

#### Analysis and Consequences of the Iridium 33-Cosmos 2251 Collision

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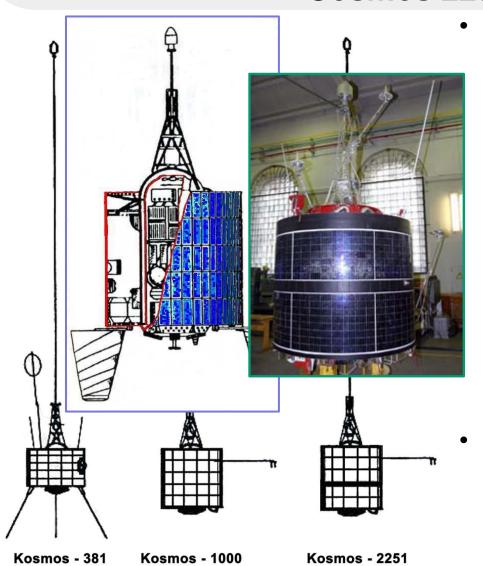
#### **Outline**



- The spacecraft
- Circumstances of the collision event
- Physical characterization of the debris clouds
  - Characteristic size, mass, and area-to-mass (A/m)
  - Directionality and ∆v distributions
  - Momentum transfer
- Comparisons with the NASA Standard Breakup Model
- Long-term evolution of the debris clouds
- Conclusions

# The Spacecraft I Cosmos 2251





The Strela-2M series utilized the versatile NPO-PM KAUR-1 standard bus (Kosmicheskiy Apparat Unversalnogo Ryada-1, (Космический Аппарат Универсального Ряда), which can be translated as Spacecraft Bus from the Standardized Line (Group)-1). In addition to the LEO communication constellation Strela-2/-2M, the KAUR-1 bus has served as the basis for navigation (Tsiklon/Parus military series and Tsikada civil series and Nadezhda civil COSPAS/SARSAT subseries), geodesy (Sfera and GEO-IK/Musson), and science (lonosfernaya, Cosmos 381 ionospheric topside sounder) spacecraft.

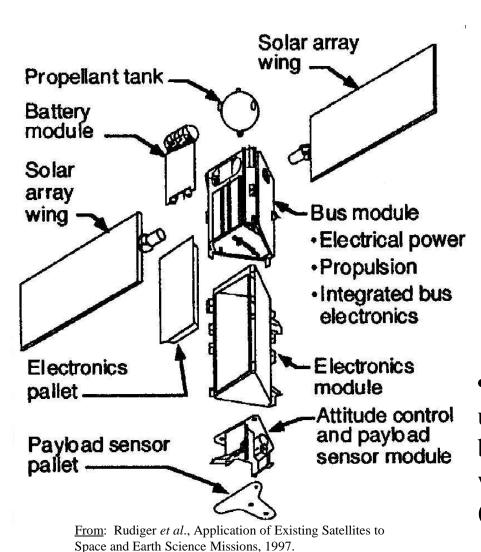
<u>Background</u>: family album of spacecraft using the KAUR-1 bus; <u>middle inset</u>: cross section of a *Nadezhda* spacecraft\*; <u>top inset</u>: a *Strela-2M* spacecraft\*\*

<sup>\*</sup> after Russian Space News, issue 24 (1994), p. 24

<sup>\*\*</sup> FROM: http://www.astronautix.com/craft/strela2m.htm

#### The Spacecraft II Iridium 33





Gateway Antennas •The Iridium first-generation constellation utilized the Lockheed-Martin LM700A bus, shown in exploded view (left) and with Iridium nadir payload module (above)

Crosslink Antennas

Battery

Module

Command Module Structure

Communication Section

Main Mission Antenna

Solar

Pon els

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# The Spacecraft III physical & operational characteristics



	Cosmos 2251	Iridium 33
bus	KAUR-1	LM700A
dry mass [kg]	900 (estimated)	556
Shape	Cylinder with boom	Triangular prism with panels
Stabilization	Gravity gradient	3 axis
Size	2 m x 2 m (body)	3.6 m long
onboard energy sources	Core cylinder may have been pressurized	Hydrazine tanks for thrusters; NiH <sub>2</sub> battery
Initial orbit	800x776 km, 74° inclination	779x776 km, 86.4° inclination
status	derelict	operational

#### The collision event



#### Estimated collision parameters:

Event time: 10 February 2009, 16<sup>h</sup> 55<sup>m</sup> 59.8<sup>s</sup> GMT

Location: 72.50° N latitude

97.86° E longitude

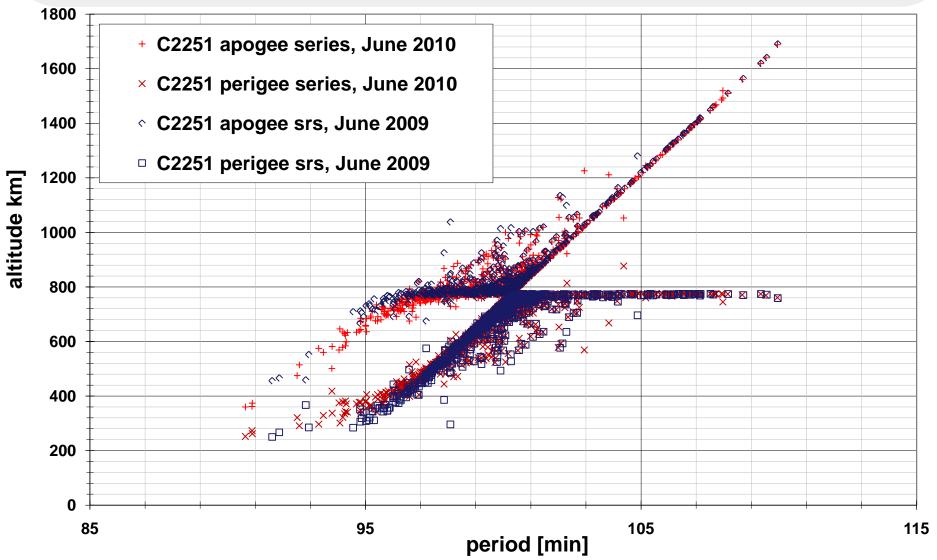
778.6 km altitude

Relative velocity: 11.647 km/s

<u>Directionality</u>: from the viewpoint of the Iridium 33 spacecraft, Cosmos 2251 approached at an elevation of -2.6° and an azimuth (measured from North) of 231° <u>Orbital distribution</u>: see *Gabbard* charts on next 2 pp.

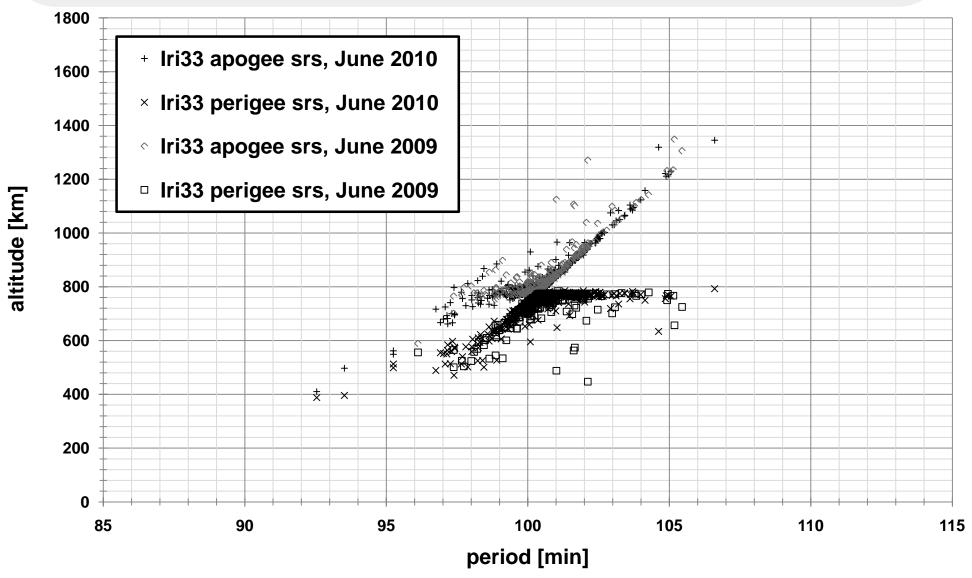
### Cosmos 2251 Gabbard diagram





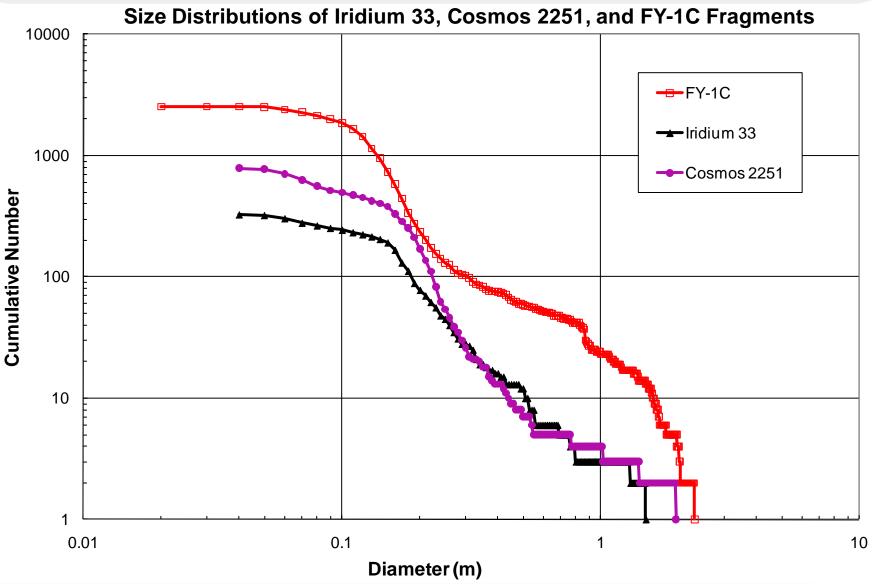
#### Iridium 33 Gabbard diagram





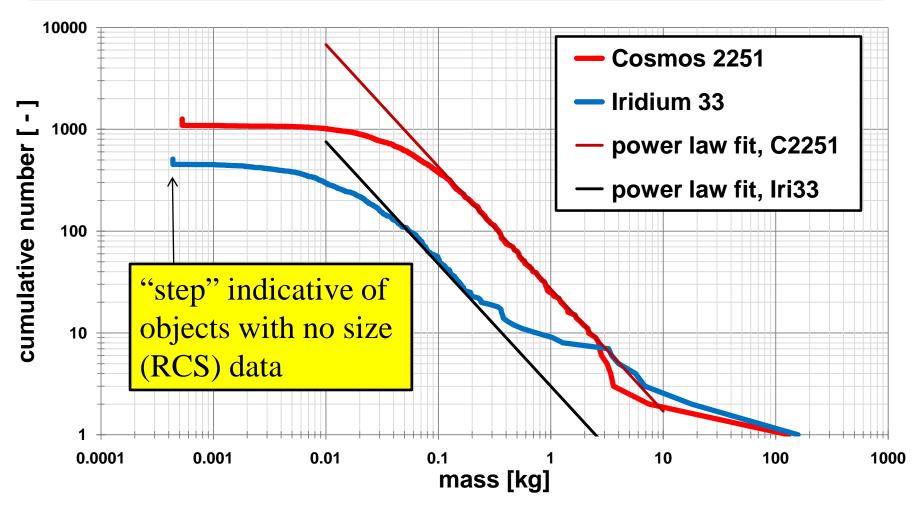
## Physical characterization of the debris clouds I characteristic size distribution



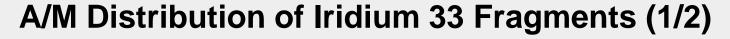


# Physical characterization of the debris clouds II mass distribution

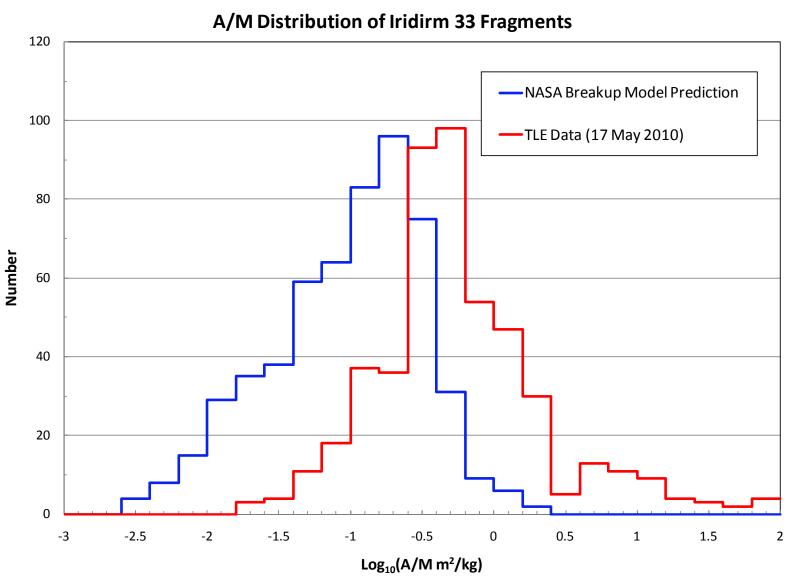




Masses estimated using median A/m and characteristic lengths; power law slope is -1.2, considerably steeper than standard breakup model.

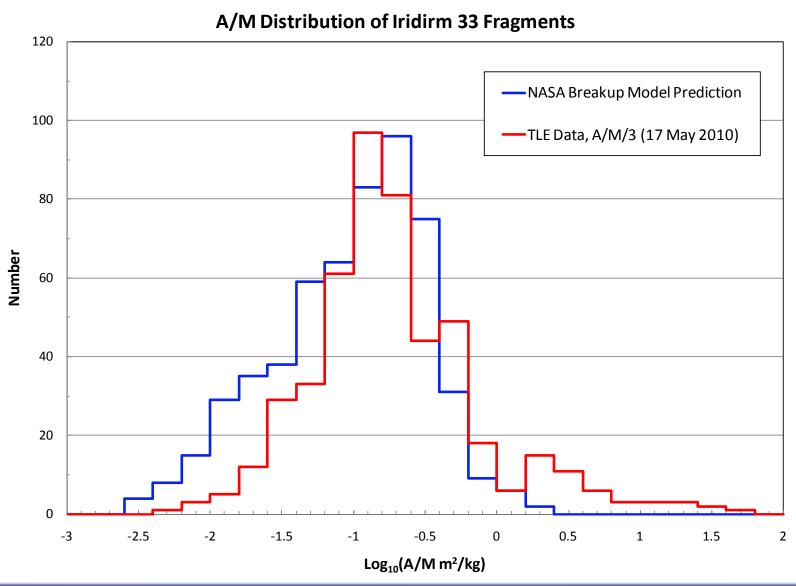






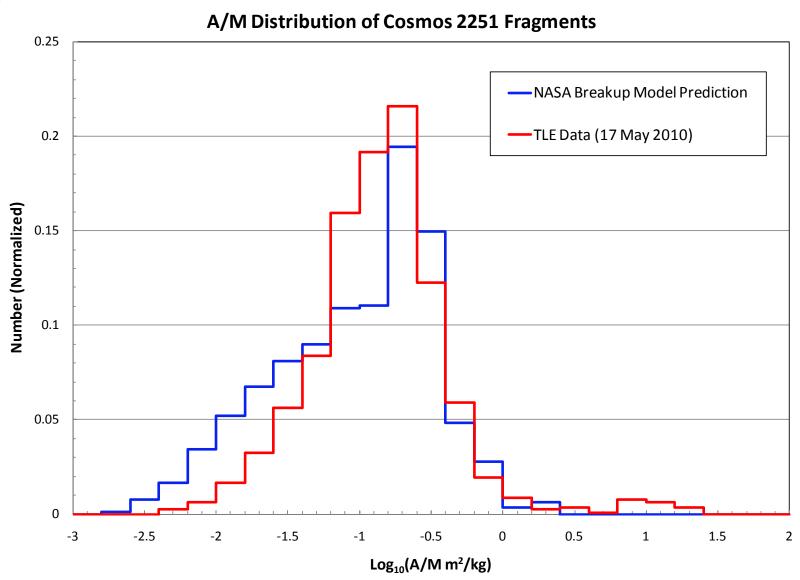






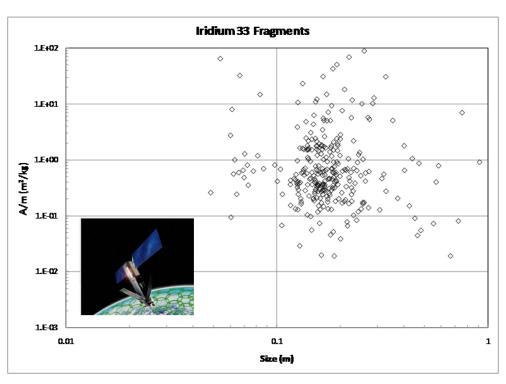


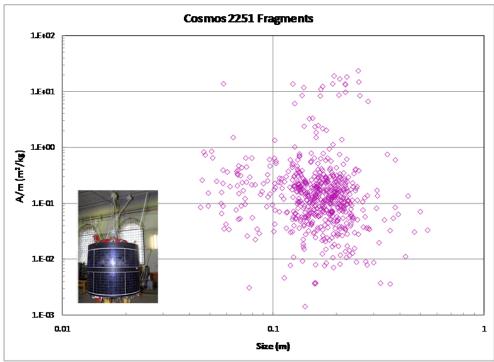












#### Physical characterization of the debris clouds IV



#### SSN Catalog data analysis

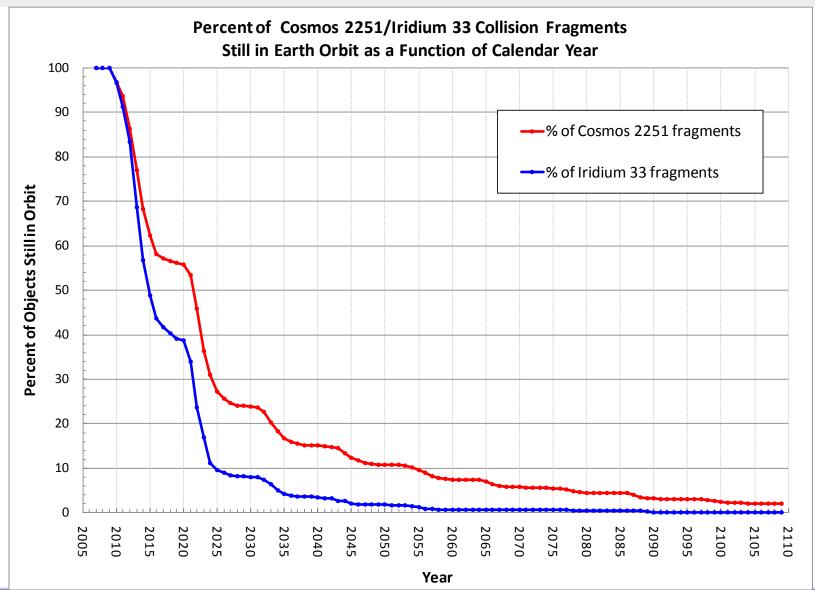
- □ ∆v and directionality distributions have been estimated for both clouds, but are currently under review
- Little or no momentum transfer observed in cataloged clouds

### Haystack/HAX data analysis

- Both clouds were observed by the Haystack and Haystack Auxiliary (HAX) radars shortly after the event
- Analysis ongoing

### Long-term evolution of the debris clouds





#### **Conclusions**



- A very large, very energetic event:
  - C2251: 1267 fragments cataloged; 1212 on orbit as of 10
     June 2010 SSN catalog
  - Iri33: 521 fragments cataloged, of which 498 are on orbit
  - History indicates that cataloging may continue for some time
  - Impact velocity highest of known intentional & accidental collisions
- These debris clouds will influence the LEO environment for decades to come
- Significant work remains to be done to understand origin of A/m distribution (Iri33), mass and size distributions in context of the NASA standard breakup model