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Carbonate alteration of ophiolitic rocks in the Arabian-Nubian Shield of Egypt: Sources and compositions of the carbonating fluid and implications for the formation of Au deposits

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3 **Carbonate alteration of ophiolitic rocks in the Arabian-Nubian Shield of**
4 **Egypt: Sources and compositions of the carbonating fluid and implications**
5 **for the formation of Au deposits**
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Abstract

Ultramafic portions of ophiolitic fragments in the Arabian-Nubian Shield (ANS) show pervasive carbonate alteration forming various degrees of carbonated serpentinites and listvenitic rocks. Despite the extent of the alteration, little is known about the processes that caused it, the source of the CO₂, or the conditions of alteration. This study investigates the mineralogy, stable (O, C) and radiogenic (Sr) isotope composition, and geochemistry of suites of variably carbonate altered ultramafics from the Meatiq area of the Central Eastern Desert (CED) of Egypt. The samples investigated include least-altered lizardite (Lz) serpentinites, antigorite (Atg) serpentinites, and listvenitic rocks with associated carbonate and quartz veins. The C, O and Sr isotopes of the vein samples cluster between -8.1‰ to -6.8‰ for δ¹³C, +6.4‰ to +10.5‰ for δ¹⁸O, and ⁸⁷Sr/⁸⁶Sr of 0.7028 to 0.70344, and plot within the depleted mantle compositional field. The serpentinites isotopic compositions plot on a mixing trend between the depleted-mantle and sedimentary carbonate fields. The carbonate veins contain abundant carbonic (CO₂±CH₄±N₂) and aqueous-carbonic (H₂O-NaCl-CO₂±CH₄±N₂) low salinity fluid, with trapping conditions of 270°C to 300°C and 0.7 to 1.1 kbar. The serpentinites are enriched in Au, As, S and other fluid-mobile elements relative to primitive and depleted mantle. The extensively carbonated Atg-serpentinites contain significantly lower concentrations of these elements than the Lz-serpentinites suggesting that they were depleted during carbonate alteration. Fluid inclusion and stable isotope compositions of Au deposits in the CED are similar to those from the carbonate veins investigated in the study and we suggest that carbonation of ANS ophiolitic rocks due to influx of mantle-derived CO₂-bearing fluids caused break down of Au-bearing minerals such as pentlandite, releasing Au and S to the hydrothermal fluids that later formed the Au-deposits. This is the first time that Au has been observed to be remobilized from rocks during the lizardite-antigorite transition.

Keywords: Arabian-Nubian Shield, serpentinite, Listvenite, Carbonation, Gold, Fluid-mobile elements, Stable isotopes, Sr isotopes

1. Introduction

Carbonation of mafic and ultramafic rocks is an important geological process with implications that include carbon sequestration (Lackner et al. 1995; Selfritz 1990; Kelemen and Matter 2008), the carbon cycle and fluxing of CO₂ to the Earth's atmosphere (Kerrick and Caldiara 1998; Lasaga et al. 2001; Skelton 2011), and formation of ore deposits (Barnes et al. 1973; Groves et al. 1998; Phillips and Evans 2004). Carbonation occurs when susceptible rocks in the crust and mantle interact with CO₂-rich fluids resulting in alteration and precipitation of carbonate and other minerals. Mafic and ultramafic rocks are particularly prone to carbonation due to abundant olivine [(Mg,Fe)₂SiO₄] and pyroxene

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3 [(Ca,Mg,Fe)₂Si₂O₆], which react with H₂O and CO₂ to form hydrous silicates such as
4 serpentine, Fe-oxides (magnetite) and carbonates (Kelemen and Matter 2008). Carbonation of
5 mafic rocks has been used to calculate fluxes of CO₂ to the Earth's atmosphere in the
6 geological past (Skelton 2011), and may be linked to formation of Au deposits worldwide
7 (Groves et al. 1998).
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10 Ophiolites of Arabian-Nubian Shield (ANS) are fragments of the Mozambique Oceanic crust
11 generated in a suprasubduction zone setting during Neoproterozoic time (Stern et al. 2004;
12 Azer and Stern 2007). The ophiolites constitute one of the most distinctive rock sequences
13 among ANS basement rocks (Fig. 1). The ophiolitic rocks are commonly altered due in part
14 to interaction with migrating carbonate-rich solutions (Stern and Gwinn 1990). Carbonation
15 of the ophiolite units is extensive and indicates a large flux of CO₂-bearing fluid. However
16 there have been few investigations into the sources and composition of carbonating fluids, or
17 the conditions at which this alteration occurred. The carbonation has been focused along
18 faults and shear zones and has resulted in formation of variably foliated talc-rich rocks
19 previously referred to as 'talc-carbonate schists', 'Barramiya rocks', 'listvenite-like rocks'
20 (Azer 2013) or 'talc-rich rocks' (Stern et al. 2004; Ali-Bik et al. 2012).
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26 The ophiolitic rocks in some parts of the ANS such as the Central Eastern Desert (CED) of
27 Egypt contain abundant Au deposits (Fig. 1) that have been mined since Pharaonic times
28 (Harraz 2000; Klemm and Klemm 2013). A significant proportion of these deposits occur in
29 or near carbonated ultramafic rocks (Osman 1995; Botros 2002, 2004; Abd El-Rahman et al.
30 2012; Zoheir and Lehmann 2011). Despite this clear spatial relationship, the genetic
31 relationship between carbonation, deformation and Au deposition remains poorly understood.
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35 This study investigates the sources and compositions of carbonate altered serpentinites in the
36 Meatiq area of the CED. We use stable (C, O) and radiogenic (Sr) isotopes to determine the
37 source of CO₂-rich fluids, and fluid inclusion microthermometry and mineral chemistry to
38 constrain the conditions of alteration. This study also investigates the element mobility that
39 has occurred during the carbonate alteration and discusses the implications of these results for
40 the formation of Au deposits in the CED.
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44 2. Geological Setting

45 The Arabian- Nubian Shield represents the northern part of East African Orogen (EAO)
46 which formed in Tonian and Cryogenian time accompanying terrane accretion around the
47 Mozambique Ocean (Stern 1994, 2002). The Egyptian segment of the ANS is exposed in the
48 Eastern Desert (ED) and can be subdivided into four main lithotectonic units: (i) a lower unit
49 "infrastructure" composed of gneiss and deformed granitoids, (ii) a structurally overlying
50 "suprastructure" composed of metamorphosed ophiolitic and island arc assemblages, (iii) a
51 sequence of non-metamorphosed to weakly metamorphosed and deformed volcanic (Dokhan
52 Volcanics) and sedimentary rocks (Hammamat Group) that unconformably overlie the
53 suprastructure in places, and (iv) post-orogenic granitic plutons that commonly intrude the
54 other units (Andresen et al. 2009, 2010). Over 30 different locations of ophiolites have been
55 mapped in the Egyptian segment of the ANS. (Fig. 1; Stern et al. 2004). These consist of
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3 tectonically disrupted fragments of mantle serpentinized ultramafic rocks and crustal gabbro,
4 sheeted dykes and pillow basalts (Stern et al. 2004). CED ophiolitic rocks are strongly
5 deformed and metamorphosed to **conditions of** lower greenschist and amphibolite facies
6 (Johnson et al. 2004, Stern et al. 2004). Radiometric ages of ANS ophiolites range between
7 880 and 740 Ma (Ries et al. 1983; Kröner 1985; Kröner et al. 1992, 1994; Stern 1994;
8 Loizenbauer et al. 2001), with two age maxima from about 850 to 780 Ma and from 750 to
9 700 Ma (Stern 1994).

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12 [Insert figure 1 near here]

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15 The study area **is known as Meatiq dome which** is located about 40 km west of **the** Red Sea
16 (Fig. 1 and 2). The Meatiq **dome** is a large, c. 500 km² quartzofeldspathic gneiss **complex**
17 surrounded by **low-grade** ophiolitic melange (Andresen et al. 2010). The eastern and western
18 dome margins are bounded by sinistral strike-slip shear zones which belong to the **Najd Fault**
19 **System** whereas the northern and southern margins are defined by prominent normal faults
20 (Fig. 2, Wallbrecher et al. 1993). Um Ba'anib Orthogneiss that comprises the core of the
21 Meatiq dome consists of coarse-grained, foliated orthogneisses that become gradually more
22 mylonitized and fine-grained at shallower levels, forming a garnet-bearing mylonitic
23 carapace (Andresen et al. 2009). The Um Ba'anib orthogneiss and the upper metasediments
24 are often considered to be "infrastructure" (Habib et al. 1985; Neumayr et al. 1996, 1998).
25 The protolith age of the Um Ba'anib Orthogneiss is 631 Ma, whereas the undeformed Arieki
26 granite **intruded the succession at** 590 Ma (Andresen et al. 2009).

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28 A variably mylonitized and metamorphosed package of ortho- and paragneisses, overlies the
29 Um Ba'anib Orthogneiss. These are grouped as mylonitized gneisses and called Abu Fananni
30 Thrust Sheet by Habib et al. (1985). A detailed description of the metamorphic evolution of
31 the Abu Fannani Thrust Sheet is given by Neumayr et al. (1996, 1998) and Loizenbauer et al.
32 (2001). The mylonitized paragneisses are intruded in south by the syn-tectonic Abu Ziran
33 diorite which yields a U-Pb zircon and titanite age of 606 Ma (Andresen et al. 2009). **The**
34 Abu Ziran diorite is generally undeformed, but the western margin of the pluton contains a
35 >50m wide ductile shear zone (Andresen et al. 2009).

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37 The structurally overlying ophiolitic and island arc assemblages are considered to be
38 "suprastructure" (Habib et al. 1985; Neumayr et al. 1996, 1998), Pan-African nappe complex
39 (El-Gaby et al. 1988; Bregar et al. 2002) or eugeoclinal allochthon (Andresen et al. 2009,
40 2010), and are characterized by greenschist- to lower amphibole facies mineral assemblages
41 (Neumayr et al. 1998). A high-strain zone separates the amphibolite-facies orthogneisses in
42 the core of the **high-grade metamorphic** domes from the surrounding lower grade ophiolitic
43 rocks (Sturchio et al. 1983; Ries et al. 1983; Habib et al. 1985). Ophiolitic rocks of the
44 Meatiq area comprise mainly serpentinites **and their altered products such as** listvenitic rocks,
45 metagabbros and metabasalts. **Strictly speaking, the term listvenite (also spelled listvanite,**
46 **listwanite, listwaenite; Kelemen et al. 2011) refers to fuchsite-quartz-carbonate lithologies**
47 **derived from ultramafic rocks by potassic and carbonate metasomatism (Halls and Zhao**
48 **1995). Fuchsite is absent in the samples we studied, and they are also carbonate-poor, and**
49 **therefore we prefer to use the terms talc-rich and/or listvenitic rocks instead of talc**
50 **carbonates or listvenites.**
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3 The contact of serpentinites with country rocks is highly sheared and deformed with the
4 development of mylonitized metasediments (Loizenbauer et al. 2001). Serpentinites
5 constitute up to 10% of the total area of suprastructure around the gneiss dome and form
6 elongated folded tabular bodies or sheets (Fig. 2). Ophiolitic serpentinites underlie and are
7 intercalated with the metavolcanosedimentary rocks. Andresen et al. (2009) reported a
8 crystallization age of 736.5 ± 1.2 Ma for zircons from an ophiolitic gabbro of the western part
9 of Meatiq (Fawakhir area), which we take to approximate the age of ophiolites in the study
10 area. A detailed age constraints on the evolution of the Neoproterozoic Meatiq dome is given
11 by (Andresen et al. 2009).

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14 A simple interpretation of the structural evolution of the Meatiq dome study area is that
15 ophiolitic suprastructure originally rested roughly horizontally above gneiss infrastructure.
16 Some of the shears in the suprastructure formed during ophiolite emplacement sometime
17 after 736 Ma. High heat flow accompanying syntectonic granitoid emplacement weakened
18 the infrastructure at roughly the same time as the Najd deformation occurred at ~ 600 Ma. The
19 heating combined with the Najd deformation caused upwelling of the infrastructure forming
20 the Meatiq Dome (Andresen et al. 2009). If this interpretation is approximately true, then
21 study areas A, B, and C are all components of the lower suprastructure, and carbonate
22 alteration occurred at one or more times during this sequence.
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27 [Insert figure 2 near here]
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30 The Fawakhir Au mine (Au = 1.5–29.7 g/ton; Hussein 1990) occurs in the western part of the
31 study area, associated with a 15m wide zone of graphitic schist in the contact of serpentinites,
32 metagabbros and granitoids. It is one of the several Au mines associated with CED ophiolitic
33 rocks that have been extensively worked since Pharaonic and Roman times (Harraz 2000).
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36 3. Sampling

37 In order to investigate the sources of fluids, the conditions of alteration, and chemical
38 changes during alteration of the ophiolitic fragments in the Meatiq area, a suite of 17 samples
39 of serpentinites, talc-rich rocks and associated quartz and carbonate veins were collected
40 (Table 1). The samples were collected from three main localities with the choice of sample
41 area controlled by the style of serpentinitization, the extent of carbonation, and crucially, the
42 availability of fresh, unweathered material. Talc-rich samples in particular are extremely
43 easily weathered, and therefore most of these samples were collected from quarries. Locality
44 A consists of large roadside exposures of serpentinite SW of the Meatiq dome (Fig. 2). The
45 serpentinite outcrops are brownish grey in colour and occur as massive units with brecciation
46 along faults with clearly developed macroscopic mesh texture and cross-cutting chrysotile
47 veinlets in hand specimen scale (Fig. 3a, b). These rocks represent the least altered
48 serpentinites in the study area. Locality B on the NE flank of the dome shows a complex
49 exposure of greenish grey serpentinite and shear zone hosted talc-rich rocks (Fig 3c, 4). The
50 serpentinites from locality B are softer and do not show the chrysotile veinlets seen at locality
51 A. These rocks are heavily sheared in places with yellowish-white to brownish-cream
52 coloured talc-rich rocks occurring in the shear zones (Fig. 3d, e). Veins, nodules and irregular
53 pockets of magnesite and dolomite with quartz occur within the talc-rich bodies (Fig. 3e).
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3 The quartz and carbonate veins generally have a NW-SE trend concordant with the foliation
4 in the talc-rich bodies (Fig. 3e, 4).

5 The talc-rich rocks are mostly very easily weathered and so despite the excellent exposures in
6 this area, sampling of these rocks was limited. Most samples from locality B were collected
7 from a talc quarry (Fig. 3f, 4). The contacts between the talc-rich rocks and the serpentinite
8 range from gradational to sharp (Fig. 3d), and some talc-rich bodies were observed to have a
9 sharp contact on one side and a gradational contact on the other (Fig. 4).
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13 [Insert figure 3 near here]
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16 Locality C, which occurs SE of locality B, shows a shear zone hosted exposure of talc-rich
17 rocks within serpentinites (Figs. 3g to 3j). The area exhibits the structural controls on
18 listvenitization with the majority of talc-rich bodies showing sharp fault contacts (Figs. 3g,
19 h). These rocks in shear zones experienced intensive mylonitization and are characterized by
20 talcose and schistose fabrics making talc-schists (Fig. 3i). At one area the talc-rich rocks have
21 been partially excavated allowing sampling of relatively fresh rocks. These bodies extend up
22 to several hundred meters in length and are up to 25 meters wide (Fig. 3j).
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26 [Insert figure 4 near here]
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28 4. Analytical methods

29 *SEM*

30 Polished thin-sections of selected serpentinites, talc-rich rocks and carbonate veins were
31 examined with a Philips XL30 FEG environmental scanning electron microscope (ESEM) at
32 Stockholm University, operating at 20kv and equipped with OXFORD energy dispersive
33 analytical X-ray spectrometer. The spectrometer detects elements with atomic number >4 (B
34 and heavier elements).
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37 *EPMA*

38 In-situ analyses of Cr-spinels of Lz-serpentinites were carried out using a field emission
39 electron probe microanalyser (FE-EPMA), JXA-8530F JEOL HYPERPROBE at Department
40 of Earth Sciences, Uppsala University, Sweden. Analytical conditions were 15kV
41 accelerating voltage, 10 nA probe current and 1 μm beam diameter, and the raw data were
42 corrected with an PAP routine. Counting time was 10s on peak and 5s on +/- background. $K\alpha$
43 spectral lines were measured. Analytical standards were natural and synthetic silicates and
44 oxides. The compositions of chromite are recalculated to cation proportions using the Fe^{3+}
45 calculation scheme of Droop (1987). The results are reported in Supplementary Table.
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51 *Gold analyses*

52 Gold analyses were carried out at Stockholm University, using Thermo Xseries 2 ICP-MS
53 following the ultra-low detection limit method described in Pitcairn et al. (2006, 2014). The
54 3σ method detection limit is 0.033 ppb Au. Analytical precision for Au analyses were
55 controlled through analyses of CANMET reference material TDB1 and USGS reference
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3 materials WMS-1, CH-4. In order to obtain representative Au concentrations in serpentinites,
4 nine hand specimen samples of serpentinite were divided into two groups, and were
5 powdered separately (total of 18 rock powder) using a hardened steel terna. The Au
6 concentrations of these two groups are reported separately in Table 1. Arsenic and antimony
7 analyses were also carried out at Stockholm University, by hydride generation-atomic
8 fluorescence spectrometry (HG-AFS) using a PSA 10.055 Millenium Excalibur instrument
9 following the methods described in Pitcairn et al. (2006, 2014). Analyses were carried out on
10 the same acid digests as those used for Au analyses. Method detection limits, using acid
11 digested blanks for 3σ , are 0.043 ppb As and 0.079 ppb Sb. Reference materials TDB-1,
12 WMS-1 and CH-4 were used to control analytical precision and accuracy. The results are
13 reported in Table 1.
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18 *Major and trace element analyses*

19 Samples were analyzed for major and trace elements by ICP-MS following lithium
20 metaborate/tetraborate fusion and total acid digestion at Activation Laboratories Ltd.
21 (Actlabs), Canada. Analyses were carried out on a Varian Vista 735 ICP-MS and calibration
22 was performed using 7 prepared USGS and CANMET reference materials. Loss on ignition
23 (LOI) is determined by weight difference after ignition at 1000°C. The detection limits are
24 between 0.001 to 0.01% for major elements, 0.005 to 5 ppm for trace elements and 0.5 ppm
25 to 0.001% for base metals. Selected trace elements were analyzed by ICP-MS at the National
26 Oceanography Centre Southampton (NOCS) following hydrofluoric, perchloric and nitric
27 acid digestion. Analyses were carried out on a ThermoFisher Scientific XSeries 2 ICP-MS
28 with accuracy and precision determined from repeated measurements of JA-2. The results are
29 reported in Table 1.
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35 *Sr isotope analysis*

36 Sr isotope analysis was carried out at the National Oceanography Centre Southampton
37 (NOCS). Sample preparation procedure is the same as explained for trace elements analyses
38 at NOCS. The mother solutions were subsampled to give approximately 1 µg Sr and the Sr
39 isolated using 50 µl Sr-Spec resin columns, the column blanks were <0.1 ng. The dried
40 samples were loaded onto a single Ta filament with a Ta activator solution. $^{87}\text{Sr}/^{86}\text{Sr}$ was
41 analyzed using multidynamic peak jumping routines on ThermoFisher Scientific Triton Plus
42 Thermal Ionisation Mass Spectrometers with a beam size of $^{88}\text{Sr} = 2\text{V}$ normalized to $^{86}\text{Sr}/^{88}\text{Sr}$
43 = 0.1194. The long term average for NBS987 on the instrument is 0.710244 ± 0.000019 (2sd)
44 on 138 analyses. The results are reported in Table 2.
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49 *Stable isotope (C and O)*

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52 Carbon and oxygen isotope analyses were carried out on the carbonate fraction of serpentinite
53 whole rock samples and 3 carbonate veins. The serpentinite samples were prepared by micro-
54 milling of carbonate bearing areas within the serpentines to produce a carbonate rich powder.
55 Isotope analyses were undertaken using an automated triple-collector gas source mass
56 spectrometer (Analytical Precision AP2003) linked to an automated gas preparation device at
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3 the Scottish Universities Environmental Research Centre (SUERC), East Kilbride. For the C
4 isotope analyses, 2mg of the powdered whole-rocks and carbonate veins samples were
5 reacted with 103% phosphoric acid to produce carbon dioxide, which was then purified
6 before analysis. Samples were reacted for 96 hours at a constant temperature of 70°C. Gas
7 mineral fractionation for dolomite was calculated following Rosenbaum and Shepherd
8 (1986). Magnesite and mixtures of magnesite and dolomite were treated as
9 dolomite. Precision and accuracy were monitored by reference to long-term analysis of
10 laboratory and international standards. Precision is better than 0.2‰ at 1σ for carbon and
11 oxygen. Results are reported as δ‰ values relative to the V-PDB and V-SMOW scales for
12 C- and O-isotopes, respectively (Table 2).
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16 17 *Microthermometry*

18 Three dolomite samples (Me2, Me12, Me28) and magnesite veinlets in serpentinite sample
19 Me9, and one quartz sample (Me3) were selected for detailed fluid inclusion studies. The
20 samples were prepared as 150 μm thick doubly polished sections. After careful
21 documentation and selection of fluid inclusions, microthermometric analyses were carried out
22 at Stockholm University using a Linkam THM 600 heating and cooling stage mounted on a
23 Nikon microscope utilising a 40x long working-distance objective. The reproducibility was
24 ±0.1°C for temperatures below +40°C and ±0.5°C for temperatures above 40°C. The stage
25 was calibrated with synthetic fluid inclusion standards (SynFlinco®) and well-defined natural
26 inclusions in Alpine quartz. The results are listed in Table 3.
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31 *Raman spectrometry*

32 Raman spectrometry analyses were performed using a laser Raman confocal spectrometer
33 (Horiba instrument LabRAM HR 800) equipped with a multichannel air cooled CCD detector
34 at Stockholm University. The same set of carbonate veins prepared for microthermometry
35 study was selected for determination of fluid composition trapped in inclusions. In addition,
36 serpentinite samples from localities A and B were analyzed to identify serpentine minerals.
37 An Ar-ion laser (λ = 514 nm) was used as the excitation source with an output power at the
38 sample of 8 mW. The instrument was integrated with an Olympus microscope and the laser
39 beam was focused to a spot of 1 μm with a 100x objective. The spectral resolution is about
40 0.3 cm⁻¹. The instrument was calibrated using a neon lamp and the Raman line of a silicon
41 wafer (520.7 cm⁻¹). Instrument control and data acquisition was made with LabSpec 5
42 software.
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47 [Insert Table 1 near here]
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50 **5. Results**

51 **5.1. Petrography and mineralogy**

52 Serpentinites from locality A (referred to from this point onwards as *Lz-serpentinite*) are
53 composed essentially of lizardite and late chrysotile veinlets with minor carbonate, opaque
54 minerals and traces of chlorite, bastite pseudomorphs, and talc. In places, **serpentine** minerals
55 retain the crystal habit of the original mafic minerals. Chrysotile occurs as long fibrous
56 veinlets cross-cutting the lizardite matrix, indicating protracted serpentinization (Fig. 5a, b).
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3 Carbonates are stained with iron-oxide and occur as scattered blocky aggregates or filling the
4 mesh texture of original mafic mineral or secondary chrysotile veinlets (Fig. 5b). The opaque
5 minerals in the serpentinites are Cr-spinel, pentlandite and magnetite. The Cr-spinels (1-3%
6 by volume) are slightly altered and occur as deep-red grains with sub-rounded outlines.
7 Locally the Cr-spinels show slight alteration at grain boundaries forming narrow rim of Cr-
8 magnetite with large unaltered core (Fig. 5c). Pentlandite (1-2% by volume) is rounded and
9 altered to garnierite and magnetite along margins. Magnetite occurs as primary scattered
10 anhedral grains or as secondary narrow rims around Cr-spinels or defines olivine and
11 orthopyroxene relicts.
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17 Serpentinites from locality B (referred to from this point onwards as *Atg-serpentinite*) are
18 composed of antigorite and carbonate with minor talc, Cr-spinel and clinochlore. Antigorite
19 is the only serpentine mineral and occurs as flakes and shreds with no evidence of late
20 chrysotile. In places, antigorite grains are intergrown with abundant magnesite and minor talc
21 (Fig. 5d). The carbonates, composed dominantly as magnesite with minor dolomite occur as
22 disseminated, cryptocrystalline veinlets, crystal clusters and infilling cracks in cases together
23 with talc (Fig. 5d, e). Similar magnesite veins have been reported in CED serpentinites by
24 previous workers (Ghoneim et al. 1999; Azer and Stern 2007; Ali-Bik et al. 2012). Based on
25 SEM analysis, clinochlore is Cr-rich type (Kammererite) and occurs as aureoles around
26 altered Cr-spinels. The opaque minerals are Cr-spinel and magnetite with no evidence of
27 sulphide minerals such as pentlandite. The Cr-spinels (1-3% by volume) in this group of
28 serpentinites occur as disseminated subhedral zoned grains, with small unaltered chromite
29 core, altered ferritchromite zone and Cr-magnetite outer rim (Fig. 5f). The relict mesh texture
30 in the *Atg-serpentinites* (Fig. 3b and d) suggests that these rocks originated as Lz-
31 serpentinites.
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37 Talc-rich rocks which in parts bear carbonate and quartz veins occur at both localities B and
38 C comprising talc, amphibole (tremolite, anthophyllite), quartz, clinochlore, altered Cr-spinel,
39 magnetite and traces of rutile. Talc mostly occurs as fine shreds, microcrystalline fibres and
40 rarely as coarse- to medium-grained flakes with tremolite/anthophyllite and clinochlore (Fig.
41 5g, h). In the study area, highly foliated talc-schists (e.g. Me15 and Me17) are distinctive
42 within the fault and shear zones. Syn-kinematic characteristics of these rocks include
43 asymmetric foliation and formation of tremolite neoblast at high angle to the mylonitic
44 textures (Fig. 5g). Chlorite occurs either as intergrowth aggregates with talc and amphibole or
45 aureoles that enclose Cr-spinels; in the latter case, composition approaches Cr-rich
46 clinochlore (Fig. 5h). The opaque minerals in talc-rich rocks are mostly relicts of primary Cr-
47 spinels which have experienced brittle brecciation and pervasive alteration to ferritchromite
48 (Fig. 5i).
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54 [Insert figure 5 near here]

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56 Carbonate veins in the study area vary from micro-scale networks infiltrating the matrix of
57 the serpentinites to metre-wide veins together with quartz in talc-rich rocks (Fig. 3e).
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Carbonate minerals are dominantly magnesite and dolomite with minor calcite. Both carbonates and quartz veins display evidence of shearing manifested by stretched grain ribbons, recrystallized bulges along crystal boundaries, and S-fabrics especially in fault and shear zones. Some of the carbonate (e.g. Me12) veins host minor magnetite which occurs as disseminated aggregates of fine to medium grain size and traces of accessory Cr-spinel, zircon, monazite, rutile, and apatite.

5.2. Geochemistry

Major elements variations

Lz-serpentinites contain 44.8 to 47.2 wt% SiO₂, 0.56 to 0.65 wt% Al₂O₃, 44.3 to 44.7 wt% MgO, 7.1 to 8.3 wt% total Fe as FeO and 0.8 to 1.3 wt% CaO (Table 1). The Atg-serpentinites show similar range of major elements and contain 42.2 to 46.7 wt% SiO₂, 0.15 to 0.62 wt% Al₂O₃, 44.6 to 50.5 wt% MgO, and 6.9 to 9.7 wt% total Fe as FeO (Table 1). Major element oxides in the talc-rich rocks range from 57.2 to 62.7 wt% SiO₂, 0.4 to 1.7 wt% Al₂O₃, 20.7 to 28.6 wt% MgO, 0.03 to 0.1 wt% Na₂O, 4.5 to 7.4 wt% total Fe as FeO and 2.8 to 12.9 wt% CaO, with other major oxides in minor amounts (Table 1). The protolith of the talc-rich rocks cannot be inferred from this composition. Both varieties of serpentinites contain abundant volatiles (H₂O and CO₂) with LOI values ranging from 13.2 to 14.2 wt% for Lz-serpentinites and 12.5 to 20.8 wt% for Atg-serpentinites (Table 1; Fig. 6a). Talc-rich rocks contain lower LOI values ranging from 2.1 to 4.1 wt% along with lower MgO and higher SiO₂, Na₂O and CaO than the serpentinites (Table 1; Fig. 6a, b, c).

[Insert figure 6 near here]

Compositional similarities between Lz-serpentinites and Atg-serpentinites along with relict olivines (in Lz-serpentinites) and Cr spinel indicate these formed from broadly similar harzburgite protoliths. In contrast, the major element compositions of the talc-rich rocks are unlike any mantle peridotite and only elevated MgO contents and Cr-spinel relicts suggest that this lithology may also have a harzburgite protolith.

Trace elements variations

The serpentinites analyzed in this study show relatively high concentrations of Cr, Ni and Co with mean values of 1735 ppm, 2217 ppm and 99 ppm, respectively in the Lz-serpentinites and 2293 ppm, 2308 ppm, and 103 ppm, respectively in the Atg-serpentinites (Table 1; Fig. 7). The talc-rich rocks show somewhat lower concentrations of these elements with average concentrations of 1430 ppm for Cr, 1152 ppm for Ni and 64 ppm for Co (Fig. 7).

The Lz-serpentinites show higher concentrations of incompatible trace elements Li, Rb, Ba, U, Pb, Sr) in comparison to Atg-serpentinites (Fig. 8a; Table 1). The Lz-serpentinites also

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3 show higher S, As and Au concentrations compared with the Atg-serpentinites and talc-rich
4 rocks (Fig. 7).
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8 [Insert figure 7 near here]
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10 Chondrite-normalized rare earth element (REE) patterns of different serpentinites, talc-rich
11 rocks and carbonate veins, using the normalizing values of McDonough and Sun (1995), are
12 presented in figure 8b, to d. The Lz-serpentinites contain higher Σ REE concentrations
13 (~2.46 ppm) than Atg-serpentinites (~0.08 ppm). The carbonate veins show the highest Σ
14 REE concentrations (~63 ppm), and the talc-rich rocks show intermediate concentrations (~
15 1.1 ppm) between the two varieties of serpentinites (Table 1; Fig. 8b-d). The Lz-serpentinites
16 are characterized by moderate enrichment of LREE/HREE ($La_N/Yb_N \sim 11$) and Atg-
17 serpentinites show slight depletion in MREE with a strong positive Eu anomaly. Sample
18 Me24 is distinct from the other Lz-serpentinites considering lower REE enrichment and
19 MREE depletion similar to Atg-serpentinites. The carbonate veins generally show flat REE
20 patterns with slight LREE enrichment ($La_N/Yb_N = 11.3$) in sample Me2. The talc-rich rocks
21 REE patterns are approximately flat with slight enrichment in LREE ($La_N/Yb_N = 3.55$ in
22 average) and a slight positive Eu anomalies (Fig. 8c). A comparison between average
23 compositions of different rock types and veins is show in figure 8d.
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36 ***5.3. Isotopic signature of Meatiq carbonate altered rocks and veins***

37 Results for carbon and oxygen isotope analysis for carbonate fraction of nine serpentinites
38 samples, Rb and Sr concentrations and $^{87}Sr/^{86}Sr$ ratios for nine whole-rock serpentinite
39 samples, and three carbonate veins are listed in Table 2 and plotted in figure (9a-c). $^{87}Sr/^{86}Sr$
40 ratios have been corrected for 600 Ma of radiogenic growth, which represents the best
41 estimate for the timing of carbonation occurred in the CED (Stern and Gwinn 1990).
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43

44 [Insert Table 2 near here]
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46 The initial $^{87}Sr/^{86}Sr$ ratios vary from 0.7028 to 0.7078 (Table 2). The low $^{87}Rb/^{86}Sr$ in all
47 samples indicates that uncertainties in the age correction are unlikely to introduce significant
48 errors, but low Sr contents of especially Atg-serpentinites makes these especially vulnerable
49 to minor alteration by groundwater or biological activity. Initial $^{87}Sr/^{86}Sr$ for samples
50 containing <10 ppm Sr should therefore be regarded with suspicion. Three of the four
51 samples that contain <10 ppm Sr show the highest initial $^{87}Sr/^{86}Sr$ (0.7067 – 0.7078); these
52 contrast with much lower initial $^{87}Sr/^{86}Sr$ for three Lz-serpentinites and carbonate veins
53 (0.7028 – 0.7046) and we interpret these high concentrations as the product of secondary
54 processes. The $\delta^{18}O$ values range from +6.4‰ to +15.1‰ while $\delta^{13}C$ values vary between -
55 8.1‰ to -4.1‰ (Table 2). The carbonate veins show tightly clustering isotopic values with
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3 $\delta^{13}\text{C}$, $\delta^{18}\text{O}$ and initial $^{87}\text{Sr}/^{86}\text{Sr}$ values ranging from -8.1‰ to -6.8‰, +6.4‰ to +10.5‰, and
4 0.7028 to 0.7034 respectively (Table 2). They contain low concentrations of Rb from 0.001
5 to 0.04 ppm (mean = 0.015 ppm) while their Sr contents are high and variable, from 503 to
6 2365 ppm (mean = 1334 ppm). The $^{87}\text{Rb}/^{86}\text{Sr}$ ratios are extremely low, invariably less than
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16 The carbonate fraction of serpentinites have higher C and O isotopic values compared to
17 carbonate veins. The serpentinites also show higher $^{87}\text{Sr}/^{86}\text{Sr}$ ratios compared to the carbonate
18 veins (Table 2; Fig. 9a-c). The carbonate fraction of Lz-serpentinites (Me21 to Me24) range
19 in $\delta^{18}\text{O}$ from +12.3‰ to +15.1‰ with uniform $\delta^{13}\text{C}$ values range between -4.2‰ to -4.1‰.
20 Their whole rock $^{87}\text{Sr}/^{86}\text{Sr}$ ratios span a modest range from 0.70421 to 0.70622. They contain
21 low concentrations of Rb from 0.033 to 0.12 ppm (mean = 0.88 ppm) and Sr concentrations
22 ranging from 5.8 to 70 ppm (mean = 37.6 ppm) with Rb/Sr ratios mostly less than 0.05. The
23 carbonate fraction of Atg-serpentinites (Me5 to Me9) show $\delta^{18}\text{O}$ values that range from
24 +10.3‰ to +11.5‰ and $\delta^{13}\text{C}$ values ranging from -5.3‰ to -5.9‰ (Table 2). Their whole
25 rock $^{87}\text{Sr}/^{86}\text{Sr}$ ratios vary from 0.70396 to 0.70585 (Table 2). The Atg-serpentinites contain
26 lower concentrations of Rb (mean = 0.023 ppm) and Sr (mean = 0.36 ppm) in comparison to
27 Lz-serpentinites. Their Rb concentrations range from 0.009 to 0.031 ppm and their Sr
28 concentrations vary from 0.23 to 0.45 ppm with Rb/Sr ratios vary from 0.22 to 0.35 (Table
29 2).
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35 **5.4. Microthermometry**

36 Five samples of carbonate and quartz veins were collected for fluid inclusion
37 microthermometry. Sample Me2, Me12 and Me28 are dolomite veins, sample Me9 is from a
38 magnesite vein in Atg-serpentinite sample and Me3 is from a quartz vein. Microthermometric
39 data are given in Table 3.
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41

42 *Fluid inclusion types*

43 Three fluid inclusion types (I to III) have been identified (Table 3; Fig. 10a-c). Some fluid
44 inclusions occur along secondary trails (Fig. 10d) and necking down is occasionally seen
45 (Fig. 10e). In cases, inclusions show shear sense due to local tectonic activity (Fig. 10f). All
46 analyses were made on fluid inclusions without clear signs of post-entrapment modifications.
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50 Type-I inclusions are aqueous two-phase inclusions ($\text{LH}_2\text{O} + \text{VH}_2\text{O}$) occurring as primary and
51 secondary inclusions with variable degrees of fill (~5-60 vol% gas). They are elongated with
52 rectangular to negative crystal shapes with a size from 4 to 40 μm (Fig. 9a). Some inclusions
53 are single phase (VH_2O). Type-II inclusions are primary three-phase ($\text{LH}_2\text{O} + \text{LCO}_2 + \text{VCO}_2$)
54 aqueous-carbonic with low degree of fill (~30 to 70 vol% gas). They are negative crystal or
55 irregular shaped and range in size from 5 to 15 μm (Fig. 10b). Type-III inclusions are
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3 carbonic ($\text{LCO}_2 + \text{VCO}_2$), less than 10 μm in size, rounded to negative crystal in shape and
4 occur in cluster or along trails (Fig. 10c). Type-III inclusions are only present in a quartz vein
5 (Me3).
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8 *Fluid inclusions hosted by carbonate veins*

9 Type I inclusions: In Me2 final ice-melting temperatures (T_{mice}) range from -2.6°C to 0.0°C
10 which gives a salinity of 0.0 to 4.3 mass% NaCl eq. (low salinity inclusions) and from $-$
11 16.4°C to -4.0°C which corresponds to 6.4 to 16.9 mass% NaCl eq. (moderate salinity
12 inclusions). Salinities are calculated using data in Bodnar (2003). The temperatures of first
13 ice-melting (T_{fmice}), around -35°C for the low salinity assemblage, and about -50°C for the
14 moderate salinity inclusions are similar to eutectic temperatures of H_2O -salt systems
15 dominated by NaCl-MgCl₂ and CaCl₂-NaCl respectively (Shepherd et al. 1985).
16 Homogenization temperature of all inclusions to liquid occurred between 161°C and 383°C .
17 In Me9, first melting of the type-I inclusions was observed around -35°C which suggests a
18 salt composition dominated by NaCl and MgCl₂ (Shepherd et al. 1985). Final ice-melting
19 temperatures range from -2.4°C to -0.5°C corresponding to a salinity of 0.9 to 5.5 mass%
20 NaCl eq. Total homogenization to liquid occurred at temperature between 210°C and 335°C .
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28 Type II inclusions: In Me12, there was no separate CO_2 phase, but CO_2 clathrate melting
29 (T_{mclath}) occurred in the range from 7.5°C to 9.8°C which indicates a salinity of 0.2 to 4.9
30 mass% NaCl (Fall et al. 2011). Total homogenization to liquid occurred in the temperature
31 range 252°C to 276°C . In Me28, melting of the frozen CO_2 (T_{mCO_2}) occurred between $-$
32 57.3°C and -56.6°C and homogenization of the CO_2 -phase to liquid took place between
33 $+26.8^\circ\text{C}$ and $+29.2^\circ\text{C}$. CO_2 clathrate melting was observed at temperatures between $+6.6^\circ\text{C}$
34 and $+10.0^\circ\text{C}$ which indicates a salinity of 0.0 to 6.8 mass% NaCl. Total homogenization to
35 liquid phase occurred between 242°C to 331°C .
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44 *Fluid inclusions hosted by quartz veins*

45 Type I inclusions: In Me3, first observed melting around -35°C suggests a predominant
46 NaCl-MgCl₂ composition. Final ice-melting from -3.5°C to -0.1°C corresponds to a salinity
47 of 0.2 to 5.7 mass% NaCl eq. Total homogenization occurred at temperatures ranging from
48 225°C to 266°C (to liquid) except for the secondary inclusions which homogenized between
49 177°C to 234°C (to liquid).
50

51 Type-II and III inclusions yield melting temperatures for CO_2 from -57.7°C to -56.6°C and
52 homogenization of CO_2 to liquid, between $+26.8^\circ\text{C}$ to $+29.5^\circ\text{C}$. The temperatures indicate
53 the presence CH_4 and/or N_2 in addition to CO_2 . Based on the data in van den Kerkhof and
54 Thiéry (2001), the corresponding molar volumes of the CO_2 inclusions range from about 65
55 to $90 \text{ cm}^3/\text{mole}$. Melting of CO_2 clathrate occurred at temperatures between $+7.7^\circ\text{C}$ to $+9.5^\circ\text{C}$
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3 which indicates a salinity of 1.0 to 4.7 mass% NaCl eq. Total homogenization to liquid phase
4 occurred between 305°C to 365°C.
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7 **6. Discussion**

8 In the following, we explore the significance of our results in relation to: 1) provenance of the
9 ophiolites; 2) condition of serpentization; 3) talc-rich rocks formation; 4) the composition
10 of the carbonating fluid and the conditions of alteration; 5) the sources of the CO₂-rich
11 carbonating fluid; 6) element mobility during serpentization and listvenitization; and 7) the
12 evolution of serpentization and carbonation and the implications for formation of Au
13 deposits.
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16 **6.1. Provenance of the ophiolites**

17 Ophiolites and ophiolitic rocks are remarkably abundant in the Arabian-Nubian Shield (Fig.
18 1). Three main tectonic settings have been proposed for formation of the ophiolites in the
19 ANS; (i) mid-ocean ridges (Zimmer et al. 1995; Khalil and Azer 2007); (ii) back-arc basins
20 (Bakor et al. 1976; Kröner 1985; Pallister et al. 1988; El-Sayed et al. 1999; Farahat et al.
21 2004; El Gaby 2005; El Bahariya 2008; Abd El-Rahman et al. 2009; Ali et al. 2009), and (iii)
22 forearcs (Stern et al. 2004; Azer and Stern 2007; Hamdy et al. 2013; Ahmed 2013; Khedr and
23 Arai 2013; Azer 2014; Khalil et al. 2014; Abu-Alam and Hamdy 2014). Serpentinized
24 peridotites investigated in this study are depleted in Al₂O₃ and CaO and plot within the field
25 of metamorphic peridotites associated with ophiolitic suits (Fig. 11a; Coleman 1977), similar
26 to harzburgites from modern intra-oceanic forearcs (Fig. 11b).
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37 In completely serpentized peridotites without relicts of primary silicate minerals, a good
38 proxy to determine original tectonic setting is the composition of Cr-spinel cores (Irvine
39 1965; Arai 1992; Dick and Bullen 1984). Cr-spinels resist alteration and are widely used to
40 constrain tectonic settings where ophiolites formed. According to Dick and Bullen (1984),
41 peridotites that contain Cr-spinel with Cr# >0.6 are commonly restricted to subduction-
42 related rocks. The Cr# of the Cr-spinel cores in the investigated Lz-serpentinites is mostly
43 >0.6 and lie in the compositional space for Cr-spinels of forearc peridotite (Supplementary
44 Table; Fig. 11.c). This is similar to recent studies on Cr-spinels from the Fawakhir
45 serpentinites in the vicinity of study area (~7 km north of study area A) and other
46 serpentinites in the ED that are considered to have formed in the forearc of a supra-
47 subduction zone (SSZ) setting (Azer and Stern 2007; Hamdy et al. 2013; Azer 2014; Abu-
48 Alam and Hamdy 2014).
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53 **6.2. Conditions of serpentization**

54 Lizardite, antigorite and talc have been widely reported in ED serpentinite bodies (Basta and
55 Kader 1969; Akaad and Noweir 1972; Abu-Alam and Hamdy 2014), but the relationships
56 between these minerals and their significance in terms of the conditions of alteration remain
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3 unclear. As noted in the results section, serpentinites investigated in this study can be
4 classified into 2 distinct groups; Lz-serpentinites, and Atg-serpentinites. In Lz-serpentinites,
5 lizardite and chrysotile assemblages become major rock forming minerals with minor
6 carbonate and relatively fresh Cr-spinel. Lizardite and chrysotile form at lower temperature
7 alteration conditions than antigorite (Mellini et al. 1987; Scambelluri et al. 1995; O'Hanley
8 1996; Trommsdorff et al. 1998; Andréani et al. 2007; Evans 2004, 2010). The recent study of
9 Schwartz et al. (2013) shows that at below 300°C, lizardite and locally chrysotile are the
10 dominant serpentine species commonly forming a mesh texture, and that between 320°C and
11 390°C, lizardite is progressively replaced by antigorite which is the sole stable serpentine
12 mineral at temperatures >390°C. The extent of spinel alteration suggests similar alteration
13 conditions, with thin outer magnetite rims in sharp contact with Cr-spinel core suggesting
14 sub- to lower-greenschist facies condition (Farahat 2008). This alteration zoning of Cr-spinel
15 together with abundant lizardite and chrysotile assemblages forming a mesh texture and lack
16 of higher temperature serpentine phase (i.e. antigorite) indicates sub-greenschist facies
17 conditions and temperatures <300°C (Abzalov 1998; Barnes 2000; Farahat 2008).
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24 In the Atg-serpentinites, antigorite and carbonate assemblages become major rock forming
25 minerals with minor talc, altered Cr-spinel, and clinocllore. The antigorite is the sole
26 serpentine mineral in these rocks and the lack of lizardite and chrysotile indicates the
27 transition of these phases to antigorite. Cr-spinels are concentrically zoned with small,
28 unaltered cores, well-developed transitional zone of ferritchromit and thick Cr-rich magnetite
29 rims (Fig. 5f). The occurrence of antigorite as the sole serpentine phase together with
30 alteration textures of Cr-spinel indicate upper greenschist conditions and temperatures
31 >450°C (Abzalov 1998; Barnes 2000).
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35 A number of different mechanisms have been suggested for the transition of lizardite to
36 antigorite. In a water-saturated open system, the transition of lizardite to antigorite can be
37 written as (following Vils et al. 2011):
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42 (Coleman 1971; Evans 2004) or alternatively in closed system with no inlet fluid, the
43 transition of lizardite to antigorite can be written as:
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49 The extensive carbonate alteration and intimate intergrowth between antigorite and magnesite
50 indicate that the generation of antigorite was most likely related to precipitation of carbonate
51 from a CO₂-bearing fluid. **Therefore, the lizardite to antigorite transition must have occurred**
52 **in the presence of CO₂. Reaction 2 would facilitate generation of magnesite due to liberation**
53 **of aqueous MgO.** The Atg-serpentinites contain higher MgO and LOI and slightly lower SiO₂
54 compared to Lz-serpentinites (Fig. 6), which is consistent with the addition of MgO and CO₂
55 to these rocks. We suggest that the lizardite to antigorite transition was contemporaneous
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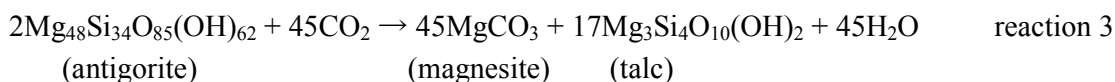
with magnesite crystallization due to influx of CO₂-rich fluid at temperatures above the stability field of lizardite (i.e. >300°C; Evans 2004; Schwartz et al. 2013).

6.3. Conditions of Listvenitization

Listvenite is formed by low to intermediate temperature hydrothermal/metasomatic alteration of mafic-ultramafic rocks (especially serpentinites), commonly associated with major fault and shear zones (Uçurum 2000; Akbulut et al. 2006). The listvenitic rocks investigated in this study occur in localities B and C and are spatially associated with shear zones. These rocks show gradational to sharp contacts with the surrounding serpentinite (Figs. 3d, 4) indicating that the listvenites formed after serpentinitization. This relationship is commonly described elsewhere (Böhlke 1989; Griffis 1972; Uçurum 2000; Hansen et al. 2005, Boschi et al. 2006) and in the ED (Ghoneim et al. 2003; Azer 2013; Zoheir and Lehmann 2011; Emam and Zoheir 2013) and is supported by several experimental studies on a MgO-SiO₂-H₂O-CO₂ system (Johannes 1969; Klein and Garrido 2011; Klein and McCollom 2013; Bjerga et al. 2015). In the study area, listvenitic rocks show variable mineral assemblages including talc-tremolite assemblages with minor ferritchromite and clinochlore in the highly foliated talc-rich rocks (Fig. 3e, 5g), and talc-tremolite-anthophyllite assemblages together with ferritchromite and clinochlore where foliation is less well-developed (Fig. 5h). The listvenitic zones also contain quartz and carbonate (with minor talc) veins (Fig. 3e).

In ultramafic rocks, the talc-tremolite (± magnesite) assemblages are indicative of upper-greenschist to low-amphibolite facies (Johannes 1969; Barns 2000; Will et al. 1990). But at mid-amphibolite facies and relatively low pressure (<2kbar), the stability limit of former assemblages is exceeded and stable assemblages are talc-tremolite-anthophyllite (Will et al. 1990; Evans and Guggenheim 1988). Anthophyllite is a common phase in the middle and upper amphibolite facies of ultramafic rocks (Will et al. 1990; Barns 2000). However, the occurrence of stable talc-tremolite-anthophyllite assemblages could also be related to higher SiO₂-H₂O activity in the fluid (Evans and Guggenheim 1988). So, the carbonate-poor nature of the talc-tremolite-anthophyllite assemblages in our shear zone samples may presumably be related to either higher temperature or higher activity of SiO₂ and H₂O in the altering fluid, or a combination of both phenomena in some parts of the shear zones.

The talc-rich rocks most likely form from breakdown of Atg-serpentinites that host these rocks. This is supported by mineralogy and mineral textures in Atg-serpentinites where cross-cutting micro-veins of magnesite and talc occur in some samples (Fig. 5e). Such a conversion has been shown to occur in the presence of CO₂-rich fluid (more than 2-6 mol% CO₂ and under fluid pressure of 2 kbar, Johannes 1969) at temperatures above 300°C (e.g. Schandl and Naldrett 1992; Hansen et al. 2005; Robinson et al. 2005):



Much of the magnesite formed by this reaction appears to occur in discrete veins and pods within the listvenitic rocks or as veinlets crosscutting Atg-serpentinites (Fig. 3e, 5e). This

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3 reaction is supported by the chemical compositions of the rocks as the talc-rich rocks are
4 strongly depleted in volatiles and MgO relative to the serpentinites (Fig. 6). Talc-rich rocks
5 contain an average of ~2.9 wt% LOI compared to average of ~15.2 wt% LOI in the
6 serpentinites indicating that, as is suggested by reaction 3, the listvenitization process
7 involves dehydration. Reaction 3 shows that MgO is balanced during the listvenitization
8 process but that much of the MgO occurs in the discrete veins of magnesite with the
9 surrounding talc-rich rocks being relatively depleted in MgO. Enrichments in SiO₂ and CaO
10 in the talc-rich rocks are in the proportions expected through residual concentration due to the
11 loss of volatiles and MgO (Fig. 6).
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15 **6.4. The composition of the carbonating fluid and the conditions of alteration**

16 The conditions at which the carbonate alteration occurs can be estimated through fluid
17 inclusion microthermometry vein carbonate. The carbonate veins were formed from aqueous
18 and carbonic fluids with a large range in relative proportions from pure aqueous to pure
19 carbonic end-members indicating coexisting, immiscible fluid phases at the time of carbonate
20 deposition. The carbonic phase has a CO₂-rich composition with small amounts (0 to 5
21 mole%) of dissolved CH₄, probably present as a result of serpentine mineral alteration
22 reactions, and N₂ of unknown origin. The aqueous phase in both carbonate and quartz veins
23 has low salinity (<7 eq. mass% Mg-Na-Cl; Table 3). Some aqueous inclusions in sample Me2
24 have higher salinities (up to 17 eq. mass% NaCl) and a more Ca-Na-Cl dominated
25 composition. These may represent local input of a more saline fluid from external source.
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36 Total homogenization temperatures of the samples range from 161°C to 383°C (Table 3).
37 Primary inclusions have a more limited range of total homogenization temperatures from
38 225°C to 383°C (Table 3), with secondary aqueous fluid inclusions in healed microfractures
39 showing lower homogenization temperatures (161°C to 255°C; Table 3). Total
40 homogenization temperature for primary fluid inclusions trapped under conditions of
41 immiscibility is interpreted to represent the original formation temperature (Diamond 2003;
42 Bodnar 2003). The trapping temperature and pressure conditions can be estimated from
43 compositions of the two end-members (Diamond 2003). Trapping conditions for immiscible
44 fluids can be determined by using the technique of intersecting isochores (Roedder 1984) and
45 at the intersecting point in the PT-diagram which two end-members have temperature and
46 pressure that are identical. The isochore to be used for the aqueous end-members (inclusions
47 with the lowest homogenization temperature) in the samples is based on the data in Bodnar
48 (2003) for aqueous fluid inclusions with homogenization temperature of about 225°C and a
49 salinity of 0-7 eq. mass% NaCl. The isochores for the carbonic end-member (after data in
50 Van den Kerkhof and Thiéry 2001) are for molar volumes of 65 and 90 cm³/mole (Fig. 12).
51 The intersection of the aqueous and carbonic isochores in the PT diagram gives trapping
52 conditions for carbonate veins of the study area within a temperature interval from 270°C to
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300°C and at a pressure of 0.7 to 1.1 kbar. The pressure corresponds to lithostatic depths of 2.6 to 4.2 km (Fig. 12).

6.5. The sources of the CO₂-rich carbonating fluid

The carbonate vein samples have $\delta^{13}\text{C}$ (mean = -7.2‰) and $^{87}\text{Sr}/^{86}\text{Sr}$ (mean = 0.7030) compositions that fall within the range of mantle-derived carbon (Taylor et al. 1967; Deines 1989; Demény et al. 1998; Fig. 9a). The $\delta^{18}\text{O}$ values of these veins (ranging from +6.4 to 10.5‰) are slightly heavier than depleted-mantle values (i.e. +5.5±0.2‰; Eiler 2001) indicating partial re-equilibration with crustal brines, groundwater or seawater (Fig. 9b). The robust $^{87}\text{Sr}/^{86}\text{Sr}$ compositions in the carbonate veins are indistinguishable from initial ratios for Neoproterozoic mantle-derived igneous rocks of the Arabian-Nubian Shield (Engel et al. 1980; Fitches et al. 1983; Stern and Hedge 1985; Stern and Gwinn 1990; Fig. 13.) Based on $\delta^{13}\text{C}$ and $^{87}\text{Sr}/^{86}\text{Sr}$ values, the carbonate veins most likely formed from mantle-derived CO₂-bearing fluids. An alternative explanation is that the veins formed from surface-derived fluids that fully equilibrated with the host rock. However in that case, one would expect full equilibration of O and perhaps Sr isotopes (e.g. Falk and Kelemen 2015), but that C isotopes would preserve the fluid signature. The data show the opposite with more crustal signatures shown in the O isotopes than in C isotopic compositions. Furthermore the fluid inclusion data shows that the carbonation was produced by a mixture of H₂O and CO₂ fluids with the primary inclusions including pure CO₂ end member fluids at around 270°C to 300°C. It is difficult to conceive how such a fluid could originate as meteoric water or seawater.

[Insert figure 13 near here]

The carbonate fraction in the Lz- and Atg-serpentinities have $\delta^{13}\text{C}$ compositions (-5.9 to -4.1‰) that also falls within the range of depleted mantle carbonate (i.e. Taylor et al. 1967; Deines 1989; Fig. 9a), and are close to the isotope ratios of mantle C ($\delta^{13}\text{C}$ ~-5‰) (Deines 2002). The $\delta^{18}\text{O}$ compositions of carbonate fraction in Lz- and Atg-serpentinities and $^{87}\text{Sr}/^{86}\text{Sr}$ ratios in whole rock Lz- and Atg-serpentinities plot outside depleted mantle values. These sample values plot on mixing lines between mantle fluid and sedimentary carbonate indicating incorporation of surface-derived O and Sr (Fig. 9c). The carbonate fraction in Lz-serpentinities shows consistently higher $\delta^{18}\text{O}$ (the highest $\delta^{18}\text{O}$ values in this study, ranging from 12.1 to 15.1‰) and more radiogenic Sr for whole rock values compared to carbonate veins, suggesting interaction with a greater proportion of surface derived fluid (Fig. 9b). A possible interpretation is that the heavier $\delta^{13}\text{C}$ values (-4.2‰) of carbonate fraction in Lz-serpentinities compared to Atg-serpentinities ($\delta^{13}\text{C}$ = -5.5‰) might reflect interactions with fluid having low C/O ratio causing a shift in $\delta^{18}\text{O}$ with little change in $\delta^{13}\text{C}$ (-1.3‰ higher). Both varieties of serpentinite contain carbonate fractions that most likely represent mixtures of mantle derived, and surface derived CO₂.

6.6. Element mobility during serpentinization and listvenitization

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3 As discussed above, the major element changes during the different styles of alteration in
4 these rocks are dominated by addition and removal of CO₂ and H₂O, and in the
5 remobilization of SiO₂ and MgO. Variations in the other major elements are most easily
6 explained by residual enrichments caused by addition of carbonate to the rock. The extent of
7 recrystallization during the alteration of serpentinite and talc-rich rock also redistributed trace
8 elements with some being locally remobilized in the rock and other elements flushed out the
9 rock. The Cr-spinels for example are recrystallized during alteration, but Cr was only locally
10 redistributed into nearby neoblasts of Cr-bearing magnetite and clinchlore (Fig. 5c, f, i).
11 Primary Ni-sulphides such as pentlandite found in the Lz-serpentinites are absent from the
12 Atg-serpentinites and the talc-rich rocks. The pentlandite broke down to Ni-bearing
13 magnetite (Fig. 14a, b), most likely during the transition from lizardite to antigorite. The Ni
14 released was locally redistributed in the rock (Fig. 7), but the S and perhaps other trace
15 elements were flushed from the rock. The removal of S from the rock is shown through the
16 large decrease in S content between the Lz-serpentinites (~447 ppm) and Atg-serpentinites
17 (~16 ppm; Fig. 7).
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25 [Insert figure 14 near here]
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28 The REE show interesting variability in the serpentinites, talc-rich rocks and carbonate veins.
29 Carbonate veins are the most enriched in REE and in fact show higher concentrations of
30 incompatible trace elements (Pb, Sr, Ba, HFSE, Th and Y) than the serpentinites (Fig. 8a to
31 d) Thin-section and SEM studies reveal that these veins contain disseminated zircon, apatite,
32 monazite and rutile that hold these elements (Fig. 14c, d). The texture and paragenetic
33 sequence of monazite and rutile (Fig. 14d) indicate they precipitated from the carbonate fluid.
34 Several studies have shown that CO₂-rich fluids promote REE mobilization mainly as REE-
35 carbonate complexes (Tsikouras et al. 2006; Tsay et al. 2014; Chen and Zho 2015). The Atg-
36 serpentinites and talc-rich rocks show lower REE concentrations than the Lz-serpentinites
37 (Fig. 8d). We suggest that REEs were remobilized by the CO₂-bearing fluids accompanying
38 formation of antigorite from lizardite and during listvenitization, as has been suggested
39 previously (e.g. Vils et al. 2011; Kodolányi and Pettke 2011; Debret et al. 2013). The higher
40 concentration of REE in talc-rich rocks relative to Atg-serpentinites (Fig. 8d) may also be
41 attributed to REE uptake in the secondary mineral phases such as tremolite and anthophyllite.
42 It has been shown that amphiboles can incorporate significant amounts of REE at
43 hydrothermal temperatures (e.g. Gillis and Meyer 2001; Boschi et al. 2006).
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50 There has been considerable recent research into the enrichments of fluid-mobile elements
51 (FME) in serpentinites formed in mantle wedge settings above subduction zones (e.g. Hattori
52 and Guillot 2003, 2007; Deschamps et al. 2011, 2012). These studies show that the FME
53 which include B, Li, S, As, Rb, Sr, Sb, Cs, Ba, Pb, and U are enriched in mantle wedge
54 serpentinites due to release from metamorphic dehydration reactions in the subducting slab
55 (Deschamps et al. 2013 and references therein), and that mantle wedge serpentinite can act as
56 a temporary residence for the FME until breakdown of antigorite occurs at temperatures
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3 >600°C (Deschamps et al. 2013 and references therein). The FME are suggested be retained
4 in the serpentinite minerals with only minor release during the lizardite to antigorite transition
5 (Deschamps et al. 2013), but our results do not support this conclusion. A comparison of
6 FME and Au concentrations between Lz- and Atg-serpentinities investigated in this study is
7 shown in figure 15. Both varieties of serpentinites are characterized by significant enrichment
8 in some FME such as As, Sb and Cs compared to primitive mantle (McDonough and Sun
9 1995) and depleted mantle values (Salters and Stracke 2004). With the exception of Sb and
10 Cs, all of the FME shown in figure 15 have lower concentrations in the Atg-serpentinities
11 compared to the Lz-serpentinite suggesting that these elements were mobilized during the
12 lizardite-antigorite transition.
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17 Gold concentrations show similar variations to some FME (Fig. 15). Gold concentrations in
18 both varieties of serpentinite range from 0.8 to 12.6 ppb and are commonly higher than Au
19 contents in serpentinitized mantle peridotite (i.e. ~3-5 ppb; Buisson and Leblanc 1987),
20 average of ophiolitic rocks (i.e. ~2.8 ppb; Foster 1991), and the depleted mantle (i.e. ~1 ppb;
21 Salters and Stracke 2004). Similar to the case for S, Lz-serpentinities contain more Au (~8.0
22 ppb) compared to Atg-serpentinities (~2.3 ppb, Fig. 7 and 15). The similar behavior of Au and
23 S suggests that both may reside in nickel-sulphides in harzburgites and Lz-serpentinities, as
24 has been suggested by previous research into ANS ophiolitic serpentinites (Takla and Surour
25 1996; Khalil et al. 2003; Emam and Zoheir 2013). A recent study by Ferraris and Lorand
26 (2015) in the Lherz peridotites suggests olivine as one of the main potential host minerals for
27 gold in the ultramafic rocks. This study showed that up to 80% of the gold in these samples
28 occurs in form of nano-inclusions within olivine, whereas only 20% of the whole-rock Au-
29 budget occurs within Cu-Fe-Ni sulphides (Ferraris and Lorand 2015). The Au in Lz-
30 serpentinites most likely sits within altered olivine grains or within Ni-sulphides with the
31 decrease of whole rock Au concentrations during lizardite-antigorite transition indicating
32 liberation of this Au by the CO₂-rich fluid. This is the first time that Au has been observed to
33 be remobilized during lizardite-antigorite transition.
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45 Listvenitic rocks have long been considered an indicator of Au mineralization, but the
46 relationship between Au and listvenite is unclear. Compared to Atg-serpentinite, the talc-rich
47 rocks in our study show lower concentrations of Au (~0.7 ppb), As (all values below
48 detection limit except for Me1) and S, suggesting that intense alteration of the serpentinites to
49 form talc-rich rocks within shear zones was accompanied by fluid migration that depleted
50 these rocks in Au, As and S. In other studies, listvenites are suggested to host Au
51 mineralization (e.g. Halls and Zhao 1995). A number of Au deposits in the ANS such as the
52 Barramiya deposit (Fig. 2) occur in part within shear-hosted listvenite (Zoheir and Lehmann
53 2011). At these localities, the listvenites are enriched in sulphides and Au (e.g. Azer 2013).
54 Our study suggests that Au is removed from Lz-serpentinities when these convert to Atg-
55 serpentinites, which may happen synchronous with carbonation and shearing events. As
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listvenitic rocks are concentrated in fault and shear zones, element mobility is prone to overprinting continuous fluxes of fluids through these channels. We suggest that the association between Au mineralization and listvenitization may be due to their soft permeable nature which provides strong rheological contrast with more coherent rocks such as granitoids that commonly occur in contact with listvenites in ANS Au deposits (e.g. Zoheir and Lehmann 2011; Azer 2013; Zoheir and Moritz 2014).

6.7. The evolution of serpentinization and carbonation and the implications for formation of Au deposits

The samples investigated in this study show a systematic evolution in the conditions and intensity of fluid-driven alteration of the ultramafic portions of the ophiolitic fragments near the Meatiq area. The Lz-serpentinite samples represent the “least altered” serpentinites. The stability of lizardite in these rocks indicates that alteration occurred at temperatures lower than 300°C. The stable and radiogenic isotope compositions in Lz-serpentinite samples suggest that their alteration was driven by influx of a mixed mantle- and surface-derived hydrous fluid. In contrast, Atg-serpentinites contain abundant intergrown antigorite and carbonates, which combined with the textures of Cr-spinel indicates higher temperature alteration. The relict mesh texture in the Atg-serpentinites (Fig. 3b and d) indicates that these rocks originated as Lz-serpentinites. The transition of lizardite to antigorite was most likely driven by influx of CO₂-bearing fluids (Fig. 16). Shear zones within the Atg-serpentinite host listvenitic rocks including the quartz-carbonate veins. The talc-rich rocks most likely formed from breakdown of antigorite in the presence of CO₂ and during shear deformation possibly related to Najd deformation ~600 Ma. Fluid inclusions from shear zone carbonate and quartz vein samples indicate precipitation at 270°C to 300°C and 0.7 to 1.1 kbar which corresponds to lithostatic depths of 2.6 to 4.2 km for talc-rich rocks. Carbonate veins preserve the clearest signature of the fluid source with C and Sr isotopes indicating a mantle-derived CO₂ source. Figure 16 summarizes the evolution of serpentinization and listvenitization due to mantle-driven CO₂ fluid and associated element mobility in forearc settings.

[Insert figure 16 near here]

The chemical effects of the alteration process are dominated by addition of volatiles and removal of some trace elements including FME, S and Au that were strongly mobilized during the alteration process. These elements are enriched during initial serpentinization but are removed during flushing by a mantle-derived CO₂-bearing fluid travelling through shear zones accompanying formation of antigorite. The extent of carbonate alteration in ANS ophiolitic fragments indicates that mass of metals mobilized during the lizardite-antigorite transition must have been significant. Atg-serpentinites thus may be the source for some ANS Au deposits. Continued flow of the metal-bearing CO₂-rich fluid along shear zones promoted listvenitization and further leaching of metals from the serpentinites. The deposition of Au from this fluid could have occurred in locations where fluid compositions changed abruptly or in regions of rheological contrast where porosity is developed, in upflow zones, or in the

vicinity of chemically reactive rocks such as carbonaceous shales (Harraz 2000; Botros 2004; Zoheir and Lehmann 2011; Azer 2013; Zoheir and Moritz 2014).

There are few stable isotope compositions of vein minerals in ANS Au deposits but what is available shows similarities to results obtained in this study. Vein carbonate from the diorite-hosted Romite deposit in the Southern Eastern Desert (Fig. 2) has $\delta^{18}\text{O}$ ranging from 7 to 9 ‰ and $\delta^{13}\text{C}$ from -4.6 to -3.1 ‰ (Zoheir 2012). The El Sid deposit occurs on the margins of a granitoid intrusion into a carbonaceous shale; a listvenite-bearing shear zone yields O and C isotope compositions of vein carbonate of between 4.9 to 7.4 ‰ $\delta^{18}\text{O}$ and -14.7 to -13.9 ‰ $\delta^{13}\text{C}$, respectively (Zoheir and Moritz 2014). These values are similar to those in the carbonate altered ophiolitic rocks investigated in this study except for the very low $\delta^{13}\text{C}$ values at El Sid which are interpreted to be due to interaction between ore fluids and the carbonaceous shale (Zoheir and Moritz 2014). Fluid inclusion compositions from these mineral deposits yield similar aqueous carbonic fluids with similar homogenization temperatures to our results.

7. Conclusions

We investigated the causes and effects of carbonate alteration of Neoproterozoic ultramafic rocks in the Central Eastern Desert of Egypt. The samples studied from three areas around the Meatiq Dome show evolution in the extent and conditions of alteration from least altered lizardite serpentinites with mesh-textured alteration and late stage chrysotile veins to antigorite serpentinite with abundant antigorite intergrown with magnesite, dolomite and minor talc to talc-tremolite and talc-tremolite-anthophyllite assemblages in foliated talc-schists and talc-rich rocks. The listvenitic rocks occur in shear zones and contain carbonate and quartz veins. The sequence from lizardite serpentinite to antigorite serpentinite and different listvenitic assemblages represents evolving conditions of alteration from below 300°C to above 400°C, and an increase in shear stress with variable activity of CO_2 , SiO_2 and H_2O in the fluid. We suggest that this evolution was synchronous with and partly controlled by influx of an aqueous carbonic fluid. The fluid composition is inferred from fluid inclusion compositions in carbonate and quartz veins in the talc-rich bodies show carbonic ($\text{CO}_2 \pm \text{CH}_4 \pm \text{N}_2$) and aqueous-carbonic ($\text{H}_2\text{O} - \text{NaCl} - \text{CO}_2 \pm \text{CH}_4 \pm \text{N}_2$) fluid inclusions with low salinity, with trapping conditions of 270°C to 300°C and 0.7 to 1.1 kbar. Stable (C, O) and radiogenic (Sr) isotope compositions suggest that the carbonating fluid was a mantle-derived CO_2 -rich fluid that mixed with surficial fluid. The extent of carbonate alteration suggests that fluxing of mantle CO_2 was significant, although we cannot yet constrain how long the carbonation event lasted. Carbonation was associated with shearing and probably related to Najd deformation at ~600 Ma.

The Lz-serpentinites are enriched in Au, As, S and other fluid-mobile elements relative to primitive and depleted mantle. The extensively carbonated Atg-serpentinites contain less of these elements than the Lz-serpentinites suggesting that they were removed during carbonate alteration. The listvenitic bodies are even more depleted. We suggest that these elements were removed from the altered serpentinites by the infiltrating CO_2 -bearing fluid. Fluid inclusion and stable isotope compositions of Au deposits in the CED are similar to those from

the carbonate veins investigated in this study and we suggest that carbonation of ultramafic rocks in the ophiolitic fragments in the ANS due to influx of mantle-derived CO₂ fluid produced Au-bearing fluids that contributed importantly to the formation of the Au-deposits in the ANS.

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Figure captions

Figure 1. Simplified geological map showing the extent of the ANS (inset), the distribution of ophiolitic bodies and selected Au deposits in the Egyptian segment of ANS.

Figure 2. (a) General geological map of Neoproterozoic basement in the Meatiq area (after El-Gaby et al. 1984; Loizenbauer et al. 2001; Andresen et al. 2009, 2010; Abd El-Rahman et al. 2010), showing important lithologies, sample localities and strike-slip structures. (b) Inset shows the location of the Meatiq area in the CED, Egypt. (c) Schematic cross-section across Meatiq Gneiss Dome showing the relationship between the Meatiq Dome lithologies, the surrounding ophiolitic rocks, Hammamat clastics and Dokhan Volcanics. The ages are from Andresen et al. (2009).

Figure 3. Field photographs of serpentinites and talc-rich rocks in the study area. (a) Mesh-textured Lz-serpentinite from locality A with visible chrysotile veinlets. (b) Lz-serpentinite exposures in locality A showing cross-cutting veinlets of fibrous chrysotile (Ctl). (c) Atg-serpentinite showing relict mesh textures from locality B. (d) Exposures of altered Atg-serpentinite from locality B. The pale layers are highly sheared talc-rich rocks. Location of photo is shown in Figure 4. (e) A carbonate (Cb) and quartz (Qz) vein hosted within sheared talc-rich rocks from locality B. Location of photo is shown in Figure 4. (f) The talc quarry

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3 from locality B showing cut surfaces of Atg-serpentinite. Location of photo is shown in
4 Figure 4 (g) talc-rich rocks in a shear zone with sharp contacts to altered serpentinites in
5 locality C. (h) An exposure of serpentinite with pale coloured shear-hosted talc-rich rocks
6 from locality C. Sharp contact with shear zone to altered serpentinite control the location of
7 the talc-rich rocks. (i) Sheared talc-schist between more coherent bodies of Atg-serpentinite
8 from Locality C. (j) The view north from locality C showing extensive outcrop of shear-
9 hosted listvenitic rocks within serpentinite.
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13 Figure 4. Simplified geological map of locality B showing patches of talc-rich rocks within
14 antigorite serpentinite, the location of the talc quarry, sampling localities, and the position
15 where photos in figure 3 were taken. The dotted line leading to the quarry is a road.
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18 Figure 5. Photomicrographs (a, b, d, e, g, h) and back-scattered electron (BSE) images (c, f
19 and i) of the investigated ophiolitic rocks in Meatiq area. (a) Lz-serpentinite with mesh
20 texture and lizardite hourglass pseudomorphs after olivine together with late chrysotile
21 veinlets, locality A-sample Me21. (b) Late chrysotile (Ctl) veinlets invaded by diffusive
22 dolomite (Dol), locality A-sample Me22. (c) Slightly altered Cr-spinel with a very thin
23 narrow Cr-rich magnetite (Cr-Mag) rim, locality A-sample Me21. (d) Antigorite flakes are
24 intergrown with magnesite (Mgs) and minor talc (Tlc) in Atg-serpentinite, locality B-sample
25 Me5. (e) Atg-serpentinite invaded by vein of talc and magnesite, locality B-sample Me9. (f)
26 Cr-spinel alteration to ferritchromite (Fe-Chr) and Cr-rich magnetite (light-grey) with
27 preservation of unaltered chromite (Chr, dark-grey) in the core, locality B-sample Me5. (g)
28 Dynamically crystallized tremolite (Tr) in highly foliated talc-schist during progressive
29 deformation, locality C-sample Me15. (h) Talc-schist rock with long anthophyllite fibres and
30 ferritchromite embedded in chromium clinocllore (Clc), locality C-sample Me17. (i)
31 Cataclastic texture in altered Cr-spinel (Fe-Chr) grain indicates intensive shearing in fault
32 zones locality C-sample Me1.
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38 Figure 6. Three-dimensional representation of LOI and major elements, showing their
39 distribution in two varieties of serpentinites and talc-rich rocks in the study area.
40 Compositions are recalculated on a volatile free basis.
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43 Figure 7. Harker-type diagrams showing variations in compatible trace elements of different
44 serpentinites and talc-rich rocks. Au values plotted from two sets of serpentinites analyses.
45 Question marks indicate values below detection limit.
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48 Figure 8. Trace element patterns for two varieties of serpentinites, talc-rich rocks and
49 carbonate veins. (a) Compositions of Lz- and Atg-serpentinites. Primitive mantle normalizing
50 values are after McDonough and Sun (1995). (b) and (c) REE compositions of two varieties
51 of serpentinites, carbonate veins, and talc-rich rocks, respectively. (d) Average compositions
52 of Lz- and Atg-serpentinites, talc-rich rocks and carbonate veins. Chondrite-normalizing
53 values after McDonough and Sun (1995).
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3 Figure 9. Plots of O, C and Sr isotope data for Meatiq serpentinites and carbonate veins. Data
4 for Central Eastern Desert (CED) and sedimentary carbonates are adopted from Stern and
5 Gwinn (1990). Field of Neoproterozoic mantle-derived carbonates (dashed-line boxes) is
6 taken from C-isotopic data of Taylor et al. (1967) and Deines (1989), and Sr-isotopic data for
7 Egyptian basement rocks (solid-line box) is from Stern and Hedge (1985). Data for Oman
8 listvenites are adopted from Falk and Kelemen (2015). (a) $\delta^{13}\text{C}$ vs. $^{87}\text{Sr}/^{86}\text{Sr}$ at $\sim 600\text{Ma}$; (b)
9 $\delta^{13}\text{C}$ vs. $\delta^{18}\text{O}$ from the same samples, with similar fields displayed; (c) $\delta^{18}\text{O}$ vs. $^{87}\text{Sr}/^{86}\text{Sr}$ at \sim
10 600Ma with similar reference fields outlined.

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14 Figure 10. Photomicrographs of typical fluid inclusions in quartz/carbonate veins: (a) type-I,
15 aqueous inclusion ($\text{LH}_2\text{O}+\text{VH}_2\text{O}$). (b) Type-II, aqueous-carbonic ($\text{LH}_2\text{O}+\text{LCO}_2+\text{VCO}_2$). (c)
16 Type-III, carbonic inclusion ($\text{LCO}_2+\text{VCO}_2$). (d) Trails of secondary inclusion crossing grain
17 boundaries. (e) Elongation and necking down of inclusions in the direction of the trails. (f)
18 Fluid inclusion with shear sense in magnesite vein (Me9).
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24 Figure 11. (a) Ternary plot of Al_2O_3 - MgO - CaO for bulk rock compositions of Meatiq
25 serpentinites compared with peridotites from other tectonic settings (after Coleman 1977). (b)
26 CaO vs. Al_2O_3 diagram comparing Meatiq serpentinites with peridotites from other tectonic
27 settings (after Ishii et al. 1992). (c) Cr# versus Mg# for the analyzed spinel cores in Meatiq
28 serpentinites. Data for different tectonic environments were compiled from the literature
29 following Stern et al. (2004). The field for core of Cr-spinel compositions occurring in some
30 of the serpentinites from the Central Eastern Desert of Egypt (CED) is also shown (Azer and
31 Stern 2007; Hamdy et al. 2013; Azer 2014). MORB = mid-ocean ridge basalt. Compositions
32 are recalculated on a volatile free basis.
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37 Figure 12. P-T diagram based on intersecting isochores of aqueous fluid inclusions with
38 homogenization of 225°C and a salinity of 5 mass% NaCl and CO_2 inclusions with molar
39 volumes of 65 and $90\text{ cm}^3/\text{mol}$. The diagram shows estimated trapping condition for fluid
40 inclusions in carbonate veins at 270°C - 300°C and 0.7-1.1 kbar.
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44 Figure 13. Sr composition of carbonate veins and altered serpentinites in the Meatiq study
45 area compared to values from Neoproterozoic intrusive carbonates from Eastern Desert of
46 Egypt. The ages of emplacement for the intrusive carbonates are estimated from the age of
47 the surrounding basement units (after Cavanagh 1979; Stern and Hedge 1985; Reischmann
48 1986; Stern and Gwinn 1990). Field of initial $^{87}\text{Sr}/^{86}\text{Sr}$ for Egyptian basement igneous rocks
49 is shown after Stern and Hedge (1985); this field represents an upper limit for the isotopic
50 composition of Sr in the depleted mantle beneath northeast Africa at that time. Fields of
51 initial $^{87}\text{Sr}/^{86}\text{Sr}$ for CED carbonate veins is shown after Stern and Gwinn (1990). Isotopic
52 composition of Neoproterozoic seawater is also shown after Halverson et al. (2007).
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56 Figure 14. Back-scattered electron (BSE) images showing different mineral phases in
57 serpentinites and carbonate veins. (a) Ni-sulphide (i.e. pentlandite-Pn) in Lz-serpentinites
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3 converted to Ni-rich magnetite (Ni-Mag) in the rims (Me23). (b) Total conversion of
4 pentlandite to Ni-rich magnetite in Atg-serpentinites (Me7). Figures (c) and (d) show REE-
5 bearing minerals (Zircon-Zr, apatite-Ap, Monazite-Mnz and rutile-Rt) in a carbonate vein
6 (Me12).
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10 Figure 15. Plots of concentrations of fluid-mobile elements (Li, As, S, Sb, Pb, U, Ba, Sr and
11 Cs) and Au in two serpentine phases from the Meatiq study area. Values of primitive mantle
12 after McDonough and Sun (1995) and depleted mantle after Salters and Stracke (2004) are
13 also reported (thick and dashed black lines). Grey boxes represent the ranges between
14 primitive and depleted mantle. Au values plotted from two sets of serpentinites analyses.
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18 Figure 16. Schematic cross section illustrating the characteristics and evolution of the
19 different serpentinite and listvenitic rocks and the tectonic setting where, based on the
20 alteration assemblages and fluid inclusion compositions, these rocks formed. The principal
21 transfer of fluid-mobile elements is also noted, as well as geochemical signature of fluids
22 released during dehydration of serpentinites, “antigorite formation” due to the prograde
23 metamorphism and later “listvenitization”. Mineral abbreviations as in Fig. 5.
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Table 1 Chemical composition of different rock types and carbonate veins in the study area

| Lithology sample | Atg-serpentinites | | | | | Lz-serpentinites | | | | Talc-rich rocks | | | | Carbonate veins | | |
|--------------------------------|-------------------|--------|--------|--------|--------|------------------|--------|--------|--------|-----------------|--------|--------|--------|-----------------|-------|-------|
| | Me5 | Me6 | Me7 | ME8 | Me9 | Me21 | Me22 | Me23 | Me24 | Me1 | Me14 | Me15 | Me17 | Me2 | Me12 | Me28 |
| SiO ₂ (wt.%) | 40.26 | 36.97 | 34.37 | 37.12 | 32.75 | 40.43 | 39.29 | 38.6 | 37.76 | 55.5 | 59.23 | 58.95 | 57.92 | n.d. | n.d. | n.d. |
| Al ₂ O ₃ | 0.48 | 0.51 | 0.31 | 0.51 | 0.12 | 0.51 | 0.5 | 0.48 | 0.55 | 1.66 | 1 | 0.38 | 1.57 | n.d. | n.d. | n.d. |
| FeO(T) | 7.63 | 9.07 | 7.36 | 6.3 | 5.86 | 6.78 | 7.28 | 7.49 | 7.81 | 7.6 | 4.8 | 5.33 | 7.98 | n.d. | n.d. | n.d. |
| MnO | 0.07 | 0.09 | 0.14 | 0.09 | 0.11 | 0.07 | 0.08 | 0.08 | 0.10 | 0.16 | 0.06 | 0.12 | 0.12 | n.d. | n.d. | n.d. |
| MgO | 38.43 | 38.52 | 39.11 | 38.81 | 39.18 | 38.22 | 37.99 | 37.47 | 37.68 | 20.16 | 27.08 | 22.52 | 23.36 | n.d. | n.d. | n.d. |
| CaO | 0.02 | 0.02 | 0.03 | 0.02 | 0.05 | 0.27 | 0.76 | 1.11 | 1.13 | 12.6 | 2.71 | 11.24 | 6.55 | n.d. | n.d. | n.d. |
| Na ₂ O | < 0.01 | 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | 0.01 | < 0.01 | 0.08 | 0.03 | 0.05 | 0.09 | n.d. | n.d. | n.d. |
| K ₂ O | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | 0.01 | < 0.01 | < 0.01 | < 0.01 | n.d. | n.d. | n.d. |
| TiO ₂ | 0.002 | 0.001 | 0.002 | 0.001 | 0.001 | 0.002 | 0.001 | 0.003 | 0.002 | 0.025 | 0.006 | 0.014 | 0.033 | n.d. | n.d. | n.d. |
| P ₂ O ₅ | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | 0.02 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | < 0.01 | n.d. | n.d. | n.d. |
| L.O.I. | 12.55 | 15.28 | 18.34 | 15.06 | 20.84 | 13.19 | 13.55 | 13.85 | 14.19 | 2.1 | 4.1 | 2.34 | 3.13 | – | – | – |
| Total | 99.43 | 100.5 | 99.67 | 97.92 | 98.92 | 99.48 | 99.46 | 99.11 | 99.22 | 99.89 | 99.01 | 101 | 100.8 | n.d. | n.d. | n.d. |
| Sc (ppm) | 7.8 | 8.7 | 8.3 | 8.5 | 2.6 | 7.2 | 6.9 | 6.2 | 5.9 | 5 | 5 | 5 | 24 | 3.9 | 6.8 | 3.3 |
| V | 29.4 | 34.5 | 23.9 | 29.7 | 12.5 | 28.1 | 36.7 | 49.8 | 20.2 | 53 | 27 | 47 | 71 | 3.0 | 0.16 | 0.08 |
| Cr | 2526 | 3145 | 2191 | 2651 | 951 | 1220 | 1924 | 2016 | 1782 | 1670 | 1990 | 230 | 1830 | 7.5 | 3.6 | 3.3 |
| Co | 93.3 | 102 | 123 | 99.1 | 99.6 | 110.6 | 101.5 | 92 | 92.9 | 74 | 75 | 46 | 60 | 0.74 | 12.5 | 9.5 |
| Ni | 2286 | 2173 | 2310 | 2344 | 2428 | 2516 | 2255 | 1951 | 2147 | 1600 | 2020 | 619 | 371 | 4.2 | 33 | 33.6 |
| Rb | 0.03 | 0.03 | 0.02 | 0.03 | 0.01 | 0.11 | 0.09 | 0.12 | 0.03 | b.d.l. | b.d.l. | b.d.l. | b.d.l. | 0.04 | 0.01 | 0.001 |
| Sr | 0.23 | 0.42 | 0.31 | 0.40 | 0.45 | 5.82 | 18.5 | 69.8 | 56.4 | 25 | b.d.l. | 5.0 | 8.0 | 1135 | 503 | 2365 |
| Y | 0.05 | 0.04 | 0.01 | 0.03 | 0.02 | 0.52 | 0.71 | 2.1 | 0.04 | 0.6 | b.d.l. | 0.5 | 0.9 | 20.9 | 46.9 | 18.1 |
| Nb | 0.03 | 0.03 | 0.1 | 0.03 | 0.04 | 0.03 | 0.03 | 0.07 | 0.01 | b.d.l. | b.d.l. | b.d.l. | b.d.l. | 0.04 | 0.02 | 0.02 |
| Sb | 1.1 | 0.76 | 1.1 | 0.92 | 0.72 | 0.78 | 0.71 | 0.52 | 0.26 | b.d.l. | b.d.l. | b.d.l. | b.d.l. | n.d. | n.d. | n.d. |
| As | 9.3 | 6.2 | 6.4 | 8.0 | 6.1 | 46.1 | 34.5 | 16.8 | 4.1 | 66.0 | b.d.l. | b.d.l. | b.d.l. | n.d. | n.d. | n.d. |
| S | b.d.l. | 30 | 20 | b.d.l. | b.d.l. | 380 | 280 | 730 | 400 | b.d.l. | 20 | 10 | 10 | n.d. | n.d. | n.d. |
| Au (ppb) | 7.7 | 1.4 | 1.8 | 1.8 | 0.86 | 11.2 | 10.3 | 3.7 | 3.8 | 1.0 | 0.30 | 1.4 | 0.24 | n.d. | n.d. | n.d. |
| Au** | 4.3 | 1.6 | 1.4 | 1.3 | 0.85 | 12.6 | 12.4 | 4.9 | 4.6 | n.d. | n.d. | n.d. | n.d. | n.d. | n.d. | n.d. |
| Li | 479 | 430 | 413 | 407 | 567 | 5697 | 5027 | 4533 | 573 | n.d. | n.d. | n.d. | n.d. | 489 | 613 | 534 |
| Cs | 157 | 63.4 | 65.4 | 83.2 | 18.9 | 44.4 | 32.3 | 44.9 | 2.7 | b.d.l. | b.d.l. | b.d.l. | b.d.l. | 3.5 | 2.7 | 6.5 |
| Ba | b.d.l. | 226 | 341 | 168 | 522 | 646 | 1512 | 2483 | 2174 | 3 | b.d.l. | b.d.l. | b.d.l. | 4781 | 973 | 1573 |
| La | 17.0 | 16.1 | 5.3 | 12.5 | 13.7 | 530 | 589 | 566 | 65.8 | 170 | 160 | 220 | 340 | 16410 | 4711 | 2358 |
| Ce | 39.1 | 31.4 | 8.7 | 24.6 | 23.7 | 1182 | 1274 | 1181 | 30.9 | 260 | 270 | 420 | 670 | 33730 | 15330 | 7715 |
| Pr | 3.6 | 3.4 | 0.42 | 2.8 | 2.0 | 166 | 180 | 161 | 0.50 | 60 | 30 | 50 | 80 | 4324 | 2631 | 1324 |
| Nd | 18.8 | 13.7 | 1.9 | 9.1 | 9.7 | 656 | 713 | 655 | 1.1 | 200 | b.d.l. | b.d.l. | 240 | 18340 | 14190 | 7088 |
| Sm | 4.5 | 3.1 | 0.01 | 1.9 | 0.91 | 104 | 119 | 118 | b.d.l. | 30 | b.d.l. | b.d.l. | 40 | 3939 | 4888 | 2489 |
| Eu | 8.1 | 4.7 | 2.8 | 4.1 | 5.9 | 26.7 | 28.8 | 31.6 | 16.3 | 73 | b.d.l. | 6 | 38 | 1464 | 1538 | 956 |
| Gd | 5.3 | 4.5 | b.d.l. | 1.4 | 1.4 | 101 | 118 | 136 | 0.31 | 90 | b.d.l. | 40 | 100 | 3783 | 6790 | 3235 |
| Tb | 0.76 | 0.95 | 0.06 | 0.44 | 0.35 | 14.5 | 16.4 | 23.5 | 0.24 | 10 | b.d.l. | b.d.l. | 20 | 508 | 1132 | 507 |
| Dy | 6.1 | 6.9 | 0.97 | 3.0 | 1.8 | 80.6 | 96.9 | 185 | 3.2 | 70 | 30 | 60 | 110 | 2663 | 7270 | 3026 |
| Ho | 1.3 | 1.4 | 0.29 | 0.70 | 0.35 | 14.6 | 18.5 | 52.3 | 1.2 | 20 | b.d.l. | 10 | 30 | 491 | 1531 | 577 |
| Er | 4.9 | 5.7 | 2.2 | 3.9 | 2.0 | 26.4 | 45.1 | 184 | 6.3 | 50 | 30 | 50 | 90 | 1253 | 4295 | 1482 |
| Tm | 1.1 | 1.4 | 0.69 | 1.0 | 0.40 | 2.6 | 5.7 | 27.9 | 1.7 | 6 | b.d.l. | 9 | 14 | 169 | 591 | 195 |
| Yb | 11 | 12.3 | 8.7 | 8.6 | 5.9 | 14.8 | 35.7 | 168 | 15.0 | 30 | 20 | 80 | 100 | 1046 | 3720 | 1159 |
| Lu | 2.4 | 2.9 | 2.6 | 2.4 | 1.1 | 2.6 | 4.4 | 22.8 | 3.7 | 3 | 3 | 16 | 17 | 140 | 524 | 156 |
| Hf | 5.0 | 7.7 | 3.7 | 7.6 | 2.0 | 8.8 | 3.7 | 7.0 | 3.6 | b.d.l. | b.d.l. | b.d.l. | b.d.l. | 8.1 | 20 | 8.3 |
| Ta | 8.0 | 4.9 | 3.0 | 9.0 | 2.1 | 3.3 | 3.9 | 5.4 | 4.2 | b.d.l. | b.d.l. | b.d.l. | b.d.l. | 3.1 | 7.8 | 2.8 |
| Pb | 52.8 | 44.4 | 41.5 | 43.4 | 28.8 | 110 | 97.2 | 103 | 440 | 4000 | 3000 | b.d.l. | 4000 | 882 | 3176 | 4908 |
| Th | 2.9 | 2.2 | 1.8 | 1.4 | 1.0 | 3.2 | 5.7 | 14.5 | 1.8 | b.d.l. | b.d.l. | b.d.l. | b.d.l. | 18.9 | 8.8 | 5.9 |
| U | 46.7 | 39.1 | 0.87 | 1.5 | 5.7 | 255 | 375 | 504 | 4.0 | 40 | 20 | 30 | 30 | 14.9 | 0.74 | 0.44 |
| Eu/Eu* | 5.1 | 3.9 | – | 7.7 | 15.8 | 0.80 | 0.74 | 0.76 | – | 4.3 | – | – | 1.8 | 1.2 | 0.82 | 1.0 |
| (La/Sm)N | 2.4 | 3.3 | 246 | 4.3 | 9.7 | 3.3 | 3.2 | 3.1 | – | 3.7 | – | – | 5.5 | 2.7 | 0.62 | 0.61 |
| (La/Yb)N | 1.1 | 0.94 | 0.44 | 1.0 | 1.7 | 25.8 | 11.8 | 2.4 | 3.2 | 4.1 | 5.7 | 2.0 | 2.4 | 11.3 | 0.91 | 1.5 |
| (La/Lu)N | 0.77 | 0.60 | 0.22 | 0.56 | 1.3 | 22.1 | 14.4 | 2.7 | 1.9 | 6.1 | 5.7 | 1.5 | 2.1 | 12.6 | 1.0 | 1.6 |
| (Tb/Yb)N | 0.31 | 0.35 | 0.03 | 0.23 | 0.27 | 4.5 | 2.1 | 0.64 | 0.07 | 1.5 | – | – | 0.91 | 2.2 | 1.4 | 2.0 |
| Total REE (ppm) | 0.12 | 0.11 | 0.03 | 0.08 | 0.07 | 2.9 | 3.2 | 3.5 | 0.15 | 1.1 | 0.54 | 0.96 | 1.9 | 88.3 | 69.1 | 32.3 |

From Sc to S in ppm, from Au to U in ppb, n.d. (not determined), b.d.l. (below detection limit), Au** values for the second group of the same serpentinites.

Table 2 Geochemical and isotopic data, Meatiq study area

| Sample | Concentrations (ppm) ^a | | ⁸⁷ Rb/ ⁸⁶ Sr | ⁸⁷ Sr/ ⁸⁶ Sr | ⁽⁸⁷ Sr/ ⁸⁶ Sr) ₆₀₀ ^b | δ ¹⁸ O (SMOW) | δ ¹³ C (PDB) |
|-------------------|-----------------------------------|-------|------------------------------------|------------------------------------|--|-----------------------------|----------------------------|
| | Rb | Sr | | | | | |
| Atg-serpentinites | | | | | | | |
| Me05 | 0.028 | 0.23 | 0.352 | 0.70697±17 | 0.70396 | 11.0 | -5.4 |
| Me06 | 0.025 | 0.42 | | | | 10.6 | -5.9 |
| Me07 | 0.022 | 0.31 | | | | 11.0 | -5.6 |
| Me08 | 0.031 | 0.40 | 0.226 | 0.70779±15 | 0.70585 | 10.3 | -5.6 |
| Me09 | 0.009 | 0.45 | | | | 11.5 | -5.3 |
| Lz-serpentinites | | | | | | | |
| Me21 | 0.11 | 5.82 | 0.054 | 0.70669±16 | 0.70622 | 15.1 | -4.2 |
| Me22 | 0.093 | 18.49 | 0.014 | 0.70461±13 | 0.70448 | 14.0 | -4.1 |
| Me23 | 0.117 | 69.75 | 0.005 | 0.70432±16 | 0.70428 | 12.3 | -4.2 |
| Me24 | 0.033 | 56.37 | 0.002 | 0.70422±14 | 0.70421 | 13.7 | -4.2 |
| Carbonate veins | | | | | | | |
| Me02 | 0.04 | 1135 | 0.0001 | 0.70344±12 | 0.70344 | 6.4 | -8.1 |
| Me12 | 0.005 | 503 | 0.00003 | 0.70281±13 | 0.7028 | 9.4 | -6.8 |
| Me28 | 0.001 | 2365 | 0.000001 | 0.70293±13 | 0.70293 | 10.5 | -6.8 |

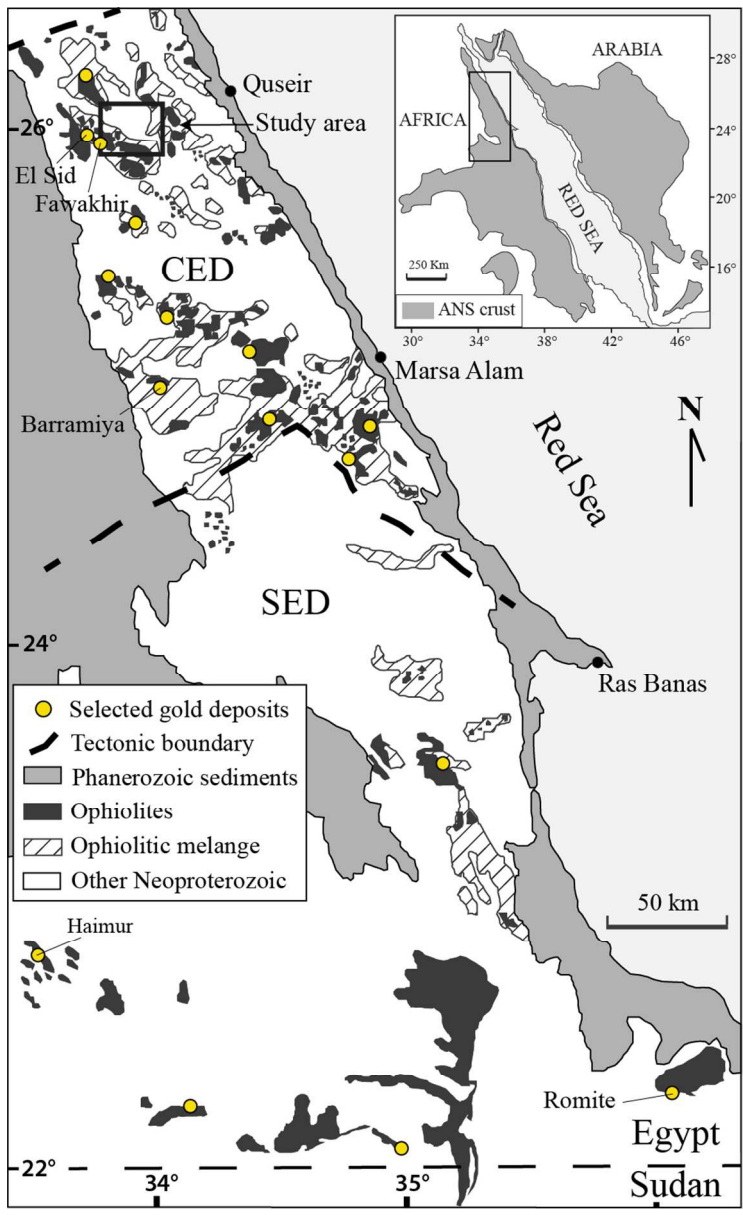
^a Blank-corrected concentrations.^b Isotopic composition at 600 Ma.

Table 3 Microthermometric data of fluid inclusions from the Meatiq study area

| Sample | Fluid inclusion Type | TmCO ₂ (°C) | Tm ice (°C) | Tm,cla (°C) | ThCO ₂ (°C) | Salinity (mass% NaCl) | Th total (°C) |
|----------|---|------------------------|---------------|-------------|------------------------|-----------------------|------------------------------|
| type-I | | | | | | | |
| Dol-Me2 | Mg-Na-Cl-H ₂ O | | -2.6 to 0.0 | | | 0.0 to 4.3 | 239 to 265 |
| Dol-Me2 | Ca-Na-Cl-H ₂ O | | -16.4 to -4.0 | | | 6.4 to 16.9* | 225 to 383 |
| Dol-Me2 | Mg-Na-Cl-H ₂ O secondary | | -1.3 to 0.0 | | | 0.0 to 2.2 | 161 to 255 |
| Qz-Me3 | Mg-Na-Cl-H ₂ O | | -3.5 to -0.1 | | | 0.2 to 5.7* | 225 to 266 |
| Qz-Me3 | Mg-Na-Cl-H ₂ O secondary | | -3.5 to -0.1 | | | 0.2 to 5.7 | 177 to 234 |
| Mgs-Me9 | Mg-Na-Cl-H ₂ O | | -2.4 to -0.5 | | | 0.9 to 5.0 | 231 to 335 |
| Dol-Me12 | Mg-Na-Cl-H ₂ O | | -2.9 to -0.8 | | | 1.4 to 4.8 | 253 to 302 |
| Dol-Me28 | Mg-Na-Cl-H ₂ O | | -3.3 to -1.0 | | | 1.7 to 5.5 | 262 to 315 |
| type-II | | | | | | | |
| Qz-Me3 | CO ₂ ±(CH ₄ /N ₂)-H ₂ O-NaCl | -57.7 to -56.6 | | 7.7 to 9.5 | 26.8 to 28.0 | 1.0 to 4.7* | 305 to 331 324 to 365 (g) |
| Dol-Me12 | CO ₂ ±(CH ₄ /N ₂)-H ₂ O-NaCl | n.d. | | 7.5 to 9.8 | | 0.2 to 4.9 | 252 to 276 |
| Dol-Me28 | CO ₂ ±(CH ₄ /N ₂)-H ₂ O-NaCl | -57.3 to -56.6 | | 6.6 to 10.0 | 26.8 to 29.2 | 0.0 to 6.8 | 242 to 331 |
| type-III | | | | | | | |
| Qz-Me3 | CO ₂ ±(CH ₄ /N ₂) | -57.3 to -56.6 | | | 26.8 to 29.5 | | |

Tm,CO₂ = melting of the CO₂; Tm,cla = melting of CO₂ hydrate; Th,CO₂ = homogenization of the CO₂ phases to liquid (l); Tm,ice = final melting of ice; Salinity = wt.% NaCl calculated from Tm,ice or Tm,cla; Th,total = total homogenization to the liquid phase or gas phase (g); Tc = critical homogenization temperature; Dol = dolomite; Qz = quartz; Mgs = magnesite; n.d. (not detected). *mass % (NaCl+CaCl₂) eq.

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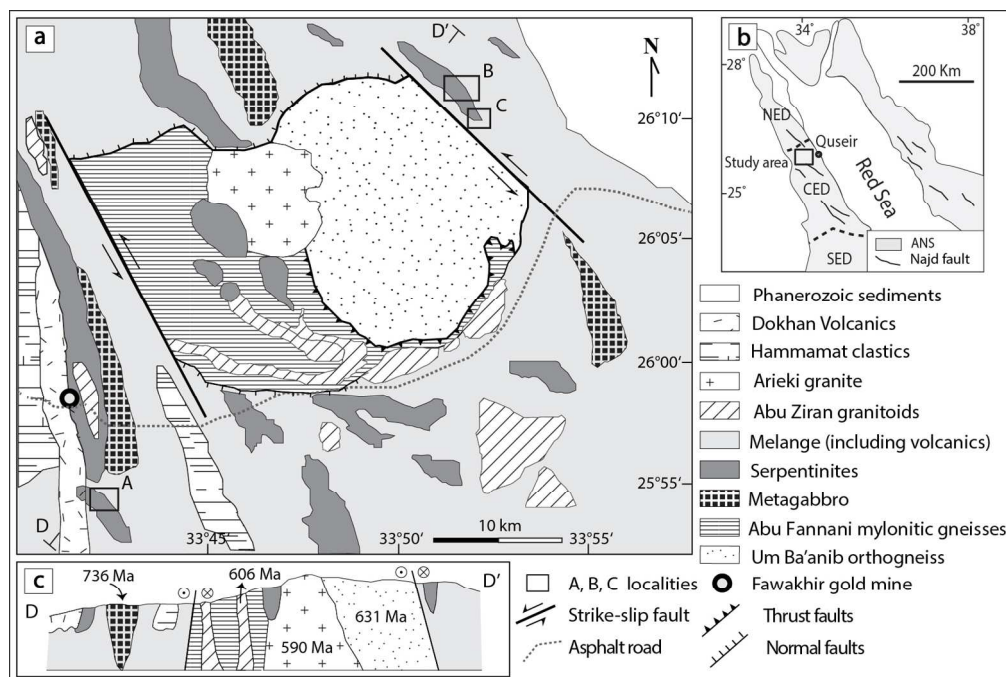


Figure 2. (a) General geological map of Neoproterozoic basement in the Meatiq area (after El-Gaby et al. 1984; Loizenbauer et al. 2001; Andresen et al. 2009, 2010; Abd El-Rahman et al. 2010), showing important lithologies, sample localities and strike-slip structures. (b) Inset shows the location of the Meatiq area in the CED, Egypt. (c) Schematic cross-section across Meatiq Gneiss Dome showing the relationship between the Meatiq Dome lithologies, the surrounding ophiolitic rocks, Hammamat clastics and Dokhan Volcanics. The ages are from Andresen et al. (2009).

144x97mm (300 x 300 DPI)



Figure 3. Field photographs of serpentinites and talc-rich rocks in the study area. (a) Mesh-textured Lz-serpentinite from locality A with visible chrysotile veinlets. (b) Lz-serpentinite exposures in locality A showing cross-cutting veinlets of fibrous chrysotile (Ctl). (c) Atg-serpentinite showing relict mesh textures from locality B. (d) Exposures of altered Atg-serpentinite from locality B. The pale layers are highly sheared talc-rich rocks. Location of photo is shown in Figure 4. (e) A carbonate (Cb) and quartz (Qz) vein hosted within sheared talc-rich rocks from locality B. Location of photo is shown in Figure 4. (f) The talc quarry from locality B showing cut surfaces of Atg-serpentinite. Location of photo is shown in Figure 4. (g) talc-rich rocks in a shear zone with sharp contacts to altered serpentinites in locality C. (h) An exposure of serpentinite with pale coloured shear-hosted talc-rich rocks from locality C. Sharp contact with shear zone to altered serpentinite control the location of the talc-rich rocks. (i) Sheared talc-schist between more coherent bodies of Atg-serpentinite from Locality C. (j) The view north from locality C showing extensive outcrop of shear-hosted listvenitic rocks within serpentinite.

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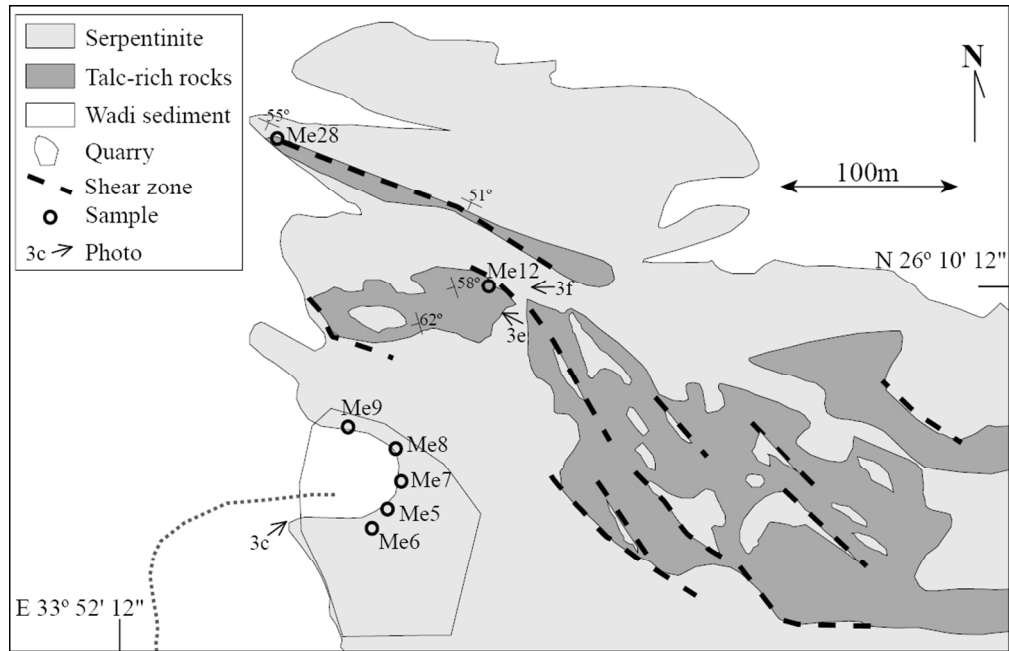


Figure 4. Simplified geological map of locality B showing patches of talc-rich rocks within antigorite serpentinite, the location of the talc quarry, sampling localities, and the position where photos in figure 3 were taken. The dotted line leading to the quarry is a road.

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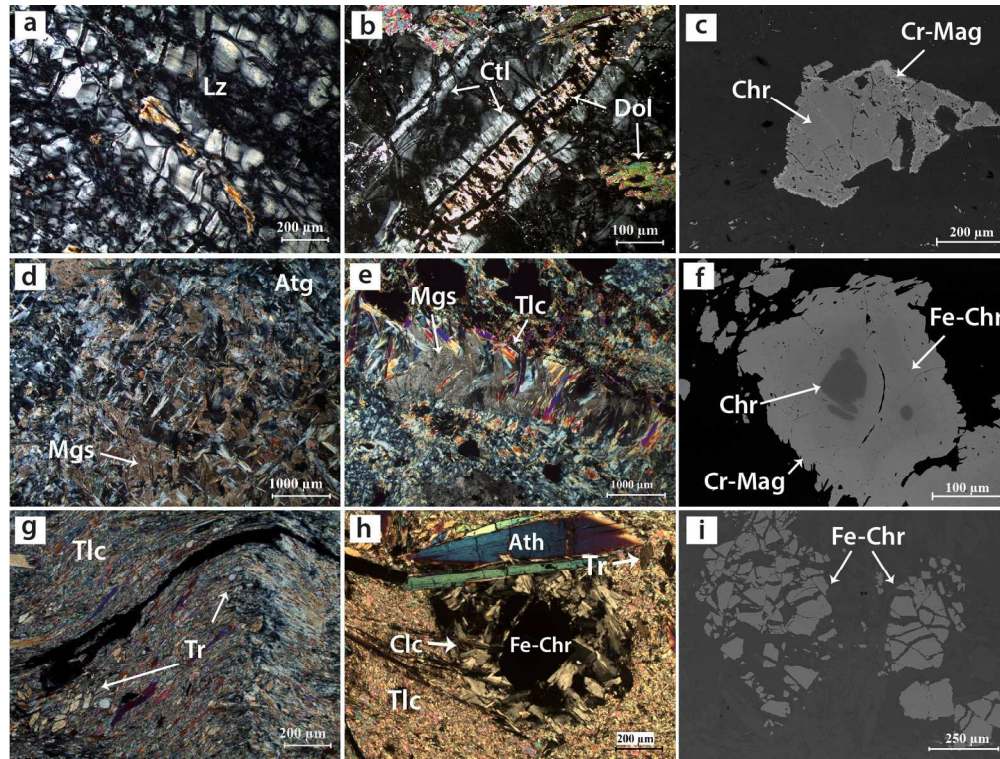


Figure 5. Photomicrographs (a, b, d, e, g, h) and back-scattered electron (BSE) images (c, f and i) of the investigated ophiolitic rocks in Meatiq area. (a) Lz-serpentinite with mesh texture and lizardite hourglass pseudomorphs after olivine together with late chrysotile veinlets, locality A-sample Me21. (b) Late chrysotile (Ctl) veinlets invaded by diffusive dolomite (Dol), locality A-sample Me22. (c) Slightly altered Cr-spinel with a very thin narrow Cr-rich magnetite (Cr-Mag) rim, locality A-sample Me21. (d) Antigorite flakes are intergrown with magnesite (Mgs) and minor talc (Tlc) in Atg-serpentinite, locality B-sample Me5. (e) Atg-serpentinite invaded by vein of talc and magnesite, locality B-sample Me9. (f) Cr-spinel alteration to ferritchromite (Fe-Chr) and Cr-rich magnetite (light-grey) with preservation of unaltered chromite (Chr, dark-grey) in the core, locality B-sample Me5. (g) Dynamically crystallized tremolite (Tr) in highly foliated talc-schist during progressive deformation, locality C-sample Me15. (h) Talc-schist rock with long anthophyllite fibres and ferritchromite embedded in chromium clinocllore (Clc), locality C-sample Me17. (i) Cataclastic texture in altered Cr-spinel (Fe-Chr) grain indicates intensive shearing in fault zones locality C-sample Me1.

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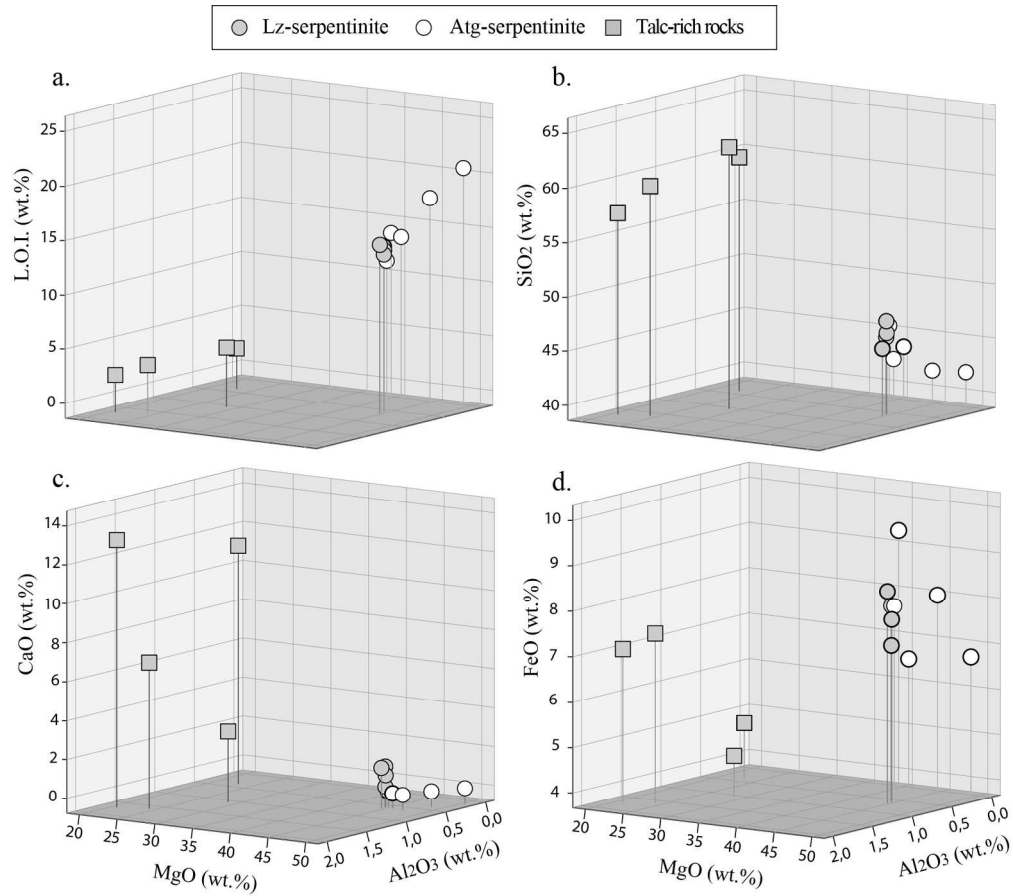


Figure 6. Three-dimensional representation of LOI and major elements, showing their distribution in two varieties of serpentinites and talc-rich rocks in the study area. Compositions are recalculated on a volatile free basis.

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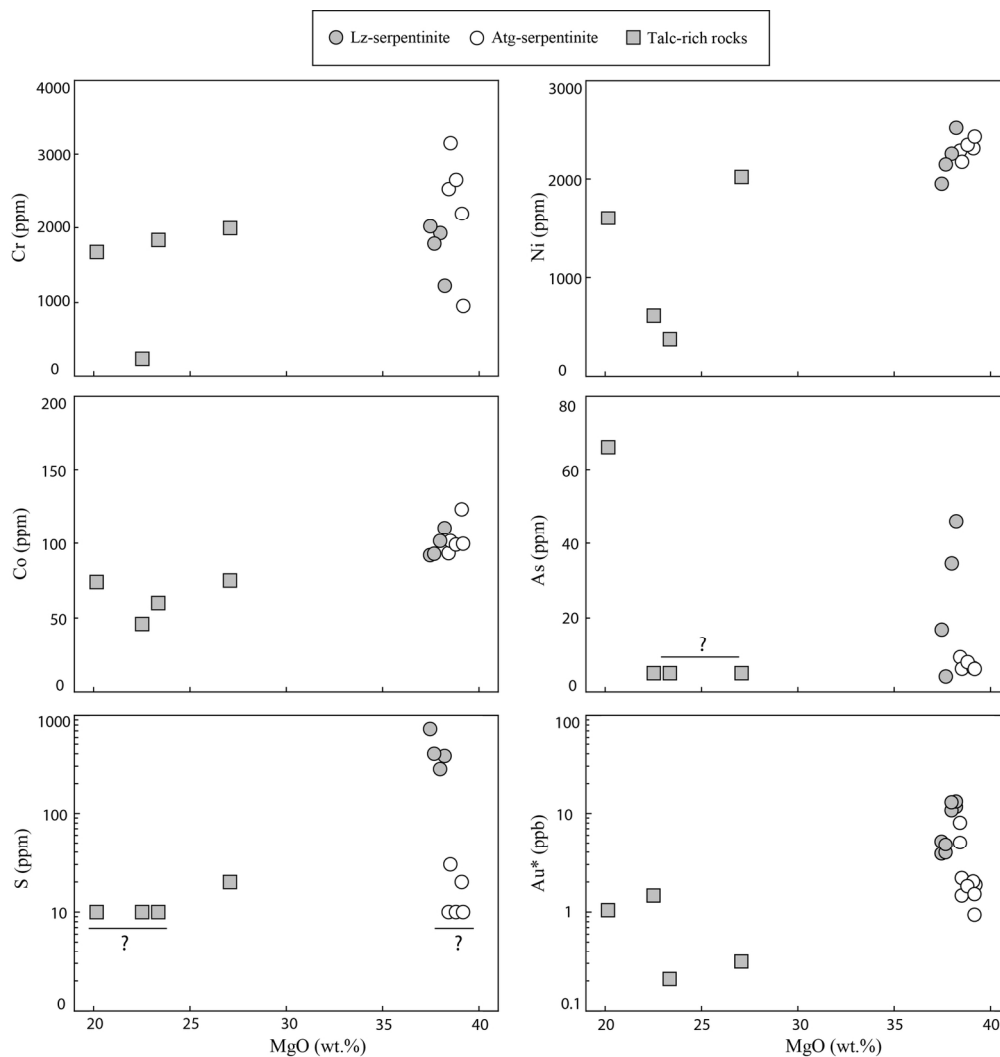


Figure 7. Harker-type diagrams showing variations in compatible trace elements of different serpentinites and talc-rich rocks. Au values plotted from two sets of serpentinites analyses. Question marks indicate values below detection limit.

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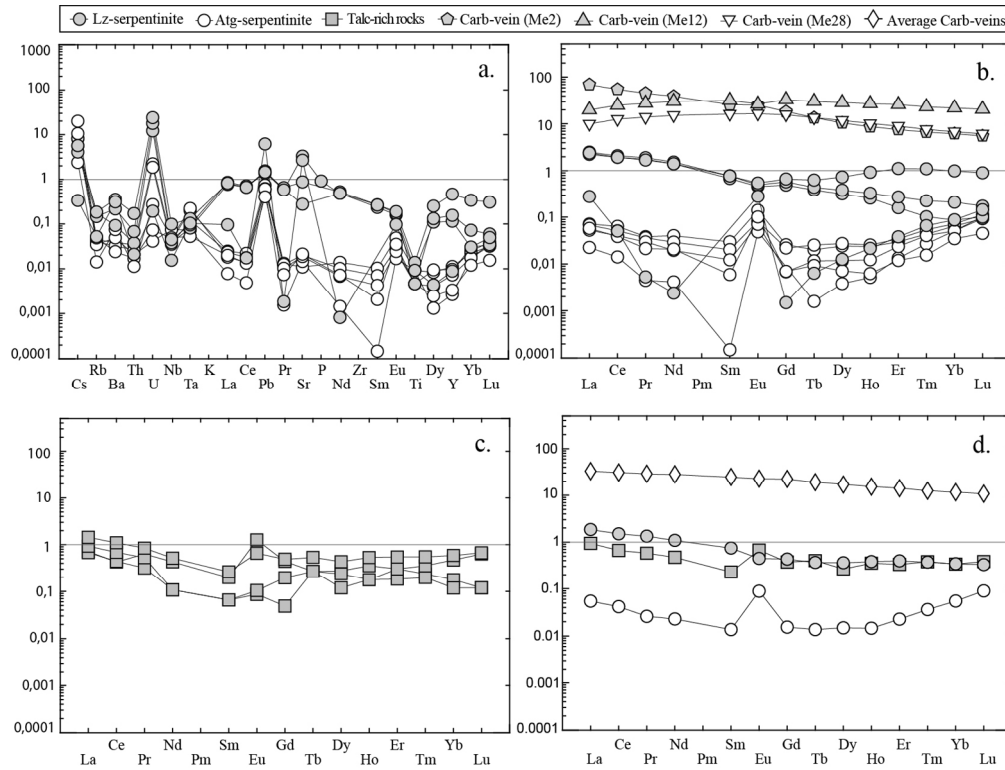


Figure 8. Trace element patterns for two varieties of serpentinites, talc-rich rocks and carbonate veins. (a) Compositions of Lz- and Atg-serpentinites. Primitive mantle normalizing values are after McDonough and Sun (1995). (b) and (c) REE compositions of two varieties of serpentinites, carbonate veins, and talc-rich rocks, respectively. (d) Average compositions of Lz- and Atg-serpentinites, talc-rich rocks and carbonate veins. Chondrite-normalizing values after McDonough and Sun (1995).

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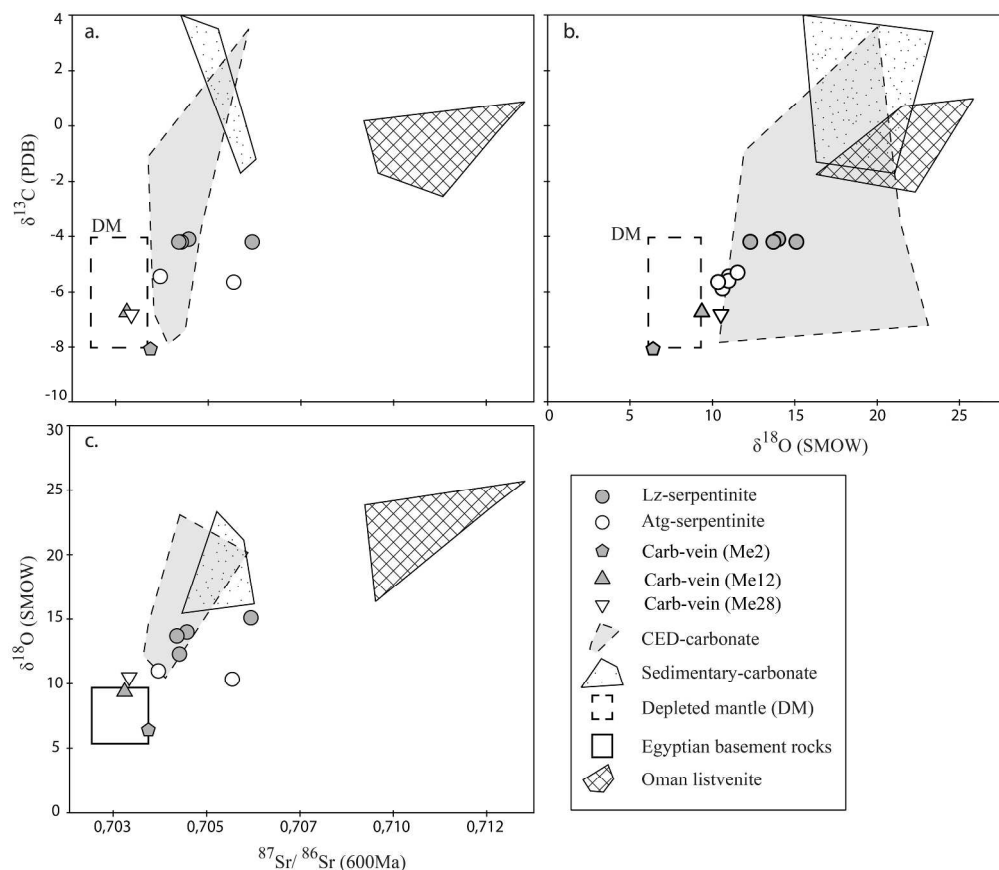


Figure 9. Plots of O, C and Sr isotope data for Meatiq serpentinites and carbonate veins. Data for Central Eastern Desert (CED) and sedimentary carbonates are adopted from Stern and Gwinn (1990). Field of Neoproterozoic mantle-derived carbonates (dashed-line boxes) is taken from C-isotopic data of Taylor et al. (1967) and Deines (1989), and Sr-isotopic data for Egyptian basement rocks (solid-line box) is from Stern and Hedge (1985). Data for Oman listvenites are adopted from Falk and Kelemen (2015). (a) $\delta^{13}\text{C}$ vs. $^{87}\text{Sr}/^{86}\text{Sr}$ at $\sim 600\text{Ma}$; (b) $\delta^{13}\text{C}$ vs. $\delta^{18}\text{O}$ from the same samples, with similar fields displayed; (c) $\delta^{18}\text{O}$ vs. $^{87}\text{Sr}/^{86}\text{Sr}$ at $\sim 600\text{Ma}$ with similar reference fields outlined.

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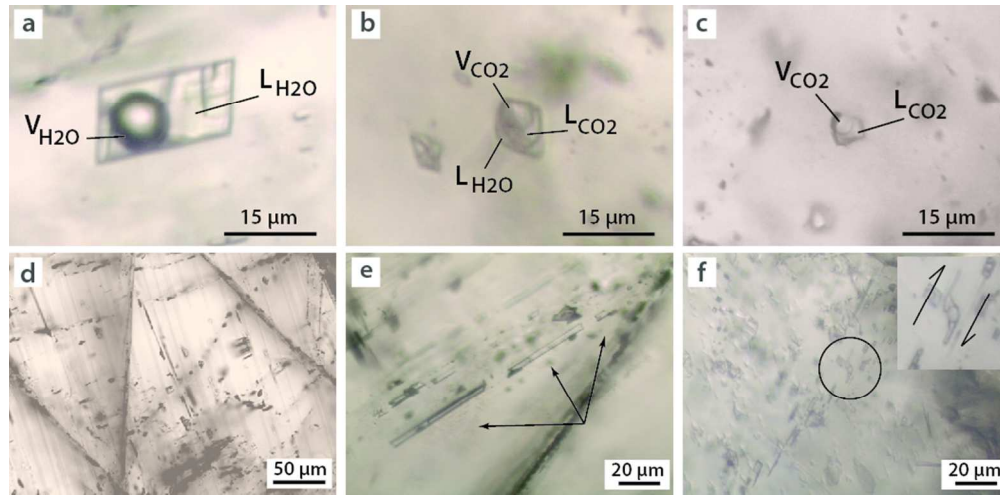


Figure 10. Photomicrographs of typical fluid inclusions in quartz/carbonate veins: (a) type-I, aqueous inclusion (LH₂O+VH₂O). (b) Type-II, aqueous-carbonic (LH₂O+LCO₂+VCO₂). (c) Type-III, carbonic inclusion (LCO₂+VCO₂). (d) Trails of secondary inclusion crossing grain boundaries. (e) Elongation and necking down of inclusions in the direction of the trails. (f) Fluid inclusion with shear sense in magnesite vein (Me9).

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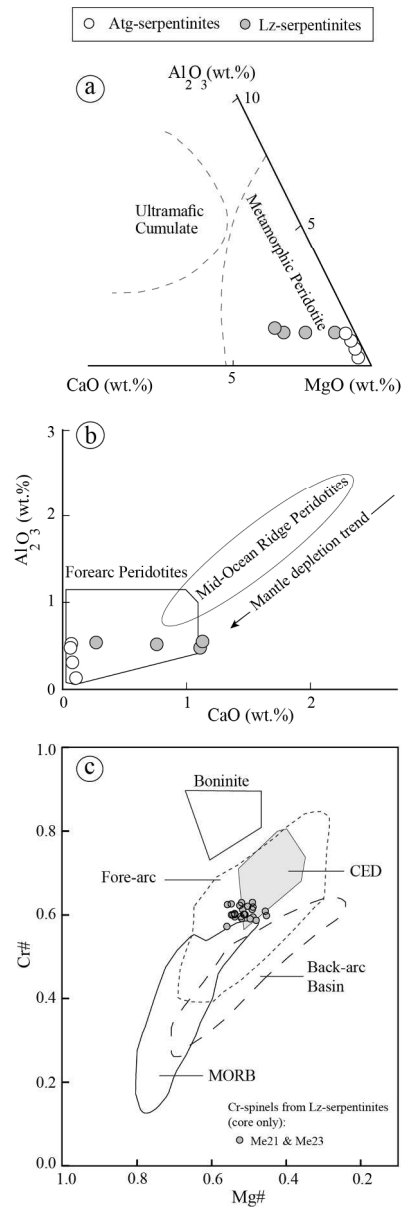


Figure 11. (a) Ternary plot of Al_2O_3 -MgO-CaO for bulk rock compositions of Meatiq serpentinites compared with peridotites from other tectonic settings (after Coleman 1977). (b) CaO vs. Al_2O_3 diagram comparing Meatiq serpentinites with peridotites from other tectonic settings (after Ishii et al. 1992). (c) Cr# versus Mg# for the analyzed spinel cores in Meatiq serpentinites. Data for different tectonic environments were compiled from the literature following Stern et al. (2004). The field for core of Cr-spinel compositions occurring in some of the serpentinites from the Central Eastern Desert of Egypt (CED) is also shown (Azer and Stern 2007; Hamdy et al. 2013; Azer 2014). MORB = mid-ocean ridge basalt. Compositions are recalculated on a volatile free basis.

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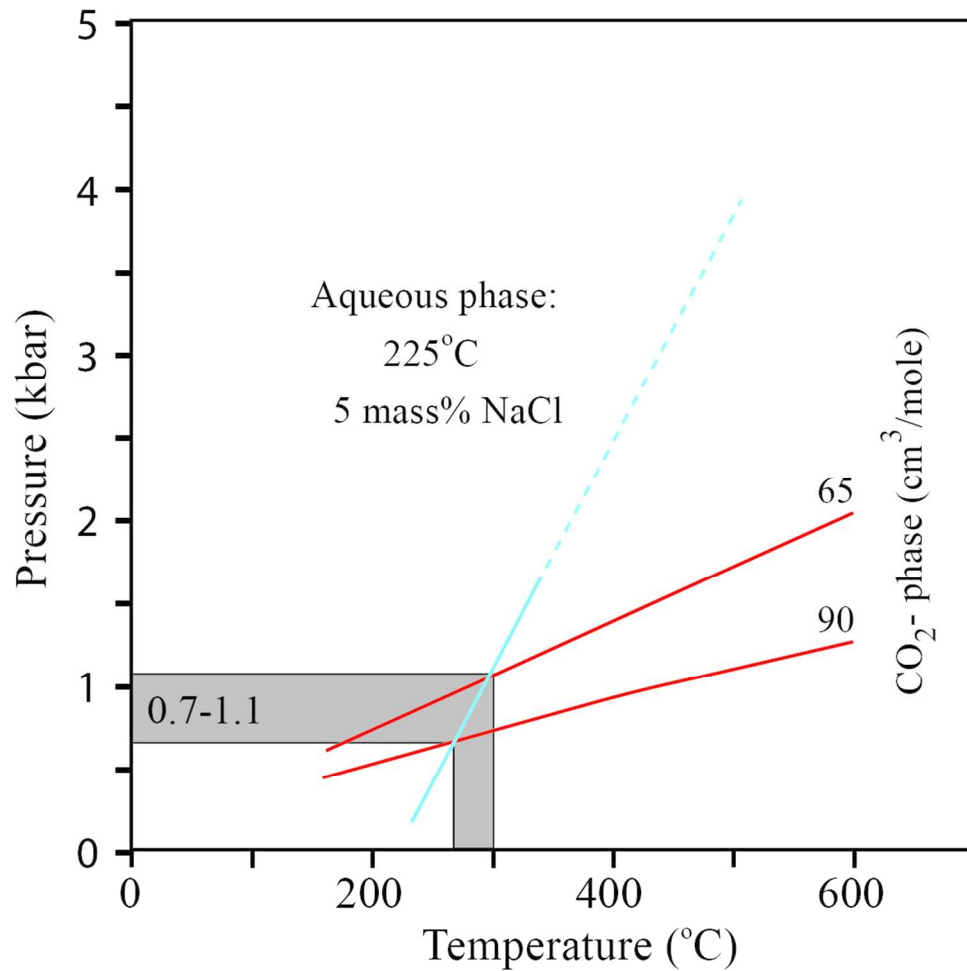


Figure 12. P-T diagram based on intersecting isochores of aqueous fluid inclusions with homogenization of 225°C and a salinity of 5 mass% NaCl and CO₂ inclusions with molar volumes of 65 and 90 cm³/mol. The diagram shows estimated trapping condition for fluid inclusions in carbonate veins at 270°C-300°C and 0.7-1.1 kbar.

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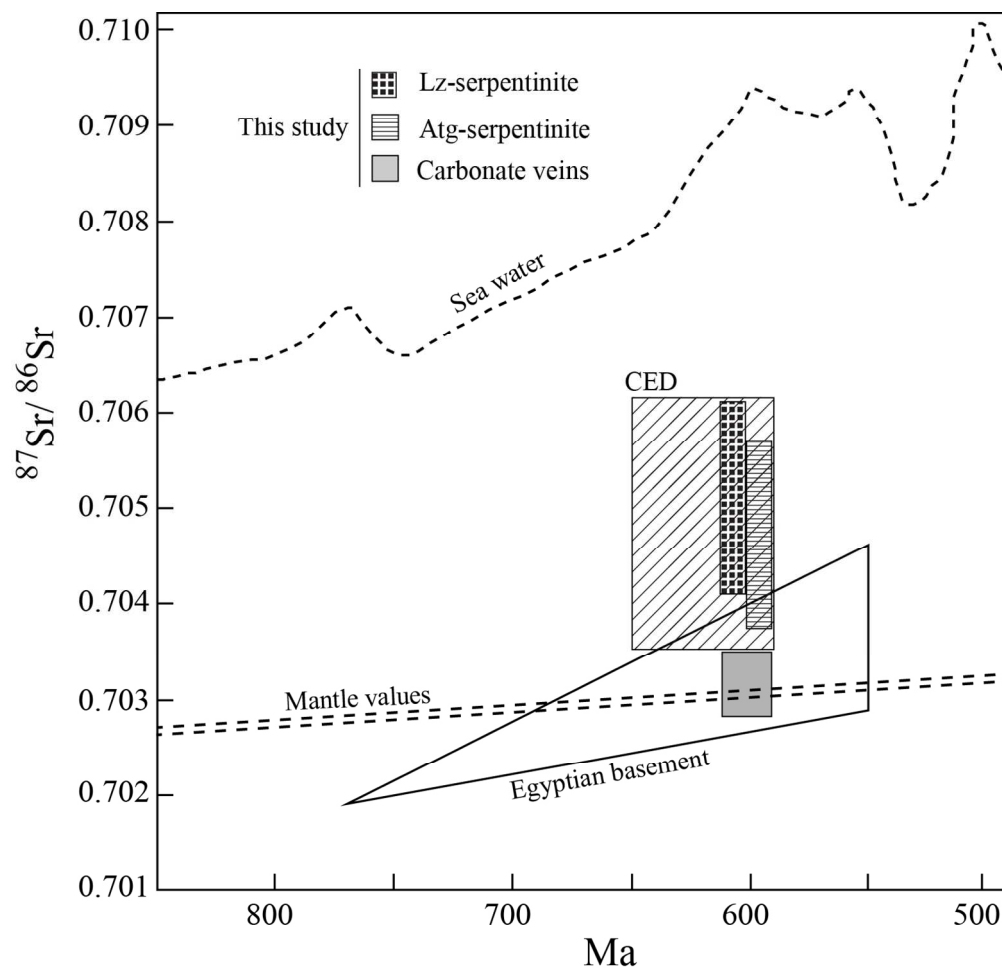


Figure 13. Sr composition of carbonate veins and altered serpentinites in the Meatiq study area compared to values from Neoproterozoic intrusive carbonates from Eastern Desert of Egypt. The ages of emplacement for the intrusive carbonates are estimated from the age of the surrounding basement units (after Cavanagh 1979; Stern and Hedge 1985; Reischmann 1986; Stern and Gwinn 1990). Field of initial $^{87}\text{Sr}/^{86}\text{Sr}$ for Egyptian basement igneous rocks is shown after Stern and Hedge (1985); this field represents an upper limit for the isotopic composition of Sr in the depleted mantle beneath northeast Africa at that time. Fields of initial $^{87}\text{Sr}/^{86}\text{Sr}$ for CED carbonate veins is shown after Stern and Gwinn (1990). Isotopic composition of Neoproterozoic seawater is also shown after Halverson et al. (2007).

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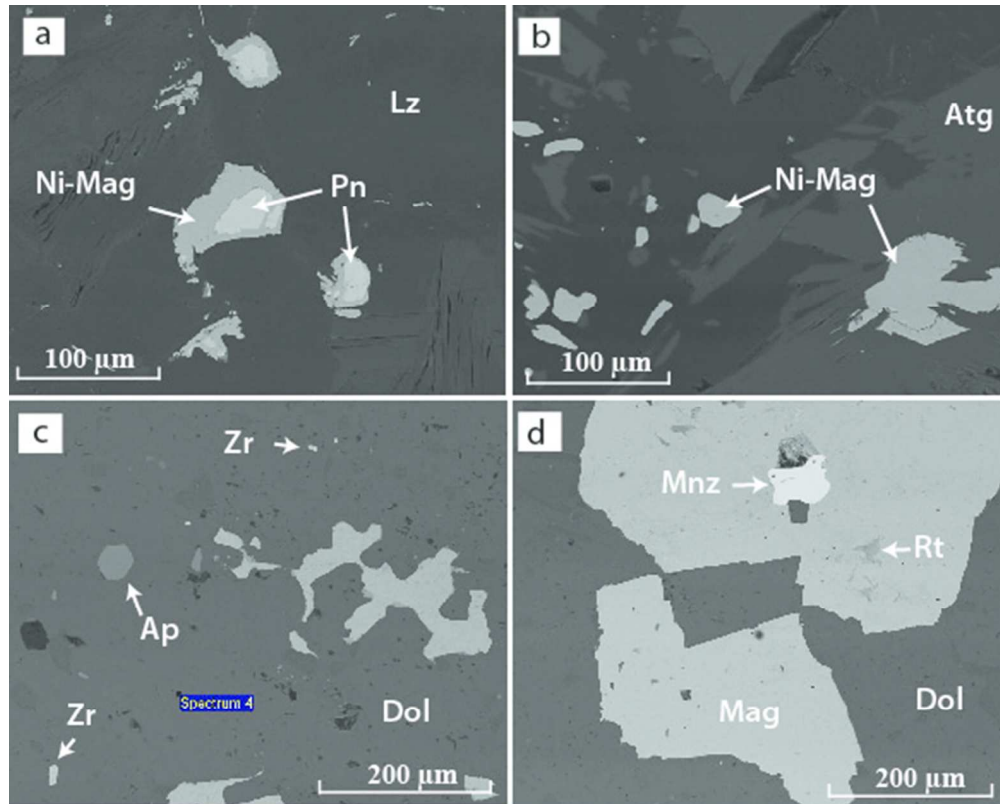


Figure 14. Back-scattered electron (BSE) images showing different mineral phases in serpentinites and carbonate veins. (a) Ni-sulphide (i.e. pentlandite-Pn) in Lz-serpentinites converted to Ni-rich magnetite (Ni-Mag) in the rims (Me23). (b) Total conversion of pentlandite to Ni-rich magnetite in Atg-serpentinites (Me7). Figures (c) and (d) show REE-bearing minerals (Zircon-Zr, apatite-Ap, Monazite-Mnz and rutile-Rt) in a carbonate vein (Me12).

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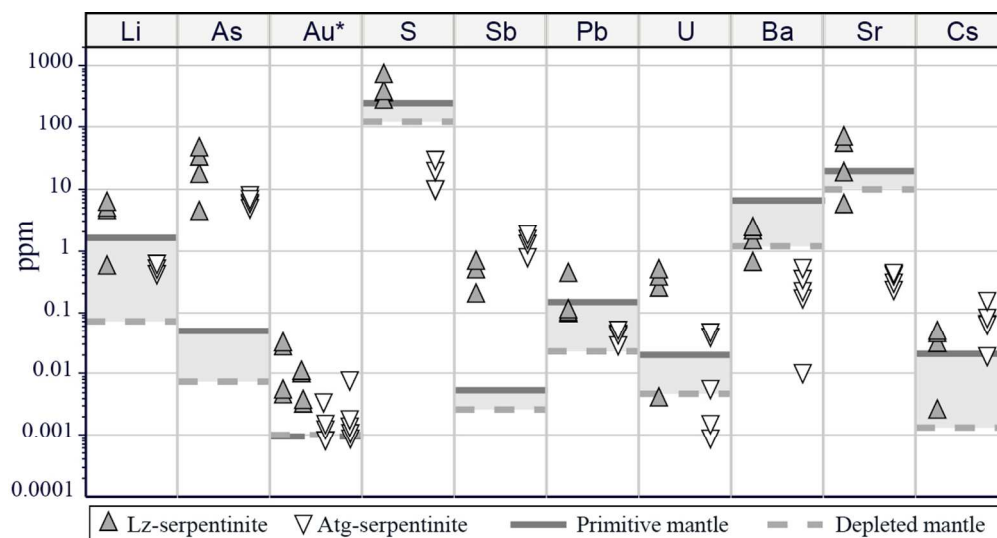


Figure 15. Plots of concentrations of fluid-mobile elements (Li, As, S, Sb, Pb, U, Ba, Sr and Cs) and Au in two serpentinite phases from the Meatiq study area. Values of primitive mantle after McDonough and Sun (1995) and depleted mantle after Salters and Stracke (2004) are also reported (thick and dashed black lines). Grey boxes represent the ranges between primitive and depleted mantle analyses.

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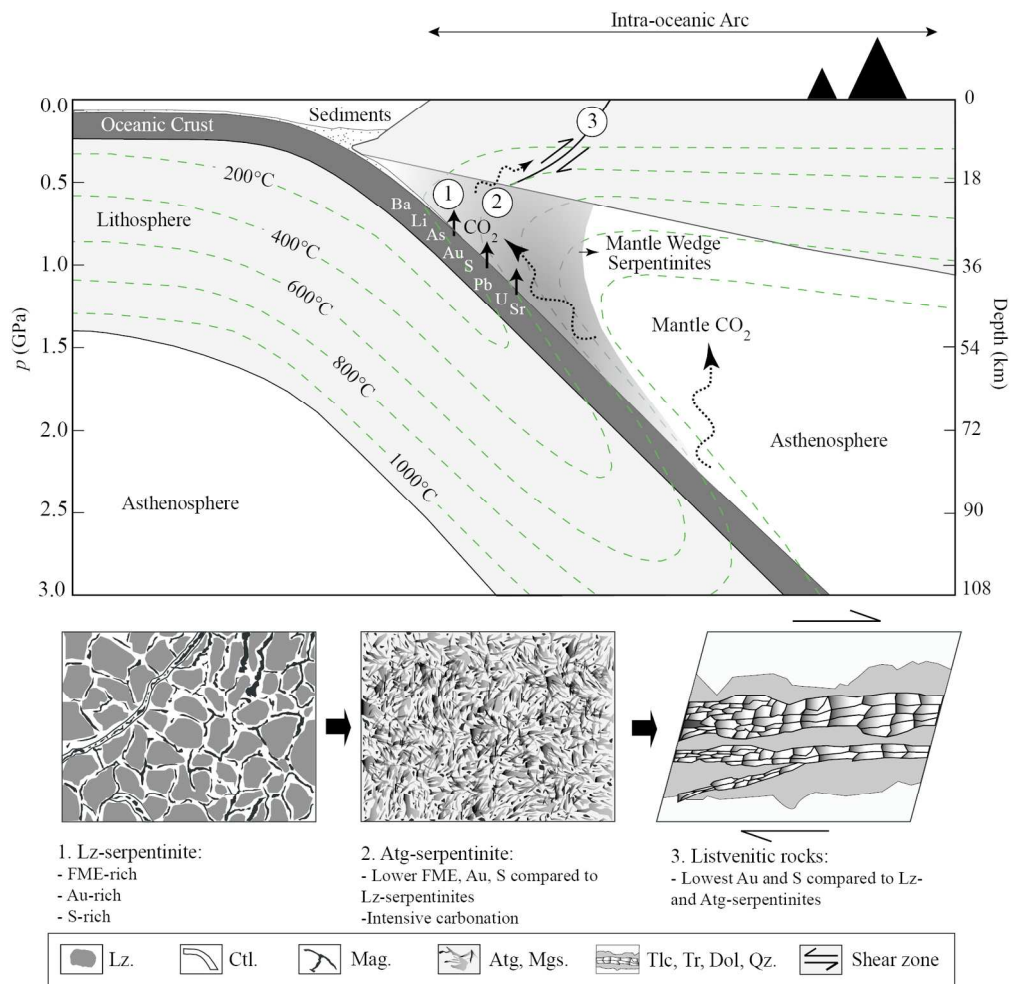


Figure 16. Schematic cross section illustrating the characteristics and evolution of the different serpentinite and listvenitic rocks and the tectonic setting where, based on the alteration assemblages and fluid inclusion compositions, these rocks formed. The principal transfer of fluid-mobile elements is also noted, as well as geochemical signature of fluids released during dehydration of serpentinites, "antigorite formation" due to the prograde metamorphism and later "listvenitization". Mineral abbreviations as in Fig. 5.

160x156mm (300 x 300 DPI)



Supplementary Table Representative electron microprobe analyses of Cr-spinel cores in Meatiq Lz-serpentinites

| Sample No. | Me21 | | | | | | | | | | Me23 | | | | | | | | | | | | | |
|--|--------|--------|--------|--------|--------|--------|--------|-------|-------|--------|--------|--------|--------|--------|--------|--------|-------|--------|--------|--------|--------|--------|--------|--------|
| | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 | 11 | 12 | 13 | 14 |
| TiO ₂ | n.d. | n.d. | 0.013 | 0.012 | n.d. | n.d. | n.d. | 0.010 | 0.022 | 0.015 | 0.008 | 0.018 | n.d. | n.d. | n.d. | n.d. | n.d. | 0.024 | n.d. | n.d. | n.d. | n.d. | n.d. | |
| Al ₂ O ₃ | 20.08 | 21.10 | 20.55 | 19.83 | 19.86 | 20.68 | 21.78 | 19.94 | 19.76 | 19.77 | 21.87 | 21.20 | 21.04 | 21.55 | 21.88 | 21.49 | 21.74 | 23.03 | 21.11 | 21.20 | 21.17 | 21.24 | 21.38 | 20.78 |
| Cr ₂ O ₃ | 50.78 | 47.89 | 49.63 | 49.91 | 50.12 | 49.81 | 48.74 | 49.50 | 50.39 | 50.95 | 47.59 | 48.21 | 48.63 | 49.29 | 48.85 | 49.09 | 47.91 | 47.34 | 47.57 | 48.81 | 48.59 | 48.93 | 48.79 | 48.98 |
| Fe ₂ O ₃ | 0.387 | 1.203 | 0.703 | 0.978 | 0.804 | 0.715 | 0.677 | 0.794 | 0.860 | 0.548 | 0.884 | 0.799 | 0.895 | 0.264 | 0.488 | 0.616 | 0.695 | 1.052 | 1.575 | 0.956 | 0.737 | 0.957 | 0.556 | 0.488 |
| FeO | 17.24 | 19.29 | 19.35 | 18.84 | 18.00 | 18.38 | 17.76 | 19.23 | 16.81 | 18.28 | 19.78 | 20.70 | 18.48 | 17.75 | 18.48 | 17.74 | 18.29 | 17.13 | 19.05 | 17.47 | 18.58 | 17.30 | 18.63 | 20.53 |
| MnO | 0.178 | 0.195 | 0.219 | 0.095 | 0.160 | 0.284 | 0.172 | 0.120 | 0.074 | 0.189 | 0.200 | 0.266 | 0.130 | 0.113 | 0.082 | 0.170 | 0.156 | 0.117 | 0.147 | 0.105 | 0.213 | 0.212 | 0.150 | 0.261 |
| NiO | 0.052 | 0.098 | 0.031 | n.d. | 0.033 | 0.117 | 0.081 | 0.085 | 0.098 | 0.035 | 0.025 | n.d. | 0.065 | 0.042 | 0.023 | 0.040 | 0.040 | 0.083 | 0.110 | 0.102 | 0.031 | 0.127 | 0.062 | n.d. |
| MgO | 11.70 | 10.36 | 10.48 | 10.71 | 11.14 | 11.03 | 11.60 | 10.32 | 11.89 | 11.10 | 10.27 | 9.620 | 10.96 | 11.61 | 11.30 | 11.60 | 11.10 | 12.20 | 10.55 | 11.65 | 10.88 | 11.72 | 10.96 | 9.66 |
| Total | 100.42 | 100.13 | 100.97 | 100.38 | 100.11 | 101.01 | 100.81 | 99.99 | 99.90 | 100.88 | 100.63 | 100.81 | 100.20 | 100.62 | 101.10 | 100.74 | 99.94 | 100.96 | 100.14 | 100.29 | 100.20 | 100.48 | 100.53 | 100.70 |
| No. of cations on the basis of four oxygen atoms | | | | | | | | | | | | | | | | | | | | | | | | |
| Ti | n.d. | n.d. | 0.000 | 0.000 | n.d. | n.d. | n.d. | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | n.d. | n.d. | n.d. | n.d. | n.d. | 0.001 | n.d. | n.d. | n.d. | n.d. | n.d. | n.d. |
| Al | 0.738 | 0.782 | 0.757 | 0.735 | 0.736 | 0.758 | 0.793 | 0.743 | 0.730 | 0.728 | 0.804 | 0.784 | 0.776 | 0.787 | 0.796 | 0.784 | 0.800 | 0.831 | 0.781 | 0.777 | 0.781 | 0.777 | 0.785 | 0.770 |
| Cr | 1.253 | 1.190 | 1.226 | 1.241 | 1.245 | 1.225 | 1.191 | 1.238 | 1.249 | 1.258 | 1.174 | 1.196 | 1.203 | 1.207 | 1.192 | 1.202 | 1.183 | 1.145 | 1.181 | 1.200 | 1.202 | 1.201 | 1.202 | 1.218 |
| Fe ³⁺ | 0.009 | 0.028 | 0.017 | 0.023 | 0.019 | 0.017 | 0.016 | 0.019 | 0.020 | 0.013 | 0.021 | 0.019 | 0.021 | 0.006 | 0.011 | 0.014 | 0.016 | 0.024 | 0.037 | 0.022 | 0.017 | 0.022 | 0.013 | 0.012 |
| Fe ²⁺ | 0.450 | 0.507 | 0.506 | 0.496 | 0.473 | 0.478 | 0.459 | 0.508 | 0.441 | 0.478 | 0.516 | 0.543 | 0.484 | 0.460 | 0.477 | 0.459 | 0.478 | 0.438 | 0.500 | 0.454 | 0.486 | 0.449 | 0.485 | 0.540 |
| Mn | 0.005 | 0.005 | 0.006 | 0.003 | 0.004 | 0.007 | 0.005 | 0.003 | 0.002 | 0.005 | 0.005 | 0.007 | 0.003 | 0.003 | 0.002 | 0.004 | 0.004 | 0.003 | 0.004 | 0.003 | 0.006 | 0.006 | 0.004 | 0.007 |
| Ni | 0.001 | 0.002 | 0.001 | n.d. | 0.001 | 0.003 | 0.002 | 0.002 | 0.002 | 0.001 | 0.001 | n.d. | 0.002 | 0.001 | 0.001 | 0.001 | 0.001 | 0.002 | 0.003 | 0.003 | 0.001 | 0.003 | 0.002 | n.d. |
| Mg | 0.544 | 0.485 | 0.488 | 0.502 | 0.522 | 0.512 | 0.534 | 0.486 | 0.556 | 0.517 | 0.478 | 0.450 | 0.511 | 0.536 | 0.520 | 0.535 | 0.517 | 0.557 | 0.494 | 0.540 | 0.507 | 0.542 | 0.509 | 0.453 |
| Mg# | 0.547 | 0.489 | 0.491 | 0.503 | 0.525 | 0.517 | 0.538 | 0.489 | 0.558 | 0.520 | 0.481 | 0.453 | 0.514 | 0.538 | 0.522 | 0.538 | 0.520 | 0.559 | 0.497 | 0.543 | 0.511 | 0.547 | 0.512 | 0.456 |
| Cr# | 0.626 | 0.595 | 0.613 | 0.621 | 0.623 | 0.613 | 0.595 | 0.619 | 0.625 | 0.629 | 0.587 | 0.598 | 0.602 | 0.604 | 0.596 | 0.601 | 0.592 | 0.573 | 0.591 | 0.600 | 0.601 | 0.600 | 0.601 | 0.609 |
| Fe ³⁺ # | 0.005 | 0.014 | 0.008 | 0.012 | 0.010 | 0.008 | 0.008 | 0.009 | 0.010 | 0.006 | 0.010 | 0.009 | 0.011 | 0.003 | 0.006 | 0.007 | 0.008 | 0.012 | 0.019 | 0.011 | 0.009 | 0.011 | 0.007 | 0.006 |

Mg# = Mg/(Mg+Fe²⁺), Cr# = Cr/(Cr+Al), Fe³⁺# = Fe³⁺/(Fe³⁺+Cr+Al), n.d. (not detected).