

Radiation Hardness Assurance: Evolving for *NewSpace*

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Acronyms

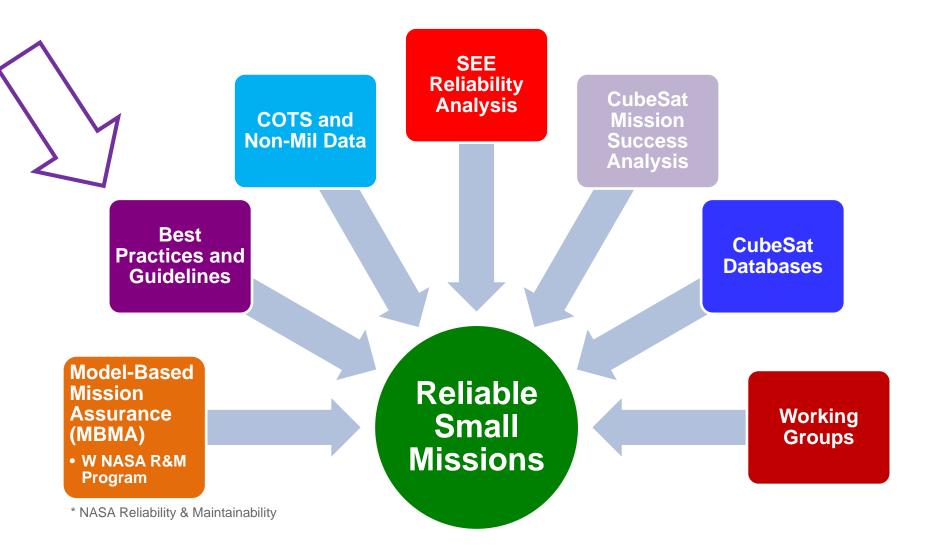
CME	Coronal Mass Ejection
COTS	Commercial Off The Shelf
DDD	Displacement Damage Dose
EEE	Electrical, Electronic, and Electromechanical
ELDRS	Enhanced Low Dose Rate Sensitivity
EP	Enhanced Performance
ESA	European Space Agency
GCR	Galactic Cosmic Ray
GOMAC	Government Microcircuits Applications and Critical Technologies Conference
GSFC	Goddard Space Flight Center
GSN	Goal Structuring Notation
HEART	Hardened Electronics and Radiation Technology
LEO	low earth orbit
LET	Linear Energy Transfer
MBMA	model based mission assurance
MRQW	Microelectronics Reliability and Qualification Workshop
NAND	Negated AND or NOT AND
NASA	National Aeronautics and Space Administration
NEPP	NASA Electronic Parts and Packaging
NEPP ETW	NASA Electronic Parts and Packaging (NEPP) Program Electronics Technology Workshop
NSREC	Nuclear and Space Radiation Effects Conference

RADECS	Radiation Effects on Components and Systems				
RHA	Radiation Hardeness Assurance				
SAA	South Atlantic Anomaly				
SEE	Single Event Effects				
SEE/MAPLD	SEE-MAPLD Single Event Effects (SEE) Symposium/				
SEE/MAPED	Military and Aerospace Programmable Logic Devices (MAPLD) Workshop				
SEGR	Single Event Gate Rupture				
SEL	Single Event Latchup				
SEP	Single Event Effects Phenomena (includes SEU, SEL, SEGR and SET)				
SERESSA	School on the Effects of Radiation on Embedded Systems for Space Applications				
SET	Single Event Transient				
SEU	Single Event Upset				
SLU	Saint Louis University				
SwaP	Size, weight, and power				
TID	Total Ionizing Dose				
TID	Total Ionizing Dose				
TMR	triple-modular redundancy				
TNID	Total Non-Ionizing Dose				
UV	Ultra-Violet				

This presentation to be published on nepp.nasa.gov, originally presented by Michael J. Campola at Radiation and its Effects on Components and Systems (RADECS), Montpellier, France, September 16, 2019.



NEPP Program- Small Mission Efforts



Outline

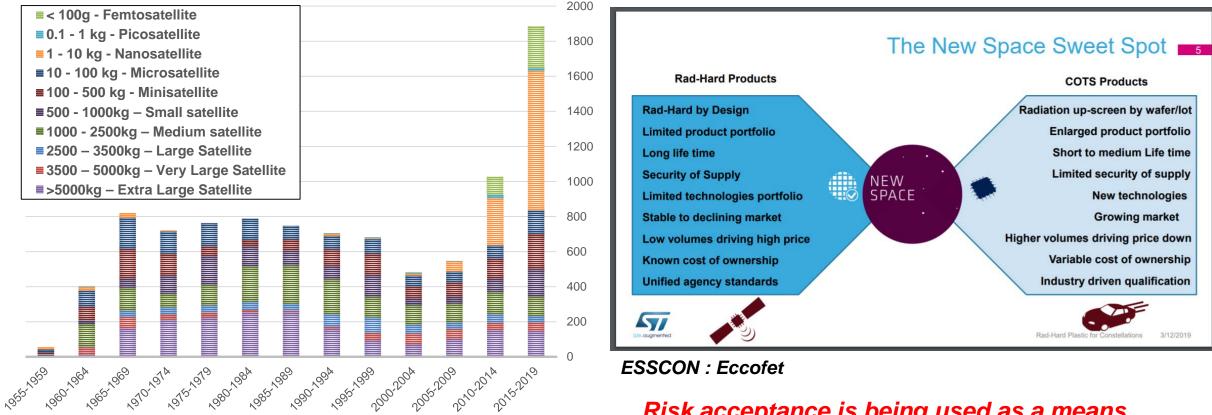


- New Space and SmallSat Considerations
- The Natural Space Radiation Environment Hazard
- Radiation Effects on Micro-Electronics
- Hardness Assurance, as a Discipline, with its Challenges
 - New Technologies
 - New Architectures
 - Unbound Risks
- Building Smart Requirements
- Risk Acceptance and Guidance

New Space – New Point of View



SmallSats Come in Many Sizes



Seradata SpaceTrak Data

Risk acceptance is being used as a means to enable innovation

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Component Grades are Merging

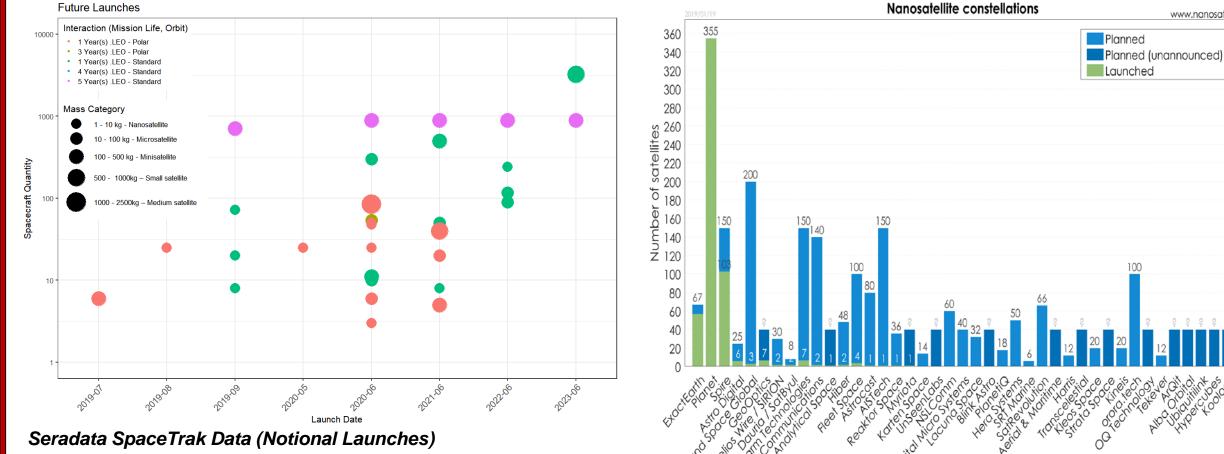
New Space – Looking Ahead



www.nanosats.eu

Constellations and Swarms

New Space = New Companies



Seradata SpaceTrak Data (Notional Launches)

New Space – Same Old Radiation

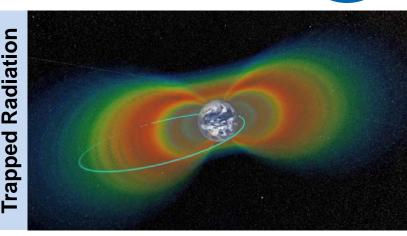
New mission concepts and SmallSat paradigm

- Radiation challenges identified in the past are here to stay; adoption of new technologies are often the risk driver
- Commercial Space, Constellations, Small missions, etc. will benefit from detailed hazard definition and mission specific requirements
- The need for Radiation Hardness Assurance (RHA)
 - Radiation effects are a mix of disciplines, evolve with technologies and techniques
 - Misinterpretation of failure modes / misuse of available data can lead to over/under design
 - RHA flow doesn't change, risk acceptance needs to be tailored
- Some Top Level Resources
 - NPR-7120.5 NASA Agency Program Management
 - GPR-8705.4 NASA Goddard Risk Classification Guidelines
 - NASA-STD-8739.10 NASA Parts Assurance Standard

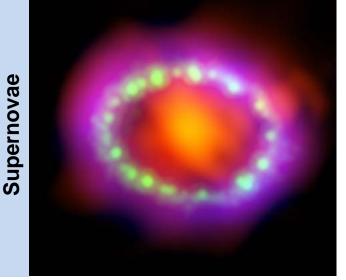
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https://www.nasa.gov/van-allen-probes



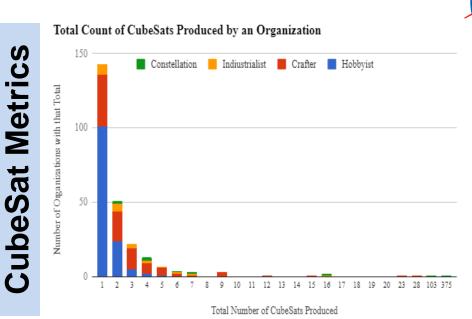


NASA, ESA, and L. Hustak (STScl)

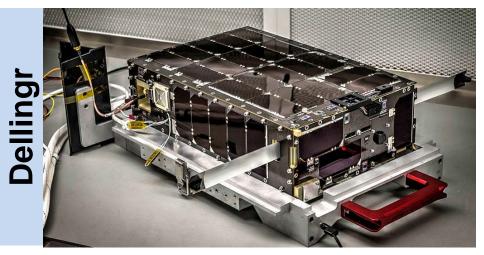
Who Needs This Guidance?

- Universities / CubeSats
 - May be first-time designers, or previous missions did not have requirements
 - Schedule driven, limited time for development
 - Rideshares could end up in multiple environments
- Space Agencies / Government
 - More compact designs in new destinations
 - Cost savings of SmallSat platform, with more reliable outcome
 - More willing to trade risk for capability
- Device / Subsystem Manufacturers
 - Product / Device offerings: Space Plastic, EP, LeanRel, radiation tolerant, modified HiRel, etc.
 - Fault tolerance in designs

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Michael Swartwout, SLU CubeSat Database



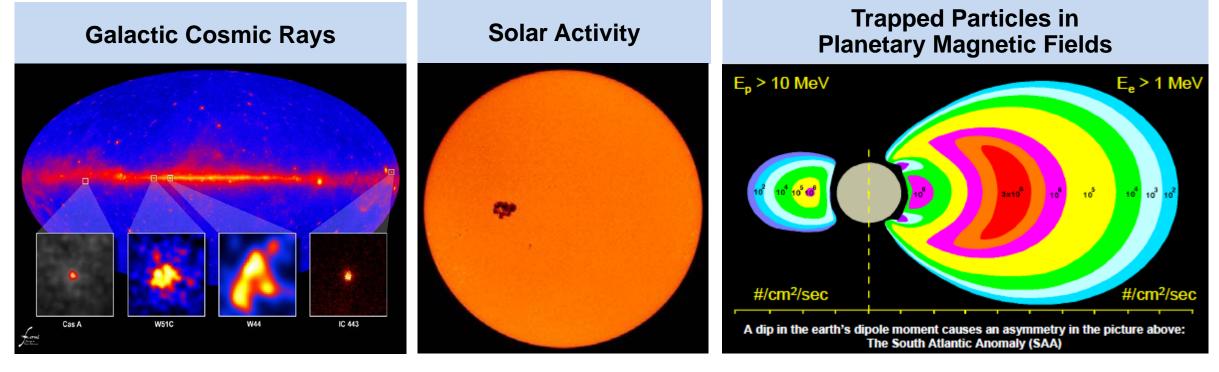
NASA's Goddard Space Flight Center/Bill Hrybyk



Natural Space Radiation Environment



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Energetic supernovae remnants (~GeV, Z=1-92) Originate outside of our solar system Solar Wind, Solar Cycle CMEs (proton rich) Flares (heavy ion rich) Fluctuate with Solar Activity and Events Not a perfect dipole Protons and Electrons trapped at different L-shell values and energies

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Images from left to right – NASA FERMI X-ray telescope, Solar Dynamics Observatory, Janet Barth (radhome.gsfc.nasa.gov)

Natural Space Radiation Environment

- Plasma
- Particle Radiation
- Neutral Gas Particles
- UV and X-Ray
- Orbital Debris

Degradation of micro-electronics Degradation of optical components Degradation of solar cells

Data corruption Noise on images System shutdowns or resets Circuit Damage Part tolerances exceeded

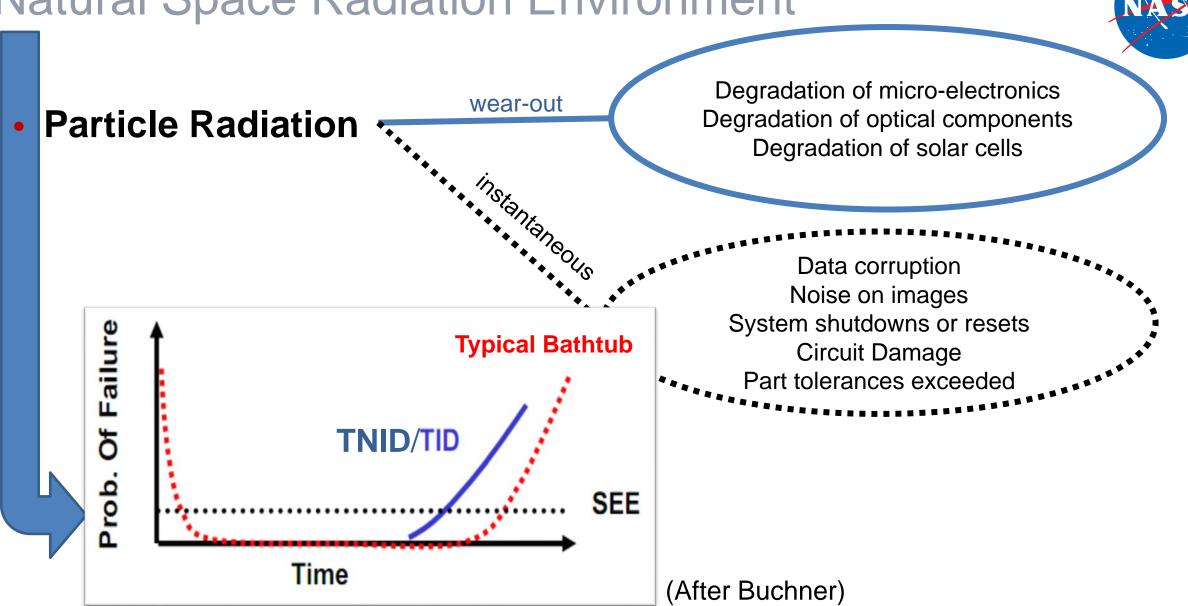
(After Barth)

Spacecraft Charging, Ionizing Dose, Non-Ionizing Dose, Single Event Effects, Drag, Surface Erosion, Debris/Micro-Meteoroid Impacts, Thermal Cycles

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wear-out

Natural Space Radiation Environment



Conventional Units Explanation

Degradation

- Total Ionizing Dose (TID)
 - Absorbed dose (rad(Si))
 - 1 rad = 100 erg/g = 0.01 J/kg; 100 rad = 1 Gy
 - Always specified for a particular material
 - 1 rad(SiO₂), 10 krad(Si), 100 Gy(H₂O)
 - This is not exposure (R), or dose equivalent (Sv)
- Total Non-Ionizing Dose (TNID)
 - Fluence (particles/cm²)
 - Number of particles per unit area
 - Displacement Damage Dose (DDD)

Single Event

- Linear Energy Transfer (LET)
 - Stopping power normalized to target material

$$S = -\frac{dE}{dx} \Rightarrow \text{LET} = -\frac{1}{\rho}\frac{dE}{dx}$$

- Units are MeV·cm²/mg
- Cross Section (σ)
 - Device particle interaction (cm²)
 - Used in calculation of rate
 - Can be /device or /bit per time interval



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Degradation Contributors vs. Single Event

1.0E+07

Shielding

۷S.

Total Ionizing Dose

Cumulative effects

- Depend highly on which contributors and duration in their presence
- Mimic wear-out/aging
- TNID and TID must be accounted for

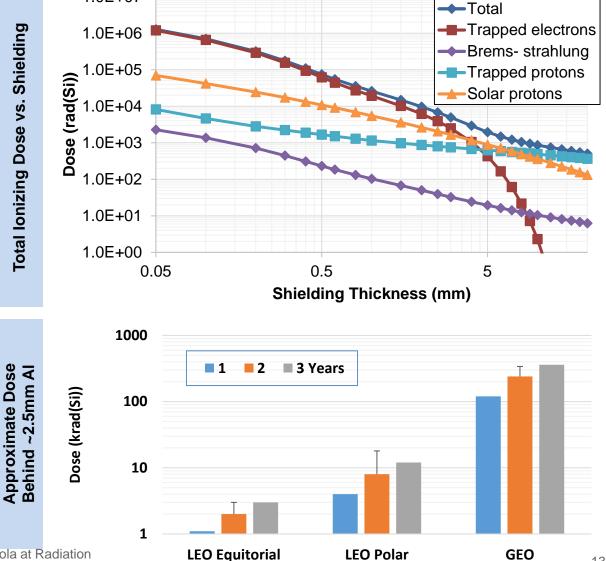
Typical destinations (LEO, GEO)

- LEO at low altitude/inclination is more protected by the Geomagnetic field
- Proximity to the poles & SAA show a large variability in dose despite short mission durations
- Electrons and their braking radiation are the big offender in Geostationary orbits (don't forget about spacecraft charging...)

Note that

- A little bit of shielding goes a long way
- Altitude plays a huge role when in/near the radiation belts (even transiting)
- Beyond Geomagnetic field, highly variable solar environment contributions (Solar cycle)

Degradation has a strong dependence on where you go, not just how long you are on orbit





Degradation vs. Single Event Contributors

Orbit

and

Altitude

Shielding Effectiveness

*

al Flux (

1.00E+00

1.00E-01

1.00E-02

1.00E-03

1.00E-04

0.1

lux (# of particles/cm²-s-MeV)

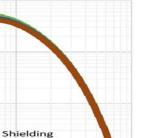
One particle causes the effect

- Random in nature, particle must traverse sensitive structure within device and have sufficient charge creation along its path
- Shielding doesn't do so much for highly energetic particles
- Device technology can be dependent on particle species
- Typical Destinations (LEO, GEO)
 - Again altitude plays a role; for some devices that is a direct threat
 - You are exposed to more GCR + Solar contribution as geomagnetic protection is reduced
 - Natural phenomena like the South Atlantic Anomaly (SAA), magnetic poles, are temporal drivers

Note that •

There will be a background rate, solar cycle dependence, solar event rate, increased rate for poles or SAA – not just one rate to consider

Single event contributors benefit very little from shielding, have dependence on where you are

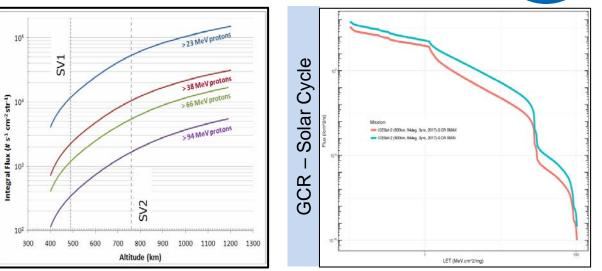


(mils of Al)

100

• 200 300

100



Trapped Proton Flux

10

Energy (MeV)

Shielding Impact

1000



Summary of Environmental Hazards



	Plasma (charging)	Trapped Protons	Trapped Electrons	Solar Particles	Cosmic Rays	Human Presence	Long Lifetime (>10 years)	Nuclear Exposure	Repeated Launch	Extreme Temperature	Planetary Contaminates (Dust, etc)
GEO	Yes	No	Severe	Yes	Yes	No	Yes	No	No	No	No
LEO (low- incl)	No	Yes	Moderate	No	No	No	Not usual	No	No	No	No
LEO Polar	No	Yes	Moderate	Yes	Yes	No	Not usual	No	No	No	No
International Space Station	No	Yes	Moderate	Yes - partial	Minimal	Yes	Yes	No	Yes	No	No
Interplanetary	During phasing orbits; Possible Other Planet	During phasing orbits; Possible Other Planet	During phasing orbits; Possible Other Planet	Yes	Yes	No	Yes	Maybe	No	Yes	Maybe
Exploration – Lunar, Mars, Jupiter	Phasing orbits	During phasing orbits	During phasing orbits	Yes	Yes	Possibly	Yes	Maybe	No	Yes	Yes

https://radhome.gsfc.nasa.gov/radhome/papers/SSPVSE05_LaBel.pdf

Radiation Hazard Contributors for Dose and SEE



Environment

		LEO Equatorial	LEO Polar (Sun Sync)	GEO / Interplanetary
	> 3 Years	Moderate Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	High Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	High Dose / High GCR, High Solar Proton Variability
	1- 3 Years	Manageable Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	Moderate Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	High Dose / High GCR, High Solar Proton Variability
2	< 1 Year	Manageable Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	Moderate Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	Moderate Dose / High GCR, High Solar Proton Variability

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Mission Lifetime

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Radiation Effects on Active Microelectronic Devices

- Cumulative effects and single event effects can <u>both</u> be permanently damaging
 - TID/DDD lead to wear-out of device operation and degrade devices beyond acceptable operations internally and externally
 - Single Event Effects can be catastrophic instantaneously by turning on parasitic devices within the semiconductor or inducing electric field across dielectrics that eventually break down
 - Synergistic effects can make ground based testing very difficult

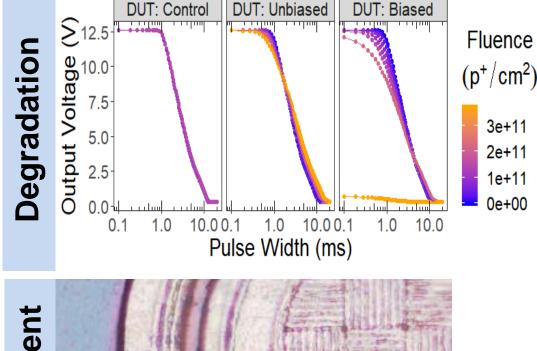
Destructive Single Event Effects (SEEs)

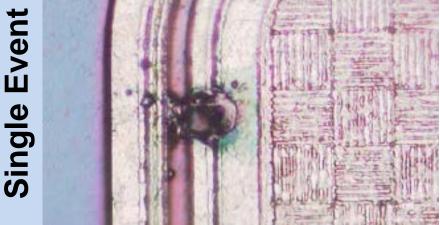
- Irreversible processes
- Terms: Latchup, Burnout, Gate Rupture

Non-Destructive SEEs

- Lead to interruptions in operation and/or errors leading to unknown state spaces or loss of science / mission if not accounted for
- Terms: Functional Interrupt, Transients, Upsets
- IEEE / Papers / Short Courses / Presentations
 - GOMAC, HEART, MRQW, NEPP ETW, NSREC, RADECS, SEE/MAPLD, SERESSA, SPWG

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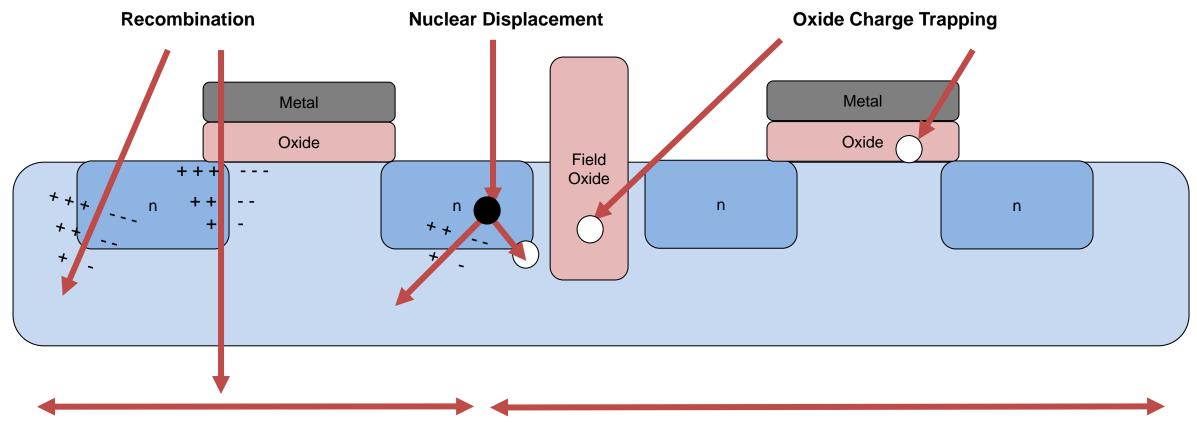
Megan Casey - https://nepp.nasa.gov/files/26196/2014-561-Casey-

Final-Web-Pres-ETW-Diodes-TN16278 v2.pdf



Device and Particle Interaction





Instantaneous

Cumulative

Brock J. LaMeres, Colin Delaney, Matt Johnson, Connor Julien, Kevin Zack, Ben Cunningham Todd Kaiser, Larry Springer, David Klumpar, "Next on the Pad: RadSat – A Radiation Tolerant Computer System," Proceedings of the 31st Annual AIAA/USU Conference on Small Satellites, Logan UT, USA, Aug. 5-10, 2017, paper: SSC17-III-11, URL: <u>http://digitalcommons.usu.edu/cgi/viewcontent.cgi?article=3618&context=smallsat</u>

Table of SEE Susceptibility



SEL	SEGR	SEB	SEDR	Stuck Bit	SEU/MCU	SET	SEFI
		POWER	One-time		Digital/bistable	bipolar	Complex
CMOS	MOSFET	MOSFET	Prog. FPGA	SRAM	technologies	technology	Microcircuits
			Bipolar			Analog	
Bipolar?	FLASH	Power JFET	Microcircuits	DRAM	Deep submicron	microcircuit	ADCs
	Schottky				CMOS more MCU	Digital	
	Diode	Power BJT		FLASH	susceptible	microcircuit	PWMs

Part-Level Consequences

How Common is Issue?

- Catastrophic failure possible
- Destructive but limited
- Nondestructive

- Common in technology
- Catastrophic failure possible
- Not seen but possible in principle

Ray Ladbury, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

List is not exhaustive, but new failure modes are found in new devices, so it would not be possible to capture all

Outline

NASA

- New Space and SmallSat Considerations
- The Natural Space Radiation Environment Hazard
- Radiation Effects on Micro-Electronics
- Hardness Assurance, as a Discipline, with its Challenges
 - New Technologies
 - New Architectures
 - Unbound Risks
- Building Smart Requirements
- Risk Acceptance and Guidance

The Job: Watch For the 'ilities



Survivability

- Must survive until needed
- Entire mission?
- Screening for early failures in components

Availability

- Must perform when necessary
- Subset of time on orbit
- Operational modes
- Environmental response

Criticality

- Impact to the system
- Part or subsystem function
- Mission objectives

Reliability

- Resultant of all
- Many aspects and disciplines
- Known unknowns

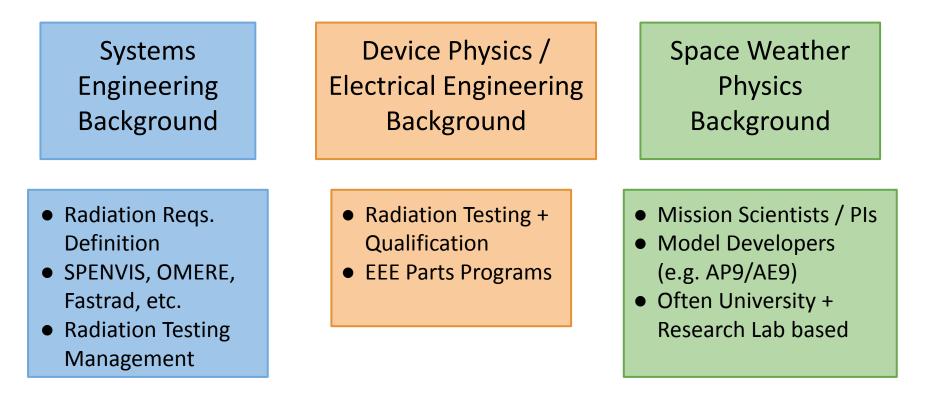
The People: Radiation Effects Engineers

Materials	Device Physics	Electrical Engineering	Systems Engineering	Space Physics
 Material Property degradations with radiation Energy loss in materials 	 Charge transport Device Process Dependencies Charge dependency of device operation 	 Part to part interconnections Understanding circuit response Device functions and taxonomy 	 Requirements System Level Impacts Understanding interconnections Understanding functionality 	 Space weather Environment models/modeling Radiation Sources and variability

Paths to Space Radiation



Space Radiation Ecosystem



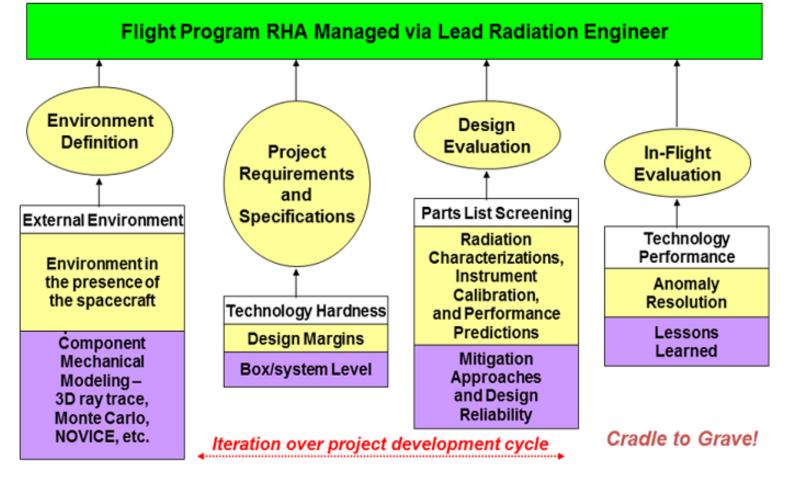
After Whitney Lohmeyer, presented at JPL meeting 2019

Radiation Hardness Assurance (RHA) Overview



RHA consists of all activities undertaken to ensure that the electronics and materials of a space system perform to their *design* specifications throughout exposure to the mission space environment

(After Poivey 2007)



(After LaBel 2004)

RHA Challenges... Not So Small

- Always in a <u>dynamic</u> environment
- New Technologies
 - Device Topology / Speed / Power
 - Increased COTS parts / subsystem usage
- New Mission Architectures
 - Profiles of mission life, objective, and cost are evolving
 - Oversight gives way to insight in some mission classifications
 - Ground systems, do no harm, hosted payloads
 - Similarity and heritage data requirements widening
- Quantifying Risk
 - Translation of system requirements to radiation trades can be problematic
 - Determining appropriate mitigation level (operational, system, circuit/software, device, material, etc.)

Unbound radiation risks are likely

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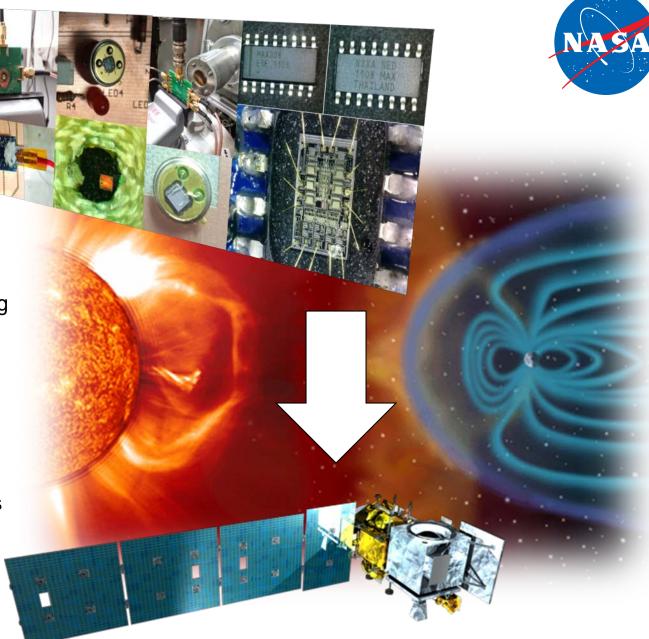


Image Credits: NASA / SOHO / ICESAT-2

New Technologies - New Susceptibilities

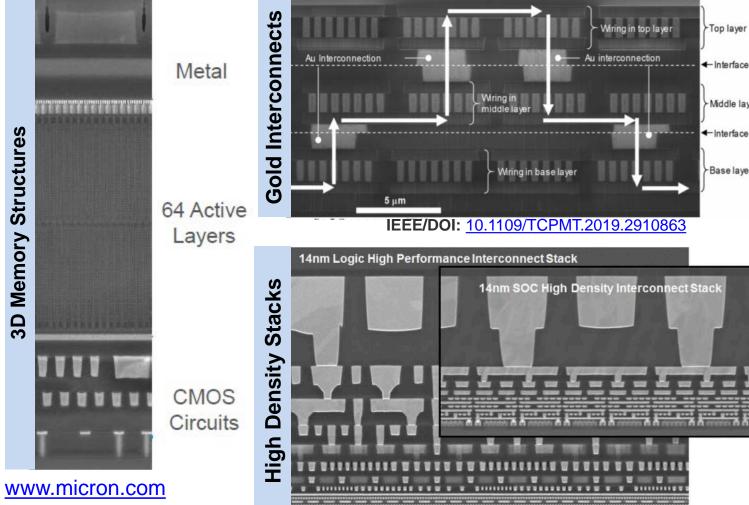
Feature Size / Critical Charge •

- Sensitivity to muons? Low energy protons?
- **3D Stacking / Structures** •
 - Deep sensitive volumes
 - New materials within structure
- Testing Challenges
 - Complexity (e.g., Systems-on-a-Chip) •
 - Speed of interfaces
 - Obfuscation of state-space
 - Flux / range of beam at facilities
- Function •
 - Integrated Photonics, MEMS, Hybrids

Without detailed part information you do not have certainty of the radiation threats

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Top layer

Middle laver

-Base layer

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New Mission Architectures - How Many to Succeed? **Allowable Losses** Single Strain Early Early

Degradation Degradation Degradation Degradation Mission Loss Mission Loss Destructive or Critical SEE Destructive or RISK Mitigated RISK Critical SEE SEE VS Mitigated SEE Non-Critical Non-Critical Manageable SEE Manageable SEE Benign Harsh Benign Harsh **Environmental Hazard Environmental Hazard**

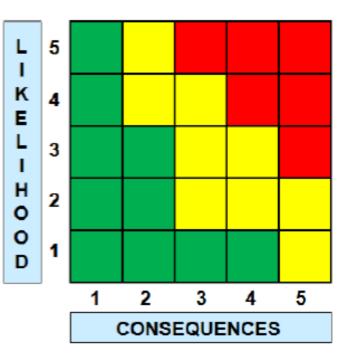
Redundancy alone does not remove the threat, adds complexity



New Challenges in Quantifying Risk

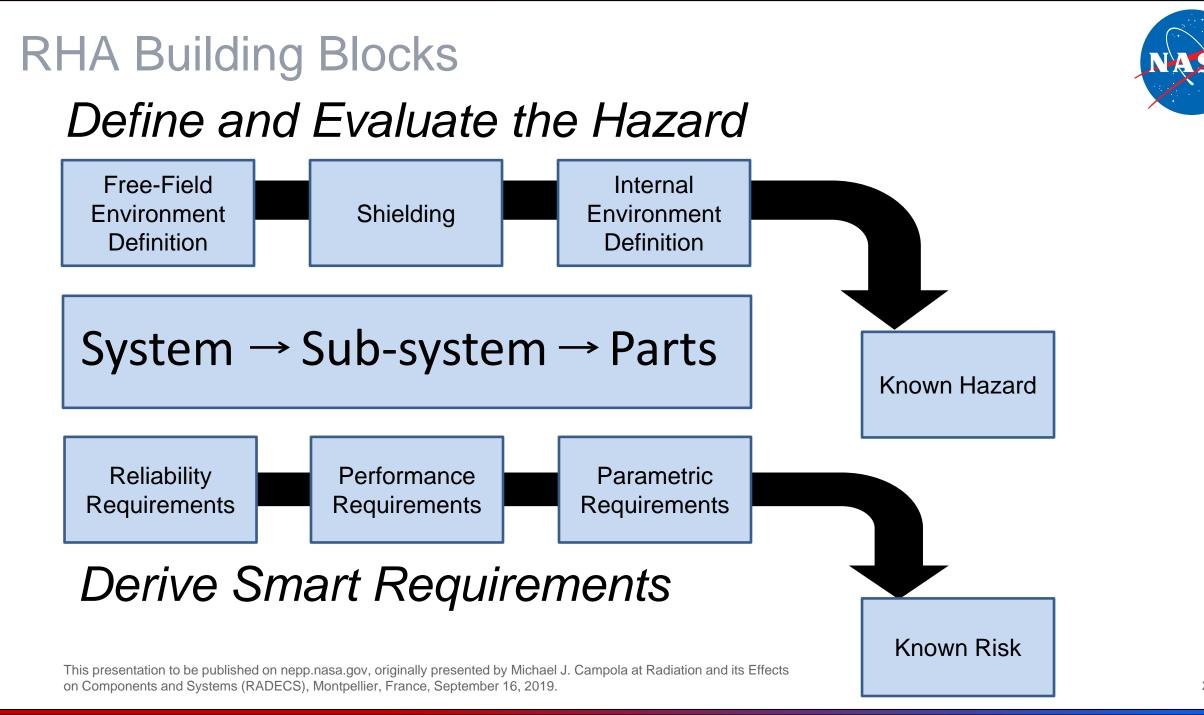
From Risk Assessment section of NASA Program Management 7120.5

Likelihood	Safety Estimated likelihood of Safety event occurrence	Technical Estimated likelihood of not meeting performance requirements	Cost Schedule Estimated likelihood of not meeting cost or schedule commitment
5 Very High	$(P_{SE} > 10^{-1})$	$(P_T > 50\%)$	(P _{CS} >75%)
4 High	$(10^{\text{-}2}\!<\!P_{SE}\!\le\!10^{\text{-}1})$	$(25\%{<}P_T{\le}50\%)$	$(50\% < P_{CS} \le 75\%)$
3 Moderate	$(10^{\text{-3}}\!<\!P_{SE}\!\le\!10^{\text{-2}})$	$(15\% < P_T \le 25\%)$	$(25\% < P_{CS} \le 50\%)$
2 Low	$(10^{\text{-5}}\!<\!P_{SE}\!\le\!10^{\text{-3}})$	$(2\% < P_T \le 15\%)$	$(10\% < P_{CS} \le 25\%)$
1 Very Low	$(10^{\text{-6}}\!<\!P_{SE}\!\le\!10^{\text{-5}})$	$(0.1\% < P_T \le 2\%)$	$(2\%\!<\!P_{CS}\!\le\!10\%\!)$



Can only get there with enough information about the system or the chosen device, need to have a known hazard and a known response

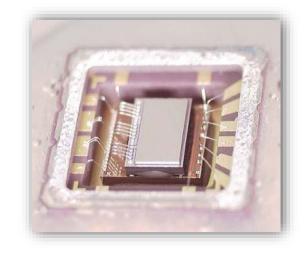




Risks Abound, What is Critical?

- Parts
 - Parametric degradation and leakage currents allowable in application?
 - Downstream/peripheral circuits considered?
 - Reset/refresh capability?
 - Mitigation within too complex?
 - Predicted radiation response unknown– loss of part functionality critical?
- Subsystem
 - Functionally required to mission that the subsystem work?
 - Interfaces allow you to get to a known state if all goes wrong?
- System
 - Increased power dissipation a mission ender?
 - Availability outweighed by error circumvention?
 - Data retention through reboots? What if there is science data loss?
 - Communications interruptions overwhelm?
 - Navigation or Attitude determination unable to deal with faults?

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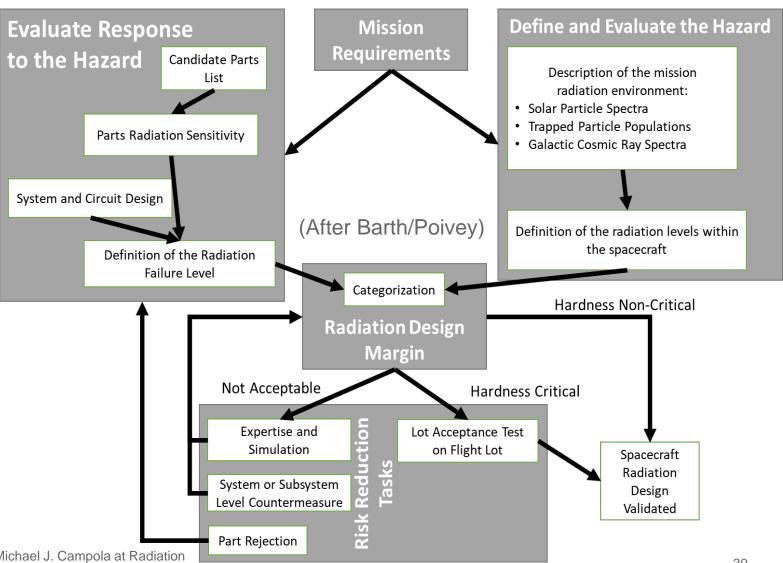
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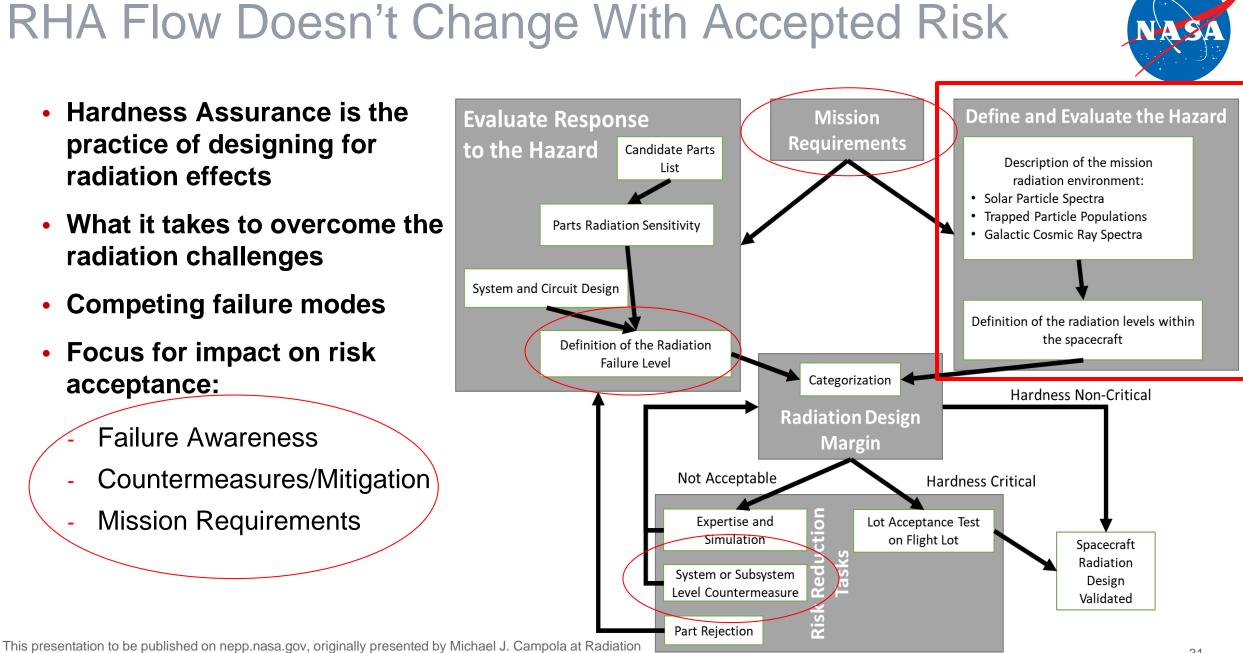


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RHA Flow Doesn't Change With Accepted Risk

- Hardness Assurance is the practice of designing for radiation effects
- What it takes to overcome the radiation challenges
- **Competing failure modes**





and its Effects on Components and Systems (RADECS), Montpellier, France, September 16, 2019.

Focus For Risk Acceptance



Failure Awareness

- Know your hazard from the natural environment
- Know your devices potential failure mechanisms or response (data)

Countermeasures and Mitigation

- Where are they necessary?
- At what level (part, card, box, mission)

Smart Requirements – and Eventually Smart Trades

Define and Evaluate the Hazard

- Define the Environment
 - External to the spacecraft
- Evaluate the Environment
 - Internal to the spacecraft
- Define the Requirements
 - Define criticality factors
- Evaluate Design/Components
 - Existing data/Testing
 - Performance characteristics
- "Engineer" with Designers
 - Parts replacement/Mitigation schemes
- Iterate Process
 - Review parts list based on updated knowledge

		Low	Medium	High
ility	High	Manageable Dose / SEE impact to survivability or availability	Moderate Dose / SEE impact to survivability or availability	High Dose / SEE impact to survivability or availability
Criticality/Availability	Medium	Manageable Dose / SEE needs mitigation	Moderate Dose / SEE needs mitigation	High Dose / SEE needs mitigation
Cri	Low	Manageable Dose / SEE do no harm	Moderate Dose / SEE do no harm	High Dose / SEE do no harm

Derive Smart Requirements

- Define the Environment
 - External to the spacecraft
- Evaluate the Environment
 - Internal to the spacecraft
- Define the Requirements
 - Define criticality factors
- Evaluate Design/Components
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 - Performance characteristics
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		Low	Medium	High
oility	High	Dose-Depth / Ray-trace GCR and Proton Spectra for typical conditions	Dose-Depth / Ray-trace GCR and proton Spectra for all conditions	Ray-Trace for subsystem / GCR and proton Spectra for all conditions
Criticality/Availability	Medium	Dose-Depth / GCR and proton spectra for background	Dose-Depth / GCR and Proton Spectra For background	Dose-Depth evaluation at shielding / All spectra conditions
Cri	Low	Similar mission dose, same solar cycle / GCR spectra	Dose-Depth / GCR spectra	Dose-Depth / GCR and Proton Spectra For background

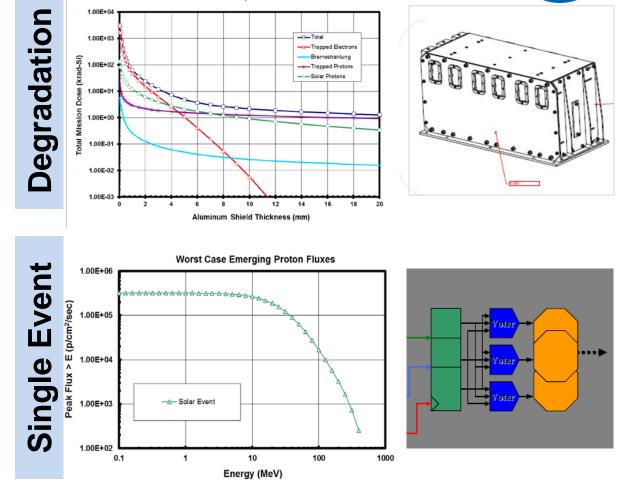


Mitigation and Countermeasure Optimization

- Define the Environment
 - External to the spacecraft
- Evaluate the Environment
 - Internal to the spacecraft
- Define the Requirements
 - Define criticality factors
- Evaluate Design/Components
 - Existing data/Testing
 - Performance characteristics
- "Engineer" with Designers
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 - Review parts list based on updated knowledge

K.A. LaBel, A.H. Johnston, J.L. Barth, R.A. Reed, C.E. Barnes, "Emerging Radiation Hardness Assurance (RHA) issues: A NASA approach for space flight programs," IEEE Trans. Nucl. Sci., pp. 2727-2736, Dec. 1998.

This presentation to be published on nepp.nasa.gov, originally presented by Michael J. Campola at Radiation and its Effects on Components and Systems (RADECS), Montpellier, France, September 16, 2019.



Dose-Depth Curves



Building Requirements

- Requirements by Environment
- Requirements by Technology
- Additional Considerations
 - **o** LET Requirements for SEE
 - o Dose Calculation
 - **o** Operation During Flare Conditions
 - o Radiation Data



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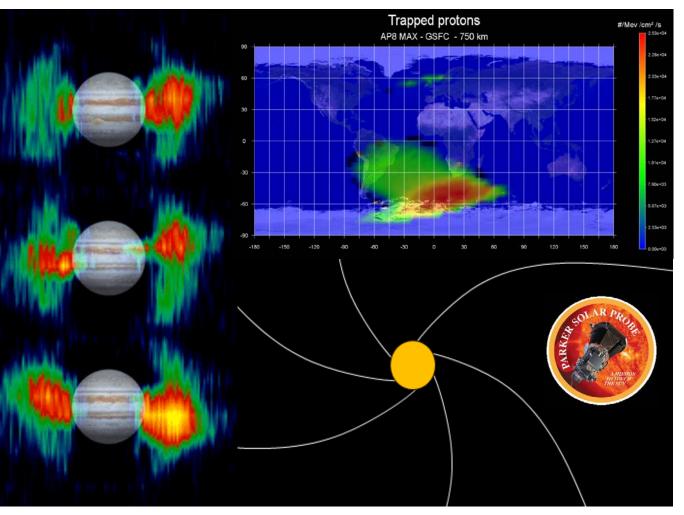
Requirements by Environment

Trapped Radiation Belts

- Can lead to high doses in a short mission:
 Jovian
- Can lead to spatially dependent SEE
 responses: South Atlantic Anomaly (SAA)
- Heliocentric Orbits
 - Solar Events, highly dynamic, energetic, directional
 - Solar Wind, will depend on the solar cycle
 - No planetary magnetic field attenuation

In essence the requirements are always driven by the environment, some more than others create a unique challenge

> NASA JPL Cassini, <u>http://saturn.jpl.nasa.gov</u>, Output from OMERE freeware <u>http://www.trad.fr/en/space/omere-software/</u>

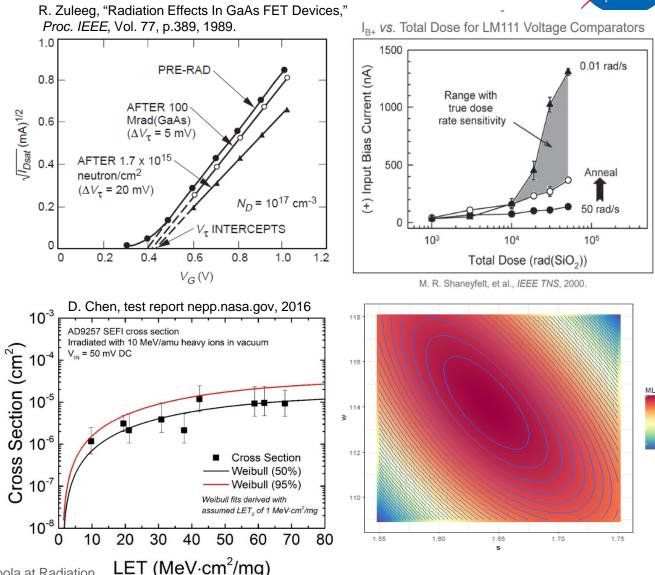




Requirements by Technology

- Technologies exhibit specific physics of failure
 - Not easy to group them all
 - Opto-electronics Displacement in the material
 - Bipolar Enhanced Low Dose Rate Sensitivity
 - Digital CMOS Latchup or SEFI
 - Power devices SEGR/SEB
 - Analog/Mixed-Signal Interruptions on PLLs, SERDES, clock dividers, etc.
- Test Data requirements
 - Failure distributions, often not enough parts
 - Destructive effects are one data point, variability from part to part
 - Statistics of the fit for rate calculations

Requirements should only be made applicable to the technologies that need to meet mission objectives and can benefit



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Considerations for SEE Requirements

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SEL •

- Environment and technology driven, risk avoidance
- Protection circuitry / diode deratings

SEGR, SEB

- Effect driven, normally incident is usually the worst 0 case
- Testing to establish Safe Operating Area (SOA) 0

SET •

- Don't harm downstream parts via 0 overvoltage/overstress on I/O, or accumulate over integrations
- Can be internal hybrids 0

SEU

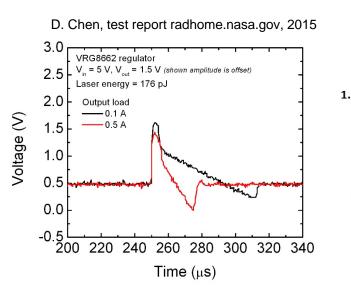
- Tailored Filtering, EDAC, or Scrubbing
- MBU, MCU, SEFI, Locked States
 - Application Voltage or Pattern dependence
 - Watchdogs / reset capability
- Proton SEE susceptible parts need evaluated in detail:
 - Low-energy proton effects:
 - May have direct ionization
 - RHA for proton sensitivity update coming:

https://nepp.nasa.gov/files/25401/Proton RHAGuide NASAAug09.pdf

FPGA Mitigation Strategies (*M. Berg*)

https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20180007760.pdf

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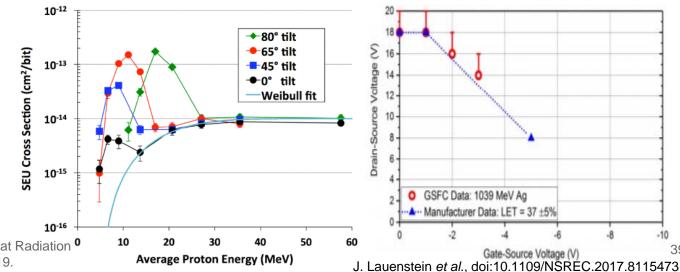


2E-10 0x00 0xAA 0xFF MLC SLC 1.5E-10 1E-10 5E-11

Pattern Dependence

LET=3.23 LET=17.2 LET=21.2 T. Wilcox, NSREC Poster DW, 2019

N. A. Dodds et al., doi: 10.1109/TNS.2015.2486763



Why You Can't Relax an LET Requirement

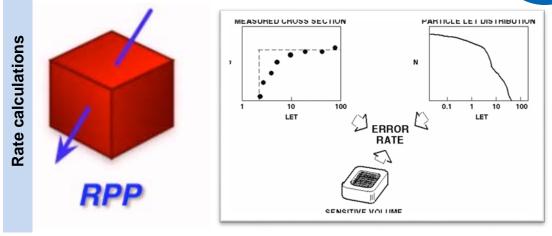
Rate calculations are not the same for Destructive vs. Non-destructive

- Data are a limiting factor, one part = one data point
- For SEE types that exist in a given technology, they present a constant risk in time domain

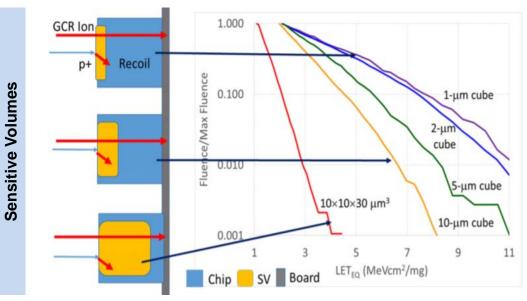
• When you require by LET:

- Spectrum from environment is imparted on sensitive volumes, where we get LET thresholds (>75 vs. 60 vs. 37 MeV·cm²/mg)
- Effective LET increases with angle critical charge is what we are trying to determine
- CRÈME calculation integrates the two
- Deep sensitive volumes won't necessarily get same LET each time with monoenergetic beams

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"Space Radiation Effects on Microelectronics," NASA Jet Propulsion Laboratory

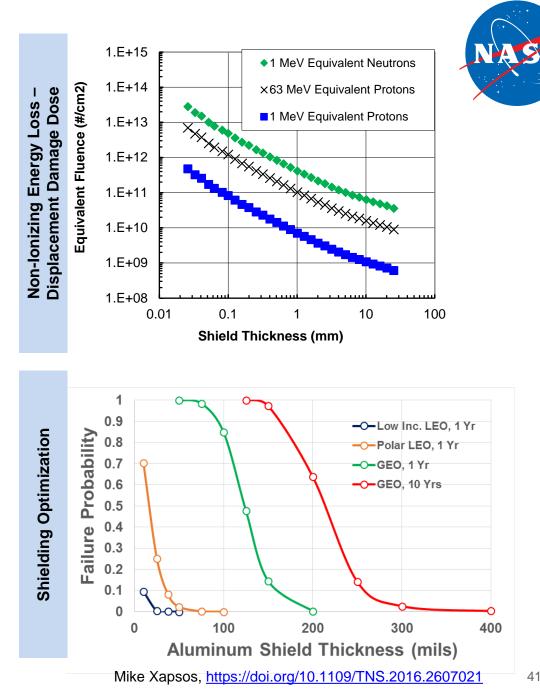


Ray Ladbury, NSREC2017 SC, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

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Appreciable Mission Doses

- Maybe degradation of a part beyond usage is okay?
 - Criticality and Application
- Did you forget about DDD?
 - External materials are susceptible as well, polymers can be bad actors and are often on commercial ground based optical systems
- Even short missions can have a common failure mode
- Low mass budget?
 - Can optimize shielding if you have failure distribution of intended components



Operation During Flare Conditions: Think Availability



Don't dose out during storm (nor the full mission)

Calculate the dose (TID/TNID) of the mission in full –

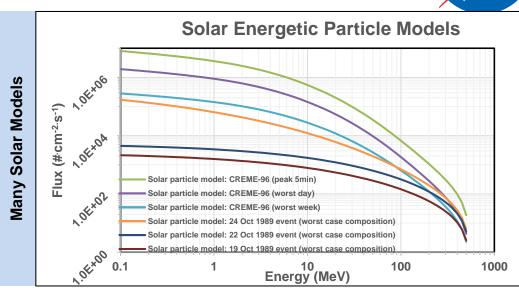
95% confidence level recommended

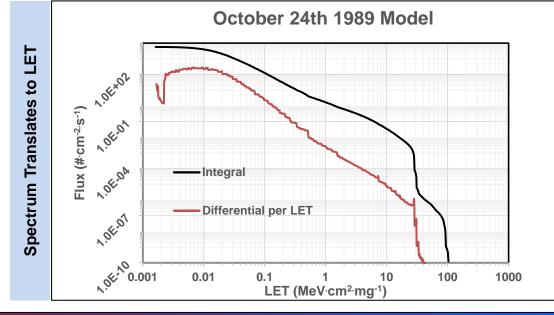
- Calculate the dose contribution from N number of events (protons & x-rays), if dose from N is > 5% of the total dose, increase confidence level of full mission model
- Don't destructively fail from a single particle during the storm (nor the full mission)
 - Standard risk-avoidant SEE approach: no destructive effects allowed
 - LET threshold for single event latchup (SEL)
 - > 75 MeV.cm²/mg (some use 60 MeV.cm²/mg)
 - LET threshold for single event burnout, gate rupture, dielectric rupture (SEB, SEGR)

> 37 MeV.cm²/mg (particles must come from normal incidence to cause effect)

If you have non-destructive single event upsets, they can't overwhelm critical instruments/systems during the storm

- Rate calculation requires part data representative of the application, looking for crosssection over LET.
- If parts' LET threshold from 20 to 75 MeV.cm²/mg, need heavy ion rate
- If parts' LET threshold is below 20, need indirect ionization from recoil ions contribution to rate (need proton data) – make sure packaging materials don't add to this, direct ionization from protons (can be built-in to heavy ion calculation) possible
- Do you need to mitigate or not confirm that event rates are not higher than mitigation (Markov process... i.e. EDAC beats the number accrued, scrub rate is faster than critical number of upset accumulation)





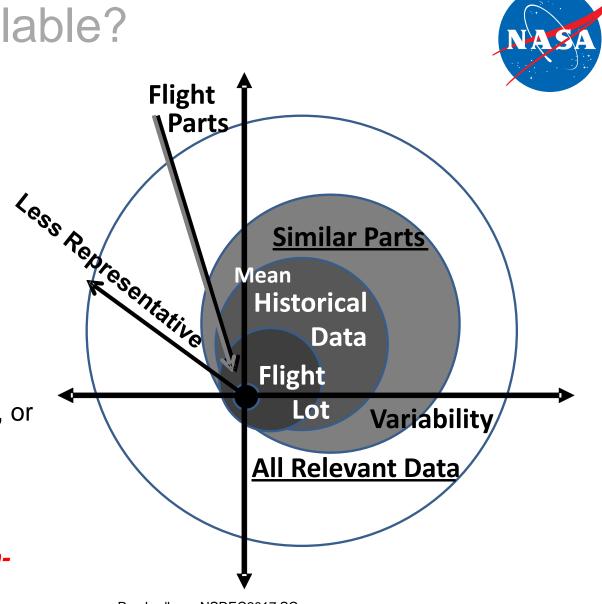
Risk Acceptance – Data Available?

- Part Classifications Growing
 - Mil/Aero vs. Industrial vs. Medical
 - Automotive vs. Commercial vs. Modified HiRel

Substitute COTS in this diagram

- Now you have another degree of separation
- Failure modes not fully understood
- Unlikely to have historical data
- Similarity data no applicable due to fab, process, or design rules
- Cost of testing usually too high

Without traceability you may be depending on nonrepresentative data.



Ray Ladbury, NSREC2017 SC, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

Notional Radiation Data Collection Guidelines



Environment

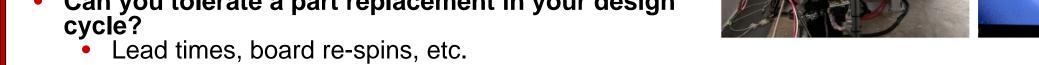
		LEO Equatorial	LEO Polar (Sun Sync)	GEO / Interplanetary
Mission Lifetime (With Assumed Risk Acceptance)	> 3 Years	Data on all SEE for critical parts, and have data on dose failure distribution on similar parts	Consider mission consequences of all SEE (Data for critical parts), have Dose failure distribution on lot	Have Data on all SEE, Have Data Dose failure distribution on lot
	1- 3 Years	Have Data on DSEE for critical parts	Consider mission consequences of all SEE (Data for critical parts), have data Dose failure distribution on similar parts	Have Data on all SEE for critical parts, Have Data on Dose failure distribution on similar parts
	< 1 Year	Look for data on DSEE for critical parts	Consider mission consequences of all SEE, and look for data on dose failure distribution on similar parts	Consider mission consequences of all SEE, and have data on dose failure distribution on similar parts

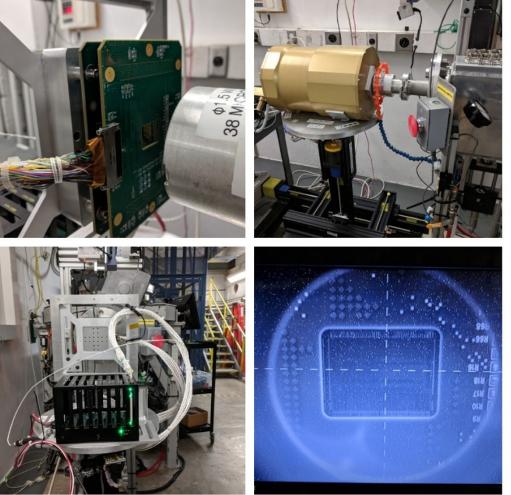
When Do You Test? When Do You Model?

- Divine your risk threshold
 - There's a doc coming for that... radhome.gsfc.nasa.gov/nepp.nasa.gov
- Unknown failure modes that would not be acceptable to the mission
 - Known unknowns can be carried as a risk if you already know that the outcome is mitigated at the board or box level
 - New technologies should be identified early on
- Fault propagation may be the problem you wish to mitigate
 - This can include cumulative effects!
 - Fault injection may not be able to cover the state space
- Destructive single event effects are an obvious target
- Can you tolerate a part replacement in your design cycle?
 - Lead times, board re-spins, etc.

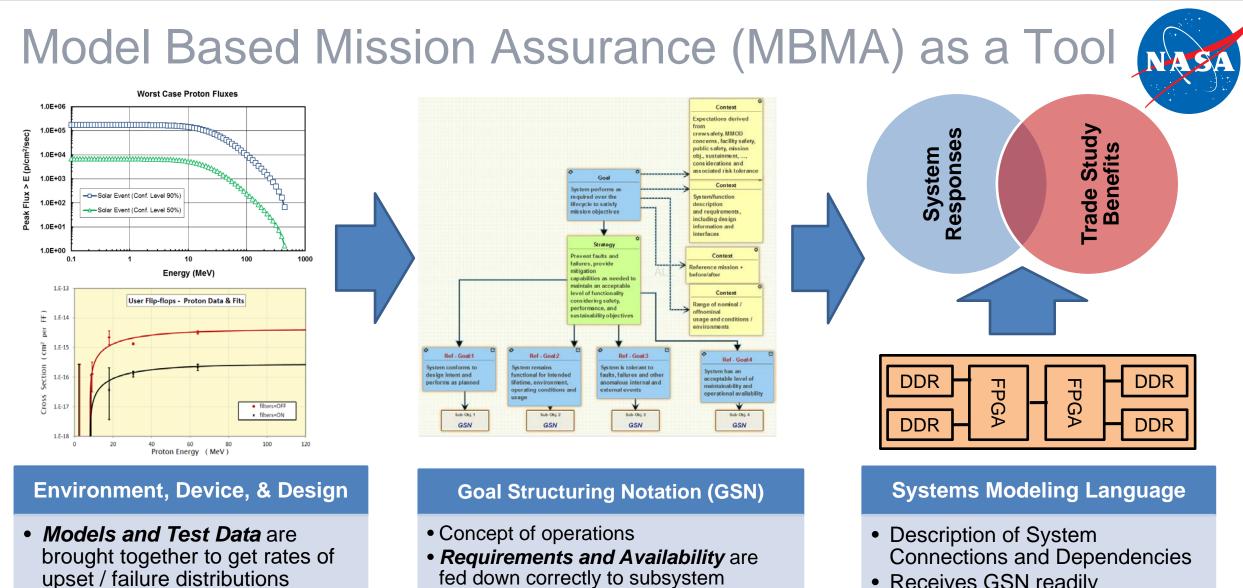
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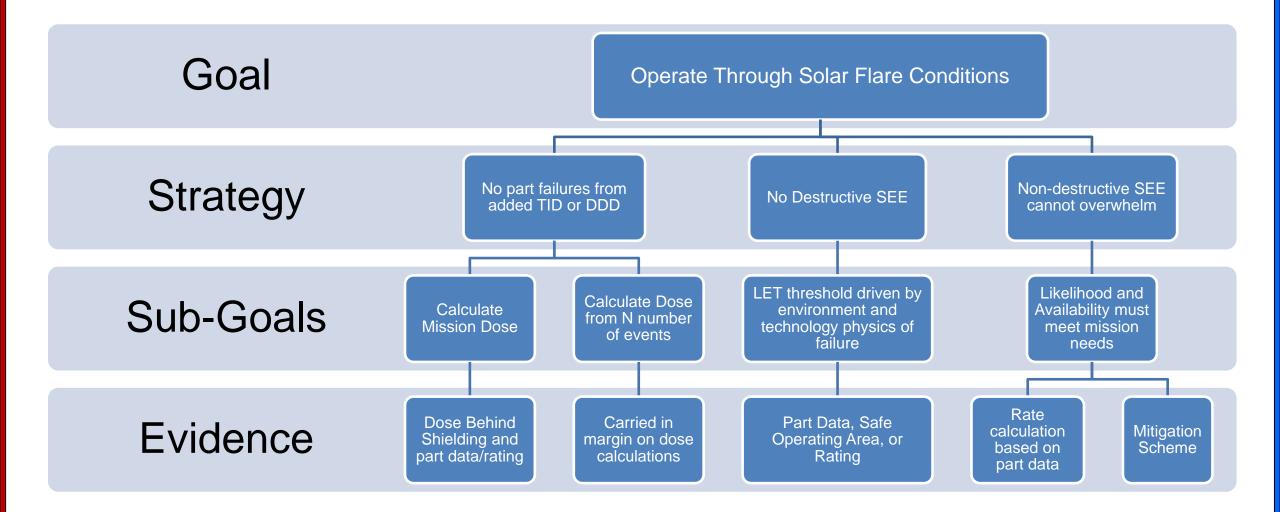


• Resources and Utilization are the scaling factors with criticality

- fed down correctly to subsystem
- Evidence is presented
- Assumptions are tracked

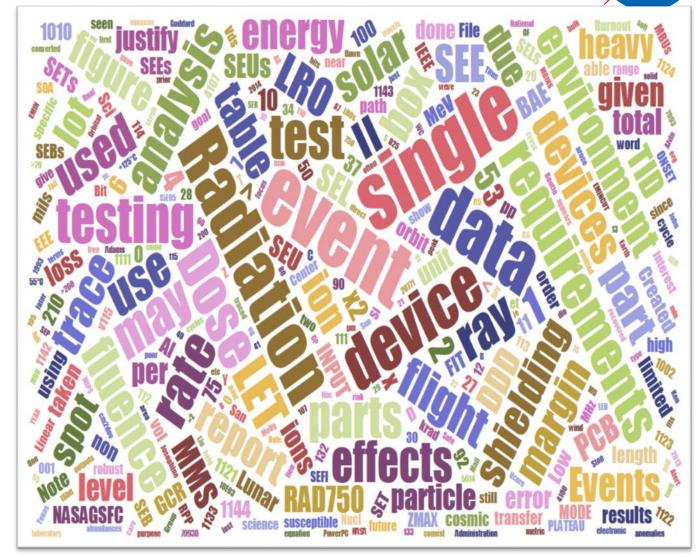
- Receives GSN readily
- Fault propagation can be identified

Goal Structuring Notation (GSN)



Questions to Keep in Mind

- What are the radiation risks:
 - What is the hazard?
 - What are the challenges?
- What can you do to reduce the risk for a given hazard?
- What does changing that radiation environment mean for success?
- Need availability throughout the mission or at specific times?
- How do *similar* systems/devices react in the space environment?







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THANK YOU