

Development of a direct fabrication technique for fullshell x-ray optics

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X-ray Optics after Chandra

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The challenge is to develop optical fabrication technology capable of producing x-ray optics with high angular resolution but with an order of magnitude lighter mirrors and at an affordable price.

The challenge can be approached from different directions:

- Very thin mirrors (0.2-0.4 mm):
 - Replicated optics
 - Pore Optics
 - Figure Correction
- Intermediate thickness mirror (0.5-2.0 mm):
 - Direct fabrication, Chandra-like
 - Figure corrections can be applied too



Direct Fabrication

Material	Density (g/cm³)	CTE (10 ⁻⁶ / K ⁻¹)	Elastic Modulus GPa	Yield Strength MPa
Fused Silica	2.2	0.5	72	48*
Beryllium	1.8	12	318	240
BeAL-162MET	2.1	24	69	276
AlSi	2.8	13.9	193	314
Duralcan F3S.30S AlSi+SiC(30% by vol)	2.8	14.6	120	210

Mechanical Properties of Potential Mirror Substrate Materials

Companies are confident they can deliver the Be, BeAl and AlSi substrates fabricated with necessary tolerances for fine-figuring and polishing

Additional Benefits of metal substrate:

- Fewer joints smaller epoxy-induced errors
- Thermal design could be simplified with the support structure fabricated from the same material

Ideally, the mirror shell has low density, low coefficient of expansion (CTE), high modulus of elasticity and high yield strength. It should also be a material that is not too difficult to figure and polish.

Substrates can be plated with the nickel phosphorous alloy:

- Be + NiP (CATS-ISS telescope)
- BeAl +NiP
- AlSi + NiP

^{*}Maximal achievable value. The 'working' value is typically much less and depends on the surface/subsurface condition.



Direct Fabrication - Metal

Diamond turning

Figuring

Super-polishing

Differential Deposition

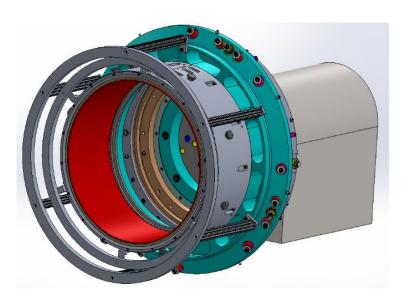
Challenge of shell thickness - fixtures



ID DT fixture

Whiffle tree station with an aluminum shell supported at 12 points.

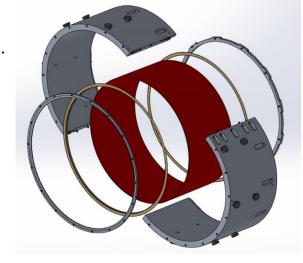
Direct Fabrication



Backing support system installed on the precision lathe for diamond turning. The precision stage (blue) permit alignment of the mirror shell (red) with the lathe.



Thin-shell backing support system. A thin layer of backing material (not shown) acts as interface between the mirror shell (red) and the stiff outer support clamshell (gray). The support rings and the gaskets shown contain the backing material.







Zeeko machine

- The machine utilizes a "bonnet" technique in which an inflated rubber hemispherical diaphragm supports the polishing medium.
- there are different "bonnet" sizes (20 mm, 40 mm and 80 mm radii of curvature)
- This computer-controlled deterministic polishing processes leads to a high convergence rate.





Parametric wear pattern simulation enables a more efficient

method of exploring the polishing parameter space.

| Velocity (m/s) vs_Surface Position (mm) | Pressure (Pa) vs_Start

Based on Preston's law:

Wear rate is proportional to

Bonnet pressure distribution

Velocity of bonnet surface

Velocity depends on

Spindle rotation

Head attack angles

Pressure depends on

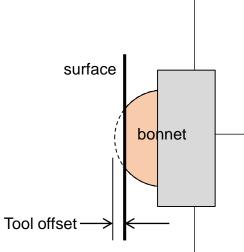
Internal pressure of bonnet

Bonnet structural and mechanical properties

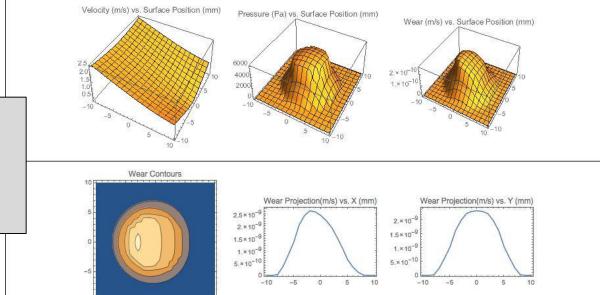
Static (current)

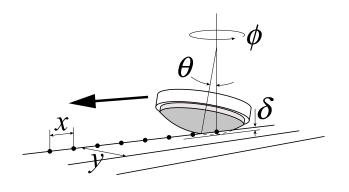
Dynamic (future)

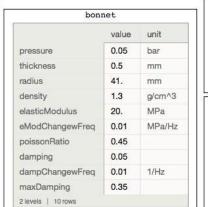
Model is validated using measured data.

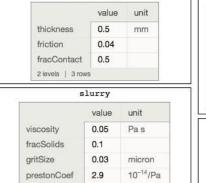


Example of Output from Model Run









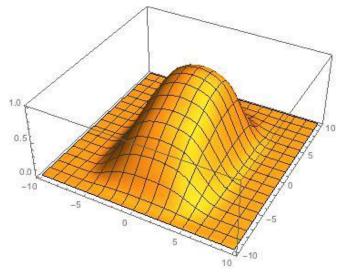
	value	
springK	100000.	N/m
rotationRateSpindle	1000.	RPM
toolOffset	0.3	mm
precessAngle	17.	deg
phiAngle	0.	deg
2 levels 5 rows		

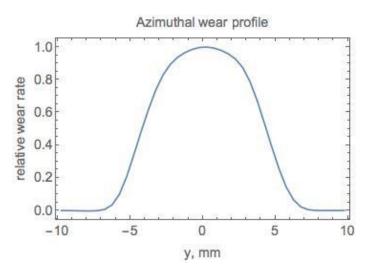
	value	unit
adiusofCurvature	20.	cm
wearGridHalfWidth	10.	mm
wearGridStep	1.	mm

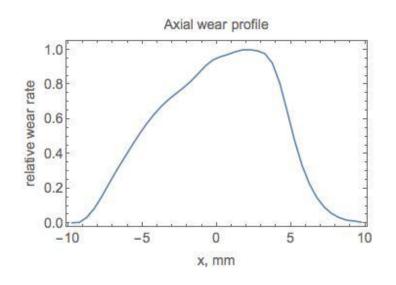
SPIE, Space Telescopes and Instrumentation 2016: Ultraviolet to Gamma Ray, Edinburg, UK, June 29, 2016

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Measured wear function

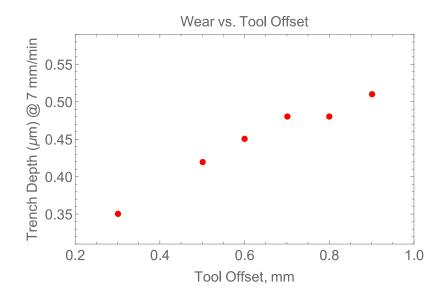






- Axial profile is asymmetric
 - Due to radial dependence of bonnet velocity relative to surface
 - Operate at 15° "precess" angle
- Axial and azimuthal directions are selected based on clearance considerations for shell polishing

- Explored polishing parameter space to balance:
 - Wear rate
 - Accuracy
- Parameter optimization
 - Bonnet pressure
 - Spindle speed
 - Tool Offset



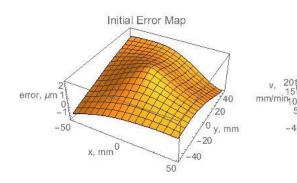
• Example: Tool Offset
Trench depth vs. tool offset flattens in 0.7-0.8
mm range
Results in minimal sensitivity to alignment in a
±50 μ m range around 0.75 mm
Selected 0.75 mm as operational tool offset

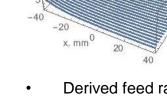


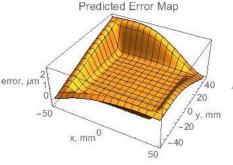


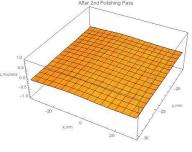
- Several options for determining the feed rate map from error map have been explored
 - Initially using a Richardson-Lucy deconvolution algorithm + small nonlinear correction
 - Tends to generate smoother edge transitions

Feed Rate Map









- Initial error map derived from metrology data.
- RMS slope errors along x-axis are 8 arcsec.
- Derived feed rate map.
- Feed rates range from 1.35-20 mm/min.
- Predicted result of polishing iteration.
- Predicted RMS slope errors are 0.6 arcsec
- Error map derived from metrology data.
- RMS slope errors were reduced to 0.5 arcsec

The polishing routine software developed is validated and permit the fabrication of high resolution x-ray optics!!



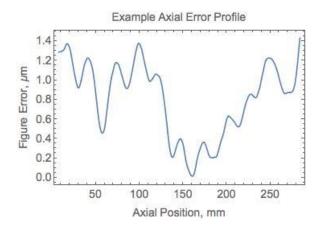


Mandrel Demonstration

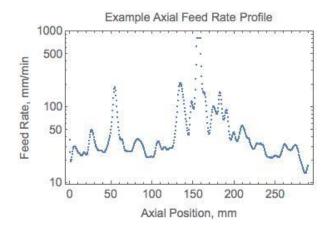
- Figure error correction process:
 - · Initial Radii vs. axial position measured with CMM
 - Axial error map generated from Zygo data (36 meridians measured).
 - Feed rates determined from error map
 - Use axial wear function as deconvolution kernel
 - Use iterative Richardson-Lucy deconvolution algorithm
 - Select iteration that produces the smallest residual RMS slope errors

Overall strategy

- Initial mandrel errors >99.5% cylindrical so derive a feed rate map from 1D deconvolution of axial error map and apply to all meridians scaled for small azimuthal variations.
- Scale the feed rates so run time fits in a work day
- Perform a series of runs over required number of days alternating CW-CCW on alternate days to minimize daily wear rate drift
- Iteratively measure(Zygo) and generate new feed rate map as required
- Continue until cylindrical errors are comparable to azimuthally incoherent errors
- Derive a feed rate map from 2D deconvolution of 2D error map
- Perform a final series of polishing runs, possibly alternating with lap polishing to reduce out-of-band errors.

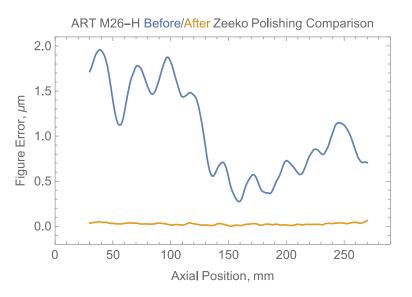


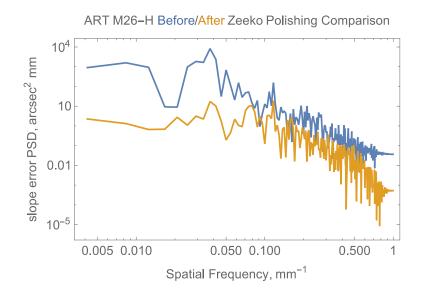












	before	after
Figure error (St. Dev.)	500 nm	10.7 nm
Slope error (> 2 cm) (RMS)	6.32 arcsec	0.30 arcsec
Low frequency (> 7 cm) slope error (RMS)	2.66 arcsec	0.09 arcsec
Mid frequency (2-7 cm) slope error (RMS)	5.73 arcsec	0.29 arcsec

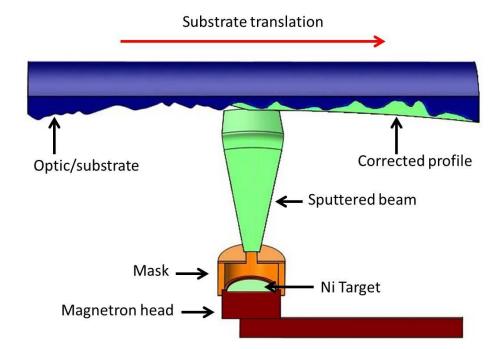


Future Work

- Demonstrate the technology on a mandrel from scratch.
- Transfer the technology to NiP plated shells and segments.
- Use the differential deposition technique for further improvement of the angular resolution (to remove features with spatial extend below 7 cm)



. A NiP plated, segmented aluminum pipe with a conical inner figure fabricated for metrology, figuring and superpolishing experiments



Conclusions



- MSFC is developing direct fabrication technology for full-shell x-ray optics made from metal substrates;
- Support fixtures for diamond-turning, polishing and metrology have been developed;
- Wear functions were determined using NiP plated flat samples;
- Tool path generation software has been developed;
- Slope error level is consistent with 1-2 arcsec optics
- The technique can be married with the Differential Deposition for further improvements