

# 1 **Diamonds from the lower mantle?**

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5 Natural diamonds, because of their great physical resiliency, can preserve information about their  
6 formation, storage and transport conditions for billions of years. Diamond samples therefore  
7 provide a unique opportunity to directly study ancient samples of the Earth's deep interior. In order  
8 to correctly interpret the information diamonds provide, it is essential to accurately constrain the  
9 depth of their origin. This depth provenance is usually identified using coexisting minerals, which  
10 are occasionally trapped as inclusions within diamonds during their growth. Comparison of an  
11 inclusion's composition and mineralogy with experimental phase equilibria allows the diamond's  
12 growth conditions to be estimated. While the majority of diamonds likely originate from depths of  
13 140-220 km in cratonic mantle, a small subset appears to have been exhumed from depths  
14 extending to > 800 km, called "superdeep" or "ultradeep" diamonds (Walter et al. 2011; Pearson et  
15 al. 2014). Inclusions of magnesiowüstite are among the most commonly described in sub-  
16 lithospheric diamonds, and have often been assumed to indicate diamond provenance in the lower  
17 mantle because [Mg,Fe]O is not stable at upper mantle conditions in a subsolidus mantle  
18 compositions (Trønnes 2009). This is despite the stability field of [Mg,Fe]O extending to ambient  
19 pressure conditions and experimental evidence of magnesiowüstite stability in equilibrium with  
20 diamond throughout the upper mantle (Brey et al. 2004; Thomson et al. 2016). A new study by  
21 Uenver-Thiele et al. (2017) in *American Mineralogist* places important new constraints on the  
22 formation and uplift history of inclusions containing magnesioferrite.

23 Studies of magnesiowüstite inclusions in diamonds from the Juina region of Brazil often report  
24 observation of nanometre-sized crystals of magnesioferrite ( $[\text{Mg,Fe}^{2+}]\text{Fe}^{3+}_2\text{O}_4$ ), which supposedly  
25 "confirm" the lower mantle origin of these samples. The magnesioferrite precipitates can occur at  
26 the interface between the diamond and [Mg,Fe]O inclusion, or as evenly distributed dislocation  
27 "necklaces" within the inclusion interior (Harte et al. 1999; Wirth et al. 2014; Palot et al. 2016).  
28 Wirth et al. (2014) describe chains of globular  $[\text{Mg}_{0.5}\text{Fe}_{0.5}]\text{Fe}_2\text{O}_4$  crystals, ~ 75 nm in size, making  
29 up 6-11 vol.% of the entire  $[\text{Mg}_{27}\text{Fe}_{71}]\text{O}$  inclusion. This suggests the original inclusion had an  
30  $\text{Fe}^{3+}/\sum\text{Fe}$  of 11-14 %, compared with  $7 \pm 2$  % in the recovered magnesiowüstite (McCammon  
31 1997). Wirth et al. (2014) also identified the magnesioferrite is accompanied by small, ~ 10-30 nm,  
32 cubic voids, Al-bearing spinel and Ni-Fe metal blebs. Palot et al. (2016) describe isolated 10-20 nm  
33 octahedra of  $\text{Mg}[\text{Fe}_{0.75}\text{Cr}_{0.17}\text{Al}_{0.08}]_2\text{O}_4$  throughout a  $[\text{Mg}_{84}\text{Fe}_{16}]\text{O}$  host with a recovered  $\text{Fe}^{3+}/\sum\text{Fe}$   
34 content of 1-2 % that also contains ~ 30 ppm H<sub>2</sub>O in brucite precipitates. The bulk inclusion

35 composition reported by Palot et al. ( $\sim [\text{Mg}_{72}\text{Fe}_{28}]\text{O}$  ignoring minor elements) implies the original  
36 magnesiowüstite must have had an  $\text{Fe}^{3+}/\sum\text{Fe}$  of approximately 10-12 %. Wirth et al. (2014) and  
37 Palot et al. (2016) both observe a topotaxial relationship between magnesioferrite lamellae and the  
38  $[\text{Mg},\text{Fe}]\text{O}$  host, confirming the magnesioferrite must have formed during exsolution from a  
39 homogenous magnesiowüstite grain. Using different arguments both studies concluded that the  
40 magnesioferrite lamellae are indicative of the lower mantle provenance of these diamonds. Wirth et  
41 al. (2014) suggested the highly non-stoichiometric magnesiowüstite inclusion sampled the high-  
42 spin-low-spin transition in the  $\epsilon$ -iron stability field, promoting high  $\text{Fe}^{3+}$  contents. This would  
43 place inclusion, and diamond, formation near the very base of the mantle. Alternatively Palot et al.  
44 (2016) interpreted the conditions of magnesioferrite exsolution using a phase diagram constructed  
45 from atmospheric-pressure experimental data in the  $\text{MgO}-\text{Fe}_2\text{O}_3$ ,  $\text{MgO}-\text{Al}_2\text{O}_3$  and  $\text{MgO}-\text{Cr}_2\text{O}_3$   
46 systems. This approach suggested that the onset of exsolution occurred at a temperature of  $\sim 1700$   
47  $^\circ\text{C}$ , which corresponds to  $\sim 25$  GPa on the mantle adiabat (Palot et al. 2016). Both approaches  
48 makes many assumptions and lack experimental verification that magnesioferrite exsolution  
49 unambiguously indicates a diamond exhumation from the lower mantle. Indeed, as outlined below,  
50 the high ferric iron contents of the inclusions and new phase relations of magnesioferrite (Uenver-  
51 Thiele et al. 2017) instead point to a much shallower origin.

52 At low pressures ( $< 5$  GPa) it is well understood that magnesiowüstite can incorporate significant  
53 ferric iron, up to  $\text{Fe}^{3+}/\sum\text{Fe}$  of 70 %, mainly charge balanced by negative cation vacancies (e.g.  
54 Hazen and Jeanloz 1984; Dobson et al. 1998). With increasing pressure and decreasing oxygen  
55 fugacity the ferric iron capacity of magnesiowüstite decreases, due to a high-pressure phase  
56 transition of  $\text{Fe}_3\text{O}_4$  (Huang and Bassett 1986; McCammon et al. 1998). Since the mantle becomes  
57 more reduced with depth, from  $\sim 1$  log unit above the nickel-nickel oxide buffer (NNO+1) at 200  
58 km to 1.5 log units below the iron-wüstite buffer (IW-1.5) at 660 km (Rohrbach and Schmidt 2011),  
59 it is expected that ferric iron concentration of  $[\text{Mg},\text{Fe}]\text{O}$  will fall rapidly with increasing formation  
60 pressure. Indeed experiments confirm at conditions just within the lower mantle the maximum  
61  $\text{Fe}^{3+}/\sum\text{Fe}$  in  $[\text{Mg}_{70}\text{Fe}_{30}]\text{O}$ , similar in composition to the inclusion observed by Palot et al. (2016), is  
62  $< 2\%$  at NNO and  $< 0.5\%$  at IW (Otsuka et al. 2013). Similarly  $[\text{Mg}_{20}\text{Fe}_{80}]\text{O}$ , similar to that  
63 observed by Wirth et al. (2014), would have a  $\text{Fe}^{3+}/\sum\text{Fe}$  capacity of  $\sim 7 - 14\%$  at IW and NNO  
64 respectively. These ferric iron capacities provide an upper bound, because “normal” lower mantle  
65 conditions are more reduced and extend to higher pressure than the experimental conditions. Thus,  
66 the bulk composition of diamond-hosted inclusions displaying magnesioferrite exsolution appears  
67 inconsistent with formation under lower mantle conditions, unless exceptionally oxidised conditions  
68 are present.

69 In this issue, Uenver-Thiele et al. (2017) experimentally determined the high-pressure phase  
70 relations of magnesioferrite ( $\text{MgFe}_2\text{O}_4$ ) using the multi anvil apparatus. Prior to this study it was  
71 believed that  $\text{MgFe}_2\text{O}_4$  had a relatively simple phase diagram, with the ambient cubic spinel  
72 structure (*Fd-3m*) stable until an isochemical phase transition to orthorhombic  $\text{CaMn}_2\text{O}_4$  structure  
73 (*Pbcm*), HP-  $\text{MgFe}_2\text{O}_4$ , at  $\sim 17$  GPa and temperatures above  $1700^\circ\text{C}$ , or breakdown to  $\text{MgO} +$   
74  $\text{Fe}_2\text{O}_3$  at lower temperatures (Levy et al. 2004). This chemography makes the interpretations of  
75 Wirth et al. (2014) and Palot et al. (2016) feasible. However, the experiments of Uenver-Thiele et  
76 al. (2017) have revealed a very different phase diagram, where the spinel-structured  $\text{MgFe}_2\text{O}_4$   
77 decomposes at  $\sim 10$  GPa. It forms a phase assemblage of  $\text{MgO} + \text{Fe}_2\text{O}_3$  at temperatures below  $1200$   
78  $^\circ\text{C}$  or  $\text{Fe}_2\text{O}_3 +$  an unrecoverable phase of  $\text{Mg}_5\text{Fe}_2\text{O}_8$ - $\text{Mg}_4\text{Fe}_2\text{O}_7$  stoichiometry at higher  
79 temperatures. At pressures beyond  $\sim 13$  GPa the unrecoverable phase(s) are replaced by  
80 orthorhombic,  $\text{CaFe}_3\text{O}_5$  structured (*Cmcm*),  $\text{Mg}_2\text{Fe}_2\text{O}_5$  (Boffa Ballaran et al. 2015). HP- $\text{MgFe}_2\text{O}_4$   
81 was not observed at any conditions up to 18 GPa and  $1300^\circ\text{C}$  in this study. Further high-pressure  
82 experiments are required in order to determine the structure(s) of the unrecoverable phase(s) using  
83 *in-situ* methods, the full extent of the  $\text{Mg}_2\text{Fe}_2\text{O}_5$  stability field and whether HP- $\text{MgFe}_2\text{O}_4$  becomes  
84 stable at higher pressures as suggested by previous studies (Andrault and Bolfan-Casanova 2001;  
85 Levy et al. 2004).

86 The phase relations determined by Uenver-Thiele et al. (2017), coupled with the low ferric iron  
87 capacity of magnesiowüstite in the lower mantle, have very significant consequences for the  
88 interpretation diamond formation pressures. Firstly, magnesioferrite is not stable at lower mantle  
89 conditions where the diamond inclusions (Wirth et al. 2014; Palot et al. 2016) were believed to have  
90 formed. Secondly, if the magnesioferrite did exsolve from  $(\text{Mg,Fe})\text{O}$  as HP- $\text{MgFe}_2\text{O}_4$  in the lower  
91 mantle, it could not have directly inverted to the spinel structure, due to the large stability field of  
92  $\text{Mg}_2\text{Fe}_2\text{O}_5 + \text{Fe}_2\text{O}_3$  as previously suggested. The presence of an additional minor phase between the  
93 magnesioferrite platelets (Wirth et al. 2014) does suggest the magnesioferrite results from inversion  
94 of lamellae of alternative stoichiometry. This idea that magnesioferrite resulted from the conversion  
95 of  $\text{Mg}_2\text{Fe}_2\text{O}_5 + \text{Fe}_2\text{O}_3$  into magnesioferrite at  $\sim 300$  km depth requires further investigation.  
96 However, the phase relations determined by Uenver-Thiele et al. (2017) demonstrate that  
97 magnesioferrite exsolution from magnesiowüstite is not an indicator of formation in the lower  
98 mantle. Instead, it suggests a maximum depth for exsolution of  $\sim 10$  GPa. While the conditions of  
99 original inclusion entrapment of the samples described previously (Wirth et al. 2014; Palot et al.  
100 2016) remain uncertain without further studies, the high ferric iron contents and magnesioferrite  
101 phase relations are consistent with formation in the upper mantle or transition zone, possibly from  
102 oxidized slab materials. The study of Uenver-Thiele et al. (2017) highlights the potentially rich and

103 unexplored chemography and importance of post-spinel phase relations for understanding the  
104 Earth's fundamental geochemical and geodynamic cycles.

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