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# Paleomagnetic field reconstruction from mixtures of titanomagnetites

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## Abstract

Stepwise thermal demagnetization and alternating field (AF) demagnetization are commonly used in paleomagnetic studies to isolate remanent magnetic components of different origins. The magnetically hardest, i.e. highest unblocking temperature/peak field component, is often interpreted as the primary magnetization and magnetically softer components as subsequent remagnetizations due to geological events posterior to the formation of the rock, such as reheating or formation of new magnetic minerals. The correct interpretation of the sequence of the geological events such as tectonic rotations from paleomagnetic data often relies on correctly attributing the observed magnetic directions to the remanence carriers and acquisition mechanisms. Using a numerical model to simulate remanence acquisition and stepwise thermal and AF demagnetization experiments, we show that the presence of mixtures of different magnetic minerals, such as magnetite and titanomagnetites of varying titanium-content can have very significant effects on Zijderveld plots. In thermal demagnetization experiments a spurious third component at intermediate temperatures or a continuous curvature may arise from an overlap of the primary remanence with a subsequent thermal or viscous remagnetization carried by small-grained iron-rich magnetite and large-grained titanium-rich titanomagnetite. AF demagnetization plots of magnetic mixtures are even more complex: primary and secondary remanences carried by different minerals may appear as either three or four components in Zijderveld plots. During alternating field demagnetization the highest coercivity component is not necessarily equivalent to the primary remanence and does not necessarily correspond to the highest temperature component in an analogous thermal demagnetization experiment, i.e., the primary remanence direction cannot be recovered. The effects are shown to be due to the different responsiveness of magnetite and titanomagnetites towards viscous or thermoviscous remanence acquisition: remanent magnetizations with long acquisition times are more effectively recorded by titanium-poor minerals, while short acquisition times are equally well recorded by titanium-rich minerals. In demagnetization experiments on laboratory

35 timescales, the relative contribution of two minerals to Zijderveld plots differs to the relative con-  
36 tribution of remanence acquisition over geological timescales, leading to overlapping components  
37 in Zijderveld plots. The model was also used to simulate paleointensity (ancient magnetic field  
38 intensity) experiments and it was found that the grain distribution affects the slope of Arai plots,  
39 but is negligible compared to the effect of the cooling rate of NRM acquisition. The simulations  
40 suggest that for slowly cooled rocks a cooling rate correction of up to 1.5 to 1.6 may be required  
41 depending on the mineralogy.

## 42 1 Introduction

43 Paleomagnetic observations continue to provide constraints on some of the most fundamental theories  
44 of the deep Earth structure, the dynamics of near surface processes and the evolution and develop-  
45 ment of the geodynamo (*Tarduno et al.*, 2015; *Biggin et al.*, 2015; *O’Rourke and Stevenson*, 2016).  
46 Reliable interpretation of paleomagnetic data can only be achieved through correct identification of  
47 the natural remanent magnetization (NRM) components and their directions; we are usually, but not  
48 always, interested in the primary remanent magnetization’s intensity and its direction carried by the  
49 magnetic minerals within rocks. Among the most common magnetic minerals occurring in rocks are  
50 both stoichiometric magnetite ( $\text{Fe}_3\text{O}_4$ ) and titanomagnetites  $x$  ( $\text{Fe}_{3-x}\text{Ti}_x\text{O}_4$ ), where titanium atoms  
51 substitute the iron atoms at varying proportions  $x$  (*Dunlop and Özdemir*, 1997). In nature, rocks do  
52 not always contain only a single type of magnetic mineral but may contain mixtures, for example of  
53 titanomagnetites of varying compositions. The grain-sizes of the magnetic mineral have been found to  
54 correlate with the titanium content in oceanic basalts (*Zhou et al.*, 1997, 2000) and the process of exso-  
55 lution can move titanium cations in the crystal lattice of the  $\text{Fe}_{3-x}\text{Ti}_x\text{O}_4$ , accumulating them in some  
56 places and depleting them in others, thereby effectively creating an amalgam of high titanium content  
57 titanomagnetite grains and pure magnetite or low-titanium content titanomagnetite grains (*Dunlop  
58 and Özdemir*, 1997). To correctly interpret paleomagnetic signals of natural rocks, it is important to  
59 understand the effect of such magnetic mineral mixtures on the paleomagnetic recording fidelity. We  
60 developed a numerical model to predict the behavior of titanomagnetite mixtures with respect to three  
61 of the most fundamental paleomagnetic studies: (1) directional analysis in thermal demagnetization  
62 experiments, (2) directional analysis in alternating field (AF) demagnetization experiments, and (3)  
63 Thellier-type paleointensity estimates (*Thellier and Thellier*, 1959).

## 64 2 Model

65 A numerical model has been built, that simulates an assembly  $f(x, V)$  of titanomagnetites of different  
66 titanium content  $x$  and different grain volumes  $V$ . The model is built on *Néel* (1949) theory of single-  
67 domain (SD) magnetic particles. The evolution of normalized magnetic moment  $\mathbf{n}$  (magnetic moment  
68 divided by the spontaneous magnetization) with time is given by the differential equation (*Néel*, 1949)

$$\frac{d\mathbf{n}}{dt} = \frac{\mathbf{n}_{eq} - \mathbf{n}}{\tau}, \quad (1)$$

69 where  $\tau$  is the relaxation time and  $\mathbf{n}_{eq}$  is the value of the normalized magnetic moment in thermody-  
70 namic equilibrium. The relaxation time is given by

$$\frac{1}{\tau} = \frac{1}{\tau_+} + \frac{1}{\tau_-}, \quad (2)$$

71 where

$$\frac{1}{\tau_{\pm}} = \frac{1}{\tau_0} \exp \left\{ -\frac{\mu_0 V H_K(T) M_s(T)}{2kT} \left( 1 \pm \frac{|\mathbf{H}_0|}{H_K} \right)^2 \right\}, \quad (3)$$

72 where  $\tau_0$  is the atomic attempt time, which was set to be  $10^{-10}$  s in the model (*Berndt et al.*, 2015),  
73  $\mu_0$  is the vacuum permeability,  $k$  is the Boltzmann constant and  $H_0$  is the applied magnetic field. The  
74 equilibrium magnetic moment is given by a Maxwell-Boltzmann distribution

$$\mathbf{n}_{eq} = \tanh \left\{ \frac{V \mathbf{H}_0 M_s(T)}{kT} \right\}. \quad (4)$$

75 The spontaneous magnetization at high temperature is modeled using the analytical approximation  
76 (*Dunlop and Özdemir*, 1997)

$$M_s(T) = M_{s0} \sqrt{1 - \frac{T}{T_C}}, \quad (5)$$

77 and the microscopic coercivity  $H_K$  is calculated assuming that shape anisotropy dominates, for which

$$H_K = \Delta N M_s, \quad (6)$$

78 using a shape anisotropy factor  $\Delta N$ . For titanomagnetites shape anisotropy and magnetocrystalline  
79 anisotropy are relatively weaker than for magnetite, but magnetostriction increases (*Dunlop and*  
80 *Özdemir*, 2007). For simplicity, however, we assume strongly elongated grains with dominant shape  
81 anisotropy for all titanium contents with a common value of  $\Delta N = 0.5$  for all grains and alignment of  
82 their elongation axis with the field.

83 The titanium content is assumed to have two effects: (1) it lowers the Curie temperature  $T_C$ ,  
84 and (2) it reduces the room-temperature spontaneous magnetization  $M_{s0}$ . The Curie temperature is  
85 modeled by the quadratic equation

$$T_C = T_{C, TM0} - ax^2 - bx, \quad (7)$$

86 where the coefficients  $a = 280$  and  $b = 500$  were found from a least-squares fit to the data published  
87 by *Dunlop and Özdemir* (1997), and  $T_{C, TM0} = 580^\circ\text{C}$  is the Curie temperature of magnetite. The  
88 spontaneous magnetization at room temperature is modeled by a linear relationship (*Stephenson*, 1969;  
89 *Dunlop and Özdemir*, 1997)

$$M_{s0} = M_{s0, TM0} - \frac{1}{0.6} (M_{s0, TM0} - M_{s0, TM60}) x, \quad (8)$$

90 where  $M_{s0, TM0} = 480 \text{ kAm}^2$  is the spontaneous magnetization of magnetite and  $M_{s0, TM60} = 125 \text{ kAm}^2$   
91 is the spontaneous magnetization of TM60 titanomagnetite (*Özdemir and O'Reilly*, 1981).

## 2.1 VRM and TRM acquisition

The grain distribution is discretized by a matrix of 1000 volumes  $V$  between  $10^{-24}$  and  $10^{-21}$  m<sup>3</sup> (being equal to cubes of 10 to 100 nm), separated on a logarithmic scale, and 100 equally spaced Curie temperatures  $T_C$  between 0°C and 580°C (corresponding to various different titanium compositions  $x$  according to eq. (7), for clarity we quote  $T_C$  rather than  $x$  values in the diagrams). The magnetization of each of these grains can take on any magnetization value representing a large number of grains, and not just  $\pm 1$ , as for a single SD grain.

For viscous remanent magnetization (VRM) acquisition at a temperature  $T_A$ , the equilibrium magnetizations  $\mathbf{n}_{eq}$  (eq. (4)) and the relaxation times (eq. (2) and (3)) are calculated for each grain set ( $V, T_C$ ) and the resulting new magnetization state  $\mathbf{n}_{new}$  is calculated from eq. (1). Thermoremanent magnetization (TRM) acquisitions are simulated by repeatedly following this procedure for 2000 temperature steps  $T_i$ , decreasing by small temperature steps  $\Delta T$  until room temperature is reached. Various scenarios of different combinations of acquired VRMs and TRMs at different times and temperatures were run. Generally, linear cooling was used, but for one case Newtonian cooling was used for a paleointensity scenario, as cooling rates are known to have a significant effect on paleointensities (Dodson and McClelland-Brown, 1980; Halgedahl et al., 1980).

## 2.2 Thermal demagnetization

Step-wise thermal demagnetization was simulated by repeatedly applying VRMs at successively higher temperatures in zero field. This simulates the time at which the sample is kept at a high temperature in a thermal demagnetizer. After each step, the total remanent magnetization vector is calculated, which is the sum the magnetization vectors  $\mathbf{n}(V, T_C)$  of all different grain sets, and the total spontaneous magnetization is calculated by summing the product of  $M_s$ , volume  $V$  and the grain distribution  $f(V, T_C)$ .

## 2.3 AF demagnetization

AF demagnetization is modeled based on the simplified assumption that all grains with a coercivity  $H_C$  less than the maximum amplitude  $\tilde{H}$  of the alternating field get demagnetized. The coercivity is given by

$$H_C = H_K - H_q, \quad (9)$$

where  $H_K$  is given by eq. (6) and  $H_q$  is the thermal fluctuation field given by (Néel, 1949)

$$H_q = \sqrt{\frac{2H_K kT \ln(t/\tau_0)}{\mu_0 V M_s(T, T_C)}}. \quad (10)$$

Using eq. (6), and approximating the time  $t$  as half the inverse of the frequency  $\tilde{f}$  of the AF field (Worm, 1998), this simplifies to

$$H_q = \sqrt{\frac{2\Delta N kT \ln(1/2\tilde{f}\tau_0)}{\mu_0 V}}. \quad (11)$$

122 The amplitude  $\tilde{H}$  is successively increased and at each step the remaining total magnetization is  
123 calculated by summing the magnetization vectors  $\mathbf{n}(V, T_C)$ .

## 124 2.4 Paleointensity

125 A series of Thellier-type paleointensity experiments (*Thellier and Thellier, 1959*) were simulated fol-  
126 lowing the methodology of *Coe (1967)*. First a TRM acquisition was simulated using either linear  
127 or Newtonian cooling. Then, Arai plots (*Nagata et al., 1963*) were produced by simulating demagne-  
128 tization steps to temperatures  $T_i$  by calculating the viscous decay in zero field during heating to  $T_i$   
129 at 1 K/s, holding the temperature for 10 min and cooling back to room temperature at 1 K/s, and  
130 calculating the remaining NRM. Each step was followed by the simulation of a heating in zero field at  
131 1 K/s, followed by a VRM acquisition in a  $30 \mu\text{T}$  field for 10 min at  $T_i$ , representing the hold time in  
132 the furnace in field in a Thellier-type experiment, followed by a TRM acquisition on cooling from  $T_i$   
133 to room temperature at 1 K/s, representing the in-field-cooling of a Thellier-type experiment.

## 134 3 Scenarios and grain distributions

135 The model was used to simulate the magnetic behavior of a number of different grain distributions for  
136 a number of different remanence acquisition scenarios.

### 137 3.1 Remanence acquisition scenarios

138 To investigate the effect of mixtures of titanomagnetites on vector demagnetization plots, the following  
139 three extreme scenarios of remanence acquisition were used:

- 140 1. A primary full TRM acquisition over 1 hour was simulated, followed by a perpendicular pTRM  
141 acquired at  $100^\circ\text{C}$  over 1 hour. This acquisition time represents fast-cooling submarine lavas  
142 (*Bowles et al., 2005*).
- 143 2. A primary full TRM acquisition over 100 ka was simulated, followed by a perpendicular pTRM  
144 acquired at  $100^\circ\text{C}$  over 100 ka. This timescale is typical of slowly cooling intrusive rocks (*Mux-  
145 worthy et al., 2013*).
- 146 3. A primary full TRM acquisition over 100 ka is simulated, followed by a perpendicular VRM  
147 acquired at room temperature ( $20^\circ\text{C}$ ) over 100 ka.

148 The ambient magnetic field was set to be  $H_0 = 30 \mu\text{T}$ . For the step-wise thermal demagnetization  
149 experiments, a hold time of 10 min was used. It was expected from theory that for pure magnetite, all  
150 of these scenarios yield two perpendicular magnetic components in thermal demagnetization *Zijderveld*  
151 (1967) plots. They are expected to only differ in the unblocking temperatures of the remagnetizations:  
152 According to *Pullaiah et al. (1975)*, scenario 1 should have an unblocking temperature close to  $100^\circ\text{C}$ ,  
153 as the timescale of the thermal demagnetization is similar to the timescale of acquisition, scenario 2  
154 should have an unblocking temperature of  $206^\circ\text{C}$  (linear cooling over 100 ka is equivalent to 5300 yr at  
155 constant temperature according to *York 1978a,b*) and scenario 3 should unblock at  $135^\circ\text{C}$ . For the AF

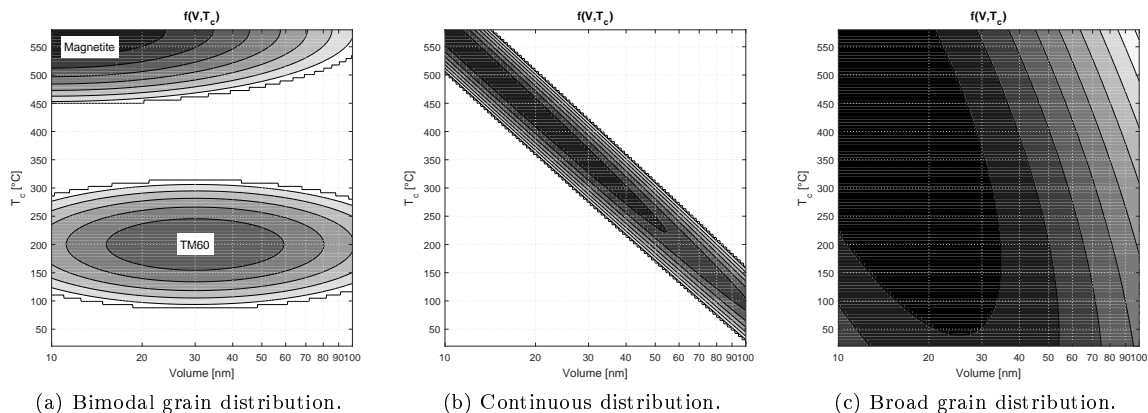


Figure 1: Plot of the grain distributions.

156 demagnetization experiments, a frequency of  $\tilde{f} = 50$  Hz was used. Throughout the following treatment,  
 157 the primary full TRM will be referred to as the characteristic remanent magnetization (ChRM), while  
 158 the secondary remagnetization will be referred to as either partial TRM (pTRM) or VRM, respectively.

159 For the paleointensity experiments, four scenarios were simulated:

- 160 1. A full TRM acquired by linear cooling over 100 ka in a  $30 \mu\text{T}$  field.
- 161 2. A full TRM acquired by linear cooling over 1 h in a  $30 \mu\text{T}$  field.
- 162 3. A full TRM acquired by linear cooling over 1 h in a  $30 \mu\text{T}$  field, followed by a VRM in the same  
 163 direction acquired over 100 ka in the same field. This scenario is meant to test if viscous overprints  
 164 of rocks formed during the Brunhes chron have an effect on paleointensity determinations.
- 165 4. A full TRM acquired by Newtonian cooling over 100 ka in a  $30 \mu\text{T}$  field. In order to avoid the  
 166 cooling process to take infinitely long, the ambient temperature was set to  $15^\circ\text{C}$  and the cooling  
 167 rate was chosen such that a target temperature of  $20^\circ\text{C}$  is reached after 100 ka.

### 168 3.2 Grain distributions

169 The model has been run with a series of different grain size and composition distributions. The first dis-  
 170 tribution investigated is a bimodal grain distribution with a magnetite ( $T_C = 580^\circ\text{C}$ ) peak at  $(10 \text{ nm})^3$ ,  
 171 and a secondary peak around  $(30 \text{ nm})^3$  large titanomagnetite TM60 with a Curie temperature of  $200^\circ\text{C}$   
 172 (Fig. 1a). Note the smallest pure magnetite grains are superparamagnetic at room temperature and  
 173 the remanence is due to only the grains larger than the peak value of 10 nm. The TM60 amounts to  
 174 approximately 22% of the volume of the magnetic material, but only to 8% of the magnetic intensity  
 175 due to the lower  $M_{s0}$  of TM60 and only 2.3% of the total number of grains due to the larger grain size.  
 176 The Curie temperature of TM60 can be seen in the simulated  $M_s(T)$  curve in Fig. 2, with a value of  
 177  $220^\circ\text{C}$  determined by the maximum second derivative method (*Ade-Hall et al., 1965*).

178 The second distribution was a continuous grain distribution with pure magnetite and titanomag-  
 179 netite with a Curie temperature close to room temperature as its end-members, with the mean grain

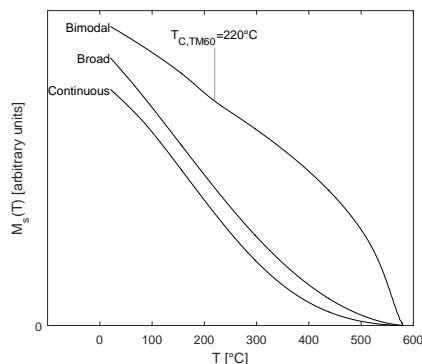


Figure 2: Simulated spontaneous magnetization as a function of temperature  $M_s(T)$  for the three distributions shown in Fig. 1.

180 size increasing with titanium content (Fig. 1b). This distribution was chosen, because distributions  
 181 with larger titanium-rich and smaller titanium-poor titanomagnetite grains have been observed in na-  
 182 ture (Zhou *et al.*, 1997, 2000). Such distributions lead to  $M_s(T)$  curves that decrease steeply at low  
 183 temperatures before slowly leveling off at high temperatures before vanishing at the Curie temperature  
 184 of magnetite (Fig. 2). Simple methods such as the maximum second derivative method (Ade-Hall *et al.*,  
 185 1965) do not allow to obtain much insight into the mineralogy in this case: doing so would yield an  
 186 intermediate value between the Curie temperatures of the end-members magnetite and low-titanium  
 187 titanomagnetite and would miss the fact that the  $M_s$  curve is due to a mixture of various minerals  
 188 with a wide range of Curie temperatures.

189 The final case was a broad grain distribution that includes grains of all sizes and all titanium con-  
 190 tents, with a slight correlation between grain-size and titanium-content, and log-normally distributed  
 191 grain-volumes (Fig. 1c). The spontaneous magnetization curve shows a similarly sharp decay as the  
 192 continuous magnetite-titanomagnetite distribution at low temperatures, leveling off at high tempera-  
 193 tures.

194 For the Thellier-type paleointensity experiments, additionally a pure magnetite grain distribution  
 195 was investigated that equaled the distribution of the grains with  $T_C = 580^\circ\text{C}$  of the bimodal grain  
 196 distribution above.

## 197 4 Results

### 198 4.1 Grain distribution 1: Bimodal distribution

199 The stepwise thermal demagnetization plot of scenario 1 (Fig. 3a), reconstructs the directions of the  
 200 two magnetic components as expected. Scenario 2, however, shows a demagnetization plot that could  
 201 incorrectly be interpreted as having three magnetic components: one unblocking around  $130\text{--}140^\circ\text{C}$ ,  
 202 with the expected direction of the pTRM, one intermediate direction unblocking around  $200\text{--}210^\circ\text{C}$   
 203 and the original ChRM. The middle component is an artifact of two different magnetic minerals,  
 204 however: A pTRM acquired at  $100^\circ\text{C}$  during cooling over 100ka should be removed at  $206^\circ\text{C}$  in a



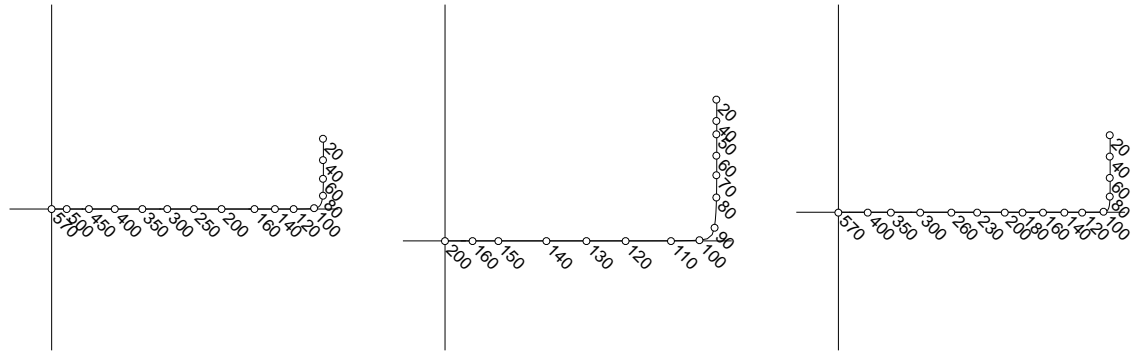
205 10 min demagnetization experiment, whereas the same pTRM acquired by titanomagnetite with a  
206 Curie temperature of 200°C should be demagnetized at 133°C (*York, 1978a,b; Dodson and McClelland-*  
207 *Brown, 1980*): The first (lowest temperature) apparent direction below 133°C in the demagnetization  
208 plot corresponds to the demagnetization of the pTRM carried by the titanomagnetite, the second  
209 apparent direction up to 206°C corresponds to the simultaneous demagnetization of the pTRM carried  
210 by magnetite and the ChRM carried by the titanomagnetite, and the third apparent direction above  
211 206°C corresponds to the demagnetization the ChRM carried (mostly) by the magnetite. A similar  
212 effect occurs in scenario 3 (Fig. 3g) with a VRM acquired over 100ka at room temperature: up to  
213 80°C the direction of the pTRM is observed, above that a curvature up to 140°C is seen and at  
214 higher temperatures the ChRM is recovered. In a 10 min demagnetization experiment, the unblocking  
215 temperature of such a VRM is 139°C for magnetite and 79°C for the titanomagnetite (*Pullaiah et al.,*  
216 *1975*). The curvature between these two temperatures is due to the overlap of the VRM carried by  
217 magnetite and the ChRM carried by titanomagnetite.

218 The AF demagnetization plots (Fig. 4a, 4d and 4g) all show four apparent components: the direction  
219 of the overprint is seen at low field  $\tilde{H}$ , after that a curvature approaching the ChRM direction is visible,  
220 followed by another section of the overprint's direction, and at highest fields the ChRM direction is  
221 again observed.

## 222 4.2 Grain distribution 2: Continuous distribution

223 As in the case of a bimodal distribution, both magnetic components are accurately recovered for  
224 scenario 1 (Fig. 3b) in stepwise thermal demagnetization. Scenario 2 has three apparent directions, the  
225 first of which unblocks around 115°C, and the second of which unblocks around 180°C (Fig. 3e). These  
226 two unblocking temperatures of the pTRM correspond to titanomagnetite with Curie temperatures  
227 of 140°C and 360°C, respectively, for a demagnetization time of 10 min. Neither of these are the  
228 end-members of the titanomagnetite grain distribution (Fig. 2). While the first inflection point at  
229 115°C is relatively clear, the second one around 180°C is curved, slowly approaching the final ChRM  
230 direction. Scenario 3 shows a similar trend with a lower blocking temperature of 45°C and a higher  
231 one of 110°C, which for the VRM of 100 ka, corresponds to titanomagnetite of Curie temperatures of  
232 80°C and 370°C (Fig. 3h). A notable difference to the bimodal grain distribution is that the intensity  
233 of the VRM is significantly weaker, although the intensity of the pTRM is of a similar order. Hence  
234 this grain distribution is less responsive to VRM acquisition than the bimodal distribution (Fig. 3g).

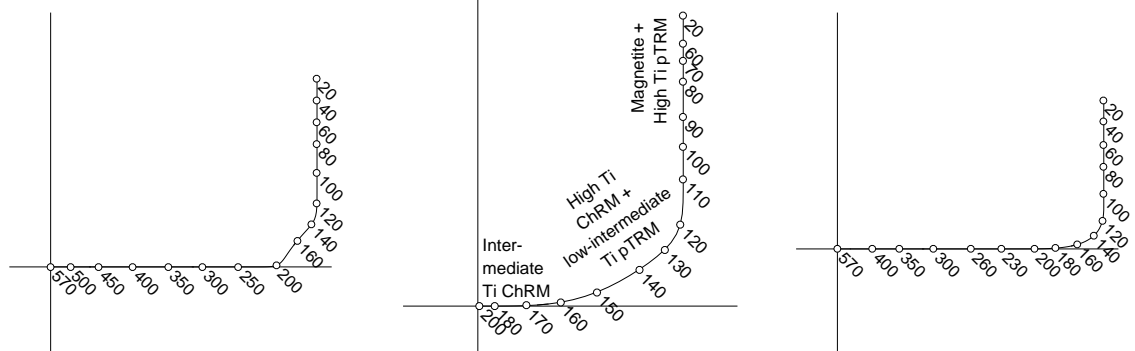
235 The AF demagnetization plot of scenarios 1 and 3 (Fig. 4b and 4h) show similar trends to the ones  
236 of the bimodal grain distribution (Fig. 4a and 4g), but with different intensities and demagnetizing  $\tilde{H}$   
237 fields: at low fields the pTRM/VRM direction is seen, but for higher fields an "S"-shape is observed,  
238 starting at approximately the ChRM direction, then bending into an intermediate direction and then  
239 bending back into the ChRM direction. The ChRM-direction can only be approximately isolated  
240 at the highest fields (>55 mT): the curvature due to overlap with the pTRM/VRM is small at the  
241 highest fields, it does not, however, necessarily vanish, such that the obtained ChRM direction may  
242 be imperfect. Compared to the bimodal distribution, the S-shape is greatly reduced in intensity in the  
243 continuous distribution and is dominated by the ChRM component. Additionally, like in the thermal  
244 demagnetization case (Fig. 3b and 3h) the VRM is strongly suppressed compared to the pTRM, almost



(a) Scenario 1 (100°C, 1 h pTRM)  
Bimodal distribution.

(b) Scenario 1 (100°C, 1 h pTRM)  
Continuous distribution.

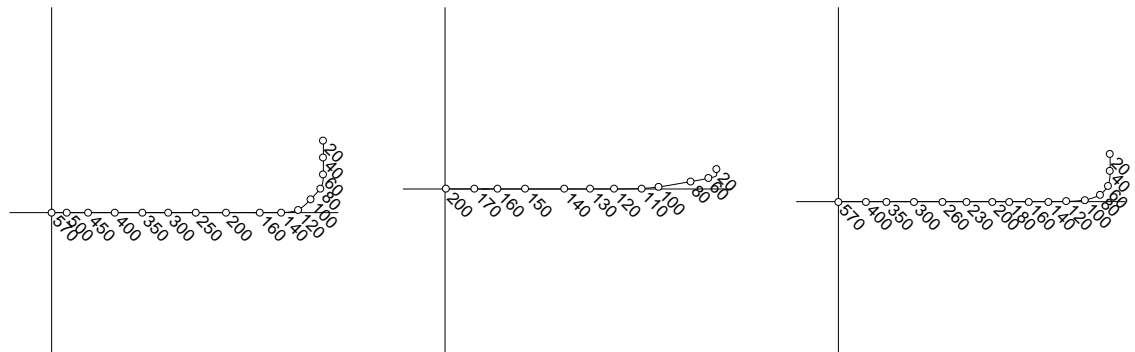
(c) Scenario 1 (100°C, 1 h pTRM)  
Broad distribution.



(d) Scenario 2 (100°C, 100 ka pTRM)  
Bimodal distribution.

(e) Scenario 2 (100°C, 100 ka pTRM)  
Continuous distribution.

(f) Scenario 2 (100°C, 100 ka pTRM)  
Broad distribution.



(g) Scenario 3 (100ka, 20°C VRM)  
Bimodal distribution.

(h) Scenario 3 (100ka, 20°C VRM)  
Continuous distribution.

(i) Scenario 3 (100ka, 20°C VRM)  
Broad distribution.

Figure 3: Vector demagnetization (*Zijderveld, 1967*) plots for stepwise thermal demagnetization. Temperatures are given in °C. Fig. 3e is annotated to highlight behaviors occurring in all scenario 2 simulations.

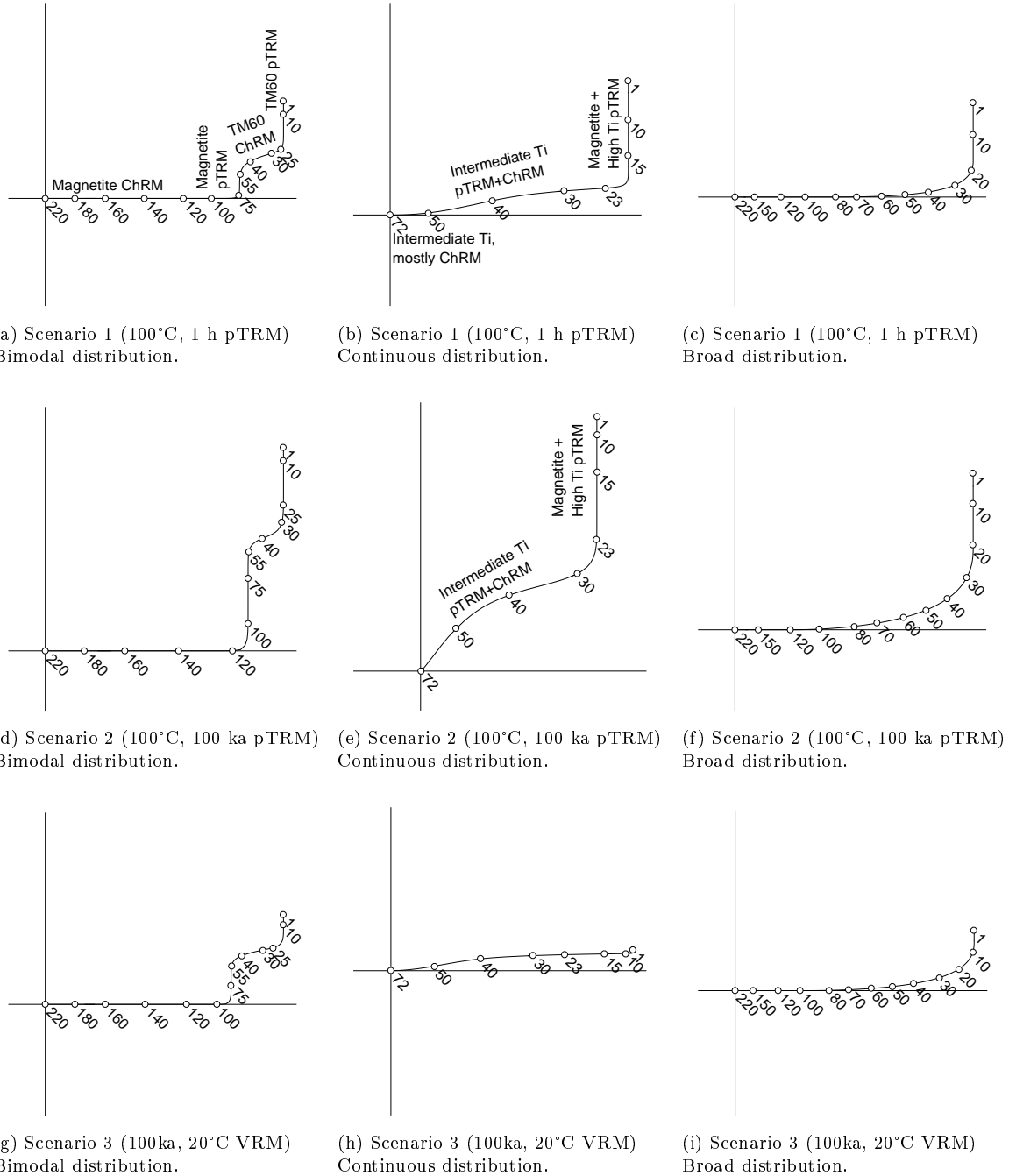


Figure 4: Vector demagnetization plots for AF demagnetization. AF peak fields  $\tilde{H}$  are given in mT. Selected plots are annotated to highlight behaviors.

245 disappearing completely at low coercivities. Therefore, in this particular grain distribution (Fig. 4b  
246 and 4h), the ChRM-direction would be better isolated at lower fields (20–30 mT), contrary to the  
247 bimodal distribution (Fig. 4a and 4g), where low fields (30–40 mT) showed an intermediate direction  
248 between the ChRM and pTRM/VRM directions. Scenario 2 (Fig. 4e) differs from the others in that  
249 the ChRM cannot be recovered: after the VRM is isolated at coercivities below 23 mT, an S-shaped curve  
250 begins that approximates the ChRM direction between  $\sim 30$ –40 mT, but turns into an intermediate  
251 direction at higher coercivities. No part of the diagram completely isolates the ChRM direction.  
252 Moreover, at the highest coercivities, the observed direction differs more from the ChRM direction  
253 than in the 30–40 mT range: the common assumption that the ChRM is best isolated at the highest  
254 demagnetization step is invalid.

### 255 4.3 Grain distribution 3: Broad grain distribution

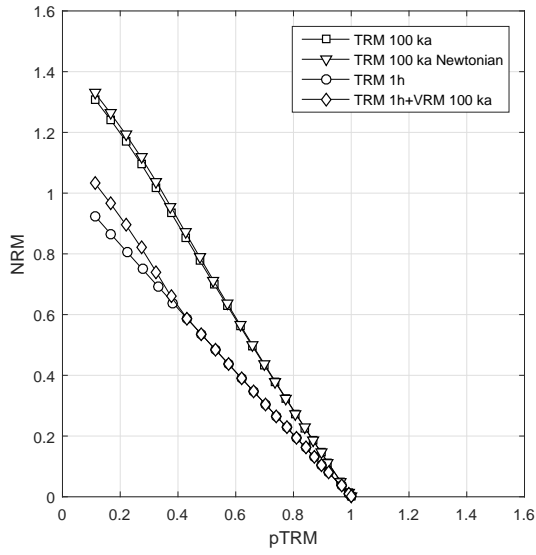
256 The demagnetization plots, both stepwise thermal demagnetization and AF demagnetization, all show  
257 a strong curvature between the ChRM and the pTRM/VRM (except scenario 1 in thermal demagne-  
258 tization, Fig. 3c), contrary to the previous cases, where more than two distinct apparent components  
259 were observed. The curvature appears in a similar temperature range as in the case of the continuous  
260 magnetite-titanomagnetite distribution (Fig. 4), with the curvature lying between 110°C and 190°C for  
261 the pTRM (scenario 2) and between 50°C and 120°C for the VRM (scenario 3). Again, the VRM ap-  
262 pears slightly weaker than the pTRM due to the presence of titanomagnetites that are less responsive  
263 to VRM acquisition.

264 The AF demagnetization plots do not show three to four apparent components as in the previous  
265 cases (Fig. 3 and 4), but rather show a strong curvature between the pTRM/VRM and the ChRM.  
266 For this grain distribution, both the pTRM/VRM and the ChRM directions can be recovered from  
267 both stepwise thermal and AF demagnetization plots.

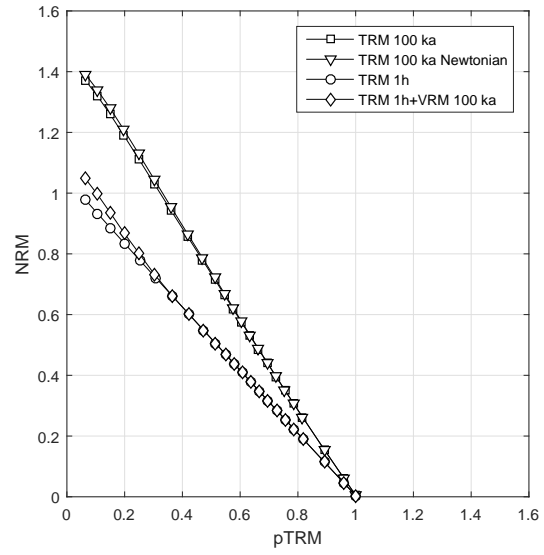
### 268 4.4 Paleointensity experiments

269 Arai plots (*Nagata et al.*, 1963) were calculated for the three grain distributions and for a pure mag-  
270 netite using the four scenarios described in section 3 (Fig. 5). All grain distributions show two types  
271 of behavior: a slow cooling behavior for linear and Newtonian cooling over 100 ka, and a fast cooling  
272 behavior for linear cooling over 1 h (with or without subsequent VRM acquisition). Within these two  
273 categories, Arai plots are almost identical for all samples, with the only exception of the 1 h TRM  
274 followed by a 100 ka VRM for the pure magnetite simulation (and, to a lesser degree, the bimodal  
275 distribution). This observation suggests that for paleointensity experiments, the effect of magnetic  
276 mineralogical mixtures is almost negligible, and the dominant factor impacting the slope of Arai plots  
277 is the cooling rate.

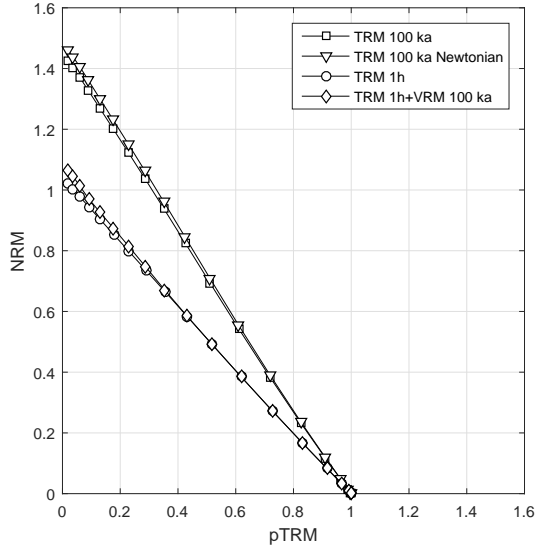
278 Arai plots are almost linear over the whole temperature range, but slight variations of the slope  
279  $dM_{pTRM}/dM_{NRM}$  occur (Fig. 6). The 1 h cooling scenario (circles) shows the most constant slope;  
280 slightly more than unity. A slope of one is expected if the NRM acquisition time equals the pTRM  
281 acquisition times in the Thellier/Coe-type experiment, but an exact comparison of the two timescales  
282 is difficult, as NRM acquisition occurs during cooling, while pTRM acquisition occurs during a hold



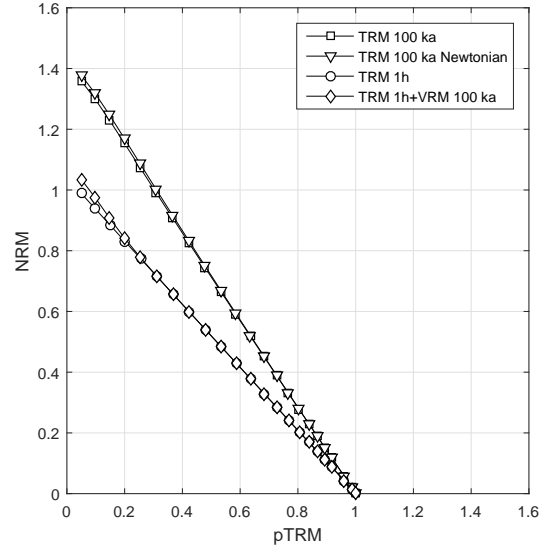
(a) Pure magnetite (model validation).



(b) Case 1: Bimodal distribution.



(c) Case 2: Continuous titanomagnetite distribution.



(d) Case 3: Broad grain distribution.

Figure 5: Simulated Arai plots (*Nagata et al., 1963*) for Thellier/Coe-type paleointensity experiments.

283 time (i.e. VRM acquisition) of 10 min at elevated temperature and the subsequent cooling. The  
284 scenario also shows that the slope tends to increase slightly at higher temperatures depending on the  
285 grain distribution.

286 In contrast, the slow cooling scenarios show a much steeper slope around 1.4–1.6 (Fig. 6). Newtonian  
287 cooling (triangles) tends to further increase slopes at intermediate temperatures compared to linear  
288 cooling (squares), but only marginally. An obvious feature of the slow cooling scenarios is the strong  
289 increase of slope at low temperatures before reaching a peak around 100°C and then slowly decreasing.

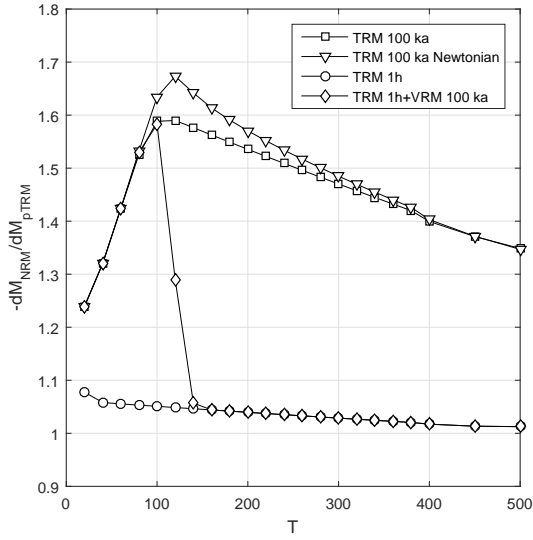
290 The scenario of a fast cooling TRM and a subsequent 100 ka VRM combines features of both the  
291 fast and slow cooling scenarios: at low temperatures, before the VRM is unblocked, the slope equals  
292 that of the slow cooling 100 ka TRM scenarios. At higher temperatures, the slope quickly approaches  
293 that of the fast cooling 1 h TRM scenario. This behavior is expected as the two parts of the Arai plots  
294 show two distinct magnetic components, that only coincide in direction, but not in intensity.

## 295 5 Discussion

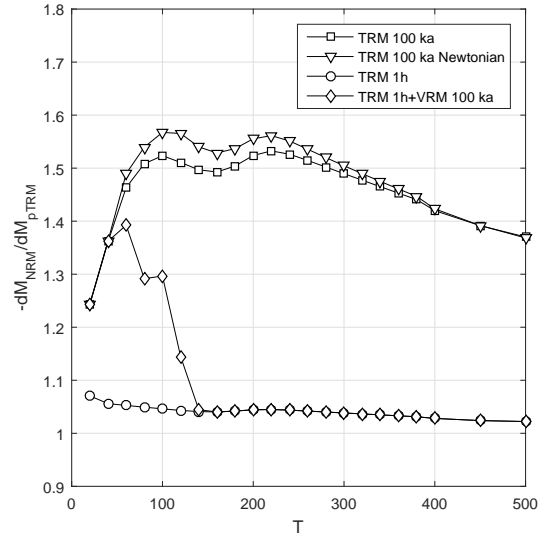
### 296 5.1 Thermal demagnetization

297 With the exception of the thermal demagnetization of the first scenario (fast-cooling TRM acquisition),  
298 all the simulated Zijderveld plots significantly deviate from the expected two-component behavior. This  
299 can be explained by considering the individual sets of grains that carry the remanence and who they  
300 get demagnetized.

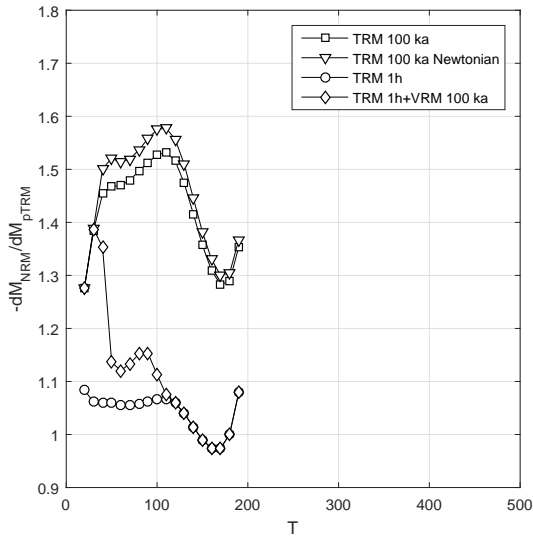
301 The grains carrying the two remanent magnetizations are indicated in Fig. 7a, 7c and 7e for the  
302 bimodal distribution. In all three scenarios, the magnetite distribution as well as the titanomagnetite  
303 both partially carry the ChRM and partially the VRM/pTRM. The line separating the two magnetic  
304 components (dashed line) depends on both the acquisition temperature and the acquisition time; the  
305 effective acquisition time in the case of the pTRMs. All grains to the bottom left of the dashed line  
306 carry the remagnetization, whereas all grains to the top right of the line preserve the ChRM. Such  
307 a line can be calculated for any time and temperature. In general, increasing either the temperature  
308 or the time shifts the line to the top right. When demagnetizing the sample, the solid line is swept  
309 from the bottom left corner of the diagram to the top right corner as the temperature is increased,  
310 demagnetizing the grains below. The remaining remanence is carried by the grains to the top right of  
311 the line and measured after each heating. If both the time and the temperature of the demagnetization  
312 experiment are identical to the acquisition (as in scenario 1, Fig. 7a), both the acquisition and the  
313 demagnetizing lines are identical, but if the timescale of the demagnetization experiment differs to the  
314 acquisition timescale, then the slope of the demagnetizing line will differ to the slope of the acquisition  
315 line (Fig. 7c). While either one, increasing temperature or increasing time shifts the line to the  
316 top right, increasing time does so while tilting it clockwise, whereas increasing temperature does so  
317 while tilting it anti-clockwise. For this reason, the titanomagnetite is more responsive to increases in  
318 temperature, while the magnetite is more responsive to increases in time. As thermal demagnetization  
319 is usually done on a shorter timescale than acquisition (minutes to hours versus days to thousands of  
320 years), the titanomagnetite tends to be demagnetized first. As the demagnetization progresses, the  
321 larger titanomagnetite grains that preserve the ChRM and the smaller magnetite grains that carry the



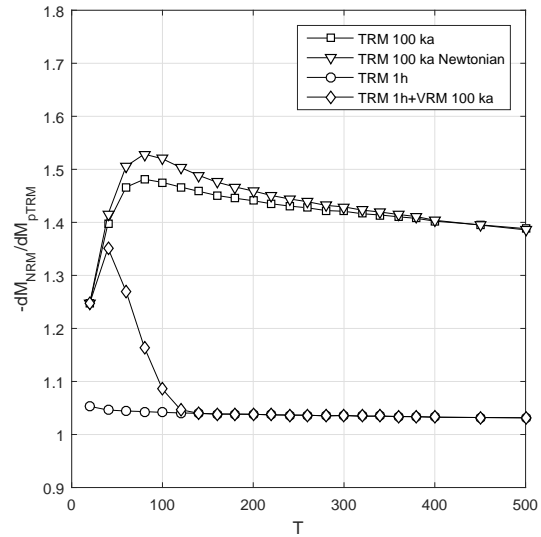
(a) Pure magnetite (model validation).



(b) Case 1: Bimodal distribution.

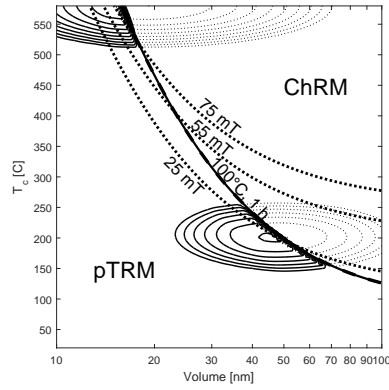


(c) Case 2: Continuous titanomagnetite distribution (this distribution is completely demagnetized at 200°C).

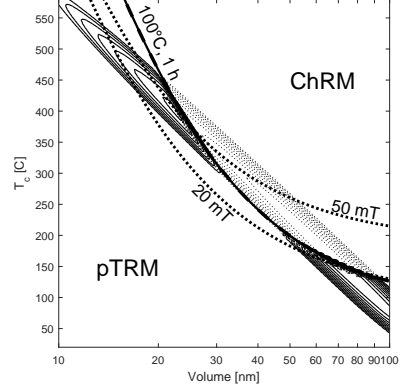


(d) Case 3: Broad grain distribution.

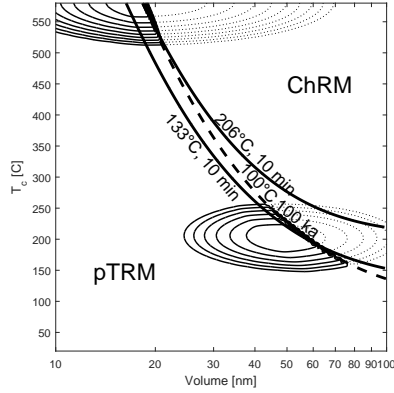
Figure 6: Slope of the Arai plots (*Nagata et al., 1963*) in Fig. 5 as a function of temperature for different distributions and different remanence acquisition scenarios. Plots are normalized by maximum pTRM (normalization by pTRM rather than NRM was done because the pTRM was independent of the remanence acquisition scenario). First data point has non-zero pTRM as it corresponds to a “heating” to room-temperature, i.e., a VRM acquisition over 10 min.



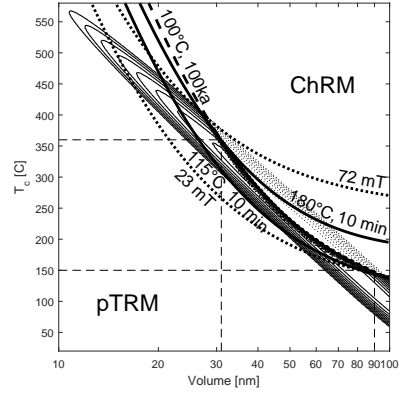
(a) Scenario 1 (pTRM at 100°C, 1 h)



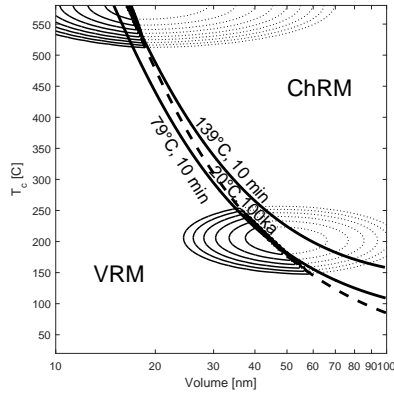
(b) Scenario 1 (pTRM at 100°C, 1 h)



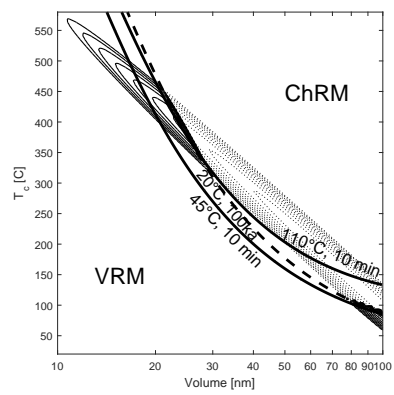
(c) Scenario 2 (pTRM at 100°C, 100 ka)



(d) Scenario 2 (pTRM at 100°C, 100 ka)



(e) Scenario 3 (100ka VRM at 20°C, 100 ka)



(f) Scenario 3 (100ka VRM at 20°C, 100 ka)

Figure 7: Plots showing the magnetic moment  $m_r$  carried by each set of grains (given by their volume  $V$  and titanium content / Curie temperature  $T_C$ ) for a bimodal distribution. The dashed lines indicate the blocking condition for the acquisition time and temperature of the respective scenarios; the black lines indicate the blocking condition for the timescale of the demagnetization experiment; the dotted lines indicate the peak AF fields  $\tilde{H}$  needed to remove the magnetization.



322 VRM/pTRM are demagnetized simultaneously and their components overlap. On further heating the  
323 larger magnetite grains that carry the ChRM are demagnetized.

324 A similar effect occurs in the continuous grain distribution (Fig. 7b, 7d and 7f). Due to the strong  
325 correlation of grain volume with titanium content, the remagnetization (pTRM and VRM) affects two  
326 distinct grain populations: small-grained low-titanium magnetite as well as large-grained high-titanium  
327 content grains. In the first scenario, the same populations are activated during demagnetization as  
328 during acquisition, which is expected for any grain distribution if the acquisition time equals the  
329 demagnetization time. Therefore, the demagnetization plot shows only two components (Fig. 3b).

330 In scenario 2 (Fig. 3e and 7d), up to 115°C both small-grained magnetite and large high-titanium  
331 content grains are demagnetized, both of which carry the pTRM. Between 115°C and 180°C, the  
332 situation is more complex: large titanium-rich grains carrying the ChRM are demagnetized, together  
333 with small-grained low-titanium magnetite carrying the pTRM as well as intermediate grain-sizes of  
334 intermediate of intermediate titanium content carrying the pTRM. Above 180°C, only the ChRM is  
335 left to demagnetize, and is carried only by intermediate sizes with intermediate titanium content. It  
336 can also be seen how the apparent unblocking temperatures of the pTRM, 115°C and 180°C relate  
337 to the Curie temperatures of the titanomagnetite spectrum: the first point where the the (solid)  
338 demagnetization line touches the ChRM, is at  $T_C = 140^\circ\text{C}$  (80 nm), and the last point touching the  
339 pTRM at the second solid line is at  $T_C = 360^\circ\text{C}$  (25 nm). These two points are strongly dependent  
340 on the grain distribution: they lie on the diagonal describing the boundary of the titanomagnetite-  
341 content/grain-volume distribution. Depending on the distribution the two points may occur at different  
342 Curie temperatures and grain-sizes.

343 Compared to scenario 2 (pTRM), scenario 3 (VRM) shows a similar picture in the small-grained,  
344 magnetite-rich half of the diagram. It is observed that less medium-sized and large titanium-rich  
345 grains acquired the VRM, compared to the pTRM. This is the reason that the demagnetization plots  
346 (Fig. 4) show a significantly weaker VRM than pTRM; the titanomagnetites are less responsive to VRM  
347 acquisition than they are to TRM acquisition. The same effect occurs for the broad grain distribution  
348 but the effect is considerably more smeared out due to the distribution shape.

## 349 5.2 AF demagnetization

350 The effects observed during thermal demagnetization are even more pronounced in the AF demagne-  
351 tization data: first, as AF demagnetization is done on a timescale of 10 ms (at 50 Hz), the difference  
352 in time between acquisition and demagnetization is even larger than in thermal demagnetization, and  
353 second, the shape of the AF demagnetization curves is given by a different equation (eq. 11); the slope  
354 of the dotted lines in Fig. 7 indicating the AF blocking condition for different peak fields  $\tilde{H}$  is shallower  
355 than the dashed lines for thermal demagnetization. Increasing the peak AF field has a similar effect  
356 as increasing the temperature: shifting the lines to the top right while rotating them anti-clockwise:  
357 titanomagnetites are more responsive to both thermal demagnetization (increases in temperature) and  
358 to AF demagnetization (increases in peak AF field) than magnetite, which is more responsive to VRM  
359 acquisition and decay (increases in time).

360 As these effects are more pronounced than for thermal demagnetization, four apparent components  
361 arise in the case of the bimodal distribution (Fig. 3): An example is shown shown for scenario 1,

362 where up to 25 mT, mostly titanomagnetite pTRM is demagnetized (Fig. 7a), and up to 55 mT,  
363 mostly titanomagnetite ChRM is demagnetized. In the latter range, small-grained magnetite is also  
364 demagnetized, causing an overlap between the two components in the Zijderveld plot. As these grains  
365 are small, however, their magnetic moment is weak, and the direction of the pTRM dominates. Above  
366 55 mT, most titanomagnetite grains have no remanence left, and the demagnetization of small-grained  
367 magnetite carrying the pTRM dominates. Above 75 mT, the larger-grained magnetite carrying the  
368 ChRM are demagnetized.

369 Similarly for the continuous grain distribution, the two minerals (high-titanium and low-titanium  
370 titanomagnetite) show completely separate components in the demagnetization plots (Fig. 4). In  
371 addition to plots showing four apparent components (Fig. 4b and 4h) similar to the bimodal grain  
372 distribution, a further effect is encountered in scenario 2 (Fig. 7d): Here the two components com-  
373 pletely overlap from 23 mT, but in different proportions. While further increasing the peak AF,  
374 the grain population is progressively demagnetized from two sides: large-grained, high-titanium and  
375 small-grained, low-titanium. While the small-grained, low-titanium magnetite carries only the pTRM,  
376 the large-grained, high-titanium titanomagnetite carries both the pTRM and the ChRM. This situ-  
377 ation continues up to 72 mT, where both components are simultaneously completely demagnetized.  
378 Therefore, the AF demagnetization plot (Fig. 4e) appears to show three components, a low-coercivity  
379 component in the pTRM direction, an intermediate coercivity component in the ChRM direction, and  
380 a high coercivity component in an intermediate direction. This interpretation is, however, incorrect, as  
381 (1) the apparent intermediate coercivity component results from the fact that in this coercivity range  
382 the magnetic moment of the large-grained titanium-rich titanomagnetite is relatively larger than that of  
383 the small-grained titanium-poor magnetite, making the direction appear close to the ChRM-direction,  
384 and (2) the high coercivity component results from an overlap of medium-sized titanium-rich grains  
385 carrying the ChRM and medium-sized iron-rich magnetite grains carrying the pTRM, that both have  
386 similar sizes and hence similar magnetic moments, yielding an intermediate direction. Both apparent  
387 directional components therefore strongly depend on the grain distribution: the directions obtained  
388 from such Zijderveld plots are equally dependent on both the grain distribution and on the directions  
389 of the magnetizing fields. Given this particular grain distribution, it becomes obvious that the highest  
390 coercivities need not represent the ChRM: in this case the intermediate coercivities are closer to the real  
391 ChRM direction due to less overlap, whereas the high coercivities are far less useful for paleodirection  
392 reconstructions.

### 393 5.3 Paleointensity

394 The *Thellier and Thellier* (1959)/*Coe* (1967) paleointensity simulations (Fig. 5 and 6) showed very  
395 little dependence on the grain distribution compared to the Zijderveld plots. They are, however, very  
396 sensitive to the cooling rate / acquisition time of the remanence, in accordance with previous studies  
397 (*Halgedahl et al.*, 1980; *Fox and Aitken*, 1980; *McClelland-Brown*, 1984; *Bowles et al.*, 2005; *Biggin*  
398 *et al.*, 2013); lower cooling rates and hence longer acquisition times lead to a steeper slope in the Arai  
399 plots. Average slopes have been calculated for a single TRM acquired at various cooling rates from  
400 10 min to 1 Ma (Fig. 8). The relationship between cooling rate (linear cooling) and Arai plot slope is  
401 approximately linear (on a logarithmic scale), but varies slightly depending on the grain distribution.

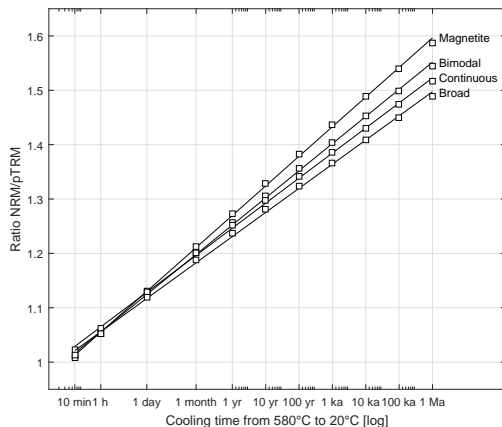


Figure 8: Plot of the cooling rate effect for linear cooling in *Thellier and Thellier* (1959) type experiments for the three grain distribution cases and model validation case of pure magnetite. Linear fit lines are indicated.

402 For cooling over a day, an NRM/pTRM ratio of  $\sim 1.12$  is obtained and for cooling over 1 Ma a  
 403 ratio of 1.5 to 1.6 depending on the grain distribution. These values correspond to an overestimate of  
 404 paleointensities of 12% and 50–60%, respectively, without appropriate cooling rate corrections. These  
 405 values are similar to those obtained by *Halgedahl et al.* (1980) analytically for slowly cooling rocks:  
 406 *Halgedahl et al.* (1980) calculated pTRM acquisition values that would lead to 10% overestimates of  
 407 paleointensities for a rock cooled over 2 days and about 45% for cooling over 1.6 Ma; the first value  
 408 coinciding closely with the one obtained here, and the second value being slightly lower.

409 The slope of the Arai plots is determined by the relative strength of NRM loss during the first  
 410 heating cycle versus pTRM gain during the second heating cycle. Both the relative independence  
 411 from mineralogy and the strong dependence on cooling rate can be explained this way: In standard  
 412 *Thellier and Thellier* (1959)/*Coe* (1967) experiments, both heating cycles have the same heating and  
 413 cooling rates and the same hold-time. Therefore, both cycles activate the same set of grains. The  
 414 slope of the Arai plot is therefore largely independent of the grain distribution. On the other hand,  
 415 the total NRM moment that is carried by the set of grains that are activated during the first heating  
 416 cycle depends on the acquisition time: longer acquisition times generally lead to a higher remanence  
 417 carried by the magnetic grains and therefore lead to a steeper Arai plot. Minor variations of the Arai  
 418 plots with mineralogy are due to the fact that the different magnetic minerals have different levels  
 419 of responsiveness to this cooling rate effect: magnetite is able to continue to acquire a thermoviscous  
 420 remanence at temperatures below its blocking temperature, thereby increasing its magnetic moment  
 421 upon slow cooling. Larger grained high-titanium titanomagnetite, on the other hand, block close their  
 422 Curie temperatures and do not significantly increase in magnetization upon slow cooling. Therefore  
 423 the cooling rate effect is slightly stronger in titanium-poor minerals, in which case a slightly larger  
 424 correction factor must be applied (Fig. 8).

425 The simulations suggest that for slowly cooling rocks a cooling rate correction of up to a factor of  
 426 1.5 (for broad grain distributions) to 1.6 (for pure magnetite) may need to be applied. The correction

427 factors obtained here (Fig. 8) agree well with those obtained by *Halgedahl et al.* (1980). These the-  
428 oretical predictions have been experimentally confirmed for SD samples by various studies: 6%–12%  
429 overestimates for archaeological baked clays refired and cooled over 7 h in the laboratory (*Fox and*  
430 *Aitken*, 1980); 15% overestimates for synthetic SD magnetite with NRM acquisition on cooling 50  
431 times slower than the *Thellier and Thellier* (1959) experiments (*McClelland-Brown*, 1984) (equiva-  
432 lent to 8 h in Fig. 8); 11–26% overestimates for remelted volcanic glass containing SD magnetite on  
433 75-fold lower NRM acquisition cooling rate (*Ferk et al.*, 2010) (equivalent to 12 h in Fig. 8); 5–10%  
434 overestimates for SD low-Ti titanomagnetite volcanic glasses at 34-fold lower NRM acquisition cooling  
435 rate (*Leonhardt et al.*, 2006) (equivalent to 6 h in Fig. 8). Similar values were obtained by others  
436 for baked clays and volcanic glasses in the SD range (*Papusoi*, 1972; *Chauvin et al.*, 2000; *Bowles*  
437 *et al.*, 2005; *Yu and Tauxe*, 2006) and the PSD range (*Yang et al.*, 1993; *Biquand*, 1994; *Genevey and*  
438 *Gallet*, 2002; *Genevey et al.*, 2003; *Morales et al.*, 2006). *Biggin et al.* (2013) found that the cooling  
439 rate effect is weaker for interacting SD, PSD and MD grains than for non-interacting SD grains, with  
440 a  $\sim 3\%$  increase in TRM magnitude per order-of-magnitude decrease in cooling rate. In summary,  
441 the cooling rate effect on paleointensities in this study coincides well both with theoretical predictions  
442 by *Halgedahl et al.* (1980) and with experimental observations and is more important than the grain  
443 size/composition distribution.

## 444 6 Conclusions

445 The simulations have shown that the presence of mixtures of titanomagnetites has very significant  
446 effects on the vector demagnetization plots in all cases except the one where the demagnetization  
447 timescale is equal to the acquisition timescale. In particular, two cases can be observed in stepwise  
448 thermal demagnetization, one that shows an apparent third component at intermediate temperatures  
449 that arises from an overlap of a remagnetization carried by small-grained iron-rich magnetite and large-  
450 grained titanium-rich titanomagnetite, and one that shows a continuous curvature between the two  
451 components. In both cases, the blocking temperatures of the “intermediate component” are a function  
452 of the grain distribution, the acquisition time and temperature and the demagnetization time. In  
453 particular, although in clearly bimodal distributions, where two clear distinct Curie temperatures can  
454 be measured in the  $M_s(T)$  curves, the upper and lower blocking temperatures can be attributed to  
455 the two grain populations with two distinct Curie temperatures, in more continuous distributions the  
456 exact mineral (Curie temperature) and grain size that the upper and lower blocking temperatures of  
457 the apparent intermediate component correspond to is not easily determined. Instead, it depends on  
458 the shape of the grain distribution, with the blocking temperatures corresponding to neither of the  
459 end-members of the distribution.

460 The simulated AF demagnetization experiments show particularly strong deviations from the case of  
461 a unique magnetic mineral, which is due to the short timescales involved in AF demagnetization and to  
462 the different blocking mechanism. For remanent magnetizations with long acquisition times, Zijderveld  
463 plots of AF demagnetization experiments may show three to four components. The highest coercivity  
464 component is not necessarily equivalent to the primary remanence and does not necessarily correspond  
465 to the highest temperature component in an analogous thermal demagnetization experiment. Although

466 the interpretation of such Zijderveld plots is not straightforward, the magnetic remanences carried by  
467 different magnetic minerals may appear completely separate in AF demagnetization, which may allow  
468 to isolate the paleomagnetic directions of interest.

469 For paleointensity experiments it was found that the grain distribution affects the slope of Arai  
470 plots, but is negligible compared to the effect of the cooling rate of NRM acquisition. The simulations  
471 suggest that for slowly cooling rocks a cooling rate correction of up to a factor of 1.5 (for broad grain  
472 distributions) to 1.6 (for pure magnetite) may need to be applied. It was also shown that VRM  
473 acquisition impacts Arai plots, even though their direction may be indistinguishable from the ChRM.  
474 Contrary to directional analysis, paleointensities can be relatively easily analyzed using the cooling  
475 rate / Arai plot slope correction factors in Fig. 8. The cooling effect in our simulations is similar  
476 in magnitude as theoretically predicted by *Halgedahl et al.* (1980) and consistent with experimental  
477 observations.

478 All this shows that it is critical to identify the presence of mixtures of different magnetic minerals  
479 when interpreting demagnetization data for paleomagnetic field reconstruction. Although the same  
480 information about the magnetic history of a sample is preserved in mixtures as in pure materials, its  
481 interpretation is significantly complicated. Mixtures of different minerals can often be identified from  
482  $M_s(T)$  curves: all curves in Fig. 2 significantly deviate from pure magnetite or single-titanomagnetite  
483 curves, either showing more than one clearly distinguishable Curie temperature or showing a strong  
484 decay at low temperatures leveling off at high temperatures. When such mixtures are identified, addi-  
485 tional information about the acquisition times is needed to correctly identify primary magnetizations  
486 in Zijderveld plots.

## 487 Acknowledgments

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489 Engineering, Imperial College London.

## 490 References

- 491 Ade-Hall, J. M., R. L. Wilson, and P. J. Smith (1965), The petrology, Curie points and natural  
492 magnetizations of basic lavas, *Geophysical Journal International*, 9(4), 323–336, doi:10.1111/j.1365-  
493 246X.1965.tb03890.x.
- 494 Berndt, T., A. R. Muxworthy, and G. A. Paterson (2015), Determining the magnetic attempt time  $\tau_0$ ,  
495 its temperature dependence, and the grain size distribution from magnetic viscosity measurements,  
496 *Journal of Geophysical Research: Solid Earth*, 120, doi:10.1002/2015JB012283.
- 497 Biggin, A. J., S. Badejo, E. Hodgson, a. R. Muxworthy, J. Shaw, and M. J. Dekkers (2013), The effect  
498 of cooling rate on the intensity of thermoremanent magnetization (TRM) acquired by assemblages  
499 of pseudo-single domain, multidomain and interacting single-domain grains, *Geophysical Journal  
500 International*, 193(3), 1239–1249, doi:10.1093/gji/ggt078.

- 501 Biggin, A. J., E. J. Piispa, L. J. Pesonen, R. Holme, G. A. Paterson, T. Veikkolainen, and L. Tauxe  
502 (2015), Palaeomagnetic field intensity variations suggest Mesoproterozoic inner-core nucleation, *Nature*,  
503 *526* (7572), 245–248, doi:10.1038/nature15523.
- 504 Biquand, D. (1994), Effet de la vitesse de refroidissement sur l’intensité de l’aimantation thermoréma-  
505 nente: étude expérimentale, conséquences théoriques, *Canadian Journal of Earth Sciences*, *31* (8),  
506 1342–1352, doi:10.1139/e94-117.
- 507 Bowles, J., J. S. Gee, D. V. Kent, E. Bergmanis, and J. Sinton (2005), Cooling rate effects on paleoin-  
508 tensity estimates in submarine basaltic glass and implications for dating young flows, *Geochemistry,*  
509 *Geophysics, Geosystems*, *6* (7), doi:10.1029/2004GC000900.
- 510 Chauvin, A., Y. Garcia, P. Lanos, and F. Laubenheimer (2000), Paleointensity of the geomagnetic  
511 field recovered on archaeomagnetic sites from France, *Physics of the Earth and Planetary Interiors*,  
512 *120* (1), 111–136, doi:10.1016/S0031-9201(00)00148-5.
- 513 Coe, R. S. (1967), The determination of paleo-intensities of the Earth’s magnetic field with emphasis  
514 on mechanisms which could cause non-ideal behavior in Thellier’s method, *Journal of geomagnetism*  
515 *and geoelectricity*, *19* (3), 157–179.
- 516 Dodson, M. H., and E. McClelland-Brown (1980), Magnetic blocking temperatures of single-  
517 domain grains during slow cooling, *Journal of Geophysical Research*, *85*, 2625–2637, doi:  
518 10.1029/JB085iB05p02625.
- 519 Dunlop, D. J., and Ö. Özdemir (1997), *Rock Magnetism: Fundamentals and Frontiers*, 595 pp., Cam-  
520 bridge University Press, Cambridge, UK.
- 521 Dunlop, D. J., and Ö. Özdemir (2007), Magnetizations in rocks and minerals, in *Geomagnetism:*  
522 *Treatise on Geophysics*, edited by M. Kono, chap. 5.08, pp. 277–336, Elsevier B.V.
- 523 Ferik, A., F. W. V. Aulock, R. Leonhardt, K.-U. Hess, and D. B. Dingwell (2010), A cooling rate  
524 bias in paleointensity determination from volcanic glass: An experimental demonstration, *Journal*  
525 *of Geophysical Research: Solid Earth*, *115* (B8), B08,102, doi:10.1029/2009JB006964.
- 526 Fox, J., and M. Aitken (1980), Cooling-rate dependence of thermoremanent magnetisation, *Nature*,  
527 *283*, 462–463, doi:10.1038/283462a0.
- 528 Genevey, A., and Y. Gallet (2002), Intensity of the geomagnetic field in western Europe over the past  
529 2000 years: New data from ancient French pottery, *Journal of Geophysical Research*, *107* (B11),  
530 2285, doi:10.1029/2001JB000701.
- 531 Genevey, A., Y. Gallet, and J.-C. Margueron (2003), Eight thousand years of geomagnetic field in-  
532 tensity variations in the eastern Mediterranean, *Journal of Geophysical Research*, *108* (B5), 2228,  
533 doi:10.1029/2001JB001612.
- 534 Halgedahl, S., R. Day, and M. Fuller (1980), The effect of cooling rate on the intensity of weak-  
535 field TRM in single-domain magnetite, *Journal of Geophysical Research*, *85* (80), 3690–3698, doi:  
536 10.1029/JB085iB07p03690.

- 537 Leonhardt, R., J. Matzka, A. Nichols, and D. Dingwell (2006), Cooling rate correction of paleointen-  
538 sity determination for volcanic glasses by relaxation geospeedometry, *Earth and Planetary Science*  
539 *Letters*, *243*(1), 282–292, doi:10.1016/j.epsl.2005.12.038.
- 540 McClelland-Brown, E. (1984), Experiments on TRM intensity dependence on cooling rate, *Geophysical*  
541 *Research Letters*, *11*(3), 205–208, doi:10.1029/GL011i003p00205.
- 542 Morales, J., et al. (2006), Cooling rate corrected paleointensities from the Xitle lava flow: Evaluation  
543 of within-site scatter for single spot-reading cooling units, *Earth, Planets and Space*, *58*(10), 1341–  
544 1347, doi:10.1186/BF03352630.
- 545 Muxworthy, A. R., M. E. Evans, S. J. Scourfield, and J. G. King (2013), Paleointensity results from the  
546 late-Archaean Modipe Gabbro of Botswana, *Geochemistry, Geophysics, Geosystems*, *14*(7), 2198–  
547 2205, doi:10.1002/ggge.20142.
- 548 Nagata, T., Y. Arai, and K. Momose (1963), Secular variation of the geomagnetic total force during  
549 the last 5000 years, *J. Geophys. Res.*, *68*, 5277–5281, doi:10.1029/JZ068i018p05277.
- 550 Néel, L. (1949), Théorie du traînage magnétique des ferromagnétiques en grains fins avec applications  
551 aux terres cuites, *Annales de Géophysique*, doi:10.1016/S0009-2509(00)00427-9.
- 552 O’Rourke, J. G., and D. J. Stevenson (2016), Powering Earth’s dynamo with magnesium precipitation  
553 from the core, *Nature*, *529*(7586), 387–389, doi:10.1038/nature16495.
- 554 Özdemir, Ö., and W. O’Reilly (1981), High-temperature hysteresis and other magnetic properties  
555 of synthetic monodomain titanomagnetites, *Physics of the Earth and Planetary Interiors*, *25*(4),  
556 406–418, doi:10.1016/0031-9201(81)90052-2.
- 557 Papusoi, C. (1972), Effet de la vitesse de refroidissement sur l’intensité de l’aimantation thermoréma-  
558 nente d’un ensemble de grains monodomaines, *An. Stiint. Univ. Al. I. Cuza Iasi Sect 1b, Tomul*, *18*,  
559 31–47.
- 560 Pullaiah, G., E. Irving, K. Buchan, and D. Dunlop (1975), Magnetization changes caused by burial  
561 and uplift, *Earth and Planetary Science Letters*, *28*, 133–143, doi:10.1016/0012-821X(75)90221-6.
- 562 Stephenson, A. (1969), The temperature dependent cation distribution in titanomagnetites,  
563 *Geophysical Journal of the Royal Astronomical Society*, *18*(2), 199–210, doi:10.1111/j.1365-  
564 246X.1969.tb03562.x.
- 565 Tarduno, J. A., R. D. Cottrell, W. J. Davis, F. Nimmo, and R. K. Bono (2015), A Hadean to  
566 Paleoproterozoic geodynamo recorded by single zircon crystals, *Science*, *349*(6247), 521–524, doi:  
567 10.1126/science.aaa9114.
- 568 Thellier, E., and O. Thellier (1959), Sur l’intensité du champ magnétique terrestre dans le passé  
569 historique et géologique, *Annales de Géophysique*, *15*, 285–376.
- 570 Worm, H. (1998), On the superparamagnetic-stable single domain transition for magnetite, and  
571 frequency dependence of susceptibility, *Geophysical Journal International*, *133*, 201–206, doi:  
572 10.1046/j.1365-246X.1998.1331468.x.

- 573 Yang, S., J. Shaw, and Q. Y. Wei (1993), Tracking a non-dipole geomagnetic anomaly using new  
574 archaeointensity results from north-east China, *Geophysical Journal International*, 115 (3), 1189–  
575 1196, doi:10.1111/j.1365-246X.1993.tb01520.x.
- 576 York, D. (1978a), Magnetic blocking temperature, *Earth and Planetary Science Letters*, 39, 94–97,  
577 doi:10.1016/0012-821X(78)90145-0.
- 578 York, D. (1978b), A formula describing both magnetic and isotopic blocking temperatures, *Earth and*  
579 *Planetary Science Letters*, 39, 89–93, doi:10.1016/0012-821X(78)90144-9.
- 580 Yu, Y., and L. Tauxe (2006), Acquisition of viscous remanent magnetization, *Physics of the Earth and*  
581 *Planetary Interiors*, 159 (1-2), 32–42, doi:10.1016/j.pepi.2006.05.002.
- 582 Zhou, W., R. Van der Voo, and D. R. Peacor (1997), Single-domain and superparamagnetic titanomag-  
583 netite with variable Ti content in young ocean-floor basalts: No evidence for rapid alteration,  
584 *Earth and Planetary Science Letters*, 150 (3-4), 353–362, doi:10.1016/S0012-821X(97)00099-X.
- 585 Zhou, W., R. Van Der Voo, D. R. Peacor, and Y. Zhang (2000), Variable Ti-content and grain size  
586 of titanomagnetite as a function of cooling rate in very young MORB, *Earth and Planetary Science*  
587 *Letters*, 179 (1), 9–20, doi:10.1016/S0012-821X(00)00100-X.
- 588 Zijdeveld, J. (1967), AC demagnetization of rocks: analysis of results, in *Methods in paleomagnetism*,  
589 edited by D. Collinson, K. Creer, and S. Runcorn, pp. 254–286.