

Large stationary gravity waves: a game changer for Venus' science

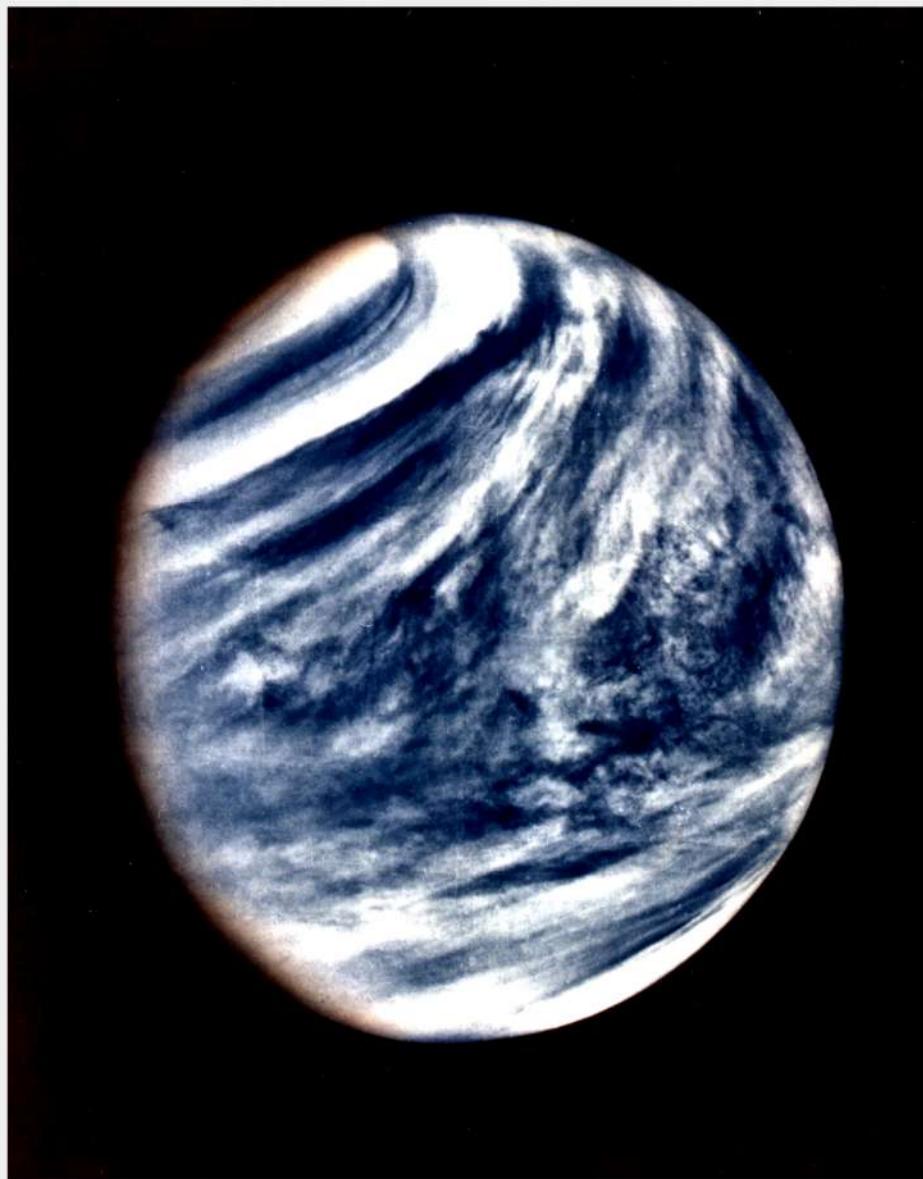
Thomas Navarro

University of California, Los Angeles

with G. Schubert & S. Lebonnois

Venera-D workshop

Moscow, October 2017



VENUS FACTS

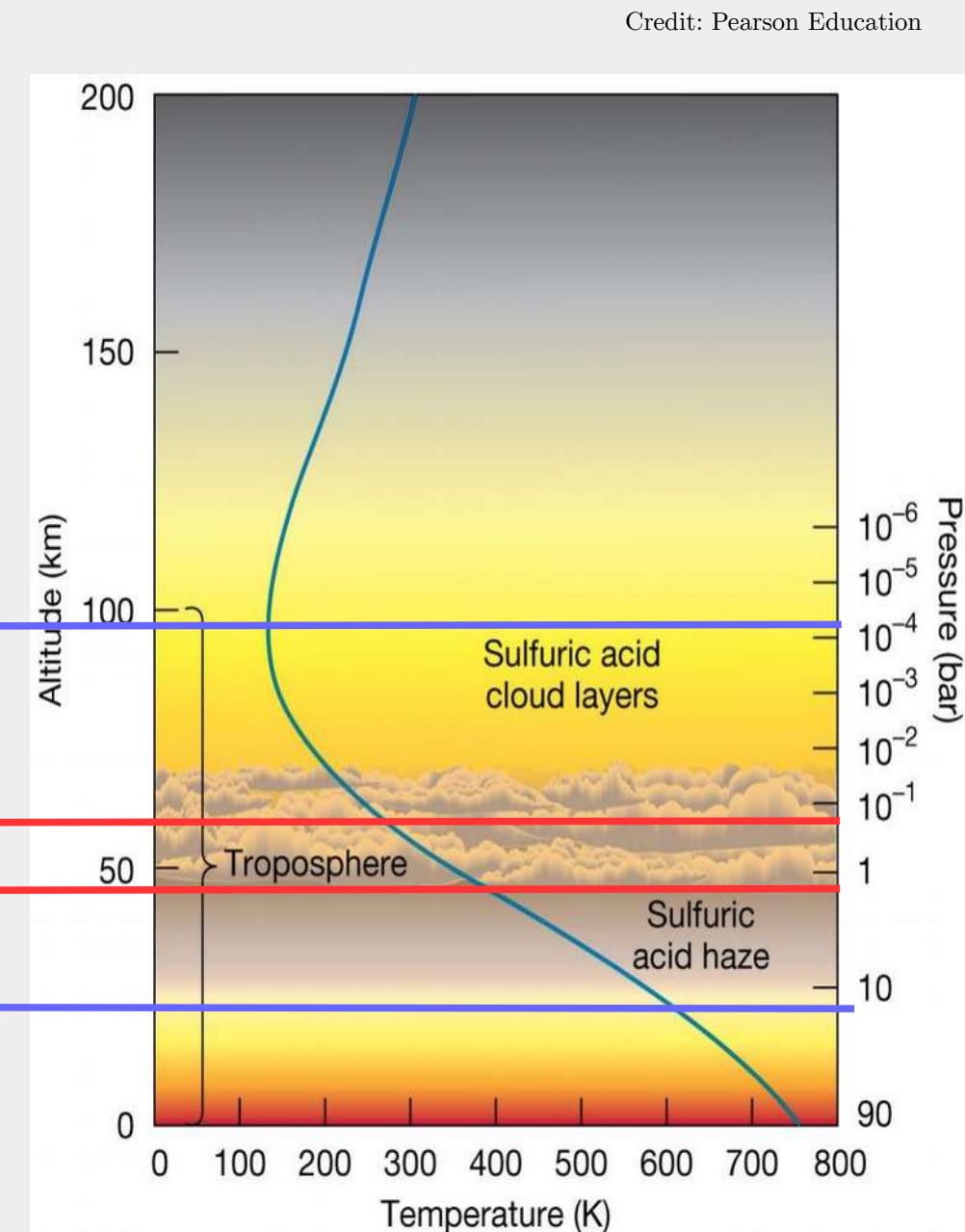
Mass	0,815 M \oplus
Radius	0,95 R \oplus
Obliquity	177,4°
Year	224 days
Sidereal day	243 days
Solar Day	117 days
Super-rotation	4 days at 60 km
Interior	???

Credit: Mariner 10/NASA

COMPOSITION

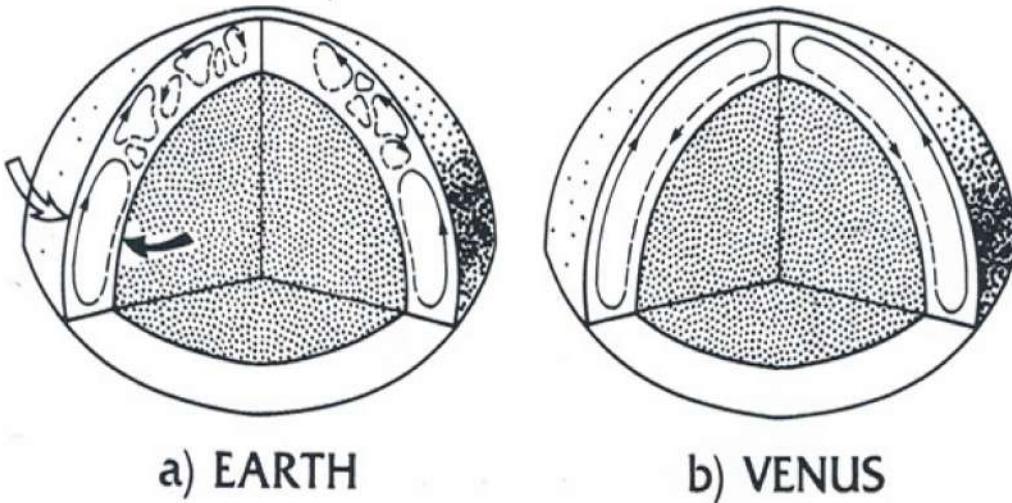
<i>Species</i>	<i>Relative</i>	<i>Compared to Earth</i>
CO₂	96,5 %	x 190,000
N₂	3,5 %	x 4
SO₂	150 ppm	x 500,000
Ar	70 ppm	x 0.6
H₂O	20 ppm	x 0.7

Super-rotating atmosphere
Convective Layer

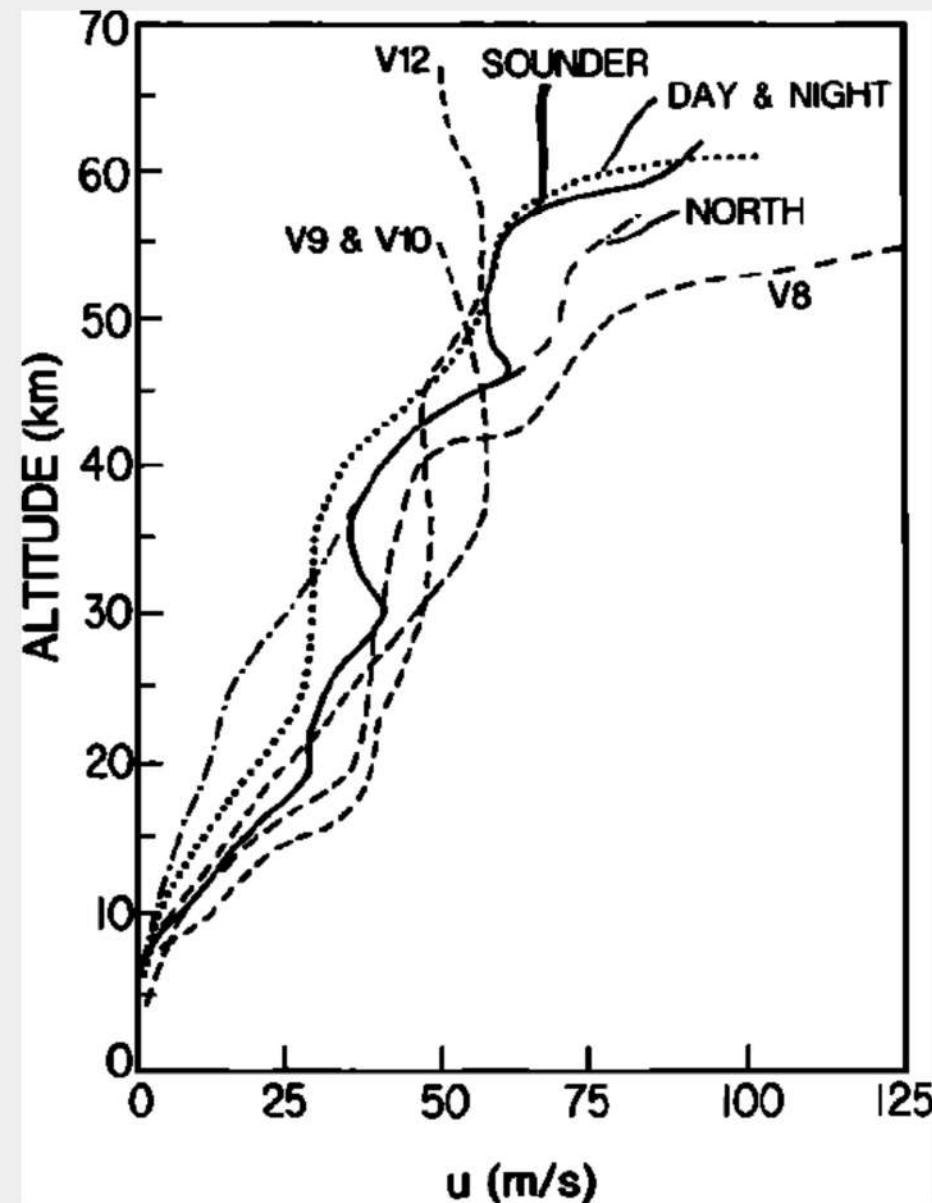


- Super-rotation :

- General circulation



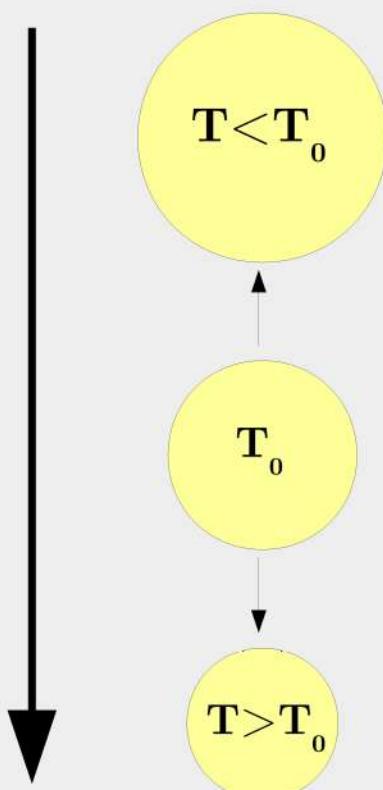
- Diurnal tide



From Schubert et al., 1980

Stability of the atmosphere

Pressure,
Temperature

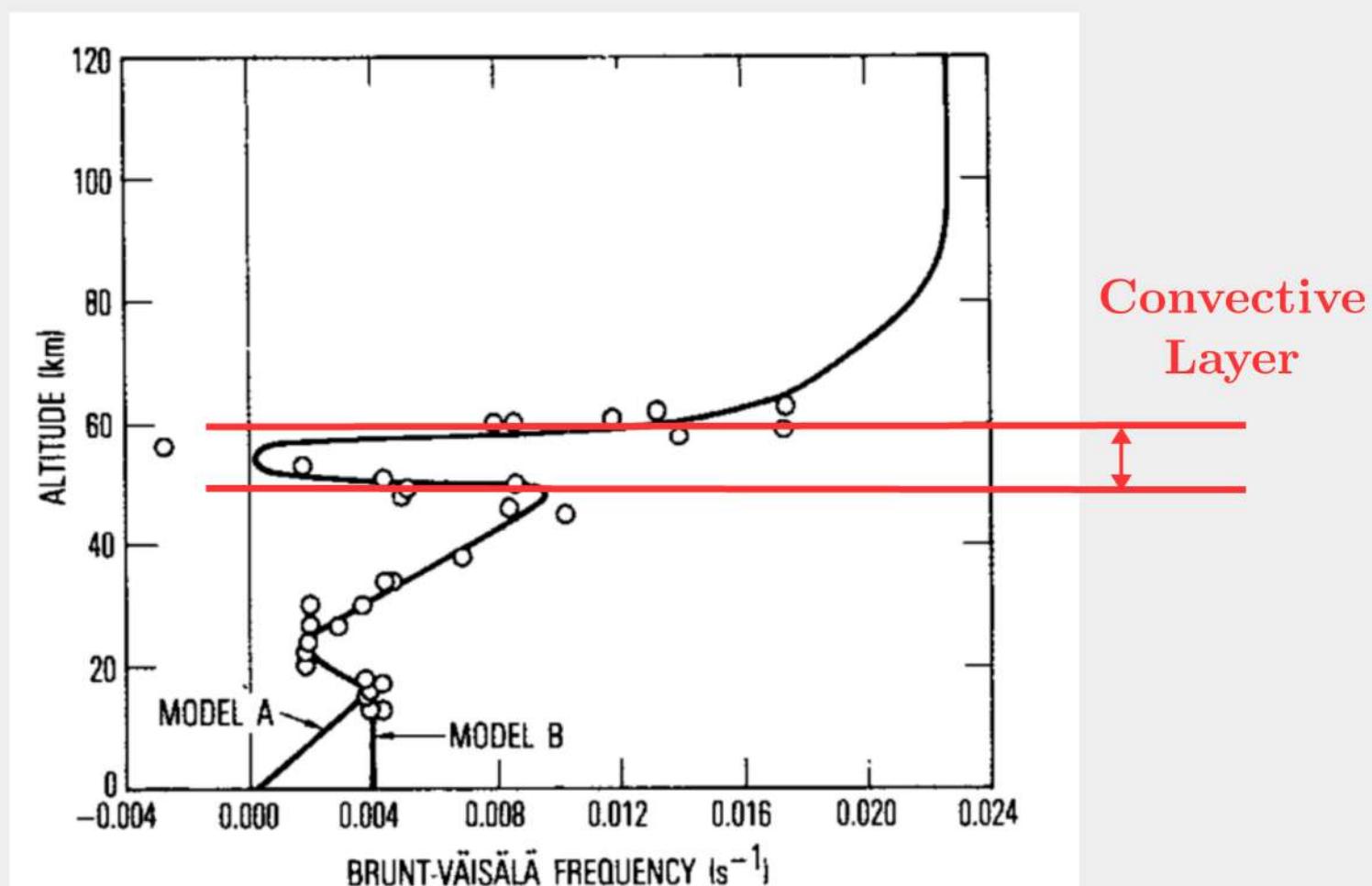


Potential Temperature

$$\theta = T \left(\frac{P_0}{P} \right)^{\frac{R}{C_p}}$$

Brunt-Vaisala Frequency

$$N^2 = \frac{g}{\theta} \frac{d\theta}{dz}$$

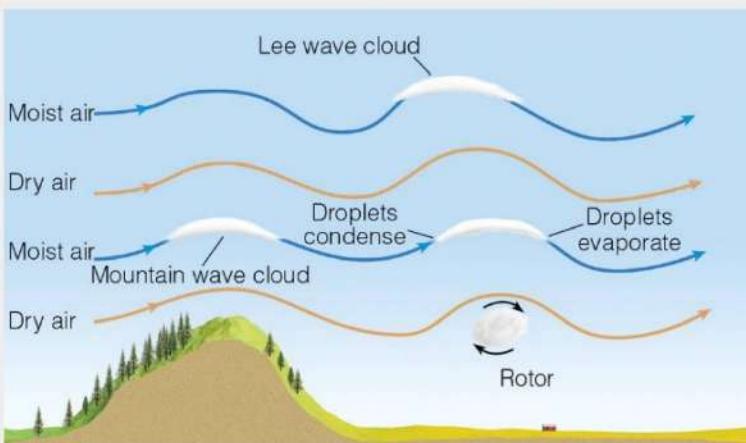


From Schubert et al., 1984

Gravity waves



Credit: WeatherFlow

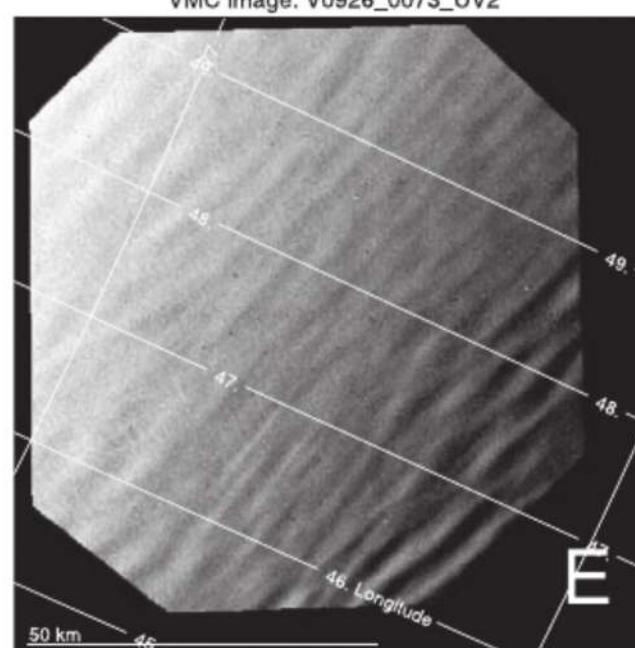
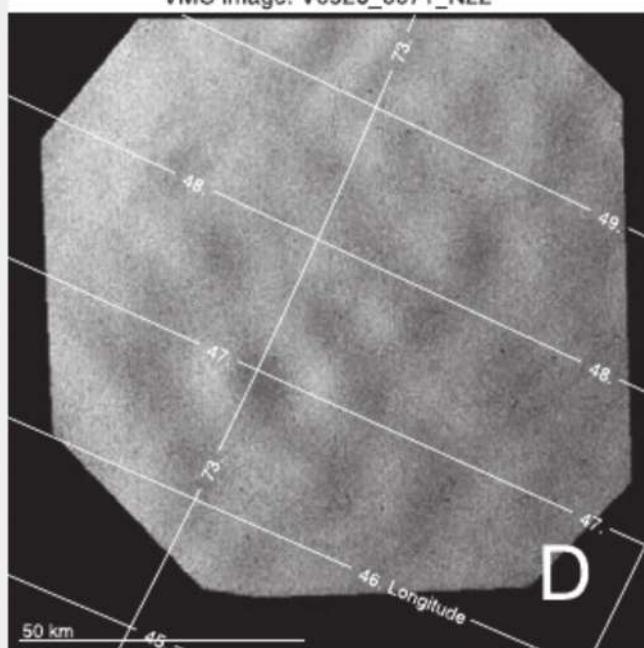
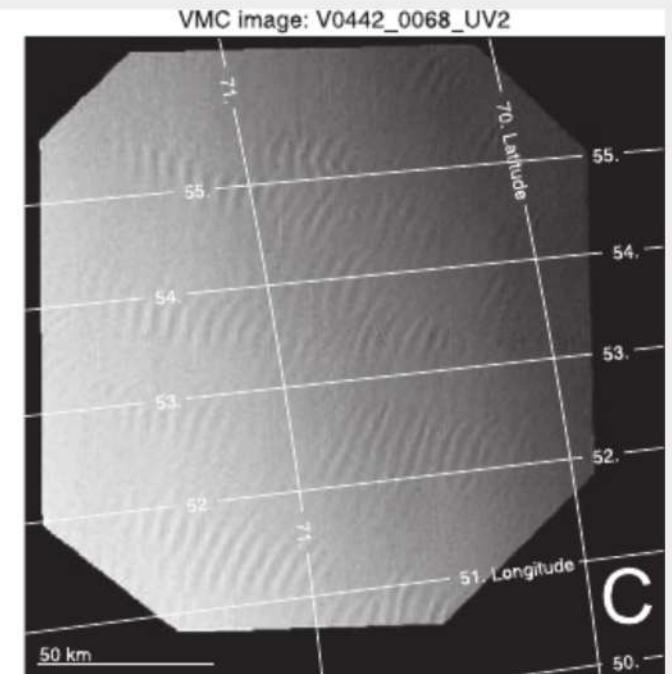
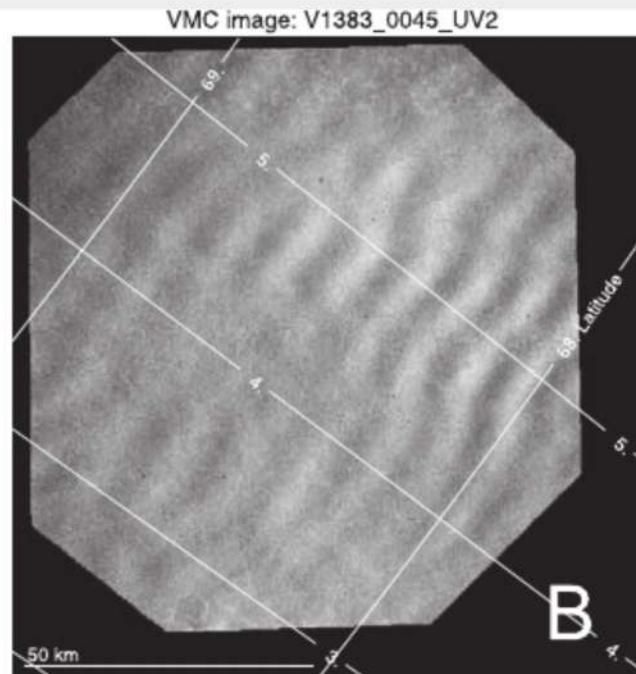
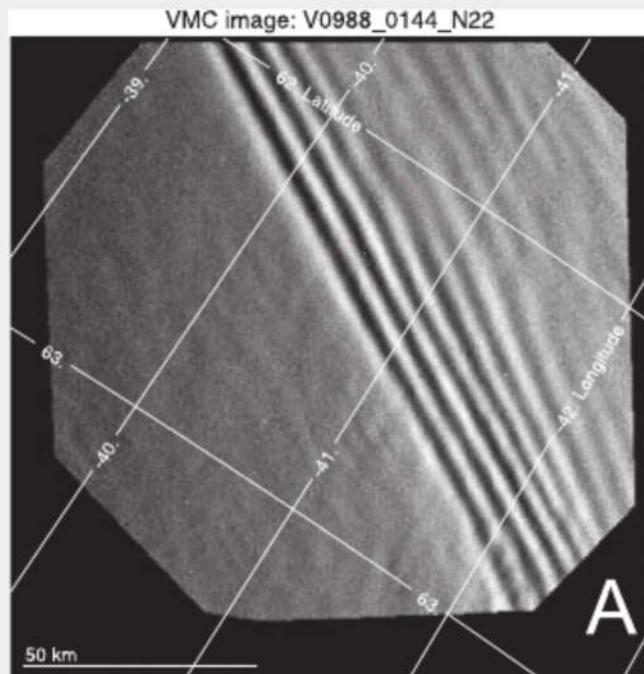


Credit: Thomson Higher Education



Credit: NASA Earth Observatory

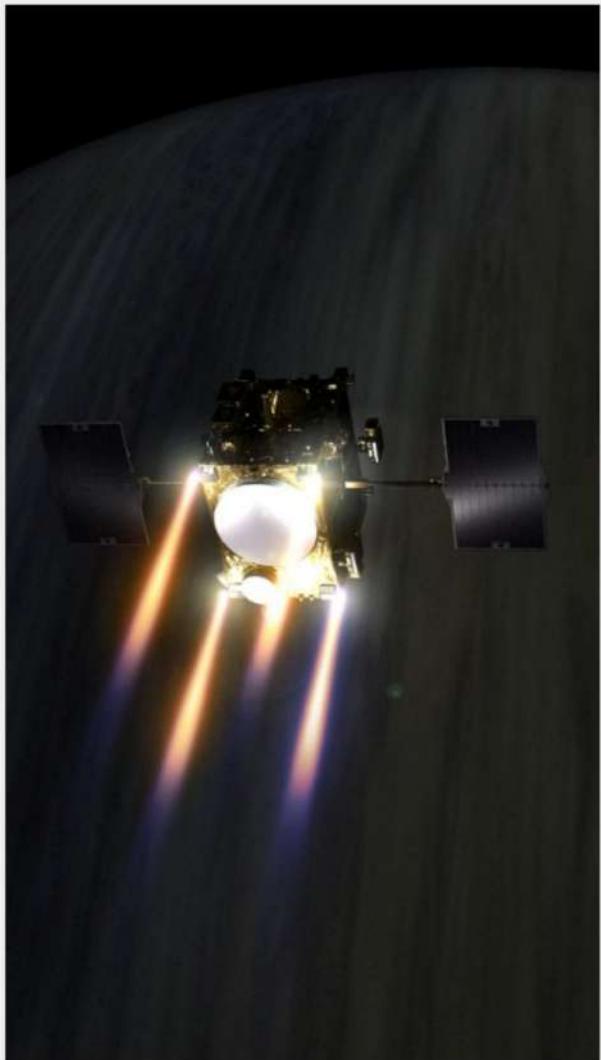
Gravity waves



50 km

Venus Express
observations from
Piccialli et al., 2014

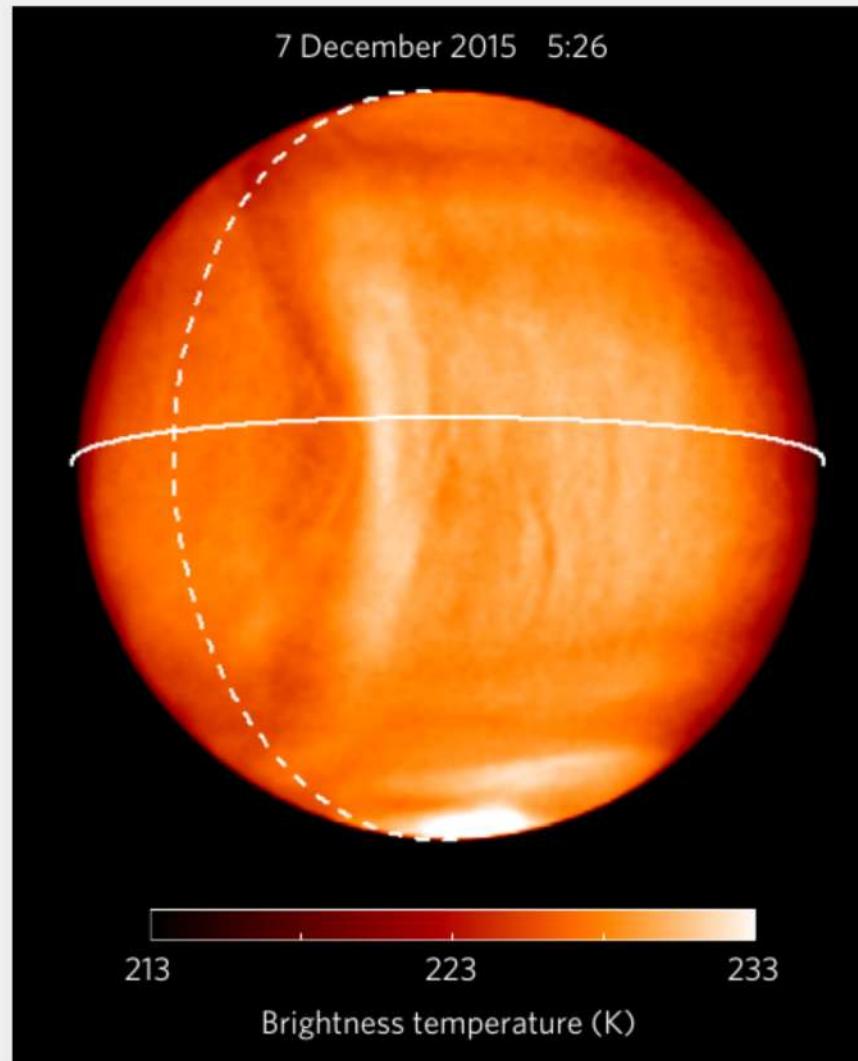
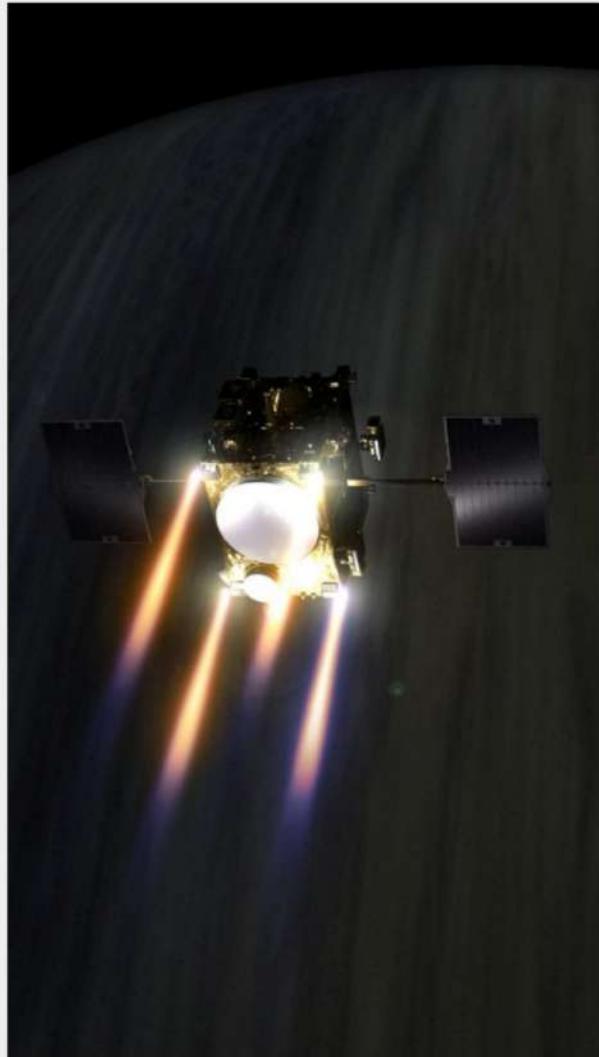
Akatsuki spacecraft



Credit: Go Miyazaki

Akatsuki spacecraft

From Fukuhara et al., 2017

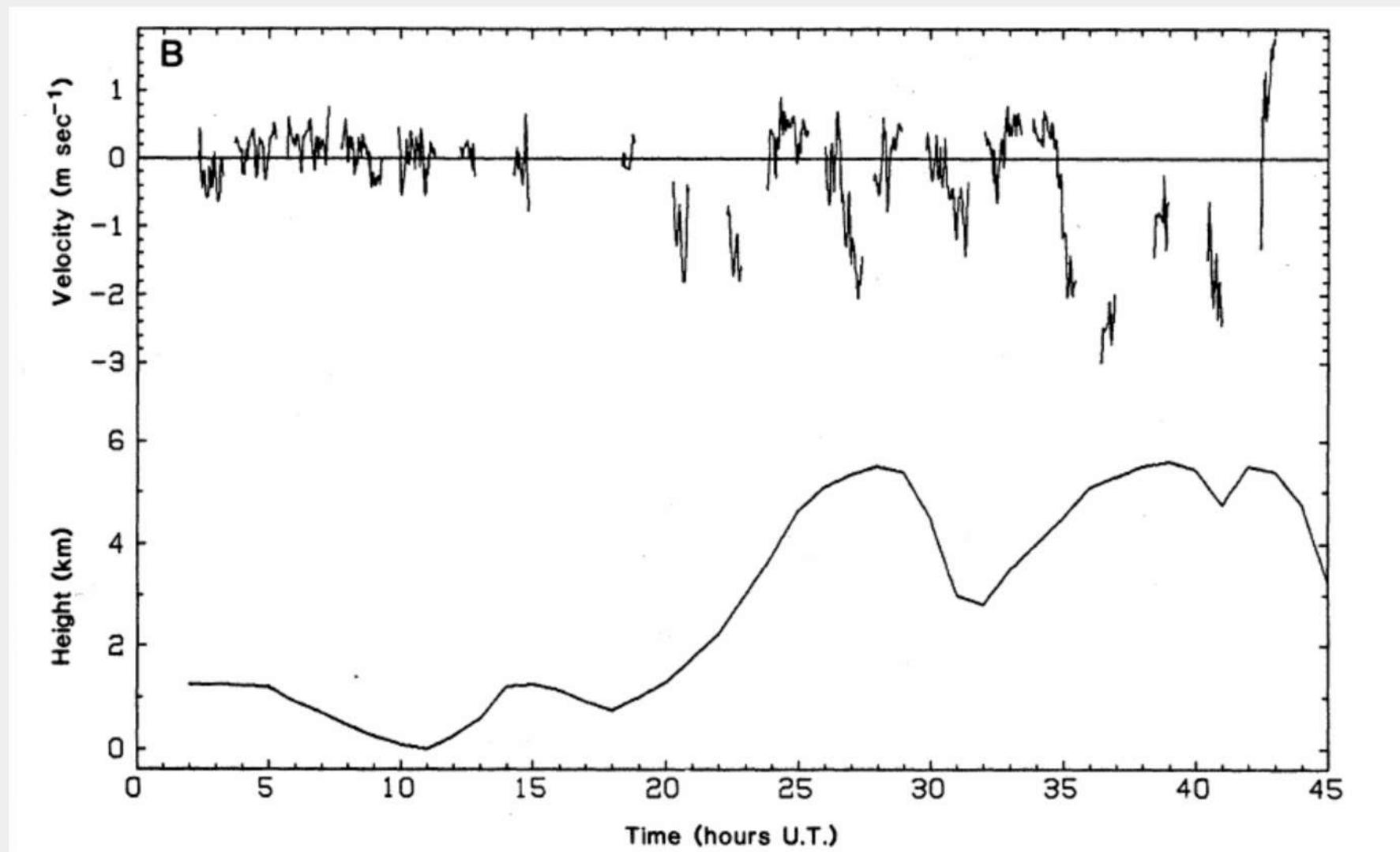


Credit: Go Miyazaki

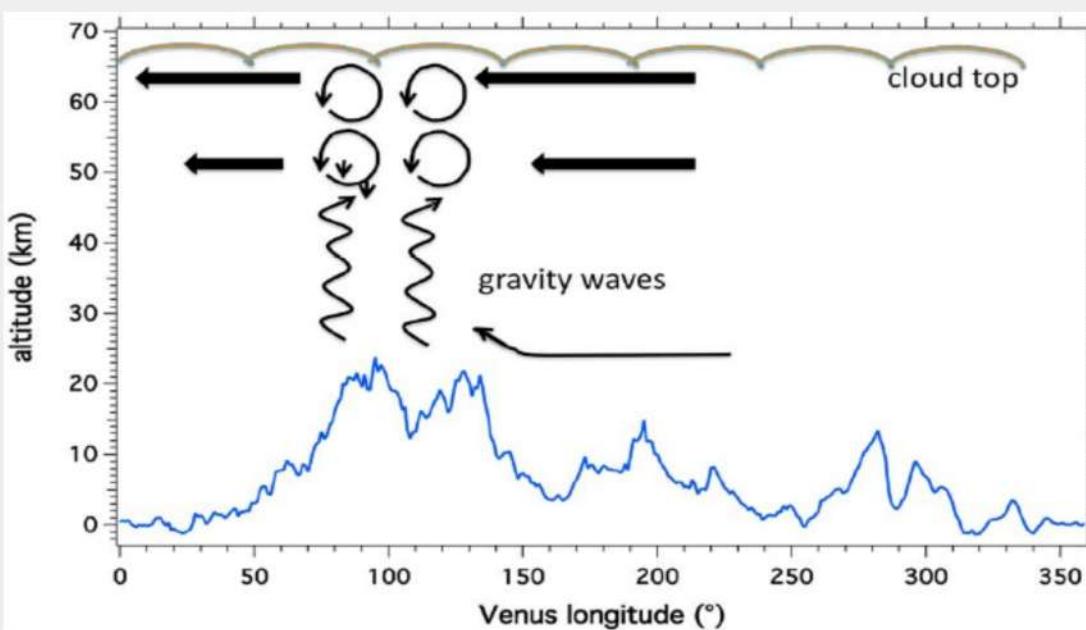
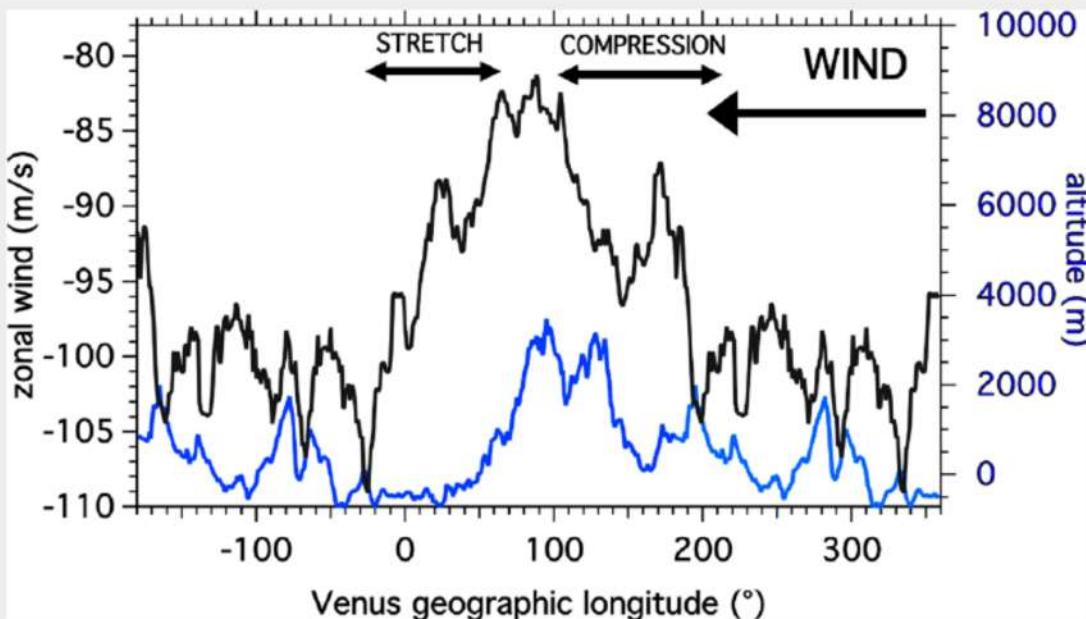
A stationary bow feature
at 70 km altitude!

VEGA Balloon

Blamont et al., 1986

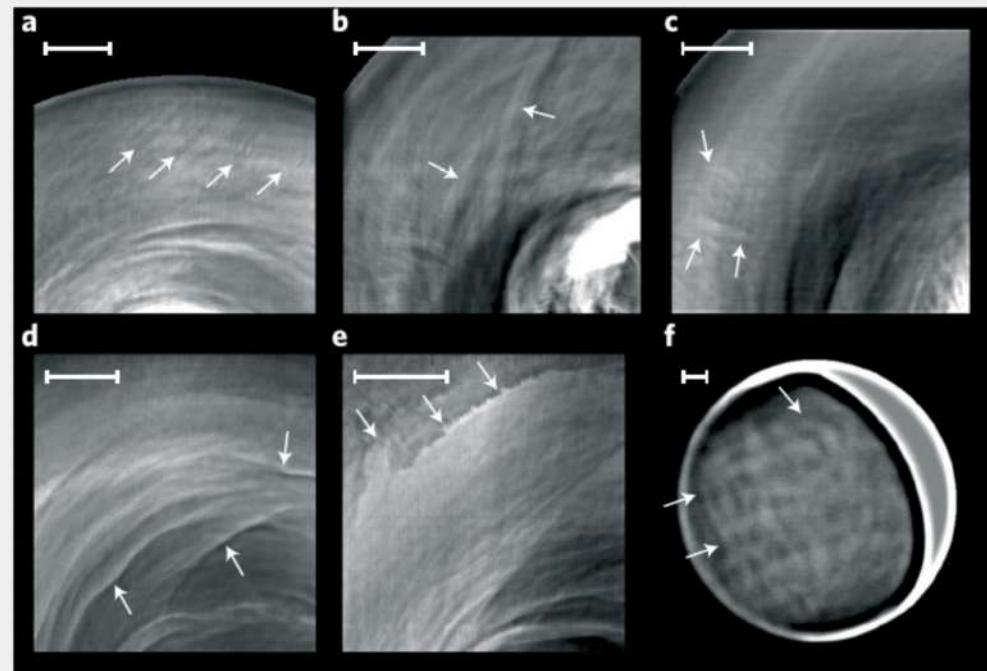
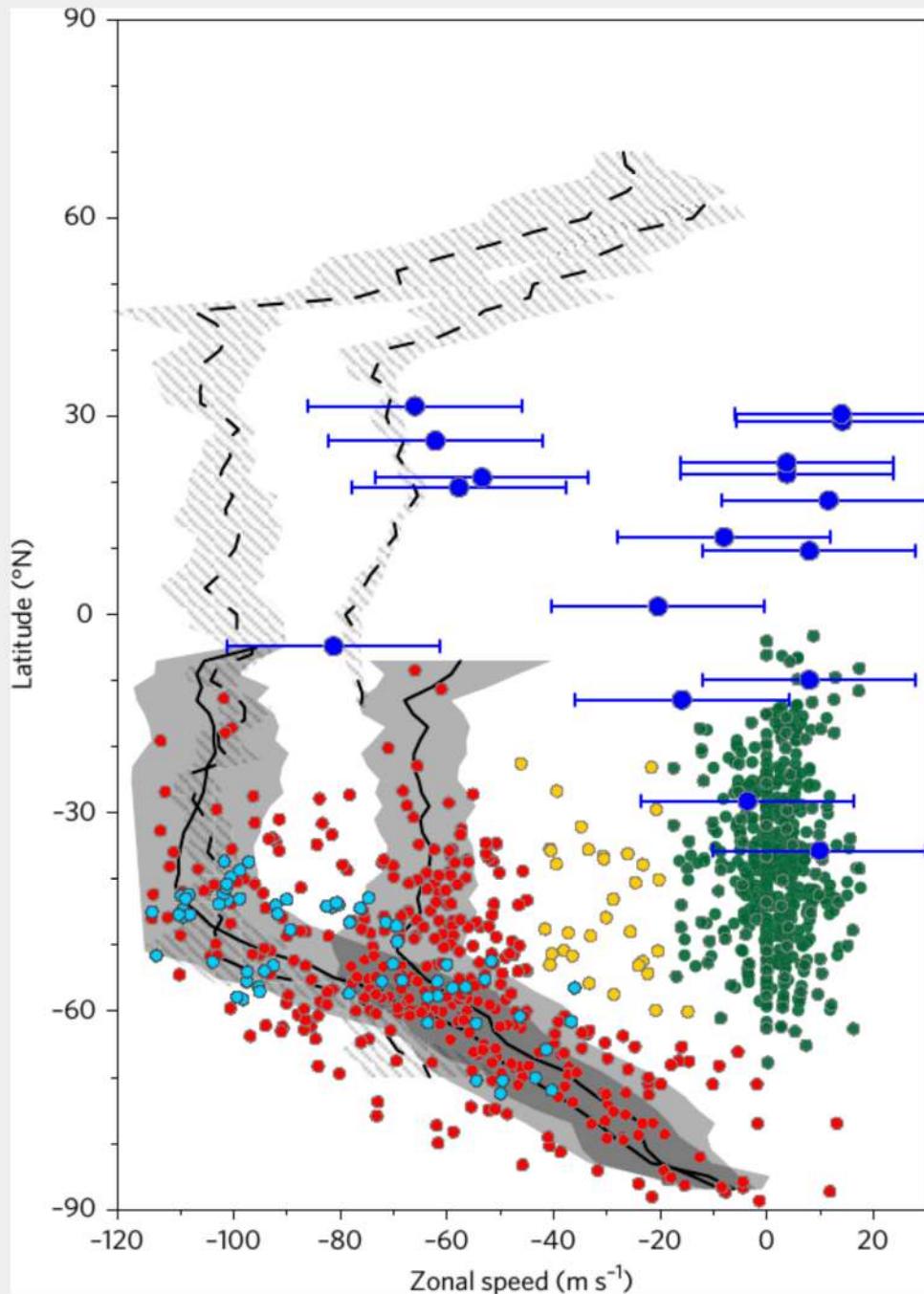


Venus Express winds

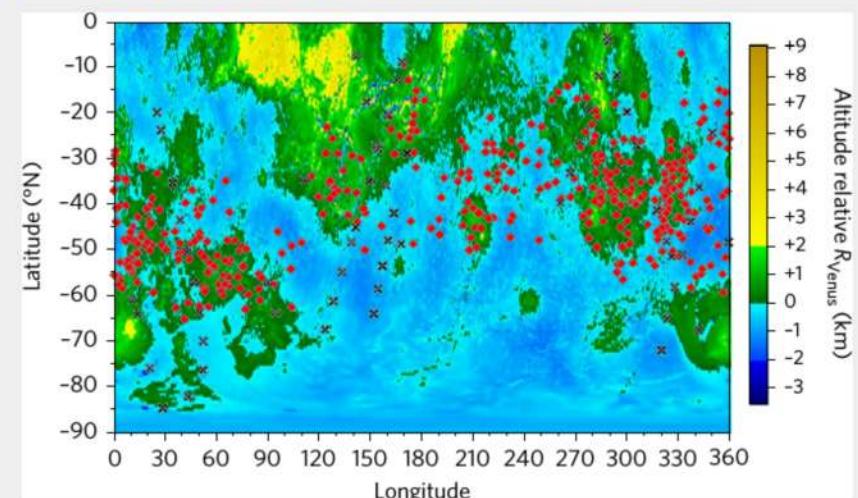


- Winds from tracking of UV images of the Venus Monitoring Camera
- There is a zonal anomaly in the averaged climatology of winds, above Aphrodite.

Venus Express features (Peralta et al. 2017)



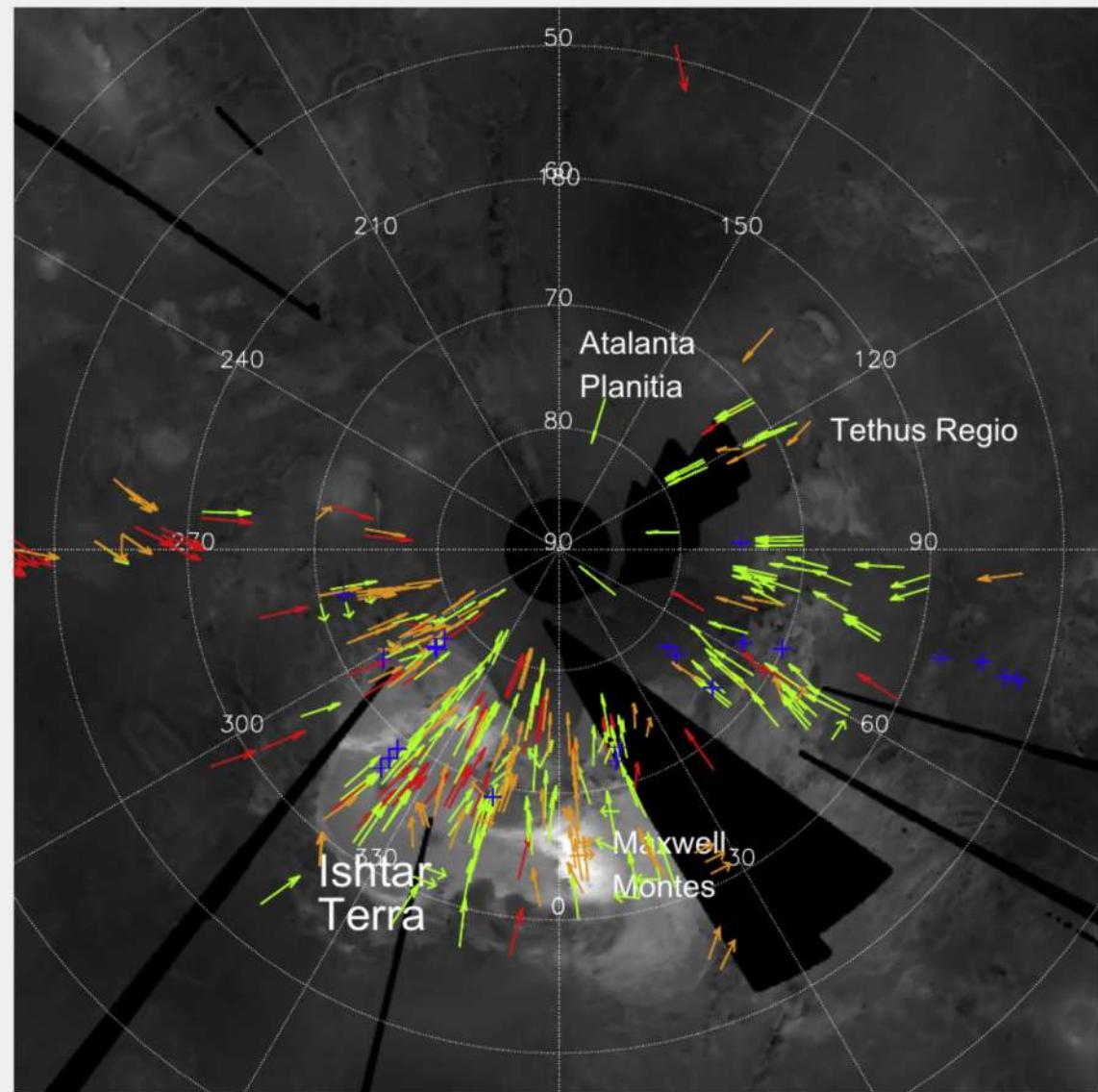
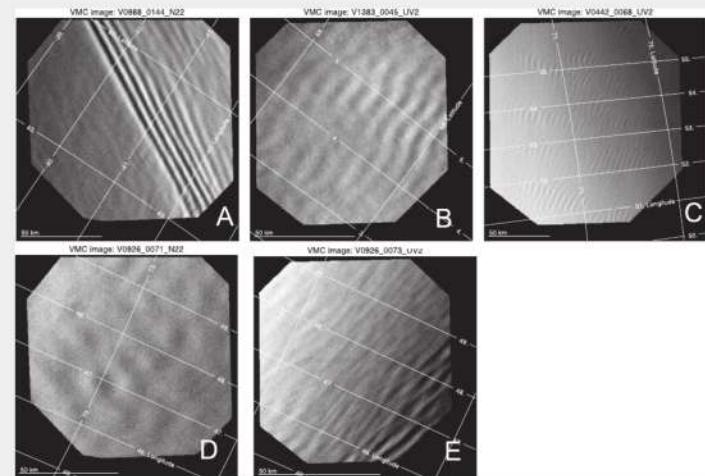
(a) "Wavy patterns" from VIRTIS
(f) "Bow-shaped" features from IRTF



Venus Express features

Green: long waves

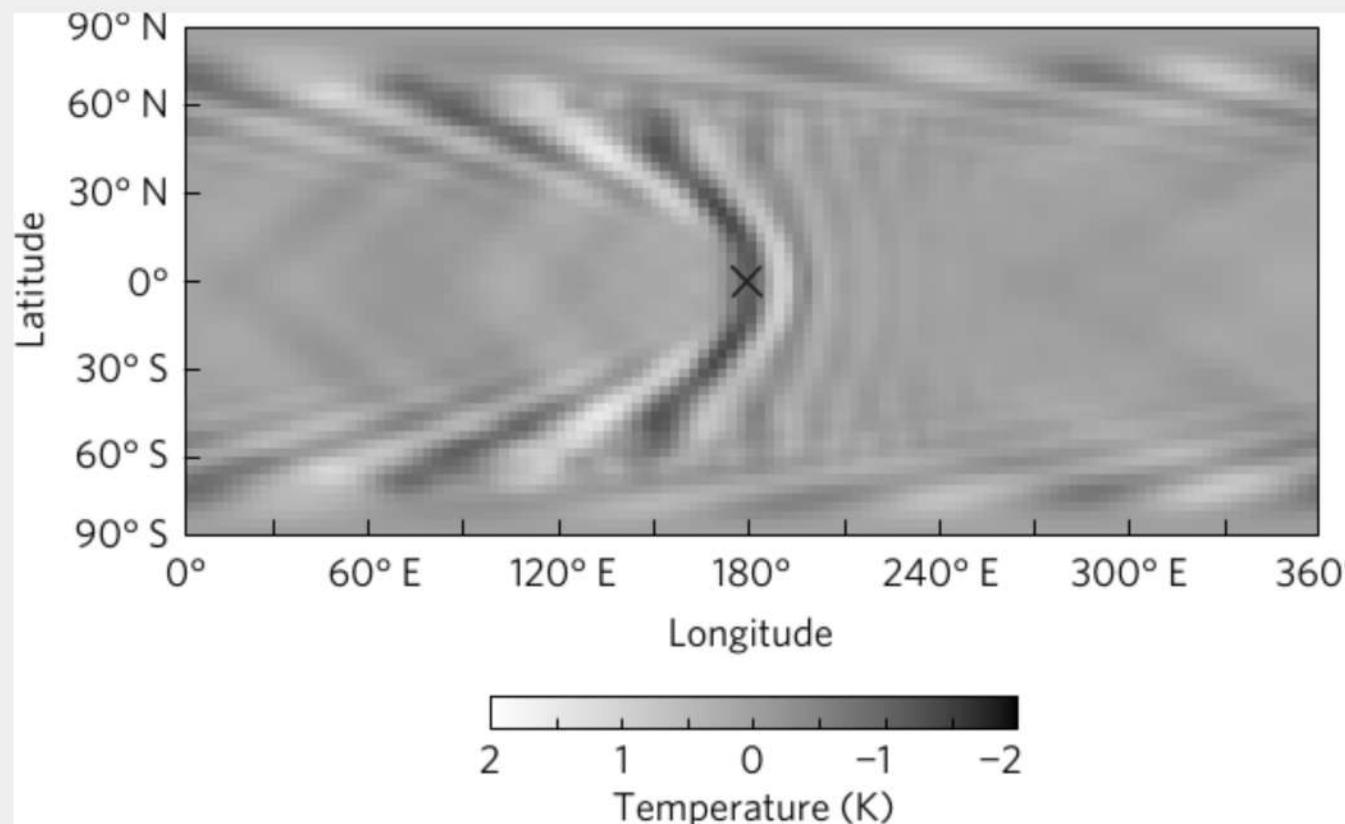
Red: short waves



From Piccialli et al., 2014

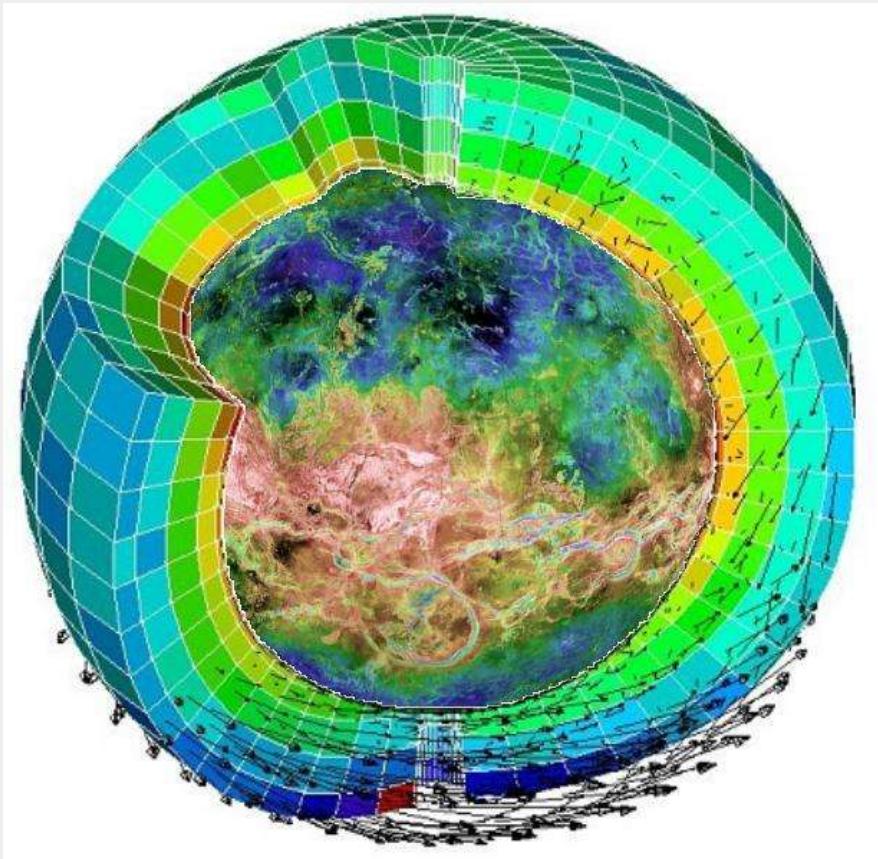
Numerical modelling

From Fukuhara et al., 2017



- Numerical simulation of the atmosphere with a perturbation at 10 km of altitude
- Very crude model: no diurnal cycle, imposed winds and temperature, no topography

LMD Model : Building a virtual planet

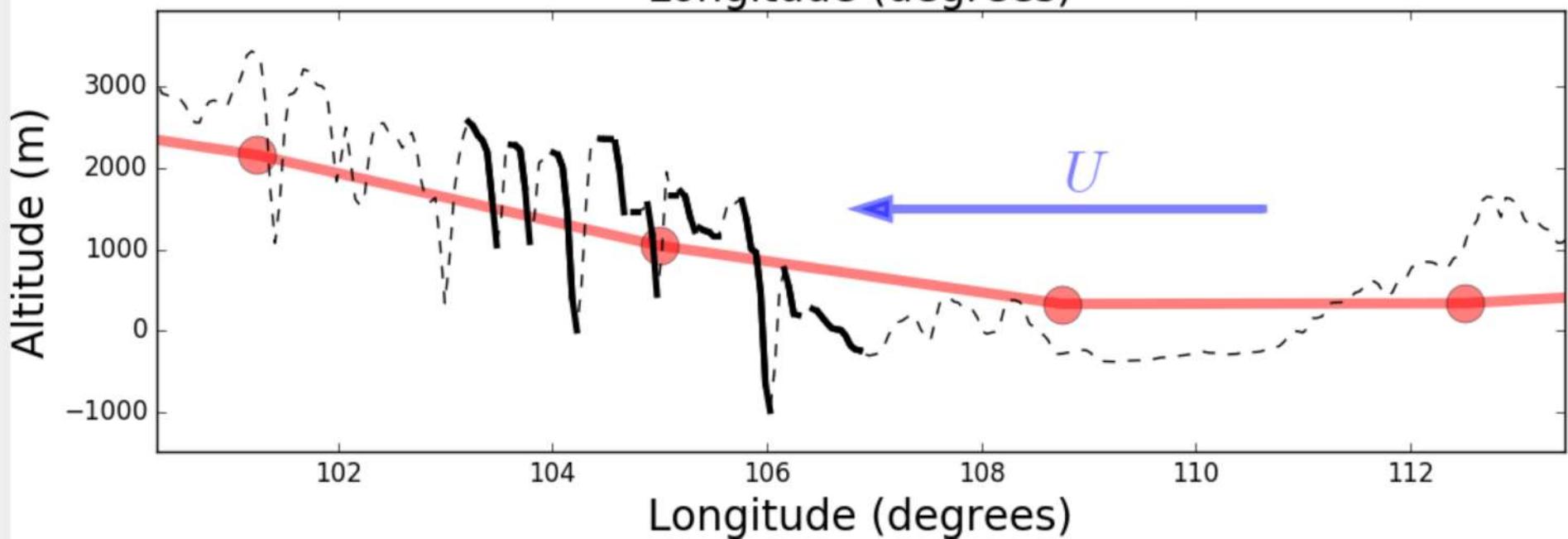
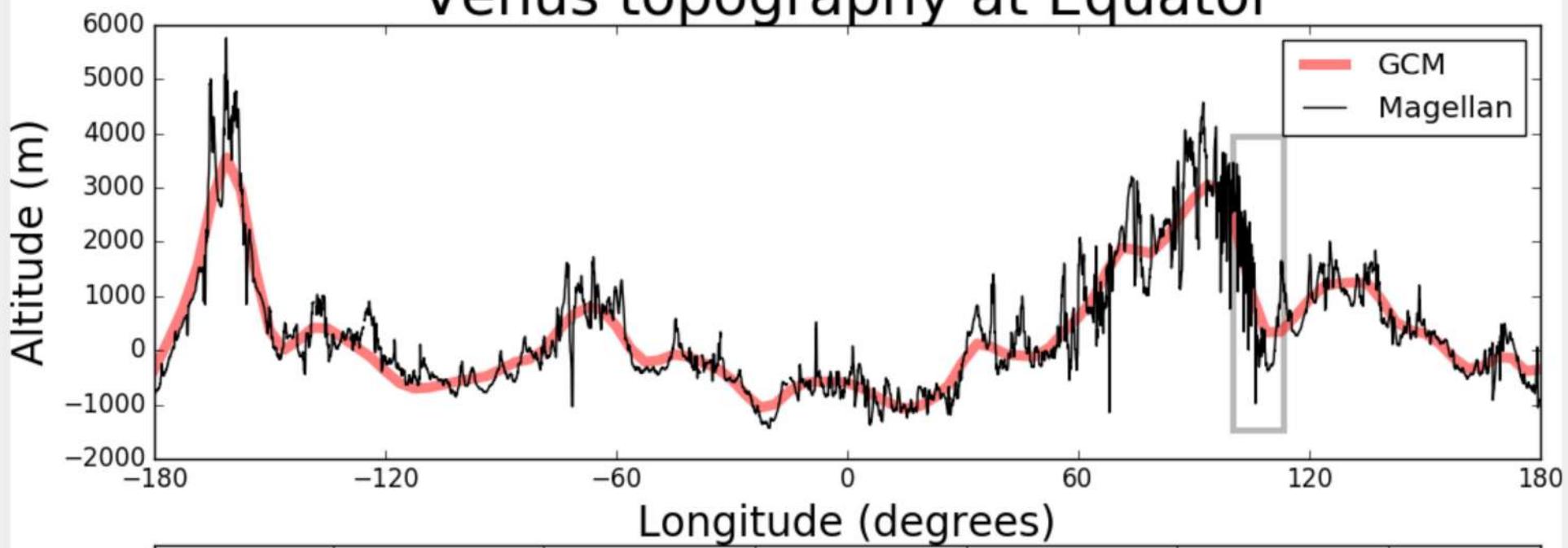


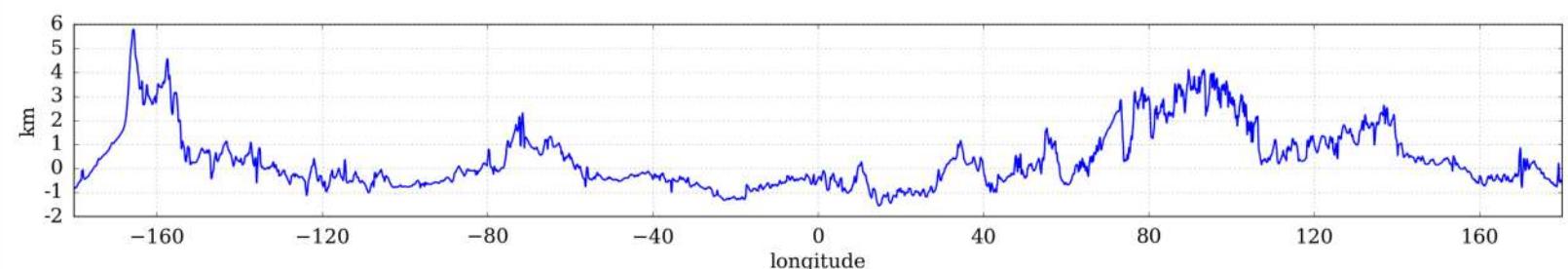
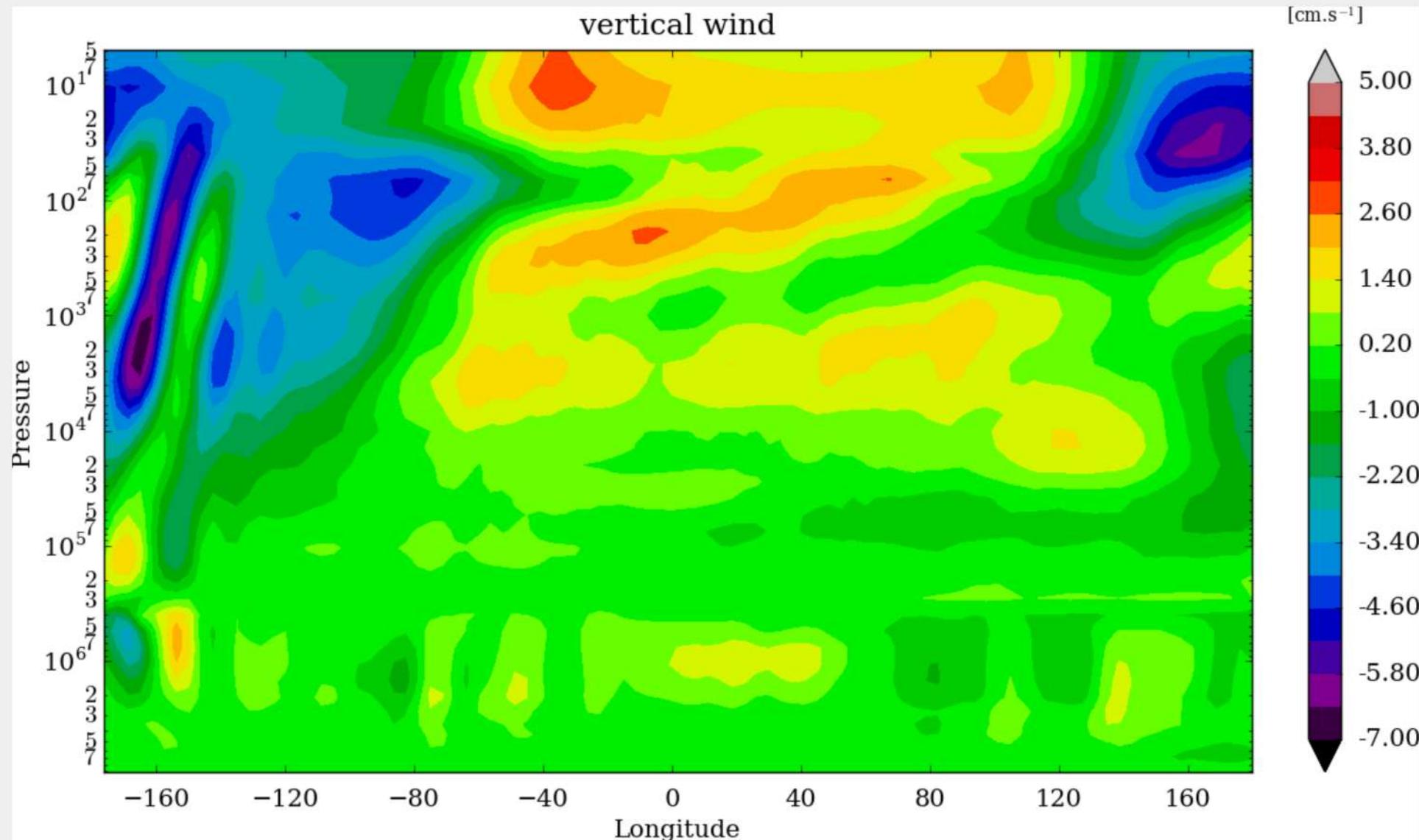
- Dynamical core
 - Radiative transfer
 - Convection
 - [Aerosols & clouds]
 - Interaction w/ surface
- + Modeler's tricks:
- Dissipation
 - Sponge layer

Subgrid parameterization

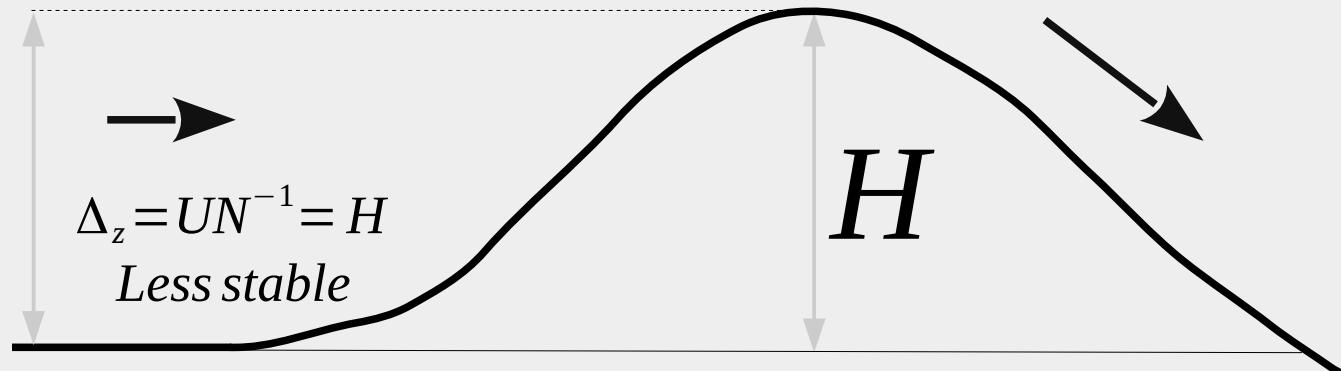
(Lott et al. 97 for Earth)

Venus topography at Equator

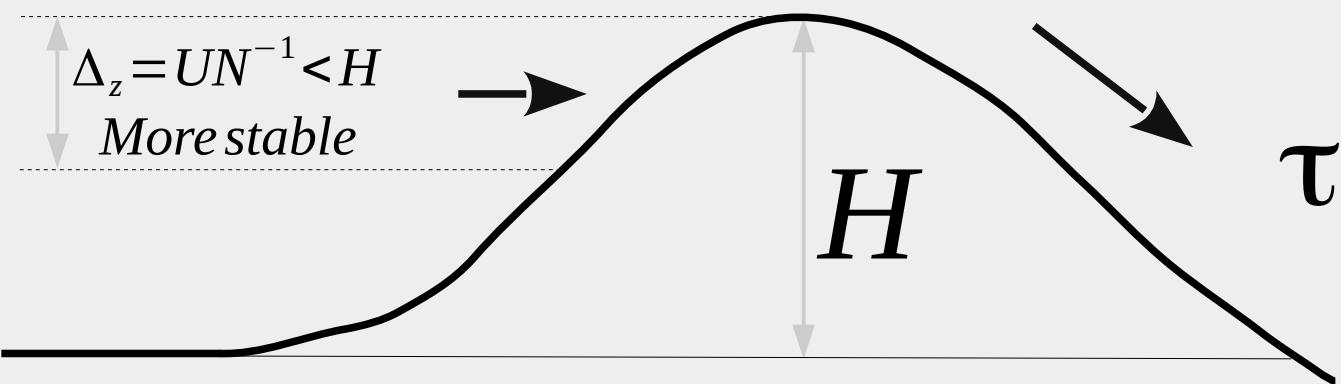




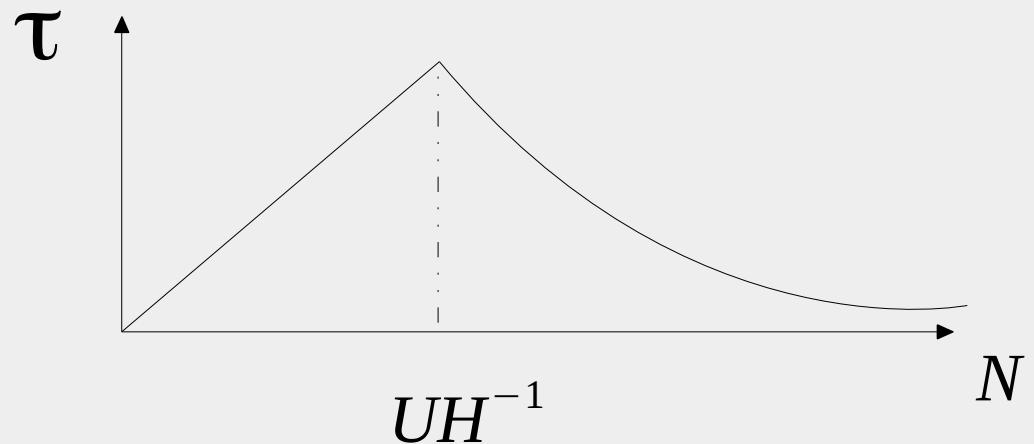
Mountain stress: $\tau = k \rho U N \Delta_z^2$



$$\tau = k \rho U N H^2$$

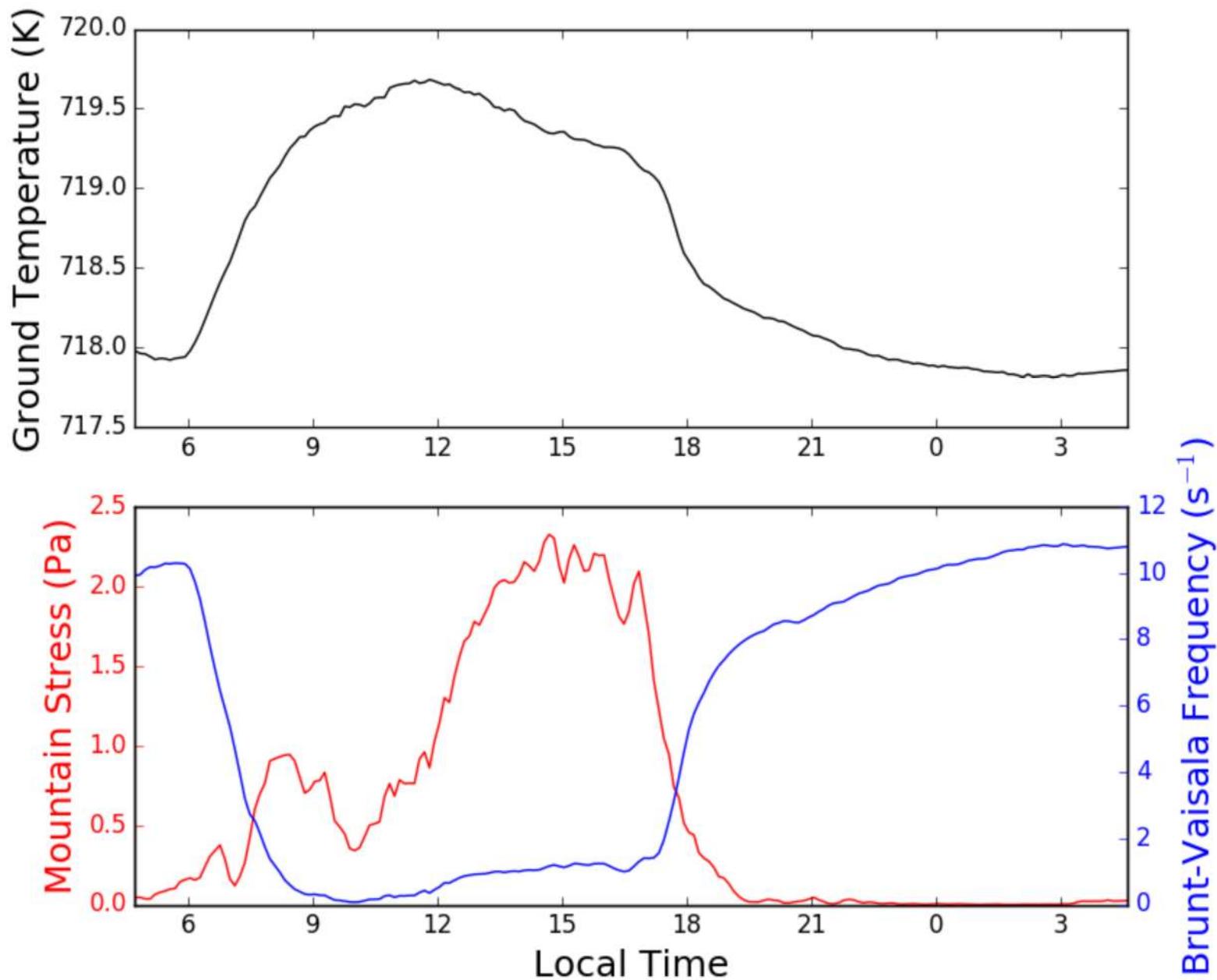


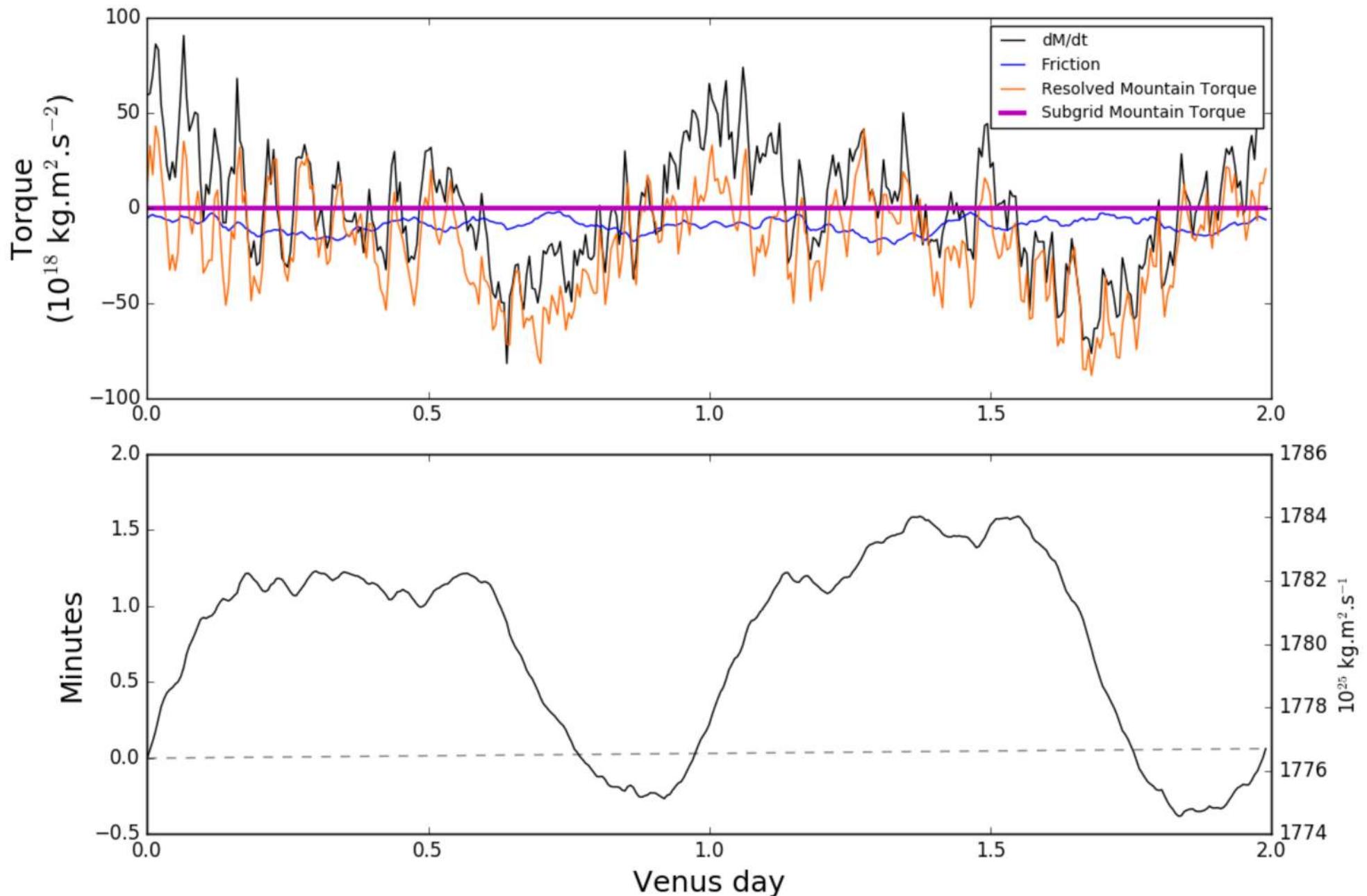
$$\tau = k \rho U^3 N^{-1}$$

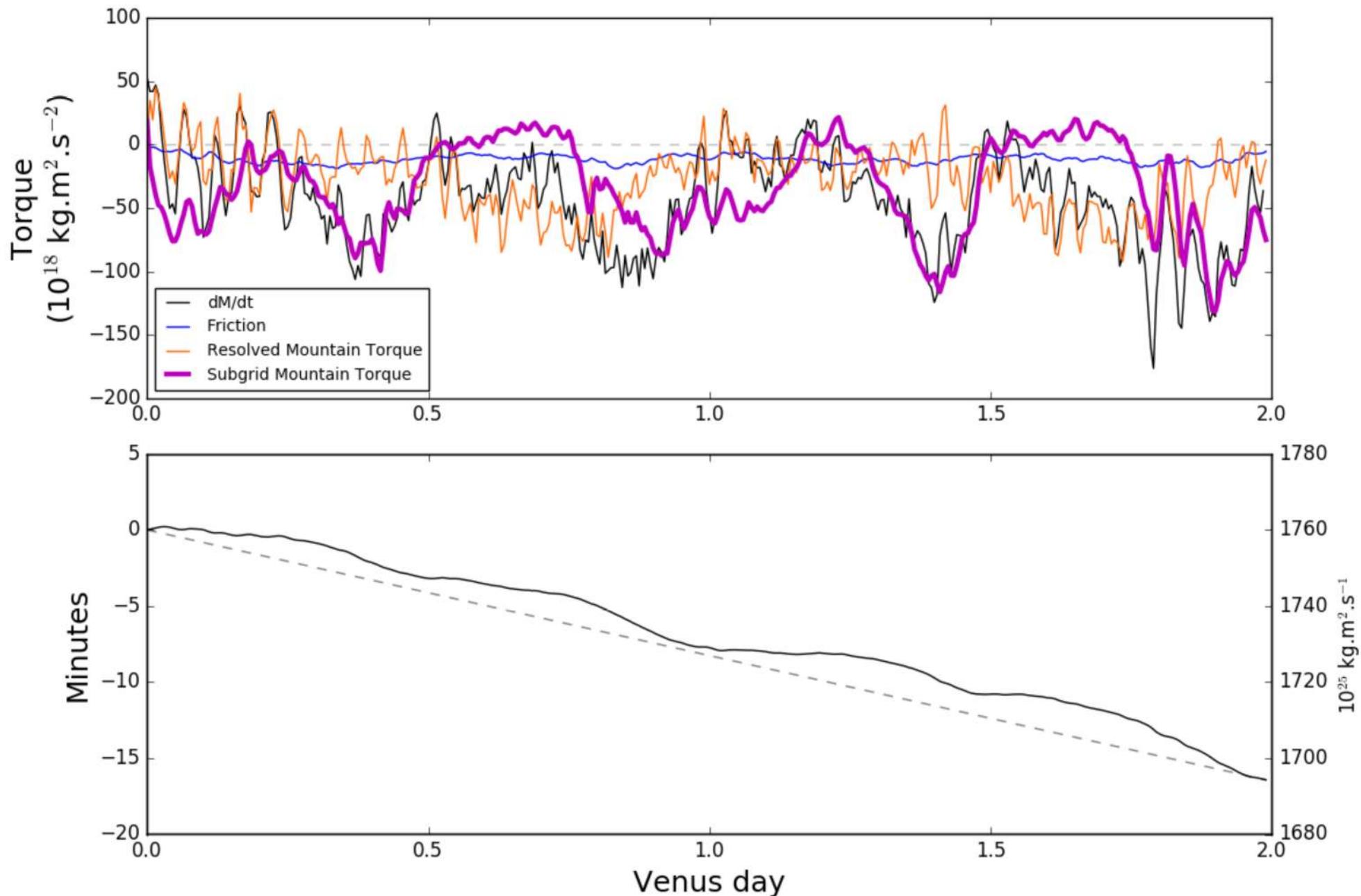


Mountain stress is
maximal for $N = UH^{-1}$

Example over Aphrodite







A fluctuating rotation rate

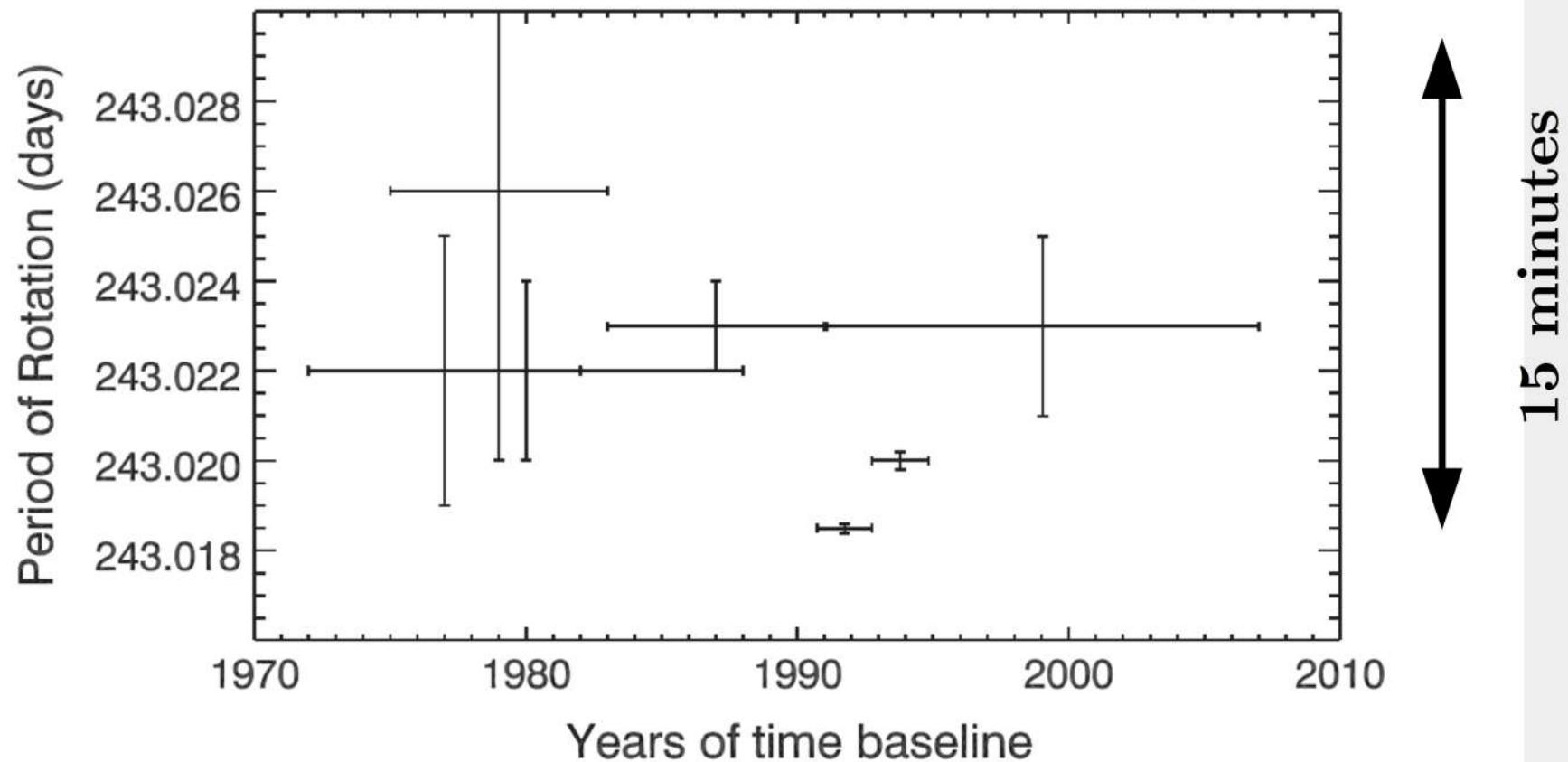


Fig. 4. The most recent estimates of the period of rotation and the time baseline of measurements. The full models and their sources are given in Table 3. The horizontal bars show the period over which the data for each estimate was acquired.

A balanced rotation rate

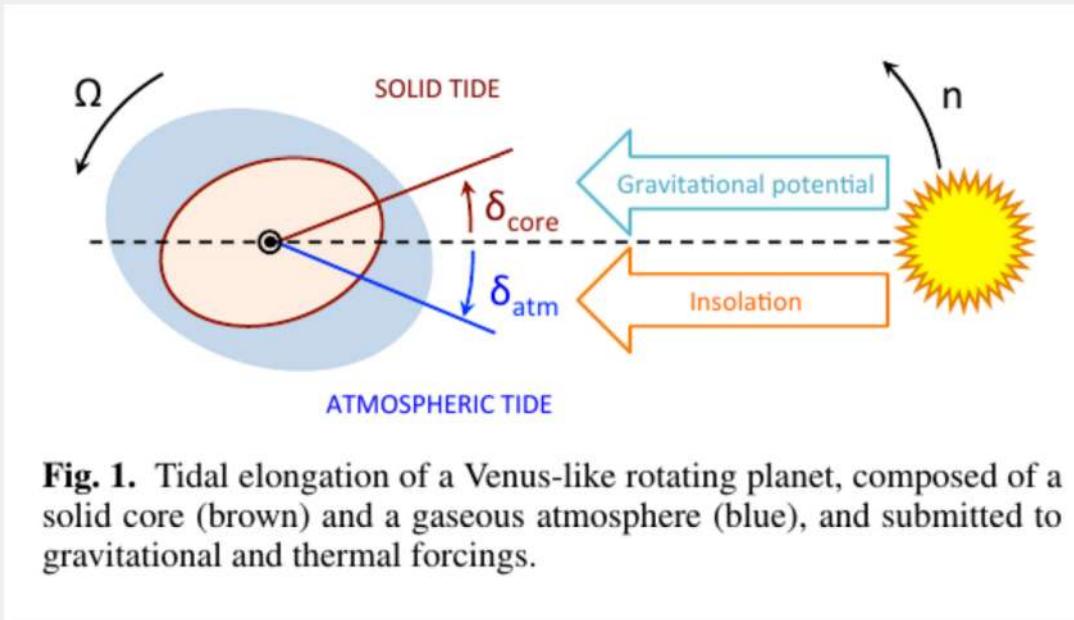
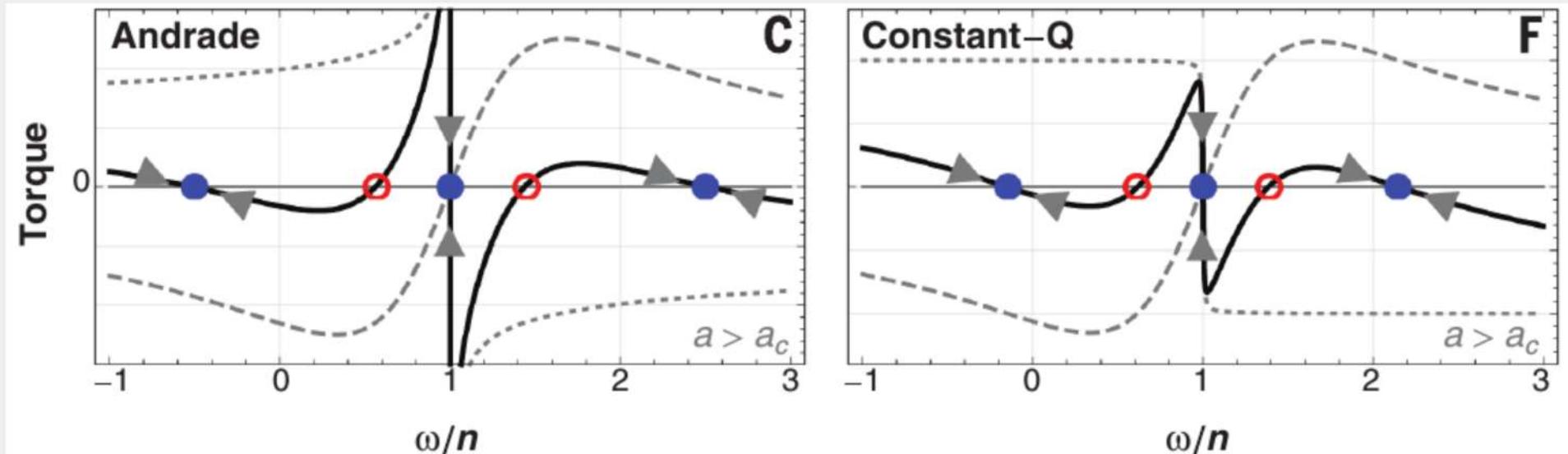


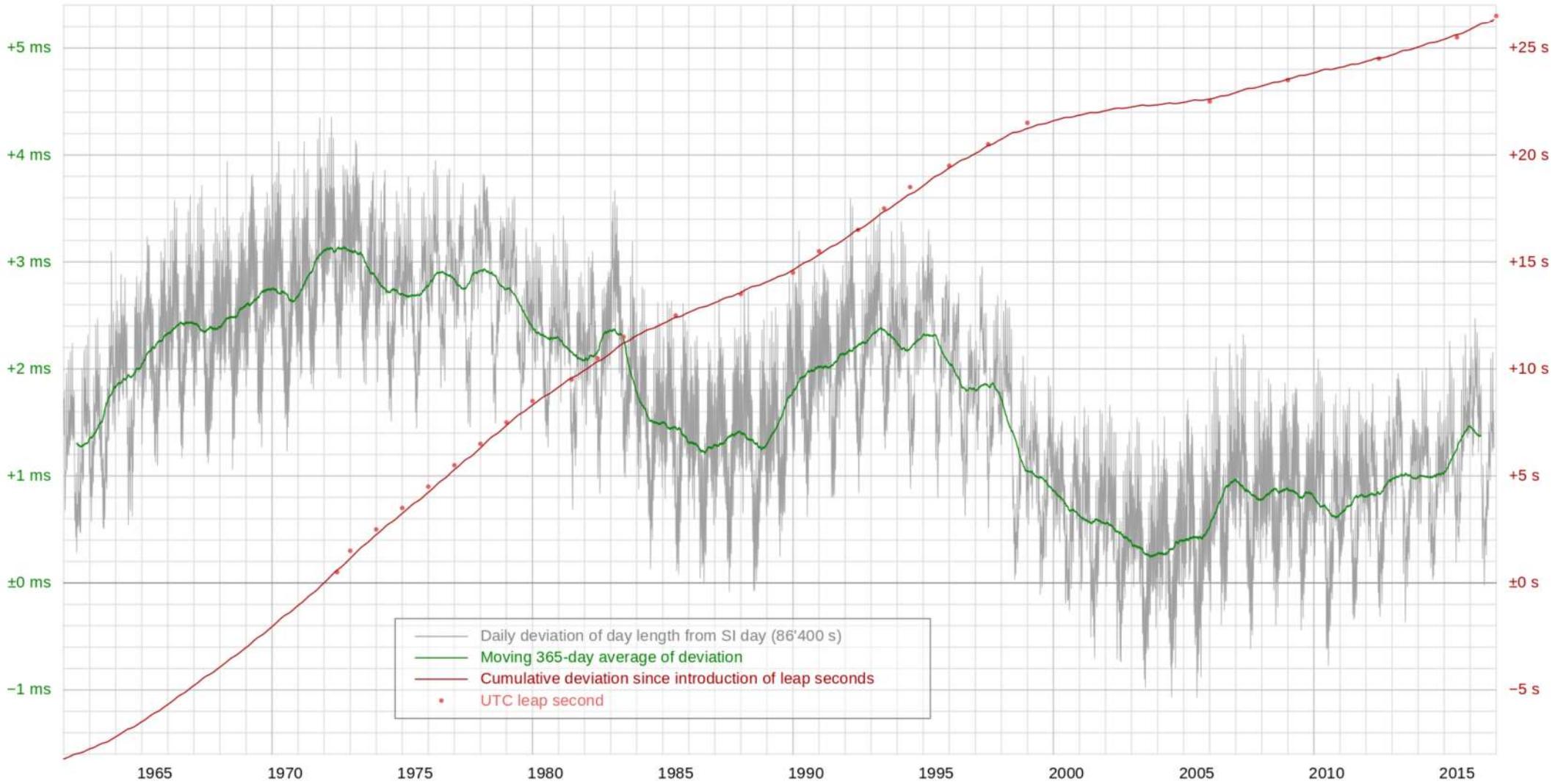
Fig. 1. Tidal elongation of a Venus-like rotating planet, composed of a solid core (brown) and a gaseous atmosphere (blue), and submitted to gravitational and thermal forcings.

Auclair-Desrotour et al., 2016

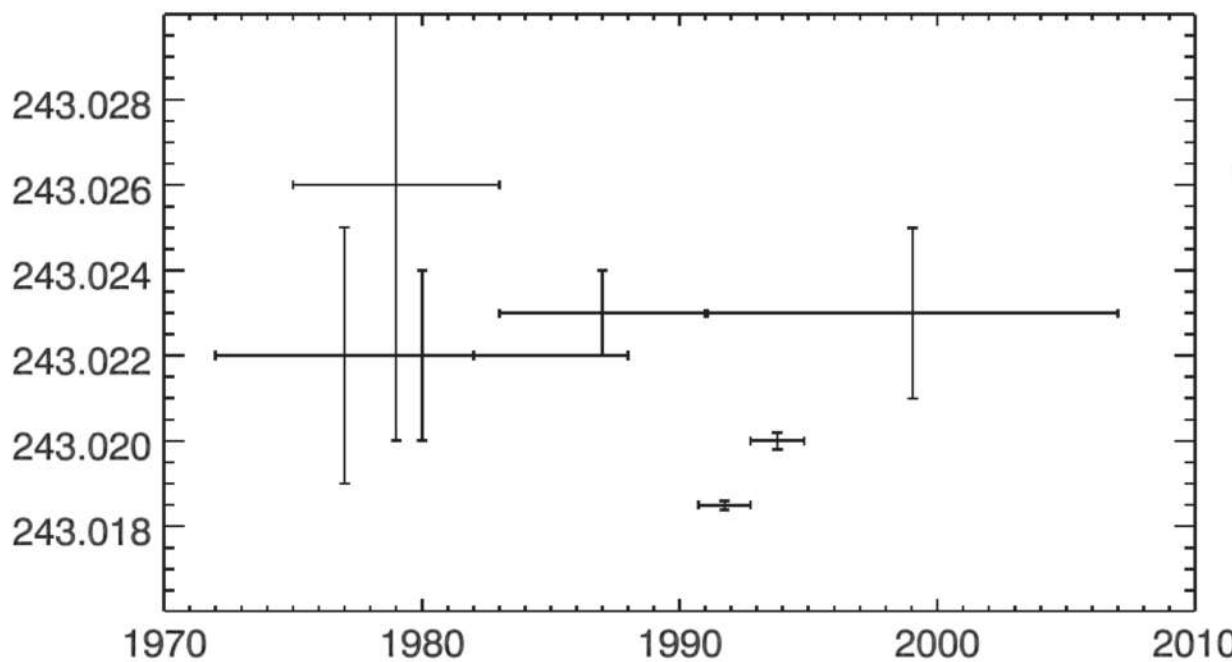


Leconte et al., 2015

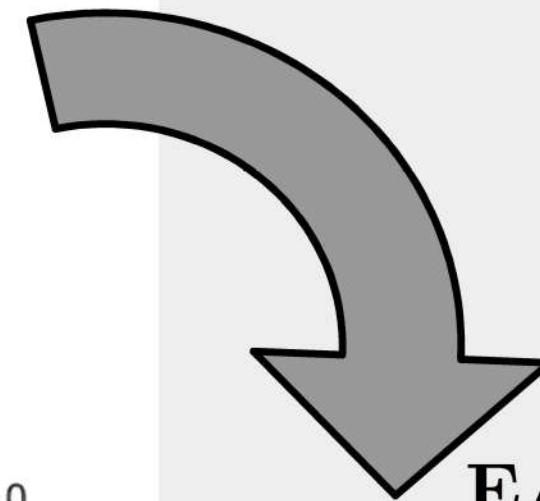
Length of day: Earth



Period of Rotation (days)

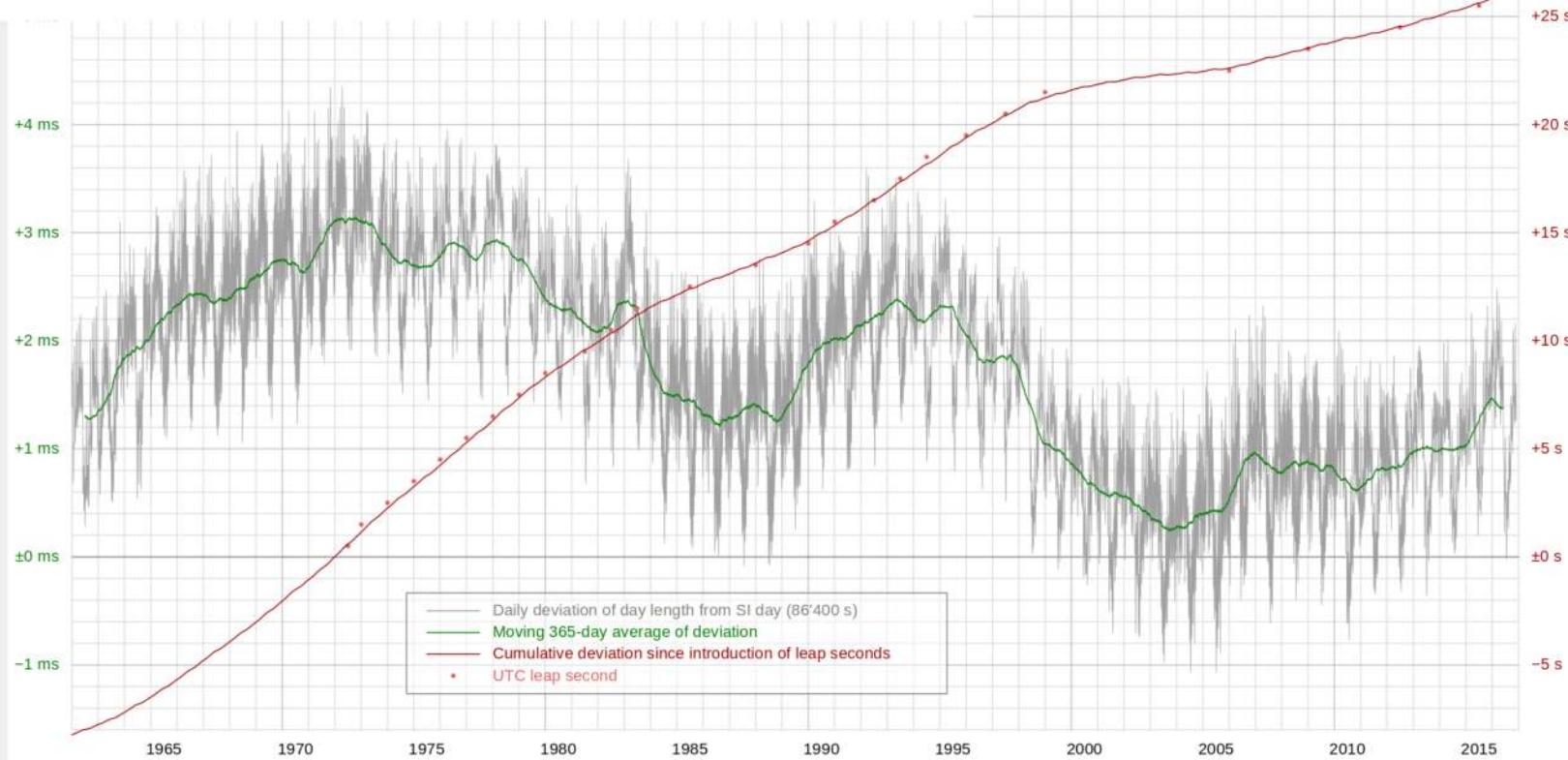


VENUS

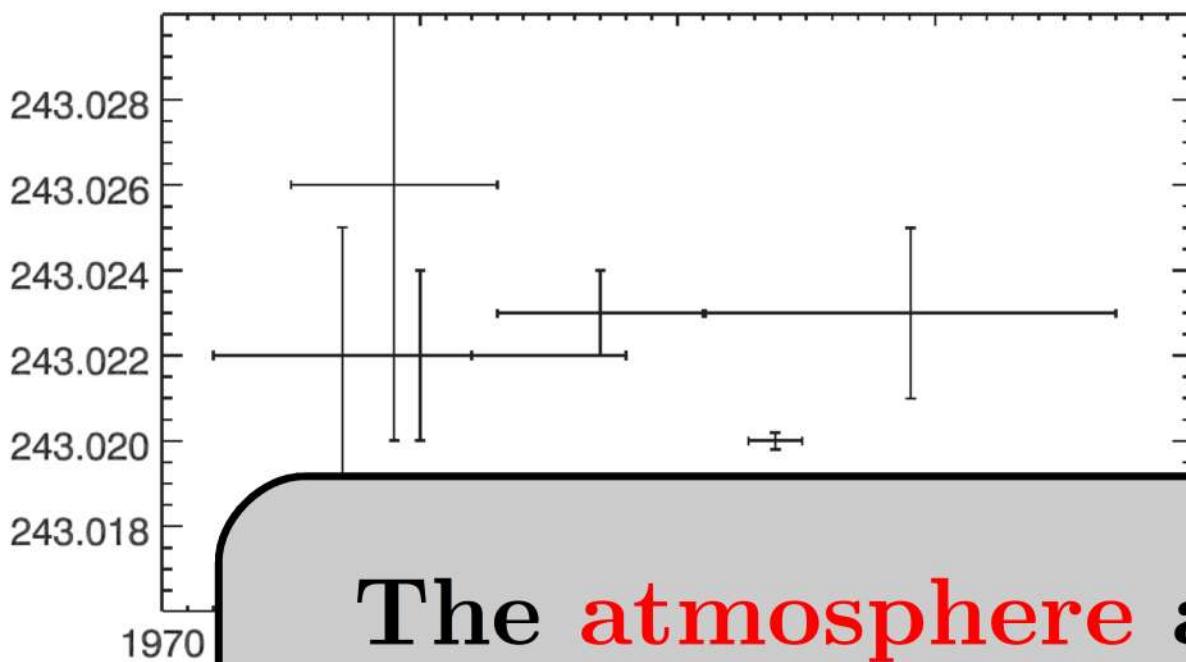


EARTH

Years of time baseline

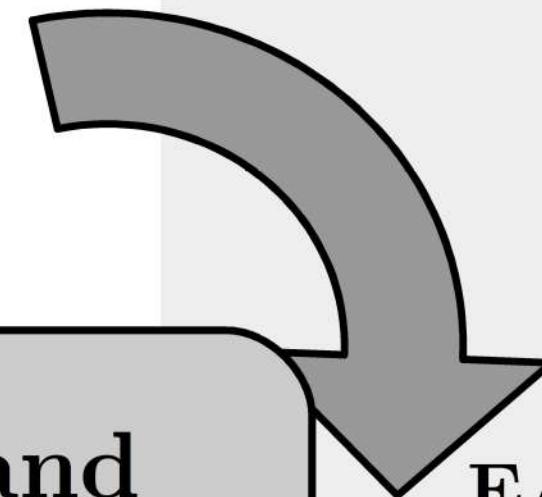


Period of Rotation (days)



VENUS

The atmosphere and
the solid body of Venus
are strongly coupled.



EARTH

