



Modeling the Space Debris Environment with MASTER-2009 and ORDEM2010

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- ESA MASTER-2009 Population Generation
(Meteoroid and Space Debris Terrestrial Environment Reference)
- NASA ORDEM2010 Population Generation
(Orbital Debris Engineering Model)
- (Software Feature Comparison)

MASTER-2009 Population Generation Process



Object Data Acquisition & Processing

- Data for tracked objects is collected from multiple sources & brought into unified format

Object Data
Acquisition & Processing

Simulation & Data Fusion

- All debris sources are simulated
- Simulation results are fused with data for tracked objects

Simulation & Data Fusion

Population Validation

- Large objects > ~10 cm: comparison of *real* and *simulated* measurement campaigns (*PROOF*)
- Small objects < ~1 mm: comparison of *real* and *simulated* impact craters (*MASTER*)

Population Validation

Final Population

MASTER-2009

Object Data Acquisition & Processing



Acquisition of object bulk:

Input: Two-Line Elements (USSTRATCOM)

Acquired data: single mean orbit parameters

Output: Quarterly orbit snapshots between 1957 and 2009

Additional objects:

Input: Satellite Catalog (Jonathan McDowell)

Acquired data: Objects not included with TLE data

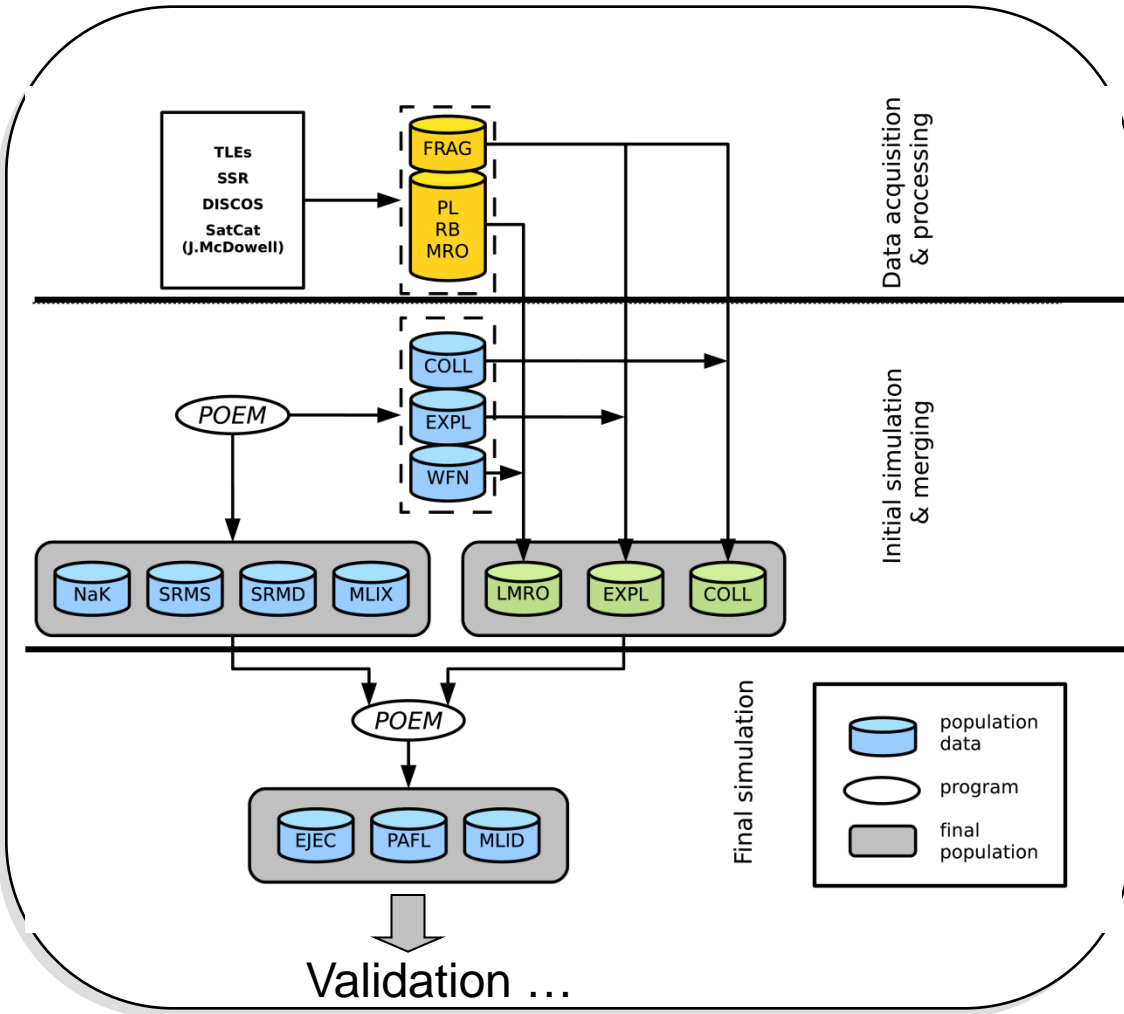
Output: Extended quarterly orbit snapshots

Object properties:

Input: Database and Information System Characterising Objects in Space (ESA: DISCOS) & Satellite Situation Report (USSTRATCOM)

Acquired data: Object size, mass and mass-to-area ratio

Output: Quarterly population snapshots subdivided into fragments & launch- and mission related objects (payloads, rocket bodies and mission debris)



POEM (Program for Orbital Debris Environment Modeling)

- Compendium of individual debris models for each source

List based debris sources:

- Individual events are simulated
- List data includes e.g.:
event epoch, orbit location, event magnitude

Continuous debris sources:

- All LMRO are analysed wrt. paint flakes, delaminated MLI and ejecta which they would have produced

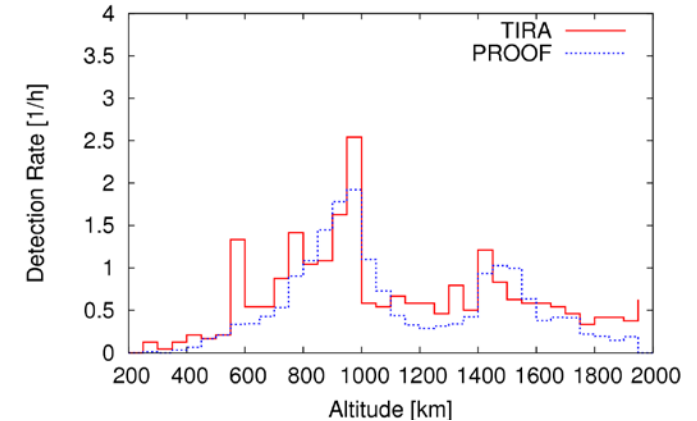
MASTER-2009

Large Object Validation

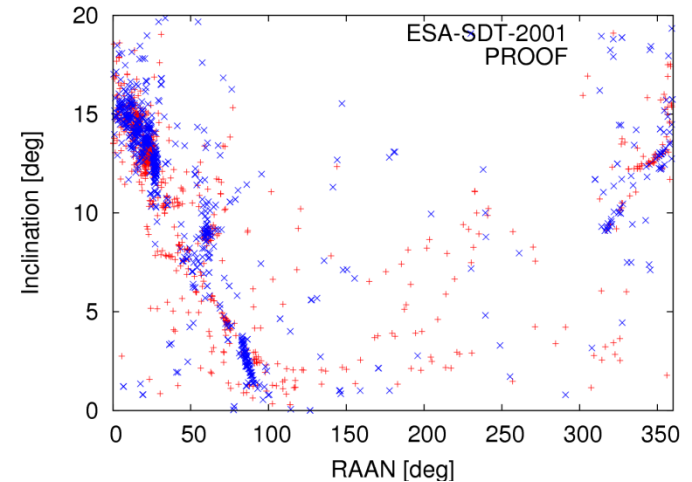


- Detection campaigns offer information on the number of objects, RCS or magnitude and orbit properties
- *PROOF* (ESA **P**rogram for **R**adar and **O**ptical **O**bservation **F**orecasting) applies filters for:
 - geometry (e.g. field of view, viewing direction)
 - performance (e.g. radar: wavelength, power; optical: CCD type, integration time)

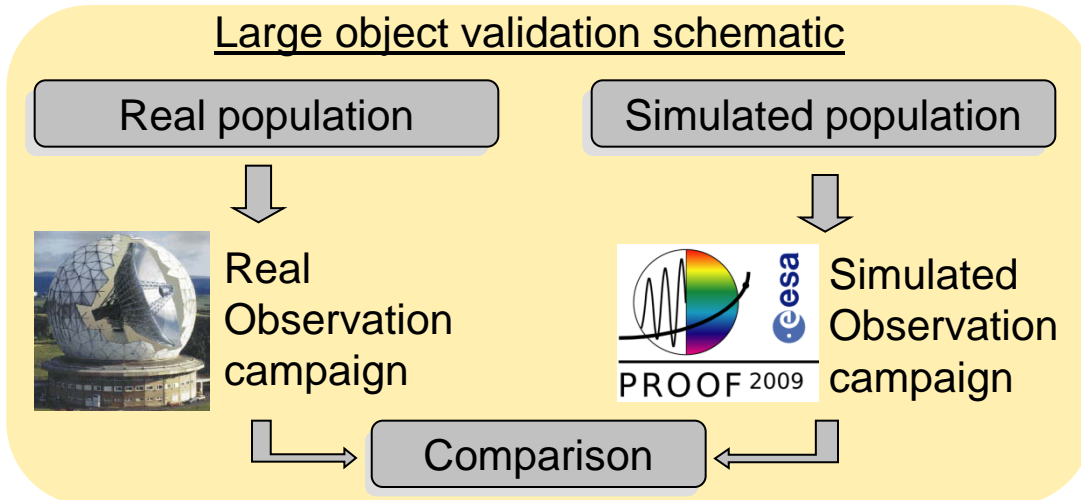
Validation of amount of debris



Validation of orbit distribution



Large object validation schematic



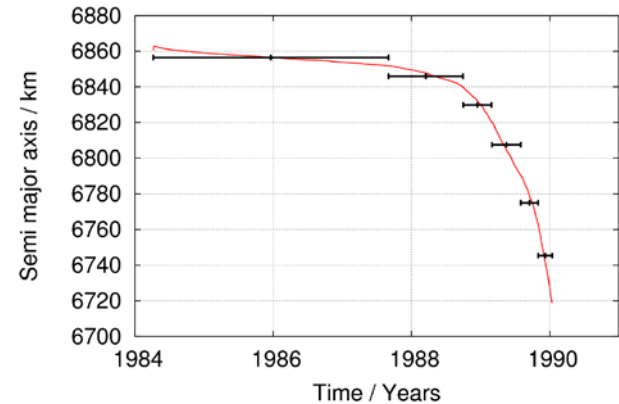
MASTER-2009

Small Object Validation

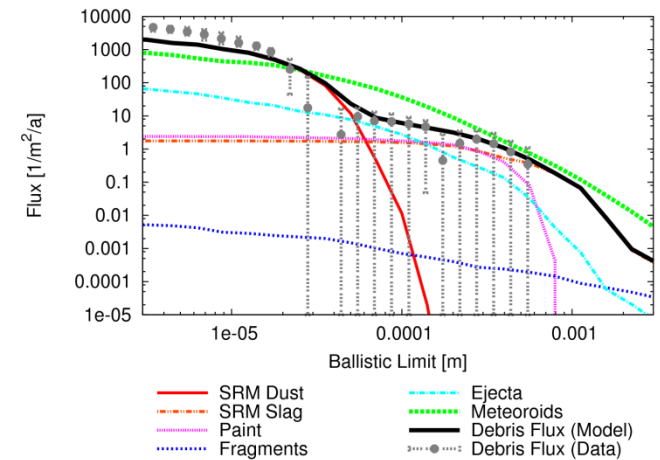


- Impact craters offer information on the impactor size/impact velocity, impact direction and total number of objects
- MASTER applies filters for e.g.:
 - target orbit evolution/maneuvers
 - rotation of target orbit line of apsides and line of nodes
 - target surface orientation
 - damage equations

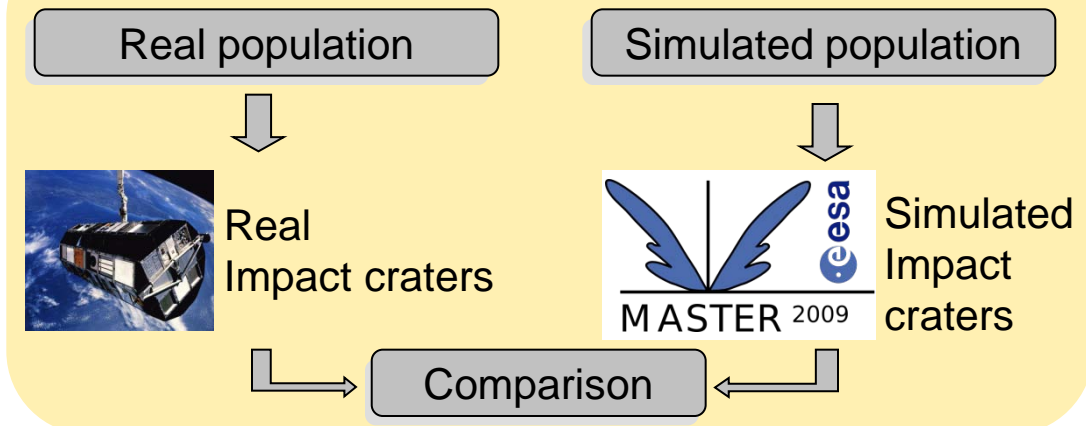
Simulating mission parameters



Validation of impact fluence



Small object validation schematic



- The Orbital Debris Engineering Model, ORDEM2010, includes,
 - **High-fidelity population file structure of the yearly debris populations from 1995 - 2035**
 - Sizes 10 μm - 1 m (LEO - GTO) ; sizes 10 cm - 1 m (GEO)
 - Stable orbital elements (i.e., those that do not randomize on a sub-year timescale)
 - LEO – GTO \rightarrow Hp, Ecc, Inc ; GEO \rightarrow MM, ECC, Inc, RAAN
 - Debris material density
 - **High-fidelity spacecraft analysis program compares the populations with a spacecraft-encompassing 'igloo' to achieve a 3-D output of flux on the spacecraft**
 - **Advanced graphical user interface (GUI) allows visualization of spacecraft flux in 2-D and 1-D**

Parameter	ORDEM2010
Spacecraft and Telescope/Radar analysis modes	YES
Time range	1995 to 2035
Altitude range with minimum debris size	200 to 34,000 km ($>10 \mu\text{m}$)* ; 34,000 to 38,000 km ($>10 \text{cm}$)
Model population breakdown	Intacts, Low-density fragments ; Medium-density fragments and degradation/ejecta ; High-density fragments and degradation/ejecta ; RORSAT NaK coolant droplets
Population material density breakdown	Low-density ($<2 \text{g/cc}$) ; Medium-density (2-6 g/cc) High-density ($>6 \text{g/cc}$) ; RORSAT NaK coolant (0.9 g/cc)
Population cumulative size thresholds	10 μm , 31.6 μm , 100 μm , 316 μm , 1 mm, 3.16 mm, 1 cm, 3.16 cm, 10 cm, 31.6 cm, 1 m
Population storage	LEO-to-GTO bins - Hp, Ecc, Inc , GEO bins - MM, Ecc, Inc, RAAN
Population extension	Bayesian statistics with ODPO models
Model S/C flux analysis method	Igloo surrounding S/C
Model T/R flux analysis method	Segments along line-of-sight

ORDEM2010

Population Generation Process



- Bayesian statistical approach to debris population analysis
 - **Ten additional years of data including,**
 - Catalog datasets → SSN
 - Statistical datasets → Haystack, HAX radars
 - Individual event datasets → FY-1C anti-satellite test, Iridium 33/Cosmos 2251 from SSN radar observation
 - **NASA Orbital debris Program Office (ODPO) models used as prior conditions**
 - LEGEND 3-D debris long-term environment model replaces the 1-D EVOLVE
 - NaK Module for RORSAT sodium potassium droplets
 - Degradation/Ejecta (D/E) for sub-millimeter particles

Model	Usage	Corroborative Data
LEGEND	LEO Fragments > 1mm ; GEO Fragments > 10cm	Haystack, HAX, SSN ; MODEST
NaKModule	NaK droplets > 1 mm	Haystack
Degradation/ejecta model	1mm > Degradation/ejecta > 10 μ m	STS windows & radiators

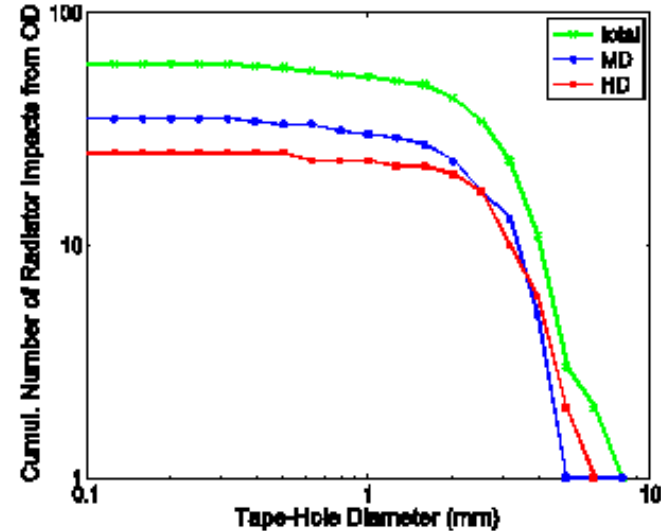
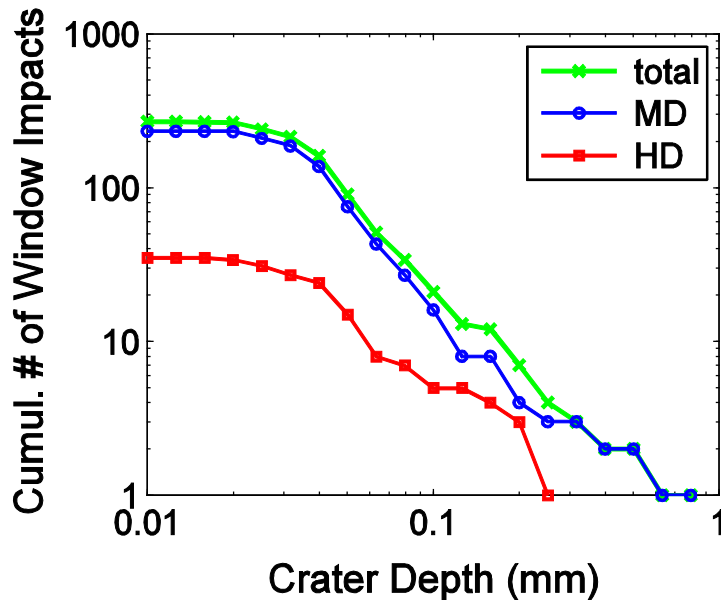
- Large object (>1mm) validation is ongoing.
- Small object (<1mm) validation
 - **A degradation/ejecta source model is constructed to provide the desired reference populations for the micro-debris population derivations.**
 - **Catalog (>10 cm) objects are taken as parent bodies of the small micron-sized particles.**
 - Number of micro-debris objects created by a surface degradation process is proportional to the surface area of a parent body.
 - Micro-debris objects created in a surface degradation process share the same orbit with its parent body *at the creation time*. Every orbit of the degradation/ejecta particles is propagated independently under the influence of solar radiation pressure and atmospheric drag, in addition to gravitational perturbations.
 - **The production rates of micro-debris are honed to be compatible with data.**

ORDEM2010

Population Validation (2)



- Window and radiator impact data from 38 STS missions. Window data is identified by the metric crater depth, radiator data by metric tape-hole diameter.
- Impactor are identified by material density when available. MD = medium density (aluminum, paint). HD = high density (steel)

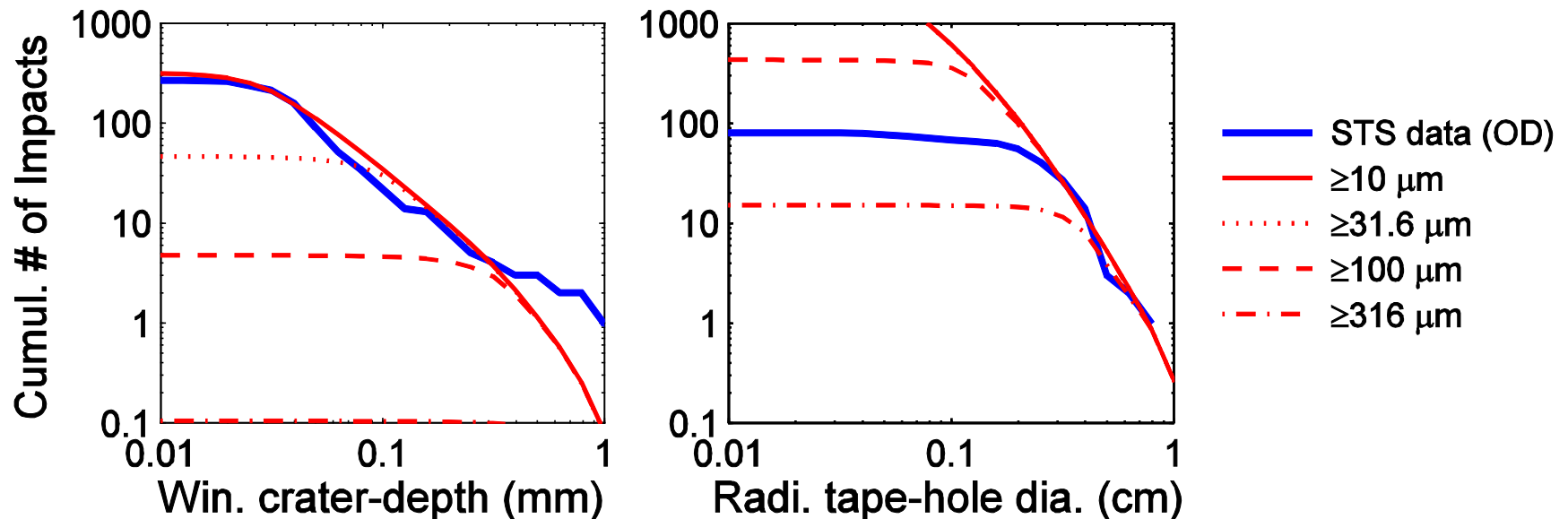


ORDEM2010

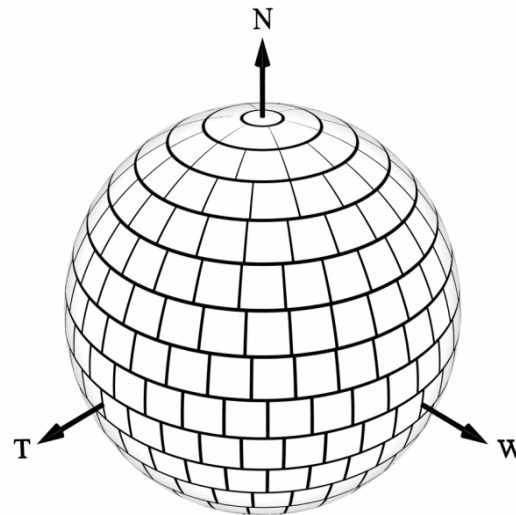
Population Validation (3)



- Degradation/ejecta model medium density population is adjusted to both window data and radiator data simultaneously.
- Detailed presentation, 'Simulation of Micron-Sized Debris Populations in Low Earth Orbit' will be given later at this conference, by Dr. Yu-Lin Xu.



- ORDEM2010 spacecraft encounters debris flux via a spacecraft-encompassing 3-D igloo
 - Population flux is tested for each igloo element in an igloo coordinate system of debris size, velocity, azimuth, and elevation with respect to spacecraft ram direction
 - Flux is summed within an element, all element fluxes are summed together for the total yearly spacecraft encounter
 - Highest fidelity igloo presently in ORDEM2010 is $10^\circ \times 10^\circ \times 1\text{km/s}$ (Az x EL x Vel)

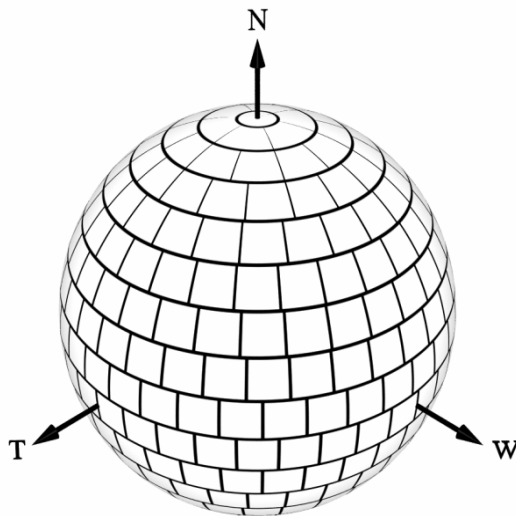


equal-area
spacecraft-encompassing igloo

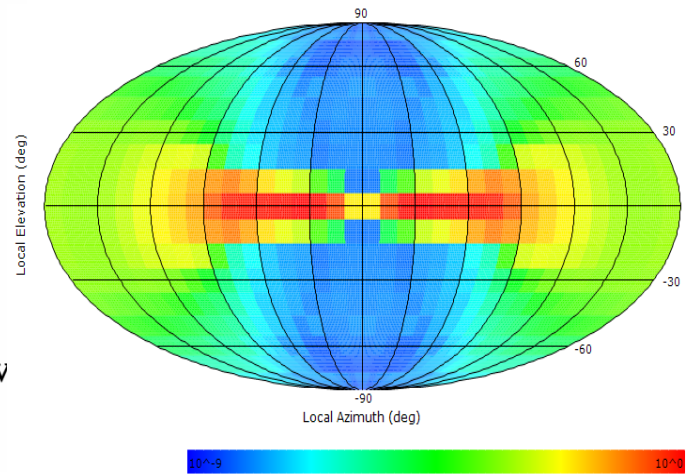
ORDEM2010 Graphical User Interface (GUI)



- This directional debris flux calculation is supported by an updated graphical user interface (GUI) package designed for ORDEM2010 that includes a 2-D directional flux chart (a.k.a. Mollweide projection, pseudo-cylindrical equal-area map projection used for global or sky maps)



equal-area
spacecraft-encompassing igloo



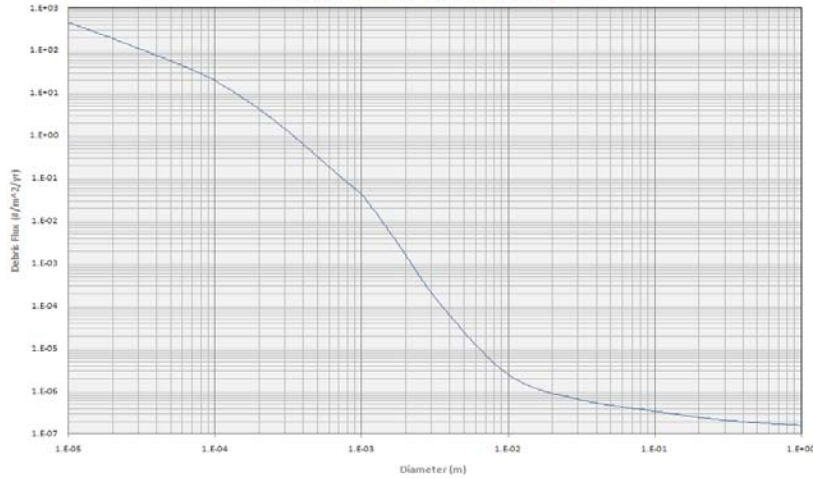
- Spacecraft velocity vector (ram direction) is defined by the azimuth, elevation coordinates $(0^\circ, 0^\circ)$
- Anti-ram is defined where $(180^\circ, 0^\circ)$ and $(-180^\circ, 0^\circ)$ meet
- Zenith is defined at $(0^\circ, 90^\circ)$, and nadir at $(0^\circ, -90^\circ)$.

ORDEM2010 GUI Example



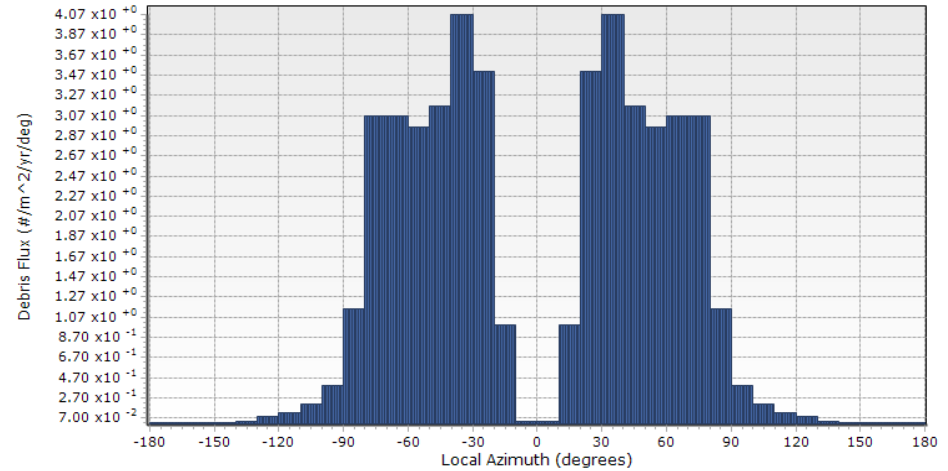
ISS ORDEM2010 GUI Outputs for Debris larger than 10 μm (Inc=51.63 $^\circ$, Hp=Ha= 400 km, yr=2010)

Average Cross-Section Flux vs. Size
Year: 2010 a = 6778.136 e = 0.000000 inc = 51.63



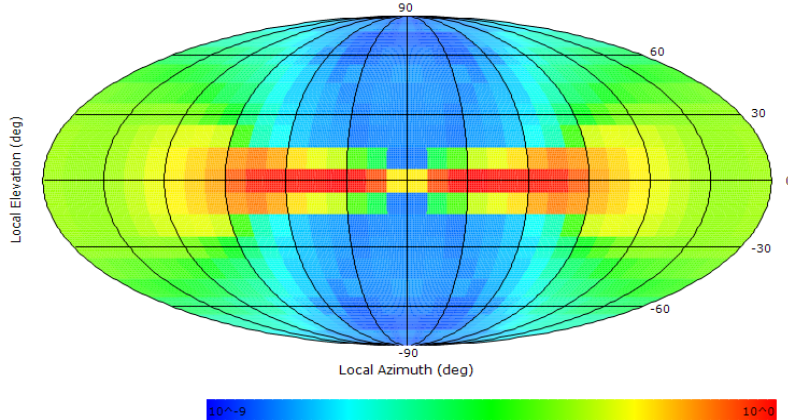
Flux vs. Local Azimuth

Year: 2010 a = 6778.136 e = 0.000000 inc = 51.63 particle size = >10um



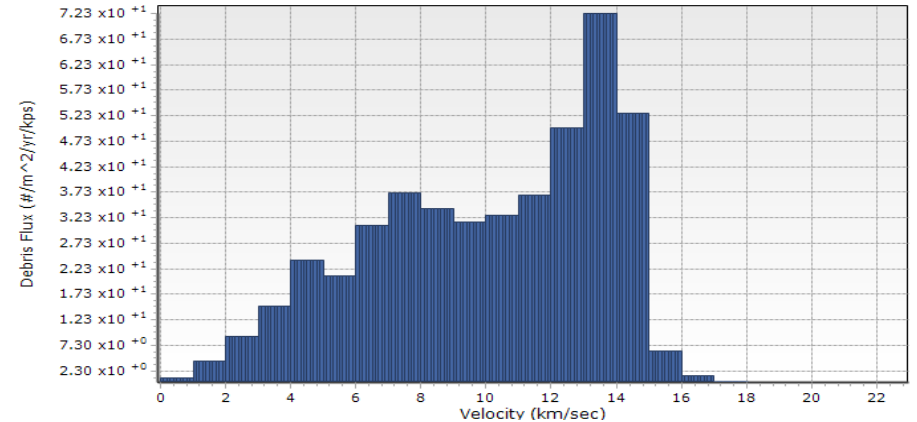
2-D Directional Flux

Year: 2010 a = 6778.136 e = 0.000000 inc = 51.63 particle size = >10um



Velocity Distribution

Year: 2010 a = 6778.136 e = 0.000000 inc = 51.63 particle size = >10um



Summary



- Spacecraft analysis using ORDEM2010 uses a high-fidelity population model to compute risk to on-orbit assets
- The ORDEM2010 GUI allows visualization of spacecraft flux in 2-D and 1-D
- The population was produced using a Bayesian statistical approach with measured and modeled environment data
- Validation of sizes $< 1\text{mm}$ were performed using Shuttle window and radiator impact measurements
- Validation of sizes $> 1\text{mm}$ is on-going

Backup Slide

MASTER-2009 Population Generation Process

