

# Terahertz spectroscopy and global analysis of the rotational spectrum of bis-deuterated amidogen radical ND<sub>2</sub>

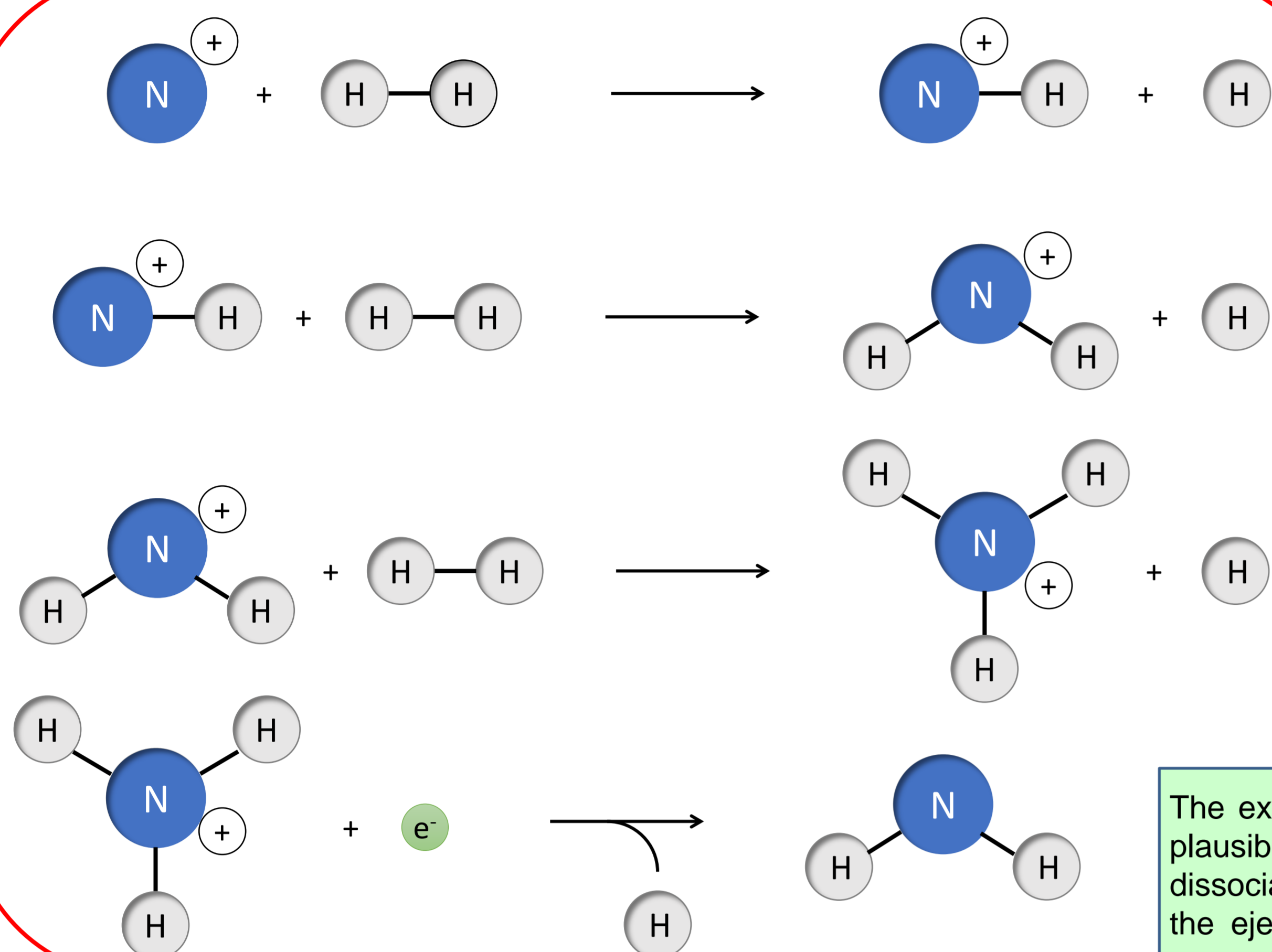
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## ABSTRACT

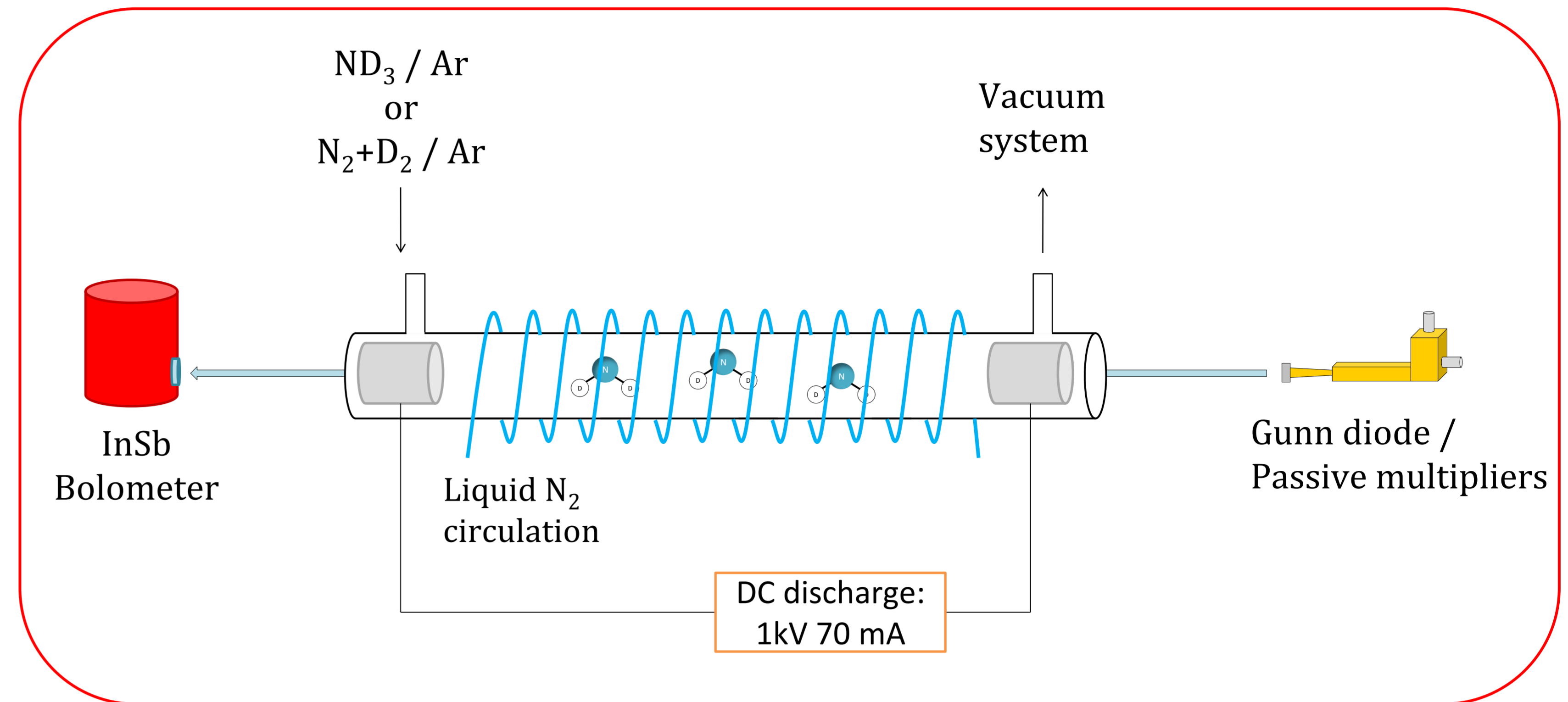
The deuteration mechanism of molecules in the interstellar medium (ISM) is still being debated. Observations of deuterium-bearing species in several astronomical sources represent a powerful tool to improve our understanding of the interstellar chemistry. The doubly-deuterated form of the astrophysically interesting amidogen radical could be a target of detection in space. In this work, the rotational spectrum of the ND<sub>2</sub> radical in its ground vibrational and electronic X<sup>2</sup>B<sub>1</sub> states has been investigated between 588 GHz and 1.131 THz using a frequency-modulation millimeter/submillimeter-wave spectrometer. The ND<sub>2</sub> has been produced in a free-space glass absorption cell by discharging a mixture of ND<sub>3</sub> and Ar. Sixty-four new transition frequencies involving *J* values from 2 to 5 and *K<sub>a</sub>* values from 0 to 4 have been measured. A global analysis including all the previous field-free pure rotational data [1,2,3] has been performed, allowing for a more precise determination of a very large number of spectroscopic parameters. Accurate predictions of rotational transition frequencies of ND<sub>2</sub> are now available from a few GHz up to several THz.

### The mechanism of formation in the ISM...



The existence of deuterated forms of amidogen in the ISM seems to be plausible as predicted by the gas-phase chemical models, where the dissociative recombination of partially deuterated intermediates favours the ejection of hydrogen atoms rather than deuterium. Roueff et al. [4] presented a steady state model of the gas phase chemistry aimed at understanding the deuterium fractionation of ammonia. They found that: (i) at high density and high depletion, deuteration of N-containing species results very efficient (ii) the deuterium fractionation is sensitive to the temperature and is large for temperatures between 5 and 20 K. Their model predicts high fractional abundances of ammonia progenitors and their deuterated isotopologues, including ND<sub>2</sub>, in dense cores.

### ... and in the laboratory



### Rotational energy levels are split by electric and magnetic interactions

This scheme is valid for symmetric rotational levels, which combine with the antisymmetric deuterium spin functions (*I<sub>D</sub>* = 1) and, thus, are split into three hyperfine sublevels.

$N_{KaKc}$

$J = N + S$

$F_1 = J + I_N$

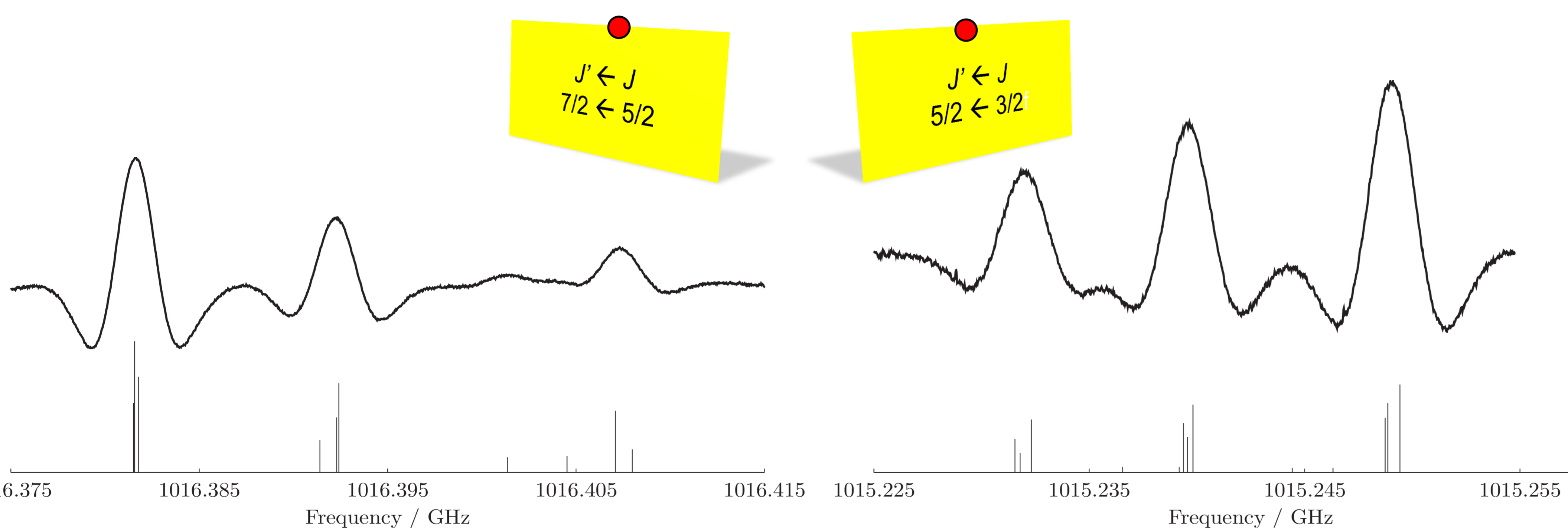
$F = F_1 + I_D$

On the other hand, antisymmetric rotational levels only combine with the symmetric deuterium spin functions (*I<sub>D</sub>* = 0, 2) and are split into six hyperfine components.

### Spectroscopic parameters

Constants	Present work	Previous MW <sup>a</sup>	Previous FIR <sup>b</sup>
<i>A</i> /MHz	399989.5534(87) <sup>c</sup>	399985.879(81)	399993.92(189)
<i>B</i> /MHz	194498.1916(150)	194488.65(16)	194498.10(102)
<i>C</i> /MHz	128610.4447(145)	128613.987(57)	128610.00(126)
Centrifugal distortion			
$\Delta_N$ /MHz	7.86074(33)	7.323(10)	7.8392(108)
$\Delta_{NK}$ /MHz	-33.48376(259)	-33.812(25)	-33.388(42)
$\Delta_K$ /MHz	198.7064(57)	196.771(33)	198.783(84)
$\delta_N$ /MHz	3.080801(201)	3.03 <sup>d</sup>	3.0858(36)
$\delta_K$ /MHz	8.5567(82)	6.300(39)	8.2788(294)
$\Phi_N$ /kHz	1.4776(80)		1.376(35)
$\Phi_{NK}$ /kHz	-6.277(306)		-5.606(35)
$\Phi_{KN}$ /kHz	-43.70(98)		-43.45(96)
$\Phi_K$ /kHz	320.34(90)		321.0(23)
$\phi_N$ /kHz	0.7224(33)		0.7402(140)
$\phi_{NK}$ /kHz	-0.834(99)		-1.01(21)
$\phi_K$ /kHz	65.90(133)		33.37(65)
$L_{KKN}$ /kHz	0.08336(141)		0.0707(38)
$L_K$ /kHz	-0.7022(85)		-0.694(26)
$l_K$ /kHz	-0.1230(33)		-0.0311(16)
$M_K$ /Hz	1.069(38)		1.08(10)
Fine interaction			
$\epsilon_{aa}$ /MHz	-5128.1262(215)	-5127.81(11)	-5127.71(41)
$\epsilon_{bb}$ /MHz	-668.6973(180)	-668.507(78)	-668.759(160)
$\epsilon_{cc}$ /MHz	3.4616(166)	3.413(63)	3.23(12)
$\Delta_{NK}^S$ /MHz	0.07613(57)	0.0469(66)	0.0779(29)
$\Delta_{NK+NK}^S$ /MHz	-0.9498(57)	-0.967(81)	-0.875(29)
$\Delta_{NK}^S$ /MHz	1.045(201)		
$\Delta_K^S$ /MHz	9.8548(76)	9.587(66)	9.476(99)
$\delta_N^S$ /MHz	0.03843(36)		0.0353(16)
$\delta_K^S$ /MHz	0.1005(59)	0.0356(45)	0.150(36)
$\Phi_{NK}^S$ /kHz	3.76(51)		
$\Phi_K^S$ /kHz	-23.10(81)		-15.36(72)
Hyperfine interaction			
$a_F(N)$ /MHz	28.0577(122)	28.055(33)	
$T_{aa}(N)$ /MHz	-43.1451(183)	-43.136(48)	
$T_{bb}(N)$ /MHz	-44.2715(213)	-44.277(63)	
$C_{aa}(N)$ /MHz	0.2660(75)	0.269(27)	
$C_{bb}(N)$ /MHz	0.0458(55)	0.045(21)	
$C_{cc}(N)$ /MHz	0.0145(64)	0.019(22)	
$a_F(D)$ /MHz	-10.2446(103)	-10.241(28)	
$T_{aa}(D)$ /MHz	2.8536(168)	2.874(45)	
$T_{bb}(D)$ /MHz	-2.0916(247)	-2.108(69)	
$X_{aa}(N)$ /MHz	0.2191(297)	0.213(84)	
$X_{bb}(N)$ /MHz	-3.744(34)	-3.75(11)	
Number of FIR lines			181
Standard deviation of the FIR data /cm <sup>-1</sup>			6.9 × 10 <sup>-4</sup>
Number of MODR lines			198
Standard deviation of the MODR data /MHz			0.376
Number of MW lines			182
Standard deviation of the MW data /kHz			51.0
Fit standard deviation			0.87
$N'_{max}, K'_{max}$			13, 10

### Portion of the Terahertz spectrum - $N'_{KaKc} \leftarrow N_{KaKc} = 3_{13} \leftarrow 2_{02}$ transition



## CONCLUSION

The rotational spectrum of the doubly-deuterated amidogen radical ND<sub>2</sub> in its ground electronic state X<sup>2</sup>B<sub>1</sub> has been investigated up to the Terahertz region. New measurements have been analyzed in a global fit including previous mmW, sub-mmW, MODR and FIR data. Some misassignments or incomplete assignments of previous data sets have been corrected, so that it has been possible to obtain a set of accurate spectroscopic constants which are compatible with the entire body of rotational data available for this radical. Accurate predictions are now available up to 8 THz for transitions with *N* < 14 and *K<sub>a</sub>* < 11, providing a set of data very useful for astronomical searches. Recently, we have also observed and analyzed the pure rotational spectrum of <sup>15</sup>ND<sub>2</sub> (manuscript in preparation).

#### REFERENCES:

[1] Cook, J. M. & Hills, G. W. 1983, J. Chem. Phys., 78, 2144. [2] Kanada, M., Yamamoto, S., & Saito, S. 1991, J. Chem. Phys., 94, 3423. [3] Morino, I. & Kawaguchi, K. 1997, J. Mol. Spectrosc., 182, 428. [4] Roueff, E., Lis, D. C., van der Tak, F. F. S., Gerin, M., & Goldsmith, P. F. 2005, A&A, 438, 585.

### Predictions

