

1                    Overview of MPLNET Version 3 Cloud Detection

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## ABSTRACT

The National Aeronautics and Space Administration Micropulse Lidar Network Version 3 cloud detection algorithm is described and its differences relative to the previous version highlighted. Clouds are identified from normalized Level 1 signal profiles using two complementary methods. The first considers signal derivatives vertically for resolving low-level clouds. The second, which resolves high-level clouds like cirrus, is based on signal uncertainties given the relatively low signal-to-noise ratio exhibited in the upper troposphere by eye-safe network instruments, especially during daytime. Furthermore, a multi-temporal averaging scheme is used to improve cloud detection under conditions of weak signal-to-noise. Diurnal and seasonal cycles of cloud occurrence frequency based on one year of measurements at the Goddard Space Flight Center (Greenbelt, MD) site are compared for the new and previous versions. The largest differences, and perceived improvement, in detection occurs for high clouds (above 5-km, mean sea level) which increase in occurrence by nearly 6%. There is also an increase in the detection of multi-layered cloud profiles from 9% to 20%. Macrophysical properties and estimates of cloud optical depth are presented for a transparent cirrus dataset. However, the limit to which molecular signal can be reliably retrieved above cirrus clouds occurs between cloud optical depths of 0.5 and 0.8.

## 31 **1. Introduction**

32 Clouds play a critical role in the Earth's climate system because they are inextricably linked  
33 to the hydrological cycle and radiation budget (Liou 1986; Ramanathan et al. 1989). Information  
34 about cloud height, thickness, occurrence, and amount are critical inputs for a host of numerical  
35 applications involving climate research. Therefore, it is important to have highly accurate and  
36 quantitative data records of cloud properties that span several years and geographic regions.  
37 Verification of even the most basic modeling processes demands compulsory observations of  
38 global cloud occurrence, if there is to be any confidence in their fidelity.

39 Various methods of determining cloud climatologies exist, each with their own advantages  
40 and limitations. Visual observations from the surface (Warren et al. 1985; Hahn et al. 1996;  
41 Hahn and Warren 1999) provide cloud fraction and morphological cloud types. However, these  
42 can be biased by the quality of technician training, underestimation of high clouds, sparse global  
43 coverage, and nighttime bias. Passive radiometric sensors aboard satellites, which are the core  
44 input of the International Satellite Cloud Climatology Project (ISCCP; Rossow and Schiffer  
45 1991, 1999), offer a true global representation and have the best (unobstructed) potential view of  
46 high clouds. However, these can undersample low-level maritime clouds and underrepresent  
47 optically-thin cirrus clouds (Holz et al. 2008).

48 Active sensors, like lidar and radar (Platt et al. 1994; Moran et al. 1998; Wang and Sassen  
49 2001), are the primary tools for observing and profiling cloud vertical structure to high accuracy.  
50 When flown aboard satellites, like Cloud Aerosol Lidar and Infrared Pathfinder Satellite  
51 Observations (CALIPSO; Winker et al. 2007) and CloudSat (Stephens et al. 2002), active  
52 sensors also provide global coverage. Even still, the relatively narrow profiling curtain of  
53 current active sensors limits observation densities. In the case of CALIPSO and CloudSat, these

54 missions provide at most two profiles per 24-hour period over most regions, which limits studies  
55 of the diurnal impact of clouds on the Earth system. Fundamentally, an array of remote sensing  
56 methods is needed in order to investigate the complexity of clouds (Schiffer and Rossow 1983).

## 57 **2. Micropulse Lidar Network**

58 The National Aeronautics and Space Administration (NASA) Micropulse Lidar Network  
59 (MPLNET; Welton et al. 2001, <http://mplnet.gsfc.nasa.gov>) is a federated network of micropulse  
60 lidar (MPL) systems deployed worldwide in support of basic science and the NASA Earth  
61 Observing System program (Wielicki et al. 1995). A benefit of MPLNET is the use of a  
62 standardized instrument employing a common data processing algorithm with thorough  
63 uncertainty characterization, which allows for straightforward comparisons between sites. With  
64 sites in polar, mid-latitude, and tropical regions and continuous day/night, high temporal  
65 resolution datasets going back as far as 1999, MPLNET datasets represent a valuable archive for  
66 improving our understanding of global cloud macrophysical properties on diurnal, season, and  
67 decadal scales.

68 There have been two versions of MPLNET data processing algorithms to date. The first,  
69 referred to as Version 1, was released in 2000. Beginning in 2006, the project transitioned to  
70 Version 2 (hereafter V2) data products which are currently available. Version 3 (hereafter V3)  
71 data processing algorithms are currently in development.

72 The V2 Level 1 MPLNET data products contain the system diagnostics (e.g. solar  
73 background counts, instrument temperature and energy, etc.), normalized relative backscatter  
74 (NRB), and NRB uncertainty. The NRB is reported at one-minute temporal resolution up to 30-  
75 km, typically at 75-m vertical resolution (some sites operate at 30-m). The NRB is defined as:

$$\text{NRB}(z) = C\beta(z)T^2(z) \quad (1)$$

76 where  $C$  is the instrument calibration constant,  $\beta$  is the backscatter coefficient from both  
77 molecules and particles,  $T^2$  is the corresponding total atmospheric two-way transmittance, and  $z$   
78 is the altitude. NRB and the calculation of its uncertainty are discussed by Campbell et al.  
79 (2002) and Welton and Campbell (2002). Level 1 data are available in real time with no quality  
80 assurance. Cloud base and top heights are identified in the V2 Level 1.5 (real time, no quality  
81 assurance) and Level 2 (not real time, quality assured) MPLNET data products. The V2 Level  
82 1.5 data products use temperature and pressure profiles from US Standard Atmospheres (COESA  
83 1976) to determine molecular calculations, while National Centers for Environmental Prediction  
84 (NCEP)/National Center for Atmospheric Research (NCAR) Reanalysis (Kalnay et al. 1996) is  
85 used for V2 Level 2 data.

86 MPLNET V2 data products have been used to distinguish cloud presence in a number of  
87 scientific investigations to date. For example, Campbell and Sassen (2008) use data from the  
88 South Pole to document polar stratospheric cloud occurrence over multiple seasons. Shupe et al.  
89 (2011) consider MPLNET measurements at Ny-Ålesund, Norway as context for evaluating Arctic  
90 cloud properties. Others have investigated cirrus contamination of Aerosol Robotic Network  
91 (AERONET) aerosol optical depth in Southeast Asia (Chew et al. 2011; Huang et al. 2011) and  
92 globally (Huang et al. 2012). Lolli et al. (2013) use collocated 355/527-nm MPLNET  
93 observations to estimate the drizzle droplet size from stratocumulus and stratus clouds.

94 A new V3 cloud detection algorithm has been developed to improve the quality of MPLNET  
95 cloud products. The new algorithm uses a combination of signal-processing techniques and a  
96 multi-resolution temporal averaging scheme to resolve cloud boundaries. Meteorological  
97 profiles provided by the Goddard Earth Observing System – Version 5 (GEOS-5) Atmospheric  
98 General Circulation Model (AGCM; Rienecker et al. 2008; Molod et al. 2012) are used for

99 molecular calculations. Specifically, the Forward Processing for Instrument Teams (FP-IT)  
100 GEOS-5 Version 5.9.1 data are utilized (<http://gmao.gsfc.nasa.gov/products>). The model data  
101 are available at 3-hour intervals over 72 pressure levels at  $0.625^\circ$  longitude and  $0.5^\circ$  latitude  
102 resolution. Modeled profiles for this study were subsampled from the GEOS-5 grid containing  
103 the Goddard Space Flight Center (GSFC) site location, and interpolated to the MPLNET range  
104 and time resolutions (75 m, 1 minute).

105 The goals of this paper are to describe the new algorithm and demonstrate performance. We  
106 outline changes relative to V2 cloud detection and describe how the new algorithm is applied to a  
107 variety of cloudy scenes. We apply one year of data collected at the GSFC MPLNET site  
108 ( $38.99^\circ$  N,  $76.84^\circ$  W, 0.05 km above mean sea level; MSL) to compare V2 and V3 results and  
109 highlight the impact of our upgraded techniques through differences in macrophysical cloud  
110 properties observed from this location.

### 111 **3. Cloud detection algorithm description**

112 Examples of daytime and nighttime NRB profiles at GSFC are shown in Fig. 1. Both  
113 profiles show high-level clouds with base heights near 10 km and top heights near 13 km, MSL.  
114 The daytime NRB profile exhibits relatively lower signal-to-noise compared with the nighttime  
115 case due to higher solar background, which makes detection of elevated layers an increasingly  
116 difficult task.

#### 117 *a. Version 2 cloud detection*

118 Layers are identified in the V2 cloud detection algorithm by a combination of two retrieval  
119 methods applied to the Level 1 data products. The first method requires that the first derivative  
120 of the lidar signal exceed a minimum threshold in order to detect a layer. The assumption of  
121 strong signal gradients makes this well suited for detecting liquid-phase clouds, which are

122 frequently at lower levels in the NRB profile and correspond with higher signal-to-noise. This is  
 123 hereafter referred to as the gradient-based cloud detection method (GCDM). The second method  
 124 is designed for use in cases of low signal-to-noise ratio (SNR) and relies on uncertainties in the  
 125 lidar signal. This method uses two tunable thresholds and one objective threshold to identify  
 126 cloud boundaries, and is hereafter referred to as the uncertainty-based cloud detection method  
 127 (UCDM).

128 Given the relatively low SNR exhibited by the MPL in the upper troposphere at base one-  
 129 minute resolution (primarily during daytime), no single procedure is used to detect all cloud  
 130 types at all times. Thus the merger of these two methods offers the possibility to retrieve the  
 131 entire cloud vertical structure to the limit of signal attenuation. We describe the basis for each  
 132 method, as follows.

133 1) GRADIENT-BASED CLOUD DETECTION

134 Autonomous methods of cloud detection using gradients in the lidar signal, such as the  
 135 differential zero-crossing method described by Pal et al. (1992), are well established. The first  
 136 step in the GCDM is to normalize the NRB using the attenuated molecular backscatter  
 137 coefficient,

$$\beta'_m(z) = \beta_m T_m^2(z), \quad (2)$$

138 which produces an attenuated scattering ratio,  $\beta'_r$ , multiplied by the instrument calibration  
 139 constant as

$$C\beta'_r(z) = \frac{C\beta(z)T^2(z)}{\beta'_m(z)} = \frac{C[\beta_m(z) + \beta_p(z)]T_m^2(z)T_p^2(z)}{\beta_m(z)T_m^2(z)} = C \left[ 1 + \frac{\beta_p(z)}{\beta_m(z)} \right] T_p^2(z). \quad (3)$$

140 Here the subscripts  $m$  and  $p$  denote contributions from molecules and particles, respectively.

141 The first derivative of  $C\beta'_r$  is used to identify clouds in the GCDM. Due to increasing  
 142 uncertainty in the profile with height, the GCDM retrieval is only performed up to a “noise

143 altitude”, defined as the altitude at which the uncertainty,  $\delta\text{NRB}$ , exceeds half of the NRB (or  
 144 conversely, analogous to an SNR of 2). Cloud presence corresponds with an increase in total  
 145 backscatter, which results in a large positive gradient in  $C\beta_r'$  with height as seen in Fig. 2. The  
 146 threshold used to identify the cloud base is defined as

$$a_{\max} = K \cdot \overline{C\beta_r'}, \quad (4)$$

147 where  $\overline{C\beta_r'}$  is the mean value of  $C\beta_r'$  up to the noise altitude and  $K$  is an empirical parameter  
 148 (unitless), set to 10 for this study. The value of  $K$  is chosen carefully so that it is high enough to  
 149 reject insignificant peaks in the first derivative of  $C\beta_r'$  (i.e. aerosol stratification in the surface-  
 150 attached layer or signal noise) while remaining sensitive enough to identify weakly-scattering  
 151 clouds.

152 The cloud base is identified at the altitude bin immediately preceding that where the first  
 153 derivative of  $C\beta_r'$  exceeds  $a_{\max}$ . Identification of cloud top is more ambiguous and is performed  
 154 using one of two processes. A negative gradient in  $C\beta_r'$  occurs near the top of a cloud,  
 155 corresponding with the decrease in total backscatter and the impact of signal attenuation through  
 156 the cloud. The first method used to identify the cloud top relies on a threshold defined as

$$a_{\min} = \overline{C\beta_r'} - a_{\max}. \quad (5)$$

157 The algorithm begins by looking for altitude bins above the cloud base where the first derivative  
 158 of  $C\beta_r'$  falls below  $a_{\min}$ . Then the altitude bin where the first derivative initially returns above  
 159  $a_{\min}$  is identified as the cloud top. However, if this condition is not met, the cloud top is chosen  
 160 as the altitude bin where the value of  $C\beta_r'$  falls below the value at the cloud base or the noise  
 161 altitude, whichever occurs first. If the lidar signal becomes significantly attenuated within the



162 cloud, the designation as an apparent cloud top is more appropriate in accordance with standards  
163 agreed upon by the Experimental Cloud Lidar Pilot Study (ECLIPS; Platt et al. 1994).

164 Two cloud layers are apparent in Fig. 2. Both cloud bases are identified at the altitudes  
165 immediately below the  $a_{\max}$  exceedances (dashed line, positive derivative). The first (lowest)  
166 cloud top can be found using the  $a_{\min}$  threshold (dashed line, negative derivative). But the  
167 derivative never falls below  $a_{\min}$  for the second cloud layer. Therefore, the alternative process is  
168 used to identify the apparent cloud top. Only true (not apparent) cloud tops are reported in V2  
169 MPLNET cloud products.

## 170 2) UNCERTAINTY-BASED CLOUD DETECTION

171 An alternative to algorithms that utilize gradients in the lidar return to identify clouds are  
172 approaches that compare cloudy lidar returns to clear sky returns (Clothiaux et al. 2007).  
173 Similarly, the UCDM uses a theoretical molecular return and the signal uncertainty to detect  
174 elevated clouds, and is fully described by Campbell et al. (2008, hereafter C08). However, a few  
175 comments regarding its implementation are warranted. The first step in the UCDM is to  
176 approximate the value of the instrument calibration constant. Level 1 NRB is divided by the  
177 attenuated molecular backscatter coefficient, resulting in an attenuated scattering ratio multiplied  
178 by the instrument calibration constant, as given in Eq. (3). Next, a clear-sky search is performed  
179 to locate a normalization region where we can approximate that  $\beta_p$  approaches zero over a  
180 certain number of range bins,  $N$ . The nature of the UCDM only allows for cloud detection at  
181 altitudes above the normalization region. The calibration constant is approximated by averaging  
182 Eq. (3) over the  $N$  bins. C08 stress that this final normalization value,  $C_f^*$ , must be distinguished  
183 from  $C$  due to unknown transmission losses below the normalization region.

184 As an example, a representation of the UCDM is shown in Fig. 3. Beginning at 1-km  
 185 above ground level (AGL) and working upward, the uncertainty in  $C\beta_r'$  is evaluated at each  
 186 altitude bin according to the criteria set forth by C08 until a “clear-air” slot is found. In this case,  
 187 the “clear-air” region used to determine the normalization value begins at  $r_1 = 6.46$  km and ends  
 188 at  $r_N = 8.71$  km. Averaging  $C\beta_r'$  between  $r_1$  and  $r_N$  gives a value of  $C_f^* = 92.33$  MHz km<sup>3</sup> sr  
 189  $\mu\text{J}^{-1}$ .

190 Once  $C_f^*$  has been calculated, a so-called pseudo-attenuated backscatter and its uncertainty  
 191 are solved as

$$\text{PAB}(z) = \frac{\text{NRB}(z)}{C_f^*}, \quad (6)$$

192 and

$$\delta\text{PAB}(z) = \text{PAB}(z) \sqrt{\left[ \frac{\delta\text{NRB}(z)}{\text{NRB}(z)} \right]^2 + \left( \frac{\delta C_f^*}{C_f^*} \right)^2}. \quad (7)$$

193 Eq. (7) is then modified by substituting the attenuated molecular backscatter for the PAB to  
 194 develop an objective threshold,

$$\alpha(z) = \beta_m(z)T_m^2(z) + \beta_m(z)T_m^2(z) \sqrt{\left[ \frac{\delta\text{NRB}(z)}{\beta_m(z)T_m^2(z)C_f^*} \right]^2 + \left( \frac{\delta C_f^*}{C_f^*} \right)^2}, \quad (8)$$

195 which is used to differentiate what are first presumed cloud returns from molecular return.

196 Range bins above  $r_N$  that meet the condition

$$\text{PAB}(z) - \delta\text{PAB}(z) > \alpha(z) \quad (9)$$

197 are then evaluated to determine whether they represent particulate layer base heights. Range

198 bins that do not meet the condition from Eq. (9) are used to establish “clear-air” slots and are

199 disqualified from the particulate base height search. In other words, a minimum detectable  
200 scattering ratio for particulates is defined as

$$\beta_{r,\min}(z) = \frac{\alpha(z) + \delta PAB(z)}{\beta'_m(z)}. \quad (10)$$

201 Running averages of PAB and  $\delta PAB$  are used in conjunction with two additional tunable  
202 thresholds,  $\phi$  and  $\kappa$  (both analogous to a SNR), to determine the actual layer base and top  
203 heights. The threshold  $\phi$  sets the minimum average value of PAB/ $\delta PAB$  for bins that exceed Eq.  
204 (10) in order to identify the layer base. At this point, and as described above, the UCDM  
205 assumes that any such layers detected are hydrometeor clouds, thereby leaving the potential for  
206 false detection of elevated aerosol layers. In the absence of supplementary information,  
207 however, such as color ratio (Liu et al. 2005) or depolarization (Cho et al. 2008, Omar et al.  
208 2009) and combined with the goal of resolving as much thin cirrus as possible in the low SNR  
209 portions of the NRB profile, this is unavoidable. Mitigation strategies are described further  
210 below.

211 The threshold  $\kappa$  sets the minimum average value of PAB/ $\delta PAB$  for bins that do not exceed  
212 Eq. (10) in order to identify clear air layers and consequently particulate layer tops. In Fig. 3, the  
213 red line indicates the threshold used to distinguish particulate from molecular returns. Bins that  
214 exceed this objective threshold are evaluated using the tunable threshold  $\phi$ , while bins with  
215 values lower than this threshold are evaluated using the tunable threshold  $\kappa$ . The sensitivity of  
216 V2 cloud detection to the tunable thresholds is evaluated in C08 and the values chosen for  $\phi$  and  
217  $\kappa$  will depend on the site location and instrument performance parameters.

218 3) V2 CLOUD RETRIEVALS

219 An example of V2 cloud retrievals is shown in the top panel of Fig. 4. Results from the  
220 GCDM and UCDM are integrated based on the noise altitude (described in Section 3.a.1).  
221 Clouds occurring below this height are reported from the GCDM. All clouds above the noise  
222 altitude are identified using the UCDM. At night, the noise altitude reaches above typical cirrus  
223 cloud heights at GSFC. Therefore, the GCDM is almost exclusively responsible for cloud  
224 detection. As a result, weakly-scattering cirrus can go undetected, since GCDM thresholds are  
225 tuned primarily with boundary layer phenomena in mind (i.e. suppression of aerosol  
226 identification). This can be observed frequently between 0300 and 0600 UTC in Fig. 4, where  
227 cirrus presence is underreported and cloud base heights are overestimated.

228 In the daytime, the noise altitude shown in Fig. 4 falls between 8–9 km and the UCDM is  
229 responsible for all cloud detection above it. In several instances, cloud bases (red markers) are  
230 shown while the corresponding cloud tops (orange markers) appear to be missing. In these cases  
231 the lidar signal is assumed to be significantly attenuated, and therefore no cloud top is reported.

### 232 *b. Version 3 cloud detection*

233 The V3 algorithm is based on V2 with a few meaningful changes to the UCDM.  
234 Consequently, the changes in V3 represent an update to C08 and how the GCDM and UCDM are  
235 merged. A schematic of the V3 cloud detection algorithm at the one-minute base NRB temporal  
236 resolution is shown in Fig. 5. Low-altitude obstructions (e.g. fog or low stratus decks) reduce  
237 SNR and limit the accuracy of cloud retrievals. Therefore, each profile is screened for these  
238 “beam-blocked” conditions by a process described further below. If no such obstructions are  
239 found, the first step in the UCDM is to calculate the normalization value.

#### 240 1) NORMALIZATION REGION

241 As mentioned in C08, it is most practical to find a normalization region to calculate  $C_f^*$   
242 nearest to the instrument in order to increase the depth of the profile analyzed for clouds. In V2,

243 the clear-sky search is initiated at 1 km AGL. However, in V3, an altitude of 5-km MSL is  
 244 chosen to reduce the likelihood of normalizing within relatively homogenous aerosol layers in or  
 245 near the boundary layer. Furthermore, based on recent global cloud and aerosol studies (Holz et  
 246 al. 2008; Sassen et al. 2008; Campbell et al. 2015), 5 km represents a reasonable height to begin  
 247 the search for high clouds in tropical and mid-latitude locations. A lower altitude may be needed  
 248 for polar sites, however. In the event that it is not possible to perform the normalization step at  
 249 5-km, the UCDM attempts to normalize lower in the atmospheric profile iteratively, to as low as  
 250 1 km AGL, as done in the V2 algorithm.

251 The normalization region also serves as the boundary between the GCDM and UCDM  
 252 retrievals in the V3 cloud algorithm, allowing the better-suited method (GCDM for low clouds  
 253 and UCDM for high clouds) to operate during both day and night. The V3 cloud retrievals in the  
 254 bottom panel of Fig. 4 can be compared with the V2 retrieval in the same figure to see the  
 255 relative apparent improvement. We also note that there are conditions when either method  
 256 (GCDM or UCDM) may be used to retrieve high-level or low-level clouds. Therefore, the  
 257 retrieval method for each cloud layer is provided as an output parameter.

## 258 2) OBJECTIVE THRESHOLD

259 The objective UCDM threshold,  $\alpha$ , defined in Eq. (8) is now adjusted for attenuation within  
 260 cloud layers in V3. In order to attenuate the  $\alpha$  threshold, we assume an initial unity transmission  
 261 at the base of the first detected cloud layer. Then the two-way transmittance is calculated at each  
 262 altitude bin  $Z$  within the cloud layer as

$$\begin{aligned}
 T_c^2(Z) &= T_c^2(Z-1) \exp[-2S_c \beta_c(Z) \Delta z] \\
 T_c^2(Z) &= T_c^2(Z-1) \exp\left[-2S_c \left(\frac{\beta_r'(Z)}{T_c^2(Z-1)} - 1\right) \beta_m(Z) \Delta z\right],
 \end{aligned}
 \tag{11}$$

263 where  $\beta_C$  is the cloud backscatter coefficient,  $S_C$  is the extinction-to-backscatter ratio (or lidar  
 264 ratio) of the cloud,  $\Delta z$  is the vertical resolution of the instrument, and  $\beta_r'$  is the attenuated  
 265 scattering ratio given by

$$\beta_r'(z) = \frac{\text{NRB}(z)}{C_f^* \beta_m'(z)} \approx \left[ 1 + \frac{\beta_C(z)}{\beta_m(z)} \right] T_C^2(z). \quad (12)$$

266 Overestimating the single unknown term in these equations, the extinction-to-backscatter  
 267 ratio, will lead to excessive attenuation of the objective threshold and eventual  
 268 mischaracterization of clear sky as cloud. Therefore, in this step, a very conservative choice for  
 269  $S_C$  equal to the molecular extinction-to-backscatter ratio ( $8\pi/3$ ) is used.

270 Once the transmittance has been determined, Eq. (8) is modified to recalculate the  $\alpha$   
 271 threshold as

$$\alpha(z) = \beta_m(z) T_m^2(z) T_C^2(z) + \beta_m(z) T_m^2(z) \sqrt{\left[ \frac{\delta \text{NRB}(z)}{\beta_m(z) T_m^2(z) C_f^*} \right]^2 + \left( \frac{\delta C_f^*}{C_f^*} \right)^2}. \quad (13)$$

272 By attenuating the  $\alpha$  threshold, we are able to better resolve cloud tops for optically-thick clouds.  
 273 Furthermore, multilayer clouds, where the upper layer was missed entirely by the V2 algorithm,  
 274 are now more effectively detected in some cases.

### 275 3) MULTI-TEMPORAL RESOLUTION

276 Cloud boundaries are only reported at one-minute temporal resolution in the V2 algorithm.  
 277 However, instances of high solar background reduce UCDM performance. So, as described by  
 278 C08, multi-temporal resolution settings are used in V3. In addition to the one-minute base  
 279 temporal resolution, the UCDM is performed for intermediate (five-minute) and long (twenty-  
 280 minute) temporal averages using a sliding window centered on a corresponding one-minute  
 281 profile.

282 A flow chart describing the process is shown in Fig. 6. Retrievals at the base temporal  
283 resolution are used to screen profiles at longer averages, thus limiting the effects of “beam-  
284 blocked” profile contamination. Within a window of  $N$  profiles,  $k$  profiles are removed from the  
285 average if an attenuating structure is detected below 5 km at the base temporal resolution. If  $k >$   
286  $N/2$ , then the entire average profile is rejected.

287 A combined cloud scene is created using cloud boundaries retrieved from the three temporal  
288 resolutions. First, the combined scene uses the cloud boundaries from the base temporal  
289 resolution. Next, the intermediate and then long temporal resolutions are used to fill in the  
290 missing gaps as needed. Cloud boundaries are always reported at the highest temporal resolution  
291 possible to reduce the effects of cloud smearing caused by averaging. This is broadly consistent  
292 with the methodology used for NASA Cloud Aerosol Lidar with Orthogonal Polarization  
293 (CALIOP) Level 2 products and their gridding of multiple spatial resolutions from the selective  
294 iterated boundary locator (SIBYL) to their vertical feature mask (Vaughan et al. 2005). A noted  
295 difference is that SIBYL also uses an intensity-clearing process to remove features detected at  
296 finer resolutions from the coarser spatial averages. No such intensity clearing is performed with  
297 the V3 algorithm.

#### 298 4) FALSE POSITIVES

299 As mentioned previously, use of the UCDM presumes first only cloud presence. Additional  
300 constraints are thus used to reduce the number of instances when noise excursions, elevated  
301 aerosol layers or poor normalizations produce false cloud retrievals. The first constraint  
302 establishes a minimum layer thickness of 150 m (i.e., two bins at 75-m resolution) in order to  
303 qualify a potential layer as a cloud. Therefore, we require all bins within a distance greater or  
304 equal to the minimum layer thickness to exceed the minimum detectable scattering ratio in Eq.  
305 (10) before a cloud base is established. Similarly, a minimum clear air distance of 150 m is used

306 to avoid falsely segmenting a single cloud into multiple layers. All bins within the minimum  
307 clear distance must fall below the minimum detectable scattering ratio in order to establish a  
308 cloud top. The second constraint requires that the standard deviation of the attenuated scattering  
309 ratio ( $\beta_r'$ ) within the detected layer exceed an empirically-determined threshold,  $\sigma_{\min}$ , which  
310 varies as a function of cloud top temperature. Cloud layers generally cause large variances in  $\beta_r'$ ,  
311 either through attenuation effects in otherwise colloidally-stable liquid water clouds or since ice  
312 crystals fall within cirrus cloud layers, creating complex structures. By contrast, aerosols in the  
313 free troposphere settle in stratified stable layers absent of convection and are expected to be  
314 homogeneous within each layer. Thus, clouds layers can be distinguished by their relatively  
315 large standard deviations of  $\beta_r'$ .

316 In order to determine  $\sigma_{\min}$ , a dataset was developed consisting of 144 days at GSFC in 2012  
317 and 27061 retrieved UCDM layers at one-minute resolution (18308 thin cirrus cloud layers, 3233  
318 non-cirrus layers, 5520 aerosol layers) when the particulate type could be reasonably identified  
319 from visual inspection (Fig. 7). Thin cirrus clouds are distinguished using a cloud top  
320 temperature threshold of  $-37$  °C (Sassen and Campbell 2001, Campbell et al. 2015) and a  
321 maximum cloud optical depth (COD) of 0.3 (Sassen and Cho 1992). The COD calculation uses  
322 a process described by Chew et al. (2011) and is discussed fully in Section 3.c.1. Non-cirrus  
323 clouds are those with cloud top temperature warmer than  $-37$  °C.

324 Several choices for  $\sigma_{\min}$  were evaluated using error matrices (Congalton and Meade 1986)  
325 and the corresponding values of accuracy and Matthews correlation coefficient (Matthews 1975;  
326 MCC) which are defined as

$$Accuracy = \frac{TP + TN}{TP + FP + FN + TN} \quad (14)$$



327 and  
328

$$MCC = \frac{TP \times TN - FP \times FN}{\sqrt{(TP + FP)(TP + FN)(TN + FP)(TN + FN)}}, \quad (15)$$

329 where  $TP$  is the number of instances when clouds were correctly identified,  $TN$  is the number of  
330 instances when aerosols were correctly identified,  $FP$  is the number of instances when aerosols  
331 were misidentified as clouds, and  $FN$  is the number of instances when clouds were misidentified  
332 as aerosols. The choice for  $\sigma_{\min}$  that resulted in the highest values of accuracy (0.92) and MCC  
333 (0.74) is given by

$$\sigma_{\min} = \begin{cases} 2, & \text{for } T > -37^\circ\text{C} \\ 10^{\frac{T+40}{10}}, & \text{for } -47^\circ\text{C} < T < -37^\circ\text{C} \\ 0.2, & \text{for } T < -47^\circ\text{C} \end{cases}. \quad (16)$$

334 The error matrix for the empirically-determined  $\sigma_{\min}$  is provided in Table 1. However, we note  
335 that the dataset contains two atypical long-range smoke transport events (with corresponding top  
336 height temperatures between  $-50^\circ\text{C}$  and  $-60^\circ\text{C}$ ), which comprised 35% of the aerosol category.  
337 If these two events are removed, the accuracy and MCC improve to 0.98 and 0.92, respectively.  
338 Similar results were observed at the five- and twenty-minute resolutions, but not shown for  
339 brevity.

340 The final constraint used to distinguish cloud from aerosol layers is that the estimated COD  
341 exceed a threshold,  $\tau_{\min}$ . Through empirical testing, we estimate COD and set  $\tau_{\min} = 0.005$  based  
342 on analysis of these subsets relative to the perception of how noise impacts these sub-samples  
343 combined with a similar analysis by Thorsen et al. (2011).

344 We briefly note here that lidars with polarization capabilities have recently been  
345 incorporated into the MPLNET project. However, because the overwhelming majority of

346 existing data (which spans from 2000 – present) was collected without polarization, the  
347 algorithm presented here does not rely on such data. It remains as a future goal to demonstrate  
348 how polarization can be used to improve aerosol-cloud discrimination, once a sufficient amount  
349 of data is collected from the new polarized sites.

### 350 *c. Version 3 algorithm output*

351 A listing of the V3 cloud detection algorithm output parameters are provided in Table 2.  
352 The output parameters from all temporal averages are gridded to one-minute temporal resolution,  
353 as previously described in the combined cloud scene. The number of cloud layers detected, day  
354 flag, and attenuation altitude are given as a single value each minute, characterizing the  
355 atmospheric column. All other cloud products and data flags correspond with individual cloud  
356 layers, and are provided each minute with dimensions equal to the number of cloud layers  
357 detected. Meteorological values at the cloud boundaries are obtained from the interpolated  
358 GEOS-5 profiles described in Section 2.

#### 359 1) CLOUD PHASE AND CIRRUS CLOUD OPTICAL DEPTH

360 In the absence of visual cloud observations, as is the case for autonomous lidar  
361 measurements made by MPLNET, Sassen and Campbell (2001) recommend using a minimum  
362 cloud top temperature of  $-37\text{ }^{\circ}\text{C}$  to identify cirrus. In the V3 cloud algorithm, we use this  
363 thermal threshold to distinguish ice clouds (i.e. cirrus) from all other cloud phases. Because  
364 depolarization capabilities are not standard for all MPLNET instruments, no attempt is made to  
365 distinguish liquid from mixed phase clouds.

366 Campbell et al. (2015) evaluate the  $-37\text{ }^{\circ}\text{C}$  cloud top temperature threshold globally versus  
367 the Level 2 CALIOP algorithms that identify ice-phase cloud layers and found that over 99% of  
368 clouds satisfying this thermal threshold were classified as ice. Furthermore, 81% of all ice  
369 clouds had cloud top temperatures less than  $-37\text{ }^{\circ}\text{C}$ . They conclude, consistent with the findings

370 of Sassen and Campbell (2001), that this thermal threshold is stable for specifically  
 371 distinguishing cirrus cloud presence in lidar studies that lack depolarization, though there is some  
 372 ambiguity in cases of “warm” cirrus that likely coincide with cloud top temperature greater than  
 373  $-37\text{ }^{\circ}\text{C}$ .

374 An estimated COD is calculated for clouds distinguished as cirrus using the procedure  
 375 described by Chew et al. (2011). Two-way cloud transmission is calculated using Eq. (11).  
 376 However, now the value of  $S_C$  is selected based on the cloud top temperature. Reported values  
 377 of  $S_C$  are on the order of 16–18 sr for liquid water clouds (Pinnick et al. 1983, Yorks et al. 2011)  
 378 and 10–40 sr for cirrus (Sassen and Comstock 2001; Chen et al. 2002; Yorks et al. 2011; Garnier  
 379 et al. 2015). A value of  $S_C = 18$  sr is chosen for layers with clouds top temperatures warmer than  
 380  $-37\text{ }^{\circ}\text{C}$  and  $S_C = 20$  sr at colder temperatures where cirrus clouds are expected. We note that due  
 381 to uncertainty in the lidar ratio for cirrus clouds, these estimates may represent the lower limit of  
 382 COD.

383 Next, Eq. (12) is used to solve for the cloud backscatter coefficient and the estimated COD  
 384 is given by

$$\tau = S_C \int_{base}^{top} \beta_C(z) dz = S_C \int_{base}^{top} \beta_m(z) \left( \frac{\beta_r'(z)}{T_C^2(z)} - 1 \right) dz. \quad (17)$$

385 The fidelity of the COD estimate is limited by the choice of  $S_C$  and accuracy of cloud boundaries  
 386 retrieved. The relative error in the lidar-derived optical depth is smallest for low optical depths  
 387 and proportional to  $\Delta S_C/S_C$  as  $\tau$  approaches zero (Winker et al. 2009). Lidar signals are unable to  
 388 penetrate through optically-thick clouds, which causes uncertainty in the value at the apparent  
 389 cloud top. In these cases, the estimated COD will be biased low. Similarly, attenuation from the  
 390 bottom-most cloud layer leads to uncertainty in corresponding retrievals of higher clouds for

391 lidar profiles containing multiple cloud layers. As a result, the estimated COD will be most  
392 reliable for single-layer, optically-thin clouds.

### 393 2) RETRIEVAL INDEX

394 While cloud boundaries are only reported at a single temporal resolution, a retrieval index is  
395 included to indicate whether the cloud was also detected at one or more of the other temporal  
396 averages. Cloud layers at different temporal resolutions are considered the same if (i) they share  
397 a common base or top height within a vertical depth of 250 m or (ii) one cloud layer is  
398 completely enveloped within the other.

399 An example of a combined cloud scene, with corresponding retrieval indices, is shown in  
400 Fig. 8. The value of the retrieval index is equal to the sum of the temporal resolutions used to  
401 identify the cloud layer. For example, if a cloud is detected at all three temporal resolutions, the  
402 value of the retrieval index is  $1+5+20 = 26$ . The advantage of the multi-temporal averaging  
403 scheme can be seen during the day between 1400–1500 UTC in Fig. 8. The elevated cloud layer  
404 (~15 km) is mostly undetected at the one-minute resolution, but can be resolved using the longer  
405 averages. The cloud layer at ~2 km produces “beam-blocked” conditions that prevent use of  
406 higher-temporal averages for much of the cirrus cloud layer above it.

### 407 3) ATTENUATION ALTITUDE

408 Because the lidar signal can become completely attenuated within optically-thick clouds, it  
409 is important to determine when a true cloud top is being reported as opposed to an apparent  
410 cloud top. Nadir-pointing lidar instruments have an advantage of using the ground return to  
411 determine if the lidar signal has been extinguished. However, with zenith-pointing lidar, that  
412 determination is more tenuous. Winker and Vaughan (1994) defined a transmittance index to  
413 determine when the lidar signal was fully attenuated based on the percentage of samples above  
414 the cloud top that exceeded the background. Other techniques used for zenith-pointing lidar

415 have included the use of a minimum threshold lidar signal along with its slope (Wang and Sassen  
416 2001) and comparisons with molecular profiles (Lo et al. 2006).

417 In V3, cloud tops (both true and apparent) are reported for all cloud layers along with the  
418 altitude at which the lidar signal is determined to be fully attenuated. This attenuation altitude is  
419 found by starting at the range bin of the highest reported cloud altitude and incrementally  
420 moving upwards in the profile until, within a depth of 2 km, (i) the percent difference between  
421 the mean pseudo-attenuated backscatter and modeled attenuated molecular backscatter falls  
422 below some threshold,  $T_1$ , and (ii) either the backscatter signal falls below a minimum value or  
423 the percentage of range bins where the backscatter signal is less than zero exceeds a threshold  $T_2$ .

424 This application pertains specifically to profiles that contain clouds or other obstructions,  
425 since the attenuation thresholds can also be satisfied by other conditions that lead to low SNR  
426 (e.g. high solar background). “Beam-block” conditions from low-altitude obstructions are found  
427 with the same search criteria, though the search is limited to the first 2 km above the surface.

#### 428 **4. Results**

429 In order to demonstrate the effects of the changes implemented in the V3 algorithm, we  
430 compare V2 and V3 cloud retrievals for one year at the GSFC MPLNET site. Table 3 and Fig. 9  
431 show data sampling statistics for 2012, including the total number of profiles and percentage of  
432 time when 1-minute NRB measurements were available monthly. Observable profiles are given  
433 as the number and percentage of available profiles that are not “beam-blocked” below 2 km,  
434 MSL. Profile attenuation was determined using the V2 method because it is the most restrictive  
435 and ensures an even comparison between the two cloud detection algorithms. The diurnal  
436 distribution of data recorded and successful V3 normalizations to calculate  $C_f^*$  are also shown.  
437 There is very little differentiation between the distributions of V2 and V3 normalizations, so only

438 the number of successful V3 normalizations are shown for simplicity. Though MPLNET  
439 measurements are continuous, there is a decrease in the number of successful normalizations  
440 near solar noon in relatively low SNR conditions.

441 Because V3 uses a merged cloud scene and V2 is only processed at one-minute resolution,  
442 V3 retrievals are evaluated using the base one-minute resolution (hereafter V3b) and the merged  
443 cloud scene (hereafter V3m). Comparisons are limited to cloud base statistics because cloud  
444 tops are not recorded for all V2 retrievals. Finally, we describe the macrophysical and optical  
445 characteristics of cirrus clouds observed during this study, again adhering to the methodology  
446 described in Campbell et al. (2015), using V3 retrievals.

#### 447 *a. Vertical dependence*

448 Figure 10 shows the cloud base distributions retrieved from the V2 and V3 algorithms,  
449 respectively, at GSFC during 2012. A bimodal distribution similar to that observed by Winker  
450 and Vaughan (1994), with peaks at ~1-2 km and ~9-10 km is apparent. The total numbers of  
451 cloud observations are 269505, 304363, and 332810 for the V2, V3b, and V3m retrievals,  
452 respectively. Compared with V2, the number of cloud observations increases by 12.9% and  
453 23.5% for V3b and V3m, respectively. The largest increase in the number of clouds observed  
454 occurs at altitudes above 5 km.

455 Because the difference in the number of clouds retrieved shows a clear vertical dependence,  
456 we examine them specifically for three sub-samples, by defining low clouds as those with base  
457 heights less than 2 km, high clouds as those with base heights greater than 5 km, and middle  
458 clouds as those with base heights between 2-5 km (WMO 1975). The number of lidar profiles  
459 for each classification, along with occurrence frequency, is shown in Table 4. Cloud occurrence  
460 frequency is defined as the number of lidar profiles containing a particular cloud classification

461 divided by the total number of observable profiles. Regardless of the retrieval method (V2, V3b,  
462 and V3m), occurrence frequency is nearly identical for low clouds, which reflects the relative  
463 consistency in GCDM application between V2 and V3 at one-minute resolution. High clouds  
464 show the largest increase in occurrence frequency. For example, comparing the V2 and V3m  
465 algorithms, the occurrence frequency of high clouds increases by 5.9% attributable to (i) the  
466 increased identification of elevated, multi-layer cloud decks using an attenuated UCDM  
467 threshold, (ii) increased use of the UCDM to identify high clouds at day and night, and (iii)  
468 multi-temporal application of UCDM to increase SNR. To (i), V2 retrievals resulted in 91% of  
469 cloudy lidar profiles containing single-layer clouds. The percentage of single-layer clouds  
470 decreases to 83% and 80% for V3b and V3m, respectively.

#### 471 *b. Seasonal dependence*

472 Figure 11 shows the annual cycle for low, middle, high, and total cloud classifications  
473 during 2012. The low cloud occurrence frequency is nearly identical for all three retrieval  
474 methods. Middle clouds retrieved using V3b and V3m exhibit a slight separation from V2. The  
475 largest differences are again seen with high-cloud retrievals. While the annual cycles for high  
476 clouds show similar patterns for all three retrievals, there is an increase in occurrence frequency  
477 of ~3% and 6% for V3b and V3m, respectively. The increase in high-cloud occurrence  
478 frequency when compared to V2 ranges from 1% to 4% using V3b and 4% to 10% for V3m.  
479 The largest differences for high-cloud occurrence frequency between V2 and V3 occurs during  
480 summer months, which is coincident with the period when the sun is at its highest elevation and  
481 thus solar background is highest.

#### 482 *c. Diurnal dependence*

483 Differences in the diurnal cycle show similar characteristics as the annual cycles for low and  
484 middle clouds. As seen in Fig. 12, V2, V3b, and V3m are nearly identical for low clouds. While  
485 V3b and V3m show slight differences from V2, they are indistinguishable from each other.  
486 High-cloud diurnal cycles follow the same trends for all three retrievals. However, the cloud  
487 occurrence frequency is higher for V3b and highest for V3m. No clear diurnal trend is apparent  
488 at GSFC because some changes (e.g. using the UCDM at all times) affect both day and night  
489 retrievals. At tropical sites, where the solar background is higher and longer temporal averaging  
490 is necessary, there may be a more obvious diurnal trend.

#### 491 *d. Macrophysical and optical cirrus properties*

492 Based on the greater detection of high clouds demonstrated above, we characterize cirrus  
493 clouds over the GSFC site as detected by the V3 algorithm. As stated earlier, cirrus presence is  
494 determined using a cloud top temperature threshold of  $-37\text{ }^{\circ}\text{C}$ . Additionally, we limit the  
495 analysis to cases when (i) only cirrus clouds (no underlying liquid water or mixed phase clouds)  
496 were detected in the profile, (ii) the estimated COD was less than 3, based on the upper-limit for  
497 cirrus clouds suggested by Sassen and Cho (1992), and (iii) the attenuation altitude was at least 2  
498 km above the cloud top. The final constraint limits the analysis to “transparent cirrus” cases for  
499 which the algorithm is more likely to identify the true cloud top.

500 The resulting dataset includes 57930 cirrus clouds. The majority of cloud detections (82%)  
501 occur at the base one-minute temporal resolution. The largest occurrence rate of the coarse  
502 temporal averages occurs at or near noon and during the summer months when the solar  
503 background is highest.

504 Table 5 summarizes the seasonal and annual mean characteristics of the transparent cirrus  
505 dataset. The monthly variation in the macrophysical properties is shown in Fig. 13. Cirrus



506 clouds over GSFC tend to be higher and thinner (geometrically and optically) in the spring and  
507 summer and lower and thicker in the fall and winter seasons. Cirrus also occur more frequently  
508 in the spring and summer months. The transparent cirrus dataset is composed almost entirely  
509 (~95%) of sub-visual ( $COD < 0.03$ ) and thin ( $COD < 0.3$ ) cirrus clouds. Uncertainties in the  
510 value of the extinction-to-backscatter ratio and cloud top height could lead to an exaggeration of  
511 this finding. However, it should be noted that when the extinction-to-backscatter ratio is  
512 increased from 20 sr to 30 sr (not shown), 86% of cirrus clouds still have a COD less than 0.3.  
513 These results are qualitatively consistent with the findings of Dupont et al. (2010), who reported  
514 50-75% of non-opaque cirrus clouds had an optical thickness less than 0.3 based on ground-  
515 based lidar and CALIPSO observations at four mid-latitude sites.

516 Frequency distributions of the optical and macrophysical properties are presented in Fig. 14.  
517 The cloud optical depth peaks in the sub-visual range and has a positive skew. The transparent  
518 cirrus dataset suggests that the limit at which we are able to resolve molecular signal above  
519 cloud, and thus reliably determine the cloud top, occurs near a COD of 0.5. However, if an  
520 extinction-to-backscatter ratio of 30 sr is used, this limit occurs near a COD of 0.8.

521 A comparison of daytime and nighttime cloud retrievals is provided in Table 6. There are  
522 only slight differences in the occurrence frequency between day and night cases. However, the  
523 geometric and optical depths are considerably lower in the daytime. The thinning of daytime  
524 cirrus may be attributable to difficulty in correctly identifying cloud boundaries due to solar  
525 background effects (Thorsen et al. 2013). However, convective cloud remnants are also likelier  
526 to occur during daytime hours. Decoupling the two, aside from seasonal influence, is outside the  
527 scope of this analysis. In the same manner, the daytime retrievals are more likely to be

528 considered as totally attenuated due to the higher solar background reducing the possibility to  
529 resolve molecular signal at cirrus heights.

## 530 **5. Summary and Discussion**

531 A new Version 3 (V3) cloud detection algorithm has been developed within the NASA  
532 Micropulse Lidar Network (MPLNET) that uses a combination of retrieval methods and a multi-  
533 temporal averaging scheme. Most V3 changes represent updates to the Version 2 (V2)  
534 uncertainty-based threshold algorithm introduced by Campbell et al. (2008). The threshold used  
535 to identify cloud presence now accounts for attenuation losses within cloud layers, which allows  
536 for better estimation of cloud tops and boundaries of overlying cloud layers in profiles where  
537 multiple cloud layers are detected. A more synergistic merging of the gradient-based cloud  
538 detection method (GCDM) and uncertainty-based cloud detection method (UCDM) improves  
539 nighttime clouds detection of tenuous high clouds. The incorporation of coarser temporal  
540 resolutions at intermediate (5-minute) and long (20-minute) averages improves detection in  
541 situations with low SNR (e.g. high solar background). One year of data at the NASA Goddard  
542 Space Flight Center (GSFC) in Greenbelt, MD is used to show the effect of these updates on  
543 cloud retrievals.

544 The largest impact of the changes to the cloud detection algorithm is evident with high  
545 clouds (those with cloud base  $> 5$  km), while the diurnal and annual cycles of low and middle  
546 clouds exhibit only slight changes from V2 to V3. The high-cloud occurrence frequency  
547 increases by nearly 6% at GSFC when using the V3 merged cloud scene compared with the V2  
548 retrieval. Furthermore, the ability to detect multi-layered cloud scenes is improved with the V3  
549 algorithm. The results show that 91% of clouds in 2012 at the NASA GSFC project site were  
550 recorded as single-layer clouds according to the V2 retrieval compared with 80% for V3.

551 A brief investigation of macrophysical properties for transparent cirrus clouds shows that  
552 the mean base and top heights at GSFC occur at  $10.17 \pm 1.63$  km and  $11.07 \pm 1.43$  km,  
553 respectively. The highest and thinnest (both geometrically and optically) cirrus are found during  
554 the spring and summer months, which was coincident with the highest cirrus occurrence  
555 frequency. There is no significant difference in occurrence frequency between daytime and  
556 nighttime retrievals. However, cirrus clouds are thinner (both geometrically and optically) in  
557 daytime than nighttime, which may be attributed to a combination of increased uncertainty due  
558 to the solar background effects and higher occurrence of convective cloud remnants during the  
559 day. Notably, the limit to which we are able to resolve molecular signal above cirrus clouds  
560 occurs between cloud optical depths of 0.5 and 0.8, allowing for uncertainty in the extinction-to-  
561 backscatter ratio.

562 The value of the MPLNET cloud datasets is in its continuous (both day and night) and long-  
563 term measurements at polar, mid-latitude, and tropical sites using a standard instrument and data  
564 processing algorithm. Incorporating the V3 cloud retrievals from MPLNET as part of a multi-  
565 instrument investigation will enhance our current knowledge of clouds, in particular cirrus. As it  
566 stands, the cloud products provide a unique validation dataset for the modeling community and  
567 satellite measurements. With some MPLNET sites now well into their second decade of  
568 continuous cloud and aerosol observations, the project has become an integral component of  
569 ground-based evaluation of atmospheric processes and verification of NASA satellite missions.  
570 This paper thus represents our continuing effort to optimize the fidelity of project datasets for the  
571 benefit of the community and in sustaining general scientific inquiry.

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730 **List of Tables**

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731

732 TABLE 1. Error matrix for cloud-aerosol discrimination using  $\sigma_{\min}$

	Predicted cloud	Predicted aerosol
True cloud	21242	299
True aerosol	1851	3669

733

734 Table 2. MPLNET V3 cloud detection algorithm output

Parameter	Description
Base and top altitudes	Cloud boundaries in km above MSL
$N_{\text{layers}}$	Number of cloud layers detected
$N_{\text{profiles}}$	Number of 1-min profiles used in average profile
Method of cloud detection	GCDM or UCDM
Cloud phase	Ice or liquid/mixed phase based on $-37^{\circ}\text{C}$ threshold
Estimated cirrus COD	Calculated using lidar ratios of 20 and 30 sr, ranging from 0.005 to 3
Base and top temperatures	From FP-IT GEOS-5 Version 5.9.1
Base and top pressures	From FP-IT GEOS-5 Version 5.9.1
Base and top wind speeds	From FP-IT GEOS-5 Version 5.9.1
Base and top wind directions	From FP-IT GEOS-5 Version 5.9.1
Day/night flag	0 = Night, 1=Day (from ephemeris)
Retrieval index	Indicates temporal average used to identify the cloud layer
Attenuation altitude	Altitude where lidar beam is deemed to be significantly attenuated

735

TABLE 3. Summary of data collected at GSFC in 2012 (percentages shown in parentheses)

	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec	Year
Profiles	44615	41690	44611	43142	32095	41479	40817	44485	43078	29081	42946	44345	536999
recorded	(100)	(100)	(100)	(100)	(72)	(96)	(91)	(100)	(100)	(65)	(99)	(99)	(93)
Observable	38022	38226	37303	39971	25870	38138	35339	39318	38104	23412	41265	35175	468165
profiles	(85)	(92)	(84)	(93)	(81)	(92)	(87)	(88)	(88)	(81)	(96)	(79)	(87)
V3 normalizations	34798	34882	35248	36090	23644	35801	34303	39408	37158	20649	36224	31041	399246
	(78)	(84)	(79)	(84)	(74)	(86)	(84)	(89)	(86)	(71)	(84)	(70)	(74)

736 TABLE 4. Number of lidar profiles and occurrence frequency at GSFC in 2012

V2	V3b	V3m
<i>Low clouds</i>		
77943	78511	78513
0.181	0.183	0.183
<i>Middle clouds</i>		
69612	75353	75462
0.162	0.175	0.175
<i>High clouds</i>		
109366	120750	134514
0.254	0.281	0.313
<i>Total clouds</i>		
246132	255079	267777
0.572	0.593	0.623

737



738 TABLE 5. Transparent cirrus cloud properties

	Spring (MAM)	Summer (JJA)	Fall (SON)	Winter (DJF)	Annual
Cloud base					
Height (km)	10.46 ± 1.44	11.17 ± 1.48	9.77 ± 1.33	9.13 ± 1.49	10.17 ± 1.63
Temperature (°C)	-50.3 ± 9.7	-47.9 ± 9.2	-43.9 ± 8.9	-45.7 ± 10.2	-47.4 ± 9.9
Pressure (mb)	262.6 ± 57.6	243.9 ± 52.0	291.0 ± 53.7	312.4 ± 66.8	275.9 ± 63.7
Wind speed (m s <sup>-1</sup> )	23.3 ± 12.5	20.3 ± 11.0	26.6 ± 10.7	38.6 ± 14.5	27.1 ± 14.2
Wind direction (°)	277.2 ± 40.2	236.8 ± 110.3	259.4 ± 28.3	279.1 ± 20.5	264.6 ± 63.3
Cloud top					
Height (km)	11.39 ± 1.35	11.93 ± 1.34	10.74 ± 1.16	10.09 ± 1.33	11.07 ± 1.48
Temperature (°C)	-57.0 ± 8.1	-53.5 ± 8.0	-51.3 ± 7.3	-52.9 ± 8.4	-54.1 ± 8.3
Pressure (mb)	225.9 ± 47.0	215.4 ± 41.6	249.9 ± 41.2	267.6 ± 52.4	238.4 ± 50.5
Wind speed (m s <sup>-1</sup> )	24.5 ± 12.9	21.8 ± 11.8	29.6 ± 11.4	42.5 ± 15.6	29.3 ± 15.4
Wind direction (°)	277.2 ± 37.4	238.3 ± 108.0	259.5 ± 29.6	279.3 ± 20.6	265.0 ± 61.7
Cloud depth (km)	0.93 ± 0.61	0.76 ± 0.51	0.97 ± 0.70	0.96 ± 0.69	0.90 ± 0.63
Estimated COD	0.07 ± 0.10	0.06 ± 0.10	0.09 ± 0.14	0.09 ± 0.12	0.08 ± 0.11
Cirrus Type (%)					
Sub-visual	47	51	45	41	46
Thin	50	46	49	53	49
Opaque	3	3	6	6	5
Occurrence (%)	17	12	10	13	13
Cloud layers	18664	13889	10769	14608	57930

739 TABLE 6. Daytime and nighttime transparent cirrus properties

	Daytime	Nighttime
Cloud base		
Height (km)	$10.15 \pm 1.69$	$10.19 \pm 1.58$
Temperature ( $^{\circ}\text{C}$ )	$-47.3 \pm 10.0$	$-47.5 \pm 9.8$
Pressure (mb)	$277.2 \pm 65.8$	$274.9 \pm 61.9$
Wind speed ( $\text{m s}^{-1}$ )	$27.1 \pm 14.0$	$27.0 \pm 14.5$
Wind direction ( $^{\circ}$ )	$258.6 \pm 70.4$	$269.6 \pm 56.2$
Cloud top		
Height (km)	$10.94 \pm 1.51$	$11.19 \pm 1.44$
Temperature ( $^{\circ}\text{C}$ )	$-53.2 \pm 8.4$	$-54.8 \pm 8.1$
Pressure (mb)	$243.5 \pm 51.4$	$234.0 \pm 49.3$
Wind speed ( $\text{m s}^{-1}$ )	$29.1 \pm 15.1$	$29.5 \pm 15.6$
Wind direction ( $^{\circ}$ )	$259.8 \pm 68.3$	$269.4 \pm 55.2$
Cloud depth (km)	$0.79 \pm 0.56$	$1.00 \pm 0.67$
Estimated COD	$0.07 \pm 0.11$	$0.09 \pm 0.12$
Cirrus Type (%)		
Sub-visual	51	42
Thin	46	52
Opaque	3	6
Occurrence (%)	12	14
Cloud layers	26360	31570

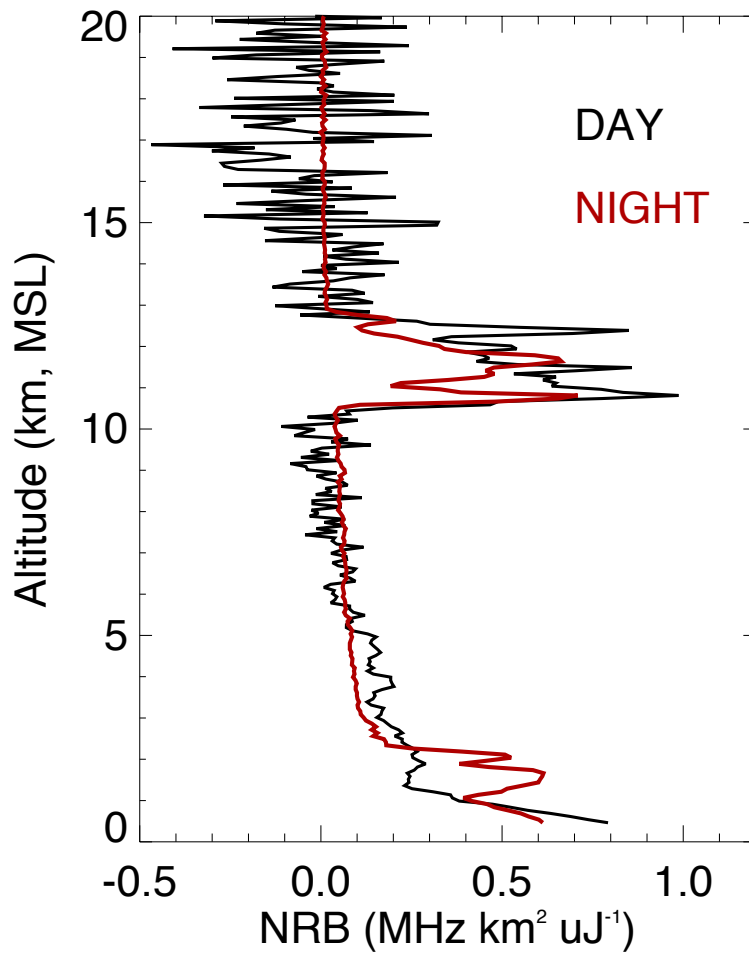
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- 5 Schematic of the V3 cloud detection algorithm at the one-minute base resolution. 49
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- 7 (Left) Representative groupings of cirrus clouds, non-cirrus clouds, and aerosols 51  
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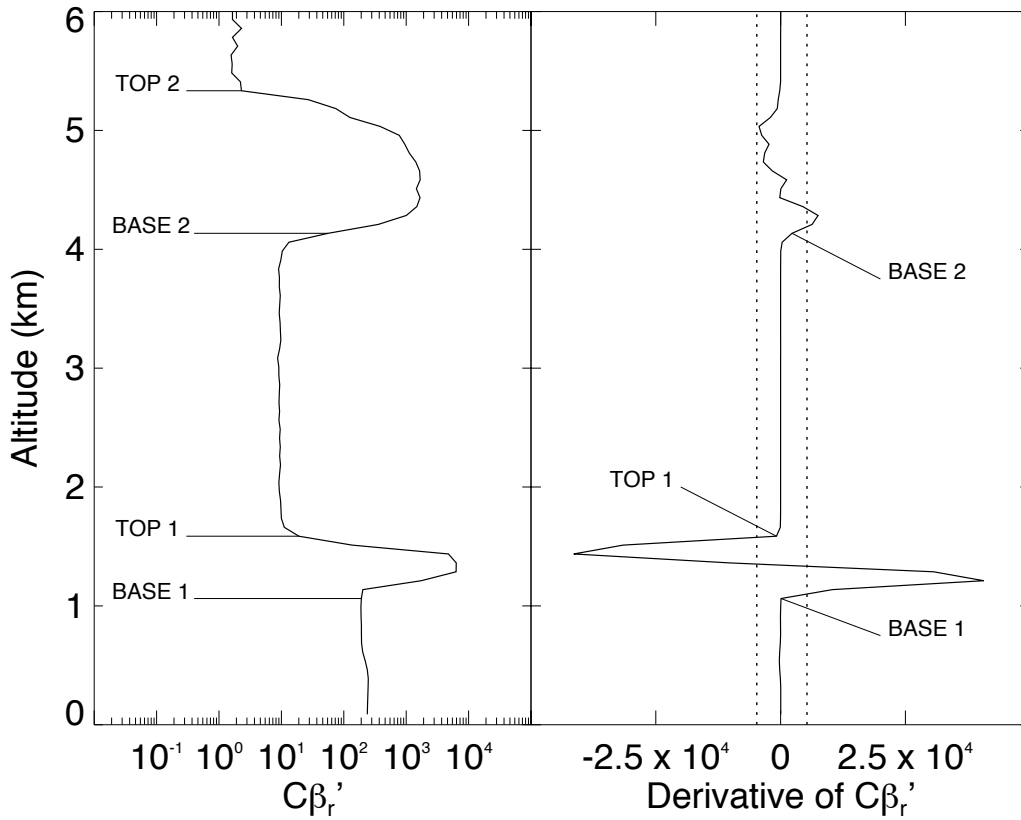
- boundaries encloses 50% of the data nearest the median and thick boundaries enclose 90% of the data. The dashed line indicates the empirically-determined  $\sigma_{\min}$ .
- 8 (Top) Example of a combined cloud seen at the GSFC MPLNET site on 7 July 2012. (Bottom) Retrieval flags showing the temporal resolutions used to detect the combined cloud scene. Square symbols are used to indicate the cloud base and the Xs are used for the cloud top. 52
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- 14 (a) Cloud base altitude, (b) cloud top altitude and (c) cloud depth for the transparent cirrus dataset. The colors indicate the cirrus type based on the estimated COD (blue for sub-visual, yellow for thin, and red for opaque cirrus). Horizontal axis bin size 58

is 0.5 km for base and top altitudes and 0.25 km for cloud depth. (d) Frequency distribution (black) and cumulative frequency distribution (red). The vertical dashed lines indicate the thresholds for sub-visual (0.03) and thin (0.3) cirrus clouds. Horizontal axis bin size equals 0.001.

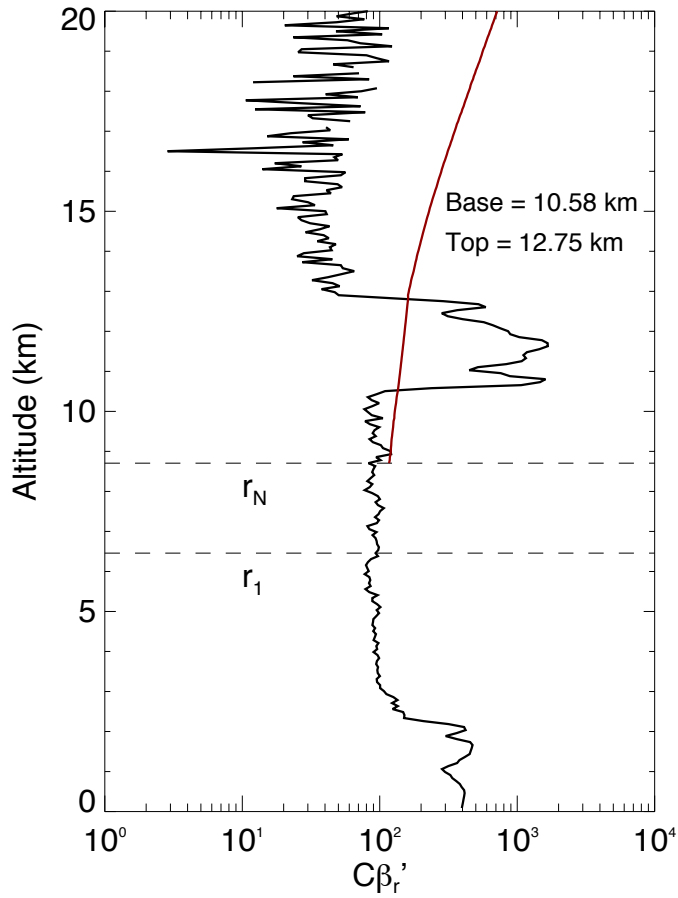
741



742  
743 FIG. 1. Examples of the NRB at GSFC on 7 July 2012 at day (1800 UTC, black) and night (0300  
744 UTC, red).

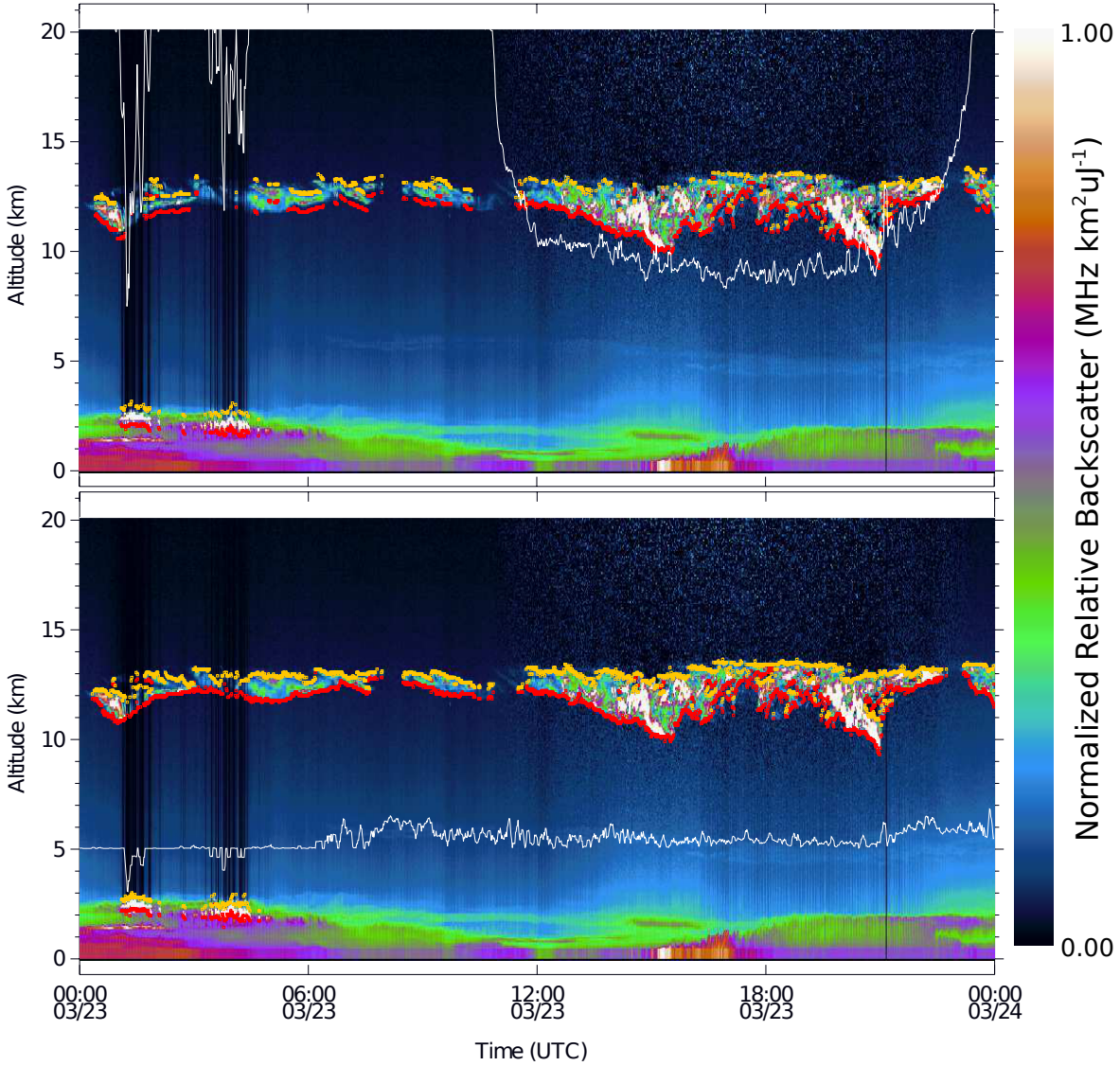


745  
 746 FIG. 2. Representation of the GCDM (GSFC; 0233 UTC 7 Nov 2012). The product of the  
 747 calibration constant and attenuated scattering ratio is shown in the left panel and its derivative is  
 748 shown on the right. The vertical dashed lines represent the  $a_{\min}$  and  $a_{\max}$  thresholds.

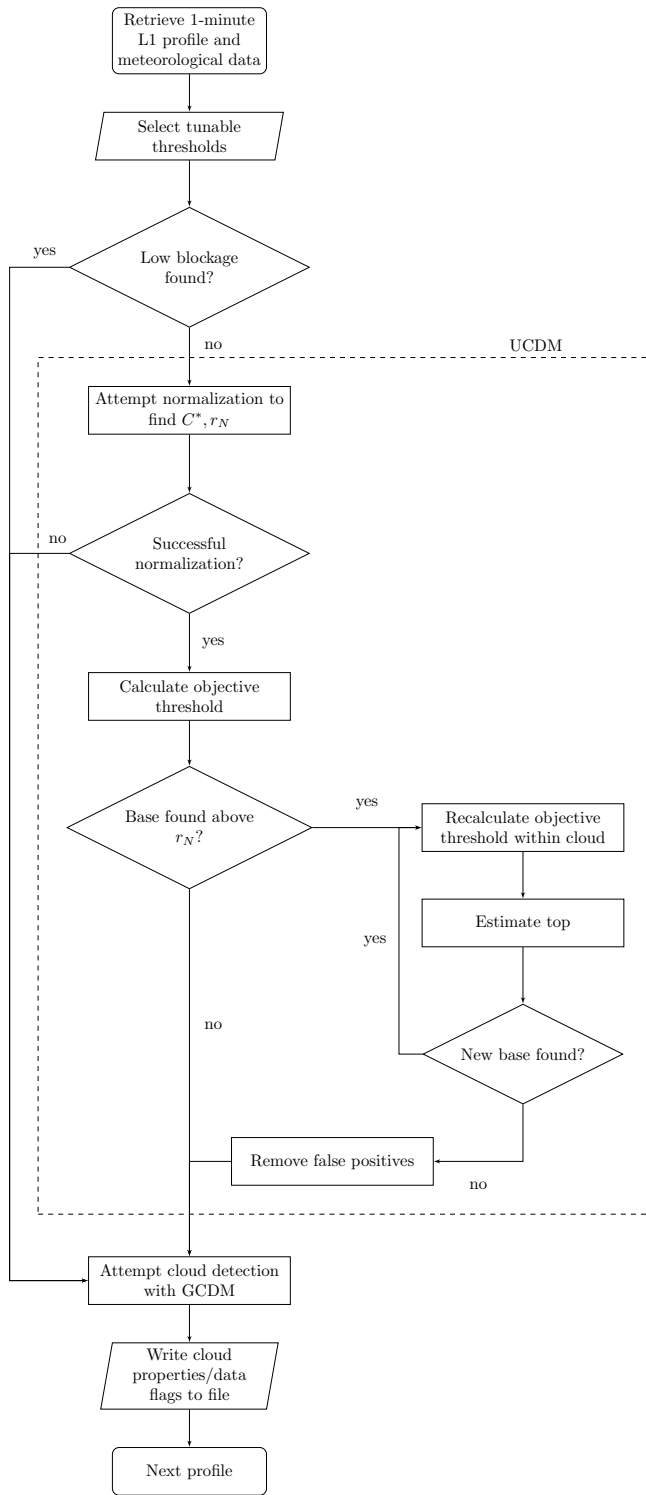


749  
 750 FIG. 3. Representation of the UCDM (GSFC; 0300 UTC 7 July 2012). The bottom ( $r_1$ ) and top  
 751 ( $r_N$ ) of the normalization region are indicated by the horizontal dashed lines. The red curve  
 752 represents the product of  $C_f^*$  and the minimum detectable scattering ratio used to determine the  
 753 cloud boundaries.

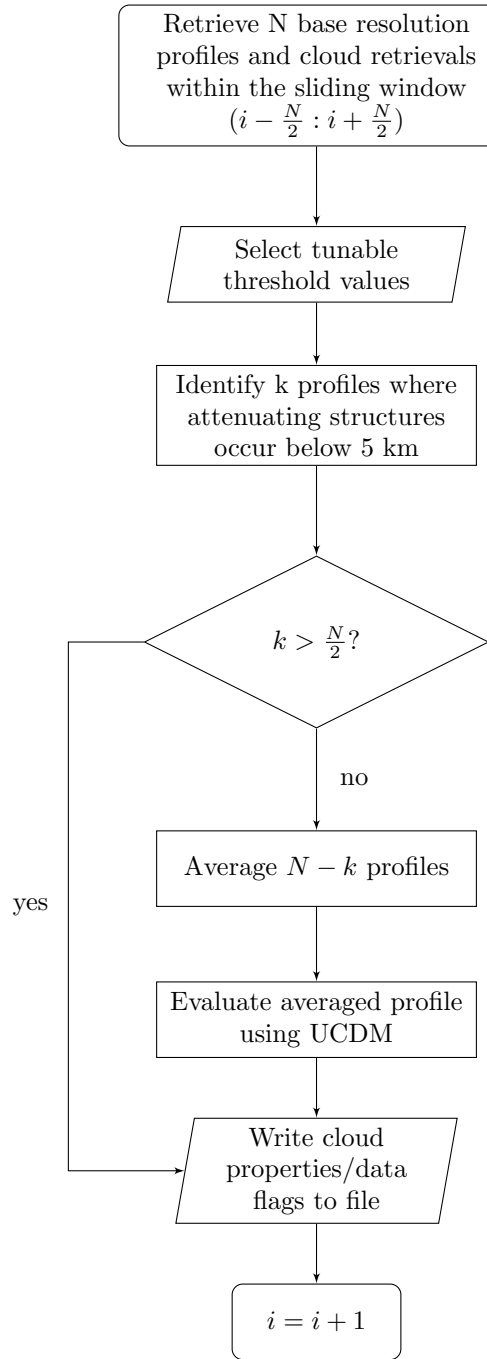




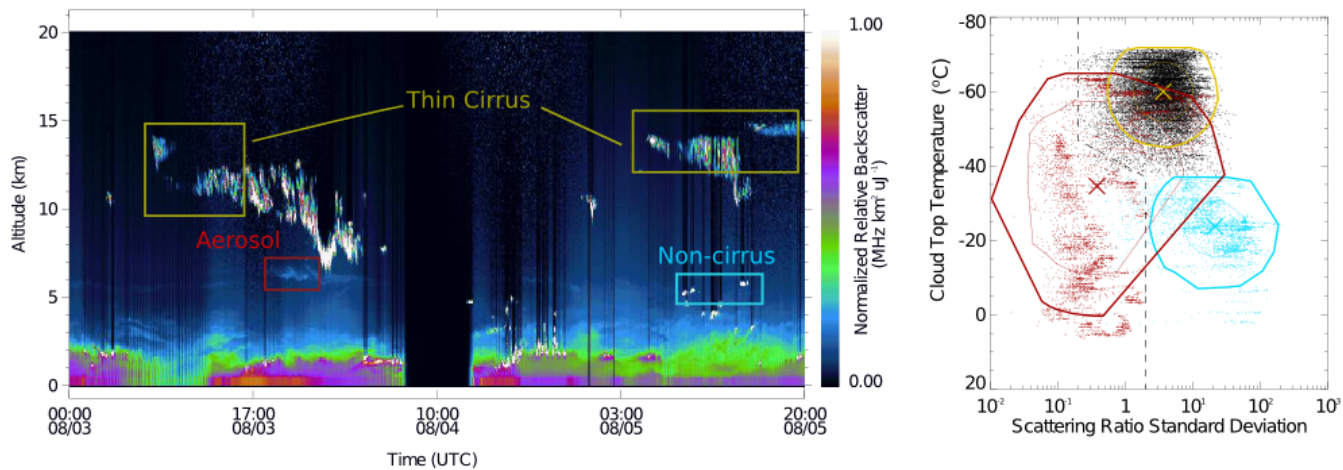
754  
 755 FIG. 4. (Top) Version 2 cloud retrievals at GSFC site on 23 March 2012 (one-minute resolution).  
 756 Cloud bases and tops are indicated by red and orange markers, respectively. The noise altitude  
 757 (i.e.  $\delta\text{NRB}/\text{NRB} > 0.5$ ) is represented by the solid white line (four-minute smoothing applied for  
 758 clarity). (Bottom) Same as top figure but using the V3 algorithm at one-minute resolution. The  
 759 bottom of the normalization region is represented by the solid white line (four-minute smoothing  
 760 applied for clarity).



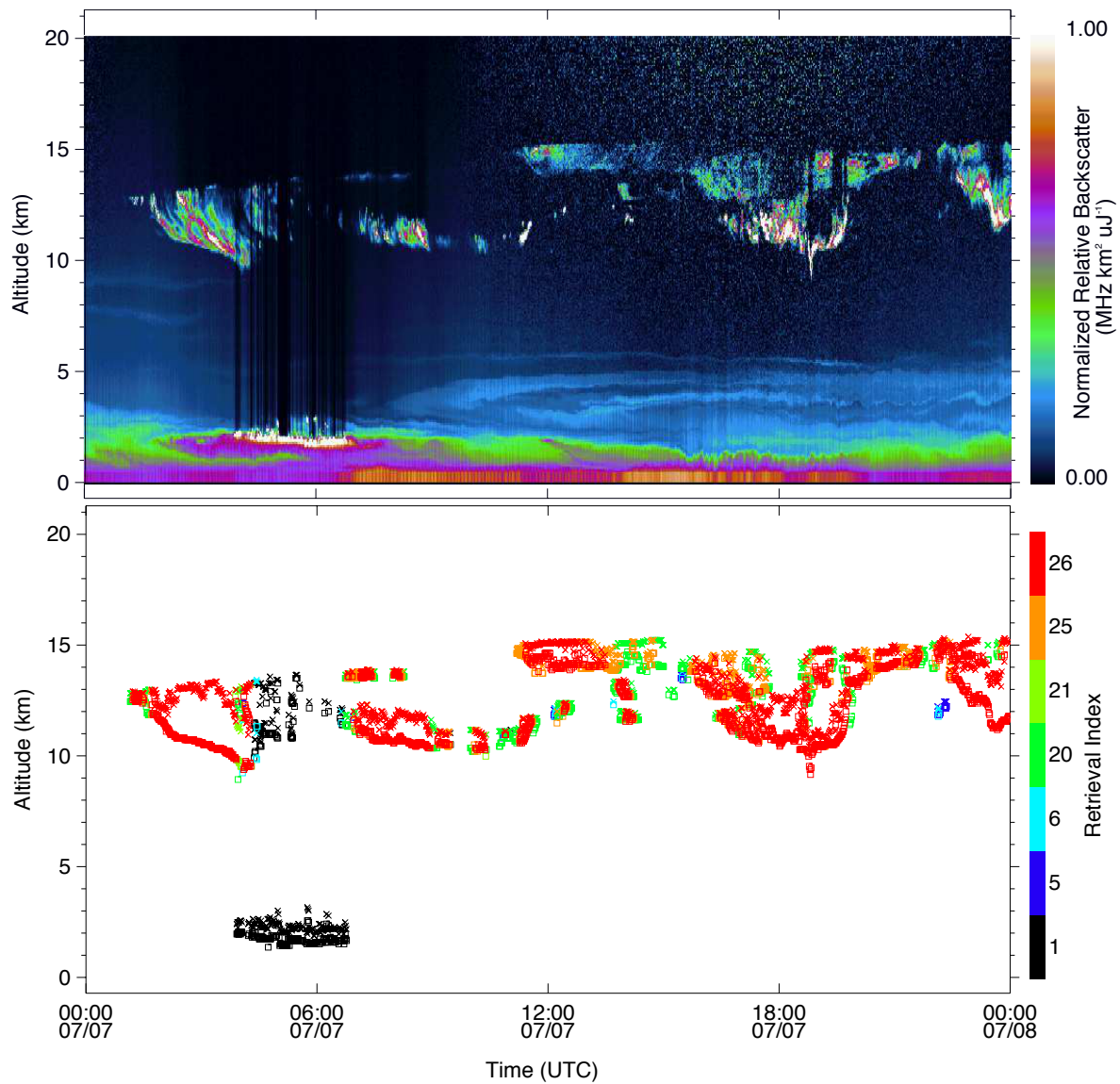
761  
762 FIG. 5. Schematic of the V3 cloud detection algorithm at the one-minute base resolution.



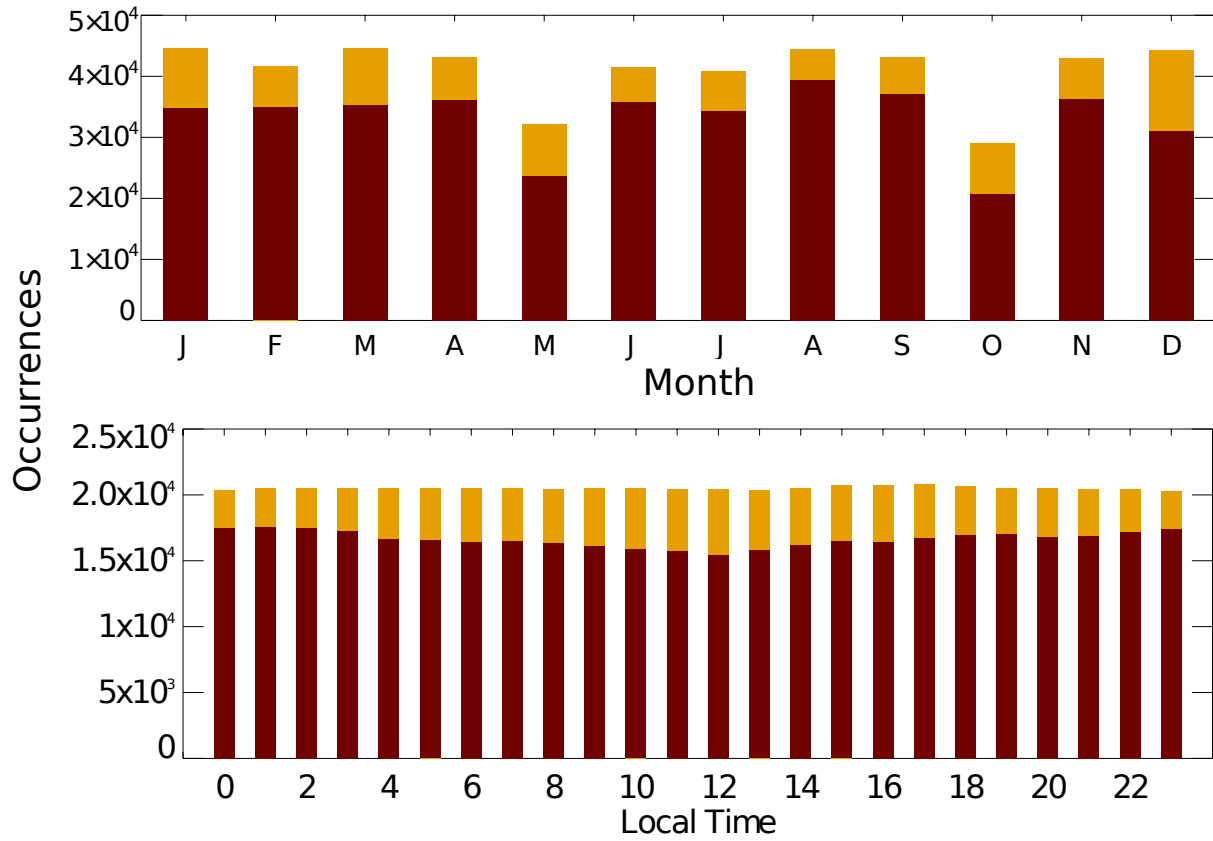
763  
764 FIG. 6. Schematic of the V3 cloud detection algorithm for higher temporal averages.



765  
 766 FIG. 7. (Left) Representative groupings of cirrus clouds, non-cirrus clouds, and aerosols from  
 767 GSFC (3-5 Aug 2012). (Right) Scatterplot of full dataset used to determine  $\sigma_{\min}$  threshold. The  
 768 Xs represent the median values of each group, the thin boundaries enclose 50% of the data  
 769 nearest the median and thick boundaries enclose 90% of the data. The dashed line indicates the  
 770 empirically-determined  $\sigma_{\min}$ .

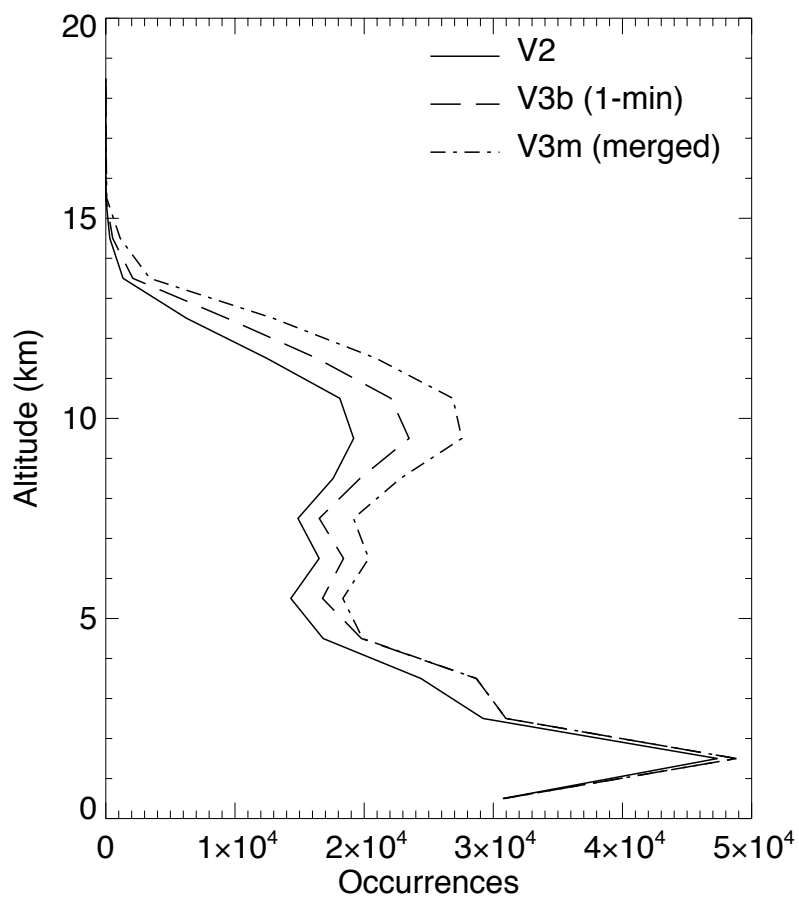


771  
 772 FIG. 8. (Top) NRB from lidar measurements at the GSFC MPLNET site on 7 July 2012.  
 773 (Bottom) Retrieval indices showing the temporal resolutions used to detect the combined cloud  
 774 scene. Square symbols are used to indicate the cloud base and the Xs are used for the cloud top.

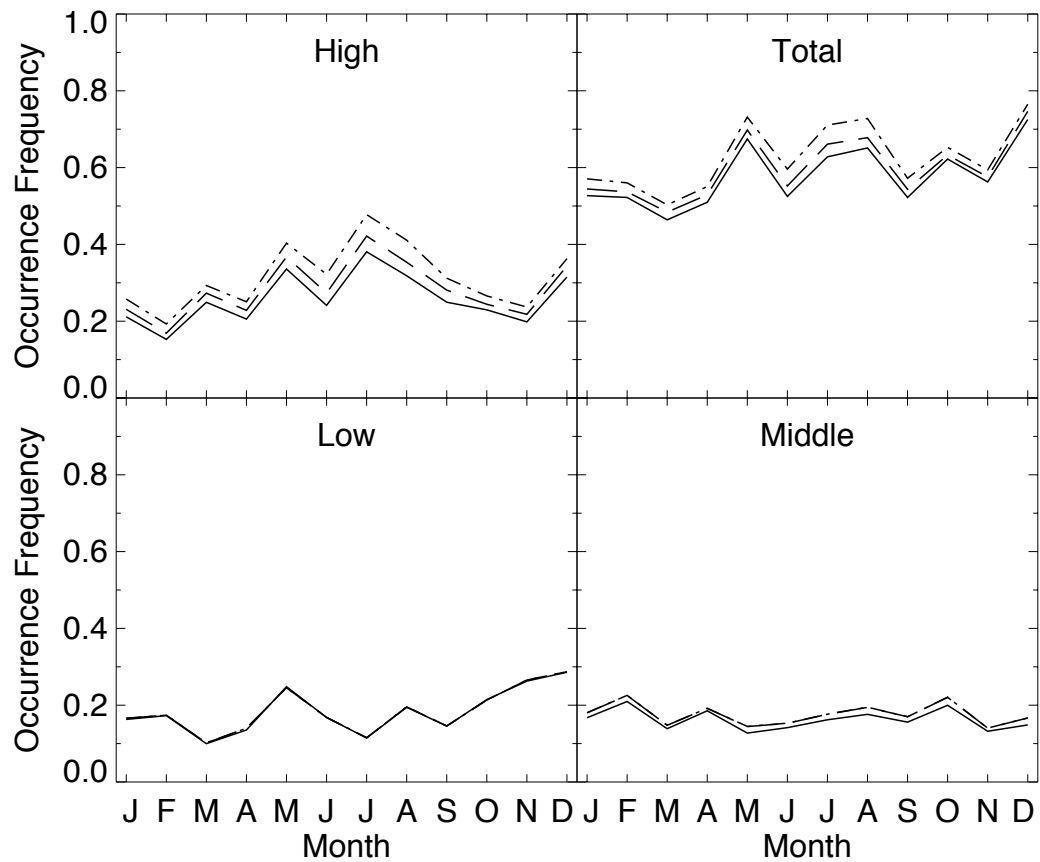


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FIG. 9. (Top) Number of minutes when data was recorded by month. (Bottom) The number of minutes data was recorded each hour. The red colors are the successful  $C_f^*$  normalizations and yellow are the total data.

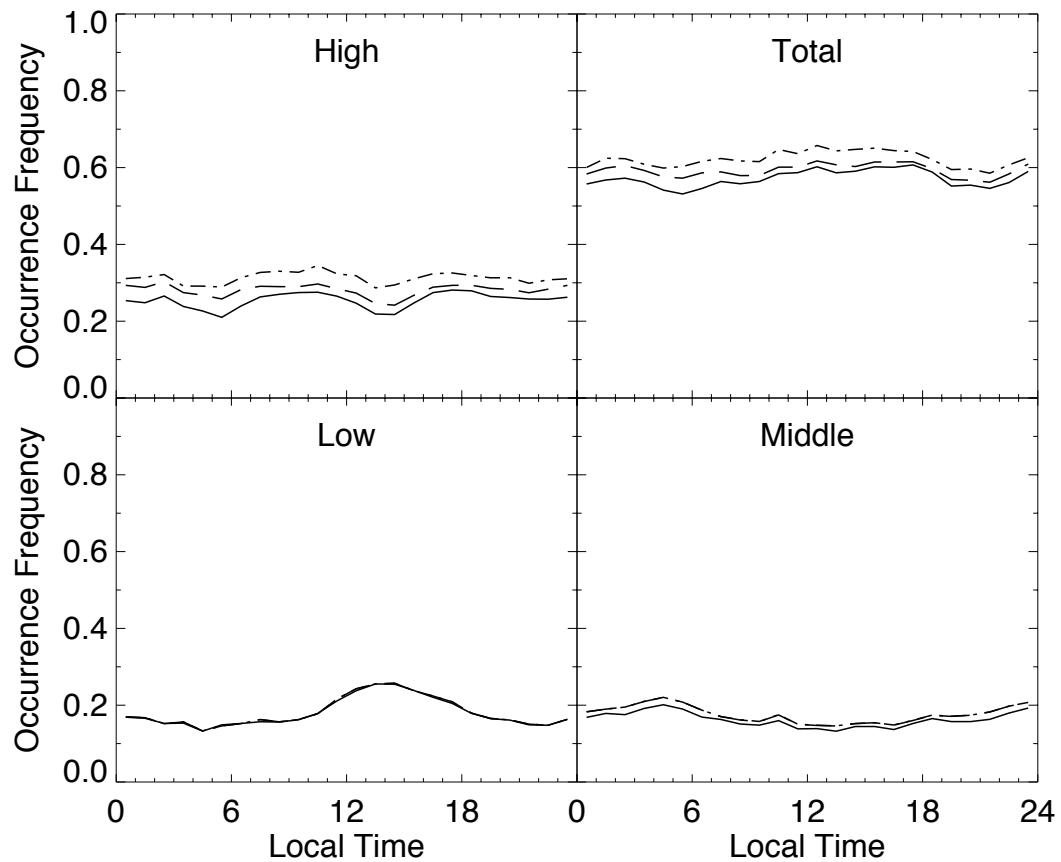


779  
 780 FIG. 10. Cloud base height distributions for GSFC during 2012 for V2 (solid line), V3b (dashed  
 781 line) and V3m (dash-dotted line). Vertical axis bin size equals 1 km.

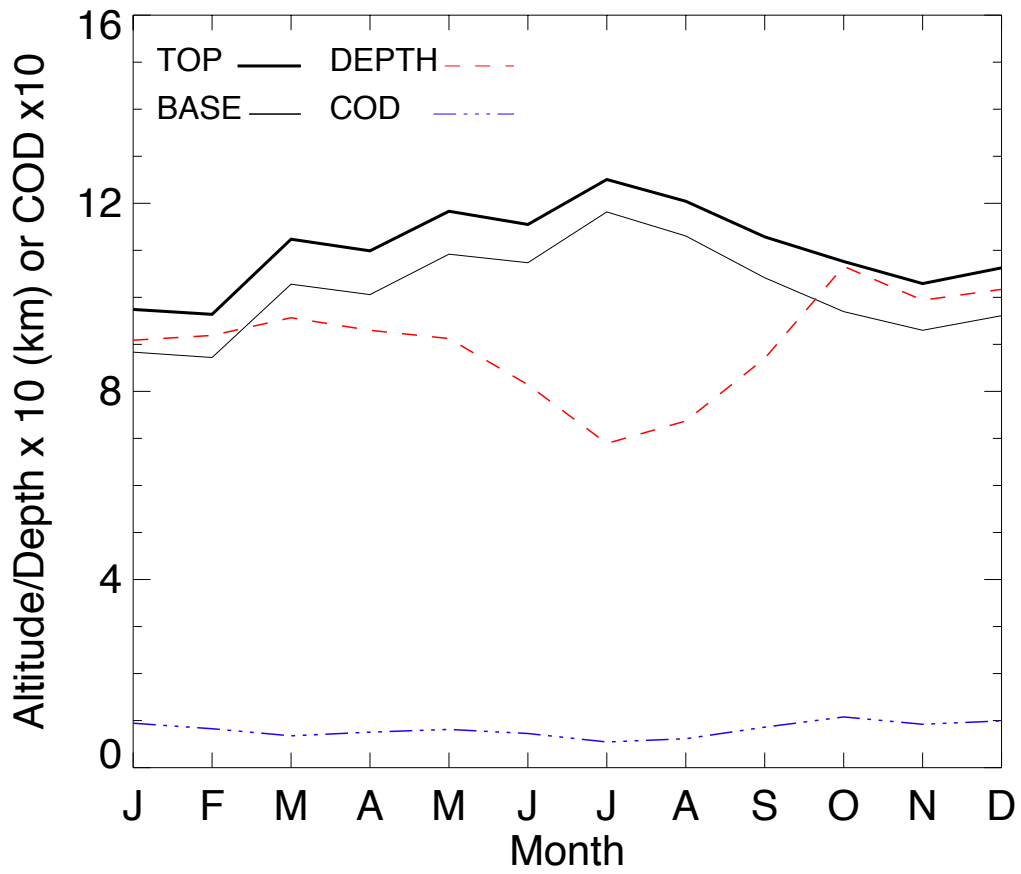


782  
 783 FIG. 11. Annual cycle of the occurrence frequency for low, middle, high and total clouds at  
 784 GSFC during 2012. The solid line represents V2, the dashed line V3b, and the dash-dotted line  
 785 V3m.

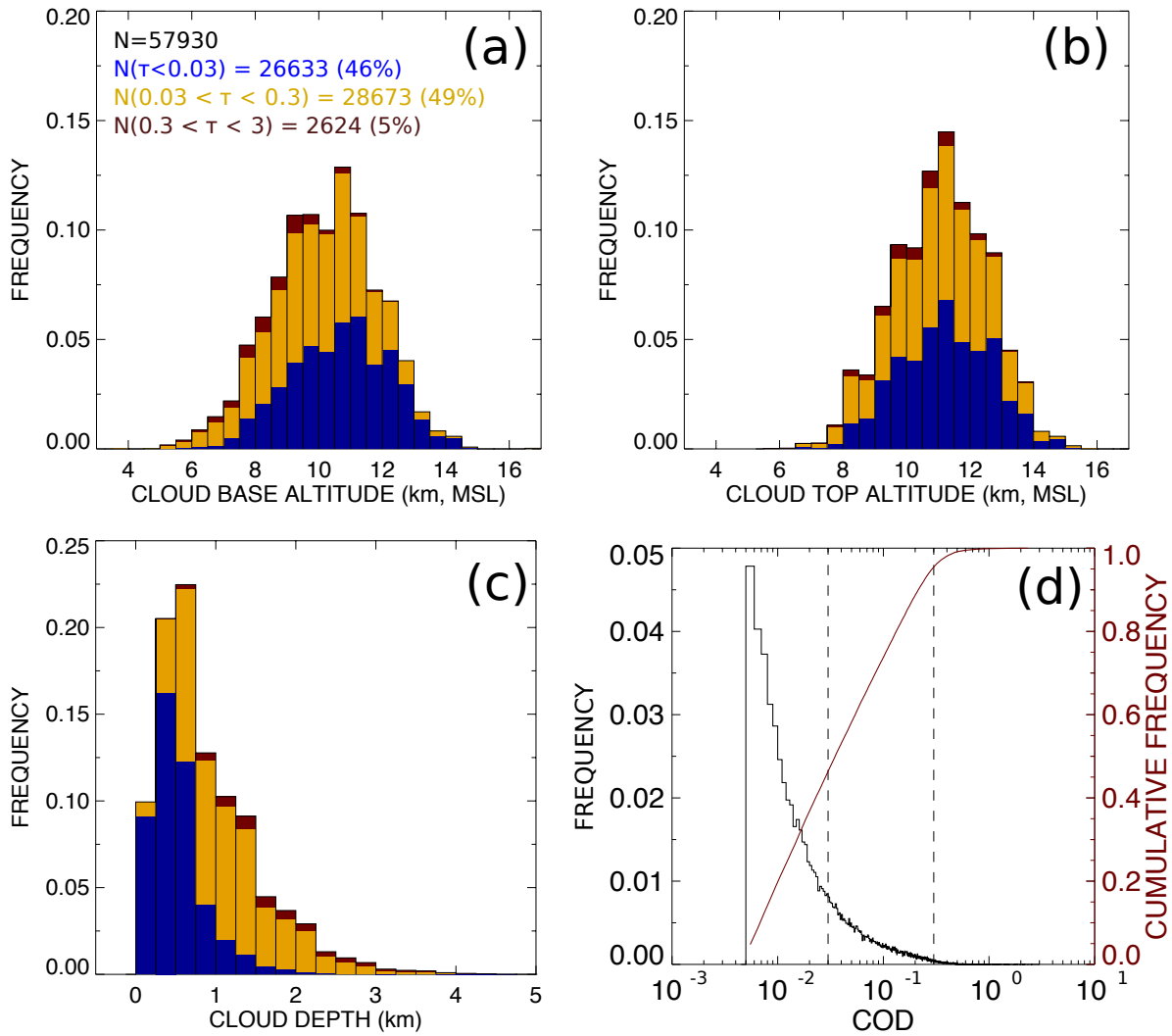




786  
 787 FIG. 12. Diurnal cycle of the occurrence frequency for low, middle, high and total clouds at  
 788 GSFC during 2012. The solid line represents V2, the dashed line V3b, and the dash-dotted line  
 789 V3m.



790  
 791 FIG. 13. Monthly averaged cloud top (thick solid line), cloud base (thin solid line), cloud depth  
 792 (dashed red line) and cloud optical depth (dashed-dotted blue line) for the transparent cirrus  
 793 dataset.  
 794



795  
 796 FIG. 14. (a) Cloud base altitude, (b) cloud top altitude and (c) cloud depth for the transparent  
 797 cirrus dataset. The colors indicate the cirrus type based on the estimated COD (blue for sub-  
 798 visual, yellow for thin, and red for opaque cirrus). Horizontal axis bin size is 0.5 km for base  
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 801 (0.03) and thin (0.3) cirrus clouds. Horizontal axis bin size equals 0.001.