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Title: Time-space focused intrusion of genetically unrelated arc magmas in the early Paleozoic Ross-Delamerian Orogen (Morozumi Range, Antarctica)

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**Abstract:** Growth of continental crust in accretionary orogenic belts takes place through repeated cycles of subduction-accretion of rock units from continental and oceanic magmatic arcs, supra-subduction zone backarcs and forearcs loaded with continent-derived materials. An ancient example relevant to magmatic arc accretion models is represented by the remnants of the Cambrian-Ordovician Ross Orogen in the Morozumi Range, Victoria Land (Antarctica). There, late Neoproterozoic phyllites host an intrusive complex which preserves a remarkably uncommon record of genetically unrelated magma pulses emplaced under a variable stress regime in a short time span: (1) a dominant K-feldspar-phyric granite, (2) fine-grained dioritic stocks and dykes, (3) a peraluminous granite; (4) a tonalitic-granodioritic dyke swarm. Laserprobe U-Pb zircon dates cluster at late Cambrian times for all these units, yet they carry differential cargoes of relict cores. Unique geochemical-isotopic signatures for both the less evolved magmas (diorite and dyke tonalite) and the most acidic ones (granite and peraluminous granite) indicate that each one of them originated from distinct sources at depth. Additionally, field relationships and chemical evolutionary trends testify for a variety of shallow level open-system processes, such as magma mingling/mixing between diorite and main granite magmas, as well as progressive incorporation of the host schists by the dyke tonalite magma. In summary, crustal growth in the Morozumi intrusive complex was contributed by fresh mantle magma issuing from the metasomatized mantle wedge, while the production of other melts did recycle different crustal portions/layers: the main granite derived from Grenville-age granulitic lower crust; the peraluminous granite from late Proterozoic upper crust; the tonalite magmas derived from subduction erosion-enriched subarc mantle and evolved by ingestion of local metasedimentary rocks. Overall, the Morozumi intrusive complex yields evidence for emplacement in the same site at the same time of magmas issuing from different sources that are usually found at different depth in the arc lithospheric section. A likely scenario to activate this specific mechanism of melt production is a subduction zone affected by subduction erosion.

1                   **Time-space focused intrusion of genetically unrelated arc magmas**  
2                   **in the early Paleozoic Ross-Delamerian Orogen (Morozumi Range, Antarctica)**

3

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10

11 ABSTRACT

12 Growth of continental crust in accretionary orogenic belts takes place through repeated cycles  
13 of subduction–accretion of rock units from continental and oceanic magmatic arcs, supra-  
14 subduction zone backarcs and forearcs loaded with continent-derived materials. An ancient  
15 example relevant to magmatic arc accretion models is represented by the remnants of the  
16 Cambrian-Ordovician Ross Orogen in the Morozumi Range, Victoria Land (Antarctica). There,  
17 late Neoproterozoic phyllites host an intrusive complex which preserves a remarkably  
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19 regime in a short time span: (1) a dominant K-feldspar-phyric granite, (2) fine-grained dioritic  
20 stocks and dykes, (3) a peraluminous granite; (4) a tonalitic-granodioritic dyke swarm.  
21 Laserprobe U-Pb zircon dates cluster at late Cambrian times for all these units, yet they carry  
22 differential cargoes of relict cores. Unique geochemical-isotopic signatures for both the less  
23 evolved magmas (diorite and dyke tonalite) and the most acidic ones (granite and  
24 peraluminous granite) indicate that each one of them originated from distinct sources at depth.  
25 Additionally, field relationships and chemical evolutionary trends testify for a variety of

26 shallow level open-system processes, such as magma mingling/mixing between diorite and  
27 main granite magmas, as well as progressive incorporation of the host schists by the dyke  
28 tonalite magma. In summary, crustal growth in the Morozumi intrusive complex was  
29 contributed by fresh mantle magma issuing from the metasomatized mantle wedge, while the  
30 production of other melts did recycle different crustal portions/layers: the main granite  
31 derived from Grenville-age granulitic lower crust; the peraluminous granite from late  
32 Proterozoic upper crust; the tonalite magmas derived from subduction erosion-enriched  
33 subarc mantle and evolved by ingestion of local metasedimentary rocks. Overall, the Morozumi  
34 intrusive complex yields evidence for emplacement in the same site at the same time of  
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37 a subduction zone affected by subduction erosion.

38

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40 erosion

41

## 42 **1. Introduction**

43

44 Magma generation in convergent settings is usually from multiple sources, including the  
45 mantle wedge, the overriding plate, and the subducting slab as well. The overriding plate can  
46 contribute as either a contaminant of uprising magmas (Davidson et al., 2005) or a direct  
47 source of melts from the upper/lower crust (Brown, 2013), and the juvenile underplates as  
48 well (Rocchi et al., 2009). The subducting lithosphere can contribute by direct partial melting  
49 (Defant and Drummond, 1990), addition to the mantle wedge of a subduction component via  
50 aqueous fluids/hydrous melts (Pearce and Peate, 1995), as well as via more massive, bulky  
51 processes such as subduction erosion (von Huene et al., 2004) and subduction of continental

52 crust (Hacker et al., 2011). Actual arc magmas are thus the outcome of a bouquet of processes  
53 (Davidson et al., 2005), including source melting degree and regime (equilibrium/  
54 disequilibrium), fractional crystallization, (possibly accompanied by assimilation), and  
55 hybridization between different magmas, from deep crustal to emplacement levels (Brown,  
56 2013). In accretionary orogens, processes of mantle modifications can occur every time  
57 subduction takes over slab rollback and backarc opening. This variety of materials and  
58 processes makes orogenic igneous complexes a rich source of information, yet difficult to  
59 disentangle.

60 A magmatic arc that underwent most of these petrogenetic processes in a convergent  
61 accretionary setting was active during the early Paleozoic at the paleo-Pacific margin of  
62 Gondwana in Antarctica. The Cenozoic uplift linked to the West Antarctic rift system led to  
63 prominent exposures of arc deep-seated terrains in clean outcrops. Among these, the  
64 Morozumi Range igneous complex in northern Victoria Land is made of a variety of intrusive  
65 rocks with well exposed mutual chronological, petrological and structural relationships. Field  
66 observations, coupled with chemical data, isotopic geochemistry and geochronology, led to the  
67 reconstruction of a scenario that shed light on potential modalities of magma formation and  
68 evolution in magmatic arcs. Additionally, new implications on the evolution of the Antarctic  
69 margin of Gondwana in Cambrian-Ordovician time are proposed to update current models.

70

71

## 72 **2. Geological setting**

73

74 Today's Transantarctic Mountains represent the roots of an orogen exposed on the shoulder  
75 of the Cretaceous–Cenozoic West Antarctic rift system (Rossetti et al., 2006a). The orogen  
76 dates back to Neoproterozoic-late Paleozoic times, when the large-scale subduction-related  
77 crustal accretion known as the Ross-Delamerian(-Tyennan) Orogeny occurred in the frame of

78 the convergence between Paleo-Pacific oceanic lithosphere and the Gondwana continental  
79 margin. Remnants of this orogen are now exposed along 3500 km of the Transantarctic  
80 Mountains, as well as in southeastern Australia, Tasmania and the South Island of New Zealand  
81 (Boger and Miller, 2004; Flöttmann et al., 1993; Foden et al., 2006; Gibson and Ireland, 1996;  
82 Glen, 2013; Stump, 1995).

83 The beginning of the Ross-Delamerian orogenic cycle, as recorded by the age of the oldest  
84 magmatism, has long been debated. A possible diachronous inception of magmatism has been  
85 postulated based on the ages of oldest igneous rocks of 514 Ma in Australia (Foden et al., 2006),  
86 of 540-520 Ma in Victoria Land (Allibone and Wysoczanski, 2002; Giacomini et al., 2007;  
87 Rocchi et al., 2004), and of around 550 Ma in central Transantarctic Mountains (Encarnación  
88 and Grunow, 1996), with detrital zircon ages and glacial clasts reaching 580-590 Ma (Goodge  
89 et al., 2012; Goodge et al., 2004). However, there is no robust evidence that the Ross Orogeny  
90 and subduction-related arc magmatism started before c. 540-530 Ma, and the diachronous  
91 starting of the orogenic cycle in Antarctica and Australia has recently been questioned (Gibson  
92 et al., 2011; Turner et al., 2009). Nevertheless, the age and geochemical patterns of intrusive  
93 rocks suggest oblique convergence along a tectonically segmented margin (Encarnación and  
94 Grunow, 1996; Goodge, 2002; Goodge et al., 2012; Goodge et al., 2004; Rocchi et al., 1998;  
95 Stump et al., 2006). The latest magmas emplaced in the Ross Orogen are about 480 Ma, and  
96 evidence for younger (460-440 Ma) amagmatic contractional tectonic activity in Victoria Land  
97 is likely related to the early stages of the formation of the Lachlan Fold Belt (Di Vincenzo et al.,  
98 2007), which is best developed in southeastern Australia (Glen et al., 2007).

99 A key segment of the Transantarctic Mountains is their Pacific Ocean-Ross Sea termination  
100 in northern Victoria Land (Fig. 1), commonly considered the along-strike continuation of  
101 Australia in Antarctica (Finn et al., 1999; Flöttmann et al., 1993; Stump et al., 1986) and  
102 described as an assembly of three fault-bounded terranes (Bradshaw et al., 1985; Gibson and  
103 Wright, 1985; Kleinschmidt and Tessensohn, 1987; Stump et al., 1983; Weaver et al., 1984),  
104 namely, from the continent toward the ocean: (i) the metamorphic Wilson terrane intruded by

105 Cambrian-early Ordovician plutonic rocks (Bomparola et al., 2007; Rocchi et al., 1998; Stump,  
106 1995), (ii) the volcanic-sedimentary Bowers terrane (Crispini et al., 2007; Weaver et al., 1984)  
107 and (iii) the Robertson Bay terrane, with thick, weakly metamorphosed turbiditic sequences  
108 intruded by Devonian granites (Di Vincenzo et al., 2014; GANOVEX TEAM, 1987; Rossetti et al.,  
109 2006b).

110 Recent works (Bracciali et al., 2009; Federico et al., 2006; Ferraccioli et al., 2009; Ferraccioli  
111 et al., 2002; Gemelli et al., 2009; Rocchi et al., 2003; Roland et al., 2004; Tessensohn and  
112 Henjes-Kunst, 2005) led to an updated model of the Ross Orogeny in Victoria Land (Rocchi et  
113 al., 2011). In this new Cambrian scenario, the convergent Paleo-Pacific margin of Gondwana  
114 consisted of a main continuous subduction zone coupled with local, transient subduction zones  
115 related to a more or less continuous ribbon of outboard pieces of stretched forearc regions.  
116 Thus, the Victoria Land portion of the Ross orogenic belt is inferred to derive neither by  
117 collision against the margin of large continents/exotic continental blocks, nor by the unique  
118 accretion of forearc oceanic lithosphere. Rather, the transient coupling between the lower and  
119 upper plates generated multiple docking of small-sized continental fragments with restricted  
120 periods of co-existence of double subduction zones. The Ross Orogen in Victoria Land was thus  
121 the result of alternate episodes of advancing and retreating subduction zone(s). Within the key  
122 segment of Victoria Land, a critical area is thus represented by the boundary zone between the  
123 Wilson arc and the forearc ribbon which underwent detachment and re-accretion to the arc  
124 (Rocchi et al., 2011). Here, the intrusive complex of the Morozumi Range (Fig. 1) recorded  
125 multiple episodes of magma emplacement in the same site, in a short time, under a variable  
126 stress regime.

127

### 128 **3. The Morozumi Range intrusive complex - field and petrographic features**

129

130 The Morozumi Range intrusive complex is hosted by the Morozumi phyllites (Fig. 2a), a

131 metasedimentary unit with a late Neoproterozoic maximum deposition age, and characterised  
132 by a regionally extensive subvertical foliation, locally overprinted by contact metamorphic  
133 effects up to 600°C at 0.2 GPa [Engel, 1984 #2916]. The Morozumi phyllites are unconformably  
134 overlain by the Permian continental-fluvial sandstones of the Beacon Supergroup, which are in  
135 turn conformably overlain by the lava flows and intruded by the columnarly jointed sills of the  
136 early Jurassic Ferrar large igneous province. The intrusive complex consists of a set of  
137 magmatic bodies with variable composition and well exposed intrusive relationships (Fig. 2).  
138 Seven different intrusive units are defined in the field based on their petrographic features (see  
139 summary in Table 1).

140 *Morozumi granite.* The Morozumi granite is by far the largest intrusive unit of the complex,  
141 extending north-south for ca. 16 km, with an east-west width of 1 km to 6 km (Fig. 1). The  
142 contacts against the country rock are subvertical, so that the intrusion has the overall aspect of  
143 a subvertical large tabular body (Fig. 2a). It mostly consists of a porphyritic monzogranite with  
144 white to pink K-feldspar megacrysts (up to 5 cm, Fig. 2b) set in a medium-grained groundmass  
145 with 10-15 vol% red-brown biotite, and allanite as a common accessory phase. K-feldspar  
146 megacrysts locally define a subvertical igneous foliation and lineation, parallel to the host  
147 schist foliation.

148 *Morozumi diorite.* The southwesternmost part of the intrusive complex include mafic stocks  
149 and dykes intruding the Morozumi granite (Fig. 2m). The contact between the mafic and felsic  
150 units is either pillowing-sharp (Fig. 2n) or diffuse (Fig. 2h), suggesting coeval emplacement  
151 relationships. The Morozumi diorite is an equigranular, fine- to medium-grained quartz-diorite  
152 characterized by the occurrence of both biotite and green amphibole totalling up to 50 vol%.

153 *Morozumi granodiorite.* This unit is a minor network of irregular bodies with fine-grained  
154 texture and high biotite content (about 20 vol%). Contacts against the Morozumi Granite are  
155 either sharp or diffuse, with evidence of progressive passage to the Morozumi granite on one  
156 side and to the Morozumi diorite on the other side (Fig. 2h).

157 *Morozumi dykes*. The intrusive complex includes several tonalitic/granodioritic dykes  
158 which, on the basis of field occurrence, have been grouped into: (i) *eastern dykes*, (ii) *western*  
159 *dykes* and (iii) *crestal dykes*. The *eastern dykes* are subvertical tabular bodies conformably  
160 intruding and interlayered with the foliation of the host Morozumi phyllites (Fig. 2a, d). Dykes  
161 are medium-grained foliated tonalites and granodiorites with 5 to 10 vol% biotite ± muscovite.  
162 At the northernmost tip of the range, the fine- to medium-grained thin peraluminous  
163 leucosyenogranite dykes with muscovite ± garnet or tourmaline intruded into the  
164 metasedimentary host are strongly foliated and boudinaged (Fig. 2d, g). Deformation style  
165 sometimes grades to mylonitic, with asymmetric boudins indicating a west-side-up movement  
166 (Fig. 2g). In other instances, fragments or streaks of Morozumi phyllites are stretched,  
167 dismembered and dispersed in the host dyke (Fig. 2f). These relationships and the parallel  
168 emplacement of eastern dykes and the main Morozumi granite body suggest the latter  
169 intruded soon after the intrusion of the dykes that were still amid the solidification process.  
170 The *western dykes* are west-dipping tabular bodies intruding the Morozumi granite at high  
171 angle with the igneous foliation, and the Jupiter granite as well. When intruding the Morozumi  
172 granite, dykes are fine-grained foliated tonalites with about 15 vol% biotite. Dykes intruding  
173 into the Jupiter granite are medium-grained granodiorites with about 20 vol% biotite and up to  
174 5 vol% muscovite. Western dikes intrude on ramps that crosscut the vertical foliation of the  
175 granite host, with some ductile deformation of the granite foliation that should thus have been  
176 not completely cooled down at the time of dyke emplacement. Kinematic indicators for this  
177 ductile deformation indicate a top-the-east relative movement. The *crestal dykes* are tabular  
178 intrusions cropping out in the summit part of the ridge: gently west-dipping tabular bodies,  
179 locally cutting the intrusive foliation of the the Morozumi granite; they are fine-to medium-  
180 grained leucomonzogranites to leucogranodiorites with less than 5 vol% biotite along with  
181 minor muscovite. In the crestal zone of the range, the dykes crosscut the subvertical igneous  
182 foliation with knife-sharp contacts (Fig. 2c) and top-to-the-east relative movement, as testified  
183 by mylonitic shear at the contact (Fig. 2e), the same relative movement observed for the



184 western dykes. In summary, in the eastern side of the complex the emplacement of the granite  
185 appears contemporaneous to slightly later than the dyke emplacement, while in the western  
186 side and the topmost outcrops the granite foliation is crosscut by the dykes that thus were  
187 emplaced when the granite was solidifying (west) or solid (crest).

188 *Jupiter granite.* This intrusive unit is named after the main spectacular outcrop in the  
189 southern part of the Jupiter Amphitheatre, where a medium- to coarse-grained massive body of  
190 peraluminous biotite+muscovite-monzogranite neatly crosscut the country rock (Fig. 2i). Field  
191 relationships with the Morozumi granite have not been observed, while peraluminous  
192 syenogranitic massive bodies of the Jupiter granite unit are intruded by tabular bodies along  
193 the western flank of the intrusive complex (western tabular intrusions). Here the Jupiter  
194 medium to coarse-grained granite contains igneous muscovite up to 5-10 vol%.

195 *Morozumi leucogranites/aplites.* Scattered minor bodies of fine- to medium-grained,  
196 equigranular peraluminous, muscovite ± garnet or tourmaline-bearing leucosyenogranites are  
197 found both within the Morozumi granite and at the contact between the granite and the host  
198 schists.

199

#### 200 **4. Analytical methodologies**

201

202 Major elements (Table 2) were determined by X-ray fluorescence (XRF-ARL 9400XP) on  
203 glass beads, at Dipartimento di Scienze della Terra, Università di Pisa, with precision between  
204 1% and 4% RSD for most elements, except TiO<sub>2</sub>, MnO, CaO, Na<sub>2</sub>O, (5–8% RSD) (Tamponi et al.,  
205 2003). Trace elements (Table 2) were determined by inductively coupled plasma-mass  
206 spectrometry (ICP-MS, VG PQII Plus) (Rocchi et al., 2002) at Dipartimento di Scienze della  
207 Terra, Università di Pisa, after dissolution with HF-HNO<sub>3</sub> mixture in screw-top PFA vessels on a  
208 hotplate at 120 °C. Analyses were performed by external calibration using the matrix-matching  
209 geochemical reference sample BE-N. The correction procedure includes blank subtraction,

210 instrumental drift correction (Rh-Re-Bi internal standardization and repeated analysis of a  
211 drift monitor), and oxide-hydroxide interference correction. Precision, evaluated by replicate  
212 dissolutions and analyses of in-house and international silicate rock reference samples, is  
213 generally between 2% and 5% RSD for most elements, except Cr, Ni, Pb (6–11% RSD).  
214 Zirconium was determined via XRF on pressed powder pellets.

215 Sr and Nd isotopic compositions were determined using a Finnigan MAT 262V  
216 multicollector mass spectrometer at Istituto di Geoscienze e Georisorse-CNR, Pisa, following  
217 separation of Sr and Nd using conventional ion exchange procedures. Measured  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios  
218 have been normalized to  $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$ ;  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios to  $^{146}\text{Nd}/^{144}\text{Nd} = 0.7219$ . During  
219 the course of this study, the external reproducibility for NIST-SRM987 and La Jolla Nd  
220 standards were  $^{87}\text{Sr}/^{86}\text{Sr} = 0.710250 \pm 24$  and  $^{143}\text{Nd}/^{144}\text{Nd} = 0.511858 \pm 14$  (2 SD), respectively.  
221 The measured Nd isotopic ratios have been adjusted to a value for La Jolla standard of  
222  $^{143}\text{Nd}/^{144}\text{Nd} = 0.511850$ .

223 Geochronological data were obtained by zircon U-Pb analyses. After crushing and sieving,  
224 zircons were concentrated from the 160-250  $\mu\text{m}$  grain sizes using standard separation  
225 techniques. About 100 zircons for each sample were cast in epoxy resin and polished to a 0.3  
226  $\mu\text{m}$  alumina paste finish to expose mid-section of crystals. Crystal for geochronological  
227 analyses have been selected on the basis of textural observations related to inclusions,  
228 occurrence/type of core and/or rim, zoning, etc carried out by Scanning Electron Microscopy-  
229 Cathodoluminescence (SEM-CL) imaging by a Philips XL30 SEM at Dipartimento di Scienze  
230 della Terra, Università di Pisa. U-Pb analyses by Laser Ablation-Inductively Coupled Plasma-  
231 Mass Spectrometry (LA-ICP-MS) were performed at Dipartimento di Fisica e Geologia,  
232 Università di Perugia (Italy) (Alagna et al., 2008) using a Thermo Electron X7 ICP-MS coupled  
233 to a New Wave UP213 frequency quintupled Nd:YAG laser. U-Pb zircon analyses were  
234 calibrated with standard zircon 91500 using a spot size of 25  $\mu\text{m}$  and zircon GJ1 has been used  
235 as quality control. The error associated with the standard reproducibility has been combined  
236 quadratically with the counting statistics of each analysis.

237

## 238 **5. Geochronological data**

239

240 In order to define the temporal evolution for the emplacement of the different intrusive  
241 units, as well as to investigate differential source inheritance in different magmas, seven  
242 samples of magmatic rocks along with one sample of the metamorphic host rock have been  
243 investigated for U-Pb zircon geochronology (Table SM1).

244 *Morozumi granite.* Data from zircons with a well-preserved oscillatory magmatic zoning  
245 give a concordia age of  $493.7 \pm 8.7$  Ma (with decay-constant errors ignored, probability of  
246 concordance = 0.995, Fig. 3a), interpreted as the emplacement age for Morozumi granite.  
247 Zircons with ghost zoning, i.e. areas where typical igneous zoning is lacking or poorly defined,  
248 yielded younger ages and are interpreted as related to post-igneous reworking. The oldest ages  
249 clustering around 1000 Ma (Grenville-like, Fig. 4) are from three zircon cores lighter than the  
250 rim, with truncated zoning, interpreted as relict cores.

251 *Morozumi diorite.* The zircon yield was poor, comprised of tiny, elongated crystals with no  
252 well-defined oscillatory zoning and the obtained data does not allow to constrain the time of  
253 emplacement. Nevertheless, the clear hot mingling relationships with Morozumi granite (Fig.  
254 2n) led to infer a coeval emplacement timing.

255 *Morozumi granodiorite.* Zircons with well-defined zoning have a concordia age, obtained  
256 from 5 points analysis and excluding reworked zircon areas, of  $485.3 \pm 5.4$  Ma (probability of  
257 concordance = 0.88, Fig. 3b). Older age clusters between 520 Ma and 650 Ma are obtained from  
258 zircons with unzoned light cores, interpreted as reworked areas. Two data points on cores give  
259 Grenvillian ages, and a single analysis yielded a Proterozoic age (Fig. 4).

260 *Morozumi dykes. Eastern dyke.* Zircon crystals are characterized by a dark core and a light  
261 rim. A few crystals have an innermost light core, and some others show oscillatory or  
262 convoluted zoned cores. Five analyses of light rims provided a Concordia Age of  $496.7 \pm 6.8$  Ma

263 (probability of concordance = 0.92, Fig. 3c), interpreted as the emplacement age. Three  
264 additional analyses of cores give ages between ca. 550 and 1900 Ma (Fig. 4). *Western dyke*. The  
265 zircon yield was modest and only 12 analyses were performed for a western dyke sample.  
266 These zircons are unzoned or show convolute/ghost zoning. Six analyses (from both cores and  
267 rims) range between  $437\pm 11$  Ma and  $573\pm 14$  Ma. The lack of any clear magmatic zoning  
268 suggests post-magmatic reworking for these zircon crystals. Five additional analyses from  
269 relict cores ( $973\pm 23$  to  $1120\pm 25$  Ma) indicate the occurrence a Grenville-age component in the  
270 magma source (Fig. 4).

271 *Jupiter granite*. Once data from zircons with ghost zoning were discarded, five zircon  
272 crystals from sample JG1 with well-defined oscillatory zoning give a concordia age of  $495.3\pm 5.7$   
273 Ma (probability of concordance = 0.78, Fig. 3d), interpreted as the emplacement age. Sample  
274 JG2 did yield just a few crystals, with two cores giving Grenvillian ages (Fig. 4).

275 *Morozumi phyllites*. The studied sample has been collected in the southern part of the  
276 Morozumi Range, far from contacts against igneous rocks, to study a pristine sample not  
277 affected by thermal effects, that are indeed not been observed in both thin sections and SEM-CL  
278 zircon images. The only overgrowth observed has a  $^{206}\text{Pb}/^{238}\text{U}$  age indistinguishable from that  
279 of the core, within the same crystal ( $2299\pm 80$  vs.  $2293\pm 81$  Ma, respectively). The pristine  
280 characters of zircons allow to infer the youngest concordant zircon ages ( $571\pm 22$  and  $601\pm 22$   
281 Ma) as the oldest limit for the deposition of this unit, at the very end of the Neoproterozoic.  
282 Younger Cambrian deposition age for the Morozumi phyllites are inferred, based on  
283 unpublished detrital zircon ages [Henjes-Kunst, 2003 #2569]. A deposition age at the boundary  
284 between latest Neoproterozoic and Cambrian (c. 545 Ma) is inferred based on detrital zircon  
285 ages for the Berg Group and the Priestley Formation of the Wilson terrane, as well as the Molar  
286 Formation of the Bowers terrane (Adams et al., 2013).. Our data, plotted in a cumulative  
287 probability plot (Fig. 4) shows a wide range of ages, with well-defined peaks slightly older than  
288 600 Ma (< 20% of zircons) and around Grenville age (~ 40%); three minor peaks occur at  
289 Paleoproterozoic-Archean times (total of ~ 40%). The lack of thermal imprints or

290 metamorphic overgrowths calls for a mechanism of cooling of the intruded magmas with no  
291 significant conductive heating of the country rock.

292 In summary, individual intrusive units provide indistinguishable emplacement ages for the  
293 main granite ( $493.7 \pm 8.7$  Ma), the eastern dykes ( $496.7 \pm 6.8$  Ma) and the peraluminous granite  
294 ( $495.3 \pm 5.7$  Ma). Geochronological data for the Morozumi diorite does not define an  
295 emplacement age, yet field relations indicate contemporaneous emplacement with Morozumi  
296 main granite (Fig. 2n). The Morozumi granodiorite provides a slightly younger nominal age of  
297  $485.3 \pm 5.4$  Ma, although still indistinguishable at  $2\sigma$  level from those of the other units, in close  
298 agreement with field observations (Fig. 2h). For the Morozumi western dykes, the  
299 emplacement age could not be defined geochronologically, nevertheless field relations with the  
300 Morozumi granite (the dykes cut the granite igneous foliation at high angle with ductile  
301 deformation of the granite) suggest once again an emplacement age almost coeval with the  
302 Morozumi granite. In summary, a short time interval of no more than a few Ma around ca. 495  
303 Ma is inferred for the emplacement of the whole Morozumi intrusive complex.

304 As for the relict zircon signatures, all the intrusive units (granite, granodiorite, dykes,  
305 peraluminous granite) contain relict cores with Neoproterozoic and Mesoproterozoic  
306 (Grenville) ages, along with rare Paleoproterozoic cores. A couple of Archean cores were found  
307 only in the peraluminous Jupiter granite.

308

## 309 **6. Geochemical and isotopic data**

310

311 The lithological diversity of the Morozumi intrusive complex is mirrored by large  
312 variabilities in the major element compositions. The Total Alkali-Silica (TAS) diagram readily  
313 shows that the whole data set splits into two main subalkaline trends with different levels of  
314 alkali enrichment (Fig. 5). The Morozumi diorite-granodiorite-granite samples define a roughly  
315 linear trend barely reaching the granite field. At lower alkali contents, the samples from

316 Morozumi dykes define a separate roughly linear trend from the granodiorite well into the  
317 granite field. Both trends have slopes that cut across the usual boundaries between magmatic  
318 associations, in contrast with the typical magmatic evolution in closed systems.

319 The trace element distributions of the most mafic rocks from the two trends share common  
320 features, such as a significant enrichment for the most incompatible elements, along with  
321 marked negative anomalies for Nb-Ta and minor for Ti (Fig. 6). Besides those similarities, the  
322 patterns also show some differences: the Morozumi diorites, despite having a silica content  
323 higher than that of the most mafic Morozumi tonalite dyke, feature a higher enrichment in rare  
324 earth elements (REE), most marked for heavy REE.

325 Among mafic igneous rocks from the early Paleozoic Ross Orogen of northern Victoria Land,  
326 the Morozumi diorite display strong affinity to potassic rocks such as the Abbott gabbro and  
327 the Vegetation lamprophyres, representing mantle-derived magmas emplaced in the late  
328 orogenic to postcollisional stage (Di Vincenzo and Rocchi, 1999; Rocchi et al., 2009). Among  
329 igneous rocks of comparable compositions, the Morozumi tonalite dyke has a trace element  
330 distribution quite similar to that of the Confusion tonalite, a calc-alkaline unit characterizing a  
331 tectonomagmatic stage earlier than the Abbott and Vegetation potassic events (Rocchi et al.,  
332 2004).

333 The Sr and Nd isotopic systematics help clarifying the petrologic affinities of the Morozumi  
334 igneous rocks (Fig. 7). The Morozumi diorite has Sr and Nd isotopic ratios higher and lower  
335 than those of the Abbott gabbro, respectively, and definitely similar to those of the Vegetation  
336 lamprophyres (Di Vincenzo and Rocchi, 1999): the Morozumi diorite has thus trace element  
337 and isotopic features typical of a potassic, post-collisional magma. The Morozumi tonalite dyke  
338 samples have scattered Sr isotopic compositions, making it difficult any direct comparison (see  
339 further on).

340 The most felsic rocks of the two trends plot separately: a gap of 5 wt% SiO<sub>2</sub> separates rocks  
341 with the same alkali (and K) contents. The Jupiter granite differs from all the other Morozumi

342 felsic rocks owing to its peraluminosity (Alumina Saturation Index: 1.1-1.2), and low Nd  
343 isotopic ratios, coupled at odd with relatively low Sr isotopic ratios (< 0.711).

344

## 345 **7. Discussion - Petrogenetic issues**

346

347 The two roughly parallel compositional trends (Fig. 5) follow possible evolutionary paths  
348 starting from two different parental magmas. The slopes of these trends, cutting across the  
349 usual boundaries of magmatic associations (Fig. 5), coupled with Sr-Nd isotope data (Fig. 7),  
350 rule out closed-system compositional evolution for both trends. The 5 wt% SiO<sub>2</sub> gap separating  
351 felsic rocks with the same alkali (and K) contents, also lend support to the hypothesis of a  
352 separate origin for the compositional groups of felsic rocks. Therefore, simple petrogenetic  
353 evolutionary relationships among the various igneous units are precluded and a minimum  
354 number of different magmas had to be involved in the genesis of the Morozumi complex,  
355 namely: a mafic potassic dioritic melt (Morozumi diorite), a mafic tonalitic melt (Morozumi  
356 mafic dykes), a felsic metaluminous granitic melt (Morozumi granite), as well as a felsic  
357 peraluminous granitic melt (Jupiter granite). Additionally, a possible high-silica metaluminous  
358 melt could have existed (Morozumi felsic dykes), yet field relationships indicate that a mass  
359 contribution of the Morozumi phyllites to the compositional variability of the dykes has to be  
360 taken into account (Fig. 2f). In the following we are going to first discuss the origin of the  
361 different melts at source depth, then their interaction/evolution at shallower levels during  
362 ascent/emplacement.

363

### 364 *7.1. What type of mantle? Origin of mafic "parental" melts:*

365

366 *Morozumi diorite.* The reconstruction of the type of mantle source activated for the  
367 production of the most mafic Morozumi magmas is hampered by the absence of any rock which

368 could be assumed as representative of a primary mantle melt. Even the least silicic Morozumi  
369 diorite samples are indeed quite evolved, having an overall andesitic composition.  
370 Nevertheless, an origin from remelting of basaltic rocks at depth in the crust is not supported,  
371 owing to the lack of any evidence for Ross basaltic magmatism in the area. The Morozumi  
372 diorites are therefore best considered as representative of mantle melts which underwent  
373 some differentiation, potentially including open-system processes. Indeed, simple fractional  
374 crystallization is not sufficient to explain the isotopic variability of the diorite samples. Also the  
375 overall compositional trend encompassing mingled rocks (the Morozumi granodiorite show  
376 evidence for mingling-mixing relationships between the Morozumi diorite and granite melts;  
377 Fig. 2h) lend neat support to the occurrence of open-system interaction processes. To back-  
378 reconstruct the possible chemical and isotopic composition of the parental basaltic melt, a  
379 modeling of assimilation and fractional crystallization (AFC) has been run out (DePaolo, 1981).  
380 In order to derive the parameters to be used for trace element and isotopic AFC modelling,  
381 preliminary mass balance calculations have been performed: the major element composition of  
382 the finest-grained and mafic Morozumi diorite (sample 23.12.05 DS7) has been obtained  
383 starting from a common high-Al Andean basalt (Wilson, 1989) via fractionation of 9 wt%  
384 orthopyroxene, 1 wt% clinopyroxene and 25 wt% plagioclase, plus the addition of about 23  
385 wt% of a migmatitic leucosome, with mineral and migmatite compositions after similar  
386 gabbroic-dioritic rocks from the Wilson terrane (Di Vincenzo and Rocchi, 1999). The  
387 parameters thus obtained (mass fraction crystallised =  $M_{FC} = 0.35$ ; ratio of the rate of  
388 assimilation to the rate of fractional crystallisation =  $a = 0.66$ ) have been used to run AFC  
389 modelling for Sr-Nd systematics, obtaining values of  $^{87}\text{Sr}/^{86}\text{Sr}_{(500 \text{ Ma})} = 0.706$  and  $\epsilon_{\text{Nd}(500 \text{ Ma})} = -1$   
390 for a possible primitive Morozumi mafic melt. These values point out that the main Morozumi  
391 original mafic melt had to have geochemical and isotopic features which are enriched  
392 compared to common primitive arc basalts. These features could either derive from (i)  
393 metasomatic enrichment in the mantle wedge, or from (ii) refilling of a fractionating-  
394 assimilating magma pond at the mantle-crust interface (R-AFC) (O'Hara, 1977), which is likely



395 to happen in a zone of underplating characterised by periodic or continuous input of fresh  
396 mantle magma; this process could potentially generate enriched compositions from non-  
397 enriched sub-arc mantle wedge. However, the Morozumi diorite has Sr and Nd contents that  
398 are above the threshold over which a metasomatically enriched source has necessarily to be  
399 invoked (Di Vincenzo and Rocchi, 1999). The zircon yield from the Morozumi diorite, although  
400 quite poor, allow to establish the absence of relict cores, further supporting the full-mantle  
401 origin of this magma. In conclusion, the original/parental Morozumi diorite melt had to derive  
402 from a metasomatized mantle wedge characterised by an  $\epsilon_{\text{Nd}(500 \text{ Ma})}$  not higher than -1 and a  
403  $^{87}\text{Sr}/^{86}\text{Sr}_{(500 \text{ Ma})}$  not lower than 0.706.

404 *Morozumi tonalite.* This unit does not include any rock which could be assumed as  
405 representative of a primary mantle melt. Even the least silicic rock samples are quite evolved,  
406 so their composition could in principle derive from several different processes. The  
407 composition of tonalite is not compatible with derivation from melting of metasedimentary  
408 material with protoliths such as pelites, greywackes or volcanoclastics (Fig. 8a). An alternative  
409 process is the chemical evolution of a mantle melt, which however should leave some traces of  
410 basic-intermediate products: the lack of products with  $\text{SiO}_2 < 65 \text{ wt}\%$  is in contrast with this  
411 hypothesis. Another alternative is the crustal contamination of the Morozumi diorite melt,  
412 which however conflicts with the observation that samples with the same  $\epsilon_{\text{Nd}(500 \text{ Ma})}$  have a  
413 difference of 10 wt% in the  $\text{SiO}_2$  content. On the other hand, the Morozumi tonalites are very  
414 similar to high-silica adakites for major elements composition as well as for their high La/Yb,  
415 low Yb, high Sr/Y and low Y (Fig. 8b). However, the adakites interpreted as products of young  
416 slab melting, as for their original definition (Defant and Drummond, 1990), usually have  
417 positive  $\epsilon_{\text{Nd}}$  (Castillo, 2012), in opposition to the low  $\epsilon_{\text{Nd}}$  of the Morozumi tonalite melt. Such a  
418 magma should therefore derive from melting of mafic material residing in the (lower?) crust.  
419 Remelting of an underplate made of Morozumi potassic diorite is not likely because the  
420 tonalite has much lower K content than the diorite; additionally, the observed  $\epsilon_{\text{Nd}}$  values

421 between -5 and -6 would require melting of evolved-hybrid diorite rather than diorite itself,  
422 but the observed hybridization process is limited, likely occurring en route to the emplacement  
423 level, and not at the underplate level. Another potential crustal basic source is represented by  
424 normal gabbros or basic amphibolites. Melting experiments (Sisson et al., 2005) indicate that  
425 the Na/K/Ca relationships of the Morozumi granodiorite/tonalite are matched by melting  
426 hornblende gabbro at high T (>925°C), yet these melts have SiO<sub>2</sub><60 wt%; using biotite- or  
427 quartz-bearing hornblende gabbro as a source and/or lowering the melting temperature yields  
428 products with higher SiO<sub>2</sub>, comparable with Morozumi granodiorite/tonalite, but with  
429 somewhat higher K/Ca ratios (Fig. 8a). Melting experiments of amphibolites with alkali  
430 basaltic and island arc tholeiitic composition (Rushmer, 1991) yield melts with lower SiO<sub>2</sub> and  
431 higher Ca/Na ratios with respect to the Morozumi tonalite (Fig. 8a); on the other hand,  
432 "synthetic amphibolites" (a mix of hornblende, albite and quartz) at high T=975 °C give melts  
433 with a composition compatible with the Morozumi tonalite (Rushmer, 1991). Also the  
434 occurrence of even a few old zircons with variable Proterozoic ages helps ruling out simple  
435 young slab or underplate melting, rather supporting addition of some crustal material to the  
436 mantle wedge. The process responsible of such an addition could be ascribed to the category of  
437 subduction erosion that can also explain strong REE fractionation coupled with high Sr content  
438 without invoking subduction of young slab at garnet depth (Kay et al., 2005). Large-scale  
439 subduction erosion or even larger-scale subduction of small continental block(s) can bring into  
440 the mantle wedge crustal materials, which then exhumed due to slab roll-back, becomes  
441 incorporated into the overriding plate (Brun and Faccenna, 2008) and can contribute to melt  
442 generation.

443

444 *7.2. What type of crust? Origin of felsic melts*

445

446 The four types of felsic intrusive units of the Morozumi complex have very different sizes:  
447 the Morozumi granite is by far the most voluminous, constituting most of the mass of the  
448 intrusive complex; the Jupiter peraluminous granites are found in two main large outcrops; the  
449 Morozumi leucogranites represent very small, vein- or pod-like bodies; the Morozumi dykes  
450 with felsic composition are leucogranitic portions of tonalite-granodiorite dykes. The chemical  
451 and isotopic compositions of these four felsic rock types point out that they also have distinct  
452 origins.

453 The Jupiter granite and the vein-like Morozumi leucogranites have compositions typical of  
454 minimum melts from metasedimentary protoliths such as greywacke or pelite rocks (Fig. 8a).  
455 These minor units have Sr-Nd compositions displaced from the main trend towards lower Sr  
456 isotopic ratios (Fig. 7), as observed elsewhere for melting occurring in disequilibrium  
457 conditions (Farina et al., 2014; Farina and Stevens, 2011; Harris and Ayres, 1998). Relict  
458 zircons show an age spectrum including scattered evidence for recycling of early Proterozoic  
459 and Archean material. In synthesis, the source that can be inferred for the Jupiter granite is a  
460 late Proterozoic metasedimentary upper crust.

461 On the other hand, the main Morozumi granite has a composition that is too Ca-rich to be  
462 derived from melting of common pelite-metagreywacke-volcaniclastic sources (Fig. 8a).  
463 Morozumi granites have Sr-Nd isotopic compositions that plot on the general trend linking  
464 mantle products and crustal magmatic rocks deriving from melting of different crustal levels,  
465 from deep granulite to upper crust metasediments (Di Vincenzo and Rocchi, 1999). Low  $\epsilon_{\text{Nd}(500}$   
466  $\text{Ma})$  and moderately high  $^{87}\text{Sr}/^{86}\text{Sr}_{(500 \text{ Ma})}$  lead to prefer a granulite source. The cargo of relict  
467 zircons of the Morozumi granite is restricted to Grenvillian and mid-Neoproterozoic ages.  
468 Overall, the origin that can be envisaged for the main Morozumi granite is from melting of a  
469 Grenville-age deep crustal granulite, isotopically similar to the enderbites and metaenderbites  
470 from northern Victoria Land, which have  $^{87}\text{Sr}/^{86}\text{Sr}_{(500 \text{ Ma})}=0.710$  to  $0.714$  and  $\epsilon_{\text{Nd}(500 \text{ Ma})}= -7.8$  to -

471 8.4, in turn geochemically similar to the Grenville-age Antarctic charnockites (Talarico et al.,  
472 1995).

473 The leucogranitic portions of the Morozumi dykes have Na-Ca-K relationships (Fig. 9) and  
474 Sr-Nd isotopic ratios (Fig. 7) potentially relating them to a source similar to that of the  
475 Morozumi granite. However, their higher silica content (Fig. 8a) and their field and  
476 geochemical relationships with host rocks suggest a different scenario (see further on).

477

### 478 *7.3. Shallow processes: Origin of the diorite-granite trend*

479

480 The rocks forming this trend (diorite, granodiorite, granite) display field evidence attesting  
481 for mingling-mixing relationships among them: dioritic melt did form pillows within the  
482 granite melt (Fig. 2n), and diffuse diorite-granite contacts are observed, with a granodiorite  
483 facies in between (Fig. 2h). Effective mixing was likely limited, as suggested by the minor  
484 volume of the granodiorite facies. Additionally, the granodiorite sometimes show angular  
485 blocks in sharp contact towards the host main granite, suggesting that the mingling process  
486 was limited, and rapidly evolved to more rigid behaviour. Chemical data (Fig. 5) and Sr-Nd  
487 isotopic systematics (Figs. 7 and 9) support field observations, with the granodiorite values  
488 fitting a simple mixing trend between the diorite and granite compositions.

489

### 490 *7.4. Shallow processes: Origin of the tonalite-leucogranite trend*

491

492 The Morozumi dykes metaluminous tonalites grade into evolved products in which white  
493 mica becomes a main mineral phase. This suggests a possible interaction with a peraluminous  
494 material. Potential candidates are both the Jupiter granite and the very minor  
495 aplite/leucogranites or the metasedimentary country rock. However, field evidence of  
496 stretched portions and streaks of mica-rich material incorporated from the host schist (Fig. 2f)

497 into the subvertical eastern dykes suggests that the tonalite melt did engulf and stretch  
498 variable amounts of schist, leading to progressively silica-richer, more peraluminous  
499 compositions. Also in this case, field evidence is supported by Sr isotopic data (Fig. 9)  
500 indicating mixing relationships between the most primitive and the most evolved dykes.  
501 However, the lack of any evidence for thermal metamorphism in the host schist suggests that  
502 the interaction process was quick, with very limited heat transfer. Therefore, the term mixing  
503 is here used to describe the chemical result of the process, yet from the petrographic point of  
504 view the schist is just incorporated, stretched, dismembered and dispersed in the melt to the  
505 point that distinguishing the origin of single crystals is actually a hard, if possible, task.

506

## 507 **8. Summary and implications**

508

509 Field evidence, coupled with geochronological and petrographic-chemical-isotopic data  
510 outline a scenario of coeval emplacement of distinct melts issuing from different sources in the  
511 mantle and from different crustal levels as well (Fig. 10). These melts were interacting  
512 between them and/or with country rock during emplacement in a swiftly changing tectonic  
513 regime.

514

### 515 *8.1. Magma origin and emplacement*

516

517 Distinct magma sources were activated at almost the same time in the metasomatized  
518 mantle, in the upper crust, in the lower crust and in a domain where mantle and crust were  
519 mixed up by subduction erosion. The first magma emplaced in the Morozumi complex was a  
520 tonalite melt, intruded as subvertical dykes, parallel to the host schist foliation (Fig. 2a, d). The  
521 almost simultaneous emplacement of the main Morozumi granitic melt did generate a local  
522 compressional (flattening) regime with subhorizontal  $\sigma_1$  leading to boudinage (Fig. 2g) of the

523 vertically emplacing tonalite magma (Fig. 2a, d) which incorporated pieces of the host schists  
524 which were further stretched and dismembered into the crystallizing melt (Fig. 2f, g). The main  
525 intrusive body of the Morozumi granite was emplaced taking the shape of a subvertical large  
526 tabular body (Fig. 2a) with vertical magmatic foliation and lineation (Fig. 2b, l), in a transient  
527 stress regime ruled by the emplacement of main granitic melt. After that, the regional stress  
528 became dominant again, and further tonalite melt was thus intruded crosscutting the vertical  
529 foliation in a top-to-the-east thrusting regime evolving from ductile (melt present, Fig. 2l, o) to  
530 mylonitic (Fig. 2c, e). The emplacement regime of the peraluminous Jupiter granite can be  
531 inferred only from its intrusive relationships with the host schist, that indicate a brittle regime  
532 with shallowly dipping contacts (Fig. 2i), suggesting an overall tectonic regime finally coming  
533 to a rest.

534 Thus, four different melts were produced from the Morozumi lithospheric section in a short  
535 time span around 495 Ma (Fig. 10): (1) peraluminous granitic melts from the upper crust, (2)  
536 significant volumes of magma produced from recycling of Grenville-age, granulitic lower crust,  
537 (3) small volumes of tonalite melt deriving from subduction erosion-enriched mantle wedge,  
538 and (4) minor mafic melt from a metasomatized mantle. The contemporaneous activation of  
539 these four different sources is a remarkable example of the complexity of magma sources  
540 involved in a magmatic arc. Overall, it suggests that physical conditions for partial melting  
541 were attained contemporaneously in domains usually found at very different depths, and that  
542 produced melts found common pathways to emplace in the same site in the arc. In a  
543 subduction arc setting, such a scenario can occur when a relevant process of subduction  
544 erosion, possibly combined with underthrusting, bring into the sub-arc mantle large slices of  
545 upper and lower continental crust, and in the mantle wedge minor amounts of crustal material.

546

547 *8.2. Geodynamic implications*

548

549 Such a scenario fits well into the space-time geological setting of the Morozumi Range  
550 intrusive complex, which is located in the area facing the classical boundary between Wilson  
551 and Bowers terranes and was built up by the latest pulses of the Wilson magmatic arc, a few  
552 Ma after the docking of the Bowers arc-backarc system against the Wilson arc, which did occur  
553 a few km to the NE (Rocchi et al., 2011). The deeply underthrust slices of continental crust  
554 and mafic material of the arc-backarc system reached ultra-high-pressure depths (Palmeri et  
555 al., 2011) and were soon exhumed, making it available in the mantle wedge and the overriding  
556 plate different sources such as metasomatized mantle, old lower crust and upper crust as well.  
557 Magma thus generated, rose and emplaced in the Morozumi Range in a NE-SW convergent  
558 tectonic regime, evolving to top-to-the NE midcrustal shearing (Rossetti et al., 2011) linked to  
559 the final docking from the NE of the Admiralty ribbon (Rocchi et al., 2011), also called  
560 Admiralty Block (Rossetti et al., 2011).

561 In a larger, orogen-scale view, this scenario leads to infer that the Antarctic margin of  
562 Gondwana in Victoria Land during Cambrian-Ordovician times was behaving in a differential  
563 way at different -current- latitudes. In the central Transantarctic Mountains, simple Sr-Nd  
564 isotope trends across the chain (Borg et al., 1990) indicate near-orthogonal subduction.  
565 Towards the north, in central Victoria Land and southernmost northern Victoria Land, the  
566 margin behaved in a more complex way, with along-strike shifting of forearc slivers (Rocchi et  
567 al., 1998). In northernmost northern Victoria Land, multiple continental and oceanic arcs were  
568 active, and convergence was not a straightforward process, rather accretion did alternate with  
569 detachment of continental material-laden forearc-backarc from the main margin (Rocchi et al.,  
570 2011). In the southernmost sectors, a single subduction zone is invoked (Goodge et al., 2012),  
571 and a tectonic lock-up finally occurred at around 490 Ma, with production of postcollisional  
572 lamprophyres and granites emplaced in a brittle regime (Rocchi et al., 2009). Differently, the  
573 northern sector was characterized by a more complex geodynamic evolution, with alternating  
574 periods of single or double SW-verging active subduction zones, and contractional-  
575 accretionary stages alternated with extensional periods of the supra-subduction forearc

576 (Rocchi et al., 2011). In this framework, the classification of granites in syn- or post-tectonic  
577 becomes increasingly meaningless towards the north, where deformation occurred at several  
578 stages, until during the emplacement of very late intrusions, and even after that (Di Vincenzo et  
579 al., 2007), in absence of coeval magmatic activity.

580

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585 as well as the Editor Nelson Eby, leading to significant improvement of the paper.

586

## 587 **Tables**

588

589 Table 1. Summary of field, petrographic and geochronological features of the lithologic  
590 units of the Morozumi Range intrusive complex..

591 Table 2. Major elements, trace elements and Sr-Nd isotopic data.

592 Table SM1. U-Pb LA-ICP-MS zircon data.

593

## 594 **Figure Captions**

595

596 Fig. 1. (a) Location map of the Transantarctic Mountains and Victoria Land in the frame of  
597 the early Paleozoic Ross-Delamerian orogen in Antarctica-Australia-Tasmania before  
598 the breakup of Gondwana (Foster et al., 2005; Glen, 2005). (b) Satellite view of  
599 Victoria Land, based on <http://rapidfire.sci.gsfc.nasa.gov>. For comparison between  
600 the representation of Victoria Land reported here (Rocchi et al., 2011) and that



601 reported in papers using the classical partition of northern Victoria Land into three  
602 terranes: (i) the Wilson terrane is here represented by the Wilson continental arc, (ii)  
603 the Bowers terrane is represented by the Bowers arc-backarc plus the southernmost  
604 part of the Tiger arc, and (iii) the Robertson Bay terrane is represented by the  
605 Admiralty crustal ribbon plus the northernmost part of the Tiger arc. IG:  
606 "postcollisional" granites and felsic dykes; VL: "postcollisional" lamprophyric dykes  
607 (Rocchi et al., 2009). (c) Magnification of the Morozumi Range, with location of field  
608 photographs reported in Figure 2.

609 Fig. 2. Field photographs of the Morozumi Range intrusive complex. (a) Panoramic view  
610 from NE of the northern Morozumi Range, vertical relief 1300 m, horizontal view ca.  
611 10 km. Dark rocks: Morozumi phyllites; light rocks: Morozumi granite. (b) Morozumi  
612 granite, width of viewfield 40 cm. (c) Morozumi crestal dyke intruding Morozumi  
613 granite with a top-to-the-east shear band at the lower contact. (d) Morozumi Range,  
614 view from NW: Morozumi granite on the right (west), Morozumi phyllites to the left  
615 (east) intruded by Morozumi eastern dykes. (e) Detail of (c): top-to-the-east shear  
616 band at the contact between Morozumi crestal dyke (above) intruding Morozumi  
617 granite (below). (f) Streaks of stretched Morozumi phyllites into Morozumi eastern  
618 dyke; person for scale. (g) Stretched portions of eastern dykes (light colour), with  
619 asymmetric boudinage structures, intruded into Morozumi phyllites (dark); hammer  
620 for scale. (h) Detail of (m): diffuse contact between Morozumi diorite, granodiorite  
621 and granite; hammers for scale. (i) Jupiter granite, light-coloured, intruding Morozumi  
622 phyllites with subhorizontal contact at Jupiter Amphitheatre; vertical relief 300 m. (l)  
623 Western dyke, below, intruding Morozumi granite; person for scale. (m) Dykes and  
624 stocks of Morozumi diorite intruding Morozumi granite; persons for scale. (n) Detail  
625 of (m): pillows of Morozumi diorite into a felsic facies of Morozumi granite at the  
626 diorite-granite contact; hammer for scale. (o) Detail of (l): western dyke (below),  
627 intruding vertically foliated Morozumi granite and aplite as well; hammer for scale.

628 Fig. 3.  $^{206}\text{Pb}/^{238}\text{U}$  vs  $^{207}\text{Pb}/^{235}\text{U}$  concordia plots for concordant ages with a  $2\sigma$  error. Plots and  
629 concordia ages obtained by Isoplot software (Ludwig, 2012). Representative SEM-CL  
630 images are reported.

631 Fig. 4. Cumulative probability distribution (Ludwig, 2012) of  $^{206}\text{Pb}/^{238}\text{U}$  ages for the  
632 Morozumi intrusive units and host rocks. Representative SEM-CL images are  
633 reported.

634 Fig. 5. Total Alkali-Silica (TAS) diagram of the Morozumi intrusive units. Also reported for  
635 comparison: (i) intrusive rocks emplaced at 490-500 Ma in the Wilson arc: (i) Abbott  
636 gabbro, Confusion tonalites, Abbott granites, Vegetation leucogranites (Di Vincenzo  
637 and Rocchi, 1999; Rocchi et al., 2004), Vegetation lamprophyres and Irizar granites-  
638 dykes (Rocchi et al., 2009), and (ii) intrusive rocks emplaced earlier in the Tiger  
639 oceanic arc, i.e. Tiger gabbro (Bracciali et al., 2009).

640 Fig. 6. N-MORB-normalized plots (Sun and McDonough, 1989) of incompatible elements. (a)  
641 Morozumi diorite samples, compared to the Abbott gabbro (Di Vincenzo and Rocchi,  
642 1999), the Vegetation lamprophyres from central Victoria Land (Rocchi et al., 2009)  
643 and the Tiger gabbro from the Tiger magmatic arc (Fig. 1) (Bracciali et al., 2009). (b)  
644 Morozumi tonalite dyke, compared to the Confusion tonalite (Di Vincenzo and Rocchi,  
645 1999; Rocchi et al., 2004) and the Tiger gabbro from the Tiger magmatic arc (Fig. 1)  
646 (Bracciali et al., 2009).

647 Fig. 7.  $\epsilon_{\text{Nd}}(500 \text{ Ma})$  vs  $^{87}\text{Sr}/^{86}\text{Sr}_{(500 \text{ Ma})}$  plot of the Morozumi mafic-intermediate intrusive rocks.  
648 Also reported for comparison: (i) intrusive complexes emplaced at around 490-500  
649 Ma in the Wilson arc: (i) the Abbott gabbro and its hybridization trend with melts  
650 from the deep crust, the Confusion tonalites, Vegetation leucogranites and their  
651 hybridization trend with melts from mesedimentary upper crust (Di Vincenzo and  
652 Rocchi, 1999; Rocchi et al., 2004), (ii) the Irizar granites and dykes (Rocchi et al.,  
653 2009), and (iii) intrusive rocks emplaced earlier in the Tiger oceanic arc, i.e. Tiger  
654 gabbro (Bracciali et al., 2009).

655 Fig. 8. (a) Ternary diagrams showing Na-K-Ca relationships for the rocks of the Morozumi  
656 intrusive complex compared to pelite-, greywacke- and metavolcanoclastics-derived  
657 fluid-absent experimental melts (Montel and Vielzeuf, 1997; Patiño Douce and Beard,  
658 1996; Patiño Douce and Johnston, 1991; Sisson et al., 2005; Skjerlie and Johnston,  
659 1996; Stevens et al., 1997; Vielzeuf and Holloway, 1988) and to adakite and TTG fields  
660 and evolutionary trends (Defant and Drummond, 1993). (b) Classical Sr/Y vs. Sr plot  
661 comparing the Morozumi mafic-intermediate intrusive rocks with the adakite and  
662 andesite-dacite-rhyolite (ADR) fields (Defant and Drummond, 1993)

663 Fig. 9.  $^{87}\text{Sr}/^{86}\text{Sr}$  vs.  $1000/\text{Sr}$  plot, showing mixing trends as linear paths.

664 Fig. 10. Idealized sketch of the reconstructed scenario from the emplacement level down to  
665 source depth.

666

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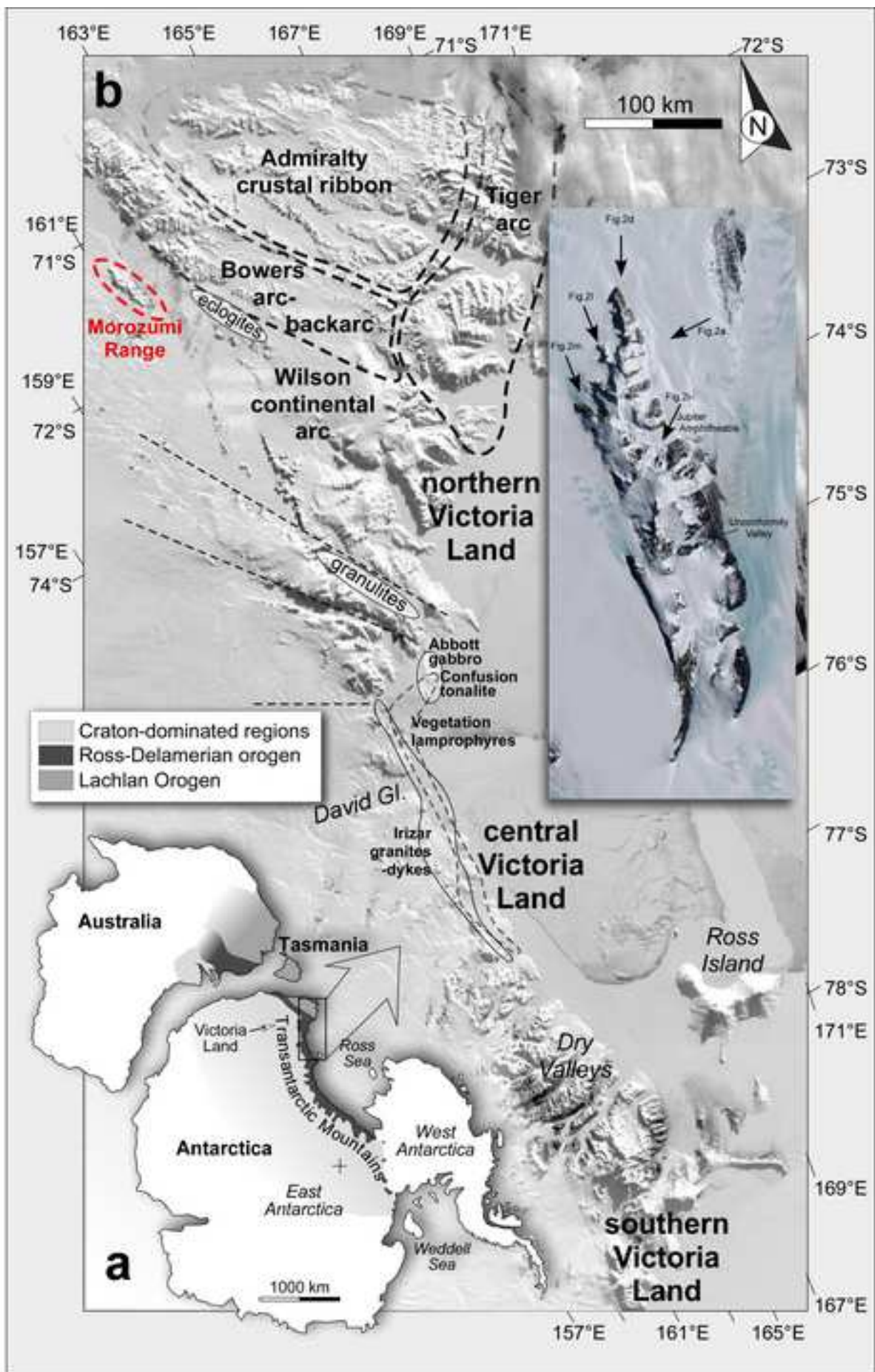


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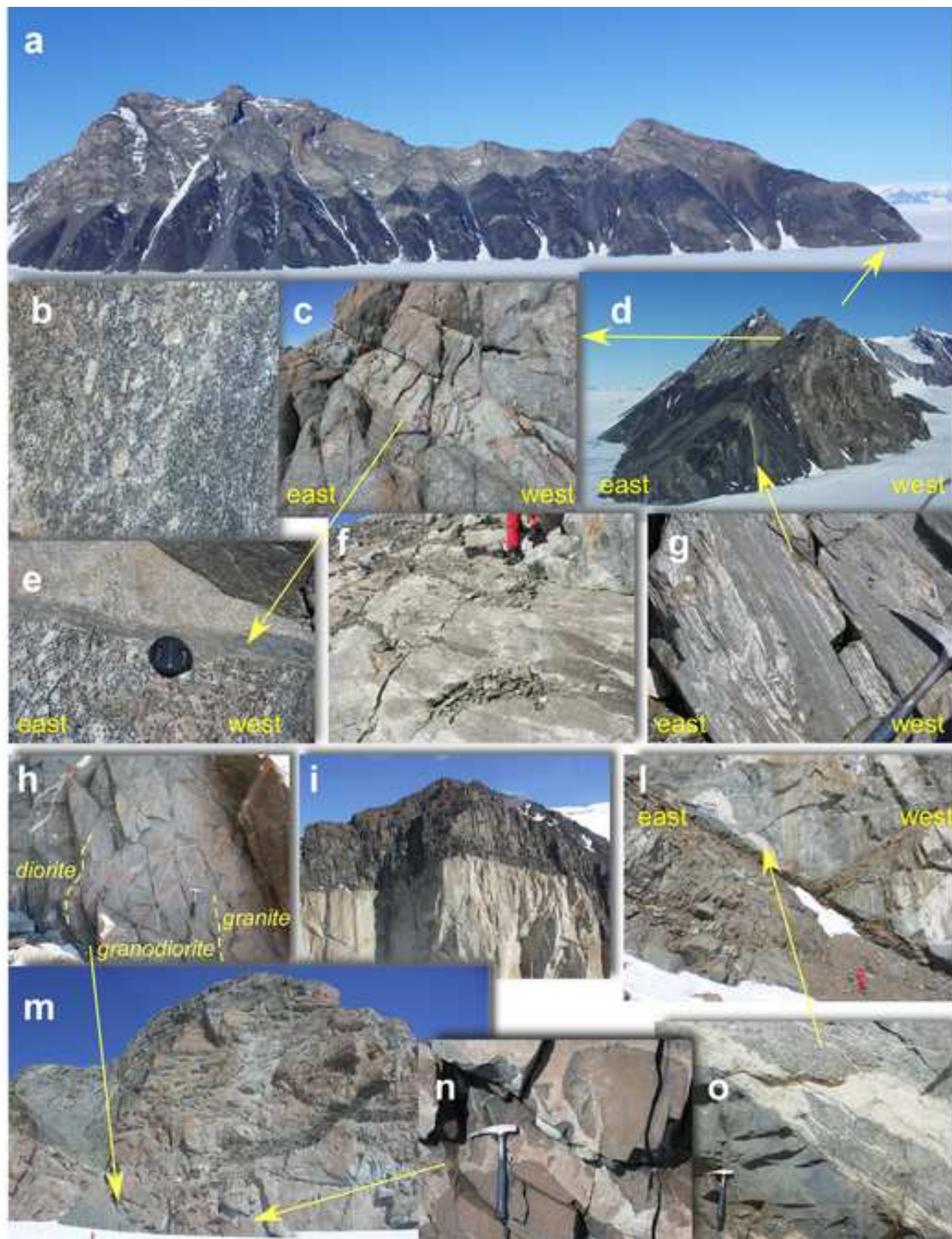


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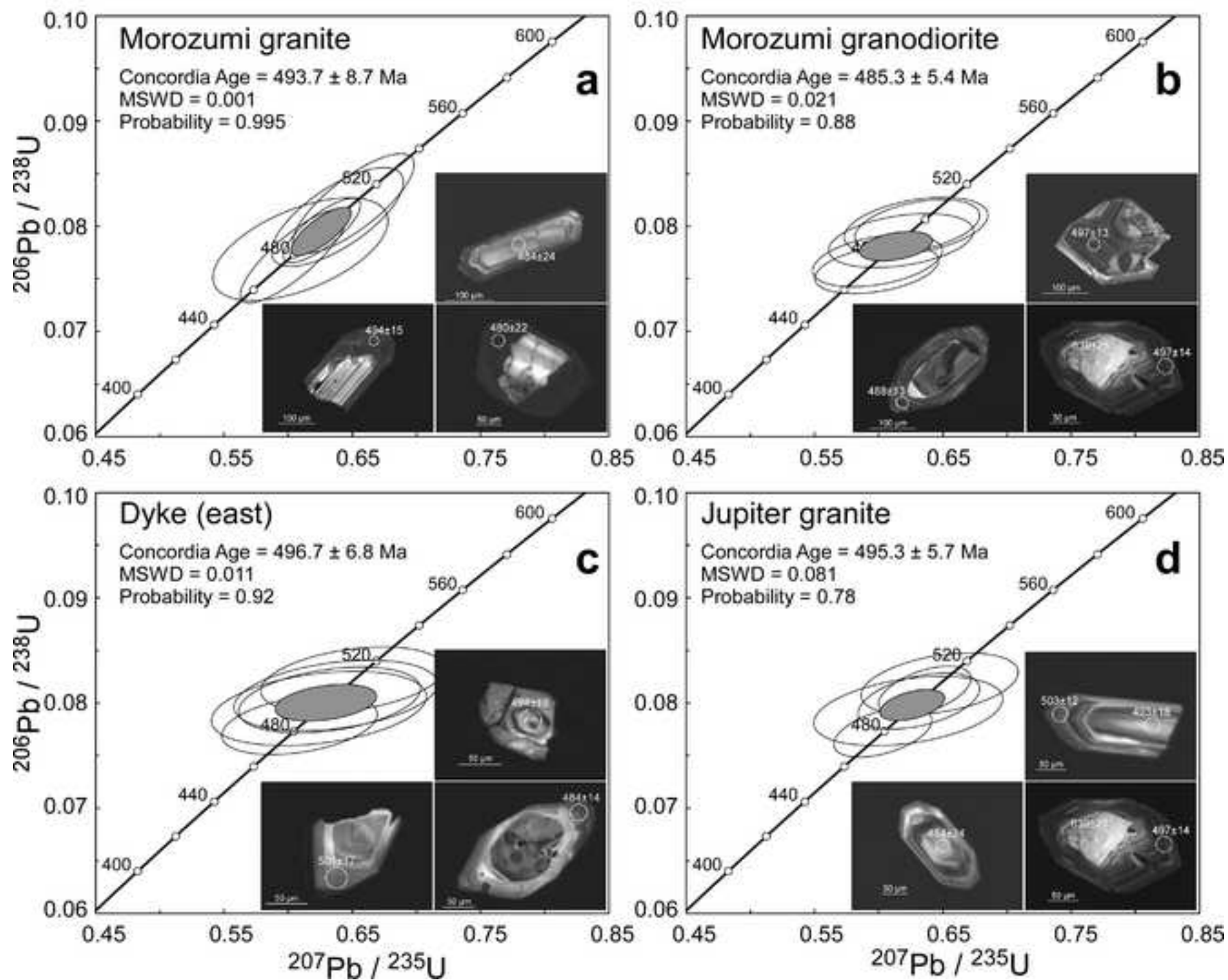


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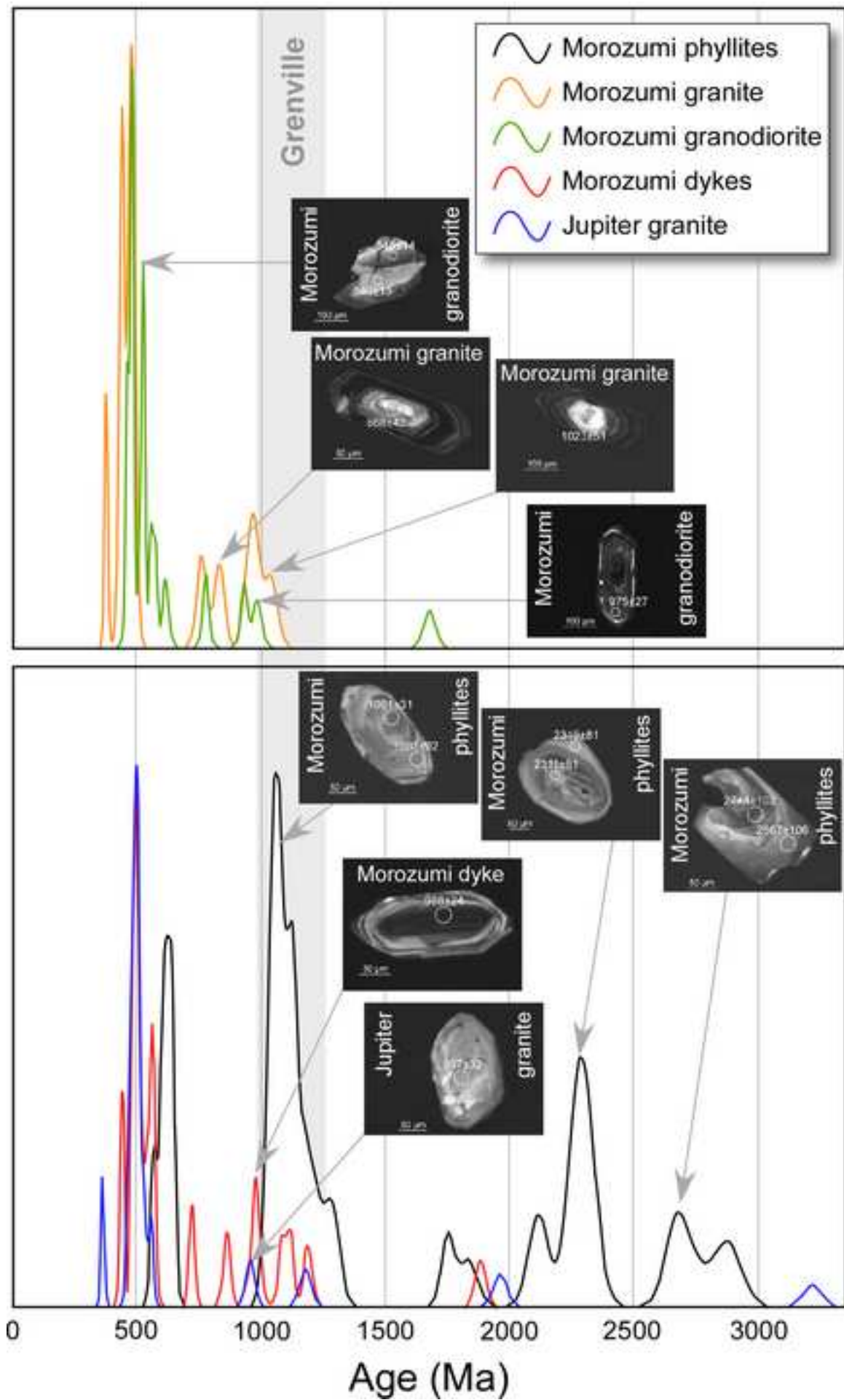


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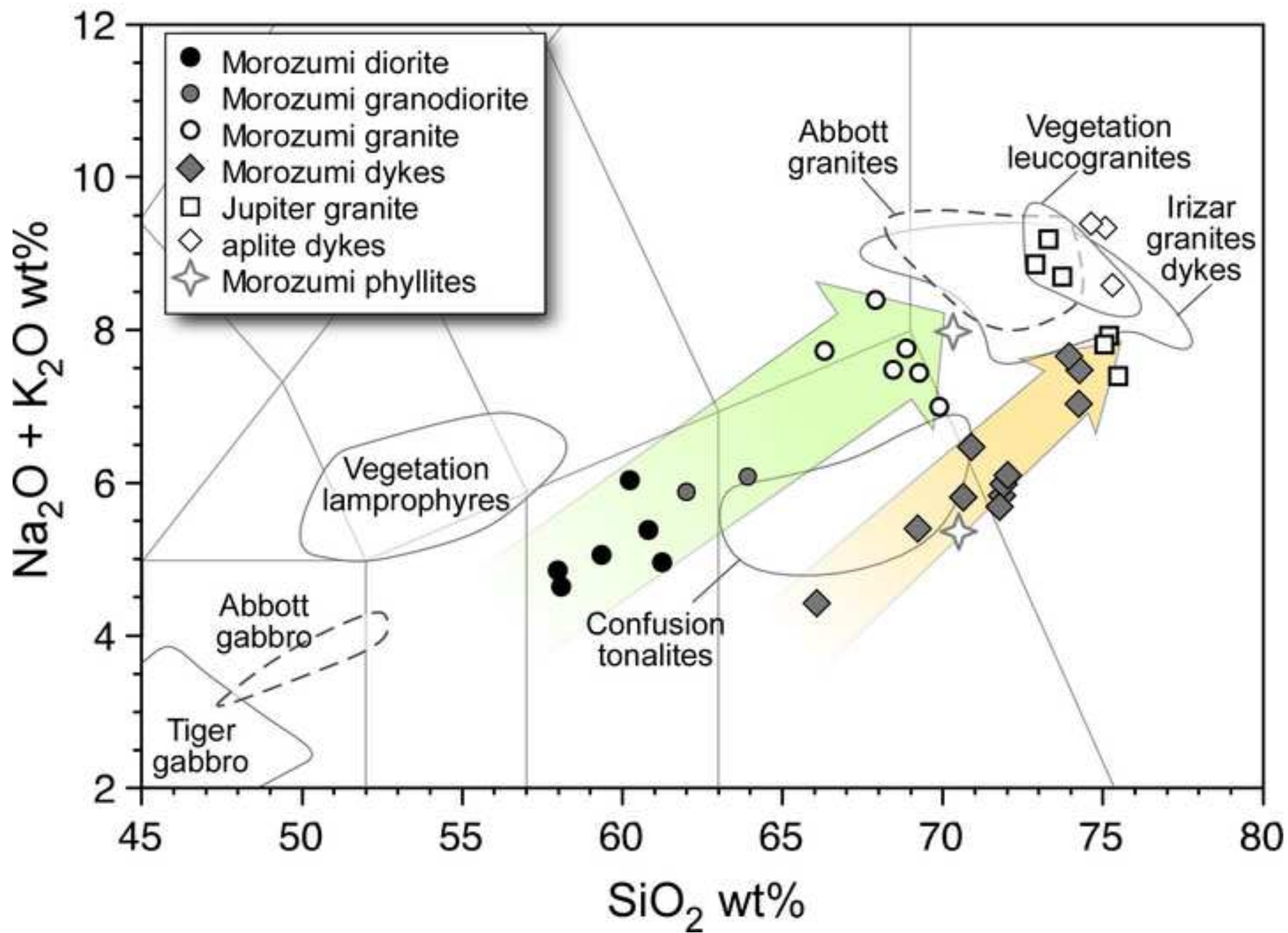


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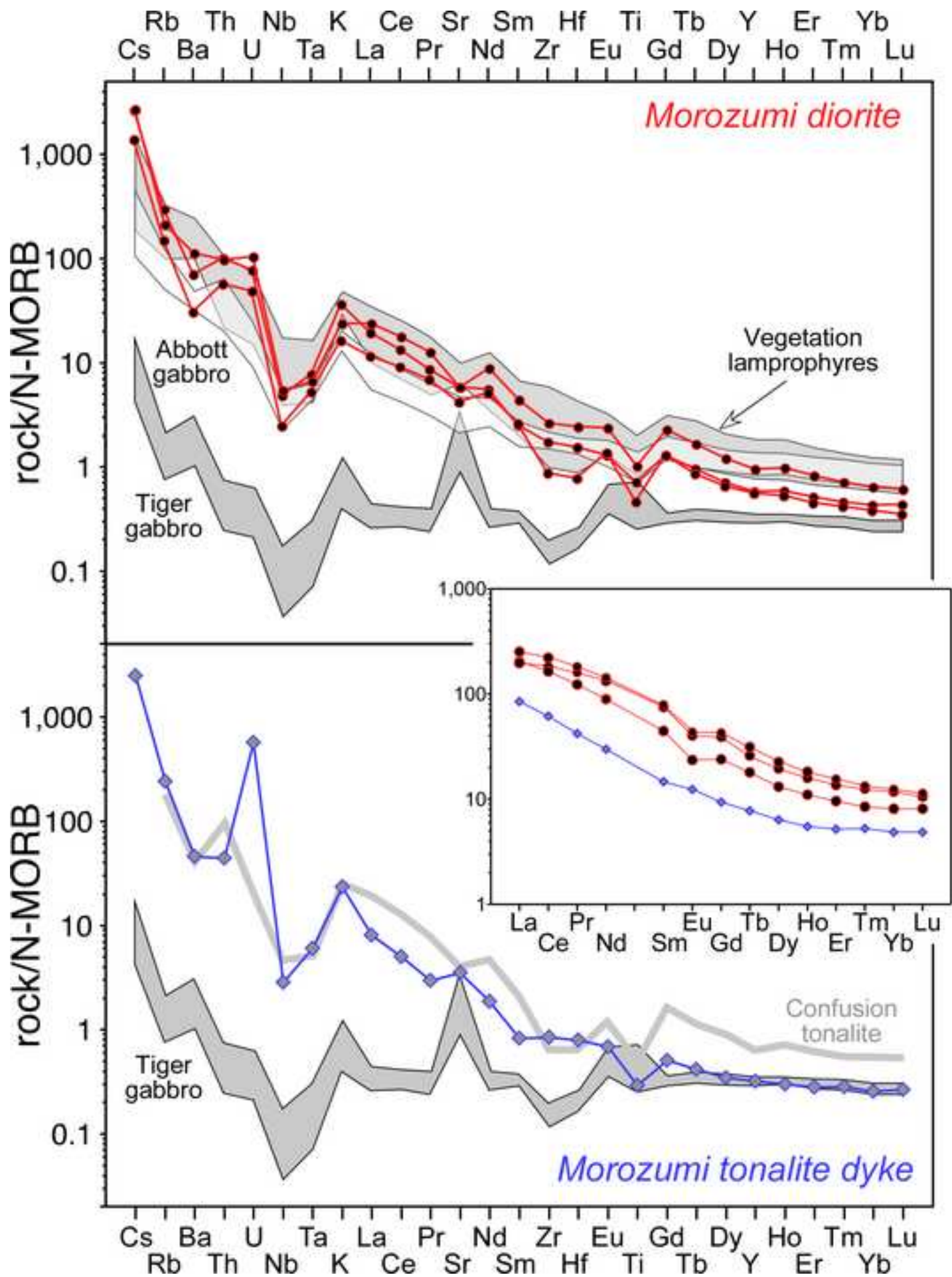


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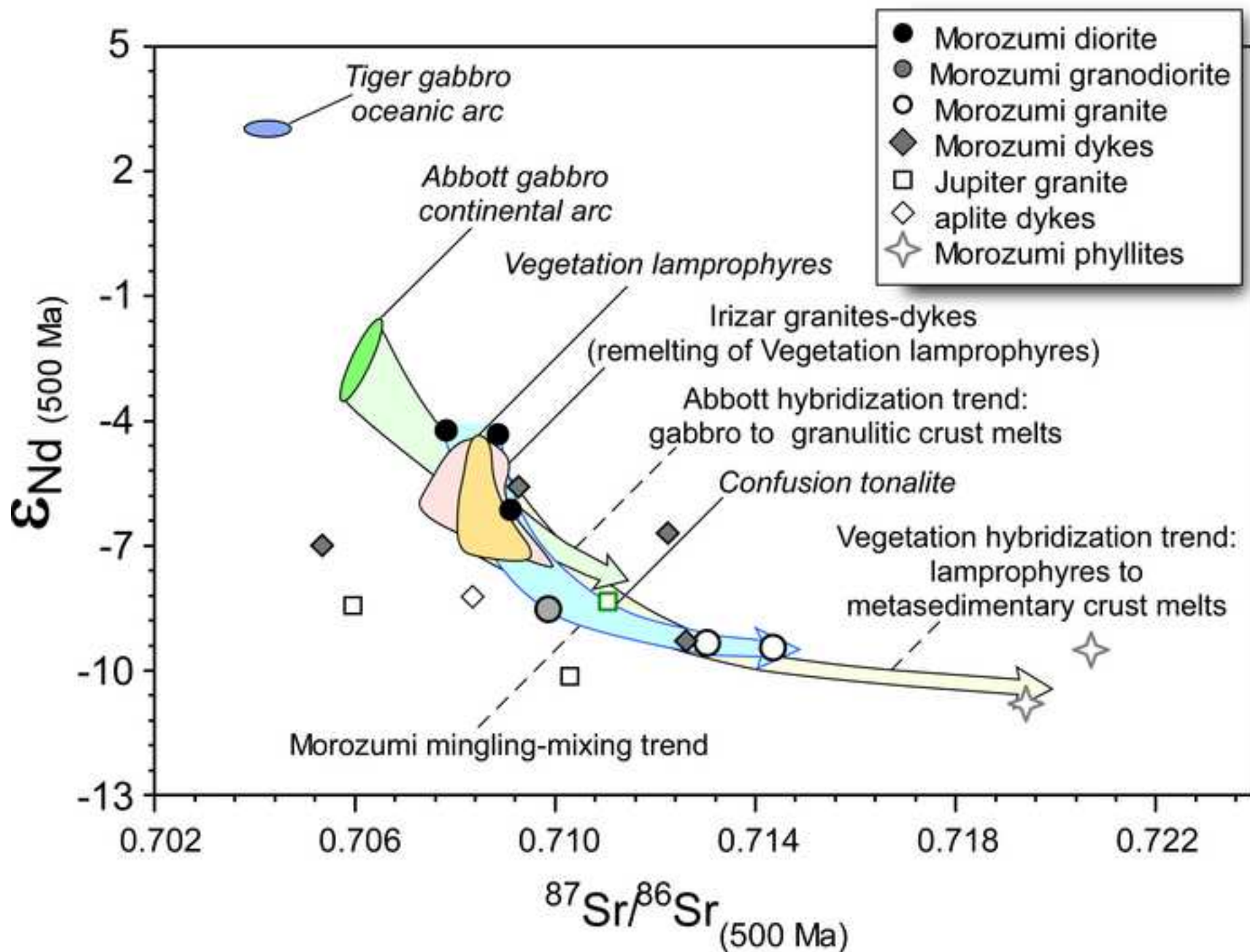


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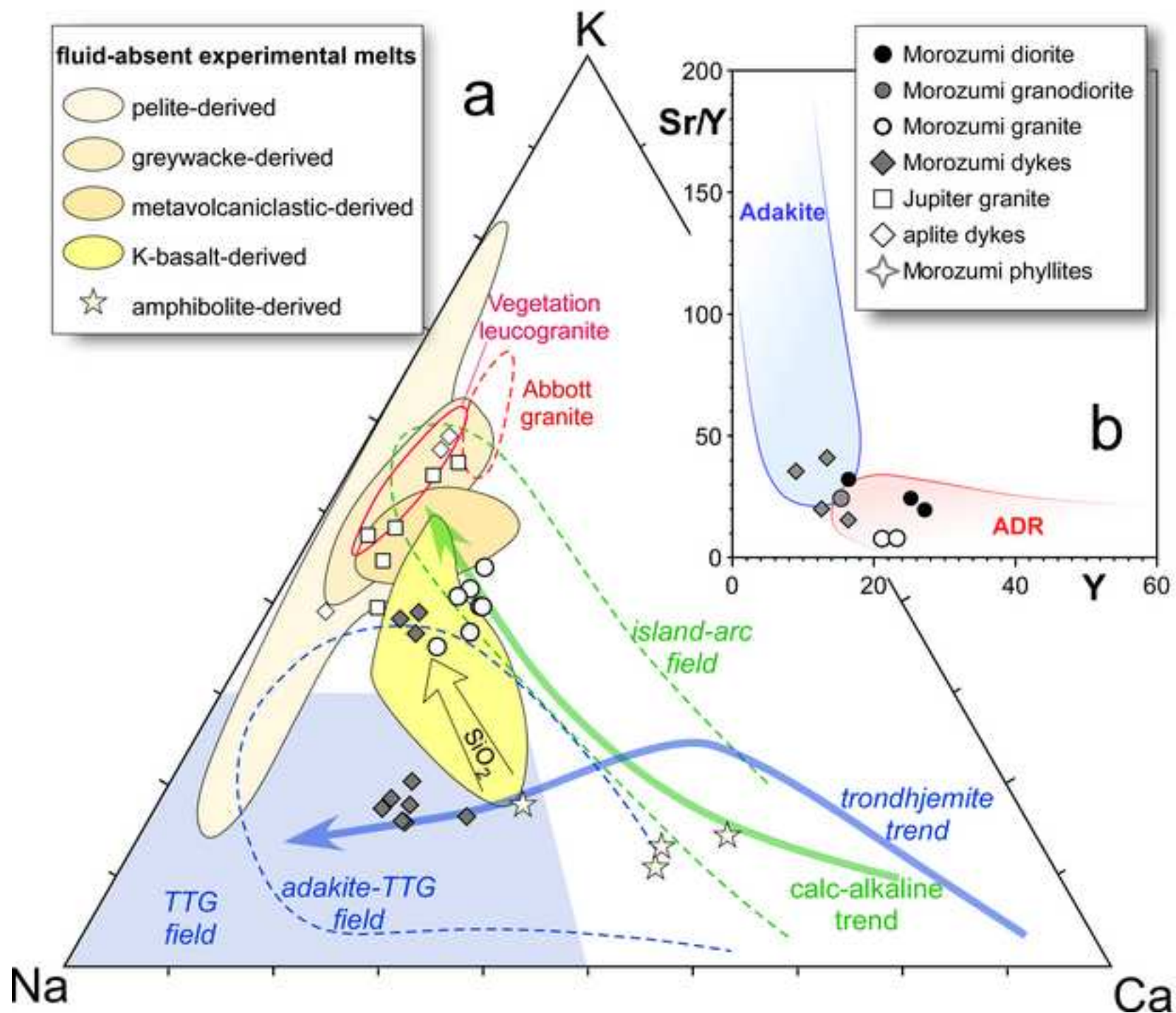




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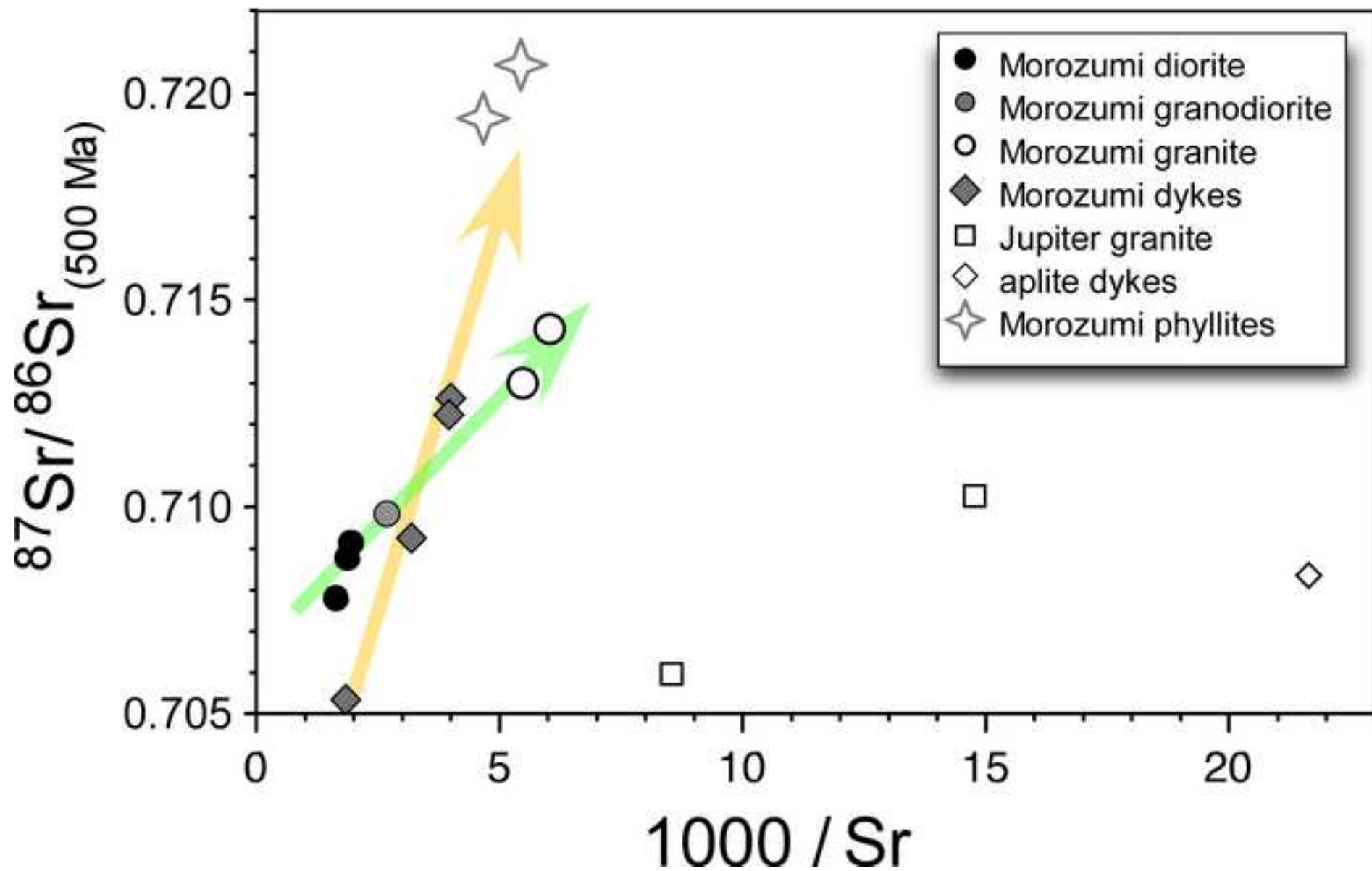


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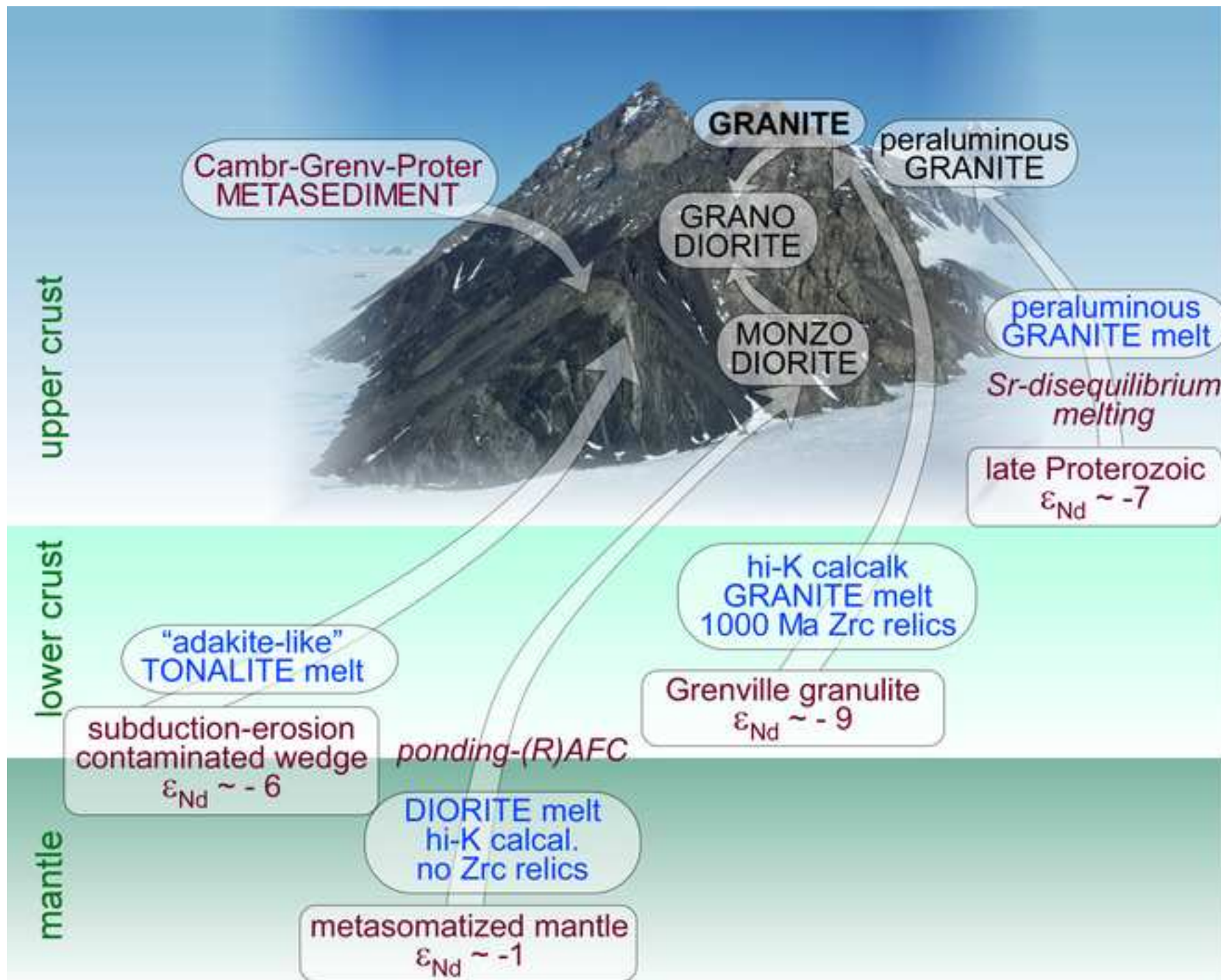


Table 1

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**Table 1.** Summary of field, petrographic and geochronological features of the lithologic units of the Morozumi Range intrusive complex.

unit	sub-unit	rock type	petrography	structure	field chronology relative to Morozumi granite	age (Ma)
<b>Morozumi phyllites</b>				vertical foliation	earlier	
<b>Morozumi granite</b>		monzogranite	Kfs-phyric, medium-grained, bt-bearing	Kfs-phyric, vertical foliation		494 ± 9
<b>Morozumi diorite</b>		quartzdiorite	fine- to medium-grained, bt+amph-bearing	isotropic	coeval (hot, ductile contacts)	
<b>Morozumi granodiorite</b>		granodiorite	fine-grained	isotropic	coeval (diffuse contact)	485 ± 5
<b>Morozumi dykes</b>	eastern	tonalite-granodiorite	medium-grained, bt+ms-bearing	foliated	early-coeval (flattened)	497 ± 7
	western	tonalite-granodiorite	fine or medium-grained, bt±ms-bearing		late (crosscut granite foliation, ductile relationships)	
	crestal	leucomonzogranites - leucogranodiorites	bt±ms-bearing		post (crosscut granite foliation, brittle relationships)	
<b>Jupiter granite</b>		monzogranite	medium- to coarse-grained, leucocratic, bt+ms-bearing	isotropic	unknown	495 ± 6
<b>Morozumi leucogranites/aplites</b>		leucogranite	fine- to medium-grained, ms ± grt/tur-bearing		late	

**Table 2.** Major elements, trace elements and Sr-Nd isotopic data.

unit sample	Morozumi diorite						Mor. granodiorite		Morozumi granite				Morozumi phyllites			
	23.12.05 DS10	23.12.05 DS7	23.12.05 DS12	23.12.05 DS9	23.12.05 DS13	23.12.05 DS11	22.12.05 DS7	22.12.05 DS6	19.12.05 DS5	20.12.05 DS1	21.12.05 BS4	23.12.05 DS6	18.12.05 DS11	23.12.05 DS5	19.12.05 DS2	23.12.05 DS14
lat	71° 32.9'	71° 32.9'	71° 32.9'	71° 32.9'	71° 32.9'	71° 32.9'	71° 32.1'	71° 32.1'	71° 27.08'	71° 26.9'	71° 28.3'	71° 32.9'	71° 29.9'	71° 32.2'	71° 46.8'	71° 33.2'
long	161° 37.7'	161° 37.7'	161° 37.7'	161° 37.7'	161° 37.7'	161° 37.7'	161° 38.7'	161° 38.7'	161° 42.0'	161° 49.9'	161° 44.3'	161° 37.7'	161° 44.9'	161° 43.0'	162° 10.9'	161° 47.8'
rock type	QD	QD	QD	MD	QMD	T	GD	GD	MG	MG	MG	MG	MG	MG		
<i>Major elements (wt%)</i>																
SiO <sub>2</sub>	56.63	56.86	57.55	59.63	59.90	60.26	60.57	62.79	64.96	66.57	67.46	67.85	68.53	69.40	69.27	69.74
TiO <sub>2</sub>	0.95	1.30	0.88	0.91	0.80	0.75	0.91	0.81	0.74	0.56	0.59	0.58	0.57	0.56	0.63	0.61
Al <sub>2</sub> O <sub>3</sub>	14.81	14.95	15.11	17.97	15.65	15.31	16.93	16.60	15.92	15.58	15.24	14.93	14.94	14.83	12.93	12.98
Fe <sub>2</sub> O <sub>3</sub>	6.82	8.52	6.52	6.10	5.88	5.84	5.88	5.19	4.62	3.19	3.80	3.71	3.52	3.58	4.01	5.36
MnO	0.11	0.13	0.09	0.09	0.09	0.10	0.07	0.07	0.07	0.06	0.07	0.06	0.06	0.06	0.05	0.07
MgO	6.32	4.82	5.46	3.04	4.80	5.17	2.76	2.30	1.51	1.13	1.22	1.23	1.16	1.17	2.18	2.68
CaO	6.81	6.22	6.22	4.92	5.84	5.84	4.54	4.27	2.27	2.59	2.64	2.40	2.64	2.59	1.36	2.02
Na <sub>2</sub> O	2.85	2.84	2.81	3.33	3.06	2.90	2.97	3.01	3.09	3.01	2.99	3.01	2.98	3.04	1.69	2.83
K <sub>2</sub> O	1.91	1.73	2.12	2.65	2.26	2.00	2.80	2.99	4.51	5.25	4.41	4.66	4.41	3.94	6.17	2.47
P <sub>2</sub> O <sub>5</sub>	0.53	0.60	0.31	0.36	0.29	0.31	0.36	0.29	0.34	0.18	0.21	0.19	0.20	0.20	0.20	0.16
LOI	1.96	1.97	1.87	1.78	1.66	1.58	1.40	1.99	1.99	0.87	1.17	0.88	0.69	1.18	1.07	1.18
<i>Trace elements (ppm)</i>																
Be	5.6	3.7		5.2			3.1					9.1	9.3		3.9	4.2
Sc	18.0	11.9		12.2			12.3					9.4	9.4		10.4	11.8
V	133.2	156.2		108.2			112.4					53.7	54.4		64.3	83.6
Cr	161.2	60.6		29.0			42.5					18.7	25.8		70.9	67.7
Co	26.7	24.8		15.6			15.2					7.2	7.4		10.0	12.3
Ni	124.7	38.9		19.1			26.1					8.6	7.9		26.5	31.1
Cu	36.9	20.3		51.5			34.8					10.9	19.7		5.0	32.0
Ga	18.4	17.7		21.3			21.8					19.8	19.3		15.9	17.3
Rb	134.6	119.6		165.5			165.0					260.4	244.4		185.5	142.4
Sr	610.6	526.0		528.1			373.8					183.0	165.7		214.6	183.8
Y	25.19	27.10		16.45			15.33					23.21	21.11		29.00	31.32
Zr	173	226		240			266					214	192			
Nb	9.33	12.75		11.42			12.91					15.92	16.31		12.72	13.10
Cs	15.64	19.06		18.60			7.53					24.69	44.92		18.35	11.76
Ba	312.3	714.9		444.0			501.5					526.5	543.1		845.2	318.9
La	46.82	59.75		48.36			61.94					58.40	46.14		50.24	31.51
Ce	110.65	135.24		100.74			129.26					123.03	97.21		102.58	65.61
Pr	14.79	16.82		11.48			14.83					14.20	11.26		11.94	7.80
Nd	60.46	65.42		40.87			53.16					50.28	40.49		44.10	29.21
Sm	11.13	11.70		6.61			8.45					9.25	7.54		7.99	6.12
Eu	2.27	2.43		1.33			1.32					1.27	1.27		1.36	1.13
Gd	7.76	8.51		4.75			5.76					6.55	5.51		6.36	5.37
Tb	0.94	1.13		0.65			0.70					0.88	0.78		0.94	0.83
Dy	4.81	5.57		3.23			3.20					4.53	3.92		5.20	5.04
Ho	0.87	1.00		0.60			0.56					0.83	0.75		1.03	1.07
Er	2.18	2.48		1.54			1.37					2.06	1.83		2.70	2.90
Tm	0.31	0.33		0.21			0.19					0.29	0.27		0.39	0.45
Yb	1.89	1.99		1.31			1.15					1.82	1.60		2.43	2.68
Lu	0.26	0.28		0.20			0.16					0.26	0.22		0.34	0.38
Hf	2.58	5.01		3.17			3.16					4.96	4.36		4.08	3.55
Ta	1.12	0.88		1.02			0.79					1.74	2.25		1.11	1.12
Tl	1.03	0.90		1.65			1.27					1.80	1.64		1.18	1.03
Pb	8.22	13.00		11.73			13.40					37.47	36.49		38.08	28.88
Th	10.98	11.73		12.06			17.11					27.04	20.76		17.90	11.90
U	3.69	4.93		3.63			1.24					4.41	2.56		2.32	2.76
<i>Sr-Nd isotope data</i>																
<sup>87</sup> Sr/ <sup>86</sup> Sr	0.712345	0.713517		0.715543			0.718944					0.742427	0.744825		0.737277	0.736718
error (2s)	0.000070	0.000011		0.000011			0.000012					0.000012	0.000012		0.000010	0.000080
<sup>87</sup> Rb/ <sup>86</sup> Sr	0.638	0.658		0.907			1.279					4.131	4.283		2.508	2.248
<sup>87</sup> Sr/ <sup>86</sup> Sr <sub>(500Ma)</sub>	0.707798	0.708826		0.709079			0.709833					0.712995	0.714309		0.719403	0.720700
e <sub>Sr(500Ma)</sub>	55	70		73			84					129	148		220	238
<sup>143</sup> Nd/ <sup>144</sup> Nd	0.512143	0.512130		0.512002			0.511873					0.511881	0.511880		0.511798	0.511922
error (2s)	0.000007	0.000007		0.000007			0.000006					0.000007	0.000005		0.000007	0.000007
<sup>147</sup> Sm/ <sup>144</sup> Nd	0.1112	0.1081		0.0977			0.0961					0.1112	0.1126		0.1095	0.1266
<sup>143</sup> Nd/ <sup>144</sup> Nd <sub>(500)</sub>	0.511779	0.511776		0.511682			0.511558					0.511517	0.511511		0.511439	0.511507
e <sub>Nd(500Ma)</sub>	-4.2	-4.3		-6.1			-8.5					-9.3	-9.4		-10.8	-9.5
T <sub>DM</sub> (Ga)	1.55	1.55		1.68			1.85					1.91	1.92		2.01	1.92

Rock name abbreviations: D: diorite, MD: monzodiorite, QD: quartz diorite, QMD: quartz monzodiorite, T: tonalite, GD: granodiorite, MG: monzogranite, SG: syenogranite, AFG: alkali feldspar granite. Igneous rock names according to Q'-ANOR diagram after Streckeisen et al. (1978). T<sub>DM</sub>: Nd model age calculated according to DePaolo et al. (1991)

**Table 2** (ctd). Major elements, trace elements and Sr-Nd isotopic data.

unit	Mor. dykes-eastern						Morozumi dykes-western						Morozumi dykes-crestal						Jupiter granite						aplites-leucogranites					
sample	18.12.05	21.12.05	20.12.05	20.12.05	23.12.05	23.12.05	22.12.05	21.12.05	21.12.05	22.12.05	21.12.05	20.12.05	23.12.05	23.12.05	22.12.05	20.12.05	23.12.05	23.12.05	18.12.05	18.12.05										
	DS7	BS1	DS4	DS2	DS2	DS3	DS1m	BS7	BS5	DS1l	BS6	DS3	DS4	DS1	DS3	DS6	DS15	DS8	DS9	DS10										
lat	71° 26.5'	71° 26.6'	71° 27.3'	71° 26.9'	71° 29.9'	71° 29.9'	71° 28.3'	71° 28.3'	71° 28.3'	71° 28.3'	71° 28.3'	71° 27.3'	71° 31.7'	71° 29.9'	71° 19.3'	71° 34.9'	71° 30.6'	71° 32.9'	71° 26.5'	71° 26.5'										
long	161° 43.0'	161° 50.6'	161° 41.9'	161° 49.9'	161° 41.9'	161° 41.9'	161° 44.3'	161° 44.3'	161° 44.3'	161° 44.3'	161° 44.3'	161° 41.9'	161° 41.1'	161° 41.9'	161° 19.7'	161° 50.6'	161° 44.7'	161° 37.7'	161° 43.0'	161° 43.0'										
rock type	GD	T	T	T	GD	T	GD	GD	MG	MG	MG	MG	SG	SG	SG	MG	SG	AFG	SG	SG										
<i>Major elements (wt%)</i>																														
SiO <sub>2</sub>	70.38	70.44	64.75	68.28	69.64	71.41	70.73	70.87	72.76	73.31	73.92	72.27	72.58	73.16	73.67	74.19	74.28	74.54	74.67	75.20										
TiO <sub>2</sub>	0.30	0.29	0.49	0.37	0.40	0.37	0.29	0.29	0.20	0.18	0.19	0.28	0.17	0.16	0.35	0.13	0.13	0.10	0.03	0.03										
Al <sub>2</sub> O <sub>3</sub>	15.29	14.92	16.53	15.92	15.31	15.07	14.77	14.95	13.82	14.10	13.99	14.08	14.37	14.23	12.26	13.91	13.67	13.86	13.82	14.49										
Fe <sub>2</sub> O <sub>3</sub>	2.61	2.53	4.48	3.41	3.25	2.99	2.72	2.86	1.79	1.50	1.69	1.68	1.25	1.29	2.41	1.23	1.33	0.97	0.48	0.31										
MnO	0.05	0.05	0.08	0.06	0.05	0.04	0.04	0.05	0.03	0.03	0.03	0.04	0.03	0.02	0.04	0.04	0.05	0.03	0.01	0.06										
MgO	1.26	0.95	2.47	1.44	1.08	0.93	0.84	0.76	0.60	0.50	0.56	0.61	0.36	0.55	0.76	0.30	0.36	0.24	0.14	0.11										
CaO	2.86	3.02	4.76	3.70	2.92	2.93	2.70	2.79	1.80	1.83	1.60	1.31	0.97	1.12	0.86	1.19	0.98	0.73	1.04	0.69										
Na <sub>2</sub> O	4.48	4.07	1.90	3.63	3.92	4.02	4.07	3.82	3.20	3.35	3.40	2.59	3.25	2.92	2.24	3.34	3.24	3.74	2.60	4.12										
K <sub>2</sub> O	1.93	1.64	2.43	1.69	1.81	1.63	1.91	2.07	3.69	4.23	4.03	6.22	5.34	6.28	5.42	3.96	4.62	5.66	6.72	4.40										
P <sub>2</sub> O <sub>5</sub>	0.11	0.10	0.10	0.11	0.17	0.10	0.12	0.11	0.09	0.10	0.10	0.14	0.20	0.14	0.10	0.08	0.18	0.08	0.08	0.10										
LOI	1.21	1.66	2.25	1.19	0.93	0.78	1.17	0.96	1.23	1.50	1.09	1.24	0.86	1.27	0.95	1.61	1.04	0.89	0.79	0.94										
<i>Trace elements (ppm)</i>																														
Be	3.1			3.2			4.3			5.5		7.3		10.6		6.3														
Sc	6.5			7.1			5.7			4.2		3.4		5.3		2.5														
V	36.7			41.2			22.1			10.5		7.2		9.4		0.6														
Cr	32.6			36.1			10.0			7.3		6.8		5.7		4.7														
Co	6.4			8.1			3.9			2.2		1.4		1.2		0.3														
Ni	14.2			11.4			4.6			2.5		2.0		1.6		0.6														
Cu	2.4			9.3			2.6			2.3		9.9		1.1		0.6														
Ga	17.6			14.9			18.3			14.9		19.7		15.8		13.0														
Rb	84.0			135.3			132.8			171.6		365.2		195.3		161.1														
Sr	547.2			313.8			251.7			251.0		67.6		117.1		46.2														
Y	13.41			8.94			12.59			16.34		13.12		25.13		12.40														
Zr	133			122			143			99		87		72		28														
Nb	5.79			6.63			15.66			10.07		17.16		11.23		7.12														
Cs	8.95			17.25			8.68			5.88		20.48		17.03		4.85														
Ba	406.3			288.5			212.6			660.1		213.7		405.8		248.9														
La	21.16			20.11			28.22			29.86		22.76		16.48		5.68														
Ce	42.83			37.60			54.40			57.03		49.34		34.65		12.41														
Pr	4.87			3.90			5.93			6.17		5.74		4.07		1.46														
Nd	18.11			13.61			20.54			21.10		20.61		15.00		5.43														
Sm	3.73			2.17			3.43			3.80		4.70		3.29		1.78														
Eu	0.66			0.70			0.76			0.80		0.54		0.60		0.16														
Gd	3.18			1.87			2.74			3.02		3.70		3.08		1.64														
Tb	0.49			0.28			0.40			0.46		0.54		0.58		0.32														
Dy	2.47			1.57			2.16			2.55		2.63		3.60		1.89														
Ho	0.46			0.30			0.41			0.53		0.42		0.80		0.39														
Er	1.20			0.83			1.13			1.50		0.95		2.30		1.14														
Tm	0.20			0.13			0.17			0.22		0.14		0.38		0.19														
Yb	1.23			0.78			1.14			1.37		0.85		2.49		1.26														
Lu	0.16			0.12			0.17			0.21		0.11		0.36		0.17														
Hf	0.77			1.63			3.14			2.19		2.62		1.91		0.99														
Ta	0.85			0.80			1.68			1.18		2.03		2.44		1.61														
Tl	0.64			1.17			0.99			1.09		2.27		1.29		0.94														
Pb	11.42			10.59			20.69			34.95		42.01		34.31		25.32														
Th	6.98			5.26			8.86			10.98		12.84		8.14		1.91														
U	0.62			26.50			2.24			2.63		3.72		2.96		1.37														
<i>Sr-Nd isotope data</i>																														
<sup>87</sup> Sr/ <sup>86</sup> Sr	0.708517			0.718166			0.723144			0.726733		0.822843		0.740428		0.780698														
error (2σ)	0.000009			0.000008			0.000011			0.000009		0.000011		0.000009		0.000010														
<sup>87</sup> Rb/ <sup>86</sup> Sr	0.444			1.249			1.529			1.982		15.801		4.842		10.157														
<sup>87</sup> Sr/ <sup>86</sup> Sr <sub>(500Ma)</sub>	0.705351			0.709270			0.712250			0.712614		0.710254		0.705927		0.708328														
ε <sub>Sr(500Ma)</sub>	20			76			118			124		90		29		63														
<sup>143</sup> Nd/ <sup>144</sup> Nd	0.512045			0.512023			0.511982			0.511873		0.511926		0.512000		0.512222														
error (2σ)	0.000007			0.000008			0.000007			0.000008		0.000009		0.000006		0.000009														
<sup>147</sup> Sm/ <sup>144</sup> Nd	0.1246			0.0965			0.1008			0.1089		0.1378		0.1325		0.1986														
<sup>143</sup> Nd/ <sup>144</sup> Nd <sub>(500)</sub>	0.511637			0.511707			0.511652			0.511516		0.511475		0.511566		0.511572														
ε <sub>Nd(500Ma)</sub>	-7.0			-5.6			-6.7			-9.3		-10.1		-8.4		-8.2														
T <sub>DM</sub> (Ga)	1.75			1.65			1.73			1.91		1.96		1.84		1.84														