

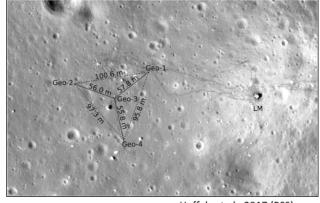
Apollo image AS17-147-Geophone 4 partial pan

Frames 22528 to 22532.

Source: NASA

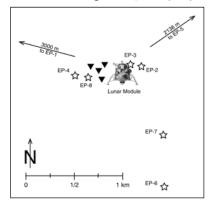
# **Background**

Lunar Seismic Profiling Experiment (LSPE) was part of the Apollo Lunar Surface Experiment Package (ALSEP)



Heffels et al., 2017 (PSS)

Sollberger et al., 2016 (GRL)



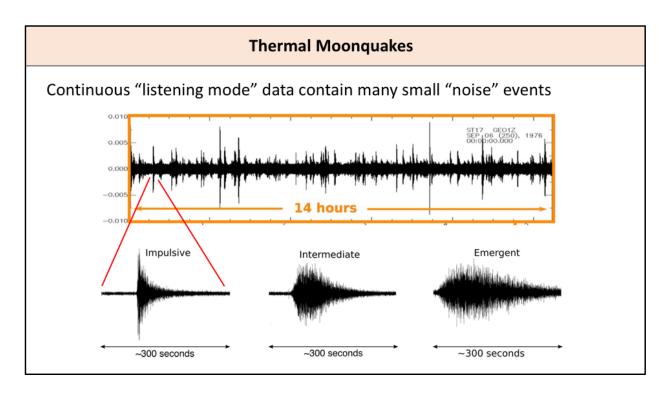
Primary modes of operation:

- 1. Active experiment
- 2. "Listening" mode

Left image: Heffels et al., 2017 "Re-evaluation of Apollo 17 Lunar Seismic Profiling Experiment data." Planetary and Space Science 135, 45-54.

Right image: Sollberger et al., 2016 "The shallow elastic structure of the lunar crust: New insights from seismic wavefield gradient analysis" Geophysical Research Letters 43, 10,078–10,087.

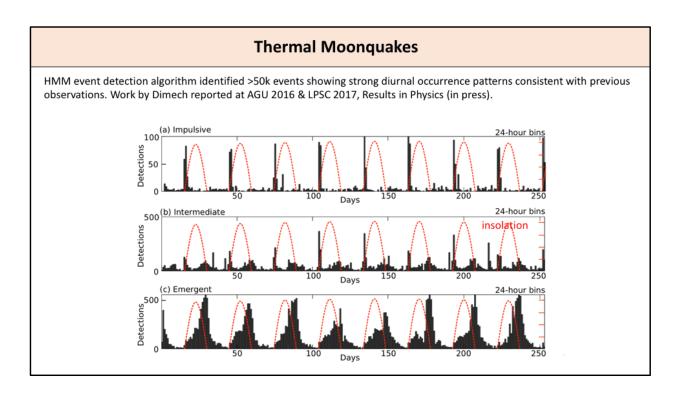
Fair use

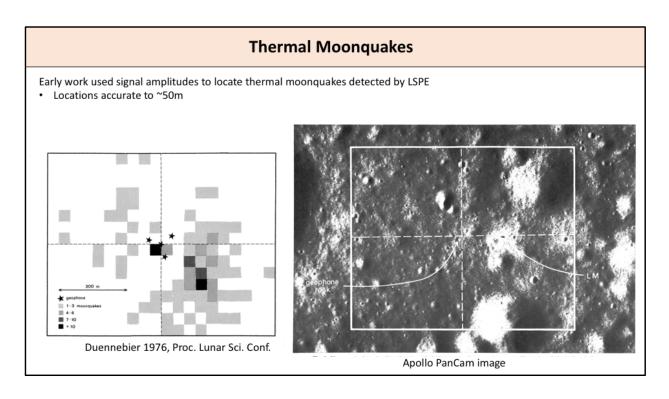


Author-generated images.

Notes: What do we know about them from previous work?

- Observed on all Apollo seismic experiments
- Occur periodically according to diurnal sunrise/sunset cycle
- Have predictable amplitudes and waveforms
  - Original event detection performed only on small portion of data
  - Original event classification performed by eye





Images from Duennebier 1976, Proc. Lunar Sci. Conf. 7<sup>th</sup>, p. 1073-1086 "Thermal Movement of the Regolith" Fair use

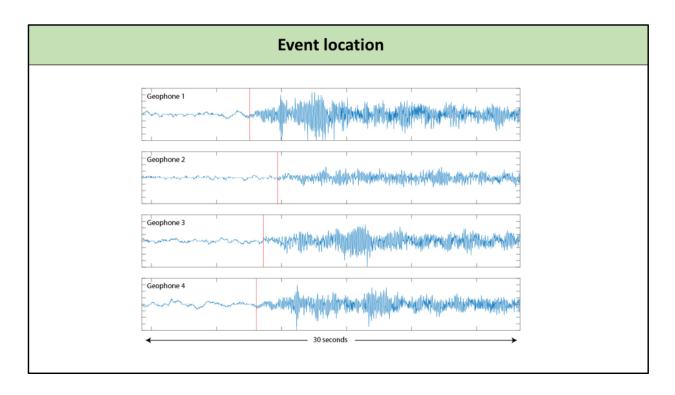
Apollo PanCam image AS 17-2309. Source: NASA

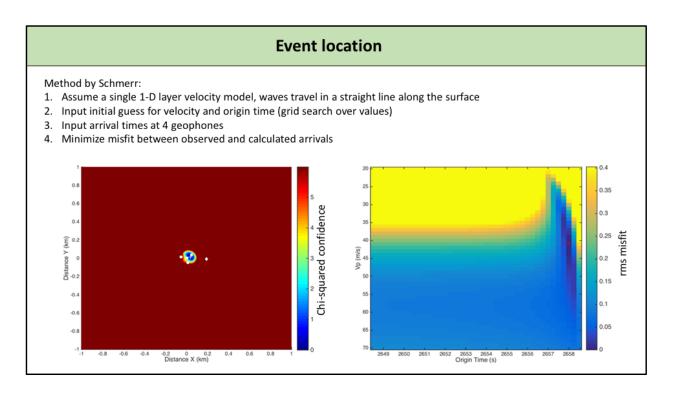
# LROC image Control of the control o

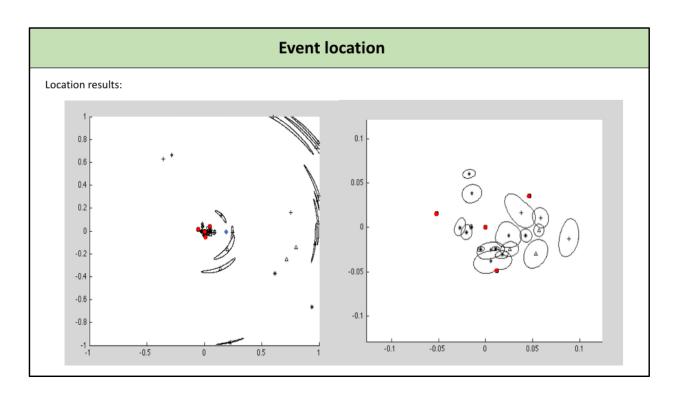
LROC image. Source: NASA

Click image: Image from Duennebier 1976, Proc. Lunar Sci. Conf. 7<sup>th</sup>, p. 1073-1086 "Thermal Movement of the Regolith" overlaid on LROC image. Source: NASA

Fair use







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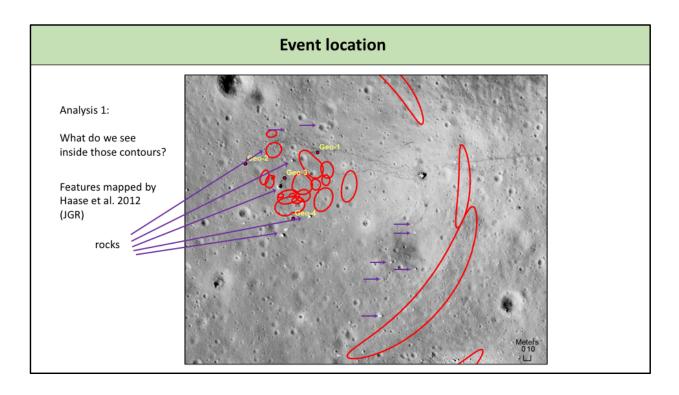
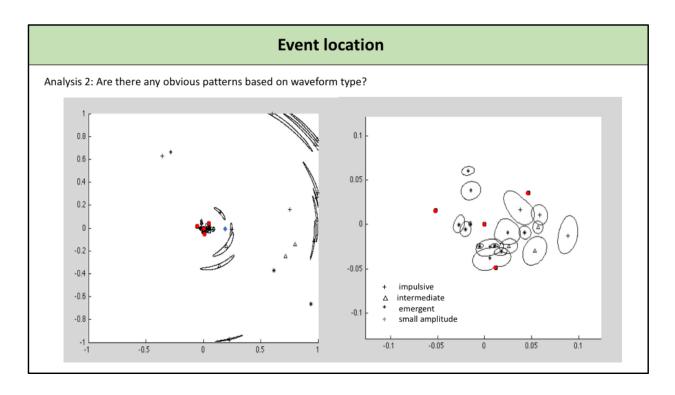
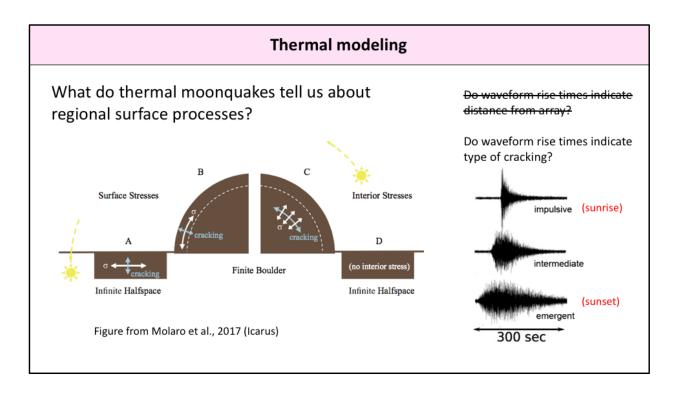


image: LROC imagery (source: NASA) with author generated points and contours Rocks mapped by Haase et al. (2012) "Mapping the Apollo 17 landing site area based on Lunar Reconnaissance Orbiter Camera images and Apollo surface photography". Journal of Geophysical Research Vol. 117. Fair use





Left: Figure from Molaro et al., 2017 (Icarus) "Thermally induced stresses in boulders on airless body surfaces, and implications for rock breakdown" Volume 294, Pages 247-261. Fair use

Right: author-generated

## **Conclusions & future work**

### Conclusions from pilot study:

- Thermal moonquake arrival times can be used to improve precision of recovered locations
- 2. Waveform type is likely indicative of failure type rather than distance from array
- Distribution suggests diffuse events throughout the lunar regolith, but because of large separations in the delay times between events, we can't rule out contributions from rocks and boulders

### Future work:

- 1. Map pick uncertainty onto location uncertainty
- Perform spatial and statistical analysis on larger event sample over a full lunation. Attempt to match event types and locations to features, time of day, temperature, and stress state.
- 3. Perform quantitative assessment of the contribution of thermal moonquakes to regolith production

