

1 **CO₂ bubble generation and migration during magma-carbonate**
2 **interaction**

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17 **1. Abstract**

18 **We conducted quantitative textural analysis of vesicles in high temperature and**
19 **pressure carbonate assimilation experiments (1200°C, 0.5 GPa) to investigate CO₂**
20 **generation and subsequent bubble migration from carbonate into magma. We**
21 **employed Mt. Merapi (Indonesia) and Mt. Vesuvius (Italy) compositions as magmatic**
22 **starting materials and present three experimental series using i) a dry basaltic-**
23 **andesite, ii) a hydrous basaltic-andesite (2 wt. % H₂O) and iii) a hydrous shoshonite**
24 **(2 wt. % H₂O). The duration of the experiments was varied from 0 to 300 s and**
25 **carbonate assimilation produced a CO₂-rich fluid and CaO-enriched melts in all**
26 **cases. The rate of carbonate assimilation, however, changed as a function of melt**
27 **viscosity, which affected the 2D vesicle number, vesicle volume, and vesicle size-**
28 **distribution within each experiment. Relatively low viscosity melts (i.e. Vesuvius**
29 **experiments) facilitated efficient removal of bubbles from the reaction site. This**
30 **allowed carbonate assimilation to continue unhindered and large volumes of CO₂ to**
31 **be liberated, a scenario thought to fuel sustained CO₂-driven eruptions at the surface.**
32 **Conversely, at higher viscosity (i.e. Merapi experiments) bubble migration became**
33 **progressively inhibited and bubble concentration at the reaction site caused localised**
34 **volatile over-pressure that can eventually trigger short-lived explosive outbursts. Melt**
35 **viscosity therefore exerts a fundamental control on carbonate assimilation rates and,**
36 **by consequence, the style of CO₂-fuelled eruptions.**

37
38 **Keywords: CO₂, carbonate assimilation, melt viscosity, bubble size distribution, eruption**
39 **style**

40

41 2. Introduction

42 Earth's long-term volatile cycle reflects mass exchange between the deep Earth and the
43 atmosphere, with volcanoes representing the major agents of volatile transfer. When
44 magma migrates through crustal rocks *en route* from the mantle to the surface, it interacts
45 thermally, chemically, and mechanically with the surrounding rocks. If these rocks are
46 volatile-rich, such as limestone and dolostone which contain structurally-bound CO₂,
47 magma-crust interaction can liberate crustal volatiles into the magmatic system, where they
48 will eventually be released to the atmosphere via eruption or degassing.

49 While long-term magmatic degassing constantly delivers volatiles to the atmosphere,
50 magma-crust interaction, in turn, may drive fluctuating peaks of "extra" CO₂ sourced from
51 the crust. The latter process has been recognised in the recent volcanic gas record at
52 Merapi volcano, Indonesia (e.g. Troll et al. 2012) and at Vesuvius volcano, Italy (Iacono-
53 Marziano et al. 2009) and is thought to have played a major role in perturbations to the
54 global carbon cycle in Earth's past (Ganino and Arndt 2009; Svensen et al. 2009; Johnston
55 et al. 2011; Lee et al. 2013).

56 Indeed, it has been suggested that CO₂ emissions were several times greater than at
57 present in, e.g., the Cretaceous, which, by consequence, had a warmer climate (Johnston et
58 al. 2011; Lee et al. 2013). Excess CO₂ outgassing in the Cretaceous notably cannot be
59 explained by pure decarbonation of subducted materials alone, especially as this process is
60 thought to be rather inefficient based on thermodynamic models (Kerrick and Connolly
61 2001). Although less pronounced than in some earlier periods of Earth's history, an
62 imbalance still exists between the amount of CO₂ released during subduction and the
63 amount of CO₂ emitted at arc volcanoes at present (Dasgupta 2013). Carbonate-rich upper
64 crustal lithologies beneath arcs have therefore been considered a missing piece in the
65 carbon-cycle puzzle (cf. Johnston et al. 2011; Troll et al. 2012; Dasgupta 2013; Lee et al.
66 2013). In addition to the repercussions for the global carbon cycle on a geological time-

67 scale, magma-carbonate interaction can also contribute to driving short-lived explosive
68 eruptions at hazardous volcanoes besides Vesuvius and Merapi, including sites such as
69 Popocatépetl, Mexico and Yellowstone, USA (Goff et al. 2001; Werner and Brantley 2003;
70 Schaaf et al. 2005; Chadwick et al. 2007; Iacono Marziano et al. 2009; Deegan et al. 2010;
71 Dallai et al. 2011; Troll et al 2012, 2013; Borisova et al 2013). In this context, carbonate
72 assimilation causes the host magma to become progressively more potassic and silica
73 undersaturated, however, the simultaneously produced CO₂-rich fluid phase is less well
74 understood (e.g. Watkinson and Wyllie 1969; Freda et al. 2008; Gaeta et al. 2009; Mollo et
75 al. 2010). The dynamic aspects of CO₂ bubble nucleation, growth, and transport during
76 carbonate assimilation and their combined effects on volcanic activity have still to be fully
77 resolved.

78 To investigate the textural and chemical evolution of CO₂ bubbles during carbonate
79 assimilation, time-constrained, high pressure-temperature experiments at conditions
80 appropriate for mid to upper crustal magma reservoirs have been carried out. Starting
81 materials from Merapi and Vesuvius volcanoes were doped with carbonate fragments and
82 demonstrated both rapid production of CaO-enriched melts and CO₂ vesiculation in the
83 vicinity of the added carbonate (Deegan et al. 2010, 2011; Jolis et al. 2013). In order to
84 better comprehend the overall dynamics of the volcanic system during CO₂-fuelled
85 eruptions and to further assess the role of magma-crust interaction in the subduction carbon
86 cycle, it is imperative to quantify how CO₂-rich fluids are produced and mobilised in
87 crustal systems (cf. Freda et al. 1997, 2011; Johnston et al. 2011; Troll et al. 2012;
88 Dasgupta 2013; Jolis et al. 2013; Lee et al. 2013). We carried out analyses of vesicle
89 diameter, vesicle number, vesicle volume, and vesicle size-distribution on previously
90 characterised time-monitored carbonate assimilation experiments for Merapi and Vesuvius
91 volcanoes (Deegan et al. 2010; Jolis et al. 2013) and compare these experimental sets

92 directly for the first time. We then present a quantitative assessment of the controls on
93 vesicle generation and mobility during crustal carbonate assimilation.

94

95

96 **3. Geological Background**

97 Merapi and Vesuvius volcanoes are characterised by carbonate-hosted magmatic
98 plumbing systems, wherein magma-carbonate interaction has been identified through
99 isotopic studies of erupted minerals (e.g. feldspar, pyroxene, and olivine) and fumarolic
100 CO₂ (e.g. Chadwick et al. 2007; Iacono Marziano et al. 2009; Dallai et al. 2011; Troll et al.
101 2012, 2013; Borisova et al. 2013). Notably, both volcanoes also have a history of explosive
102 behaviour and are in proximity to major population centres, i.e., Yogyakarta in central Java
103 (3.5 M inhabitants) and Naples in central Italy (> 4 M inhabitants), making Merapi and
104 Vesuvius significant threats to modern society.

105 Merapi is situated on the Eurasian plate, beneath which the Indian ocean plate subducts
106 at a rate of 6 to 7 cm/yr (Tregoning et al. 1994). In central Java, crustal thickness is ~ 25
107 km (Curry et al. 1977) and Merapi sits at the margin of the Kendeng basin, which
108 contains ≥10 km of upper crustal sedimentary strata (van Bemmelen 1949; Smyth et al.
109 2005). Tertiary marine strata, carbonates, volcanics, and volcanoclastics, probably similar
110 to those outcropping in the nearby Djiwo hills, underlie Merapi volcano (van Bemmelen
111 1949; Hamilton 1979; Chadwick et al. 2007). The Merapi plumbing system consists of a
112 plexus of magma pockets dispersed throughout the crust (Chadwick et al. 2013; Troll et al.
113 2013; van der Zwan et al. 2013) and calc-silicate xenoliths are frequent amongst the
114 eruptive products, providing direct petrological evidence for magma-carbonate interaction
115 in the mid to upper parts of the Merapi supply system (Camus et al. 2000; Chadwick et al.
116 2007; Deegan et al. 2010; Troll et al. 2013, 2015). Fumarole chemistry indicates that
117 carbonate assimilation is ongoing (Allard 1983; Troll et al. 2012), and what is more, post-

118 earthquake fumarole emissions from the summit crater sampled in 2006 demonstrate
119 exacerbated assimilation, likely as a consequence of crustal basement rupture (Troll et al.
120 2012). Notably, the highly explosive 2010 Merapi eruption lacks abundant carbonate and
121 skarn xenoliths, but shows strong geochemical evidence for intense carbonate digestion
122 (Borisova et al. 2013). This “100-year” event was associated with an enormous CO₂ output
123 during the eruption (Surono et al. 2012), thus implying very efficient carbonate
124 assimilation on that occasion.

125 Vesuvius is located in the Campania Province, Italy, where carbonate rocks extend
126 from ≥ 2 km to at least 8 km depth (Zollo et al. 1996; Bruno et al. 1998; Auger et al. 2001;
127 Peccerillo 2005). The current magma plumbing system is distributed over small chambers
128 throughout much of the crust, e.g., between 8 and 22 km (Auger et al. 2001, Civetta et al.
129 2004). Geophysical data place the top of a magma chamber, which is of considerable
130 horizontal extent, at ~ 8 km depth, i.e., at the base of carbonate crustal strata (Auger et al.
131 2001). A shallower (ca. 2 to 4 km) magma chamber fed older Plinian eruptions (e.g.
132 Pompeii, AD 79; Avellino, 3,500 BP), making shallow crustal magma-carbonate
133 interaction a distinct possibility (Barberi et al. 1981; Civetta et al. 2004; Scaillet et al.
134 2008). Skarn xenoliths are a common feature in Vesuvius erupted products and provide
135 independent petrological evidence for magma-carbonate interaction (Fulignati et al. 2000,
136 2004 a,b; Del Moro et al. 2001; Iacono Marziano et al. 2009; Dallai et al. 2011). The
137 addition of volatiles sourced from carbonate must therefore play a significant role in
138 magma chamber dynamics, especially since the current magma residence depth is assumed
139 to be 3 to 4 km, e.g., for the 1944 eruption (cf. Fulignati et al. 2004b; Scaillet et al. 2008).
140 In line with this inference, present day fumarole emissions appear to record a considerable
141 carbonate-derived CO₂ component (Iacono Marziano et al. 2009).

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143

144 **4. Methodology**

145 *4.1 Experimental methods and analytical techniques*

146 Piston cylinder experiments were previously carried out at the High Pressure-High
147 Temperature Laboratory of Experimental Volcanology and Geophysics at the Istituto
148 Nazionale di Geofisica e Vulcanologia (INGV) in Rome, Italy (Deegan et al. 2010; Jolis et
149 al. 2013). The experiments consisted of a carbonate fragment (ca. 6 to 10 mg) enclosed in
150 pre-fused, powdered magmatic starting material (ca. 24 to 42 mg) that were placed
151 together in a platinum capsule. In all charges, the carbonate clast and magmatic material
152 were ~ 20 % and 80 % of the total volume in the capsule, respectively (Table 1). Analyses
153 of all starting materials are reported in Table 2. The experiments were subsequently
154 pressurised at room temperature to 0.5 GPa and then heated in two stages, from ambient
155 temperature to 1180°C at a rate of 200°C/min, and at 20°C/min until the experimental
156 temperature was reached. Three experimental series are compared here: i) a dry Merapi
157 basaltic-andesite, ii) a hydrous Merapi basaltic-andesite (2 wt. % H₂O), and iii) a hydrous
158 Vesuvius shoshonite (2 wt. % H₂O). Note that a total of 2 wt. % H₂O was added to the
159 hydrous experiments to account for the likely range of hydrous subduction zone
160 compositions. The starting materials were then allowed to react at the target conditions for
161 0, 60, 90, and 300 s. The experiments were terminated by shutting down the power source,
162 which caused quenching at a rate of ~2000°C/min and allowed the textural interplay
163 between limestone and melt to be preserved. For a full account of starting materials,
164 experiment preparation, experimental procedure, and rationale, see Deegan et al. (2010,
165 2011) and Jolis et al. (2013).

166 After the experiments were terminated, the capsules were mounted in epoxy resin,
167 abraded to remove the capsule, and then polished to expose the surface of the experimental
168 products at approximately mid-capsule depth. Back-scattered electron images of the
169 polished surface were acquired using a JEOL JSM-6500F Field Emission Gun Scanning

170 Electron Microscope (FEG-SEM) at INGV, Rome. As vesicle size varied over three orders
171 of magnitude, image magnification was varied accordingly.

172

173 *4.2 Image analysis and data treatment*

174 The aim of this paper has been to generate quantitative textural data to complement
175 previous studies that employ quantification of the vesicle characteristics of volcanic rocks
176 (e.g. Gardner et al. 1999; Shea et al. 2010; Heap et al. 2014). Low magnification FEG-
177 SEM images (as low as 15x magnification), covering the whole area of the exposed
178 experimental charges, were used to measure the size of the largest vesicles. Conversely,
179 images of increasingly higher magnification (up to 1500x) were used to measure smaller
180 vesicles over progressively smaller areas. The number of images and the magnifications
181 employed were modified for each sample to best capture its vesicle size distribution. First,
182 a visual inspection of the entire experiment at variable magnifications ensured the absence
183 of clusters of small bubbles. This step was followed by taking the highest magnification
184 images from randomly selected portions of the experiments. We then measured the area of
185 each high magnification image, including the areas of all chemical domains that contained
186 vesicles, but excluded any area with remaining carbonate or calcite crystals. The area of all
187 individual vesicles present in each high magnification image was also measured separately
188 using the ImageJ software (Abramoff et al. 2004). Vesicles attached to the edge of the
189 capsule were included in the count, whereas those on the edge of an image were excluded.
190 The Heywood diameter (equivalent diameter of a non-circular object) or Feret diameter
191 (largest diameter obtainable from the area of a near-circular object) was calculated for each
192 vesicle. Since these two measurements are equivalent for near-spherical vesicles, we will
193 refer to the Heywood diameter only from here. These diameter values were ordered with
194 the largest first, and were assigned a cumulative number (C_n). To compensate for the
195 smaller vesicle areas measured at high magnification, we corrected the number of bubbles

196 to represent the whole capsule using a correction factor (F_{corr}). In this process, bubbles with
197 a large area in the lower magnification groups were removed from all successively higher
198 magnification groups to prevent smaller bubbles being extrapolated over areas that were
199 occupied by larger bubbles. This correction factor (F_{corr}) and the cumulative correction
200 (C_{corr}) were then calculated for each magnification class as follows:

$$F_{\text{corr}} = \text{Area}_{\text{capsule}} / \Sigma \text{Area}_{\text{image}} \quad (1)$$

$$C_{\text{corr}} = C_n \cdot F_{\text{corr}} \quad (2)$$

203 The cumulative correction value (C_{corr}) represents the theoretical number of bubbles
204 present in the whole sample that are larger than any specific diameter (Figure 1) and
205 accounts for any over- or under-estimation of bubble numbers produced in the
206 measurement procedure. We used a Heywood diameter vs. cumulative correction plot for
207 each magnification group to identify where bubble measurements from each subsequently
208 higher magnification set appear. Since bubbles of almost all sizes were measured at least at
209 two or three different magnifications, overlap between the magnification groups ensured
210 the effectiveness of the area correction (Figure 1). Finally, one single curve was created for
211 each experimental charge from all magnification groups by using (i) representative
212 magnifications to define the different sections of the curve and (ii) favouring higher
213 magnification curves for smaller vesicle domains. Unfortunately, during polishing of the
214 experimental charge 445-V7 (Vesuvius, 300 s), part of the sample was lost and the data are
215 thus incomplete for this run. To compensate for this shortcoming, the missing area was not
216 included in higher magnification area measurements, but was included in the whole sample
217 area measurement. For extrapolation of vesicles over the whole sample area, we assumed
218 that the lost area simply had the same vesicle arrangement as that observed in the
219 remaining sample.

220 This method of vesicle size and number estimation employs a 2D surface and assumes
221 that the largest diameter of each vesicle was intersected to derive an “ideal” diameter that

222 forms the basis of our vesicle volume calculations. Our approach would introduce error
223 into a 3D volume calculation as the largest diameter of each vesicle is not always
224 intersected. Furthermore, as larger vesicles dominate the samples, the potential for smaller
225 vesicles to be underrepresented in 3D is increased (e.g., Sahagian and Proussevitch 1998).
226 The method employed here, however, produces an internally consistent data set for the
227 experiments that is applicable for our purpose, but does not lend itself to accurate volume
228 determinations.

229

230

231 **5. Results**

232 *5.1 Textural observations*

233 Here we provide a summary of the textural features recorded in the respective Merapi
234 and Vesuvius experiments. This direct comparison of the two experimental studies is
235 intended to consolidate the available information in Deegan et al. (2010, 2011) and Jolis et
236 al. (2013).

237 All three experimental series record chemical interaction between carbonate and melt
238 as a result of partial or total assimilation of the carbonate (Figure 2). Three compositional
239 glass domains are identified: i) a Ca-normal glass, compositionally similar to the starting
240 material, ii) a Ca-rich glass, which arises due to breakdown of carbonate in the melt, and
241 iii) a contamination front of varying extent that usually marks the transition between the
242 Ca-normal and Ca-rich domains (see Deegan et al. 2010 and Jolis et al. 2013 for full
243 details). Assimilation of carbonate in these experiments is always accompanied by rapid
244 vesiculation and the size and spatial distribution of the vesicles are related to the stage of
245 the reaction. Notably, in the dry series experiments, the Ca-normal domain contains
246 crystal-rich regions, which affects melt viscosity (see discussion also).

247

248 *Merapi Dry Series*

249 In the Merapi dry series, the degree of carbonate assimilation increases with increasing
250 run time (i.e. from 0 s to 90 s), with the most extensive carbonate incorporation into the
251 melt at 300 s (Figure 2 a-d). At 0 to 90 s, the remaining carbonate is usually highly
252 fractured and rounded and can show spallation of small fragments into the surrounding Ca-
253 rich glass. The contamination front is well defined in all experiments of this series.
254 Vesicles are primarily located in the Ca-rich glass with small vesicles frequently found
255 near the carbonate while larger vesicles occupy generally more distant regions in the
256 capsules (e.g. Figure 3 a) and sometimes appear to concentrate at the contamination front.
257 In fact, the contamination front frequently follows the profile of large bubbles (e.g. in the
258 60 and 300 s experiments; Figure 2 b, d), indicating that bubble growth and migration
259 progressively drives the compositional interface away from the carbonate clast. Notably, a
260 sub-set of smaller vesicles is also located in the crystal-rich domain (Figure 2 d).

262 *Merapi Wet Series*

263 In the Merapi wet series, the carbonate was no longer observed in the capsule by 90 s.
264 The edges of the remaining carbonate in the 0 and 60 s runs are irregular, rounded, and
265 also show spallation of carbonate clasts from the main carbonate fragment. Calcite crystals
266 are present surrounding the remnant carbonate fragments in the 0 and 60 s experiments and
267 display a dendritic texture, which probably relates to rapid quenching of the experiments
268 (Figure 4 a).

269 The extent of the Ca-rich glass generally increases at the expense of the Ca-normal one
270 with increasing run time (Figure 2 e-h). Ca-normal glass is observed enclosing areas of Ca-
271 rich glass in the advanced time series, implying mingling of the two melts. The
272 contamination front becomes wider and less defined at increasing experimental duration
273 (cf. Deegan et al. 2010). Larger vesicles are located in both the Ca-rich and Ca-normal

274 glasses, with an overall preference for the capsule margins. Smaller bubbles also occur in
275 both glass types. However, in the Ca-normal glass, the smaller bubbles are arranged at a
276 distance from the contamination front in semi-parallel bands, but maintain the shape of the
277 contamination front (Figure 4 b). In addition, as in the Merapi dry experiments, the
278 contamination front frequently follows the profile of large bubbles (e.g. 60 s experiment;
279 Figure 2 f), highlighting the effect of vesicle growth and migration on the distribution of
280 Ca-enriched (contaminated) melt domains.

281

282 *Vesuvius Wet Series*

283 In the Vesuvius wet series, the carbonate fragment was completely digested at 0 s.
284 Although part of the 300 s experiment disintegrated during sample preparation, but
285 because no carbonate was present at 0, 60 and 90 s, we assume that the carbonate was also
286 fully dissolved in the 300 s run. In comparison with the Merapi series, the Ca-rich and Ca-
287 normal glasses are no longer defined and the contamination front is generally more diffuse,
288 increasing in width as the experimental duration progresses (Jolis et al. 2013). Notably, the
289 majority of vesicles at 0s, 60 s and 90 s are now in the Ca-normal glass. At 0 s there are
290 some larger bubbles located in the Ca-rich glass, though the majority of smaller vesicles
291 appear to be in the Ca-normal glass (Figure 5 a). At 300 s, Ca-rich glass is dominant in the
292 experimental charge (see Jolis et al. 2013) and almost all vesicles are located there. Calcite
293 crystals are present in the 0, 60, and 90 s experiments and show a dendritic texture due to
294 quenching (Figure 5 b).

295

296 *5.2 Bubble nucleation and growth*

297 Vesicle diameter and the cumulative vesicle number (C_{CORR}) were obtained for each of
298 the experimental series. In the Merapi dry series (Figure 6 a), the 300 s experiment
299 contains the largest amount of small bubbles ($< 5 \mu\text{m}$) and overall, the size of the smallest

300 bubbles decreases with increasing experimental duration (from 2.5 μm at 0 s to 0.2 μm at
301 300 s). Conversely, the size of the largest bubbles increases with increasing experiment
302 duration (from 330 μm at 0 s to 880 μm at 300 s) and only the 90 s experiment does not
303 follow these trends.

304 In the Merapi wet series (Figure 6 b), the 300 s experiment contains the largest amount
305 of bubbles less than 5 μm in size. The diameter of the smallest and largest bubbles
306 increases from 0.1 μm at 60 s run to 0.4 μm at 300 s and from ~ 135 μm at 0 s to >1000
307 μm at 300 s, respectively. Both Merapi series show the total number of bubbles to increase
308 with time, i.e., the 300 s experiments contain significantly more bubbles than the 0 s ones.

309 The Vesuvius wet series shows an analogous increase in maximum bubble size with
310 time, but also a decreasing number of bubbles with increasing run time (Figure 6 c). The
311 smallest bubble size in the Vesuvius wet series ranges from 0.1 μm at 0 s to 0.3 μm at 300
312 s, whereas the largest bubble size at 60 s is ~ 470 μm and ~ 610 μm at 300 s. The variation
313 of the largest bubble size for the Vesuvius series (Figure 6 c) is therefore much less than in
314 the two Merapi series (Figure 6 a, b). In this respect, vesicle diameter versus cumulative
315 vesicle volume shows a general trend towards larger cumulative volumes associated with
316 longer experimental duration for both of the Merapi series, as expected from the increasing
317 degree of carbonate assimilation (Figure 7 a, b). In contrast, complete assimilation of
318 carbonate is confirmed in the Vesuvius wet experiments by approximately equal
319 cumulative bubble volumes in all experiments (Figure 7 c) and is consistent with the
320 textural similarities to those Merapi experiments where extensive carbonate assimilation
321 occurred (e.g. the 300 s Merapi experiments).

322

323

324 **6. Discussion**

325 *6.1 Water and carbon dioxide solubility in Merapi and Vesuvius magmas*

326 Given the presence of water and the addition of CO₂ to the experiments, it is important
327 to consider volatile solubility in the starting materials to understand volatile exsolution and
328 migration in the corresponding experiments (e.g. Dixon 1997; Vetere et al. 2011). For the
329 Vesuvius wet series, the shoshonite melt can dissolve up to 7.9 wt. % H₂O at 1250 °C and
330 0.4 GPa (Vetere et al. 2011). In the case of Merapi, a basalt can contain 8.5 wt. % H₂O at
331 1200 °C and 0.5 GPa (Newman and Lowenstern 2002) and an andesite up to 10 wt. % H₂O
332 at 1200 °C and 0.5 GPa (Botcharnikov et al. 2006). With respect to these data, all
333 experimental glasses are H₂O under-saturated (H₂O ca. 2 wt. %). However, the presence of
334 CO₂, whose quantity increases with progressive carbonate assimilation, will reduce water
335 solubility in the melts (e.g. Holloway and Blank 1994; Botcharnikov et al. 2005; Deegan et
336 al. 2010; Jolis et al. 2013). Using the volatile solubility model of Papale et al. (2006), we
337 calculated H₂O-CO₂ solubilities for the Merapi wet experimental series. For the given
338 temperature and pressure conditions (1200 °C and 0.5 GPa) the Merapi wet experiments
339 show a water solubility limit at 2.12 wt. % and a CO₂ solubility of 0.36 wt. %. As
340 determined by Jolis et al. (2013), the Vesuvius wet experiments can dissolve up to 1.9 wt.
341 % H₂O and 0.34 wt. % CO₂. The water solubility for the Merapi wet and Vesuvius
342 experiments in the presence of CO₂ is thus very close to the initial ~ 2 wt. % H₂O of the
343 hydrous starting material, implying that when volatile saturation is reached the evolving
344 vapour phase is dominated by CO₂. This can also be shown by employing the method of
345 Deegan et al. (2010) and Jolis et al. (2013) to calculate the maximum amount of CO₂
346 produced from carbonate breakdown in the experiments. Taking the average masses of the
347 magmatic starting material and the carbonate used in each experimental series, we
348 calculate that 4.29 mg CO₂ (7.02 wt.%), 4.28 mg CO₂ (7.0 wt.%), and 2.9 mg CO₂ (4.75
349 wt.%) can be produced for the Merapi dry, Merapi wet, and Vesuvius wet experiments,
350 respectively. Carbon dioxide is thus more abundant in the experiments than H₂O
351 (maximum 2 wt. %) and represents the critical vapour phase in all series presented here.

353 *6.2 Carbonate dissolution and the influence of melt viscosity*

354 The rate of CaCO₃ breakdown in magma is controlled by the removal of CO₂ from the
355 reaction site. Studies of natural systems such as the Hortavær Igneous Complex (Norway)
356 and Etna (Italy) have suggested that prolonged magma-carbonate interaction is permitted
357 only if CO₂ is efficiently and continually removed from the reaction site (Barnes et al.
358 2003; 2005; 2009; Mollo et al. 2012; Heap et al. 2013). This is because decarbonation in a
359 closed system would eventually slow down due to an excess of the reaction product.
360 Although the removal of CO₂ in the experiments would eventually be constrained by the
361 capsule, those experiments where the Ca-rich glass reaches the edge of the capsule show
362 that bubbles were still found to grow larger with increased experiment duration, which
363 suggests that outgassing is not yet limited by the capsules (cf. Figure 2).

364 In contrast, the experiments show that under the investigated conditions, carbonate
365 dissolution began before the experimental target temperature was reached and thereafter
366 progressed with increasing experimental run-time. We note that the break-down of
367 carbonate is accelerated in the experiments over that of pure decarbonation because of
368 reaction with magma. Bubble nucleation occurred heterogeneously, i.e., at the carbonate-
369 melt interface, and homogeneously, i.e., within the melt. Heterogeneous nucleation also
370 occurred on other available nucleation surfaces, such as microlites even at very low
371 degrees of oversaturation (cf. Masotta et al. 2014) and probably explains the formation of
372 small vesicles in crystal-rich domains (Figure 2d). Homogeneous nucleation in certain melt
373 domains, in turn, implies that volatiles were added (likely by diffusion) and that volatile
374 saturation was then locally achieved (cf. Hurwitz and Navon 1994).

375 Indeed, Ca-rich glass and vesicles are present in all 0 s experiments, with the Vesuvius
376 experiments representing the most extreme case wherein the carbonate was completely
377 dissolved at 0 s. Comparing (i) the Merapi dry series, wherein carbonate assimilation

378 required 300 s to reach an advanced stage and (ii) the Merapi wet series, wherein the
379 carbonate was largely dissolved by 90 s, it becomes apparent that compositionally-
380 controlled melt viscosity exerts a key control on the reaction rate.

381 For example, at 1200 °C, the Merapi dry, Merapi wet, and Vesuvius wet experiment
382 starting materials (equivalent to the Ca-normal domains) have viscosities of $10^{2.7}$, $10^{1.4}$ and
383 $10^{1.1}$ Pa.s, respectively (calculated after Giordano et al. 2008). If we take microcrystals in
384 in the Merapi dry series into account, we obtain a viscosity value of $10^{3.5}$ Pa.s. On the other
385 hand, the Ca-rich melts have calculated viscosities between $10^{0.3}$ and $10^{0.2}$ Pa.s for the
386 Merapi and Vesuvius series, respectively (assuming 25 wt. % CaO), which is an order of
387 magnitude lower than the results for the Ca-normal melts and confirms that the rate of
388 carbonate dissolution is inversely related to melt viscosity.

389 In addition, the spatial distribution of vesicles in our time lapse experiments provides a
390 record of bubble growth and migration. For all experiments, bubbles tend to move away
391 from the carbonate clast as run time proceeds, with greater vesicle mobility in lower
392 viscosity systems (Figure 8 and 9). Bubbles migrate away from the carbonate clast because
393 a pressure and composition gradient is rapidly established in the experiments once the
394 decarbonation reaction begins. The experiments quickly become oversaturated with CO₂
395 close to the reaction site and bubble nucleation and migration become the main mass
396 transport mechanisms. In the experiments, fluid pressure is greater at the reaction site than
397 in the distal parts of the capsule, which causes bubbles to move away from the reaction site
398 in a broadly radial fashion (Figure 8 and 9). Importantly, throughout the whole Vesuvius
399 series, larger bubbles tend to be positioned preferentially at the capsule margins (Figure 2
400 i-l), reflecting fast volatile migration into the uncontaminated areas. Vesicle migration is
401 sluggish in more viscous systems, however. For instance, in the Merapi dry series, where
402 bubbles and Ca-rich melt remain relatively close to the carbonate clast and/or the reaction
403 site even at 300 s, virtually all bubbles are contained in the Ca-rich melt (Figure 2 d).

404 Conversely, bubbles occur in both the Ca-rich and Ca-normal melt in the Merapi wet series
405 and there, the Ca-rich glass area is much wider. Notably, in the Merapi wet series, the
406 boundary between Ca-rich and Ca-normal glasses follows the profile of large bubbles
407 (Figure 2 f) or is aligned with many smaller bubbles (Figure 2 h and 4 b).

408 In the Vesuvius wet 0 s experiment, mixing/physical mingling between the Ca-rich and
409 the Ca-normal melts is indicated by chemical gradients across the interface between the
410 two glass domains. The interfaces between Ca-rich and Ca-normal melts roughly follow
411 the shape of the bubbles, indicating that bubble migration facilitates melt mingling (Figure
412 2 and 5). We therefore suggest that the viscosity contrast between the Ca-rich and Ca-
413 normal melts (see calculated values above) probably represents an initial hindrance for
414 bubble migration and magma mixing, temporarily locking the bubbles in the Ca-rich melt
415 as evident from the concentration of small bubbles along the contamination front in the
416 Merapi dry series (Figure 3 b). In contrast, in the Merapi wet series, vesicles also grew in
417 the Ca-normal zones. In this case, CO₂ pulsing from the carbonate resulted in CO₂ and
418 bubble migration from the Ca-rich melt, as suggested by bubble trails in the Ca-normal
419 glass that parallel the contamination front (e.g. Merapi wet 300 s; Figure 4 b). For the
420 Vesuvius wet series, bubble nucleation also occurred in both Ca-rich and Ca-normal melts.
421 The relatively low viscosity contrast between these two melts appears to have promoted
422 rapid bubble migration and efficient CO₂ diffusion into the Ca-normal domains (Figure 5
423 b).

424 Prolonged nucleation dominates the Merapi dry and wet runs where both small and
425 large bubbles increase in number with time (Figure 6 a, b). The similar cumulative vesicle
426 number for Merapi wet 60 s and Vesuvius wet 0 s experiments (Figure 6), together with
427 the increasing cumulative vesicle volume at longer run times for both Merapi series (not
428 observed in Vesuvius experiments; Figure 7), demonstrates the different rates of
429 assimilation-induced CO₂ exsolution. The slower assimilation rate causes prolonged
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430 bubble nucleation in both the Merapi dry and wet series, relative to the Vesuvius wet series
431 where nucleation is already exhausted as the carbonate is fully dissolved.

432 Fast digestion of carbonate in both Merapi and Vesuvius wet experiments hence causes
433 sudden release of volatiles from the reaction site but ambient melt viscosity controls their
434 migration style. Bubble release likely occurs in a pulsatory fashion as multiple vesiculation
435 events (or “bubble blow-outs”) are recognised from the sharp increases in vesicle number
436 over a given diameter range in all of our experimental series (Figure 6). In fact, the number
437 of vesiculation pulses increases with experimental duration, and for example, the Vesuvius
438 and Merapi wet experiments record two vesiculation pulses at 0 s, but three are seen at 300
439 s (see Figure 6 b and c). Although evidence of coalescence is found in all experimental
440 series, in the form of, e.g., partially retracted bubble walls and films (e.g. Merapi dry
441 series; Figure 3 b), coalescence becomes the dominant growth mechanism in the low
442 viscosity Vesuvius runs, where the overall number of bubbles (especially in the small size
443 fraction) decreases with time (Figure 6 c).

444 Chemical gradients and “pile-up” of Ca and Sr between the Ca-rich and Ca-normal
445 glass domains are found in both Merapi and Vesuvius hydrated experiments (Deegan et al.
446 2010; Jolis et al. 2013) and are of similar magnitude in all experiments (except for
447 Vesuvius wet 300 s where homogenisation becomes increasingly relevant). The alignment
448 of bubbles, both at and parallel to the contamination front, indicates a similar pile-up for
449 CO₂, and underlines that volatile migration is influenced by viscosity through slowing
450 down diffusive volatile migration and the movement of bubbles that have already formed.

451 Bubble formation affects both Ca-rich and Ca-normal domains, indicating that some
452 CO₂ can diffuse through the rheological barrier at the melt interface(s) (Figure 2). We also
453 note that in low viscosity systems bubbles are efficiently transported away from the
454 carbonate, causing a high frequency of bubble blowouts that can create a steady state
455 dissolution system, analogous to an effervescent tablet in water that dissolves in very rapid

456 pulses, usually almost too fast for us to see. In high viscosity systems, in contrast, bubbles
457 are initially locked in the melt close to the carbonate where they continue to grow, thus
458 generating fluid overpressure and potentially episodic bubble blowout events. Initiation of
459 a bubble blowout in such systems could be caused by several factors including, (i) the
460 bubble number density gradient between the contaminated and normal melt, (ii) a
461 progressive change of the chemical gradient, i.e., an increased tendency towards chemical
462 homogenisation, or (iii) external factors such as earthquakes, dyking, or magma
463 replenishments (e.g. Deegan et al. 2011; Troll et al. 2012).

465 *6.3 Implications for natural volcanic systems*

466 The experiments represent laboratory simulations of magma-carbonate interaction, a
467 process that has been shown to occur at both Merapi and Vesuvius volcanoes (Fulignati et
468 al. 2004b; Chadwick et al. 2007; Iacono Marziano et al. 2009; Dallai et al. 2011; Borisova
469 et al. 2013). The experiments are particularly relevant because they mirror features found
470 in natural rocks. Micro-analyses of calc-silicate xenoliths from Merapi, for example, have
471 been shown to contain extremely CaO-enriched, sometimes glassy domains (up to 61 wt.%
472 CaO, Deegan et al. 2010). At Vesuvius, Mg-rich olivine demonstrates elevated $\delta^{18}\text{O}$ values
473 relative to typical, mantle-derived mafic magmas due to carbonate-derived CO_2 fluxing in
474 the roots of the volcanic system (Dallai et al. 2011). While the isothermal conditions of the
475 experiments may not be fully applicable to assimilation at magma chamber walls, i.e., with
476 a strong thermal contrast (cf. Del Moro et al. 2001; Spera and Bohrson 2001), the
477 experimental conditions replicate magmatic digestion of a stoped carbonate block or
478 xenoliths from a “warm” chamber roof or wall (i.e. we assume the mass of the assimilated
479 clast to be much less than that of the magma that provides heat to drive the reaction; Figure
480 9). Specifically, we found that melt viscosity exerts a major control on the carbonate
481 dissolution reaction by regulating the efficiency of bubble migration from the reaction site,

482 which in turn determines how fast the carbonate can be digested. Low viscosity systems
483 such as Vesuvius will likely permit frequent vesiculation pulses where bubbles can migrate
484 freely from the Ca-rich melt. Efficient migration of bubbles from the dissolving carbonate
485 can thus trigger more frequent new vesiculation pulses and promote higher assimilation
486 rates (Figure 9). This style of vesiculation would lead to steady state CO₂-production for as
487 long as the reactant is available. A potential outcome of this process would be an increase
488 in the buoyancy of Vesuvius magmas that would trigger rapid ascent and lead to sustained
489 explosive eruptions (cf. Dallai et al. 2011).

490 For a more viscous magma (e.g. Merapi), bubble accumulation at the reaction site will
491 progressively retard the carbonate dissolution reaction and the magma will locally behave
492 as a closed system. If bubbles are not transported away, the reaction will simultaneously
493 increase local fluid pressure until it terminates due to a lack of melt at the reaction site. If
494 the confining pressure is exceeded during this process, however, a bubble blow-out may be
495 triggered in a broadly radial fashion from the reaction site, i.e., the carbonate (Figure 8 and
496 9), thus relieving localised fluid over-pressure and allowing the decarbonation reaction to
497 proceed again. This type of vesiculation would lead to an explosive, gas-driven eruption
498 once the overpressure threshold is exceeded and may be repeated with successive CO₂
499 build up events. This may apply to Merapi, especially the pre-2010 eruptions, as the latter
500 was a “100-year event” (Surono et al. 2012) caused by a large contribution of fresh, low
501 viscosity basaltic magma from depth, rather than a dome-forming eruption as typical of the
502 20th century activity at Merapi. Notably, the bubble-blow out mechanism seems to be also
503 recorded in plutonic systems that underwent magma-carbonate interaction, such as the
504 Hortavaer intrusive complex, Norway, where fragmented structures have been interpreted
505 in this context (Barnes et al. 2009).

506 Regardless of the vesiculation style presented by a magmatic system undergoing
507 magma-carbonate interaction, a fundamental finding from the experiments is that CO₂

508 fluxing from carbonate is coupled to magma viscosity. Where vesiculation becomes
509 progressively sluggish, the bubble blow-out mechanism described here can re-ignite
510 vesiculation once the local fluid pressure threshold has been exceeded. In contrast, efficient
511 CO₂ fluxing is permitted in relatively low viscosity melts, which will allow carbonate
512 assimilation to proceed unhindered and will likely lead to steady-state degassing.

513 The wider implication arising from these observations is that carbonate assimilation in
514 a variety of different types of magmatic systems with differing melt properties can
515 nevertheless contribute substantial quantities of CO₂ to the atmosphere via volcanism. It is
516 thus imperative that carbonate-rich upper crustal lithologies beneath arcs are considered in
517 mass balance models concerning Earth's deep carbon-cycle (cf. Johnston et al. 2011; Troll
518 et al. 2012; Dasgupta 2013; Lee et al. 2013).

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520

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789 **Figure Captions**

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4 791 **Fig 1** Heywood diameter vs corrected cumulative number (C_{CORR}) of vesicles for Vesuvius
5
6 792 wet 0 s experiment. The cumulative vesicle correction extrapolates the number of vesicles
7
8 793 measured over the area of the whole capsule. The data are firstly divided into respective
9
10 794 magnification groups at which the vesicles were measured. Subsequently, the most
11
12 795 representative magnification group for vesicles of different diameters was identified. The
13
14 796 0-80 magnification group is most representative of, e.g., the largest vesicles, characterising
15
16 797 those to ca. 16 μm in size, while the other curves are more representative of smaller sizes.
17
18 798 A complete experimental curve can thus be reconstructed from the individual
19
20 799 magnification curves (see Fig 6c).
21
22
23
24

25 800
26

27 801 **Fig 2** Back-scatter electron images of textures and vesicle distributions produced in the
28
29 802 three different experimental series and at four experimental durations. **a) to d)** Merapi dry
30
31 803 series; **e) to h)** Merapi wet series; and **i) to l)** Vesuvius wet series. Annotations show the
32
33 804 remnant carbonate, the Ca-normal and crystal-rich domain, the Ca-rich glass and, where
34
35 805 appropriate, the contamination front.
36
37
38
39

40 806
41

42 807 **Fig 3** Merapi dry 60 s experiment **a)**. Vesicles form at the interface between melt and
43
44 808 carbonate (the reaction site). Smaller vesicles are generally located closer to the carbonate
45
46 809 remnant. The contamination front, at the interface between the Ca-rich and the Ca-normal
47
48 810 glasses, contours the outside edge of the larger vesicles. This suggests that bubble growth
49
50 811 and bubble migration plays a role in driving the contamination process in lower viscosity
51
52 812 systems. **b)** Merapi dry 300 s experiment. Smaller vesicles collect at the contamination
53
54 813 front between the crystal-rich domain and the Ca-normal glass. Two large vesicles can be
55
56 814 seen to coalesce on the right hand side of the image. Both quench crystals and smaller
57
58
59
60
61

815 vesicles are present in the crystal-rich domain, while there are almost no small vesicles or
816 crystals in the Ca-rich glass.

817

818 **Fig 4** Merapi wet 60 s experiment. **a)** Example of dendritic texture present in some
819 samples. This texture results from quenching of the Ca-rich melt that has digested calcite.
820 Vesicles are located near the edge of the carbonate clast, some surrounded by microcrystals
821 of calcite. **b)** Merapi wet 300s experiment. Vesicles of similar size are aligned parallel to
822 the contamination front, but are located ahead of it in the Ca-normal glass. The lower
823 viscosity of the Merapi wet experiments relative to the Merapi dry ones allows volatiles to
824 diffuse away from the reaction site more quickly. Here, all volatiles produced in one single
825 vesiculation event seem to move away perpendicular to the reaction site at a similar rate.
826 However, the viscosity is still relatively high in the Merapi wet series and substantial
827 mingling between the Ca-rich and Ca-normal melts is not yet observed in this experimental
828 series.

829

830 **Fig 5** Vesuvius wet 0 s experiment. **a)** At 0 s, just as experimental temperature is reached,
831 intermingling of Ca-rich and Ca-normal glass has already commenced, causing complex
832 textures in comparison to the Merapi dry and wet series. Smaller vesicles in the Ca-normal
833 glass align themselves along the contamination front (CF). Beyond the CF, i.e., further into
834 the Ca-normal glass, smaller vesicles occur, but do not form a regular pattern (e.g. Figure
835 4b). The black line around the dendritic textured area highlights the location of the higher
836 magnification image shown in Figure 5b. **b)** The dendritic texture demonstrates a strong
837 CaO enrichment but is not associated with any solid carbonate remnants.

838

839 **Fig 6** Cumulative vesicle correction (C_{CORR}) vs vesicle diameter (μm) for **a)** Merapi dry, **b)**
840 Merapi wet, and **c)** Vesuvius wet experiments. All four experiment durations (0 s, 60 s, 90

841 s and 300 s) are plotted. The number of vesiculation events that can be identified increases
842 from two in the Merapi wet 0 s to three in the Merapi wet 300 s. A similar sequence of
843 events is also observed in the Vesuvius wet experiments. Vesicle coalescence can be
844 identified where the number of bubbles in the smaller size fraction decreases with longer
845 experimental duration (compare Figure 7c).

846

847 **Fig 7** Cumulative vesicle volume ($C_{VOL} \mu\text{m}^3$) vs vesicle diameter (μm) for **a)** Merapi dry,
848 **b)** Merapi wet, and **c)** Vesuvius wet experiments. All four experiment durations (0 s, 60 s,
849 90 s and 300 s) are plotted. For **a)** and **b)** as experiment duration increases from 0 s to 300
850 s, a general increase in the cumulative volume of bubbles is observed, demonstrating that
851 decarbonation has not reached completion in the shorter duration runs, and that more CO_2
852 is exsolved in the longer duration experiments as the reaction proceeds. In **c)**, the
853 cumulative volume of vesicles is the same for all experimental durations. This indicates
854 that by 0 s (i.e. just as experimental temperatures are reached) the decarbonation reaction is
855 almost complete. As experimental duration increases for this series, homogenisation of the
856 reaction products is the only change observed, i.e., no new bubbles are generated.

857

858 **Fig 8** Schematic diagram of carbonate dissolution processes. **a)** Start of carbonate
859 dissolution. Bubbles nucleate mostly in the Ca-rich melt close to the dissolving carbonate
860 fragment, although some appear also in the Ca-normal melt. **b)** The carbonate clast
861 decreases in size, and the Ca-rich melt increases in volume. Larger vesicles grow by
862 diffusion and coalescence, pushing earlier bubbles and the contamination front away from
863 the carbonate clast. Small vesicles parallel the contamination front, but are also found in
864 Ca-normal melt where they formed as a result of volatile diffusion beyond the
865 contamination front (Merapi wet and Vesuvius cases). Stages **a)** and **b)** will be repeated
866 multiple times as the carbonate clast dissolves in vesiculation pulses. The frequency and

867 style of these pulses is controlled by the rate of removal of dissolved and exsolved CO₂
868 from the reaction site, which itself is controlled by melt viscosity. Once CO₂ has migrated
869 from the reaction site, dissolution can proceed further, thus allowing for the next
870 vesiculation pulse. c) Eventually, the whole carbonate clast dissolves, and the majority of
871 smaller bubbles coalesce into a small number of larger ones. Most vesicles have moved
872 away from the centre reaction site and bubble growth and migration promotes significant
873 mingling and mixing of melts.

874

875 **Fig 9** Model depicting magma-carbonate interaction processes prior to an eruption. a) The
876 plumbing systems beneath both Vesuvius and Merapi consist of magma pockets distributed
877 through the crust, which provide ample surface area for extensive magma-crust interaction.
878 b) Carbonate xenoliths are entrained into a slowly crystallising pocket of magma through,
879 e.g., stoping, dyking, or earthquake activity. Reaction between carbonate blocks and the
880 surrounding host melt produces CO₂ bubbles and a contaminated Ca-rich melt. The
881 reaction area is initially localised and confined to the proximity of the carbonate. c) As the
882 reaction continues, multiple vesiculation pulses cause CO₂ to penetrate a larger volume of
883 host melt. Vesicle growth, coalescence, and migration thus promote mingling of Ca-rich
884 and Ca-normal melts. If melt viscosity is relatively high (e.g. Merapi), volatiles are unable
885 to freely escape and localised over-pressurisation can be achieved, resulting in an
886 explosive “bubble blowout” event. Alternatively, if melt viscosity is relatively low (e.g.
887 Vesuvius), efficient removal of CO₂ bubbles allows for a higher reaction rate, which
888 results in more sustained volatile escape.

889

Figure 1

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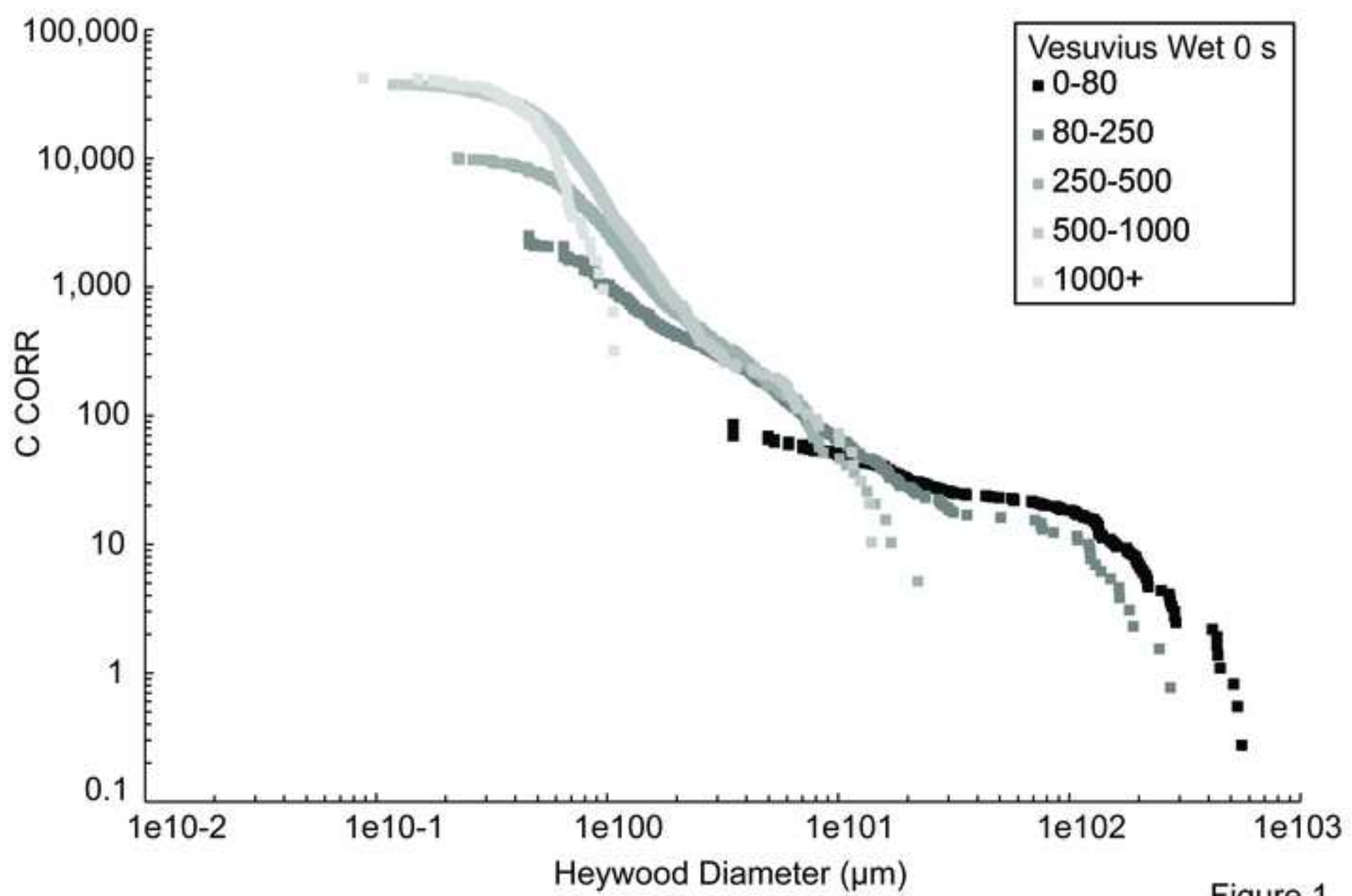


Figure 1

Figure 2

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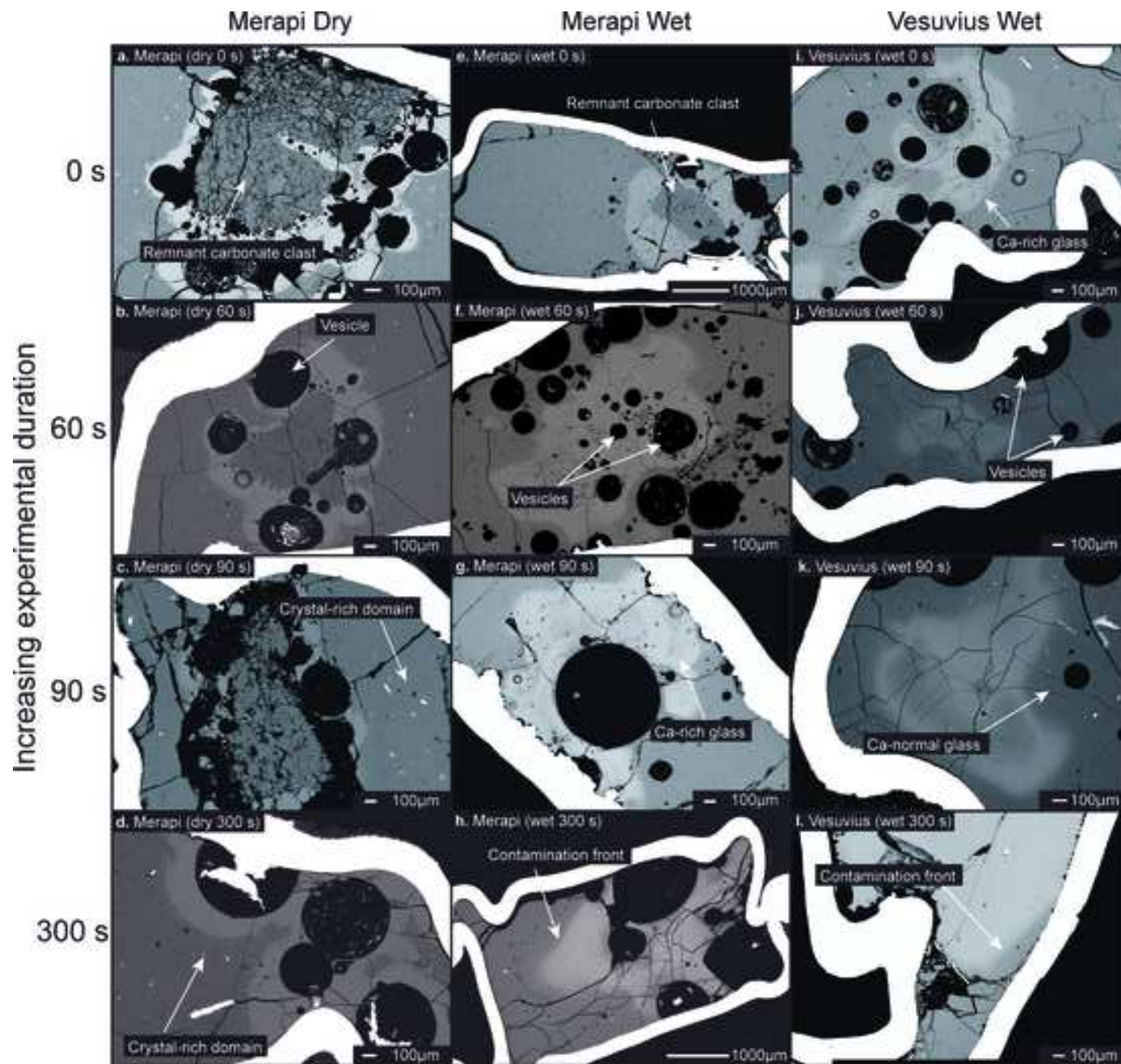


Figure 2

Figure 3

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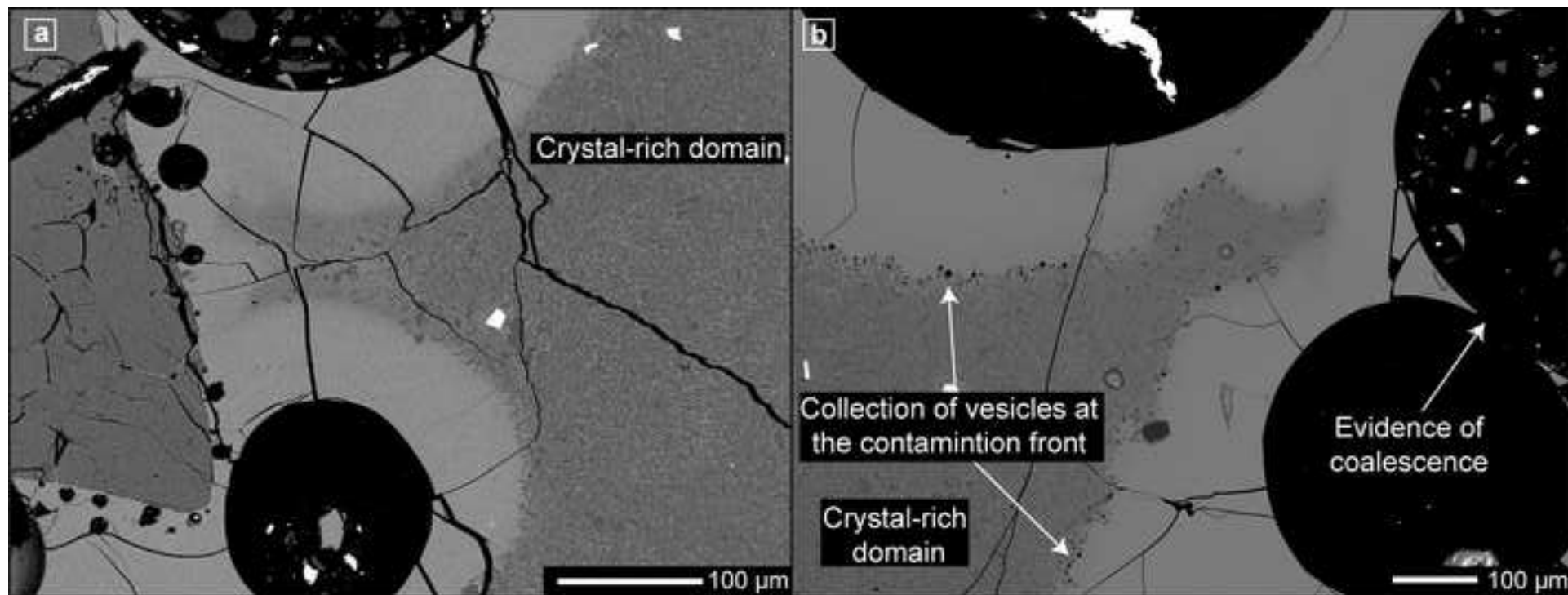


Figure 3

Figure 4

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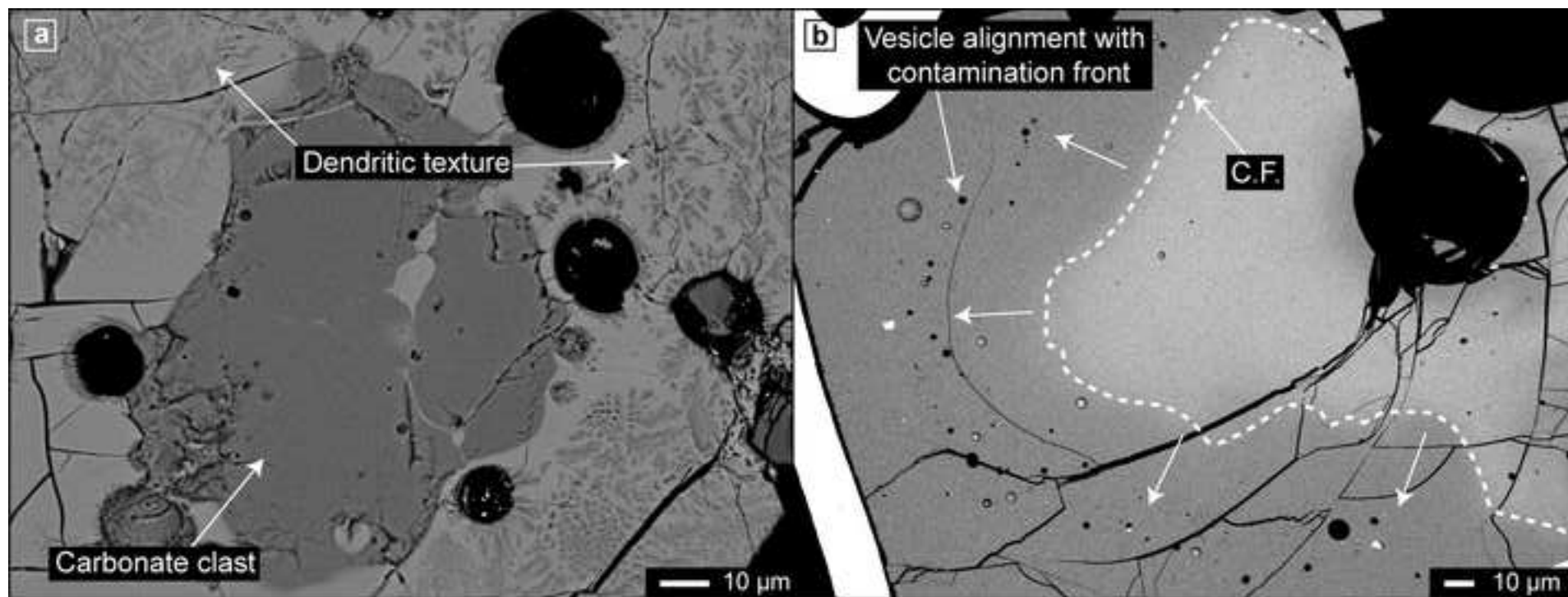


Figure 4

Figure 5

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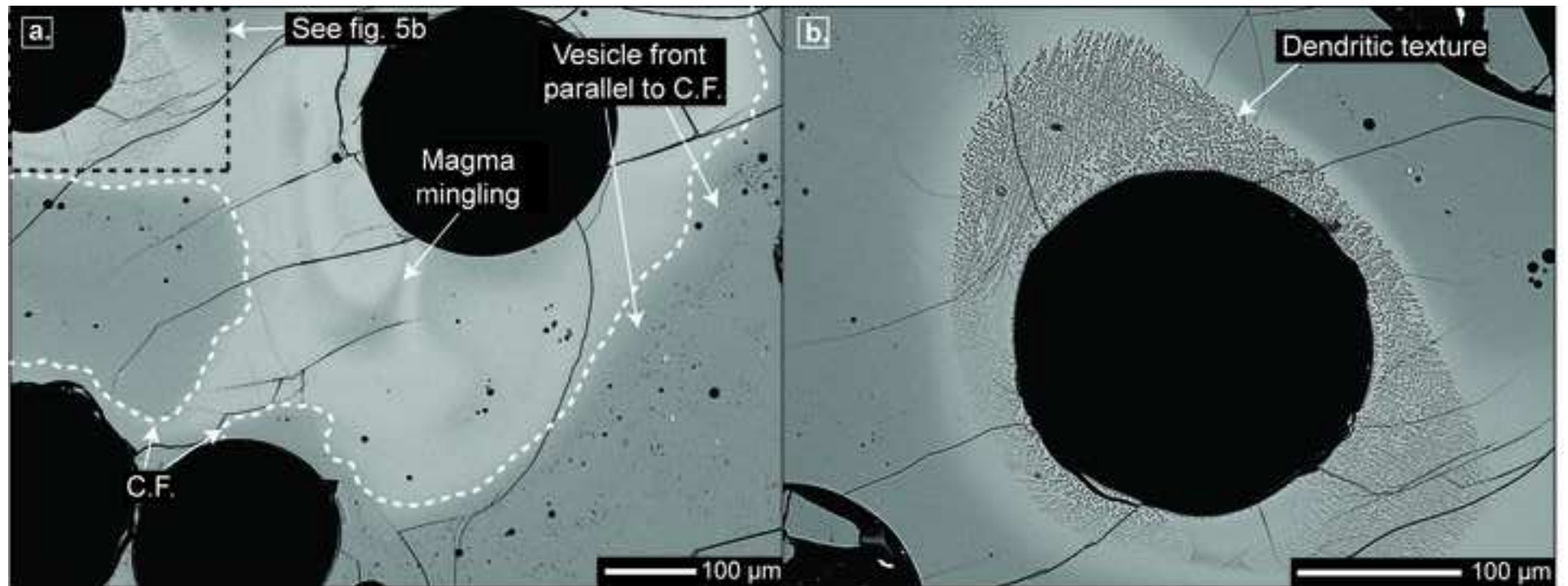


Figure 5

Figure 6

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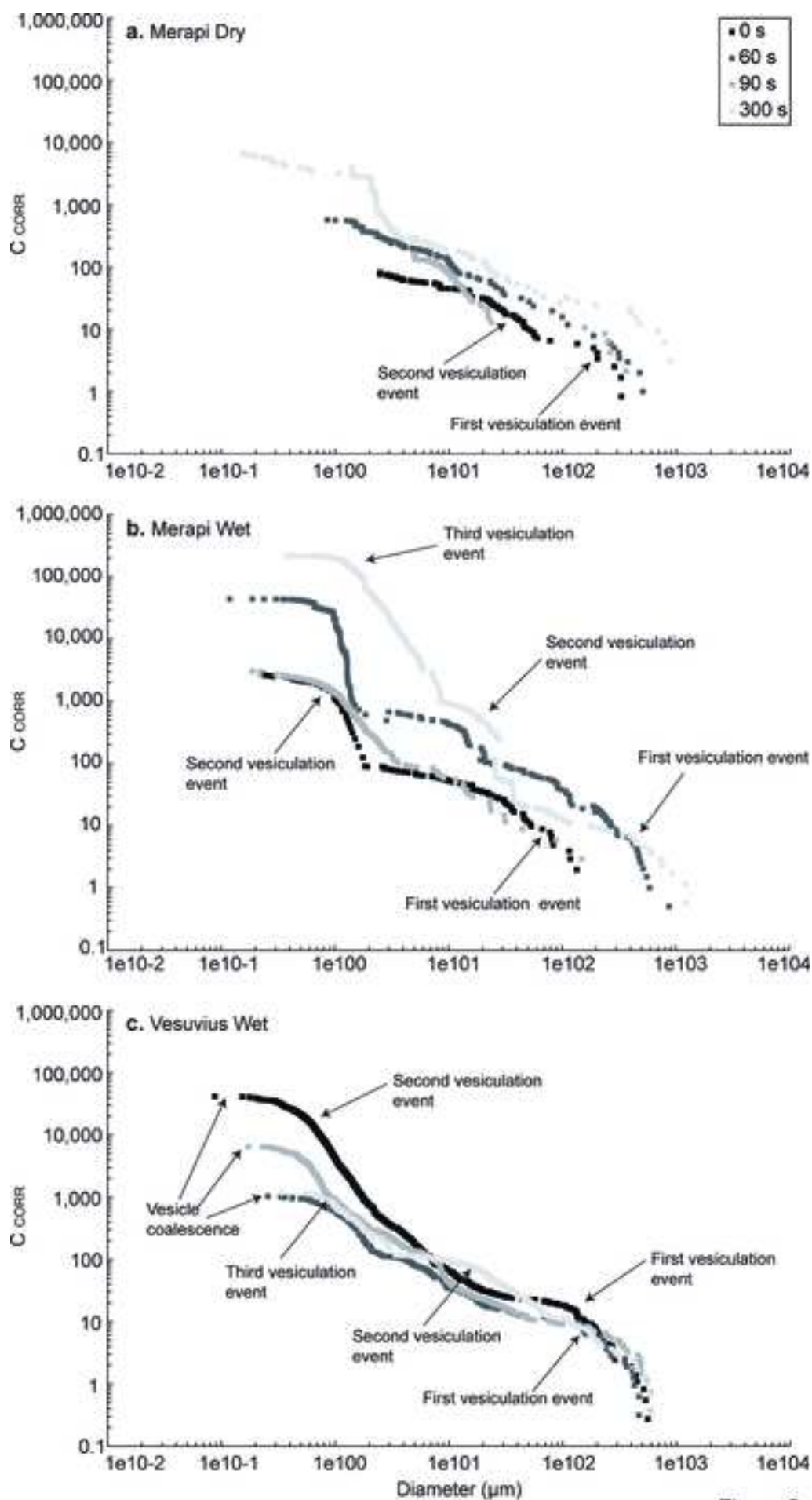


Figure 6

Figure 7
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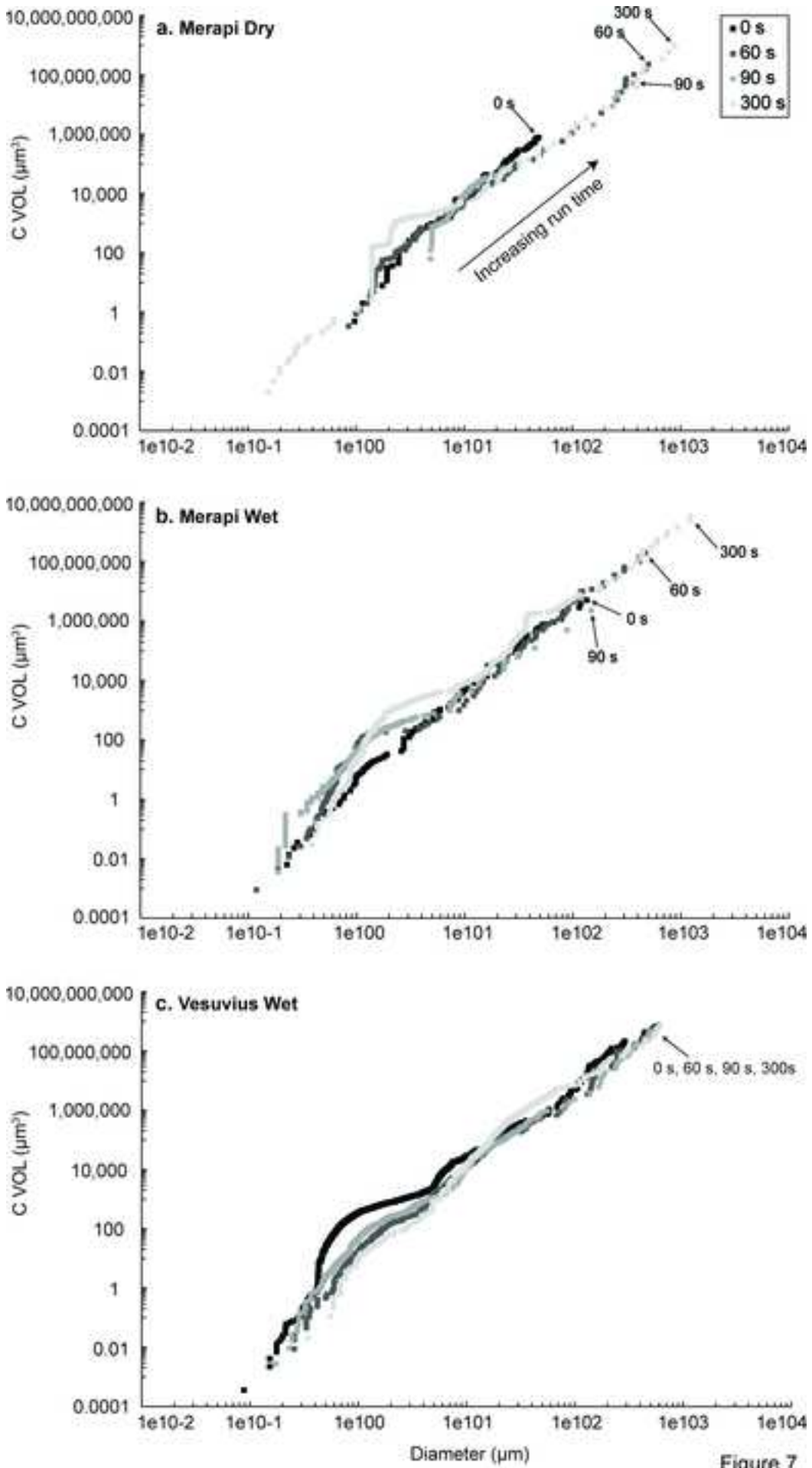


Figure 7

Figure 8

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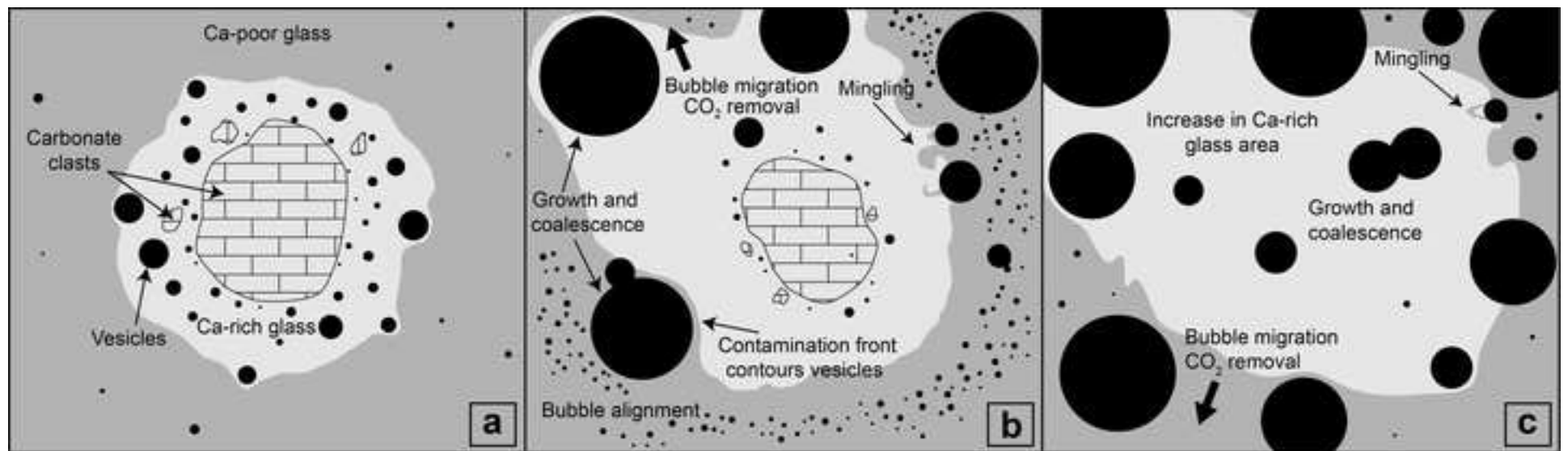


Figure 8

Figure 9

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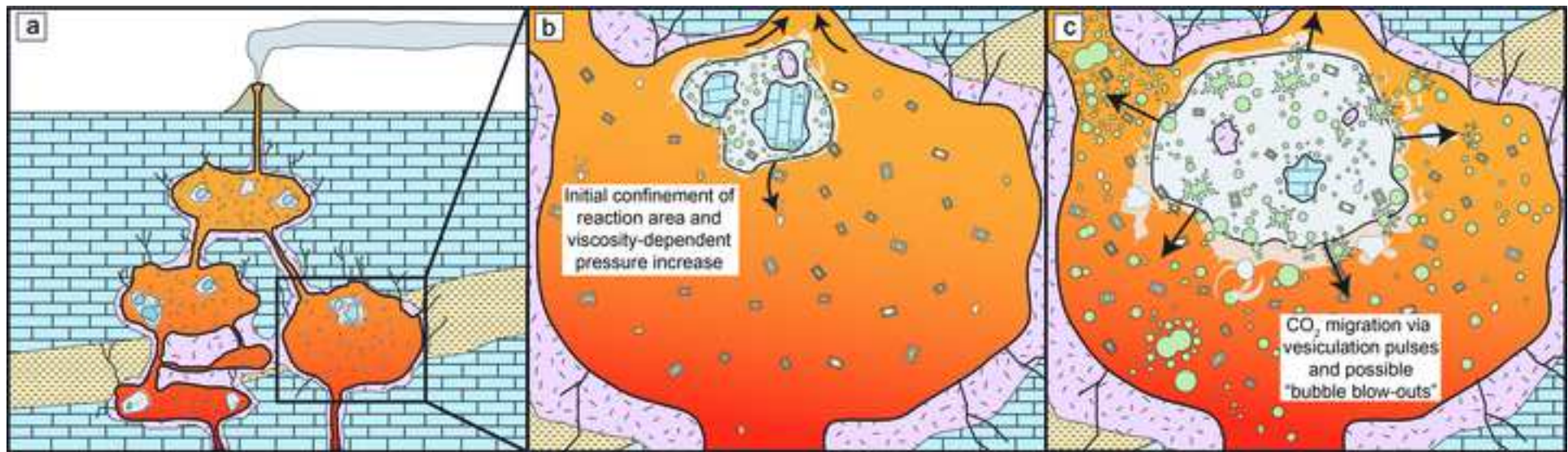


Figure 9

Table 1. *Experimental conditions used for each series*

Run ^{1,2}	Starting Material	H ₂ O (wt. %)	Carbonate (wt. %) ³	T (°C)	P (MPa)	t (s)
Merapi Dry Series ¹						
379-16	Basaltic-Andesite	0	19.6	1200	500	0
387-20	Basaltic-Andesite	0	20.0	1200	500	60
376-10	Basaltic-Andesite	0	19.0	1200	500	90
386-18	Basaltic-Andesite	0	19.4	1200	500	300
Merapi Wet Series ¹						
379-17	Basaltic-Andesite	2.23	19.7	1200	500	0
387-21	Basaltic-Andesite	2.23	19.3	1200	500	60
376-11	Basaltic-Andesite	2.23	17.5	1200	500	90
386-19	Basaltic-Andesite	2.23	19.3	1200	500	300
Vesuvius Wet Series ²						
443-V5	Shoshonite	2.02	17.1	1200	500	0
442-V3	Shoshonite	2.02	17.9	1200	500	60
441-V1	Shoshonite	2.02	21.9	1200	500	90
445-V7	Shoshonite	2.02	21.7	1200	500	300

¹ Deegan et al. (2010)² Jolis et al. (2013)³ Carbonate (wt. %) = Assimilant weight (mg) / (Assimilant weight (mg) + Glass weight (mg))

Table 2. *Chemical compositions of starting materials.*

Oxide (wt. %)	Average composition of basaltic- andesite glass (anhydrous) ¹	1 σ (10) ^{1†}	Average composition of basaltic- andesite glass (hydrated) ¹	1 σ (9) ^{1†}	Average composition of shoshonite glass (hydrated) ²	1 σ (9) ^{2†}	Merapi limestone ¹	Vesuvius limestone ²
SiO ₂	54.11	0.60	51.83	0.43	49.85	0.21	0.28	0.02
TiO ₂	0.85	0.09	0.89	0.05	1.03	0.08	0.01	-
Al ₂ O ₃	18.98	0.17	18.08	0.24	15.95	0.67	0.13	-
FeO*	7.89	0.56	8.17	0.16	7.98	0.25	0.01	0.01
MnO	0.24	0.03	0.20	0.03	0.14	0.02	-	0.01
MgO	2.98	0.13	2.97	0.08	6.03	0.33	0.40	0.79
CaO	8.89	0.17	9.19	0.16	9.98	0.29	56.72	54.99
Na ₂ O	3.56	0.13	3.48	0.09	2.35	0.07	0.12	-
K ₂ O	2.05	0.10	2.05	0.03	4.01	0.19	-	-
P ₂ O ₅	0.29	0.04	0.34	0.04	0.70	0.09	0.03	0.02
Total	99.85		97.20		98.02		57.70	55.84
H ₂ O	-		2.23		2.00		0.15	-
CO ₂	-		-		-		44.93	44.11

¹Deegan et al. (2010).²Jolis et al. (2013).

* FeO reported as total Fe.

† 1 σ standard deviation; the number in parentheses represents the number of analyses for each sample.