

## TEMPERATURE AND WAVELENGTH DEPENDENCE OF RECOMBINATION PROCESSES IN 1.5 $\mu\text{m}$ InGaAlAs/InP-BASED LASERS

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### Abstract

The improved thermal stability of 1.5 $\mu\text{m}$  InGaAlAs- compared with InGaAs-based lasers is investigated using a combination of low temperature and high pressure techniques. The results indicate that this is due to lower non-radiative Auger recombination in the InGaAlAs devices because of the higher conduction band offset made possible with the InGaAlAs system which results in a lower hole density in the quantum wells at threshold.

### I. Introduction

Semiconductor lasers operating in the 1400-1600nm band are key components of optical fibre based communications systems both as signal and pump sources. InGaAs(P)/InP is traditionally used for the production of devices at these wavelengths. However, due to the strong temperature dependence of the threshold current ( $I_{th}$ ) of these devices there is ongoing activity in searching for an alternative, more thermally stable material. The development of such materials requires a detailed understanding of the thermal factors influencing device characteristics over their operating temperature range. InGaAlAs/InP-based lasers have exhibited improved thermal performance with higher  $T_0$  values compared with InGaAsP/InP devices at 1.3 $\mu\text{m}$  (1). It is of great interest to extend this benefit to longer wavelengths for pump and signal applications. In order to achieve this, it is vital to understand which processes limit the thermal performance of InGaAlAs/InP lasers and how these vary with wavelength and temperature.

In this paper, we describe low temperature and high pressure measurements performed on 1.5 $\mu\text{m}$  InGaAlAs/InP-based devices to determine the important recombination processes and their variation with wavelength and temperature. The results are compared with similar measurements performed on conventional InGaAs/InP-based 1.5 $\mu\text{m}$  lasers. We find that InGaAlAs lasers exhibit ideal radiative behaviour up to a much higher temperature than the InGaAs devices. Furthermore,

we find that  $I_{th}$  in InGaAlAs lasers has a lower dependence on wavelength than in InGaAs devices. These results are consistent with a reduced contribution of Auger recombination to  $I_{th}$  for the InGaAlAs devices at room temperature (RT) compared with InGaAs-based 1.5 $\mu\text{m}$  lasers. Specifically, we find that whilst Auger recombination accounts for 80% of  $I_{th}$  in 1.5 $\mu\text{m}$  InGaAs devices at RT, it accounts for only ~55% of  $I_{th}$  in the 1.5 $\mu\text{m}$  InGaAlAs devices. We further demonstrate that this is a consequence of the increased conduction band offset offered by the InGaAlAs system, which results in a more even balance of electrons and holes in the quantum wells at threshold, thereby reducing Auger processes which depend strongly on the hole density.

### II. Carrier recombination mechanisms

In long wavelength semiconductor lasers for which we can assume high quality material and therefore may neglect defect related recombination,  $I_{th}$  may be written in terms of the carrier density,  $n$ , (for equal concentrations of electrons and holes) as

$$I_{th} = eV(Bn^2 + Cn^3) + I_{leak} \quad (1)$$

where  $e$  is the electronic charge and  $V$  is the active region volume. The  $Bn^2$  term corresponds to the spontaneous, radiative recombination of electrons and holes giving rise to the radiative current,  $I_{rad}$ .  $Cn^3$  is due to Auger recombination whereby the energy of a recombining electron-hole pair promotes a third electron (hole) into the conduction (valence) band, resulting in the Auger

recombination current,  $I_{Aug}$ . Finally  $I_{leak}$  describes the thermal escape of carriers from the quantum wells (QWs) forming a leakage current.

For each of the current paths discussed above, the extent to which they contribute to the threshold current depends strongly on the band gap and temperature. Therefore, in order to determine their relative importance we have undertaken measurements of  $I_{th}$  as a function of temperature and used hydrostatic pressure to vary the band gap. These measurements are described in the following sections.

### III. Samples

The InGaAlAs lasers studied here consist of 6, compressively strained InGaAlAs QWs within unstrained InGaAlAs barriers ( $\lambda_g=1.1\mu\text{m}$ ) and separate confinement layers grown on an InP substrate. For comparison, standard InGaAs devices were measured containing 4, compressively strained QWs within unstrained InGaAsP barriers ( $\lambda_g=1.3\mu\text{m}$ ) and separate confinement layers grown on an InP substrate.

### IV. Temperature Dependence

In Figure 1 we plot the normalised temperature dependence of  $I_{th}$  for (a) InGaAlAs and (b) InGaAs based devices. The measurements were performed under pulsed conditions using a liquid nitrogen cryostat. At low temperatures, for both devices, we observe a linear increase in  $I_{th}$  with temperature consistent with the ideal temperature dependence  $I_{rad}\propto T^2$  (2) as indicated by the dotted line. However, with increasing temperature,  $I_{th}$  departs from the ideal radiative line and increases super-linearly. The temperature at which this occurs is higher in the InGaAlAs lasers than in the InGaAs devices.

The Auger current has a strong temperature dependence given by  $I_{Aug}\propto T^3 \exp(-E_a/kT)$  (2) where  $k$  is the Boltzmann constant and  $E_a$  is the activation energy of the Auger process. The lines in Figure 1 are the fitted  $I_{th}(T)$  assuming the above temperature variations of  $I_{rad}$  and  $I_{Aug}$ . Details of the fitting technique can be found in Ref. (2). From these data we observe that InGaAlAs devices have a

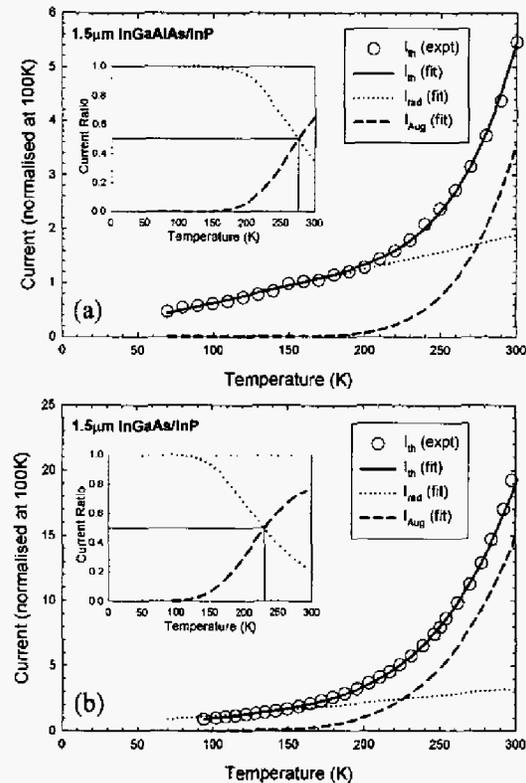


Figure 1 shows the temperature dependence of  $I_{th}$  for the InGaAlAs (a) and InGaAs (b) devices. Also shown are the fitted variations of  $I_{rad}$  and  $I_{Aug}$ . Inset shows the fraction of each current path at  $I_{th}$  as a function of temperature.

considerably lower temperature dependence of  $I_{th}$  which increases by a factor  $\sim 6$  from 100K to room temperature (RT) compared with a factor  $\sim 20$  for the InGaAs devices. Whilst clearly the strong temperature dependence for both devices is consistent with Auger recombination, at this stage, carrier leakage cannot be discounted although in section V we will provide further strong evidence for Auger recombination. In the inset of Figure 1 we plot the ratios  $I_{rad}/I_{th}$  (dotted line) and  $I_{Aug}/I_{th}$  (dashed line) as a function of temperature. At low temperature,  $I_{rad}$  dominates both devices but with increasing temperature,  $I_{Aug}$  becomes dominant. The temperature at which  $I_{rad}$  and  $I_{Aug}$  become equal is considerably higher for the InGaAlAs devices (276K) compared with the InGaAs devices (231K). This gives rise to a lower Auger fraction at RT for the InGaAlAs devices of  $\sim 60\%$  compared with  $\sim 80\%$  for the InGaAs devices.

## V. Pressure (Wavelength) dependence

The application of hydrostatic pressure to semiconductor lasers is a very useful way of varying the emission wavelength because the direct band gap ( $E_g$ ) of semiconductors increases with increasing pressure. High pressure affects each current path in very different ways. For an ideal QW laser,  $I_{rad} \propto E_g^2$  (3), thus one would expect to observe an increase in  $I_{rad}$  with pressure. In contrast,  $I_{Aug} \propto \exp(-E_g/E_o)$  where  $E_o$  is approximately independent of pressure, thus  $I_{Aug}$  decreases strongly with pressure. Finally, due to the fact the pressure coefficients of the band gaps of the QW and barrier layers are approximately equal, the QW-barrier band offset is approximately independent of pressure, and hence,  $I_{leak}$  remains constant as a function of pressure. In Figure 2 we show the normalised pressure variation of  $I_{th}$  for 1.5 $\mu\text{m}$  InGaAs and InGaAlAs devices. For both devices  $I_{th}$  decreases with increasing pressure (decreasing wavelength). This can be understood if Auger recombination contributes strongly to the threshold current in both of the devices. The dotted line shows the expected variation of the radiative current whilst the dashed line corresponds to the theoretically predicted variation of the Auger current for the InGaAs devices. From this we deduce that, under

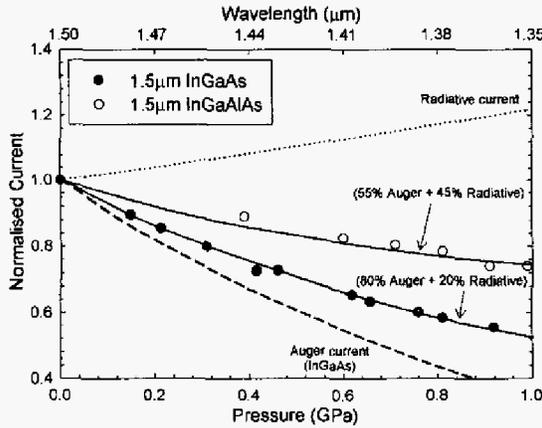


Figure 2 Measured pressure (wavelength) dependence of  $I_{th}$  for InGaAs (solid circles) and InGaAlAs (open circles) based 1.5 $\mu\text{m}$  lasers. These measurements indicate a smaller contribution from Auger recombination to  $I_{th}$  in the InGaAlAs devices.

ambient conditions,  $I_{th}$  consists of 80% $I_{Aug}$ +20% $I_{rad}$  for the InGaAs devices and, assuming the same pressure dependence of  $I_{Aug}$  for the InGaAlAs devices, we find that  $I_{th}$  is composed of 55% $I_{Aug}$ +45% $I_{rad}$ . These values are in very good agreement with those obtained from the temperature dependence measurements in section IV.

These results indicate that non-radiative Auger recombination is important in both InGaAs and InGaAlAs devices but that it plays a smaller role in the InGaAlAs devices.

## VI. Effect of Conduction Band Offset

It is of considerable interest to understand the physical origin of the smaller role of Auger recombination in InGaAlAs devices. In Figure 3 we plot the band alignments for the InGaAs and InGaAlAs QW structures used in the experimental studies. The InGaAlAs material offers a significantly larger band offset for electrons ( $\Delta E_c=142\text{meV}$ ) in the conduction band QW which would be beneficial for the suppression of electron leakage. However, measurements performed on 1.3 $\mu\text{m}$  InGaAlAs devices, which have a lower  $\Delta E_c$  of 109meV, show significantly *less* temperature sensitivity and *reduced* non-radiative current at threshold when compared with the 1.5 $\mu\text{m}$  InGaAlAs devices such that, at RT,  $I_{rad}$  forms  $\sim 70\%$  of  $I_{th}$  (c.f.  $\sim 40\%$  for 1.5 $\mu\text{m}$  InGaAlAs). Along with the pressure dependence data this suggests that the *operating wavelength* rather than the band offset is of most importance to the non-radiative current. This is consistent with Auger recombination dominating the non-radiative current since it is a strongly wavelength dependent process (3).

With reference to equation (1), the Auger current depends upon *both* the Auger coefficient,  $C$ , and the cubed carrier density,  $n^3$ . The  $n^3$  dependence of Auger recombination is, in fact, a generalisation where it is assumed that the QW electron density,  $n$ , is equal to the QW hole density,  $p$ .

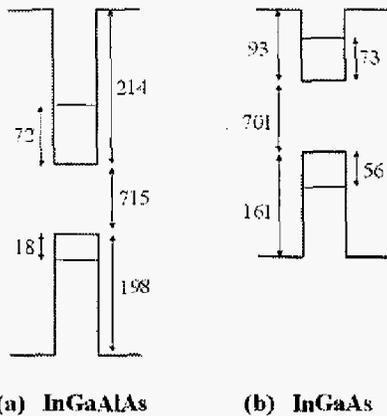


Figure 3 Calculated band alignments at RT for the (a) 1.5 $\mu\text{m}$  InGaAlAs and (b) 1.5 $\mu\text{m}$  InGaAs lasers. Energies are in meV (not to scale).

More correctly, for Auger processes which lead to the excitation of an electron, the corresponding Auger current,  $I_{Aug}^e = eVC^e n^2 p$ . Similarly for hole excitation,  $I_{Aug}^h = eVC^h n p^2$  where  $C^e$  and  $C^h$  are the electron- and hole- Auger coefficients respectively. Previously we demonstrated that for InGaAs-based devices, Auger hole excitation dominates (4) and therefore  $I_{th}$  is most strongly dependent on  $p$ . Thus, minimising  $p$  will reduce the Auger current. In Figure 4, for the structures shown above, we plot the ratio of hole to electron density,  $p/n$ , within the QW as a function of material gain. The calculations were based upon an  $8 \times 8$   $k \cdot p$  theory accounting for the effects of strain together with a self-consistent solution of Schrödinger's and Poisson's equations, accounting for carrier spill-over into the barrier layers. Further details may be found in Ref. (2). These calculations show a striking difference in the electron and hole densities as a function of gain for the two materials. Over this range of material gain,  $n=p$  for the InGaAlAs QWs, whilst  $p>n$  for the InGaAs QWs. Furthermore, the *absolute* value of  $p$  is consistently higher for the InGaAs QWs, where, for a material gain of  $2500\text{cm}^{-1}$  for example ( $\approx$  threshold gain for the devices considered here),  $p(\text{InGaAs})=1.13p(\text{InGaAlAs})$ . This will have a profound effect on the hot hole Auger process which, assuming the same value of  $C$ , will be a factor of  $(1.13)^2=1.27$  (*i.e.* 27%) greater for the

InGaAs devices compared with the InGaAlAs devices. This is consistent with the lower non-radiative current contribution in the InGaAlAs devices and can help to explain the improved thermal characteristics of the InGaAlAs-based 1.5 $\mu\text{m}$  devices.

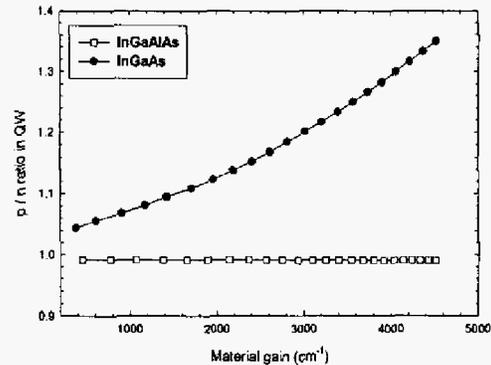


Figure 4 Calculated variation of the QW hole to electron ratio as a function of gain for 1.5 $\mu\text{m}$  InGaAlAs and InGaAs single QWs (at RT).

## VII. Conclusions

In summary, using a combination of high pressure and low temperature techniques we find that the improved thermal stability of 1.5 $\mu\text{m}$  InGaAlAs-based devices is due to a reduced Auger recombination current compared with InGaAs-based devices. This is a direct consequence of the higher conduction band offset afforded by the InGaAlAs material system which suppresses Auger processes which result in the production of hot holes.

## Acknowledgements

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