

UNIVERSITY OF BIRMINGHAM

University of Birmingham
Research at Birmingham

Scaling properties of the hot gas in early-type galaxies

Ponman, Trevor J.; O'sullivan, J.

DOI:

[10.1017/S1743921310009336](https://doi.org/10.1017/S1743921310009336)

License:

None: All rights reserved

Document Version

Publisher's PDF, also known as Version of record

Citation for published version (Harvard):

Ponman, TJ & O'sullivan, J 2009, 'Scaling properties of the hot gas in early-type galaxies', *Proceedings of the International Astronomical Union*, vol. 5, no. H15, pp. 291-292. <https://doi.org/10.1017/S1743921310009336>

[Link to publication on Research at Birmingham portal](#)

Publisher Rights Statement:

Copyright © International Astronomical Union 2010

General rights

Unless a licence is specified above, all rights (including copyright and moral rights) in this document are retained by the authors and/or the copyright holders. The express permission of the copyright holder must be obtained for any use of this material other than for purposes permitted by law.

- Users may freely distribute the URL that is used to identify this publication.
- Users may download and/or print one copy of the publication from the University of Birmingham research portal for the purpose of private study or non-commercial research.
- User may use extracts from the document in line with the concept of 'fair dealing' under the Copyright, Designs and Patents Act 1988 (?)
- Users may not further distribute the material nor use it for the purposes of commercial gain.

Where a licence is displayed above, please note the terms and conditions of the licence govern your use of this document.

When citing, please reference the published version.

Take down policy

While the University of Birmingham exercises care and attention in making items available there are rare occasions when an item has been uploaded in error or has been deemed to be commercially or otherwise sensitive.

If you believe that this is the case for this document, please contact UBIRA@lists.bham.ac.uk providing details and we will remove access to the work immediately and investigate.

Scaling properties of the hot gas in early-type galaxies

Trevor J. Ponman¹ and Ewan J. O’Sullivan^{1,2}

¹School of Physics & Astronomy, University of Birmingham, Birmingham B15 2TT, UK.

²Harvard-Smithsonian Center for Astrophysics, 60 Garden St, Cambridge, MA02138, USA
email: tjp@star.sr.bham.ac.uk

Abstract. We examine the processes responsible for the large scatter in the X-ray/optical luminosity scaling relation of early-type galaxies.

1. The L_X - L_{opt} relation in early-type galaxies

Optically brighter galaxies tend to have higher X-ray luminosity. The largest sample yet presented, by Ellis & O’Sullivan (2006), is shown in Fig. 1. It is apparent that: (a) there is a factor ~ 100 scatter in L_X around the mean relation; (b) the relation steepens at high L_B , where most galaxies lie well above the discrete source line – their X-ray emission dominated by a diffuse hot gas component; (c) many of these luminous galaxies are the bright central members of groups or clusters (BGGs or BCGs). Here we briefly examine a number of possible causes for the large scatter apparent in this relation.

2. Sources of scatter in L_X - L_{opt}

Star formation history – Recent star formation can considerably raise the blue luminosity of a galaxy, at a given stellar mass. However, Ellis & O’Sullivan (2006) showed that replacing L_B by K-band luminosity in Fig.1, results in a very similar scatter.

X-ray data quality – The X-ray data used by Ellis & O’Sullivan (2006) came from ROSAT. However, smaller samples of higher quality datasets from Chandra and XMM (e.g. Diehl & Statler (2007), Sun (2009)) show similar scatter in the relation.

Galaxy age – It has been recognised for some years (e.g. O’Sullivan *et al.* (2001)) that spectroscopically young ellipticals have low X-ray luminosity. The compilation of Sansom *et al.* (2006) shows a decline in L_X/L_B for 1-2 Gyr after a substantial burst of star formation (e.g. following a galaxy merger), presumably due to hot gas being dispersed by the energetic galaxy winds known to be associated with starbursts. However, galaxies with larger spectroscopic ages show no age trend, and a wide scatter in L_X/L_B .

Galaxy environment – Ellis & O’Sullivan (2006) found no clear environmental trend in L_X/L_B , apart from the high values apparent in BGGs and BCGs. However, recent studies with Chandra, which is better able to resolve galaxy emission, show that L_X/L_B is suppressed by a factor ~ 3 -4 in non-central galaxies within both clusters (Sun *et al.* (2007)) and groups (Jeltema *et al.* (2008)). This presumably results from stripping of hot gas within dense environments. The properties of *isolated* early-type galaxies provide a probe of whether environmental effects can account for the bulk of the scatter in L_X/L_B . Memola *et al.* (2009) assembled L_X - L_B data for 27 isolated ellipticals, and found a good deal of scatter. However, excluding probable ‘fossil groups’ (Ponman *et al.* (1994)) and upper limits, the remaining sample is small. Preliminary results from a Chandra study of isolated ellipticals currently underway indicate (Humphrey private communication) that significant L_X/L_B scatter is still present, but may be limited to a factor ~ 10 .

Halo mass – Mathews *et al.* (2006) suggested that the primary factor driving L_X is the mass of the dark halo surrounding a galaxy. For a sample of ellipticals (mostly BGGs) they found that L_X/L_K increased with X-ray derived virial mass. This is consistent with the earlier result, that

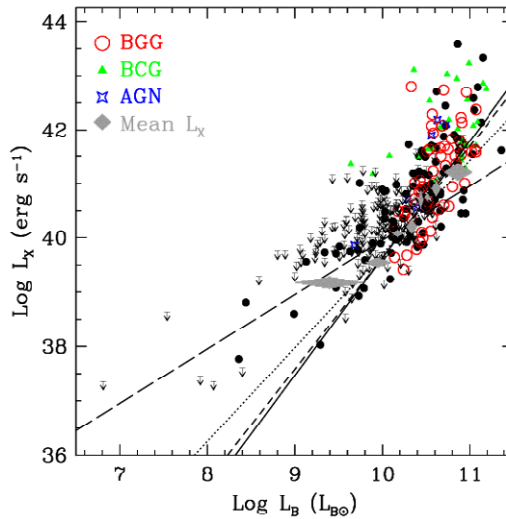


Figure 1. L_X - L_B plot for 401 early-type galaxies from Ellis & O'Sullivan (2006). The long dashed line is the discrete source contribution estimated by Kim & Fabbiano (2004), other lines show fits to various subsets of the sample, and the grey diamonds represent the mean L_X in a series of L_B bins.

BGGs have L_X/L_B ratios much higher than other elliptical galaxies. The study of NGC 4555 by O'Sullivan & Ponman (2004) shows that isolated ellipticals can also possess substantial dark matter halos.

Feedback – For a given dark halo mass, the gas content can be affected by feedback processes. Injection of energy raises the entropy, resulting in low density and X-ray luminosity. The low L_X/L_B ratio seen in ellipticals with recent star formation is probably an example of this. AGN feedback may have similar effects, if it can couple effectively to the surrounding gas. Sun (2009) finds that all radio-bright elliptical galaxies contain hot gas coronae – either compact or extensive. He postulates that AGN jets can couple to the extended hot halos found in groups or clusters, but not to compact gas cores.

References

- Diehl S. & Statler T. S. 2007, *ApJ* 668, 150
 Ellis S. J. & O'Sullivan E. J. 2006, *MNRAS* 367, 627
 Jeltama T. E. *et al.* 2008, *ApJ* 679, 1162
 Kim D.-W. & Fabbiano G. 2004, *ApJ* 611, 846
 Mathews W. G. *et al.* 2006, *ApJ* 652, L17
 Memola E. *et al.* 2009, *A&A* 497, 359
 O'Sullivan E. J., Forbes D. A., & Ponman T. J. 2001, *MNRAS* 324, 420
 O'Sullivan E. J. & Ponman T. J. 2004, *MNRAS* 354, 935
 Ponman T. J. *et al.* 1994, *Nature* 369, 462
 Sansom A. *et al.* 2006, *MNRAS* 370, 1541
 Sun M. *et al.* 2007, *ApJ* 657, 197
 Sun M. 2009, *arXiv:0904.2006*