Self-learning Monte Carlo method and cumulative update in fermion systems

Junwei Liu, ¹ Huitao Shen, ¹ Yang Qi, ¹ Zi Yang Meng, ² and Liang Fu^{1,*}

¹Department of Physics, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA

²Institute of Physics, Chinese Academy of Sciences, Beijing 100190, China

(Received 8 December 2016; published 7 June 2017)

We develop the self-learning Monte Carlo (SLMC) method, a general-purpose numerical method recently introduced to simulate many-body systems, for studying interacting fermion systems. Our method uses a highly efficient update algorithm, which we design and dub "cumulative update", to generate new candidate configurations in the Markov chain based on a self-learned bosonic effective model. From a general analysis and a numerical study of the double exchange model as an example, we find that the SLMC with cumulative update drastically reduces the computational cost of the simulation, while remaining statistically exact. Remarkably, its computational complexity is far less than the conventional algorithm with local updates.

DOI: 10.1103/PhysRevB.95.241104

Monte Carlo (MC) method is an unbiased numerical tool that obtains statistically exact results by sampling configurations according to a probability distribution [1–7]. Configurations may be generated sequentially through the reversible Markov process obeying the detailed balance principle (DBP). In order for an MC simulation to be efficient, the process of generating a new configuration from the current one should be fast, and consecutive configurations should be uncorrelated. The performance of conventional MC method is severely impeded when either of the two conditions fails.

Recently, we introduced a new method dubbed "self-learning Monte Carlo" (SLMC), to speed up configuration updates in MC simulations [8]. SLMC consists of two stages: learning and simulating. First, we perform trial simulations with the conventional local update method to generate a large set of configurations along with their weights, which serve as the training data. By fitting the configuration distribution, we learn an effective Hamiltonian $H_{\rm eff}$ for the system, which can be simulated faster than the original Hamiltonian H. Next, in performing the actual simulation, we use $H_{\rm eff}$ to propose smart moves in configuration space. The acceptance of proposed moves is set properly to satisfy the detailed balance condition of the original Hamiltonian, ensuring the simulation is statistically exact.

SLMC is a general-purpose method rooted in the philosophy "first learn, then earn". In our previous work [8], SLMC is implemented for classical statistical models near second-order phase transitions. In such systems, conventional MC simulation suffers from critical slowing down because successive configurations are highly correlated [9,10]. We showed that SLMC method can significantly reduce the autocorrelation time. As an example, SLMC simulation on a generalized Ising model is found to be 10–20 times faster.

In this work, we develop a generic SLMC method for simulating interacting fermions as well as mixed Bose-Fermi systems, and demonstrate its enormous power. By theoretical analysis and numerical simulation, we show that SLMC method *generally* reduces the complexity of simulating fermion systems, thus achieving a tremendous speedup that

grows with system size. The central component of our method is a highly efficient "cumulative update" algorithm, which we introduce for updating field configurations to which fermions couple.

Below we first present SLMC method with cumulative update in full generality and theoretically analyze its complexity, i.e., determine the scaling of the computational cost with system size. Next, we apply our method to the double exchange model, which describes itinerant electrons interacting with localized spins. Our method achieves a speedup of at least $O(L^d)$, where L is the system size and d is the spatial dimension. For the double exchange model on $8 \times 8 \times 8$ cubic lattice, SLMC simulation is numerically shown to be over 10^3 times faster than conventional method.

In quantum MC simulations of interacting fermion systems, we may employ the Hubbard-Stratonovich transformation [11,12] to write the partition function in the form of fermions coupled to fluctuating fields [13–16]

$$Z = \sum_{\phi(\tau)} \det \left[\mathbf{I} + \prod_{\tau} e^{-\Delta \tau H_f[\phi(\tau)]} \right] \equiv \sum_{\phi(\tau)} W[\phi(\tau)], \quad (1)$$

where ϕ is a space- and time-dependent field, $H_f[\phi]$ is the single-particle Hamiltonian of fermions moving in the background of the field ϕ , $\Delta \tau = \beta/N$ is the duration of time slice in Trotter decomposition [17,18] and $\tau = n\Delta \tau$ with $n=1,\ldots,N$. In the MC simulation, we sample configurations of ϕ with weights $W[\phi]$, which is determined by integrating out fermion degrees of freedom. Partition functions of the form Eq. (1) also appear in systems of fermions interacting with dynamical boson fields such as spins or phonons [19–24].

The partition function Eq. (1) can also describe systems of fermions coupled to classical spins or other classical degrees of freedom. A well-known example is the double exchange model [25–27] describing itinerant electrons coupled to a lattice of localized spins, which are represented by classical vectors of unit length. In such cases, the field ϕ in Eq. (1) is space-dependent but *time-independent*. The partition function (1) then simplifies to

$$Z = \sum_{\phi} \det[\mathbf{I} + e^{-\beta H_f[\phi]}] \equiv \sum_{\phi} W[\phi]. \tag{2}$$

^{*}liangfu@mit.edu

Computing the weight W in both Eqs. (1) and (2) involves calculating the fermion determinant. This task is very time consuming and its computational cost grows polynomially with system size, which is a major bottleneck of MC simulations in fermion systems.

In this work, we focus on models of fermions coupled to classical fields, where ϕ is time-independent. Our treatment here can be straightforwardly applied to models with timedependent fields, which we leave to a forthcoming work [28]. For the partition function in Eq. (2), the weight Wcan be calculated as $\prod_{n} (1 + e^{-\beta E_n(\phi)})$, where $\{E_n(\phi)\}$ is the single-particle energy spectrum obtained by the exact diagonalization (ED) of the fermion Hamiltonian $H_f[\phi]$ for a given ϕ configuration. Performing this ED has the complexity of $O(L^{3d})$. (Algorithms that compute $W[\phi]$ approximately can be faster [29–34].) In the conventional MC simulation using the Metropolis algorithm, the ED needs to be performed every time ϕ is updated on a single site. Therefore the computational cost for each full sweep of the lattice in the simulation grows as $O(L^{4d})$. Assuming the configurations become uncorrelated after τ_0 iterations of such sweeps, the total cost of generating two successive, statistically independent configurations is $O(\tau_0 L^{4d})$.

In contrast, in SLMC simulation, we first generate a set of configurations using the local update, and fit their weights using an effective model $H_{\text{eff}}[\phi]$ for the field ϕ such that

$$W[\phi] \simeq e^{-\beta H_{\text{eff}}[\phi]}.$$
 (3)

Typically, we take $H_{\rm eff}[\phi]$ to have the form of a power series of ϕ , and determine the coefficients from multilinear fitting. $H_{\rm eff}[\phi]$ is to be viewed as an approximation to the exact Hamiltonian for the ϕ field after integrating out the fermions $H[\phi] \equiv -\frac{1}{\beta} \ln W[\phi]$, because by construction the Boltzmann distribution of $H_{\rm eff}[\phi]$ approximately reproduces the desired distribution $W[\phi]$ for dominant ϕ -field configurations. Importantly, for a given ϕ , evaluating $H_{\rm eff}[\phi]$ whose explicit expression has been learned from the fitting is much faster than numerically computing $W[\phi]$ exactly.

Next, in performing the actual simulation, we design a "cumulative update" algorithm to update the ϕ -field configurations efficiently by the guidance of the effective model $H_{\rm eff}[\phi]$. Starting from the last configuration reached in the Markov chain of $H[\phi]$, denoted by ϕ_A , we propose a global move by performing a sequence of local updates as one would do in simulating the effective model $H_{\mathrm{eff}}[\phi]$: each local move attempts to change the value of ϕ on a random site, and its acceptance probability is determined by the detailed balance condition of $H_{\text{eff}}[\phi]$. A sequence of such local updates is performed to generate a new field configuration ϕ_B that is sufficiently uncorrelated with ϕ_A . Then we propose ϕ_B as the candidate configuration for the next state in the Markov chain of $H[\phi]$. As we will describe below, the probability of accepting this move $\phi_A \to \phi_B$ should be designed properly to ensure that SLMC method is statistically exact.

The advantage of the cumulative update algorithm comes from replacing many iterations of "expensive" local updates according to the exact weight $W[\phi]$, which is necessary to obtain two statistically independent configurations in the conventional MC simulation, with cumulative local updates

according to the approximate weight $e^{-\beta H_{\text{eff}}[\phi]}$, which is numerically much faster.

We now derive the desired probability of accepting the proposed move $\phi_A \to \phi_B$ in a single step of the cumulative update. As in the general Metropolis-Hastings algorithms [2], our update scheme consists of two stages: first, a candidate configuration ϕ_B is generated from ϕ_A through a sequence of local updates; second, the update $\phi_A \to \phi_B$ is accepted with a probability $p(A \to B)$. The Markov-chain transition probability $P(A \to B)$ is the product of the probability of generating the particular configuration ϕ_B among all the possibilities in the first stage, denoted by $S(A \to B)$, and the probability of accepting ϕ_B , $p(A \to B)$. The detailed balance condition for $P(A \to B)$ requires

$$\frac{P(A \to B)}{P(B \to A)} = \frac{S(A \to B)}{S(B \to A)} \frac{p(A \to B)}{p(B \to A)} = \frac{W(B)}{W(A)}, \quad (4)$$

where $W(A) \equiv W[\phi_A]$ is the exact weight of the configuration ϕ_A . Moreover, by construction, the ratio $S(A \to B)/S(B \to A)$ is set by the detailed balance condition of the effective model

$$\frac{S(A \to B)}{S(B \to A)} = \prod_{i=0}^{l_c - 1} \frac{\tilde{P}(C_i \to C_{i+1})}{\tilde{P}(C_{i+1} \to C_i)}$$

$$= \prod_{i=0}^{l_c - 1} e^{-\beta(\tilde{E}_{i+1} - \tilde{E}_i)} = e^{-\beta(\tilde{E}_B - \tilde{E}_A)}, \quad (5)$$

where $\phi_{C_0} \equiv \phi_A$ and $\phi_{C_n} \equiv \phi_B$. $\tilde{P}(C_i \rightarrow C_{i+1})$ is the transition probability for *i*-th local update of the effective Hamiltonian $H_{\text{eff}}[\phi]$, and $\tilde{E}_i \equiv H_{\text{eff}}(\phi_{C_i})$ is the energy of the configuration ϕ_{C_i} in the effective model. Combining Eqs. (4) and (5), we find the desired probability of accepting the candidate configuration found through cumulative update,

$$p(A \to B) = \min\{1, e^{-\beta(E_B - \tilde{E}_B) - (E_A - \tilde{E}_A)}\},$$
 (6)

where E is the energy of the exact Hamiltonian $H[\phi]$. In the ideal case when $H_{\text{eff}}[\phi] = H[\phi]$, p = 1.

The computational cost of one complete cumulative update step contains two parts: (1) the cost of local updates based on H_{eff} , which generally has the complexity $O(l_c)$, where l_c denotes the number of local updates; and (2) the cost of ED to compute the acceptance probability in Eq. (6), which has the complexity $O(L^{3d})$. Hence the total cost is $O(l_c) + O(L^{3d})$. In order to obtain two uncorrelated configurations in a single cumulative update step, we take l_c to be of the order of $\tau_0 L^d$, with τ_0 to be the autocorrelation time (a unit τ_0 corresponds to a single full sweep of L^d sites). For systems away from critical points, τ_0 is taken as a constant, and close to critical points τ_0 scales as L^z with z empirically found to be around 2 [9,10]. Thus the cost of the ED operation almost always dominates in the cumulative update process. Assuming that the effective model is accurate enough such that the proposed global move $\phi_A \rightarrow \phi_B$ is accepted with a probability of the order of 100%, a cumulative update generates a statistically independent configuration with a computational cost of a single ED step, i.e., $O(L^{3d})$. Compared to the complexity of $O(\tau_0 L^{4d})$ for the conventional local update, this gives a speedup of $O(\tau_0 L^d) = O(L^{d+\eta})$ where $\eta = 0$ for noncritical systems and $\eta = z$ for critical systems.

We notice that there are faster methods that evaluate the weight in Eq. (2) *approximately* without performing an ED [29–34]. These methods are not exclusive from the SLMC methods, and by incorporating these approximations, we can further reduce the complexity of SLMC.

In the rest of this work, we numerically demonstrate the power of SLMC by simulating the double exchange model [25–27]:

$$\hat{H} = -t \sum_{\langle ij \rangle \alpha} (\hat{c}_{i\alpha}^{\dagger} \hat{c}_{j\alpha} + \text{h.c.}) - \frac{J}{2} \sum_{i \alpha \beta} \vec{S}_{i} \cdot \hat{c}_{i\alpha}^{\dagger} \vec{\sigma}_{\alpha\beta} \hat{c}_{i\beta}, \quad (7)$$

where $\langle ij \rangle$ is the summation over the nearest-neighbor pairs of sites. $\hat{c}_{i\sigma}$ is the fermion annihilation operator of spin σ at site i, and $\vec{\sigma}$ are the Pauli matrices for the fermion spin operator. \vec{S}_i is a classical vector of unit length on site i, representing the localized spin that couples to the fermion on the same site. Despite the absence of direct interactions in this model between localized spins on different sites, the presence of itinerant fermions effectively generates a RKKY-type [35–37] interaction $g_{ij}\vec{S}_i \cdot \vec{S}_j$, where g_{ij} decays with the distance between site i and j.

Below we simulate the double exchange model on the threedimensional (3D) cubic lattice. Previous numerics have shown that for J/t=16 and at filling $\nu=1/4$, this model exhibits a phase transition from paramagnet to ferromagnet at a critical temperature $T_c/t \sim 0.12$ [31]. We focus on simulations around T_c , where the speedup of SLMC is expected to be maximum.

Following the general procedure of the SLMC, we first perform trial simulations to train an effective model that includes *all* two-body interactions between two localized spins, preserving translational and spin-rotational symmetries

$$H_{\text{eff}} = E_0 - J_1 \sum_{\langle ij \rangle_1} \vec{S}_i \cdot \vec{S}_j - J_2 \sum_{\langle ij \rangle_2} \vec{S}_i \cdot \vec{S}_j - \cdots,$$
 (8)

where $\langle ij \rangle_n$ is the summation over nth nearest-neighbor pairs of localized spins. Denoting $C_n^{\alpha} = \sum_{\langle ij \rangle_n} \vec{S}_i \cdot \vec{S}_j$, we train E_0 and J_n through a multi-linear regression with $E^{\alpha} = E_0 + \sum_n J_n C_n^{\alpha}$. To speed up and improve the fitting, we have exploited the reinforcement learning strategy here, i.e., we use the SLMC method on the trained model to generate more uncorrelated spin configurations efficiently, and then further optimize parameters in the effective model using these new training data. This reinforced learning process can be iterated a few times until the desired parameters converge.

The fitting result for L=4 is shown in the Fig. 1. We see a clear RKKY profile of the spin-spin interaction mediated by itinerant fermions: the coupling strength J_n decays and oscillates with the distance between two localized spins. To demonstrate how good the fitting is, we sample 10^5 independent configurations on the Markov chain according to the Boltzmann distribution of the original model and of the effective model. As shown in the insets of Fig. 2, these two Boltzmann distributions look nearly identical.

To quantify the accuracy of the effective model Eq. (8), we measure the differences in the energies of $H_{\rm eff}$ (denoted by \tilde{E}) and of H (denoted by E) for sample configurations (denoted by S) of localized spins drawn from the Boltzmann

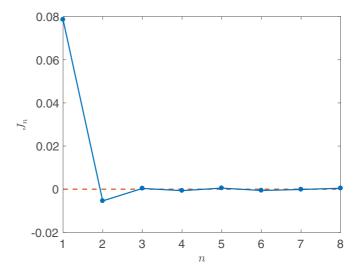


FIG. 1. The trained parameters J_n for the effective model in Eq. (8) for L=4.

distribution $e^{-\beta H}$, known as residuals in statistical analysis: $\mathcal{E}_S \equiv \tilde{E}_S - E_S$. As shown in Fig. 2, the distribution of \mathcal{E} is Gaussian, with a peak centered at $\mathcal{E} = 0$. The narrow width of the peak demonstrates the high accuracy of our effective model. In the ideal case where $H_{\rm eff} = H$, the distribution of \mathcal{E} becomes a delta function.

We further use the statistical measure R^2 , also called the coefficient of determination, as a figure of merit ("score") characterizing how well the effective model replicates the Boltzmann distribution of the original model. This score is defined through the sum of squares of residuals:

$$R^2 = 1 - \frac{\sum_{S} \mathcal{E}_S^2}{\sum_{S} (E_S - \langle E \rangle)^2},$$
 (9)

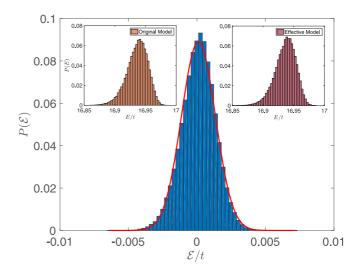


FIG. 2. For 10^5 independent spin configurations on the Markov chain of H, the blue histogram is the distribution of the energy difference $\mathcal{E}(S)$ of these configurations between the original model H(S) and the effective model $H_{\rm eff}(S)$. The red curve is the fitted Gaussian distribution with $\sigma=0.012$. The upper-left and upper-right insets are distributions of E(S) and $E_{\rm eff}(S)$, respectively.

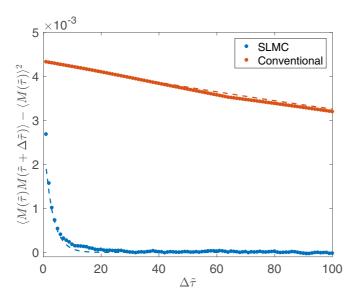


FIG. 3. The autocorrelation function $C(\tilde{\tau}) = \langle M(\tilde{\tau})M(\tilde{\tau} + \Delta \tilde{\tau}) \rangle - \langle M(\tilde{\tau}) \rangle^2$ of the magnetization with $\tilde{\tau}$ (number of EDs) as the time unit for both the SLMC and the conventional method on the system with L=4. Here, for SLMC we choose $l_c=16\times 4^3$. The dotted curves are the fitted $\tilde{\tau}_0$ based on the equation $C(\tilde{\tau})=C(0)\exp(-\tilde{\tau}/\tilde{\tau}_0)$.

and ranges from 0 to 100%. A perfect effective model $H_{\rm eff} = H$ has $R^2 = 100\%$. The scores of our effective model for the double exchange Hamiltonian with L = 4,6,8 are $R^2 = 99.5\%$, 99.8%, and 99.9% respectively.

Finally, we show the actual speedup of our SLMC simulation. We use the autocorrelation time of SLMC and conventional method to gauge their performance. As well known, the statistical error of a MC simulation is proportional to $1/\sqrt{N_{\rm MC}}$, where $N_{\rm MC}$ is the number of statistically independent samples. Hence, to achieve a given level of statistical error, the total number of MC steps needed in the simulation is proportional to the autocorrelation time, which measures the average number of MC steps needed to obtain another independent sample. Here, since ED calculations dominate the computational cost in both the conventional and the SLMC method, we use the number of EDs denoted by $\tilde{\tau}$, instead of the number of sweeps τ , as the unit to measure the autocorrelation time. Figure 3 shows the decay of autocorrelation functions with $\tilde{\tau}$ as the time unit for both methods. From these we extract the autocorrelation time $\tilde{\tau}_0$, which precisely characterizes the actual computational cost to obtain two uncorrelated spin configurations. For L=4, $\tilde{\tau}_0$ is over 150 times shorter in SLMC than in the conventional method.

In Fig. 4, we demonstrate the remarkable efficiency of the SLMC method as the system size increases. As discussed previously, in the conventional MC method $\tilde{\tau}_0 \sim O(L^{d+z})$ scales extremely fast with system size L due to the cost of EDs, as confirmed by our expensive numerics for L=4 and 6. In contrast, in SLMC method with an accurate effective

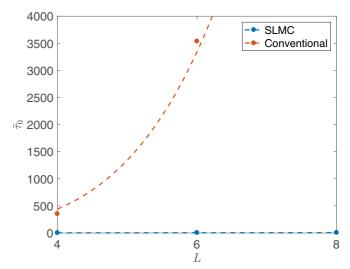


FIG. 4. The autocorrelation time $\tilde{\tau}_0$ for both the SLMC and the conventional method on systems with L=4, 6, and 8. The red dashed curve, obtained through extrapolating the two data points with the well-known dynamical exponent $z\simeq 2$, shows the scaling behavior of the conventional method [38].

model, we expect $\tilde{\tau}_0 \sim O(1)$. Indeed, our SLMC simulation shows $\tilde{\tau}_0$ hardly scales with the system size. For L=8, $\tilde{\tau}_0$ in the conventional method is estimated to be $\sim 10^4$ from extrapolation based on the scaling, whereas in SLMC it is numerically found to be ~ 6 , leading to an enormous speedup by $\sim 10^3$ times.

To conclude, we developed a general-purpose, statistically exact SLMC method with a cumulative update algorithm for simulating fermion systems, whose computational complexity is significantly reduced compared to the conventional MC method. We believe our method holds great promise for solving a wide class of challenging many-body problems of fundamental interest.

We thank Xiao Yan Xu for helpful discussions. This work is supported by the DOE Office of Basic Energy Sciences, Division of Materials Sciences and Engineering under Award DE-SC0010526. L.F. is supported partly by the David and Lucile Packard Foundation. J.L. is supported partly by S3TEC Solid State Solar Thermal Energy Conversion Center, an Energy Frontier Research Center funded by the U.S. Department of Energy (DOE), Office of Science, Basic Energy Sciences (BES), under Award No. DE-SC0001299/DE-FG02-09ER46577. H.S. is supported by MIT Alumni Fellowship Fund For Physics. Z.Y.M. is supported by the Ministry of Science and Technology (MOST) of China under Grant No. 2016YFA0300502, the National Natural Science Foundation of China (NSFC Grants No. 11421092 and No. 11574359), and the National Thousand-Young-Talents Program of China. J.L. and H.S. contributed equally to this work.

^[1] N. Metropolis, A. W. Rosenbluth, M. N. Rosenbluth, A. H. Teller, and E. Teller, J. Chem. Phys. **21**, 1087 (1953).

^[2] W. K. Hastings, Biometrika 57, 97 (1970).

^[3] The Monte Carlo Method in Condensed Matter Physics, edited by K. Binder (Springer-Verlag Berlin Heidelberg, 1995).

- [4] M. E. J. Newman and G. T. Barkema, *Monte Carlo Methods in Statistical Physics* (Oxford University Press, 1999).
- [5] The Monte Carlo Method in the Physical Sciences: Celebrating the 50th Anniversary of the Metropolis Algorithm, edited by J. E. Gubernatis, AIP Conference Proceedings Vol. 690 (AIP Publishing, Los Alamos, New Mexico, USA, 2003).
- [6] D. Landau and K. Binder, A Guide to Monte Carlo Simulations in Statistical Physics (Cambridge University Press, 2005).
- [7] A. W. Sandvik, AIP Conf. Proc. 1297, 135 (2010).
- [8] J. Liu, Y. Qi, Z. Y. Meng, and L. Fu, Phys. Rev. B 95, 041101(R) (2017).
- [9] R. H. Swendsen and J.-S. Wang, Phys. Rev. Lett. 58, 86 (1987).
- [10] U. Wolff, Phys. Rev. Lett. 62, 361 (1989).
- [11] R. Stratonovich, Soviet Physics Doklady (1957), Vol. 2, p. 416.
- [12] J. Hubbard, Phys. Rev. Lett. 3, 77 (1959).
- [13] R. Blankenbecler, D. J. Scalapino, and R. L. Sugar, Phys. Rev. D 24, 2278 (1981).
- [14] J. E. Hirsch, Phys. Rev. B **31**, 4403 (1985).
- [15] J. E. Hirsch and R. M. Fye, Phys. Rev. Lett. **56**, 2521 (1986).
- [16] S. R. White, D. J. Scalapino, R. L. Sugar, E. Y. Loh, J. E. Gubernatis, and R. T. Scalettar, Phys. Rev. B 40, 506 (1989).
- [17] H. F. Trotter, Proc. Am. Math. Soc. 10, 545 (1959).
- [18] M. Suzuki, Commun. Math. Phys. 51, 183 (1976).
- [19] E. Berg, M. A. Metlitski, and S. Sachdev, Science 338, 1606 (2012)
- [20] S. Lederer, Y. Schattner, E. Berg, and S. A. Kivelson, Phys. Rev. Lett. 114, 097001 (2015).

- [21] Z.-X. Li, F. Wang, H. Yao, and D.-H. Lee, arXiv:1512.04541.
- [22] Y. Schattner, M. H. Gerlach, S. Trebst, and E. Berg, Phys. Rev. Lett. 117, 097002 (2016).
- [23] Y. Schattner, S. Lederer, S. A. Kivelson, and E. Berg, Phys. Rev. X 6, 031028 (2016).
- [24] X. Y. Xu, K. S. D. Beach, K. Sun, F. F. Assaad, and Z. Y. Meng, Phys. Rev. B 95, 085110 (2017).
- [25] C. Zener, Phys. Rev. 82, 403 (1951).
- [26] P. W. Anderson and H. Hasegawa, Phys. Rev. 100, 675 (1955).
- [27] P. G. de Gennes, Phys. Rev. 118, 141 (1960).
- [28] X.-Y. Xu, Y. Qi, J. Liu, L. Fu, and Z.-Y. Meng (unpublished).
- [29] Y. Motome and N. Furukawa, J. Phys. Soc. Jpn. **68**, 3853 (1999).
- [30] J. Alonso, L. Fernández, F. Guinea, V. Laliena, and V. Martín-Mayor, Nucl. Phys. B 596, 587 (2001).
- [31] G. Alvarez, C. Şen, N. Furukawa, Y. Motome, and E. Dagotto, Comput. Phys. Commun. 168, 32 (2005).
- [32] S. Kumar and P. Majumdar, Eur. Phys. J. B **50**, 571 (2006).
- [33] S. Kumar and J. van den Brink, Phys. Rev. Lett. 105, 216405 (2010).
- [34] A. Mukherjee, N. D. Patel, C. Bishop, and E. Dagotto, Phys. Rev. E 91, 063303 (2015).
- [35] M. A. Ruderman and C. Kittel, Phys. Rev. 96, 99 (1954).
- [36] T. Kasuya, Prog. Theor. Phys. 16, 45 (1956).
- [37] K. Yosida, Phys. Rev. 106, 893 (1957).
- [38] We cannot generate more data points because the computation is too costly with the conventional method. Fortunately, the computational cost of the conventional method is well-established and can be estimated quite reliably.