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**Dust source identification using MODIS: a comparison of techniques applied to the Lake Eyre Basin, Australia**

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**Abstract**

The impact of mineral aerosol (dust) in the Earth’s system depends on particle characteristics which are initially determined by the terrestrial sources from which the sediments are entrained. Remote sensing is an established method for the detection and mapping of dust events, and has recently been used to identify dust source locations with varying degrees of success. This paper compares and evaluates five principal methods, using MODIS Level 1B and MODIS Level 2 aerosol data, to: (a) differentiate dust (mineral aerosol) from non-dust, and (2) determine the extent to which they enable the source of the dust to be discerned. The five MODIS L1B methods used here are: (1) un-processed false colour composite (FCC), (2) brightness temperature difference, (3) Ackerman’s (1997: *J.Geophysical Res.*, 102, 17069-17080) procedure, (4) Miller’s (2003:*Geophysical Res. Lett.* 30, 20, art.no.2071) dust enhancement algorithm and (5) Roskovensky and Liou’s (2005: *Geophysical Res. Lett.* 32, L12809) dust differentiation algorithm; the aerosol product is MODIS Deep Blue (Hsu et al., 2004: *IEEE Trans. Geosci. Rem. Sensing*, 42, 557-569), which is optimised for use over bright surfaces (i.e. deserts). These
are applied to four significant dust events from the Lake Eyre Basin, Australia. OMI AI was also examined for each event to provide an independent assessment of dust presence and plume location. All of the techniques were successful in detecting dust when compared to FCCs, but the most effective technique for source determination varied from event to event depending on factors such as cloud cover, dust plume mineralogy and surface reflectance. Significantly, to optimise dust detection using the MODIS L1B approaches, the recommended dust/non-dust thresholds had to be considerably adjusted on an event by event basis. MODIS L2 aerosol data retrievals were also found to vary in quality significantly between events; being affected in particular by cloud masking difficulties. In general, we find that OMI AI and MODIS AQUA L1B and L2 data are complementary; the former are ideal for initial dust detection, the latter can be used to both identify plumes and sources at high spatial resolution. Overall, approaches using brightness temperature difference (BT10-11) are the most consistently reliable technique for dust source identification in the Lake Eyre Basin. One reason for this is that this enclosed basin contains multiple dust sources with contrasting geochemical signatures. In this instance, BTD data are not affected significantly by perturbations in dust mineralogy. However, the other algorithms tested (including MODIS Deep Blue) were all influenced by ground surface reflectance or dust mineralogy; making it impossible to use one single MODIS L1B or L2 data type for all events (or even for a single multiple-plume event). There is, however, considerable potential to exploit this anomaly, and to use dust detection algorithms to obtain information about dust mineralogy.
1. Introduction

Atmospheric mineral aerosols (termed here dust) play an important role in the land-atmosphere-ocean system (Ridgwell, 2002; Jickells et al., 2005; Waeles et al., 2007). For example, they affect soil nutrients at source and sink (McTainsh & Strong, 2007; Muhs et al., 2007; Li et al., 2007; Reynolds et al., 2006; Soderberg & Compton, 2007; Swap et al., 1992; Wang et al., 2006), the radiative forcing of the atmosphere (Haywood & Boucher, 2000; Hsu et al., 2000; Satheesh & Moorthy, 2005; Yoshioka et al., 2007) and may regulate phytoplankton activity of oceans (de Baar et al., 2005; Erickson et al., 2003; Mackie et al., 2008; Piketh et al., 2000; Wolff et al., 2006). The impact of dust in the Earth's system depends on characteristics such as particle size, shape and mineralogy (in particular iron content: Jickells et al., 2005; Mahowald et al., 2005). Whilst these characteristics can change during dust transport (Desboeufs, 2005; Mackie et al., 2005) they are initially determined by the terrestrial sources from which the particles are entrained.

The detection and mapping of dust events and dust transport pathways has benefited greatly from the use of remote sensing, and at the global scale major dust source regions have been identified using satellite data, such as from the Total Ozone Mapping Spectrometer (TOMS; Prospero et al., 2002; Washington et al., 2003). The passage of dust along specific regional transport pathways over land and ocean and the behavior of individual dust events have also been tracked using TOMS and OMI (Ozone Monitoring Instrument; e.g. Alpert et al. 2004) and at higher temporal and spatial resolutions using data from, amongst others, AVHRR (Advanced Very High Resolution Radiometer; e.g., Evan et al., 2006; Zhu et al., 2007), GOES-VISSR (Geostationary Operational Environmental Satellite, Visible Infra-Red Spin-Scan Radiometer, e.g., MacKinnon et al., 1996), METEOSAT (e.g., Moorby et al., 2007), MODIS (Moderate Resolution Imaging Spectroradiometer, e.g., Badarinath et al., 2007; Gassó & Stein, 2007;
Systematic determination of both the geomorphological and geochemical variability of dust sources, and hence the variability of the sediments which are entrained and transported, requires as accurate and precise an identification of the upwind (source) end of the dust plume as possible. Researchers have recently started to use remote sensing data to achieve this (e.g., Bullard et al., 2008; Lee et al., 2008; Zhang et al., 2008), but with varying levels of success. The ability to use remotely-sensed data both to detect a dust plume and identify the location from which it has originated is affected by several factors including the radiative transfer properties of the material emitted, the radiative properties of the ground/ocean surface over which the plume is transported, the size and density of the dust plume, the time of satellite overpass relative to dust emission, the presence or absence of cloud, the horizontal and vertical plume trajectory, and the sensor characteristics and radiative transfer model used to detect dust. In many respects, the relative impacts of these factors on dust source determination are hard to determine without close reference to surface meteorological data (e.g. wind speed and visibility records) and ground-based aerosol determination records (e.g. AERONET – Aerosol Robotic Network) which can allow comparative characterisation of individual dust events (e.g. Bullard et al., 2008; Mahowald et al., 2007). Even where these records exist, the direct comparison of ground and remote sensing data retrievals to determine dust sources can be problematic, with some remote sensing data products being unable consistently to detect dust events due to the factors listed above; particularly the presence of cloud, and the existence of low contrast between...
dust plume and ground/ocean surface (e.g. Gassó & Stein, 2007; Bullard et al., 2008). The principal aim of this paper is to evaluate in detail the use of MODIS data, one of the most widely and successfully-used sensors, for improved identification of dust source locations. This paper varies in emphasis from many previous studies because the focus is on the precision with which the upwind (source) location of the plume can be discerned, rather than on the simple determination of plume location, density and trajectory. Specifically, we compare five methods of using MODIS Level 1 band data and one MODIS Level 2 aerosol product and evaluate them in terms of: (a) how well they enable the differentiation of dust and non-dust (cloud, smoke, volcanic aerosols) and, (b) the extent to which it is possible to discern the location of the dust source (i.e. the upwind part of the dust plume - or ‘dust head’) and how much this varies from method to method. The influence of environmental factors such as plume density and mineralogy on source detection by MODIS will also be evaluated.

2. Data and Methods

2.1 Data

Mineral aerosol (dust) can be detected and mapped through remote sensing via inversion of radiative transfer models which operate in the following wavelengths: (a) ultraviolet (UV 0.315-0.4 μm) via absorption (e.g. TOMS AI; Torres et al., 1998), (b) visible (VIS 0.38-79 μm) via scattering (e.g. Tanré and Legrand, 1991), and (c) thermal infrared (TIR 8-15 μm) via contrasting land/aerosol emissivity and/or temperature (e.g. Ackerman, 1997). Due to constraints of sensor design, observations by remote sensing systems operating in VIS wavelengths can be determined at higher resolution (pixel size = x) than those made in the TIR (pixel size = x*2-4) and UV (pixel size = x*100-200), and this has implications for both plume and source detection using these approaches. Radiative transfer model inversion of aerosol observations made within (or via combinations of) each of the three wavelength ranges often provides either a relative indication of aerosol concentration (e.g. via TOMS AI), or a calibrated (e.g. through comparison...
with AERONET observations) measure of wavelength-dependent total aerosol optical thickness/depth (AOT/D). The success of the radiative transfer model inversion in each case is often complicated by factors such as the non-spherical nature of the mineral aerosol, changes in the chemical/physical nature of the material, and location within the atmosphere during transport. In addition, over very bright surfaces (e.g. desert regions and urban areas), in the presence of cloud, and at night, mineral aerosol detection using UV/VIS/TIR wavelengths can become increasingly uncertain (e.g. Kaufman et al., 2000). The short-term nature of some mineral aerosol events (often <1 day) also means that an understanding of any bias associated with mineral aerosol detection at the time of satellite over-passes and temporal sampling (i.e. either am or pm data collection time) is needed in order to characterize fully the emission and transport process. In order to evaluate, compare and contrast mineral aerosol detection approaches, a range of remote sensing data are used here (see Table 1).

2.1.1 MODIS Data

Data from the Moderate Resolution Imaging Spectroradiometer (MODIS) were used to make comparisons of retrievals using VIS and TIR (often combined) approaches. MODIS makes observations using 36 spectral bands with wavelengths from 0.41 to 14.4 µm and nadir spatial resolutions of 0.25 km, 0.5 km, and 1 km. It is currently operating onboard the NASA Earth Observing System (EOS) Terra and Aqua satellites, launched in December 1999 and May 2002, respectively. Daily MODIS Level 1B (L1B) 1 km data (MOD021KM = Terra, and MYD021KM = Aqua) used in this work have been processed to convert the sensor’s on-orbit responses in digital numbers to radiometrically calibrated and geo-located data products (v5.06 processing for Terra and v5.07 for Aqua). Data were obtained from the Level 1 and Atmosphere Archive and Distribution System (LAADS; http://ladsweb.nascom.nasa.gov/). Details of images dates and subsequent processing of MODIS L1B data are outlined below.
Daily MODIS Level 2 Aerosol data are produced at the spatial resolution of a 10 x 10 km (at nadir) pixel array. There are two MODIS Aerosol data product file types: MOD04_L2, containing data collected from the Terra platform and MYD04_L2, containing data collected from the Aqua platform. Here we only use the MYD04 Aqua product because to date Deep Blue (see below) retrievals are not yet available for MOD04 Terra data. Aerosol properties within MYD04_L2 are derived by the inversion of MODIS observed reflectances at 500 m resolution using pre-computed radiative transfer look-up tables based on dynamical aerosol models (Kaufman et al., 1997; Remer et al., 2005). Derivation of aerosol from these data is far from straightforward and, in initial versions of the MODIS aerosol product, the ability to retrieve aerosol optical thickness (AOT) and single scattering albedo over bright-reflecting surfaces has been problematic because the algorithm relies in part on the initial detection of dark surfaces or targets (Kaufman et al., 2000). In addition, the cloud screening has been shown to have problems where mis-identification of some dust plumes as cloud has led to artifacts in the final data (e.g. as noted by Brindley and Ignatov, 2006). These products have been under continued and careful evaluation and development, and product MYD04 (see http://modis-atmos.gsfc.nasa.gov/C005_Changes/C005_Aerosol_5.2.pdf) has recently received an improved aerosol determination (via reprocessing to collection 5.1/2; Levy et al., 2006, 2007; Remer et al., 2006) over bright surfaces through the integration of a revised determination of AOT over land (Levy et al., 2007), and inclusion of the Deep Blue algorithm (Hsu et al., 2004; Hsu et al., 2006). Here we evaluate the Deep Blue algorithm, which relies on the blue wavelengths and libraries of surface reflectance to make retrievals over bright surfaces (Hsu et al., 2004).

The Deep Blue processing approach involves the following processing elements: (1) Rayleigh Correction for Terrain Elevation in the following MODIS channels: R8 (0.405-0.42 µm), R3 (0.459-0.479 µm) and R1 (0.62-0.67 µm); (2) Cloud Screening using: R8 (3 x 3 pixel spatial variance) and R3/R8 AI; (3) the surface reflectance for a given pixel is determined from a clear-scene database based upon its geo-location; (4) R8, R3 and R1
reflectances are then compared to radiances contained in a lookup table with
dimensions consisting of solar zenith, satellite zenith, and relative azimuth
angles, surface reflectance, AOT, and single scattering albedo; (5) a
maximum likelihood method is used to compute a mixing ratio between dust
and smoke models until the calculated spectral reflectances make the best
match with those that are measured; and (6) for mixed aerosol conditions,
once the aerosol models and the mixing ratio that produce the best match are
determined, the values of AOT and Ångström exponent are reported. For
dust-dominant cases, the values of single scattering albedo are retrieved in
addition to these parameters. MODIS Deep Blue data within MYD04_L2
includes AOT ($\tau$) determination at 0.412, 0.47, 0.55 and 0.66 µm, although
only the 0.412 µm data are used here. MYD04_L2 data were obtained from
the Level 1 and Atmosphere Archive and Distribution System (LAADS;
http://ladsweb.nascom.nasa.gov/). The typical aerosol optical thickness for
visible light in clear air is 0.1, very hazy skies have AOTs of $\geq$0.3. During initial
processing, typical scale (0.001) and offset (0) values were applied to
MYD04_L2 AOT data prior to display and subsequent data processing.

2.1.2 AURA OMI

This paper focuses on an evaluation of MODIS data but for each case
study, in addition to MODIS L1B and L2 aerosol data, co-incident data from
an independent sensor, the Ozone Monitoring Instrument (OMI) were also
acquired. OMI is on the Aura satellite (launch date: July 2004) which flies as
part of the NASA A-Train constellation (http://aqua.nasa.gov/doc/pubs/A-
Train_Fact_sheet.pdf) a few minutes behind the Aqua satellite. OMI is
designed to continue the Total Ozone Mapping Spectrometer (TOMS) record
for total ozone and other atmospheric parameters related to ozone chemistry
and climate. OMI measurements are sensitive to aerosol absorption in UV
wavelengths, thus providing an independent source of information relating to
mineral aerosol detection in the scene under observation. In addition, and
unlike MODIS, OMI AAI (Absorbing Aerosol Index: e.g. Torres et al., 2007) is
sensitive to aerosol absorption even when the particles are above cloud and
AAI is therefore derived successfully in both cloudless and cloudy conditions
(although see Ahn et al., 2008). OMI has a ground resolution of 13 x 24 km
(nadir) and uses a retrieval algorithm similar to the one used by TOMS (Torres et al., 1998). The OMI AI is defined as follows:

\[
OMI \text{ AI} = 100 \log_{10}\left( \frac{I_{360}^{\text{Meas}}}{I_{360}^{\text{Calc}}} \right)
\]

(Eq.1)

where \(I_{360}^{\text{Meas}}\) is the measured 360 nm OMI radiance and \(I_{360}^{\text{Calc}}\) is the calculated 360 nm OMI radiance for a Rayleigh atmosphere. Under most conditions, the AI (Eq.1) is positive for absorbing aerosols and negative for non-absorbing aerosols (pure scattering). An AI >1 is typical of absorbing aerosols such as smoke or dust (Gassó & Stein, 2007; Kubilay et al., 2005; Washington et al., 2003). In this instance, we have chosen to use the OMI-Aura_OMTO3E data, which is a daily Level 3 global gridded product which is generated by binning the original pixels from the Level 2 data products (15 orbits per day; 13 x 24 km spatial resolution at nadir) into a 0.25 x 0.25 degree global grid.

2.2 Methods

2.2.1 Study region and event selection

The performance of different MODIS dust detection methods in identifying source locations involved the analysis of four dust events which all originate in the same drainage basin. The Lake Eyre Basin (LEB), Australia was chosen for several reasons. First, it has been identified as a persistent and significant southern hemisphere dust source on the basis of surface observations (Middleton, 1986) and using TOMS AI (Washington et al., 2003). Second, it is the only inland basin dust source region in Australia, a geographically-isolated continent distant from other dust sources. Consequently, within the LEB there is less potential for interaction with other major dust sources than would be the case, for example, in the Sahara (Prospero et al. 2002) or China (Shao & Wang, 2003). Third, the basin is large enough to give rise to several major dust events each year, but not such an intense dust source as to make it difficult to discern individual plumes.
The LEB covers 1.14 million km$^2$, with mean annual rainfall of less than 125 mm and annual potential evaporation in excess of 2500 mm. There are several different sedimentary environments in the LEB, all of which emit dust. The most significant of these are: (1) aeolian deposits covering 33% of the basin area and accounting for 37% of the dust plumes, (2) alluvial deposits and floodplains (11.55% area, 30% dust plumes), and (3) ephemeral lakes and playas which cover only 2.26% of the basin area but from which originate 29% of the dust plumes making these the most intense dust sources (figures averaged over 2003-6; Bullard et al., 2008). Inter- and intra-annual variability of dust storm frequency in the LEB is high, responding to changes in synoptic pressure distributions across the continent (Ekstrom et al., 2004). The Sprigg Model, which characterizes dust transporting wind systems in Australia (Sprigg, 1982), suggests that as frontal systems pass over the LEB, pre-frontal northerly and post-frontal southerly winds can entrain dusts which travel southeast or northwest respectively. It is important to note, however, that the estimated total annual number of dust events in the LEB varies not only in response to climate but also as a result of differences in how events are defined. In a previous study (Bullard et al., 2008), we examined MODIS imagery for all days (between July 2003 and June 2006) where at least one meteorological station in the LEB (or within 250 km of the catchment boundary) recorded a dust-induced reduction in visibility to $\leq$1 km (which corresponds to the WMO definition of a dust storm). Whilst there are some inconsistencies in the relationship between visibility records from meteorological stations and other indicators of dust emissions, (including AERONET, TOMS AI and TOMS AOD: Mahowald et al., 2007) and the spatial distribution of meteorological stations across the arid LEB is sparse which means a number of events will be missed, visibility remains a useful criteria for identifying days on which significant dust events have occurred. From the 43 days on which dust events were identified four case studies were chosen to illustrate key types of event that occur in the LEB, and also to include factors which can significantly affect dust plume and source identification (i.e. single/multiple dust plumes and varying amounts of cloud; see Figure 1; Table 2). Although there are versions of some dust detection algorithms designed to work at night (e.g. Wald et al., 1998), we focus on daytime events.
here so that the influence of surface reflectance on dust source identification can be explored.

<Insert Figure 1> -
<Insert Table 2>

2.2.2 MODIS Level 1B Processing Algorithms

As outlined earlier, mineral aerosol is sometimes detectable on un-adjusted VIS satellite images (particularly over the ocean), but because mineral aerosol can have similar reflectivity to the desert surfaces from which it is entrained it can be difficult to detect over land. In addition, mineral aerosol is often hard to differentiate from cloud, sea salt and anthropogenic pollution. As a result of this, and also due to problems with the performance of MODIS L2 aerosol products (section 2.1.1), a number of studies have used changes in brightness temperature (TIR) to detect mineral aerosol over land surfaces. Initial attempts using single TIR channel data, such as that by Shenk & Curran (1974) using Nimbus-THIR (Temperature Humidity Infrared Radiometer) 11 µm data, had limited success because changes in surface emissivity at this wavelength can be misinterpreted as dust (Roskovensky & Liou, 2003; 2005). As a result of observed variability in the emissive and transmissive nature of mineral aerosols within multiple TIR wavelength ranges, other researchers have used methods based on brightness temperature difference (BTD) in either two or three wavelength ranges, typically 11-12 µm bands (bi-spectral split window technique) or near 8, 11 and 12 µm bands (tri-spectral) (e.g. Ackerman, 1997). BTD values from this method reveal temperature differences that exist between the ground surface and cooler mineral aerosol while at the same time are largely unaffected by absorption from other atmospheric gases (Darmenov & Sokolik, 2005). In addition to detecting dust over land, these approaches may also allow discrimination between cloud and dust when both exist in the vicinity of each other.

Here we initially apply the simple BTD approach detailed by Ackerman (1997) to MODIS L1B data (Table 3). Using this methodology it has been inferred that (BTD; 11.03-12.02 µm or MODIS BT_{31}-BT_{32}) values <0 K signify
the presence of mineral aerosol (dimensionless) and BTD values ≥0 K indicate no mineral aerosol. While developing the MODIS cloud mask, Ackerman et al. (2002) have also placed the mineral aerosol detection threshold at < -1 K. Although Ackerman’s (1997) analysis implied that the 0 K threshold could be widely used over a range of land surfaces, it is likely that this will vary slightly according to variability in the emissive/transmissive nature of the mineral aerosol. This in turn is determined by factors such as mineralogy as well as processes acting upon the aerosol as it is transported in the atmosphere. Mineralogical composition is an important control on the TIR radiative properties of mineral aerosol and can vary significantly from region to region (e.g. Claquin et al., 1999; Caquineau et al., 2002; Satheesh & Moorthy, 2005). Darmenov and Sokolik (2005) investigated the TIR radiative signature of dust transported over oceans from 7 different regions and located the BTD (11.03 -12.02 µm) aerosol detection threshold at 0.5, -0.2, -1.0 and -0.4 K for the Nubian, Thar, Gobi/Taklimakan and Australian deserts respectively; but could not locate a clear threshold to distinguish mineral aerosol from cloud for dust over oceans sourced from NW Africa, Libya or the Iranian desert. It may also be the case that the threshold varies for a single geographical region, the precise value being dependent on factors such as the density of the dust plume (Darmenov and Sokolik, 2005) or local variation in dust source mineralogy (e.g. iron-rich sources versus illite-rich sources). This simple bi-spectral split window approach will be applied here to identify appropriate aerosol detection thresholds over land for the Lake Eyre Basin.

Using BTD as a basis, a range of more complex algorithms has been developed that combine BTD and VIS wavelengths to detect mineral aerosol over land and remove the effects of dense cloud cover, which can obscure dust, and cirrus clouds which have similar reflectance and BTD properties to fine dust particles. In this paper we evaluate two of these cloud-removal approaches applying them to MODIS L1B data for the LEB. The first is the multispectral dust enhancement algorithm of Miller (2003) which exploits the fact that dust particles can have contrasting VIS reflective properties when compared to cloud (Table 3). In this model, an inverse brightness temperature difference is used (BTD; 12.02-11.03 µm or MODIS BT\textsubscript{32} – BT\textsubscript{31})
which is rescaled/normalized to lie within the $-2$ to $+2$ K range. Based on Miller's (2003) algorithm the mineral aerosol output (D) has values constrained between 1.3 and 2.7 (dimensionless). In addition this approach, through manipulation of the red (R), green (G) and blue (B) display, enables mineral aerosol to be visually differentiated from cloud using colour (D is loaded on the red color gun). The second approach is that of Roskovensky and Liou (2005) and Hansell et al., (2007) which focuses on the differentiation of mineral aerosol from cirrus clouds by combining BTD (11.03-12.02 µm) and VIS wavelengths (reflectance ratio of 0.54 µm/0.86 µm). In the final output image, values of D$>1$ (dimensionless) indicate mineral aerosol is present and values ≤1 indicate cirrus cloud or non-mineral aerosol in the scene (Table 3). Inclusion of the reflectance ratio in this case reduces the amount of false detection of dust over land observed by Ackerman et al. (2002).

Although the majority of studies cited above have used data from MODIS it is worth noting that similar approaches have been explored using data from other sensors such as AVHRR, HIRS/2, GOES-8 and MSG-SEVIRI with varying degrees of success (e.g. Legrand et al. 1989, Sokolik, 2002, Schepanski et al., 2007).

Table 3 includes a summary of the default threshold values used to differentiate dust from non-dust in each of the original algorithms used here. The threshold values used in these algorithms are sensitive to varying atmospheric conditions, surface reflectance, dust density and dust mineralogy, but are formulated in a manner such that they allow a certain degree of tuning to adjust for specific conditions such as regional variability (Darmonov & Sokolik, 2005), or for dust blowing over land or ocean (Roskovensky & Liou (2005). In this study, we verified the published models by using the authors' original data and study-events both to check the set up of the algorithms and to ensure we could reproduce the initial values of dust/non-dust threshold and coefficients used. For the four case studies presented here, we therefore established event-specific thresholds using the
approach suggested by each author (Table 4). The definition of thresholds therefore involved the interrogation of pixel histograms for each scene-algorithm combination (see below). In each case, peaks were found to be attributable to specific scene components (e.g. densities and types of cloud and aerosol), and thresholds were chosen to represent the value which was best able to identify dust in the scene, as judged by the user (Figure 2a). In each case, peaks representing dust were relatively easy to identify and disinter from other scene components, and the effects of the choice of dust threshold in each case is outlined below. Given the rather inflexible nature of this approach, data from other study regions, where atmospheric conditions and water vapor concentrations vary more significantly, may pose a challenge to the straightforward identification and threshold determination for dust peaks outlined here.

Figure 2 shows the histograms used to derive BTD thresholds and Miller’s D for each event. For the Roskovensky and Liou (2005) output the dust/non-dust threshold remained at 1, but the D-parameter scaling factor ‘a’ and BTD offset ‘b’ were adjusted for each event by using the midpoints between the clear sky and dust histograms of the reflectance ratio and BTD respectively (Figure 3).

2.2.3 Evaluation of output images

To evaluate the different MODIS dust detection algorithms it is necessary to have a common reference against which to compare the output data. For each of the four dust events examined, an eight panel figure was produced. In each case, panel (a) represents the MODIS VIS image (where red = band 1, green = band 4, blue = band 3). Panel (b) represents the bi-spectral brightness temperature difference (\(\text{BTD} = BT_{31} - BT_{32}\)) with no dust threshold applied. This is the principal image against which the outputs from the different algorithms outlined in panels (c) Ackerman (1997), (d) Roskovensky & Liou (2005) and (e) Miller (2003) were compared, because a
bi-spectral BTD is a common component of each of these Level 1B MODIS algorithms. What is evaluated therefore is the extent to which the additional components of the algorithms actually led to improved dust source detection. In addition, a simple objective comparison of the outputs from each of the Ackerman (1997), Roskovensky & Liou (2005) and Miller (2003) algorithms is shown in panel (f). To produce this, boolean outputs from panels (c), (d) and (e), where pixels were categorized as dust (=1) or non-dust (=0) were colored red, green and blue respectively and combined to create a color composite output image. For example, if a pixel was categorized as dust following Ackerman’s (1997) procedure it will appear red; if categorized as dust by both Ackerman (1997) and Miller (2003) it will appear pink; if all three algorithms categorize it as dust it will appear white; if all three categorize it as non-dust it will appear black (Figure 4).

Panels (g) and (h) represent the two dust products, MODIS L2 aerosol (MYD04) Deep Blue AOT and OMI AI respectively. Although the spatial resolution of the data is lower than the MODIS, OMI AI provides an independent check on the spatial extent and intensity of aerosol retrieval for all panels because it is not derived from MODIS.

Whilst a comparison of the different approaches to dust detection will help to understand how MODIS can best be used to identify the presence or absence of dust, the main aim of this paper is to evaluate the use of MODIS for identifying dust sources. This means that the way in which the upwind edge of the dust plume is depicted is of most interest. In these comparison figures, the areas highlighting the active sources for each event (denoted by coloured squares; e.g. figure 5), which were used to compare the outcome of the techniques, were determined by an informed approach. Since the BTD principle is a component of all of the evaluated algorithms (panels c-e), sources were determined from a combination of the scene BTD plus other readily available information, including the use of wind direction data to ascertain the upwind side of plumes. From the companion dust source
inventory work of Bullard et al. (2008), certain sources in the LEB region could be identified as relatively recurrent points of emission; locations which had a record of acting as source areas for several different dust events across the three year study period. The case study events here were therefore chosen to ensure that several of the plume origins used for comparison were from ‘proven’ dust sources. Consequently, it is worth noting that the persistence of certain key sources allowed their identification with a further confidence when flagged as active in each BTD scene. One such example is the point-source located at the south east margin of Lake Callabonna, South Australia (centred on 140°15’0E, 30°S.). This well-studied source location was seen to be active in both the third and fourth case studies, where BTD data are able accurately to indicate it as an area where a plume has originated (Figures 7 and 8). The relative performance of the dust enhancement models was evaluated on this basis. It is worth noting that extensive background knowledge of dust events in the LEB was used to help verify source locations determined from BTD data and this may not be possible in areas where comparable auxiliary data are unavailable. Nevertheless, in this study, we were able to carefully and thoroughly assess model performance for detection of dust emanating from known source locations.

As noted earlier, there are some additional caveats to the use of remote sensing to determine (or infer) dust source locations; these include the relative timings of the satellite overpass and the onset of dust emissions (which might affect not only the location of the plume head, but also the density of the dust), and the fact that only the upwind dust source can be located with any additional contributing sources lying under the dust plume possibly going undetected. To test the likely impact of some of these issues, HYSPLIT (http://www.arl.noaa.gov/ready) was used to calculate possible trajectories and plume concentrations for each event. These data are not presented here but confirm that the dust emitted during the events was close to, or at, source at the time of data capture and rarely reached an altitude of more than 500 m. This suggests that overpass timings for remote sensing data capture were likely to have provided data suitable for source identification.
3. Results

3.1 Event 1: 7th October 2005

For this event, raised dust is quite difficult to observe in the visible scene (Figure 5a) and sources are not at all apparent regardless of the level of contrast enhancement applied. However, BTD analysis (Figure 5b) reveals the presence of dust plumes, which appear as dark streaks at the centre of the image. Much of the cloud in the lower left of Figure 5b also appears dark, indicating some overlap in the thermal signature of cloud and dust in this scene. This is further highlighted in Figure 5c where a dust/non-dust threshold of 0 (dust < 0, cloud > 0) has been applied. Here, not only are parts of the dust plumes categorized as dust, but so too are some of the patches of cloud. Although this simple threshold effectively separates cirrus cloud (white/light grey in Figure 5b) from the dust plumes, the thicker areas of cloud (which have similar BTD values to the dust) are mis-identified. Adjustment of the dust/non-dust threshold for this scene highlighted that there was no single BTD value that could differentiate these two components. In terms of identification of the sources of dust in this scene, there was an observed offset between the upwind dust heads shown in Figure 5c and the dust sources identified using Figure 5b. This is most likely to be because the dust at source is still close to the ground surface and therefore has a less pronounced thermal contrast with the ground surface than airborne dust further downwind of the source which will have risen to a higher atmospheric level. Given the dust/non-dust threshold applied to the scene is 0, this also indicates that some of the dust from this location can have a BTD value of >0 (Figure 5b).

![Figure 5]

Application of the Roskovensky and Liou (2005) algorithm (Figure 5d) effectively removes the cirrus cloud from the scene, and only a small area of the remaining cloud is included when a dust threshold (using event-specific ‘a’)
and ‘b’ coefficients; Table 4) is applied. The upwind ends of the main dust
plumes map on to the same source locations inferred from BTD (Figure 5b),
with the exception of the most northwestern plume which is not detected.
Miller’s (2003) algorithm clearly differentiates the cloud from the main dust
plume (Figure 5e) which is picked out in red. With the exception of the
northernmost plume, source detection is comparable to those in Figure 5b.
However, there are some parts of this scene where dust is likely to have been
mis-identified. These areas (marked ‘FS’ in Figure 5e) are patches on the
ground surface where fires have changed the ground surface reflectance
characteristics significantly (Jacobberger-Jellison, 1994), and suggest that the
Miller algorithm is sensitive to ground reflectance variability. Given that the fire
scar is clearly discernible in the visible image (Figure 5a), but not when the
other techniques (which rely more heavily on BTD to detect dust) are used,
this implies that the component within the Miller (2003) algorithm that uses
VNIR wavelengths is slightly over weighted in this application. Figure 5f
shows the extent to which the three MODIS L1B algorithms agree and
highlight co-incident pixels containing dust. Whilst all three pick out the main
central plumes of dust, there are considerable differences elsewhere in the
scene. In particular, the Miller (2003) algorithm suggests a much more
extensive plume of dust than the other two approaches, especially in the
northeast. All of the techniques misidentify some of the most dense cloud as
dust, with Ackerman (1997) and Miller (2003) performing particularly poorly.

The MODIS Deep Blue AOT image (Figure 5g) shows that the cloud
mask applied in this instance is effective in separating the clear or dusty sky
from the clouds, and some dust is detected (AOT values close to 0). The dust
plumes in the far right of the scene that are highlighted in previous panels
(Figure 5 b, c, and d) are clearly defined, but the main central plume is less
obvious. In comparison, despite the relatively coarse resolution, the OMI AI
image (Figure 5h) depicts clearly the central and far right plumes, and with
similar AI values. This suggests that the inability of Deep Blue to detect both
of these plumes is not likely to be due to vastly different aerosol densities in
each plume. Instead, one possibility (explored further in section 4.6) is that
MODIS Deep Blue data are actually more sensitive to variations in dust colour
(mineralogy) than OMI. There is also an area in Figure 5g that is excluded by the Deep Blue cloud mask (marked ‘FP’ in Figure 5g) where cloud is not apparent. On the ground, this is a floodplain and the high surface reflectance characteristics also cause confusion for the cloud mask when applied to event 3 (Figure 7g). Overall, the location of the dust plumes outlined by OMI AI data correspond very closely to the position of the plumes in the MODIS 1B algorithm outputs. In addition, the gridded AI data capture the extent and variation of aerosol density apparent in the other panels. Although plume identification is acceptable, the coarse resolution of the aggregated OMI AI data (0.25°x0.25°) mean these data are less able to define the dust source location or the nature of the surface sedimentary environments with the same precision as can been achieved using the combination of outputs from the MODIS L1B algorithms (e.g., Bullard et al., 2008; Lee et al., 2008).

3.2 Event 2: 24th September 2006

In the second event, the downwind (northerly) limit of the advancing dust is very distinct in the MODIS VIS (Figure 6a). The upwind edges of the plume are, however, not distinct and are in places difficult to differentiate from the underlying bright desert surface. For Event 1, the dust/non-dust thresholds or coefficients chosen are the same as those recommended in the published techniques. However, if these values are applied to Event 2 some problems become evident. Figure 6 shows the visible MODIS (panel a) and BTD values (panel b) for Event 2 and the results of applying the Event 1 thresholds (Table 4; Figure 6 panels c-f). There is no possibility of identifying dust sources using the Roskovensky and Liou (2005) or Miller (2003) algorithms with these thresholds. The Roskovensky and Liou (2005) output suggests that the dust plume fills most of the panel, whilst no dust is highlighted using the Miller (2003) approach. This is emphasized in Figure 6f which shows there are no areas of the image that all the different approaches identify as containing dust.

<Figure 6>
For this reason, thresholds and parameters appropriate to this event were determined using the histogram approach (see Table 4 for values). The results of applying these event-specific thresholds are shown in Figure 7. The dust source areas for this event and extent of the plume are reasonably well discerned using BTD (Figure 7b), which also reveals several other minor plumes that are not evident in the VIS. When an event-specific threshold is applied to the BTD (Figure 7c) most of the plume is highlighted but some of the thin, discrete plumes are not identified as dust or are foreshortened. This again suggests the use of the dust/non-dust threshold can affect the accurate identification of dust sources. The main plume is successfully identified using Roskovensky and Liou’s (2005) algorithm (Figure 7d) but the source areas are poorly represented. Despite extensive experimentation with the ‘a’ and ‘b’ coefficients to improve dust detection, it was not possible to pick out the westernmost dust plumes without introducing a significant component of the ground surface reflectivity to the dust determination.

Using Miller’s (2003) algorithm (Figure 7e), the maximum D value for this event falls below the published dust >1.3 threshold (see Table 3) necessitating an adjustment of this threshold such that dust >-0.55. Although this adjustment enhances the dust visualization significantly, it does not do so without introducing further artifacts. First, some areas of the plume evident in Figure 7b were not highlighted, for example the thin streaks to the left of the main plume that were also not identified using the default Ackerman (1997) or Roskovensky and Liou (2005) thresholds (Figures 7c and d). Second, whilst detection of airborne dust is improved, some areas of the ground surface are also mis-identified as dust. Some of these are the same areas (marked FS in Figure 7e) that caused difficulties in event 1 due to changes in surface reflectivity caused by fire scars, and can clearly be seen in other panels (Figure 7 a-d).

Figure 7f shows that the agreement between the MODIS L1B enhancement methods for dust (white) is restricted to the most dense part of
the plume. Ackerman (1997) performs best at highlighting the more subtle
(perhaps less dense) dust plumes in the west of the scene. The Roskovensky
and Liou (2005) threshold mis-identifies dust not only to the west but also to
the south and so under-represents its spatial extent; with obvious implications
for upwind source detection.

The central dust plume is shown clearly in the MODIS Deep Blue
aerosol product for this event (Figure 7g). Not only is the sharp advancing
dust front apparent, but these data also indicate higher dust concentrations in
the southerly, upwind source area of the plume. At best, however, these data
are only able to provide a broad regional indication of the plume’s source
because the thin, discrete plumes in the west are not detected. The MODIS
Deep Blue product also suggests that the highest AOT values in the scene
are associated with the plume at the extreme eastern edge of the scene
(marked ‘X’). This contrasts with the MODIS BTD analyses where dust
appears to have a higher concentration at the furthest downwind edge of the
plume (marked ‘Y’). Similarly, OMI AI data clearly outline the main plume, and
also suggest higher aerosol density in the east (Figure 7h). Comparison of the
Deep Blue and OMI data highlight the performance of the cloud masks used
in these products. The Deep Blue cloud mask only removes the cloud in the
top left of the scene, which is clearly present in MODIS VIS, whilst the OMI
cloud-mask obscures as much as 15% of the scene.

3.3 Event 3: 2nd February 2005

Event 3 was an extensive dust event during which 35 meteorological
stations recorded a reduction in visibility to ≤1 km. Here we concentrate on
two areas in the central LEB: the first is where two parallel dust plumes can
clearly be seen blowing northwards out of Lake Eyre North and the second is
in the lower right corner of the image and is difficult to see in the VIS (Figure
8a), but is clearly shown on the BTD (Figure 8b). The previous two events
used a BTD threshold of <0 with varying success. For this event (Figure 8c), a
significantly lower threshold for dust detection (dust <-1.2) was required to
identify dust plumes (Table 4). Using this value, most of the pixels highlighted
contain dust although the area marked ‘GS’ on Figure 8c is not dust, but is the ground surface. Crucially, the <-1.2 threshold in this instance does allow the observed dust plumes to be traced all the way back to the source areas.

<Figure 8>

The Roskovensky and Liou (2005) algorithm (Figure 8d) is effective at picking out the two main dust areas, but the origin of the twin plumes is situated north of the known dust source (the bed of Lake Eyre North). The parallel plumes are also visible when Miller’s (2003) algorithm (Figure 8e) is applied with an adjusted dust threshold, but this dense dust is only enhanced by the model (coloured red) at the downwind end of the plume, and not in the source locations. Furthermore, in comparison to BTD (Figure 8b), the origin of the dust in Figure 8e (in white) would be placed approximately 70km north of the actual lake bed source. The ‘best’ dust threshold that could be determined for the Miller (2003) algorithm in this instance also seems to divorce the apparent upwind boundary of the plume from the source area marked ‘X’ in the right of the image (Figure 8e). However, the entire aerosol outbreak to the lower right is highlighted in red using this approach, and the plume clearly extends back to the assumed source. This demonstrates that there can be significant variability in the performance of this enhancement approach within a single scene. The composite image (Figure 8f) for the MODIS L1B enhancement techniques highlights the problem of surface reflectance evident when applying the Miller algorithm, as to the northwest of the twin plumes the fire scar (FS) is clearly shown in blue.

In Figure 8g, the distinctive parallel dust plumes are mainly excluded by the Deep Blue cloud mask, and only the downwind end of the plume is detected. Other parts of the image, where no cloud is present (cf. Figure 8a), are also excluded by the cloud mask. For example, the dry bed of Lake Eyre and small patches across the whole area of the scene are flagged as no-data. In Figure 8h, the OMI AI data show dust over the majority of the scene, suggesting a widespread dust haze. The AI maximum of 5.4 is very high for Australia but, whilst all the main areas of dust are identified in a manner
broadly consistent with the MODIS L1B algorithms, the spatial resolution is insufficient to illustrate the detail of the parallel plumes or the specific source locations. Indeed, although these data clearly have limited utility for determining the specific point-sources in this scene, the OMI data do suggest the presence of diffuse raised dust across the scene which would be expected given the number of meteorological stations recording the event.

3.4 Event 4: 30th August 2005

The weather systems that promote LEB dust storms (thunderstorms, pre- and post-frontal winds; Sprigg, 1982) mean that dust events are often associated with cloud cover. Event 4 was selected to explore further the extent to which dust and cloud can be distinguished. Dust is visible in the centre of the VIS scene between the bands of cloud (Figure 9a) and can also be identified using BTD (Figure 9b); although large areas of thicker cloud can be seen which exhibit a similar BTD as the dust, making initial interpretation of this scene using BTD alone problematic. A dust threshold of < -0.35 was applied to BTD in this instance (Figure 9c), and was effective in isolating the major dust plumes that exist between the clouds, but at the expense of the thinner plumes which are removed when this particular threshold is applied.

Most of the cloud is removed from the image by application of Roskovensky and Liou’s (2005) algorithm (Figure 9d) and the inferred sources for the two main dust plumes clearly map on to those identified from BTD (Figure 9b). For this event, the Miller (2003) algorithm (Figure 9e) was scaled using dust>0.45, and can be seen to be very effective as the principal dust plumes are easily discernable and source determination is possible. The enhancements discussed here are unable to ameliorate the blanketing effect of thick cloud when it obscures active dust sources or plumes; they can only enhance the dust that can be ‘seen’ between the cloud banks. Interestingly, none of the BTD-based methods pick up the thin plume which is best seen on the MODIS visible panel (marked ‘X’, Figure 9a). The most notable feature,
other than the general agreement of spatial extent and source location for the
two plumes shown in Figure 9h, is the appearance of the blue areas where
the Miller (2003) output has confused the thickest cloud with dust.

The cloud masking of the Deep Blue (Figure 9g) scene seems to work
well for much of the cloud coverage in the image, but does also remove much
of the northernmost dust evident in the other panels (Figures 9b-f). The shape
and downwind extent of the central plume is indicated by raised AOT values,
but the MODIS Deep Blue data also suggest an origin for the dust that is
some distance removed downwind from the source when compared with the
BTD-based approaches. The OMI data (Figure 9h) again show more of the
dust plume than the MODIS Deep Blue product, and less of the image is
affected by cloud masking.

4. Discussion

The main aim of this paper is to evaluate the use of MODIS for
detecting dust sources. In some instances dust plumes may be discernible on
the MODIS VIS (e.g. Figure 6); but this certainly is not always the case (e.g.
Figure 4). From the results presented above, we can confirm that all the dust
enhancement techniques used here make it easier to detect dust. However,
with respect to source determination, the results suggest that, of the MODIS
L1 processing techniques, the ‘best’ approach varies from event to event. For
events 1 and 2 arguably the best source detection came from the simple
brightness temperature difference calculation (BTD), often with no dust
threshold applied. Of the more complex processing techniques, that of
Roskovensky & Liou (2005) works well for event 1 (Figure 5) as it is very
effective at removing cloud cover, whereas the Ackerman (1997) is better for
events 2 and 3, where cloud cover is less of an issue. The cloud cover in
event 4 (Figure 8) makes it much harder to determine sources from the BTD
alone, but both Roskovensky & Liou (2005) and Ackerman (1997) work well.
For these four events, the Miller (2003) algorithm is extremely useful for
visualizing dust, but there are significant problems with precise source
identification and determination of dust plume extent in all cases except event
4. For the majority of events and algorithms the published, or indicative, thresholds under-perform and the values vary from event to event. This makes it difficult to suggest appropriate regional scale thresholds. Whilst some of this variation is due to factors specific to the algorithms or individual events, other causes such as diurnal and seasonal variations in surface temperature/dust contrast (which affect BTD) will affect all the methods. The advantages and disadvantages of each of the approaches from an operational perspective are discussed in detail below.

4.1 Brightness Temperature Difference (bispectral split window)

Calculating BTD is straightforward, and keeping the full range of values (rather than applying a dust threshold) is often preferable for both dust plume and source detection. The procedure does not appear to be very sensitive to observed mineralogical variability either within or between plumes, and so all dust, regardless of source, is enhanced provided it can be differentiated thermally from the ground surface. The main disadvantages are that because no definition of dust/non-dust is applied the interpretation of the BTD data becomes subjective and data retrieval can suffer through lack of cloud cover elimination. With the exception of event 4, this is not a major problem in the case studies presented here, but it is likely to be important for anyone interpreting the data, to have a good understanding of how and why ground surface characteristics may vary.

4.2 Ackerman (1997)

Although Ackerman (1997) did not explicitly present a dust/non-dust threshold of zero, he observed negative differences in BT$_{11}$-BT$_{12}$ for dust storms and a universal threshold of dust<0 could be implied. Darmenov & Sokolik (2005) demonstrated that this dust threshold was in fact variable when applied to dust over oceans and suggested that dust sourced from the Lake Eyre Basin and travelling southeast over the Tasman Sea had a value of <-0.4 K. All the dust plumes examined here are over land, and whilst the threshold of zero worked effectively for events 1 and 2, adjustments had to be
made for events 3 and 4. For event 3, in order to eliminate interference from
the ground surface, the threshold had to be lowered to <-1.2 K; for event 4 the
threshold was <-0.35 to eliminate cloud. Interestingly, whilst it was possible to
find a dust/non-dust threshold for event 4 where all cloud could be removed
this was not possible for event 1. Here (Figure 5c), a dust threshold lower
than zero removed more dust in the scene so the threshold was left at 0.
Another factor affecting the BTD threshold is likely to be the thickness of the
dust plume. Where there is low AOT (as confirmed by comparison with
MODIS Deep Blue) we have determined BTD differences of >0 for pixels
populated by dust. One possibility is that where the dust plume
thickness/density is low, BTD becomes increasing affected by the ground
surface temperature signal. Using this approach, therefore, may involve a
compromise between dust detection and the elimination of cloud. If both dust
and cloud are dense/opaque then it is straightforward to identify and
implement a dust threshold. If the dust is thin and cloud cover is dense (as in
event 1), then it can be hard to identify an appropriate dust/non-dust
threshold. Where the cloud cover is sparse and the dust plume is
dense/opaque (as in event 4) then their differentiation through threshold
adjustment is straightforward. From this study it is also apparent that the
thickness/density of the dust plume also affects the degree to which the dust-
head can be pinpointed, and an inappropriate threshold value may
foreshorten plumes.

4.3 Roskovensky & Liou (2005)

This approach was designed explicitly as a simple method for the
differentiation of dust from cirrus cloud, and is very effective at doing so in
both of the cloudy scenes examined here (events 1 and 4). For all events it
was necessary to adjust the scaling factor ‘a’ and BTD offset value ‘b’.
Roskovensky and Liou (2005) calculated these to be 1.1 and 0 over ocean
(around the Korean peninsula) and 3 and 0 over land (the Gobi desert). All of
the events examined here occurred over land and the coefficients determined
were variable (values of ‘a’ ranged from 0.25 to 1.2 and values of ‘b’ ranged
from –0.5 to +1) and made a significant difference to both the number of
pixels classified as dust and the inferred location of the dust sources. Although the algorithm is slightly more computationally complex to calculate than simple BTD, it is easy to tune it for specific events, and certainly worth the extra effort. Overall this model worked best on dense dust; there was little observed confusion with ground surface reflectance, and the inferred upwind plume source locations compared well with those suggested by BTD alone.

4.4 Miller (2003)

The Miller (2003) algorithm is designed to provide improved differentiation of dust from water/ice clouds over bright desert surfaces and was found to be visually very effective for all events observed in this study. In particular, there is generally a clear distinction between dust and cloud. In a similar manner to optimizing the Ackerman (1997) data, the Miller (2003) dust threshold also had to be tuned for each event to be effective (Table 4) and in most cases it was necessary to decrease the lower threshold value to well below Miller’s suggested +1.3 (as low as -0.55 in one instance). Unfortunately tuning this algorithm was not straightforward, and although it worked very well for some events, this was not always the case.

4.5 The Aerosol Products (Deep Blue and OMI)

The main utility of the aerosol products is in the detection of dust because for dust source identification the coarse spatial resolution of the products is a limitation. Average AI values over a long time series have been used to detect persistent, regional scale dust sources (e.g. Washington et al., 2003) but accurate and event-specific source identification requires the clear delineation of the upwind margin of the plumes. We have presented values of Deep Blue AOT and OMI AI which, whilst not comparable in terms of absolute values can be compared relatively. There are occasions where OMI agrees well with the much higher resolution MODIS, but often the upwind margin of a dust plume is difficult to detect using AI.
In some cases, it was difficult to determine not only the upwind dust sources but also the extent of the dust plumes because the in-built cloud masks of the products eliminated them. For example, in Figure 8g, the origin of the parallel twin plumes was not discernible because MYD04 returned no data from the bright dry lake surface, which was classified as cloud. This is a recognised limitation and is probably due to the colour and density of the dust or reflectance of the surface. The twin parallel plumes comprise white coloured dense dust sourced from Lake Eyre and are sufficiently bright to saturate the pixels causing misidentification as cloud (a known problem with Level 2 aerosol product http://modis-atmos.gsfc.nasa.gov/MOD04_L2/ga.htm). There are obvious implications for dust source identification – ephemeral lakes are often very bright surfaces and have been seen in this study to be routinely masked out as cloud even in cloud-free and dust-free scenes, yet are common dust sources not only in the LEB (Bullard et al. 2008), but also in the USA (e.g. Reynolds et al. 2007), southern Africa (Mahowald et al. 2003) and other dryland regions. The authors are also evaluating the Aura-OMI Aerosol Data Product; OMAERUV (V003) which provides aerosol extinction and optical depth via swath data (rather than global gridded) at the native 13 x 24 km pixel (see: http://daac.gsfc.nasa.gov/Aura/OMI/omaeruv.shtml). This may offer further potential for dust source identification, but requires further validation.

4.6 Impacts of Dust Mineralogy and Surface Reflectance on Data Retrieval

One issue that can be explored briefly here, and will be developed further as a future project, is the impact of dust mineralogy on both the Miller (2003) dust thresholds and MODIS Deep Blue retrievals. A range of different sedimentary environments emit dust within the Lake Eyre Basin (Bullard et al., 2008) and these have different mineralogical compositions which in turn control the infrared radiative properties of the dust (Claquin et al., 1999; Sokolik, 2002). An example can be seen in event 3 (Figure 8) where a dust plume was observed originating from the bed of Lake Eyre (which is illite-rich), in the west of the scene and a dust plume originating from dune sands with iron-rich coatings in the southeast of the scene. Using threshold values
where $0.6 < \text{dust} < 1.98$ it was possible to pick out the dust sourced from the
dunes in red, but not the near-source dust from the lake (although part of the
distal end of the plume is highlighted). When a threshold was selected to
highlight both plumes large areas of ground surface were also included.

The effect of dust mineralogy on detection is also apparent in the
MYD04 Deep Blue data presented here. In the Deep Blue algorithm, surface
reflectivity over desert regions is assumed to be low in blue wavelengths and,
dust aerosol slightly more reflective (Hsu et al., 2004). However, dust
reflectivity is variable depending on its chemistry and can decrease
significantly with increased iron concentration (Dubovic et al., 2002; Arimoto
et al., 2002). Consequently red, or iron-rich, dust will have relatively low blue
reflectivity and therefore potentially lower contrast relative to background
reflectance, whereas white dust (composed of carbonates, bleached quartz or
evaporite minerals) will have higher reflectance in the VIS and significantly
higher contrast with respect to underlying soils and vegetation. This is
illustrated in Figures 7g and 8g where the white, dry lake bed sources of dust
are more distinct than the iron-oxide rich dust from dune sources (noting
however that the OMI product suggests these red plumes are also less dense
than the white dust plumes which will affect AOT).

For the Lake Eyre Basin events there are some problems with the
differentiation of dust from the underlying desert surface. Specifically, large
areas of the LEB dunefields comprise red brown to orange, highly reflective
sands with a partial vegetation cover; although sand color is variable
especially between the redder Simpson dunefield and the less red Strzelecki
dunefield (Pell et al., 2000; Bullard and White, 2002). Where the vegetation
cover has not been disturbed by fire, the Miller (2003) algorithm works well.
However, where vegetation has been removed by fire and large areas of
bright sand are exposed the algorithm cannot distinguish dust from the
surface. Fire changes both surface reflectance characteristics, through the
removal of absorbing vegetation, and sediment reflectance characteristics,
through fire-induced reddening or changes in mineralogy (Jacobberger-
algorithm is affected by these changes has important implications for dust provenance because areas which have been de-vegetated through fire may be mis-identified as a dust source. A complicating issue is that fire scars can act as sources of dust under certain conditions (Bullard et al. 2008; McGowan & Clark, 2008). This highlights two issues. First, it is important to deploy event-specific dust/non-dust thresholds to limit the influence of surface reflectance as far as possible. Second, familiarity with the underlying surface characteristics is essential, and examination of more than one scene, including known dust-free images is desirable. The latter means that permanent or semi-permanent ground characteristics can be discerned. For example, the firescars in the LEB have distinctive shapes that can be identified on most of the VIS images taken post-2001 when widespread burning occurred. Cross-referencing with other MODIS-derived output can also be useful. For example, in event 1 the fire scar is only clearly discernible on the VIS and Miller (2003) panels which suggests that it was not a dust source. In contrast, in event 2 the fire scars are discernable in all panels (Figure 7 a-e) which may suggest that they acted as dust sources at this time.

4.7. Tools for dust source detection

Figure 10 presents a summary of the how the different data sources and analyses used here complement one another. At the global-regional scale, dust events can be detected using visibility criteria, as in this paper, or OMI AI. An alternative that we have not discussed in detail here is the MOD/MYD08 Global Gridded Atmospheric Product (1 x 1 deg) which could also be used (Bryant et al., 2007), although it has some limitations over bright desert surfaces (Chu et al., 2002). If none of these three indicates dust it does not necessarily mean that dust is not present as the relative timing of satellite overpass or visibility observation may result in no dust being recorded (see quality control indicators), but the events missed are likely to be minor. If any one of these indicates the presence of dust then there is the potential for determining dust sources at higher resolution. The choice of higher resolution technique depends on the precise research question to be answered. The MOD/MYD04 aerosol products give data processed to a common standard
that enables comparison from one region to another. The versatile MOD/MYD02 data can be processed simply using brightness temperature difference to enhance the dust signal. Where cloud is present, or if it is necessary to highlight the dust plume then one of the methods for employing a dust/non-dust threshold can be used, but it is recommended that event-specific thresholds are calculated where possible (as opposed to using global or regional thresholds).

5. Conclusions

This paper set out to evaluate the use of MODIS data for identifying dust sources. Through objective and subjective comparison of several different approaches to dust detection several conclusions can be drawn. Whilst these conclusions have implications for regional and global scale studies of dust, the clear outcomes are that even within a single drainage basin, dust events should be examined on an event by event basis, and that the ‘best’ algorithm for identifying dust sources varies considerably.

(1) For the region examined here (the Lake Eyre Basin), no single MODIS technique was found to be ideal for source determination.

(2) MODIS VIS full colour composite data can be useful but are often insufficient to discern dust plumes over reflective desert surfaces. In particular, for event detection MODIS VIS (particularly via quicklooks) should be used with caution, or in combination with other data sources, because not all dust plumes will be visible over the bright surfaces. An ideal combination for rapid detection of dust activity is OMI AI (or equivalents) and MODIS but the former can not be used to derive source location due to its low spatial resolution.

(3) BTD data are simple to calculate and very effective at highlighting dust that is not seen in VIS data. If the user is familiar with the underlying ground surface characteristics and has the VIS image available to discern cloud cover, this can be the most simple and possibly most accurate method of source determination. From the four example events presented here, BTD is the approach that is least sensitive to dust mineralogy.
(4) For scenes where cloud is present, there is potential for confusion between cloud and dust using just BTD and the algorithms designed to differentiate these have been demonstrated as effective for screening out cirrus cloud (for which they were designed) although thick cloud can remain.

(5) Whilst there are published dust/non-dust thresholds for each method compared here, for each method thresholds may need adjusting on a regional and/or event scale. Although there are clear guidelines for positioning thresholds, a considerably amount of informed (but subjective) judgement can be required. It is important therefore for users of these techniques to consider the effects of not adjusting thresholds for each event. This paper suggests that for an event-based study it is essential to derive event-specific thresholds. However it seems likely that for global or longer-term studies of this nature to be effective, it may be pragmatic to use regional thresholds.

(6) In the Lake Eyre Basin, it is not possible to use a single dust/non-dust threshold for all events for any of the algorithms tested here. One possible reason is that it is a basin with multiple potential dust sources with different mineralogies; where different sources (e.g. iron-rich dunes, illite-rich lake beds) can emit dust simultaneously it is necessary to use event-specific, or even plume-specific thresholds. In regions with a single definable source (for example a large playa such as the Magkadigadki, Botswana) it may be possible to discern a single dust/non-dust threshold, however any use of a regional (or global) threshold is likely to result in errors or inconsistencies.

(7) Some sedimentary environments are more intense dust sources than others. In particular, previous studies have noted ephemeral lakes and areas devegetated by fire as prominent dust sources. Significantly, these are two of the ground surface types that have proved most problematic for establishing dust/non-dust thresholds due to confusion caused by their bright surfaces which cause false positive dust signals or may be falsely identified as cloud.

(8) Whilst the findings of this research may be challenging in some respects, what is clear is that there is considerable potential for using
MODIS data to obtain information about dust mineralogy by interpreting the shifts to the thresholds (or coefficients) that need to be made. Possible mineral aerosol information that could be gained can not only assist in identifying the dust source, but also its radiative properties.

Using data from sensors such as MODIS will inevitably mean that some dust activity is missed due to the relative timing of overpass and dust emissions or cloud cover. Other sources of data such as MSG can minimize the problem of overpass timings but cloud cover is still a problem and coverage is not yet global. This paper has focused on daytime dust emissions only, but analysis of night time dust emissions and source identification are the subject of future research. It is worth noting that the products and data used here are subject to continual development and improvement and consequently some of the issues raised here may be resolved or change.
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**Figure Captions**

**Figure 1.** The Lake Eyre Basin of Australia, and the coverage extents of the four dust event case studies by event number.

**Figure 2.** Histograms illustrating (a) the basic method for determining dust/non-dust threshold; (b) dust/non-dust thresholds derived for BTD (Ackerman, 1997) (dust < threshold; and (c) dust/non-dust threshold derived for Miller (2003) algorithm outputs (dust > threshold) for events 1-4.

**Figure 3.** Histograms showing the derivation of scaling factor ‘a’, based on reflectance ratio, for events 1-4 (graphs a-d), and BTD offset ‘b’ for events 1-4 (graphs e-h) necessary for applying Roskovensky & Liou’s (2005) algorithm. For event 4 the scaling factor = 0.25 (not shown). Note: vertical axes vary.

**Figure 4.** Key to interpreting the colour composite combining dust/non-dust detection using Ackerman (1997), Miller (2003) and Roskovensky & Liou (2005) and shown in Figures 5-8, panel (f).

**Figure 5.** Event 1: 7th October 2005. Across all panels, squares highlight the prominent active dust source areas, as identified from the BTD split window product. The dust/non-dust thresholds used are detailed in Table 4. Panels (a) to (e) are derived from MODIS (Aqua) data. (a) MODIS visible (wind direction marked with arrow); (b) BTD (BT_{11}-BT_{12}); (c) threshold applied to BTD (after Ackerman, 1997); (d) Roskovensky & Liou (2005) dust enhancement algorithm; (e) Miller’s (2003) dust enhancement algorithm (dust is red); (f) Composite image of Ackerman, Miller and Roskovensky and Liou output for key see Fig.4. Panels (g) and (h) are aerosol products for the same event where (g) MYD04 Deep Blue AOT 0.1°x 0.1° spatial resolution; (h) OMI AAI 0.25°x 0.25° spatial resolution.
Figure 6. Event 2: 24th September 2006. Dust/non-dust thresholds used are those recommended in the published literature and the same as those used for Event 1.

Figure 7. Event 2: 24th September 2006. Dust/non-dust thresholds used are specific to Event 2 and adjusted using the histogram approach (Figure 3).

Figure 8. Event 3: 2nd February 2005. For panel explanations see Figure 5. Dust/non-dust thresholds used are specific to Event 3 and adjusted using the histogram approach (Figure 3)

Figure 9. Event 4: 30th August 2005. For panel explanations see Figure 5. Dust/non-dust thresholds used are specific to Event 4 and adjusted using the histogram approach (Figure 3)

Figure 10. Summary of the relationships between the methods and approaches discussed in this paper.
Table 1. Spatial and temporal characteristics of remote sensing data used in this study.

<table>
<thead>
<tr>
<th>Data Type</th>
<th>Spatial Resolution at Nadir (km)</th>
<th>Scenes used per day</th>
<th>Archive Length</th>
<th>Typical Overpass: (am/pm)</th>
</tr>
</thead>
<tbody>
<tr>
<td>MOD02 [Terra]</td>
<td>0.25 x 0.25 (VIS&lt;sup&gt;1&lt;/sup&gt;)</td>
<td>1</td>
<td>1999 - date</td>
<td>am (10:30) ect†</td>
</tr>
<tr>
<td></td>
<td>0.5 x 0.5 (VIS + NIR&lt;sup&gt;2&lt;/sup&gt;)</td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td>1 x 1 (TIR&lt;sup&gt;3&lt;/sup&gt; + all bands)</td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>MYD02 [Aqua]</td>
<td>0.25 x 0.25 (VIS)</td>
<td>1</td>
<td>2002 - date</td>
<td>pm (13:30) ect</td>
</tr>
<tr>
<td></td>
<td>0.5 x 0.5 (VIS + NIR)</td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td></td>
<td>1 x 1 (TIR + all bands)</td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>MYD04 [Deep Blue]</td>
<td>10 x 10</td>
<td>1</td>
<td>2002 - date</td>
<td>pm (13:30) ect</td>
</tr>
<tr>
<td>OMI</td>
<td>13 x 24</td>
<td>1</td>
<td>2004 - date</td>
<td>pm (13:38) ect</td>
</tr>
</tbody>
</table>

<sup>1</sup>Visible; <sup>2</sup>Near-infrared; <sup>3</sup>Thermal-infrared; †Equatorial crossing time
Table 2. Summary of dust event case studies used in this study to evaluate MODIS dust detection algorithms

<table>
<thead>
<tr>
<th>Event</th>
<th>Date</th>
<th>Satellite overpass time (UTC)</th>
<th>BoM(^\dagger) stations recording visibility ≤1km</th>
<th>Selection Criteria</th>
<th>Synoptic Conditions</th>
</tr>
</thead>
<tbody>
<tr>
<td>1</td>
<td>07/10/05 (2005:280)</td>
<td>AQUA (01:35) AURA (01:46)</td>
<td>1</td>
<td>Dust and 30% cloud cover</td>
<td>Pre-frontal northerly</td>
</tr>
<tr>
<td>2</td>
<td>24/09/06 (2006:267)</td>
<td>AQUA (04:20) AURA (04:31)</td>
<td>2</td>
<td>Dust - Single frontal plume</td>
<td>Post-frontal southerly</td>
</tr>
<tr>
<td>3</td>
<td>02/02/05 (2005:033)</td>
<td>AQUA (04:10) AURA (04:21)</td>
<td>35</td>
<td>Dust - Multiple plumes and source types</td>
<td>Post-frontal southerly</td>
</tr>
<tr>
<td>4</td>
<td>30/08/05 (2005:242)</td>
<td>AQUA (04:50) AURA (05:01)</td>
<td>1</td>
<td>Dust and 50% cloud cover</td>
<td>Pre-frontal northerly</td>
</tr>
</tbody>
</table>

\(^\dagger\)Bureau of Meteorology
Table 3. Summary of dust detection algorithms applied to MODIS L1B data in this study

<table>
<thead>
<tr>
<th>Algorithm</th>
<th>Display parameters</th>
<th>Source</th>
</tr>
</thead>
<tbody>
<tr>
<td>$D = (BT_{31} - BT_{32})$</td>
<td>Dust &lt; 0.0</td>
<td>Ackerman (1997) Sokolik (2002)</td>
</tr>
</tbody>
</table>

where:
- $BT_{31} = BT_{10.780-11.280 \mu m}$
- $BT_{32} = BT_{11.770-12.27 \mu m}$

$D = \exp \left(-\frac{R_{4}}{R_{16}} \times a + ((BT_{31} - BT_{32}) - b)\right)$


where:
- $a =$ Scaling Factor (0.8) $R_{16} =$ R0.862–0.877 µm
- $b =$ bt0 offset (2.0) $BT_{31} =$ BT10.780-11.280 µm
- $R_{4} =$ R0.545–0.565 µm $BT_{32} =$ BT11.770-12.27 µm

$D = \left[ \left( BT_{32} - BT_{31} \right)^{a} + \left( 2R_{1} - R_{3} - R_{4} - BT_{31} \right)^{b} - \left( R_{26} \right)^{c} + \left( 1 - BT_{31} \right)^{d} \right]^{1.3 < Dust < 2.7}$ Miller (2003)

where:
- $R_{1} =$ R0.620–0.670 µm $a = \{-2 \text{ to } +2\}$
- $R_{3} =$ R0.459–0.479 µm $b = \{-5 \text{ to } +0.25\}$
- $R_{4} =$ R0.545–0.565 µm $c =$ (if $M_{32} > 0.05$, c=0, otherwise c=1)
- $R_{26} =$ R1.360 - 1.390 µm $d =$ occurrences of $M_{31}$ are normalized in this equation as $(M_{31max} - 21K$ if $M_{31max} < 301K$) or $(M_{31max} - 273)/4 + 273$ otherwise.
- $BT_{31} =$ BT10.780-11.280 µm
- $BT_{32} =$ BT11.770-12.27 µm

Normalization:
- $a =$ Scaling Factor (0.8) $R_{16} =$ R0.862–0.877 µm
- $b =$ bt0 offset (2.0) $BT_{31} =$ BT10.780-11.280 µm
- $R_{4} =$ R0.545–0.565 µm $BT_{32} =$ BT11.770-12.27 µm
Table 4. Comparison of recommended dust/non dust thresholds and thresholds identified for events described in this study.

<table>
<thead>
<tr>
<th>Technique</th>
<th>Suggested threshold values</th>
<th>Dust/non-dust thresholds</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td></td>
<td>Event 1</td>
</tr>
<tr>
<td>Ackerman (1997)</td>
<td>Dust &lt; 0.0</td>
<td>Dust &lt; 0.0</td>
</tr>
<tr>
<td>Miller (2003)</td>
<td>1.3 &lt; Dust &lt; 2.7</td>
<td>1.3 &lt; Dust &lt; 2.7*</td>
</tr>
<tr>
<td>Roskovensky &amp; Liou (2005)</td>
<td>Dust &gt; 1.0, a = 0.8, b = 2</td>
<td>Dust &gt; 1.0, a = 0.8, b = 1</td>
</tr>
</tbody>
</table>

*upper value is scene maximum
Figure 1

[Map of Australia showing the location of Lake Eyre Basin]
Figure 2

A) 

B) 

C) 

Non dust (cloud & ground surface)

Increasing ground surface signal

Cloud peak

Threshold value

Dust peak

Number of pixels (x1000)

Data scaling value

Event 1
Event 2
Event 3
Event 4

Number of pixels (x1000)

Data scaling value (K)

Event 1
Event 2
Event 3
Event 4

Number of pixels (x1000)

Data scaling value

Event 1
Event 2
Event 3
Event 4
Figure 3

Reflectance ratio

![Graph A) Reflectance ratio for Event 1](image)

![Graph B) Reflectance ratio for Event 2](image)

![Graph C) Reflectance ratio for Event 3](image)

![Graph D) Reflectance ratio for Event 4](image)

BTD

![Graph E) BTD for Event 1](image)

![Graph F) BTD for Event 2](image)

![Graph G) BTD for Event 3](image)

![Graph H) BTD for Event 4](image)
### Figure 4

<table>
<thead>
<tr>
<th>Dust pixels</th>
<th>A</th>
<th>M</th>
<th>R &amp; L</th>
</tr>
</thead>
<tbody>
<tr>
<td>Dust pixels</td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>A</td>
<td><img src="#" alt="Red" /></td>
<td></td>
<td></td>
</tr>
<tr>
<td>M</td>
<td><img src="#" alt="Pink" /></td>
<td><img src="#" alt="Blue" /></td>
<td></td>
</tr>
<tr>
<td>R &amp; L</td>
<td><img src="#" alt="Yellow" /></td>
<td><img src="#" alt="Cyan" /></td>
<td><img src="#" alt="Green" /></td>
</tr>
</tbody>
</table>

<table>
<thead>
<tr>
<th>ALL indicate dust pixels</th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>ALL indicate non-dust pixels</td>
<td><img src="#" alt="Black" /></td>
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</table>
Figure 8