university of groningen

## University of Groningen

## AdS-Carroll branes

Clark, T. E.; ter Veldhuis, T.

Published in:
Journal of Mathematical Physics

DOI:
10.1063/1.4967969

IMPORTANT NOTE: You are advised to consult the publisher's version (publisher's PDF) if you wish to cite from it. Please check the document version below.

Document Version
Publisher's PDF, also known as Version of record

Publication date:
2016

Link to publication in University of Groningen/UMCG research database

Citation for published version (APA):
Clark, T. E., \& ter Veldhuis, T. (2016). AdS-Carroll branes. Journal of Mathematical Physics, 57(11), [112303]. https://doi.org/10.1063/1.4967969

## Copyright

Other than for strictly personal use, it is not permitted to download or to forward/distribute the text or part of it without the consent of the author(s) and/or copyright holder(s), unless the work is under an open content license (like Creative Commons).

Take-down policy
If you believe that this document breaches copyright please contact us providing details, and we will remove access to the work immediately and investigate your claim.

Downloaded from the University of Groningen/UMCG research database (Pure): http://www.rug.nl/research/portal. For technical reasons the number of authors shown on this cover page is limited to 10 maximum.

## AdS-Carroll branes

T. E. Clark, and T. ter Veldhuis

Citation: Journal of Mathematical Physics 57, 112303 (2016); doi: 10.1063/1.4967969
View online: https://doi.org/10.1063/1.4967969
View Table of Contents: http://aip.scitation.org/toc/jmp/57/11
Published by the American Institute of Physics

## Articles you may be interested in

Oscillator versus prefundamental representations
Journal of Mathematical Physics 57, 111702 (2016); 10.1063/1.4966925
Inverse eigenvalue problems
Journal of Mathematical Physics 57, 112102 (2016); 10.1063/1.4964390
All-order bounds for correlation functions of gauge-invariant operators in Yang-Mills theory Journal of Mathematical Physics 57, 122301 (2016); 10.1063/1.4967747

Exactly solvable Hermite, Laguerre, and Jacobi type quantum parametric oscillators Journal of Mathematical Physics 57, 122107 (2016); 10.1063/1.4972293

Bender-Wu singularities
Journal of Mathematical Physics 57, 122109 (2016); 10.1063/1.4972290
Cosmic anisotropic doomsday in Bianchi type I universes
Journal of Mathematical Physics 57, 112502 (2016); 10.1063/1.4967954

| MANAGER'S GUIDE | READ NOW |
| :---: | :---: | :---: |
| WHITEPAPERS | Accelerate R\&D with |
| Multiphysics Simulation | PRESENTED BY |

## AdS-Carroll branes

T. E. Clark ${ }^{1, \mathrm{a})}$ and T. ter Veldhuis ${ }^{2, \mathrm{~b})}$<br>${ }^{1}$ Department of Physics and Astronomy, Purdue University, West Lafayette, Indiana 47907-2036, USA<br>${ }^{2}$ Van Swinderen Institute for Particle Physics and Gravity, University of Groningen, Nijenborgh 4, 9747 AG Groningen, The Netherlands and Department of Physics and Astronomy, Macalester College, Saint Paul, Minnesota 55105-1899, USA

(Received 18 May 2016; accepted 6 November 2016; published online 30 November 2016)

Coset methods are used to determine the action of a co-dimension one brane (domain wall) embedded in $(d+1)$-dimensional AdS space in the Carroll limit in which the speed of light goes to zero. The action is invariant under the nonlinearly realized symmetries of the AdS-Carroll spacetime. The Nambu-Goldstone field exhibits a static spatial distribution for the brane with a time varying momentum density related to the brane's spatial shape as well as the AdS-C geometry. The AdS-C vector field dual theory is obtained. Published by AIP Publishing. [http://dx.doi.org/10.1063/1.4967969]

## I. INTRODUCTION

The symmetries of spacetime delimit the form of the action for fields on it. The familiar case of Poincaré symmetric spacetime results in particle motion being restricted to the local forward lightcone. This lightcone opens up to be the forward time half space in the Galilean limit in which the speed of light $c \rightarrow \infty$ and instantaneous interaction is possible. On the other hand, as the speed of light vanishes, $c \rightarrow 0$, the causal lightcone closes to be just the forward time half-line. Such a contraction of spacetime is known as Carroll spacetime with symmetries generated by the Wigner-Inönü contracted Poincaré algebra, $c \rightarrow 0$, to the Carroll algebra. ${ }^{1,2}$ A particle in such a spacetime must remain stationary as the time axis is the lightcone. This lack of motion can be found by considering the $c \rightarrow 0$ limit of its Poincaré geodesic action. For a free particle moving in $1+1$ dimensional Minkowski spacetime its action is given by

$$
\begin{align*}
\Gamma & =-m c^{2} \int d \tau=-m c^{2} \int \sqrt{d t^{2}-d x^{2} / c^{2}} \\
& =-m c^{2} \int d t \sqrt{1-\dot{x}(t)^{2} / c^{2}} \tag{1.1}
\end{align*}
$$

Introducing a Lagrange multiplier auxiliary velocity $v(t)$

$$
\begin{equation*}
\dot{x}(t) / c=\tanh v(t) \tag{1.2}
\end{equation*}
$$

the action becomes

$$
\begin{equation*}
\Gamma=-m c^{2} \int d t \cosh v(t)\left[1-\frac{1}{c} \dot{x}(t) \tanh v(t)\right] \tag{1.3}
\end{equation*}
$$

In order to take the Carroll limit $c \rightarrow 0$, the velocity is scaled by the speed of light $v(t)=2 c w(t)$ yielding the action

$$
\begin{equation*}
\Gamma=-m c^{2} \int d t \cosh 2 c w(t)\left[1-\frac{1}{c} \dot{x}(t) \tanh 2 c w(t)\right] \tag{1.4}
\end{equation*}
$$

[^0]Letting $c \rightarrow 0$ the Carroll limit for the action $\Gamma_{C}=\Gamma / m c^{2}$ is obtained

$$
\begin{equation*}
\Gamma_{C}=-\int d t[1-2 w(t) \dot{x}(t)] . \tag{1.5}
\end{equation*}
$$

As expected there is no causal relation between different events along the particle's trajectory and it remains stationary $\dot{x}(t)=0=\dot{w}(t)$. ${ }^{3,4}$

Extending the limiting procedure to membranes inserted into ( $\mathrm{d}+1$ )-dimensional Minkowski spacetime, the $c \rightarrow 0$ contraction yields the Carrollian Nambu-Goto action for the brane. Such a limit occurs in the case of effective field theory of tachyon brane condensation in which the tachyon field rolls to the Carrollian limit. ${ }^{5}$ The Carroll brane action can be obtained from the contraction of the Nambu-Goto action for a one-codimensional brane

$$
\begin{align*}
\Gamma_{\mathrm{NG}} & =-\sigma \int d^{d} x \sqrt{-(-1)^{d} \operatorname{det} g}=-\sigma \int d^{d} x \sqrt{-(-1)^{d} \operatorname{det}\left(\eta_{\mu \nu}-\partial_{\mu} \phi \partial_{\nu} \phi\right)} \\
& =-\sigma \int d^{d} x \sqrt{1-\partial_{\mu} \phi \partial^{\mu} \phi} \tag{1.6}
\end{align*}
$$

where the $(d+1)$-dimensional spacetime has been spontaneously broken to that of a d-dimensional world volume by the formation of a domain wall in the additional dimension. These Poincaré symmetries are compactly described by the invariant interval $d s^{2}=d x^{\mu} \eta_{\mu \nu} d x^{\nu}-d z^{2}$ with $z$ denoting the one-dimensional covolume coordinate and $x^{\mu}$ with $\mu=0,1, \ldots, p=(d-1)$ denoting the $d$-dimensional world volume coordinates. Replacing $z=\phi(x)$ it is obtained that $d s^{2}=d x^{\mu} g_{\mu \nu} d x^{\nu}=$ $d x^{\mu}\left(\eta_{\mu \nu}-\partial_{\mu} \phi \partial_{\nu} \phi\right) d x^{\nu}$ resulting in the $(d+1)$ dimensional space-time invariance of the NambuGoto action (1.6) with brane tension $\sigma$.

In order to implement the Carrollian contraction, Lagrange multiplier auxiliary fields $V_{\mu}(x)$ are introduced so that

$$
\begin{equation*}
\partial^{\mu} \phi=-\eta^{\mu \nu} V_{\nu} \frac{\tanh \sqrt{V^{2}}}{\sqrt{V^{2}}} \tag{1.7}
\end{equation*}
$$

with $V^{2}=V_{\mu} \eta^{\mu \nu} V_{\nu}$ and $x_{\mu}=\left(c t, x_{m}\right)$ while $\partial^{\mu}=\frac{\partial}{\partial x_{\mu}}$. The action becomes

$$
\begin{equation*}
\Gamma_{\mathrm{NG}}=-c \sigma \int d t d^{p} x \cosh \sqrt{V^{2}}\left[1+\left(V_{\mu} \frac{\tanh \sqrt{V^{2}}}{\sqrt{V^{2}}}\right) \partial^{\mu} \phi\right] . \tag{1.8}
\end{equation*}
$$

Making the speed of light explicit $V_{0}=2 c w$ and $V_{m}=2 v_{m}$ for $m=1,2, \ldots, p=(d-1)$, the Carrollian limit $c \rightarrow 0$ of the action $\Gamma_{C}=-(1 / c \sigma) \Gamma_{\mathrm{NG}}=\int d t d^{p} x \mathcal{L}_{C}$ is obtained

$$
\begin{equation*}
\Gamma_{C}=\int d t d^{p} x \mathcal{L}_{C}=\int d t d^{p} x \cos \sqrt{4 v^{2}}\left[1+\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\left(2 w \dot{\phi}+2 v_{m} \partial^{m} \phi\right)\right], \tag{1.9}
\end{equation*}
$$

where, after separation of space and time coordinates, the spatial metric is just $\delta^{m n}$. Only subscripts $x_{m}, v_{m}$ and superscript derivatives $\partial^{m}=\frac{\partial}{\partial x_{m}}$ will be used, with $\dot{\phi}=\frac{\partial}{\partial t} \phi$. In the case of a Goldstone field, as is $\phi$, the leading term in the derivative expansion of the action (thin wall limit) is uniquely fixed. The field equations are found to be

$$
\begin{align*}
& 0=\frac{\delta \Gamma_{C}}{\delta w(x, t)}= 2 \dot{\phi}\left[\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right], \\
& 0=\frac{\delta \Gamma_{C}}{\delta v_{m}(x, t)}= 2 \cos \sqrt{4 v^{2}}\left(\left[2 \delta^{n r} v_{r} \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}-\partial^{n} \phi\right] \times\right. \\
&\left.\times\left[\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} P_{T n s}(v)+P_{L n s}(v)\right] \delta^{s m}\right) \\
&+2 \dot{\phi} \frac{w v_{s} \delta s^{s m}}{v^{2}} \cos \sqrt{4 v^{2}}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}-1\right), \\
& 0=\frac{\delta \Gamma_{C}}{\delta \phi(x, t)}=-\frac{\partial}{\partial t}\left[2 w \frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right]-\frac{\partial}{\partial x_{m}}\left[2 v_{m} \frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right], \tag{1.10}
\end{align*}
$$

where the transverse $P_{T m n}(v)$ and longitudinal $P_{L m n}(v)$ projection matrices are defined as

$$
\begin{align*}
& P_{T m n}(v)=\delta_{m n}-\frac{v_{m} v_{n}}{v^{2}}, \\
& P_{L m n}(v)=\frac{v_{m} v_{n}}{v^{2}} . \tag{1.11}
\end{align*}
$$

Although informally obtained in the Introduction, these results also follow from the coset method of Section II for the brane embedded in $A d S-C$ space when the flat (Minkowski) space-time limit $m^{2} \rightarrow 0$ is taken and $A d S-C \xrightarrow{m^{2} \rightarrow 0} C$, Carroll spacetime.

The field equations reflect the Carroll spacetime symmetries yielding $\dot{\phi}=0$ from the $w$ equation of motion, $\delta \Gamma_{C} / \delta \omega=0$ so that the brane's initial spatial shape does not evolve as expected from the collapse of the light cone in this limit. Having set $\dot{\phi}=0$, the spatial components of $v_{m}$ obey the constraint $2 v_{m} \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}=\partial^{m} \phi$ as dictated by the $v_{m}$ field equation, $\delta \Gamma_{C} / \delta v_{m}=0$. Both field equations are consistent with the initial auxiliary velocity field equation (1.7) for $c \rightarrow 0$. Finally the broken space translation symmetry in the $(p+1)$ direction yields the time variation of the momentum density as given by the $\phi$ field equation $\delta \Gamma_{C} / \delta \phi=0$. The momentum density $\Pi$ is defined by

$$
\begin{equation*}
\Pi \equiv \partial \mathcal{L}_{C} / \partial \dot{\phi}=2 w \frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} \tag{1.12}
\end{equation*}
$$

while the derivatives of $\mathcal{L}_{C}$ with respect to the spatial derivatives of $\phi$ are denoted

$$
\begin{equation*}
\Pi_{m}=\partial \mathcal{L}_{C} / \partial \partial^{m} \phi=2 v_{m} \frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} \tag{1.13}
\end{equation*}
$$

Thus the $\phi$-field equation has the form of a current conservation equation. Indeed, the field equation is the spontaneously broken translation current conservation equation for the Carroll spacetime. The corresponding Noether current has the conserved form as above

$$
\begin{equation*}
\frac{\partial}{\partial t} \Pi+\frac{\partial}{\partial x_{m}} \Pi_{m}=0 . \tag{1.14}
\end{equation*}
$$

The action, Equation (1.9), is invariant under the Carroll transformations, obtained by contracting the Poincaré transformations as $c \rightarrow 0$, of the ( $\mathrm{d}+1$ )-dimensional Carroll spacetime which include the unbroken d-dimensional worldvolume time and space translations, space rotations and boosts with respective parameters $\epsilon, a_{m}, \alpha_{m n}, \beta_{m}$ and additionally, now non-linearly realized, the broken space translation of the covolume which is just a $\phi$ shift symmetry, boosts in that direction and rotations in a worldvolume-covolume plane with respective parameters $\zeta, \lambda, \kappa_{m}$. Exploiting the invariance of the Minkowski interval $d s^{2}=d x^{\mu} \eta_{\mu \nu} d x^{\nu}-d z^{2}$ under $(d+1)$-dimensional Poincaré transformations

$$
\begin{equation*}
x_{M}^{\prime}=x_{M}+\lambda_{M N} \eta^{N P} x_{P}+a_{M} \tag{1.15}
\end{equation*}
$$

where $x_{M}=\left(c t, x_{m}, z\right)$ with $z=\phi(x)$ and $\lambda_{M N}=-\lambda_{N M}$, the coordinate and field Poincaré transformations have the form

$$
\begin{align*}
t^{\prime} & =t-\frac{1}{c} \lambda_{0 n} x_{n}-\frac{1}{c} \lambda_{0 z} \phi+\frac{1}{c} a_{0}, \\
x_{m}^{\prime} & =x_{m}+c \lambda_{0 m} t-\lambda_{m n} x_{n}-\lambda_{m z} \phi+a_{m}, \\
\phi^{\prime}\left(x^{\prime}, t^{\prime}\right) & =\phi(x, t)+c \lambda_{z 0} t-\lambda_{z m} x_{m}+a_{z} . \tag{1.16}
\end{align*}
$$

Contracting the Poincaré symmetry transformations to those of the Carroll symmetries requires a rescaling of the time components of the transformation parameters so that

$$
\begin{equation*}
a_{0}=c \epsilon \quad, \quad \lambda_{0 m}=c \beta_{m} \quad, \quad \lambda_{0 z}=c 2 \lambda \tag{1.17}
\end{equation*}
$$

while the purely spatial components are unchanged and are denoted as

$$
\begin{equation*}
\zeta=a_{z} \quad, \quad \alpha_{m n}=-\lambda_{m n} \quad, \quad 2 \kappa_{m}=\lambda_{m z} \tag{1.18}
\end{equation*}
$$

The new parameters denote the Carroll transformation parameters. The $c \rightarrow 0$ Carroll spacetime transformations of the coordinates and field $\phi$ are thus obtained and have the non-linear form (see Appendix A for the coset method derivation and the $A d S-C$ to Carroll spacetime $C$ limit to obtain Equations (1.19) and (1.21))

$$
\begin{align*}
t^{\prime} & =t+\epsilon-2 \lambda \phi-\beta_{m} x_{m}, \\
x_{m}^{\prime} & =x_{m}+a_{m}+\alpha_{m n} x_{n}-2 \phi \kappa_{m}, \\
\phi^{\prime}\left(x^{\prime}, t^{\prime}\right) & =\phi(x, t)+\zeta+2 \kappa_{m} x_{m} . \tag{1.19}
\end{align*}
$$

Applying these transformations to the auxiliary Lagrange multiplier field definition, Equation (1.7), so that

$$
\begin{equation*}
\partial_{\mu}^{\prime} \phi^{\prime}\left(x^{\prime}, t^{\prime}\right)=-V_{\mu}^{\prime}\left(x^{\prime}, t^{\prime}\right) \frac{\tanh \sqrt{V^{\prime 2}}}{\sqrt{V^{\prime 2}}} \tag{1.20}
\end{equation*}
$$

with $V_{\mu}^{\prime}\left(x^{\prime}, t^{\prime}\right)=V_{\mu}(x, t)+\Delta V_{\mu}(x, t)$, where $\Delta V_{\mu}=\left(2 c \Delta w, 2 \Delta v_{m}\right)$ yields the $w$ and $v_{m}$ auxiliary fields' Carroll transformations

$$
\begin{align*}
w^{\prime}\left(x^{\prime}, t^{\prime}\right)=w(x, t)[1 & \left.+\frac{v_{m} \kappa_{m}}{v^{2}}\left(1-\sqrt{4 v^{2}} \cot \sqrt{4 v^{2}}\right)\right] \\
& -\beta_{m} v_{m}+\lambda \sqrt{4 v^{2}} \cot \sqrt{4 v^{2}}, \\
v_{m}^{\prime}\left(x^{\prime}, t^{\prime}\right)=v_{m}(x, t)+ & \alpha_{m n} v_{n}+\left(\sqrt{4 v^{2}} \cot \sqrt{4 v^{2}} P_{T m n}(v)+P_{L m n}(v)\right) \kappa_{n} \tag{1.21}
\end{align*}
$$

Since the time and space transformations involve functions thereof, the differential form of Equation (1.19) yields the general coordinate transformation $G=\partial\left(t^{\prime}, x^{\prime}\right) / \partial(t, x)$. That is, recombining $t$ and $x_{m}$ in the matrix $X_{M}=\left(t, x_{m}\right)$ where now $M, N=0,1, \ldots, p$, the transformations are given by

$$
\begin{align*}
d X_{M}^{\prime} & =\left(d t^{\prime}, d x_{m}^{\prime}\right) \\
& =d X_{N} G_{M}^{N}=\left(d t G_{0}^{0}+d x_{n} G_{0}^{n}, d t G_{m}^{0}+d x_{n} G_{m}^{n}\right) \tag{1.22}
\end{align*}
$$

with

$$
\begin{align*}
G_{M}^{N} & =\left(\begin{array}{cc}
\frac{\partial t^{\prime}}{\partial t} & \frac{\partial x_{m}^{\prime}}{\partial t} \\
\frac{\partial t^{\prime}}{\partial x_{n}} & \frac{\partial x_{m}^{\prime}}{\partial x_{n}}
\end{array}\right)_{M}^{N} \\
& =\left(\begin{array}{cc}
1-2 \lambda \dot{\phi} & -2 \dot{\phi} \kappa_{m} \\
\left(-\beta_{n}-2 \lambda \partial^{n} \phi\right) & \left(\delta_{n m}+\alpha_{n m}-2 \partial^{n} \phi \kappa_{m}\right)
\end{array}\right)_{M}^{N} . \tag{1.23}
\end{align*}
$$

The spacetime transformation Jacobian is $d t^{\prime} d^{p} x^{\prime}=d t d^{p} x \operatorname{det} G$. On the other hand, the action $\Gamma_{C}$ is invariant under the Carrollian symmetry transformations, thus

$$
\Gamma_{C}^{\prime}=\int d t^{\prime} d^{p} x^{\prime} \mathcal{L}_{C}^{\prime}\left(x^{\prime}, t^{\prime}\right)=\int d t d^{p} x \operatorname{det} G \mathcal{L}_{C}^{\prime}\left(x^{\prime}, t^{\prime}\right)
$$

TABLE I. Carroll spacetime transformations and Noether currents.

| Transformation | Parameter | Noether currents |
| :---: | :---: | :---: |
| Time translations | $\epsilon$ | $\begin{gathered} \mathcal{H}=\Pi \dot{\phi}-\mathcal{L}_{C} \\ =\cos \sqrt{4 v^{2}}+\Pi_{m} \partial^{m} \phi \\ h_{m}=\Pi_{m} \dot{\phi} \end{gathered}$ |
| Space translations | $a_{n}$ | $\begin{gathered} T^{n}=\Pi \partial^{n} \phi \\ T_{m}^{n}=\frac{\partial \mathcal{L}_{C}}{\partial \partial^{m} \phi} \partial^{n} \phi-\delta_{m}^{n} \mathcal{L}_{C} \\ =\delta_{m}^{n}\left[\cos \sqrt{4 v^{2}}+\Pi \dot{\phi}+\Pi_{r} \partial^{r} \phi\right]-\Pi_{m} \partial^{n} \phi \end{gathered}$ |
| Broken space translations | $\zeta$ | $\begin{aligned} z & =\Pi \\ z_{m} & =\Pi_{m} \end{aligned}$ |
| Broken boosts | $\lambda$ | $\begin{gathered} l=\phi \mathcal{H} \\ l_{m}=\phi h_{m} \end{gathered}$ |
| Broken rotations | $\kappa_{n}$ | $\begin{gathered} K^{n}=2 \phi T^{n}+2 \delta^{n r} x_{r} \Pi \\ K_{m}^{n}=2 \phi T_{m}^{n}+2 \Pi_{m} \delta^{n r} x_{r} \end{gathered}$ |
| Unbroken rotations | $\alpha_{r s}$ | $\begin{aligned} \mathcal{M}^{r s} & =x_{r} T^{s}-x_{s} T^{r} \\ \mathcal{M}_{m}^{r s} & =x_{r} T_{m}^{s}-x_{s} T_{m}^{r} \end{aligned}$ |
| Unbroken boosts | $\beta_{n}$ | $\begin{aligned} B^{n} & =\delta^{n r} x_{r} \mathcal{H} \\ B_{m}^{n} & =\delta^{n r} x_{r} h_{m} \end{aligned}$ |

$$
\begin{equation*}
=\int d t d^{p} x \mathcal{L}_{C}(x, t)=\Gamma_{C} \tag{1.24}
\end{equation*}
$$

so that

$$
\begin{equation*}
\mathcal{L}_{C}^{\prime}\left(x^{\prime}, t^{\prime}\right)=\operatorname{det} G^{-1} \mathcal{L}_{C}(x, t) \tag{1.25}
\end{equation*}
$$

For these Carrollian transformations, the Noether currents take the couplet form of time and spatial component currents

$$
\begin{equation*}
J_{M}=\binom{\Pi \delta \phi+\delta t \mathcal{L}_{C}}{\Pi_{m} \delta \phi+\delta x_{m} \mathcal{L}_{C}}_{M=(0, m)} \tag{1.26}
\end{equation*}
$$

with $\delta t=t^{\prime}-t$ and $\delta x_{m}=x_{m}^{\prime}-x_{m}$, where the intrinsic transformation $\delta \phi$ is defined as

$$
\begin{equation*}
\delta \phi \equiv \phi^{\prime}(x, t)-\phi(x, t)=\Delta \phi(x, t)-\delta t \dot{\phi}-\delta x_{m} \partial^{m} \phi \tag{1.27}
\end{equation*}
$$

with the total variation given by

$$
\begin{equation*}
\Delta \phi(x, t) \equiv \phi^{\prime}\left(x^{\prime}, t^{\prime}\right)-\phi(x, t) \tag{1.28}
\end{equation*}
$$

Thus Noether's theorem is $\left(\varphi_{i}=\left\{\phi, w, v_{m}\right\}\right)$

$$
\begin{align*}
\Delta \mathcal{L}_{C} & =\mathcal{L}_{C}^{\prime}\left(x^{\prime}, t^{\prime}\right)-\mathcal{L}_{C}(x, t) \\
& =\frac{\partial}{\partial t} J_{0}+\frac{\partial}{\partial x_{m}} J_{m}-\mathcal{L}_{C} \frac{\partial \delta t}{\partial t}-\mathcal{L}_{C} \frac{\partial \delta x_{m}}{\partial x_{m}}+\frac{\delta \Gamma_{C}}{\delta \varphi_{i}} \delta \varphi_{i}, \tag{1.29}
\end{align*}
$$

where the last term vanishes by the field equations $\delta \Gamma_{C} / \delta \varphi_{i}=0$. The conserved currents (before use of the field equation constraints) are given by the pairs displayed in Table I. The action is invariant as reflected by the vanishing or explicit cancellation of the $\Delta \mathcal{L}_{C}+\mathcal{L}_{C}\left(\frac{\partial \delta t}{\partial t}+\frac{\partial \delta x_{m}}{\partial x_{m}}\right)=0$ terms in Noether's theorem for each of the Carrollian symmetry transformations. Hence $\partial^{M} J_{M}=$ $\frac{\partial}{\partial t} J_{0}+\frac{\partial}{\partial x_{m}} J_{m}=-\frac{\delta \Gamma_{C}}{\delta \varphi_{i}} \delta \varphi_{i}$. The associated conserved charges are given by $Q=\int d^{p} x J_{0}$, where $\dot{Q}=-\int_{S \rightarrow \infty} \vec{J} \cdot d \vec{S} \longrightarrow 0$. Once again the time evolution of the momentum density is contained in the broken translation current conservation equation and the invariance of the Lagrangian

$$
\begin{equation*}
0=\dot{\Pi}+\partial^{m} \Pi_{m}=-\frac{\delta \Gamma_{C}}{\delta \phi} \tag{1.30}
\end{equation*}
$$

In summary, the $w$-field equation of motion, Equation (1.10), yields the frozen spatial distribution of the domain wall as expected from the collapse of the lightcone to the positive time half-line
in the Carrollian $c \rightarrow 0$ limit: $\frac{\partial}{\partial t} \phi(x, t)=0$. Along with this the Lagrange multiplier $v_{m}$-field equation of motion simply reproduces the constraint of the "inverse Higgs mechanism" ${ }^{13}$

$$
\begin{equation*}
2 v_{m} \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}=\partial^{m} \phi . \tag{1.31}
\end{equation*}
$$

These field equation constraints also follow directly from the Lagrange multiplier equation (1.7) in the $c \rightarrow 0$ limit. Finally, although the brane is stationary, the momentum must vary in time in order to balance the tension due to the domain wall's local spatial shape where using the Lagrange multiplier constraints so that $\dot{\phi}=0$ and $2 v_{m} \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}=\partial^{m} \phi$, it is found that

$$
\begin{equation*}
\dot{\Pi}=-\partial^{m}\left[\frac{\partial^{m} \phi}{\sqrt{1+\partial^{s} \phi \partial^{s} \phi}}\right] \tag{1.32}
\end{equation*}
$$

The purpose of this paper is to determine the Carrollian limit for branes in $\operatorname{AdS}$ spacetime. ${ }^{6,7}$ The $D=d+1$ dimensional $A d S$ spacetime symmetry algebra is contracted in the Carroll limit, $c \rightarrow 0$. For a p-brane action in the alternate string Carrollian limit of Minkowski space see Ref. 8. Application of the Carrollian limit to gravity and electromagnetism is discussed in Refs. 9 and 10.

In Section II coset methods ${ }^{11-13}$ are applied to the $A d S-C$ algebra for the case of an embedded co-dimension one p-brane (domain wall). The induced vielbeine, covariant derivatives, and spin connections are determined using the Maurer-Cartan one-form associated with the p-brane coset element. The action is constructed and shown to be invariant under the non-linearly realized $A d S-C_{d+1}$ broken to $A d S-C_{d}$ symmetries by the brane embedding. The symmetry transformations are detailed in Appendix A. Alternatively, the $A d S-C$ action can be obtained by making the speed of light $c$ dependence explicit in the $A d S_{d+1} \rightarrow A d S_{d}$ case and taking the $c \rightarrow 0$ limit. Using the results of Ref. 6, this approach is demonstrated in Appendix B.

From the action the field equations are determined. As expected due to the collapse of the forward light cone to the positive time half-line, the spatial shape of the brane is stationary. However, the spatial shape of the brane as well as the $A d S-C$ geometry requires its conjugate momentum density to be time dependent. Finally Noether's theorem is applied to the broken space translation symmetry in order to calculate the current and its conservation equation. Section III presents the action in terms of a product of the background $A d S-C_{d}$ world volume vielbein and the Nambu-Goto-Carrollian vielbein. This is then used to express the action in terms of its dual vector theory. The results of the brane embedding are reviewed in Section IV.

## II. AdS-CARROLL SPACE AND THE COSET METHOD

The AdS-Carroll spacetime is defined by the Wigner-Inönü contraction of the AdS symmetry algebra for the speed of light vanishing, $c \rightarrow 0$. The isometry group of the $D$-dimensional $\operatorname{AdS}$ space is given by the $S O(2, D-1)$ algebra of symmetry generators with the commutation relations

$$
\begin{align*}
{\left[M^{M N}, M^{R S}\right] } & =-i\left(\eta^{M R} M^{N S}-\eta^{M S} M^{N R}+\eta^{N S} M^{M R}-\eta^{N R} M^{M S}\right), \\
{\left[M^{M N}, P^{L}\right] } & =i\left(P^{M} \eta^{N L}-P^{N} \eta^{M L}\right), \\
{\left[P^{M}, P^{N}\right] } & =-i m^{2} M^{M N}, \tag{2.1}
\end{align*}
$$

where $L, M, N, R, S=0,1,2, \ldots, D-1$ and the metric $\eta_{M N}=(+1,-1,-1, \ldots,-1)$ with $m^{2}=1 / R^{2}$ and $R$ the curvature of the AdS hyperboloid.

Introducing the explicit factors of the speed of light for the time related components, the time component involved generators $H$ and $B^{A}$ are defined as

$$
\begin{align*}
P^{0} & =\frac{1}{c} H \\
M^{A 0} & =\frac{1}{c} B^{A} \tag{2.2}
\end{align*}
$$

while the spatial components remain unscaled $P^{A}, M^{A B}$,

$$
\begin{align*}
P^{A} & =P^{A}, \\
M^{A B} & =M^{A B} \tag{2.3}
\end{align*}
$$

for $A, B, C, D=1,2, \ldots,(D-1)$ denoting the spatial indices. The $S O(2, D-1)$ algebra contracts to the AdS-Carroll algebra in the $c \rightarrow 0$ limit

$$
\begin{align*}
{\left[M^{A B}, M^{C D}\right] } & =+i\left(\delta^{A C} M^{B D}-\delta^{A D} M^{B C}+\delta^{B D} M^{A C}-\delta^{B C} M^{A D}\right), \\
{\left[M^{A B}, B^{C}\right] } & =-i\left(B^{A} \delta^{B C}-B^{B} \delta^{A C}\right), \\
{\left[M^{A B}, P^{C}\right] } & =-i\left(P^{A} \delta^{B C}-P^{B} \delta^{A C}\right), \\
{\left[B^{A}, P^{B}\right] } & =+i \delta^{A B} H, \\
{\left[H, P^{A}\right] } & =+i m^{2} B^{A}, \\
{\left[P^{A}, P^{B}\right] } & =-i m^{2} M^{A B}, \tag{2.4}
\end{align*}
$$

with remaining commutators vanishing.
A brane embedded in this AdS-Carroll spacetime will break its $A d S-C_{D}$ symmetries down to those of the $d$-dimensional worldvolume $A d S-C_{d}$ and its complementary covolume with the remaining symmetries being spontaneously broken. In the case considered here, the insertion of a domain wall results in a $d=(1+p)=(D-1)$-dimensional worldvolume and 1 -dimensional covolume. Choosing the $(p+1)^{\text {th }}$ spatial direction as the broken translation symmetry direction, the AdS-Carroll algebra can be expressed in terms of broken and unbroken generators with the generators $H, P^{m}, M^{m n}, B^{m}$, with $m, n=1,2, \ldots, p$, as unbroken generators and $P^{p+1} \equiv Z, M^{p+1, m} \equiv$ $\frac{1}{2} K^{m}, B^{p+1} \equiv \frac{1}{2} L$ as the broken generators. $M^{m n}$ are the $S O(p)$ worldvolume spatial rotation generators while the worldvolume spatial translation generators $P^{m}$ form an $\operatorname{SO}(p)$ vector with time translations generated by $H$. The $S O(p)$ vector $B^{m}$ generates worldvolume boosts in the $m$-direction. The translations in the covolume spatial direction are generated by $Z$ while boosts in that direction are generated by $L$. Finally broken rotations in the covolume- $m$ worldvolume plane are generated by the $S O(p)$ vector $K^{m}$. Consequently the $A d S-C_{D=d+1}$ algebra can be expressed in terms of these worldvolume and domain wall charges. The $A d S-C_{d=p+1}$ worldvolume isometries are given by the $H, P^{m}, M^{m n}, B^{m}$ algebra (only nontrivial commutators listed)

$$
\begin{align*}
{\left[M^{m n}, M^{r s}\right] } & =+i\left(\delta^{m r} M^{n s}-\delta^{m s} M^{n r}+\delta^{n s} M^{m r}-\delta^{n r} M^{m s}\right), \\
{\left[M^{m n}, B^{l}\right] } & =-i\left(B^{m} \delta^{n l}-B^{n} \delta^{m l}\right), \\
{\left[M^{m n}, P^{l}\right] } & =-i\left(P^{m} \delta^{n l}-P^{n} \delta^{m l}\right), \\
{\left[B^{m}, P^{n}\right] } & =+i \delta^{m n} H, \\
{\left[H, P^{m}\right] } & =+i m^{2} B^{m}, \\
{\left[P^{m}, P^{n}\right] } & =-i m^{2} M^{m n} . \tag{2.5}
\end{align*}
$$

The broken symmetry generators $Z, L, K^{m}$ commute with the unbroken generators above according to their unbroken subgroup representation

$$
\begin{aligned}
{\left[M^{m n}, Z\right] } & =0, \\
{\left[M^{m n}, L\right] } & =0, \\
{\left[M^{m n}, K^{l}\right] } & =-i\left(K^{m} \delta^{n l}-K^{n} \delta^{m l}\right), \\
{\left[B^{m}, Z\right] } & =0, \\
{\left[B^{m}, L\right] } & =0 \\
{\left[B^{m}, K^{n}\right] } & =+i \delta^{m n} L, \\
{\left[P^{m}, Z\right] } & =\frac{i}{2} m^{2} K^{m}, \\
{\left[P^{m}, L\right] } & =0, \\
{\left[P^{m}, K^{n}\right] } & =+2 i \delta^{m n} Z,
\end{aligned}
$$

$$
\begin{align*}
{[H, Z] } & =\frac{i}{2} m^{2} L, \\
{[H, L] } & =0, \\
{\left[H, K^{n}\right] } & =0 . \tag{2.6}
\end{align*}
$$

Finally the broken charges $Z, L, K^{m}$ commute amongst themselves to yield the charges of the unbroken subalgebra

$$
\begin{align*}
{[Z, L] } & =-2 i H, \\
{\left[Z, K^{m}\right] } & =-2 i P^{m} \\
{\left[L, K^{m}\right] } & =-4 i B^{m} \\
{\left[K^{m}, K^{n}\right] } & =4 i M^{m n} \tag{2.7}
\end{align*}
$$

The domain wall spontaneously breaks the $A d S-C_{D}$ spacetime symmetries down to those of the $A d S-C_{d}$ worldvolume. As a derivative expansion, the leading form of the brane action is uniquely determined. The Goldstone boson fields $\phi(x, t)$ corresponding to the long wavelength oscillations of the domain wall parameterize the coset coordinates along with the fields associated with the broken boost and rotations, $w(x, t)$ and $v_{m}(x, t)$, respectively. The geometry of the underlying $A d S-C_{d}$ worldvolume spacetime is described by the time $t$ and space $x_{m}$, $m=1,2, \ldots, p=(d-1)$, coordinate group elements. Overall these fields and spacetime coordinates parameterize the $\operatorname{AdS}-C_{D} / I S O(p)$ coset element $\Omega$ (note generators are defined with superscripts)

$$
\begin{equation*}
\Omega \equiv e^{i t H+i x_{m} P^{m}} e^{i \phi(x, t) Z} e^{i \omega(x, t) L+i v_{m}(x, t) K^{m}} \tag{2.8}
\end{equation*}
$$

where $\operatorname{ISO}(\mathrm{p})$ is the unbroken subgroup with generators $M^{m n}$ and $B^{m}$. The background worldvolume coset $\bar{\Omega} \in A d S-C_{d} / I S O(p)$

$$
\begin{equation*}
\bar{\Omega} \equiv e^{i t H+i x_{m} P^{m}} \tag{2.9}
\end{equation*}
$$

is used to determine the $A d S-C_{d}$ background vielbeine and spin connections via the MaurerCartan 1-form $\tilde{\bar{\omega}}$

$$
\begin{equation*}
\tilde{\tilde{\omega}} \equiv-i \bar{\Omega}^{-1} d \bar{\Omega}=\bar{\omega}_{H} H+\bar{\omega}_{P a} P^{a}+\frac{1}{2} \bar{\omega}_{a b} M^{a b}+\bar{\omega}_{B a} B^{a} . \tag{2.10}
\end{equation*}
$$

Expanding the 1 -forms in terms of the coordinate differentials, the Maurer-Cartan 1-form becomes (with tangent space indices denoted $a, b=1,2, \ldots, p$ and world volume indices denoted $m, n=1,2, \ldots, p$ )

$$
\begin{align*}
\tilde{\bar{\omega}}=\left(d t \bar{e}_{0}^{0}\right. & \left.+d x_{m} \bar{e}_{0}^{m}\right) H+\left(d t \bar{e}_{a}^{0}+d x_{n} \bar{e}_{a}^{n}\right) P^{a} \\
& +\frac{1}{2}\left(d t \bar{\omega}_{a b}^{t}+d x_{r} \bar{\omega}_{a b}^{r}\right) M^{a b}+\left(d t \bar{\omega}_{0 a}^{t}+d x_{r} \bar{\omega}_{0 a}^{r}\right) B^{a}, \tag{2.11}
\end{align*}
$$

where the background vielbeine are found to be

$$
\begin{align*}
& \bar{e}_{0}^{0}=\frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}, \\
& \bar{e}_{0}^{m}=\frac{x_{m} t}{x^{2}}\left(1-\frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\right), \\
& \bar{e}_{a}^{n}=\left(\frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\right) P_{\text {Tna }}(x)+P_{\text {Lna }}(x), \\
& \bar{e}_{a}^{0}=0, \tag{2.12}
\end{align*}
$$

with $x^{2}=x_{m} \delta^{m n} x_{n}=x_{m} x_{m}$. The background spin connections are also obtained as

$$
\begin{aligned}
& \bar{\omega}_{a b}^{t}=0, \\
& \bar{\omega}_{a b}^{r}=\left(\frac{1-\cosh \sqrt{m^{2} x^{2}}}{x^{2}}\right)\left(\delta_{a}^{s} \delta_{b}^{r}-\delta_{b}^{s} \delta_{a}^{r}\right) x_{s},
\end{aligned}
$$

$$
\begin{align*}
& \bar{\omega}_{0 a}^{t}=\left(\frac{1-\cosh \sqrt{m^{2} x^{2}}}{x^{2}}\right) x_{a}, \\
& \bar{\omega}_{0 a}^{r}=-\left(\frac{1-\cosh \sqrt{m^{2} x^{2}}}{x^{2}}\right) t \delta_{a}^{r} . \tag{2.13}
\end{align*}
$$

The $A d S-C_{d}$ background vielbein $\bar{E}_{A}^{M}$ is defined as the matrix relating the coordinate differentials $d X_{N}=\left(d t, d x_{n}\right)$, with $M, N=0,1, \ldots, p$, to the covariant coordinate differentials $\tilde{\tilde{\omega}}_{A}=$ $\left(\bar{\omega}_{H}, \bar{\omega}_{P a}\right)$, with $A, B=0,1, \ldots, p$ as well as $\tilde{\bar{\omega}}_{0}=\bar{\omega}_{H}$ and $\tilde{\bar{\omega}}_{a}=\bar{\omega}_{P a}$, thus

$$
\begin{align*}
\tilde{\bar{\omega}}_{A} & =\left(\begin{array}{ll}
\bar{\omega}_{H} & \bar{\omega}_{P a}
\end{array}\right)=d X_{M} \bar{E}_{A}^{M} \\
& =\left(\begin{array}{ll}
d t & d x_{m}
\end{array}\right)\left(\begin{array}{cc}
\bar{e}_{0}^{0} & 0 \\
\bar{e}_{0}^{m} & \bar{e}_{a}^{m}
\end{array}\right) \tag{2.14}
\end{align*}
$$

that is

$$
\bar{E}_{A}^{M}=\left(\begin{array}{cc}
\bar{e}_{0}^{0} & \bar{e}_{a}^{0}=0  \tag{2.15}\\
\bar{e}_{0}^{m} & \bar{e}_{a}^{m}
\end{array}\right)=\left(\begin{array}{cc}
\bar{E}_{0}^{0} & \bar{E}_{a}^{0}=0 \\
\bar{E}_{0}^{m} & \bar{E}_{a}^{m}
\end{array}\right),
$$

with $\operatorname{det} \bar{E}=\bar{e}_{0}^{0} \operatorname{det} \bar{e}_{a}^{m}$, where $a, b, m, n=1,2, \ldots, p$.
On the other hand, the Maurer-Cartan 1-form for the domain wall breakdown of $A d S-C_{D} \rightarrow$ $A d S-C_{d}$ can be constructed using the coset element $\Omega$

$$
\begin{align*}
\tilde{\omega} & =-i \Omega^{-1} d \Omega \\
& =\omega_{H} H+\omega_{P a} P^{a}+\omega_{Z} Z+\omega_{L} L+\omega_{K a} K^{a}+\frac{1}{2} \omega_{M a b} M^{a b}+\omega_{B a} B^{a} \tag{2.16}
\end{align*}
$$

This yields the vielbeine, covariant derivatives of the fields, and spin connections. The vielbeine are given in terms of the coordinate differentials according to

$$
\begin{align*}
\omega_{H} & =d t e_{0}^{0}+d x_{m} e_{0}^{m} \\
\omega_{P a} & =d t e_{a}^{0}+d x_{m} e_{a}^{m} \tag{2.17}
\end{align*}
$$

with $v^{2}=v_{a} \delta^{a b} v_{b}=v_{a} v_{a}$ and

$$
\begin{align*}
& e_{0}^{0}=\bar{e}_{0}^{0} \cosh \sqrt{m^{2} \phi^{2}}+2 w \dot{\phi}\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right), \\
& e_{0}^{m}=\bar{e}_{0}^{m} \cosh \sqrt{m^{2} \phi^{2}}+2 w \partial^{m} \phi\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right) \\
& +\cosh \sqrt{m^{2} \phi^{2}}\left(\frac{\cos \sqrt{4 v^{2}}-1}{v^{2}}\right) w v_{a} \bar{e}_{a}^{m}, \\
& e_{a}^{0}=2 \dot{\phi} v_{a}\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right), \\
& e_{a}^{m}=2 \partial^{m} \phi v_{a}\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right) \\
& +\cosh \sqrt{m^{2} \phi^{2}}\left[P_{T a b}(v)+\left(\cos \sqrt{4 v^{2}}\right) P_{L a b}(v)\right] \bar{e}_{b}^{m} . \tag{2.18}
\end{align*}
$$

The $A d S-C_{D}$ spacetime vielbein $E_{A}^{M}$ is defined as the matrix relating the coordinate differentials $d X_{M}=\left(d t, d x_{m}\right)$ to the covariant coordinate differentials $\tilde{\omega}_{A}=\left(\omega_{H}, \omega_{P a}\right)$, that is $\tilde{\omega}_{0}=\omega_{H}$ and $\tilde{\omega}_{a}=\omega_{P a}$, thus

$$
\begin{align*}
\tilde{\omega}_{A} & =\left(\begin{array}{ll}
\omega_{H} & \omega_{P a}
\end{array}\right)=d X_{M} E_{A}^{M} \\
& =\left(\begin{array}{ll}
d t & d x_{m}
\end{array}\right)\left(\begin{array}{ll}
e_{0}^{0} & e_{a}^{0} \\
e_{0}^{m} & e_{a}^{m}
\end{array}\right) \tag{2.19}
\end{align*}
$$

that is

$$
E_{A}^{M}=\left(\begin{array}{ll}
E_{0}^{0} & E_{a}^{0}  \tag{2.20}\\
E_{0}^{m} & E_{a}^{m}
\end{array}\right)=\left(\begin{array}{cc}
e_{0}^{0} & e_{a}^{0} \\
e_{0}^{m} & e_{a}^{m}
\end{array}\right) .
$$

The brane field's covariant derivatives, $\nabla^{t} \phi$ and $\nabla^{m} \phi$, are given by the $\omega_{Z}$ one-form

$$
\begin{equation*}
\omega_{Z}=d t \nabla^{t} \phi+d x_{m} \nabla^{m} \phi, \tag{2.21}
\end{equation*}
$$

where

$$
\begin{align*}
& \nabla^{t} \phi=\dot{\phi} \cos \sqrt{4 v^{2}}, \\
& \nabla^{m} \phi=\bar{e}_{a}^{m} \cosh \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left(\frac{\bar{e}_{n}^{-1 a} \frac{\partial}{\partial x_{n}} \phi}{\cosh \sqrt{m^{2} \phi^{2}}}-2 \delta^{a b} v_{b}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)\right) . \tag{2.22}
\end{align*}
$$

Likewise the auxiliary fields $w$ and $v_{a}$ have covariant derivatives determined by $\omega_{L}$ and $\omega_{K a}$,

$$
\begin{align*}
\omega_{L} & =d t \nabla^{t} w+d x_{m} \nabla^{m} w, \\
\omega_{K a} & =d t \nabla^{t} v_{a}+d x_{m} \nabla^{m} v_{a}, \tag{2.23}
\end{align*}
$$

where the derivatives are found to be

$$
\begin{align*}
& \nabla^{t} w=\dot{w}+\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}-1\right)\left(\dot{w}-w \frac{v_{a} \dot{v}_{a}}{v^{2}}\right)-\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} \bar{\omega}_{0 a}^{t} v_{a} \\
&-\frac{1}{2} m^{2} \phi \frac{\sinh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \cos \sqrt{4 v^{2}} \bar{e}_{0}^{0}, \\
& \nabla^{m} w= \partial^{m} w+\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}-1\right)\left(\partial^{m} w-w \frac{v_{a} \partial^{m} v_{a}}{v^{2}}\right)-\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} \bar{\omega}_{0 a}^{m} v_{a} \\
&-\frac{1}{2} m^{2} \phi \frac{\sinh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}}\left[\cos \sqrt{4 v^{2}} \bar{e}_{0}^{m}+\frac{\left(1-\cos \sqrt{4 v^{2}}\right)}{v^{2}} w v_{a} \bar{e}_{a}^{m}\right], \\
& \begin{aligned}
\nabla^{t} v_{a}= & \dot{v}_{a}+\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}-1\right) P_{\text {Tab }}(v) \dot{v}_{b}+\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} \bar{\omega}_{a b}^{t} v_{b}, \\
\nabla^{m} v_{a}= & \partial^{m} v_{a}
\end{aligned}+\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}-1\right) P_{T a b}(v) \partial^{m} v_{b}+\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} \bar{\omega}_{a b}^{s} v_{b} \\
&-\frac{1}{2} m^{2} \phi \frac{\sinh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}}\left[\cos \sqrt{4 v^{2}} P_{T a b}(v)+P_{L a b}(v)\right] \bar{e}_{b}^{m} .
\end{align*}
$$

Finally the spin connections are obtained from $\omega_{M}$ and $\omega_{B}$

$$
\begin{align*}
\omega_{M a b}= & d t \omega_{a b}^{t}+d x_{m} \omega_{a b}^{m} \\
= & \bar{\omega}_{a b}+m^{2} \phi \frac{\sinh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} \\
& +\left(1-\cos \sqrt{4 v^{2}}\right)\left[P_{L a c}(v) \bar{\omega}_{b c}-P_{L b c}(v) \bar{\omega}_{a c}-\left(\frac{d v_{a} v_{b}-d v_{b} v_{a}}{v^{2}}\right)\right], \\
\omega_{B a}= & d t \omega_{a}^{t}+d x_{m} \omega_{a}^{m} \\
= & \bar{\omega}_{B a}+\left(\frac{1-\cos \sqrt{4 v^{2}}}{v^{2}}\right)\left[d w v_{a}-w d v_{a}-\bar{\omega}_{a b} w v_{b}-\bar{\omega}_{B b} v_{b} v_{a}\right] \\
& +2 m^{2} \phi \frac{\sinh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\left[w \bar{e}_{a}^{m} d x_{m}-v_{a}{ }_{0}^{0} d t-v_{a} d x_{m} \bar{e}_{0}^{m}\right] . \tag{2.25}
\end{align*}
$$

The $A d S-C_{D}$ invariant action is constructed in terms of the vielbein $E$

$$
\begin{equation*}
\Gamma_{A d S-C_{D}}=\int d t d^{p} x \mathcal{L}_{A d S-C_{D}}(x, t)=\int d t d^{p} x \operatorname{det} E . \tag{2.26}
\end{equation*}
$$

The $A d S-C_{D}$ transformations are non-linearly realized according to the group multiplication properties involving the coset $\Omega$ as detailed in Appendix A. The invariance of the action follows from the transformation properties of the vielbein $E$. The Maurer-Cartan one-forms transform according to which representation of the local $\operatorname{ISO}(p)$ tangent space transformations that the associated operator belongs given by the unbroken subgroup element $h(x, t)$ obtained in Appendix A. Using the coset transformation law $g \Omega(x, t)=\Omega^{\prime}\left(x^{\prime}, t^{\prime}\right) h(x, t)$, the one-forms transform as $\omega^{\prime}\left(x^{\prime}, t^{\prime}\right)=h(x, t) \omega(x, t) h^{-1}(x, t)-i h(x, t) d h^{-1}(x, t)$, yielding

$$
\begin{align*}
\omega^{\prime}= & \omega_{H}^{\prime} H+\omega_{P a}^{\prime} P^{a}+\omega_{Z}^{\prime} Z+\omega_{L}^{\prime} L+\omega_{K a}^{\prime} K^{a}+\frac{1}{2} \omega_{M a b}^{\prime} M^{a b}+\omega_{B a}^{\prime} B^{a} \\
= & h \omega h^{-1}-i h d h^{-1} \\
= & \omega_{H} h H h^{-1}+\omega_{P a} h P^{a} h^{-1}+\omega_{Z} h Z h^{-1}+\omega_{L} h L h^{-1}+\omega_{K a} h K^{a} h^{-1} \\
& +\frac{1}{2} \omega_{M a b} h M^{a b} h^{-1}+\omega_{B a} h B^{a} h^{-1}-\frac{1}{2} d \theta_{a b} M^{a b}-d \theta_{a} B^{a} . \tag{2.27}
\end{align*}
$$

Hence the one-forms' variations are obtained

$$
\begin{align*}
\omega_{H}^{\prime} & =\omega_{H}-\omega_{P a} \theta_{a}, \\
\omega_{P a}^{\prime} & =\omega_{P b}\left(\delta_{b a}-\theta_{b a}\right) \equiv \omega_{P b} R_{b a}^{-1}, \\
\omega_{Z}^{\prime} & =\omega_{Z}, \\
\omega_{L}^{\prime} & =\omega_{L}-\omega_{K a} \theta_{a}, \\
\omega_{K a}^{\prime} & =\omega_{K b} R_{b a}^{-1}, \\
\omega_{M c d}^{\prime} & =\omega_{M a b} R_{a c}^{-1} R_{d b}^{-1}-d \theta_{c d}, \\
\omega_{B a}^{\prime} & =\omega_{B b} R_{b a}^{-1}-d \theta_{a}+\frac{1}{2} \omega_{M c d}\left(\theta_{c} \delta_{d a}-\theta_{d} \delta_{c a}\right) . \tag{2.28}
\end{align*}
$$

The covariant coordinate differentials and vielbeine transform as

$$
\left.\begin{array}{rl}
\tilde{\omega}_{A}^{\prime} & =\left(\begin{array}{ll}
\omega_{H}^{\prime} & \omega_{P a}^{\prime}
\end{array}\right)=\left(\begin{array}{ll}
\omega_{H} & \omega_{P b}
\end{array}\right)\left(\begin{array}{cc}
1 & 0 \\
-\theta_{b} & R_{b a}^{-1}
\end{array}\right) \\
& =d X_{M}^{\prime} E_{A}^{\prime M}=\left(d t^{\prime}\right. \\
d x_{m}^{\prime}
\end{array}\right)\left(\begin{array}{cc}
e^{\prime 0} & e^{\prime \prime}  \tag{2.29}\\
e_{a}^{\prime m} \\
e_{0} & e^{\prime m}
\end{array}\right) .
$$

with letters from the beginning of the alphabet denoting the tangent space transformation properties. From Appendix A the coordinate differentials transform according to the general coordinate transformation

$$
d X_{M}^{\prime}=\left(d t^{\prime} d x_{m}^{\prime}\right)=\left(\begin{array}{ll}
d t & d x_{n}
\end{array}\right)\left(\begin{array}{ll}
G_{0}^{0} & G_{m}^{0}  \tag{2.30}\\
G_{0}^{n} & G_{m}^{n}
\end{array}\right)=d X_{N} G_{M}^{N}
$$

where the complicated general coordinate transformation matrix is denoted by $G^{N}{ }_{M}$, with letters in the middle of the alphabet indicating world volume coordinate transformations. The spacetime differentials have the Jacobian $d t^{\prime} d^{p} x^{\prime}=d t d^{p} x \operatorname{det} G$. Thus the vielbein $E_{A}^{M}$ transforms as

$$
\begin{equation*}
E_{A}^{\prime M}=G^{-1 M}{ }_{N} E_{B}^{N} \Lambda_{A}^{B}, \tag{2.31}
\end{equation*}
$$

where the tangent space transformations have been denoted by

$$
\Lambda_{A}^{B}=\left(\begin{array}{cc}
1 & 0  \tag{2.32}\\
-\theta_{b} & R_{b a}^{-1}
\end{array}\right) .
$$

Noting that $\operatorname{det} \Lambda=1$ so that $\operatorname{det} E^{\prime}=\operatorname{det} G^{-1} \operatorname{det} E$, the action is invariant

$$
\begin{align*}
\Gamma_{A d S-C_{D}}^{\prime} & =\int d t^{\prime} d^{p} x^{\prime} \operatorname{det} E^{\prime}=\int d t d^{p} x \operatorname{det} G \operatorname{det} G^{-1} \operatorname{det} E \\
& =\int d t d^{p} x \operatorname{det} E=\Gamma_{A d S-C_{D}} \tag{2.33}
\end{align*}
$$

With the vielbeine in Equation (2.18) the $A d S-C_{D}$ invariant action is found

$$
\begin{equation*}
\Gamma_{A d S-C_{D}}=\int d t d^{p} x \operatorname{det} E=\int d t d^{p} x \operatorname{det}\left(e_{a}^{m}\right)\left[e_{0}^{0}-e_{a}^{0} e^{-1 a} e_{0}^{n}\right] \tag{2.34}
\end{equation*}
$$

with (noting that $\bar{e}^{-1 a}{ }_{m} \bar{e}_{0}^{m}=\delta^{a}{ }_{m} \bar{e}_{0}^{m}=\bar{e}_{0}^{a}$ )

$$
\begin{align*}
& \operatorname{det} E=\operatorname{det} \bar{E} \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left\{\cosh \sqrt{m^{2} \phi^{2}}+\mathcal{D}^{a} \phi 2 v_{a} \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right. \\
&\left.+2 \bar{e}_{0}^{-0}-\dot{\phi}\left(w-v_{a} \bar{e}_{0}^{a}\right) \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right\} \tag{2.35}
\end{align*}
$$

The background $A d S-C_{d}$ spacetime measure is given by $\operatorname{det} \bar{E}=\bar{e}_{0}^{0} \operatorname{det} \bar{e}_{a}^{m}$ and the partially covariant spatial derivative $\mathcal{D}^{a} \phi=\bar{e}_{m}^{-1 a} \frac{\partial}{\partial x_{m}} \phi$. Note that the covariant derivatives of $\phi$, Equation (2.22), are given by

$$
\begin{align*}
& \nabla^{t} \phi=\dot{\phi} \cos \sqrt{4 v^{2}} \\
& \nabla^{m} \phi=\bar{e}_{a}^{m} \cosh \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left(\frac{\mathcal{D}^{a} \phi}{\cosh \sqrt{m^{2} \phi^{2}}}-2 \delta^{a b} v_{b}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)\right) \tag{2.36}
\end{align*}
$$

and can be used to covariantly constrain (as $\omega_{Z}^{\prime}=\omega_{Z}$ is invariant, Equation (2.28)) the field $v_{a}$ equivalent to the constraint obtained from the $v_{a}$ field equation as well as constrain $\phi$ to be static as obtained from the $w$ field equation.

Indeed the field equations are obtained directly from the $A d S-C_{D}$ invariant action. The $w$-equation of motion is obtained as

$$
\begin{align*}
0 & =\frac{\delta}{\delta w(x, t)} \Gamma_{A d S-C_{D}} \\
& =\left(\operatorname{det} \bar{e}_{a}^{m}\right) \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left[2\left(\frac{\partial}{\partial t} \phi\right)\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)\right] . \tag{2.37}
\end{align*}
$$

The $v_{a}$-field equation yields

$$
\begin{aligned}
& 0=\frac{\delta}{\delta v_{a}(x, t)} \Gamma_{A d S-C_{D}} \\
& \begin{array}{l}
=2 \operatorname{det} \bar{E} \cosh ^{p+1} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left(\frac{\mathcal{D}^{b} \phi}{\cosh \sqrt{m^{2} \phi^{2}}}-2 \delta^{b d} v_{d}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)\right) \times \\
\\
\quad \times\left[-2 v_{b}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right) \frac{\mathcal{D}^{a} \phi}{\cosh \sqrt{m^{2} \phi^{2}}}\right. \\
\\
\quad+\left(\frac{\mathcal{D}^{c} \phi}{\cosh \sqrt{m^{2} \phi^{2}}} 2 v_{c}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)-\tan ^{2} \sqrt{4 v^{2}}+\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right) P_{T b a}(v) \\
\\
\left.\quad+\left(\frac{\mathcal{D}^{c} \phi}{\cosh \sqrt{m^{2} \phi^{2}}} 2 v_{c}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)+1\right) P_{L b a}(v)\right]
\end{array}
\end{aligned}
$$

$$
\begin{align*}
+2\left(\frac{\partial}{\partial t} \phi\right)\left(\operatorname{det} \bar{e}_{b}^{n}\right) \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}} & {\left[-\bar{e}_{0}^{a}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)\right.} \\
& +\left(w-v_{f} \bar{e}_{0}^{f}{ }_{0} \frac{v_{a}}{v^{2}}\left(1-\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)\right] \tag{2.38}
\end{align*}
$$

Finally the $\phi$-equation of motion is obtained

$$
\begin{align*}
0= & \frac{\delta}{\delta \phi(x, t)} \Gamma_{A d S-C_{D}} \\
= & \operatorname{det} \bar{E} \cosh ^{p+1} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}(1+p) m^{2} \phi\left(\frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}}\right) \\
& -\operatorname{det} \bar{E} \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left(2 \bar{e}_{m}^{-1 a} \partial^{m} v_{b}\right)\left[\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right) P_{T b a}(v)+P_{L b a}(v)\right] \\
& -\left[\partial^{m}\left(\operatorname{det} \bar{E} \bar{e}_{m}^{-1 a}\right)\right] \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left[2 v_{a}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)\right] \\
& -\left(\operatorname{det} \bar{e}_{c}^{m}\right) \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}} 2\left(\dot{w}-v_{a} \dot{\bar{e}}_{0}^{a}\right)\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right) \\
& +\left(\operatorname{det} \bar{e}_{c}^{m}\right) \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left(2 \dot{v}_{a} \bar{e}_{0}^{a}\right)\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right) \\
& -\left(\operatorname{det} \bar{e}_{c}^{m}\right) \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left(\frac{v_{a} \dot{v}_{a}}{v^{2}}\right) 2\left(w-v_{b} \bar{e}_{0}^{b}\right)\left[1-\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right)\right] . \tag{2.39}
\end{align*}
$$

Introducing the momentum density $\Pi(x, t)$ as

$$
\begin{equation*}
\Pi(x, t) \equiv \frac{\partial \operatorname{det} E}{\partial \dot{\phi}}=\operatorname{det} \bar{E} \bar{e}_{0}^{0-1} \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}} 2 \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\left(w-v_{a} \bar{e}_{0}^{a}\right) \tag{2.40}
\end{equation*}
$$

and likewise defining the derivative of the Lagrangian with respect to the spatial derivatives of $\phi$ as $\Pi_{m}(x, t)$

$$
\begin{equation*}
\Pi_{m}(x, t) \equiv \frac{\partial \operatorname{det} E}{\partial \partial^{m} \phi}=\operatorname{det} \bar{E} \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}} \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} 2 v_{a} \bar{e}_{m}^{-1 a} \tag{2.41}
\end{equation*}
$$

the $\phi$-field equation is expressed as

$$
\begin{align*}
0= & \frac{\delta \Gamma_{A d S-C_{D}}}{\delta \phi(x, t)} \\
= & -\dot{\Pi}-\partial^{m} \Pi_{m}+\operatorname{det} \bar{E} \cosh ^{p} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\left[\frac{\sinh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} m^{2} \phi\right] \times \\
& \times\left\{p+1+p\left[\frac{\mathcal{D}^{a} \phi}{\cosh \sqrt{m^{2} \phi^{2}}} 2 v_{a} \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}+\frac{\bar{e}_{0}^{-0-1} \dot{\phi}}{\cosh \sqrt{m^{2} \phi^{2}}} 2\left(w-v_{a} \bar{e}_{0}^{a}\right) \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right]\right\} . \tag{2.42}
\end{align*}
$$

Applying the $w$-field equation implies that $\dot{\phi}=0$. As expected there is no causal connection so the field has a static spatial distribution. Applying this to the $v_{m}$-field equation yields the "inverse Higgs mechanism" ${ }^{13}$ for the (spatial) components of the $S O(p)$ vector field $v_{a}$

$$
\begin{equation*}
\frac{\mathcal{D}^{a} \phi}{\cosh \sqrt{m^{2} \phi^{2}}}=2 \delta^{a b} v_{b}\left(\frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right) \tag{2.43}
\end{equation*}
$$

Since the $\phi$ covariant derivatives have the same form, Equation (2.22), the static nature of the spatial distribution of the $\phi$ field and the inverse Higgs constraint for the spatial vector field $v_{a}$
could equivalently be covariantly imposed on the fields by $\omega_{Z}=0$. Finally applying the first two field equations to the $\phi$-field equation the momentum density time dependence is obtained

$$
\begin{gather*}
\dot{\Pi}+\partial^{m} \Pi_{m}=\frac{\operatorname{det} \bar{E} \cosh ^{(p+1)} \sqrt{m^{2} \phi^{2}}}{\sqrt{\cosh ^{2} \sqrt{m^{2} \phi^{2}}+\mathcal{D}^{a} \phi \mathcal{D}^{a} \phi}}\left[m^{2} \phi \frac{\sinh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}}\right] \times \\
\times\left\{p+1+p\left[\frac{\mathcal{D}^{b} \phi \mathcal{D}^{b} \phi}{\cosh ^{2} \sqrt{m^{2} \phi^{2}}}\right]\right\}, \tag{2.44}
\end{gather*}
$$

with the spatial momentum $\Pi_{m}$ becoming

$$
\begin{equation*}
\Pi_{m}=\frac{\operatorname{det} \bar{E} \cosh ^{(p+1)} \sqrt{m^{2} \phi^{2}}}{\sqrt{\cosh ^{2} \sqrt{m^{2} \phi^{2}}+\mathcal{D}^{b} \phi \mathcal{D}^{b} \phi}} e_{m}^{-1 a} \mathcal{D}^{a} \phi \tag{2.45}
\end{equation*}
$$

The Noether current for the broken translation symmetry with parameter $\zeta$ is more complex in the background $A d S-C_{d}$ case and no longer simply produces the $\phi$ field equation as in the Carroll space case but involves time variation of a related composite operator. Noether's theorem has the same form as Equation (1.29) (with $\mathcal{L}_{A d S-C_{D}}$ replacing $\mathcal{L}_{C}$ ) with the broken space translation variations given in Appendix A. In addition to these variations, the induced general coordinate transformation matrix $G=\partial\left(t^{\prime}, x^{\prime}\right) / \partial(t, x)$ is obtained and hence the $\operatorname{det} G^{-1}$

$$
\begin{equation*}
\operatorname{det} G^{-1}=1+\left(1+p+t \frac{\partial}{\partial t}+x_{m} \partial^{m}\right)\left[m^{2} \zeta \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\right] . \tag{2.46}
\end{equation*}
$$

This yields the relation $\Delta \mathcal{L}_{A d S-C}+\mathcal{L}_{A d S-C}\left(\frac{\partial}{\partial t} \delta t+\partial^{m} \delta x_{m}\right)=0$ and Noether's theorem is obtained

$$
\begin{equation*}
0=\dot{D}+\partial^{m} D_{m} \tag{2.47}
\end{equation*}
$$

where

$$
\begin{align*}
& D= \Pi \delta \phi+\delta t \mathcal{L}_{A d S-C_{D}} \\
&= \Pi \cosh \sqrt{m^{2} x^{2}}+m^{2} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\left\{\Pi x_{m} \partial^{m} \phi-t \Pi_{m} \partial^{m} \phi\right. \\
&\left.\quad+t \operatorname{det} \bar{E} \cosh ^{(p+1)} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\right\}, \\
& D_{m}= \Pi_{m} \delta \phi+\delta x_{m} \mathcal{L}_{A d S-C_{D}} \\
&= \Pi_{m} \cosh \sqrt{m^{2} x^{2}}+m^{2} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\left\{t \Pi_{m} \dot{\phi}-x_{m} \Pi \dot{\phi}\right. \\
&\left.\quad+\left(\Pi_{m} x_{n}-\Pi_{n} x_{m}\right) \partial^{n} \phi-x_{m} \operatorname{det} \bar{E} \cosh ^{(p+1)} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}\right\} . \tag{2.48}
\end{align*}
$$

Applying the stationary constraint, $\dot{\phi}=0$, the broken translation symmetry currents become

$$
\begin{array}{r}
D=\Pi\left[\cosh \sqrt{m^{2} x^{2}}+m^{2} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}} x_{m} \partial^{m} \phi\right] \\
+t m^{2} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}} \times \\
\times\left(\operatorname{det} \bar{E} \cosh ^{(p+1)} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}-\Pi_{m} \partial^{m} \phi\right), \\
D_{m}=\Pi_{m}\left[\cosh \sqrt{m^{2} x^{2}}+m^{2} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}} x_{n} \partial^{n} \phi\right]
\end{array}
$$

$$
\begin{align*}
& -x_{m} m^{2} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}} \times \\
& \quad \times\left(\operatorname{det} \bar{E} \cosh ^{(p+1)} \sqrt{m^{2} \phi^{2}} \cos \sqrt{4 v^{2}}-\Pi_{n} \partial^{n} \phi\right) . \tag{2.49}
\end{align*}
$$

The non-linear broken translation symmetry of $A d S-C_{D}$ space no longer results in the $\phi$ field equation directly but involves a composite operator current. In the Carroll spacetime limit $m^{2} \rightarrow 0$, as applied to the above currents, $D \rightarrow \Pi$ and $D_{m} \rightarrow \Pi_{m}$ and the broken translation current conservation equation is just the $\phi$ field equation as the broken translation transformation is a simple $\phi$ field shift symmetry (1.19).

Similarly combining the $L$ and $K^{a}$ one-forms in matrix notation as was done for the $H$ and $P^{a}$ oneforms, Equations (2.19) and (2.20), the covariant derivatives are given as $\nabla^{M} V_{A}$ with $V_{A}=\left(w, v_{a}\right)$ and

$$
\tilde{\omega}_{K A}=\left(\begin{array}{ll}
\omega_{L} & \omega_{K a}
\end{array}\right)=d X_{M} \nabla^{M} V_{A}=\left(\begin{array}{ll}
d t & d x_{m}
\end{array}\right)\left(\begin{array}{cc}
\nabla^{t} w & \nabla^{t} v_{a}  \tag{2.50}\\
\nabla^{m} w & \nabla^{m} v_{a}
\end{array}\right) .
$$

Noting that under the non-linear transformations (2.30) of the coordinates and the transformations of the covariant derivative one-forms with Equation (2.32) so that

$$
\tilde{\omega}_{K A}^{\prime}=\left(\begin{array}{ll}
\omega_{L}^{\prime} & \omega_{K a}^{\prime}
\end{array}\right)=\left(\begin{array}{ll}
\omega_{L} & \omega_{K b}
\end{array}\right)\left(\begin{array}{cc}
1 & 0  \tag{2.51}\\
-\theta_{b} & R_{b a}^{-1}
\end{array}\right)=\tilde{\omega}_{K B} \Lambda_{A}^{B}
$$

implying that

$$
\begin{equation*}
\left(\nabla^{M} V_{A}\right)^{\prime}=G_{N}^{-1 M}\left(\nabla^{N} V_{B}\right) \Lambda_{A}^{B} . \tag{2.52}
\end{equation*}
$$

Combining the vielbeine into the matrix vielbein $E$ as in Equation (2.19), its variation involves $G$ and $\Lambda$ as shown in Equation (2.31). Thus calculating the trace of $\nabla V$ using $E^{-1}$

$$
\begin{equation*}
\left(E_{M}^{-1 A} \nabla^{M} V_{B}\right)^{\prime}=\Lambda_{D}^{-1 A}\left(E_{M}^{-1 D} \nabla^{M} V_{C}\right) \Lambda_{B}^{C}, \tag{2.53}
\end{equation*}
$$

an invariant is obtained

$$
\begin{equation*}
\left(\operatorname{Tr}\left[E^{-1} \nabla V\right]\right)^{\prime}=\operatorname{Tr}\left[E^{-1} \nabla V\right] . \tag{2.54}
\end{equation*}
$$

Applying the $w$ and $v_{a}$ field equation constraints, $\dot{\phi}=0=\dot{v}_{a}$ and implicitly the Equation (2.43), the remaining $\phi$ field equation (2.39) can be shown to be given by the trace of the covariant derivative of $w$ and $v_{a}$. The inverse vielbein with the $\dot{\phi}=0$ constraint explicit so that $e_{a}^{0}=0$ is found to be

$$
E_{M}^{-1 A}=\left(\begin{array}{cc}
e_{0}^{0-1} & 0  \tag{2.55}\\
-e_{n}^{-1 a} e_{0}^{n} e_{0}^{0-1} & e_{m}^{-1 a}
\end{array}\right),
$$

where $e_{0}^{-10}=e_{0}^{0-1}$ with $e_{0}^{0}=\bar{e}_{0}^{0} \cosh \sqrt{m^{2} \phi^{2}}$ and the inverse submatrix $e_{m}^{-1 a}$ is found to be

$$
\begin{equation*}
e_{m}^{-1 a}=\frac{1}{\cosh \sqrt{m^{2} \phi^{2}}}\left[P_{T a b}(v)+\cos \sqrt{4 v^{2}} P_{L a b}(v)\right] \bar{e}_{m}^{-1 b}, \tag{2.56}
\end{equation*}
$$

both with the field constraints having been used. Multiplying these matrices together yields the vector field covariant derivative trace

$$
\begin{equation*}
\operatorname{Tr}\left[E^{-1} \nabla V\right]=e_{0}^{0-1} \nabla^{t} w+e_{m}^{-1 a} \nabla^{m} v_{a} \tag{2.57}
\end{equation*}
$$

where again the constraints have been used, particularly $\dot{v}_{a}=0$. Applying the covariant derivatives which are recalled in Equations (2.24) with again the constraints applied finally yields the trace

$$
\begin{align*}
\operatorname{Tr}\left[E^{-1} \nabla V\right]=- & \frac{1}{2} m^{2} \phi(1+p) \cos \sqrt{4 v^{2}} \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \\
+ & \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}}\left(\bar{e}_{0}^{0-1}\left(\dot{w}-p \bar{\omega}_{0 a}^{t} v_{a}\right) \frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right. \\
& \left.+\bar{e}_{n}^{-1 b} \partial^{n} v_{a}\left[\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} P_{T a b}(v)+\cos \sqrt{4 v^{2}} P_{L a b}(v)\right]\right) \tag{2.58}
\end{align*}
$$

Using the background one-forms, Equation (2.12), and background spin connections, Equation (2.13), it is found that

$$
\begin{equation*}
\partial^{m}\left[\operatorname{det} \bar{e} \bar{e}_{0}^{0} \bar{e}_{m}^{-1 a}\right]=\operatorname{det} \bar{e}\left[\dot{\bar{e}}_{0}^{a}-p \bar{\omega}_{0 b}^{t} \delta^{b a}\right], \tag{2.59}
\end{equation*}
$$

with $\bar{e}_{0}^{a}=\bar{e}^{-1 a}{ }_{r} \bar{e}_{0}^{r}$ so that $\dot{\bar{e}}_{0}^{a}=\bar{e}_{r}^{-1 a} \dot{\bar{e}}_{0}^{r}$. Applying this and the field constraints to the $\phi$ field Equation (2.39) and exploiting the background spin connections $\bar{\omega}_{a b}^{n}$ and $\bar{\omega}_{0 a}^{t}$, it is found that (2.39) (with $\dot{\phi}=0=\dot{v}_{a}$ ) can be written as

$$
\begin{equation*}
0=\frac{\delta \Gamma_{A d S-C_{D}}}{\delta \phi}=-2 \operatorname{det} \bar{E} \cosh ^{(1+p)} \sqrt{m^{2} \phi^{2}} \operatorname{Tr}\left[E^{-1} \nabla V\right] \tag{2.60}
\end{equation*}
$$

## III. THE NAMBU-GOTO CARROLLIAN VIELBEIN AND THE DUAL VECTOR ACTION

Returning to the unconstrained fields and combining the vielbeine in matrix form $E_{A}^{M}$, it and the Lagrangian, det $E$, can be factorized into a product of the background $A d S_{d}$ world volume vielbein, $\bar{E}_{A}^{M}$, times the Nambu-Goto Carrollian vielbein, $N_{A}^{B}$, as such

$$
\begin{equation*}
E_{A}^{M}=\bar{E}_{B}^{M} N_{A}^{B}, \tag{3.1}
\end{equation*}
$$

with Equation (2.20)

$$
E_{A}^{M}=\left(\begin{array}{cc}
e_{0}^{0} & e_{a}^{0}  \tag{3.2}\\
e_{0}^{m} & e_{a}^{m}
\end{array}\right)
$$

and background $A d S_{d}$ vielbein

$$
\bar{E}_{A}^{M}=\left(\begin{array}{cc}
\bar{e}_{0}^{0} & \bar{e}_{a}^{0}=0  \tag{3.3}\\
\bar{e}_{0}^{m} & \bar{e}_{a}^{m}
\end{array}\right)
$$

and correspondingly the inverse background world volume vielbein

$$
\bar{E}_{M}^{-1 A}=\left(\begin{array}{cc}
\bar{e}_{0}^{-0-1} & 0  \tag{3.4}\\
-\bar{e}_{r}^{-1 a} \bar{e}_{0}^{r} \bar{e}_{0}^{-0-1} & \bar{e}_{m}^{-1 a}
\end{array}\right) .
$$

This yields the Nambu-Goto Carrollian vielbein $N_{A}^{B}=\bar{E}^{-1 B}{ }_{M} E_{A}^{M}$

$$
N_{A}^{B}=\left(\begin{array}{cc}
N_{0}^{0} & N_{a}^{0}  \tag{3.5}\\
N_{0}^{b} & N_{a}^{b}
\end{array}\right)
$$

with

$$
\begin{align*}
N_{0}^{0} & =\bar{e}_{0}^{0-1} e_{0}^{0} \\
& =e^{A}\left[1+2 w e^{-A} \mathcal{D}^{t} \phi \frac{\sin 2 v}{2 v}\right], \\
N_{0}^{b} & =\bar{e}_{n}^{-1 b}\left[e_{0}^{n}-\bar{e}_{0}^{n} e_{0}^{0-1} e_{0}^{0}\right] \\
& =2 w \frac{\sin 2 v}{2 v}\left[\mathcal{D}^{b} \phi-\bar{e}_{0}^{b} \mathcal{D}^{t} \phi\right]+e^{A} w v_{b}\left(\frac{\cos 2 v-1}{v^{2}}\right), \\
N_{a}^{0} & =\bar{e}_{0}^{0-1} e_{a}^{0} \\
& =2 \mathcal{D}^{t} \phi v_{a} \frac{\sin 2 v}{2 v}, \\
N_{a}^{b} & =\bar{e}_{m}^{-1 b}\left[e_{a}^{m}-\bar{e}_{0}^{m} \bar{e}_{0}^{0-1} e_{a}^{0}\right] \\
& =2 v_{a} \frac{\sin 2 v}{2 v}\left[\mathcal{D}^{b} \phi-\bar{e}_{0}^{b} \mathcal{D}^{t} \phi\right]+e^{A}\left[P_{\text {Tba }}(v)+\cos 2 v P_{L b a}(v)\right], \tag{3.6}
\end{align*}
$$

where the notation uses the simplifying expressions

$$
e^{A}=e^{\ln \cosh \sqrt{m^{2} \phi^{2}}}=\cosh \sqrt{m^{2} \phi^{2}}=\cosh m \phi
$$

$$
\begin{align*}
\mathcal{D}^{t} \phi & =e_{0}^{-0-1} \frac{\partial}{\partial t} \phi=\bar{e}_{0}^{-0-1} \dot{\phi}, \\
\mathcal{D}^{a} \phi & =e_{m}^{-1 a} \frac{\partial}{\partial x_{m}} \phi=\bar{e}_{m}^{-1 a} \partial^{m} \phi, \\
x^{2} & =x_{m} \delta^{m n} x_{n}=x_{m} x_{m}, \\
v & =\sqrt{v^{2}}=\sqrt{v_{a} v_{a}}, \tag{3.7}
\end{align*}
$$

where recall the shorthand expressions such as $v_{a}=\delta^{a b} v_{b}$ and likewise $x_{m}=\delta^{m n} x_{n}$ are used and recall the background component vielbeine

$$
\begin{align*}
\bar{e}_{m}^{-1 a} \bar{e}_{0}^{m} & =\frac{x_{a} t}{x^{2}}\left(1-\bar{e}_{0}^{0}\right)=\bar{e}_{0}^{a}, \\
\bar{e}_{m}^{-1 a} & =\bar{e}_{0}^{-1} P_{\text {Tam }}(x)+P_{\text {Lam }}(x), \\
\bar{e}_{0}^{0} & =\frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}} . \tag{3.8}
\end{align*}
$$

Hence the $A d S-C_{D}$ action is given by

$$
\begin{equation*}
\Gamma_{A d S-C_{D}}=\int d t d^{p} x \operatorname{det} E=\int d t d^{p} x \operatorname{det} \bar{E} \operatorname{det} N, \tag{3.9}
\end{equation*}
$$

where $\operatorname{det} \bar{E}=\bar{e}_{0}^{0} \operatorname{det} \bar{e}_{a}^{m}$ and

$$
\begin{equation*}
\operatorname{det} N=\left(\operatorname{det} N_{d}^{c}\right)\left[N_{0}^{0}-N_{a}^{0} N_{b}^{-1 a} N_{0}^{b}\right] . \tag{3.10}
\end{equation*}
$$

Letting $N_{b}^{a}=u_{a} v_{b}+e^{A}\left(P_{T a b}(v)+\cos 2 v P_{\text {Lab }}(v)\right)$ the inverse submatrix $N_{b}^{-1 a}$ is

$$
\begin{equation*}
N_{b}^{-1 a}=\alpha u_{a} v_{b}+e^{-A}\left[P_{T a b}(v)+\beta P_{L a b}(v)\right], \tag{3.11}
\end{equation*}
$$

where

$$
\begin{align*}
& \alpha=\frac{-e^{-A}}{\left[u_{a} v_{a}+e^{A} \cos 2 v\right]}, \\
& \beta=\frac{\left(e^{A}+u_{b} v_{b}\right)}{\left[u_{a} v_{a}+e^{A} \cos 2 v\right]} . \tag{3.12}
\end{align*}
$$

For the case at hand

$$
\begin{equation*}
u_{a}=2\left(\mathcal{D}^{a} \phi-\bar{e}_{0}^{a} \mathcal{D}^{t} \phi\right) \frac{\sin 2 v}{2 v}, \tag{3.13}
\end{equation*}
$$

and so $u_{a} v_{a}=2\left(\mathcal{D}^{a} \phi-\bar{e}_{0}^{a} \mathcal{D}^{t} \phi\right) v_{a} \frac{\sin 2 v}{2 v}$. The submatrix determinant, det $N_{b}^{a}$, is also determined to be

$$
\begin{equation*}
\operatorname{det} N_{b}^{a}=e^{p A} \cos 2 v\left[1+2 e^{-A}\left(\mathcal{D}^{a} \phi-\bar{e}_{0}^{a} \mathcal{D}^{t} \phi\right) v_{a} \frac{\tan 2 v}{2 v}\right] . \tag{3.14}
\end{equation*}
$$

Putting all these expressions together yields the Nambu-Goto Carrollian determinant $\operatorname{det} N=\cosh ^{p} m \phi\left[\cosh m \phi \cos 2 v+2\left(\mathcal{D}^{a} \phi v_{a} \frac{\sin 2 v}{2 v}\right)+2\left(w-v_{a} \bar{e}_{m}^{-1 a} \bar{e}_{0}^{m}\right) \mathcal{D}^{t} \phi \frac{\sin 2 v}{2 v}\right]$.
Thus with $\operatorname{det} \bar{E}=\bar{e}_{0}^{0} \operatorname{det} \bar{e}_{a}^{m}=\left(\bar{e}_{0}^{0}\right)^{p}$, the $\operatorname{det} E=\operatorname{det} \bar{E} \operatorname{det} N$ Lagrangian is in agreement with the Lagrangian obtained in Equation (2.35). The invariant action is obtained

$$
\begin{align*}
& \Gamma_{A d S-C_{D}}=\int d t d^{p} x \operatorname{det} E=\int d t d^{p} x \operatorname{det} \bar{E} \operatorname{det} N \\
&=\int d t d^{p} x e_{0}^{-0}(\operatorname{det} \bar{e}) e^{p A}\left[e^{A} \cos 2 v+2\left(\mathcal{D}^{a} \phi v_{a} \frac{\sin 2 v}{2 v}\right)\right. \\
&\left.+2\left(w-v_{a} \bar{e}_{m}^{-1 a} \bar{e}_{0}^{m}\right) \mathcal{D}^{t} \phi \frac{\sin 2 v}{2 v}\right] . \tag{3.16}
\end{align*}
$$

As in the $A d S$ case $^{6}$ a vector field dual formulation of the $A d S-C_{D}$ action can be obtained by introducing a vector field $F_{M}$, with $M=0,1,2, \ldots, p$

$$
\begin{equation*}
F_{M} \equiv 2 \operatorname{det} \bar{E} \frac{\sin 2 v}{2 v} V_{A} \bar{E}_{M}^{-1 A}, \tag{3.17}
\end{equation*}
$$

with space-time component fields

$$
\begin{equation*}
V_{A}=\left(w, v_{a}\right) \tag{3.18}
\end{equation*}
$$

and

$$
\begin{equation*}
F_{M}=\left(f, f_{m}\right) \tag{3.19}
\end{equation*}
$$

It is useful to introduce the singular tangent space metric

$$
h^{A B}=\left(\begin{array}{cc}
0 & 0  \tag{3.20}\\
0 & \delta^{a b}
\end{array}\right)
$$

as well as the singular background $A d S-C_{d}$ metric

$$
\begin{align*}
\bar{G}^{M N} & \equiv \bar{E}_{A}^{M} h^{A B} \bar{E}_{B}^{N} \\
& =\left(\begin{array}{cc}
0 & 0 \\
0 & \bar{E}_{a}^{m} \delta^{a b} \bar{E}_{b}^{n}
\end{array}\right) \\
& \equiv\left(\begin{array}{cc}
0 & 0 \\
0 & \bar{g}^{m n}
\end{array}\right) . \tag{3.21}
\end{align*}
$$

Note that $\operatorname{det} \bar{G}=0$ and hence has no inverse. Using the singular metric it is found that

$$
\begin{equation*}
F_{M} \bar{G}^{M N} F_{N}=f_{m} \bar{g}^{m n} f_{n} \tag{3.22}
\end{equation*}
$$

In terms of component fields it is obtained that

$$
\begin{equation*}
F_{M}=\binom{f}{f_{m}}=2 \operatorname{det} \bar{E} \frac{\sin 2 v}{2 v}\binom{\left(w-v_{a} \bar{e}^{-1 a}{ }_{r} \bar{e}_{0}^{r}\right) \bar{e}_{0}^{-0-1}}{v_{a} \bar{e}^{-1 a}{ }_{m}}, \tag{3.23}
\end{equation*}
$$

Inverting Equation (3.17)

$$
\begin{equation*}
V_{A}=\left(\frac{1}{2 \operatorname{det} \bar{E} \frac{\sin 2 v}{2 v}}\right) F_{M} \bar{E}_{A}^{M}, \tag{3.24}
\end{equation*}
$$

so that the component fields are related by

$$
\begin{equation*}
V_{A}=\binom{w}{v_{a}}=\frac{1}{2 \operatorname{det} \bar{E} \frac{\sin 2 v}{2 v}}\binom{f \bar{e}_{0}^{-0}+f_{m} \bar{e}_{0}^{m}}{f_{m} \bar{e}_{a}^{m}} . \tag{3.25}
\end{equation*}
$$

Hence the useful expressions are found

$$
\begin{align*}
v^{2} & =v_{a} \delta^{a b} v_{b}=V_{A} h^{A B} V_{B}=\frac{1}{\left(2 \operatorname{det} \bar{E} \frac{\sin 2 v}{2 v}\right)^{2}} F_{M} \bar{G}^{M N} F_{N} \\
& =\frac{1}{\left(2 \operatorname{det} \bar{E} \frac{\sin 2 v}{2 v}\right)^{2}} f_{m} \bar{g}^{m n} f_{n}, \tag{3.26}
\end{align*}
$$

while

$$
\begin{equation*}
\operatorname{det} \bar{E} \cos 2 v=\sqrt{(\operatorname{det} \bar{E})^{2}-F_{M} \bar{G}^{M N} F_{N}} \tag{3.27}
\end{equation*}
$$

and

$$
\begin{align*}
V_{A} \bar{E}_{M}^{-1 A} \partial^{M} \phi & =\left(w-v_{a} \bar{e}_{m}^{-1 a} \bar{e}_{0}^{m}\right) \mathcal{D}^{t} \phi+v_{a} \mathcal{D}^{a} \phi \\
& \equiv V_{A} \hat{\mathcal{D}}^{A} \phi, \tag{3.28}
\end{align*}
$$

where $\hat{\mathcal{D}}^{A} \phi=\bar{E}_{M}^{-1 A} \partial^{M} \phi$ with $\partial^{M}=\left(\partial^{t}, \partial^{m}\right)$ and the partial covariant derivatives $\mathcal{D}^{t} \phi=\bar{e}_{0}^{0-1} \dot{\phi}$ and $\mathcal{D}^{a} \phi=\bar{e}_{m}^{-1 a} \partial^{m} \phi$.

The determinant of the Nambu-Goto Carrollian vielbein, Equation (3.15), can be written in terms of $\phi$ and $F_{M}$ as

$$
\begin{equation*}
\operatorname{det} \bar{E} \operatorname{det} N=e^{p A}\left(e^{A} \sqrt{(\operatorname{det} \bar{E})^{2}-F_{M} \bar{G}^{M N} F_{N}}+F_{M} \partial^{M} \phi\right) \tag{3.29}
\end{equation*}
$$

Following Ref. 6, introduce the function $h(\phi)$ so that

$$
\begin{equation*}
\partial^{M} \phi e^{p A(\phi)} \equiv \partial^{M} h(\phi)=\frac{d h}{d \phi} \partial^{M} \phi \tag{3.30}
\end{equation*}
$$

and hence $\frac{d h(\phi)}{d \phi}=e^{p A(\phi)}$, the $A d S-C_{D}$ action now has the form, after integration by parts,

$$
\begin{equation*}
\Gamma_{A d S-C_{D}}=\int d t d^{p} x\left[e^{(p+1) A(\phi)} \sqrt{(\operatorname{det} \bar{E})^{2}-F_{M} \bar{G}^{M N} F_{N}}-h(\phi) \partial^{M} F_{M}\right] \tag{3.31}
\end{equation*}
$$

The scalar field $\phi$ equation of motion follows $(d=(1+p))$

$$
\begin{equation*}
\frac{\delta \Gamma_{A d S-C_{D}}}{\delta \phi}=0=\frac{d h}{d \phi}\left[d m^{2} \phi \frac{\sinh m \phi}{m \phi} \sqrt{(\operatorname{det} \bar{E})^{2}-F_{M} \bar{G}^{M N} F_{N}}-\partial^{M} F_{M}\right] \tag{3.32}
\end{equation*}
$$

The $\phi$ equation of motion can be enforced by introducing the Lagrange multiplier field $L$ yielding the action
$\Gamma_{A d S-C_{D}}=\int d t d^{p} x\left[\sqrt{(\operatorname{det} \bar{E})^{2}-F_{M} \bar{G}^{M N} F_{N}}\left(T(\phi)+L d m^{2} \phi \frac{\sinh m \phi}{m \phi}\right)-L \partial^{M} F_{M}\right]$,
with

$$
\begin{equation*}
T(\phi)=e^{(1+p) A(\phi)}-h(\phi) d m^{2} \phi \frac{\sinh m \phi}{m \phi} \tag{3.34}
\end{equation*}
$$

Thus we see that the previous $\phi$ equation of motion, now coming from the $L$ field equation, results in the $L$ dependent terms cancelling and the $\phi$ field itself being expressed in terms of $F_{M}$. Thus using this equation of motion

$$
\begin{align*}
\sinh m \phi & =\frac{\partial^{M} F_{M}}{d m \sqrt{(\operatorname{det} \bar{E})^{2}-F_{R} \bar{G}^{R S} F_{S}}} \\
& =\frac{\dot{f}+\partial^{m} f_{m}}{d m \sqrt{(\operatorname{det} \bar{E})^{2}-f_{r} \bar{g}^{r s} f_{s}}} \tag{3.35}
\end{align*}
$$

the $T(\phi)$ can be written in terms of $F_{M}$, so adopting the notation $T(\phi)=T\left(\phi\left(F^{M}\right)\right) \rightarrow T(F)$, the dual vector form of the action is

$$
\begin{align*}
\Gamma_{A d S-C_{D}} & =\int d t d^{p} x T(F) \sqrt{(\operatorname{det} \bar{E})^{2}-F_{M} \bar{G}^{M N} F_{N}} \\
& =\int d t d^{p} x T\left(f, f_{m}\right) \sqrt{(\operatorname{det} \bar{E})^{2}-f_{m} \bar{g}^{m n} f_{n}} \tag{3.36}
\end{align*}
$$

along with Equations (3.35) and (3.34) to determine $\phi=\phi(F)$ and $T(F)$.
The equivalence runs in reverse as well, introducing a Lagrange multiplier $L$ to enforce Equation (3.35) yields the action of Equation (3.33) where $T(\phi)$ is given in Equation (3.34). The fields $\phi, F_{M}$, and $L$ are independent. The $\phi$ equation of motion allows $L$ to be eliminated as $\delta \Gamma_{A d S-C_{D}} / \delta \phi=0$ implies

$$
\begin{equation*}
L=-\frac{1}{d m^{2}} e^{-A} \frac{d T}{d \phi}=h(\phi) \tag{3.37}
\end{equation*}
$$

Upon substitution into $\Gamma_{A d S-C_{D}}$, Equation (3.33), yields Equation (3.31). The definition of $F_{M}$ in terms of $V_{A}$, Equation (3.17), can be applied along with (recalling that $V^{2}=V_{A} h^{A B} V_{B}=v^{2}$ )

$$
\begin{equation*}
\cos 2 V=\cos 2 v=\frac{1}{(\operatorname{det} \bar{E})} \sqrt{(\operatorname{det} \bar{E})^{2}-F_{M} \bar{G}^{M N} F_{N}} \tag{3.38}
\end{equation*}
$$

Thus substituting this into Equation (3.31) and integrating by parts results in the Nambu-GotoCarrollian action, note Equation (B58),

$$
\begin{equation*}
\Gamma_{A d S-C_{D}}=\int d t d^{p} x \operatorname{det} \bar{E} e^{p A(\phi)}\left[e^{A(\phi)} \cos 2 V+\frac{\sin 2 V}{2 V} 2 V_{A} \bar{E}_{M}^{-1 A} \partial^{M} \phi\right] \tag{3.39}
\end{equation*}
$$

This can be expanded in terms of component fields $w$ and $v_{a}$ to obtain the Nambu-Goto Carrollian action of Equation (3.16).

## IV. CONCLUSION

A p-brane with codimension one was embedded in $D=d+1$-dimensional AdS-Carroll space by means of the coset method. The vanishing speed of light, $c \rightarrow 0$, Wigner-Inönü contraction of the $A d S$ space $S O(2, D-1)$ symmetry algebra was obtained and the $A d S-C_{D} / I S O(p)$ coset element, Equation (2.8), was formed along with the unbroken background $A d S-C_{d} / I S O(p)$ coset element Equation (2.9). The non-linearly realized spontaneously broken $A d S-C_{D} \rightarrow A d S-C_{d}$ symmetry transformations were obtained in Appendix A. The invariant brane action was found using the Maurer-Cartan one-forms. Expanding the one-forms in terms of coordinate differentials, the vielbeine and background vielbeine were obtained. These component vielbeine were re-assembled as a Carroll spacetime matrix vielbein, Equations (2.14) and (2.19). The $A d S-C_{D} S O(2, D-1)$ invariant action was shown to be given by the determinant of this matrix vielbein

$$
\begin{equation*}
\Gamma_{A d S-C_{D}}=\int d t d^{p} x \operatorname{det} E \tag{4.1}
\end{equation*}
$$

with $\operatorname{det} E$ expressed in terms of the component fields in Equation (2.35).
The $w, v_{a}$, and $\phi$ field equations followed directly from the action. The $w$ and $v_{a}$ equations of motion implied the inverse Higgs mechanism constraints yielding the static nature of the spatial shape of the $\phi$ field, $\dot{\phi}=0$, as expected from the contraction of the light cone to the time axis as $c \rightarrow 0$, and the auxiliary vector field $v_{a}$ and the spatial derivatives of $\phi$ were related, Equation (2.43). Both of these constraints can alternatively be imposed by the invariant $\phi$-covariant derivative constraint $\omega_{Z}=0$. The canonical momentum density defined in Equation (2.40) on the other hand exhibits time variation related to the shape of the brane, $\phi\left(x_{m}\right)$, as well as the $A d S-C$ geometry. Finally the $\phi$-field equation can be written in terms of the covariant derivatives of the auxiliary fields $w$ and $v_{a}$ as Equation (2.60).

In the flat Minkowski space limit, $m^{2} \rightarrow 0$, the $A d S-C_{D}$ results describe the $p$-brane embedded in a Carroll spacetime, $A d S-C_{D} \rightarrow C_{D}$. These results agree with those of the more informally derived results discussed in the Introduction. In addition the broken translation symmetry Noether current in the $A d S-C$ case, Equation (2.47), with currents Equation (2.49), goes over to the Carroll space currents $z=\Pi, z_{m}=\Pi_{m}$, found in Table I, and the Carrollian component vielbeine can be obtained as the $m^{2} \rightarrow 0$ limit of Equation (2.18) and correspondingly the $m^{2} \rightarrow 0$ action $\Gamma_{C}=\int d t d^{p} x \mathcal{L}_{C}$ with $\operatorname{det} E \xrightarrow{c \rightarrow 0} \mathcal{L}_{C}$ and action $\Gamma_{C}$ of Equation (1.9).

The $A d S-C_{D}$ vielbein $E$ has the product form of the $A d S-C_{d}$ background vielbein $\bar{E}$ times the Nambu-Goto-Carroll vielbein $N, E=\bar{E} N$, as expressed in Equations (3.1)-(3.7). The $p$-brane action can be re-formulated in terms of its dual vector field $F_{M}$ action, Equation (3.36), with functions $\phi=\phi(F)$, Equation (3.35), and $T(F)$, Equation (3.34). Likewise the dual action can be reformulated to yield the brane Nambu-Goto-Carrollian action Equation (3.39) with component form Equation (3.16).

An equivalent approach to obtain the $p$-brane action is to expose the speed of light in the already known $A d S_{(d+1)} \rightarrow A d S_{d}$ brane action results and take the $c \rightarrow 0$ Carrollian limit thereof.

This method was presented in Appendix B where the Maurer-Cartan one-forms and NambuGoto vielbein of Ref. 6 were used to obtain the $A d S$ action with the speed of light parameter, Equation (B58). The Carrollian limit was then taken to obtain the coset method component action (3.16) results.

## ACKNOWLEDGMENTS

The work of T.t.V was supported in part by the NSF under Grant No. PHY- 1102585. T.t.V gratefully acknowledges the hospitality of E. Bergshoeff and the Van Swinderen Institute for Particle Physics and Gravity at the University of Groningen while on sabbatical leave from Macalester College. T.t.V also thanks Joaquim Gomis for useful discussions.

## APPENDIX A: AdS - $C$ TRANSFORMATIONS

Using the group multiplication laws as applied to the coset $\Omega$, the non-linearly realized $A d S-$ $C_{D}$ transformations are determined from

$$
\begin{equation*}
g \Omega(x, t)=\Omega^{\prime}\left(x^{\prime}, t^{\prime}\right) h(x, t), \tag{A1}
\end{equation*}
$$

where the infinitesimal $A d S-C_{D}$ transformations form the group elements

$$
\begin{equation*}
g=e^{i \epsilon H} e^{i a_{m} P^{m}} e^{i \zeta Z} e^{i \lambda L} e^{i \kappa_{m} K^{m}} e^{\frac{i}{2} \alpha_{m n} M^{m n}} e^{i \beta_{m} B^{m}}, \tag{A2}
\end{equation*}
$$

while the transformed coset element is given by

$$
\begin{equation*}
\Omega^{\prime}\left(x^{\prime}, t^{\prime}\right)=e^{i t^{\prime} H+i x_{m}^{\prime} P^{m}} e^{i \phi^{\prime}\left(x^{\prime}, t^{\prime}\right) Z} e^{i \omega^{\prime}\left(x^{\prime}, t^{\prime}\right) L+v_{m}^{\prime}\left(x^{\prime}, t^{\prime}\right) K^{m}} \tag{A3}
\end{equation*}
$$

The $h(x, t)$ is an element of the invariant $I S O(p)$ subgroup

$$
\begin{equation*}
h(x, t)=e^{\frac{i}{2} \theta_{m n}(x, t) M^{m n}} e^{i \theta_{m}(x, t) B^{m}} \tag{A4}
\end{equation*}
$$

with parameters $\theta_{m n}$ and $\theta_{m}$ that also depend on $g$. The transformations of the spacetime coordinates and fields are found to be non-linearly realized

$$
\begin{aligned}
t^{\prime}= & t\left[1+\frac{a_{m} x_{m}}{x^{2}}\left(1-\sqrt{m^{2} x^{2}} \operatorname{coth} \sqrt{m^{2} x^{2}}\right)\right. \\
& -m^{2} \zeta \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}} \\
& \left.-\frac{\kappa_{m} x_{m}}{x^{2}} 2 \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}}\left(\cosh \sqrt{m^{2} x^{2}}-\frac{\sqrt{m^{2} x^{2}}}{\sinh \sqrt{m^{2} x^{2}}}\right)\right] \\
+ & +\sqrt{m^{2} x^{2}} \operatorname{coth} \sqrt{m^{2} x^{2}}-\lambda 2 \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sqrt{m^{2} x^{2}}}{\sinh \sqrt{m^{2} x^{2}}}-\beta_{m} x_{m}, \\
x_{m}^{\prime}= & x_{m}\left[1-m^{2} \zeta \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\right]+\alpha_{m n} x_{n} \\
& +\left(\sqrt{m^{2} x^{2}} \operatorname{coth} \sqrt{m^{2} x^{2}} P_{T m n}(x)+P_{L m n}(x)\right) a_{n} \\
& -2 \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}}\left[\frac{\sqrt{m^{2} x^{2}}}{\sinh \sqrt{m^{2} x^{2}}} P_{T m n}(x)+\cosh \sqrt{m^{2} x^{2}} P_{L m n}(x)\right] \kappa_{n}, \\
\phi^{\prime}\left(x^{\prime}, t^{\prime}\right)= & \phi(x, t)+\zeta \cosh \sqrt{m^{2} x^{2}}+2 \kappa_{m} x_{m} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}, \\
w^{\prime}\left(x^{\prime}, t^{\prime}\right)= & w(x, t)+\epsilon m^{2} x_{m} v_{m} \frac{\tanh \sqrt{m^{2} x^{2}} / 2}{\sqrt{m^{2} x^{2}}}-\beta_{m} v_{m}
\end{aligned}
$$

$$
\begin{align*}
&-m^{2} t \frac{\tanh \sqrt{m^{2} x^{2}} / 2}{\sqrt{m^{2} x^{2}}} a_{m} v_{m}+\lambda \frac{\sqrt{4 v^{2}} \cot \sqrt{4 v^{2}}}{\cosh \sqrt{m^{2} \phi^{2}}} \\
&+\lambda 2 \phi \frac{v_{m} x_{m}}{x^{2}}\left(\sqrt{m^{2} x^{2}} \tanh \sqrt{m^{2} x^{2}} / 2\right)\left(\frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}}\right) \\
&+\frac{1}{2} m^{2} \zeta \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\left[w \frac{x_{m} v_{m}}{v^{2}}+\sqrt{4 v^{2}} \cot \sqrt{4 v^{2}}\left(t-w \frac{x_{m} v_{m}}{v^{2}}\right)\right] \\
&+w \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \frac{\left[1-\sqrt{4 v^{2}} \cot \sqrt{4 v^{2}}\right]}{v^{2}} v_{m}\left[P_{T m n}(x)+\cosh \sqrt{m^{2} x^{2}} P_{L m n}(x)\right] \kappa_{n} \\
&-m^{2} t \kappa_{m} v_{m} \phi\left(\frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}}\right)\left(\frac{\tanh \sqrt{m^{2} x^{2}} / 2}{\sqrt{m^{2} x^{2}} / 2}\right) \\
&+t \kappa_{m} x_{m}\left(\frac{\sqrt{4 v^{2}} \cot \sqrt{4 v^{2}}}{\cosh \sqrt{m^{2} \phi^{2}}}\right)\left(\frac{\cosh \sqrt{m^{2} x^{2}}-1}{x^{2}}\right) \\
& v_{m}^{\prime}\left(x^{\prime}, t^{\prime}\right)= v_{m}(x, t)+\alpha_{m n} v_{n}+\frac{m^{2}}{2}\left(a_{m} x_{n}-a_{n} x_{m}\right) \frac{2}{\sqrt{m^{2} x^{2}}} \tanh \sqrt{m^{2} x^{2}} / 2 v_{n} \\
&+\frac{m^{2}}{2} \zeta \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\left[\sqrt{4 v^{2}} \cot \sqrt{4 v^{2}} P_{T m n}(v)+P_{L m n}(v)\right] x_{n} \\
&-m^{2} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{2}{\sqrt{m^{2} x^{2}}} \tanh \sqrt{m^{2} x^{2}} / 2\left(x_{m}\left(\kappa_{n} v_{n}\right)-\kappa_{m}\left(x_{n} v_{n}\right)\right) \\
&+\left(\sqrt{4 v^{2}} \cot \sqrt{4 v^{2}} P_{T m n}(v)+P_{L m n}(v)\right) \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \\
&  \tag{A5}\\
&\left(P_{T n r}(x)+\cosh \sqrt{m^{2} x^{2}} P_{L n r}(x)\right) \kappa_{r} . \\
&
\end{align*}
$$

The invariant $I S O(p)$ subgroup parameters $\theta_{m n}$ and $\theta_{m}$ are also obtained

$$
\begin{aligned}
\theta_{m n}= & \alpha_{m n}+\frac{m^{2}}{2}\left(a_{m} x_{n}-a_{n} x_{m}\right) \frac{2}{\sqrt{m^{2} x^{2}}} \tanh \sqrt{m^{2} x^{2}} / 2 \\
& -m^{2} \zeta \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\left(x_{m} v_{n}-x_{n} v_{m}\right) \frac{\tan \sqrt{v^{2}}}{\sqrt{v^{2}}} \\
& -m^{2} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{2}{\sqrt{m^{2} x^{2}}} \tanh \sqrt{m^{2} x^{2}} / 2\left(x_{m} \kappa_{n}-x_{n} \kappa_{m}\right) \\
& -2 \frac{\tan \sqrt{v^{2}}}{\sqrt{v^{2}}} \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}}\left[\left(P_{T m r}(x)+\cosh \sqrt{m^{2} x^{2}} P_{L m r}(x)\right){\kappa_{r} v_{n}}^{\theta_{m}=}\right. \\
& \beta_{m}-\epsilon m^{2} x_{m} \frac{\tanh \sqrt{m^{2} x^{2}} / 2}{\sqrt{m^{2} x^{2}}}+m^{2} \operatorname{ta}_{m} \frac{\tanh \sqrt{m^{2} x^{2}} / 2}{\sqrt{m^{2} x^{2}}} \\
& +m^{2} \zeta \frac{\tan \sqrt{v^{2}}}{\sqrt{v^{2}}} \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\left(t v_{m}-x_{m} w\right) \\
& -\lambda m^{2} x_{m} \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{2}{\sqrt{m^{2} x^{2}}} \tanh ^{m^{2} x^{2}} / 2+2 \lambda v_{m} \frac{\tan \sqrt{v^{2}}}{\sqrt{v^{2}}} \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \\
& +m^{2} t \phi \frac{\tanh \sqrt{m^{2} \phi^{2}}}{\sqrt{m^{2} \phi^{2}}} \frac{2}{\sqrt{m^{2} x^{2}}} \tanh _{L n r}(x){\left.\kappa_{r} v_{m}\right]}_{\tan }^{m^{2} x^{2} / 2 \kappa_{m}}
\end{aligned}
$$

$$
\begin{align*}
& +2 t v_{m}\left(\kappa_{n} x_{n}\right) \frac{\tan \sqrt{v^{2}}}{\sqrt{v^{2}}} \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \frac{\cosh \sqrt{m^{2} x^{2}}-1}{x^{2}} \\
& -2 w \frac{\tan \sqrt{v^{2}}}{\sqrt{v^{2}}} \frac{1}{\cosh \sqrt{m^{2} \phi^{2}}}\left(P_{T m n}(x)+\cosh \sqrt{m^{2} x^{2}} P_{L m n}(x)\right) \kappa_{n} \tag{A6}
\end{align*}
$$

The symmetry transformations for $D=d+1$ Carrollian spacetime as given in Equations (1.19) and (1.21) and as well as the induced local rotations and boosts

$$
\Lambda=\left(\begin{array}{rr}
1 & 0  \tag{A7}\\
-\theta_{n} & R_{n m}^{-1}
\end{array}\right)
$$

with $R_{n m}^{-1}=\delta_{n m}-\theta_{n m}$, where the induced infinitesimal rotation has parameter $\theta_{m n}$,

$$
\begin{equation*}
\theta_{n m}=\alpha_{n m}-2 \frac{\tan \sqrt{v^{2}}}{\sqrt{v^{2}}}\left(\kappa_{n} v_{m}-\kappa_{m} v_{n}\right) \tag{A8}
\end{equation*}
$$

while the unbroken induced boosts have parameter $\theta_{n}$,

$$
\begin{equation*}
\theta_{n}=\beta_{n}+2\left(\lambda v_{n}-w \kappa_{n}\right) \frac{\tan \sqrt{v^{2}}}{\sqrt{v^{2}}} \tag{A9}
\end{equation*}
$$

are obtained as the $m^{2} \rightarrow 0$ limit of these $A d S-C_{D} \rightarrow A d S-C_{d}$ transformations.

## APPENDIX B: AdS $\xrightarrow{c \rightarrow 0} A d S-C$

The purpose of this appendix is to make the speed of light $c$ explicit in the $A d S_{d+1} \rightarrow A d S_{d}$ isometry algebra and associated coset elements in order to implement the $c \rightarrow 0$ limit directly, reproducing the action of Sections II and III. Returning to the $S O(2, d)$ symmetry algebra for $A d S_{d+1}$, Equation (2.1), where now the $S O(2, d) \rightarrow S O(2, d-1)$ isometry algebra for $A d S_{d+1} \rightarrow A d S_{d}$ is denoted with hatted operators so that

$$
\begin{align*}
P^{M} & =\hat{P}^{M} \text { for } M=0,1,2, \ldots, p \\
P^{p+1} & =-\hat{Z} \tag{B1}
\end{align*}
$$

and

$$
\begin{align*}
M^{M N} & =\hat{M}^{M N} \text { for } M, N=0,1,2, \ldots, p \\
M^{p+1 M} & =\hat{K}^{M} \quad \text { for } M=0,1,2, \ldots, p \tag{B2}
\end{align*}
$$

where now $M, N=0,1,2, \ldots, p$ labelling only the $A d S_{d}$ components while the $(p+1)^{\text {th }}$ components are separated into $\hat{Z}$ and $\hat{K}^{M}$. The $S O(2, d)$ algebra becomes that used in Equation (B.5) of Ref. 6

$$
\begin{align*}
{\left[\hat{M}^{M N}, \hat{M}^{R S}\right] } & =-i\left(\eta^{M R} \hat{M}^{N S}-\eta^{M S} \hat{M}^{N R}+\eta^{N S} \hat{M}^{M R}-\eta^{N R} \hat{M}^{M S}\right) \\
{\left[\hat{M}^{M N}, \hat{P}^{L}\right] } & =i\left(\hat{P}^{M} \eta^{N L}-\hat{P}^{N} \eta^{M L}\right) \\
{\left[\hat{M}^{M N}, \hat{K}^{L}\right] } & =i\left(\hat{K}^{M} \eta^{N L}-\hat{K}^{N} \eta^{M L}\right), \\
{\left[\hat{M}^{M N}, \hat{Z}\right] } & =0 \quad, \quad\left[\hat{P}^{M}, \hat{K}^{N}\right]=i \eta^{M N} \hat{Z} \\
{\left[\hat{P}^{M}, \hat{P}^{N}\right] } & =-i m^{2} \hat{M}^{M N} \quad, \quad\left[\hat{P}^{M}, \hat{Z}\right]=-i m^{2} \hat{K}^{M} \\
{\left[\hat{K}^{M}, \hat{K}^{N}\right] } & =i \hat{M}^{M N} \quad, \quad\left[\hat{Z}, \hat{K}^{M}\right]=i \hat{P}^{M} \tag{B3}
\end{align*}
$$

To make the speed of light explicit introduce the generators

$$
\begin{aligned}
\hat{P}^{0} & =\frac{1}{c} H \\
\hat{P}^{m} & =P^{m} \\
\hat{K}^{0} & =\frac{1}{2 c} L
\end{aligned}
$$

$$
\begin{align*}
\hat{K}^{m} & =\frac{1}{2} K^{m}, \\
\hat{Z} & =-Z, \\
\hat{M}^{m 0} & =\frac{1}{c} B^{m}, \\
\hat{M}^{m n} & =M^{m n}, \tag{B4}
\end{align*}
$$

where the spatial indices are labelled by $m, n=1,2, \ldots, p=(d-1)$. Hence the $S O(2, d)$ algebra of Equation (B3) is as given in Equations (2.5)-(2.7) except for the four commutators involving the explicit factor of the speed of light, which are now

$$
\begin{align*}
& {\left[B^{m}, B^{n}\right]=-i c^{2} M^{m n} \quad, \quad\left[B^{m}, L\right]=i c^{2} K^{m},} \\
& {\left[B^{m}, H\right]=i c^{2} P^{m} \quad, \quad[H, L]=-2 i c^{2} Z .} \tag{B5}
\end{align*}
$$

Define the operators $\mathcal{P}^{M}, \mathcal{M}^{M N}, \mathcal{Z}$, and $\mathcal{K}^{M}$ with the explicit speed of light factors removed as

$$
\begin{align*}
\mathcal{P}^{M} & =\binom{H}{P^{m}}, \\
\mathcal{K}^{M} & =\binom{L}{K^{m}}, \\
\mathcal{Z} & =Z, \\
\mathcal{M}^{M N} & =\left(\begin{array}{cc}
0 & -B^{n} \\
B^{m} & M^{m n}
\end{array}\right) . \tag{B6}
\end{align*}
$$

The relation to the hatted operators is given succinctly by

$$
\begin{align*}
\mathcal{P}^{M} & =C_{N}^{M} \hat{P}^{M} \\
\mathcal{K}^{M} & =2 C_{N}^{M} \hat{K}^{M}, \\
\mathcal{Z} & =-\hat{Z}, \\
\mathcal{M}^{M N} & =C_{R}^{M} \hat{M}^{R S} C_{S}^{N}, \tag{B7}
\end{align*}
$$

with

$$
C_{N}^{M}=\left(\begin{array}{cc}
c & 0  \tag{B8}\\
0 & \delta_{m n}
\end{array}\right)
$$

In terms of these operators, the $S O(2, d)$ algebra of Equation (B3) becomes

$$
\begin{array}{rlrl}
{\left[\mathcal{M}^{M N}, \mathcal{M}^{R S}\right]} & =-i\left(n^{M R} \mathcal{M}^{N S}-n^{M S} \mathcal{M}^{N R}+n^{N S} \mathcal{M}^{M R}-n^{N R} \mathcal{M}^{M S}\right), \\
{\left[\mathcal{M}^{M N}, \mathcal{P}^{L}\right]} & =i\left(\mathcal{P}^{M} n^{N L}-\mathcal{P}^{N} n^{M L}\right), \\
{\left[\mathcal{M}^{M N}, \mathcal{K}^{L}\right]} & =i\left(\mathcal{K}^{M} n^{N L}-\mathcal{K}^{N} n^{M L}\right), \\
{\left[\mathcal{M}^{M N}, \mathcal{Z}\right]} & =0 & & {\left[\mathcal{P}^{M}, \mathcal{K}^{N}\right]=-2 i i^{M N} \mathcal{Z},} \\
{\left[\mathcal{P}^{M}, \mathcal{P}^{N}\right]} & =-i m^{2} \mathcal{M}^{M N} & , & {\left[\mathcal{P}^{M}, \mathcal{Z}\right]=+\frac{i}{2} m^{2} \mathcal{K}^{M},} \\
{\left[\mathcal{K}^{M}, \mathcal{K}^{N}\right]} & =4 i \mathcal{M}^{M N} \quad, & & {\left[\mathcal{Z}, \mathcal{K}^{M}\right]=-2 i \mathcal{P}^{M},} \tag{B9}
\end{array}
$$

where the metric has the form of a $(p+1) \times(p+1)$ diagonal matrix denoted $n^{M N}$

$$
n^{M N} \equiv C_{R}^{M} \eta^{R S} C_{S}^{N}=\left(\begin{array}{cc}
c^{2} & 0  \tag{B10}\\
0 & -\delta_{m n}
\end{array}\right)
$$

Rather than use the coset method directly with this form of the algebra, the Maurer-Cartan one-forms found using the hatted form of the algebra can be converted to one-forms with the explicit powers of $c$ exhibited and then the $c \rightarrow 0$ limit performed. First the coset elements for the
two sets of operators are identified. Consider the coordinates

$$
\begin{equation*}
\hat{x}_{M} \equiv\left(x_{0}, x_{m}\right)=\left(c t, x_{m}\right) \tag{B11}
\end{equation*}
$$

and

$$
\begin{equation*}
X_{M} \equiv\left(t, x_{m}\right) \tag{B12}
\end{equation*}
$$

that is $X_{M}=\hat{x}_{N} C^{-1 N}$. Hence the coset elements

$$
\begin{equation*}
e^{i \hat{X}_{M} \hat{P}^{M}}=e^{i X_{M} \mathcal{P}^{M}} . \tag{B13}
\end{equation*}
$$

Likewise let $\hat{\phi}(\hat{x}) \hat{Z}=\phi(X) \mathcal{Z}$ so that

$$
\begin{equation*}
e^{i \hat{\phi} \hat{Z}}=e^{i \phi Z}, \tag{B14}
\end{equation*}
$$

and $\phi=-\hat{\phi}$. Also define the components of $\hat{v}_{M}$ as

$$
\begin{equation*}
\hat{v}_{M}=\left(\hat{v}_{0}, \hat{v}_{m}\right)=\left(2 c w, 2 v_{m}\right) \tag{B15}
\end{equation*}
$$

and those of $V_{M}$ as

$$
\begin{equation*}
V_{M}=\left(w, v_{m}\right) \tag{B16}
\end{equation*}
$$

Thus $V_{M}=\frac{1}{2} \hat{v}_{N} C^{-1 N}$ so that

$$
\begin{equation*}
e^{i \hat{\vartheta}_{M} \hat{K}^{M}}=e^{i V_{M} \mathcal{K}^{M}} . \tag{B17}
\end{equation*}
$$

Finally equating the unbroken subgroup operators $\hat{\theta}_{M N} \hat{M}^{M N}=\Theta_{M N} \mathcal{M}^{M N}$ with

$$
\begin{equation*}
\Theta_{R S}=C_{R}^{-1 M} \hat{\theta}_{M N} C_{S}^{-1 N} \tag{B18}
\end{equation*}
$$

so that

$$
\Theta_{R S}=\left(\begin{array}{cc}
0 & -\theta_{s}  \tag{B19}\\
\theta_{r} & \theta_{r s}
\end{array}\right)
$$

while

$$
\hat{\theta}_{M N}=\left(\begin{array}{cc}
0 & -\hat{\theta}_{n}  \tag{B20}\\
\hat{\theta}_{m} & \hat{\theta}_{m n}
\end{array}\right)=\left(\begin{array}{cc}
0 & -c \theta_{n} \\
c \theta_{m} & \theta_{m n}
\end{array}\right)
$$

and the subgroup elements are equal

$$
\begin{equation*}
e^{\frac{i}{2} \hat{\theta}_{M N} \hat{M}^{M N}}=e^{\frac{i}{2} \hat{\theta}_{M N} \mathcal{M}^{M N}} . \tag{B21}
\end{equation*}
$$

These coset elements so identified,

$$
\begin{equation*}
\hat{\Omega}=e^{i \hat{x}_{M} \hat{P}^{M}} e^{i \hat{\phi} \hat{Z}} e^{i \hat{v}_{M} \hat{K}^{M}}=\Omega=e^{i X_{M}{ }^{\rho M}} e^{i \phi Z} e^{i V_{M} \mathcal{K}^{M}}, \tag{B22}
\end{equation*}
$$

allow their respective Maurer-Cartan one-forms to be related, recalling the one-forms $\omega=-i \Omega^{-1} d \Omega$ and $\hat{\omega}=-i \hat{\Omega}^{-1} \hat{d} \hat{\Omega}$ with $\hat{d}=d \hat{x}_{M} \frac{\partial}{\partial \hat{x}_{M}}=d t \frac{\partial}{\partial t}+d x_{m} \frac{\partial}{\partial x_{m}}=d X_{M} \frac{\partial}{\partial X_{M}}=d$, the one-forms are equal $\omega=\hat{\omega}$. Expanding them in terms of the generators with tangent space indices $A, B=0,1, \ldots, p$ (recall world indices $M, N=0,1, \ldots, p$ also)

$$
\begin{equation*}
\omega=\omega_{A} \mathcal{P}^{A}+\omega_{\mathcal{Z}} \mathcal{Z}+\omega_{K A} \mathcal{K}^{A}+\frac{1}{2} \omega_{A B} \mathcal{M}^{A B} \tag{B23}
\end{equation*}
$$

and

$$
\begin{equation*}
\hat{\omega}=\hat{\omega}_{A} \hat{P}^{A}+\hat{\omega}_{Z} \hat{Z}+\hat{\omega}_{K A} \hat{K}^{A}+\frac{1}{2} \hat{\omega}_{A B} \hat{M}^{A B}, \tag{B24}
\end{equation*}
$$

and utilizing Equation (B7) the Maurer-Cartan one-forms are related

$$
\begin{aligned}
\omega_{A} & =\hat{\omega}_{B} C_{A}^{-1 B}, \\
\omega z & =-\hat{\omega}_{Z},
\end{aligned}
$$

$$
\begin{align*}
& \omega_{K A}=\frac{1}{2} \hat{\omega}_{K B} C_{A}^{-1 B}, \\
& \omega_{A B}=\hat{\omega}_{C D} C_{A}^{-1 C} C_{B}^{-1 D} . \tag{B25}
\end{align*}
$$

These yield the relations for the component one-forms and the eventual $c \rightarrow 0$ relation to the one-forms of Section II. The explicit factors of $c$ relating the component one-forms are, with $a, b=1,2, \ldots, p$,

$$
\begin{align*}
\omega_{0} & =\frac{1}{c} \hat{\omega}_{0} \\
\omega_{a} & =\hat{\omega}_{a} \\
\omega_{z} & =-\hat{\omega}_{Z} \\
\omega_{K 0} & =\frac{1}{2 c} \hat{\omega}_{K 0} \\
\omega_{K a} & =\frac{1}{2} \hat{\omega}_{K a} \\
\omega_{a 0} & =\frac{1}{c} \hat{\omega}_{a 0} \\
\omega_{a b} & =\hat{\omega}_{a b} \tag{B26}
\end{align*}
$$

The relation to the $A d S-C$ one-forms of Section II is found in the $c \rightarrow 0$ limit of the above, for example, $\omega_{0}=\frac{1}{c} \hat{\omega}_{0} \xrightarrow{c \rightarrow 0} \omega_{H}$.

Similarly for the background one-forms for which $\hat{\bar{\Omega}}=e^{i \hat{x}_{A} \hat{P}^{A}}=\bar{\Omega}=e^{i X_{A} \rho^{A}}$ and so $\bar{\omega}=\hat{\bar{\omega}}$. Expanding in terms of the generators

$$
\begin{equation*}
\bar{\omega}=\bar{\omega}_{A} \mathcal{P}^{A}+\frac{1}{2} \bar{\omega}_{A B} \mathcal{M}^{A B} \tag{B27}
\end{equation*}
$$

and

$$
\begin{equation*}
\hat{\bar{\omega}}=\hat{\bar{\omega}}_{A} \hat{P}^{A}+\frac{1}{2} \hat{\bar{\omega}}_{A B} \hat{M}^{A B} \tag{B28}
\end{equation*}
$$

and using the relations for the one-forms

$$
\begin{align*}
\bar{\omega}_{A} & =\hat{\bar{\omega}}_{B} C_{A}^{-1 B}, \\
\bar{\omega}_{A B} & =\hat{\bar{\omega}}_{C D} C_{A}^{-1 C} C_{B}^{-1 D}, \tag{B29}
\end{align*}
$$

these yield the component background one-form equalities

$$
\begin{align*}
\bar{\omega}_{0} & =\frac{1}{c} \hat{\bar{\omega}}_{0} \\
\bar{\omega}_{a} & =\hat{\bar{\omega}}_{a} \\
\bar{\omega}_{a 0} & =\frac{1}{c} \hat{\bar{\omega}}_{a 0} \\
\bar{\omega}_{a b} & =\hat{\bar{\omega}}_{a b} \tag{B30}
\end{align*}
$$

with the $A d S-C$ background one-forms of Section II found in the $c \rightarrow 0$ limit, for example $\bar{\omega}_{0}=\frac{1}{c} \hat{\bar{\omega}}_{0} \xrightarrow{c \rightarrow 0} \bar{\omega}_{H}$.

Applying these $c$-factor conversions to the Maurer-Cartan one-form $\hat{\omega}_{A}$ found in Equation (2.10) of Ref. 6 for charges defined with upper indices, as is the convention here,

$$
\begin{equation*}
\hat{\omega}_{A}=-\frac{\sinh \sqrt{\hat{v}^{2}}}{\sqrt{\hat{v}^{2}}} \hat{v}_{A} \hat{d} \hat{\phi}+\cosh \sqrt{m^{2} \hat{\phi}^{2}}\left[P_{T A B}(\hat{v})+\cosh \sqrt{\hat{v}^{2}} P_{L A B}(\hat{v})\right] \eta^{B C} \hat{\bar{\omega}}_{C} \tag{B31}
\end{equation*}
$$

with the corresponding background one-form $\hat{\bar{\omega}}_{A}$ of Ref. 6

$$
\begin{equation*}
\hat{\bar{\omega}}_{C}=\left[\frac{\sin \sqrt{m^{2} \hat{x}^{2}}}{\sqrt{m^{2} \hat{x}^{2}}} P_{T C D}(\hat{x})+P_{L C D}(\hat{x})\right] \eta^{D E} d \hat{x}_{E} \tag{B32}
\end{equation*}
$$

yields the resulting $\omega_{A}$ one-form $\omega_{A}=\hat{\omega}_{B} C_{A}^{-1 B}$

$$
\begin{equation*}
\omega_{A}=\frac{\sinh \sqrt{4 V^{2}}}{\sqrt{4 V^{2}}} 2 V_{A} d \phi+\cosh \sqrt{m^{2} \phi^{2}}\left[\mathcal{P}_{T A}^{B}(V)+\cosh \sqrt{4 V^{2}} \mathcal{P}_{L A}^{B}(V)\right] \bar{\omega}_{B}, \tag{B33}
\end{equation*}
$$

with the projection operators

$$
\begin{align*}
& \mathcal{P}_{L A}^{B}(V)=P_{L A C}(V) n^{C B}=\frac{V_{A} V_{C} n^{C B}}{V_{D} n^{D E} V_{E}}, \\
& \mathcal{P}_{T A}^{B}(V)=\delta_{A}^{B}-\mathcal{P}_{L A}^{B}(V), \tag{B34}
\end{align*}
$$

and where $\hat{v}^{2}=\hat{v}_{A} \eta^{A B} \hat{v}_{B}=4 V_{D} n^{D E} V_{E} \equiv 4 V^{2}$. In analogous fashion the background one-form is derived $\bar{\omega}_{A}=\hat{\bar{\omega}}_{B} C_{A}^{-1 B}$

$$
\begin{equation*}
\bar{\omega}_{A}=\left[\frac{\sin \sqrt{m^{2} X^{2}}}{\sqrt{m^{2} X^{2}}} \mathcal{P}_{T A}^{M}(X)+\mathcal{P}_{L A}^{M}(X)\right] d X_{M}, \tag{B35}
\end{equation*}
$$

with $\hat{x}^{2}=\hat{x}_{M} \eta^{M N} \hat{x}_{N}=X_{M} n^{M N} X_{N} \equiv X^{2}$. Likewise from Equation (2.10) of Ref. 6

$$
\begin{equation*}
\hat{\omega}_{Z}=\cosh \sqrt{\hat{v}^{2}}\left[\hat{d} \hat{\phi}-\cosh \sqrt{m^{2} \hat{\phi}^{2}} \hat{\bar{\omega}}_{A} \eta^{A B} \hat{v}_{B} \frac{\tanh \sqrt{\hat{v}^{2}}}{\sqrt{\hat{v}^{2}}}\right], \tag{B36}
\end{equation*}
$$

from which it is found that

$$
\begin{equation*}
\omega_{Z}=-\hat{\omega}_{Z}=\cosh \sqrt{4 V^{2}}\left[d \phi+\cosh \sqrt{m^{2} \phi^{2}} \bar{\omega}_{A} n^{A B} 2 V_{B} \frac{\tanh \sqrt{4 V^{2}}}{\sqrt{4 V^{2}}}\right] . \tag{B37}
\end{equation*}
$$

The vielbein $\mathcal{E}_{A}^{M}$ is defined by relating the covariant differentials $\omega_{A}$ to the coordinate differentials $d X_{M}$

$$
\begin{equation*}
\omega_{A} \equiv d X_{M} \mathcal{E}_{A}^{M} \tag{B38}
\end{equation*}
$$

Likewise

$$
\begin{equation*}
\hat{\omega}_{B} \equiv d \hat{x}_{M} \hat{E}_{A}^{M} \tag{B39}
\end{equation*}
$$

and so the vielbeine are related through the one-forms $\omega_{A}=\hat{\omega}_{B} C_{A}^{-1 B}$ as

$$
\begin{equation*}
\mathcal{E}_{A}^{M}=C_{N}^{M} \hat{E}_{B}^{N} C_{A}^{-1 B} \tag{B40}
\end{equation*}
$$

Similarly for the background one-forms

$$
\begin{align*}
& \bar{\omega}_{A}=d X_{M} \overline{\mathcal{E}}_{A}^{M}, \\
& \hat{\bar{\omega}}_{A}=d \hat{x}_{M} \hat{E}_{A}^{M}, \tag{B41}
\end{align*}
$$

and hence the related vielbeine

$$
\begin{equation*}
\overline{\mathcal{E}}_{A}^{M}=C_{N}^{M} \hat{\bar{E}}_{B}^{N} C_{A}^{-1 B} . \tag{B42}
\end{equation*}
$$

Since the one-forms $\omega_{A}$ and $\bar{\omega}_{A}$ are already obtained, the vielbeine can be read off from their forms. From Equation (B35) the background vielbein $\overline{\mathcal{\delta}}_{A}^{M}$ is seen to be equal to

$$
\begin{equation*}
\overline{\mathcal{E}}_{A}^{M}=\frac{\sin \sqrt{m^{2} X^{2}}}{\sqrt{m^{2} X^{2}}} \boldsymbol{P}_{T A}^{M}(X)+\mathcal{P}_{L A}^{M}(X) . \tag{B43}
\end{equation*}
$$

Equation (B33) with $d \phi=d X_{M} \frac{\partial}{\partial X_{M}} \phi=d X_{M} \partial^{M} \phi$ and $\bar{\omega}_{B}=d X_{M} \overline{\mathcal{E}}_{B}^{M}$ provides the vielbein $\mathcal{E}_{A}^{M}$

$$
\begin{equation*}
\mathcal{E}_{A}^{M}=\frac{\sinh \sqrt{4 V^{2}}}{\sqrt{4 V^{2}}} 2 V_{A} \frac{\partial}{\partial X_{M}} \phi+\cosh \sqrt{m^{2} \phi^{2}}\left[\mathcal{P}_{T A}^{B}(V)+\cosh \sqrt{4 V^{2}} \mathcal{P}_{L A}^{B}(V)\right] \overline{\mathcal{E}}_{B}^{M} . \tag{B44}
\end{equation*}
$$

The speed of light can be taken to zero to obtain the results of Sections II and III. Displaying the component one-forms and vielbeine, it is found for the background one-forms and vielbeine
that

$$
\begin{align*}
\bar{\omega}_{0} & =d t \overline{\mathcal{E}}_{0}^{0}+d x_{m} \overline{\mathcal{\delta}}_{0}^{m} \\
& \xrightarrow{c \rightarrow 0} d t\left(\frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\right)+d x_{m}\left(\frac{t x_{m}}{x^{2}}\left(1-\frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}}\right)\right) \\
& =\bar{\omega}_{H}=d t \bar{e}_{0}^{0}+d x_{m} \bar{e}_{0}^{m}, \\
\bar{\omega}_{a} & =d t \overline{\mathcal{E}}_{a}^{0}+d x_{m} \overline{\mathcal{\delta}}_{a}^{m} \\
& \xrightarrow{c \rightarrow 0} d x_{m}\left(\frac{\sinh \sqrt{m^{2} x^{2}}}{\sqrt{m^{2} x^{2}}} P_{\text {Tam }}(x)+P_{\text {Lam }}(x)\right) \\
& =\bar{\omega}_{P a}=d t \bar{e}_{a}^{0}+d x_{m} \bar{e}_{a}^{m} . \tag{B45}
\end{align*}
$$

Thus the vielbeine components of Section II, Equations (2.12) and (2.14), have been obtained. In short this $c \rightarrow 0$ limit is

$$
\begin{equation*}
\overline{\mathcal{E}}_{A}^{M} \xrightarrow{c \rightarrow 0} \bar{E}_{A}^{M} . \tag{B46}
\end{equation*}
$$

Proceeding in a similar manner for the one-forms and vielbeine their $c \rightarrow 0$ limits are obtained as those of Section II

$$
\begin{align*}
& \omega_{0}=d t \mathcal{E}_{0}^{0}+d x_{m} \mathcal{E}_{0}^{m} \\
& \xrightarrow{c \rightarrow 0} d t\left[\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} 2 \omega \partial^{t} \phi+\cosh \sqrt{m^{2} \phi^{2}} \overline{\mathcal{E}}_{0}^{0}\right] \\
& +d x_{m}\left[\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} 2 w \partial^{m} \phi+\cosh \sqrt{m^{2} \phi^{2}} \overline{\mathcal{E}}_{0}^{m}\right. \\
& \left.+\cosh \sqrt{m^{2} \phi^{2}}\left(\cos \sqrt{4 v^{2}}-1\right) \frac{w v_{b}}{v^{2}} \overline{\mathcal{E}}_{b}^{m}\right] \\
& =\omega_{H}=d t E_{0}^{0}+d x_{m} E_{0}^{m}, \\
& \omega_{a}=d t \mathcal{E}_{a}^{0}+d x_{m} \mathcal{E}_{a}^{m} \\
& \xrightarrow{c \rightarrow 0} d t\left[\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} 2 v_{a} \partial^{t} \phi\right]+d x_{m}\left(\frac{\sin \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}} 2 v_{a} \partial^{m} \phi\right. \\
& \left.+\cosh \sqrt{m^{2} \phi^{2}}\left[P_{T a b}(v)+\cos \sqrt{4 v^{2}} P_{L a b}(v)\right] \overline{\mathcal{E}}_{b}^{m}\right) \\
& =\omega_{P a}=d t E_{a}^{0}+d x_{m} E_{a}^{m} . \tag{B47}
\end{align*}
$$

Thus the vielbeine components of Section II, Equations (2.18)-(2.20), have been obtained. In short this $c \rightarrow 0$ limit is

$$
\begin{equation*}
\mathcal{E}_{A}^{M} \xrightarrow{c \rightarrow 0} E_{A}^{M} . \tag{B48}
\end{equation*}
$$

The $c \rightarrow 0$ limit of the $\phi$ covariant derivatives is found from the $\omega_{\mathcal{Z}}$ one-form Equation (B37)

$$
\begin{align*}
& \omega_{Z} \equiv d t \tilde{\nabla}^{t} \phi+d x_{m} \tilde{\nabla}^{m} \phi \\
& \xrightarrow{c \rightarrow 0} d t\left(\dot{\phi} \cos \sqrt{4 v^{2}}\right) \\
&+d x_{m} \cos \sqrt{4 v^{2}}\left[\partial^{m} \phi-\cosh \sqrt{m^{2} \phi^{2}} 2 v_{a} \bar{e}_{a}^{m} \frac{\tan \sqrt{4 v^{2}}}{\sqrt{4 v^{2}}}\right] \\
&= \omega_{Z}=d t \nabla^{t} \phi+d x_{m} \nabla^{m} \phi, \tag{B49}
\end{align*}
$$

which agree with the $\phi$ covariant derivatives in Section II Equation (2.22). Thus the same $A d S-C$ results are obtained as in the use of the coset method.

The Nambu-Goto vielbein, $\hat{N}_{A}^{B}$, is defined by factoring the background vielbein from $\hat{E}_{A}^{M}$ so that

$$
\begin{equation*}
\hat{E}_{A}^{M}=\hat{\bar{E}}_{B}^{M} \hat{N}_{A}^{B} \tag{B50}
\end{equation*}
$$

and so

$$
\begin{equation*}
\hat{N}_{A}^{B}=\hat{\bar{E}}^{-1 B}{ }_{M}^{B} \hat{E}_{A}^{M} \tag{B51}
\end{equation*}
$$

Correspondingly $\mathcal{E}_{A}^{M}=\overline{\mathcal{E}}_{B}^{M} \mathcal{N}_{A}^{B}$ and $\mathcal{N}_{A}^{B}=\overline{\mathcal{E}}^{-1 B}{ }_{M} \mathcal{E}_{A}^{M}$. Exploiting the relations between $\hat{E}, \hat{\bar{E}}$ and $\mathcal{E}$, $\overline{\mathcal{E}}$, it is obtained that

$$
\begin{equation*}
\mathcal{N}_{A}^{B}=C_{D}^{B} \hat{N}_{C}^{D} C_{A}^{-1 C} . \tag{B52}
\end{equation*}
$$

Thus $\operatorname{det} \mathcal{N}=\operatorname{det} \hat{N}$ and from Equations (B40) and $(\mathrm{B} 42) \operatorname{det} \mathcal{E}=\operatorname{det} \hat{E}$ as well as $\operatorname{det} \overline{\mathcal{E}}=\operatorname{det} \hat{\bar{E}}$. Consequently the $A d S$ invariant action, Equation (3.18) of Ref. 6 rescaled by $-\sigma / c, \Gamma_{A d S} \equiv$ $\int d t d^{p} x \operatorname{det} \hat{E}$, is written in terms of $\operatorname{det} \mathcal{E}$ as

$$
\begin{equation*}
\operatorname{det} \hat{E}=\operatorname{det} \hat{\bar{E}} \operatorname{det} \hat{N}=\operatorname{det} \overline{\mathcal{E}} \operatorname{det} \mathcal{N}=\operatorname{det} \mathcal{E} \tag{B53}
\end{equation*}
$$

Utilizing Equation (3.19) or (3.20) of Ref. 6 for the $\operatorname{det} \hat{N}$

$$
\begin{equation*}
\operatorname{det} \hat{N}=\cosh ^{d} \sqrt{m^{2} \hat{\phi}^{2}} \cosh \sqrt{\hat{v}^{2}}\left[1-\left(\hat{v}_{A} \frac{\tanh \sqrt{\hat{v}^{2}}}{\sqrt{\hat{v}^{2}}}\right)\left(\frac{\hat{\bar{E}}_{M}^{-1 A} \hat{\partial}^{M} \hat{\phi}}{\cosh \sqrt{m^{2} \hat{\phi}^{2}}}\right)\right] \tag{B54}
\end{equation*}
$$

and converting $\hat{\phi}, \hat{v}_{A}$ and $\hat{x}_{M}$ to $\phi, V_{A}$ and $X_{M}$ as well as using the relation

$$
\begin{equation*}
\overline{\mathcal{E}}_{M}^{-1 A}=C_{B}^{A} \hat{\bar{E}}_{N}^{-1 B} C_{M}^{-1 N} \tag{B55}
\end{equation*}
$$

in order to find that

$$
\begin{equation*}
C_{A}^{B} \hat{\bar{E}}_{M}^{-1 A} \frac{\partial}{\partial \hat{x}_{M}} \hat{\phi}=-\overline{\mathcal{E}}^{-1 B} \frac{\partial}{\partial X_{M}} \phi \tag{B56}
\end{equation*}
$$

the $\operatorname{det} \mathcal{N}$ is found

$$
\begin{equation*}
\operatorname{det} \mathcal{N}=\cosh ^{d} \sqrt{m^{2} \phi^{2}} \cosh \sqrt{4 V^{2}}\left[1+\left(2 V_{A} \frac{\tanh \sqrt{4 V^{2}}}{\sqrt{4 V^{2}}}\right)\left(\frac{\overline{\mathcal{E}}_{M}^{-1 A} \partial^{M} \phi}{\cosh \sqrt{m^{2} \phi^{2}}}\right)\right] \tag{B57}
\end{equation*}
$$

Thus the $A d S_{d+1} \rightarrow A d S_{d}$ brane embedded action $\Gamma_{A d S}$ is obtained (note Equation (3.20) of Ref. 6)

$$
\begin{align*}
\Gamma_{A d S}= & \int d^{d} X \operatorname{det} \overline{\mathcal{E}} \operatorname{det} \mathcal{N} \\
= & \int d^{d} X \operatorname{det} \overline{\mathcal{E}} \cosh ^{d} \sqrt{m^{2} \phi^{2}} \cosh \sqrt{4 V^{2}}\left[1+\left(2 V_{A} \frac{\tanh \sqrt{4 V^{2}}}{\sqrt{4 V^{2}}}\right)\right. \\
& \left.\left(\frac{1}{\cosh \sqrt{m^{2} \phi^{2}}} \overline{\mathcal{E}}^{-1 A} \frac{\partial}{\partial X_{M} \phi}\right)\right] \tag{B58}
\end{align*}
$$

in which the explicit factors of $c$ are in the background vielbein $\overline{\mathcal{E}}_{A}^{M}$ and its inverse $\overline{\mathcal{E}}^{-1 A}{ }_{M}$ (note the form of Equation (3.39) in which the $c \rightarrow 0$ limit is already taken). Further, taking the $c \rightarrow 0$ limit, the explicit component fields can be exhibited from Equations (B12) and (B16) to obtain Equation (3.15) of Section III for $\operatorname{det} \mathcal{N} \xrightarrow{c \rightarrow 0} \operatorname{det} N$ and likewise $\operatorname{det} \overline{\mathcal{E}} \xrightarrow{c \rightarrow 0} \operatorname{det} \bar{E}=\bar{e}_{0}^{0} \operatorname{det} \bar{e}_{a}^{m}$. Thus $\Gamma_{A d S} \xrightarrow{c \rightarrow 0} \Gamma_{A d S-C_{D}}$ and Equation (3.16) is obtained.

[^1]${ }^{4}$ E. Bergshoeff, J. Gomis, and L. Parra, "The Symmetries of the Carroll Superparticle," J. Phys. A 49(18), 185402 (2016).
${ }^{5}$ G. Gibbons, K. Hashimoto, and P. Yi, J. High Energy Phys. 2002, 061; e-print arXiv:hep-th/0209034.
${ }^{6}$ T. E. Clark, S. T. Love, M. Nitta, and T. ter Veldhuis, J. Math. Phys. 46, 102304 (2005); e-print arXiv:hep-th/0501241.
${ }^{7}$ T. E. Clark, S. T. Love, M. Nitta, T. ter Veldhuis, and C. Xiong, Phys. Rev. D 76, 105014 (2007); e-print arXiv:hep-th/ 0703179 [HEP-TH].
${ }^{8}$ B. Cardona, J. Gomis, and J. M. Pons, "Dynamics of Carroll Strings," JHEP 1607, 050 (2016).
${ }^{9}$ J. Hartong, J. High Energy Phys. 1508, 069 (2015); e-print arXiv:1505.05011 [hep-th].
${ }^{10}$ C. Duval, G. W. Gibbons, P. A. Horvathy, and P. M. Zhang, Classical Quantum Gravity 31, 085016 (2014); e-print arXiv: 1402.0657 [gr-qc].
${ }^{11}$ S. R. Coleman, J. Wess, and B. Zumino, Phys. Rev. 177, 2239 (1969); C. G. Callan, Jr., S. R. Coleman, J. Wess, and B. Zumino, Phys. Rev. 177, 2247 (1969).
${ }^{12}$ D. V. Volkov, Sov. J. Part. Nucl. 4, 3 (1973); V. I. Ogievetsky, in Proceedings of the X-th Winter School of Theoretical Physics in Karpacz (Wroclaw, 1974), Vol. 1, p. 227.
${ }^{13}$ E. A. Ivanov and V. I. Ogievetsky, Teor. Mat. Fiz. 25, 164 (1975).


[^0]:    ${ }^{\text {a) }}$ e-mail address: clarkt@purdue.edu
    b) e-mail address: terveldhuis@ macalester.edu

[^1]:    ${ }^{1}$ J.-M. Levy-Leblond, Ann. I.H.P.: Phys. Theor. 3, 1 (1965), http://eudml.org/doc/75509.
    ${ }^{2}$ H. Bacry and J. Levy-Leblond, J. Math. Phys. 9, 1605 (1968).
    ${ }^{3}$ E. Bergshoeff, J. Gomis, and G. Longhi, Classical Quantum Gravity 31(20), 205009 (2014); e-print arXiv:1405.2264 [hepth].

