

Exploring XMM-ATLAS with the ARCHES tools

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Paris, France, 30-November- 2015



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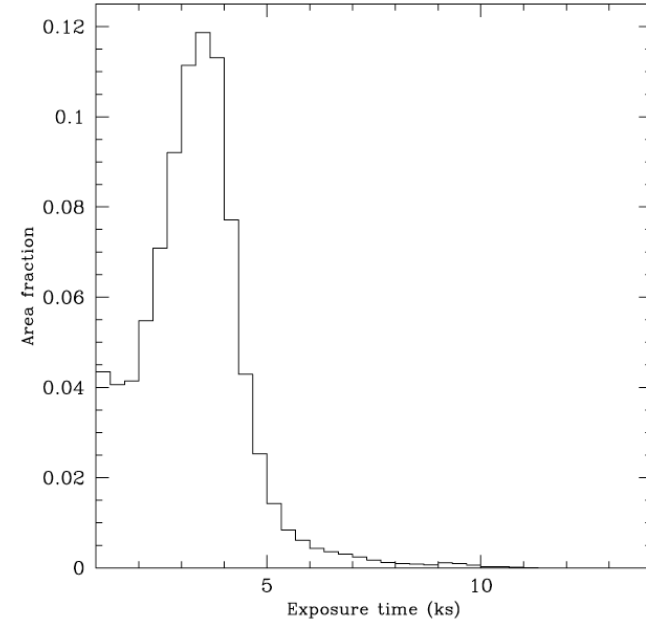
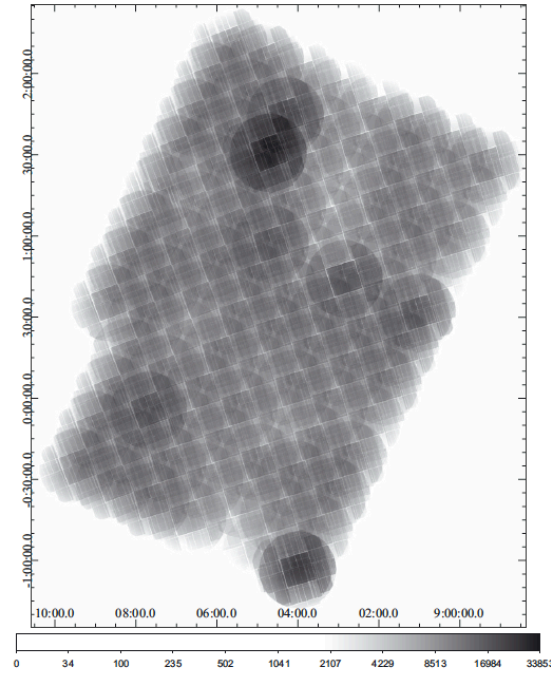
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Outline

- Introduction
 - XMM-ATLAS
 - Other catalogues: SDSS, VIKING, WISE
 - Astrometry
 - The ARCHES cross-correlation tool
- SDSS in full area:
 - Improvements
 - Checks
- Conclusions

XMM-ATLAS

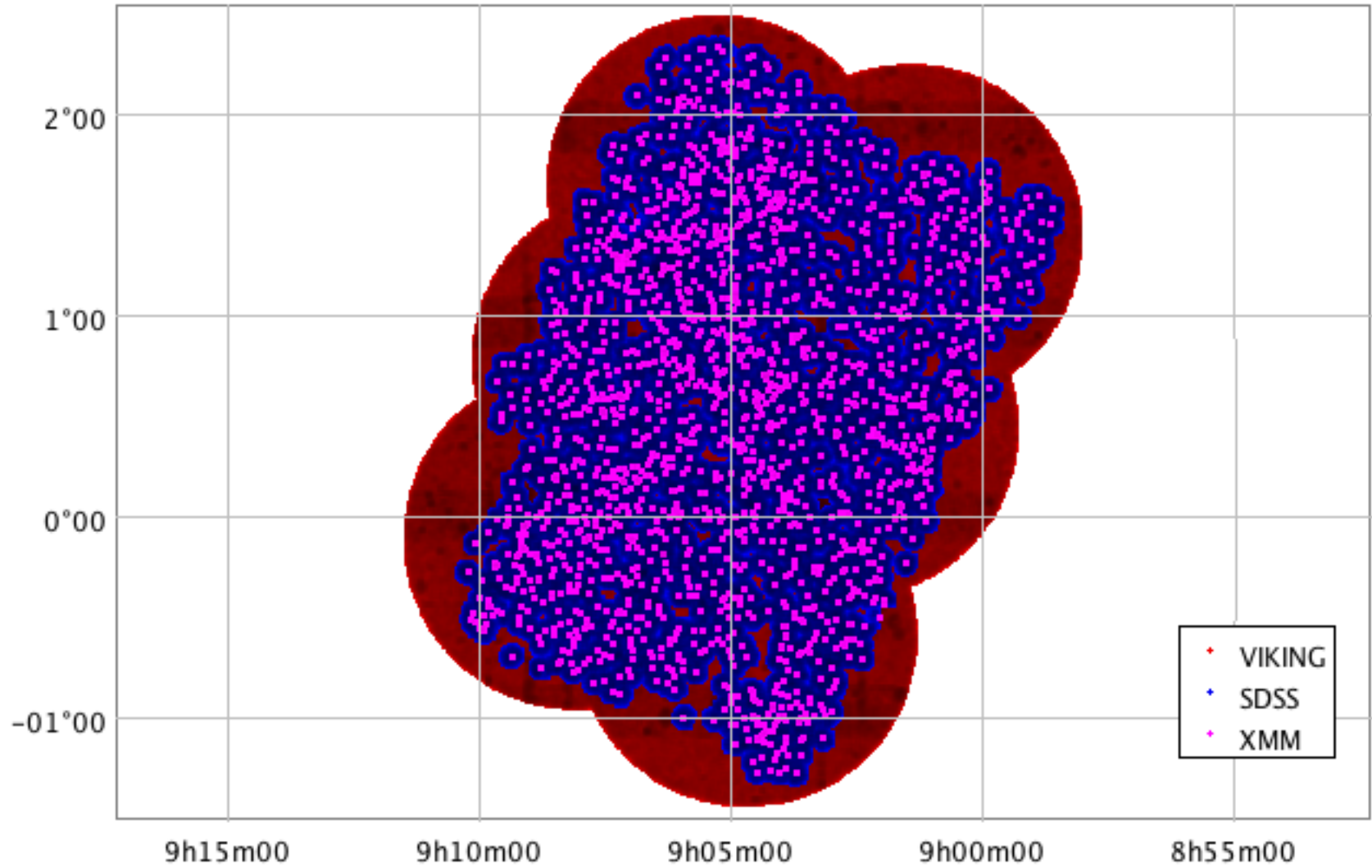


- Centered in Herschel-ATLAS SDP ([Rigby+11](#))
 - (09:04:30,+00:34:00)
- [Ranalli+15](#): source catalogue
 - 7.1deg², 336ks total: mode ~3.5ks
 - Wavelet+em1detect source search: **1816 sources**
 - in three bands: 0.5-2, 2-8, 0.5-2keV

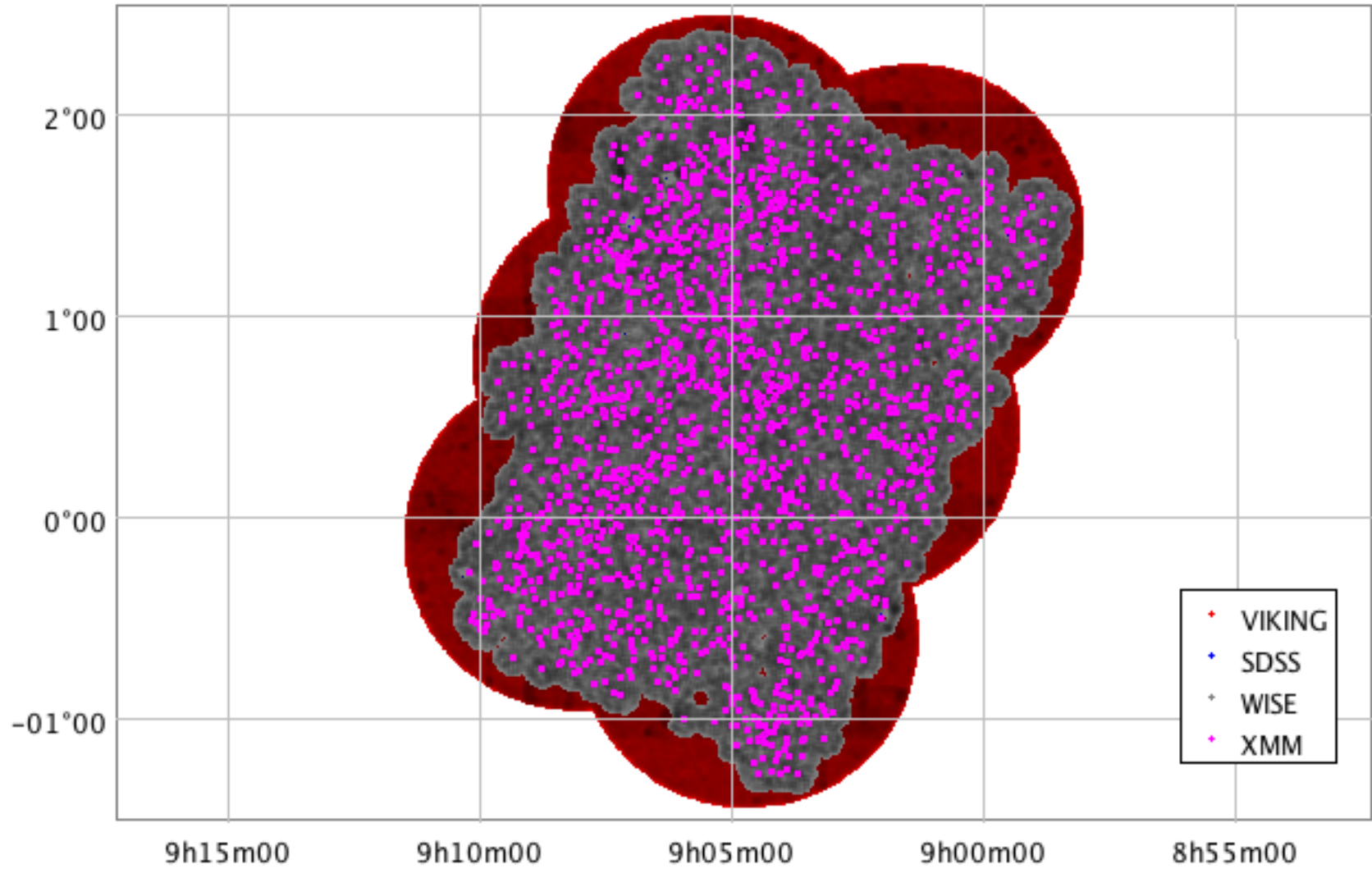
Other catalogues

- SDSS DR9 ([Adelman-McCarthy+09](#)): ugriz
 - 161131 sources in overlapping area (111404 `clean=1`)
- VIKING DR1 ([Edge & Sutherland 2014](#)): zYJHKs
 - band-merged source catalogue
 - 736187 sources in \sim overlapping area (`421850 pNoise<0.5 && nBands \geq 2`)
- WISE ([Cutri+2012](#)): 3.4, 4.6, 12 & 22 μ m
 - 68147 sources in overlapping area (all det. 5σ in ≥ 1 band)

Other catalogues



Other catalogues



Astrometry

- SDSS DR9: adding 0.1" in quadrature to the RA,Dec pos. errors
- For **XMM-ATLAS** using SDSS DR12 QSO as reference
 - Filtering QSOs on good z quality
 - Filtering XMM-ATLAS pointlike $ext058 < 0.001$
 - 71 pairs within 5": $\langle dRA \rangle = -0.02''$ $\langle dDec \rangle = -0.2''$ $\sigma \sim 1.5''$
 - Ranalli+15 rectified w.r.t. SDSS DR7 QSOs, residual difference?
 - Adding 1.5" in quadrature to `radec_err` from `emldetect`
- For **VIKING** using SDSS DR9 point sources as reference
 - Filtering on VIKING point sources
 - 51014 pairs within 1": $\langle dRA \rangle \sim \langle dDec \rangle \sim 0.035''$ $\sigma \sim 0.1''$
 - Setting `errpos=0.1''` (no positional error in catalogue)
- **WISE** used as in catalogue

The ARCHES cross-correlation tool

- `archesxmatch`: Full N-dimensional symmetric cross-correlation of catalogues (see F.-X. Pineau's talk)
- It uses:
 - Source positions
 - Source positional errors (1σ)
 - Catalogue area
 - Sky densities of pairs, triplets...
- It provides:
 - List of tuples within some user-chosen 3σ distance
 - Probabilities of all combinations of catalogues for that tuple
 - List of five highest probabilities (when defined):
`maxProbaVal1...maxProbaVal5`
 - Marginalised 2D probabilities for tuples in ≥ 3 catalogues:
 - e.g.: `margProba_AB` for A,B,C: $P_{\text{marg}}(\text{AB})=P(\text{ABC})+P(\text{AB}_C)$

ARCHES xcorr. tool: X and 3 other cats

- For simplicity, let's consider $N=4$ (XMM-ATLAS+3 cats.)
 - $A=XMM$, $B=SDSS$, $C=VIKING$, $D=WISE$
- "Left-join": each XMM-ATLAS source will be considered afresh with each new catalogue: **X-ray-centric**
- **Output tuples could have up to four dimensions:**
 - No source in any other catalogue within limit $nPos=1$
 - A source in 1 other catalogue, $nPos=2$: AB, A_B ; AC, A_C ; AD, A_D
 - One source in each of 2 other catalogues, $nPos=3$:
 - $ABC, AB_C, AC_B, A_BC \equiv (A)(BC), A_B_C \equiv (A)(B)(C)$
 - $ABD, AB_D, AD_B, A_BD, A_B_D$
 - $ACD, AC_D, AD_C, A_CD, A_C_D$
 - $BCD, BC_D, BD_C, B_CD, B_C_D$ (remember, X-ray-centric)
 - One source in each of 3 other catalogues, $nPos=4$:
 - $ABCD$
 - ABC_D, ABD_C, ACD_B
 - $AB_CD, AB_C_D, AC_BD, AC_B_D, AD_BC, AD_B_C$
 - $A_BCD, A_BC_D, A_BD_C, A_B_CD, A_B_C_D$

ARCHES xcorr. tool: X and 3 other cats

- **X-ray-centric**: Considering only relationship with XMM source
 - Using for classification (for the time being) only maximum probability $\max[P(*)]=P_{\max} \equiv \max \text{ProbaVal1}$

		nPos			
		1	2	3	4
Only X	X___	All	A_B, A_C, A_D	A_BC,A_B_C, A_CD,A_C_D, A_BD,A_B_D	A_BCD,A_BC_D,A_B_CD,A_B_C_D, A_BD_C
X+1cat	XS__	-	AB	AB_C,AB_D	AB_CD,AB_C_D
	X_V_	-	AC	AC_B,AC_D	AC_BD,AC_B_D
	X__W	-	AD	AD_B,AD_C	AD_BC,AD_B_C
X+2cat	XSV_	-	-	ABC	ABC_D
	X_VW	-	-	ACD	ACD_B
	XS_W	-	-	ABD	ABD_C
X+3cat	XSVW	-	-	-	ABCD

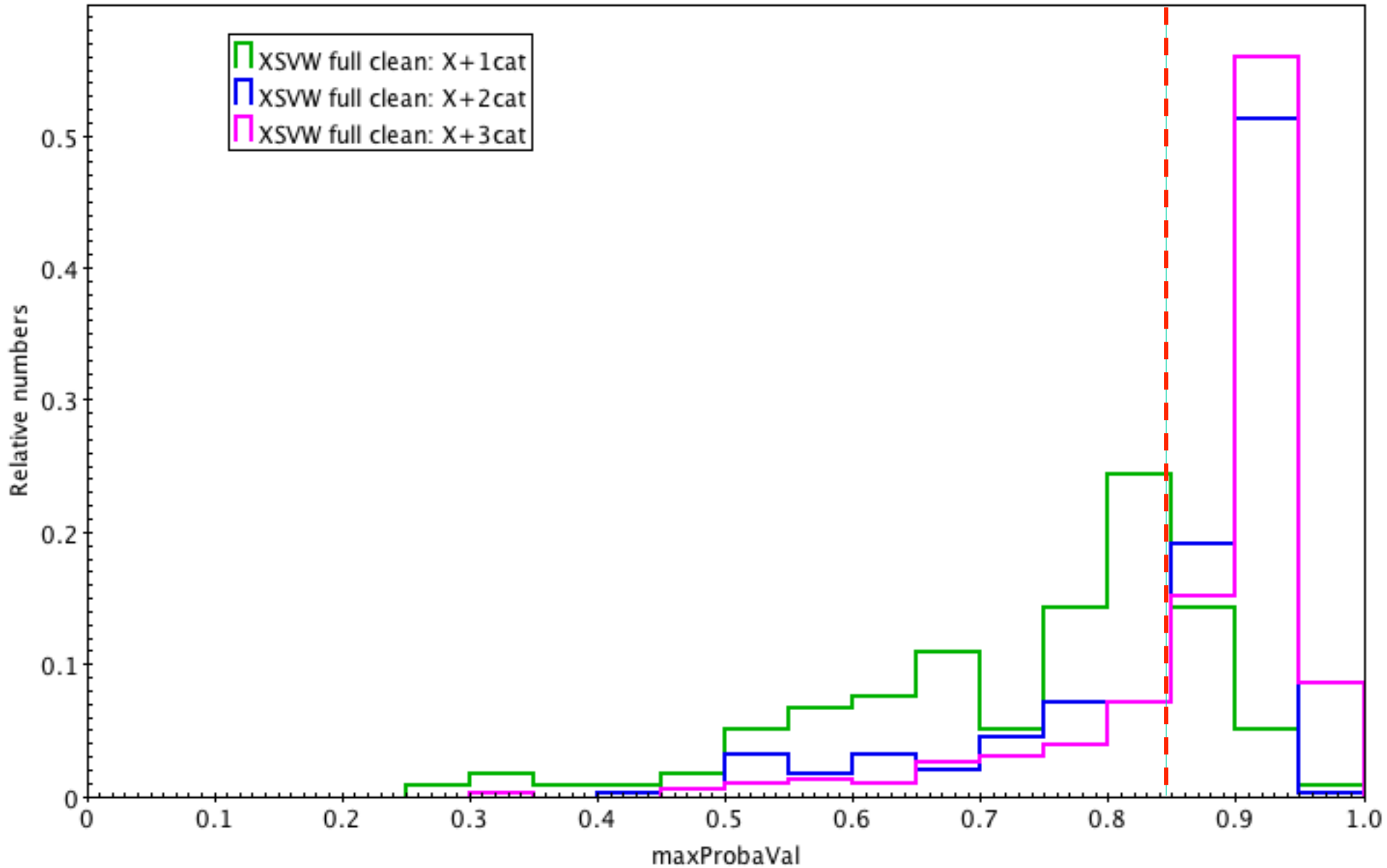
SDSS in full area: "clean" sample

- Full sample 1816 XMM sources in 7.1deg^2
- For each XMM sou. keeping only tuple with highest P_{max}
- Using "clean" sample:
 - SDSS: `clean=1`
 - VIKING: `pNoise<0.9` & present in >1 band
- "Clean" xcorr better overall: more XMM with counterparts
 - Because probabilities also depend on density of X+1,2,3cat pairs

	Subsample	SDSS		SDSS clean	
Only X	X__	803	803	717	717
X+1cat	XS__	77	26	119	11
	X_V_		40		92
	X__W		11		16
X+2cat	XSV_	360	324	419	310
	X_VW		34		108
	XS_W		2		1
X+3cat	XSVW	576	576	516	561

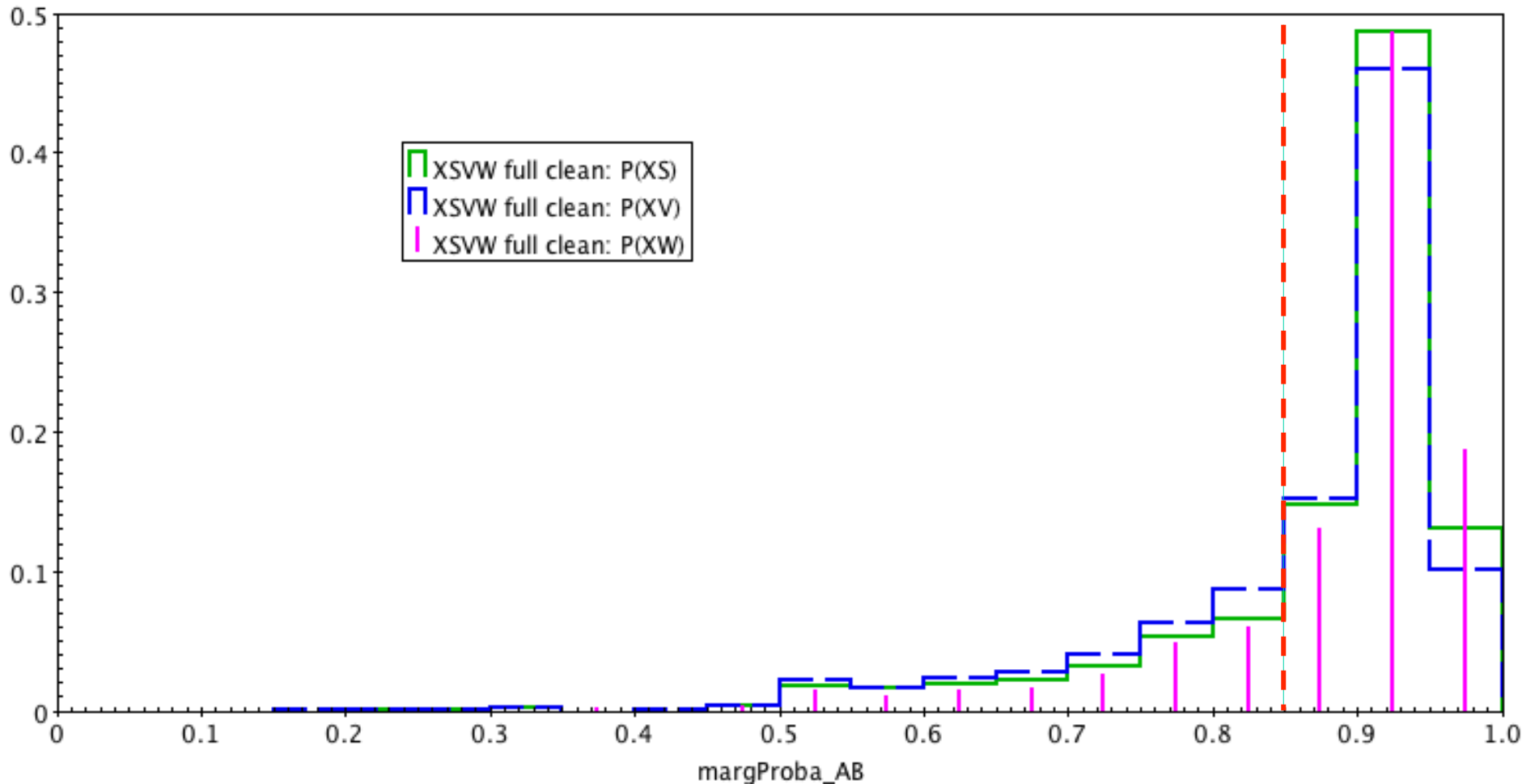
SDSS full clean: probability threshold

- Choosing threshold **85%**: drop in P_{\max} X+2,3cat



SDSS full clean: probability threshold

- Choosing threshold **85%**: drop in marginalised probabilities too



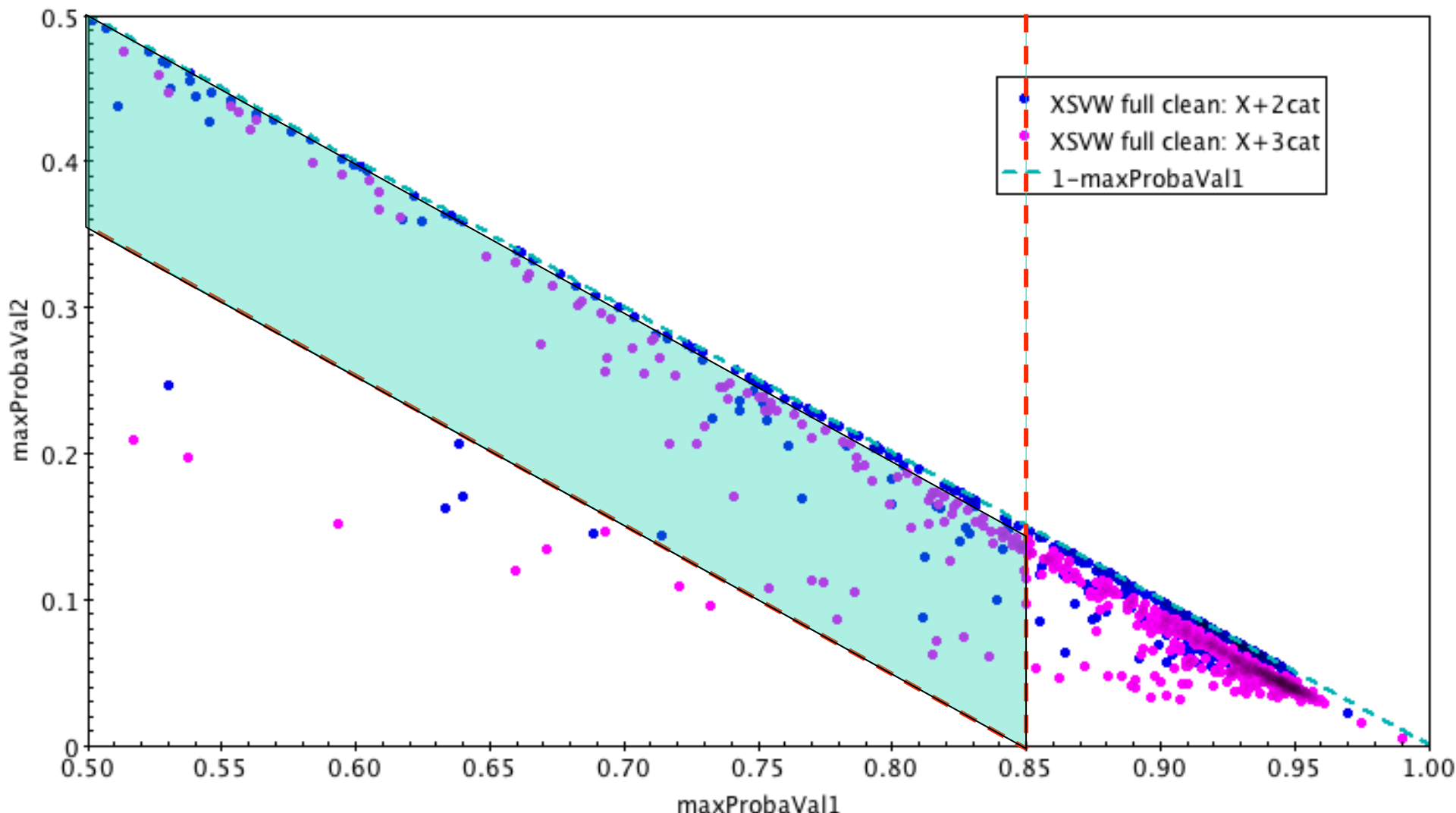
SDSS full clean: prob. $\geq 85\%$

- $\sim 2/3$ of the X+2,3cat survive (by design)

	Subsample	SDSS clean prob. $\geq 85\%$		SDSS clean	
Only X	X__	1049	1049	717	717
X+1cat	XS__	24	5	119	11
	X_V_		15		92
	X__W		4		16
X+2cat	XSV_	296	216	419	310
	X_VW		80		108
	XS_W		0		1
X+3cat	XSVW	447	447	516	561

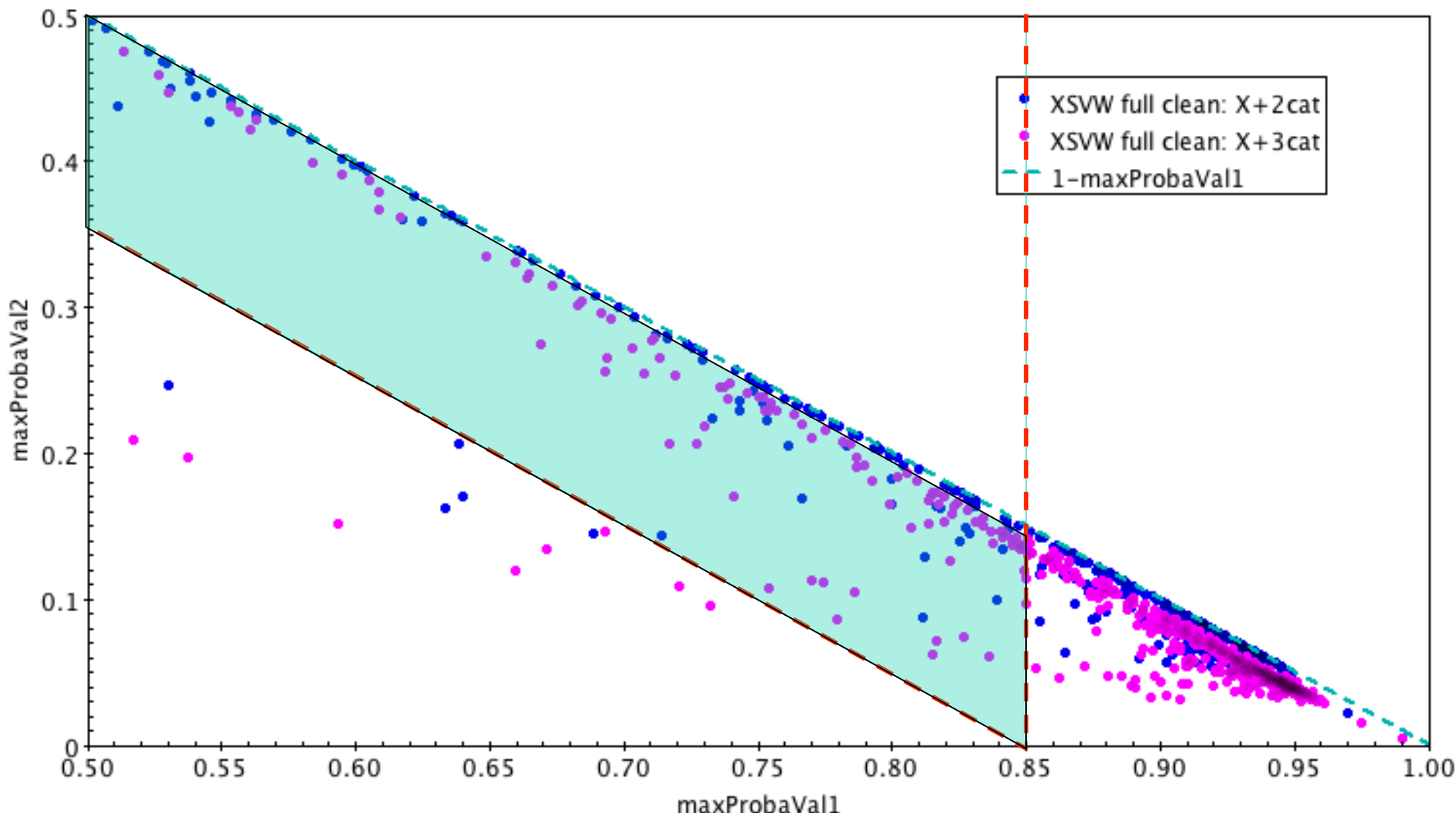
SDSS full clean: best tuple, 1st & 2nd

- For the tuple with the highest P_{\max} , comparing P_{\max} and second highest prob. $P_{2\text{nd}} = \max \text{ProbaVal}2$
- What do we do about those with $P_{\max} + P_{2\text{nd}} \geq 0.85$?



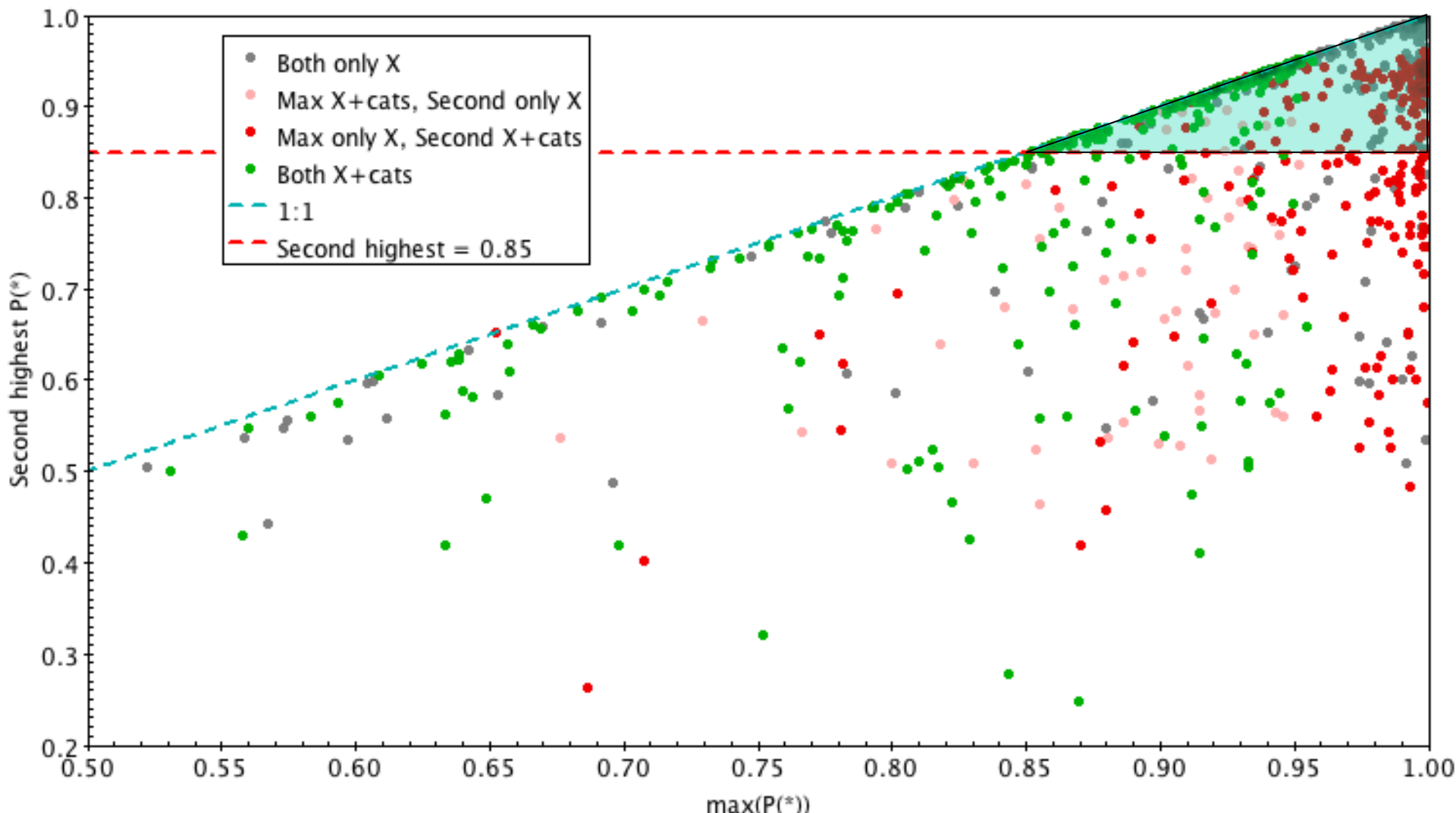
SDSS full clean: best tuple, 1st & 2nd

- What do we do about those with $P_{\max} + P_{2\text{nd}} \geq 0.85$?
 - If compatible, could boost total probability above threshold: 0 cases
 - If not compatible, could change P_{\max} combination: always $\sim ABC, A_BC$



SDSS full clean: 1st & 2nd best tuples

- For each XMM source, comparing tuples with 2 highest P_{\max} : 801 unique XMM-ATLAS sources ($1015 \leq 1$ tuple)
- What do we do about those with $P_{\max}(2^{\text{nd}}) \geq 0.85$?

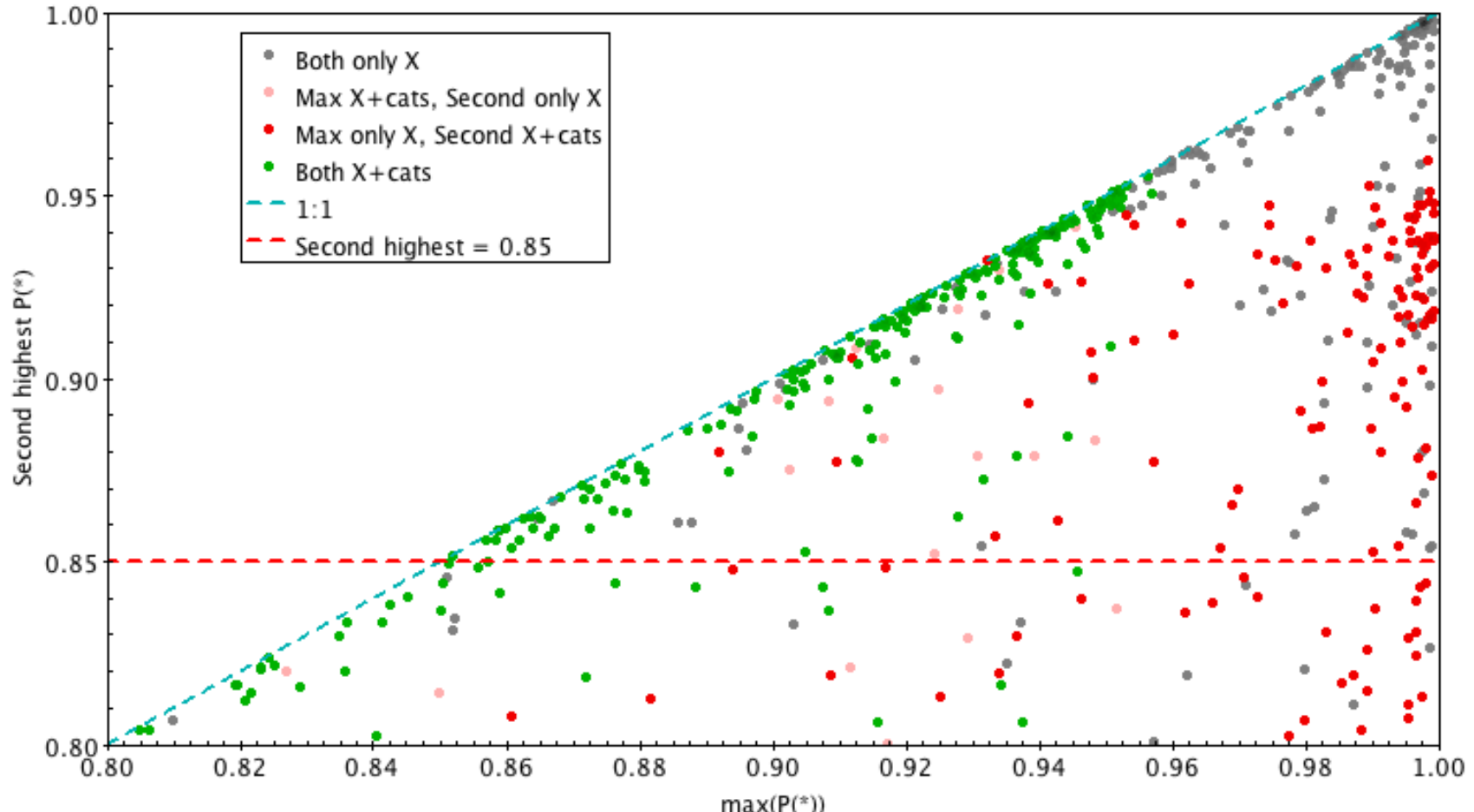


SDSS full clean: 1st & 2nd best

1st	2nd	N	$P_{\max}(2^{\text{nd}}) \geq 0.85$
Only X	Only X	230	-
X+cat	Only X	60	13
Only X	X+cat	192	97
X+cat	X+cat	319	192
Total		801	302

• Counting:

- Remember $P_{\max} \geq P_{\max}(2^{\text{nd}})$
- Worry about $P_{\max}(2^{\text{nd}}) \geq 0.85$?



SDSS in full area: questions/issues

- Areas not quite matched for VIKING
- At the moment, for each tuple, using just $\max[P(*)]$, should we use marginalised probabilities instead? e.g.
 - For probs. from 3 catalogues: margProba_A^*
 - For probs. from 4 catalogues: $P(\text{XSV_}) = P(\text{ABC_D}) + P(\text{ABCD})$
 - In how many these marginalised probs. would change combination?
 - For XMM-ATLAS/SDSS/VIKING/WISE: 0
- At the moment, for each XMM source, using just the tuple with the highest $\max[P(*)]$, should we worry about the other tuples?
 - How many of those share SDSS/VIKING/WISE sources?
 - Each source from each catalogue can only belong to one tuple!
 - For XMM-ATLAS/SDSS/VIKING/WISE: most change VIKING, a few change also SDSS
 - In how many the marginalised probs. above would change order?
 - For XMM-ATLAS/SDSS/VIKING/WISE: 1 from X_VW to X__W, diff. VIKING

Conclusions

- Tool works well
- The input needs to be worked on:
 - Matched astrometry
 - Matched sky coverage
- The output needs to be worked on:
 - Understand the probabilities and their meaning
 - Choose the one(s) that best suit what is needed
- Future for XMM-ATLAS cross-correlations:
 - Match XMM astrometry?
 - Match sky coverages (XMM-SDSS-VIKING-WISE, add KiDS)
 - Repeat cross-correlations
 - Marginalise 3,4 catalogue probabilities?
 - Get SEDs (ARCHES tool too)
 - Get photo-z (+ U. Napoli: machine learning)

4 catalogues: marginalised probs.

- X-ray-centric**: Considering only relationship with XMM source

		Probabilities for nPos			
		1	2	3	4
Only X	X___	All	$P(A_B)$ $P(A_C)$ $P(A_D)$	$P(A_{BC})+P(A_{B_C})$ $P(A_{CD})+P(A_{C_D})$ $P(A_{BD})+P(A_{B_D})$	$P(A_{BCD})+P(A_{BC_D})$ $+P(A_{B_CD})+P(A_{B_C_D})+$ $P(A_{BD_C})$
X+1cat	XS__	-	margProba_AB		
	X_V_	-	margProba_AC		
	X__W	-	margProba_AD		
X+2cat	XSV_	-	-	$P(ABC)$	$P(ABC_D)+P(ABCD)$
	X_VW	-	-	$P(ACD)$	$P(ACD_B)+P(ABCD)$
	XS_W	-	-	$P(ABD)$	$P(ABD_C)+P(ABCD)$
X+3cat	XSVW	-	-	-	$P(ABCD)$