



In-situ K-Ar dating based on UV-Laser ablation coupled with a LIBS-QMS system development, calibration and application

LPSC 50

Abstract ID: 1311

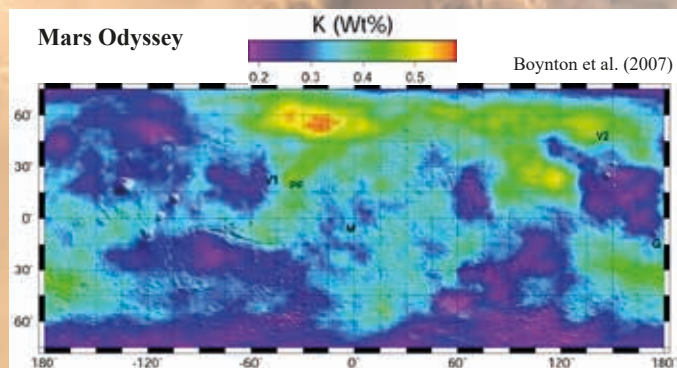
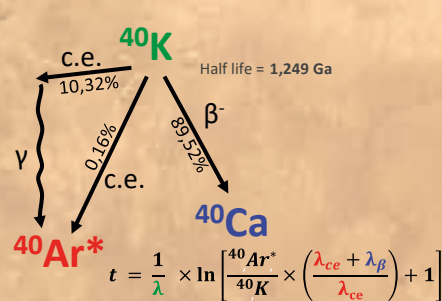
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Principle

Absolute age determination is necessary to check and calibrate the relative Martian chronology presently available from meteoritic crater counting.

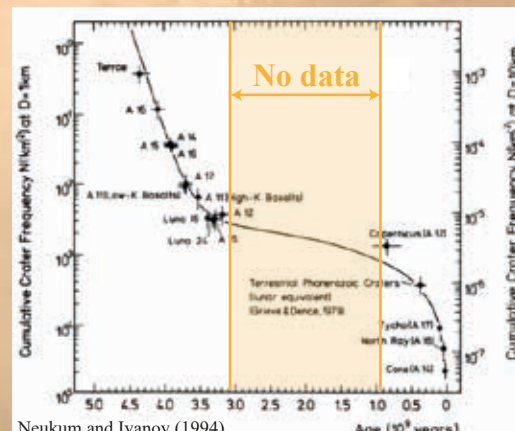


K-Ar dating method is particularly suitable for such purpose.

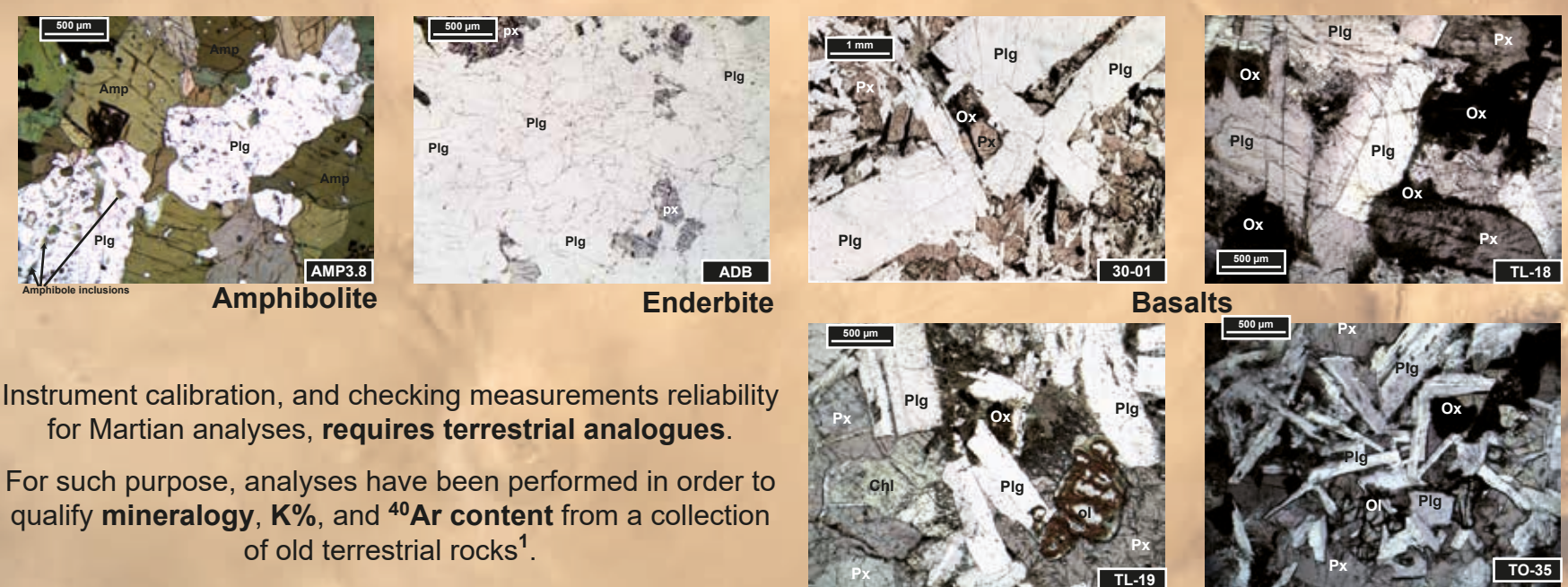
It is based on the radioactive decay of ⁴⁰K, a major element universally distributed. Its daughter, ⁴⁰Ar, noble gas, which accumulates as a function of time.

The measurement of K and ⁴⁰Ar can be used for *in-situ* dating planetary surface rocks as long as the sample mass from which they were extracted is known.

We present here an *in-situ* K-Ar dating prototype, **KArMars**¹, based on UV-laser ablation, K measurement by LIBS, and argon analysis by QMS.



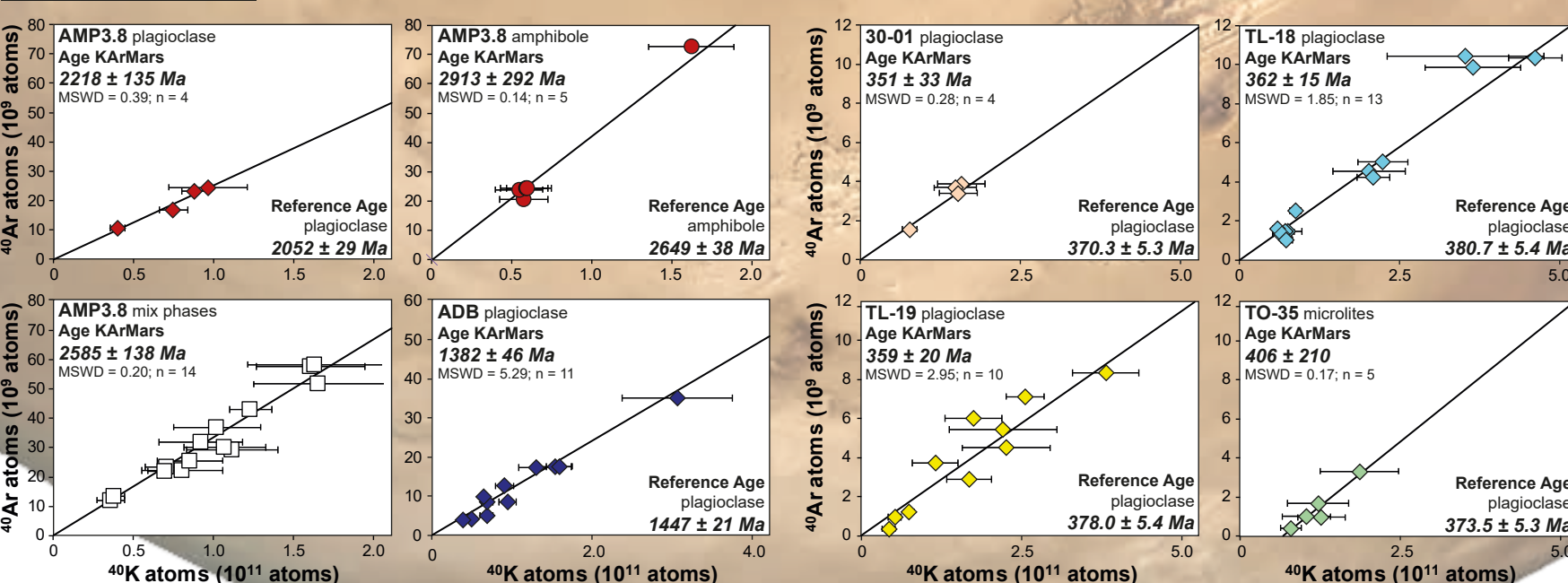
Reference samples



Instrument calibration, and checking measurements reliability for Martian analyses, requires terrestrial analogues.

For such purpose, analyses have been performed in order to qualify mineralogy, K%, and ⁴⁰Ar content from a collection of old terrestrial rocks¹.

Application



UV laser ablations were performed on AMP3.8, ADB, and 30-01, TL-18, TL-19, TO-35 (basalt from Viluy) used here as unknowns. The error bars on each point are defined by the uncertainties on K% (between 5 and 40%), mass (around 4%) and ⁴⁰Ar (less than 2%) measurements¹.

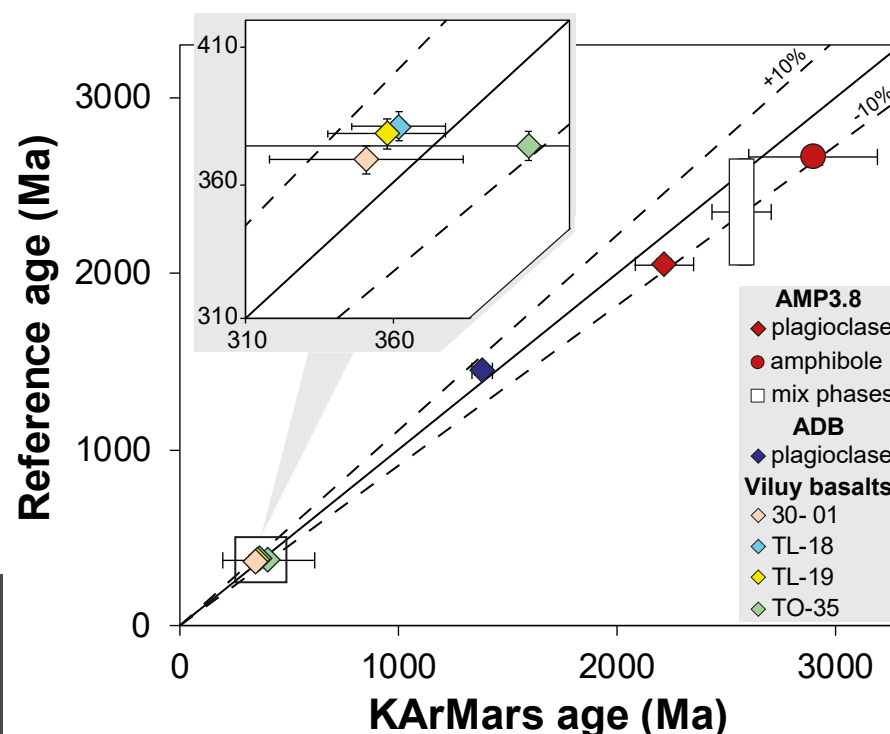
In order to reduce this uncertainty, KArMars ages of these samples have been determined using the **isochron approach**⁸ with the following equation:

$$^{40}\text{Ar} = ^{40}\text{Ar}_i + (\lambda_e / \lambda) \times ^{40}\text{K} (e^{\lambda t} - 1)$$

Where λ is the total decay constant of ⁴⁰K, and λ_e is the decay constant of ⁴⁰K to ⁴⁰Ar

here ⁴⁰Ar_i represents the intercept on the y axis (b) and $(\lambda_e / \lambda)(e^{\lambda t} - 1)$ represent the slope (m) which is a function of the age. Thus $t = 1/\lambda \times \ln [m (\lambda / \lambda_e) + 1]$.

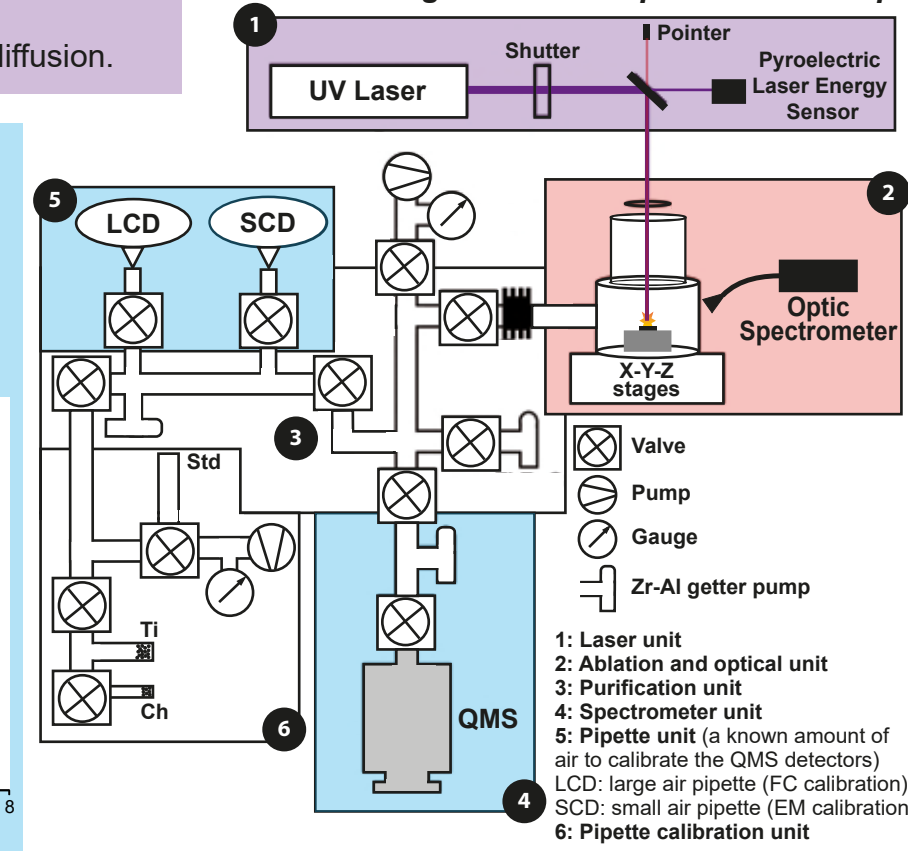
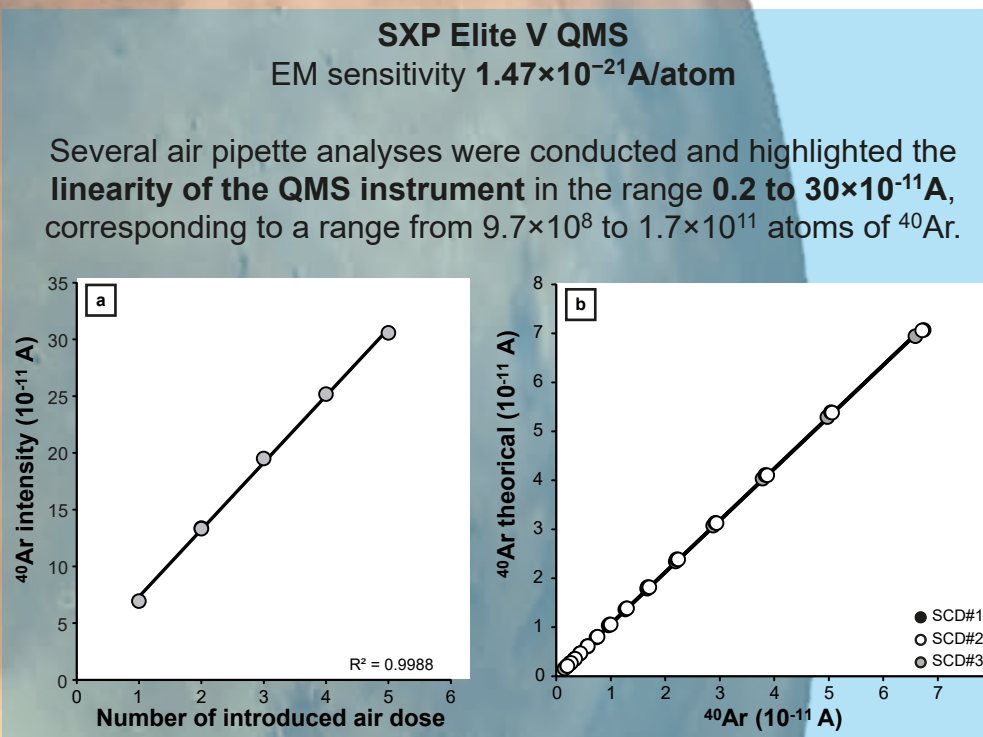
That allows to obtain an uncertainty and an accuracy lower than 10%. This is even observed for relatively young and low K samples from Viluy basalts, which are at the limit of detection of the instruments¹.



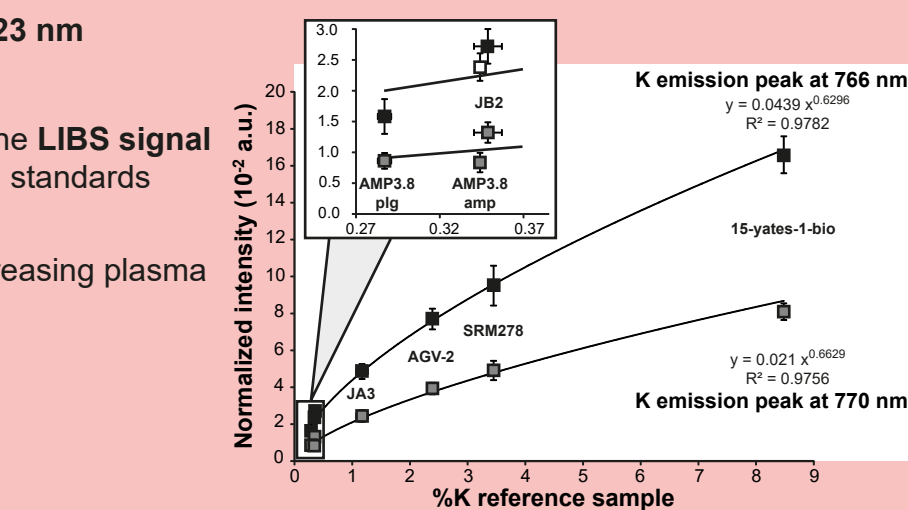
Features and Calibration

Laser unit with a Q-switched Spectra Physics Nd-YAG / 14 ns / 10 Hz / 266 nm
Evapore a specified amount of sample without significant argon diffusion.

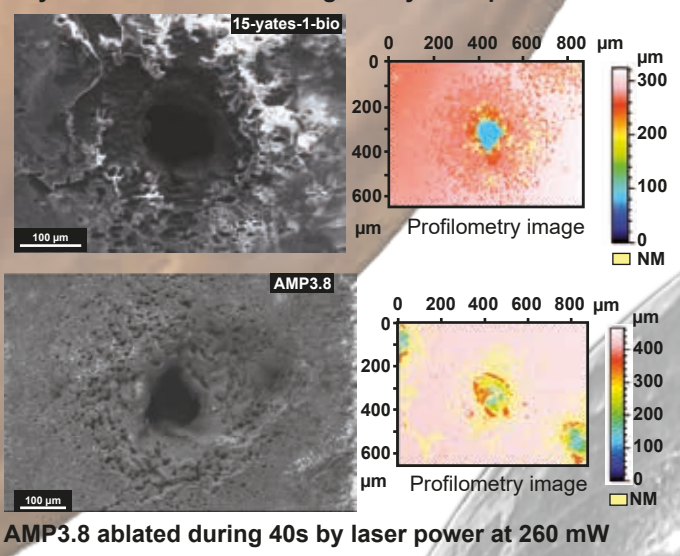
Schematic diagram of the experimental setup



Ocean Optics HR2000+ / gate: 385-835 nm / resolution: 0.23 nm
Univariate analysis was selected to define a correlation between the LIBS signal and a known concentration using calibration curves based on standards analyses (glasses and reference samples).
The non-linear trend can be explained by self-absorption, the decreasing plasma luminosity and/or matrix effects²⁻⁴.
Limits of detection and quantification
LOD766: 0.14% LOD770: 0.05%
LOQ766: 0.52% LOQ770: 0.19%



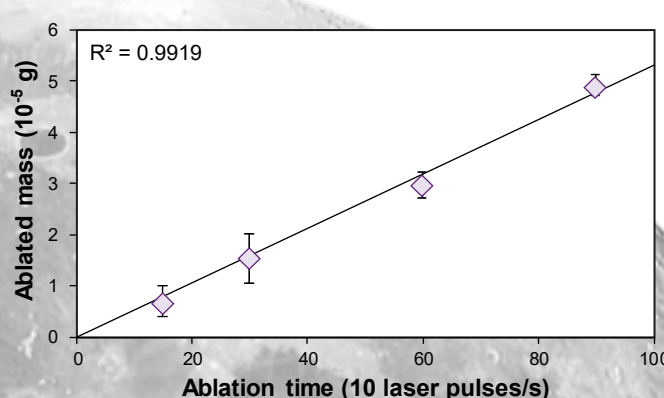
15-yates-1-bio ablated during 60s by laser power at 260 mW



The ablated mass determination depends on laser parameters, and on the mineral analyzed. The profilometry technique can define this mass⁵⁻⁷ but this technique can be affected by several parameters¹.
In this study the mass determination is based on the QMS measurements.

Several analyses of reference plagioclases allowed us to define the relation between the ablation time and the ablated mass.

Indeed, the relation is obtained by comparison between QMS signal and the known content of ⁴⁰Ar atoms per gram of the reference.



B.A. Cohen's team develops a device based on the same principle, **KARLE**⁵, at NASA GSFC

Perspectives

DALI project

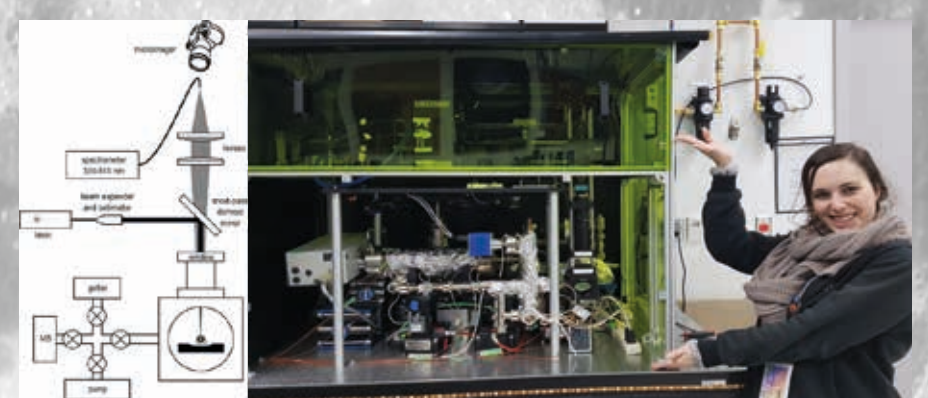
The agency's Development and Advancement of Lunar Instrumentation, or DALI, program recently awarded funding to mature spacecraft-based instruments for use in future lander missions.

Knowing the precise ages afforded through potassium-argon dating would help scientists understand the Moon's history, its formation, the effects of bombardment, and by extension, the history of the solar system. KARLE is an especially good match to the DALI call because all of the necessary components have been proven on flight missions.

As part of the DALI project, we need to miniaturise KARLE and calibrate the instruments for the moon.

PLS method⁹

The intensity of each elemental peak is not a simple function of elemental abundance. Furthermore, LIBS data is sensitive to the composition of the sample. PLS method helps understand the bulk composition of the sample and refine the K calibration curve better than univariate analysis.



KARLE device

[1] Cattani F. et al. (2019) Chem. Geol. [2] Cho Y. et al. (2015) Spectrochim. Acta B At. Spectrosc. 106, 28–35. [3] Cremers D.A. and Radziemski L.J., "Handbook of LIBS", 2006. [4] Guzenoc J. et al. (2017) Spectrochim. Acta B At. Spectrosc. 134, 6–10. [5] Cohen B.A. et al. (2014) Geostand. Geoanal. Res. 38, 421–439. [6] Cho Y. et al. (2016) Planet. Space Sci. 128, 14–29. [7] Cho Y. and Cohen B.A. (2018) Rap. Com. Mass Spectrom. 32, 1755–1765. [8] Bogard D.D. (2009) Meteorit. Planet. Sci. 44, 3–14. [9] Clegg S. et al. (2017) Spectrochim. Acta B At. Spectrosc. 129, 64–85.