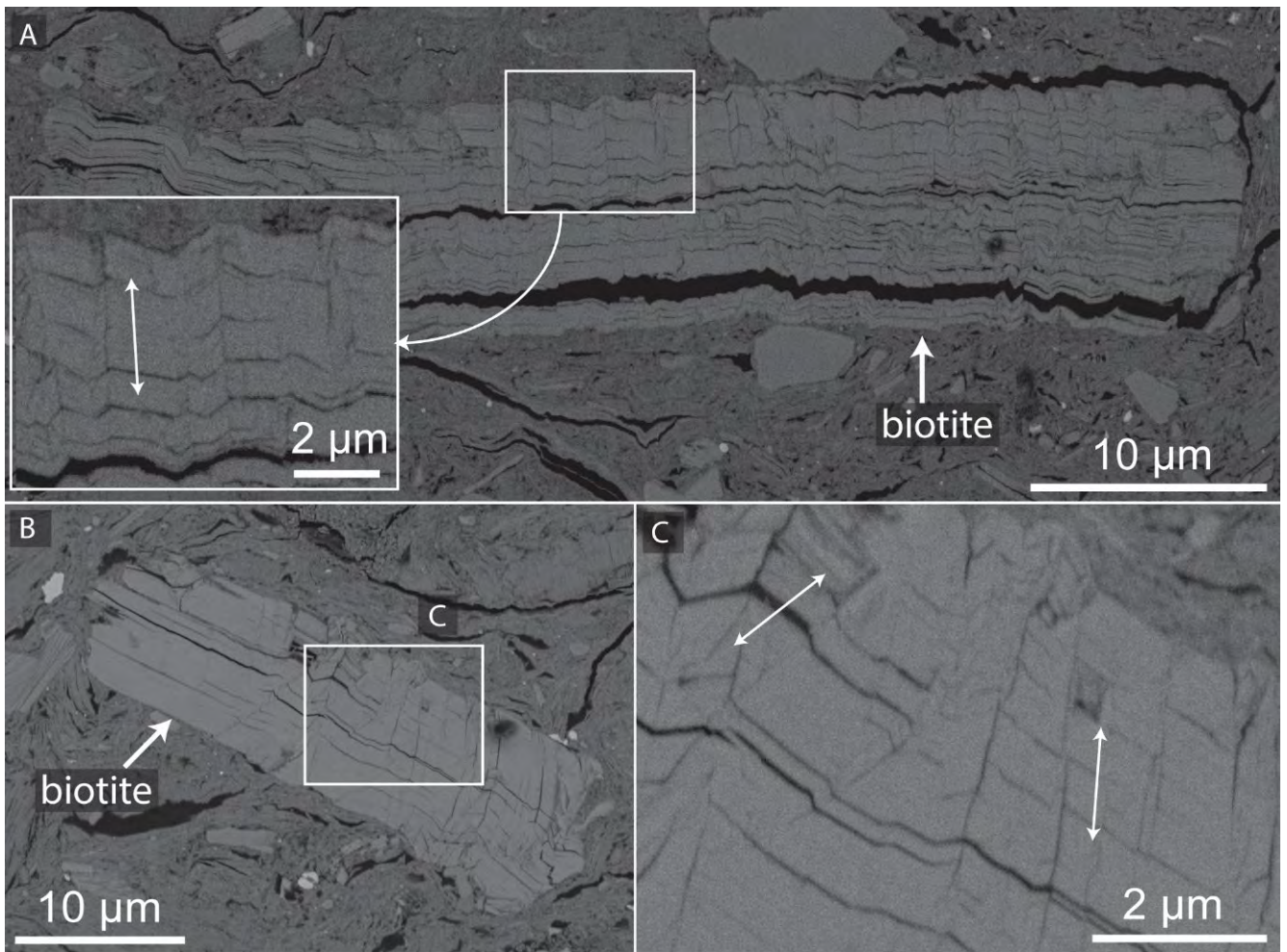


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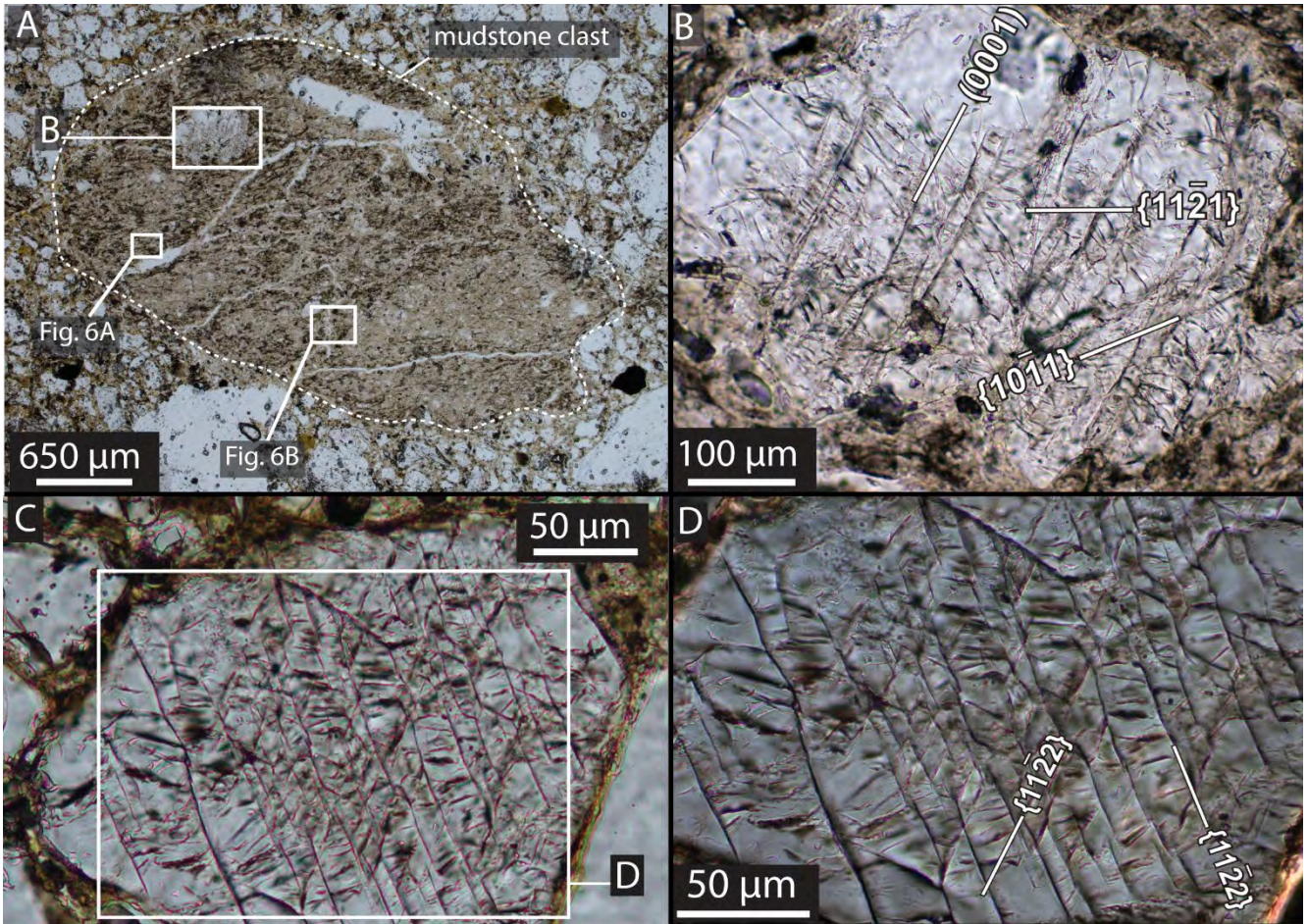


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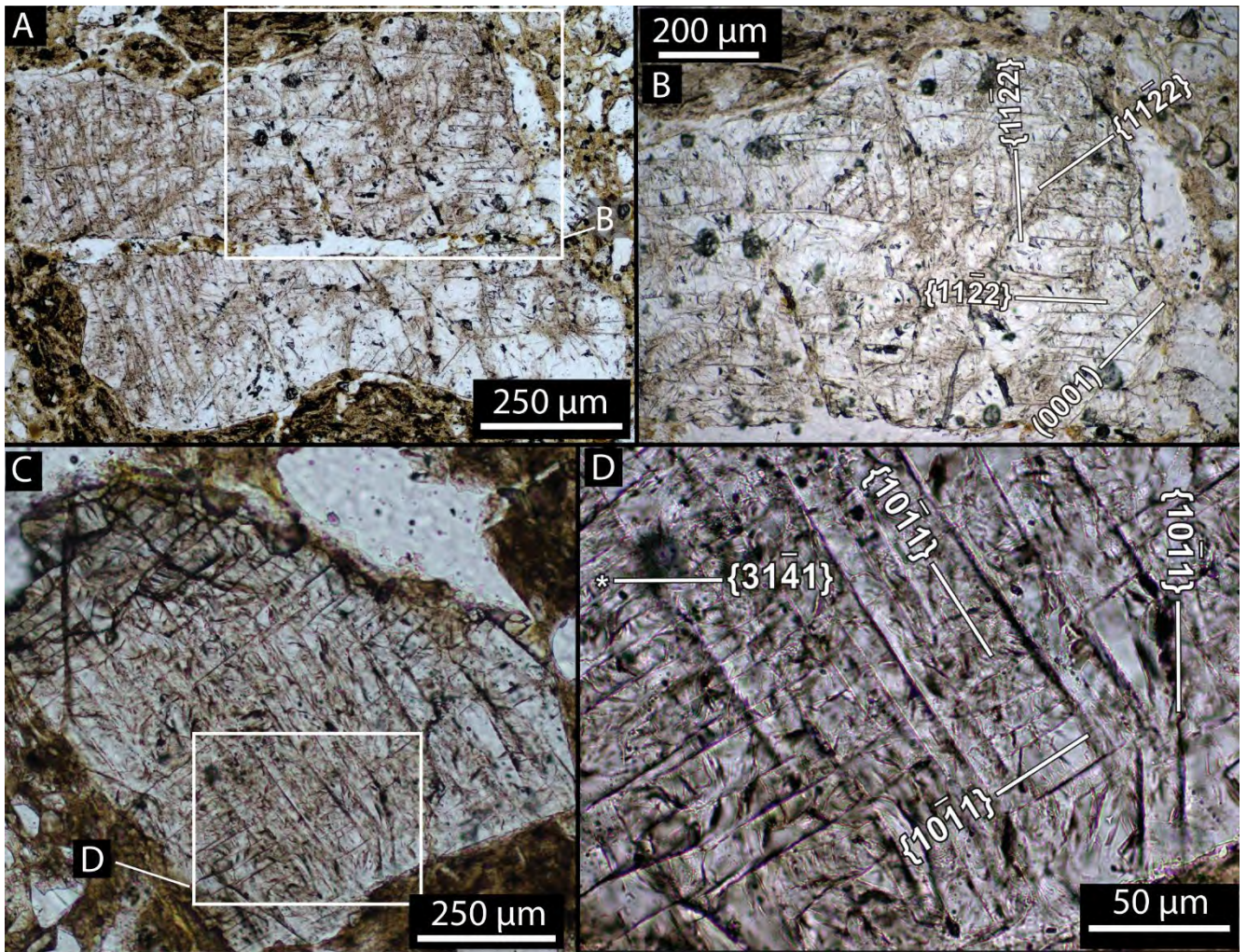


Figure 8

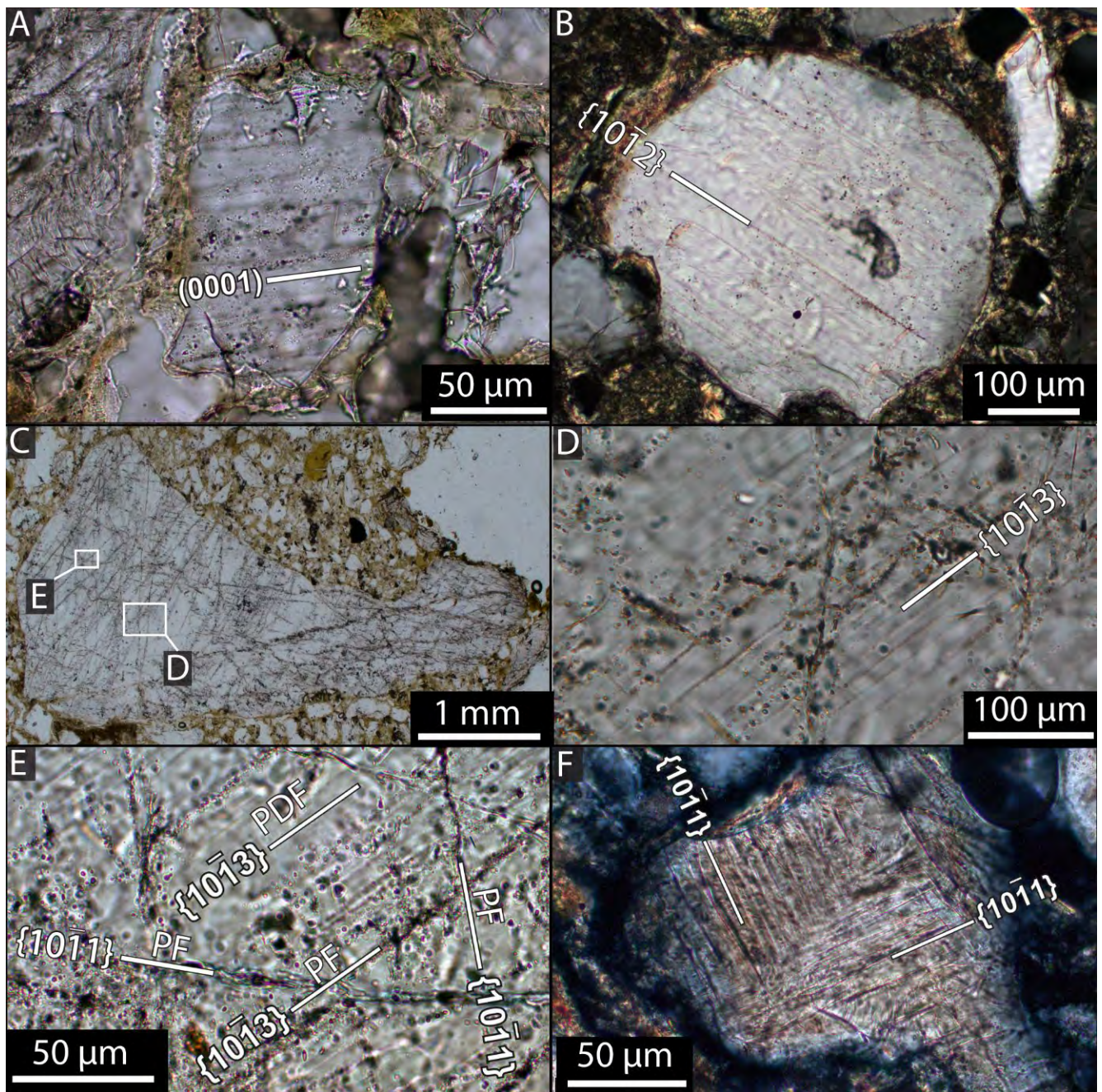


Figure 9

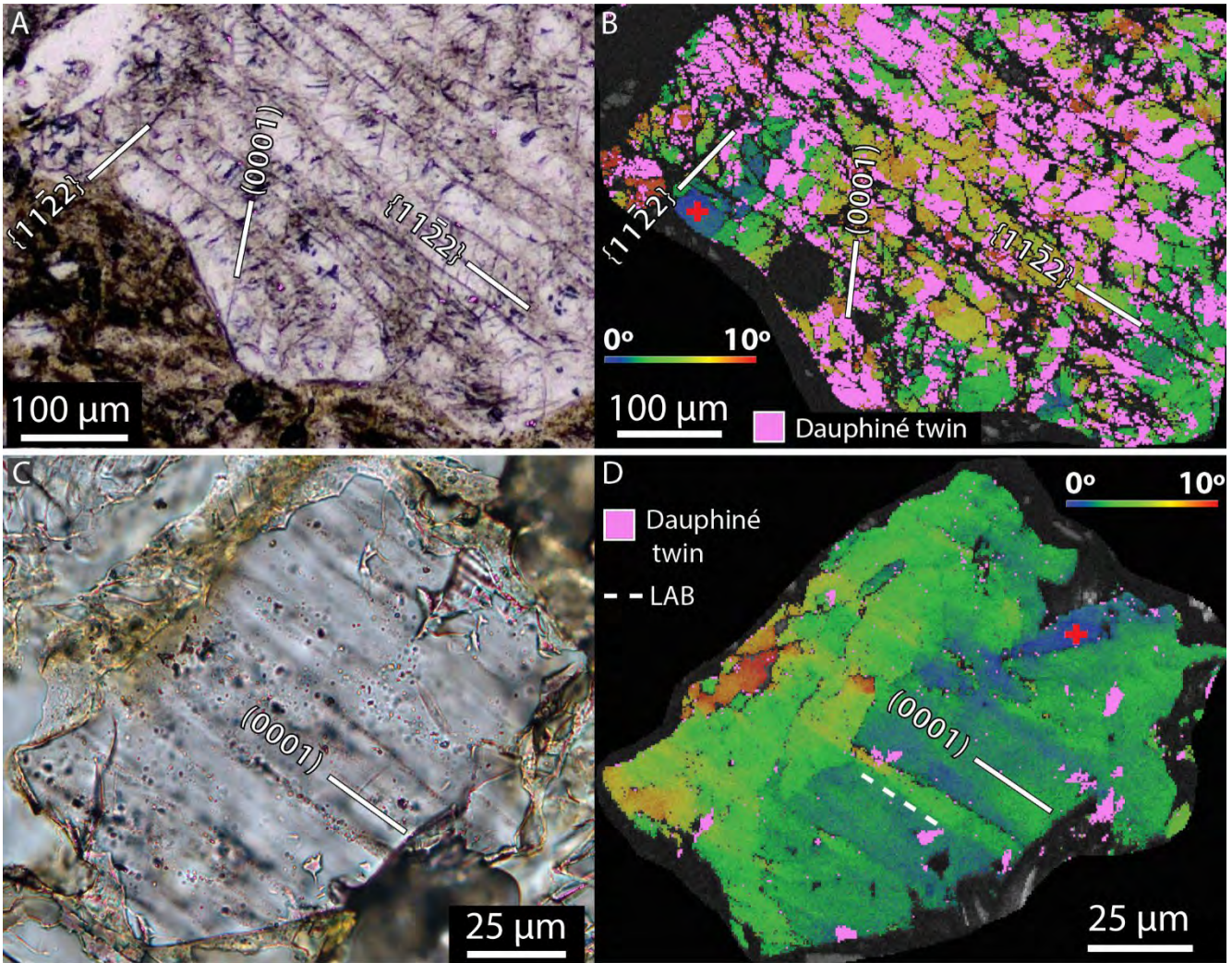


Figure 10

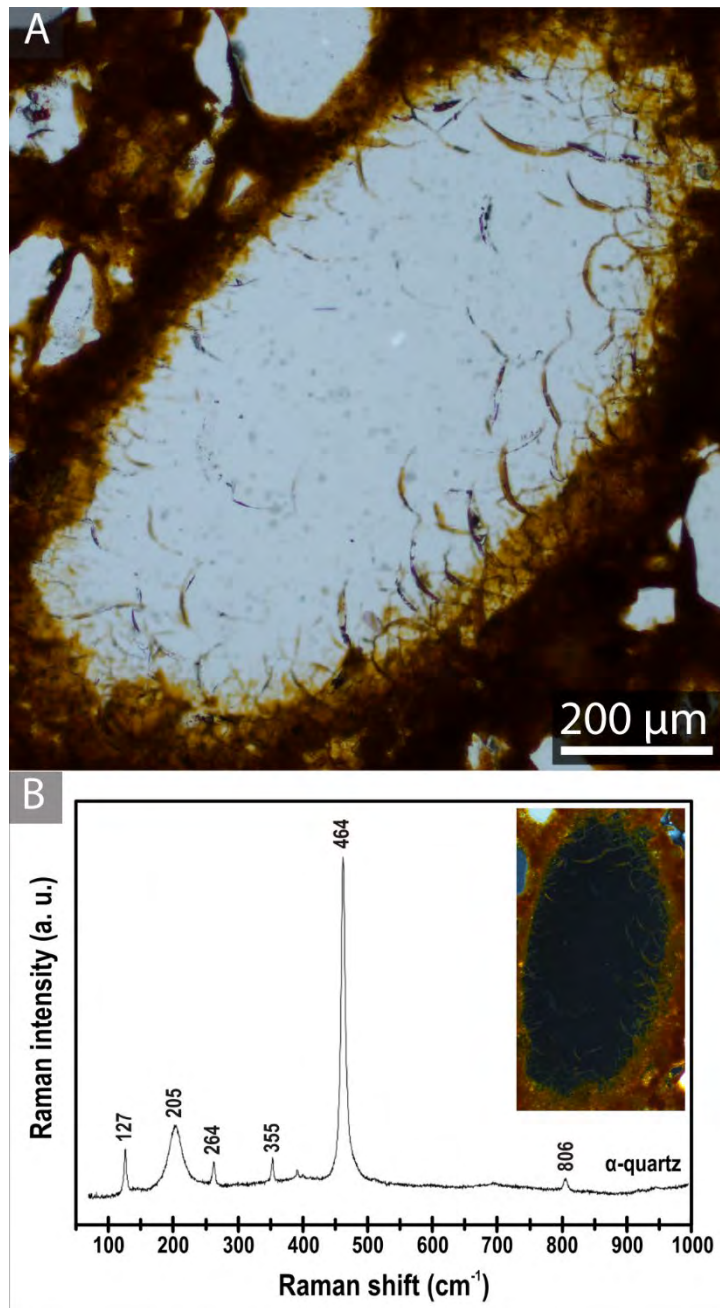


Figure 11

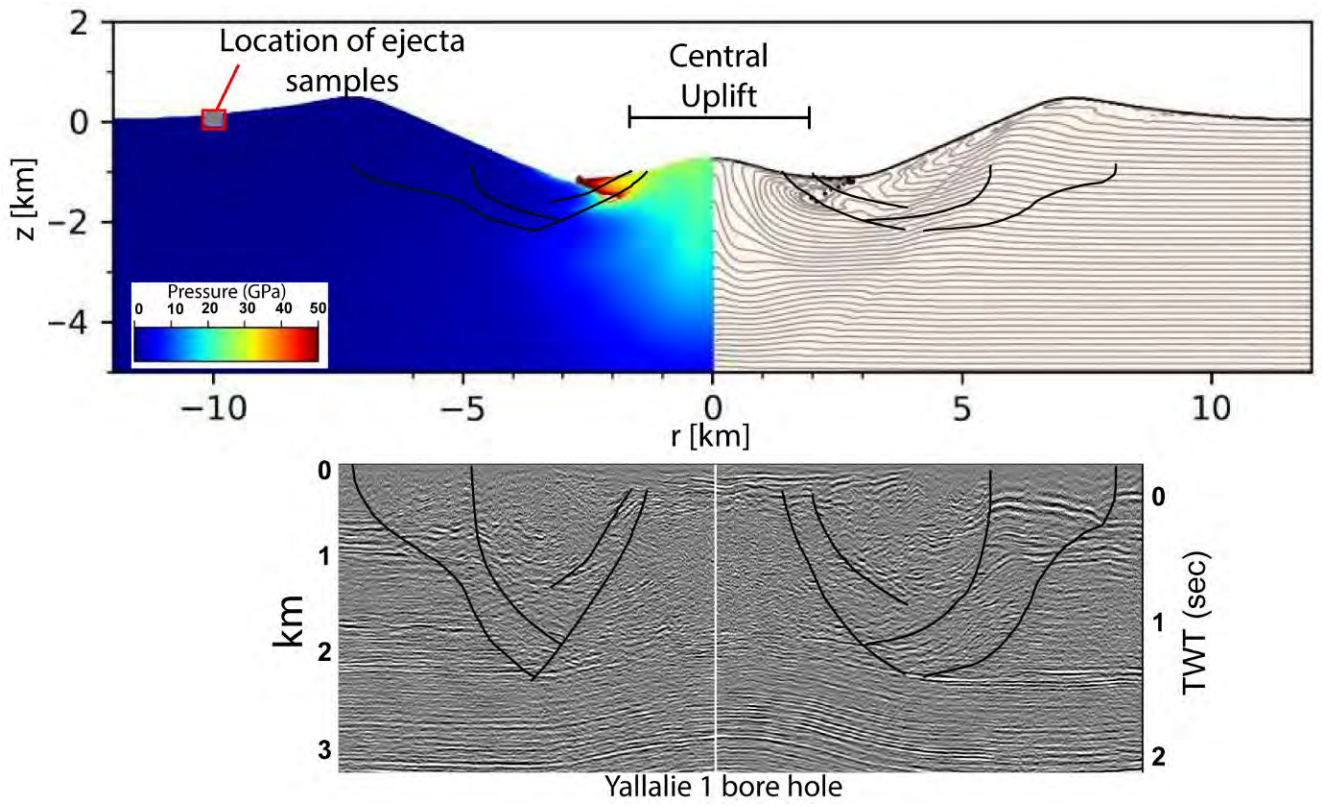


Figure 12