

The magnetic flow on the manifold of oriented geodesics of a three dimensional space form

Yamile Godoy and Marcos Salvai *

Abstract

Let M be the three dimensional complete simply connected manifold of constant sectional curvature $0, 1$ or -1 . Let \mathcal{L} be the manifold of all (unparametrized) complete oriented geodesics of M , endowed with its canonical pseudo-Riemannian metric of signature $(2, 2)$ and Kähler structure J . A smooth curve in \mathcal{L} determines a ruled surface in M .

We characterize the ruled surfaces of M associated with the magnetic geodesics of \mathcal{L} , that is, those curves σ in \mathcal{L} satisfying $\nabla_{\dot{\sigma}}\dot{\sigma} = J\dot{\sigma}$. More precisely: a time-like (space-like) magnetic geodesic determines the ruled surface in M given by the binormal vector field along a helix with positive (negative) torsion. Null magnetic geodesics describe cones, cylinders or, in the hyperbolic case, also cones with vertices at infinity. This provides a relationship between the geometries of \mathcal{L} and M .

*Key words and phrases:*¹ manifold of oriented geodesics, Hermitian symmetric space, magnetic flow, ruled surface, horospherical distribution

1 Introduction

For $\kappa = 0, 1, -1$, let M_κ be the three dimensional complete simply connected manifold of constant sectional curvature κ , that is, \mathbb{R}^3 , \mathbb{S}^3 and the hyperbolic space \mathbb{H}^3 . Let \mathcal{L}_κ be the manifold of all (unparametrized) complete oriented geodesics of M_κ . We may think of an element c in \mathcal{L}_κ as the equivalence class of unit speed geodesics $\gamma : \mathbb{R} \rightarrow M_\kappa$ with image c such that $\{\dot{\gamma}(s)\}$ is a positive basis of $T_{\gamma(s)}c$ for all s .

Let γ be a complete unit speed geodesic of M_κ and let \mathcal{J}_γ be the space of all Jacobi fields along γ which are orthogonal to γ . There exists a well-defined canonical isomorphism

$$T_\gamma : \mathcal{J}_\gamma \rightarrow T_{[\gamma]}\mathcal{L}_\kappa, \quad T_\gamma(J) = \left. \frac{d}{dt} \right|_0 [\gamma_t], \quad (1)$$

*Partially supported by CONICET, FONCYT, SECYT (UNC).

¹*Mathematics Subject Classification:* 53C22, 53C35, 53C55; 53B30, 53C50.

where γ_t is any variation of γ by unit speed geodesics associated with J (see [8]).

A pseudo-Riemannian metric of signature $(2, 2)$ can be defined on \mathcal{L}_κ as follows [9]: For $X \in T_{[\gamma]}\mathcal{L}_\kappa$, the square norm $\|X\| = \langle X, X \rangle$ is well defined by

$$\|X\| = \langle \dot{\gamma} \times J, J' \rangle, \quad (2)$$

where $X = T_\gamma(J)$, the cross product \times is induced by a fixed orientation of M_κ and J' denotes the covariant derivative of J along γ . Indeed, the right hand side of (2) is a constant function. In the following, for any vector X , we will denote $\|X\| = \langle X, X \rangle$ and $|X| = \sqrt{|\langle X, X \rangle|}$. Recall that X is null, time-like or space-like if $\|X\| = 0$, $\|X\| < 0$ or $\|X\| > 0$, respectively.

Let $[\gamma] \in \mathcal{L}_\kappa$ and let R_γ be the rotation in M_κ fixing γ through an angle of $\pi/2$. This rotation induces an isometry \tilde{R}_γ of \mathcal{L}_κ whose differential at $[\gamma]$ is a linear isometry of $T_{[\gamma]}\mathcal{L}_\kappa$ squaring to $-\text{id}$. This yields a complex structure J on \mathcal{L}_κ . With the metric defined above, \mathcal{L}_κ is Kahler.

A *magnetic geodesic* σ of \mathcal{L}_κ is a curve satisfying $\nabla_{\dot{\sigma}}\dot{\sigma} = J\dot{\sigma}$. These curves have constant speed, so they will be null, time-like or space-like.

A smooth curve in \mathcal{L}_κ determines a ruled surface in M_κ . For $\kappa = 0, -1$, a generic geodesic of \mathcal{L}_κ describes a helicoid in M_κ [5, 4, respectively]. Our purpose is to characterize the ruled surfaces in M_κ associated with the magnetic geodesics of \mathcal{L}_κ . For $v \in TM_\kappa$, γ_v denotes the geodesic of M_κ with initial velocity v .

Theorem 1 *A generic magnetic geodesic σ of \mathcal{L}_κ describes the ruled surface in M_κ given by the binormal vector field of a helix. More precisely, σ is a time-like (space-like) magnetic geodesic of \mathcal{L}_κ if and only if σ has the form*

$$\sigma(t) = [\gamma_{B(t)}], \quad (3)$$

where B is the binormal vector field of a helix in M_κ with curvature k , speed $1/k$ and positive (negative) torsion, for some $k > 0$.

Now we study null magnetic geodesics in $\mathcal{L}_{-1} = \mathcal{L}(\mathbb{H}^3)$. We recall some concepts related with the hyperbolic space (see for instance [3]).

Two unit speed geodesics γ and α of \mathbb{H}^3 are said to be asymptotic if there exists a positive constant C such that $d(\gamma(s), \alpha(s)) \leq C, \forall s \geq 0$. Two unit vectors $v, w \in T^1\mathbb{H}^3$ are said to be asymptotic if the corresponding geodesics γ_v and γ_w have this property.

A point at infinity for \mathbb{H}^3 is an equivalence class of asymptotic geodesics of \mathbb{H}^3 . The set of all points at infinity for \mathbb{H}^3 is denoted by $\mathbb{H}^3(\infty)$ and has a canonical differentiable structure diffeomorphic to the 2-sphere. The equivalence class represented by a geodesic γ is denoted by $\gamma(\infty)$, and the equivalence class represented by the oppositely oriented geodesic $s \mapsto \gamma(-s)$ is denoted by $\gamma(-\infty)$.

Given $v \in T^1\mathbb{H}^3$, the *horosphere* $H(v)$ is the limit of metric spheres $\{S_n\}$ in \mathbb{H}^3 that pass through the foot point of v as the centers $\{p_n\}$ of $\{S_n\}$ converge to $\gamma_v(\infty)$. Below we present a more precise definition.

Let $\psi^\pm : \mathcal{L}(\mathbb{H}^3) \rightarrow \mathbb{H}^3(\infty)$ be the smooth functions given by $\psi^\pm([\gamma]) = \gamma(\pm\infty)$ and let \mathcal{D}^\pm be the distributions on $\mathcal{L}(\mathbb{H}^3)$ given by $\mathcal{D}_{[\gamma]}^\pm = \text{Ker}(d\psi_{[\gamma]}^\pm)$. These distributions are called *the horospherical distributions* on $\mathcal{L}(\mathbb{H}^3)$.

Cones with vertices at infinity: Let $x \in \mathbb{H}^3(\infty)$ and let $v_o \in T^1\mathbb{H}^3$ such that $\gamma_{v_o}(\pm\infty) \in x$. Let $t \mapsto v(t)$ be a curve in $T^1\mathbb{H}^3$ such that $v(0) = \pm v_o$, $v(t)$ is asymptotic to $\pm v_o$ for all $t \in \mathbb{R}$ and the foot points of $v(t)$ lie on a circle of geodesic curvature $\pm k$ (with $k > 0$) and speed $1/k$ in the horosphere determined by $\pm v_o$. Under these conditions we say that the curve in $\mathcal{L}(\mathbb{H}^3)$ given by $t \mapsto [\gamma_{\pm v(t)}]$ describes a *forward cone with vertex at x (for $+$)* or a *backward cone with vertex at x (for $-$)*. These cones can be better visualized in the upper half space model of \mathbb{H}^3 (in particular $\mathbb{H}^3(\infty) = \{z = 0\} \cup \{\infty\}$): Let $\gamma_t^\pm(s) = (\frac{1}{k} \cos(t), \pm \frac{1}{k} \sin(t), e^{\pm s})$. A curve σ in $\mathcal{L}(\mathbb{H}^3)$ describes a cone with forward (respectively, backward) vertex at ∞ if it is $Sl(2, \mathbb{C})$ -congruent to $t \mapsto [\gamma_t^+]$ (respectively, to $t \mapsto [\gamma_t^-]$).

Theorem 2 *A null magnetic geodesic of $\mathcal{L}(\mathbb{H}^3)$ describes in \mathbb{H}^3 a cylinder, a cone with vertex at $p \in \mathbb{H}^3$ or a cone with vertex at infinity. More precisely, if σ is a curve in $\mathcal{L}(\mathbb{H}^3)$, then*

- a) σ is a null magnetic geodesic with $\dot{\sigma}(0) \in \mathcal{D}_{\sigma(0)}^\pm$ if and only if σ describes a cone with vertex at $\sigma(0)(\pm\infty)$ (forward for $+$ and backward for $-$);
- b) σ is a null magnetic geodesic with $\dot{\sigma}(0) \notin \mathcal{D}_{\sigma(0)}^\pm$ if and only if σ either has the form

$$\sigma(t) = [\gamma_{B(t)}], \quad (4)$$

where B is the binormal vector field of a helix h in \mathbb{H}^3 with curvature k , speed $1/k$ and zero torsion (in particular, h is contained in a totally geodesic surface S and B is normal to S and parallel along h), or σ has the form

$$\sigma(t) = [\gamma_{v(t)}], \quad (5)$$

where v is a curve with geodesic curvature k and speed $1/k$ in $T_p^1\mathbb{H}^3$, for some $p \in \mathbb{H}^3$, for certain $k > 0$.

Theorem 3 *The ruled surfaces associated with null magnetic geodesics of \mathcal{L}_κ for $\kappa = 0, 1$ are described in an analogous manner as in the previous theorem, except that case a) is empty. Besides, for $\kappa = 1$, a null magnetic geodesic has simultaneously the forms (4) and (5).*

2 Preliminaries

For the simultaneous analysis of the three cases $\kappa = 0, 1, -1$, we consider the standard presentation of M_κ as a submanifold of \mathbb{R}^4 . That is, $\mathbb{R}^3 = \{(1, x) \in \mathbb{R}^4 \mid x \in \mathbb{R}^3\}$, $\mathbb{S}^3 = \{x \in \mathbb{R}^4 \mid |x|^2 = 1\}$ and $\mathbb{H}^3 = \{x \in \mathbb{R}^4 \mid -x_0^2 + x_1^2 + x_2^2 + x_3^2 = -1 \text{ and } x_0 > 0\}$.

Let G_κ be the identity component of the isometry group of M_κ , that is, $G_0 = SO_3 \times \mathbb{R}^3$, $G_1 = SO_4$ and $G_{-1} = O_o(1, 3)$. We consider the usual presentation of G_0 as a subgroup of $GL_4(\mathbb{R})$. The group G_κ acts on \mathcal{L}_κ as follows: $g \cdot [\gamma] = [g \circ \gamma]$. This action is transitive and smooth.

If we denote by \mathfrak{g}_κ the Lie algebra of G_κ we have that

$$\mathfrak{g}_\kappa = \left\{ \begin{pmatrix} 0 & -\kappa x^t \\ x & B \end{pmatrix} \mid x \in \mathbb{R}^3, B \in so_3 \right\}.$$

Let γ_o be the geodesic in M_κ with $\gamma_o(0) = e_0$ and initial velocity $e_1 \in T_{e_0}M_\kappa$, where $\{e_0, e_1, e_2, e_3\}$ is the canonical basis of \mathbb{R}^4 . For $A, B \in \mathbb{R}^{2 \times 2}$, let $\text{diag}(A, B) = \begin{pmatrix} A & 0_2 \\ 0_2 & B \end{pmatrix}$, where 0_2 denotes the 2×2 zero matrix. Then the isotropy subgroup of G_κ at $[\gamma_o]$ is

$$H_\kappa = \{\text{diag}(R_\kappa(t), B) \mid t \in \mathbb{R}, B \in SO_2\},$$

where

$$R_0(t) = \begin{pmatrix} 1 & 0 \\ t & 1 \end{pmatrix}, \quad R_1(t) = \begin{pmatrix} \cos t & -\sin t \\ \sin t & \cos t \end{pmatrix}, \quad R_{-1}(t) = \begin{pmatrix} \cosh t & \sinh t \\ \sinh t & \cosh t \end{pmatrix}. \quad (6)$$

Let $j = \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix}$. The Lie algebra of H_κ is

$$\mathfrak{h}_\kappa = \{\text{diag}(r_\kappa(t), sj) \mid s, t \in \mathbb{R}\},$$

where $r_\kappa(t) = \begin{pmatrix} 0 & -\kappa t \\ t & 0 \end{pmatrix}$. We may identify \mathcal{L}_κ with G_κ/H_κ via the diffeomorphism

$$\phi : G_\kappa/H_\kappa \rightarrow \mathcal{L}_\kappa, \quad \phi(gH_\kappa) = g \cdot [\gamma_o]. \quad (7)$$

For $x, y \in \mathbb{R}^2$ we denote $Z(x, y) = \begin{pmatrix} 0_2 & (-\kappa x, -y)^t \\ (x, y) & 0_2 \end{pmatrix}$. Let

$$\mathfrak{p}_\kappa = \{Z(x, y) \in \mathfrak{g}_\kappa \mid x, y \in \mathbb{R}^2\},$$

which is an $\text{Ad}(H_\kappa)$ -invariant complement of \mathfrak{h}_κ .

For $\kappa = 0, 1$, we consider on \mathfrak{g}_κ the inner product such that $\mathfrak{h}_\kappa \perp \mathfrak{p}_\kappa$, $\|Z(x, y)\| = \det(x, y)$ and

$$\|\text{diag}(r_\kappa(t), sj)\| = -ts.$$

(for $\kappa = 0$, we have learnt of this inner product from [6, page 499]). On \mathfrak{g}_{-1} we consider the Killing form ($\mathfrak{h}_\kappa \perp \mathfrak{p}_\kappa$ also holds). For $\kappa = 0, 1, -1$, this inner product on \mathfrak{g}_κ induces on G_κ a bi-invariant metric. Thus, there exists a unique pseudo-Riemannian metric on $\mathcal{L}_\kappa \simeq G_\kappa/H_\kappa$ such that $\pi : G_\kappa \rightarrow G_\kappa/H_\kappa$ is a pseudo-Riemannian submersion. For $\kappa = 0, 1$, this metric on \mathcal{L}_κ coincides with the given in (2), see Lemma 5 b). For $\kappa = -1$, the metric on \mathcal{L}_{-1} associated with the Killing form is different from the one in (2).

However, the magnetic geodesics of either metric on \mathcal{L}_{-1} are the same. This follows since the geodesics are the same (see [8]), so the Levi-Civita connections coincide.

Let us call $A = \text{diag}(0_2, j)$, which is in the center of \mathfrak{h}_κ . We have that ad_A is orthogonal and $\text{ad}_A^2 = -\text{id}$ in \mathfrak{p}_κ . Hence, ad_A induces a complex structure on G_κ/H_κ . A straightforward computation shows that it coincides, via ϕ in (7), with the complex structure given in the introduction. With the metric above and this complex structure, \mathcal{L}_κ is a Hermitian symmetric space.

As a direct application of a result by Adachi, Maeda and Udagawa in [1] (see also [2], Remark 1) we have

Theorem 4 *Let σ be a magnetic geodesic of G_κ/H_κ with initial conditions $\sigma(0) = H_\kappa$ and $\dot{\sigma}(0) = X \in \mathfrak{p}_\kappa$. Then $\sigma(t) = \pi(\exp t(X + A))$.*

As we saw in (1), \mathcal{J}_{γ_o} is isomorphic to $T_{[\gamma_o]}\mathcal{L}_\kappa \cong \mathfrak{p}_\kappa$. In the next Lemma we relate \mathfrak{p}_κ and \mathcal{J}_{γ_o} explicitly, involving the matrix A .

Lemma 5 *Let $Z = Z(x, y) \in \mathfrak{p}_\kappa$.*

a) *The Jacobi field $J(s) = \frac{d}{dt}\Big|_0 \exp t(Z + A) \cdot \gamma_o(s)$ in \mathcal{J}_{γ_o} is the unique one that satisfies $J(0) = (0, 0, x)^t$ and $J'(0) = (0, 0, y)^t$.*

b) *$T_{\gamma_o}(J) = d(\phi \circ \pi)Z$ and its norm is $\|d(\phi \circ \pi)Z\| = \det(x, y)$.*

Proof. For each κ , we consider the following parameterization of γ_o :

$$\begin{aligned} \gamma_o(s) &= (1, s, 0, 0), & \text{if } \kappa = 0; \\ \gamma_o(s) &= (\cos s, \sin s, 0, 0), & \text{if } \kappa = 1; \\ \gamma_o(s) &= (\cosh s, \sinh s, 0, 0), & \text{if } \kappa = -1. \end{aligned}$$

Given $Z = Z(x, y) \in \mathfrak{p}_\kappa$, the Jacobi field along γ_o defined by $J(s) = \frac{d}{dt}\Big|_0 \exp t(Z + A) \cdot \gamma_o(s)$ belongs to \mathcal{J}_{γ_o} , because for all $s \in \mathbb{R}$,

$$\langle J(s), \dot{\gamma}_o(s) \rangle = \langle (Z + A)(\gamma_o(s)), \dot{\gamma}_o(s) \rangle = 0,$$

since $(Z + A)(\gamma_o(s))$ is orthogonal to e_0 and e_1 , while $\dot{\gamma}_o(s)$ has non zero components only in these two directions.

One verifies easily that $J(0) = (Z + A)(e_0) = (0, 0, x)^t$. On the other hand,

$$\begin{aligned} J'(0) &= \frac{D}{ds}\Big|_0 \frac{\partial}{\partial t}\Big|_0 \exp t(Z + A) \cdot \gamma_o(s) \\ &= \frac{D}{\partial t}\Big|_0 \exp t(Z + A)(e_1) = (Z + A)(e_1) = (0, 0, y)^t. \end{aligned}$$

Besides,

$$\begin{aligned} T_{\gamma_o}(J) &= \frac{d}{dt}\Big|_0 [\exp t(Z + A) \cdot \gamma_o] = \frac{d}{dt}\Big|_0 \phi(\exp t(Z + A)H_\kappa) \\ &= \frac{d}{dt}\Big|_0 \phi(\pi(\exp t(Z + A))) = d\phi \circ d\pi Z, \end{aligned}$$

where the last equality holds since $A \in \mathfrak{h}_\kappa$. Finally, the norm (2) of $d(\phi \circ \pi)Z$ equals

$$\|d(\phi \circ \pi)Z\| = \langle \dot{\gamma}_o(0) \times J(0), J'(0) \rangle = \det(x, y)$$

and the assertions of b) are verified. \square

Let $Z(x, y) \in \mathfrak{p}_\kappa$ and let $h = \text{diag}(R_\kappa(t), B) \in H_\kappa$, where $B \in SO_2$ and

$$R_\kappa(t) = \begin{pmatrix} c_\kappa(t) & -\kappa s_\kappa(t) \\ s_\kappa(t) & c_\kappa(t) \end{pmatrix}$$

is as in (6). Then $\text{Ad}(h)Z(x, y) = Z(Bx_t, By_t)$, where

$$x_t = c_\kappa(t)x - s_\kappa(t)y, \quad y_t = \kappa s_\kappa(t)x + c_\kappa(t)y.$$

We denote by ϵ_1 and ϵ_2 the vectors of the canonical basis of \mathbb{R}^2 .

Lemma 6 *Let $Z(x, y) \neq 0$ in \mathfrak{p}_κ .*

- a) *If $\{x, y\}$ is a linearly independent set of \mathbb{R}^2 , then there exists $h \in H_\kappa$ such that $\text{Ad}(h)Z(x, y) = Z(a\epsilon_1, b\epsilon_2)$, with $a > 0$ and $b \neq 0$, for $\kappa = 0, \pm 1$.*
- b) *If $\kappa = 0, 1$ and $\{x, y\}$ is a linearly dependent set of \mathbb{R}^2 , then there exists $h \in H_\kappa$ such that either $\text{Ad}(h)Z(x, y) = Z(0, b\epsilon_2)$, with $b \neq 0$, or $\text{Ad}(h)Z(x, y) = Z(a\epsilon_1, 0)$, with $a > 0$. This is true for $\kappa = -1$ if in addition $|x| \neq |y|$.*
- c) *For $\kappa = 1$, there exists $h \in H_\kappa$ such that $\text{Ad}(h)Z(\epsilon_1, 0) = Z(0, \epsilon_2)$.*

Proof. For the proof of a), as $\{x, y\}$ is a linearly independent set, then for $\kappa = 0, \pm 1$ there exists $t \in \mathbb{R}$ such that $\langle x_t, y_t \rangle = 0$. Indeed, for each κ , this is equivalent to fact that the equation

$$\begin{aligned} c_3 - c_2 t &= 0 && \text{if } \kappa = 0; \\ \frac{1}{2}(c_1 - c_2) \sin(2t) + c_3 \cos(2t) &= 0 && \text{if } \kappa = 1; \\ -\frac{1}{2}(c_1 + c_2) \sinh(2t) + c_3 \cosh(2t) &= 0 && \text{if } \kappa = -1 \end{aligned}$$

has a real solution, where $c_1 = \langle x, x \rangle$, $c_2 = \langle y, y \rangle$ and $c_3 = \langle x, y \rangle$. But the linear independence of x and y determines the existence of the solution in each case. Then, we can take $B \in SO_2$ such that $Bx_t = a\epsilon_1$, with $a > 0$ and $By_t = b\epsilon_2$, with $b \neq 0$. Therefore the isometry $h = \text{diag}(R_\kappa(t), B) \in H_\kappa$ satisfies $\text{Ad}(h)Z(x, y) = Z(a\epsilon_1, b\epsilon_2)$.

For the proof of b), first we suppose that $x = 0$ or $y = 0$ (but not both zero since $Z(x, y) \neq 0$). Let $B \in SO_2$ such that $Bx = a\epsilon_1$ with $a > 0$, if $x \neq 0$, and in the case that $y \neq 0$, let $B \in SO_2$ such that $By = b\epsilon_2$, with $b \neq 0$. Then we can take $h = \text{diag}(I, B) \in H_\kappa$.

Now, let $x \neq 0$ and $y \neq 0$. So $x = \lambda y$ or $y = \lambda x$, with $\lambda \neq 0$. We suppose that $y = \lambda x$ (for $x = \lambda y$ the argument is similar). In the cases $\kappa = 0, 1$ there exists $t \in \mathbb{R}$ such that $x_t = 0$. In fact, from the hypothesis and some computations, $t \in \mathbb{R}$ is obtained by solving

$$1 - \lambda t = 0, \quad \text{if } \kappa = 0 \quad \text{and} \quad \cos t - \lambda \sin t = 0, \quad \text{if } \kappa = 1.$$

Thus, taking $B \in SO_2$ such that $By_t = b\epsilon_2$ (with $b \neq 0$ as $y_t \neq 0$), we have that $h = \text{diag}(R_\kappa(t), B) \in H_\kappa$ satisfies $\text{Ad}(h)Z(x, y) = Z(0, b\epsilon_2)$.

For $\kappa = -1$, as in the cases $\kappa = 0, 1$, we find $t \in \mathbb{R}$ such that either $x_t = 0$ or $y_t = 0$ by solving

$$\cosh t - \lambda \sinh t = 0, \quad \text{and} \quad -\sinh t + \lambda \cosh t = 0,$$

respectively. But these equations have a solution if and only if $\lambda \neq \pm 1$. That is, if and only if $|x| \neq |y|$. Hence, taking $B \in SO_2$ such that either $By_t = b\epsilon_2$ or $Bx_t = a\epsilon_1$ (with $a > 0$; here again we have that $x_t \neq 0$), as appropriate. Then $h = \text{diag}(R_{-1}(t), B) \in H_{-1}$ is as desired in this case.

For part c), we observe that $h = \text{diag}(R_1(\pi/2), B) \in H_1$, where $B \in SO_2$ takes ϵ_1 to ϵ_2 , satisfies $\text{Ad}(h)Z(\epsilon_1, 0) = Z(0, \epsilon_2)$. \square

Remark. The previous lemma corresponds, geometrically, with the fact of finding $s \in \mathbb{R}$ at which the Jacobi field associated with $Z(x, y)$ (given by Lemma 5) and its covariant derivative are orthogonal.

Recall that if h is a regular curve in M_κ of constant speed a , then the Frenet frame of h is

$$T(t) = \frac{1}{a} \dot{h}(t), \quad N(t) = \dot{h}'(t) / \left| \dot{h}'(t) \right|, \quad B(t) = T(t) \times N(t) \quad (8)$$

(here the prime denotes the covariant derivative along h), and its curvature and torsion are given by

$$k(t) = \frac{1}{a^2} \left| \dot{h}'(t) \right|, \quad \tau(t) = -\frac{1}{a} \langle B'(t), N(t) \rangle. \quad (9)$$

For each $g \in G_\kappa$ we have that g is an isometry of \mathcal{L}_κ and preserves the Hermitian structure. Hence, g takes magnetic geodesics to magnetic geodesics.

3 Time- and space-like magnetic geodesics

Proof of Theorem 1. Let $Z \in \mathfrak{p}_\kappa$ be the initial velocity of σ , with $\|Z\| \neq 0$. First, we consider the case $Z = Z(a\epsilon_1, b\epsilon_2)$, with $a > 0$ and $b \neq 0$.

For each $t \in \mathbb{R}$, let $\alpha(t) = \exp t(Z + A)$. By Theorem 4 and the diffeomorphism ϕ in (7), we know that $\sigma(t) = \alpha(t) \cdot [\gamma_o]$, that is, $\sigma(t) = [\alpha(t) \cdot \gamma_o]$.

Let h be the curve in M_κ given by $h(t) = \alpha(t)(e_0)$. As α is a one-parameter subgroup of isometries of M_κ , we have that h is a curve with constant curvature and torsion, thus h is a helix in M_κ .

Let us see that $\sigma(t) = [\gamma_{B(t)}]$, where $B(t)$ is the binormal field of h . For each $t \in \mathbb{R}$, the initial velocity of the geodesic $\alpha(t) \cdot \gamma_o$ is $d(\alpha(t))(e_1)$, hence $\sigma(t) = [\gamma_{d(\alpha(t))(e_1)}]$. Then, we have to verify that $B(t) = d(\alpha(t))(e_1)$, for all $t \in \mathbb{R}$. Since $\alpha(t)$ is an isometry that preserves the helix and takes the Frenet frame at $t = 0$ to the Frenet frame at t , it suffices to show that $B(0) = e_1$.

By the usual identifications, since $\alpha(t)$ is a linear transformation, we can write $d(\alpha(t))(e_1) = \alpha(t)(e_1)$, so

$$\dot{h}(t) = \alpha(t)((Z + A)e_0) \quad \text{and} \quad \dot{h}'(t) = [\alpha(t)((Z + A)^2 e_0)]^T,$$

where T denotes the tangent projection. Since

$$\dot{h}(0) = (Z + A)e_0 = ae_2,$$

$$\dot{h}'(0) = [(Z + A)^2 e_0]^T = [-\kappa a^2 e_0 + ae_3]^T = ae_3$$

and $\alpha(t)$ is an isometry, we have $|\dot{h}(t)| = a = |\dot{h}'(t)|$. By the computation before and (8) we obtain

$$B(0) = \frac{1}{a^2} \dot{h}(0) \times \dot{h}'(0) = e_1.$$

Since $B'(t) = [\alpha(t)((Z + A)e_1)]^T$ (recall that we have just proven that $B(t) = \alpha(t)(e_1)$), then $B'(0) = be_3$. Besides, using (8) and the previous computations it follows that $N(0) = e_3$. Therefore, by (9) we have that the curvature and torsion of h are equal to

$$k = 1/a, \quad \tau = -b/a. \quad (10)$$

The assertion regarding the sign of the torsion is immediate from Lemma 5 b) and (10). Thus, the theorem is proved in this particular case.

Now, let σ be a magnetic geodesic with $\sigma(0) = [\gamma]$ and initial velocity with non zero norm. Since G_κ acts transitively on \mathcal{L}_κ , there is an isometry g such that $g \cdot [\gamma] = [\gamma_0]$. So, the magnetic geodesic $g \cdot \sigma$ also has initial velocity with non zero norm and $g \cdot \sigma(0) = [\gamma_0]$. By Lemma 5 b), if $d(\phi \circ \pi)Z(x, y)$ is the initial velocity of $g \cdot \sigma$, we have that the vectors $\{x, y\}$ are linearly independent. Then, by Lemma 6 a), there exists $h \in H_\kappa$ such that $\text{Ad}(h)Z(x, y) = Z(ae_1, be_2)$, with $a > 0$ and $b \neq 0$. Since $((h \circ g) \cdot \sigma)'(0) = d(\phi \circ \pi)(\text{Ad}(h)Z(x, y))$, the curve $(h \circ g) \cdot \sigma$ is a magnetic geodesic of the type studied above. Therefore, σ has the form (3).

Conversely, let h be a helix in M_κ with curvature $k > 0$, non zero torsion τ and speed $1/k$. Let $\{T, B, N\}$ be the Frenet frame of h . As M_κ is a simply connected manifold of constant curvature, we have that there exists an isometry g of M_κ preserving the orientation such that $g(h(0)) = e_0$ and its differential at $h(0)$ takes $B(0)$ to e_1 , $T(0)$ to e_2 and $N(0)$ to e_3 .

Let $a = 1/k$ and $b = -\tau/k$. Let $Z = Z(ae_1, be_2) \in \mathfrak{p}_\kappa$. We consider, for each $t \in \mathbb{R}$, $\alpha(t) = \exp t(Z + A)$. According to computations from the first part of the proof, both helices have initial position e_0 , curvature k , torsion τ , speed $1/k$ and the same Frenet frame at $t = 0$. Hence $(g \circ h)(t) = \alpha(t)e_0$. So, if we call \bar{B} the binormal field of $g \circ h$, we have that $\bar{B}(t) = d(\alpha(t))e_1$, for all t . Finally, since the curve $[\gamma_{\bar{B}(t)}]$ is a magnetic geodesic in \mathcal{L}_κ and

$$[\gamma_{B(t)}] = [\gamma_{dg^{-1}\bar{B}(t)}] = g^{-1} \cdot [\gamma_{\bar{B}(t)}],$$

we obtain that $[\gamma_{B(t)}]$ is a magnetic geodesic. □

4 Null magnetic geodesics

We deal first with the hyperbolic case. We use the notation given in the introduction and we recall from [3] certain properties of horospheres and related concepts. To simplify the notation we omit the subindex $\kappa = -1$.

Let γ be a geodesic of \mathbb{H}^3 . Then, for each $p \in \mathbb{H}^3$ there exists a unique unit speed geodesic α of \mathbb{H}^3 such that $\alpha(0) = p$ and α is asymptotic to γ . Let $v \in T^1\mathbb{H}^3$. If p is any point of \mathbb{H}^3 , then $v(p)$ denotes the unique unit tangent vector at p that is asymptotic to v . The Busemann function $f_v : \mathbb{H}^3 \rightarrow \mathbb{R}$ is defined by

$$f_v(p) = \lim_{t \rightarrow +\infty} d(p, \gamma_v(t)) - t,$$

and satisfies $\text{grad}_p(f_v) = -v(p)$. The *horosphere* determined by v is given by

$$H(v) = \{q \in M : f_v(q) = 0\}.$$

The Jacobi vector fields orthogonal to $\dot{\gamma}_o$ have the form

$$J(s) = e^s U(s) + e^{-s} V(s), \quad (11)$$

where U and V are parallel vector fields along γ_o and orthogonal to $\dot{\gamma}_o$.

A Jacobi vector field Y along a geodesic γ of \mathbb{H}^3 is said to be *stable* (*unstable*) if there exists a constant $c > 0$ such that

$$|Y(s)| \leq c \quad \forall s \geq 0 \quad (\forall s \leq 0).$$

In what follows we shall denote by $\hat{\pi}$ the canonical projection from $T\mathbb{H}^3$ onto \mathbb{H}^3 . We recall that in the introduction we have defined the smooth maps $\psi^\pm : \mathcal{L}(\mathbb{H}^3) \rightarrow \mathbb{H}^3(\infty)$ by $\psi^\pm[\gamma] = \gamma(\pm\infty)$ and the distributions \mathcal{D}^\pm in $\mathcal{L}(\mathbb{H}^3)$ given by $\mathcal{D}_{[\gamma]}^\pm = \text{Ker}(d\psi_{[\gamma]}^\pm)$. We need to relate the distributions \mathcal{D}^\pm with distributions $\bar{\mathcal{E}}^\pm$ and \mathcal{E}^\pm on G and $T^1\mathbb{H}^3$, respectively.

Let $\bar{\mathcal{E}}^\pm$ be the left invariant distribution on G defined at $I \in G$ by

$$\bar{\mathcal{E}}_I^\pm = \{Z(u, \mp u) \in \mathfrak{p} \mid u \in \mathbb{R}^2\}.$$

As the canonical action of G on $T^1\mathbb{H}^3$ is transitive, the projection $\bar{p} : G \rightarrow T^1\mathbb{H}^3$ given by $\bar{p}(g) = dg_{e_0}e_1$ is a submersion. Since given $v \in T^1\mathbb{H}^3$ there exists $g \in G$ such that $\bar{p}(g) = v$, we define:

$$\mathcal{E}^\pm(v) = (d\bar{p} \bar{\mathcal{E}}^\pm)(\bar{p}(g)) = d\bar{p}_g(\bar{\mathcal{E}}_g^\pm).$$

We have that \mathcal{E}^\pm determines a well defined distribution on $T^1\mathbb{H}^3$, which is called the *horospherical distribution* on $T^1\mathbb{H}^3$. This distribution has the following property: if $t \mapsto v(t)$ is a curve in $T^1\mathbb{H}^3$ tangent to the distribution \mathcal{E}^\pm , then $\hat{\pi}(v(t))$ is in the horosphere $H(\pm v(0))$.

Lemma 7 *Let $Z \in \bar{\mathcal{E}}_I^\pm$. For each $t \in \mathbb{R}$, let $\gamma_t^\pm(s) = \exp t(Z + A) \cdot \gamma_o(\pm s)$. Then the geodesics γ_t^\pm are asymptotic to each other for all $t \in \mathbb{R}$.*

Proof. Let J be the Jacobi vector field associated with the variation by geodesics $t \mapsto \gamma_t^\pm$. By Lemma 5 a), $J(0) = -J'(0)$. Hence, by (11) we have that $J(s) = e^{-s}U(s)$, where U is a parallel vector field along γ_o orthogonal to $\dot{\gamma}_o$. Thus, J is a stable vector field, that is, there exists $c > 0$ such that $|J(s)| \leq c \forall s \geq 0$.

We have to show that given $t_0, t_1 \in \mathbb{R}$ with $t_0 < t_1$, there exists $N > 0$ such that

$$d(\gamma_{t_0}^\pm(s), \gamma_{t_1}^\pm(s)) \leq N \quad \forall s \geq 0.$$

For fixed s ,

$$d(\gamma_{t_0}^\pm(s), \gamma_{t_1}^\pm(s)) \leq \text{length}([t_0, t_1] \ni t \mapsto \gamma_t^\pm(s)) = \int_{t_0}^{t_1} \left| \frac{d}{dt} \gamma_t^\pm(s) \right| dt.$$

For each $t \in \mathbb{R}$, let $J_t(s) = \frac{d}{dt} \gamma_t^\pm(s)$. We observe that $J_{t'+t}(s) = d \exp(t'Z) J_t(s)$ for all t, t' . Since $\exp(t'Z)$ is an isometry, we have $|J_t(s)| = |J(s)|$. Therefore,

$$\int_{t_0}^{t_1} |J_t(s)| dt = \int_{t_0}^{t_1} |J(s)| dt \leq c(t_1 - t_0)$$

for all $s \geq 0$. Then, we may take $N = c(t_1 - t_0) > 0$. \square

We consider the projection $p : T^1\mathbb{H}^3 \rightarrow \mathcal{L}(\mathbb{H}^3)$, $p(v) = [\gamma_v]$. We call $\bar{\mathcal{D}}^\pm$ the distribution on $\mathcal{L}(\mathbb{H}^3)$ p -related with \mathcal{E}^\pm (well defined). More specifically, given $[\gamma] \in \mathcal{L}(\mathbb{H}^3)$ and $v \in T^1\mathbb{H}^3$ such that $p(v) = [\gamma]$,

$$\bar{\mathcal{D}}^\pm([\gamma]) = dp_v \mathcal{E}_v^\pm.$$

Proposition 8 *Let \mathcal{D}^\pm and $\bar{\mathcal{D}}^\pm$ be the distributions on $\mathcal{L}(\mathbb{H}^3)$ defined above. Then $\mathcal{D}^\pm = \bar{\mathcal{D}}^\pm$.*

Proof. Since \mathcal{D}^\pm and $\bar{\mathcal{D}}^\pm$ are G -invariant, it is enough to show $\mathcal{D}_{[\gamma_o]}^\pm = dp_{(e_0, e_1)}(\mathcal{E}_{(e_0, e_1)}^\pm)$ (we observe that $\bar{p}(I) = (e_0, e_1)$ and $p(e_0, e_1) = [\gamma_o]$).

Let $Z \in \bar{\mathcal{E}}_I^\pm$. We take the curve in $\mathcal{L}(\mathbb{H}^3)$ given by $\alpha(t) = \exp tZ \cdot [\gamma_o]$. As $\alpha(t) = p \circ \bar{p}(\exp tZ)$, we have that $\alpha(0) = [\gamma_o]$ and $\dot{\alpha}(0) = d(p \circ \bar{p})_I Z$. That is, $\dot{\alpha}(0) \in dp_{(e_0, e_1)}(\mathcal{E}_{(e_0, e_1)}^\pm)$. Besides,

$$\left. \frac{d}{dt} \right|_0 \exp tZ \cdot \gamma_o(s) = \left. \frac{d}{dt} \right|_0 \exp t(Z + A) \cdot \gamma_o(s), \quad (12)$$

since both Jacobi fields have the same initial conditions. Hence, Lemma 7 applies to the geodesics $\gamma_t^\pm(s) = \exp tZ \cdot \gamma_o(\pm s)$. Thus, $\psi^\pm \circ \alpha$ is constant. Then $(d\psi^\pm)_{[\gamma_o]}(\dot{\alpha}(0)) = 0$, that is, $\dot{\alpha}(0) \in \mathcal{D}_{[\gamma_o]}^\pm$.

On the other hand, let $\varphi : T_{e_0}^1\mathbb{H}^3 \rightarrow \mathcal{L}(\mathbb{H}^3)$, $\varphi(v) = [\gamma_v]$, be the submanifold whose image $\mathcal{L}_{e_0}(\mathbb{H}^3)$ consists of all the oriented geodesics passing through e_0 . Besides, $H(\infty)$ is a manifold with the differentiable structure (well defined) such that $F_{e_0} : T_{e_0}^1\mathbb{H}^3 \rightarrow$

$H(\infty)$ given by $F_{e_0}(v) = \gamma_v(\infty)$ is a diffeomorphism. Then, since $\psi^+|_{\mathcal{L}_{e_0}(\mathbb{H}^3)} \circ \varphi = F_{e_0}$, we have that $(d\psi^+)_{[\gamma_o]}$ is surjective. Now, $(d\psi^-)_{[\gamma_o]}$ is also surjective because ψ^- is the composition of ψ^+ with the diffeomorphism of $\mathcal{L}(\mathbb{H}^3)$ assigning $[\gamma^{-1}]$ to $[\gamma]$. Therefore, $\dim \mathcal{D}_{[\gamma_o]}^\pm = \dim \bar{\mathcal{D}}_{[\gamma_o]}^\pm$ and equality follows. \square

The word *cylinder* in the statement of Theorem 2 refers to a ruled surface determined by a parallel vector field along a curve c of constant geodesic curvature k contained in a totally geodesic surface in M_κ (and normal to it), as explained. For $\kappa = -1$, this ruled surface is diffeomorphic to $S^1 \times \mathbb{R}$ if $|k| > 1$; otherwise it is diffeomorphic to a plane.

Proof of Theorem 2 a). By Lemma 5 b), we have that every element of $\mathcal{D}_{[\gamma]}^\pm$ is null. As G acts transitively on $\mathcal{L}(\mathbb{H}^3)$ and by the G -invariance of the horospherical distributions, we may suppose without loss of generality that $\sigma(0) = [\gamma_o]$, hence $\dot{\sigma}(0) \in \mathcal{D}_{[\gamma_o]}^\pm$. By Proposition 8, there exists $Z \in \bar{\mathcal{E}}_I^\pm$ such that $\dot{\sigma}(0) = (dp)_{(e_0, e_1)}(d\bar{p})_I Z$. Thus, by Theorem 4, $\sigma(t) = [\exp t(Z + A) \cdot \gamma_o]$.

We assume that $Z \in \bar{\mathcal{E}}_I^+$. Let us show that σ describes a forward cone with vertex at $\gamma_o(+\infty)$. In a similar way, if $Z \in \bar{\mathcal{E}}_I^-$, then σ describes a backward cone with vertex at $\gamma_o(-\infty)$.

We consider the geodesics $\gamma_t(s) = \exp t(Z + A) \cdot \gamma_o(s)$ of \mathbb{H}^3 . As $Z \in \bar{\mathcal{E}}_I^+$, by Lemma 7, we have that the geodesics γ_t are asymptotic to each other for all t . Hence, $z(t) = \dot{\gamma}_t(0)$ is a curve in $T^1\mathbb{H}^3$ of asymptotic vectors to e_1 .

Let $c(t) = \hat{\pi}(z(t)) = \exp t(Z + A)(e_0)$. In order to see that $c(t) \in H(e_1)$ for all t , we observe that

$$\frac{d}{dt} f_{e_1}(c(t)) = (df_{e_1})_{c(t)} \dot{c}(t) = \langle \text{grad}_{c(t)}(f_{e_1}), \dot{c}(t) \rangle. \quad (13)$$

Since $\text{grad}_p(f_v) = -v(p)$ we have that

$$\text{grad}_{c(t)}(f_{e_1}) = -z(t) = -d(\exp t(Z + A))e_1.$$

On the other hand,

$$\dot{c}(t) = d(\exp t(Z + A))(Z + A)e_0.$$

Since $\exp t(Z + A)$ is an isometry and observing that $(Z + A)e_0$ and e_1 are perpendicular ($Z \in \bar{\mathcal{E}}_I^+$), it follows that the expression in (13) is equal to $-\langle e_1, (Z + A)(e_0) \rangle = 0$. Then, $f_{e_1}(c(t)) = f_{e_1}(e_0) = 0$ for all t , that is, $c(t) \in H(e_1)$ for all t .

Now, as c is the orbit through e_0 of a one-parameter subgroup of isometries of G preserving $H(e_1)$, its geodesic curvature and speed are constant. If $Z = Z(u, -u)$ for certain $0 \neq u \in \mathbb{R}^2$, we obtain that the speed of c is $|u|$. For each $v \in T^1H^3$ we consider on $H(v)$ the orientation given by $-\text{grad } f_v$. The geodesic curvature of c is then

$$k = \langle -\text{grad}_{e_0}(f_{e_1}), \dot{c}(0) \times \dot{c}'(0) \rangle / |u|^3 = 1/|u|,$$

since $\dot{c}(0) = (Z + A)e_0$ and $\dot{c}'(0) = ((Z + A)^2 e_0)^\top$. As for each $v \in T^1\mathbb{H}^3$, $H(v)$, with the induced metric of \mathbb{H}^3 , is isometric to \mathbb{R}^2 , we have that $c(t)$ runs along a circle on $H(e_1)$ of geodesic curvature $k = 1/|u| > 0$ and speed $1/k = |u|$.

Besides, $\sigma(t) = [\gamma_{z(t)}]$. Thus we have that all conditions are satisfied in order to assert that σ describes a forward cone with vertex at $\gamma_o(+\infty)$.

Conversely, let σ be a curve in $\mathcal{L}(\mathbb{H}^3)$ that describes a forward cone with vertex at infinity. As G acts transitively on the positively oriented frame bundle, and also each element of G takes horospheres to horospheres, preserving their orientation, we may suppose that $\sigma(t) = [\gamma_{v(t)}]$, where $v(t)$ is a curve in $T^1\mathbb{H}^3$ of asymptotic vectors to $v(0) = e_1$ and $c(t) = \hat{\pi}(v(t))$ is a curve of geodesic curvature k and speed $1/k$ in $H(e_1)$ with $\dot{c}(0) = \frac{1}{k}e_2$, for some $k > 0$. Let $Z = Z(\frac{1}{k}\epsilon_1, -\frac{1}{k}\epsilon_1) \in \bar{\mathcal{E}}_I^+$. We define

$$\bar{c}(t) = \exp t(Z + A)(e_0) \quad \text{and} \quad \bar{v}(t) = d(\exp t(Z + A))(e_1).$$

We showed above that $\bar{c}(t)$ is a curve of geodesic curvature k and speed $\frac{1}{k}$ in $H(e_1)$. Moreover, $\bar{c}(0) = e_0$ and the initial velocity of \bar{c} is $\frac{1}{k}e_2$. So, we obtain that $\bar{c} = c$. This implies, together with the identities $\hat{\pi} \circ \bar{v} = \bar{c}$ and $\hat{\pi} \circ v = c$, that $\hat{\pi} \circ \bar{v} = \hat{\pi} \circ v$.

According to the first part of the proof, \bar{v} and v are curves of asymptotic vectors to e_1 . Hence, $-\bar{v}(t) = \text{grad}_{\bar{c}(t)}(f_{e_1}) = -v(t)$. Therefore, $[\gamma_{\bar{v}(t)}] = [\gamma_{v(t)}]$, which is a null magnetic geodesic with initial velocity in the horospherical distribution since $[\gamma_{v(t)}] = [\exp t(Z + A) \cdot \gamma_o]$. \square

Proof of Theorem 2 b). We suppose first that σ is a null magnetic geodesic such that $\sigma(0) = [\gamma_o]$ and $\dot{\sigma}(0) = d(\phi \circ \pi)Z(a\epsilon_1, 0)$, with $a > 0$. The expression (4) and the relation between the speed and curvature of h are obtained as in the prove of Theorem 1. By (10) we know that the torsion of h is $\tau = -b/a = 0$ (since $b = 0$). Thus h is contained in a totally geodesic surface S of \mathbb{H}^3 and B is normal to S .

Now, we suppose that $\dot{\sigma}(0) = d(\phi \circ \pi)Z$, where $Z = Z(0, b\epsilon_2)$ with $b \neq 0$. By Theorem 4 we have that $\sigma(t) = [\alpha(t) \cdot \gamma_o]$, where $\alpha(t) = \exp t(Z + A)$. Since $Z + A$ is in the Lie algebra of the isotropy subgroup H of G at $e_0 \in \mathbb{H}^3$, we get that $\alpha(t)$ fixes e_0 . Moreover, if v is the curve in $T_{e_0}^1\mathbb{H}^3$ given by $v(t) = d(\alpha(t))e_1$, then

$$\sigma(t) = [\alpha(t) \cdot \gamma_o] = [\gamma_{v(t)}],$$

since the initial velocity of the geodesic $\alpha(t) \cdot \gamma_o$ is $v(t)$, for each $t \in \mathbb{R}$.

Furthermore, as v is the orbit through e_1 of a one-parameter subgroup of H (the canonical differential action of G on $T_{e_0}^1\mathbb{H}^3$), then v has constant speed and constant geodesic curvature in $T_{e_0}^1\mathbb{H}^3 \cong \mathbb{S}^2$. Easy computations yield

$$\dot{v}(0) = (0, 0, b)^t \quad \text{and} \quad \ddot{v}(0) = (-b^2, -b, 0)^t.$$

So, the speed of v is $|b|$ and its geodesic curvature is

$$k = \langle v(0), \dot{v}(0) \times \ddot{v}(0) \rangle / |b|^3 = 1/|b|$$

(we consider the orientation of the sphere given by the unit normal field pointing outwards). Thus, v is a curve in $T_{e_0}^1\mathbb{H}^3$ of geodesic curvature $k > 0$ and speed $1/k$. Consequently, σ has the form (5).

Now, let σ be a null magnetic geodesic such that $\sigma(0) = [\gamma]$ and $\dot{\sigma}(0) \notin \mathcal{D}_{[\gamma]}^\pm$. As G acts transitively on $\mathcal{L}(\mathbb{H}^3)$ and by the G -invariance of the horospherical distributions, we may suppose that $\sigma(0) = [\gamma_o]$ and $\dot{\sigma}(0) \notin \mathcal{D}_{[\gamma_o]}^\pm$. Let $Z = Z(x, y) \in \mathfrak{p}$ such that $\dot{\sigma}(0) = d(\phi \circ \pi)Z$. By Lemma 5 b), as the norm of the initial velocity of σ is zero, we have that x and y are linearly dependent, and since $d(\phi \circ \pi)Z \notin \mathcal{D}_{[\gamma_o]}^\pm$, we also have $|x| \neq |y|$. Now, the isometries in Lemma 6 b) take σ to magnetic geodesics of the particular types studied above. Therefore, σ has the form (4) or has the form (5), as desired.

Conversely, given a helix h in \mathbb{H}^3 with curvature k , speed $1/k$ and torsion $\tau = 0$, the proof that the expression (4) is a magnetic geodesic is identical to the proof of the converse of Theorem 1. As h has zero torsion, the initial velocity of the magnetic geodesic in (4) is not in the distributions \mathcal{D}^\pm .

Let v be a curve in $T_p^1\mathbb{H}^3$ with geodesic curvature $k > 0$ and speed $1/k$. Let g be the isometry of \mathbb{H}^3 preserving the orientation such that $g(p) = e_0$, $dg(v(0)) = e_1$ and $dg(\dot{v}(0)) = be_3$, for certain $b > 0$. Hence, $g \cdot v$ is a curve in $T_{e_0}^1\mathbb{H}^3$ having the same geodesic curvature and the same speed as v , and also $b = 1/k$. As we showed above, \bar{v} is a curve in $T_{e_0}^1\mathbb{H}^3$ with $\bar{v}(0) = g \cdot v(0)$ and with the same initial velocity and geodesic curvature that $g \cdot v$. By uniqueness, we have that $\bar{v} = g \cdot v$. To complete the proof we observe that $g \cdot [\gamma_{v(t)}] = [\gamma_{g \cdot v(t)}] = [\gamma_{\bar{v}(t)}]$. \square

Proof of Theorem 3. Lemma 6 b) implies that the analogue of Theorem 2 a) is empty for the cases $\kappa = 0, 1$. The proof of the fact that every curve σ in \mathcal{L}_κ is a null magnetic geodesic if and only if σ has the form (4) or (5) is similar to that of Theorem 2 b).

We check the last statement of the theorem. Without loss of generality, we consider only null magnetic geodesics passing through $[\gamma_o]$ at $t = 0$. We observe that if, in particular, σ is a magnetic geodesic with initial velocity $d(\phi \circ \pi)Z(ae_1, 0)$, with $a > 0$, (that is, σ has the form (4)), then by Lemma 6 c) there exists $h \in H_1$ such that $\text{Ad}(h)Z(ae_1, 0) = Z(0, ae_2)$. Hence, $h \cdot \sigma$ is a null magnetic geodesic with initial velocity $d(\phi \circ \pi)Z(0, ae_2)$, and then it has the form (5). So, σ also has this form. \square

References

- [1] T. Adachi, S. Maeda, S. Udagawa: *Simpleness and closedness of circles in compact Hermitian symmetric spaces*, Tsukuba J. Math. **24** (2000), 1–13.
- [2] A. Bolsinov and B. Jovanovic: *Magnetic flows on homogeneous spaces*, Comment. Math. Helv. **83** (2008), 679–700.
- [3] P. Eberlein: *Geometry of nonpositively curved manifolds*, Chicago Lectures in Mathematics, 1996.

- [4] N. Georgiou and B. Guilfoyle: *On the space of oriented geodesics of hyperbolic 3-space*, Rocky Mountain J. Math. **40** (2010), 1183–1219.
- [5] B. Guilfoyle and W. Klingenberg: *An indefinite Kähler metric on the space of oriented lines*, J. London Math. Soc. **72** (2005), 497–509.
- [6] M. Kapovich and J. Millson: *The symplectic geometry of polygons in Euclidean space*, J. Diff. Geom. **44** (1996), 479–513.
- [7] M. Salvai: *On the geometry of the space of oriented lines of Euclidean space*, Manuscr. Math. **118** (2005), 181–189.
- [8] M. Salvai: *On the geometry of the space of oriented lines of the hyperbolic space*, Glasgow Math. J. **49** (2007), 357–366.
- [9] M. Salvai: *Global smooth fibrations of \mathbb{R}^3 by oriented lines*, Bull. London Math. Soc. **41** (2009), 155–163.

FaMAF - CIEM, Ciudad Universitaria, 5000 Córdoba, Argentina
ygodoy@famaf.unc.edu.ar, salvai@famaf.unc.edu.ar