Human Missions to Mars Key Challenges

National Aeronautics and Space Administration

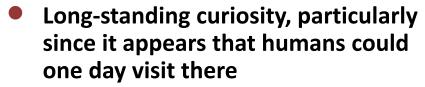




Bret G. Drake NASA Lyndon B. Johnson Space Center February 2013



Why Do We Want To Explore Mars?



- A NASA chartered group, Mars Exploration Program Analysis Group, has organized a set of four primary goals:
 - Determine if life ever arose on Mars
 - Understand the processes and history of climate on Mars
 - Determine the evolution of the surface and interior of Mars
 - Prepare for human exploration
- Two additional goals considered as well:
 - Preparing for sustained human presence
 - Ancillary science such as heliophysics, space weather, astrophysics

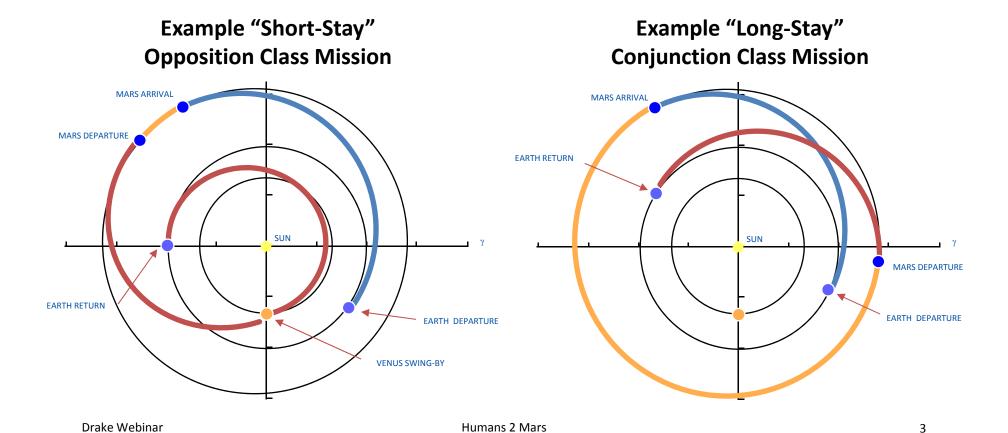
Goals and Objectives Summary Implications

- The first three human missions to Mars should be to three different geographic sites
- Maximize the amount of time that the astronauts spend exploring the planet
- Maximize mobility to extend the reach of human exploration beyond the landing site
- Provide subsurface access
- Return a minimum of 250 kg of samples to Earth



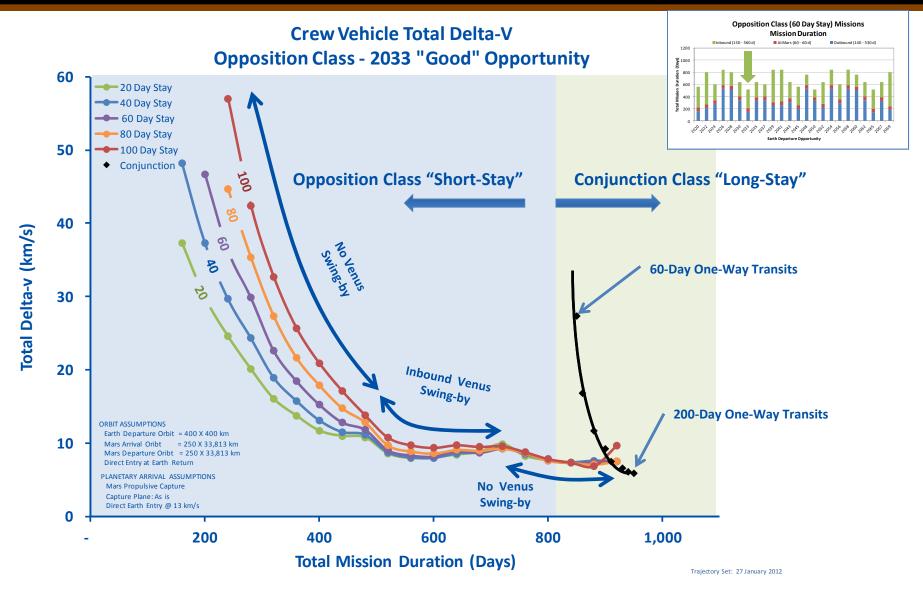
Mars Trajectory Classes

- A trip to Mars with a return back to Earth is a double rendezvous problem
 - Mars round-trip missions are flown in heliocentric space
 - Relative planetary alignment is a key driver in the mission duration and propulsion required





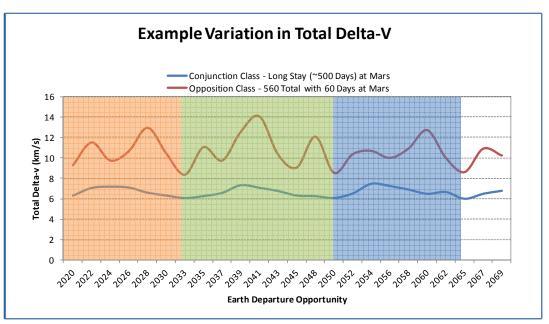
Example Delta-v versus Mission Duration





Synodic Period – Variation in Delta-V

- The difference in orbits of the Earth and Mars influence the mission delta-v and timing
 - Earth departure opportunities occur approximately ever 26 months
 - The Earth departure "window" lasts a few weeks and is highly dependent on the propulsion system choice
 - The round-trip mission delta-v varies over a 15-year cycle (the Synodic Cycle)
 - Although "good"
 opportunities occur in
 2018, 2033, and 2047,
 the ability to conduct
 missions in any
 opportunity across the
 Synodic Cycle will
 reduce programmatic
 risk





Advanced In-Space Transportation

Options, options, options....

High Thrust: Chemical Propulsion

<u>Advantages</u>:

- More "state of the art"
- Multiple destinations

Challenges:

- High Mass / Lots of Launches
- Long-term storage of cryogenic propellants, particularly H₂
- Configuration and integration challenges
- Long-stay missions only

High Thrust: Nuclear Thermal Propulsion (NTP)



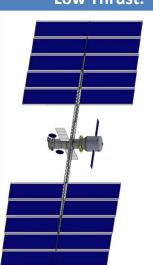
Advantages:

- Good combination of high thrust and high efficiency (Isp)
- Low architectural mass
- Both long and short stay missions
- Has been demonstrated (NERVA)

Challenges:

- Long-term storage of cryogenic H₂
- Large launch volume (due to H₂)
- Nuclear regulatory compliance/testing

Low Thrust: Solar Electric Propulsion (SEP)



Advantages:

- Low architectural mass
- Multiple destinations

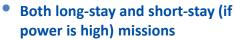
Challenges:

- Limited to long-stay missions
- Configuration and integration challenges (large solar arrays)
- Long operating times (spirals)

Low Thrust: Nuclear Electric Propulsion (NEP)

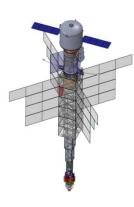
Advantages:





Challenges:

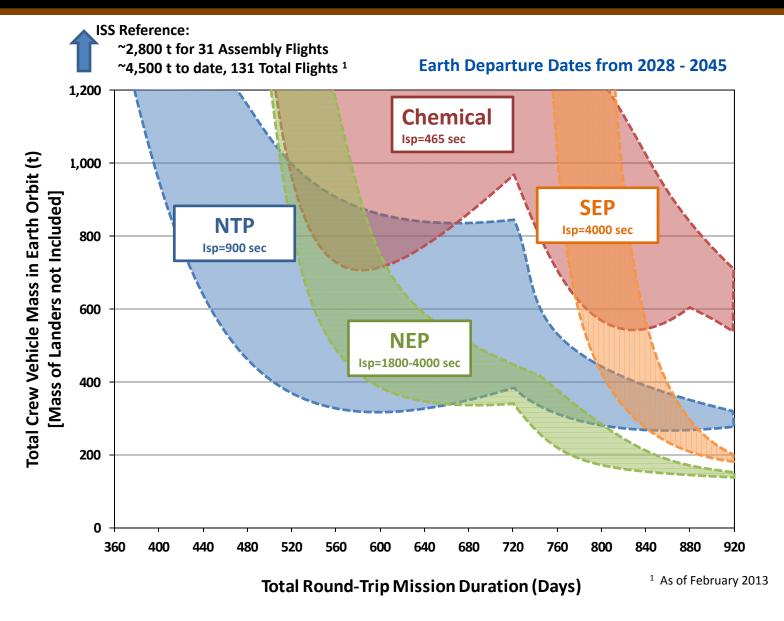
- No experience base for space based high power, high efficiency, nuclear reactors
- Configuration and integration challenges (large radiators)
- Nuclear regulatory compliance/testing
- Long operating times (spirals)





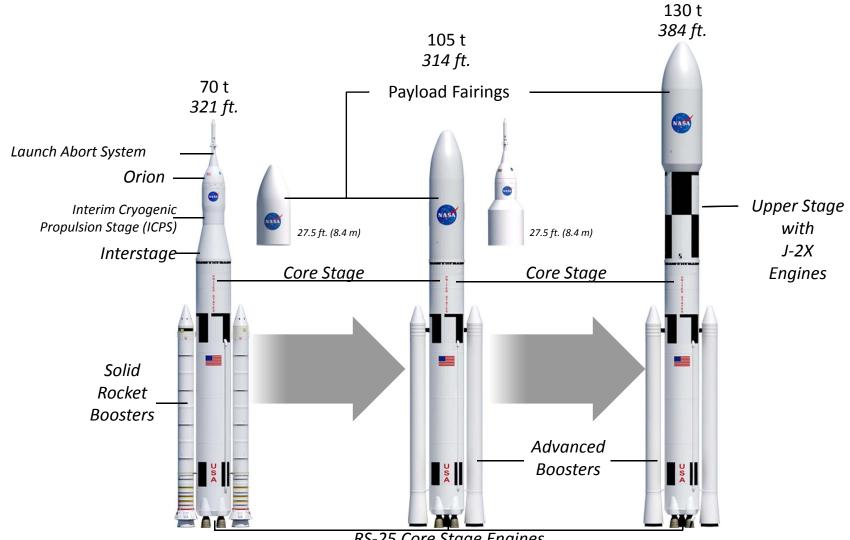
Propulsion Technology Comparisons

Crew Vehicle Mass as a Function of Trip Time – Short Stay Opposition Missions





SLS Architecture Block Upgrade Approach



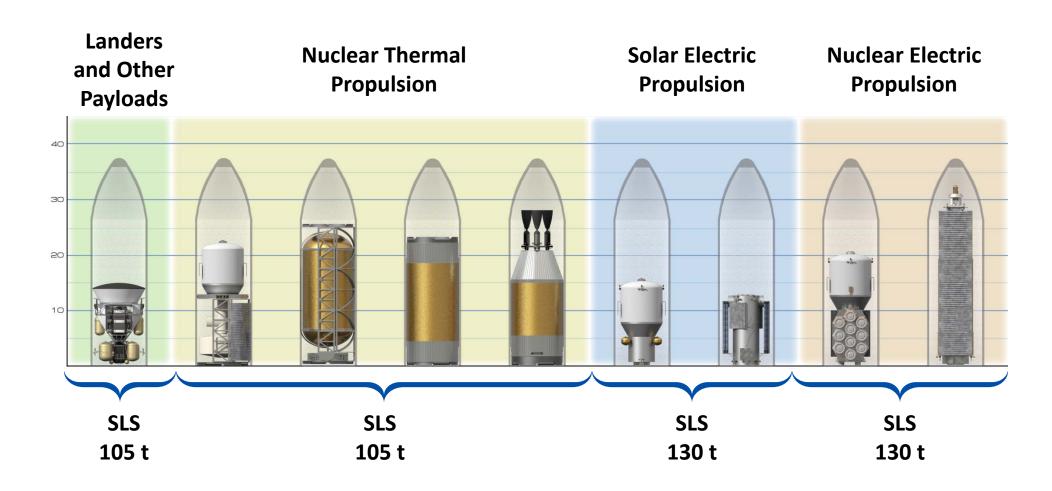
RS-25 Core Stage Engines (Space Shuttle Main Engines)

Starting with Available Assets and Evolving the Design



Example Launch Packaging

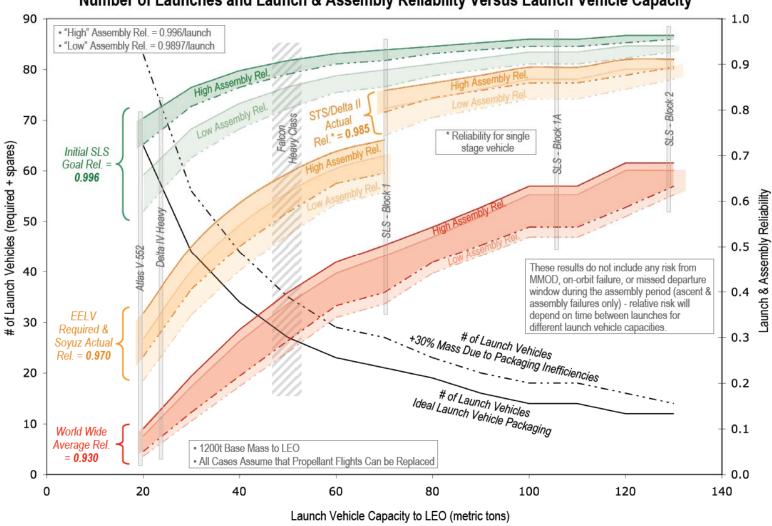
Diameter and Volume are also Key





Example Relationship Between Launches and Reliability

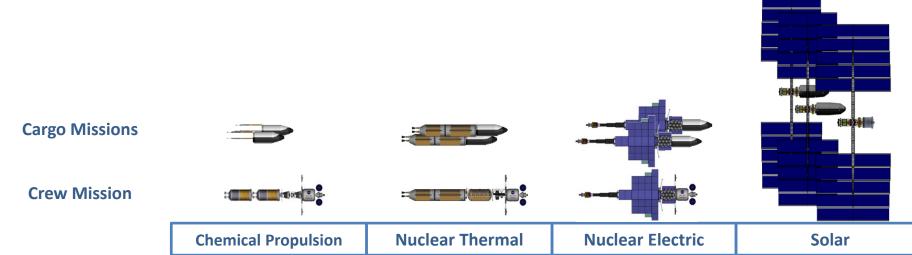




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Launch Campaign and Transportation Options



	Circinical Fropulsion		Tracical Internal		Mucical Electric		30101	
Parameter	"Stressing" ¹ Long-Stay	"Easy" Short-Stay						
Total Mass (mt)	~1,250	~1,460	~890	~860+	~770	~960	~780	X
# SLS Launches	~12	~13+	~9	~9+	~7	~9+	~7	X
SLS delivery to LEO (mt)	105 & 130	105 & 130	105	105	105 & 130	105 & 130	105 & 130**	X
SLS Shroud Dia./Barrel	10 / 22	10 / 22	10 / 25	10 / 29	10 / 25	10 / 25	10 / 15	X
Launch Spacing (days)*	50-120	10-110+	70-150	70-150	90-200	70-150	90-200	X

[&]quot;Hard" Long Stay - Represents most stressing conjunction class (2037 long-stay) mission. Typical mission values will be less for other opportunities.

[&]quot;Easy" Short-Stay - Represents the easiest opposition class (2033 short-say) mission. Values for other opportunities will vary greatly and will be much more stressing.

^{*} Launch spacing lower/upper values represent spacing required for crew missions every opportunity (26 months) & every-other opportunity (52 months) respectively + 6 mo schedule margin.

^{**}Depending upon SLS performance 1-2 ATV launches using a Ariane 5 class vehicle are required to provide consumables



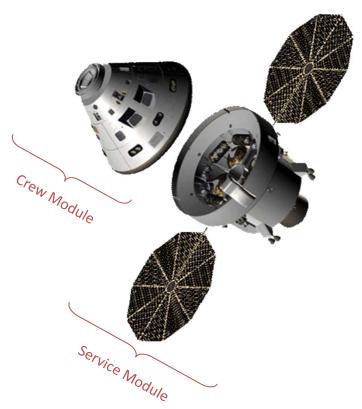
Orion Crew Transfer / Earth Return Vehicle

Crew Delivery to Earth Departure Point

- Provide safe delivery of 4-6 crew to Earth departure point for rendezvous with the Mars Transfer Vehicle
 - Delivery and return of checkout crew prior to the mission
 - Delivery of the mission crew

End of Mission Crew Return (Mars Block)

- Provide safe return of 4-6 crew from the Mars-Earth transfer trajectory to Earth at the end of the mission
 - 12 km/s entry speed (13+ km/s for short-stay mission)
 - 900 day dormant operations
 - 3 day active operations
 - Much smaller service module (~300 m/s delta-v) for retargeting and Earth entry corridor set-up



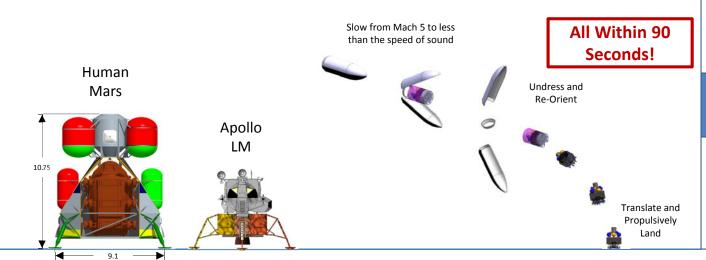




Challenges of Landing on Mars

The Atmosphere of Mars

- The Good: Mars has an atmosphere that can help slow the entry vehicle down
- The Bad: The atmosphere is thick enough that it requires a heat shield, but not thick enough to provide substantial drag (density 1% of Earth's)
- Atmospheric dust may prohibit ability or timing of landing at designated landing sites
- The Current Mars Science Laboratory Landing Strategy is Limited
 - ~ 1 mt payload to the surface (target 40 mt)
- Key for Human Missions Challenge: <u>Supersonic Transition</u>



Technology Options

Hypersonic Inflatable Aerodynamic Decelerator (HIAD)



Rigid Aeroshells (mid L/D)



Supersonic Retropropulsion

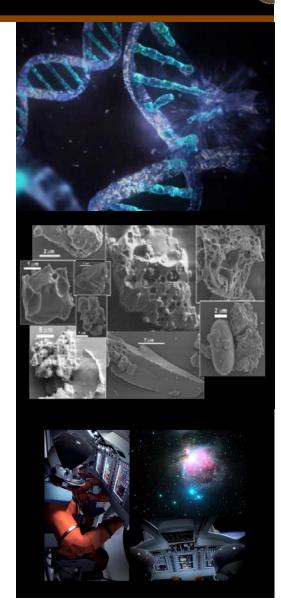




Challenges of Supporting Humans in Deep Space



- Long-Duration: 600 days minimum, 900 days most probable
- <u>Deep-Space</u>: Micro-gravity and harsh environment
- Remote: No logistics train, no fast return aborts
- Categories of Key Human Support Challenges
 - Ocular Syndrome: Intercranial pressure
 - Toxicity: Dust and other hazards
 - <u>Autonomous Emergency</u>: Response to system emergencies (e.g. life support system failure)
 - Radiation: Solar Proton (solutions exist), Galactic Cosmic Radiation (currently no standards for exploration)
 - <u>Behavioral Health and Performance</u>: Remote isolated missions with no real-time communications.
 - <u>Autonomous Medical Care</u>: Response to medical issues
 - <u>Nutrition</u>: Food with adequate nutrition for long missions
 - <u>Hypogravity</u>: Adjusting to the gravity of Mars
 - Musculoskeletal: Muscle atrophy and bone decalcification
 - <u>Sensorimotor</u>: Sensory changes/dysfunctions





Living off of the Land: In-Situ Resources

Atmosphere

- Atmospheric resources found globally with slight change in pressure/concentration
- Primary product: oxygen (O₂) bound in carbon dioxide (CO₂)
- Oxygen can be used for propulsion, life support, and extra vehicular activity (EVA) applications
- Production of O₂ only from CO₂ makes over 75% of ascent propellant mass
- Production of O₂ and CH₄ (or other hydrocarbon fuel)
 possible with hydrogen (H₂) brought from Earth

Soil Processing for Water

- Water resources found globally with large variations in concentration, form, and depth.
- Water can be used for life support, EVA, and radiation shielding
- Water can be processed into O₂ and H₂ or with CO₂ to make fuels for propulsion and power
- Production of O₂ and methane (CH₄) from CO₂ and H₂O allows for 100% of ascent propellant mass

Leverage

 Producing oxygen from the atmosphere provides significant leverage in terms of mass (32%) and volume (lander packaging)

Resources

Atmosphere	Quantity			
Carbon Dioxide (CO ₂)	95.5%			
Nitrogen (N ₂)	2.7%			
Argon (Ar)	1.6%			
Oxygen (O ₂)	0.15%			
Water (H ₂ O)	<0.03%			

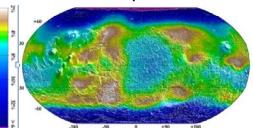
Global

Dependent

Site

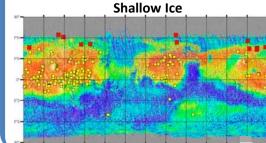
Landing !

Water in Top 1 meter



Hydrated Minerals



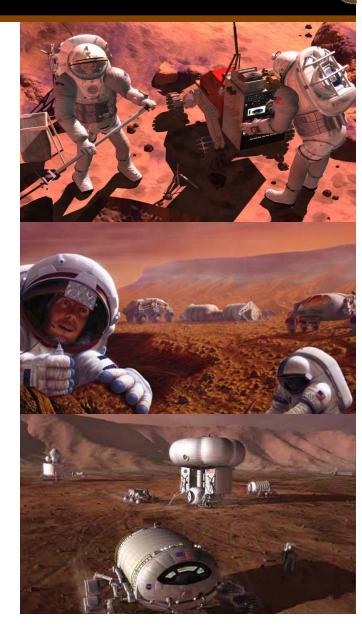


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Surface Exploration and Discovery

- Long surface stays with visits to multiple sites provides scientific diversity thus maximizing science return
- Sustainability objectives favor return missions to a single site (objectives lend themselves best to repeated visits to a specific site on Mars)
- Mobility at great distances (100's km) from the landing site enhances science return (diversity)
- Subsurface access of 100's m or more highly desired
- Advanced laboratory and sample assessment capabilities necessary for highgrading samples for return





Human Exploration of Mars Capability Needs

Launch

- Multiple launches
- Short spacing
- Large mass: 130 t
- Large Volume 10 x 30 m

Space Transportation

- Advanced propulsion to reduce mass
- Fast Transits for Crew (180 days)
- Limited / lack of quick aborts

Entry Descent and Landing

- Large mass (40 t) / Large volume
- Abort to surface
- Precision landing

Crew Surface Health and Support

- Crew acclimation post landing
- Human Support (radiation, hypogravity, dust, behavior)
- Planetary protection

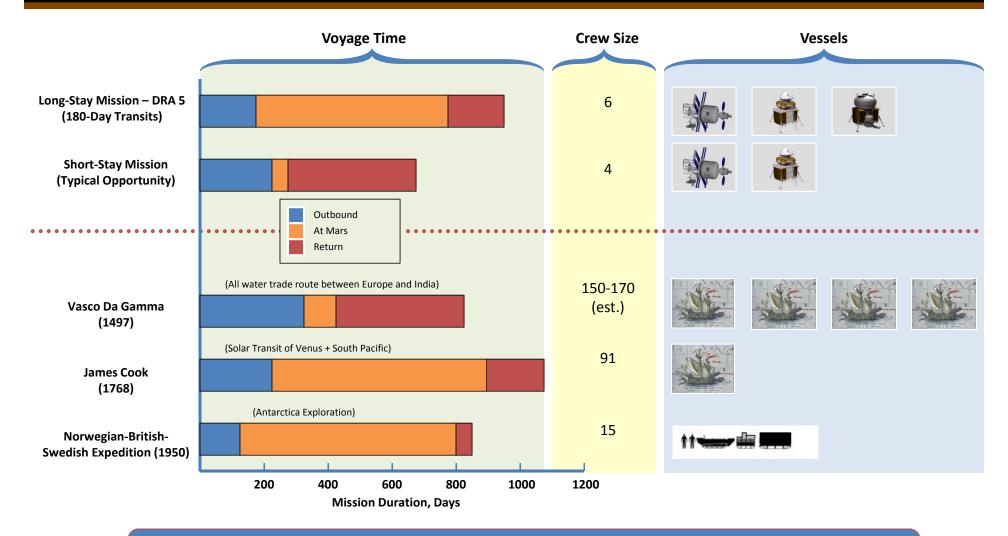
Operations

- Automated, rendezvous and docking
- Pre-deploy cargo
- No logistics
- Reliability, maintenance and repair
- Autonomous operations post landing
- Infrastructure emplacement (power)
- High continuous power (40 kWe)
- ISRU oxygen production atmosphere
- Multiple EVAs, long-range roves, routine exploration



Human Exploration: A Historical Perspective





Human mission to Mars will be long and complex, but the round trip duration is within the experience of some of the past successful exploration missions with significantly far fewer crew