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CHARGE-COUPLED DEVICE FOR LOW BACKGROUND OBSERVATIONS

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- $(*)$ Notice: Subject to any disclaimer, the term of this patent is extended or adjusted under 35 U.S.C. 154(b) by 0 days.
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- (22) Filed: **Jul. 8, 1999**

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- (51) **Int. Cl.**⁷ **H01L 27/148**; **H01L 31/0232**
- (52) **U.S. C1.** **250/208.1;** 2571437; 3481311
- (58) Field of Search 250/208.1, 214 R, 2501214.1, 214 A, 338.4; 2571225, 294, 432, 435, 436, 437; 3481298, 311, 314

(56) **References Cited**

U.S. PATENT DOCUMENTS

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(57) **ABSTRACT**

A charge-coupled device with a low-emissivity metal layer located between a sensing layer and a substrate provides reduction in ghost images. In a typical charge-coupled device of a silicon sensing layer, a silicon dioxide insulating layer, with a glass substrate and a metal carrier layer, a near-infrared photon, not absorbed in the first pass, enters the glass substrate, reflects from the metal carrier, thereby returning far from the original pixel in its entry path. The placement of a low-emissivity metal layer between the glass substrate and the sensing layer reflects near infrared photons before they reach the substrate so that they may be absorbed in the silicon nearer the pixel of their points of entry so that the reflected ghost image is coincident with the primary image for a sharper, brighter image.

9 Claims, 5 Drawing Sheets

CCD With Extra Metal Layer

FIG. 1b

 $\hat{\mathcal{A}}$

CCD Without Extra Metal Layer

FIG. 1a
Prior Art

FIG. 4

FIG. 5

HG. 6

40

CHARGE-COUPLED DEVICE FOR LOW BACKGROUND OBSERVATIONS

CROSS-REFERENCE TO RELATED APPLICATIONS

This application claims the benefit of priority under 35 U.S.C. \$119 (e) of U.S. Provisional patent application 601092, 489 filed Jul. 8, 1998, entitled "CCD System Technology for Extremely Low Background Observations," which is hereby incorporated by reference.

GOVERNMENT RIGHTS

The invention described herein was made in the performance of work under a NASA grant NAG5-3218 and by an 15 employee of the United States Government and is subject to Public Law 96-517 (35 U.S.C. 200 et seq.) and may be manufactured and used by or for the Government for governmental purposes without the payment of any royalty thereon or therefor.

BACKGROUND

FIELD OF THE INVENTION

The field of the invention is charge-coupled devices.

Charge-coupled devices (CCDs) have proven to be exceptionally versatile and effective detectors from the nearultraviolet through the near infrared. **As** demands for detector format or size and sensitivity increase, a number of problems have become the limiting factor for lowbackground observations. With larger detector formats, for example 4096x4096 photosensors or pixels, the problem of cooling to reduce dark current from bulk silicon layers of a typical CCD has become challenging because the exposed area of the device presents a large surface for absorption of ³⁵ thermal infrared photons from the environment.

CCD readout noise, which is introduced by the readout amplifier and the electronics outside the detector when reading a pixel without any photoelectrons, is a second limitation for low-background observations, especially when the devices can be cooled to their optimal operating temperature.

Charge-coupled devices (CCDs) have been used for astronomical observations since the 1970s. Their high quantum efficiency, low noise, relative stability and ease of use have brought about dramatic improvements in observational efficiency for ground based and space observations, culminating in the unprecedented discoveries made by missions such as the Hubble Space Telescope (HST). Space-based telescopes are more likely to be photon starved than their ground-based counterparts so that detector quantum efficiency is a greater factor in discovery potential. Typical requirements for space instruments such as the Space Telescope Imaging Spectrograph (STIS) on the HST are $\lt 25$ ₅₅ electrons/hr/pixel dark current and <4 electrons/pixel readout noise.

Three sources of noise in a typical CCD detector are photon noise, dark current and readout noise.

65 Dark current is caused by thermally excited electrons, 60 which are collected by the CCD pixel wells in the same way as photoelectrons are collected. [FIG.](#page-2-0) **3** shows the relationship between dark current and temperature. The dark current drops very rapidly with decreasing temperature. The emissivity of a typical silicon backside-illuminated, thinned CCD can be high (i.e. 0.66). Although silicon has no absorption in the thermal infrared, the silicon is attached to a glass

substrate, and $SiO₂$, the main constituent in glass, has a strong absorption band from $200-2000 \text{ cm}^{-1}$ that covers the thermal infrared band. Also, there may be a 1 um-thick layer of SiO, for electrical insulation, and this layer also absorbs some thermal radiation. A typical photon in the thermal infrared is absorbed within tens of micrometers of entering the glass substrate. For longer integration times $(i.e. >1000$ seconds), the dark current becomes significant.

The total noise in a CCD system can be modeled as:

10 $n_r=(n_r+(Dt)+n_r^2)^{1/2}$

Where n_t is the total noise for a pixel, n_s is the number of $photelectrons$, D is the dark current in electrons/pixel/ second, t is the integration time in seconds, and n_r is the readout noise in electrons. Additional noise may also be generated by the environment of the detector, such as cosmic rays.

Because of its simple operation, the CCD detectors are used for many observations between 800 and 1000 nm but ghost images in these wavelength ranges reduce quantum *20* efficiency. In the near infrared, 800-1000 nm, a sizable fraction of the light passes through the CCD without being absorbed and converted into free electrons. The absorption length in silicon is greater than the thickness of a typical CCD silicon layer of 13 um so that for wavelengths longer 25 than 800 nm, less than 1/e of the light is absorbed in the first pass. At λ <600 nm, all of the light that enters the CCD is detected (some of the light is reflected off of the siliconvacuum interface), but at 900 nm, the quantum efficiency is much reduced, typically about 30%. The result is that the light that is not absorbed during the first pass through the *³⁰*silicon forms a ghost image, which is enlarged and, if the focal plane is tilted, displaced from the pixel nearest its point of entry into the CCD detector.

It is desired to decrease the dark current and readout noise to increase overall quantum efficiency of a charge-coupled device system in low background or low photon environments, and in particular, to reduce ghost images caused by light not immediately absorbed by the silicon.

SUMMARY

The addition of a low emissivity reflecting layer to a typical charge-coupled device has improved overall quantum efficiency. The present invention is a charge-coupled device comprising a substrate, a sensing layer and a reflecting layer wherein the reflecting layer is positioned between 45 the sensing layer and the substrate. The reflecting layer may be a metal such as aluminum for example. Typically the substrate will be a glass substrate. The sensing layer will be silicon. In addition, an insulating layer of silicon dioxide will typically be positioned between the sensing layer and SO the metal reflecting layer.

Aparticular embodiment of the invention would comprise a charge-coupled device having a glass substrate have a front side and a back side. On its back side would be located a metal carrier layer and on its front side, the metal reflecting layer, typically aluminum. The other side of the metal reflecting layer would be adjacent to the silicon insulating layer which is positioned between the sensing layer and the metal reflecting layer.

It is also an object of the invention to provide a system that combines the above embodiments of the charge-couple device with a correlated double sampler that controls a readout rate for the device so that the rate may be slowed down to reduce readout noise.

BRIEF DESCRIPTION OF THE DRAWINGS

FIG. la shows the absorption of a near infrared photon **(12)** and a thermal infrared photon **(14)** in a cross-section of a typical charge-coupled device (CCD) without the metal reflecting layer **(10).** TABLE 1-continued

FIG. **lb** shows the absorption of a near infrared photon **(12)** and a thermal infrared photon **(14)** in a cross-section of layer **(20).** a typical charge-coupled device with the metal reflecting 5

(32) and a metal-backed CCD **(34).** [FIG.](#page-2-0) 2 shows the infrared absorbance of a typical CCD

(38), normalized to -100° **C., vs. temperature.**

FIG. 4 shows the comparison between the magnitudes of a point source magnitude (PS) and a large source magnitude (LS) for a standard CCD $(+)$ and a metal-backed CCD (X)

[FIG.](#page-3-0) 5 shows the image of a white-light point source

FIG. **6** is a plot of the intensity of the first order [+for

FIG. **7** is a block diagram for the integrating correlated double sampler readout electronics.

DETAILED DESCRIPTION

(10) comprising an upper layer of silicon sensing material photon (12) makes a first pass (23) through the upper layers expected to reach the glass substand reflects from the metal carrier. The near infrared photon have a large infrared absorbtivity. and reflects from the metal carrier. The near infrared photon

entered so that it contributes to a ghost image.

FIG. 1*b* shows an embodiment of the present invention is changed from n₀ to wherein a low-emissivity metal reflecting layer (30) is $n=[n_0^2-(\omega_0^2\tau/\omega(\omega\tau+i)]^{1/2}$. inserted between the SiO₂ insulating layer (18) and the glass Where ω is the angular frequency, ω_p is the plasma substrate (22). In this embodiment, a thermal infrared phowherein a low-emissivity metal reflecting layer (30) is that it will not be absorbed by the glass substrate, thereby

Table 1 shows an example of CCD layer structure comprising the metal reflecting layer. Use of the term metal In further analysis, it was discovered that the periodic backed CCD or modified CCD refers to embodiments of the 55 structure of the CCD causes diffraction at wavelengths present invention. longer than 700 nm. Therefore, the reflection spectropho-

	Thickness (um)		
Material Layer	Standard CCD	Modified CCD	
Silicon sensing region	13	13	
Silicon dioxide insulator	0.1	0.1	
Polysilicon gate	0.5	0.5	
Silicon dioxide insulator	1.5	1.5	6
Metal reflecting layer	none	0.1	

3 4

			Thickness (um)	
	Layer Material	Standard CCD	Modified CCD	
6 7	Glass substrate Metal Carrier	700	700.	

[FIG.](#page-2-0) 3 shows the relationship between CCD dark current 10 For the embodiment shown in Table 1, initial results indicated that there was no essential difference measured for thermal infrared emissivity between the modified and a standard CCD. However, it was discovered that red light reflected from the CCD is diffracted by the periodic structure (15) for a standard CCD (+) and a mean-backed CCD ($^{(2)}$) 15 of shift registers in the CCD. Apparently a reflection spec-
trophotometer used in the measurement only measured the zeroth order reflection and missed the other orders.

reflected from a CCD. In the initial measurements, the infrared absorbance of the standard and modified CCDs from $500-4000$ cm⁻¹ (20-2.5) $(1,0)$; X for $(-1,0)$] diffraction peaks relative to the zeroth 20 **u**) were measured with a fourier-transform spectrophotomorder ueak for a metal-backed CCD. eter. The results are shown in [FIG.](#page-2-0) **2** showing the absorbance, normalized to $-log_{10}$ (Reflected flux /Incident flux), versus v $(cm⁻¹)$ for a standard CCD (32) in the top portion of the graph and for the modified CCD **(34)** in the 2s bottom portion of the graph. The gross features of the absorbance are essentially the same for the two CCDs. The FIG, $1a$ shows layers in a typical charge-coupled device spacing between the interference fringes is about $\delta v=100$
(a) comprising an upper layer of silicon sensing material cm^{-1} , which indicates the main reflection i (16) here 0.013 mm thick. Beneath this layer is a silicon back and front of the silicon. At v>3000 cm⁻¹, a finer insulating layer of SiO. (18) which is 0.001 mm thick in this ³⁰ spacing is evident for the standard CCD. insulating layer of SiO_2 (18) which is 0.001 mm thick in this 30 spacing is evident for the standard CCD. Here, the light version is located on a glass substrate (22) shown to be passes through the glass substrate and g version, is located on a glass substrate (22) shown to be passes through the glass substrate and gives a spacing of $(0.750 \text{ mm} \text{ in this example})$. The glass substrate has a front $\delta v = 5.3 \text{ cm}^{-1}$. In the modified CCD, the meta (0.750 mm in this example). The glass substrate has a front $6V=5.3 \text{ cm}^{-1}$. In the modified CCD, the metal layer prevents side (26) and a back side (28). Beneath the glass substrate is the light from entering the glass, side **(26)** and a back side **(28)**. Beneath the glass substrate is the light from entering the glass, and the fine fringes are not a metal carrier layer **(24)**. A typical thermal infrared photon seen. The modified CCD was a metal carrier layer (24). A typical thermal infrared photon seen. The modified CCD was expected to have a much lower
(14) is absorbed in the glass substrate. Also a near infrared 35 absorbance because much of the infrare (14) is absorbed in the glass substrate. Also, a near infrared ³⁵ absorbance because much of the infrared light was not photon (12) makes a first pass (23) through the upper layers expected to reach the glass substrate,

may then be absorbed on a return path (25) in the silicon It was observed that the polysilicon gates, which had been
sensing layer but far from an original pixel near where it ignored, also had a high infrared absorbtivity sensing layer, but far from an original pixel near where it ignored, also had a high infrared absorbtivity, because of the entered so that it contributes to a ghost image 40 free charges. With the free charges, the index o

ton is shown having a path (27*a*) ending at the metal ⁴⁵ hequency, and the damping time. The mass of entantion reflecting layer where it is reflected in a return path $(27b)$, so because n^2 has an imaginary part. Furthermore, at some that it will not be absorbed by the giass substrate, thereby
decreasing the dark current in the CCD. Furthermore, a near
infrared solution was to make an embodiment wherein the metal
infrared photon (12) that is not absor infrared photon (12) that is not absorbed in a first pass $(21a)$ so reflecting layer be inserted between the silicon sensing reflects back for a second pass $(21b)$ nearer to the original region and the polygilian gate pixel (29) it passed through so as to reinforce the image. system was updated to include the absorption by the poly- 45 frequency, and τ is the damping time. The index of refraction region and the polysilicon gate layer. Modeling of the silicon layer.

> tometer does not measure the total reflectivity and underes-TABLE 1 timates the reflectivity.

> > Although the results for the thermal infrared light was not 60 as expected, the metal-backed CCD did show improvement with respect to alleviating ghost images.

> > Awhite-light point source was imaged that reflected off of a CCD. In the picture of FIG. 5, there are many images, which can be labeled by the order numbers (n,m) , where n 55 and m count orders that are in the direction of the parallel and serial shift registers, respectively. The (0,0) image (52) is sharp and saturated as the angles of incidence and reflec-

The (1,0) order (54) to the southeast is also particularly and not as the usual expression $(\epsilon_{MB}/\epsilon_{STD})^{1/2}$. At 950 nm, the bright. The other orders are elongated because of the spread detection limit improves by *0.75* mag.

strongest diffraction. A weaker diffraction pattern with m changing is due to a structure perpendicular to the parallel The readout noise, that which arises by reading a pixel

There is a substantial amount of light in the diffracted orders. The (1,0) order contains as much as 60% of the light 20 For early CCD's, the amplifier noise power was indeed
in the zeroth order at near-infrared wavelengths accessible to f^{-1} , and the readout noise was inde the CCD. Therefore, it is likely that the thermal infrared amplifiers have improved in the readout noise. For the reflectivity is substantially underestimated with the reflec-
amplifier on a typical CCD, the f-1 noise and

long wavelengths $(\lambda > 750 \text{ nm})$ passes through the silicon detecting layer without being absorbed. That light passes readout rate.

into the glass substrate, reflects off of bubbles in the glass Alower readout noise vields a substantial improvement in into the glass substrate, reflects off of bubbles in the glass $\frac{1}{\sqrt{2}}$ and the metal carrier, and returns through the sensing layer performance. If the sampling rate were slowed from *50* KHz as a diffuse ghost image. In a metal-backed **CCD,** in 30 to *20* KHz, the noise would be reduced by a factor of 1.4. reflects back at essentially the same pixel and gets a second in the readout rate, but the exact improvement depends on chance to be absorbed, which results in a higher quantum the detailed noise characteristics of the readout amplifier. efficiency. For long, photon-starved exposures where the readout noise

was measured in the following way. A point source and a increase, and the required integration time would be lowered large source with a diameter of 50 pixels were imaged. The by up to a factor of 2 (depending on the sky b sources were made by placing a pin hole or a larger hole in dark current contributions). For short integrations, the rate front of a diffused light. The point source was photometered can be commanded to a large value (up to 100 Khz) since the with an aperture of diameter of 8 pixels. Ideally, a detector 40 source would be bright and noise is not an issue. should have the same efficiency regardless of the size of the The proposed integrating correlated double sampler (FIG. image. For such a detector, the ratio of the fluxes of a point **7)** is similar to the standard "correlated double sampling" source and a large source should be the ratio of the areas of method, but has the additional feature that the noise is the two holes, and therefore should be independent of color. processed in a true integrator during each of the two signal The measurements of the flux ratio expressed in magnitudes, 4s phases. This allows for easy implementation of a commandmag(PS)-mag (LS), as shown in FIG. 4, for the standard able readout rate to optimize an observation. The sequence (42) and metal-backed CCD's (44) show the differences. For for reading the charge on a picture element is as **(42)** and metal-backed CCD's **(44)** show the differences. For the standard CCD, mag(PS)-mag(LS) is independent of the on-chip capacitor (94) is discharged by the reset switch color at λ <700 nm but falls at longer wavelengths. At 950 (92) , and the integrator capacitor (98) nm, it drops by *0.75* mag which is a factor of *2.* **As** shown *SO* Before the charge packet is shifted onto the on-chip in FIG. 4, for the metal-backed CCD, mag (PS)-mag(LS) is capacitor, the signal level is integrated on the integrator independent of color.

standard and metal-backed CCD at λ <700 nm as the photons arrives on the on-chip capacitor, the signal level is integrated are absorbed before they reach the metal backing. At λ >700 55 on the integrator capacitor wit are absorbed before they reach the metal backing. At λ >700 55 nm, the effective quantum efficiency depends on the size of ing to the positive sign input **(100)** for the same time. 4) The the image for a standard CCD but not for a metal-backed input to the integrator is removed, and the output is digitized CCD. For some astronomical problems, this is a serious via analog to digital converter (110). Thus, th problem for standard CCDs because a calibration made with both filters and takes the difference between the signal level
a defocused star does not apply to in-focus stars. For 60 with and without the charge packet. With th a defocused star does not apply to in-focus stars. For 60 background limited observations with the metal-backed readout rate can easily be optimized for an observation CCD, there is a substantial improvement in detection limit without changing the fixed component values that are typi-
beyond the gain in quantum efficiency. The additional gain cally used to set the bandwidth of the standa arises because in a standard CCD, the photons that reenter double sampler.
the sensing layer far from where they entered contribute for 65 Obviously, many modifications and variations of the the sensing layer far from where they entered contribute for 65 the background but not for a star. Therefore, the improve- present invention are possible in light of the above teachment in detection limit scales as the ratio of the efficiencies ings. It is therefore to be understood that within the scope of

tion are equal. Plain white indicates very bright intensities. for the metal backed and standard respectively, $\epsilon_{MB}/\epsilon_{STD}$,

in wavelength.
The diffracted peaks (n,m) with n changing are due to the s correlated double sampler readout electronics. The detailed correlated double sampler readout electronics. The detailed structure making the parallel shift registers. This is the schematics of the basic integrator and circuitry for introduc-
strongest diffraction. A weaker diffraction pattern with m ing dead time are not shown.

shift registers. Perhaps, this is caused by the structure that without any photoelectrons, can be minimized by sampling defines the channel in which charge is confined. 10 or filtering the signal suitably. The problem is t 10 or filtering the signal suitably. The problem is to adjust the The intensity of the $(1,0)$ (62) and $(-1,0)$ (64) peaks, as filter to the noise power spectrum of the on-chip amplifier. shown in FIG. 6, is complex. The diffraction cannot be An added constraint is that the time allowed for sampling is detected at λ <700 nm. The reason is that the diffraction fixed, and it is essential to take the differ fixed, and it is essential to take the difference between the arises from the light that reaches the back. There are local signal with and without the detected photoelectrons to avoid maxima at 780 nm and 920 nm and a minimum at 840 nm. 15 a large "capacitance noise," f^{-1} . For a noise power spectrum These are probably due to the interference caused by the step that is proportional to f^1 , the readout noise is independent of structure of the polysilicon electrodes. The (1,O) order is the sample frequency *s.* On the other hand, for a noise power much brighter than the $(-1, 0)$ order.
There is a substantial amount of light in the diffracted is proportional to $s^{1/2}$.

 f^{-1} , and the readout noise was independent of sampling rate; amplifier on a typical CCD, the f-1 noise and white noise are tion spectrophotometer. comparable at 9 KHz. For the correlated double sampler, In a standard CCD, a substantial fraction of the light at 25 which is described below, the readout noise with these ng wavelengths $(\lambda > 750 \text{ nm})$ passes through the silicon amplifiers can be improved substantially by slow

Additional improvements are possible with further reduction The improvement in efficiency of the metal-backed **CCD** *3s* dominates, the extra readout time is a small fractional by up to a factor of 2 (depending on the sky background and

(92), and the integrator capacitor **(98)** is discharged **(104)**. 2) capacitor with a negative sign for a fixed time by connecting The results indicate that there is no difference between a to the negative sign input **(102).** 3) After the charge packet via analog to digital converter **(110)**. Thus, the integrator cally used to set the bandwidth of the standard correlated

the appended claims the invention may be practiced otherwise than as specifically described.

1. A charge-coupled device comprising a substrate having reflecting layer is aluminum. a front side and a back side, a sensing layer, a metal carrier 5 **8,** Asystem comprising a charge-coupled device having a layer, and a reflecting layer wherein said reflecting layer is substrate having a front side and a back side, sensing layer, positioned between the sensing layer and the front side of the a metal carrier layer, and a refle substrate and said metal carrier layer is positioned on the reflecting layer is positioned between the sensing layer and back side of the substrate.

4. The device of claim 1 wherein the sensing layer is

5. The device of claim 1 further comprising an insulating coupled device, say is readout capacitor. layer positioned between the sensing layer and the reflecting layer.
 9. The system of claim **8** wherein the correlated double

layer.

substrate, the sensing layer is comprised of silicon, the 20 reflecting layer is comprised of metal and further comprising a silicon insulating layer positioned between the sensing

layer and the reflecting layer, the reflecting layer being positioned on the front side of the glass substrate.

What is claimed is: **7.** The device of claim **6** wherein the metal of the

a metal carrier layer, and a reflecting layer wherein said EX side of the substrate.

2. The device of claim 1 wherein the reflecting layer is 10^{10} and the substantial fields whether the substrate of the subs comprised of metal.
 3. The device of claim 1 wherein the substrate is com-

prised of glass.
 3. The device of claim 1 wherein the substrate is com-

prised of glass.
 3. The device of claim 1 wherein the substrate The device of claim 1 where the sensing layer is
comprised to sampler that controls a readout rate of the charge-

S The device of claim 1 further comprising an insulating

S The device of claim 1 further comprising an ins positioned on the back side of the substrate; said charge-

6. The device of claim 1 wherein the substrate is a glass sampler has an integrator having positive and negative sign
between the sensing layer is comprised of silicon, the 20 inputs and an analog to digital converter.