

Modern Hardness Assurance: A Brand New Game Except When it Isn't

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Acronyms



COTS	Commercial Off The Shelf
DD	Displacement Damage
GEO	Geostationary Earth Orbit
GSFC	Goddard Space Flight Center
LEO	Low Earth Orbit
LET	Linear Energy Transfer
MBU	Multi-Bit Upset
MCU	Multi-Cell Upset
NEPP	NASA Electronic Parts and Packaging

RDM	Radiation Design Margin
RHA	Radiation Hardness Assurance
SEB	Single Event Burnout
SEDR	Single Event Dielectric Rupture
SEE	Single Event Effects
SEFI	Single Event Functional Interrupt
SEGR	Single Event Gate Rupture
SEL	Single Event Latchup
SOA	Safe Operating Area
TID	Total Ionizing Dose

NEPP - Small Mission Efforts





Outline



- New Space and SmallSat Considerations
- The Natural Space Radiation Environment Hazard
- Radiation Effects on Micro-Electronics
- Hardness Assurance, as a Discipline, with its Challenges
 - New Technologies
 - New Architectures
 - Unbound Risks
- Building Smart Requirements
- Risk Acceptance and Guidance

New Space & SmallSats – Same Old Radiation



The need for Radiation Hardness Assurance (RHA)

- Radiation effects are a mix of disciplines, evolve with technologies and techniques
- Misinterpretation of failure modes / misuse of available data can lead to over/under design
- New mission concepts and SmallSat paradigm •
 - Challenges identified in the past are here to stay; adoption of new technologies are often the risk driver
 - Commercial Space, Small missions, Constellations will benefit from detailed hazard definition and mission specific requirements
 - RHA flow doesn't change, risk acceptance needs to be tailored
- Some Top Level Resources
 - NPR 7120.5 NASA Agency Program Management
 - GPR 8705.4 Goddard Risk Assessments



https://sdo.gsfc.nasa.gov



https://www.nasa.gov/van-allen-probes

Supernovae



New Space – New Point of View

SmallSats / Constellations / Swarms



Component Grades are Merging



ESSCON : Eccofet

Risk acceptance is being used as a means to enable innovation

Who Needs This Guidance?

Universities / CubeSats

- May be first-time designers, or previous missions did not have requirements
- Schedule driven, limited time for development
- Rideshares could end up in multiple environments
- Space Agencies / Government
 - More designs in new destinations
 - Cost savings of SmallSat platform, with more reliable outcome
 - More risk acceptance
- Device / Subsystem Manufacturers
 - Product / Device offerings (middle of the road seems to be the target)
 - Fault tolerance in designs



Michael Swartwout, SLU CubeSat Database



NASA's Goddard Space Flight Center/Bill Hrybyk

Notional Questions to Keep in Mind

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- What are the radiation risks:
 - What is the hazard?
 - What are the challenges?
- What can you do to reduce the risk for a given hazard?
- How do *similar* systems/devices react in the space environment?
- What does changing that radiation environment mean for success?
- Need availability throughout the mission or at specific times?



Natural Space Radiation Environment





Energetic supernovae remnants (~GeV, Z=1-92) Originate outside of our solar system Solar Wind, Solar Cycle CMEs (proton rich) Flares (heavy ion rich) Fluctuate with Solar Activity and Events Not a perfect dipole Protons and Electrons trapped at different L-shell values and energies

Natural Space Radiation Environment

- NASA
- Plasma Degradation of micro-electronics wear-out Particle Radiation Degradation of optical components Degradation of solar cells Neutral Gas Particles UV and X-Ray Data corruption Noise on images **Orbital Debris** System shutdowns or resets Circuit Damage Part tolerances exceeded

Spacecraft Charging, Ionizing Dose, Non-Ionizing Dose, Single Event Effects, Drag, Surface Erosion, Debris/Micro-Meteoroid Impacts, Thermal Cycles

Natural Space Radiation Environment





Units explanation

Degradation

- Total Ionizing Dose (TID)
 - Absorbed Dose (rad(Si))
 - 1 rad = 100 erg/g = 0.01 J/kg; 100 rad = 1 Gy
 - Always specified for a particular material
 - 1 rad(SiO2), 10 krad(Si), 100 Gy(H2O)
 - This is not exposure (R), or dose equivalent (Sv)
- Non Ionizing Energy Loss (NIEL)
 - Fluence (p/cm2)
 - Number of particles per unit area

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Single Events

- Linear Energy Transfer (LET)
 - Stopping Power Normalized to target material

$$S = -\frac{dE}{dx} \Rightarrow \text{LET} = -\frac{1}{\rho}\frac{dE}{dx}$$

- Units are MeV·cm²/mg
- Rate (/device or /bit per time interval)

Degradation Contributors vs. Single Event

Shielding

۷S.

Total Ionizing Dose

Approximate Dose

Cumulative effects

- Depend highly on which contributors and duration in their presence
- Mimic wear-out/aging
- NIEL and TID must be accounted for

Typical destinations (LEO, GEO)

- LEO at low altitude/inclination is more protected by the Geomagnetic field
- Proximity to the poles & SAA show a large variability in dose despite short mission durations
- Electrons and their braking radiation are the big offender in Geostationary orbits (don't forget about spacecraft charging...)

Note that

- A little bit of shielding goes a long way
- Altitude plays a huge role when in/near the radiation belts (even transiting)
- Beyond Geomagnetic field, highly variable solar environment contributions (Solar cycle)

Degradation has a strong dependence on where you go, not just how long you are on orbit





To be presented by Michael J. Campola at the 2019 SEE-MAPLD Single Event Effects (SEE) Symposium and Military and Aerospace Programmable Logic Devices (MAPLD) Workshop, La Jolla, CA, May 20-23, 2019.

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Orbit

and

Altitude a

104 IO4

Flu

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Degradation vs. Single Event Contributors

• One particle causes the effect

- Random in nature, particle must traverse sensitive structure within device and have sufficient charge creation along its path
- Shielding doesn't do so much for highly energetic particles
- Device technology can be dependent on particle species

• Typical Destinations (LEO, GEO)

- Again altitude plays a role; for some devices that is a direct threat
- You are exposed to more GCR + Solar contribution as geomagnetic protection is reduced
- Natural phenomena (SAA, magnetic poles) are temporal drivers

Note that

- There will be a background rate, solar cycle dependence, solar event rate, increased rate for poles or SAA – not just one rate to consider
- Always dependent on mission

Single event contributors benefit very little from shielding, have dependence on materials near the sensitive volume







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Summary of Environmental Hazards



	Plasma (charging)	Trapped Protons	Trapped Electrons	Solar Particles	Cosmic Rays	Human Presence	Long Lifetime (>10 years)	Nuclear Exposure	Repeated Launch	Extreme Temperature	Planetary Contaminates (Dust, etc)
GEO	Yes	No	Severe	Yes	Yes	No	Yes	No	No	No	No
LEO (low- incl)	No	Yes	Moderate	No	No	No	Not usual	No	No	No	No
LEO Polar	No	Yes	Moderate	Yes	Yes	No	Not usual	No	No	No	No
International Space Station	No	Yes	Moderate	Yes - partial	Minimal	Yes	Yes	No	Yes	No	No
Interplanetary	During phasing orbits; Possible Other Planet	During phasing orbits; Possible Other Planet	During phasing orbits; Possible Other Planet	Yes	Yes	No	Yes	Maybe	No	Yes	Maybe
Exploration – Lunar, Mars, Jupiter	Phasing orbits	During phasing orbits	During phasing orbits	Yes	Yes	Possibly	Yes	Maybe	No	Yes	Yes

https://radhome.gsfc.nasa.gov/radhome/papers/SSPVSE05_LaBel.pdf

Radiation Hazard Contributors for Dose and SEE



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Environment

	LEO Equatorial	LEO Polar (Sun Sync)	GEO / Interplanetary
> 3 Years	Moderate Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	High Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	High Dose / High GCR, High Solar Proton Variability
1- 3 Years	Manageable Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	Moderate Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	High Dose / High GCR, High Solar Proton Variability
< 1 Year	Manageable Dose / Attenuated GCR, Trapped Proton, SAA, Some Solar Proton dependence for variation	Moderate Dose / Higher GCR, High Energy Trapped Protons in SAA and Poles, Some Solar Proton dependence for variation	Moderate Dose / High GCR, High Solar Proton Variability

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Mission Lifetime

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Radiation Effects on Active Microelectronic Devices

- Cumulative effects and single event effects can <u>both</u> be permanently damaging
 - TID/DDD lead to wear-out of device operation and degrade devices beyond acceptable operations internally and externally
 - Single Event Effects can be catastrophic instantaneously by turning on parasitic devices within the semiconductor or inducing electric field across dielectrics that eventually break down
 - Synergistic effects can make ground based testing very difficult

Destructive Single Event Effects (SEEs)

- Irreversible processes
- Terms: Latchup, Burnout, Gate Rupture

Non-Destructive SEEs

- Lead to interruptions in operation and/or errors leading to unknown state spaces or loss of science / mission if not accounted for
- Terms: Functional Interrupt, Transients, Upsets
- Short Courses / Presentations / Papers / IEEE
 - NSREC, RADECS, SEE/MAPLD, NEPP ETW, HEART, GOMAC, SPWG, MRQW, SERESSA

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Final-Web-Pres-ETW-Diodes-TN16278 v2.pdf



Device and Particle Interraction





Brock J. LaMeres, Colin Delaney, Matt Johnson, Connor Julien, Kevin Zack, Ben Cunningham Todd Kaiser, Larry Springer, David Klumpar, "Next on the Pad: RadSat – A Radiation Tolerant Computer System," Proceedings of the 31st Annual AIAA/USU Conference on Small Satellites, Logan UT, USA, Aug. 5-10, 2017, paper: SSC17-III-11, URL: http://digitalcommons.usu.edu/cgi/viewcontent.cgi?article=3618&context=smallsat

Table of SEE susceptibility



SEL	SEGR	SEB	SEDR	Stuck Bit	SEU/MCU	SET	SEFI
		POWER	One-time		Digital/bistable	bipolar	Complex
CMOS	MOSFET	MOSFET	Prog. FPGA	SRAM	technologies	technology	Microcircuits
			Bipolar			Analog	
Bipolar?	FLASH	Power JFET	Microcircuits	DRAM	Deep submicron	microcircuit	ADCs
	Schottky				CMOS more MCU	Digital	
	Diode	Power BJT		FLASH	susceptible	microcircuit	PWMs

Part-Level Consequences

How Common is Issue?

- Catastrophic failure possible
- Destructive but limited
- Nondestructive

- Common in technology
- Catastrophic failure possible
- Not seen but possible in principle

Ray Ladbury, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

List is not exhaustive, but new failure modes are found in new devices, so it would not be possible to capture all

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 - New Technologies
 - New Architectures
 - Unbound Risks
- Building Smart Requirements
- Risk Acceptance and Guidance

The Job: Watch out for the 'ilities



Survivability

- Must survive until needed
- Entire mission?
- Screening for early failures in components

Availability

- Must perform when necessary
- Subset of time on orbit
- Operational modes
- Environmental response

Criticality

- Impact to the system
- Part or subsystem function
- Mission objectives

Reliability

- Resultant of all
- Many aspects and disciplines
- Known unknowns

The People: Radiation Effects Engineers

Materials	Device Physics	Electrical Engineering	Systems Engineering	Space Physics
 Material Property degradations with radiation Energy loss in materials 	 Charge transport Device Process Dependencies Charge dependency of device operation 	 Part to part interconnections Understanding circuit response Device functions and taxonomy 	 Requirements System Level Impacts Understanding interconnections Understanding functionality 	 Space weather Environment models/modeling Radiation Sources and variability

Radiation Hardness Assurance (RHA) Overview



RHA consists of all activities undertaken to ensure that the electronics and materials of a space system perform to their *design* specifications throughout exposure to the mission space environment



(LaBel)

(Poivey)

Radiation Hardness Assurance Flow



RHA consists of all activities undertaken to ensure that the electronics and materials of a space system perform to their *design* specifications throughout exposure to the mission space environment

Define the Environment					
 External to the spacecraft 					
Evaluate the Environment					
 Internal to the spacecraft 					
Define the Requirements					
 Define criticality factors 					
Evaluate Design/Components					
 Existing data/Testing 					
 Performance characteristics 					
"Engineer" with Designers					
 Parts replacement/Mitigation 	schemes				
Iterate Process					
 Review parts list based on updated knowledge 					

K.A. LaBel, A.H. Johnston, J.L. Barth, R.A. Reed, C.E. Barnes, "Emerging Radiation Hardness Assurance (RHA) issues: A NASA approach for space flight programs," IEEE Trans. Nucl. Sci., pp. 2727-2736, Dec. 1998.

RHA Challenges... Not So Small

- Always in a <u>dynamic</u> environment
- New Technologies
 - Device Topology / Speed / Power
 - Increased COTS parts / subsystem usage
- New Mission Architectures
 - Profiles of mission life, objective, and cost are evolving
 - Oversight gives way to insight in some mission classifications
 - Ground systems, do no harm, hosted payloads
 - Similarity and heritage data requirement widening
- Quantifying Risk
 - Translation of system requirements to radiation trades can be problematic
 - Determining appropriate mitigation level (operational, system, circuit/software, device, material, etc.)

Unbound radiation risks are likely



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25 IEEE/DOI:10.1109/IITC-AMC.2016.7507637

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New Technologies - New Susceptibilities

- Feature Size / Critical Charge
 - Will we now have sensitivity to muons? Low energy protons?
- 3D Stacking/Structures
 - Deep sensitive volumes
 - New materials
- Testing Challenges
 - Complexity (e.g. SoCs)
 - Speed of interfaces
 - Obfuscation of state space
 - Flux / Range of beam @ facilities

Without a lot of part information you may not have a representative characterization of the radiation threats to the device or technology.



Interconnects

Au Interconnection



Top layer

Interface

Middle lave

Interface

Base layer

Au interconnection

14nm SOC High Density Interconnect Stack

Wring in



Single Strain

Allowable Losses



Redundancy alone does not remove the threat, adds complexity

Quantifying Risk – Likelihood vs. Consequence



From Risk Assessment GPR 7120.5

Likelihood	Safety Estimated likelihood of Safety event occurrence	Technical Estimated likelihood of not meeting performance requirements	Cost Schedule Estimated likelihood of not meeting cost or schedule commitment
5 Very High	$(P_{SE} > 10^{-1})$	$(P_T > 50\%)$	$(P_{CS} > 75\%)$
4 High	$(10^{\text{-}2}\!<\!P_{SE}\!\le\!10^{\text{-}1})$	$(25\% < P_T \le 50\%)$	$(50\% < P_{CS} \le 75\%)$
3 Moderate	$(10^{\text{-3}}\!<\!P_{SE}\!\le\!10^{\text{-2}})$	$(15\% < P_T \le 25\%)$	$(25\% < P_{CS} \le 50\%)$
2 Low	$(10^{\text{-5}}\!<\!P_{SE}\!\le\!10^{\text{-3}})$	$(2\% < P_T \le 15\%)$	$(10\% < P_{CS} \le 25\%)$
1 Very Low	$(10^{\text{-6}}\!<\!P_{SE}\!\le\!10^{\text{-5}})$	$(0.1\% < P_T \le 2\%)$	$(2\%\!<\!P_{CS}\!\le\!10\%\!)$



Can only get there with enough information about the system or the chosen device, need to have a known hazard and a known response

radiation effects Solar Particle Spectra Trapped Particle Populations Parts Radiation Sensitivity

Candidate Parts List



Hardness Assurance is the

practice of designing for

Competing failure modes

Galactic Cosmic Ray Spectra What it takes to overcome the



Mission Requirements

Define and Evaluate the

Hazard

Description of the mission radiation Environment:

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RHA Flow Doesn't Change With Accepted Risk

Evaluate the

to the Hazard

Circuit Response



Focus For Risk Acceptance



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Failure Awareness

- Know your hazard from the natural environment
- Know your devices potential failure mechanisms or response (data)

Countermeasures and Mitigation

- Where are they necessary?
- Where are they effective?
- At what level (part, card, box, mission)

Smart Requirements – and Eventually Smart Trades

Failure Awareness

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Define and Evaluate the Hazard



Aerospace Programmable Logic Devices (MAPLD) Workshop, La Jolla, CA, May 20-23, 2019.

Define and Evaluate the Hazard

- Define the Environment
- External to the spacecraft
- Evaluate the Environment
 - Internal to the spacecraft
- Define the Requirements
 - Define criticality factors
- Evaluate Design/Components
 - Existing data/Testing
 - Performance characteristics
- "Engineer" with Designers
 - Parts replacement/Mitigation schemes
- Iterate Process
 - Review parts list based on updated knowledge

Environment Severity/Mission Lifetime





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Derive Smart Requirements

- Define the Environment
 - External to the spacecraft
- Evaluate the Environment
 - Internal to the spacecraft
- Define the Requirements
 - Define criticality factors
- Evaluate Design/Components
 - Existing data/Testing
 - Performance characteristics
- "Engineer" with Designers
 - Parts replacement/Mitigation schemes
- Iterate Process
 - Review parts list based on updated knowledge

		Low	Medium	High
ticality/Availability	High	Dose-Depth / Ray-trace GCR and Proton Spectra for typical conditions	Dose-Depth / Ray-trace GCR and proton Spectra for all conditions	Ray-Trace for subsystem / GCR and proton Spectra for all conditions
	Medium	Dose-Depth / GCR and proton spectra for background	Dose-Depth / GCR and Proton Spectra For background	Dose-Depth evaluation at shielding / All spectra conditions
Cri	Low	Similar mission dose, same solar cycle / GCR spectra	Dose-Depth / GCR spectra	Dose-Depth / GCR and Proton Spectra For background

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Mitigation and Countermeasure Optimization

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- Review parts list based on updated knowledge



Dose-Depth Curves

Building Requirements



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- Requirements by Environment
- Requirements by Technology
- Cases that may need additional considerations

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Requirements by Environment

Van Allen Belts

- Can lead to high doses in a short mission: Jovian
- Can lead to spatially dependent SEE responses: South Atlantic Anomaly
- Solar Orbits
 - Solar Events, highly dynamic, energetic, directional
 - Solar Wind, will depend on the solar cycle

In essence the requirements are always driven by the environment, some more than others create a unique challenge





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Requirements by Technology

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- Technologies exhibit specific physics of failure
 - Not easy to group them all
 - Opto-electronics Displacement in the material
 - Bipolar Enhanced Low Dose Rate Sensitivity
 - Digital CMOS Latchup and SEFI
 - Power devices SEGR/SEB

Test Data requirements

- Failure distributions, often not enough parts
- Destructive effects are one data point, variability from part to part
- Statistics of the fit for rate calculations

Requirements should only be made applicable to the technologies that need to meet mission objectives and can benefit



Why you can't relax an LET requirement

- Not like wear-out, flat-line risk
- Rate calculations are not the same for DSEE vs. Non-destructive
 - Data are a limiting factor
 - One part = one data point

When you require by LET:

- Spectrum from environment is then imparted on sensitive volumes
- LET increases at angle critical charge is what we are trying to determine
- Deep SV doesn't get same LET each time
- CRÈME Calculation integrates





Ray Ladbury, NSREC2017 SC, https://ntrs.nasa.gov/archive/nasa/casi.ntrs.nasa.gov/20170006865.pdf

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Tailored Filtering, EDAC, or Scrubbing SEL

• Environment and technology driven, risk avoidance

Don't harm downstream parts, or accumulate

Protection circuitry / diode deratings

• SEGR, SEB

SEE, SET

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- Effect driven, normally incident is worst case
- Testing to establish Safe Operating Area (SOA)
- MBU, MCU, SEFI, Locked States
 - Application Voltage or Pattern dependence
 - Watchdogs / reset capability

• Proton SEE susceptible parts need evaluated in detail:

https://nepp.nasa.gov/files/25401/Proton_RHAGuide_NASAAug09.pdf





So you don't care about dose?

- Maybe degradation of a part beyond usage is okay?
- Did you forget about DDD?

Maybe you do!

- Short Mission, common failure mode
- Low mass budget, can optimize shielding if you have failure distribution of parts.



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Say you want to survive a flare? Think Availability

- Don't dose out during storm (nor the full mission)
 - Calculate the dose (TID/DDD) of the mission in full 95% confidence level recommended
 - Calculate the dose contribution from N number of events (protons & x-rays)
 - If dose from N is > 5% of the total dose, increase confidence level of full mission model
- Don't destructively fail from a single particle during the storm (nor the full mission)
 - Standard risk-avoidant SEE approach: no destructive effects allowed
 - LET threshold for single event latchup (SEL) > 75 MeV.cm2/mg
 - LET threshold for single event burnout, gate rupture, dielectric rupture (SEB, SEGR, SEDR) > 37 MeV.cm2/mg (particles must come from normal incidence to cause effect)
- If you have non-destructive single event upsets, they can't overwhelm critical instruments/systems during the storm
 - Rate calculation requires part data representative of the application, looking for crosssection over LET.
 - If a parts' LET threshold is anywhere from 20 to 75 MeV.cm2/mg, need heavy ion rate
 - If a parts' LET threshold is below 20, need direct ionization from protons (can be builtin to heavy ion calculation) and indirect ionization from recoil ions contribution to rate (need proton data) – make sure packaging materials don't add to this
 - Do you need to mitigate or not confirm that event rates are not higher than mitigation (Markov process... i.e. EDAC beats the number accrued, scrub rate is faster than critical number of upset accumulation)







Risk Acceptance – Data Available?

Part Classifications Growing

- Mil/Aero vs. Industrial vs. Medical
- Automotive vs. Commercial vs. Modified HiRel

Substitute in COTS

- Now you have another degree of separation
- Failure modes not fully understood
- Unlikely to have historical data
- Similarity data no applicable due to fab, process, or design rules
- Cost of testing usually too high

Without traceability you may be depending on nonrepresentative data.



Structure of Constraining Data



Risks abound, would you know the root cause?

Parts

- Parametric degradation and leakage currents allowable in application?
- Downstream/peripheral circuits considered?
- Reset/refresh capability?
- Mitigation within too complex?
- Predicted radiation response unknown– loss of part functionality critical?

Subsystem

- Criticality to mission that the subsystem work?
- Interfaces allow you to get to a known state if all goes wrong?

System

- Increased power dissipation a mission ender?
- Availability outweighed by error circumvention?
- Data retention through reboots? What if there is science data loss?
- Communications interruptions overwhelm?
- Navigation or Attitude determination unable to deal with faults?





When do you test?

- Divine your risk threshold
 - There's a doc coming for that... radhome.gsfc.nasa.gov/nepp.nasa.gov
- Unknown failure modes that would not be acceptable to the mission
 - Known unknowns can be carried as a risk if you already know that the outcome is mitigated at the board or box level
 - New technologies should be identified early on
- Fault propagation may be the problem you wish to mitigate
 - This can include cumulative effects!
 - Fault injection may not be able to cover the state space
- Destructive single event effects are an obvious target
- Can you tolerate a part replacement in your design cycle?



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Notional SmallSat Radiation Guidelines



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Environment

		LEO Equatorial	LEO Polar (Sun Sync)	GEO / Interplanetary
ission Lifetime umed Risk Acceptance)	> 3 Years	Data on all SEE for critical parts, and have data on dose failure distribution on similar parts	Consider mission consequences of all SEE (Data for critical parts), have Dose failure distribution on lot	Have Data on all SEE, Have Data Dose failure distribution on lot
	1- 3 Years	Have Data on DSEE for critical parts	Consider mission consequences of all SEE (Data for critical parts), have data Dose failure distribution on similar parts	Have Data on all SEE for critical parts, Have Data on Dose failure distribution on similar parts
M (With Assı	< 1 Year	Look for data on DSEE for critical parts	Consider mission consequences of all SEE, and look for data on dose failure distribution on similar parts	Consider mission consequences of all SEE, and have data on dose failure distribution on similar parts



- Models and Test Data are brought together to get rates of upset / failure distributions
- Resources and Utilization are the scaling factors with criticality

- Concept of operations
- Requirements and Availability are fed down correctly to subsystem
- Evidence is presented
- Assumptions are tracked

 Description of System **Connections and Dependencies**

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- Receives GSN readily
- Fault propagation can be identified





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THANK YOU