

Detector Modeling and CMB Polarimetry Technology Development at GSFC

David T. Chuss, Edward J. Wollack, S. Harvey Moseley
NASA Goddard Space Flight Center

Stafford Withington, George Saklatvala
University of Cambridge



Introduction

- Planar detector modeling
 - Goal: To investigate the consequences of using planar bolometers in the limit in which pixel size is comparable to wavelength
 - We use the k-domain dyadic (Withington et al. 2003) to propagate the second-order statistical correlations of radiation through a model optical system
 - Model is general and preliminary, but it is unlikely more realistic bolometers will be better.
- CMB Polarization Technology Options
 - Modulators
 - Antenna-coupled detectors



Single-Mode

- Horn-coupled detectors
- Coherent across horn aperture
- Diffraction-limited resolution of optical system is dependent upon horn illumination of primary

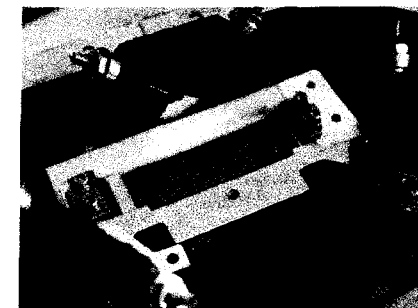
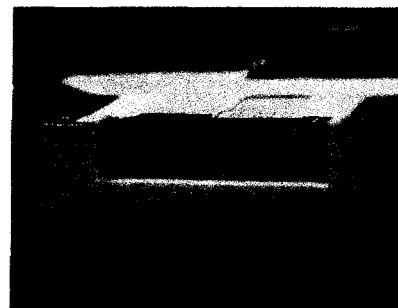
Multi-Mode

- Geometric Limit
- Incoherent across imaging element
- Diffraction-limited resolution normally determined by size of the primary

This work explores the intermediate case- Incoherent techniques in the limit where the Geometric limit is not strictly valid (few-mode limit)

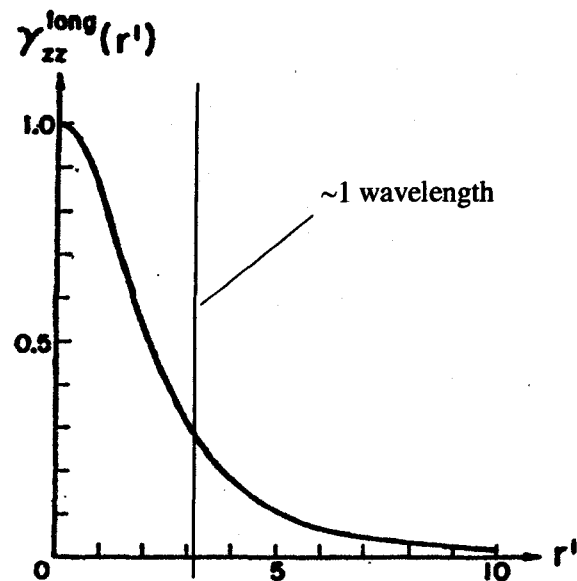


Bolometer Arrays

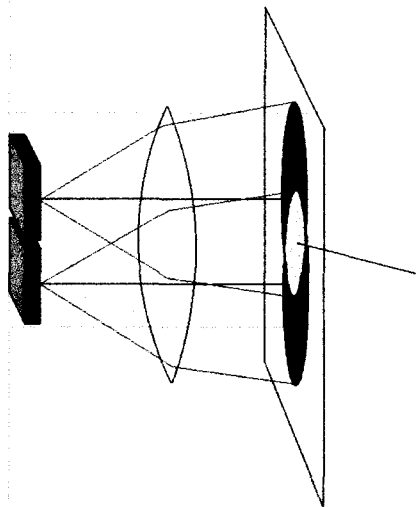


Selected Filled Focal Planes

Instrument	Array Size	Detector Type	λ (mm)	Pixel pitch (mm)	p/λ
HAWC/SOFIA	12x32	Semiconducting Bolometer	0.053	1.00	18.87
	12x32	Semiconducting Bolometer	0.088	1.00	11.36
	12x32	Semiconducting Bolometer	0.155	1.00	6.45
	12x32	Semiconducting Bolometer	0.215	1.00	4.65
SHARC II	12x32	Semiconducting Bolometer	0.350	1.00	2.86
	12x32	Semiconducting Bolometer	0.450	1.00	2.22
	12x32	Semiconducting Bolometer	0.850	1.00	1.18
SCUBA 2	64x64	TES	0.450	1.135	2.52
	32x32	TES	0.850	1.135	1.34
GBT	8x8	TES	3.00	3.00	1.00
GISMO	8x16	TES	2.00	2.00	1.00
ACT	32x32	TES	1.13	1.00	0.88
	32x32	TES	1.33	1.00	0.75
	32x32	TES	2.07	1.00	0.48



Beam Overlap



Space-Domain Dyadic

$$\bar{\bar{E}}(\bar{r}_1, \bar{r}_2) = \langle \bar{E}(\bar{r}_2) \bar{E}^*(\bar{r}_1) \rangle$$



K-domain Dyadic

$$\bar{\bar{A}}(\bar{k}_t', \bar{k}_t) = \frac{1}{(2\pi)^2} \iint \bar{\bar{E}}(\bar{r}_1, \bar{r}_2) e^{-j\bar{k}_t' \cdot \bar{r}_{t2}} e^{j\bar{k}_t \cdot \bar{r}_{t1}} e^{-jk_z z_2} e^{jk_z z_1} d^2\bar{r}_{t1} d^2\bar{r}_{t2}$$

$$\bar{\bar{E}}(\bar{r}_1, \bar{r}_2) = \frac{1}{(2\pi)^2} \iint \bar{\bar{A}}(\bar{k}_t', \bar{k}_t) e^{j\bar{k}_t \cdot \bar{r}_{t2}} e^{-j\bar{k}_t' \cdot \bar{r}_{t1}} e^{jk_z z_2} e^{-jk_z z_1} d^2\bar{k}_t d^2\bar{k}_t'$$



Stokes Parameters

$$I = E_{xx}(\bar{r}, \bar{r}) + E_{yy}(\bar{r}, \bar{r})$$

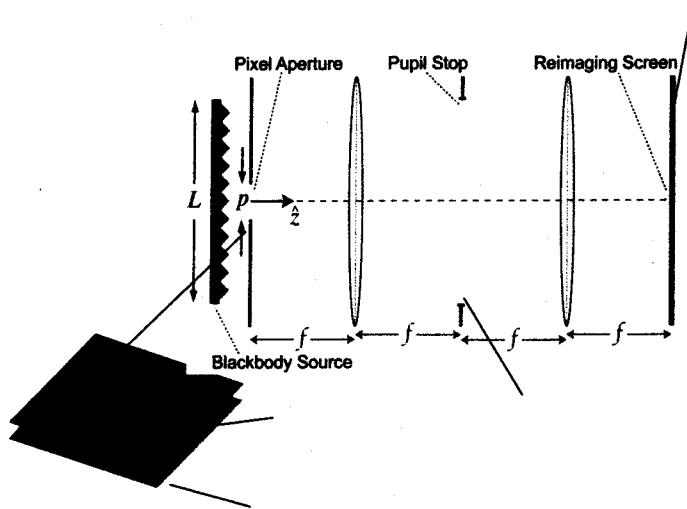
$$Q = E_{xx}(\bar{r}, \bar{r}) - E_{yy}(\bar{r}, \bar{r})$$

$$U = \Re E_{xy}(\bar{r}, \bar{r})$$

$$V = \Im E_{xy}(\bar{r}, \bar{r}).$$



Withington et al. 2003 Bolometer Modeling Method

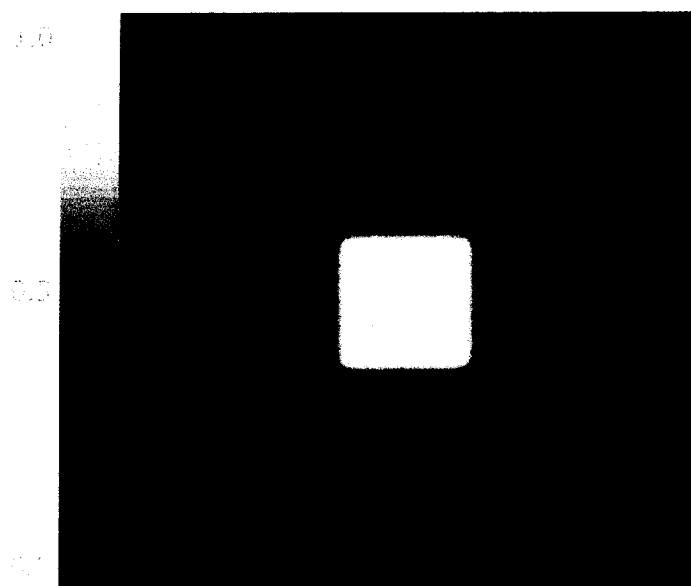


Technique

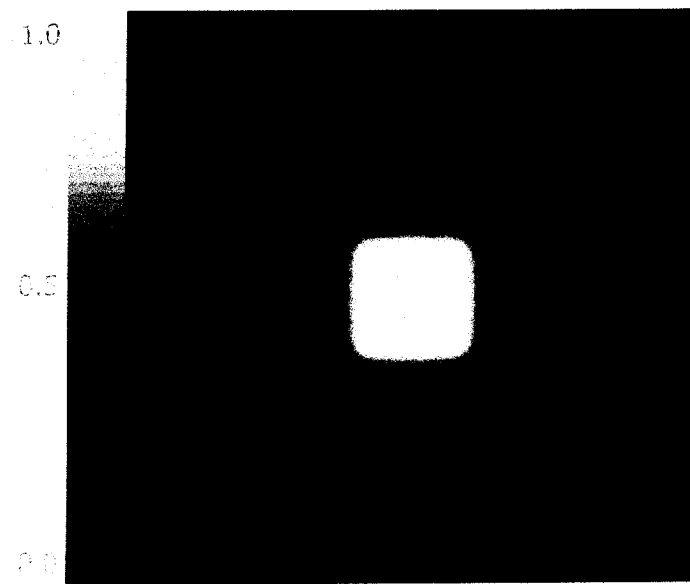
- Develop space domain correlation dyadic for blackbody radiation using plane wave expansion
- Transform to k-domain and scatter through aperture(pixel)
- Limit number of modes (pupil)
- Reconstruct the 2-D space domain correlation dyadic
- Construct the real Stokes parameters from the complex space domain correlation dyadic



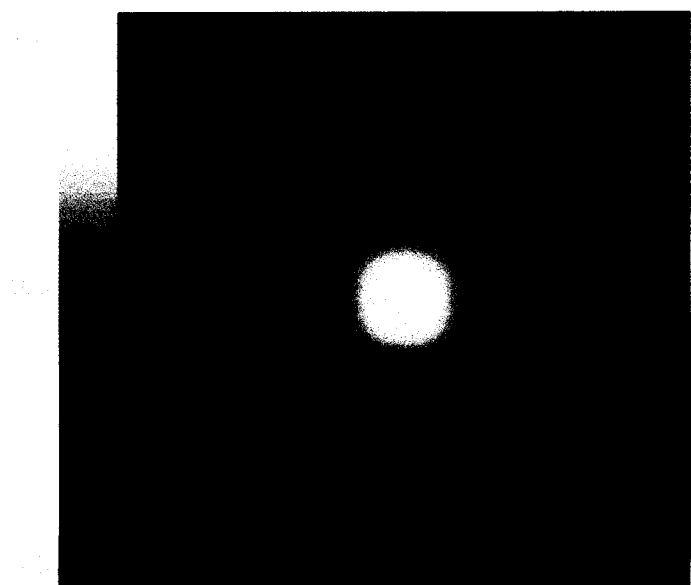
Stokes I, $p/\lambda=4.0$



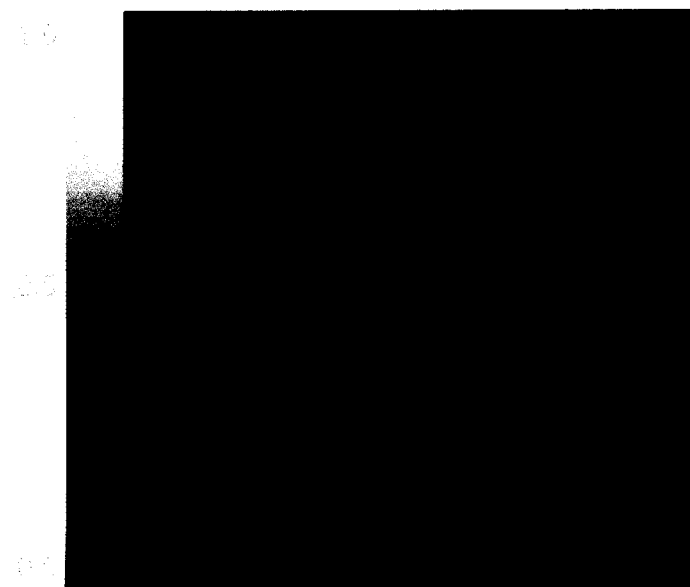
Stokes I, $p/\lambda=2.0$



Stokes I, $p/\lambda=1.0$



Stokes I, $p/\lambda=0.5$



Stokes I, $p/\lambda=0.25$

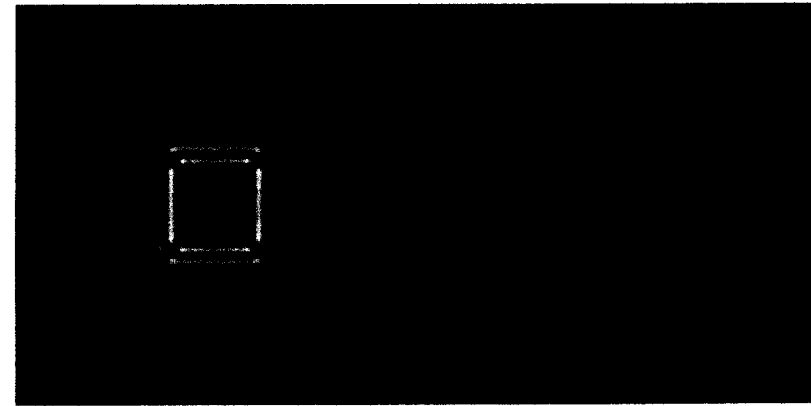
1.0

0.5

0.0



$P/\lambda=4$

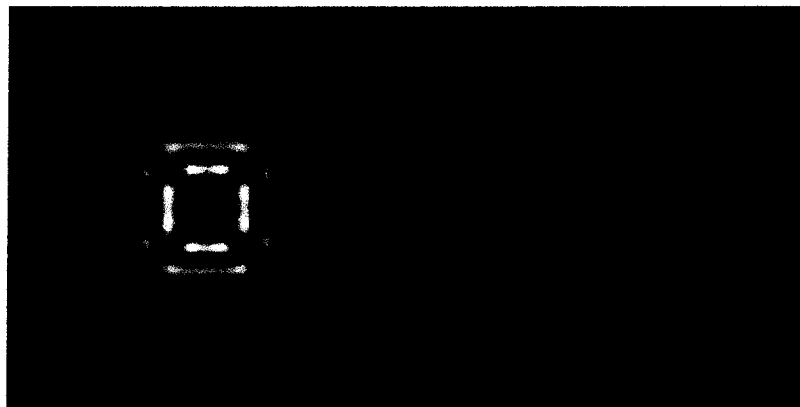


14

-4



$P/\lambda=2$



14

-4



$P/\lambda=1$

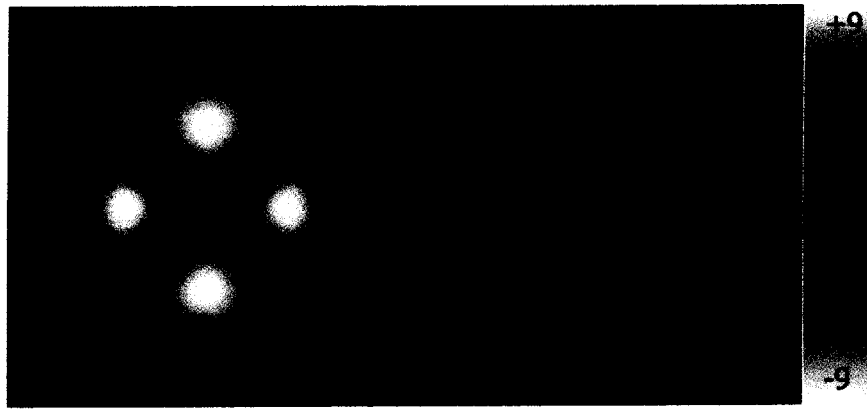


14

-4



$P/\lambda=0.5$

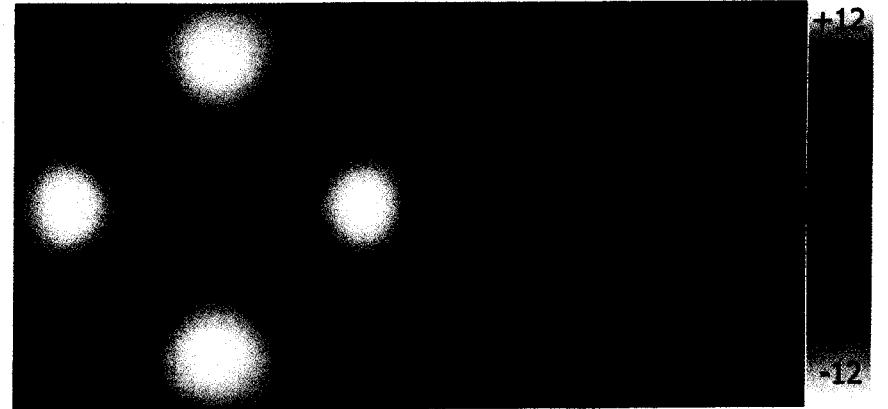


Q

U



$P/\lambda=0.25$

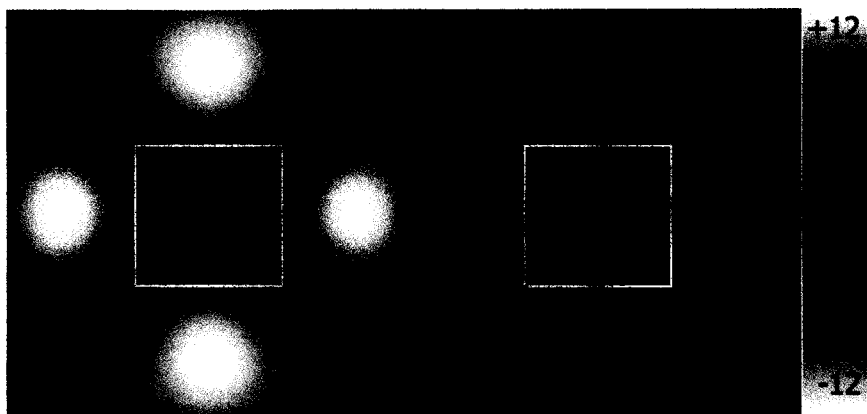


Q

U



$P/\lambda=0.25$

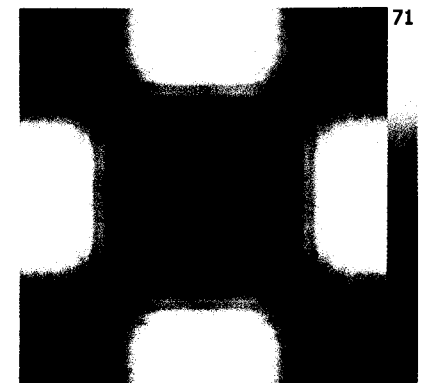
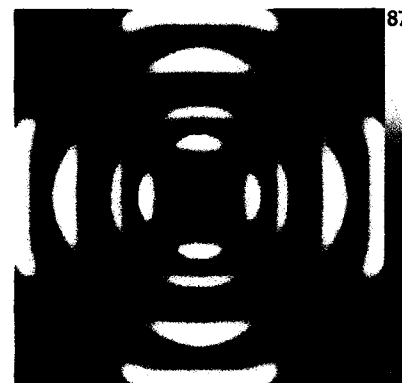


Q

U



Polarization



Summary

- Pixel size limits the resolution in the focal plane. This should be accounted for in optical design. Alternatively, this reduces the effective number of independent detectors.
- Polarization and scattering are intrinsically related, and both are more severe at low p/λ .
- Future work: Quantification of the pixel cross-coupling- calculate a theoretical covariance matrix to predict performance of future detector arrays.



0.01	0.08	0.01
0.08	1.00	0.08
0.01	0.08	0.01

$p/\lambda = 4.00$

0.02	0.13	0.02
0.13	1.00	0.13
0.02	0.13	0.02

$p/\lambda = 2.00$

0.07	0.25	0.07
0.25	1.00	0.25
0.07	0.25	0.07

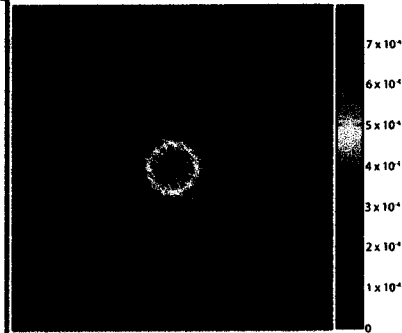
$p/\lambda = 1.00$

0.02	0.04	0.07	0.04	0.02
0.04	0.24	0.45	0.24	0.04
0.07	0.45	1.00	0.45	0.07
0.04	0.24	0.45	0.24	0.04
0.02	0.04	0.07	0.04	0.02

$p/\lambda = 0.50$

0.01	0.02	0.04	0.06	0.06	0.04	0.02	0.01
0.02	0.05	0.10	0.15	0.18	0.15	0.10	0.05
0.04	0.10	0.22	0.38	0.46	0.38	0.22	0.10
0.06	0.15	0.38	0.67	0.82	0.67	0.38	0.15
0.06	0.18	0.46	0.82	1.00	0.82	0.46	0.18
0.06	0.15	0.38	0.67	0.82	0.67	0.38	0.15
0.04	0.10	0.22	0.38	0.46	0.38	0.22	0.10
0.02	0.05	0.10	0.15	0.18	0.15	0.10	0.05
0.01	0.02	0.04	0.06	0.06	0.04	0.02	0.01

$p/\lambda = 0.25$



$$\beta(\Delta x, \Delta y) = \int \rho(x, y, \Delta x, \Delta y) dx dy$$

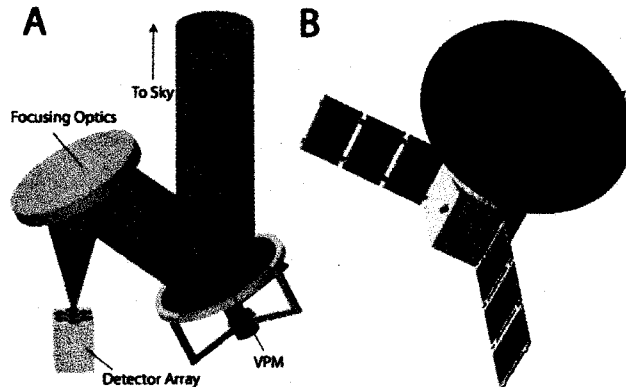
$$\Gamma = \int [I(x', y')]^2 dx' dy'$$

$$\rho(x, y, \Delta x, \Delta y) = \frac{1}{\Gamma} I(x, y) I(x + \Delta x, y + \Delta y)$$

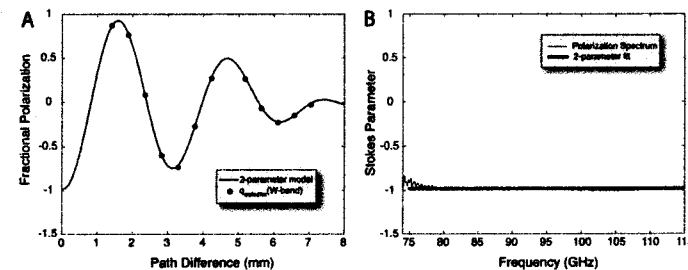
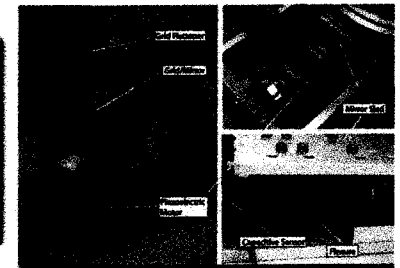
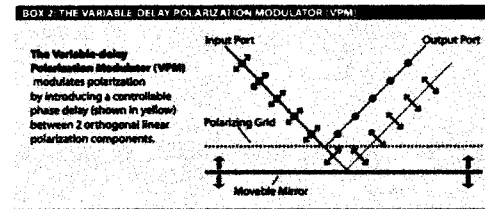


GSFC CMBPol Detector Effort

Harvey Moseley, Ed Wollack, Dave Chuss,
Gary Hinshaw, Al Kogut, Chuck Bennett (JHU)



Modulators



Laboratory
Tests at 90
GHz

Detectors

