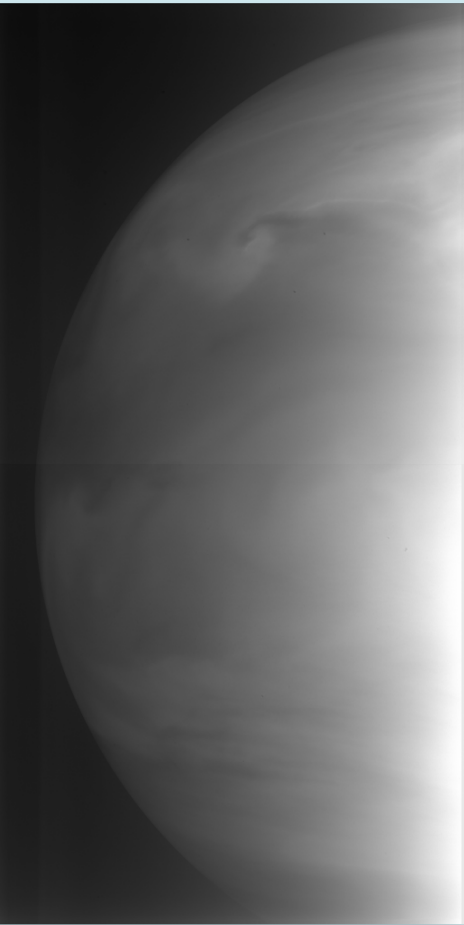


MESOSCALE VORTEX CIRCULATIONS ON VENUS OBSERVED IN AKATSUKI IR2 IMAGES

S.S. Limaye¹, T. Satoh², J. Peralta², T. Horinouchi³, T. Imamura

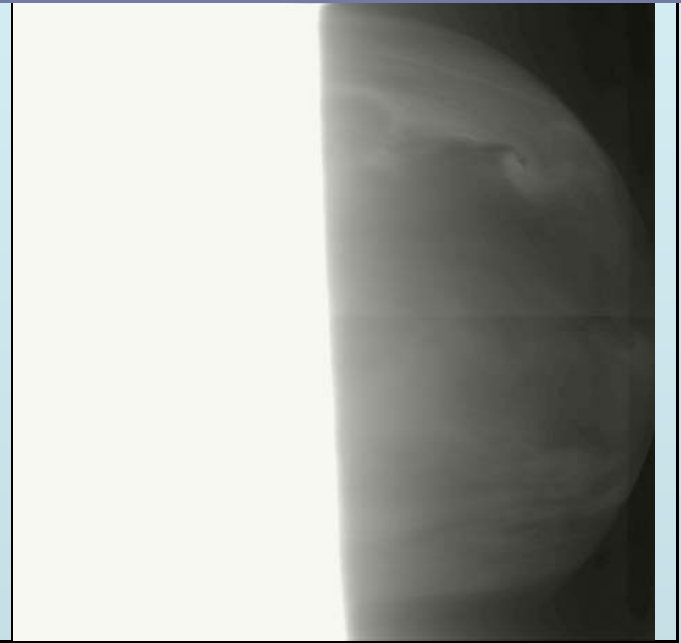
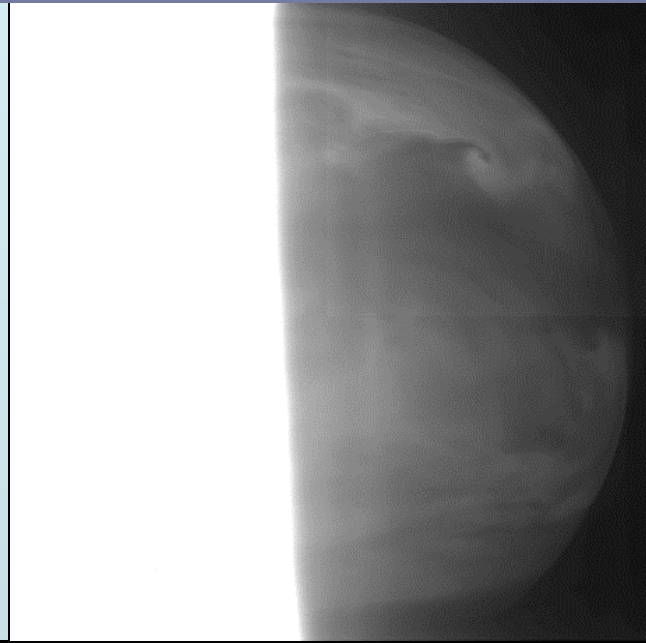
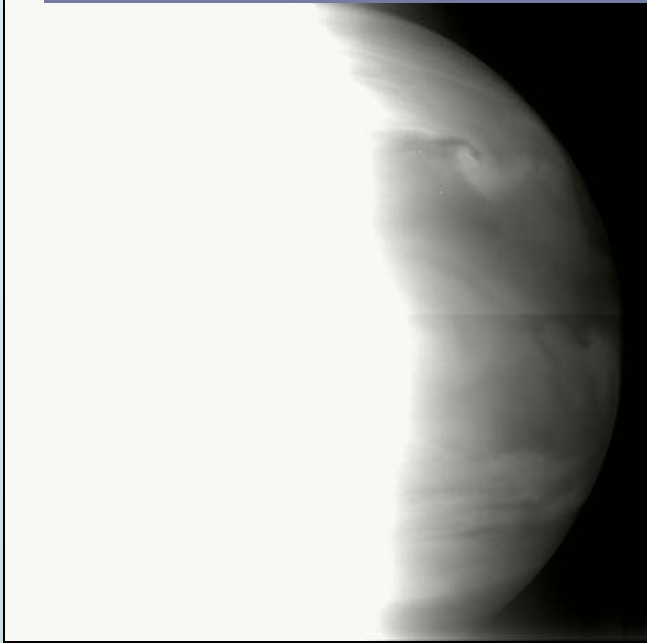
8th Moscow Solar System Symposium, Space Research Institute, Moscow,
Russia
9-13 October 2017



7 May 2016 04:02:11 1.74 μm

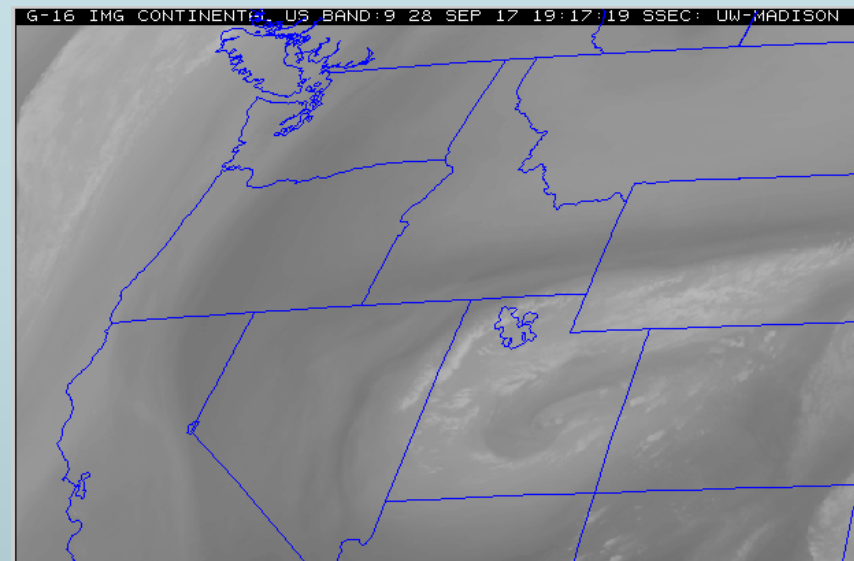
7 May 2016 04:03:33 2.26 μm

7 May 2016 04:10:36 2.32 μm



A similar “mushroom” feature, but more developed, was seen two orbits earlier at the same latitude, suggesting that it may be the same feature evolving and migrating mostly in longitude. Such features have been shown to be pairs of cyclonic and anti-cyclonic circulations

Similar vortex pair features are routinely seen in water vapor images of Earth obtained from weather satellites



Circulation aspects of the mushroom feature on Earth

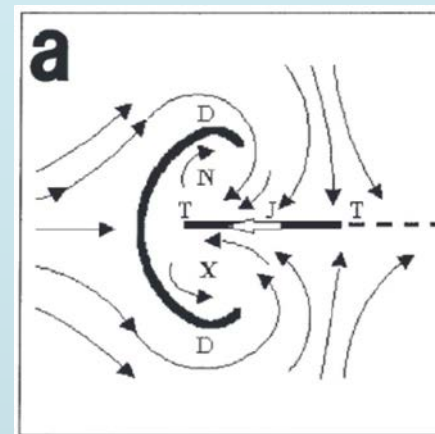
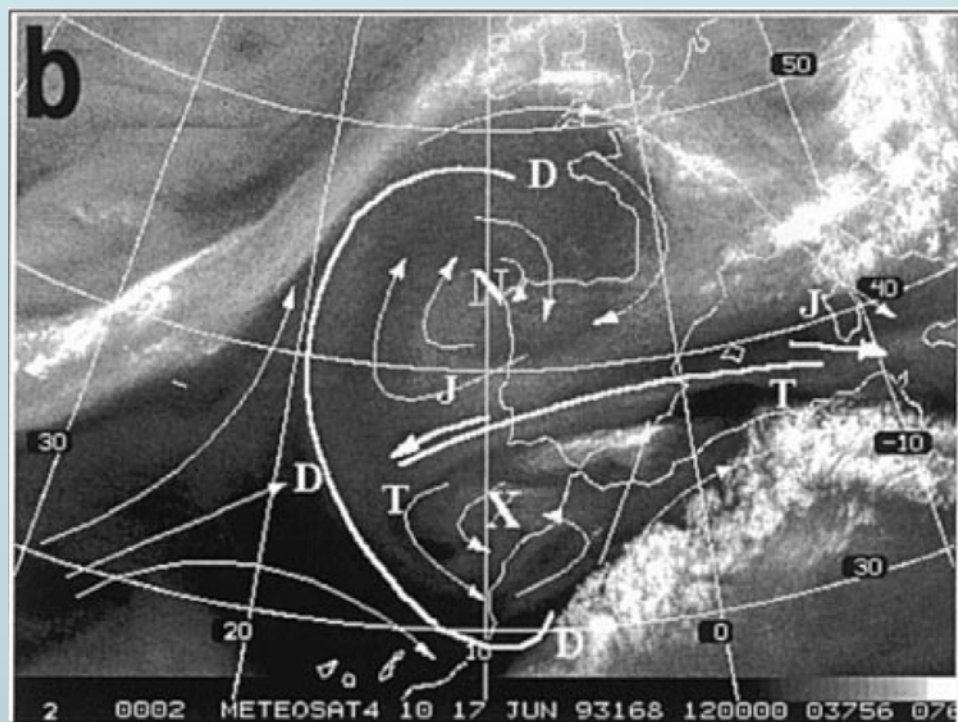
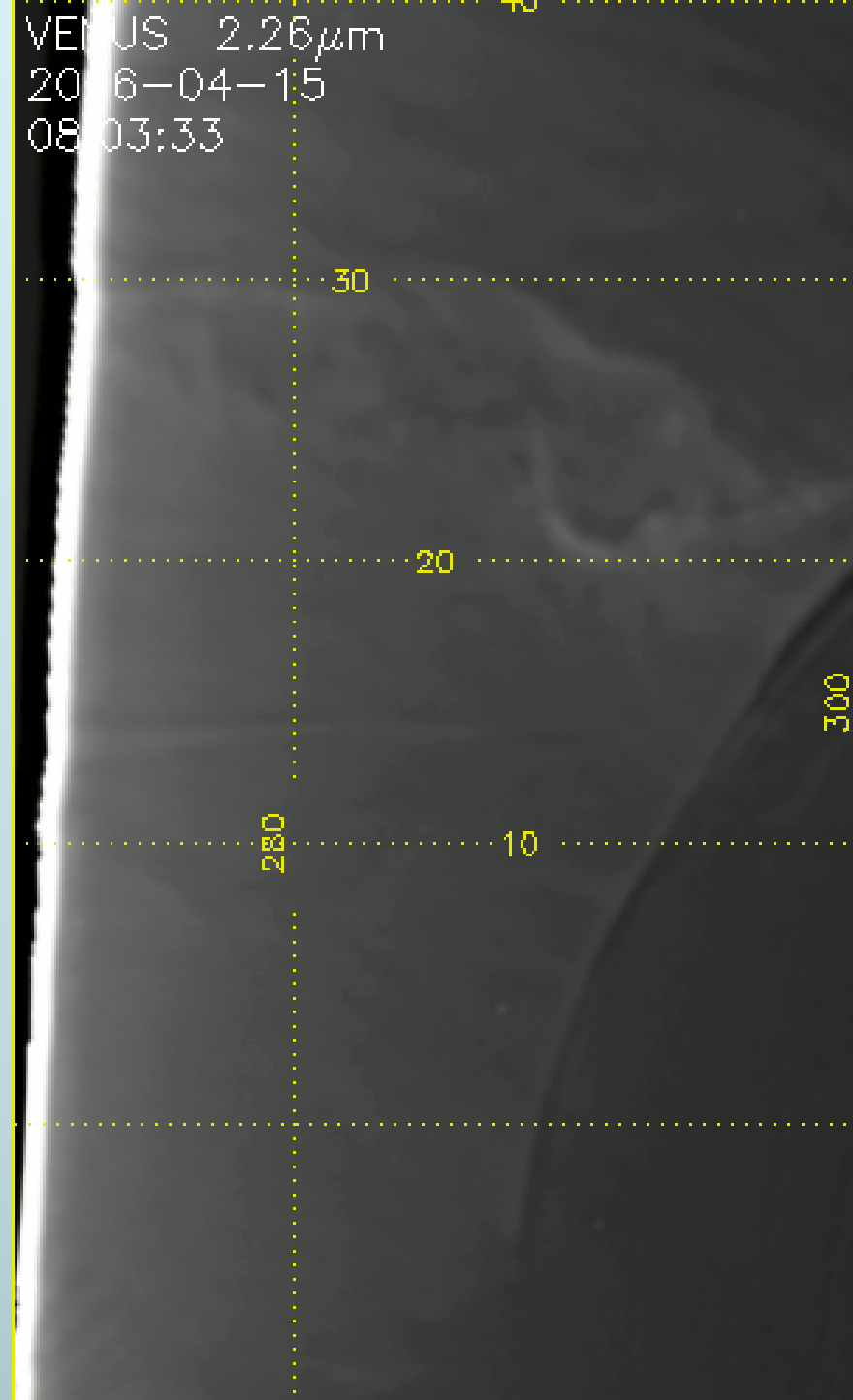


Figure 1, . (a) Idealized mushroom-like system as seen in WV imagery for an easterly flow. The two thick lines represent the deformation axes. DD represents the leading deformation zone, TT is the trailing deformation zone, X and N mark the negative and positive vorticity centers, respectively, and J indicates the local jet stream circulation. Thin solid lines are the streamlines of the relative motion of the flow. (b) METEOSAT WV imagery at 1200 UTC on 17 June 1993 showing a mushroom-like pattern close to the Iberian peninsula. A sketch of the main kinematic features, derived from a satellite imagery sequence, has been drawn.

“The mushroom configuration in water vapor imagery and operational applications”, Martin et al., *Meteorol. Appl.* 6, 143–154 (1999)



300° 290° 280°

Are these circulations triggered by topography?

The vertical shear of the horizontal wind can lead to transformation of vorticity through forced vertical motion over elevated features on the surface.

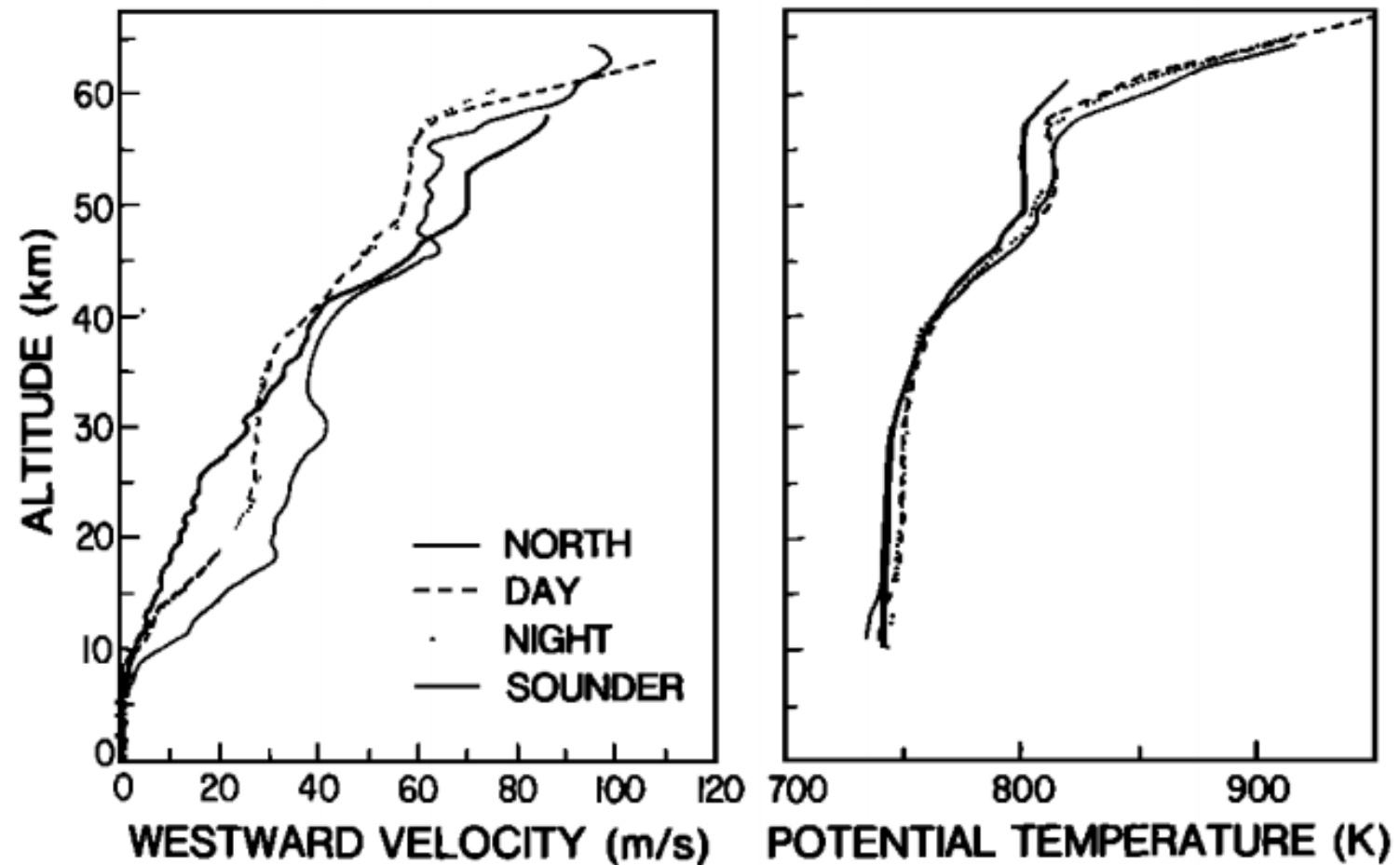
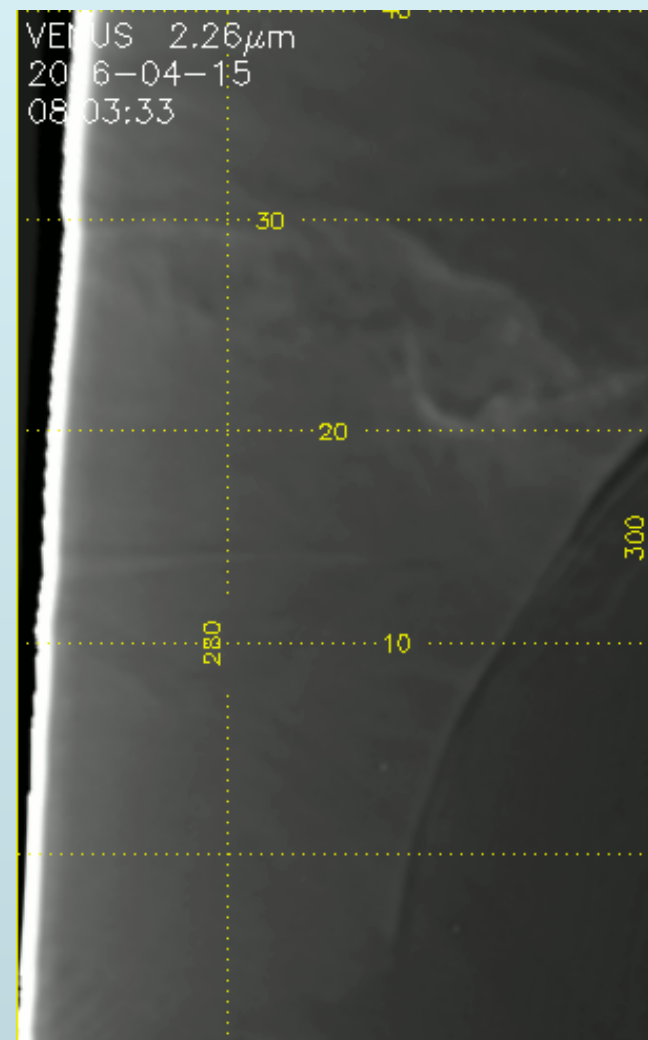


Fig. 4. Zonal wind velocity and potential temperature profiles. The wind velocities are derived from the Pioneer Venus DLBI and Doppler tracking data. Potential temperatures are derived from temperature measurements by the probes, assuming an atmosphere of 96.5% CO₂ and 3.5% N₂. In the potential temperature representation, adiabats are vertical lines, positive slopes indicate statically stable subadiabatic regions, and negative slopes indicate unstable superadiabatic regions.

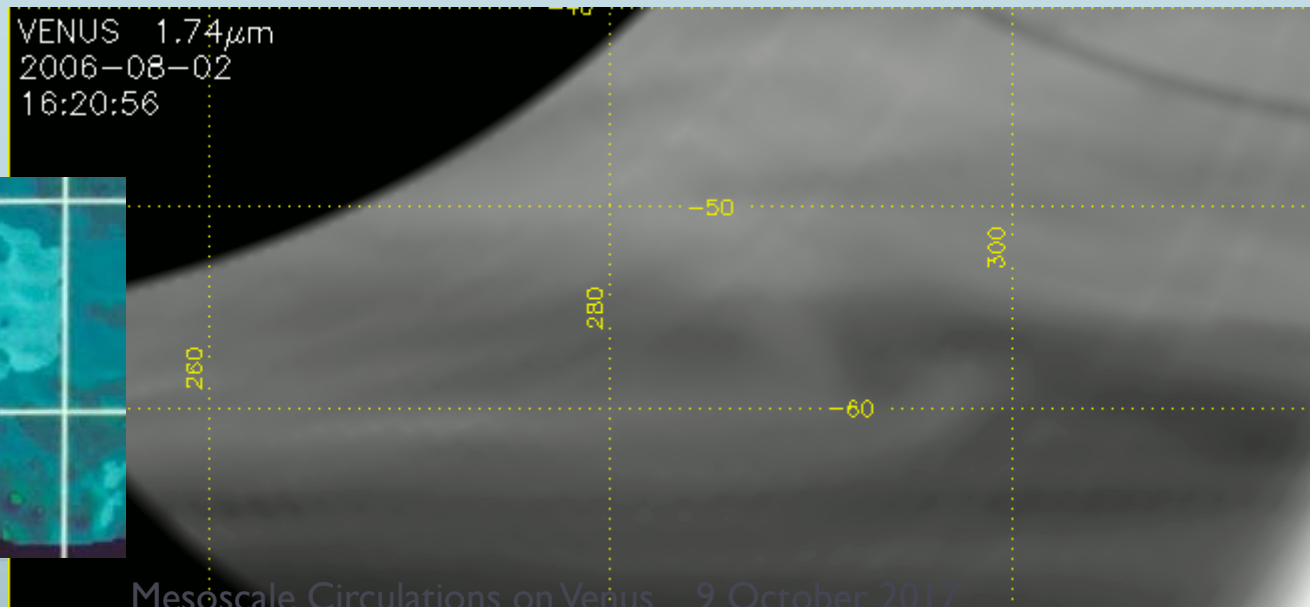


260° 280° 300°

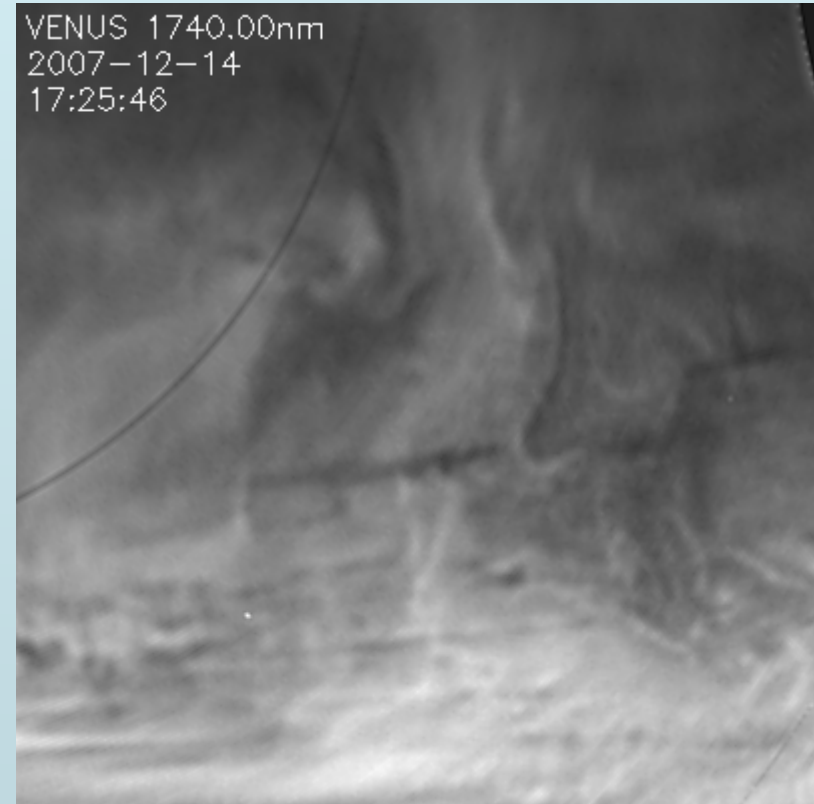
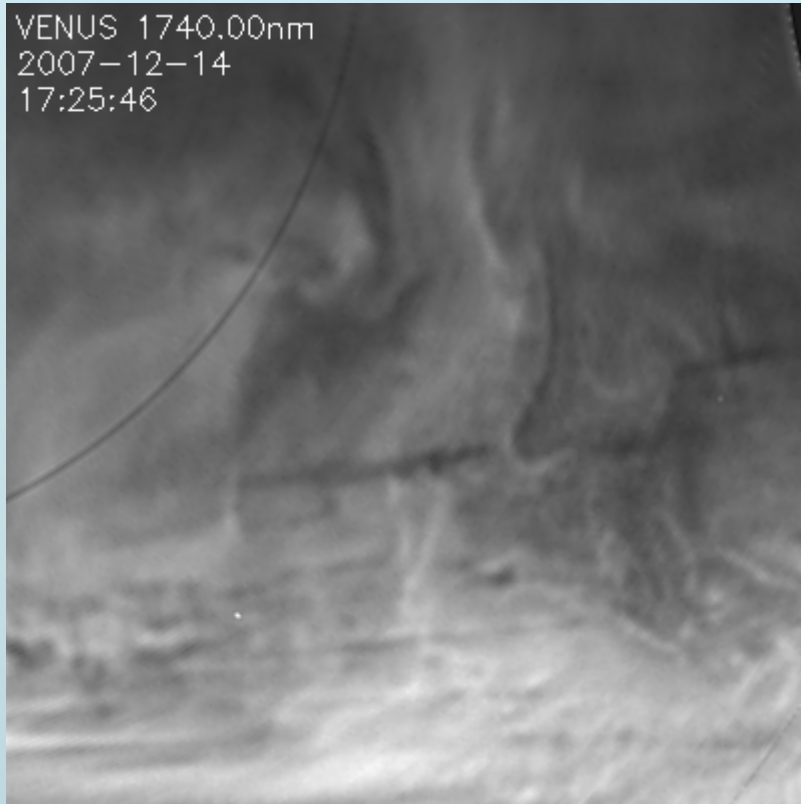
Similar features in VIRTIS Data



-60°



Mesoscale Circulations on Venus - 9 October 2017



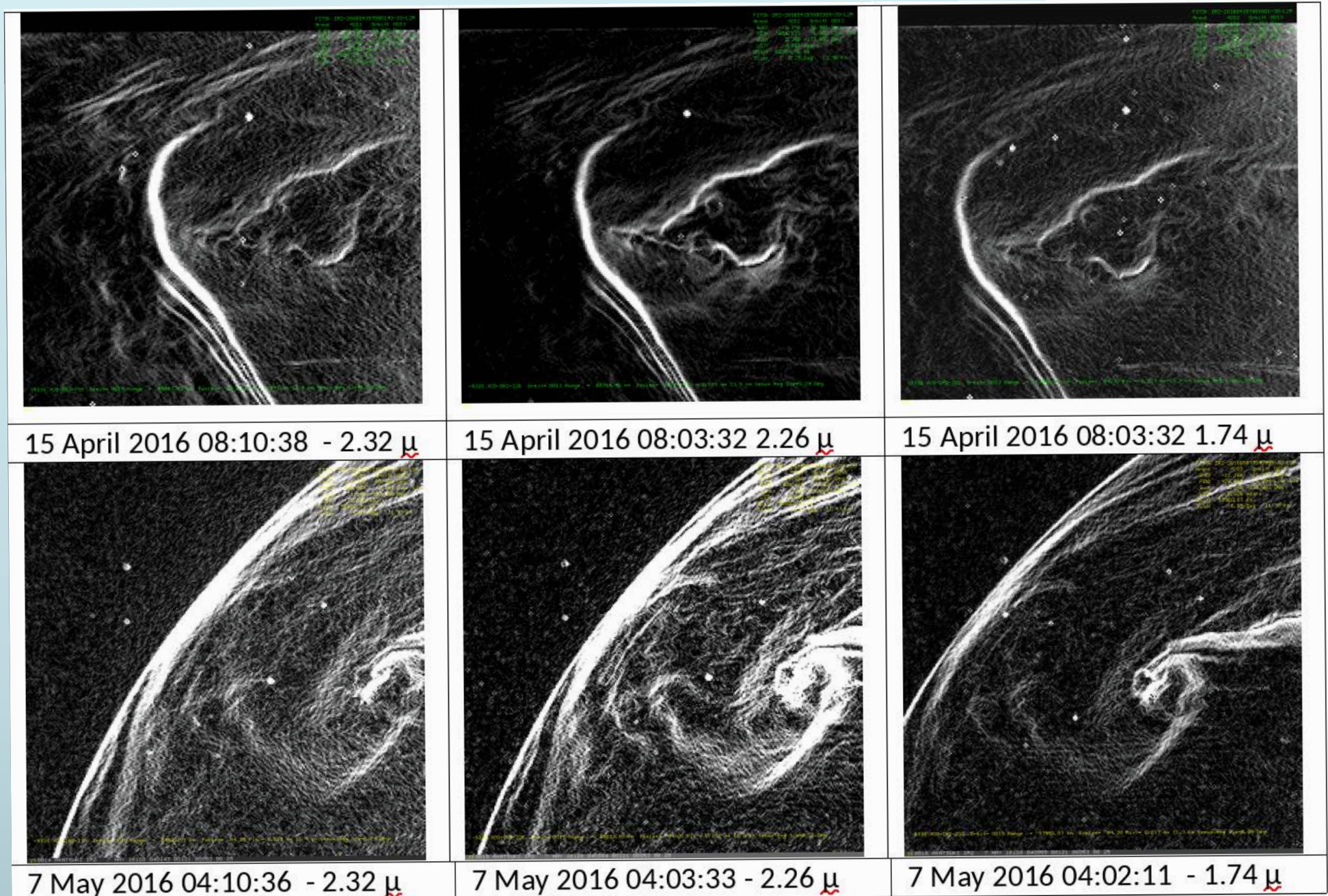
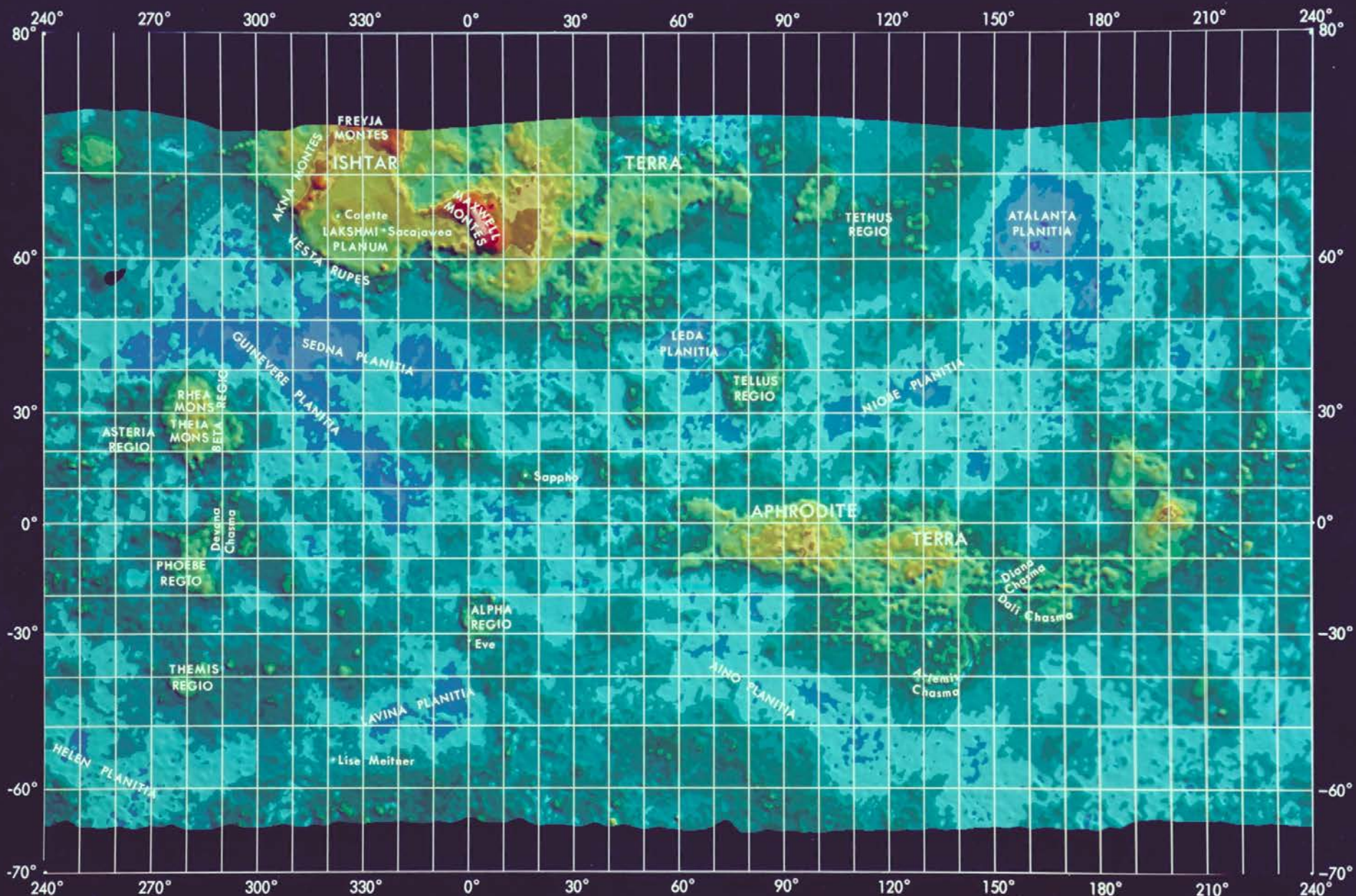


Figure 4. A sequence of gradient filtered images shown in Figure 1 reveals the intricate structure within and around the mushroom feature.

Mesoscale Circulations on Venus 9 October 2017



NASA, Ames Research Center
 U.S. Geological Survey
 Massachusetts Institute of Technology

JUNE 1980

VENUS

Mesoscale Circulations on Venus 9 October 2017

NOTES ON BASE

This chart is one of a series of maps covering the planet of Venus at a scale of 1:50,000,000. It is prepared for the Venus Radar Mapper (VRM) mission. Maps in this series are based on radar altimetry data from the Pioneer Venus Orbiter and are not intended to be used as a basis for navigation. The data were obtained from the Colorado State University, Fort Collins, the Pioneer Venus Orbiter, and the Venus Radar Mapper (VRM) and Marsden and others (1977). The scale is indicated by the scale bar at the bottom of the chart.

ANOMALY FIGURE

The map of Venus used for the construction of this chart is based on a surface with a mean radius of 6052 km. The map is based on the Pioneer Venus Orbiter (PVO) and Marsden and others (1977) data. A primary grid is shown with a radius of 6052 km.

PROJECTION

The map is based on a surface with a mean radius of 6052 km. The map is based on the Pioneer Venus Orbiter (PVO) and Marsden and others (1977) data. A primary grid is shown with a radius of 6052 km.

CONTROL

Planimetric control is derived from the tracked paths of the orbiter. The map is based on the Pioneer Venus Orbiter (PVO) and Marsden and others (1977) data. A primary grid is shown with a radius of 6052 km.

MAPPING TECHNIQUES

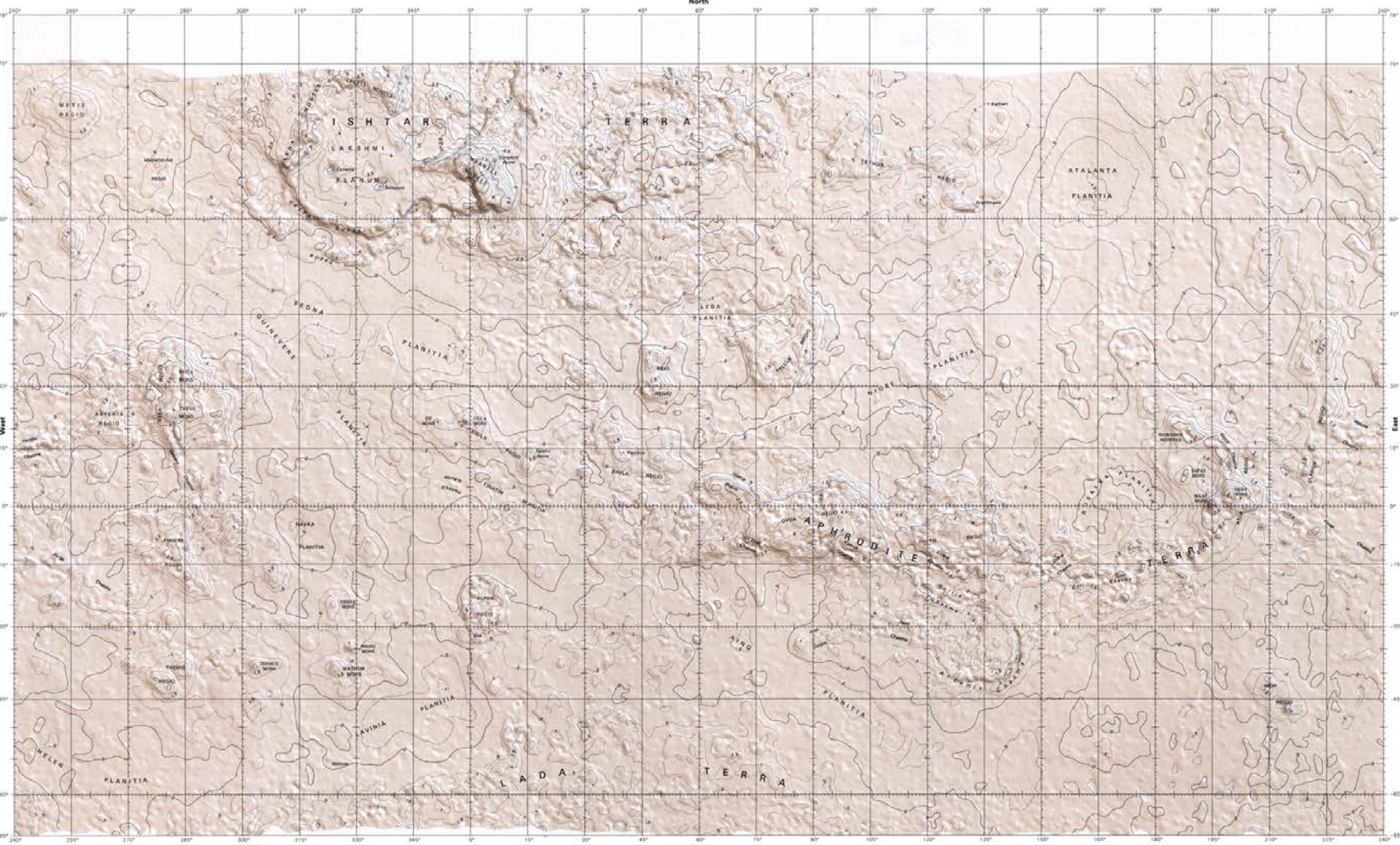
Data for the surface relief and elevation were derived from altimetry measurements of radar altimetry and interferometry altimetry. The altimetry data were obtained from the Pioneer Venus Orbiter (PVO) and Marsden and others (1977) data. A primary grid is shown with a radius of 6052 km.

SCHEMATIC LINES

Maples on this chart are prepared by the International Astronomical Union (IAU), 1961.

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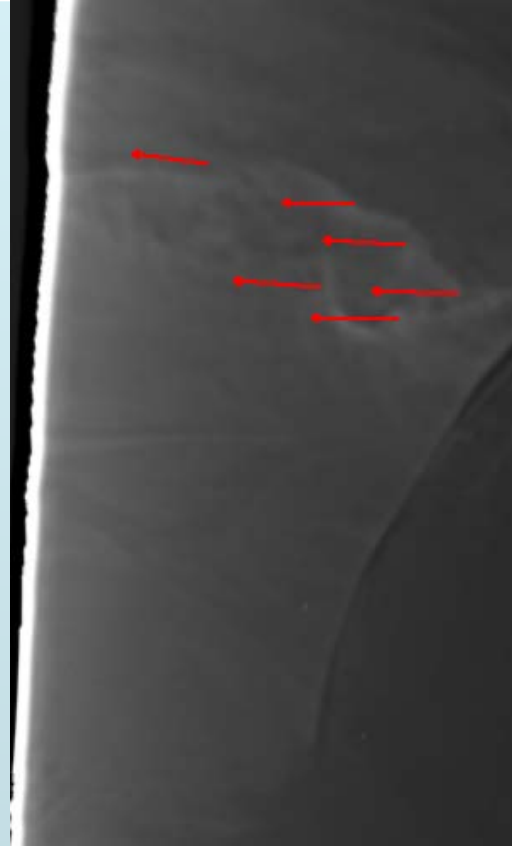
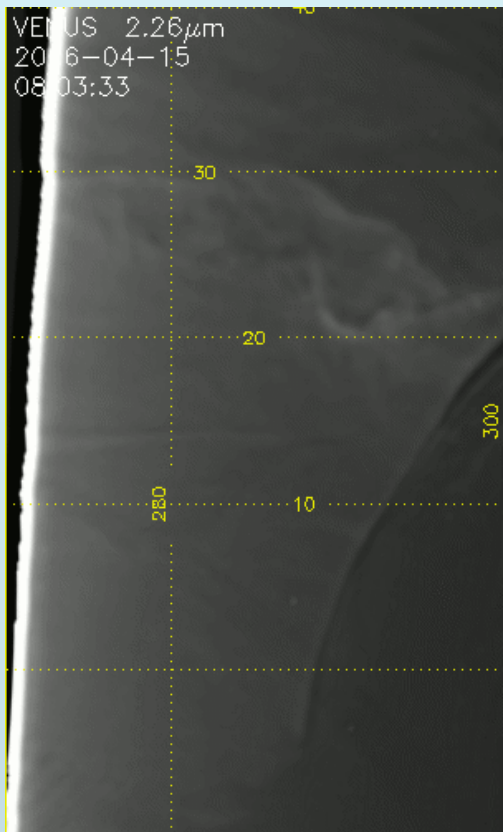


TOPOGRAPHIC MAP OF VENUS
VRM PLANNING CHART
V 50M 6/60 RT
1984

Prepared on behalf of the Planetary Geology Program, Planetary Division, Office of Space Science, National Aeronautics and Space Administration under contract N 11,705.
Control on this map was primarily obtained in order to facilitate plotting of VRM data.

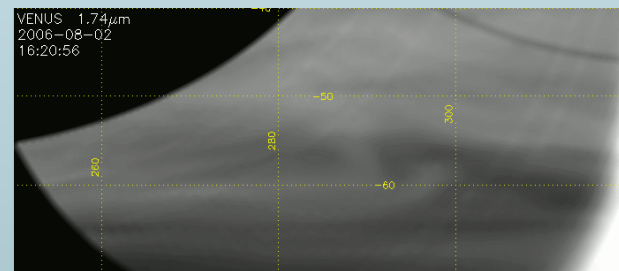
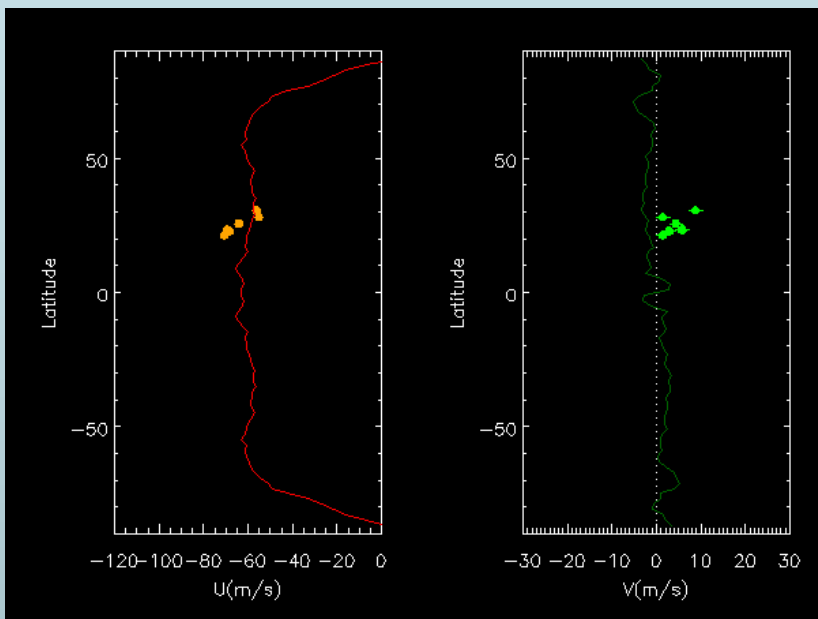
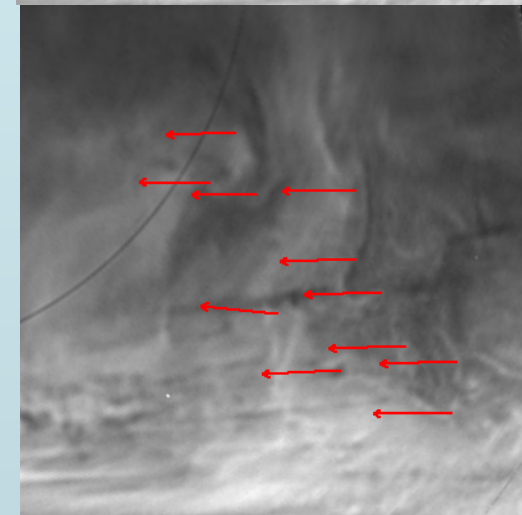
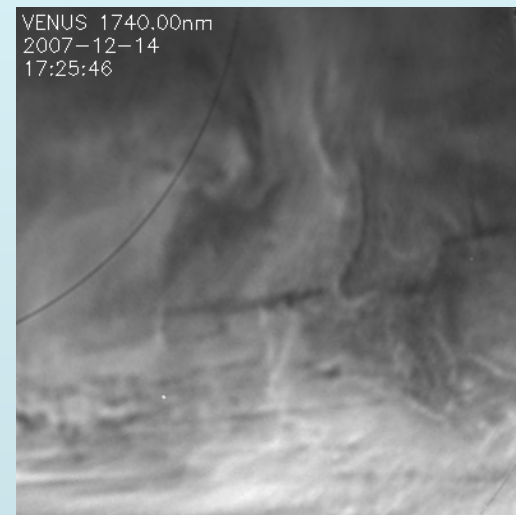
NOTE TO USERS
Users making prints or reproductions are urged to indicate when the map was prepared in the U.S. Geological Survey Working Paper 1162-A, Venus: A Planet of the Future, a publication series. Science, v. 281, no. 4411, p. 412-418.

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Good temporal and coverage is not available at high resolution to measure rotational motions.

Only zonal displacement is measurable due to low resolution



Puzzles

- **Some questions to ponder:**
 - Is there any difference in radiances of the cyclonic and anti-cyclonic features?
 - What makes them detectable?
 - What is the sharp boundary seen in the April 15 discovery frame?

Summary

- This is the first instance of meso-scale circulations on Venus in low and mid latitude regions, as only in the “eye” of the hemispheric vortex have meso-scale features arising out of dynamical instability have been observed
- The morphology indicates strong meridional flows and opposite to the ambient zonal flows.
- The vortex pair are often found in stratospheric extrusion events on Earth (cold dry air fingers sinking fast) which run into “blocking” high pressure regions causing the flows to bifurcate and form a vortex pair with opposite vorticities
- The same location suggests that local topography could induce such flows by creating a blocking feature over mountain tops.
- No time sequences are available on those two orbits, only three IR 2 filter images, so measuring cloud motions is not very feasible
- IR2 images show many amazing and perplexing non zonal flows
- Earth vortex pairs are short lived – only about a day or less, but the Venus vortex pairs may last longer?