



*Beyond Einstein: From the Big Bang to Black Hole*

## *LISA Technology Development at NASA/GSFC*

J.I. Thorpe

37<sup>th</sup> COSPAR Scientific Assembly

Montréal, Québec

July 16<sup>th</sup>, 2008



- 🪐 Laser Frequency Stabilization
  - Optical Cavities with frequency tuning
  - Molecular Iodine
- 🪐 Stable Environments
  - Stable test-bed for formation flying
  - Fused-silica fibers for torsion pendula
- 🪐 Surface Effects
  - Kelvin Probe
- 🪐 Laser Study

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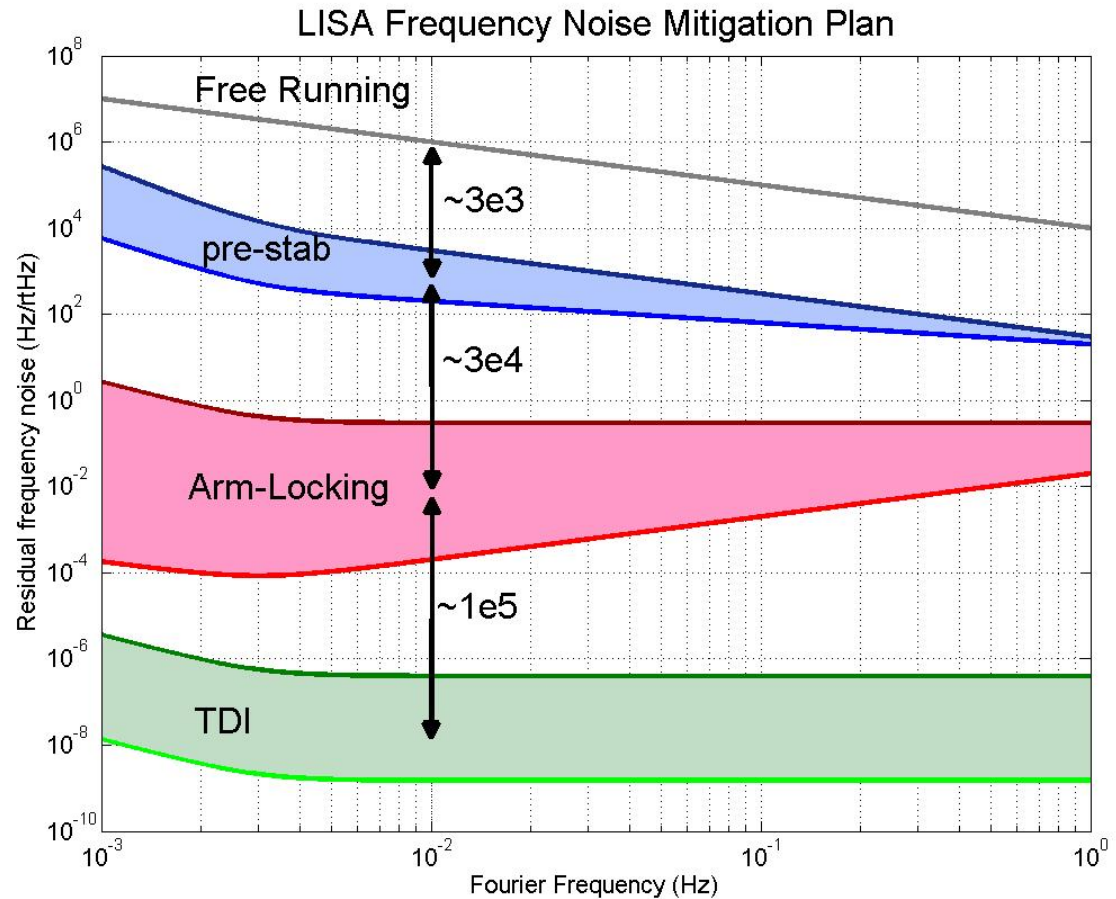
# Laser Frequency Noise in LISA



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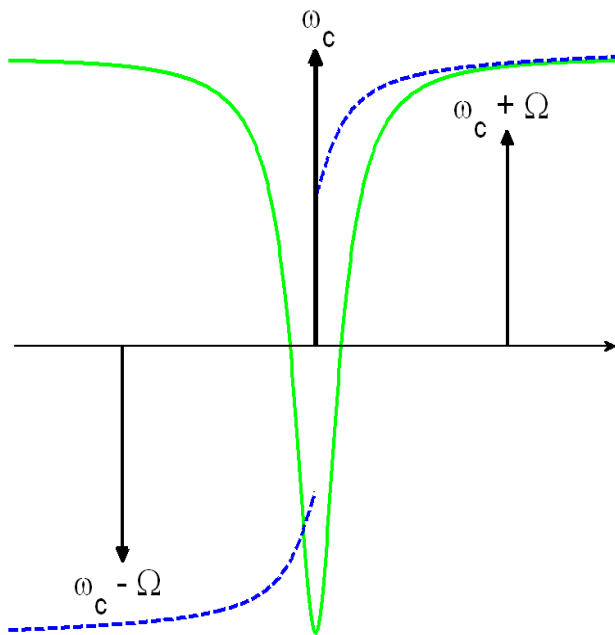
Laser frequency noise is a major potential noise source for LISA

- Three-stage system (two active one passive) to achieve overall suppression of  $\sim 10^{13}$
- Running pre-stabilization and arm-locking in series reduces gain (bandwidth) requirements on arm-locking.
- Serial arrangement *requires frequency-tunable pre-stabilization*

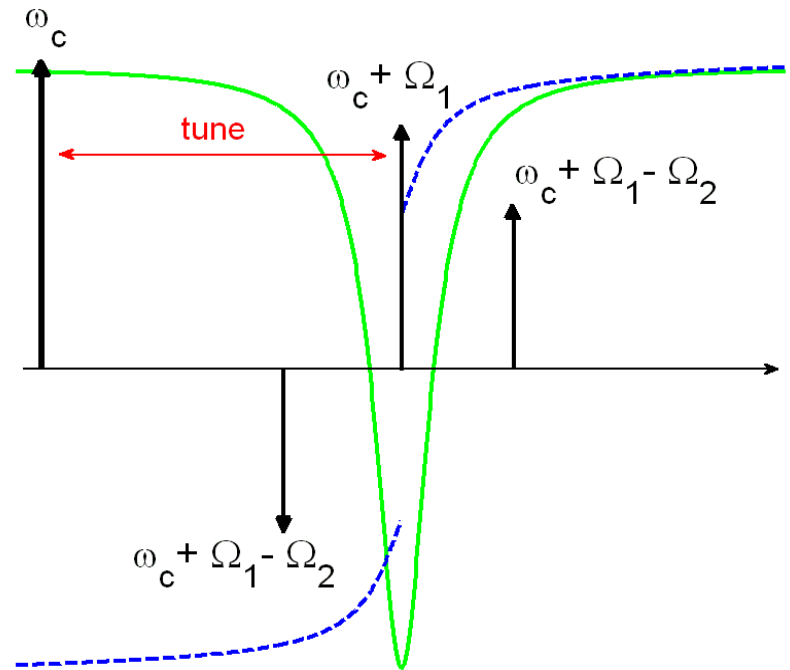


**Concept:** Lock phase-modulation sidebands to cavity resonance and tune central frequency by adjusting modulation frequency.

## Normal Pound-Drever-Hall Lock



## Sideband Lock



Thorpe, Numata, Livas

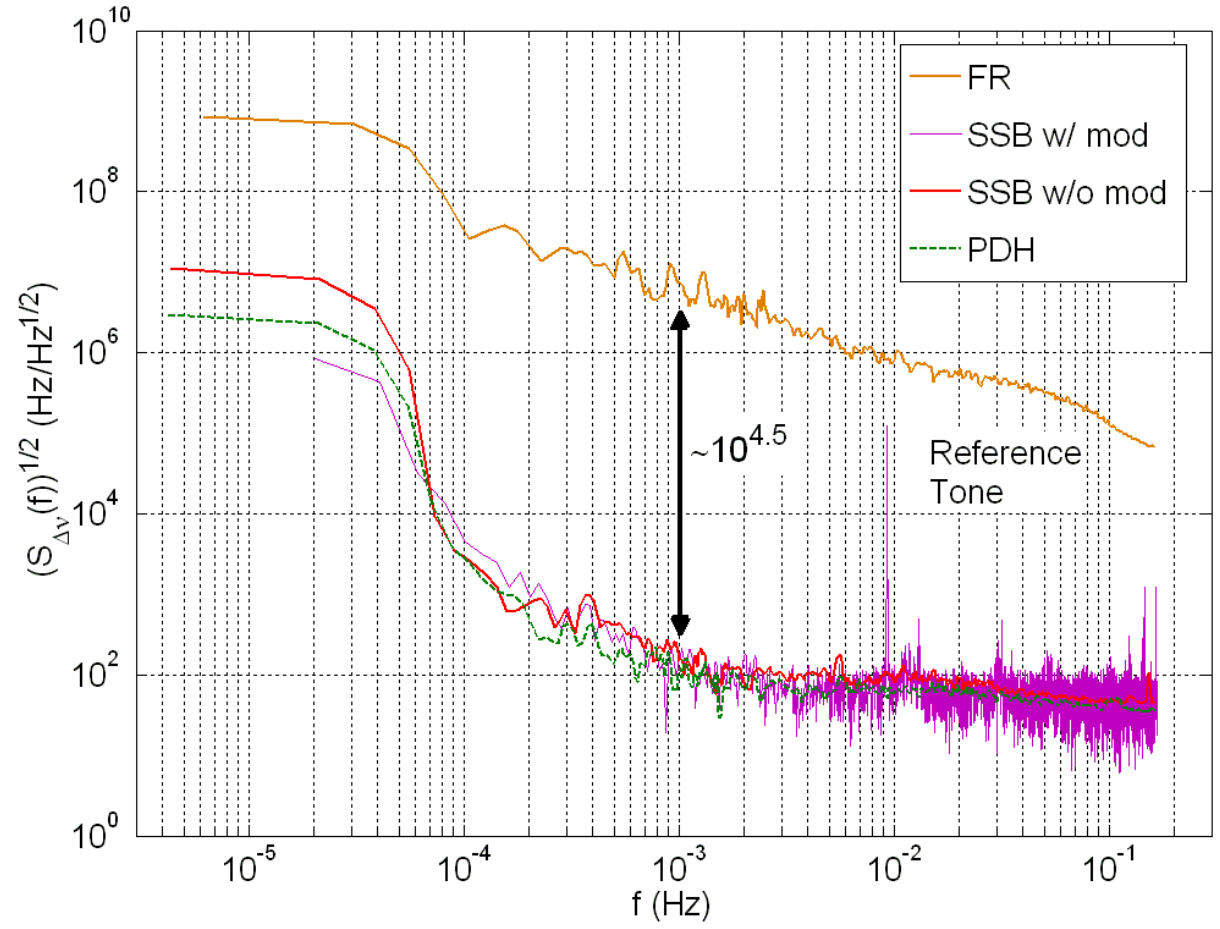


# Offset Sideband Locking with Optical Cavity



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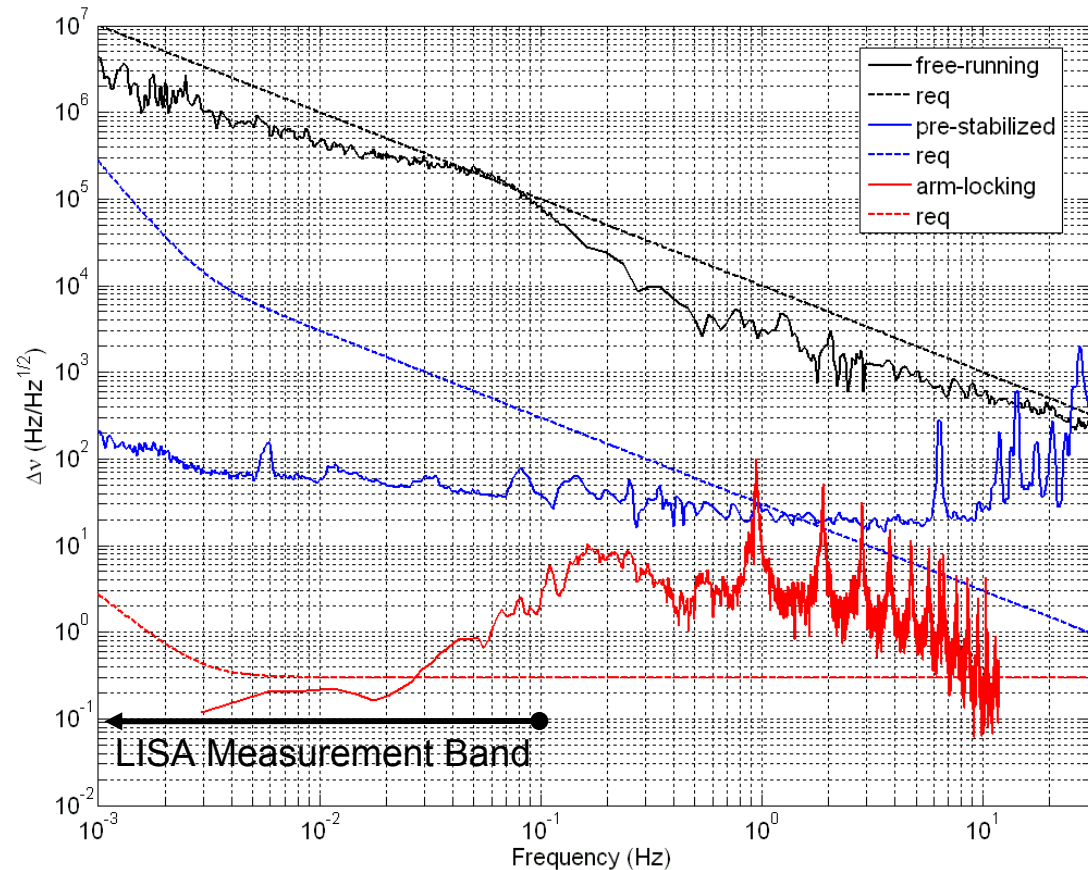
- Standard PDH and sideband locking have identical noise performance
- Common technical noises limit both systems.
- Adding modulation tone does not disturb the broadband noise floor.



Thorpe, Numata, Livas



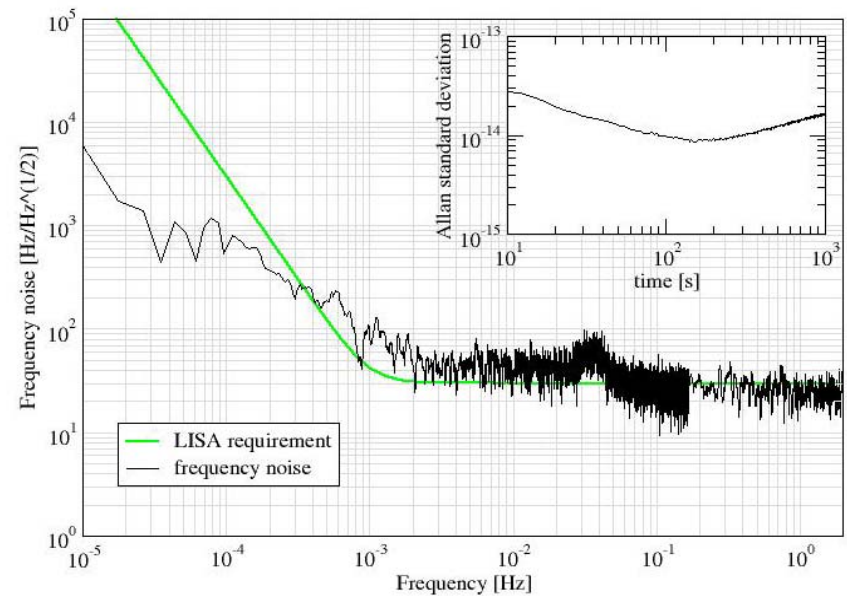
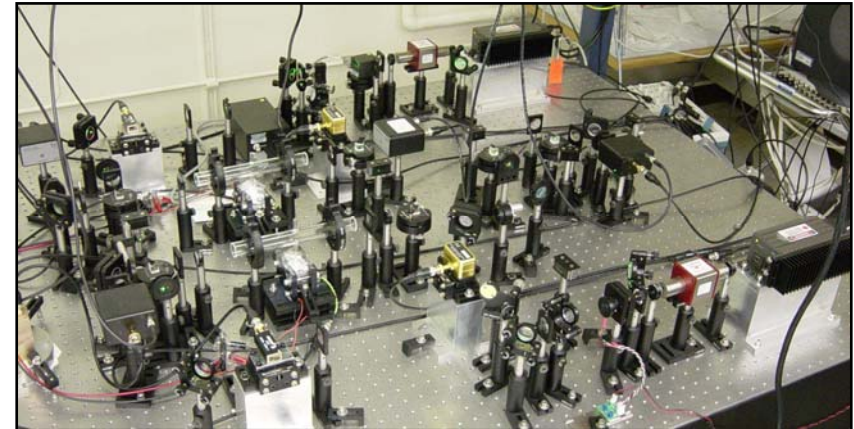
- Free-running and pre-stabilized lasers *meet LISA requirements in band*.
- Arm-locking system behaves as predicted. (noise spikes at  $n/\tau$  frequencies)
- Progress towards demonstration of 2/3 of LISA frequency mitigation plan.



Thorpe, Mitryk, Wand



- 🌐 Spectroscopic reference provides **Absolute reference frequency**
- 🌐 Laboratory study of frequency stability using two independent Nd:YAG lasers stabilized to hyperfine transition in  $I_2$
- 🌐 Slightly worse than cavities for  $f > 1\text{mHz}$
- 🌐 Better performance below  $0.1\text{ mHz}$



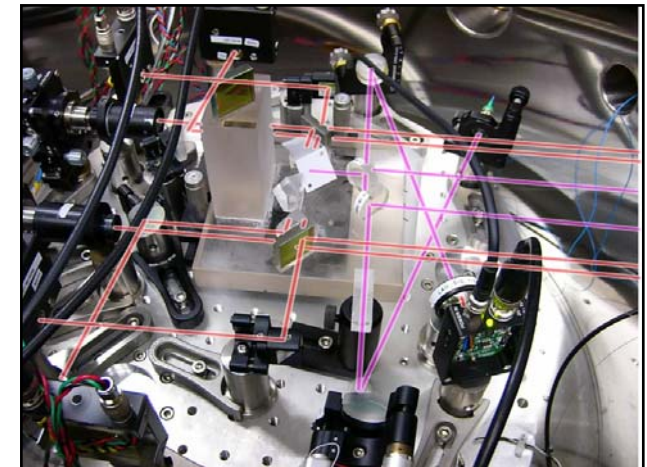
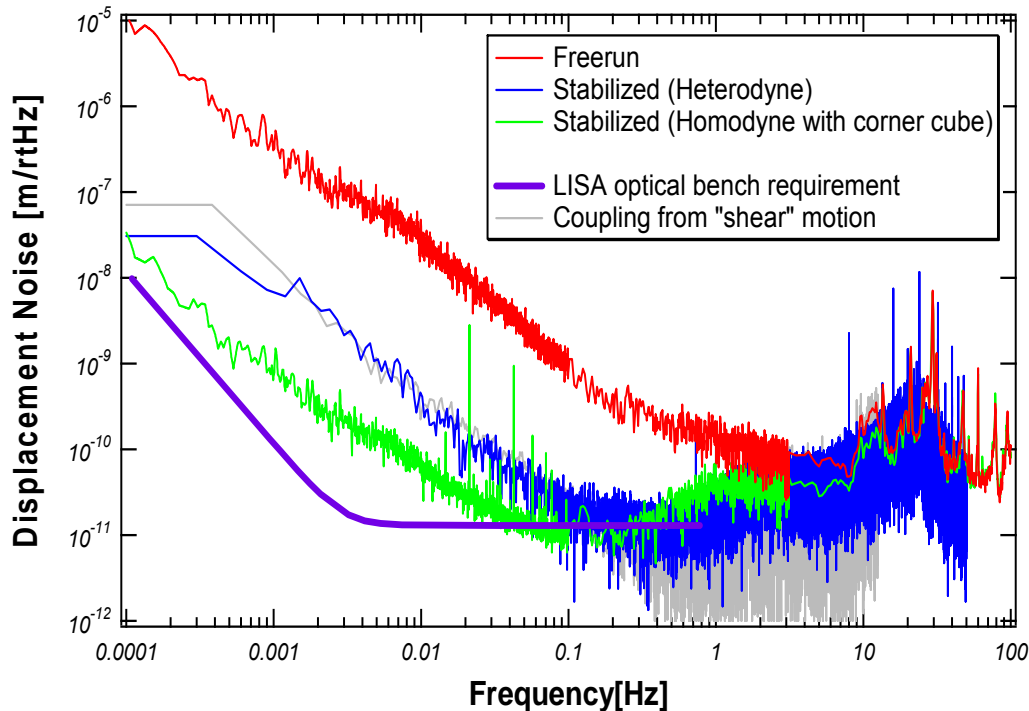
Leonhart & Camp

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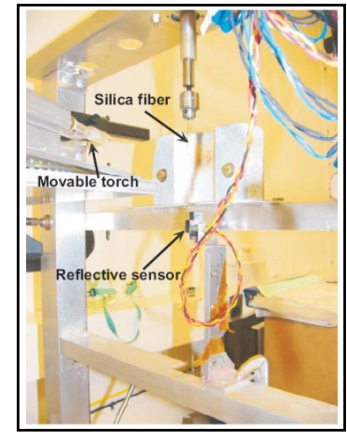
## Testing LISA's inter-spacecraft interferometer on stable platforms

Numata & Camp

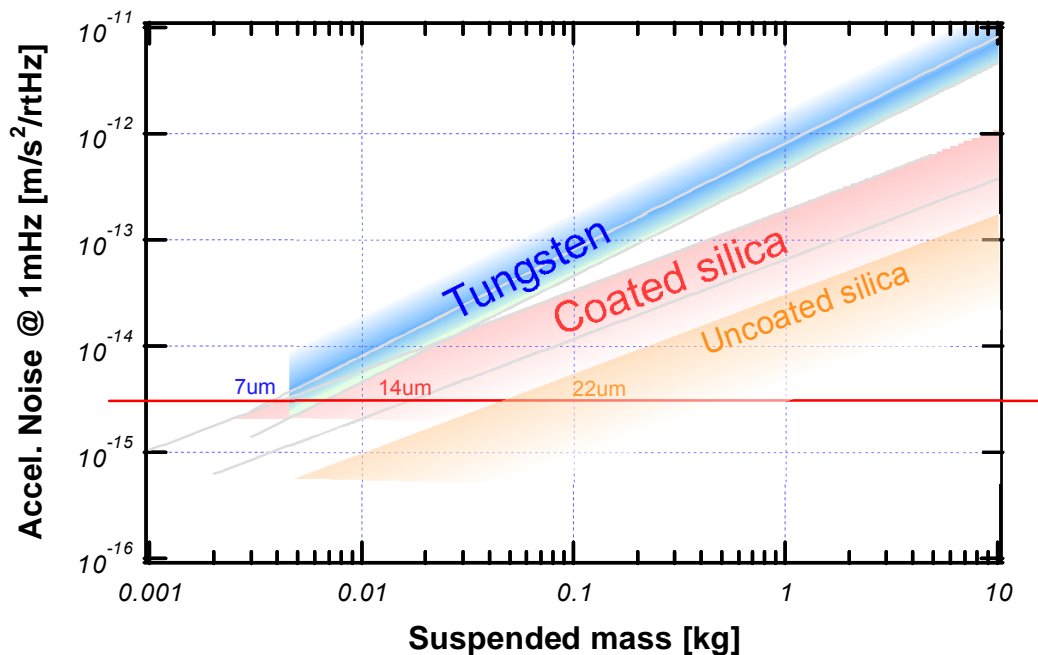
- 2 optical benches with 2 independent pre-stabilized lasers
  - Silicate bonded optical bench, heterodyne interferometer with phasemeter
- 2 degree-of-freedom active control
  - Intended to kill unwanted ground & thermal motion.
  - PZT-based hexapod provides actuation capability.
  - Noise suppression factor: 100~500
    - Performance limited by mechanical coupling from uncontrolled other 4 DoFs.



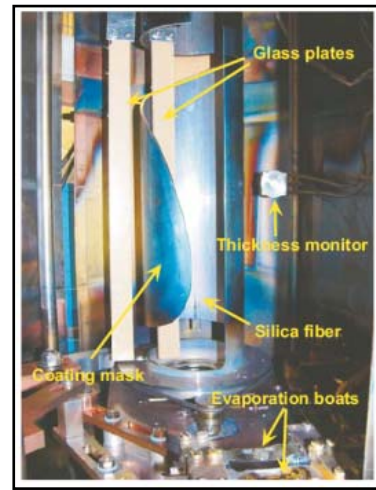
- For lowering fundamental noise limit of torsion pendulum
  - Our methodology
    - Fiber puller, coater, pendulum for loss measurement
    - Thin coating technique development
- Significant advantages confirmed
  - LISA requirement should be reachable with silica
    - Test started in LISA torsion pendula in Univ. of Trento & Univ. of Washington



Fiber puller



Numata & Camp



Fiber coater

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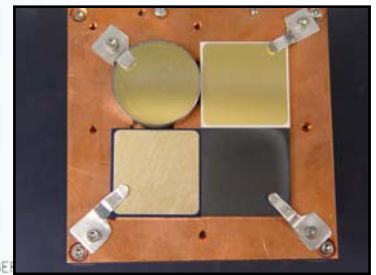
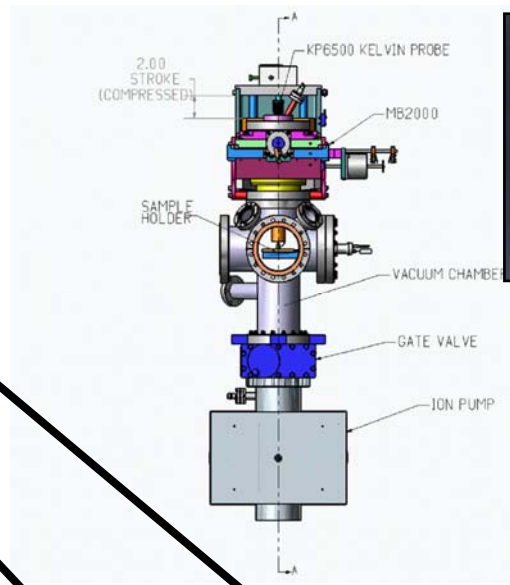


# KP measurements of LISA gold surface

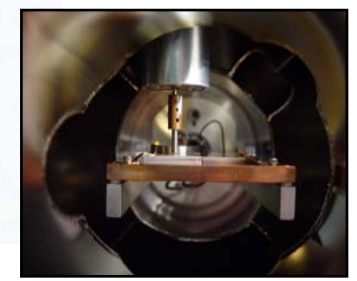


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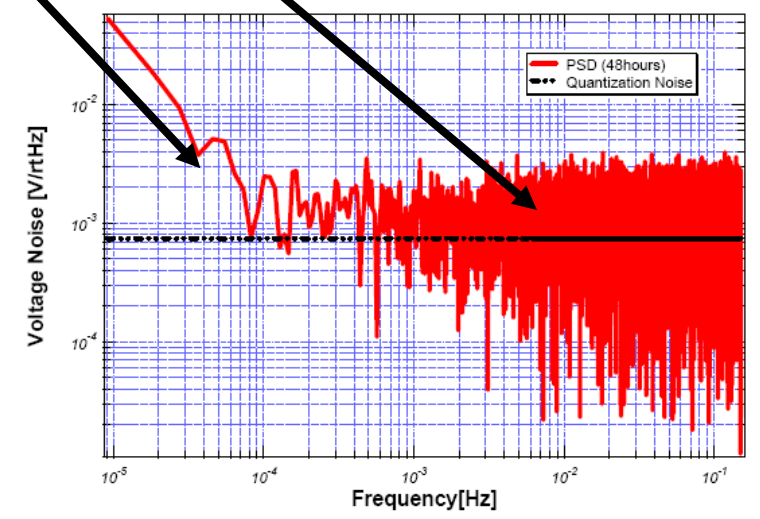
- 🌀 Vibrating probe induces current proportional to surface potential
- 🌀 KP limited by ADC quantization noise (recently upgraded)
- 🌀 Excess low frequency voltage noise of gold surface measured with KP
- 🌀 Magnitude barely OK for LISA, but cause unknown
- 🌀 LISA Advantages for patch-effect problem
  - Gold coatings are non-reactive
  - Test mass kept at room temperature



Coated Samples

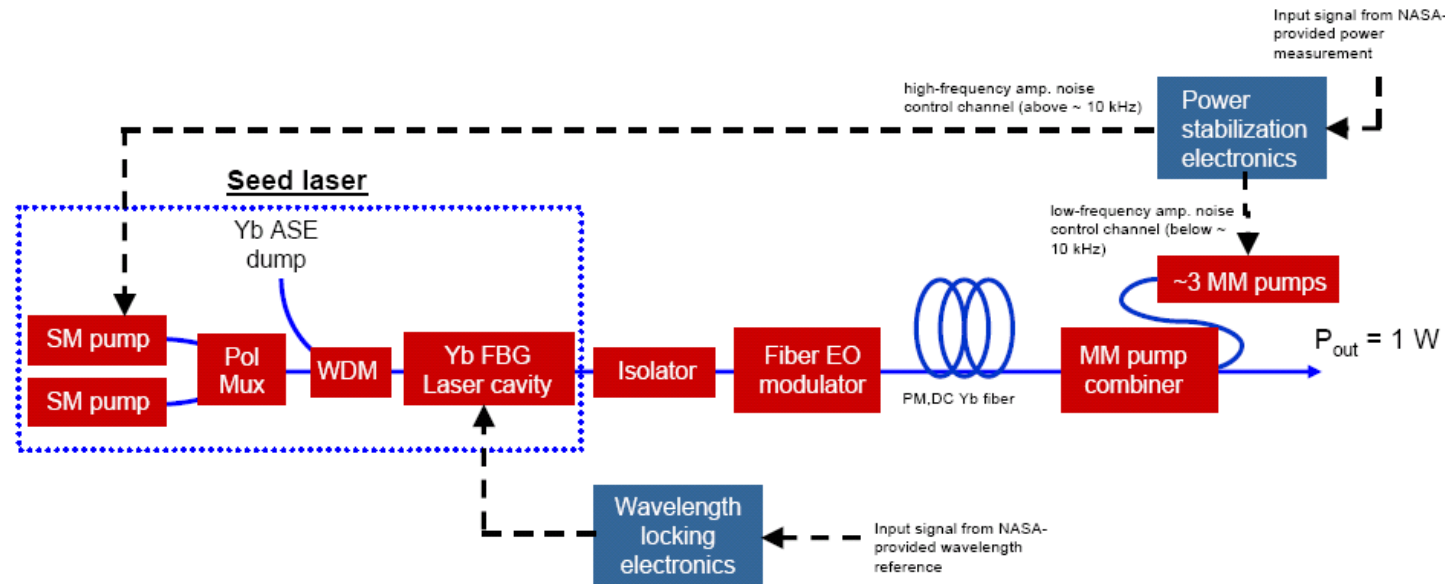


ST-7 sample under test



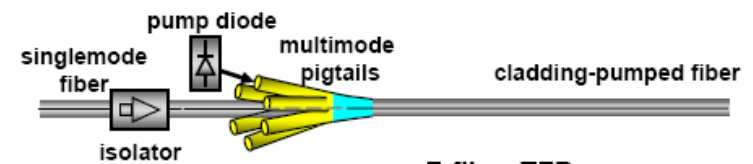
Camp

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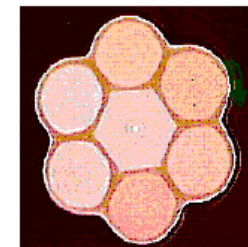


## Testing of pump combiner

- optical characterization (insertion loss and PER stability) from 5 - 70 C
- thermal screening under high power in vacuum
- temperature cycling in air

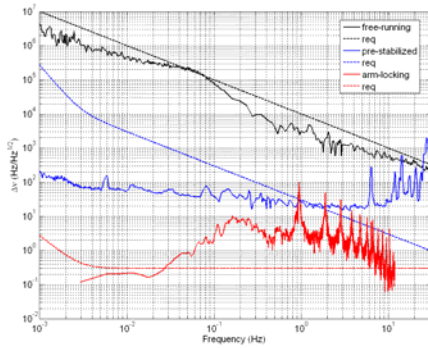
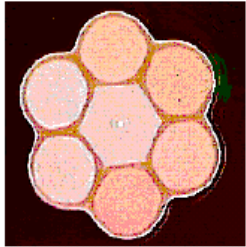


7-fiber TFB



J. Camp





## Contributors

Jordan Camp

Volker Leonhart

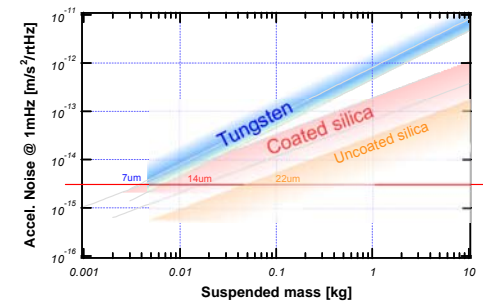
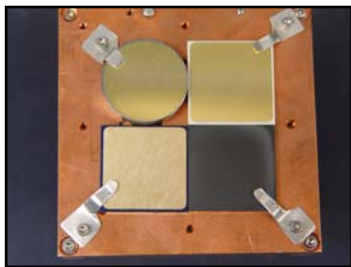
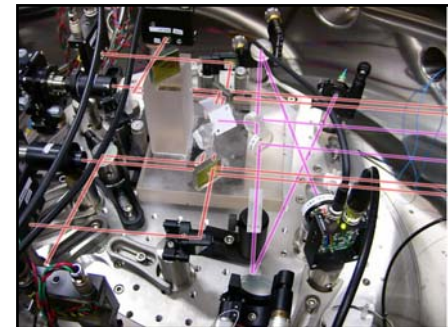
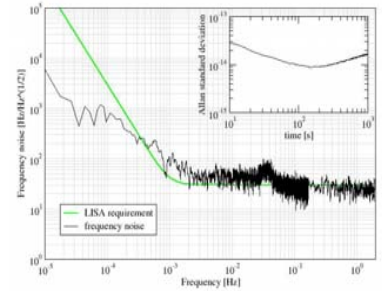
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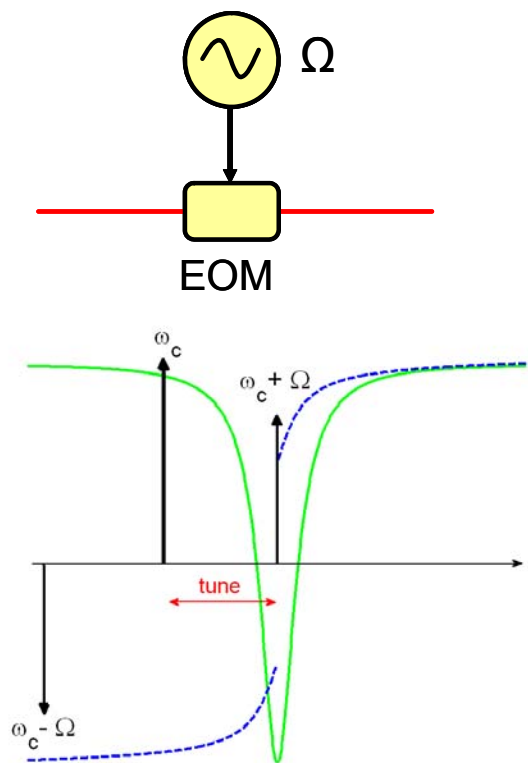
*Backup Slides*



GSFC - JPL

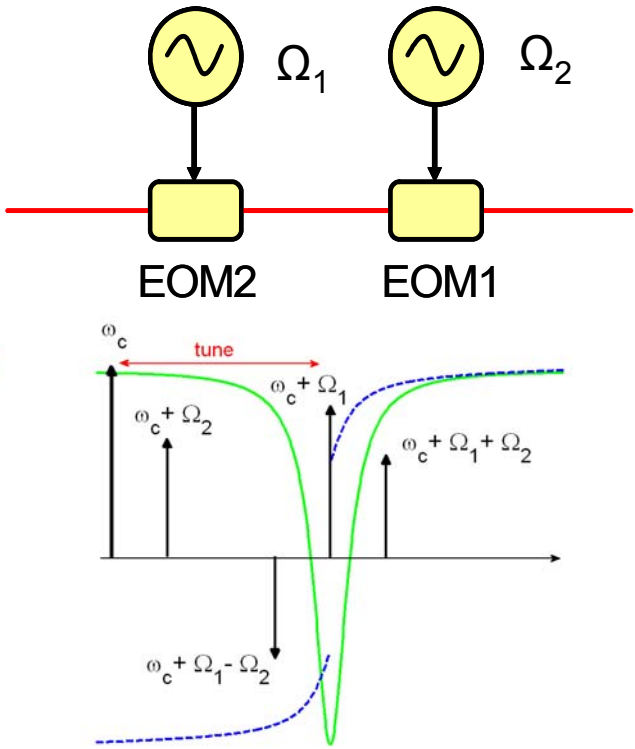


## Single Sideband (SSB)



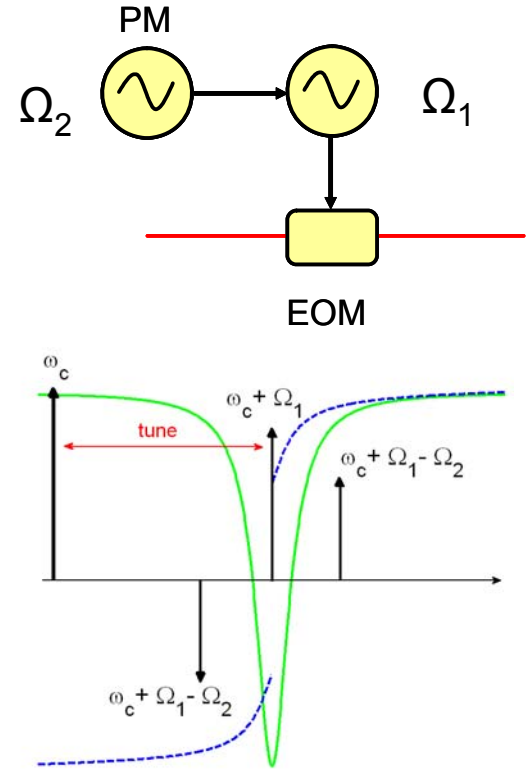
- Simplest to implement
- Some noise coupling due to asymmetry

## Dual Sideband (DSB)



- Restores PDH symmetry
- Complex modulation pattern

## Electronic Sideband (ESB)



- Simple, symmetric modulation pattern
- Requires phase modulation capability on LO



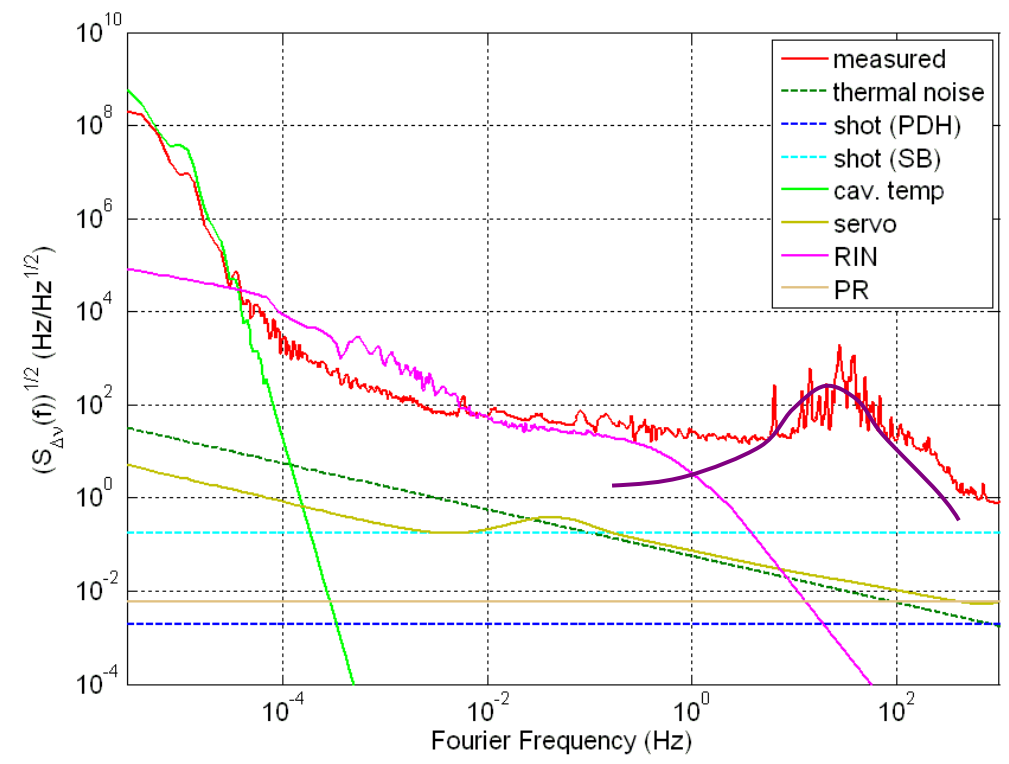
# Preliminary Noise Model



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- 📍 Fundamental Noise
  - Shot noise
  - Cavity thermal noise

- 📍 Technical Noise
  - Temperature Fluctuations
  - Servo Noise
  - Photoreceiver noise
  - RIN
    - via RFAM
    - via absorption
  - Vibration Noise/Acoustic
  - Pointing
  - ???





# Arm-Locking Transfer Function



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- Measured noise suppression matches expectations
- ~40dB at 100mHz

