The Location of the CO₂ Fundamental in Clathrate Hydrates and its Application to Infrared Spectra of Icy Solar System Objects S. A. Sandford, R. M. E. Mastrapa, M. P. Bernstein, D. P. Cruikshank NASA Ames Research Center

 CO_2 is present on the surface of many Solar System objects, but not always as a segregated, pure ice. In pure CO_2 -ice, the fundamental absorption is located near 4.268 μ m (2343.3 wavenumbers). However, on several objects, the CO_2 fundamental is shifted to higher frequency, (Table 1). This shift may be produced by CO_2 gas trapped in another material, or adsorbed onto minerals. We have seen that a mixture of H₂O, CH₃OH, and CO₂ forms a type II clathrate when heated to 125 K and produces a CO_2 fundamental near 4.26 μ m (Blake, et al. 1991). The exact location of the feature is strongly dependent on the initial ratio of the three components, (Table 1).

We are currently exploring various starting ratios relevant to the Solar System to determine the minimum amount of CH_3OH needed to convert all of the CO_2 to the clathrate, i.e. eliminate the splitting of the CO_2 fundamental. We are testing the stability of the clathrate to thermal processing and UV photolysis, and documenting the changes seen in the spectra in the wavelength range from 1-5 μ m. We acknowledge financial support from the Origins of Solar Systems Program, the Planetary Geology and Geophysics and the NASA Postdoctoral Program.

Object	Lab Sample	Shift of CO_2 From 2343.3 cm ⁻¹
	H ₂ O:CO ₂	-4.3,+15.7*
Ganymede		+5.8
Callisto		+5.2
Phoebe		+3.7
Iapetus		+3.7
	H ₂ O:CH ₃ OH:CO ₂ (100:2.5:1)	+3.3, -3.3*
	H ₂ O:CH ₃ OH:CO ₂ (100:10:5)	+3.5, -6.1*
·	H ₂ O:CH ₃ OH:CO ₂ (100:50:1)	+2.7

Table 1. Location of CO2 Fundamental in the Solar System and the Laboratory

* Splitting of fundamental absorption.

Blake, D. F., L. Allamandola, et al. (1991). "Clathrate hydrate formation in amorphous cometary ice analogs in vacuo." <u>Science</u> 254: 548-551.