

Magneto-acoustic resonance in a non-uniform current carrying plasma column

By J. VACLAVIK

Department of Physics, University of Fribourg

(Received 17 March 1971)

The forced radial magneto-acoustic oscillations in a plasma column with non-uniform mass density and temperature are investigated. It turns out that the oscillations have a resonant character similar to that of the magneto-acoustic oscillations in a uniform plasma column. The properties of the axial and azimuthal components of the oscillating magnetic field are discussed in detail.

1. Introduction

An investigation of magneto-acoustic resonance is of considerable interest in problems connected with the development of methods of the high frequency heating of plasma. It was shown theoretically by Frank-Kamenetskii (1960*a, b*) that the forced radial oscillations of a uniform plasma column immersed in a uniform axial magnetic field have a resonant character when the forcing frequency approaches one of the natural frequencies of the plasma column. This phenomenon, which is called magneto-acoustic resonance, has actually been observed and investigated (by e.g. Kovan *et al.* 1962; Cantieni & Schneider 1963; Borodin *et al.* 1963; Faessler, Vaclavik & Schneider 1969). All the investigations have been made for a plasma where the mass density is uniform throughout.

In this paper, the properties of magneto-acoustic resonance in a plasma column possessing non-uniform density and temperature distributions are treated. The column is immersed in a uniform axial magnetic field and carries a uniform axial current. On the surrounding walls an azimuthal oscillating current is excited.

2. Basic equations

We now proceed to write the basic equations needed to describe the character of the oscillations inside the column. Let $\rho_0, p_0, \mathbf{B}_0, \mathbf{j}_0$ represent the equilibrium values of the mass density, pressure, magnetic field and current density. The linearized equations of magneto-hydrodynamics can then be written in the form,

$$\rho_0 \frac{\partial \mathbf{v}}{\partial t} = -\nabla p + \frac{1}{c} ([\mathbf{j}, \mathbf{B}_0] + [\mathbf{j}_0, \mathbf{B}]), \quad (1)$$

$$\frac{\partial p}{\partial t} + (\mathbf{v} \cdot \nabla) p_0 + \gamma p_0 (\nabla \cdot \mathbf{v}) = 0, \quad (2)$$

$$\mathbf{j} = \sigma \left(\mathbf{E} + \frac{1}{c} [\mathbf{v}, \mathbf{B}_0] \right), \quad (3)$$

$$[\nabla, \mathbf{B}] = \frac{4\pi}{c} \mathbf{j}, \tag{4}$$

$$[\nabla, \mathbf{E}] = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}, \tag{5}$$

where \mathbf{E} is the electric field, \mathbf{v} the plasma velocity, γ the adiabaticity index and $\boldsymbol{\sigma}$ the conductivity tensor.

In what follows we adopt a cylindrical co-ordinate system r, ϕ, z , and assume that $\partial/\partial\phi \equiv \partial/\partial z \equiv 0$ throughout. For the equilibrium quantities of the plasma, we can then use the simplified model of pinched plasma column. Thus

$$\mathbf{B}_0 = i_z \mathbf{B}_0 + \mathbf{i}_\phi \left\{ \begin{array}{l} \frac{2\pi}{c} j_0 r \quad (r \leq R), \\ \frac{2\pi}{c} j_0 \frac{R^2}{r} \quad (r \geq R), \end{array} \right\} \tag{6}$$

$$\mathbf{j}_0 = \mathbf{i}_z j_0,$$

$$p_0 = \left(\frac{j_0 R}{c}\right)^2 \pi \left(1 - \left(\frac{r}{R}\right)^2\right),$$

where R is the radius of the column, and \mathbf{i}_ϕ and \mathbf{i}_z are the unit vectors in the azimuthal and axial directions, respectively.

Finally, (1)–(5) must be completed by a boundary condition. Consistently with our assumptions, we may write it as

$$\mathbf{B} = \mathbf{i}_z B_{ez} \cos \omega t \quad (r = R), \tag{7}$$

where ω is the forcing frequency.

3. Uniform temperature case

Equations (1)–(5) are too hard to solve in general. Therefore, we only consider some special cases. We study first the case when the equilibrium plasma temperature is uniform over the column cross-section. From (6) it then follows that

$$p_0 = \delta^2 c_A^2 \rho_0, \quad \rho_0 = \rho_0^* \left(1 - \left(\frac{r}{R}\right)^2\right), \tag{8}$$

ρ_0^* being the mass density at the centre of the column, $c_A = B_0/(4\pi\rho_0^*)^{1/2}$ the Alfvén velocity and $\delta = B_{0\phi}(R)/B_0$.

Let us, for a while, assume that $\boldsymbol{\sigma} \rightarrow \infty$. Thus, setting $\mathbf{j} = 0$ in (3), introducing the time dependence $\exp(-i\omega t)$ and eliminating the quantities \mathbf{j} and \mathbf{j}_0 by means of (4) and (6), respectively, we obtain the following equations:

$$i\omega\rho_0 v = \frac{dp}{dr} + \frac{B_0}{4\pi} \left\{ \frac{dB_z}{dr} + \frac{\delta}{R} \left(r \frac{dB_\phi}{dr} + 3B_\phi \right) \right\}, \tag{9}$$

$$i\omega p = v \frac{dp_0}{dr} + \gamma p_0 \frac{1}{r} \frac{d}{dr}(rv), \tag{10}$$

$$E_\phi = \frac{vB_0}{c}, \tag{11}$$

$$E_z = -\frac{v}{c} \cdot \frac{\delta r}{R} B_0, \tag{12}$$

$$B_z = \frac{c}{i\omega r} \frac{d}{dr} (rE_\phi), \tag{13}$$

$$B_\phi = -\frac{c}{i\omega} \frac{dE_z}{dr}, \tag{14}$$

where $v \equiv v_r$.

Although the quantities of interest are the magnetic field components B_z and B_ϕ , from the mathematical point of view it is more convenient to solve the equations governing the electric field components E_ϕ and E_z . On making use of (8), (9)–(14) can be reduced to one equation involving the quantity E_ϕ . This is accomplished by eliminating p, v, E_z, B_z, B_ϕ by (10)–(14), respectively. Now, (9) becomes

$$\left\{ \frac{d}{d\bar{r}} \frac{1}{\bar{r}} \frac{d}{d\bar{r}} \bar{r} + \bar{\omega}^2 (1 - \bar{r}^2) \right\} E_\phi = -\delta^2 \frac{d}{d\bar{r}} \{ \gamma(1 - \bar{r}^2) + \bar{r}^2 \} \frac{1}{\bar{r}} \frac{d}{d\bar{r}} \bar{r} E_\phi, \tag{15}$$

where $\bar{r} = r/R$ and $\bar{\omega} = \omega R/c_A$.

Since the solution of (15) for arbitrary values of δ is difficult, we confine ourselves to the limiting case $\delta^2 \ll 1$. The electric field component E_ϕ , with the accuracy up to the first order of the parameter δ , is then given by the solution of the equation,

$$\frac{d}{d\bar{r}} \frac{1}{\bar{r}} \frac{d}{d\bar{r}} \bar{r} E_\phi + \bar{\omega}^2 (1 - \bar{r}^2) E_\phi = 0. \tag{16}$$

With the same accuracy, the electric field component E_z is obtained by combining (11) and (12) as

$$E_z = -\delta \bar{r} E_\phi, \tag{17}$$

E_ϕ being the solution of (16).

We now return to the case when the conductivity is finite and assume that $\omega c^2/(4\pi\sigma_\perp c_A^2) \ll 1$. Hence, by means of (3)–(5), it can easily be shown that (16) is to be replaced by

$$\frac{d}{d\bar{r}} \frac{1}{\bar{r}} \frac{d}{d\bar{r}} \bar{r} E_\phi + \bar{\omega}^2 (1 - \bar{r}^2) \{ 1 + ia\bar{\omega}(1 - \bar{r}^2) \} E_\phi = 0, \tag{18}$$

where $a = c^2/(4\pi\sigma_\perp Rc_A) \ll 1$. As far as the quantity E_z is concerned, it is sufficient to use the approximate formula (17).

Equation (18) can be reduced to simpler form by making the transformation,

$$x = \bar{\omega}\bar{r}^2, \quad E_\phi(x) = x^{\frac{1}{2}} e^{-\frac{1}{2}x} \psi(x). \tag{19}$$

This transformation reduces (18) to

$$x \frac{d^2\psi}{dx^2} + (2-x) \frac{d\psi}{dx} - \alpha\psi \left\{ 1 - \frac{ia}{4\alpha} (\bar{\omega} - x)^2 \right\} = 0, \tag{20}$$

where $\alpha = 1 - \frac{1}{4}\bar{\omega}$.

In order to solve (20), we use a perturbation method. In particular, we let

$\psi = \psi^{(0)} + \alpha\psi^{(1)} + \dots$, and confine ourselves to the first two terms of the expansion. It is evident that the functions $\psi^{(0)}$ and $\psi^{(1)}$ satisfy

$$x \frac{d^2\psi^{(0)}}{dx^2} + (2-x) \frac{d\psi^{(0)}}{dx} - \alpha\psi^{(0)} = 0, \tag{21}$$

$$x \frac{d^2\psi^{(1)}}{dx^2} + (2-x) \frac{d\psi^{(1)}}{dx} - \alpha\psi^{(1)} = -\frac{1}{4}i(\bar{\omega} - x)^2 \psi^{(0)}. \tag{22}$$

Equations (21) and (22) are homogeneous and inhomogeneous confluent hypergeometric differential equations, respectively. The only solution of (21), which is regular at $x = 0$, is

$$\psi^{(0)} = AM(\alpha, 2, x), \tag{23}$$

A being an arbitrary constant. $M(\alpha, 2, x)$ is the Kummer's function (Abramowitz & Stegun 1965), defined by the series

$$M(\alpha, 2, x) = \sum_{n=0}^{\infty} \frac{\Gamma(\alpha + n)}{\Gamma(\alpha)} \cdot \frac{x^n}{(n+1)!n!},$$

where Γ is the gamma function.

The solution of (22) can be found by variation of constants. The fundamental system of solutions of the considered differential equation is represented by the functions $M(\alpha, 2, x)$ and $U(\alpha, 2, x)$ (Abramowitz & Stegun), the latter being defined by the relation,

$$U(\alpha, 2, x) = \left\{ M(\alpha, 2, x) \lg x + \sum_{n=0}^{\infty} \frac{\Gamma(\alpha + n)}{\Gamma(\alpha)} \cdot \frac{x^n}{(n+1)!n!} \times \left[\frac{\Gamma'(\alpha + n)}{\Gamma(\alpha + n)} - 2 \frac{\Gamma'(2 + n)}{\Gamma(2 + n)} + \frac{1}{1 + n} \right] \right\} + \frac{1}{x\Gamma(\alpha)}.$$

The prime indicates the differentiation with respect to argument. The Wronskian $W\{M, U\}$ of the functions M and U is given by

$$W\{M, U\} = -\frac{e^x}{x^2\Gamma(\alpha)}. \tag{24}$$

On taking into account (23) and (24), the solution of (22) is obtained in the form,

$$\psi^{(1)} = \frac{1}{4}iA\Gamma(\alpha) \{U(\alpha, 2, x)I(M, x) - M(\alpha, 2, x)I(U, x)\}, \tag{25}$$

$$I(M, x) = \int_0^x M^2(\alpha, 2, y) e^{-y(y-\bar{\omega})^2} y dy,$$

$$I(U, x) = \int_0^x M(\alpha, 2, y) U(\alpha, 2, y) e^{-y(y-\omega)^2} y dy.$$

The complete solution of (20), with the considered accuracy, is given by the pertinent superposition of the solutions $\psi^{(0)}$ and $\psi^{(1)}$. Reapplying the transformation (19), and using (13) and the boundary condition (7), we determine the constant A . The final solution for the electric field component E_ϕ is then found as

$$E_\phi(\bar{r}) = B_{ex} i\bar{\omega} \frac{cA}{c} \exp \left\{ \frac{1}{2}\bar{\omega}(1 - \bar{r}^2) \right\} \bar{r} \times \frac{M(\alpha, 2, \bar{\omega}\bar{r}^2) \{1 - \frac{1}{4}ia\Gamma(\alpha) I(U, \bar{\omega}\bar{r}^2) + \frac{1}{4}ia\Gamma(\alpha) U(\alpha, 2, \bar{\omega}\bar{r}^2) I(M, \bar{\omega}\bar{r}^2)\}}{D(\bar{\omega}) \{1 - \frac{1}{4}ia\Gamma(\alpha) I(U, \bar{\omega})\} + \frac{1}{4}ia\Gamma(\alpha) F(\bar{\omega}) I(M, \bar{\omega})}, \tag{26a}$$

$$D(x) = (2-x)M(\alpha, 2, x) + 2x \frac{dM(\alpha, 2, x)}{dx}, \tag{26b}$$

$$F(x) = (2-x)U(\alpha, 2, x) + 2x \frac{dU(\alpha, 2, x)}{dx}. \tag{26c}$$

The magnetic field component B_z is easily obtained from (26) by means of (13). Thus,

$$B_z(\bar{r}) = B_{ez} \exp\left\{\frac{1}{2}\bar{\omega}(1-\bar{r}^2)\right\} \times \frac{D(\bar{\omega}\bar{r}^2)\left\{1 - \frac{1}{4}ia\Gamma(\alpha)I(U, \bar{\omega}\bar{r}^2)\right\} + \frac{1}{4}ia\Gamma(\alpha)F(\bar{\omega}\bar{r}^2)I(M, \bar{\omega}\bar{r}^2)}{D(\bar{\omega})\left\{1 - \frac{1}{4}ia\Gamma(\alpha)I(U, \bar{\omega})\right\} + \frac{1}{4}ia\Gamma(\alpha)F(\bar{\omega})I(M, \bar{\omega})}. \tag{27}$$

Finally, combining (13), (14) and (17) we find

$$B_\phi(\bar{r}) = \delta\bar{r}B_z(\bar{r}). \tag{28}$$

It is obvious that the amplitudes of the magnetic field components B_z and B_ϕ are given by $|B_z|$ and $|B_\phi|$, respectively. Figure 1 shows the variation of the B_z amplitude with the frequency $\bar{\omega}$, at the centre of the plasma column. The dashed lines represent the B_z amplitude in the case when the mass density is uniform over the column cross section and equals ρ_0^* (Cantieni & Schneider). We see that, in the non-uniform column, the magneto-acoustic resonance occurs at the frequency $\bar{\omega} \approx 3.5$ which is higher than the resonant frequency in the case of the uniform column ($\bar{\omega} \approx 2.4$). Moreover, we note that in the resonance, the magnetic field amplitudes are higher as compared with those corresponding to the uniform plasma. Figures 2 and 3 show the variation of the B_z and B_ϕ amplitudes with the radius \bar{r} , in the resonance. Here again the dashed lines correspond to the case of the uniform plasma. It is seen that the magnetic field component B_z is confined

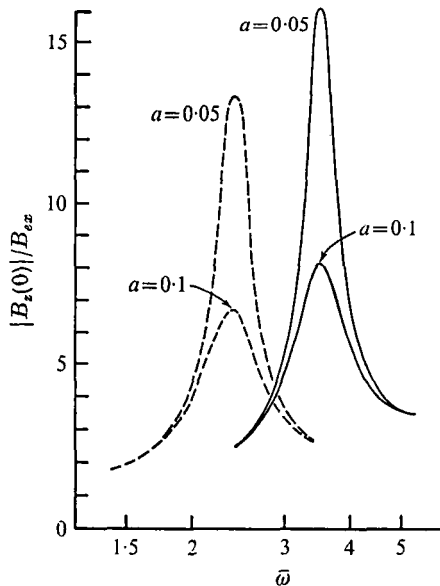


FIGURE 1. Variation of the axial component of magnetic field with frequency.

largely to the central region of the column, decaying quite rapidly towards the edges. On the other hand, the B_ϕ component is confined largely to the neighbourhood of the point $\bar{r} = 0.4$.

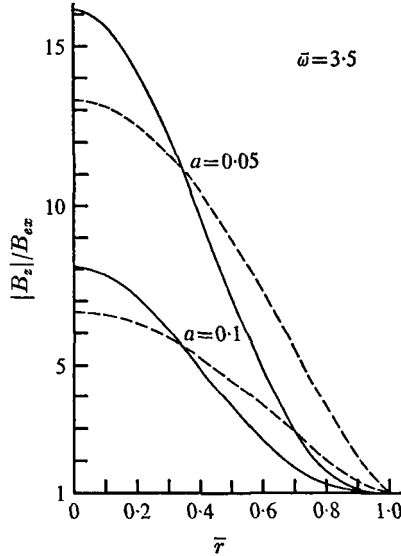


FIGURE 2. Variation of the axial component of magnetic field with radius.

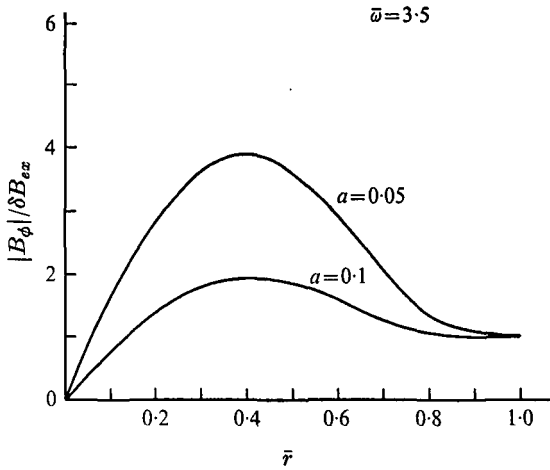


FIGURE 3. Variation of the azimuthal component of magnetic field with radius.

4. Non-uniform temperature case

In § 4 we consider the effects of a non-uniform temperature. For simplicity, we set $\delta = 0$ and assume the equilibrium plasma temperature to possess the following radial distribution:

$$T = T^*(1 - \epsilon \bar{r}^2) \quad (0 \leq \epsilon < 1), \tag{29}$$

T^* being the temperature at the centre of the column. According to Spitzer's (1956) formula for the conductivity, we can then write the approximate relation

$$\sigma_\perp = \sigma_\perp^*(1 - \epsilon \bar{r}^2)^{\frac{3}{2}}. \tag{30}$$

It is easily seen that (18) is now replaced by

$$\frac{d}{d\bar{r}} \frac{1}{\bar{r}} \frac{d}{d\bar{r}} \bar{r} E_\phi + \bar{\omega}^2 (1 - \bar{r}^2) \{1 + ia^* \bar{\omega} (1 - \bar{r}^2) (1 - \epsilon \bar{r}^2)^{-\frac{1}{2}}\} = 0, \tag{31}$$

where $a^* = c^2 / (4\pi\sigma_\perp R c_A)$.

On re-applying the procedure used in § 3 to (31), we find the following expression for the magnetic field:

$$B_z(\bar{r}) = B_{ez} \exp\left\{\frac{1}{2}\bar{\omega}(1 - \bar{r}^2)\right\} \times \frac{D(\bar{\omega}\bar{r}^2) \left\{1 - \frac{1}{4}ia^*\Gamma(\alpha) K(U, \bar{\omega}\bar{r}^2)\right\} + \frac{1}{4}ia^*\Gamma(\alpha) F(\bar{\omega}\bar{r}^2) K(M, \bar{\omega}\bar{r}^2)}{D(\bar{\omega}) \left\{1 - \frac{1}{4}ia^*\Gamma(\alpha) K(U, \bar{\omega})\right\} + \frac{1}{4}ia^*\Gamma(\alpha) F(\bar{\omega}) K(M, \bar{\omega})}, \tag{32a}$$

$$K(M, x) = \int_0^x M^2(\alpha, 2, y) e^{-\nu(y - \bar{\omega})^2} (1 - \epsilon y/\bar{\omega})^{-\frac{1}{2}} y dy, \tag{32b}$$

$$K(U, x) = \int_0^x M(\alpha, 2, y) U(\alpha, 2, y) e^{-\nu(y - \bar{\omega})^2} (1 - \epsilon y/\bar{\omega})^{-\frac{1}{2}} y dy. \tag{32c}$$

Figures 4 and 5 show the variation of the magnetic field amplitude with the frequency $\bar{\omega}$ and with the radius \bar{r} , respectively, for some typical values of the parameter ϵ . We note that the greater the temperature inhomogeneity, the smaller is the resonant amplitude of the magnetic field. Furthermore, we see that the effects of the non-uniform temperature are largely manifested in the central region of the column.

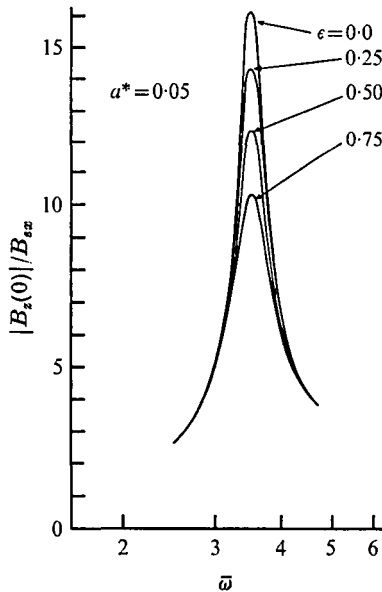


FIGURE 4. Variation of magnetic field with frequency in the case of non-uniform temperature.

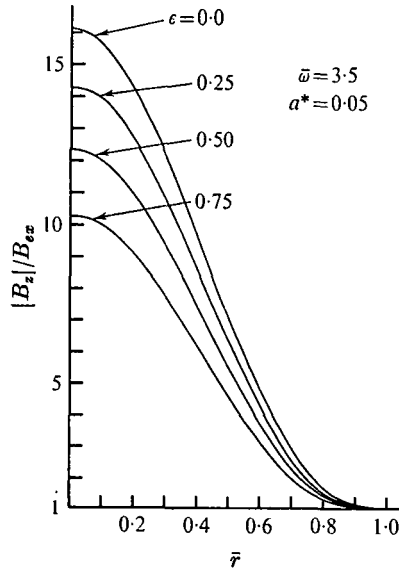


FIGURE 5. Variation of magnetic field with radius in the case of non-uniform temperature.

The author is grateful to Prof. O. Huber and Prof. H. Schneider for their continued interest in this work. The work was supported by the Swiss National Science Foundation.

REFERENCES

- ABRAMOWITZ, M. & STEGUN, I. A. 1965 *Handbook of Mathematical Functions*. Dover.
- BORODIN, A. V., KOVAN, I. A., RAKHIMBABAYEV, Y. R., RUSSANOV, V. D., SMIRNOV, V. P., FRANK-KAMENETSKII, D. A. 1963 *Nuc. Fus.* **3**, 38
- CANTIENI, E. & SCHNEIDER, H. 1963 *Helv. Phys. Acta*, **36**, 993.
- FAESSLER, K., VACLAVIK, J. & SCHNEIDER, H. 1969 *Helv. Phys. Acta*, **42**, 23.
- FRANK-KAMENETSKII, D. A. 1960a *Zh. exp. teor. fiz.* **39**, 669.
- FRANK-KAMENETSKII, D. A. 1960b *Zh. tekh. fiz.* **30**, 899.
- KOVAN, I. A., PATRUSHEV, B. I., RUSSANOV, V. D., TILININ, G. N., FRANK-KAMENETSKII, D. A. 1962 *Zh. exp. teor. fiz.* **43**, 16.
- SPITZER, L. 1956 *Physics of Fully Ionized Gases*. Interscience.