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Petrogenesis and geochronology of the Arkasani Granophyre and felsic Dalma volcanic rocks: implications for the evolution of the Proterozoic North Singhbhum Mobile Belt, east India

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> Abstract – The North Singhbhum Mobile Belt (NSMB) is a 200 km long, curved Proterozoic foldthrust belt that skirts the northern margin of the Archean Singhbhum Craton of NE India. The Singhbhum Shear Zone (SSZ) developed between the Dhanjori and Chaibasa formations near the southern margin of the NSMB and represents an important Cu-U-P metallotect. A SHRIMP U–Pb zircon date of 1861 ± 6 Ma, obtained for the syn- to post-kinematic Arkasani Granophyre that has intruded the SSZ, provides a minimum age for the prolonged tectonic activity and mineralization along the SSZ and for the time of closure of the Chaibasa and Dhanjori sub-basins. The Dalma Volcanic Belt, a submarine rift-related bimodal mafic-felsic volcanic suite, forms the spine of the NSMB. A SHRIMP U–Pb zircon igneous crystallization date of 1631 ± 6 Ma was obtained for an unfoliated felsic volcanic rock from the base of the Dalma volcanic sequence. These new findings suggest that the different sub-basins in the NSMB evolved diachronously under contrasting tectonic environments and were juxtaposed during a later orogenic movement.

> Keywords: Arkasani Granophyre, Dalma volcanism, North Singhbhum Mobile Belt, Singbhum Shear Zone.

1. Introduction

The eastern segment of the Indian Precambrian Shield, commonly referred to as the Singhbhum Protocontinent, preserves a record of geological events spanning early Archaean - late Proterozoic time (Saha, 1994; Sarkar & Gupta, 2012). The early Archaean Singhbhum Craton (Sarkar & Saha, 1983) comprises tonalite-trondhjemite-granodiorite and middle-late Archaean supracrustal volcano-sedimentary greenstone belt successions (IOG) that have been intruded by late Archaean granites. A Proterozoic thrust-fold belt, the North Singhbhum Mobile Belt (cf. Gupta, Basu & Ghosh, 1980; Mukhopadhyay, 1984; Sarkar, Gupta & Basu, 1992; Fig. 1), occurs between the Archean Singhbhum Craton in the south and the Chhotanagpur Granite Gneissic Complex in the north. The North Singhbhum Mobile Belt (NSMB) is a nearly 200 km long and 50 km wide curvilinear orogenic belt (Fig. 1). In the east, the belt has a general north-south trend, which gradually changes to east-west in the central part and NE-SW in the west (Fig. 1). The NSMB is regarded as a Proterozoic rift basin (Bhattacharya & Mahapatra, 2008) that developed along a major suture zone (Sarkar & Saha, 1983). Rifting produced Dhanjori, Chaibasa, Dalma and North Dalma sub-basins with distinctly different petrotectonic evolutionary histories, juxtaposed as linear domains during a terminal orogeny (Mukhopadhyay, 1984; Sarkar, Gupta & Basu, 1992). The contacts between such linear domains are sheared with north-dipping shear planes. A simplified stratigraphy of the terrane is provided in Table 1.

The southern margin of the NSMB is marked by a curved zone of intense shearing, 1-10 km wide and over 200 km long, known as the Singhbhum Shear Zone (SSZ) (Fig. 1) which separates the Chaibasa Formation of the Singhbhum Group and the Dhanjori Group (Saha, 1994). The SSZ becomes bifurcated in the west, with the Chakradharpur Granite (CKPG) coming in between the two branches (Fig. 1). The shear planes in the shear zone dip at 45° or less towards the north. Several granite lenses were emplaced along the shear zone (Dunn & Dey, 1942). The shear zone was active during a prolonged period and is characterized by repeated cycles of mylonitization, folding and down-dip rotation/stretching of pre-existing structures, indicating ductile shearing (Sengupta & Ghose, 1997). Truncation of stratigraphic units in the zone lying between Kudada and Saraikela indicates thrusting along the Singhbhum Shear Zone (Bhattacharya & Mahapatra, 2008). Metamorphism close to the shear zone further suggests crustal thickening due to thrusting (Mahato et al. 2008). The SSZ therefore records both ductile shearing and thrusting during the closure of the Dhanjori and Chaibasa sub-basins.

Five different tectonic models have been proposed to explain the evolution of the NSMB. Three of these models are based on principles of plate tectonic (Sarkar &

1.6–0.7 Ga	Younger granitoids	
3 1 Ga	Proterozoic fold-belts and supracrustals Singhbhum Granite-B	Dhanjori Group. Singhbhum Group, Dalma volcanic rocks and Kolhan Group
	Iron Ore Group	Ultramafic to felsic volcanic rocks, conglomerate, quartzites, phyllites, dolomite, tuffs and banded iron formation

Table 1. Generalized stratigraphy of Singhbhum, east India.

3.3 GaSinghbhum Granite-A3.4 GaOlder Metamorphic Tonalitic Gneiss

> 3.4 Ga Older Metamorphic Group

Foliated granitoids Quartzites, mica schists, calc-silicates and amphibolites Basement not identified



Figure 1. Geological map of North Singhbhum Mobile Belt (NSMB), showing the occurrences of Arkasani Granophyre, Dalma volcanic rocks and Singhbhum Shear Zone. Sample locations for geochronological analyses and other important stratigraphic units are also shown in the map. The location of the study area is shown in the map of India.

Saha, 1977; Bose & Chakraborti, 1981; Sarkar, 1982; Bose, Chakrabarti & Saunders, 1989), drawing analogues from Phanerozoic crustal segments. The fourth model proposes intracratonic extension and ensialic orogenesis with or without reference to an early stage of a Wilson Cycle (Gupta, Basu & Ghosh, 1980; Mukhopadhyay, 1984; Sarkar, Gupta & Basu, 1992; Gupta & Basu, 2000; Sarkar, 2000). The fifth model (Bhattacharya & Mahapatra, 2008)emphasizes sedimentolo-gical aspects of the sub-basins and proposes the form-ation of younger basins along the margins of the main rift and their collapse during a terminal orogeny. As a result of this south-directed terminal orogenic move-ment, the Chhotonagpur Granite Gneissic Complex in the north overrode the NSMB along a thrust, and the NSMB was driven onto the Archean cratonic block lying to the south.

The Dhanjori and Chaibasa sub-basins of the NSMB may have developed during reactivation of regional curved faults that were formed during doming of the Archean craton as a result of the emplacement of granite sheets (Roy & Bhattacharya, 2012).

The findings of the present study indicate that the NSMB does not form a single regional rift basin, as was conceived in proposed models, but instead consists of separate basins that developed over a considerable time period. The present paper integrates a stratigraphic, sedimentologic, structural and tectonic database with precise SHRIMP U–Pb zircon dates to present a new and comprehensive model for the evolution of the NSMB.

2. Arkasani Granophyre

At least seven isolated lens-shaped granitic bodies, known as the Arkasani Granophyre, were emplaced well within the mylonitized metasediments of the NSMB, forming the northern branch of the forked Singhbhum Shear Zone in the north of CKPG-Gneissic Complex (Fig. 1). The largest body forms the Arkasani Hill. These granitic bodies were first mapped by Dunn (1929) and later studied by Banerjee, Bhattacharya & Chattaopadhyay (1978), Sengupta et al. (1984) and Chattopadhyay (1990). The cores of these granite bodies are generally massive and porphyritic (Fig. 2a), but they are schistose near their peripheries (Fig. 2b). Schistosity is more pervasive in the smaller bodies. Near their margins, these granitic bodies commonly contain enclaves of mica-schist and schistose amphibolites. Sengupta et al. (1984) reported a Rb-Sr whole-rock isochron date of 1.0 Ga, interpreted as the time of a post-emplacement thermal event, for the Arkasani Granophyre.



Figure 2. (Colour online) Field photographs of (a) porphyritic variety and (b) schistose variety of Arkasani Granophyre. Schistosity is defined by biotite/muscovite flakes and flattened and recrystallized quartz grains forming quartz ribbons (QR).

The Arkasani Granophyre is a medium- to coarsegrained porphyritic rock with phenocrysts of plagioclase (An₂₂-An₃₅) commonly occurring as clusters or cumulates. The groundmass contains sodic plagioclase (An_2-An_{15}), K-feldspar, grevish quartz, biotite and muscovite as the major minerals, with accessory tourmaline, epidote, zircon, apatite, titanite and magnetite. The grain size is commonly reduced towards the body margins. The groundmass shows intergrowth texture, mostly granophyric with myrmekitic and micropegmatitic variations. The proportion of granophyric intergrowth in these granite bodies commonly increases towards their margins. Along the margins the granite bodies are schistose, with the schistosity defined by flattened and recrystallized quartz grains with wavy extinction forming quartz ribbons, and parallel-aligned biotite and muscovite flakes (Fig. 2b).

Major, trace and rare Earth element (REE) chemical compositions of the Arkasani Granophyre, obtained at the Wadia Institute of Himalayan Geology by inductively coupled plasma mass spectrometry (ICP-MS; Khanna et al. 2009), are presented in Table 2. Bulk chemical compositions of granitic rocks that have intruded the Older Metamorphic Group at Champua, Garumahisani Granite and Mayurbhanj Granite are also listed in Table 2 for comparison. The SiO₂ content of the Arkasani Granophyre lies within the range 71.79–73.36 wt %, K₂O exceeds Na₂O and the Ba and U contents are exceptionally high. Plots in the SiO₂ v. $Na_2O + K_2O$ classification diagram (after Cox, Bell & Pankhurst, 1979) show the granitic character of the Arkasani Granophyre (Fig. 3a), whereas plots in the A/ NK (molecular proportion of $Al_2O_3/Na_2O + K_2O$) v. A/CNK (molecular proportion of $Al_2O_3/CaO + Na_2O$ + K₂O) diagram (after Maniar & Piccoli, 1989) reveal its metaluminous character (Fig. 3b). Plots of chondrite-normalized trace elements show prominent negative Sr, P, Ti and V anomalies (Fig. 4a).

Chondrite-normalized REE plots (Fig. 4b) show an unfraction-ated pattern ($[La/Yb]_{uc} = 2.04-2.42$) with slight en-richment in the light REEs (LREEs) relative to heavy REEs (HREEs), and a prominent negative Eu anomaly (Eu/Eu^{*} = 0.59-0.67).

The porphyritic texture of the Arkasani Granophyre reflects early growth of feldspar crystals during slight cooling of the magma, followed by finer grains resulting from more rapid cooling due to shallow emplacement. The granophyric texture indicates rapid and simultaneous crystallization of quartz and K-feldspar from an under-cooled magma at shallow depth (Clarke, 1992). Flattened quartz grains with wavy extinction and parallel-aligned biotite and muscovite flakes defining schistosity are likely responses to synto post-kinematic emplacement (see also Sarkar & Gupta, 2012). Apatite, zircon and magnetite grains are euhed-ral, consistent with their primary nature (Clarke, 1992). These minerals form the majority of inclusions in other minerals.

The relatively constant K_2O/Na_2O and $K_2O > Na_2O$ of the Arkasani Granophyre indicate differentiation of magma generated from a single source (Jung et al. 1999), whereas Ti/P ratios near 2 and the negative Nb anomaly suggest a shallow crustal source. Rb/Sr (2.48-2.61) and low Sr/Ba (0.05-0.07) ratios are the result of early crystallization of plagioclase from the melt dur-ing fractional crystallization. The negative Sr and Eu anomalies (Fig. 4a, b) are consistent with early plagio-clase fractionation (Rollinson, 1993). Prominent₂Zr, P₂O₅, T i O and V anomalies possibly reflect crystal-lization/removal of accessory minerals such as rutile, apatite, zircon and magnetite from the melt.

High K contents, low Ti/P ratios, the negative Nb anomaly and unfractionated HREE (and Y) lend strong support in favour of the melting of a metaluminous shallow crustal source for the generation of Arkasani Granophyre (see also Sengupta *et al.* 1984). Rb

Table 2. Major element, trace and rare Earth element composition of Arkasani Gra	ranophyre and other granites o	of the Singhbhum region.
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$\begin{array}{c c c c c c c c c c c c c c c c c c c $		A1	A2	A3	D1	D2	D3	CSG	GG	MG
SiO ₂ matrix of the set of the	Maior element (wt %)									
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	SiO	73.36	72.55	71.79	74.89	74.81	75.36	73.01	71.56	77.18
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	TiO ₂	0.30	0.30	0.31	0.27	0.3	0.26	0.10	0.21	0.30
$ Fe_{0}^{-1} , 3.56 3.49 3.56 3.06 3.29 3.04 0.94 2.21 3.88 Mon 0 0 0.01 0.01 0.06 0.06 0.06 0.00 0.01 0.04 0.05 MgO 0.37 0.33 0.39 0.34 0.34 0.37 0.11 0.41 0.07 CaO 0.93 0.85 0.98 0.40 0.37 0.4 1.04 1.83 0.80 KgO 2.29 2.74 3.13 3.78 3.70 3.79 3.92 4.91 3.54 KgO 2.429 2.74 4.36 3.96 3.87 3.8 5.01 3.23 3.89 SUM 99.01 1.5 0.14 0.15 0.6 0.6 0.7 0.03 0.08 0.03 SUM 99.01 97.80 97.71 99.56 99.98 100.18 98.07 99.06 100.95 Loss on ignition (LOI) 1.08 1.11 1.08 0.52 0.3 0.55 0.67 0.89 0.64 Na_0VK_SO 0.68 0.64 0.72 0.95 0.96 1.00 0.78 0.57 0.89 0.64 Na_0VK_SO 0.68 0.64 0.72 0.95 0.96 1.00 0.78 1.52 0.91 Trace element (ppm) Trace element (ppm) 100 1028 1175 1158 1154 1270 307 388 1823 Cr 2 9 10 10 9 9 8 8 6 27 <5.0 Co 150 135 145 62 58 67 104 129 78.0 310 82 Ni 1 17 15 18 4 4 4 3 <5.0 <5.0 <5.0 <5.0 <5.0 Co 2.50 <5.0 <5.0 <5.0 Co 2.50 <5.5 5.5 6 6 .5 6.8 5.9 <5.0 <5.0 <5.0 <5.0 Co 2.50 <5.5 5.5 5.5 6 6 .5 6.8 5.9 <5.0 <5.0 <5.0 <5.0 Co 2.50 <5.5 5.5 5.5 6 6 .5 6.8 5.9 <5.0 <5.0 <5.0 <5.0 Co 2.50 <5.5 5.5 5.5 6 6 .5 6.8 5.9 <5.0 <5.0 <5.0 <5.0 Co 2.50 <5.5 5.5 5.5 6 6 .5 6.8 5.9 <5.0 <5.0 <5.0 <5.0 Co 2.50 Sr 2.50 <5.5 5.5 5.5 6 6 .5 6.8 5.9 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0$	Al ₂ O ₃	13.52	13.11	13.03	12.20	12.34	11.9	13.90	14.58	11.21
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Fe ₂ O ₂	3.56	3.49	3.56	3.06	3.29	3.04	0.94	2.21	3.88
	MnO	0.01	0.01	0.01	0.06	0.06	0.06	0.01	0.04	0.05
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	MgO	0.37	0.33	0.39	0.34	0.34	0.37	0.11	0.41	0.07
	CaO	0.93	0.85	0.98	0.40	0.37	0.4	1.04	1.83	0.80
	Na ₂ O	2.99	2.74	3.13	3.78	3.70	3.79	3.92	4.91	3.54
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	K ₂ O	4.42	4.27	4.36	3.96	3.87	3.8	5.01	3.23	3.89
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	P_2O_5	0.15	0.14	0.15	0.6	0.6	0.7	0.03	0.08	0.03
Loss on ignition (LOI) 1.08 1.11 1.08 0.52 0.3 0.5 0.67 0.89 0.64 Na2OK2O 0.68 0.64 0.72 0.95 0.96 1.00 0.78 1.52 0.91 Tace element (ppm) Tace element (ppm) Tace element (ppm) Tace element (ppm) 1.10 1.08 1.17 1158 1154 1270 307 388 1823 Cr 28 25 29 14 14 16 8 12 19 V 9 10 10 9 8 6 27 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <td>SUM</td> <td>99.61</td> <td>97.80</td> <td>97.71</td> <td>99.56</td> <td>99.98</td> <td>100.18</td> <td>98.07</td> <td>99.06</td> <td>100.95</td>	SUM	99.61	97.80	97.71	99.56	99.98	100.18	98.07	99.06	100.95
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Loss on ignition (LOI)	1.08	1 11	1.08	0.52	0.3	0.5	0.67	0.89	0.64
Trace element (ppm) Ba 1100 1028 1175 1158 1154 1270 307 388 1823 Cr 28 25 29 14 14 16 6 8 12 19 V 9 10 10 9 9 8 6 27 <5.0 Co 150 135 145 62 58 67 104 129 78 Ni 17 15 18 4 4 4 3 <5.0 <5.0 <5.0 <5.0 Cu 23 21 24 18 177 20 3 10 8 Zn 14 15 12 34 36 31 18 41 97 Ga 15 16 14 15 15 13 18 18 41 97 Ga 15 16 14 15 15 13 18 19 19 Pb 9 10 8 26 24 27 26 18 22 Th 28 29 26 32 35 32 5 9 10 Rb 169 180 165 114 105 113 124 94 73 U 5.5 5.5 6 6.5 6.8 5.9 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0 <5.0	Na ₂ O/K ₂ O	0.68	0.64	0.72	0.92	0.96	1.00	0.78	1.52	0.01
The constant (ppm) is a second secon	Trace element (nnm)	0.00	0.01	0.72	0.55	0.20	1.00	0.70	1.02	
LaLi <td>Ba</td> <td>1100</td> <td>1028</td> <td>1175</td> <td>1158</td> <td>1154</td> <td>1270</td> <td>307</td> <td>388</td> <td>1823</td>	Ba	1100	1028	1175	1158	1154	1270	307	388	1823
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Co150105105105105105105105105106107105107108Ni171518443<5.0	V	0	10	10	0	0	8	6	27	~5.0
Co 150 153 143 02 153 07 144 152 163 Ni 17 15 18 4 4 3 <5.0	Čo.	150	135	145	62	58	67	104	120	78
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Ni	17	155	145	02	1	3	<5.0	~5.0	-50
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Cu Zn	23	15	12	24	26	20	18	10	07
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$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	KD U	109	180	105	114	105	50	124	94	/3
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	U Sa	3.3	3.3 72	62	0.3	0.8	20	< 3.0	< 5.0	< 3.0
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Sr	08	26	03	40	43	39	244	303	82
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Nb2021201719176820Sc14.81413.613.213.612.90.11.80.8Sk/Sr2.492.502.622.852.442.900.510.260.89Sr/Ba0.060.070.050.030.040.030.790.940.04Rare Earth element (ppm)La878095666171173663Ce1681701871241301202858157Pr18161812.812.611.92.55.811.8Nd74.682.772.75351.654.29.321.149Sm15.816.515.61211.613.21.83.59.5Eu1.451.331.51.61.641.720.770.772.62Gd12.611.612.310.99.810.91.42.88.5Dy1110.511.511.910.912.30.91.57.4Ho2.342.692.42.822.652.880.170.31.64Er6.15.15.47.37.57.30.40.84.3Dy1110.511.511.910.91.020.050.090.58The2.34<		259	240	284	254	233	264	101	139	300
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Rb/Sr2.492.502.622.852.442.900.510.260.89Sr/Ba0.060.070.050.030.040.030.790.940.04Rare Earth element (ppm)La878095666171173663Ce1681701871241301202858157Pr18161812.812.611.92.55.811.8Nd74.682.772.75351.654.29.321.149Sm15.816.515.61211.613.21.83.59.5Eu1.451.331.51.61.641.720.770.772.62Gd12.611.612.310.99.810.91.42.88.5Tb2.011.912.021.921.761.930.190.341.32Dy1110.511.511.910.912.30.91.57.4Ho2.342.692.42.822.652.880.170.31.64Er6.15.15.47.37.57.30.40.84.3Tm0.80.80.71.010.991.020.050.090.58Yb5.15.45.46.46.276.720.30.664Lu <td>Sc DI /G</td> <td>14.8</td> <td>14</td> <td>13.6</td> <td>13.2</td> <td>13.6</td> <td>12.9</td> <td>0.1</td> <td>1.8</td> <td>0.8</td>	Sc DI /G	14.8	14	13.6	13.2	13.6	12.9	0.1	1.8	0.8
St/Ba 0.06 0.07 0.05 0.03 0.04 0.03 0.79 0.94 0.04 Rare Earth element (ppm)La 87 80 95 66 61 71 17 36 63 Ce 168 170 187 124 130 120 28 58 157 Pr 18 16 18 12.8 12.6 11.9 2.5 5.8 11.8 Nd 74.6 82.7 72.7 53 51.6 54.2 9.3 21.1 49 Sm 15.8 16.5 15.6 12 11.6 13.2 1.8 3.5 9.5 Eu 1.45 1.33 1.5 1.6 1.64 1.72 0.77 0.77 2.62 Gd 12.6 11.6 12.3 10.9 9.8 10.9 1.4 2.8 8.5 Tb 2.01 1.91 2.02 1.92 1.76 1.93 0.19 0.34 1.32 Dy 11 10.5 11.5 11.9 10.9 12.3 0.9 1.5 7.4 Ho 2.34 2.69 2.4 2.82 2.65 2.88 0.17 0.3 1.64 Er 6.1 5.1 5.4 7.3 7.5 7.3 0.4 0.8 4.3 Tm 0.8 0.8 0.7 1.01 0.99 1.02 0.05 0.09 0.58 Yb 5.1 5.4 5	Rb/Sr	2.49	2.50	2.62	2.85	2.44	2.90	0.51	0.26	0.89
Rare Earth element (ppm)La878095666171173663Ce1681701871241301202858157Pr18161812.812.611.92.55.811.8Nd74.682.772.75351.654.29.321.149Sm15.816.515.61211.613.21.83.59.5Eu1.451.331.51.61.641.720.770.772.62Gd12.611.612.310.99.810.91.42.88.5Tb2.011.912.021.921.761.930.190.341.32Dy1110.511.511.910.912.30.91.57.4Ho2.342.692.42.822.652.880.170.31.64Er6.15.15.47.37.57.30.40.84.3Tm0.80.80.71.010.991.020.050.090.58Yb5.15.45.46.46.276.720.30.64Lu0.740.810.780.960.90.970.050.090.64Eu/Eu*0.630.590.660.430.470.442.961.501.78 <td>Sr/Ba</td> <td>0.06</td> <td>0.07</td> <td>0.05</td> <td>0.03</td> <td>0.04</td> <td>0.03</td> <td>0.79</td> <td>0.94</td> <td>0.04</td>	Sr/Ba	0.06	0.07	0.05	0.03	0.04	0.03	0.79	0.94	0.04
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Rare Earth element (ppm))								
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	La	87	80	95	66	61	71	17	36	63
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Ce	168	170	187	124	130	120	28	58	157
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Pr	18	16	18	12.8	12.6	11.9	2.5	5.8	11.8
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Nd	74.6	82.7	72.7	53	51.6	54.2	9.3	21.1	49
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Sm	15.8	16.5	15.6	12	11.6	13.2	1.8	3.5	9.5
	Eu	1.45	1.33	1.5	1.6	1.64	1.72	0.77	0.77	2.62
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Gd	12.6	11.6	12.3	10.9	9.8	10.9	1.4	2.8	8.5
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Tb	2.01	1.91	2.02	1.92	1.76	1.93	0.19	0.34	1.32
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Dy	11	10.5	11.5	11.9	10.9	12.3	0.9	1.5	7.4
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Но	2.34	2.69	2.4	2.82	2.65	2.88	0.17	0.3	1.64
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Er	6.1	5.1	5.4	7.3	7.5	7.3	0.4	0.8	4.3
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Tm	0.8	0.8	0.7	1.01	0.99	1.02	0.05	0.09	0.58
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Yb	5.1	5.4	5.4	6.4	6.27	6.72	0.3	0.6	4
Eu/Eu* 0.63 0.59 0.66 0.43 0.47 0.44 2.96 1.50 1.78	Lu	0.74	0.81	0.78	0.96	0.9	0.97	0.05	0.09	0.64
	Eu/Eu*	0.63	0.59	0.66	0.43	0.47	0.44	2.96	1.50	1.78

A1-A3 - Arkasani Granophyre; D1-D3 - Dalma felsic volcanic rocks; CSG - Champua syn-tectonic granite; GG - Garumahisani granite; MG - Mayurbhanj granite.

v. Y + Nb plots (Fig. 5a) after Pearce, Harris & Tindale (1984) show that the Arkasani Granophyre has some geochemical similarities to recent volcanic arc granites with proximity to within-plate granite. The R_1 (4Si – 11(Na + K) – 2(Fe + Ti) v. R_2 (6Ca + 2Mg + Al) diagram (Fig. 5b) after Batchelor & Bowden (1985) shows that the Arkasani Granophyre samples plot in the syn-collisional to post-orogenic granite field.

In summary, field, petrographic and geochemical characteristics of the Arkasani Granophyre indicate that the parent magma was derived by low degrees of partial melting of upper crustal material and was emplaced along the Singhbhum Shear Zone during a late phase of an orogenic movement.

3. Dalma volcanic rocks

The Dalma volcanic suite and intercalated sediments constitute the Dalma Belt which forms an east–west range along the middle of the NSMB. The belt is composed of interlayered carbonaceous shales, cherts and felsic and mafic volcanic rocks that grade upwards to high-Mg komatiitic flows and pyroclastics. This volcano-sedimentary succession is followed upwards by pillowed low-K, high-Mg tholeiites. The succession is deformed into a synclinal structure, with the lower horizons exposed along the flanks of the east–west range of hills. The felsic vol-canic rocks exposed near the downstream side of the Chandil Dam on the Subarnarekha River (Fig. 6) w e r e



Figure 3. (a) $Na_2O + K_2O/SiO_2$ variation plots of Arkasani Granophyre (after Cox *et al.* 1979), showing the granitic character of the intrusion. (b) Alumina saturation (Shand index) diagram for the studied Arkasani Granophyre (modified after Maniar & Piccoli, 1989).



Figure 4. (a) Trace element spider diagram, normalized by chondrite values of Arkasani Granophyre, after Sun & McDonough (1989). (b) REE distribution patterns normalized by chondrite values of Arkasani Granophyre, after Boynton (1984).



Figure 5. (a) Geotectonic discrimination diagram after Pearce, Harris & Tindale (1984). (b) Geotectonic regimes discrimination diagram after Batchelor & Bowden (1985). $R_1 = 4Si - 11(Na + K) - 2(Fe + Ti)$; $R_2 = 6Ca + 2Mg + Al$. Symbols as for Figure 3.



Figure 6. (Colour online) Field photograph of felsic volcanic rocks of Dalma volcanic suite, southern bank of Subarnarekha River, near Chandil dam.



Figure 7. SiO₂ v. $Zr/TiO_2 \times 0.0001$ plots of felsic volcanics, Dalma volcanic suite, showing their rhyolitic character (after Winchester & Floyd 1977).

sampled for geochronological study (see the following section).

Major, trace and REE chemical compositions of the felsic volcanic rocks, obtained at the Wadia Institute of Himalayan Geology by X-ray fluorescence (XRF) and ICP-MS (Khanna et al. 2009), are listed in Table 2. The felsic volcanic rocks have high concentrations of SiO₂ and $K_2O + Na_2O$ and lower abundances of TiO₂, Al₂O₃, MgO, CaO and P₂O₅. The felsic volcanic rocks fall in the rhyolite field of SiO₂ v. $Zr/TiO_2 \times 0.0001$ classification diagram of Winchester & Floyd (1977) (Fig. 7). In general, the felsic volcanic rocks display the typical geochemical characteristic of A-type granitoids with high SiO₂, N a $_{2}O + K_{2}O$, Fe₂O₃, MgO, Ga/Al, Zr, Nb, Ga, Y and REE (except Eu), and low CaO and Sr (Eby, 1990; Bonin, 2007). A chondritenormalized trace element plot (Fig. 8a) shows enrichment in Ba, Th, U, Nb and Zr and depletion in Pb, Sr, Co and Ni. Chondrite-normalized REE patterns of the felsic vol-canic rocks (Fig. 8b) show large negative Eu anomalies (Eu/Eu * = 0.43–0.47) and fractionated and enriched LREE over HREE. Rb/Sr (2.44–2.90), low Sr/Ba, low

Sr and large Eu anomalies suggest early separation of plagioclase or retention of these elements in feldspar at the source during partial melting (Rollinson, 1993), whereas low contents of Zr, P_2O_5 , T i O₂ and V supports early crystallization and separation of rutile, apatite, zircon and magnetite from the melt. The felsic vol-canic rocks also plot in the within-plate tectonic setting field in Nb v.Y tectonic discrimination diagram (after Pearce, Harris & Tindale, 1984; Fig. 9).

The A-type granitoid chemistry of the felsic volcanic rocks lends support to a non-orogenic rift setting (Bar-barin, 1990). Further, their overall geochemical signa-tures favour their generation by low degrees of crustal melting in an extensional tectonic setting under the in-fluence of advecting mafic–ultramafic magma (also see Roy *et al.* 2002; Sarkar & Gupta, 2012).

4. Geochronology

Sample SC07/5 was taken from the Arkasani Granophyre for geochronology from a 1 m diameter boulder located c. 20 m up on the SE side of the forested slopes of Arkasani Hill, on the track to the Arkasani Temple (sampling site coordinates 22°46'11''N, 85° 51' 20'' E). Dark grev to black, unfoliated felsic volcanic rocks occur at the base of the Dalma volcanic succession. Sample SC06/3 was taken from a steep rocky outcrop (Fig. 6) on the left side of the access road to the Subernarekha River road bridge crossing, 250 m SSE (downstream) of the Chandil Dam wall (22° 58' 11.5 "N; 86° 01' 16.0" E). The felsic volcanic rocks at this site are homogeneous and aphanitic and contain rare, 1 mm thick, whitish quartz veins (not included in the sample processed). The sample is microfelsitic (in places, spherulitic) to vitrophyric and contains subhedral to anhedral phenocrysts of quartz and sodic plagioclase with accessory biotite, partly replaced by iron oxides, zircon, titanite and tourmaline. The phenocrysts are set in a cryptocrystalline groundmass, at places glassy with variable degree of devitrification. No evidence of flow layering is evident in thin-section.

Crushing and screening of samples was undertaken at Presidency University, Kolkata. Heavy minerals were isolated from 0.5 kg of sample using conventional heavy-liquid and magnetic techniques. Representative zircons were hand picked, mounted in epoxy and sec-tioned approximately in half, and the mount surface was then polished to expose the grain interiors. U, Th and Pb isotopic measurements were taken using the Perth Consortium SHRIMP II instrument employing operat-ing and data processing procedures similar to those described by Compston, Williams & Meyer (1984) and Williams et al. (1984). Pb/U ratios were determ-ined relative to that of the zircon standards CZ3 and Temora, which have been assigned ²⁰⁶Pb/²³⁸U values of 0.0914 and 0.066783 corresponding to ²⁰⁶Pb/²³⁸U dates of 564 Ma and 417 Ma, respectively. Data pro-cessing and plot generation were undertaken using the



Figure 8. (a) Trace element spider diagram, normalized by chondrite values of the studied felsic volcanic rocks, after Sun &McDonough (1989). (b) REE distribution patterns normalized by chondrite values of the studied felsic volcanics, after Boynton (1984).



Figure 9. Log Nb v. Log Y plots of Dalma felsic volcanic rocks, after Pearce, Harris & Tindale (1984). Syn-Colg – syn-collision granite; VAG – volcanic arc granite; WPG – within-plate granite; ORG – ocean ridge granite.

generalized ion-microprobe data processing software package CONCH (Nelson, 2006).

For the Arkasani Granophyre (sample SC07/05), 21 analyses were obtained of 21 zircons (Table 3; Fig. 10a). Most analyses are concordant within uncertainty, although two analyses indicating high common-Pb contents (12.1 and 16.1) plot along a zero-age discordia line. The population weighted mean 207 Pb/²⁰⁶Pb ratio and uncertainty correspond to a date of 1861 ± 8 Ma (chi squared = 0.83). This is interpreted as corresponding to the time of igneous crystalliza-tion of the granophyre. As the Arkasani Granophyre occurs as massive pods and lenses within and aligned along the Singhbhum Shear Zone, and these lenses are only weakly deformed around their margins, the date of 1861 ± 8 Ma is interpreted as providing the time of the final stages of major tectonic displacement along the Singhbhum Shear Zone. All 20 analyses obtained for zircons from the felsic volcanic rocks from the base of the Chandil Formation (SC06/03) are concordant within uncertainty (Table 4; Fig. 10b). The population weighted mean 207 Pb/ 206 Pb ratio and uncertainty correspond to a date of 1631 ± 6 Ma (chi squared = 0.71). This is interpreted as corresponding to the time of igneous crystallization of the felsic volcanic rocks and the timing of initiation of volcanism in the Dalma sub-basin.

5. Discussion

It has been established that NSMB is a collage of sedimentary and volcano-sedimentary successions that were developed in fault-controlled sub-basins in a regional rift basin that was closed during the Northern Singhbhum Orogeny (Bhattacharya & Mahapatra, 2008; Sarkar & Gupta, 2012). The Dhanjori, Chaibasa and Dhalbhum formations, along with the Dalma volcanic and North Dalma volcano-sedimentary successions, have been juxtaposed as curved linear belts within the NSMB. The contacts of these sub-basin ensembles are tectonized (see Saha, 1994; Bhattacharya & Mahapatra, 2008; Sarkar & Gupta, 2012) and the times of opening and closure of these sub-basins are not yet firmly established. The Singhbhum Shear Zone separates the Chaibasa Formation from the Dhanjori Formation and movement along the shear zone may also have resulted in the closure of the Dhanjori and Chaibasa sub-basins. The history of sedimentation (Bhattacharya & Mahapatra, 2008), deformation (Sengupta & Ghose, 1997) and metamorphism (Mahato et al. 2008) indicates development of this shear zone over a prolonged period of time that must have begun long before the intrusion date of 1861 ± 8 M a o f t h e Arkasani Granophyre.

The Arkasani Granophyre evolved through fractional crystallization of a metaluminous felsic magma generated at shallow crustal level and emplaced as



Figure 10. (a) Wetherill Concordia diagram for sample SC07/05: Arkasani Granophyre. (b) Wetherill Concordia diagram for sample SC06/03: felsic volcanic rocks, Dalma volcanic suite.

numerous isolated bodies aligned along the Singhbhum Shear Zone. Development of schistosity in the Arkasani Granophyre, particularly along the intrusion margins, and the geochemical character of Arkasani Granophyre indicates its syn- to post-kinematic emplacement. The SHRIMP U–Pb zircon date of 1861 ± 6 Ma for the Arkasani Granophyre therefore provides a minimum age for the shearing/thrusting along the Singhbhum Shear Zone and for the time of closure of the Chaibasa and Dhanjori sub-basins.

Singhbhum Shear Zone hosts minable Cu–P and U mineralizations. Cu–P and U mineralization occurrences are genetically unrelated and U mineralization might have been superimposed on Cu–P mineralization (Sarkar & Gupta, 2012). The 1.882 ± 23 Ma uraninite and 1.885 ± 31 Ma monazite dates (U–Pb dates by ICP-MS laser ablation) reported by Pal *et al.* (2011) and the date of 1861 ± 6 Ma reported here for the Arkasani Granophyre do not support any genetic linkage between granophyre emplacement and ore mineralization. Arkasani Granophyre emplacement might have been preceded by Cu–P and U mineralizations.

The SHRIMP U–Pb zircon date of 1631 ± 6 Ma obtained for the felsic volcanic rocks of the Dalma subbasin in the north of the Chaibasa sub-basin corresponds to a younger basin formation event that postdates the orogenic movement responsible for closure of the Chaibasa and Dhanjori sub-basins.

The new age data indicate that the evolution of adjacent sub-basins within the NSMB is complex, occurring at different times and by unrelated orogenic movements. Further, it seems evident that the younger subbasins formed further north away from the Archaean basement, and that the accretion of these basin ensembles onto Archaean basement was episodic. Domal exhumation of the Archean granitoid basement due to episodic emplacement of granite sheets during 3.4– 3.0 Ga may have resulted in development of peripheral curved fault systems and down-sagging of banded-ironformation (BIF) -bearing IOG basins (Roy & Bhattacharya, 2012). Movements along such Archean shear planes might have been responsible for the formation of Proterozoic basins skirting the Archean craton (see Roy and Bhattacharya, 2012). The Dhanjori and Chaibasa subbasins to the north of the Archean basement may be manifestations of such a process. The curved outcrop pattern of these two sub-basins is consistent with their development by the reactivation of ancient (Archean) curved fault systems. Closure and accretion of these two sub-basins on the Archean basement complex took place around 1861 \pm 6 Ma, as indicated by the age of emplacement of syn- to post-kinematic emplacement of Arkasani Granophyre. The SSZ, presently defining the contact between Chaibasa and Dhanjori sub-basins, might be related to ancient fault systems that were reactivated during closure of the two sub-basins. The mafic-felsic Dalma volcanic sequence and associated sedi-ments developed in an extensional basin that formed at least 200 Ma after the closure of the Dhanjori-Chaibasa sub-basin and its accretion onto the Archean basement complex. An evolution model for the NSMB is sum-marized in Figure 11.

The results obtained in the present study suggest that the different sub-basins in the NSMB did not evolve simultaneously within a regional rift system, as conceived in earlier models. Instead, the formation and closure of these basins took place diachronously. Late orogenic movement was responsible for juxtaposition of the sub-basins to form the NSMB. The present model also explains the curved pattern of the orogenic belt and the distinctly different petrotectonic evolutionary his-tories of the sub-basins.

Acknowledgements. We are grateful to Dr Subhasis Sengupta, Dr Saumitra Misra and Dr Nilanjan Dasgupta for their help during the different phases of fieldwork. We also thank



Figure 11. The evolution of the NSMB. (a) During stage 1 (late Archean – early Proterozoic), Dhanjori and Chaibasa sub-basins were formed as a result of reactivation of the early fault system developed skirting the central Archean block due to granite emplacement and doming. (b) During stage 2 (c. 1861 Ma), the Dhanjori and Chaibasa sub-basins were collapsed and accreted on the southern Archean block with the development of the Singhbhum Shear Zone and emplacement of the Arkasani Granophyre along the shear zone. Development of a zone of crustal extension in the north, under the influence of an uprising plume (see Roy *et al.* 2002), led to the compression which caused the collapse of the Dhanjori and Chaibasa sub-basins. (c) During stage 3 (c. 1631 Ma), the Dalma volcanic suite and the associated sediments were deposited in the extensional rift basin after the plume burst. (d)The ensembles of the Dalma sub-basin were juxtaposed on the Dhanjori and Chaibasa sub-basinal rocks after a later terminal orogeny during stage 4.

Table 3. Ion microprobe analytical results for sample SC07/5: Arkasani Granophyre.

Grain .spot	U (ppm)	Th (ppm)	Pb (ppm)	f206%	207Pb/206Pb	$\pm l\sigma$	208Pb/206Pb	$\pm l\sigma$	206Pb/238U	$\pm l\sigma$	207Pb/235U	$\pm l\sigma$	% concordance	207Pb/206Pb age	±lσ
1.1	92	42	33	0.026	0.11396	0.00117	0.12950	0.00216	0.3331	0.0039	5.234	0.087	99	1863	19
2.1	139	64	51	0.000	0.11383	0.00087	0.12631	0.00151	0.3454	0.0039	5.421	0.078	103	1861	14
3.1	445	290	166	-0.032	0.11517	0.00057	0.18404	0.00123	0.3334	0.0035	5.295	0.064	99	1883	9
4.1	86	46	31	0.116	0.11295	0.00161	0.15697	0.00346	0.3235	0.0038	5.039	0.099	98	1847	26
5.1	144	77	52	0.011	0.11154	0.00097	0.15010	0.00192	0.3311	0.0037	5.092	0.076	101	1825	16
6.1	133	69	50	0.139	0.11392	0.00124	0.14578	0.00257	0.3447	0.0039	5.415	0.090	103	1863	20
7.1	121	65	45	-0.026	0.11386	0.00122	0.15390	0.00254	0.3424	0.0039	5.375	0.089	102	1862	19
8.1	123	76	46	0.256	0.11431	0.00163	0.16943	0.00362	0.3329	0.0038	5.247	0.101	99	1869	26
9.1	161	113	61	-0.092	0.11427	0.00096	0.19480	0.00212	0.3392	0.0037	5.345	0.078	101	1868	15
10.1	269	126	97	0.000	0.11288	0.00064	0.13306	0.00114	0.3354	0.0036	5.221	0.066	101	1846	10
11.1	97	46	36	0.112	0.11494	0.00266	0.12904	0.00590	0.3405	0.0040	5.396	0.147	101	1879	42
12.1	210	158	72	1.050	0.11579	0.00177	0.20126	0.00404	0.2910	0.0031	4.646	0.092	87	1892	28
13.1	92	40	33	-0.043	0.11373	0.00155	0.12161	0.00315	0.3406	0.0040	5.341	0.101	102	1860	25
14.1	111	61	40	0.173	0.11244	0.00144	0.14959	0.00306	0.3309	0.0038	5.131	0.093	100	1839	23
15.1	114	54	43	0.225	0.11234	0.00140	0.14709	0.00299	0.3479	0.0039	5.389	0.096	105	1838	23
16.1	129	130	32	1.142	0.11379	0.00245	0.15486	0.00546	0.2224	0.0025	3.490	0.089	70	1861	39
17.1	139	75	51	-0.074	0.11387	0.00133	0.15216	0.00286	0.3349	0.0037	5.258	0.090	100	1862	21
18.1	140	47	50	-0.017	0.11411	0.00100	0.09533	0.00173	0.3420	0.0038	5.381	0.081	102	1866	16
19.1	124	69	46	0.160	0.11499	0.00166	0.15448	0.00364	0.3369	0.0038	5.342	0.104	100	1880	26
20.1	127	74	48	0.091	0.11416	0.00140	0.15944	0.00304	0.3404	0.0038	5.358	0.095	101	1867	22
21.1	93	43	33	0.218	0.11340	0.00156	0.12420	0.00320	0.3324	0.0038	5.197	0.099	100	1855	25

Table	4.	Ion micropro	be analytical	results for	sample	e SC06/3:	felsic vo	olcanic roc	ks of D	alma vo	lcanic su	ite
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Grain .spot	U (ppm)	Th (ppm)	Pb (ppm)	f206%	207Pb/206Pb	$\pm l\sigma$	208Pb/206Pb	$\pm l\sigma$	206Pb/238U	$\pm l\sigma$	207Pb/235U	$\pm l\sigma$	% concordance	207Pb/206Pb age	$\pm l\sigma$
1.1	420	145	125	0.056	0.10047	0.00047	0.09837	0.00071	0.2878	0.0041	3.987	0.062	100	1633	9
2.1	152	82	47	0.092	0.10039	0.00092	0.15092	0.00178	0.2854	0.0042	3.951	0.072	99	1631	17
3.1	454	204	135	0.095	0.10072	0.00051	0.12628	0.00091	0.2810	0.0040	3.902	0.062	97	1638	9
4.1	197	175	65	0.086	0.09983	0.00072	0.25255	0.00165	0.2831	0.0041	3.896	0.066	99	1621	13
5.1	241	100	73	-0.041	0.10068	0.00063	0.11773	0.00100	0.2898	0.0042	4.023	0.066	100	1637	12
6.1	218	110	68	-0.032	0.10098	0.00113	0.14500	0.00240	0.2900	0.0042	4.038	0.078	100	1642	21
7.1	138	93	44	0.053	0.10194	0.00089	0.19149	0.00178	0.2860	0.0042	4.020	0.072	98	1660	16
8.1	112	72	35	0.081	0.09993	0.00204	0.18122	0.00457	0.2816	0.0042	3.880	0.104	99	1623	38
9.1	278	147	87	0.110	0.10010	0.00062	0.14995	0.00115	0.2899	0.0042	4.002	0.066	101	1626	12
10.1	207	145	66	-0.041	0.10038	0.00134	0.20247	0.00301	0.2865	0.0042	3.966	0.083	100	1631	25
11.1	168	96	52	0.005	0.10052	0.00074	0.16514	0.00135	0.2876	0.0042	3.986	0.068	100	1634	14
12.1	244	160	77	0.040	0.10022	0.00065	0.18797	0.00130	0.2857	0.0041	3.948	0.066	100	1628	12
13.1	358	156	107	0.024	0.09987	0.00052	0.12489	0.00087	0.2830	0.0041	3.897	0.062	99	1622	10
14.1	192	70	58	0.080	0.10016	0.00079	0.10268	0.00135	0.2889	0.0042	3.990	0.069	101	1627	15
15.1	191	147	62	0.236	0.09944	0.00086	0.21215	0.00187	0.2829	0.0041	3.879	0.069	100	1614	16
16.1	291	143	90	0.020	0.10067	0.00054	0.14117	0.00089	0.2891	0.0042	4.013	0.064	100	1636	10
17.1	194	99	60	0.049	0.10022	0.00070	0.14410	0.00122	0.2871	0.0042	3.967	0.067	100	1628	13
18.1	143	112	48	0.020	0.10175	0.00077	0.22597	0.00165	0.2907	0.0043	4.078	0.071	99	1656	14
19.1	241	148	75	0.094	0.09909	0.00071	0.17200	0.00142	0.2828	0.0041	3.864	0.066	100	1607	13
20.1	363	202	112	0.148	0.10028	0.00059	0.15977	0.00114	0.2835	0.0041	3.919	0.064	99	1629	11

Mr Kaushik Kiran Ghosh, who drafted some of the figures. We wish to acknowledge the technical support provided by Tarak Pradhan at Presidency University. Finally, we extend our sincere gratitude to anonymous reviewers for their constructive suggestions.

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