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Using Lunar Apatite to Assess the Volatile Inventory of the Lunar Interior

N.J.Potts^{1,2}, M.Anand¹, W.Van Westrenen², R.Tartèse¹, I.A.Franchi¹

1. Department of Physical Sciences, The Open University, Walton Hall, Milton Keynes, MK7 6AA
2. Faculty of Earth and Life Sciences, Vrije University, De Boelelaan, Amsterdam.

Lunar petrology, most notably the absence of hydrous minerals (such as micas and amphiboles) and the lack of Fe₂O₃, imply a low oxygen activity for the Moon [1]. The anhydrous nature of the Moon is consistent with observed depletions in volatile elements compared to the Earth [2]. Recent analytical developments have led to the re-investigation of lunar samples. In volcanic products, heterogeneous water contents in volcanic glass beads [3] olivine-hosted melt inclusions [4] and in the accessory phase apatite [5] [6] [7] [8] [9] indicate a wetter lunar interior than previously thought. Analysis of lunar apatite has produced OH contents as high as ~12000 ppm [5] and volatile contents of olivine-hosted melt inclusions appear to be similar to terrestrial mid-ocean ridge basalts values [4]. However, analysis of Cl isotope compositions from a range of lunar rocks (basalts, glasses, apatite grains) identified a Cl fractionation 25 times larger than on Earth [10]. This has been interpreted as reflecting a relatively dry lunar interior. The coupled nature of Cl and H, together with this high fractionation of Cl has been used to suggest the Moon's mantle has H values as low as 10 ppb [10].

To calculate the volatile contents of lunar melts, the partitioning behaviour of volatiles into apatite must be considered. Very little work has been done on the partition of volatiles under lunar conditions, however to fully constrain the H content of the magmatic source regions based on apatite grain measurements, determination of accurate partition coefficients is required.

Experimental work using a piston-cylinder assembly at VU, University Amsterdam, is being carried out to derive these partition coefficients for volatiles (F, OH, Cl) between apatite and melts. Measurements of the volatile contents in experimental synthesised apatites are being carried out using a Cameca NanoSIMS 50L ion probe at the Open University. Primary experiments have looked at the temperature effect of F partitioning into apatite. This experimental work will be combined with measurements of Cl, F, and OH concentrations as well as Cl and H isotope compositions in mare basalts. This will provide better constraints on the volatile budget of the lunar magmatic source regions.

[1] Albee A. L. (2003) Lunar Rocks. Encyclopedia of Physical Science and Technology (Third Edition), 825–837. [2] Humayun M. and Clayton R. N. (1995) Potassium isotope cosmochemistry: Genetic implications of volatile element depletion. *Geochimica et Cosmochimica Acta* 59, 2131–2148. [3] Saal A. E., Hauri E. H., Cascio M. L., Van Orman J. a., Rutherford M. C. and Cooper R. F. (2008) Volatile content of lunar volcanic glasses and the presence of water in the Moon's interior. *Nature* 454, 192–5. [4] Hauri E. H., Weinreich T., Saal A. E., Rutherford M. C. and Van Orman J. a (2011) High pre-eruptive water contents preserved in lunar melt inclusions. *Science* (New York, N.Y.) 333, 213–5. [5] Greenwood J. P., Itoh S., Sakamoto N., Warren P., Taylor L. and Yurimoto H. (2011) Hydrogen isotope ratios in lunar rocks indicate delivery of cometary water to the Moon. *Nature Geoscience* 4, 79–82. [6] Barnes J. J., Franchi I. a., Anand M., Tartèse R., Starkey N. a., Koike M., Sano Y. and Russell S. S. (2013) Accurate and precise measurements of the D/H ratio and hydroxyl content in lunar apatites using NanoSIMS. *Chemical Geology* 337–338, 48–55 [7] McCubbin F. M., Jolliff B. L., Nekvasil H., Carpenter P. K., Zeigler R. a., Steele A., Elardo S. M. and Lindsley D. H. (2011) Fluorine and chlorine abundances in lunar apatite: Implications for heterogeneous distributions of magmatic volatiles in the lunar interior. *Geochimica et Cosmochimica Acta* 75, 5073–5093. [8] McCubbin F. M., Steele A., Hauri E. H., Nekvasil H., Yamashita S. and Hemley R. J. (2010) Nominally hydrous magmatism on the Moon. *Proceedings of the National Academy of Sciences of the United States of America* 107, 11223–8. [9] Boyce J. W., Liu Y., Rossman G. R., Guan Y., Eiler J. M., Stolper E. M. and Taylor L. a (2010) Lunar apatite with terrestrial volatile abundances. *Nature* 466, 466–9. [10] Sharp Z D, Shearer C. K., McKeegan K. D., Barnes J D and Wang Y. Q. (2010) The chlorine isotope composition of the moon and implications for an anhydrous mantle. *Science (New York, N.Y.)* 329, 1050–3.