

# **City Research Online**

## City, University of London Institutional Repository

**Citation**: He, Y-H., Jejjala, V. & Pontiggia, L. (2017). Patterns in Calabi-Yau Distributions. Communications in Mathematical Physics, 354(2), pp. 477-524. doi: 10.1007/s00220-017-2907-9

This is the published version of the paper.

This version of the publication may differ from the final published version.

Permanent repository link: http://openaccess.city.ac.uk/18200/

Link to published version: http://dx.doi.org/10.1007/s00220-017-2907-9

**Copyright and reuse:** City Research Online aims to make research outputs of City, University of London available to a wider audience. Copyright and Moral Rights remain with the author(s) and/or copyright holders. URLs from City Research Online may be freely distributed and linked to.

City Research Online:	http://openaccess.city.ac.uk/	publications@city.ac.uk



### Patterns in Calabi–Yau Distributions

Yang-Hui He<sup>1,2,3</sup>, Vishnu Jejjala<sup>4</sup>, Luca Pontiggia<sup>4</sup>

<sup>1</sup> School of Physics, NanKai University, Tianjin 300071, People's Republic of China

<sup>2</sup> Department of Mathematics, City University, London EC1V 0HB, UK

<sup>3</sup> Merton College, University of Oxford, Oxford OX1 4JD, UK. E-mail: hey@maths.ox.ac.uk

<sup>4</sup> NITheP, School of Physics, and Mandelstam Institute for Theoretical Physics, University of the Witwatersrand, 1 Jan Smuts Avenue, Johannesburg 2050, South Africa. E-mail: vishnu@neo.phys.wits.ac.za; lucatpontiggia@gmail.com

Received: 18 December 2015 / Accepted: 17 April 2017 Published online: 30 May 2017 – © The Author(s) 2017. This article is an open access publication

**Abstract:** We explore the distribution of topological numbers in Calabi–Yau manifolds, using the Kreuzer–Skarke dataset of hypersurfaces in toric varieties as a testing ground. While the Hodge numbers are well-known to exhibit mirror symmetry, patterns in frequencies of combination thereof exhibit striking new patterns. We find pseudo-Voigt and Planckian distributions with high confidence and exact fit for many substructures. The patterns indicate typicality within the landscape of Calabi–Yau manifolds of various dimension.

#### Contents

1.	Introduction	478
2.	Calabi–Yau Threefolds	480
	2.1 Analysis of $h^{1,1} - h^{1,2}$	481
	2.1.1 Å pseudo-Voigt fit	483
	2.2 Analysis of $h^{1,1} + h^{1,2}$	488
	2.2.1 A Planckian fit	488
	2.3 The distribution of the Euler number	495
	2.4 Goodness-of-fit	496
	2.5 Implications for physics	501
3.	Calabi–Yau Twofolds: K3 Surfaces	502
4.	Calabi–Yau Fourfolds	502
5.	Conclusions and Outlook	505
A.	Appendix	507
	A.1 Supplementary plots for the $h^{1,1} - h^{1,2}$ distribution	507
	A.1.1 Plots for the odd distribution as counterparts to the even ones	507
	A.1.2 Comparative plots	507
	A.1.3 A first approximation to the data	510
	A.1.4 Table of parameter values and statistics	510

A.2 Supplementary plots for the $h^{1,1} + h^{1,2}$ distribution	511
A.2.1 Plots for the odd distribution as counterparts to the even ones	511
A.2.2 Table of parameter values, coefficient values and statistics	512
A.3 Supplementary plots for the fourfold data	515

#### 1. Introduction

A Calabi–Yau *n*-fold is a Kähler manifold of *n* complex dimensions with a trivial canonical bundle. In superstring theory, it serves as a compactification manifold wherein a ten dimensional theory at high energies reduces to an effective theory in four spacetime dimensions. In particular, global SU(n) holonomy ensures that  $2^{1-n}$  of the original supersymmetry is preserved. Thus, confronted by the vacuum selection problem, Calabi–Yau compactifications present an avenue for Standard Model building, especially in the context of the heterotic string [1–4]. Indeed, the basis of the landscape is to consider flux compactifications on these geometries [5,6].

To facilitate this approach to a low-energy phenomenology derived from string theory, mathematicians and physicists have constructed large datasets of Calabi–Yau threefolds [7,9–22] as well as various refined analyses of properties thereof [28–35]. By far the largest database was constructed in a *tour de force* of algebraic geometry, combinatorics, physics, and computer algorithms by Kreuzer and Skarke based on the theorems of Batyrev and Borisov [9–14,36,37]. In short, these Calabi–Yau *n*-manifolds  $X_n$  are realized as a smooth hypersurface embedded in a toric variety  $A_{n+1}$  of complex dimension n + 1; the Calabi–Yau condition simply translates to the requirement that the polytope defining  $A_{n+1}$  be **reflexive**. We will henceforth consider only such Calabi–Yau manifolds, of which there are a plethora.

Let us briefly recollect what all this means. The (possibly singular) toric variety  $A_{n+1}$  is specified by an integer polytope  $\Delta$  in  $\mathbb{R}^{n+1}$ , which is a collection of vertices (dimension 0) each of which is an (n + 1)-vector with integer entries and such that each pair of neighboring vertices defines an edge (dimension 1), each pair of edges defines a face (dimension 2), etc., all the way up to a facet (dimension *n*). Alternatively,  $\Delta$  can be defined by a set of integer linear inequalities, each of which slices a facet. The polytope is then the convex body in  $\mathbb{R}^{n+1}$  enclosed by these facets. We will always include the origin as being contained in  $\Delta$ . Using the usual dot product  $\langle , \rangle$  inherited from  $\mathbb{R}^{n+1}$ , the dual polytope is defined by

$$\Delta^{\circ} := \left\{ v \in \mathbb{R}^{n+1} | \langle m, v \rangle \ge -1, \forall m \in \Delta \right\}.$$
(1.1)

The polytope  $\Delta$  is *reflexive* if all the vertices of  $\Delta^{\circ}$  are integer vectors. In this case, we can define the Calabi–Yau hypersurface  $X_n$  explicitly as the polynomial equation

$$\sum_{m \in \Delta} c_m \prod_{r=1}^k x_r^{\langle m, v_r \rangle + 1} = 0 , \qquad (1.2)$$

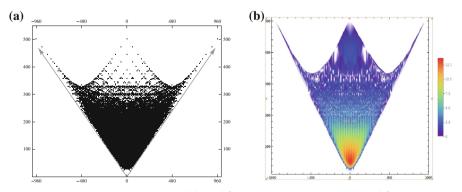
where  $v_{r=1,...,k}$  are the vertices of  $\Delta^{\circ}$  with k being the number of vertices of  $\Delta^{\circ}$  (or equivalently the number of facets of  $\Delta$ ),  $x_r$  are the coordinates of  $A_{n+1}$ , and  $c_m$  are numerical coefficients parameterizing the complex structure of  $X_n$ . Indeed, the reflexivity of  $\Delta$  ensures that the exponents are integral thereby making the hypersurface polynomial as required.

The classification of these Calabi–Yau manifolds thus amounts to that of reflexive polytopes in various dimensions, and the intense computer work of Kreuzer and Skarke was to combinatorially find such polytopes. For n = 1, there are 16 such polytopes in  $\mathbb{R}^2$ , and we have Calabi–Yau onefolds, or elliptic curves. For n = 2, there are 4319 such polytopes in  $\mathbb{R}^3$ , and we have Calabi–Yau twofolds, or K3 surfaces. For n = 3, there are 473, 800, 776 such polytopes (which was a formidable computer task!), and we have the Calabi–Yau threefolds. This sequence

$$\{1, 16, 4319, 473800776, \ldots\}$$
(1.3)

of remarkable growth rate can be found in the Online Encyclopedia of Integer Sequences [38]. The numbers in higher dimension are still not known, nor has there been an asymptotic analysis of their growth. It should be emphasized that generically a reflexive polytope corresponds to a *singular* toric variety even though the hypersurface is chosen (by generic coefficients  $c_m$ ) to miss the singularities and hence ensuring the smoothness of the Calabi–Yau  $X_n$ . For example, of the some half-billion reflexive polytopes in  $\mathbb{R}^4$ , only 136  $A_4$  are in fact smooth [39]. As we desingularize the toric variety by various star-triangulations of  $\Delta$ , we are led to potentially *inequivalent* Calabi–Yau manifolds. In principle, the *same* Calabi–Yau geometry can arise from different reflexive polytopes or triangulations of a given reflexive polytope. Whereas K3 is essentially unique, we do not know how many Calabi–Yau threefolds there are. A systematic study to classify the desingularizations, to compute the necessary topological data, and to build an interactive online database [19] is under way. The moral is that there are almost certainly far more than half a billion Calabi–Yau threefolds!

Luckily, the Hodge numbers depend only on the polytope and not on the choice of desingularization. (The intersection numbers, however, do depend on the choice.) For Calabi–Yau threefolds, the pair of Hodge numbers  $(h^{1,1}, h^{1,2})$  is a famous quantity. Indeed, the plot in Part (a) of Fig. 1 has become iconic. Here, the sum  $h^{1,1} + h^{1,2}$  is plotted against the Euler number  $\chi = 2(h^{1,1} - h^{1,2})$ , and the left-right symmetry supplies "experimental evidence" for *mirror symmetry*. There is enormous redundancy in this data: of the some half a billion reflexive polytopes, there are only 30, 108 distinct pairs of Hodge numbers and the pair (27, 27) dominates the multiplicity, totaling almost one million. In Part (b) of Fig. 1 we have attempted to visualize the distribution of the



**Fig. 1. a** The cumulative plot of  $\chi = 2(h^{1,1} - h^{1,2})$  on the abscissa versus  $h^{1,1} + h^{1,2}$  on the ordinate for Calabi–Yau threefolds as hypersurfaces in toric fourfolds; **b** marking also the natural logarithm of the multiplicity of the Hodge pair with a *color* grading (color figure online)

multiplicity by having a color density plot of the logarithm of the number over each Hodge pair.

Understanding this multiplicity forms the inspiration for the present work. While there have been analyses on the *shape* of the funnel-like plot [28, 33, 35], there has not been much work on its *density*, i.e., the distribution of the multiplicity of Hodge data for the Calabi–Yau manifolds of various dimension. Of course, fundamentally, this is entirely due to the combinatorics of reflexive polytopes and might in principle be analytically determined. However, given the complexity of the problem it is expedient to analyze the available data which have been compiled over the years, observe intriguing patterns, and draw statistical inferences before turning to analytic treatments. This is what we achieve in this work.

The organization of the paper is as follows. We perform a detailed analysis on the structure and behavior of the threefold data in Sect. 2. This is motivated by looking for an exact function describing the relationship of the distribution of the Hodge pairs  $(h^{1,1}, h^{1,2})$  with frequency.

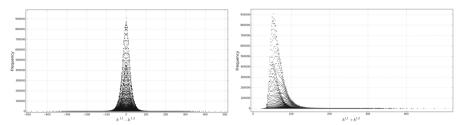
In Sect. 2.1, we study the distribution of  $(h^{1,1} - h^{1,2}, f)$ . We find that this distribution is composed of a family of curves, for which each curve can be described using a modified pseudo-Voigt model. Although an approximation, the model is able to describe the general trend of the data, as well as some additional fine structure within each individual data point. Performing an analysis on the parameter relationships shows that three out of the five parameters can be expressed as a single variable, but we conclude that additional modifications need to be introduced in the model to overcome certain shortfalls.

Subsequently, Sect. 2.2 performs an analysis on the structure of  $(h^{1,1} + h^{1,2}, f)$ . Similarly, this distribution is composed of a family of curves for which each curve can be described using a Planckian profile. Combining the regression analysis for each curve within the distribution, we construct a single function able to approximately model the entire distribution of  $(h^{1,1} + h^{1,2}, f)$  with only two variables. Section 2.3 uses the model developed in Sect. 2.1 to describe the distribution of the Euler number  $\chi$ .

Section 2.4 is dedicated to the description of model validation in our context, as the usual statistical tests are inadequate. Section 2.5 discusses possible implications to physics by referencing recent advancements in F theory and further investigations of structures within the Kreuzer–Skarke database. In Sects. 3 and 4, we perform primary analyses of Calabi–Yau twofolds (Picard number and multiplicity) and Calabi–Yau fourfolds. Due to the lack of a complete data set, we are unable to provide a thorough analysis of the fourfolds as with threefolds. Finally, the Appendix presents many supplementary plots and figures for the various sections. We conclude with a summary and outlook in Sect. 5.

#### 2. Calabi–Yau Threefolds

As advertised in the Introduction, we will begin with the analysis of threefolds and identify patterns within this rich distribution of Hodge numbers and their frequency as plotted in Fig. 1. It turns out striking patterns do exist, pointing to a definite structure within the threefold data, which consists of the triple  $(h^{1,1}, h^{1,2}, f)$ , where f is the number of reflexive polytopes in the Kreuzer–Skarke database with the given Hodge pair. Here,  $h^{1,1}$  and  $h^{1,2}$  respectively count the Kähler and complex structure moduli of the Calabi–Yau obtained from the reflexive polytope. More precisely [8], we have that



**Fig. 2. a** Frequency f plotted against  $\frac{1}{2}\chi = h^{1,1} - h^{1,2}$ ; **b** frequency f plotted against the sum of Hodge numbers  $h^{1,1} + h^{1,2}$ 

$$h^{1,1}(X) = \ell(\Delta^*) - \sum_{\operatorname{codim}\theta^*=1} \ell^*(\theta^*) + \sum_{\operatorname{codim}\theta^*=2} \ell^*(\theta^*)\ell^*(\theta) - 5;$$
  
$$h^{1,2}(X) = \ell(\Delta) - \sum_{\operatorname{codim}\theta=1} \ell^*(\theta) + \sum_{\operatorname{codim}\theta=2} \ell^*(\theta)\ell^*(\theta^*) - 5.$$
(2.1)

In the above,  $\Delta$  is the defining polytope for the Calabi–Yau threefold X and  $\Delta^*$  is its dual. Moreover,  $\theta$  and  $\theta^*$  are the faces of specified codimension of these polytopes respectively;  $\ell()$  is the number of integer points of the polytope while  $\ell^*()$  is the number of integer points. Indeed, our analysis of the distribution of Hodge numbers ultimately reduces to counting these integer points.

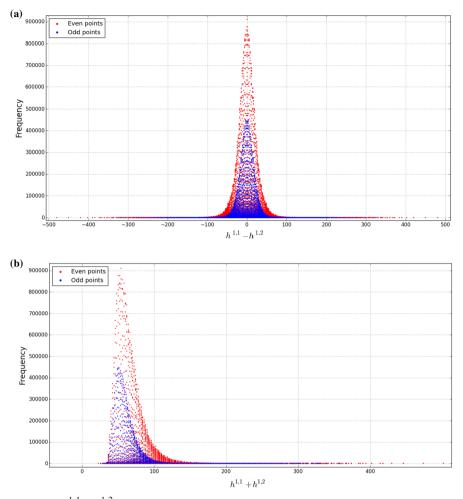
To facilitate the analysis, we plot  $(h^{1,1} - h^{1,2}, f)$  and  $(h^{1,1} + h^{1,2}, f)$  as shown in (a) and (b) of Fig. 2, respectively. Recall that the Euler number  $\chi = 2(h^{1,1} - h^{1,2})$ . We will use the difference  $h^{1,1} - h^{1,2}$  rather than the Euler number. In the simplest heterotic constructions,  $|h^{1,1} - h^{1,2}|$  corresponds to the index of the Dirac operator and gives the number of generations of particles in the low-energy spectrum [1].

By inspection, these plots already exhibit two patterns. Firstly, in both the  $h^{1,1} - h^{1,2}$  and  $h^{1,1} + h^{1,2}$  plots, there appears to be an inner distribution contained within the outer distribution. We find that these inner and outer distributions are related to the parity of  $h^{1,1} \pm h^{1,2}$ . Figure 3 elucidates this point by having the odd and even values in different colors.

Though this parity structure may be a result of the Kreuzer–Skarke algorithm, its consistent appearance means we need to treat the distributions of even and odd distinctly for now.

The second evident structure which can been seen by inspection, is that the outer edge of the distribution of  $h^{1,1} - h^{1,2}$  (Fig. 3a) appears to follow a normal like curve, whereas the edge of  $h^{1,1} + h^{1,2}$  (Fig. 3b) follows a Planck like curve. It is through the analysis of these distributions that we deduce their characteristic behavior and underlying structure. In the main body of this paper, we outline the results and analysis of only the even distributions for  $h^{1,1} - h^{1,2}$  and  $h^{1,1} + h^{1,2}$ , except where it is important to present both. It turns out that any structure and patterns which are found in the even distributions for  $h^{1,1} - h^{1,2}$  and  $h^{1,1} + h^{1,2}$  are found identically in the odd distribution (see "Appendix" for various plots).

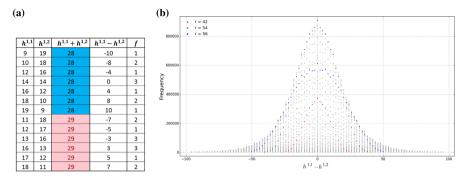
2.1. Analysis of  $h^{1,1} - h^{1,2}$ . Before we can present the results, it is important to explain some notation. When working with the distribution of  $h^{1,1} - h^{1,2}$ , we find that it is composed of many curves, whose individual structure is the same as the "edge" or boundary of the distribution mentioned earlier. As a consequence of this, we refer to



**Fig. 3. a** The  $h^{1,1} - h^{1,2}$  distribution for threefolds, highlighting the two sub-distributions, where *red* and *blue* data points correspond to even and odd values of  $h^{1,1} - h^{1,2}$ , respectively; **b** the same, but for  $h^{1,1} + h^{1,2}$  (color figure online)

 $h^{1,1} - h^{1,2}$  as being composed of a "family of curves." Each curve is then classified by its *r*-value, where  $r = h^{1,1} + h^{1,2}$ . It is important to be clear that in this analysis, although  $h^{1,1} - h^{1,2}$  is just half the Euler number, we are not summing over all the possible values of  $h^{1,1} + h^{1,2}$ . We are keeping these values distinct: hence, the *r*-curves we obtain. Later on in Sect. 2.3 we sum over all possible values of  $h^{1,1} + h^{1,2}$  to get two plots representing the full Euler number distribution.

Consider the example in Fig. 4a. By ordering the data in terms of  $h^{1,1} + h^{1,2}$ , one can classify data sets within  $h^{1,1} - h^{1,2}$  by an *r*-value. Holding *r* fixed, we can plot the frequency *f* versus the difference  $h^{1,1} - h^{1,2}$ . We call each value of *r* a curve, which we can overlay on the same plot. In this example, we tabulate data for curves identified by r = 28 and r = 29. As a further illustration, we show explicitly the curves of the even distribution within  $h^{1,1} - h^{1,2}$  for r = 42, 54, 66 in Fig. 4b. By mirror symmetry, the curve is symmetric about the vertical axis, where  $h^{1,1} - h^{1,2} = 0$ .



**Fig. 4.** a Example of repeated values of the sum  $h^{1,1} + h^{1,2}$  being 28 and 29; **b** three highlighted curves (r = 42, 54, 66) within the even  $h^{1,1} - h^{1,2}$  distribution. The transparent *grey data dots* are all the data plots for the distribution. Refer to Fig. 23 for the corresponding odd plot

We can now perform a regression analysis for each individual curve, in the quest of obtaining a function describing the distribution. In the analysis, we indeed find an approximate function predicting the fine structure of the data. We operate with one caveat: we ignore data points which have a frequency lower than 2000. At large r, the data, whose frequency is below 2000, begins to deviate from our model. The reason for such deviations, comes down to the fact that our model, though remarkably accurate, is still an approximation. We suspect that with further modifications, such deviations can be accounted for and that consequently, it may be possible to find an exact function to map the frequency distribution of  $h^{1,1} - h^{1,2}$ . Such statements also apply to the distribution of  $h^{1,1} + h^{1,2}$ .

2.1.1. A pseudo-Voigt fit Due to the normally-distributed, peak-like nature of these curves, we performed a regression analysis using the following models: Gaussian; Cauchy (Lorenztian); Pearson7; Breit–Wigner; Voigt; and pseudo-Voigt. In the "Appendix A.1.2", we perform a side by side comparison. It turns out that both the Voigt model (25e) as well as the pseudo-Voigt model (25f) give excellent fits.

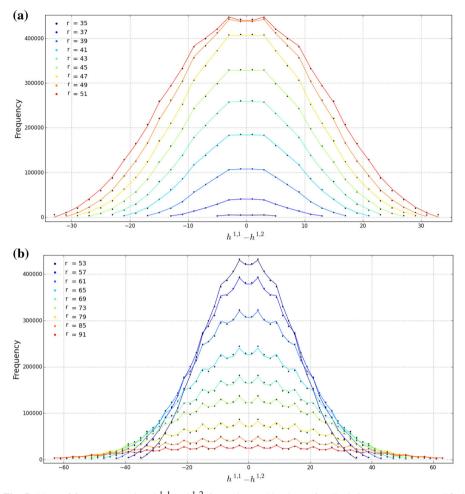
We focus on the **pseudo-Voigt model** as it gives the best fits. This is a linear combination of a Gaussian and Lorentzian (Cauchy) distribution:

$$f(x, A, \mu, \sigma, \alpha) = (1 - \alpha) \frac{A}{\sigma \sqrt{2\pi}} e^{\frac{-(x - \mu)^2}{2\sigma^2}} + \alpha \frac{A}{\pi} \left[ \frac{\sigma^2}{(x - \mu)^2 + \sigma^2} \right],$$
(2.2)

with amplitude (*A*), center ( $\mu$ ), Gaussian width ( $\sigma$ ), and fractional parameter alpha ( $\alpha$ ). However, we can modify the above distribution slightly so that the amplitude *A* of the distribution has an oscillating component

$$A(x, A_0, a, b) = A_0 + a\cos(2\pi b \cdot x),$$
(2.3)

where  $A_0$  is the original amplitude of a particular curve described by the pseudo-Voigt distribution, *a* is the amplitude of oscillations, and *b* represents the period. By doing a regression analysis one curve at a time using this modified pseudo-Voigt model, we are almost able to replicate not just the basic structure of each curve, but even the individual behavior of each data point in the entire distribution. (See "Appendix A.1.3" for a comparative plot of the all the regression curves using the standard, unmodified, pseudo-Voigt model.)



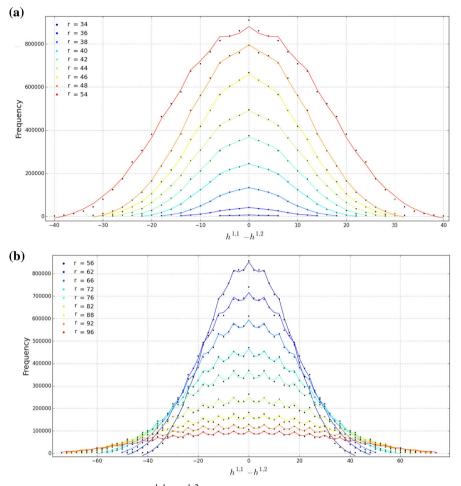
**Fig. 5.** Plots of frequency against  $h^{1,1} - h^{1,2}$  for various odd values of *r*. Each *line* represent a modified pseudo-Voigt profile based on the regression analysis for each curve. See Fig. 28a for a plot of all even curves. **a** Regression lines for all odd *r* valued curves, with  $r \in [35, 51]$ . **b** regression lines for few select odd *r* values, with r > 51

We plot the frequency against  $h^{1,1} - h^{1,2}$  for various values of r (odd and even). Figures 5 and 6 are striking in their accuracy.

As these figures illustrate, each curve follows a pseudo-Voigt profile, however the individual data points seem to "jump" up and down, as if oscillating. It is this behavior of the data points which can be accounted for by the modified pseudo-Voigt model. To do the regression analysis, we used Python *lmfit* with a custom model which is just the modified pseudo-Voigt model. The parameters that were fitted are  $(A_0, a, b, \sigma, \alpha)$ . Due to mirror symmetry,  $\mu = 0$ . In "Appendix A.1.4", one can find a table with the value of every parameter for every curve as well as their reduced  $\chi^2$  values.

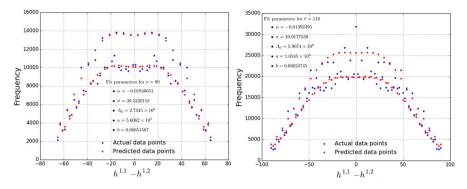
A few comments explicate the regression lines and the behavior of the distributions.

1. When we refer to the model as being an "excellent fit," it is principally a statement made by inspection of the curves and the data. If one inspects the reduced  $\chi^2$  values



**Fig. 6.** Plots of frequency against  $h^{1,1} - h^{1,2}$  for various even values of *r*. Each *line* represent a modified pseudo-Voigt profile based on the regression analysis for each curve. See Fig. 28b for a plot of all odd curves. **a** Regression lines for few select even *r* values, with  $r \le 54$ . **b** Regression lines for few select even *r* values, with r > 54

(Fig. 29), the numbers are large, which statistically does not refer to a good fit. This is misleading however. Firstly, we need to consider that the number of parameters used in the model is five. This allows for a larger  $\chi_R^2$  value. Secondly, the distribution is based on a discrete set of data. When doing a regression analysis using the modified pseudo-Voigt model, one obtains an equation which describes a continuous curve. Lastly, the frequency values span over several orders of magnitude. The tiniest deviation from a parametric model—in this case, the modified pseudo-Voigt profile—will be detected in cases where there is such a huge sample size. Typically the predicted model gives data points which are in the range of 0.02–3% accuracy from the actual data point. The tail behavior of the model is less accurate however, here the predicted values can be off from between 60 and 80%. For cases with a very poor fit, the last data point (large value of  $h^{1,1} - h^{1,2}$ ) can have an error of up to 300%—this is another example of the model being less accurate at lower frequency. When one is dealing with such



**Fig. 7.** These two plots serve two purposes. The first is to show how the modeled data should really look by using data points (*red points*) instead of the (perhaps misleading) lines (refer to Comment 1 below). The second purpose is to illustrate that as *r* becomes large (*left plot* has r = 99, *right plot* has r = 118), the actual data points deviate more and more from the modeled data, implying that there is a missing function in the modified pseudo-Voigt model which would allow one to describe the data at much lower frequencies (color figure online)

sample sizes, even a 1% error can give a difference of up to a couple of thousand. This difference summed over all the data points for a particular curve result in a large  $\chi_R^2$  value. Due to the discussion in Sect. 2.4 we from now on ignore the  $\chi_R^2$  as a test for model validation. Instead we opt for probability plots—which can also be seen in Sect. 2.4.

- 2. One obtains a continuous model to describe the discrete data, in reality, we should not be plotting fitted curves, but rather fitted data points—as can be seen in Fig. 7. It is just illustratively more clear to display the curves. One could in principal work out what the discrete approximation is to our continuous model.
- 3. Although the modified pseudo-Voigt distribution does a good job to model the behavior of the data, one still needs to address the problems experienced with our model at low frequency. A problem which is hidden, by virtue of our cut-off frequency, is that the tail of our models predicts negative values, Fig. 8. There is a possibility that by having different variances  $\sigma_g$ ,  $\sigma_c$  for the mixing of the two distributions (Gaussian, Cauchy), one could adjust the tail behavior. Introducing more and more parameters however does not always resolve the problem, as it is possible to over-fit the data. Yes, the model may be more accurate, but one loses physical significance. In a situation like ours, where one does not have any physical backing for choice in models, this line between fitting and over fitting is not so clear.
- 4. The odd distribution's behavior is more regular. In comparison to the even distribution, as one increases in *r* value, the behavior of the individual data points remain somewhat constant relative to the fitted curve. The even distribution becomes more and more irregular as one increases the *r* value. This suggests that there is an added parameter which seems as if it should be function of *r*. By regular and irregular we are referring to how well the data point is described by the model.
- 5. Both distributions become very irregular as the value of r becomes large (r > 100 and r > 120 for odd and even distributions respectively—see Fig. 7). A large r value refers to curves which have a relatively low frequency. Again this suggests that the pseudo-Voigt model needs to some how have some function of r which "distorts" the behavior of the curves as r increases (by the looks of how the real data deviates from the modeled one, it seems that the missing functions is also oscillating in nature).

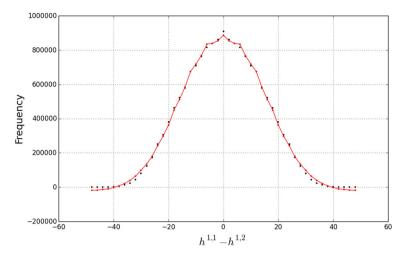
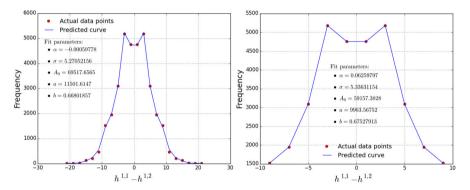


Fig. 8. By considering the entire frequency range, the model is not able to adequately describe the tail behavior. The model goes into the negative frequency range instead of tapering off to 0



**Fig. 9.** Left plot shows the modeled line according to the modified pseudo-Voigt distribution with no cutoff frequency. We obtain a good fit to the data. The *right plot* has a cutoff frequency of 460, which is equivalent to a percentage cut off of 9.68% (calculated relative to the peak frequency for that *r*-curve). This curve is exact

There exist, however, certain cases where the model is exact. In other words predicted values are the same as the actual values. This happens when one adjusts the frequency cutoff for each r curve individually. That is to say, we only examine data points with at least  $f_0$  reflexive polytopes with a given value of r and  $h^{1,1} - h^{1,2}$ . If there are fewer than  $f_0$  cases, the data is ignored.

This trend persists for all values of r, however what becomes apparent is that it's not the percentage cutoff frequency that determines whether or not one gets an exact fit, but rather, the number of data points that remains after the percentage cut of has been effected. Figure 30 gives a table of how many data points remain after an appropriate cut off percentage has been chosen to achieve a perfect fit. From this table we see that for even curves, one almost always requires 7 data points to achieve a perfect fit; for the odd curves, the number of data points is 10. The reason for this constant number throughout all the curves is that the centers of all the distributions for the various curves are all similar. As soon as one includes a larger number of data points we cannot achieve exact

fits, and the model becomes approximate. At very low r values the number of data points remaining after cutoff are not too different to the total number of points. As r increase, the total number of points increase—the fact that we can achieve exact fits becomes less meaningful. The other models—even when including an oscillatory component were unable to give exact fits.

The model is thus much more accurate at low r values, and as r increases the actual data deviates more and more from the fit. This reinforces the statements from the comments that the pseudo-Voigt model can be modified further with some function g(r, x) such that it will greatly improve the accuracy of the fit, and perhaps even become exact.

After the above analysis, we return to our goal of finding a single function describing the distributions. It is clear from the above that the function has to be a function of at least two variable, f = f(x, r). We thus continue the analysis by plotting all the parameters versus r, in search for any relationships. We find that three parameters  $\sigma$ , b and  $\alpha$  can be expressed in terms of r, the other parameters, while they show trends, do not give a precise relationship with r. For the even distribution of  $h^{1,1} - h^{1,2}$ , the r values range from 36 to 110, whereas for the odd distribution (see Fig. 24a, b) the r values range from 37 to 99. By looking at Fig. 10a, it turns out that:

$$\alpha(r) = c_{\alpha}, \quad b(r) = c_b, \quad \sigma(r) = c_{\sigma_1}r + c_{\sigma_2}.$$
 (2.4)

Our model of  $h^{1,1} - h^{1,2}$  now looks as follows:

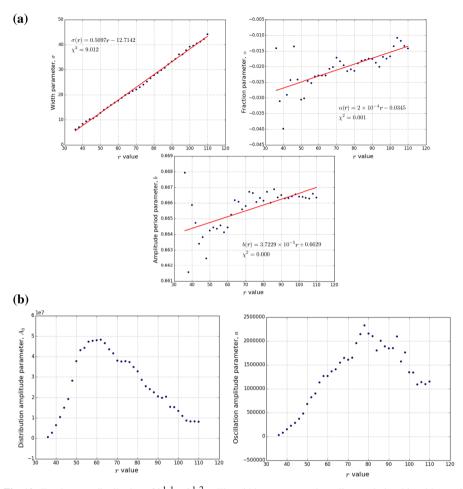
$$f(x, r, A_0, a) = (1 - c_{\alpha}) \frac{A_0(r) + a(r)\cos(2\pi c_b \cdot x)}{\sqrt{2\pi}(c_{\sigma_1}r + c_{\sigma_2})} e^{\frac{-(x)^2}{2(c_{\sigma_1}r + c_{\sigma_2})^2}} + c_{\alpha} \frac{A_0(r) + a(r)\cos(2\pi c_b \cdot x)}{\pi} \left[ \frac{(c_{\sigma_1}r + c_{\sigma_2})^2}{x^2 + (c_{\sigma_1}r + c_{\sigma_2})^2} \right], \quad (2.5)$$

where  $A_0(r)$  and a(r) are two unknown functions yet to be determined (see Fig. 10b for relationship plots). For replicating the plots as precisely as possible, one would need to keep the parameters, as they are, up to their 17 decimal values, without excluding terms as we have done. If one wants to reproduce the data from the model, one has to use the exact expressions. Making an approximation from an already approximate model leads to large errors.

The first plot in Fig. 10a in particular evinces a sinusoidal fluctuation about the mean. This again indicates the possibility of refining the plots by adding an extra function.

2.2. Analysis of  $h^{1,1} + h^{1,2}$ . We begin by classifying the curves within the  $h^{1,1} + h^{1,2}$  distribution (Fig. 2) in an analogous way to how it was explained before. This time, we order the data by  $h^{1,1} - h^{1,2}$  such that a single curve within  $h^{1,1} + h^{1,2}$  can be identified by its q-value, where  $q = h^{1,1} - h^{1,2}$ . Due to mirror symmetry, the curve for q = -a is the same curve as q = a, thus within our two-dimensional plots will only have q > 0. In continuation to the analysis on  $h^{1,1} - h^{1,2}$ , we use a cutoff frequency of 2000 and only present results from the even distribution within  $h^{1,1} + h^{1,2}$ , unless stated otherwise. As an example, illustrating the classification of curves within  $h^{1,1} + h^{1,2}$ , consider the curves q = 0, 18, 30 in Fig. 11.

2.2.1. A Planckian fit Each curve within the  $h^{1,1} + h^{1,2}$  distribution behaves the same. Just like in the  $h^{1,1} - h^{1,2}$  distribution, we do a regression analysis for each curve within

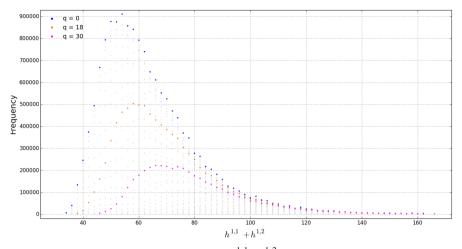


**Fig. 10.** For the even distribution of  $h^{1,1} - h^{1,2}$ . **a** The width parameter  $\sigma$  has a linear relationship with *r* such that  $\sigma(r) = 0.5097r - 12.7142$ . The amplitude period parameter, *b*, also has a linear relationship, however, since *r* is at most order 3 in magnitude, we can regard it as a constant such that  $b(r) = 0.6629 \sim 2/3$ . The same goes for the fraction parameter,  $\alpha$ ; we can regard it as a constant such that  $\alpha(r) = -0.0345$ . For odd parameter fit statistics see Fig. 24a; **b** plots of  $A_0$  versus *r* (*left*) and *a* versus *r* (*right*). Both exhibit a similar pattern, however it is difficult to discern any nice relationships. For odd parameter plots see Fig. 24b

the distribution independently, in the quest to describe the entire  $h^{1,1} + h^{1,2}$  with a single function. The model we chose to describe  $h^{1,1} + h^{1,2}$  is the simplest possible Planckian model

$$f(x, A, n, b) = \frac{A}{x^n} \frac{1}{e^{b/(x-22)} - 1}$$
(2.6)

The parameter names in the fit results are the amplitude *A*, the power *n*, and some real constant *b*. The shift in *x*-axis is so that the distribution begins at 0 as the smallest  $h^{1,1} + h^{1,2}$  above the cutoff is 22. The choice of a Planckian model in the above form is greatly motivated by the blackbody distribution  $f(T, \lambda)$ . The *q* curves within  $h^{1,1} + h^{1,2}$  appear to behave in a manner analogous to the curves of constant *T* within the blackbody



**Fig. 11.** Three curves (q = 0, 18, 30) within the even  $h^{1,1} + h^{1,2}$  distribution. The transparent *grey data dots* are all the data plots for the distribution. Refer to Fig. 31 to see the same example for the classification of odd curves within the odd distribution

distribution. This is an initial trial. Later, we will discover additional structure in the distribution by trying to mimic the blackbody distribution exactly. It turns out that the general behavior of the distribution is modeled very well, cf. Fig. 12a.

Consider the maximum of each of the curves. As indicated in Fig. 12a, we can fit the maxima to a curve as indicated using the data plotted for the given values of q. From the above analysis, the  $h^{1,1} + h^{1,2}$  distribution behaves analogously to a blackbody spectrum—except for one small subtlety. It is in this subtlety that the added structure within  $h^{1,1} + h^{1,2}$  is observed.

Just as was seen in Fig. 2,  $h^{1,1} + h^{1,2}$  appears to split up into two smaller distributions based on the parity of  $h^{1,1} + h^{1,2}$ . One can then further break up both the even and odd distributions into three further sets. The manner we observed this added fine structure is again motivated by a blackbody spectrum. In a true blackbody distribution, the curves of constant *T* never overlap. However, if you consider the lines of best fit only, when looking at our distribution one sees an overlap of certain curves. For example, observe the following plot of curves which clearly cross in Fig. 12b.

It turns out that this overlapping occurs consistently to the point where one can classify the curves (defined by their q value) into residue classes  $q_n$  distinguished by  $n \mod 6$ . On the left hand side of the  $h^{1,1} + h^{1,2}$  axis, the curves are ordered with red (residue class  $q_2$ ) above yellow (residue class  $q_4$ ) above blue (residue class  $q_0$ ), whereas on the right hand side of the axis, the order is reversed. Similar behavior is observed in the odd distribution of  $h^{1,1} + h^{1,2}$  with the curves in the residue classes  $q_1$ ,  $q_3$ , and  $q_5$  (see Fig. 32b).

The clusters of curves constitute an entire set of mod 6 residue classes. These classes now define a set of curves which belong to very "nice" distributions that behave exactly like a blackbody distribution.<sup>1</sup> Compare, for example, a plot of the all the curves for even distribution of  $h^{1,1} + h^{1,2}$ , separated into their residue classes, Fig. 13

<sup>&</sup>lt;sup>1</sup> Of course  $h^{1,1} + h^{1,2}$  is not continuous. It is discrete. However, the structure of the best fit curve to the data points appears very similar to that of a continuous blackbody distribution.

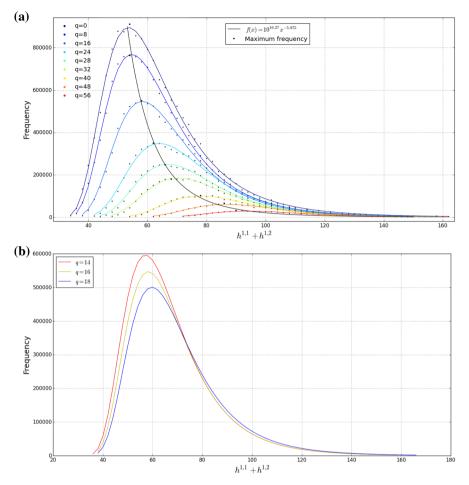
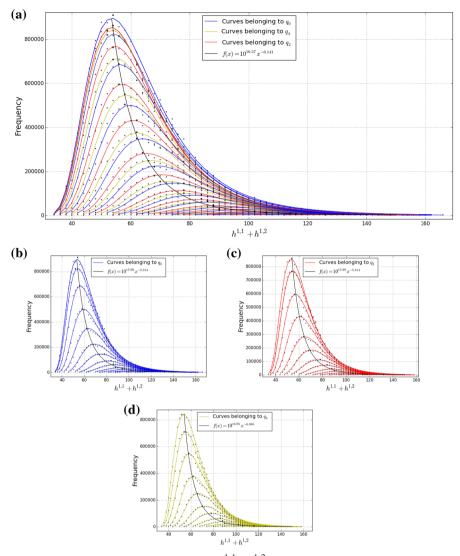


Fig. 12. In the attempt to describe the data analogously to a blackbody distribution (a), we discover some subtle structure (b). a Lines of best fit from a regression analysis for a few select curves. The *black data* points represent the maximum frequency for that particular q-curve. The *Black line* is a line of best fit to describe the points of maximum frequency—this is analogous to a blackbody spectrum. See Fig. 32a for the curves within the odd distribution. b The curves segregate into three classes determined by the value of the even integer modulo 6. A similar pattern occurs in the odd distribution; see Fig. 32b

As a first approximation we have successfully modeled the general trend of the data. There is, however, a fine structure to the individual data points that we would like to model. Introducing an oscillating term in the amplitude, as seen in the analysis of  $h^{1,1} - h^{1,2}$ , unfortunately did not seem to improve the fits.

Again, it appears that the least number of variables our functions can have is two, f = f(x, q). This function will be slightly different in the values of coefficients, depending on which residue class one is modeling.

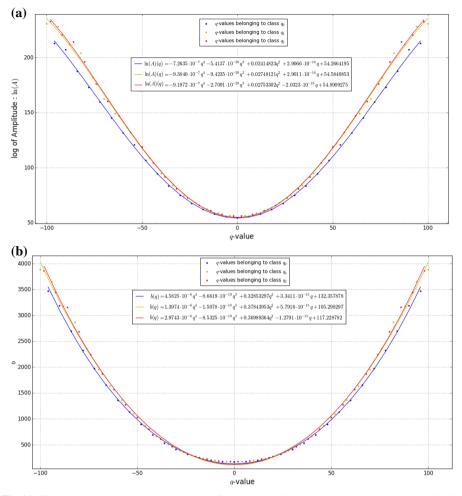
Just as for  $h^{1,1} - h^{1,2}$ , we wish to express the parameters for the  $h^{1,1} + h^{1,2}$  model (2.6) in terms of q. We therefore write A = A(q), b = b(q), n = n(q) and seek to find expressions for the coefficients.



**Fig. 13.** We illustrate the added structure for even  $h^{1,1} + h^{1,2}$  data, by displaying how the regression curves can be divided into residue classes. For the list of odd curves, refer to Fig. 33. **a** All the curves *color* coded according to what residue class their curves  $q_n$  belongs to. **b** Family of curves all belonging to  $q_0$ . **c** Family of curves all belonging to  $q_2$ . **d** Family of curves all belonging to  $q_4$  (color figure online)

While the x-axis of  $h^{1,1} + h^{1,2}$  has only positive q values—due to the fact the data points will overlap—when plotting them against the parameter values, we also have to consider the negative values of q. We present the various relationships (see Fig. 34 for the plots for the odd distribution of  $h^{1,1} + h^{1,2}$  analogous to Fig. 14).

Each distribution has an equation with different parameter values. However, the fact that we can express all the parameters in terms of q means we are able to get a generalized formula to describe the entire  $h^{1,1} + h^{1,2}$  distribution—as long as the frequency is above 2000. For succinctness we use the following notation for the coefficients



**Fig. 14.** The parameter plots are *color* coded according to what residue class their q value belong to. **a** Plotting the q-value parameter versus the  $\log(A)$  parameter. **b** Plotting the q-value parameter versus the b parameter. **c** Plotting the q-value parameter versus the power n parameter (color figure online)

$$A_{k,i}, n_{k,i}, b_{k,i},$$
 (2.7)

where the subscript k = 0, 1, 2, 3, 4, 5 refers to residue class  $q_k$ , and i = 0, 1, 2, 3, 4 refers to the coefficient of the  $i^{th}$  power of q. Thus, we have:

$$A_k(q) = \exp(\sum_{i=0}^4 A_{k,i}q^i), \quad n_k(q) = \sum_{i=0}^4 n_{k,i}q^i, \quad b_k(q) = \sum_{i=0}^4 b_{k,i}q^i, \quad (2.8)$$

where the matrix of coefficient values for  $A_{k,i}$ ,  $n_{k,i}$  and  $b_{k,i}$  can be found in "Appendix A.2.2".<sup>2</sup> Our function (2.6) now is able to approximately describe the entire  $h^{1,1} + h^{1,2}$  distribution:

<sup>&</sup>lt;sup>2</sup> Perhaps it is important to state explicitly—due to potential confusion—that the coefficients  $A_{k,i}$  refers to the natural logarithm of the amplitude values while  $A_k$  is the actual amplitude seen in the model.

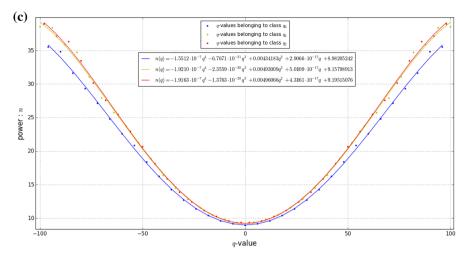


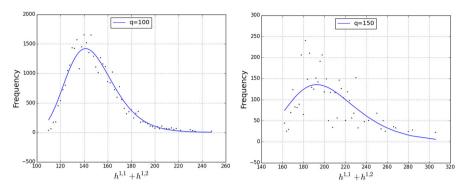
Fig. 14. continued

$$f_k(x,q) = \frac{e^{\sum_{i=0}^4 A_{k,i}q^i}}{x^{\sum_{i=0}^4 n_{k,i}q^i}} \frac{1}{\left(e^{\frac{\sum_{i=0}^4 b_{k,i}q^i}{(x-22)}} - 1\right)},$$
(2.9)

Of course there are certain constraints on the values of q. For a given k, q has to be an integer which falls within the residue class  $q_k$ . For even values of k, x = 2m, and for odd k, x = 2m + 1. We have m > 12.

A few comments about the analysis on the  $h^{1,1} + h^{1,2}$  distribution are in order.

- 1. The Planckian model used in (2.6) could be modified in some manner such that there is some oscillating behavior in the amplitude. Any kind of oscillatory term we introduce, only has a mild effect on the model's behavior. As the q values exceed 100, the model is not able to describe the data very well.
- 2. Assuming one adds an oscillatory component to the model, the module used in python to do the regression analysis called *lmfit* is sensitive to the initial conditions set by the user. Since the model is a custom model, it is difficult to find the correct initial conditions such that the best fit line oscillates close to every point (as with  $h^{1,1} h^{1,2}$ ).
- 3. It is possible that the model used does not have the features required to describe the oscillatory "up and down" behavior of the data points. The Planckian model was chosen in that the  $h^{1,1} + h^{1,2}$  distribution resembled a blackbody distribution.
- 4. In choosing a polynomial model for Fig. 14a–c, we picked the lowest order polynomial that gave the best fit. Choosing the order to be four for all the plots appeared to be convenient. However, it is apparent that the parameter relationship plot in Fig. 14b would be better described by a polynomial of order 6. One could use an order 6 polynomial for all the other relationships plots too, but doing so might not have any physical significance. One can achieve an arbitrarily good fit the larger the order of the polynomial used, but that does not necessarily mean the chosen model is the correct model.



**Fig. 15.** Left figure is the fitted model(blue line) for a q value of 100 and right has a q value of 150. As the q-value increases, the scattering of the data points within  $h^{1,1} + h^{1,2}$  increases to the point where the model works no longer. For an example of how the model begins to break down at large q, see Fig. 35 (color figure online)

2.3. The distribution of the Euler number. The Euler number for Calabi–Yau threefolds is

$$\chi = 2(h^{1,1} - h^{1,2}). \tag{2.10}$$

As mentioned previously, we are summing over all the various *r*-curves to obtained the full-Euler number distribution. A plot of  $\chi$  versus frequency yields the pseudo-Voigt distribution. In particular, we can model the behavior of the distribution almost perfectly using the modified pseudo-Voigt curve (2.11) and (2.12), which is repeated here for convenience:

$$f(x, A, \sigma, \alpha) = (1 - \alpha) \frac{A}{\sigma\sqrt{2\pi}} e^{\frac{-(x)^2}{2\sigma^2}} + \alpha \frac{A}{\pi} \left[\frac{\sigma^2}{x^2 + \sigma^2}\right],$$
(2.11)

where

$$A(x, A_0, a, b) = A_0 + a\cos(2\pi b \cdot x).$$
(2.12)

The results of the regression analysis for the Euler number distribution is presented in Fig. 16a.

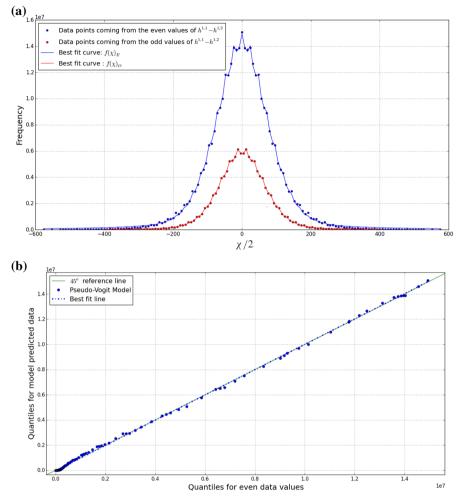
The fitted parameter values for  $f(\chi)_E$  corresponding to even values of  $h^{1,1} - h^{1,2}$  are:

$$(A_0, \sigma, \alpha, b, a) = (1.9032 \times 10^9, 75.8305889, 0.00718459, 0.58347826, 8.7427 \times 10^7).$$
(2.13)

Likewise, the fitted parameter values for  $f(\chi)_0$  corresponding to odd values of  $h^{1,1} - h^{1,2}$  are:

$$(A_0, \sigma, \alpha, b, a) = (7.6043 \times 10^8, 64.9735680, 0.00549425, 0.83357720, 3.6881 \times 10^7).$$
(2.14)

Although  $\chi$  is only even, the two curves originate from the fact that if you take  $\chi/2$  you get even and odd values. The two curves arise from the parity of  $\chi/2$  and are presented in Fig. 16a.



**Fig. 16.** Various plots illustrating the actual fit of the modified pseudo-Voigt model. We can tell we have a good fit by looking at the probability plots for the quantiles of the standard pseudo-Voigt distribution versus quantiles for the actual data. The  $R^2$  values in (**b**) and (**c**) are given relative to the line y = x. **a** The distribution of Euler numbers fitted to a modified pseudo-Voigt curve. The *blue* curve  $f(\chi)_E$  represents even values of  $\chi/2$ . The *red* curve  $f(\chi)_O$  represents odd values. **b** Probability plot for the even values of  $\chi/2$ . The model fits the data with  $R^2 = 0.99944$ . **c** Probability plot for the odd values of  $\chi/2$ . The model fits the data with  $R^2 = 0.99965$  (color figure online)

2.4. Goodness-of-fit. A goodness-of-fit test is implemented as a means of testing how well a given model describes some given data. Typically the model validation process consists of only quoting a single statistically generated number like the  $R^2$ ,  $\chi^2$  or p values. Based on the size of this number, one then makes inferences on how well the chosen model fits the observation. One needs to be careful however of misusing such indicators as an absolute measure for assessing goodness-of-fit.

For a structural equation model (SEM)—in our case, the modified pseudo-Voigt and Planckian models—this assessment is not so straight forward as it would be for a simple regression analysis. To quantify the predictive power of an SEM, a single statistical

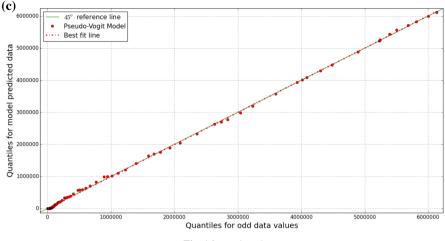


Fig. 16. continued

test does not suffice - in fact, there is no single test. According to [41], the best one can do is assess three different aspects of what it means to have a good fit, these are: overall fit, comparative fits to a test model and model parsimony.<sup>3</sup> The only real test available is the chi-squared ( $\chi^2$ ) test, when it comes to overall fit, this  $\chi^2$  statistic is the most popular test. The  $\chi^2$  test compares observed and predicted correlation matrices with each other, and so, statistical significance is evaluated based on the value of  $\chi^2$ . A large  $\chi^2$  value signifies a considerable difference between the correlation matrices. A low value indicates there is little statistical difference between matrices. Since the  $\chi^2$  test is between actual and predicted matrices only, when looking for overall fit, one searches for non-significant differences between the correlation matrices. Often, rather than presenting the  $\chi^2$  or  $\chi^2_R$  (the chi-squared value relative to the degrees of freedom for the model) value, a p value is given instead. The p value, in a way, informs us whether one should reject a null hypothesis or not. A small *p*-value suggests that the differences in observed versus predicted are too large to be consistent with the null-hypothesised model i.e. assuming the null-hypothesised model, the probability of observing what we did is relatively small, suggesting either an absolutely fluke experimental outcome or an incorrect model null-hypothesis. The p-values can be determined by a p-value calculator by inputting the  $\chi_R^2$  value. There is no standard way of choosing a significance level for the *p*-value, but typically p < 0.05 is considered statistically significant.

In general, statistical non-significance given by appropriate values of the  $\chi^2$  fit statistics is adequate. However, one must be careful of drawing similar conclusions for structural equation modeling. The fit statistic makes a statement of the correlation matrices only, not about whether or not the correct model is identified. This is largely due to the sensitivity to sample size of the  $\chi^2$  test. In our analysis, the sample size (number of reflexive polytopes) is enormous—almost one billion! For large samples (> 200) the  $\chi^2$  test will give significant differences for any model used. This sensitivity to a sample size, together with an *effect size* and *alpha value*, is related to what one calls the power of a test - the probability of not incorrectly accepting a null hypothesis that is actually false.

 $<sup>^3</sup>$  Parsimony refers to the ability of a model to give a certain degree of fit whilst having the least required number of predictor variables.

Without worrying too much about what an effect size and alpha value is; for any alpha value, the greater the sample size, the greater the power of the statistical test. However, increasing the sample size beyond a certain amount, can result in the test having "too much" power.<sup>4</sup> Perceived effects in very large sample sizes, will always become significant.<sup>5</sup> Observe how in Figs. 29 and 36 the  $\chi^2_R$  values for all the different curves is extremely large, naively indicating that we have a horrible fit—which would be an incorrect conclusion.

It is clear from the above discussion that we cannot use the  $\chi^2$  or p values in validating our choice in model. What is not so clear, is the additional subtlety in using purely statistical means to asses goodness-of-fit for our data. This subtlety lies at the heart of almost all statistical tests—the construction of a null hypothesis. The term frequency, as used in the statistical sense, refers to the number of outcomes for a certain event. The measurement of this outcome will often have certain known or unknown factors affecting it. These tests check for the probability that the errors found are too significant to be solely due to random variations in the data. For example, assume that statistical tests give non-significant results. If the residuals are small enough to be considered random errors in the measurement of the frequency, we could say that the model is appropriate. If however, the residuals are too large or present additional structure, we could say the model is good, but not quite the correct one as the residual errors are not "random enough". In our case, there is no notion of measured frequency and error in measurement of frequencies. Our frequencies are generated as a result of a combinatoric calculation. Statistical tests assume that the input is from measurement and observations (obeying some null-hypothesis), thus they are inherently constructed with this notion in mind. By inputting our data, the tests are trying to calculate something from a data set which does not obey the very assumption they use in their calculations. We are not exactly clear how much this affects statistical outcomes, but it is important to keep in mind.

How do we validate then, that our chosen models are a good fit, or that our model is the best one at describing the data? We implement graphical methods. The first graphical method is obviously through pure inspection—this is not quite statistically quantifiable. There is a statistically based graphical method to asses goodness-of-fit called probability plots, Q-Q plots or P–P<sup>6</sup> plots. These plots were initially constructed to test the "normality" of a data set when the sample size is too large to depend on the  $\chi^2$  and p values. In principle, a standard probability plot tells you the likelihood that the a sample's distribution of data obeys a normal distribution—hence checking for normality. The answer to the question is not given by a statistical value, but rather by a graphical representation from which one can extract statistical numbers. If the plotted data on this probability plot is a straight line, then we can determine that the sample set is normally distributed.

We can extend this concept further: we can take two different samples, and take a probability plot to determine if two data sets come from populations with a common distribution. Such a probability plot is referred to as a Q–Q (quantile–quantile) plot. Extending this concept one more time—as for our use—we will take the quantiles of our theoretical distribution (the modified pseudo-Voigt and Planckian profiles) as our

<sup>&</sup>lt;sup>4</sup> Power is the probability that you do detect deviations from your null-hypothesised model, when the null-hypothesised model is, in fact, incorrect.

<sup>&</sup>lt;sup>5</sup> Conversely is also true, for extremely small sample sizes, any effect which should be significant, becomes insignificant.

<sup>&</sup>lt;sup>6</sup> A P–P plot is the plot of the cumulative distribution frequency of the one data set against the CDF of the other. P-P plots are not as useful as Q–Q plots, thus are seldom used.

"first sample" and plot them against the quantiles of our data as our "second sample"; this will give us our probability plot. In all the probability plots, it is the quantiles of the respective data sets which are plotted against each other.

Quantiles are basically just a generalization of quartiles. For example, the *k*th percentile of a set of values divides them, such that the number of values which lie below is k%, and the number of values which lie above is (100 - k)%. The 25th percentile is the lower quartile or the  $\frac{1}{4}$  quantile. Quantiles are the same as percentiles, but indexed by sample fractions rather than by sample percentages. Suppose that  $p \in [0, 1]$ , the aim is to find the value that is the fraction p of the way through the ordered data set. As an example, if  $p = \frac{1}{2} = 0.5$ , we want to know what is the value that sits at p = 0.5 of the way through i.e. half way. The value that sits there (this value may have to be interpolated) will be called the quantile for the fraction p = 0.5. There are many different algorithms for generating the quantiles for a given data set, we use python to generate the quantiles in a manner similar to that discussed above. For an ordered data set,  $x_1 \le x_2 \le x_1 \ldots \le x_{n-1} \le x_n$ , the most common way of calculating quantiles is to first compute the empirical distribution function:

$$F(x) = \frac{1}{n} \sum_{i=1}^{n} = 1(x_i \le x), \quad x \in \mathcal{R},$$
(2.15)

and then define the quantile function to be the inverse of F(x):

$$F^{-1}(p) = \min\{x \in \mathcal{R} : F(x) \ge p, \ p \in (0, 1)\}.$$
(2.16)

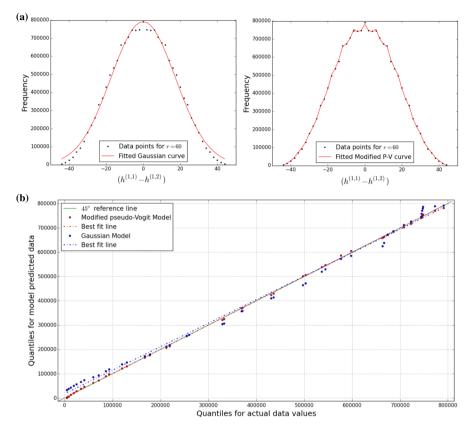
By generating the quantiles of some theoretical model and comparing them to the quantiles of a given data set of equal length, one can determine if the data set belongs to the same distribution as the data set belonging to the theoretical model—i.e., does the data fit the model. If the quantiles are roughly equal the plots will all be more or less on a straight line.

In probability plots:

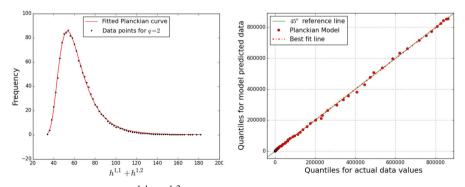
- 1. The length of data set needs to be equal. For unequal lengths, one must perform an interpolation of data.
- 2. If two identical data sets were compared to one another, the points would lie exactly on a 45 degree line. Thus, for two different data sets, the deviation from this reference line determines the likelihood that the sets belong to similar distributions. To quantify this likelihood, one can calculate the  $R^2$ -value of the data, relative to the y = x reference line.
- 3. Q–Q plots are not only limited to determining similarity in data sets. By analyzing the deviations which occur, one can determine how the scale and location of the data is shifted the data would follow some line y = mx + c, where m, c would be the estimates of these shifts in scale and location. Also, from the distribution of points above or below the reference line, one can infer aspects of the tails and skewness in the data.

Consider the following curves for the  $h^{1,1} - h^{1,2}$  distribution with r = 60 in Fig. 17a, b.

For the  $h^{1,1} + h^{1,2}$  distribution we just plot the data of q = 2 together with the corresponding probability plot in Fig. 18.



**Fig. 17.** Using probability plots, we are able to statistically see which model provides the better fit. We employ such graphical methods as standard goodness-of-fit tests such as the  $\chi^2$  fail to give meaningful results. **a** Best fit curve for r = 60 based on the *left* Gaussian model, *right* modified pseudo-Voigt model. **b** Probability plot for Fig. 17a. The x-axis represents the quantiles for the actual data, the y-axis represents the theoretically predicted quantiles—dependent on the model chosen (*red* modified pseudo-Voigt model ( $R^2 = 0.99974$ ); *blue* Gaussian model ( $R^2 = 0.99334$ ). The  $R^2$  values are not relative to the best fit lines, but are relative to the 45° reference line y = x. The closer the  $R^2$  value is to 1, the more similar the predicted quantiles are to the actual ones, thus, the better the model describes the data (color figure online)



**Fig. 18.** Left best fit curve of  $h^{1,1} - h^{1,2}$  distribution for curve q = 2 based on the Planckian model. Right probability plots of our fitted theoretical Planck model versus the q = 2,  $h^{1,1} - h^{1,2}$  distribution

In its current form, the probability plots do not allow us to calculate *p*-values of the various models. This due to the same issue encountered previously. If one however standardizes the data according to the *Z*-standardization:

$$Z = \frac{X - \mu}{\sigma},\tag{2.17}$$

where  $\mu$  and  $\sigma$  are the mean and standard deviation, it is possible to calculate the *p*-values since the magnitude of each sample gets rescaled. The probability plot of all the models is displayed in the "Appendix", with the relative *p*-values for each model— Fig. 25g, h. What we see is that the modified pseudo-Voigt is statistically the model which provides the best fit.

2.5. Implications for physics. Calabi–Yau threefold compactifications of string theory have been the traditional approach to obtaining interesting phenomenological models. The plethora of geometries and configurations, ranging from heterotic strings on Calabi–Yau threefolds endowed with stable bundles, to D-brane probes on local Calabi–Yau varieties, to F-theory compactification on elliptic fibrations, has over the years justified the landscape and inspired various statistical analyses of the space of vacua.

Of particular interest has been the investigation of further structures in the Kreuzer-Skarke database, including identification of "the tip" where Hodge numbers are small [21,35,46], the top bounding curves where Hodge numbers are large [43], identifying elliptically fibered threefolds [28,29,42,44], finding further fibrations such as K3-fibers [33,45], or a step-by-step construction of all possible smooth Calabi–Yau hypersurfaces from the reflexive polytope data [19], etc. Now, it should be emphasized that each of the some 473 million reflexive polytopes admits, as an ambient toric variety, many<sup>7</sup> so-called maximal projective crepant partial (MPCP) desingularization, each of which gives rise to a different Calabi-Yau threefold. Therefore, the actually number of Calabi-Yau threefolds from the Kreuzer–Skarke database is many orders of magnitude larger than  $10^{10}$ . While manifolds coming from the same reflexive polytope have different geometrical data such as triple intersection numbers, which in the standard embedding in heterotic compactification correspond to Yukawa couplings, they do share the same Hodge numbers because these, by virtue of (2.1), depend only on the combinatorics of the polytope. We need to wait for significant theoretical and/or computational advances to have the full data of the Hodge pairs in view of the Calabi–Yau manifolds themselves, which might give new statistics. It would be perhaps even more interesting if the statistics remain largely the same, thereby hinting at some universality in the distribution of such topological data.

In the context of the recent works on F-theory, it is an important fact the vast majority of the Kreuzer–Skarke threefolds are elliptic fibrations over some complex surface, and in fact birational to [42,44,45] a Weierstrass model. For example, some 10<sup>6</sup> alone [42] come from elliptic fibrations over  $\mathbb{P}^2$ . Therefore the Kreuzer–Skarke dataset is directly relevant to F-theory. In the more classical context of heterotic strings, the Hodge numbers dictate the number of (anti-)generations in the standard embedding. In our above plots, the Euler number  $\pm 6$  indicate the three generation models. The generic paucity of  $\chi = \pm 6$  manifolds led to the industry of non-standard embedding where extra vector bundle and Wilson line information is needed. The advantage of F-theory models is that

<sup>&</sup>lt;sup>7</sup> The actual numbers are not yet known, but even up to  $h^{1,1} = 7$ , we already see from tens to thousands and with the number increasing potentially exponentially as we go up in Hodge number [19].

the compactification data comes only from the Calabi–Yau manifold. In particular, the intersection theory of the cycles and fiber-degeneration structure determine the gauge group, anomaly cancellation, matter content, and Yukawa couplings. Much of this can be extracted from the polytope data.

F-theory compactifications on threefolds, resulting in six dimensional gauge theories have been considered from the point of view of systematically classifying the base complex surfaces [44] and the statistics have been performed therein. Non-toric bases were considered and a number of Calabi–Yau threefolds beyond the Kreuzer–Skarke data were found. It is remarkable that the overall distribution of Hodge numbers remains largely unchanged. Indeed, in unpublished work of Kreuzer–Skarke, where they extended the hypersurface in toric fourfolds to double hypersurfaces in fivefolds, obtaining some 10<sup>10</sup> more manifolds and the shape of Fig. 1 persists. All these point to the Kreuzer–Skarke data being a robust representative in the space of Calabi–Yau threefolds. Our distribution subsequently seems a representative sample, and we speculate that analyses of string vacua, in any context, should be thus weighted. For example, in study of the "typical" number of generations in four dimensional heterotic compatification, or of charged matter in six dimensional F-theory compactification, one should superpose our pseudo-Voigt profile.

#### 3. Calabi-Yau Twofolds: K3 Surfaces

As noted in the Introduction, there are 4319 data points, corresponding to hypersurfaces as Calabi–Yau twofolds, i.e., K3 surfaces, in reflexive three dimensional polytopes. Being algebraic K3 surfaces, there is only one relevant topological invariant, the Hodge number,  $h^{1,1} = 19$ . However, there is a further refined algebraic quantity for the K3 surface X, the rank of the Neron–Severi lattice  $H^2(X; \mathbb{Z}) \cap H^{1,1}(X)$ , which is the **Picard Number**  $\rho(X)$  and which enumerates the number of divisors on the surface up to algebraic equivalence. The Picard numbers of the 4319 K3 surfaces were computed in [12]. We present the distribution thereof in Fig. 19a.

We only used the standard pseudo-Voigt profile as the modified one did not change the fit significantly. Here are the fit statistics for best fit curve:  $(A, \mu, \sigma, \alpha) = (4517.45, 10.76, 2.97, -0.031)$ , as shown in Fig. 19.

What is interesting about Fig. 19a is that the "oscillations" of the actual data points above and below the modeled curve is very apparent, yet modifying the pseudo-Voigt profile is unable to give any significant improvement. This leads to two potential conclusions: (a) the pseudo-Voigt profile is not the best profile to use in combination with an oscillatory component; (b) the manner in which the oscillations occur is not so straight forward as introducing simple cosine function. An interesting exercise would be to superimpose a cosine function along the distribution, by rotating it as one traverses the profile. As long as the wavelength, amplitude and angle of rotation are all small enough, the continuously rotated cosine function should remain a function everywhere along the profile.

#### 4. Calabi-Yau Fourfolds

The analysis of the four fold data is performed in the same spirit as the threefold data. We aim to look for patterns in the frequency plots. Due to complex conjugation and Poincaré duality, the only topological invariants of fourfolds that vary are  $h^{1,1}$ ,  $h^{1,2}$ ,  $h^{1,3}$ , and  $h^{2,2}$ . Three of these are independent [15]:

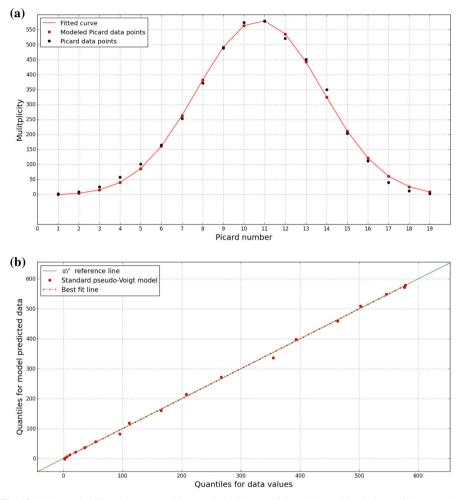


Fig. 19. Using probability plots, we are able to statistically see which model provides the better fit. We employ such graphical methods as standard goodness-of-fit tests, such as the  $\chi^2$  test, fail to give meaningful results. **a** For K3 surfaces, the multiplicity is plotted against Picard number with a pseudo-Voigt fit. **b** Probability plot for the multiplicity quantiles versus the fitted standard pseudo-Voigt quantiles. The  $R^2$  value is 0.99908

$$h^{2,2} = 44 + 4h^{1,1} - 2h^{1,2} + 4h^{1,3}.$$
(4.1)

We compiled a database for the frequency of the triplets  $(h^{1,1}, h^{1,2}, h^{1,3})$  to then obtain the following data structure

$$(h^{1,1}, h^{1,2}, h^{1,3}, f)$$

Since one expects mirror symmetry within the invariants  $(h^{1,1} \pm h^{1,3})$  [40], a plot of  $h^{1,1} - h^{1,3}$  against  $h^{1,1} + h^{1,3}$  (Fig. 20) should be symmetric about the line  $h^{1,1} - h^{1,3} = 0$ . Doing a quick analysis of the data yields the following observations: only partial mirror symmetry is found. For 69.77% of data points, the point  $(h^{1,1} - h^{1,3}, h^{1,1} + h^{1,3})$  is accompanied by the point  $(-h^{1,1} + h^{1,3}, h^{1,1} + h^{1,3})$ . Taking frequency into account,

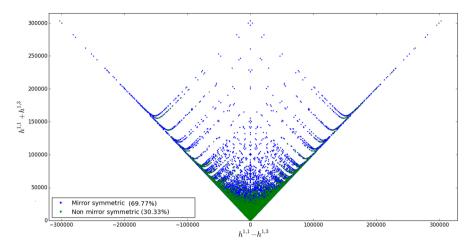


Fig. 20. The *blue points* correspond to manifolds with a mirror symmetric counterpart in the data set (color figure online)

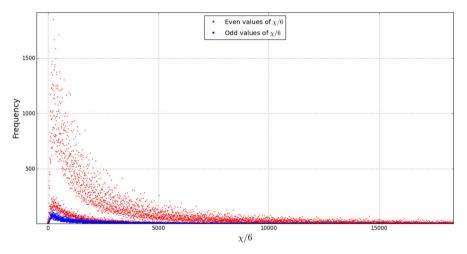


Fig. 21. Frequency of Calabi-Yau fourfolds with a given Euler number

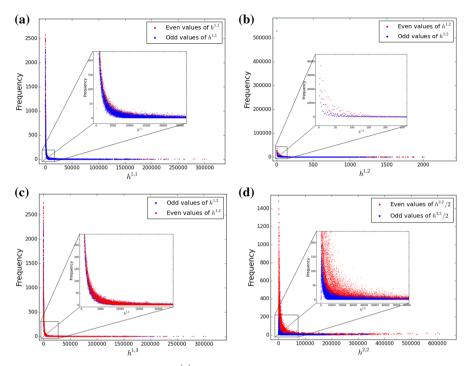
the percentage drops to 27.35%—see Fig. 37 in the "Appendix". This is most likely due to an incomplete data base.

For now, we have performed a primary analysis on the Euler distribution only. The Euler number for fourfolds is [15]:

$$\chi = 6(8 + h^{1,1} - h^{1,2} + h^{1,3}). \tag{4.2}$$

Interestingly enough, the distinction between even and odd distributions persist in the fourfold data base. For illustrative purposes, we show the distribution of  $\chi/6$  against frequency.

It is not immediately clear what is the reason for the gap, presumably it could be a cluster of data points which is missing from the data base. Until one obtains the complete fourfold data base of Hodge numbers, one can't say much else. We also preset plots of the individual Hodge numbers  $h^{i,j}$  versus frequency.



**Fig. 22.** The frequency for all the hodge  $h^{i,j}$  numbers. *Red points* and *blue* are odd and even points respectively for the various Hodge numbers. The data points are very dense close to the origin making it difficult to properly illustrate the mixing of odd and even Hodge numbers. Only  $h^{2,2}$  **c** has a clear separation between of an even values. **a**  $h^{1,1}$  versus frequency. **b**  $h^{1,2}$  versus frequency. **c**  $h^{1,3}$  versus frequency. **d**  $h^{2,2}$  versus frequency (color figure online)

#### 5. Conclusions and Outlook

By examining the distribution of Hodge numbers of Calabi–Yau manifolds of complex dimension two, three and four, realized as hypersurfaces in toric varieties of one higher dimension as constructed by Kreuzer and Skarke based on the results of Batyrev and Borisov, we have found many hithertofore undiscovered patterns. We summarize our key points as follows.

- For threefolds, there are 30108 distinct pairs of Hodge numbers  $(h^{1,1}, h^{1,2})$  from 473800776 reflexive polytopes, the frequency of both the half-Euler number  $h^{1,1} h^{1,2}$  and the sum  $h^{1,1} + h^{1,2}$  are distributed according to whether the value is odd or even;
  - The half-Euler number  $h^{1,1} h^{1,2}$  follows a modified pseudo-Voigt distribution

$$f(x) = (1 - \alpha) \frac{A'}{\sigma\sqrt{2\pi}} e^{\frac{-(x)^2}{2\sigma^2}} + \alpha \frac{A'}{\pi} \left[\frac{\sigma^2}{x^2 + \sigma^2}\right]$$

where the modification is made in the amplitude A of the distribution, such that

$$A' = A_0 + b\cos(2\pi \cdot b).$$

There is fine periodic substructure in terms of curves indexed by an integer r. Our model is accurate for low r-values ( $r \in [36, 110]$  and  $r \in [37, 99]$ ); using probability plots as test for goodness of fit, this modified pseudo-Voigt model is indeed the best one out of several standard candidates (cf. Fig. 29 for all the  $R^2$  and p values).

Among  $A, \sigma, \alpha, b, a$ , the parameters  $\sigma, b, \alpha$  have a strong linear relationship with r:

Even r	Odd r
$\sigma(r) = 0.5097r - 12.7142$	0.51379r - 13.2494
$\alpha(r) = 2 \times 10^{-4} r - 0.0345$	$2.25 \times 10^{-4} r - 0.0388,$
$b(r) = 3.7299 \times 10^{-5}r + 0.6629$	$7.9101 \times 10^{-5}r + 0.65956$

For a small subset of curves with a low *r*-value and an appropriate cut-off frequency, it is extraordinary that the model *exactly fits the data*. That is, it appears that the number of data points for each curve required, such that the model will result in a perfect fit is: 7 for even *r*-valued curves and 10 for the odd valued *r*-curves, see Fig. 30.

- The quantity  $h^{1,1} + h^{1,2}$  follows a Planckian distribution

$$f(x) = \frac{A}{x^n} \frac{1}{e^{b/(x-22)} - 1}$$

There is a substructure of curves, indexed by an integer q, each Planckian and with some periodic behavior. The curves  $q_n$  appear clustered into groups of residue classes distinguished by  $n \mod 6$ , and the parameters  $\log(A)$ , n, b all have extremely strong relationships with the q value.

By substituting this relationship into the model, we have a function  $f_k(x, q)$  that approximately describes the entire  $h^{1,1} + h^{1,2}$  distribution up to a q value of 69, 100:

$$f_k(x,q) = \frac{e^{\sum_{i=0}^4 n_{k,i}q^i}}{x^{\sum_{i=0}^4 n_{k,i}q^i}} \frac{1}{\left(e^{\frac{\sum_{i=0}^4 b_{k,i}q^i}{(x-22)}} - 1\right)},$$
(5.1)

with  $k = 0, 1, \dots 5$  and the coefficients given in A.8, A.9, A.10.

- The Euler number  $\chi = 2(h^{1,1} - h^{1,2})$  follows the modified pseudo-Voigt distribution composed with a sinusoidal  $A + A_0 + a\cos(2\pi b \cdot x)$  which is almost an exact fit, with the coefficients given by  $(A_0, \sigma, \alpha, b, a) = (1.9032 \times 10^9, 75.8305889, 0.00718459, 0.58347826, 8.7427 \times 10^7)$ , at  $R^2 = 0.99944$  for even  $\chi$  and

 $(1.9032 \times 10^9, 75.8305889, 0.00718459, 0.58347826, 8.7427 \times 10^7)$  at  $R^2 = 0.99965$  for odd  $\chi$ ,

The modified pseudo-Voigt distribution is remarkably accurate in predicting the overall and fine sub-structure of the Euler number distribution.

- For K3 surfaces, we have looked at the distribution of the multiplicity with Picard number. We find that this distribution follows a standard pseudo-Voigt profile. Adding in the sinusoidal modification does not significantly increase the overall fit. The parameters are given by  $(a, \mu, \sigma, \alpha) = (4517.45, 10.76, 2.97, -0.031)$  with  $R^2 = 0.99908$ .
- For Calabi–Yau fourfolds, there is no exact mirror symmetry, due to incompleteness of available data. Nevertheless, by breaking up the data into three groups, we have
  - Mirror symmetric partners with the same frequency: 27.35%

- Mirror symmetric partners without the same frequency: 42.22%
- Non mirror symmetric partners: 30.33%

By plotting the various  $h^{i,j}$  versus frequency we see there is no distinction between even and odd data values for  $h^{i,j}$ , expect for  $h^{2,2}/2$ . This distinction is carried out further in the Euler number distribution where odd points are clustered on a band with much lower frequencies. The even values of  $\chi/6$  appear to be distributed along to separate bands.

It is remarkable how well the pseudo-Voigt distribution, modified with a sinusoidal component, fits the distribution of topological numbers of toric Calabi–Yau manifolds, often giving an exact fit. Of course, what we are studying at heart is the number of integer points inside (cf. (2.1)) reflexive polytopes. This is a highly non-trivial counting problem whose answer will ultimately give full analytic results for our distributions and we suspect that the answer should be some generalized pseudo-Voigt function.

Now, in addition of Calabi–Yau manifolds, stable vector bundles over various such manifolds in a variety of construction beyond Kreuzer–Skarke have also been studied algorithmically over the years in the context of heterotic compactification (cf. e.g., [23–26]). One can see a somewhat pseudo-Voigt profile in these as well, even though there is no underlying polytope and the counting problem is dictated by certain Diophantine system. It would be interesting to see why this shape is universal in such classifications.

Acknowledgements. We are grateful to Cyril Matti for collaboration during the early stages of this project. We thank Mark Dowdeswell for his input with regards to the goodness-of-fits for the various plots. YHH is indebted to the Science and Technology Facilities Council, UK, for grant ST/J00037X/1, the Chinese Ministry of Education, for a Chang-Jiang Chair Professorship at NanKai University, and the city of Tian-Jin for a Qian-Ren Award. YHH is also perpetually indebted to Merton College, Oxford for continuing to provide a quiet corner of Paradise for musing and contemplations. VJ and LP are supported by the South African Research Chairs Initiative of the Department of Science and Technology and the National Research Foundation.

**Open Access** This article is distributed under the terms of the Creative Commons Attribution 4.0 International License (http://creativecommons.org/licenses/by/4.0/), which permits unrestricted use, distribution, and reproduction in any medium, provided you give appropriate credit to the original author(s) and the source, provide a link to the Creative Commons license, and indicate if changes were made.

#### A. Appendix

Here we include all additional plots to supplement the main body. This includes the relevant plots for the odd distributions—since in the main text we only presented the plots for even distributions—as well as the regression analysis statistics and parameter values for both distributions.

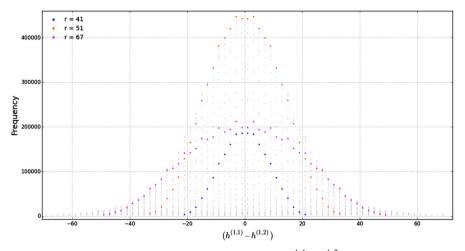
A.1. Supplementary plots for the  $h^{1,1} - h^{1,2}$  distribution. All even plot counterparts will be referenced in the figures. The plots appear in the same order as in the main body, with descriptions only if necessary.

#### A.1.1. Plots for the odd distribution as counterparts to the even ones

A.1.2. Comparative plots Here we present a comparison of various models we used, by plotting them side by side with the relevant fit-statistics. We choose a single even curve, r = 54, and odd curve, r = 51, to illustrate the difference between models.

#### **Gaussian Model**

$$f(x, A, \mu, \sigma) = \frac{A}{\sigma\sqrt{2\pi}}e^{-(x-\mu)^2/2\sigma^2}$$
 (A.1)



**Fig. 23.** Three highlighted curves (r = 41, 51, 67) within the odd  $h^{1,1} - h^{1,2}$  distribution. The transparent grey data dots is the rest of the distribution. Refer to Fig. 4 for the even plot

Lorentzian Model

$$f(x, A, \mu, \sigma) = \frac{A}{\pi} \left[ \frac{\sigma}{(x - \mu)^2 + \sigma^2} \right]$$
(A.2)

Pearson7 Model

$$f(x, A, \mu, \sigma, m) = \frac{A}{\sigma\beta(m - \frac{1}{2}, \frac{1}{2})} \left[ 1 + \frac{(x - \mu)^2}{\sigma^2} \right]^{-m},$$
 (A.3)

where  $\beta$  is the Beta function.

#### **Breit-Wigner Model**

This model is based on the Breit-Wigner function.

$$f(x, A, \mu, \sigma, t) = \frac{A(t\sigma/2 + x - \mu)^2}{(\sigma/2)^2 + (x - \mu)^2}$$
(A.4)

Voigt Model

$$f(x, A, \mu, \sigma, \gamma) = \frac{a \operatorname{Re}[(z)]}{\sigma \sqrt{2\pi}}$$
(A.5)

where

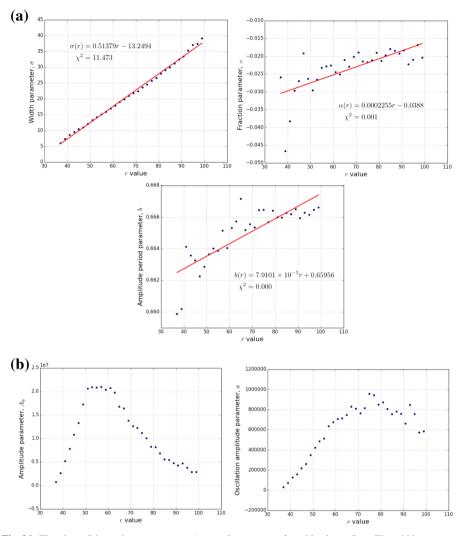
$$z = \frac{x - \mu + i\gamma}{\sigma\sqrt{2}}, \quad w(z) = e^{-z^2} \operatorname{erfc}(-iz)$$
(A.6)

The Voigt model is a convolution of the Gaussian and Lorentzian models.

#### **Pseudo-Voigt Model**

$$f(x, A, \mu, \sigma, \alpha) = (1 - \alpha) \frac{A}{\sigma\sqrt{2\pi}} e^{\frac{-(x-\mu)^2}{2\sigma^2}} + \alpha \frac{A}{\pi} \left[ \frac{\sigma^2}{(x-\mu)^2} + \sigma^2 \right]$$
(A.7)

We present the standardized and shifted probability plots for the above comparisons:



**Fig. 24.** The plots of the various parameters A,  $\sigma$ ,  $\alpha$ , b, a versus r for odd values of r. **a** The width parameter  $\sigma$  has a linear relationship with r such that  $\sigma(r) = 0.51379r - 13.2494$ . The amplitude period parameter, b, also has a linear relationship, however, since r is at most order 3 in magnitude, we can regard it approximately as a constant such that  $b(r) = 0.65956 \sim 2/3$ . The same goes for the fraction parameter, $\alpha$ , we can regard it as a constant such that  $\alpha(r) = -0.0388$ . For even parameter fit statistics see Fig. 10. **b** Plots of  $A_0$  versus r (*left*) and a versus r (*right*). Both exhibit a similar pattern, however it is difficult to find any nice relationships. For even parameter plots see Fig. 10

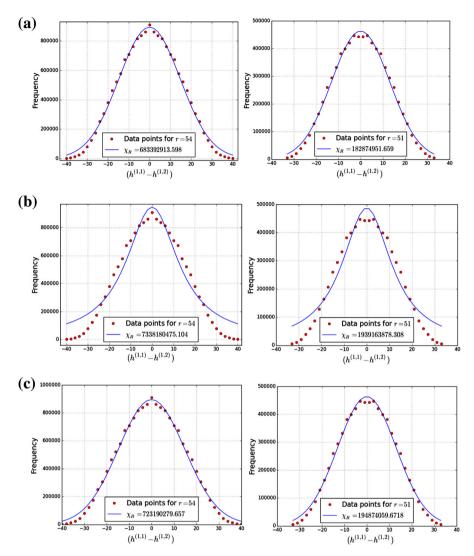


Fig. 25. For all models, the *left* hand graph is for r = 54 and the right is for r = 51. The probability plot presents all the models together. All the above mentioned modeled are included to compare their resemblance with the actual data. The larger the *p* value the better the line y = x fits the data, implying the better the model is at describing the data. **a** Gaussian model. **b** Lorentzian (Cauchy) model. **c** Pearson7 model. **d** Breit–Wigner model. **e** Voigt model. **f** Pseudo-Voigt model. **g** The probability plot for r = 51. **h** The probability plot for r = 54

A.1.3. A first approximation to the data The overall behavior of the data across each curve is modeled extremely well using the pseudo-Voigt model. Here we present a few plots illustrating a first approximation to the data. A second approximation can be made by introducing an oscillating amplitude as described in Sect. 2.1

A.1.4. Table of parameter values and statistics Here we present the parameter values as well as the reduced  $\chi$  value,  $\chi_R$ , in a tabular format for all even r curves— $r \in [34, 120]$ —and for all odd r curves— $r \in [35, 99]$ .

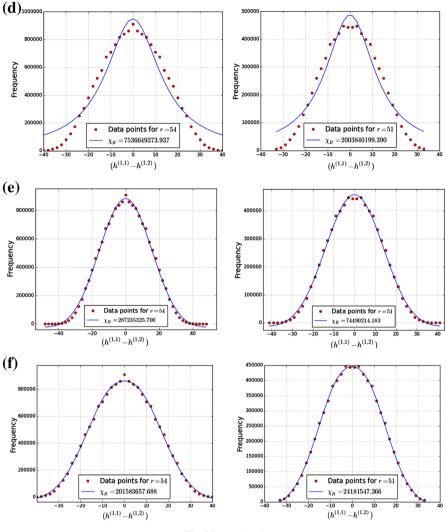
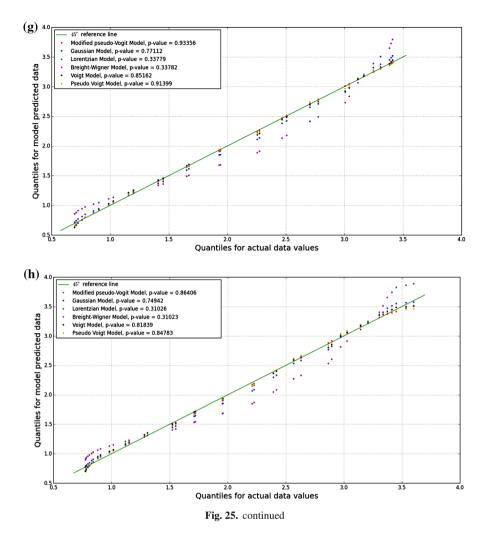


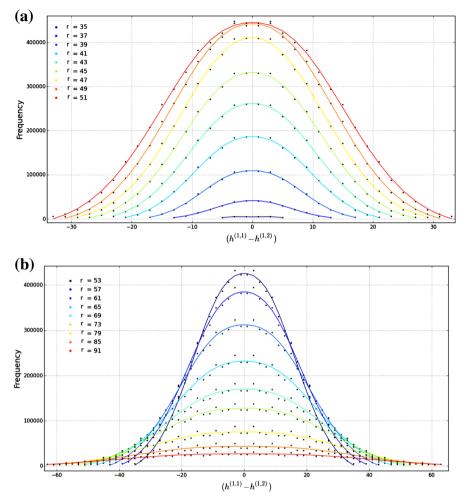
Fig. 25. continued

A.2. Supplementary plots for the  $h^{1,1} + h^{1,2}$  distribution.

A.2.1. Plots for the odd distribution as counterparts to the even ones All even plot counterparts will be referenced in the figures. The plots appear in the same order as in the main body, with descriptions only if necessary.



A.2.2. Table of parameter values, coefficient values and statistics Coefficient values for the description of the entire  $h^{1,1} + h^{1,2}$  distribution



**Fig. 26.** Best fit curve based on the pseudo-Voigt model for the same sets of curves as seen in Fig. 5. **a** Regression lines for few select even r values, with  $r \in [35, 51]$ . **b** Regression lines for few select even r values, with r > 51

$$A_{k,i} = \begin{pmatrix} 54.2664195 & 2.9066 \times 10^{-16} & 0.02414823 & -5.4137 \times 10^{-20} & -7.2635 \times 10^{-7} \\ 65.0676835 & -2.0296 \times 10^{-16} & 0.03354614 & 3.7552 \times 10^{-19} & -3.1443 \times 10^{-7} \\ 54.8909275 & -2.0323 \times 10^{-16} & 0.02753302 & -2.7091 \times 10^{-20} & -9.1972 \times 10^{-7} \\ 62.6423777 & 1.2736 \times 10^{-16} & 0.03020535 & -1.1234 \times 10^{-19} & -8.6929 \times 10^{-7} \\ 54.5840853 & 2.9011 \times 10^{-16} & 0.02748121 & -9.4235 \times 10^{-20} & -9.3840 \times 10^{-7} \\ 64.2001359 & -1.3980 \times 10^{-16} & 0.03700128 & 8.3795 \times 10^{-20} & -1.3712 \times 10^{-7} \end{pmatrix}$$

$$(A.8)$$

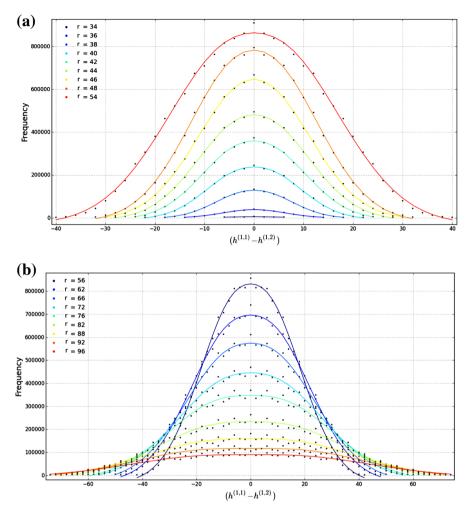


Fig. 27. Best fit curve based on the pseudo-Voigt model for the same sets of curves as seen in Fig. 6. **a** Regression lines for few select even r values, with  $r \le 54$ . **b** Regression lines for few select even r values, with r > 54

$$b_{k,i} = \begin{pmatrix} 132.357878 & 3.3411 \times 10^{-15} & 0.32753297 & -8.6619 \times 10^{-19} & 4.5825 \times 10^{-6} \\ 184.853063 & -5.7999 \times 10^{-17} & 0.31981034 & 1.0014 \times 10^{-18} & 3.9052 \times 10^{-5} \\ 117.228782 & -1.2791 \times 10^{-15} & 0.36989364 & -8.5325 \times 10^{-20} & 2.9743 \times 10^{-6} \\ 173.033950 & -1.1829 \times 10^{-15} & 0.31584408 & 8.9872 \times 10^{-19} & 2.5454 \times 10^{-5} \\ 105.298297 & 5.7916 \times 10^{-15} & 0.37843953 & -1.5078 \times 10^{-18} & 1.3974 \times 10^{-6} \\ 171.521189 & 1.5811 \times 10^{-15} & 0.36410293 & -2.5726 \times 10^{-19} & 2.5139 \times 10^{-5} \end{pmatrix}$$
(A.9)

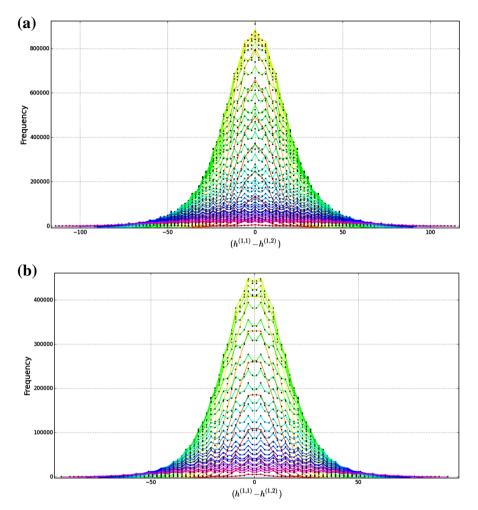


Fig. 28. This is what the entire distribution looks like using our modified pseudo-Voigt model. See Fig. 29 for the fitted coefficients as well as the fits for every curve given by the probability plots. **a** Every fitted even curve from r = 34 until r = 120. **b** Every fitted even odd from r = 35 until r = 99

$$n_{k,i} = \begin{pmatrix} 8.98205242 & 2.9066 \times 10^{-17} & 0.00434183 & -6.7671 \times 10^{-21} & -1.5512 \times 10^{-7} \\ 11.6018246 & 5.1148 \times 10^{-17} & 0.00644305 & 0 & -1.7241 \times 10^{-7} \\ 9.19515076 & 4.3161 \times 10^{-17} & 0.00496066 & -1.3763 \times 10^{-20} & -1.9163 \times 10^{-7} \\ 11.0620173 & -1.1446 \times 10^{-18} & 0.00570064 & 2.8085 \times 10^{-20} & -2.4813 \times 10^{-7} \\ 9.15798913 & 5.0109 \times 10^{-17} & 0.00493009 & -2.3559 \times 10^{-20} & -1.9210 \times 10^{-7} \\ 11.4578629 & -6.0813 \times 10^{-18} & 0.00705818 & 9.2055 \times 10^{-21} & -3.5862 \times 10^{-7} \end{pmatrix}$$
(A.10)

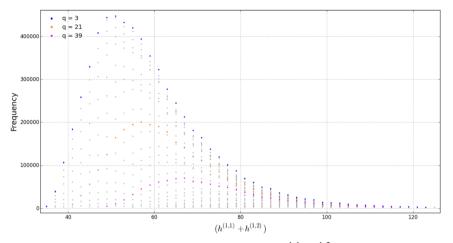
A.3. Supplementary plots for the fourfold data. When looking for mirror symmetry in the fourfold data, we only observed partial mirror symmetry. Below is the full break down of the data set.

r	A <sub>0</sub>	σ	α	b	а	$\chi^2_R$	$R^2$	р	r	A <sub>0</sub>	σ	α	b	a	$\chi^2_R$	$R^2$	p
34	74808.00828	5.61029	0.003766498	0.671247693	11882.85554	1913.323108	1	1	35		5.27052174	-0.00059798	0.66801823	11501.6207	2615.83922	0.98471088	0.83735891
36	621112.5048	6.14542	+0.009876003	0.667458363	28438.58633	40004.88553	0.99902292	0.917177648	37	666812.118	5.8927625	-0.01836059	0.66078722	27241.5063	53480.6329	0.99993762	0.97888572
38		7.04214	-0.021320726	0.661106908	70029.4258	775992.3274	0.99989369	0.967870386	39	2416867.36	7.26453848	-0.03399152	0.66024643	67114.7518	1572652.39	0.99986798	0.97276103
40	5997498.444	8.38473	-0.027896601	0.664313042	135241.8977	7016236.151	0.999812973		41		8.55598954			118674.653		0.99969021	
42	10051959.39	9.39476	-0.023578526	0.664331865	214365.7566	11248381.6	0.999606558	0.944633818	43		9.48971721			149610.971			
44	14383706.27	10.1952	-0.019800045	0.663363561	275921.3615	10356248.68	0.999910944				10.3532866			209654.283		0.99987931	
46	18900236.24		-0.011402813	0.663897086	363905.1143	12630489.6	0.999822533	0.958642702		13000374.3				254340.624			
48	26936446.43	11.52	-0.019075394		456942.6269	48618344.01	0.99980274		49		11.9906222			334507.932			
50	35415476.39		-0.02505046		634805.0051	159409130.3	0.999167385	0.888518617	51		13.1819995			402205.527			
52	40513641.09		-0.025150833	0.663741398	770677.2752	156093179.3	0.999322987	0.887202979	53		14.1510506			465061.332			0.93661201
54	42054878.16		-0.020889662	0.664242039	851781.3562	177830377.5	0.998742804	0.864062454			14.9970179			497654.277			
56		15.9308	-0.022451431	0.664639342	1081188.801	91014311.13	0.999616346			20461751.5				618719.277			
58	45777655.84	16.8075	+0.020439012	0.66390829	1216222.825	79515308.6	0.999550544	0.915805654			16.8390468			650103.969			
60	45383436.12	17.6159	-0.019455309	0.664461299	1195317.789	67324781.93	0.99975046		61		17.818317			671348.429			
62	45890243.65		-0.020089061	0.664969685	1299727.161	95590289.64	0.999225833				18.9977095			679837.657			
64	44629202.3	19.8429	-0.020615871	0.665932096	1347466.72	78628169.68	0.998988401	0.871831349		16183229.9				721485.873			
66	41517968.02	20.5755	-0.018305682	0.666138254	1468293.568	54603587.95	0.999239136			15604477.1	20.8972797			789587.968			
68	39712672.75		-0.017963577	0.66544129	1569245.184	40212010.61	0.999379453	0.892005972		13104503.1		-0.01784441				0.99891504	
70	36807367.68		-0.015684425	0.665873362	1557320.642	33793439.97	0.999174607	0.878072158		12181331.7				737223.294			
72	36162771.81		-0.016476545	0.666763067	1581985.228	21554913.66	0.999683961	0.917179602			23.5912589			778331.602			
74			-0.017822108		1872368.976	28640120.59	0.999322336				24.4411893			888710.412			
76			-0.018605896		1980563.649	44636083.29	0.998671434				25.6570194			893712.308			
78			-0.018823052	0.666381088	2189453.136	33663175.17	0.998301995				26.6109387			817061.631			
80		27.8144	-0.019346889	0.665996548	2025935.144	27318925.61	0.998478381	0.808972237			27.9445846	-0.01901739		808353.444			
82	27351655.4	28.6931	-0.017490104		2011512.915	26284425.4	0.99726603			6530766.47				771351.758			
84	24566921.31	29.8261	-0.016927049		1732875.478	23309744.54	0.996555935	0.706834711			29.9134628	-0.01668256		721999.658			
86	22906614.56		-0.016442317	0.666905358	1911979.009	14429329.24 18088956.91	0.997492504			4543976.97	31.1360788			749182.031			
88	21528402.71 19886629.72	32.153 33.3369	-0.016214982 -0.016681365	0.666381087 0.666516895	1804104.196 1783587.312	18088956.91	0.995744097 0.996471494	0.678346825 0.693352255		4543976.97				724406.102 645509.413			
90	19886629.72	33.3369	-0.016681365		1783587.312	1152/968.91 5377588.556	0.996471494	0.769625304		4114525.48 4317572.14				783101.682			
94	18809829.16		-0.017242588	0.666336104	1925176.842	8355841.233	0.998485757	0.561674719		3525255.74		-0.02087123			531215.258		
94	14889894.87	36.1253	-0.018385807		1520828.275	4989263.018	0.987709885	0.711953875		2748721.76				558435.856			
96	14869694.87	37.7735	-0.016184516	0.666516294	1693173.472	4989263.018	0.996861999				39.2526357			558530.974			
100		39.1882	-0.016641154	0.666363484	1333642.914	3437131.53	0.994423193	0.5603671	33	2721320.91	39.2320337	-0.01300372	0.00030073	338330.374	130104.307	0.99011499	0.0092328
100	11130677.16		-0.013892005	0.66646458	1359292.476	4680926.074	0.990953211										
102	9339364.392	40.191	-0.013852003	0.666411434	1159602.308	3377262.44	0.993594806										
106		41.0922	-0.012755565	0.666256708	1185946.732	2646190.089	0.993791084										
108	8380821.944	42.139	-0.013639004	0.666558606	1104233.414	1209490.828	0.996572623	0.62437074									
110	8195376.057	42.139	-0.013639004	0.666306562	1162489.945	922920.416	0.996020933	0.58859257									
112	7991589.586		-0.014513112	0.749095555	-170845.2098	5798070.695	0.996020933										
114	7502725.304	44.9131	-0.016013329	0.666363214	1206184.755	1116881.881	0.993672141	0.622571007									
114	6781922.133	48.3831	-0.015129142	0.000303214	-72768.74536	4292917.226	0.993672141										
118	6003445.42	46.5651	-0.015577227	0.666241038	1072317.828	4292917.228	0.991182119										
120	5081179.349		-0.014286545		907092.9689												
120	30011/9.349	20.3332	-0.01231933	0.000554981	301032.9089	920033.0370	0.303746662	0.2308/138/									

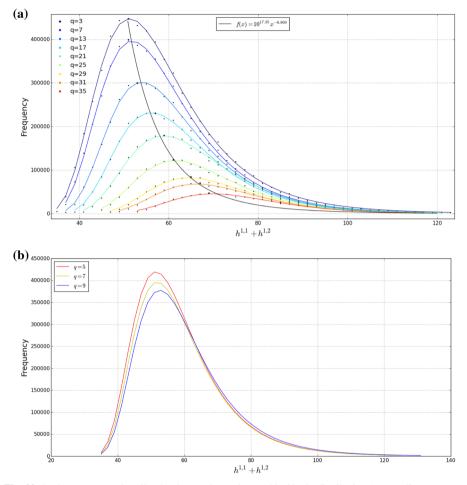
**Fig. 29.** Left list of best fit coefficients for all even curves  $r \in [34, 120]$ . Right list of best fit coefficients for all odd curves  $r \in [35, 99]$ . In both figures, the last two columns represent the  $R^2$  and p values for the probability plot for each curve. The p-values were obtained by first performing a Z-Standardization on the data

		Eve	n		Odd							
		N. Cut at	Number of da			01 C	Number of data points					
r-value	Max F	% Cut off	Total	At cut off	r-value	Max F	% Cut off	Total	At cut off			
28	3	0	7	7	29	3	0	6	6			
30	99	13.13	11	9	31	22	9.09	12	8			
32	768	9.6	23	9	33	553	4.88	20	10			
34	6258	15.1	25	9	35	5180	19.3	22	10			
36	40739	24.35	27	9	37	40607	16.25	24	10			
38	133355	35.99	31	9	39	108236	32.34	28	10			
40	244716	50.26	35	9	41	185481	46.9	30	10			
42	373126	69.68	33	7	43	259859	53.49	34	10			
44	494185	76.89	37	7	45	330009	59.99	36	10			
46	666992	73.76	41	7	47	408797	61.89	38	10			
48	793852	80.74	43	7	49	443162	69.95	40	10			
50	877191	82.42	43	7	51	447109	74.45	42	10			
52	875275	86.6	45	7	53	432081	76.37	46	10			
54	910113	84.6	49	7	55	419456	77.24	46	10			
56	816288	92.86	49	7	57	393842	86.33	48	10			
58	793170	92.54	51	7	59	354495	81.52	52	10			
60	791325	89.72	55	7	61	322535	89.91	54	10			
70	495068	94.53	65	7	71	164257	84.63	64	10			
80	278120	89.89	75	7	81	69757	86.01	76	10			
90	278120	48.5	85	7	91	31675	82.08	82	10			
100	78244	88.18	93	7	99	13812	86.88	90	10			
110	45370	88.16	105	7								
120	22840	87.56	113	9	]							

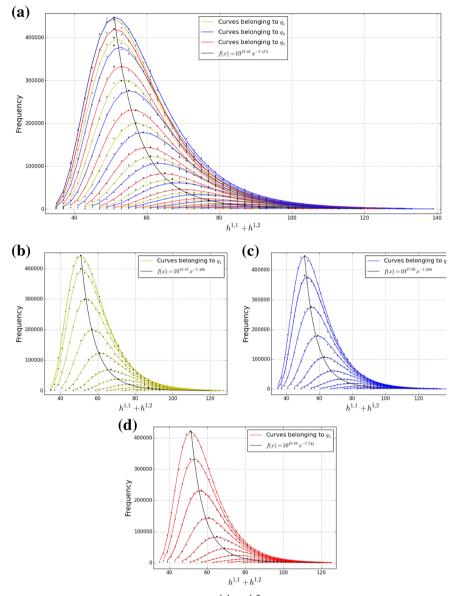
**Fig. 30.** A list showing the number of data points left after increasing the cut off frequency to achieve a perfect fit. Conversely, one may state is as, the number of data points for each curve required such that the model will result in a perfect fit



**Fig. 31.** Three highlighted curves (q = 3, 19, 31) within the odd  $h^{1,1} + h^{1,2}$  distribution. The transparent *grey data dots* are all the data plots for the distribution. Refer to Fig. 11 for the even plot



**Fig. 32.** In the attempt to describe the data analogously to a blackbody distribution (**a**), we discover some subtle structure (**b**). These are the odd counterparts to Fig. 12. **a** Lines of best fit from a regression analysis for a few select curves. The *black data* points represent the maximum frequency for that particular q - curve. The *black line* is a line of best fit to describe the points of maximum frequency—this is analogous to a blackbody spectrum. See Fig. 12a for the curves within the even distribution. **b** The curves segregate into three classes determined by the value of the even integer modulo 6. A similar pattern occurs in the even distribution; see Fig. 12b



**Fig. 33.** We illustrate the added structure for odd  $h^{1,1} + h^{1,2}$  data, by displaying how the regression curves can be divided into residue classes. For the list of even curves, refer to Fig. 13. **a** All the curves color coded according to what residue class their curves  $q_n$  belongs to. **b** Family of curves all belonging to  $q_1$ . **c** Family of curves all belonging to  $q_3$ . **d** Family of curves all belonging to  $q_5$ 

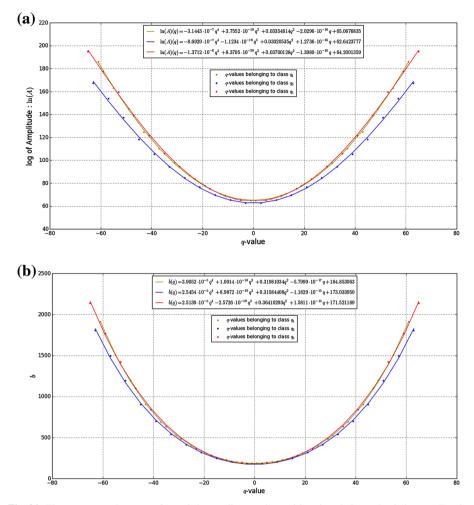
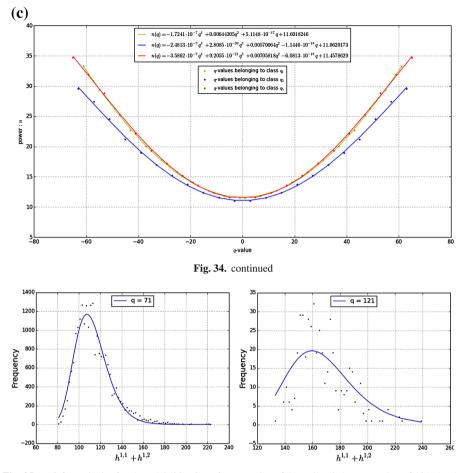


Fig. 34. The parameter plots are *color* coded according to what residue class their q value belong to. For the relationships in the even distribution, see Fig. 14. **a** Plotting the q-value parameter versus the log(A) parameter. **b** Plotting the q-value parameter versus the power n parameter (color figure online)



**Fig. 35.** *Left* figure is the fitted model (*blue line*) for a *q* value of 71 and *right* has a *q* value of 121. As the *q*-value increases, the scattering of the data points within  $h^{1,1} + h^{1,2}$  increases to the point where the model works no longer. For an example of how the model begins to break down at large *q*, see Fig. 15

is 9308131             163.22244             4.4402667             11538787.4             0.9994545             0.9995593             4             3.9308132             11.482869             183.5728             62.2667             71.519942             62.3555             66.23555             66.23555             66.23555             66.23555             66.23555             66.23555             66.23555             66.2355             57.24242.3             0.9994529             9.152719             20.43851             65.2355             66.2355             57.24242.4             0.999652             9.152719             20.43851             65.2187             51.9328             0.999652             9.152719             20.43851             65.2187             51.9328             65.2667             75.24242.4             0.999652             9.152719             20.43851             65.2187             51.9328             66.2667             75.9386             86.46438             7485382             74.5998             48.472718             0.999452             9.152719             20.43851             65.2187             51.9328             44.60440             0.999652             11             12.3880             24.0392             24.0392             24.0392             24.0392             24.0392             24.0392             24.0392             24.0392             24.0392             24.039             24.0392             24.039             24.03             24.039             24.039             24.039             24.039             24.039             24.039             24.049	q	n	b	$\ln(A)$	$\chi^2_R$	R <sup>2</sup>	p	$\left[ q \right]$	n	b	$\ln(A)$	$\chi^2_R$	$R^2$	p
2         33100737         17.619423         56.2365573         86744223.88         0.9994632         0.9913829         3         11.006489         19.352232         62.616043         7073000         0.9996632           6         3.55714274         17.468966         56.641275 0.99667426         0.99946322         0.99394222         0.99946322         0.9996521         57.51529         79.31231         0.9996521         57.51529         79.31231         75.348314         72.448388.42         0.99944329         0.9935217         9         11.527199         20.498551         66.21677         75.319311         65.61677         59.939621           10         9.7151432         155.738630         65.438677         17.31557661         78.20092         70.62285         4510106.2         0.9995631           11         12.586304         22.436645         0.9132766         0.91327647         17.155778         75.393111         8516578         0.999581           11         11.3557861         22.6310778         0.99937441         0.9995284         0.9912787         70.9356613         77.850065         37.43700         0.9993741         13.1278997         75.049858         30.99472.0         0.9992541         23.1515737         13.33775         0.7491804         0.99937446         0.9912721 </th <th><u> </u></th> <th></th> <th>-</th> <th>· · ·</th> <th></th> <th></th> <th></th> <th></th> <th></th> <th></th> <th>· · ·</th> <th></th> <th></th> <th>•</th>	<u> </u>		-	· · ·							· · ·			•
4         35921321         174.43346         6.418379         766507426         0.9994528         0.9994527         1.531627         194.73374         62.35556         6642755.4         0.9996529           6         9.57442978         186.106521         57.5571627         7812235.5         0.99944227         0.90345178         7         1.729208         202.3355         6.6226267         57.82482.4         0.9996529           10         9.7745477         18.105621         207.33557         57393127         75.83777         0.9994577         1.513126         6.037244         660440.01         0.9996529           11         12.58638         207.12867         57393127         5582775         56139405         0.999577         1.513171         3551578         0.999577         1.513171         3551578         0.999577         1.513171         3551578         0.999577         1.513171         35515748         0.999577         1.513171         35515746         0.999577         1.513171         35515748         0.999579         1.513171         35515740         0.999579         1.513171         35515740         0.9991714         0.991724         1.3599493         31.604077         5549468         5394772         0.9992448         0.9991274         231.5199359         347.40264522														0.9243267 0.9315442
6         15714724         174.68966         55.605124         781300.89         0.9994371           9         15746729         186.105617         55571639         7981235         0.9990421         11.272028         20.4953871         19.1322719         20.495319         55.127467         572842         0.09996371           10         9.79154152         105.73855         58.648384         72485389.44         0.0996371         10.248105         6.0053744         6460440.1         0.9996371           11         12.2385604         20.71267         57.3336112         75.30482         6.0053744         6405421         0.9995703           12         13.226607         274.7068         65.60531         4053641         0.9995703           12         13.076764         26.3386772         59905548         0.9995703         11.3557861         28.00507         7.530865         304472         0.999381           12         13.122807         27.838167         7.430864         3923710.23         0.9982541         21         1.3069783         33.44726         0.9991272           15         13.55047         7.998864         3923700.33         0.9985703         3825451         21.5557474         3999471         292456         2923726         21.555														0.9315442
Image: solution         Statestyne: 16:00:021         Statestyne: 10:00:021         S	-													0.9327556
10         9.79154152         195.73855         58.6438354         7.2485389.34         0.99948539         0.91392681           11         12.358534         225.46685         69.057348         4660440.1         0.999503           13         0.2491052         20.310734         0.9995703         15         12.36072         47.47068         69.05073         45.05003         405.65003         405.65003         405.641.0         0.9996528           14         10.4291162         26.3581723         587095548         0.99957466         0.92151757         11         13.55861         20.65003         405.65013         77.550081         31.4447         0.99995412           11         11.018075         27.455030         57.498074         4884280.37         0.9991988         68006012         21         15.699437         15.64097         75.50081         31.4447         0.9999712           21         11.365725         298.819564         76.491806         2937102.23         0.9982710         427244         55.69074         42.692424         23         17.228911         483.6516         4.48822.6         0.9991714           31         13.6816516         54.581954         7.667051         7.06902240         0.8982423         31.779867         25.21112.1264148 </td <td></td> <td>0.9276872</td>														0.9276872
12       3.588096;       200.712867;       57.395112;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33612;       75.33611;       0.50526;       75.135111;       0.55626;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.9995746;       0.999748;       0.9995746;       0.9995746;       0.9994852;       21.1366927;       75.50948;       304477;       0.9994852;       0.9994852;       21.155953]       34.64057;       75.50948;       304473;       0.999246;       21.155953]       45.657810;       10.9994852;       21.155953]       45.657810;       10.9994852;       21.155953]       45.6184;       22.41766;       0.999246;       29.17248;       31.778664;       6.365781;       0.999246;       29.17248;       31.778664;       45.85781;       10.9994852;       25.157874741;       10.994266;       0.9950561;       31.516566;       7.85081;       10.4626;       8.560377;       10.57065;       27.75071														0.9336964
14       10.2491103       220.819009       61.032495       6407713.03       0.9995486       0.92157       17       15.338067       247.47068       69.650053       4053524       0.9994172         16       10.3760463       246.531143       62.092735       58119944.17       0.9995688       0.9163263       17       15.557661       28.099673       0.9994172         20       11.52527       24.851340       C.9928640       4885420.37       0.99951748       0.8991748       0.8997428       21       1.659283       36.40697       75.504981       30.99475.0       0.9994524         21       11.552572       31.330475       67.4918064       32571022       20.9991748       0.8991748       0.8991726       22       15.59593       36.4.06697       75.50498       30.99475.0       0.9992464       21       15.59595       36.4.2067.67       10.00294       35.59764       0.9992464       29       17.228914       48.50714       0.9994576       29.99154.0       8894272       27.1578667       25.80169       9.6.597610       10.902415.7       0.9998576       29.1728967       25.80178       9.4.50741       0.5992460       0.8916243       31       1.728967       25.80178       9.6.507747       0.9986276       31.2.206761       31.2.206761       31														0.9281643
16       10.491492       256.074532       62.368732       5870955485       0.9991786       0.991786       0.911177         18       10.3760463       246.531143       62.0927375       58119944.17       0.99956689       0.91632683         21       11.123672642       246.5531143       62.0927375       58119944.17       0.99951868       0.8906012         21       11.5532872       28881967       67.9886889       42481778.28       0.9991784       0.8912726         21       11.5532872       31.304765       67.9886873       0.8993110       0.8706776         22       11.553287       31.304767       6.9994872       29988272       0.87228524       21       15.00676       41.00209       85.57001       0.9994174         30       12.6894483       406.767631       74.076260       17629976.3       0.8999314       0.88949423       31       10.778966       32.73858       94.12770       9807737       0.9984275         31       14.447555       55.557474       3937400.0       0.99922061       0.8916331       33       16.306018       33       16.33601       33.78858       94.12770       9807737       0.9984274         31       14.441527       52.387648       83.72864       83.78875														0.9234965
18       10.3760.463       246.5314.43       62.0927375       581194.417       0.9995689       0.98632693       19       14.076773       0.560155       77.8500981       3174.437       0.9993681         20       11.1218077       274.958037       65.7459807       4885420.37       0.99919888       0.88001226         21       11.555273       313.307475       67.4918064       3223710.23       0.99982721       0.8725617         21       12.4616.617       75.504985       365.56615       77.8500981       3174.437       0.9994452         22       11.555539       364.72264       85.57414       0.9992446       0.8991256       21       1.555539       364.72264       85.5741       0.9992461       0.8992764       21       1.555539       36.4507.66       0.9992465       0.99927666       0.8916381       31       1.7798067       525.80198       97.650175       16.2968.6       0.9986274         31       15.4166105       668.57248       91.404742       55897426       0.89917261       0.8916381       31       12.7198967       525.80198       97.650175       116.2968.6       0.9986274         31       15.416626       6652.748       91.404742       558974748       0.99917261       0.8916381       31 <td< td=""><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td>0.9199323</td></td<>														0.9199323
20       11.1218075       27.4956303       65.79807       48852803       70.99917848       0.89910266         22       11.5532872       298.881967       67.9886289       2481778.28       0.99917848       0.8910276         24       11.663725       31.3.07475       67.4918064       3923710.23       0.5998311       0.8706107         24       11.665725       31.3.07475       67.4918064       3923710.23       0.59983121       0.8706107         24       12.765105       384.515947       47.6674372       2223365617       0.9992446       382352442         31       12.6894483       40.767631       74.0706202       1762987.3       0.99927066       0.88164331         31       12.6894483       40.767631       74.0706202       1762987.3       0.99927061       0.8816331         31       12.6894483       40.767631       74.0706202       1762987.63       0.99927061       0.88163331         33       15.810616       60.8272.14       83.591016       52.805.98       9.1125.75       0.9984455         33       15.410515       65.607.6272.44       9.99917318       18.9393393       68.7232.61       0.5927.448       0.897937         34       14.2413274       52.3870488       3.70107.63<														0.9254928
12       11.55327/2       29.881967       67.9886289       42481782.8       0.9991311       0.870251         24       11.3663725       31.30745       67.4918064       39237109.23       0.9989311       0.8702572         25       15.729403       397.96698       86.578101       1902413.7       0.9994512         26       12.6161621       355.56044       72.6497065       28.982723       0.8722852         27       11.530676       41.00298       84.560741       2064262       0.9992264         29       12.729165       38.45154       47.6674572       22243666.17       0.9992566       0.8916381         34       14.4765555       505.577447       83.9371435       9337609.09       0.99922066       0.8916381         36       13.213724       22.337648       83.372010       88947231       0.8891653       35       19.278011       10.63270127       105.81396       691125.75       0.9984491         37       15.8169165       606.52248       9.404742       556997.245       0.99921301       0.88910501       35       19.278601       632.70127       105.81396       691125.75       0.9984491         42       16.912135       691.261106       9.4223259       9.69972301       0.8896753 </td <td></td> <td>0.9312652</td>														0.9312652
24       11.3663725       33.307475       67.4918064       3923709.23       0.99982732       0.99982732       25       15.2720403       397.56688       85.78101       1902413.7       0.99992464         26       12.7611565       384.581954       74.6674572       2224366.17       0.99924444       0.8932544       21       27       15.200676       411.0209       84.580741       2.064263.2       0.99932464       0.8932544       21       17.2182411       48.65516       94.488721       1.448852.8       0.9991744       31       17.738956       552.8018       97.650175       11626864.6       0.9986273       31       16.93601       553.78358       94.17770       99.980773       0.9984245       0.9991743       31       16.93601       537.8358       94.17770       99.980773       0.9984245       0.99912132       35       15.2780018       97.650175       116.2966.6       0.9982454         36       15.31636       608.652248       93.14047442       5569077.245       0.99924541       0.8991633       31       12.175717       9.9984373       39       18.39393       688.72264       105.348046       0.9990344       42       12.017718       130.7175751       10.728704       130.7175751       10.728704       130.7175751       14.34733       26.6														0.918997
26       12.4166129       355.560944       72.6497065       28790825       0.9982423       0.87228524         28       12.7691656       384.581954       74.6674572       22243686.17       0.9993144       0.88924423       21       17.289715       52.80198       74.488827       1448828.6       0.99991714         30       12.6894483       406.76751       74.7062620       1762987.3       0.9993154       0.8894423       31       16.73607       552.80198       97.66017.3       182.97666       0.9982561       0.8911557       53       16.93606       0.9992561       0.8911517       53       16.93606       632.0212       10.8339369       0.99925081       0.8911557       53       97.6517.3       0.9984497         36       15.8169165       608.6325248       9.1404742       556907.7744       0.99921301       0.88390679       31       19.31381       11.91311       1942938       0.9990131       12.31511       1942938       0.9990131       14.210521       15.4566       3650.7752       90.63277       14.64212       15.7143.8       0.9990131       18.31641641       12.210757       89.76713       19.7383       15.816451       19.44245       0.9990184       14.8100508       12.815691       12.162269       12.51112.511141       19.214114														0.917291
28       12.769165       334.891954.       74.6674572       22243686.17       0.9993154       0.89326445         30       12.6894483.       406.767631       74.7062062       17629876.3       0.9993154       0.891249423       31       17.728967       52.80198       97.650175       1162966.6       0.9998576         31       13.4815554       50.577447       83.3731435       9337600.90       0.99922056       0.99212182       37       10.638001       55.78855       94.12770       0.9988226         36       14.4765555       50.577447       83.3731435       933760.99942056       0.99942056       0.9921182       37       10.42868       94.12770       939.897731       0.9990213         40       16.349038       658.037252       94.494412       5569077.45       0.9991338       0.881540184       42       2.107573       83.97631       121.39181       194299.88       0.9990218         44       16.1012135       69.127110       104.222499       3575959.522       0.99911189       0.84339057       124.64212       151.148.40       0.9990183         45       21.162261       96.21717       106.074662       47       125.1371       110.07267       183.4099       675.13.0984578       0.9990581         46 <td></td> <td>0.9002134</td>														0.9002134
30         12.6894483         406.767631         74.702876.3         0.99932766         0.8994233         31         17.799967         525.80198         97.650175         1162966.6         0.9988245         33         1633601         557.75865         94.127709         98077737         0.9988245           34         14.4765555         505.577447         83.9716135         0.99922061         0.89116317         33         16.39601         657.75865         94.127709         98077737         0.9988497           36         14.2413274         523.837648         83.27010.2         8819447.781         0.99922301         0.88976733         31         18.333939         698.72231         10.834939         698.72321         10.59867         39         10.04235         514349.44         0.9998339           40         16.349038         658.03765         99.757954         0.99912031         0.833939         99.73231         116.349039         698.72331         112.19181         194299.88         0.9998439           41         16.1010213         691.25.71964         0.99912081         0.84339097         47         22.162799         90.33577         124.44212         15174.543         0.9997081         0.991244         12.162296         0.5231411515         1537.05610         1537.755														0.8994929
32       13.881504       45.268749       80.709756       12509144.76       0.9992056       0.8916331         33       14.4765555       505.57747       83.9731435       9337609.9       0.9992056       0.8916127       35       19.278601       632.70127       106.81339       691125.75       0.9984497         34       14.4765555       505.57747       83.9731435       9337609.9       0.9992056       0.60221328       37       10.2425       51.439.44       0.9990738       39       15.8369187       10.04225       51.439.44       0.9990738       39       18.8339968.72236       105.34806       364057.66       0.9990318         42       16.1912135       691.2161106       94.2923259       4679157.964       0.9990389       0.8339069       43       22.637039       901.93577       124.64212       15213.48       0.9990488         44       18.005802       P36.1931       104.122439       384528.484       0.99605247       0.80146662       47       2.5.1371       110.05777       104.64212       15213.48       0.9990488         50       20.6727131       105.37914473       384528.484       0.9961784       0.80146621       47       2.5.1371       110.5079       9.914488       0.9971183       0.804746622       1155.917105 </td <td></td> <td>0.8730177</td>														0.8730177
34         14.4765555         505.577447         83.9731435         9337609.09         0.99925081         0.89116175           36         14.2413274         523.937648         33.9731002         8819477.81         0.99922001         689.78187         110.04235         513.4339         691.125.75         0.9984393           40         16.349038         658.037252         94.4944182         487847.4443         0.99912301         0.8897633         91         18.933393         698.72321         105.3499.44181         194299.88         0.99906349           42         16.1912135         691.261106         94.292329         477915.7964         0.99906349         0.88339057         146.4712         1523.484         0.9990143           48         10.832493         86.994217         106.51792         94.492327         105.1792         94.8239077         146.4712         1523.484         0.9990143         142.16276         90.253125         118.15491         153776.51         0.999704057           48         18.324437         86.994271         10.95514726         0.9966327         0.80731875         101.0797         138.4009         67751.3         0.99984393           52         20.6272191         100.263688         118.64632         2550656         0.997571.35														0.8660956
36         14.241274         523.937648         83.720102         881964.7781         0.99922300         0.88057633         39         15.8169165         608.625248         91.4047442         5569077.245         0.99923201         0.88057633         39         18.33393         698.72236         105.34006         364507.66         0.9983439           40         16.349038         658.07272         94.4944182         467847.443         0.9991338         0.88154016         41         12.105773         89.76313         121.391611         19429.88         0.9990618           42         16.1912135         691.26110         94.293259         467915.7964         0.99913980         0.88396059         43         22.007573         89.76313         121.99161         19429.88         0.99910810           44         18.1005027         74.09813981         0.880746862         47         25.2137         1101.0979         138.94099         67751.3         0.9994051           50         20.677714         110.745086         1156.04622         250085.401         0.99402921         0.75114473         158.34344         6.3946418         0.9974051           51         20.757715         110.743086         120.49214         124.525682         1192.4421         124.525483         3928277.														0.8745987
38         15.8169165         608.825248         91.4077425         5569077.245         0.9992330         0.88967633           40         16.349038         658.037252         94.494412         487874.443         0.9991338         0.88154018           41         16.1912135         691.251106         94.292259         4679157.964         0.98906349         0.88396651         41         22.107573         839.7631         121.39181         194299.88         0.9990818           44         18.1005802         796.219314         104.222499         3575559.522         0.9991189         0.84339079         122.63703         90133577         124.64212         151314.8         0.9990183           45         18.3204437         88.9940271         105.17192         374238.4478         0.99663247         0.80146621           52         20.2757187         117.0790         121.92727         2068004.81         0.99402924         0.75114473         53         28.790335         141.5498         193.3434         39928.712         109.0955425         0.7584744         52         2.82458661         1527.1161.510.379265         129.275778         0.9975542         0.7781154         157         2.3245424         193.1484         59.9926271         0.5834726         159.31452         193.82474 <td></td> <td>0.900835</td>														0.900835
40         16.349038         658.037252         94.4944182         48747.443         0.9991338         0.88154018         41         22.16773         839.76313         121.39181         194299.88         0.9990818           42         16.1912135         691.261106         94.292235         4679157946         0.99910891         0.84339657         124.464212         15214.84         0.99900818         43         22.63703         901.33577         124.464212         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         15376.55         163.34397         1155.3353.41         15376.55         109705         153.2452682         1192.442         137.1453         92767.708         0.9913448           52         2.05775175         1127.34981         120.49721         121.3838.21         0.9951652         757837484         53         28.70033         142.5488         3992.821         0.9931448         53         28.70033         135.24242         33.91421.5488         3992.821         0.9936516         52         2.3.320741         150.0433 <t< td=""><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td></td><td>0.843742</td></t<>														0.843742
42       16.1912.135       691.2611.06       94.2923259       4679157.964       0.99904390       0.88398659         44       18.1005802       796.21931.4       104.225499       3575959.582       0.99819991       0.84339807       45       21.162296       902.53125       118.15491       153776.51       0.99974057         45       18.3294437       0.863994271       106.517192       3742836.478       0.99663247       0.80146862       47       25.2137       101.00797       138.4099       653246.618       0.9974057         50       20.6272191       102.63881       118.60463       25.20685.400       0.99422924       0.7814876       53       28.50355       115.09346       63294.618       0.9913448       53       28.20335       1421.5498       199.33483       39282.21       0.9913448       53       28.20335       1421.5498       199.33483       39282.21       0.9936361       55       20.62774875       151.0419       162.4653       7.9946452       0.9936361       55       29.32074       151.0419       162.4653       0.9936351       56       2.675666       1257.21615       130.979251       107.8756776       138.247979       158.43242       40851.655       0.9935876       59       31.85477       126.563791       168.7756														0.8920139
44       18.1005802       796.219314       104.225499       3575959.582       0.98819891       0.8433097         45       18.837612       864.065993       108.413983       3465429.4849       0.99711189       0.80746862         48       18.3294437       86.5942.11       105.11792       3742836.478       0.99653247       0.8046201         52       21.1759554       1091.39760       318.34699       67757.13       0.9985178         54       20.6272191       1026.3688       118.64632       2550085.404       0.99402921       0.75114475         54       22.7571875       1127.43808       120.497547       2068604.81       0.99402921       0.75114475         55       22.6875666       1257.21615       130.798265       120045.569       0.9955482       0.77381147         56       22.6875666       1257.21615       130.798265       120045.569       0.9955536       0.76067776         61       23.23074       151.0.0419       152.84224       0.99315482       0.9355736       0.76067776         62       2.5.313715       155.85041       148.28866       670967.3101       0.99562740       0.7364239         63       2.4.6357215       16.83.37621       144.6682359       69273517       0														0.8941551
46       18.8376152       866.069993       108.413983       3485429.849       0.9971189       0.80746862         48       18.3294437       186.54069993       106.517192       742836.478       0.99663247       0.80148621         49       16.3294437       1105.3588       118.604632       2550085.404       0.9942224       0.7611887         50       20.6272191       1002.5888       118.604632       2550085.404       0.9942224       0.75114473         54       20.7571875       1127.43968       120.422       137.14593       32767.708       0.9915852         54       20.7571875       1127.43968       120.4891       123.734743       33726452       0.9935852         54       20.7571875       1127.43968       120.442       137.14593       377.4562       9393651         56       22.6875666       1257.21615       130.798265       1200445.963       0.775607       53       3.84324       40851.635       0.9936571         52       2.323207       156.342565       130.910755       156.342565       0.99362714       7.7361737318       158.97421       20.99362714       7.73676776       13.857173       158.29445       128.1790.99362714       7.73676776       13.857173       168.24635715       168.368261 <td></td> <td>0.8071796</td>														0.8071796
48       18.3294437       886.994271       105.17192       374283.478       0.99663247       0.80148621       49       26.28397       1195.3354       145.03946       63294.618       0.99793         50       20.2772191       1026.3688       118.60432       2550085.404       0.9942224       0.75114473       53       24.57568       1192.442       137.1453       3276.7708       0.9913448         52       21.575554       1091.7970       12.927527       0.68604.81       0.9954223       0.75114473       53       22.8.790335       142.1548       159.3483       39928.21       0.9936578         54       20.7571875       1127.43508       120.47932       150.049556       0.776507       150.997578       59       31.857.7497       20935782       0.99355723       0.77650776       6       53       3.291403       110.1974       2051.9768       0.9938572       6.76207776       6       33.291403       110.936.7578       6.993883170       1618.4565       0.9911659       6       3.321403       110.9376.185.87455       1618.4565       0.9911659       6       3.482451       1618.4565       0.9916597       6       3.482451       1618.4565       0.9916597       6       3.483431       12.499170       3.28620.007       0.99434220       <	46													0.8710315
50         20.62772191         1026.3688         118.604632         255008.404         0.99492291         0.75118475           52         21.1759554         1091.79709         121.925727         2068604.81         0.99402921         0.75118475           54         20.7571875         1127.43808         120.4975277         2068604.81         0.999518652         0.75814747           55         22.6875666         1257.21615         130.798265         1200445.969         0.99554622         0.77318115           56         22.6875666         1257.21615         130.798265         120045.969         0.9955462         0.7731815           58         32.638021         1359.24221         13.912342         13.912472         135.84324         400851.635         0.9903882           60         22.4580953         1352.48226         130.910755         1267334.281         0.9955536         0.76607776         61         33.291403         1910.9736         185.8455         1618.4565         0.9903882           64         25.324221         163.32141         148.32466         67097.7810         0.99362740         0.7364239           70         27.7560774         1938.97103         161.69022         3293721         0.7800742         0.78345333         0.99055126														0.8479883
52       21.175954       1091,79709       121.927527       2066604.81       0.9940221       0.75114473         54       20.7571875       1127.43808       120.497481       2213288.382       0.99518652       0.7384174         55       20.257666       1257.12615       130.798265       1200445.560       0.77384115       55       29.32074       151.0.0419       162.8653       37196.452       0.99936581         56       2.2675666       1257.12615       130.798265       1200445.560       0.9955462       0.77884784       55       29.32074       151.0.0419       162.8653       37196.552       0.99936576         57       3.2628300       1359.4262.0       130.91755       126734.281       0.99955362       0.7765067       59       31.8271705       185.87452       1618.355       0.9916559       63       2.9515581       1805.3579       167.28047       2001.7018       0.9984542       0.9961755       0.78642333       63       2.9515581       1805.3579       167.28047       20047.013       0.9886505         64       2.524428       160.312461       140.083556       69238.170       0.9961755       0.78642333       77.286477       738.837453       140.88356       0.9201755       0.78642333       727.560774       138.971403	50													0.7268201
54       20.7571875       1127.43808       120.47481       211288.382       0.9951862       0.7584748         56       22.687566       1257.21615       130.798265       1200845.969       0.99554623       0.77318115         57       27.365459       1494.7997       153.48324       40851.635       0.9935716         58       23.6283802       1359.92622       13.017324       1071346.75       0.99605953       0.7765676       57       31.8577       176.57928       177.4976       0.5913.452       0.9910582         62       22.4580751       163.817623       146.824885       647021.379       0.99305716       63       32.91033       1910.9736       185.87455       1618.4565       0.9911685         64       25.324228       1603.12161       146.824885       64721.379       0.99362740       0.7182371       63       29.51581       180.53579       167.28047       2047.013       0.9884544         64       25.324289       1603.12161       157.9921.735       0.7823735       157.84725804       7.78245944       0.9993571       0.78247594       7.495.1455       0.986527       0.5801974       7.6333357       2.984246       2291.1684       171.035076       9.9937518       0.5873776       0.89807807       0.5363038														0.7553077
56         22.6875666         1257.21615         130.798265         120045.966         0.99555462         0.77318115           58         23.6283802         1359.392622         136.31234         1171384.578         0.99605563         0.7765076           60         22.4580053         1352.48226         130.910755         126733.4281         0.9955536         0.76667776           61         25.53137153         1558.90413         146.324466         670967.8101         0.99505786         0.76067776         63         2.9515581         1805.3579         167.28047         24047.013         0.988454           62         25.53137153         1558.90413         146.324668         670967.8101         0.99362740         0.7364229           64         25.324289         1603.12416         146.068359         699328.179         0.99434529         0.7364229           66         24.6357215         1638.37623         144.068359         699238.179         0.99434529         0.7364229           70         27.750747         138.57101         16.69022         32651756         0.58615763           72         2.6900085         1955.18548         158.266559         0.99805827         0.56315763           74         29.43732         172.63049														0.7295352
58         23.6283802         1359.92622         131.2334         1171384.578         0.9906826         0.778506         59         31.8577         776.5728         177.4976         2051.9768         0.9908825           60         22.4580953         1352.48226         130.010755         1267334.281         0.9995536         0.76067776         61         33.291403         1910.9766         185.87455         16184.565         0.99101659         61         33.291403         1910.9766         185.87455         16184.565         0.9901659         61         23.317153         158.90413         146.324866         607967.8101         0.99500786         0.7602774         63         25.317153         145.87455         16184.565         0.9911659         62         25.317153         146.324866         60792.17776         0.9934.6720         0.7602774         63         25.317153         146.823813         144.82485         64712.17779         0.99434252         0.766243335         70         27.745074         138.97103         1516.9022.20         2429.131584         158.26555         0.6865575         0.76233335         70         29.433332         222.222.222.22549         124.84859         0.2037.7676         0.587657         78         28.984246         231.16584         154.306976         0.587657         0	56													0.7339095
60         22.4580953         1352.48226         130.910755         1267334.281         0.9955536         0.76067776           62         25.313715         1558.90413         146.324866         6709778100         0.99502786         0.76027774           64         25.3242829         1603.1214         146.324866         67077740101         0.99362734         0.7182791           66         24.6357215         1583.37623         144.6824866         67077740101         0.99362734         0.7182791           67         24.6357215         1633.37623         144.082359         69932331.79         0.99434751         0.7264239           68         27.1759004         1938.97103         161.69025         422971.033         0.9961755         0.7264339           72         2.696005         1955.18548         158.266559         64280.509         0.99956221         0.7263335           72         2.960054         155.18548         158.266559         0.99055221         0.5301974           76         20.75113         0.7486739         0.99055221         0.5301974         0.9935748         0.5586308           82         3.2755107         171.05076         149357.371         0.98667807         0.5586308           82         3.2755107<	58	23.6283802						59	31.8577	1765.7928	177.4976	20519.768		0.6964478
62       25.3137133       155.88.90413       146.3248658       670967.8101       0.99500786       0.76027754         64       25.3147153       155.88.90413       146.324885       647121.3779       0.99346273       0.78642239         65       24.657125       1638.37623       144.068359       699238.170       0.993436229       0.73642239         66       24.657125       1638.37623       144.068359       699238.170       0.99434529       0.73642239         67       27.755004       1836.1188       157.949175       328620.4071       0.99434529       0.73642333         70       27.755074       1938.97101       161.69022       34251.0330       0.99617755       0.76233335         72       26.960055       1955.18548       158.266559       642806.509       0.98750524       0.587467         74       29.943382       222.22549       174.848859       202372.2104       0.99055632       0.5861076         74       29.943382       223.5593772       19.787352       20551.44666       0.69750424       0.587467         78       28.984246       221.16584       171.036976       349357.371       0.98670870       0.53279776         80       32.2575970       2931.558.2728       180.790451	60	22.4580953	1352.48226	130.910755	1267334.281	0.9955536	0.76067776	61	33.291403	1910.9736	185.87455	16184.565	0.9911659	0.7134993
66         24.6357215         168.37623         144.068359         699238175         0.99424629         0.73644239           68         27.1759004         1836.21188         157.949175         326820.4071         0.99439751         0.72455049           70         27.7560774         1938.97103         161.69022         32671.033         0.99617755         0.76233335           72         25.960075         1955.18548         158.266959         64280.509         0.99905827         0.5801074           74         29.433382         2222.22349         174.894859         9.09055827         0.5615753           74         29.43382         2222.22349         174.894859         9.09055827         0.5631074           76         30.7510953         233.259771         179.797525         206551.4666         0.98970807         0.53279776           80         32.2657309         2971.130509         193.774326         9238552151         0.98870807         0.5383038           82         32.951907         271.130509         193.774326         9238552151         0.98316783         0.5360360           86         32.2223115         2807.7888         196.32844         70.633147         0.69653737         0.48653732         0.48653372														0.6685204
66         24.6357215         168.337623         144.068359         699238179         0.99434629         0.73644239           68         27.1759004         1836.21188         157.949175         25.8620.4071         0.99439751         0.72645049           70         27.7560774         1938.97103         16.169022         342571.303         0.99617755         0.7263335           72         27.560774         1938.97103         16.169022         342571.303         0.9905520         0.5815185           74         29.493382         2222.2249         1.7384.056595         64280.5390         0.9905520         0.58301974           76         30.7510953         2332.98771         179.797525         2003571.4666         0.9905520         0.58301974           78         28.944249         221.1584         171.103676         349357.371         0.98070807         0.53279776           93         252.521         19.302077         125820.2085         0.98870807         0.5383038           82         32.951907         271.130509         193.774326         9238552151         0.9837638         0.57063060           84         30.4719125         2588.5228         105.050514         6.9131476         0.96513710         0.46635329	64	25.3244289	1603.12416	146.824885	647121.3779	0.99362734	0.71823791	65	34.683819	2134.8346	194.79778	7495.1455	0.9866505	0.675547
70         27.7560774         1938.97103         161.69022         342571.3033         0.99611775         0.76233335           72         22.636008         1955.18548         158.266555         64280.509         0.98965827         0.63615763           74         29.9433382         2222.2249         174.884855         0.9905552         0.8301974           76         30.751053         2332.98771         179.977252         209551.4666         0.9905552         0.5801974           78         28.9842496         2291.16584         171.036976         349357.371         0.98067809         0.553279776           80         32.2551907         2711.30509         193.774326         92385.53151         0.9837680         0.55363038           82         32.951907         2711.30509         193.774326         9238552.5151         0.9837680         0.55360308           84         30.471912         2588.5228         180.790451         11555.2120         0.98337680         0.55360308           86         33.2223315         2870.76888         196.32384         67083.31487         0.9831580         0.5580301           90         3.425371         29.565345         541.448199         0.97812326         0.56580301           90         3.	66													
72         26.960085         1955.18548         158.266959         642806.509         98968827         0.65615763           74         29.943382         2222.22549         174.848859         202372.2104         0.99055632         0.65801974           78         30.7510953         232.29771         179.797552         206551.4666         0.89750424         0.587467           78         28.9842496         221.16584         171.036976         349357.371         0.98670670         0.53279776           80         32.265790         273.15323         198.320271         125882.0585         0.9878007         0.535363038           82         32.951907         271.1 30509         193.774326         92385.52151         0.9887067         0.535363038           84         30.4719125         2585.82228         180.790451         161559.2102         0.98337638         0.5503608           66         33.2223315         2870.76888         196.23844         0.08311012         0.9815071         0.53162425           88         30.015223         2905.88625         198.460534         5413.48199         0.78143256         0.55589001           90         3.2427476         2956.596548         192.49148         48845.94672         0.91545439         0.3442344	68	27.1759004	1836.21188	157.949175	326820.4071	0.99439751	0.72455049							
72         26.960085         1955.18548         158.266955         642806.590         0.98965837         0.65615763           74         29.9433382         2222.22549         174.848859         202372.2104         0.99055632         0.63801974           76         30.7510553         2332.9971         179.797525         206551.4666         0.58750424         0.587464           78         28.9842496         2291.16584         170.103676         349357.371         0.98670007         0.53279776           80         32.2657302         259.15523         189.302077         125882.0585         0.9878007         0.535363038           82         32.951907         2711.30509         193.774326         92385.52151         0.9878007         0.55363038           83         3.25223315         2807.08881         196.232012         0.98337638         0.5536308           84         30.4719125         2585.82228         180.790451         161559.2102         0.98337638         0.552603668           84         32.227315         2807.58824         6708.31487         0.98310760         0.39162425           88         33.015223         2905.68625         194.6966537         0.49653373         0.49653373         0.49653373           90	70	27.7560774												
74         29.9433382         2222.22549         174.848859         202372.2104         0.99055632         0.63801974           76         30.751055         232.99771         179.797325         206551.4666         0.98750424         0.587467           78         28.942494         220.15584         171.035974         293957.371         0.98075040         0.587467           78         28.942494         220.15584         171.035976         39357.371         0.98070800         0.55363038           82         32.951907         7211.30509         193.77426         2385.5215         0.98856611         0.51710224           43         30.4719125         258.52228         180.790451         161559.2102         0.98337638         0.5706008           86         33.2023315         2870.76888         196.53344         6708331487         0.96513716         0.39162425           88         33.0152923         295.655648         194.134.98199         0.97813256         0.55630301           90         3.4253478         295.565548         124.33.7688         0.9655377         0.46956336           92         3.22748776         295.565548         192.49448         4884.594672         0.9156439         0.34423447           94         30	72													
76         30.7510953         2332.98771         127.977525         206551.46666         0.98750424         0.587467           78         28.9842496         2291.16584         171.036976         349357.371         0.98607809         0.553279776           08         32.657369         2579.15523         189.320277         125882.0858         0.98870807         0.55363038           82         32.951907         2711.30509         193.774326         92385.52151         0.98870807         0.55363038           84         30.4719125         2588.2228         180.790451         116559.2102         0.98337688         0.5506300           86         33.2223315         2870.76888         196.32384         67083.31487         0.98337688         0.5506300           86         33.2223315         2870.76888         196.32384         67083.31487         0.96310176         0.39162425           88         3.015223         2905.88625         194.5065345         6.9655373         0.46563732         0.4656373         0.4656373           90         32.4257476         2956.596548         192.249148         48845.49672         0.91954939         0.34423447           94         30.5994413         2867.18956         183.016328         60329.22018         0.	74													
80         32.2657369         2579.15523         189.320277         125882.0585         0.98870807         0.55363038           82         32.951907         2711.30509         193.774326         92385.52151         0.98837681         0.55363038           84         30.4719125         2585.22281         180.790451         116155.2120         0.98337681         0.552063060           86         33.222315         2870.76888         196.32344         67083.31487         0.96317681         0.552063061           90         24.52978         2953.68556         194.434.98199         0.97813256         0.565630301           90         32.452978         2955.965548         192.4934348         0.9655377         0.46936396           92         30.5994413         2867.18956         183.016328         0.924292.2018         0.79414806         0.22700301           94         30.5994413         2867.18956         183.063286         1027.74244         0.48637432         0.2130179	76	30.7510953												
82         32.951907         2711.30509         193.774326         92385.52151         0.98586611         0.51710224           84         30.4719125         2585.82228         180.790451         161559.2102         0.98337638         0.52003608           63         32.223315         2870.76888         196.32344         67083.31487         0.9631076         0.39162425           88         33.015223         2905.88625         195.605348         5413.49199         0.97813256         0.55580301           90         32.45278         293.6055346         5413.49199         0.97813256         0.55580301           90         32.45278         293.605346         5124.39198         9.0955373         0.46933396           92         32.748776         2965.956548         192.429148         48845.94672         0.91954939         0.34423447           94         30.5994413         2867.18956         183.016328         60329.22018         0.79416806         0.22700301           94         30.57937576         2945.66088         183.699561         12777.4424         0.84637422         0.2130179	78	28.9842496	2291.16584	171.036976	349357.371	0.98607809	0.53279776							
84         30.4719125         2585.82228         180.790451         161559.2102         0.98337638         0.52603608           86         33.2223315         2870.76588         196.32384         67083.31487         0.96310176         0.39162425           88         33.0152292         2905.886255         193.495666         128633.7698         0.96655373         0.46936396           90         32.452978         2953.68556         193.495666         128633.7698         0.96655373         0.46936396           92         32.7429776         2955.96548         192.49148         48845.94672         0.9191526         0.55680301           94         30.5994413         2867.189566         183.016328         60323.22018         0.79416806         0.22700301           94         30.5934716         2945.66688         183.699965         12777.4424         0.84637422         0.22130179	80	32.2657369	2579.15523	189.320277	125882.0585	0.98870807	0.55363038							
86         33.2223315         2870.76888         196.32384         67083.31487         0.96310176         0.39162425           88         33.015223         2905.88653         195.605346         6413.49199         0.97813256         0.5658030           90         32.452974         293.68565         193.49566         128633.7689         0.97813256         0.5658030           92         32.2748776         2956.596548         192.499148         48845.94672         0.91956439         0.34423447           94         30.5994413         2867.18956         183.016328         0.3292.2018         0.7941806         0.22700301           96         30.537376         2945.66088         183.699661         12777.4424         0.48637432         0.21310179	82	32.951907	2711.30509	193.774326	92385.52151	0.98586611	0.51710224							
88         33.0152923         2905.88625         195.605348         54134.98199         0.97813256         0.56580301           90         32.45278         2953.68556         193.495666         12663.7698         0.96655373         0.46936396           92         32.2748776         2965.96548         192.49148         48845.94672         0.91956493         0.34423447           94         30.5994413         2867.18956         183.016328         6023.22018         0.79416806         0.22700011           96         30.5373576         2945.666088         183.699991         126777.4424         0.84637432         0.22130179	84	30.4719125	2585.82228	180.790451	161559.2102	0.98337638	0.52603608							
88         33.0152923         2905.88625         195.605348         54134.98199         0.97813256         0.56580301           90         32.452978         2953.68556         193.495666         12683.7698         0.96655373         0.46936396           92         32.2748776         2965.96548         192.49148         48845.94672         0.91956493         0.34423447           94         30.5994413         2867.18956         183.016328         6023.22018         0.79416806         0.22700011           96         30.5373576         2945.666088         183.699991         126777.4424         0.84637432         0.2130179	86	33.2223315	2870.76888	196.32384	67083.31487	0.96310176	0.39162425							
92 32.2748776 2965.96548 192.249148 48845.96472 0.91956493 0.34423447 94 30.5994413 2867.18956 183.016328 60329.22018 0.79416806 0.22700301 96 30.5373576 2945.66088 183.699961 126777.4424 0.84637432 0.22130179	88	33.0152923	2905.88625				0.56580301							
94         30.5994413         2867.18956         183.016328         60329.22018         0.79416806         0.22700301           96         30.5373576         2945.66088         183.699961         126777.4424         0.84637432         0.22130179	90	32.452978	2953.68556	193.495666	128633.7698	0.96655373	0.46936396							
96 30.5373576 2945.66088 183.699961 126777.4424 0.84637432 0.22130179	92	32.2748776	2965.96548	192.249148	48845.94672	0.91956493	0.34423447							
	94	30.5994413	2867.18956	183.016328	60329.22018	0.79416806	0.22700301							
	96	30.5373576	2945.66088	183.699961	126777.4424	0.84637432	0.22130179							
98  29./580503  2914.9165  1/9.028421  43017.60215  0.64681657  0.28617484	98	29.7580503	2914.9165	179.028421	43017.60215	0.64681657	0.28617484							
100 28.0712553 2800.34637 169.674959 31972.1718 0.5910797 0.36058935														

**Fig. 36.** Left: list of best fit coefficients for all even curves  $q \in [0, 100]$ . Right list of best fit coefficients for all odd curves  $q \in [1, 65]$ 

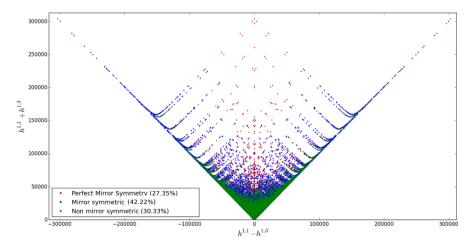


Fig. 37. Mirror symmetry is incomplete in the fourfold data set

## References

- 1. Candelas, P., Horowitz, G.T., Strominger, A., Witten, E.: Vacuum configurations for superstrings. Nucl. Phys. B 258, 46 (1985)
- Candelas, P., Dale, A.M., Lutken, C.A., Schimmrigk, R.: Complete intersection Calabi–Yau manifolds. Nucl. Phys. B 298, 493 (1988)
- Candelas, P., Lutken, C.A., Schimmrigk, R.: Complete intersection Calabi–Yau manifolds. 2. Three generation manifolds. Nucl. Phys. B 306, 113 (1988)
- Gagnon, M., Ho-Kim, Q.: An exhaustive list of complete intersection Calabi–Yau manifolds. Mod. Phys. Lett. A 9, 2235 (1994)
- 5. Hitchin, N.: Generalized Calabi-Yau manifolds. Quart. J. Math. 54, 281. arXiv:math.DG/0209099
- 6. Douglas, M.R.: The statistics of string/M theory vacua. JHEP 0305, 046 (2003). arXiv:hep-th/0303194
- 7. Candelas, P., Lynker, M., Schimmrigk, R.: Calabi-Yau manifolds in weighted P(4). Nucl. Phys. B 341, 383 (1990)
- Batyrev, V.: Dual Polyhedra and Mirror Symmetry for Calabi–Yau Hypersurfaces in Toric Varieties. arXiv:alg-geom/9310003
- Batyrev, Victor V., Borisov, Lev A.: On Calabi–Yau complete intersections in toric varieties. In: Andreatta, M., Peternell, T. (eds.) Higher Dimensional Complex Varieties, Proceedings of the International Conference, pp. 39–65. Waller de Gruyter, Trento, Italy, Berlin (1996). arXiv:alg-geom/9412017
- Kreuzer, M., Skarke, H.: On the classification of reflexive polyhedra. Commun. Math. Phys. 185, 495 (1997). arXiv:hep-th/9512204
- Avram, A.C., Kreuzer, M., Mandelberg, M., Skarke, H.: The web of Calabi–Yau hypersurfaces in toric varieties. Nucl. Phys. B 505, 625 (1997). arXiv:hep-th/9703003
- Kreuzer, M., Skarke, H.: Classification of reflexive polyhedra in three-dimensions. Adv. Theor. Math. Phys. 2, 847 (1998). arXiv:hep-th/9805190
- 13. Kreuzer, M., Skarke, H.: Reflexive polyhedra, weights and toric Calabi–Yau fibrations. Rev. Math. Phys. **14**, 343 (2002). arXiv:math/0001106 [math-ag]
- Kreuzer, M., Skarke, H.: Complete classification of reflexive polyhedra in four-dimensions. Adv. Theor. Math. Phys. 4, 1209 (2002). arXiv:hep-th/0002240
- Kreuzer, Maximilian, Skarke, Harald: Calabi–Yau 4-folds and toric fibrations. J. Geom. Phys. 26, 272– 290 (1998). arXiv:hep-th/9701175v1
- Gray, J., Haupt, A., Lukas, A.: Calabi–Yau fourfolds in products of projective space. Proc. Symp. Pure Math. 88, 281 (2014)
- Gray, J., Haupt, A., Lukas, A.: All complete intersection Calabi–Yau four-folds. JHEP 1307, 070 (2013). arXiv:1303.1832 [hep-th]
- Anderson, L.B., Apruzzi, F., Gao, X., Gray, J., Lee, S.J.: A new construction of Calabi–Yau manifolds: generalized CICYs. Nucl. Phys. B 906, 441–496 (2016). arXiv:1507.03235 [hep-th]
- Altman, R., Gray, J., He, Y.H., Jejjala, V., Nelson, B.D.: A Calabi–Yau database: threefolds constructed from the Kreuzer–Skarke list. JHEP 1502, 158 (2015). arXiv:1411.1418 [hep-th]
- Davies, R.: The expanding zoo of Calabi-Yau threefolds. Adv. High Energy Phys. 2011, 901898 (2011). arXiv:1103.3156 [hep-th]
- Candelas, P., Davies, R.: New Calabi-Yau manifolds with small Hodge numbers. Fortsch. Phys. 58, 383 (2010). arXiv:0809.4681 [hep-th]
- 22. He, Y.H.: Calabi–Yau geometries: algorithms, databases, and physics. Int. J. Mod. Phys. A 28, 1330032 (2013). arXiv:1308.0186 [hep-th]
- Anderson, L.B., He, Y.H., Lukas, A.: Heterotic compactification, an algorithmic approach. JHEP 0707, 049 (2007). doi:10.1088/1126-6708/2007/07/049. arXiv:hep-th/0702210 [hep-th]
- Gabella, M., He, Y.H., Lukas, A.: An abundance of heterotic vacua. JHEP 0812, 027 (2008). doi:10.1088/ 1126-6708/2008/12/027. arXiv:0808.2142 [hep-th]
- Gao, P., He, Y.H., Yau, S.T.: Extremal Bundles on CalabiYau Threefolds. Commun. Math. Phys. 336(3), 1167 (2015). doi:10.1007/s00220-014-2271-y. arXiv:1403.1268 [hep-th]
- Anderson, L.B., Gray, J., Lukas, A., Palti, E.: Heterotic line bundle standard models. JHEP 1206, 113 (2012). doi:10.1007/JHEP06(2012)113. arXiv:1202.1757 [hep-th]
- Braun, V., He, Y.H., Ovrut, B.A., Pantev, T.: The exact MSSM spectrum from string theory. JHEP 0605, 043 (2006). doi:10.1088/1126-6708/2006/05/043. arXiv:hep-th/0512177
- Taylor, W.: On the Hodge structure of elliptically fibered Calabi–Yau threefolds. JHEP 1208, 032 (2012). arXiv:1205.0952 [hep-th]
- Taylor, W., Wang, Y.N.: A Monte Carlo exploration of threefold base geometries for 4d F-theory vacua. JHEP 01, 137 (2016). arXiv:1510.04978 [hep-th]
- Gao, X., Shukla, P.: On classifying the divisor involutions in Calabi–Yau threefolds. JHEP 11, 170 (2013). arXiv:1307.1139 [hep-th]
- Blumenhagen, R., Jurke, B., Rahn, T.: Computational tools for cohomology of toric varieties. Adv. High Energy Phys. 2011, 152749 (2011). arXiv:1104.1187 [hep-th]

- Gray, J., He, Y.-H., Jejjala, V., Jurke, B., Nelson, B.D., Simon, J.: Calabi–Yau manifolds with large volume vacua. Phys. Rev. D 86, 101901 (2012). arXiv:1207.5801 [hep-th]
- Candelas, P., Constantin, A., Skarke, H.: An abundance of K3 fibrations from polyhedra with interchangeable parts. Commun. Math. Phys. 324(3), 937–959 (2013). arXiv:1207.4792 [hep-th]
- Braun, V.: On free quotients of complete intersection Calabi–Yau manifolds. JHEP 1104, 005 (2011). arXiv:1003.3235 [hep-th]
- Candelas, P., de la Ossa, X., He, Y.H., Szendroi, B.: Triadophilia: a special corner in the landscape. Adv. Theor. Math. Phys. 12, 429 (2008). arXiv:0706.3134 [hep-th]
- Kreuzer, M., Skarke, H.: PALP: a package for analyzing lattice polytopes with applications to toric geometry. Comput. Phys. Commun. 157, 87 (2004). arXiv:math/0204356 [math-sc]
- 37. Braun, A.P., Knapp, J., Scheidegger, E., Skarke, H., Walliser, N.O.: PALP-a User Manual. arXiv:1205.4147 [math.AG]
- 38. The On-Line Encyclopedia of Integer Sequences. http://oeis.org, Number A090045
- He, Y.H., Lee, S.J., Lukas, A.: Heterotic models from vector bundles on toric Calabi–Yau manifolds. JHEP 1005, 071 (2010). arXiv:0911.0865 [hep-th]
- 40. Lynker, M., Schimmrigk, R., Wisskirchen, A.: Landau–Ginzburg vacua of string, M theory and F theory at c = 12. Nucl. Phys. B **550**, 123 (1999). arXiv:hep-th/9812195
- 41. Stamatis, D.H.: Six Sigma and Beyond: Statistics and Probability, vol. 3, 1st edn. CRC Press (2002)
- Braun, V.: Toric elliptic fibrations and F-theory compactifications. JHEP 1301, 016 (2013). doi:10.1007/ JHEP01(2013)016. arXiv:1110.4883 [hep-th]
- Johnson, S.B., Taylor, W.: Calabi–Yau threefolds with large h<sup>2,1</sup>. JHEP 1410, 23 (2014). doi:10.1007/ JHEP10(2014)023. arXiv:1406.0514 [hep-th]
- Taylor, W., Wang, Y.N.: Non-toric Bases for Elliptic Calabi–Yau Threefolds and 6D F-Theory Vacua. arXiv:1504.07689 [hep-th]
- Anderson, L.B., Gao, X., Gray, J., Lee, S.J.: Multiple fibrations in Calabi–Yau geometry and string dualities. JHEP 1610, 105 (2016). doi:10.1007/JHEP10(2016)105. arXiv:1608.07555 [hep-th]
- Candelas, P., Constantin, A., Mishra, C.: Calabi–Yau Threefolds With Small Hodge Numbers. arXiv:1602.06303 [hep-th]
- Bianchi, M., Ferrara, S.: Enriques and octonionic magic supergravity models. JHEP 0802, 054 (2008). doi:10.1088/1126-6708/2008/02/054. arXiv:0712.2976 [hep-th]

Communicated by Y. Yin