## ORIGINAL PAPER

Timothy L. Grove · Linda T. Elkins-Tanton Stephen W. Parman · Nilanjan Chatterjee Othmar Müntener · Glenn A. Gaetani

# Fractional crystallization and mantle-melting controls on calc-alkaline differentiation trends

Received: 1 August 2002 / Accepted: 8 January 2003 / Published online: 21 June 2003 © Springer-Verlag 2003

Abstract The phase relations of primitive magnesian andesites and basaltic andesites from the Mt. Shasta region, N California have been determined over a range of pressure and temperature conditions and H<sub>2</sub>O contents. The experimental results are used to explore the influence of H<sub>2</sub>O and pressure on fractional crystallization and mantle melting behavior in subduction zone environments. At 200-MPa H<sub>2</sub>O-saturated conditions the experimentally determined liquid line of descent reproduces the compositional variation found in the Mt. Shasta region lavas. This calc-alkaline differentiation trend begins at the lowest values of FeO\*/MgO and the highest SiO<sub>2</sub> contents found in any arc magma system and exhibits only a modest increase in FeO\*/MgO with increasing SiO<sub>2</sub>. We propose a two-stage process for the origin of these lavas. (1) Extensive hydrous mantle melting produces H<sub>2</sub>O-rich (>4.5-6 wt% H<sub>2</sub>O) melts that are in equilibrium with a refractory harzburgite (olivine + orthopyroxene) residue. Trace elements and H<sub>2</sub>O are contributed from a slab-derived fluid and/or melt. (2) This mantle melt ascends into the overlying crust and undergoes fractional crystallization. Crustallevel differentiation occurs under near-H2O saturated conditions producing the distinctive high SiO2 and low FeO\*/MgO characteristics of these calc-alkaline andesite and dacite lavas. In a subset of Mt. Shasta region lavas, magnesian pargasitic amphibole provides evidence of high pre-eruptive  $H_2O$  contents (>10 wt%  $H_2O$ ) and lower crustal crystallization pressures (800 MPa). Igneous rocks that possess major and trace element characteristics similar to those of the Mt. Shasta region lavas are found at Adak, Aleutians, Setouchi Belt, Japan, the Mexican Volcanic Belt, Cook Island, Andes and in Archean trondhjemite–tonalite–granodiorite suites (TTG suites). We propose that these magmas also form by hydrous mantle melting.

#### Introduction

In subduction zone settings, fractional crystallization of primitive magmas derived from the mantle wedge above the subducted oceanic lithosphere is often called upon to produce andesitic lavas that characterize the common eruptive rock type of arc volcanoes (see Gill 1981; Thorpe 1982, for reviews). Alternatively, a subset of subduction zone magmas have been proposed to be direct melts of the subducted oceanic lithosphere. Notable examples of these lavas include the Adak-type high-Mg andesites from the Aleutians (Kay 1978; Yogodzinski et al. 1995) and the Setouchi high-Mg andesites from SW Japan (Tatsumi and Ishizaka 1982; Ishizaka and Carlson 1983; Shimoda et al. 1998). These lavas are found in subduction zone environments where the age of the subducted oceanic crust is young. The lavas of the Mt. Shasta region provide an unparalleled opportunity to examine the processes of subduction zone magma generation in a young, hot-slab environment. Mt. Shasta lies above the young (12-14 million years old) subducting Juan de Fuca plate (Wilson 1988; Green and Harry 1999) and, therefore, represents an environment of hot oceanic lithosphere subduction. Partial melting of mantle wedge peridotite fluxed by volatiles from the subducted oceanic lithosphere has played an important role in the origin of Mt. Shasta region primitive lavas. In this paper, we present experimental evidence relevant to the conditions of mantle melting and crustal level

T. L. Grove  $(\boxtimes)\cdot L.$  T. Elkins-Tanton  $\cdot$  S. W. Parman N. Chatterjee

Department of Earth, Atmospheric and Planetary Sciences, Massachusetts Institute of Technology, Cambridge, MA 02139, USA

E-mail: tlgrove@mit.edu

O. Müntener

Institut de Géologie, Université de Neuchâtel, Neuchâtel, Switzerland

G. A. Gaetani Department of Geology and Geophysics, Woods Hole Oceanographic Institution, Woods Hole, MA 02543, USA

Editorial responsibility: J. Hoefs

fractional crystallization experienced by Mt. Shasta region lavas. We will demonstrate that the major element characteristics of these lavas are controlled by a hydrous mantle melting process that has extended to shallow mantle depths.

The model developed in this paper (Fig. 1) explores a melting process that is initiated when H<sub>2</sub>O is released from subducted oceanic lithosphere by dehydration of hydrous minerals and/or vapor-saturated anatexis (I, Fig. 1). This H<sub>2</sub>O-rich component rises into the overlying mantle, interacts chemically with mantle peridotite (II, Fig. 1, Zhang and Frantz 2000), and ultimately vapor-saturated mantle melting takes place near the base of the wedge (III, Fig. 1). This vapor-saturated melt segregates from its residue and ascends to shallower mantle depths where it encounters progressively higher mantle temperatures. Melting progresses by a reactive

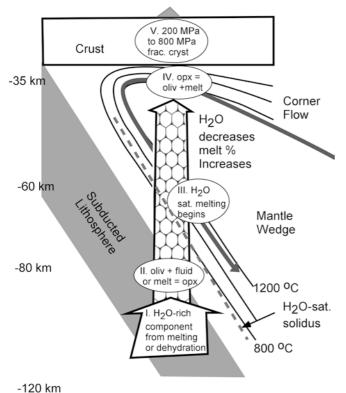


Fig. 1 Model of melt generation and subsequent modification of hydrous magmas produced in the mantle wedge. An H<sub>2</sub>O-rich component is derived by dehydration of minerals in the subducted lithosphere and/or by low-degree vapor saturated melting of subducted sediment or basalt (I). This H<sub>2</sub>O-rich component ascends into the overlying mantle wedge, where it is modified by reaction with the descending mantle and SiO2 is stripped from the fluid-rich component (II). The modified fluid-rich component ascends into shallower, hotter mantle that exceeds the vaporsaturated solidus of peridotite and melting begins (III). Melting takes place throughout the mantle wedge by reactive porous flow (Grove et al. 2002; Gaetani and Grove 2003) and continues into the shallow part of the mantle wedge, where the melt becomes SiO<sub>2</sub>rich by dissolving opx and precipitating oliv (IV). Finally the magma reaches the overlying crust where assimilation of deep and shallow crust, fractional crystallization and magma mixing operate to modify most subduction zone magmas (V)

porous flow process (flux melting). As the melt traverses the inverted thermal gradient in the mantle wedge and encounters hotter, shallower mantle, it finds itself out of equilibrium with its surroundings. In order to equilibrate with the surrounding peridotite, the temperature of the melt increases and the melt dissolves silicate minerals, resulting in a decrease in melt H<sub>2</sub>O content as equilibration pressure decreases. This process continues to shallow mantle depths where a hydrous melt is segregated (IV, Fig. 1). This flux-melting model (Grove et al. 2002; Gaetani and Grove 2003) results in a dilution of the slab-contributed trace element and fluid components through significant (>20 wt%) melting of the mantle. By the time it has reacted its way to the top of the mantle wedge, the melt is far from vapor saturation. This H<sub>2</sub>O-undersaturated mantle melt passes into the crust and ultimately undergoes fractional crystallization (and other crustal interaction processes) in shallow magma reservoirs (V, Fig. 1). Under low-pressure conditions the H<sub>2</sub>O content of the melt is sufficiently high that crystallization takes place at near vapor-saturated conditions (200 MPa). The focus of this paper is the establishment of the depth and H<sub>2</sub>O contents of the mantle melt at step IV, and the modification of that melt during step V.

## **Experimental methods**

#### Starting materials

Synthesis experiments were performed using two of the most primitive lavas from the Mt. Shasta region, northern California, USA, a basaltic andesite (85-44) and a primitive magnesian andesite (85-41c). Compositions of the starting materials are reported in Table 1. The bulk composition of basaltic andesite 85-44 is in equilibrium with the most magnesian olivine present in the lava (Fo<sub>89.6</sub>, Baker et al. 1994). Sample 85-41c contains Fo<sub>93.6</sub> olivine, but has experienced a small amount of contamination by a SiO<sub>2</sub>-rich liquid (Anderson 1974, 1979; Baker et al. 1994). When the bulk composition is corrected for contamination and brought into equilibrium with the Fo<sub>93.6</sub> olivine (oliv), the resulting composition has ~55 wt% SiO<sub>2</sub>. Composition A was calculated as an analog to 85-41c to represent a melt in equilibrium with the Fo<sub>90</sub> oliv, and this composition (A) was also used for experiments (Table 1). Both 85-44 and 85-41c have been used in experimental studies by Baker et al. (1994) and Müntener et al (2001).

#### Experimental procedures

Three starting materials were used: natural powders of 85-44 and 85-41c and an H<sub>2</sub>O-bearing oxide mix similar to 85-41c (Composition A). Samples 85-44 and 85-41c were reduced to powders by grinding in a SPEX shatterbox for 3 min, and then subjected to grinding in an agate mortar and pestle for an additional 15 min under ethanol. Composition A was prepared from a mix of high purity oxides (Gaetani and Grove 1998). Composition A was conditioned by melting it in a gas pressure vessel at 200 MPa, H<sub>2</sub>O-saturated conditions. This produced a starting composition with 4.5 wt% H<sub>2</sub>O, and this material was powdered in an agate mortar and used as the starting mix. Experiments were carried out under a variety of conditions with respect to H<sub>2</sub>O content: dry at 0.1 MPa and 1 GPa, H<sub>2</sub>O-saturated at 200 and 800 MPa, and H<sub>2</sub>O-under-saturated at 1 to 1.3 GPa.

Table 1 Compositions of Mt. Shasta region lavas used as starting materials in this study and other compositions used in  $H_2O$ -saturated melting experiments of subduction zone magmas

Sample	SiO <sub>2</sub>	TiO <sub>2</sub>	Al <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	$P_2O_5$	Sum	Mg#
85-41c <sup>a</sup>	57.79	0.60	14.46	5.74	0.11	9.14	8.17	3.11	0.71	0.15	100.0	0.74
85-44 <sup>a</sup>	51.68	0.60	16.40	7.93	0.16	10.79	9.67	2.24	0.42	0.11	100.0	0.71
$A^a$	57.03	0.56	14.16	7.20	0.07	9.16	8.02	3.00	0.69	0.11	100.0	0.70
79-35g <sup>b</sup>	48.2	0.66	18.2	8.39	0.16	9.86	12.0	2.29	0.09	0.05	99.9	0.68
82-66 <sup>b</sup>	51.2	1.01	17.3	8.66	0.16	7.47	10.2	3.11	0.70	0.16	100.0	0.61
87S35 <sup>b</sup>	51.12	1.28	19.3	8.78	0.17	4.36	8.94	4.27	1.01	0.37	99.6	0.47
$1140Mf^{c}$	56.3	0.95	17.4	7.17	0.12	4.42	7.88	3.54	1.26	0.18	99.2	0.47
1471Mb <sup>d</sup>	52.1	1.13	17.9	9.03	0.16	5.98	9.70	3.44	0.48	0.18	100.7	0.54
Mas-22 <sup>e</sup>	55.83	0.75	17.59	6.05		6.75	7.36	4.01	1.19	0.27	99.8	0.67
MAS-911 <sup>f</sup>	59.28	0.91	14.8	4.82	0.08	6.64	7.40	3.83	1.77	0.19	99.7	0.71

<sup>&</sup>lt;sup>a</sup>Analyses reported in Baker et al. (1994) and Grove et al. (2002). Primitive magnesian andesite corrected for crustal contamination and calculated to be in equilibrium with Fo<sub>90</sub> olivine <sup>b</sup>Medicine Lake basalts and Sierra Nevada mafic dikes investigated

Synthesis experiments at 0.1 MPa were performed on 85-41c in Deltech DT31-VT rapid-quench gas mixing furnaces. Experiments were performed at the quartz–fayalite–magnetite (QFM) buffer using  $\rm CO_2$ – $\rm H_2$  gas mixtures at flow rates of 0.1 ml/s or less. Experimental methods were similar to those described in Tormey et al. (1987) with the following exceptions. Pellets of starting material (0.07–0.09 g) were pressed using elvanol as a binder, and sintered to 0.008" FePt alloy loops preconditioned to minimize Fe exchange between the silicate sample and the loop (Grove 1981). Oxygen fugacity was monitored using  $\rm CaO$ – $\rm ZrO_2$  electrolyte cells calibrated at the Fe–FeO, Ni–NiO, and Cu–Cu<sub>2</sub>O buffers. Temperature was monitored using Pt-10Rh thermocouples calibrated against the melting points of NaCl, Au, and Pd on the IPTS 1968 temperature scale.

Synthesis experiments were carried out at 200 MPa on both 85-44 and 85-41c under H<sub>2</sub>O-saturated conditions in ZHM (Zirconium-Hafnium carbide-Molybdenum) cold seal pressure vessels following experimental procedures described in Sisson and Grove (1993a, 1993b) with the following exceptions. Experiments performed at temperatures < 1,050 °C used Au inner and outer capsules while experiments performed above 1,050 °C used Au<sub>90</sub>Pd10 alloy inner and outer capsules. Oxygen fugacity was controlled at the nickel-nickel oxide (NNO) buffer. The buffer assemblage was isolated from contact with the Au by placing it into two or three unsealed Pt capsules. Temperature measurement, calibration, pressure application, and gas mixtures are as described in Sisson and Grove (1993a, 1993b). The pressure vessel was positioned vertically in a Deltech DT31VT furnace and held at pressure and temperature of the experiment for its duration. Experiments were terminated by removing the vessel from the furnace, inverting it and rapping on the hot portion of the vessel with a wrench. The capsule dropped to the water-cooled pressure seal and quenched rapidly with no growth of quench crystals.

Synthesis experiments at 800 MPa were carried out on 85-44 under  $\rm H_2O$ -saturated conditions in a 0.5" solid medium piston cylinder apparatus (Boyd and England 1960). Experiments were pressurized to 800 MPa at room temperature, and then the temperature was raised to 865 °C at 100 °C/min. The experiment was held at these conditions for 6 min, and then the temperature was raised to the final run conditions at 50 °C/min. The sample was held at isothermal conditions for the duration of the experiment (Table 2). At the end of the experiment, the sample was quenched by turning off the power. The pressure medium was BaCO<sub>3</sub>, and pressure was calibrated against the reaction: anorthite + gehlenite + corundum = Ca-tschermakite (Hays 1966). Pressures are thought to be accurate to  $\pm$ 50 MPa. Temperature was monitored and controlled using  $W_{97}Re3-W_{75}Re25$  thermocouples with no correction applied for the effect of pressure on thermocouple EMF.

<sup>d</sup>Medicine Lake basalts and Sierra Nevada mafic dikes investigated by Wagner et al. (1995)

<sup>e</sup>Basaltic andesite from the western Mexican volcanic belt (Moore and Carmichael 1998)

<sup>f</sup>Andesite from the western Mexican volcanic belt (Blatter and Carmichael 1998)

The sample was placed in the center of the furnace, and a correction of 20 °C was applied to account for the temperature difference between the thermocouple and the hot spot. This temperature difference was measured using offset thermocouples. Temperatures are thought to be accurate to  $\pm\,10$  °C.

A modified two-capsule technique similar to that used for the ZHM experiments at 200 MPa was developed to achieve  $\rm H_2O$ -saturation and oxygen buffering in the piston cylinder at 800 MPa. A shorter outer Au capsule ( $\sim\!0.75''$ ) was triple crimped on the bottom, welded, and flattened by hammering with a drill blank. An inner Au capsule containing the silicate starting material was loaded into the outer capsule along with two Pt capsules containing the Ni–NiO buffer. These capsules were welded on one end and crimped on the other. A micro-syringe was used to add  $\sim\!15$ –20  $\mu l$   $\rm H_2O$  and the outer capsule was crimped flat and welded shut.

After termination of the 200 and 800 MPa H<sub>2</sub>O-saturated experiments by quenching, the outer capsules were punctured to verify the presence of water. The buffer capsules were examined for the presence of both Ni and NiO, and to verify that there was no leakage of silicate material into the buffer capsules. Experiments that passed both of these tests were considered successful and were mounted for analysis by electron microprobe. Experimental temperatures and durations are reported in Table 2 and chemical analyses of experimental products are presented in Table 3.

Anhydrous and H<sub>2</sub>O-undersaturated experiments on composition A were carried out in a piston cylinder apparatus. Anhydrous experiments used oxide mix samples packed into a graphite capsule that was then placed in a Pt capsule and welded shut. H<sub>2</sub>O-bearing sample powders of composition A (4.5 wt% H<sub>2</sub>O) were used for H<sub>2</sub>O-undersaturated melting experiments. These powders were placed in an Au<sub>80</sub>Pd<sub>20</sub>-Fe alloy capsule and welded shut. The capsules were placed inside graphite capsule machined from graphite rod. The experimental setup is essentially the same as developed by Gaetani and Grove (1998). The Au<sub>80</sub>Pd<sub>20</sub> capsules were pre-conditioned with Fe before the high-pressure experiments. This was done by filling them with an andesite composition and placing them in a 1-atm furnace at  $\log f_{\rm O2} = \rm QFM$ -1 and 1,250 °C for 72 h. The silicate glass was then removed using HF + HNO<sub>3</sub>. Fluoride residues after HF dissolution were first removed manually with a pick and then by soaking in HCl. Experimental conditions are given in Table 2 and compositions are reported in Table 3.

Experimental durations varied widely depending on limitations imposed by the experimental geometry and the time duration inferred from previous experimental studies required for establishment of exchange equilibrium. In the 0.1-MPa experiments the duration varied from 40 h at near liquidus conditions to 230 h at high degrees of crystallinity. These durations were based on the time that Grove and Juster (1989) found necessary for establishment

<sup>&</sup>lt;sup>b</sup>Medicine Lake basalts and Sierra Nevada mafic dikes investigated by Sisson and Grove (1993a)

<sup>&</sup>lt;sup>c</sup>Medicine Lake basalts and Sierra Nevada mafic dikes investigated by Grove et al. (1997)

**Table 2** Experimental conditions and products. Experiments at 0.1 MPa were buffered at QFM using CO<sub>2</sub>-H<sub>2</sub> gas mixing. Experiments on 85-41c and 85-44 at 200 and 800 MPa were buffered at NNO with P<sub>H2O</sub> = P<sub>lotal</sub>. Experiments on composition A are H<sub>2</sub>O-undersaturated and contain 0 or 4.5 wt% H<sub>2</sub>O added. See Table 3 for phase compositions

Run no.	T (°C)	P (MPa)	Time (h)	Phases <sup>a</sup>	$\Sigma R^2$	% Fe loss <sup>b</sup>	K <sub>D</sub> ol/liq	K <sub>D</sub> pl/liq	K <sub>D</sub> pl/liq K <sub>D</sub> cpx/liq	K <sub>D</sub> opx/liq	K <sub>D</sub> amp/liq
85-41c											
#1	1,250	0.1	4	gl(100), ol(<1)	98.0	12.0	0.266				
#2	1,225	0.1	4	gl (98), ol (3), sp (-1)	0.56	5.6	0.289				
#3	1,200	0.1	53	gl (89), ol (-3), opx (16), pl (-2)	0.07	< \ \	0.272	1.27		0.256	
#4	1,175	0.1	40	gl (90), ol (-4), opx (14), sp ( $<$ 1)	0.76	12.0	0.247			0.196	
8#	1,160	0.1	174	$g_1(57), cpx(5), opx(18), p_1(20)$	69.0	-4.8		1.29	0.250	0.256	
#2	1,150	0.1	20	gl (63), cpx (6), opx (16), pl (15)	0.02	\ \ -		1.57	0.233	0.210	
9#	1,125	0.1	98	gl (40), cpx (11), opx (20), pl (29)	0.13	3.5		1.80	0.203	0.203	
47	1,120	0.1	230	gl (44), cpx (9), opx (20), pl (27)	0.15	2.3		1.71	0.219	0.222	
#2	1,040	200	56	(82), ol (8), cpx (1	0.21	-7.2	0.275		0.247		
47	1,020	200	4	(75), ol (12), cpx (	0.02	<u>~</u>	0.293		0.244	0.281	
8#	086	200	20	(46), cpx (15), opy	0.11	-4.6		2.34	0.216	0.284	
6#	940	200	48	gl (40), cpx (16), opx (21), pl (24), sp	0.16	1.9		2.83	0.221	0.272	
85-44											
47	1,090	200		gl (96), ol (3), cpx (1)	0.31	\ \ -	0.291		0.249		
9#	1,070	200		gl (77), ol (4), cpx (19), sp ( $<$ 1)		1.5	0.309				
#3	1,030	200		gl (55), ol (17), cpx (10), sp ( $<$ 1), pl (18)		\ \ -	0.298	4.03	0.246		
#4	1,010	200	19	gl $(47)$ , ol $(18)$ , cpx $(13)$ , sp $(<1)$ , pl $(23)$	0.26	\ \ 	0.328	4.42	0.250		
#2	066	200		gl (41), ol (-5), cpx (28), pl (21), amp (16), sp $(<1)$		\ \ 	0.292	5.55	0.223		0.297
B372	1,090	800		gl, ol							
B375	1,077	800		gl, ol, cpx							
B373	1,065	800		gl (52), amph (62), opx (-3), cpx (-9)	0.28	8.4			0.284	0.332	0.340
B380	1,045	800		gl (47), amph (52), cpx (1), sp (2)	0.24	3.1			0.252		0.313
Ą				wt. % $H_2O$ added + Phases							
A1.2	1,300	1,500		0.0  gl  (77),  opx  (9),  cpx  (14)	0.62	-1.1			0.33	0.35	
A1.7	1,175	1,000	30	4.5 gl, ol							
A1.8	1,200	1,300		4.5 gl, opx, cpx							
A1.9	1,200	1,150		4.5 gl, opx, cpx							
A1.10	1,175	1,000		4.5 gl, opx, cpx							

 $^a Phase\ proportions\ calculated\ by\ mass\ balance,\ ignoring\ MnO\ and\ P_2O_5.\ Abbreviations\ for\ phases\ are\ in\ Table\ 3$   $^b Apparent\ loss\ or\ gain\ of\ FeO\ estimated\ as\ 100*(FeO_{calc}-FeO_{starting\ material})/FeO_{starting\ material}$ 

of exchange equilibrium between crystals and melt for major elements. At 200 and 800 MPa experiment duration is limited by the rate of hydrogen diffusion through the Au outer capsule. Experiments at 1,070 and 1,090 °C could be continued for  $\sim$ 24 h. At temperatures below 1,000 °C durations of 48 h were possible.

#### Analytical methods

Compositions of minerals and quenched glasses in the experimental products were obtained with the MIT five-spectrometer JEOL 733 Superprobe using on-line data reduction. For the 200- and 800-MPa experiments analyses were made with automation software that used the matrix correction procedures of Bence and Albee (1968) with modifications of Albee and Ray (1970). For the 0.1-MPa and 1-1.3-GPa experiments on 85-41c newer automation software was employed that utilized the CITZAF correction package of Armstrong (1995). The atomic number correction of Duncumb and Reed, Heinrich's tabulation of absorption coefficients, and the fluorescence correction of Reed were used to obtain a quantitative analysis (Armstrong 1995). A comparison of the two techniques on a basalt standard glass showed no appreciable difference between the results obtained by the two methods. Analyses of mineral and quenched liquid products are presented in Table 3. All analyses were performed using a 15-kV accelerating voltage and a 10-nA beam current. For the glass phase, the beam size was  $\sim$ 20  $\mu$ m and Na was counted at the beginning of the analysis for 5 s. Other elements were measured for up to 40 s, depending on abundance level. Analytical precision can be inferred from replicate analyses of an andesite glass working standard (38b-129) from a 0.1-MPa experiment (Grove and Juster 1989). One standard deviation of replicate glass analyses expressed as relative percent of oxides are SiO<sub>2</sub>:0.4%, Al<sub>2</sub>O<sub>3</sub>:0.9%, CaO:1.5%, MgO:1.5%, FeO:1.4%, MnO:8.1%, P<sub>2</sub>O<sub>5</sub>:5%, Na<sub>2</sub>O:1.9%, K<sub>2</sub>O:1.1%, based on 289 individual measurements over 28 analytical sessions. The mean sum of analyses of the anhydrous glass is 99.4%. Analytical precision for NiO analyses in olivine in Table 3 are poor, but are presented as an indication of the absence of contamination of the silicate sample by the buffer.

#### **Experimental results**

## Approach to equilibrium

The experiments presented here are all synthesis experiments where crystals and melt have grown from starting materials that consist of a powdered volcanic rock or a melted oxide mix. Phase appearance temperatures have not been reversed because past experience (Grove and Bence 1977) indicates that direct synthesis is sufficient to recover equilibrium appearance temperature when significant (>40 wt%) melt is present. For the reasons outlined below, we conclude that the experiments approached equilibrium sufficiently so that they can be used to understand crystallization behavior under the conditions of this investigation. Maintenance of constant sample composition with little or no loss or gain of iron is an important prerequisite for an approach to equilibrium. A materials-balance technique (Bryan et al. 1969) was used to estimate phase proportions in the experimental products and to estimate iron loss or gain from the starting material. Results are summarized in Table 2, and the loss or gain of iron is considered to be acceptably low, generally less than  $\pm 5$  wt% relative. Two exceptions are the 0.1 MPa experiments at 1,250 and 1,175 °C experiments that gained iron from the FePt alloy loop.

We include these experiments despite the iron gain because they provide useful information of the low-pressure liquidus phase relations. The reason for the iron gain is that the loops used for these experiments contained more Fe in the alloy than the desired equilibrium amount (7–9 wt%) and this was transferred to the silicate portion of the charge.

Another important criterion for equilibrium is the achievement of regular and consistent partitioning of major elements between crystalline phases and melt. The mineral-melt Fe-Mg exchange distribution coefficient (K<sub>D</sub><sup>Fe-Mg</sup>) between olivine (oliv) and melt and pyroxene and melt are constant within each bulk composition. For 85-44, the oliv–melt Fe–Mg exchange distribution coefficients ( $K_D^{Fe-Mg} = [X_{Fe}^{Ol}X_{Mg}^{Liq}]/[X_{Mg}^{Ol}X_{Fe}^{Liq}]$ ) have an average value of 0.30 ( $\pm$ 0.02 at 2 $\sigma$ ) for the 200-MPa experimental pairs. For 85-41c, the oliv-melt Fe-Mg exchange distribution coefficients the average value is 0.27  $(\pm 0.04 \text{ at } 2\sigma)$  for the 200-MPa experimental pairs. Even though these averages are within  $2\sigma$  of each other there appears to be a systematic difference in the oliv-melt  ${
m K_D^{Fe-Mg}}$ , possibly related to the higher SiO<sub>2</sub> content of 85-41c. The high-Ca pyroxene (cpx)  ${
m K_D^{Fe-Mg}}$  is 0.24  $(\pm 0.02 \text{ at } 2\sigma)$  for the 85-44 and 85-41c pyroxenes and the orthopyroxene (opx)-melt  $K_D^{\text{Fe-Mg}} = 0.24(\pm 0.04 \text{ at } 2\sigma)$ for the eight opx-bearing experiments in 85-41c. The average of the three amphibole (amph)-melt  $K_D^{\text{Fe-Mg}}$ values from the 200- and 800-MPa experiments on 85-44 has a value of 0.32 ( $\pm$  0.04 at  $2\sigma$ ). In composition 85-41c, the plagioclase (plag)—melt Ca–Na exchange distribution coefficient ( $K_D^{Ca-Na} = [X_{Ca}^{Pl} X_{Na}^{Liq}]/[X_{Na}^{Pl} X_{Ca}^{Liq}]$ ) for the 0.1-MPa experiments is 1.6 ( $\pm$ 0.1 at 2 $\sigma$ ) and the term of the  $\Delta$ 1 at 2 $\Delta$ 2 and the  $\Delta$ 2 at 2 $\Delta$ 3 and the  $\Delta$ 3 at 2 $\Delta$ 4 at 2 $\Delta$ 5 and the  $\Delta$ 3 at 2 $\Delta$ 4 at 2 $\Delta$ 5 and the  $\Delta$ 4 at 2 $\Delta$ 5 at 2 $\Delta$ 5 at 2 $\Delta$ 6 at 2 $\Delta$ 6 at 2 $\Delta$ 6 at 2 $\Delta$ 6 at 2 $\Delta$ 7 at 2 $\Delta$ 8 at 2 $\Delta$ 9 at plag-saturated 200 MPa experiments give a  $K_D^{Ca-Na}$ = 2.6. In the basaltic andesite (85-44), the 200-MPa hydrous plag-melt  $K_D^{\text{Ca-Na}}$  for the three plag-melt pairs is 4.0  $(\pm 0.1 \text{ at } 2\sigma).$ 

The values for the exchange K<sub>D</sub>s and their variations provide an indication of the extent to which a mineral has established Fe–Mg or Ca–Na exchange equilibrium with its enclosing melt. Olivine–melt, cpx–melt, and opx–melt pairs show a regular and nearly constant value for K<sub>D</sub>. Plagioclase–melt K<sub>D</sub><sup>Ca–Na</sup> at 200 MPa is similar to the values found by Gaetani et al. (1994) for Lau Basin basaltic andesite, and lower than the value of 5.5 determined by Sisson and Grove (1993a) for basaltic melts and Moore and Carmichael (1998) for a Mexican volcanic belt basaltic andesite. These values and variations are generally similar to those achieved in experiments of much longer duration where Fe–Mg and Ca–Na exchange reactions have been reversed (Grove and Juster 1989) and are taken to indicate a sufficiently close approach to equilibrium.

#### Mineral appearances

#### Basaltic andesite 85-44

The phase appearance sequence for the primitive basaltic andesite is summarized in Fig. 2. At 0.1 MPa,

Table 3 Electron microprobe analyses of phases produced in experiments

Run no.	Phase <sup>a</sup>	SiO <sub>2</sub>	TiO <sub>2</sub>	Al <sub>2</sub> O <sub>3</sub>	$Cr_2O_3$	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	$K_20$	$P_2O_5$	NiO	Total
85-41c	0.1 MPa													
83-416#1	gl (9) <sup>b</sup> ol (4)	56.5 (3)° 40.9 (9)	0.68 (2)	14.8 (1) 0.05 (2)	0.08 (1) 0.03 (4)	6.31 (9) 9.45 (12)	0.08 (2) 0.16 (2)	9.01 (8) 50.7 (7)	7.98 (8) 0.23 (2)	3.08 (6)	0.69 (2)	0.09 (3)	0.29 (1)	99.28 101.81
65-41C#2	gl (7) ol (4) sp (3)	57.0 (7) 40.1 (7) 0.15 (26)	0.74 (4) 0.02 (2) 0.33 (6)	15.4 (1) 0.05 (1) 21.0 (2)	0.07 (1) 0.15 (7) 46.6 (3)	5.91 (8) 10.7 (2) 15.5 (4)	0.07 (2) 0.15 (1) 0.03 (3)	7.78 (9) 48.6 (5) 15.7 (1)	8.23 (12) 0.26 (2) 0.18 (5)	3.30 (8)	0.73 (3)	0.15 (3)	0.13 (3)	99.34 100.16 99.49
85-41c#3	gl (10) ol (5) pl (7) opx (4)	56.7 (3) 40.4 (5) 53.7 (7) 56.5 (8)	0.74 (3)	16.4 (6) 0.11 (6) 28.9 (5) 1.11 (25)	0.04 (2) 0.10 (9) 0.48 (12)	5.51 (9) 11.6 (2) 0.54 (3) 7.31 (50)	0.06 (2) 0.17 (2) 0.15 (2)	6.22 (8) 48.0 (7) 0.34 (4) 32.2 (2)	8.73 (10) 0.29 (5) 12.2 (4) 2.02 (19)	3.65 (11) 4.02 (25)	0.79 (4) 0.15 (1)	0.15 (3)	0.15 (2)	99.00 100.82 99.85 99.88
85-41c#4	gl (10) ol (5) opx (4) sp (4)	56.8 (2)) 39.4 (6) 56.4 (3) 0.18 (37)	0.93 (3) 0.02 (2) 0.09 (2) 0.43 (5)	15.6 (1) 0.05 (2) 1.18 (18) 18.1 (9)	0.06 (1) 0.03 (7) 0.61 (22) 48.2 (9)	6.71 (8) 11.5 (2) 6.26 (67) 17.8 (9)	0.07 (2) 0.17 (1) 0.13 (2) 0.10 (3)	6.89 (5) 47.8 (7) 32.8 (5) 14.1 (4)	8.48 (10) 0.28 (3) 1.68 (22) 0.10 (5)	3.33 (10)	0.73 (3)	0.15 (3)	0.18 (2)	99.69 99.43 99.15 99.01
65-41c#3	gl (8) cpx (10) opx (8) pl (10)	58.7 (7) 53.5 (5) 56.1 (5) 54.6 (6)	1.01 (4) 0.38 (7) 0.16 (3)	15.5 (2) 2.21 (47) 1.25 (44) 28.7 (5)	0.03 (1) 0.60 (28) 0.20 (5)	6.16 (7) 5.89 (49) 8.70 (67) 0.74 (19)	0.06 (2) 0.14 (2) 0.19 (2)	4.63 (23) 19.0 (4) 31.2 (8) 0.39 (16)	7.60 (10) 18.9 (5) 2.24 (17) 12.1 (3)	3.93 (16) 0.19 (7) 3.98 (23)	1.11 (4) 0.17 (3)	0.23 (3)		98.97 100.81 100.04 100.68
85-41¢#6	gl (6) cpx (8) opx (8) pl (7)	63.0 (3) 53.2 (3) 55.8 (2) 55.0 (4)	1.33 (9) 0.50 (16) 0.21 (4)	14.8 (2) 2.16 (30) 0.96 (40) 28.7 (5)	0.23 (12) 0.16 (5)	6.21 (45) 7.56 (57) 11.7 (12) 1.14 (31)	0.06 (2) 0.22 (1) 0.24 (4)	3.12 (10) 18.7 (6) 29.0 (8) 0.18 (11)	6.03 (12) 17.6 (7) 2.40 (31) 11.4 (4)	4.02 (9) 0.12 (6) 0.03 (2) 4.24 (27)	1.60 (5)	0.33 (4)		100.49 100.29 100.5 100.86
65 41 c 110	gl (17) cpx (5) opx (6) pl (5)	63.1 (3) 53.4 (3) 55.3 (6) 54.4 (3)	1.16 (4) 0.49 (16) 0.22 (6)	14.5 (1) 1.66 (16) 1.27 (51) 28.4 (3)	0.47 (13)	5.98 (15) 7.34 (21) 11.2 (4) 0.60 (4)	0.11 (2) 0.19 (3) 0.22 (3)	3.49 (4) 19.6 (5) 29.4 (4) 0.40 (2)	6.55 (16) 17.5 (6) 2.29 (26) 12.0 (2)	3.80 (20) 0.11 (8) 0.02 (2) 4.07 (15)	1.47 (4)	0.37 (2)	0.05 (2)	100.65 100.76 100.09 100.05
85-41c	gl (24) cpx (7) opx (7) pl (7) 200 MPa	60.0 (4) 53.4 (4) 56.2 (4) 54.0 (5)	0.93 (2) 0.34 (6) 0.16 (5)	14.9 (1) 1.95 (38) 1.14 (26) 28.8 (3)	0.46 (11)	5.77 (15) 5.95 (69) 9.74 (64) 0.53 (5)	0.10 (2) 0.17 (2) 0.21 (4)	4.66 (8) 19.2 (7) 30.7 (7) 0.28 (3)	7.80 (12) 18.3 (8) 2.38 (22) 12.0 (3)	3.43 (11) 0.12 (3) 4.08 (13)	2.34 (6) 0.32 (3)	0.25 (3)		100.18 99.89 100.74 100.01
65-41C#3	gl (17) ol (9) cpx (7)	60.4 (5) 39.7 (3) 53.6 (3)	0.66 (3)	17.3 (2) 0.03 (1) 1.79 (27)	0.01 (2) 0.70 (15)	4.62 (9) 13.3 (3) 4.72 (59)	0.11 (2) 0.21 (2) 0.14 (1)	4.44 (5) 46.5 (4) 18.4 (5)	7.36 (16) 0.21 (1) 20.9 (6)	3.81 (18) 0.14 (5)	0.95 (4)	0.24 (2)	0.03 (2)	94.41 <sup>d</sup> 100.04 100.66
1#51#-50	gl (15) ol (9) cpx (5) opx (6) sp (4)	61.7 (5) 39.1 (3) 52.7 (4) 55.5 (3) 0.31 (33)	0.70 (3) 0.02 (1) 0.42 (2) 0.19 (2) 0.74 (15)	18.4 (3) 0.03 (3) 2.36 (21) 1.64 (18) 13.0 (20)	0.04 (5) 0.49 (8) 0.25 (7) 46.4 (3.4)	4.18 (14) 17.2 (2) 5.69 (12) 11.5 (4) 29.4 (1.8)	0.09 (2) 0.24 (3) 0.17 (2) 0.23 (2) 0.28 (4)	3.07 (19) 43.1 (2) 17.1 (3) 30.0 (4) 8.41 (34)	6.33 (27) 0.18 (3) 21.5 (2) 1.37 (8) 0.26 (3)	4.13 (17)	1.09 (4)	0.23 (3)	0.04 (2) 0.11 (4) 0.05 (2)	93.45 100.02 100.60 100.68 98.85

96.86 100.28 99.80 100.36	93.8 99.5 100.2 100.4	93.42 99.45 100.43 100.82 99.23	93.19 99.79 99.88 100.21 99.69	93.54 97.64 100.39 100.77 100.46 99.72	94.63 100.12	94.84 99.81 100.35 99.92	100.3	85.76 96.8 100.0 99.9
0.04 (2)		0.16 (3)	0.14 (2)	0.14 (2)	0.01 (1)	0.01 (1)	0.04 (2)	
0.33 (5)	0.23 (3)	0.21 (3)	0.21 (3)	0.23 (2)				
1.46 (4)	1.50 (2)	0.59 (4)	0.64 (3)	0.75 (3) 0.22 (2) 0.06 (3)	0.58 (2)	0.43 (3)		0.29 (2) 0.21 (2)
4.60 (17) 0.17 (4) 4.26 (33) 0.03 (2)	4.70 (19) 0.28 (3) 0.08 (5) 4.35 (23)	4.07 (15) 2.06 (10) 0.28 (3)	4.62 (7) 2.30 (15) 0.28 (4)	5.20 (18) 2.75 (8) 2.39 (12) 0.29 (4)	4.17 (19)	3.24 (16) 0.24 (13) 0.25		1.21 (12) 2.07 (12) 0.22 (3)
5.39 (15) 19.8 (6) 11.7 (6) 2.04 (20)	4.63 (13) 20.3 (3) 1.33 (9) 12.1 (4)	8.19 (24) 0.27 (2) 16.7 (3) 20.9 (6) 0.23 (10)	7.31 (18) 0.27 (2) 16.1 (6) 20.9 (9) 0.27 (9)	6.27 (26) 11.5 (1) 0.25 (2) 16.0 (6) 20.9 (7) 0.25 (8)	8.19 (14) 0.28 (3)	10.3 (1) 21.9 (7) 0.30 (5) 0.25	0.14 (3)	8.85 (13) 11.1 (1) 20.2 (1.0) 1.29 (4)
2.59 (11) 17.9 (7) 0.26 (11) 28.0 (4)	1.75 (11) 16.1 (2) 27.2 (3) 0.21 (4)	4.05 (13) 40.9 (7) 0.38 (14) 16.4 (4) 9.10 (1.92)	3.51 (7) 38.8 (6) 0.19 (17) 16.0 (6) 9.09 (1.28)	2.86 (4) 15.0 (3) 38.8 (8) 0.17 (11) 15.8 (4) 8.60 (1.80)	4.16 (10) 41.9 (8)	5.64 (9) 16.5 (5) 45.5 (4) 10.7	47.0 (7)	4.98 (15) 17.0 (2) 16.9 (6) 28.2 (4)
0.08 (2) 0.16 (4) 0.23 (2)	0.04 (2) 0.16 (5) 0.28 (5)	0.16 (7) 0.30 (3) 0.15 (3) 0.41 (4)	0.20 (5) 0.31 (4) 0.18 (4) 0.40 (2)	0.14 (2) 0.13 (2) 0.34 (2) 0.19 (3) 0.42 (5)	0.11 (4) 0.26 (2)	0.15 (3) 0.11 (4) 0.18 (4) 0.31	0.14(1)	0.14 (2) 0.09 (1) 0.17 (2) 0.28 (2)
4.21 (12) 6.28 (38) 0.78 (12) 13.0 (3)	3.55 (6) 7.22 (45) 15.0 (4) 0.64 (9)	6.35 (14) 19.1 (7) 0.74 (12) 6.23 (38) 32.9 (6.9)	6.01 (10) 21.8 (3) 0.64 (21) 6.86 (88) 32.4 (3.8)	5.68 (12) 8.85 (9) 22.5 (5) 0.65 (13) 6.99 (49) 32.6 (6.1)	5.98 (11) 18.6 (7)	6.31 (14) 4.60 (33) 14.8 (1) 30.1	12.5 (4)	6.80 (11) 7.90 (18) 6.56 (48) 12.8 (2)
0.44 (19)	0.18 (4)	0.04 (2) 0.04 (3) 0.63 (11) 38.1 (4.2)	0.03 (4) 0.04 (1) 0.54 (8) 37.4 (1.8)	0.07 (4) 0.92 (20) 0.03 (2) 0.46 (10) 37.5 (2.6)	0.02 (2) 0.03 (3)	0.01 (1) 0.77 (13) 0.10 (3) 37.9	0.03 (2)	0.01 (1) 0.25 (11) 0.25 (25) 0.20 (4)
17.7 (2) 2.23 (31) 28.8 (4) 1.28 (36)	17.6 (2) 2.39 (35) 1.49 (21) 29.4 (5)	18.8 (3) 0.07 (3) 32.1 (5) 3.14 (51) 16.7 (3.7)	18.6 (2) 0.10 (3) 32.3 (8) 3.02 (44) 19.1 (3.4)	18.3 (2) 12.4 (3) 0.11 (3) 32.4 (1.0) 3.06 (32) 18.8 (4.8)	19.0 (2) 0.05 (6)	19.0 (1) 3.59 (33) 0.10 (4) 17.9	0.03 (1)	17.1 (2) 12.4 (2) 2.38 (48) 1.61 (10)
0.89 (3) 0.50 (5) 0.25 (6)	0.47 (4) 0.58 (7) 0.24 (2)	0.86 (5) 0.01 (1) 0.59 (9) 1.55 (1.53)	0.92 (4) 0.03 (2) 0.63 (10) 0.80 (28)	0.96 (2) 3.07 (12) 0.02 (2) 0.67 (11) 1.23 (1.75)	0.80 (3) 0.10 (3)	0.70 (3) 0.40 (7) 0.07 (3) 1.33	0.02 (1)	0.57 (2) 0.71 (5) 0.24 (8) 0.06 (1)
62.7 (5) 52.8 (5) 53.8 (6) 55.4 (6)	65.6 (2) 52.3 (3) 54.6 (4) 53.5 (3)	56.7 (4) 38.6 (7) 48.4 (5) 52.5 (5)	58.0 (4) 38.3 (3) 48.3 (6) 51.8 (6)	59.5 (3) 42.8 (5) 38.2 (6) 49.1 (7) 52.1 (6) 0.09 (24)	57.0 (4) 38.9 (2)	54.2 (2) 51.7 (4) 39.3 (2) 0.09	40.4 (4)	57.2 (2) 45.2 (4) 53.1 (5) 55.5 (6)
gl (21) cpx (4) pl (5) opx (5)	gl (8) cpx (6) opx (6) pl (4) 200 MPa	gl (9) ol (5) pl (7) cpx (10) sp (5)	gl (8) ol (5) pl (11) cpx (11) sp (5)	gl (7) amp (6) ol (6) pl (8) cpx (8) sp (7)	gl (8) ol (8)	gl (8) cpx (13) ol (9) sp (1) <b>800 MPa</b>	ol (5)	gl (6) amp (10) cpx (9) opx (7)
85-41c#8 85-41c#9	85-44 85-44#3	2 × × × × × × × × × × × × × × × × × × ×	24.4.5	0 3HV 20	0#44-50	85-44 R372	B373	

Table 3 (Contd.)	Contd.)													
Run no.	$Phase^{\mathrm{a}}$	$SiO_2$	${ m TiO}_2$	$Al_2O_3$	$Cr_2O_3$	FeO	$M_{nO}$	$_{ m MgO}$	CaO	$\mathrm{Na}_2\mathrm{O}$	$K_20$	$\mathbf{P}_2\mathbf{O}_5$	NiO	Total
B380														
	gl (4)	59.1 (2)	0.39(2)	19.7 (1)	0.04 (3)	5.86 (4)	0.14(2)	3.69 (6)	8.22 (8)	2.24 (18)	0.72(2)			85.66
	amp (4)	44.4 (4)	0.70 (4)	12.9 (4)	0.12(10)	8.20 (82)	0.10(2)	16.5 (4)	11.3 (4)	2.34 (3)	0.25(1)			8.96
	cpx(5)	52.1 (6)	0.25(9)	2.58 (55)	0.11 (11)	6.71 (63)	0.21(3)	16.8 (6)	20.9	0.25(2)				9.66
	sp (4)	0.11(3)	0.28 (4)	24.1 (2.2)	39.2 (1.5)	23.9 (4.6)	0.30(9)	11.5 (2.3)	0.29(5)					8.66
A1.2	gl (13)	57.8 (3)	0.73(1)	17.6 (1)	0.00 (1)	4.07 (16)	0.04(1)	6.16 (6)	9.03 (11)	3.19 (14)	0.58(2)	0.12(2)	0.00(1)	99.32
	opx (15)	54.3 (4)	0.14(2)	4.69 (53)	0.24(2)	6.70 (14)	0.09 (1)	30.9 (4)	2.16 (8)	0.1 (1)				99.32
	cpx (23)	52.2 (5)	0.24(2)	5.84 (38)	0.32 (2)	4.59 (14)	0.10(2)	19.6 (3)	15.8 (6)	0.62 (4)				69.86

Cone standard deviation of replicate analyses in terms of least unit cited. Thus 56.5 (3) should be read as  $56.5 \pm 0.3$  danalyses of hydrous glasses are normalized to 100% with all Fe reported as FeO. Unnormalized total is reported gl Glass; ol olivine; cpx clinopyroxene; pl plagioclase; sp spinel; amp amphibole Number of microprobe analyses

1300 85-44 H<sub>2</sub>O-saturated 12501 NNO buffer 1200 Oliv in 1100 1050 1000 950 600 800 1000 0 200 400 P in MPa

**Fig. 2** Pressure–temperature diagram showing the results of phase equilibrium experiments performed on 85-44 under H<sub>2</sub>O-saturated conditions at the NNO buffer at 200 and 800 MPa (see Tables 2 and 3 for durations and compositions). The 0.1-MPa experiments were carried out at the QFM buffer and results from 85-44 (near liquidus) and a compositionally similar primitive basaltic andesite (82-94a) are reported in Baker at al. (1994). *Black squares* indicate an experimental data point. *Lines* indicate the effect of pressure and H<sub>2</sub>O content on phase appearance temperature. The oliv in curve is inferred from results of Sisson and Grove (1993a). The plag in boundary at 800 MPa is inferred from the experiments of Holloway and Burnham (1972). The phase appearance boundaries reported here are very similar to those obtained by Holloway and Burnham (1972) on the 1921 Kilauea basalt using 58% CO<sub>2</sub>-42%H<sub>2</sub>O and NNO (see their Fig. 3)

experimental results on 85-44 (Baker et al. 1994) place the liquidus above 1,280 °C. The Baker et al. experiments were performed under anhydrous conditions at the QFM buffer. Lower temperature phase appearances were taken from experiments on 82–94a (Baker et al. 1994), which is similar in composition to 85-44. Also shown are the 200- and 800-MPa H<sub>2</sub>O-saturated experiments carried out at the Ni-NiO buffer. At the highest temperature investigated at 200 MPa (1,090 °C, Table 2), the experimental charge is  $\sim$ 96% melt with a small amount of olivine (oliv) and high-Ca pyroxene (cpx). Based on direct measurements of the effect of H<sub>2</sub>O on the melting relations of basalts (Sisson and Grove 1993a), the liquidus is inferred to be at  $\sim 1.150$  °C. At 1,070 °C, oliv, cpx, and an iron-rich Al-Cr spinel (sp) coexist. At 1,030 °C, plagioclase (plag) has joined the crystallization sequence, and the liquid has evolved to a high-Al andesite with 18.8% Al<sub>2</sub>O<sub>3</sub> at  $\sim$ 57 wt% SiO<sub>2</sub>. Amphibole (Mg#=Mg/[Mg+Fe<sup>2+</sup>] molar = 0.75), oliv, cpx, sp, and plag coexist at the lowest temperature investigated (990 °C). Under these conditions the charge is >55% crystalline and the liquid is andesitic (59.7%) SiO<sub>2</sub> and 18.3% Al<sub>2</sub>O<sub>3</sub>). Based on phase relations at 0.1 and 800 MPa, orthopyroxene (opx) is inferred to appear at a lower temperature, where it would be in reaction relation with oliv and liquid.

At 800 MPa, the high temperature experiment (1,090 °C) contained >95% quenched melt and oliv. The liquidus temperature was not determined. Both cpx and oliv are present in the 1,077 °C experiment. At 1,065 °C, oliv is no longer a part of the mineral assemblage, and opx

and a magnesian (Mg#=0.80) pargasite amphibole (amph) join cpx. The higher Mg# of amph is a result of its appearance closer to the liquidus. At the lowest temperature investigated (1,045 °C) opx reacts with amph, leading to the disappearance of opx and a liquid that coexists with amph, cpx, and sp. At these low temperatures and high  $\rm H_2O$  contents the 800 MPa derivative liquids are high alumina andesites (59%  $\rm SiO_2$  and 20%  $\rm Al_2O_3$ ). Müntener et al. (2001) observed similar enrichments in  $\rm Si$  and  $\rm Al$  in their 800-MPa experiments on 85-44. At a lower temperature, plagioclase is inferred to join the crystallizing assemblage.

The pressure–temperature relations found for 85-44 are similar to those found by Holloway and Burnham (1972) for the 1921 Kilauea basalt. The Holloway and Burnham (1972) study used a mixed CO<sub>2</sub>–H<sub>2</sub>O (42% CO<sub>2</sub> and 58% H<sub>2</sub>O, molar) fluid, and their experiments were also buffered at Ni–NiO. In both sets of experiments the appearance temperature of amph increases with increasing pressure by approximately 100 °C from 200 to 800 MPa, and amph becomes more magnesian because it appears closer to the liquidus. Plagioclase appearance temperature decreases by >150 °C over the 0.1–800-MPa pressure interval.

# Primitive magnesian andesite 85-41c

In comparison to the basaltic andesite, this composition contains much higher  $SiO_2$  (57.9 wt%) and lower  $Al_2O_3$  (14.6 wt%). At 0.1 MPa, under anhydrous conditions, oliv is the liquidus phase at  $\sim$ 1,250 °C and sp joins oliv by 1,225 °C. Orthopyroxene and plag have joined the crystallization assemblage by 1,200 °C, and oliv + liquid react to form opx over the temperature range of 1,200 to 1,167 °C. After oliv has been removed by reaction, opx, cpx, and plag co-crystallize to the lowest temperature investigated.

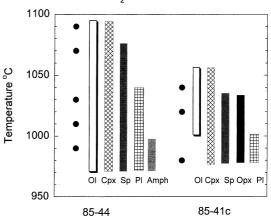
At 200 MPa, the primitive magnesian andesite has oliv and cpx as liquidus phases at the highest temperature investigated (1,040 °C). Figure 3 shows the 200-MPa phase appearance sequence for both 85-41c and 85-44. In 85-41c, opx and sp appear as crystallizing phases at  $\sim\!1,030$  °C, and oliv reacts with melt and disappears from the crystallizing assemblage at  $\sim\!1,000$  °C. In the lower temperature experiments (980 and 940 °C) plag crystallizes along with opx, cpx, and sp.

The difference in the phase appearance sequences at 200 MPa,  $H_2O$ -saturated in 85-44 and 85-41c, are a result of the higher  $SiO_2$  and lower  $Al_2O_3$  contents of 85-41c. This leads to the appearance of opx in 85-41c before plag (Fig. 3). In contrast, opx has not appeared at the lowest temperature investigated in 85-44 and plag saturation occurs at a higher temperature  $\sim$ 50 °C higher in 85-44.

## Composition A and 85-44

Experiments on the modified primitive andesite composition (A) with 4.5 wt% H<sub>2</sub>O added at 1 to 1.3 GPa

## 200 MPa, H<sub>2</sub>O-saturated, Ni-NiO buffer



**Fig. 3** Phase appearance sequences for the primitive basaltic andesite (85-44) and primitive magnesian andesite (85-41c) at 200 MPa, H<sub>2</sub>O-saturated at the NNO buffer. See Tables 2 and 3 for experimental conditions and phase compositions. In 85-44, plag appears after oliv + cpx + sp, and amphibole (amph) has just appeared in the lowest temperature experiment. In contrast, the phase appearance sequence in the high-SiO<sub>2</sub>, low-Al<sub>2</sub>O<sub>3</sub> primitive magnesian andesite is oliv, followed by cpx and sp. Opx is the next phase to appear in reaction relation with oliv. Plag is the last phase to appear in the lowest temperature experiment. The lowest temperature experiment at 940 °C contains the same assemblage present at 980 °C

(Table 2) are shown in Fig. 4 along with the results of Baker et al. (1994) for 85-44 at 1 GPa and variable  $H_2O$  content. Both compositions have pyroxene as the liquidus phase under anhydrous conditions at 1.0 to 1.5 GPa. Under  $H_2O$ -undersaturated conditions, the addition of 4.5 to 6 wt%  $H_2O$  is sufficient to bring oliv and opx simultaneously to the liquidus at 1 GPa. For composition A, the liquid is saturated with oliv, opx, and cpx at  $\sim$ 1,175 °C and 1 GPa. For 85-44, the 1-GPa experiments performed by Baker et al. (1994) indicate that  $\sim$ 4.5 wt%  $H_2O$  would be necessary to saturate with oliv and opx on the liquidus.

These H<sub>2</sub>O-undersaturated results indicate that the primitive basaltic andesite and magnesian andesite must contain at least 4.5 wt% H<sub>2</sub>O to be in equilibrium with a mantle harzburgite (oliv + opx) residual assemblage at a depth of 30 km beneath Mt. Shasta. This is the measured depth to the Moho beneath Mt. Shasta (Fuis et al. 1987), and represents a minimum estimate of pre-eruptive H<sub>2</sub>O content necessary to saturate these primitive melts with a mantle residue assemblage. These minimum pre-eruptive H<sub>2</sub>O contents coincide with pre-eruptive H<sub>2</sub>O estimates based on mineral compositional evidence (Baker et al. 1994) and measured H<sub>2</sub>O contents in melt inclusions (Anderson 1979; Sisson and Layne 1993). It is possible for these two compositions to be in equilibrium with of + opx over a range of pressure, temperature, and H<sub>2</sub>O contents. Olivine and opx could be sustained as liquidus phases as pressure increases by an increasing melt H<sub>2</sub>O content and decreasing temperature. This effect can be inferred by comparing the H<sub>2</sub>O-saturated liquidus phase relation for 85-44 (Fig. 2) with the

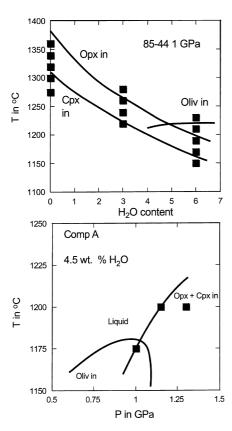


Fig. 4 Results of high-pressure experiments on the Mt. Shasta region primitive lavas at  $H_2O$ -undersaturated conditions. *Upper diagram* shows the effect of variable  $H_2O$  at 1 GPa on the liquidus phase relations of 85-44 (Baker et al. 1994). Approximately 4.5 wt%  $H_2O$  is required to saturate 85-44 with a harzburgitic mantle residual assemblage. *Lower diagram* shows the results of variable pressure on the liquidus phase appearance sequence of 85-41c. At 1 GPa, oliv + opx + cpx are near liquidus phases with 4.5 wt%  $H_2O$ 

 $\rm H_2O$ -undersaturated boundaries (Fig. 4). Under  $\rm H_2O$ -saturated conditions ol + opx should be liquidus phases at 1 GPa and  $\sim$ 1,100 °C (Fig. 2). Melt  $\rm H_2O$  contents under  $\rm H_2O$ -saturated conditions would be  $\sim$ 16 wt% (Hamilton et al. 1964 and 85-44 would be 800 MPa for glass analyses, Table 3). If the ol + opx saturation boundary in Fig. 4 is extended to higher melt  $\rm H_2O$  contents, a similar value of liquid  $\rm H_2O$  content would be inferred.

## **Discussion**

Role of fractional crystallization in producing calc-alkaline differentiation trends

Compositional variations in FeO\*/MgO vs. SiO<sub>2</sub>

Miyashiro (1974) introduced the FeO\*/MgO vs. SiO<sub>2</sub> diagram and used it to discriminate calc-alkaline from tholeiitic differentiation trends and to distinguish among different types of calc-alkaline trends that were found in various arc environments. On the Miyashiro (1974)

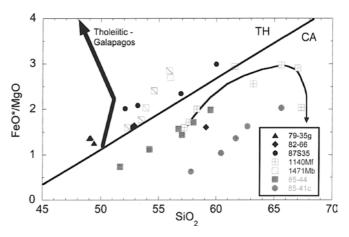


Fig. 5 Experimental liquids projected on the FeO\*/MgO vs. SiO<sub>2</sub> diagram. All compositions are renormalized to an anhydrous basis. The tholeiitic (TH)-calc-alkaline (CA) dividing line is from Miyashiro (1974). The black symbols (79–35 g, 82-66 and 87S35) are liquid lines of descent followed by low-SiO2 high-Al2O3 basalts at 200 MPa and NNO (Sisson and Grove 1993a). These liquid lines of descent define trends similar to the differentiation trends found in many subduction related magmatic suites. Near vertical arrow shows tholeiitic differentiation trend exhibited by Galapagos MORB lavas (Grove and Baker 1984; Juster et al. 1989). Also shown are liquid lines of descent followed by the 200 MPa experiments on the primitive high-SiO<sub>2</sub>, low Al<sub>2</sub>O<sub>3</sub> lavas from the Mt. Shasta region (85-44 and 85-41c, gray symbols). These liquid lines of descent follow calc-alkaline trends defined by many subductionrelated high-magnesian andesite suites. The results of experiments on high-alumina basalt, 1471Mb (Wagner et al. 1995), at 100 MPa and NNO also show more FeO enrichment as a result of lower H<sub>2</sub>O contents. The 100-MPa experiments on andesite 1140Mf (Grove et al. 1997) are included to show the differentiation path for SiO<sub>2</sub>-rich residual liquids at low temperatures and high-SiO<sub>2</sub> contents. An increase in the proportion usp/Fe-Mg silicate results in rapid FeO depletion at the NNO buffer (arrow illustrates trend)

diagram, an anhydrous fractional crystallization trend (tholeiitic trend) is nearly vertical, with FeO\*/MgO increasing continuously at constant SiO<sub>2</sub>. The tholeiitic trend (Fig. 5) is followed during fractional crystallization of basalt at low pressure, anhydrous conditions, and an excellent example is the differentiation trend of Galapagos mid-ocean ridge lavas (Grove and Kinzler 1986; Juster et al. 1989). The strong near vertical increase of FeO\*/MgO at constant SiO2 results from the early appearance of plag as a crystallizing phase, a high proportion of plag/Fe-Mg silicates in the crystallizing assemblage, the late appearance of sp as a crystallizing phase, and the high SiO<sub>2</sub> content of the Na-rich crystallizing plag. Sisson and Grove (1993a) demonstrated that liquids produced by crystallization of hydrous basaltic magmas at modestly oxidizing conditions (Ni-NiO buffer) define crystallization paths similar to those found in many calc-alkaline suites. The presence of  $\sim$ 6 wt% H<sub>2</sub>O in the 200-MPa liquids has a significant effect on the phase appearance sequence and, therefore, on the liquid line of descent followed during crystallization. The influence of H<sub>2</sub>O is to expand the primary phase volumes for oliv and cpx and shrink the plag primary phase volume, causing ferromagnesian minerals to crystallize early and plag to crystallize late. In addition, H<sub>2</sub>O lowers the temperature at which silicate minerals crystallize, but has little effect on the appearance temperature of sp. Thus, an iron-rich Cr–Al sp appears at or near the liquidus of a hydrous basalt melt (Sisson and Grove 1993a). When plag appears as a crystallizing phase, it is Ca- and Al-rich and SiO<sub>2</sub> poor, which promotes enrichment of SiO<sub>2</sub> and alkalis in the residual liquids. Therefore, liquids produced by fractional crystallization become depleted in FeO and enriched in SiO<sub>2</sub>; the calc-alkaline differentiation trend.

The liquids from the 200-MPa H<sub>2</sub>O-saturated crystallization experiments for the primitive basaltic andesite (85-44) and andesite (85-41c) are plotted in Fig. 5 along with the 200-MPa data of Sisson and Grove (1993a) on basaltic compositions. The crystallization paths followed by these primitive basaltic andesite and andesite compositions parallel those of the Sisson and Grove (1993a) experiments, and also follow a calc-alkaline differentiation trend. These differentiation trends begin at a lower value of FeO\*/MgO, and extend along a path that parallels the Sisson and Grove 200 MPa H<sub>2</sub>O-saturated experimental trend. The 85-44 and 85-41c crystallization paths plot in this portion of the FeO\*/MgO vs. SiO<sub>2</sub> diagram because the parent magmas contain lower FeO\*/MgO and higher SiO<sub>2</sub> than the parental melts that were examined by Sisson and Grove (1993a). Otherwise, the slopes of the trends are fairly similar, reflecting similar proportions of minerals in the crystallizing assemblage and similar influences of mineral composition on the crystallization paths.

Also shown in Fig. 5 are 100- and 200-MPa H<sub>2</sub>Osaturated experiments on basalts and basaltic andesites from Medicine Lake volcano (Wagner et al. 1995, 1471Mb; Grove et al. 1997, 1140Mf). The 1471Mb experiments have a slightly steeper slope than the calcalkaline trends of Sisson and Grove (1993a), reflecting the lower pressure (100 MPa) of the 1471Mb experiments. The effect of lower H<sub>2</sub>O-saturated pressure is to cause plag to appear earlier in the crystallization sequence and to increase the proportion of plag relative to Fe-Mg silicates in the crystallizing assemblage. Both effects lead to a steeper slope for the 1471Mb trend. In 1471Mb, the proportions of plag/Fe–Mg silicates is  $\sim 1.5$ whereas the proportion of plag/Fe-Mg silicates in the 200 MPa H<sub>2</sub>O-saturated experiments on 79-35g is 0.6. The shallower slopes of the 200 MPa H<sub>2</sub>O-saturated differentiation trends on the FeO\*/MgO vs. SiO<sub>2</sub> diagram are caused by three factors: the decrease in plag/ Fe–Mg silicate proportions, the earlier appearance of sp and the effect of higher melt H<sub>2</sub>O that lowers the SiO<sub>2</sub> in the crystallizing plag.

Sample 1140Mf is petrologically and geochemically similar to Mt. Shasta primitive basaltic andesites (Kinzler et al. 2000) and can be produced from a magma similar to 85-44 by fractional crystallization of oliv + cpx + plag at 100 to 200 MPa. Since 1140Mf has 56.3 wt% SiO<sub>2</sub> (Table 1), experiments could be extended to lower temperatures allowing the higher SiO<sub>2</sub> portion of the liquid line of descent to be examined. The trend is

illustrated in Fig. 5 with the curved arrow for the 100-MPa experiments. The vertical descent of residual liquids to low FeO\*/MgO coincides with the disappearance of oliv from the crystallization assemblage, and the appearance of opx, along with cpx, plag, and magnetite—ulvöspinel (usp). The abrupt decrease in FeO\*/MgO is a consequence of the increase in the proportion of spinel in the opx + cpx + plag + usp assemblage, relative to the higher temperature oliv + cpx + plag + usp assemblage.

Compositional variations in mineral component space

The bulk compositions of 85-44 and 85-41c and the liquids produced in the 200-MPa  $H_2O$ -saturated experiments are projected into the Oliv (forsterite + fayalite)-Plag (anorthite + albite)-Qtz (SiO<sub>2</sub>) and Oliv-Cpx (diopside + hedenbergite)-Qtz pseudoternaries in Fig. 6 using the projection scheme of Tormey et al.

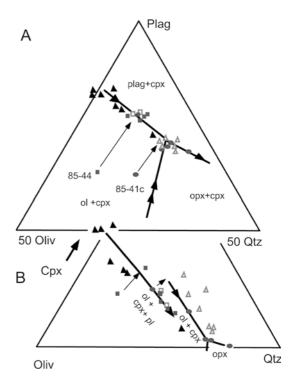


Fig. 6 Compositions of 200-MPa, NNO, H<sub>2</sub>O-saturated liquids on the pseudoternary projections: A Olivine (Oliv)-Plagioclase (Plag)-Quartz (Qtz) projected from Clinopyroxene (Cpx) showing only the upper half of the triangle from 50% Oliv-50% Qtz to 100% Plag. B Oliv-Cpx-Qtz projection scheme from plag. Projection scheme is from Tormey et al. (1987) and Grove (1993). The upper half of the triangle is not shown. Experimental liquids from Sisson and Grove (1993a) are shown as black triangles. The three triangles that plot near the 50 oliv-50 qtz base belong to projection **B**. Experimental liquids from 85-44 are shown as gray squares and 85-41c experiments are plotted as gray ovals. Also shown are experiments of Moore and Carmichael (1998; open squares) and Blatter and Carmichael (1998; open triangles). Dark lines are saturation and reaction boundaries. Arrows show direction of liquid evolution with decreasing temperature. Double arrows indicate a reaction boundary (oliv + opx + cpx in  $\mathbf{A}$ ). The dark lines in  $\mathbf{A}$  join at the oliv + opx + cpx + plag + sp reaction. Lighter arrows in  $\bf A$  and  $\bf B$ indicate crystallization paths followed by the 85-44 and 85-41c from liquidus conditions

Table 4 Mass balance models of mantle melting reactions

1. MBC <sup>a</sup> Mantle Melt <sup>a</sup>	SiO <sub>2</sub> 47.8	TiO <sub>2</sub> 0.66	Al <sub>2</sub> O <sub>3</sub> 19.2	FeO 7.61	MnO 0.0	MgO 11.0	CaO 11.2	Na <sub>2</sub> O 2.42	K <sub>2</sub> O 0.11	Sum 100	Mg# 0.72	FeO/MgO 0.692
G&G98 <sup>b</sup>										2		
	$MM^{a}$	$\pm$	Cpx	$\pm$	Opx	$\pm$	Oliv			$\Sigma R^2$	Opx <sup>c</sup>	Oliv <sup>c</sup>
B392	0.96	_	0.08	+	0.29	_	0.19			1.1	+1	-0.65
B384	0.86	+	0.0	+	0.23	_	0.11			3.1	+1	-0.48
B388	0.84	+	0.0	+	0.27	_	0.12			2.2	+1	-0.44
2. PRM <sup>d</sup>												
%melt	$SiO_2$	$TiO_2$	$Al_2O_3$	FeO	MnO	MgO	CaO	$Na_2O$	$K_2O$	Sum	Mg#	FeO/MgO
0.18	51.5	0.62	16.6	6.83	0.0	12.2	10.1	2.05	0.09	100.	0.76	0.557
0.21	53.7	0.61	15.1	6.33	0.0	12.8	9.48	1.83	0.08	100.	0.78	0.494
0.26	55.6	0.59	13.9	5.90	0.0	13.4	8.96	1.61	0.08	100.	0.80	0.441
3. NLM <sup>e</sup>												
Shasta										2		
85-41c	$MM^{a}$	$\pm$	Cpx	$\pm$	Opx	±	Oliv	±	Fluid	$\Sigma R^2$	Opx <sup>c</sup>	Oliv <sup>c</sup>
	0.51	+	0.0	+	1.1	_	0.61	+	0.02	7.1	+1	-0.55
85-44	$MM^{a}$											
	0.74	+	0.05	+	0.41	_	0.22	+	0.01	3.9	+1	-0.54

<sup>a</sup>1. *MBC* mass balance calculation. Mantle melt (MM) is the average of three liquids (B304, B330, and B329) reported in Gaetani and Grove (1998) that are saturated with the mantle residue assemblage (oliv + opx + cpx + spinel at 1.2 GPa under H<sub>2</sub>O-undersaturated conditions ∼6 wt% H<sub>2</sub>O)

<sup>b</sup>Compositions B392, B384, and B388 are liquids from basalt + peridotite mixtures reported in Gaetani and Grove (1998). B392 is saturated with oliv + opx + cpx. B384 and B388 are saturated with oliv + opx. Results of mass balance calculations are shown for creating these three liquids from the lower temperature 4-phase saturated mantle melt (MM). Negative coefficients indicate that the phase crystallizes, positive coefficients indicate dissolution

<sup>c</sup>Stoichiometric coefficients of the reaction normalized to one unit of liquid in weight units

<sup>d</sup>2. PRM Peridotite reaction melting model calculates liquid compositions of harzburgite melt. Melt compositions are calculated

(1987). Also shown are multiply saturated liquids of high-alumina basalts from Medicine Lake volcano (Sisson and Grove 1993a) and experiments that established high  $\rm H_2O$  contents in Mexican Volcanic Belt lavas (Blatter and Carmichael 1998; Moore and Carmichael 1998). Figure 6a shows the projection through Cpx and is especially useful because most of the 200 MPa liquids are saturated with clinopyroxene. The projection from Plag shown in Fig. 6b is also useful, as it illustrates the silica enrichment exhibited by these lavas during crustal level crystallization under  $\rm H_2O$ -saturated conditions.

The 200-MPa experimentally produced liquids for the 85-44 composition that coexist with oliv + plag + cpx + sp define a multiple saturation boundary in Fig. 6a, b that coincides with the oliv + plag + cpx boundary defined by Sisson and Grove (1993a). The principal difference in primitive basaltic andesite 85-44 is that it has an appreciable interval of oliv +cpx crystallization, before it becomes saturated with plag. The oliv + plag + cpxsaturated liquids obtained by Moore and Carmichael (1998) on a magnesian basaltic andesite from the Mexican volcanic belt (Mas-25, Table 1) are coincident with the 85-44 liquids. For 85-41c, oliv + cpx  $\pm$  sp are the liquidus phases at 200-MPa H<sub>2</sub>O-saturated conditions, followed by the appearance of opx, and then followed by plag saturation (Fig. 6a). In the Oliv-Cpx-Qtz pseudoternary (Fig. 6b), the oliv + cpx liquids from

from the average mantle melt (MM) by removing olivine and adding orthopyroxene in weight proportions 1.0:0.48. Modeling was carried out in 2 wt% increments adding phases that were calculated to be in Fe–Mg exchange with the liquid at each step assuming an oliv–melt  $K_D^{\mbox{Fe-Mg}}(K_D^{\mbox{Fe-Mg}}=[X_{Fe}^{\mbox{Ol}}X_{Mg}^{\mbox{Liq}}]/[X_{Mg}^{\mbox{Cl}}X_{Fe}^{\mbox{Liq}}])=0.29$  and an opx–melt  $K_D^{\mbox{Fe-Mg}}=0.27$ . Numbers below % melt indicate weight fraction of total solid reacted per initial mass of liquid. Final value represents the extreme of the mantle melting/reaction trends plotted in Figs. 7 and 8

<sup>e</sup>3. *NLM* Natural lava melting model is a calculation using experimental liquids and Shasta region lavas. Model mass balance calculation uses MM, cpx, opx, oliv, and the major elements (Na<sub>2</sub>O and K<sub>2</sub>O) from the subduction zone fluid-rich component (fluid) from Table 5of Grove et al. (2002) to make primitive Shasta lavas 85-41c and 85-44

85-41c plot at higher qtz values than those from 85-44. and do not coincide with the oliv + plag + cpx saturation boundary. Olivine reacts with liquid to form opx over a 40 °C temperature interval in 85-41c, oliv is exhausted from the phase assemblage, and the liquid leaves the reaction boundary and follows the opx + cpx + plag-saturation boundary. Experimental studies have also been carried out by Blatter and Carmichael (1998) on a high H<sub>2</sub>O, primitive magnesian andesite from the western Mexican Volcanic Belt (MAS-911, Table 1). This composition shows a similar crystallization sequence (Fig. 6) to the one observed for 85-41c. Note that the oliv + cpx saturated liquids in MAS-911 plot at higher SiO<sub>2</sub> values than the 85-41c liquids saturated with a similar phase assemblage. This is most likely a consequence of the higher alkali content of MAS-911 (in MAS-911 Na<sub>2</sub>O +  $K_2O = 5.5$  wt% vs. 3.8 wt% in 85-41c), which has the effect of expanding the olivine primary phase volume (Grove and Juster 1989; Longhi 1991).

Role of hydrous mantle melting in producing calc-alkaline differentiation trends

Both Shasta region lavas are very primitive despite their high SiO<sub>2</sub> contents. Composition A is multiply saturated

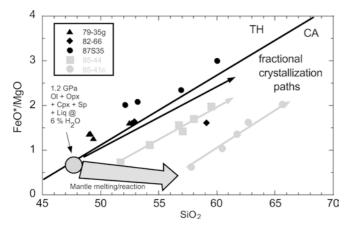


Fig. 7 Experimental liquids projected on the FeO\*/MgO vs. SiO<sub>2</sub> diagram. The tholeiitic (TH)-calc-alkaline (CA) dividing line is from Miyashiro (1974). Symbols are the same as those used in Fig. 5. Lines with arrows illustrate the liquid lines of descent followed by primitive melts as they experience crystal fractionation under hydrous conditions at 200 MPa and the NNO buffer. The large circle plotting on the TH-CA dividing line marks the locus of primitive mantle hydrous melts derived at 1.2 GPa (Gaetani and Grove 1998) and coexisting with oliv + opx + cpx + sp. These liquids follow a differentiation trend similar to that defined by the Sisson and Grove (1993a) liquids (black arrow) at 200 MPa H<sub>2</sub>Osaturated. The gray arrow that trends to high-SiO<sub>2</sub> and low FeO\* MgO is the path followed during mantle melting at 1.2-GPa and 6 wt% H<sub>2</sub>O (Table 4) after cpx and sp are exhausted from the mantle residue. This melting path lies on the oliv + opx reaction boundary

on its liquidus with oliv + opx at 1 GPa with 4.5 wt% H<sub>2</sub>O added. The inferred pre-eruptive H<sub>2</sub>O content for 85-41c (Anderson 1979) is 6.5 wt% H<sub>2</sub>O, so the multiple saturation depth and temperature may be higher than that given by the 4.5 wt% H<sub>2</sub>O experiments. Baker et al. (1994) showed that 85-44 is saturated with oliv + opx on its liquidus at  $\sim 4.5$  wt%  $H_2O$  and 1 GPa. This magmatic H<sub>2</sub>O content also corresponds with the preeruptive H<sub>2</sub>O content estimated for 85-44 (Anderson 1979; Baker et al. 1994). The pressure of 1.0 GPa is equivalent to the depth to the top of the mantle wedge beneath Mt Shasta. Both of the magmas are in equilibrium with a mantle residue assemblage representative of high extents of melting under hydrous (H<sub>2</sub>O-undersaturated) conditions (Fig. 4). The basaltic andesite (85-44) contains oliv phenocrysts of Fo<sub>89,6</sub>. Sample 85-41c contains Fo<sub>93.6</sub> olivine. Thus, 85-41c represents a higher extent of mantle melting than 85-44.

Melting experiments on mantle peridotite under  $H_2O$ -undersaturated conditions have been carried out by Kushiro (1990), Hirose and Kawamoto (1995), Hirose (1997), and Gaetani and Grove (1998). Gaetani and Grove (1998) provide the most complete characterization of the mantle-melting reaction when the residue contains oliv + opx + cpx + sp from 1.2 to 1.6-GPa. They present experimental data on 60% peridotite: 40% 85-44 mixture that can be used to calculate the mantle melting reaction extended beyond the point of exhaustion of sp and cpx. The results of the Gaetani and Grove (1998) experiments at 1.2-GPa have been used to

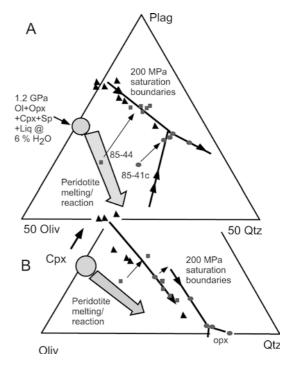


Fig. 8 Compositions of 200-MPa, NNO,  $H_2O$ -saturated liquids on the pseudoternary projections: A Olivine (Oliv)-Plagioclase (Plag)-Quartz (Qtz) projected from Clinopyroxene (Cpx) showing only the upper half of the triangle from 50% Oliv-50% Qtz to 100% Plag. B Oliv-Cpx-Qtz projection scheme from Plag. Projection scheme is from Tormey et al. (1987) and Grove (1993). Symbols and saturation boundaries are the same as in Fig. 6. Thin arrows pointing away from 85-44 and 85-41c starting compositions indicate fractionation paths. Also shown are the 1.2-GPa, 6 wt%  $H_2O$  liquids saturated with ol + opx + cpx + sp from Gaetani and Grove (1998). The oliv + opx melting reaction followed during hydrous mantle melting after cpx + sp are exhausted is shown by the dark gray arrow (Table 4) and is defined by crystallization of oliv and melting of opx

calculate this melting reaction ( Table 4). The first set of results in Table 4 are least squares mass balance calculations that relate an average of the experimental liquids (MM in Table 4) saturated with oliv + opx + cpx + sp at 1.2-GPa and containing  $\sim$ 6 wt% H<sub>2</sub>O. This average liquid is used in the mass balance to calculate the phase assemblage necessary to relate it to liquids saturated with ol + opx + cpx (B392 in Table 4), or ol + opx (B384 and B388 in Table 4). The calculated coefficients are given in Table 4 (1. MBC).

These coefficients are consistent with the melting reaction:

$$1.0 \text{ Opx} = 0.5 \text{ Oliv} + 0.5 \text{ Liq} \tag{1}$$

In this melting reaction, opx dissolves as oliv + melt are produced. This is the well-known incongruent melting of opx to oliv + liquid that operates at low pressure and/or high magmatic H<sub>2</sub>O contents. This reaction leads to significant silica enrichment in the high-extent mantle melts produced after cpx and sp are exhausted as residual phases. The path followed by liquids as melting proceeds has been calculated (Table 4, 2.

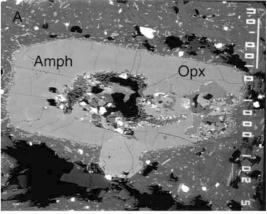
PRM) using the Gaetani and Grove multiply saturated melt (MM in Table 4), and the phase proportions from the mass balance from Table 4. The path is plotted in Figs. 7 and 8.

An additional mass balance calculation (Table 4, 3. NLM) was performed using the compositions of 85-44 and 85-41c, the Gaetani and Grove mantle multiply saturated (oliv + opx + cpx + sp) melts and the major element composition (Na<sub>2</sub>O + K<sub>2</sub>O) of the slab-derived fluid-rich component from Grove et al. (2002) to calculate the melting reaction. The results are shown in Table 4 and require similar proportions of oliv and opx. The major elements (Na<sub>2</sub>O + K<sub>2</sub>O) in the fluid-rich component that were used in this calculation were assumed to be added at the bottom of the wedge during vapor-saturated melting of peridotite, and they have been significantly diluted during flux melting (see Grove et al. 2002; Gaetani and Grove 2003).

The three different approaches discussed above all point to reaction [1], the incongruent melting of opx on the oliv + opx reaction boundary, as the process responsible for moving primitive subduction zone magmas to the low FeO\*/MgO and high SiO<sub>2</sub> portion of the Miyashiro diagram. This reaction operates once cpx + sp are exhausted as mantle residue phases and harzburgite becomes the mantle source rock. In the pseudoternary projections schemes, the influence of opx melting and oliv crystallization is to increase SiO<sub>2</sub> and decrease Al<sub>2</sub>O<sub>3</sub>, driving the mantle melts first to basaltic andesites like 85-44, and, with further melting, to hypersthene-normative compositions like that of 85-41c (Fig. 8).

An alternative mechanism for producing these SiO<sub>2</sub>rich liquids in the mantle wedge is melt-wall-rock interaction. Kelemen (1990) and Kelemen et al. (1997) propose that melts rising through the mantle will become saturated with oliv and undersaturated in opx as the oliv + opx boundary shifts to higher SiO<sub>2</sub> contents with decreasing pressure. This mode of reaction is identical to reaction [1] and it now results in assimilation of opx with precipitation of oliv, raising the SiO<sub>2</sub> content of a magma without significantly affecting the other major and trace elements. Because the assimilation process is controlled by the same saturation boundaries and crystal-melt reactions that govern melting, it is difficult to distinguish assimilation from a progressive melting and re-equilibration process that takes place at low pressures. However, both processes are restricted to shallow mantle depths and high H<sub>2</sub>O contents.

Miyashiro (1974) recognized that there is not a single calc-alkaline or tholeiitic suite, but that a continuum exists from tholeiitic to calc-alkaline, and that each arc has its own distinctive signature. In addition, many arc volcanoes possess both tholeiitic and calc-alkaline suites (e.g., Hakone region, Japan and the Nasu volcanic zone, northeast Japan, Cascades, USA). Miyashiro (1974) also found that many arcs (particularly notable is the largest arc on Earth, the Mariana arc) display FeO\*/MgO vs. SiO<sub>2</sub> variations that lie on the tholeiitic side of the plot. The trends defined by Sisson and Grove (1993a),



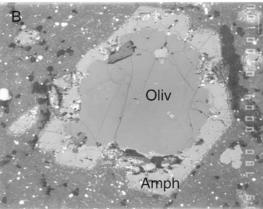


Fig. 9 Amphibole and associated overgrowth minerals from Mt. Shasta andesite 82-85 (Table 5). Backscattered electron images with *scale bars* of 100  $\mu$ m (*upper right*). A Euhedral amphibole exhibiting internal skeletal growth form. On the right side of the crystal is an inclusion of high Mg# opx (Mg# = 0.87). B Rounded and amoeboid shaped olivine crystal (Fo<sub>88</sub>) surrounded by high Mg# amphibole

therefore, are characteristic of a large volume of the Earth's current output of arc magmatism, and represent fractional crystallization trends produced by high magmatic  $H_2O$  content melts under crustal level conditions. Based on the reasoning presented above, these hydrous fractional crystallization trends are the ones expected to result when the starting point for fractionation is a  $H_2O$ -rich mantle melt produced at high pressures and saturated with oliv + opx + cpx + sp or oliv + opx + cpx + garnet. The mantle melt compositions that define the starting point for fractionation from these liquids are represented by the large gray dot in Figs. 7 and 8, and the calc-alkaline fractional crystallization paths followed by these liquids coincide with the TH-CA boundary.

The calc-alkaline differentiation trends delineated by the Shasta region primitive lavas lie in a portion of the diagram that is not as commonly represented by modern arc environments. Similar lavas are found in other hotslab environments, where young oceanic lithosphere is being subducted. Localities include Shasta, the Adak region of Alaska (Kay 1978; Yogodzinski et al. 1995), the Setouchi belt of SW Japan arc (Ishizaka and Carlson 1983; Tatsumi and Ishizaka 1983; Shimoda et al. 1998),

Table 5 Compositions of minerals in Mt. Shasta magnesian andesites

Sample no.	Mineral	$SiO_2$	$TiO_2$	$Al_2O_3$	Cr <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	$K_2O$	$P_2O_5$	Sum	Mg# or An
82-85														
inc.a	Oliv	40.3		0.11		12.8	0.24	48.3	0.22				102.0	0.870
og <sup>a</sup>	Amph	45.4	1.71	10.7	0.25	6.75		18.9	11.64	2.64	0.67		98.6	0.833
inc	Cpx	52.7	0.39	2.15	0.10	6.01	0.24	16.9	20.9	0.31			99.8	0.834
og	Amph	44.8	1.92	11.1	0.46	7.42	0.03	18.1	11.3	2.55	0.62		98.4	0.813
og ig <sup>a</sup>	Cpx	52.2	0.47	3.42	0.31	5.91	0.17	16.8	21.2	0.56			101.1	0.835
ig 82-88a	Plag	56.6		27.5		0.25		0.02	9.84	6.13	0.24		100.5	0.463
	011	20.5		0.02		150	0.05	44.0			0.22		1000	0.01.5
inc	Oliv	39.7		0.03		17.8	0.25	44.2			0.33		102.0	0.815
og 85-59	Amph	45.1	1.47	11.5		9.55	0.21	17.7	11.4	2.51			99.9	0.768
inc	Oliv	40.1		0.03	0.04	16.7	0.21	44.0	0.11				101.7	0.823
inc	Opx	55.9	0.13	1.19	0.03	12.0	0.30	30.4	1.32	0.02			101.3	0.818
og	Amph	43.8	2.10	11.8	0.02	8.40	0.10	17.0	11.6	2.37	0.53		97.7	0.782
pheno	Plag	56.4		27.6		0.27			10.5	5.43	0.19		100.4	
ig	Usp-mag	0.12	6.05	1.73	0.30	83.8	0.38	2.24					94.7	
ig ig 83-56	Ilm-hem	0.07	30.5	0.33	0.14	62.8	0.38	3.31					97.6	
inc	Opx	55.2	0.14	2.09	0.14	11.3	0.24	31.3	1.13				101.5	0.813
og 96-5	Amph	46.6	1.55	10.5	0.04	7.88	0.12	17.7	11.3	2.12	0.24		98.1	0.800
inc	Opx	55.3	0.17	1.34	0.45	9.58	0.23	30.3	1.13	0.04			98.5	0.849
inc	Срх	53.2	0.19	1.79	0.21	5.62	0.22	16.2	21.5	0.32			99.2	0.837
	Amph	46.6	1.08	9.44	0.41	7.80	0.11	17.6	11.2	2.95	0.18		97.3	0.800
og 96-11b	•										0.10			
inc	Cpx	53.0	0.17	1.74	0.75	4.16	0.14	16.7	22.0	0.29			99.0	0.878
og 97-1f	Amph	47.0	1.05	9.10	0.38	7.18	0.08	17.4	11.4	2.51	0.20		96.3	0.812
inc	Oliv	39.8		0.01	0.02	13.8	0.16	45.2	0.11				99.1	0.854
inc	Cpx	53.1	0.30	2.04	0.70	4.64	0.07	16.9	21.1	0.38			99.3	0.866
og	Amph	45.5	2.12	10.7	0.09	7.90	0.14	17.4	11.3	2.48	0.52		98.2	0.797
Melt inclusion	on 246-4-1 ii	n opx, S	76 of (A	nderson	1979)									$H_2O^b$
		61.1	0.32	16.8	0.09	0.71	0.15	0.38	3.42	3.02	1.65	0.21	99.82	10.7
Lavas <sup>c</sup>														Mg#
82-85	SR	61.46	0.66	16.99		4.58		3.83	5.87	4.40	1.82	0.30		0.598
82-88a	SR	63.53	0.59	16.89		4.02		2.81	5.16	4.69	1.95	0.28		0.555
85-59	SR	60.29	0.73	16.78		4.86		4.34	6.14	4.60	1.84	0.34		0.614
83-56	M	65.62	0.44	16.64		2.94		2.16	4.86	4.60	1.36	0.12		0.567
96-5	SH	64.36	0.49	16.89		3.47		3.47	6.07	4.07	0.99	0.13		0.641
96-11b	SH	64.90	0.51	16.99		3.41		3.02	5.96	3.90	1.08	0.15		0.612
97-1f	SR	57.64	0.92	16.84		5.57		5.11	7.31	4.23	1.81	0.48		0.621
S76	SH	63.62	0.60	16.71		3.86		3.55	6.08	4.16	1.15	0.21		0.621

<sup>a</sup>inc. Inclusion in mineral; og overgrowing mineral; ig intergrowth <sup>b</sup>Melt inclusion in opx (Mg#0.84-0.89, Anderson 1979) in Shastina andesite S76. Resampled as 97-7 by TL Grove;  $H_2O$  estimated from quantitative analysis of oxygen by electron microprobe (see Gaetani and Grove 1998)

<sup>c</sup>Mt. Shasta region lavas. *SR* Sargents Ridge; *M* Misery; *SH* Shastina. These are eruptive stages of the Mt. Shasta stratocone. See Christiansen et al. (1977)

the Mexican Volcanic Belt (Blatter and Carmichael 2001), and the Chilean arc (Stern and Kilian 1996). These lavas plot in the low FeO\*/MgO and high SiO<sub>2</sub> portion of Fig. 5 because they represent high degree melts of refractory mantle with only oliv + opx as residual phases (Fig. 4). The increased extent of hydrous mantle melting at shallow mantle depths has produced SiO<sub>2</sub>-rich primitive parent melts with low FeO at elevated MgO contents (Figs. 7 and 8).

Evidence for high water contents in primitive Mt. Shasta region magmas

The petrologic evidence preserved in the primitive Mt. Shasta region lavas is suggestive of pre-eruptive  $H_2O$ 

contents of 4.5 to 6.5 wt%. These values are significantly higher than those inferred by early experimental studies of the influence of  $H_2O$  on mantle melting (e.g., Kushiro 1968, 1969, 1972). These studies argued that thermodynamic effects of  $H_2O$  on melt composition produced the high  $SiO_2$  by addition of only 2 to 3 wt%  $H_2O$  to the melt.

There is petrologic evidence in Mt. Shasta lavas suggestive of even higher pre-eruptive H<sub>2</sub>O contents. The evidence is found in several high-Mg# andesite and dacite lavas of the Mt. Shasta stratocone. These lavas preserve petrologic signatures that indicate complex mixing processes between primitive and evolved melt components. These mixed lavas may record a fractionation event that occurred in a primitive magnesian

andesite under high pressure (~800 MPa) and high H<sub>2</sub>O contents (>10 wt%). In some lavas, a magnesian pargasitic amph is preserved in overgrowth relation with magnesian oliv, opx, and cpx (Fig. 9, Table 5). This magnesian pargasite is similar in composition to the amph produced in the 800-MPa, H<sub>2</sub>O-saturated experiments on 85-44 (Table 3). As  $P_{\rm H2O}$  increases along an oxygen buffer curve (NNO in our experiments), amph appears earlier in the crystallization assemblage until, at 800 MPa, it appears after oliv along with opx and cpx at 1,070 °C. In contrast, amph is the last phase to crystallize in the 200-MPa experiments, appearing by 980 °C. This change in the appearance temperature of amph manifests itself in a systematic variation in Mg# of the amph. This variation is illustrated in Fig. 10, along with experimental data of Blatter and Carmichael (1998) and Moore and Carmichael (1998), who worked on compositions from the Mexican Volcanic Belt at pressures up to 300 MPa and at higher  $f_{\rm O2}$  conditions. The Blatter and Carmichael (1998) experiments were performed at NNO+3 over the pressure range of 150 to 300 MPa, and the Moore and Carmichael (1998) experiments were performed at  $f_{O2}$  values between NNO + 1 and NNO + 2. The experiments on the Mexican Volcanic Belt lavas show the important additional influence of variations in  $f_{O2}$  on amph appearance. Increasing  $f_{O2}$  increases the appearance temperature of amph by decreasing the amount of  $Fe^{2+}$  in the melt, thereby increasing liquid Mg#.

A line has been extended between the two 85-44 experimental data at 200 and 800 MPa (Fig. 10). The line approximates the effect of increasing H<sub>2</sub>O content at vapor-saturated conditions along the NNO buffer curve on the composition of the first-appearing amph. Also shown are amph compositions from the Shasta lavas. Amphibole is often found in Mt. Shasta andesites (Baker 1988), but it most commonly occurs intergrown with plag. When associated with plag the amph has an Mg# of 0.75 or lower, and the amph crystal overgrows and encloses the plag. The coincidence of Mg# < 0.75 and included plag is consistent with crystallization of andesite liquid at 200 MPa at NNO and H<sub>2</sub>O contents of ~6 wt%. The magnesian pargasite assemblages tabulated in Table 5 do not occur with plag, but exist in an overgrowth relation with magnesian oliv, opx, and/or cpx. This association is consistent with the phase relations (Fig. 2) that show amph appearing as a crystallizing phase with oliv, opx, and cpx before plag at pressures > 400 MPa. The magnesian pargasite assemblages in Shasta andesites are positioned on Fig. 10 according to their Mg#. If this Mg# is used to indicate  $P_{H2O}$  of crystallization, values of 500 to >800 MPa are obtained. The magnesian pargasite amph from 96-5 and 97-1f resemble the amph produced at 800 MPa H<sub>2</sub>Osaturated at NNO and one would infer H<sub>2</sub>O contents of  $\sim$ 14 wt% (Hamilton et al. 1964) to stabilize this assemblage at NNO. The highest Mg# amph found in the Shasta lavas (82-85) would require H<sub>2</sub>O-saturated crystallization at pressures of > 1 GPa, indicating H<sub>2</sub>O

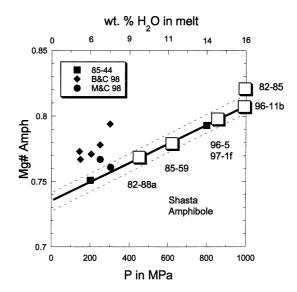


Fig. 10 Mg# of first amphibole to appear in H<sub>2</sub>O-saturated experiments on 85-44 at the NNO buffer at 200 and 800 MPa is plotted against pressure (solid squares). Dark line was placed through these two amphibole compositions and assumes a linear relation between pressure and Mg# of amphibole, as amphibole appears as an earlier phase in the crystallizing assemblage (Fig. 2). Upper scale shows H<sub>2</sub>O contents of H<sub>2</sub>O-saturated liquids using the method of Hamilton et al. (1964). Also shown are experiments performed by Blatter and Carmichael (1998) on a high-SiO<sub>2</sub> andesite at NNO+3 and from Moore and Carmichael (1998) at NNO+1 and NNO+2. Crystallization conditions for Shasta lavas is inferred to be NNO+1.8 or lower. Large open squares are compositions of high Mg# amphiboles that contain included opx, cpx, or oliv (Table 5). Plot has not been extended to pressures greater than 1 GPa. The high Mg# of 82-85 would indicate higher pressure, but we are not confident of this extrapolation. Dashed boundary on either side of the Mg#-P line is the  $2\sigma$  uncertainty of amphibole composition from replicate analyses reported in Table 3

contents of > 16 wt%, if the simplified extrapolation of Mg# vs. pressure applies.

Additional support for high pre-eruptive  $H_2O$  contents comes from the  $H_2O$  content of melt inclusions in Mt. Shasta primitive lavas. Anderson (1979) estimated  $\sim 13$  wt%  $H_2O$  in melt inclusions in opx (Mg#=0.84–0.89) in Shastina lavas. We reanalyzed these melt inclusions quantitatively for oxygen and obtained  $H_2O$  contents of 10.5 to 10.7 wt% (Table 5, 246-4-1).

Oxygen fugacity variations are an additional important variable in controlling amphibole composition. Also present in the andesites are coexisting usp and ilmenite (Table 5, 85-59). The QUILF barometer of Andersen and Lindsley (1988) was applied to this assemblage to estimate temperature and  $f_{\rm O2}$ . An  $f_{\rm O2}$ value of NNO+1.8 at 890 °C is obtained from the coexisting oxides, and this assemblage crystallized in the evolved rhyodacite liquid that constitutes one mixing end member in the Shasta andesite. If this value of  $f_{O2}$  was also relevant for the crystallization of the magnesian end member in the mixed lava, it would lower the pressure of crystallization inferred from the NNO experimental data by  $\sim 150$  MPa, and the maximum pre-eruptive H<sub>2</sub>O content would be  $\sim$ 14 wt%.

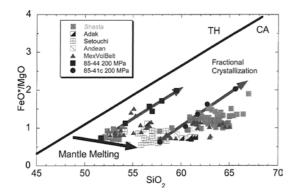


Fig. 11 Compositions of selected suites of high-Mg, high-SiO<sub>2</sub> lavas plotted on the FeO\*/MgO vs. SiO<sub>2</sub> diagram. The tholeiitic (TH)calc-alkaline (CA) dividing line is from Miyashiro (1974). Symbols (black) and lines (dark gray) for 85-44 and 85-41 show crystallization paths followed under 200 MPa H<sub>2</sub>O-saturated conditions at NNO. Solid black line labeled mantle melting is the path followed by mantle melts generated by melting of harzburgite. The left-handside of the line begins where cpx + sp are exhausted from the mantle residue at ~20% melting at 1.2 GPa (Gaetani and Grove 1998). At the right-hand arrow, an additional 20% melting of oliv + opx has occurred. Also shown are lavas from the Mt. Shasta region (Baker 1988; Baker et al. 1994; Grove et al. 2002), the Adak Island region of Alaska (Kay 1978; Yogodzinski et al. 1995), the Setouchi volcanic belt (Tatsumi and Ishizaka 1982; Shimoda et al. 1998), the Andean Austral Volcanic Zone (Stern and Killian 1996) and the Mexican Volcanic Belt (Luhr and Carmichael 1985; Carmichael et al. 1996; Blatter and Carmichael 2001)

With the realization that amph composition may be a complex indicator of pre-eruptive H<sub>2</sub>O content and  $f_{\rm O2}$  variations, we present these results with caution. However, the amphibole and melt inclusion data are consistent with the existence of Mt. Shasta lavas that contain pre-eruptive H<sub>2</sub>O contents > 10 wt% H<sub>2</sub>O. The magnesian amph assemblages are present in ~10% of the andesite and dacite lavas that have been sampled at Mt. Shasta. They are represented in three (Sargents, Misery and Shastina) of the four Shasta eruptive units described by Christiansen et al. (1977). High-Mg amph are present in greatest abundance in the Shastina lavas, where nearly every dacite sampled contains the assemblage magnesian amph enclosing opx, cpx, and/or oliv. The opx containing melt inclusions with 10.7 wt% H<sub>2</sub>O is also from a Shastina dacite (S76, Anderson 1979).

Role of fractional crystallization in the generation of Adakites, Setouchi andesites, and Archean TTG suites.

A number of subduction zone lavas exhibit trace element geochemical signatures that are consistent with a significant trace element contribution from the subducted oceanic lithosphere. These include the Adak-type high-Mg andesites from the Aleutians (Kay 1978; Yogodzinski et al. 1995), the Setouchi high-Mg andesites from SW Japan (Tatsumi and Ishizaka 1982; Ishizaka and Carlson 1983; Shimoda et al. 1998), the andesites of the Mexican Volcanic Belt (Luhr and Carmichael 1985;

Carmichael et al. 1996), and high-Mg andesites from Cook Island in the Andean Austral volcanic zone (Stern and Kilian 1996). These lavas are shown in Fig. 11 on the Miyashiro (1974) FeO\*/MgO vs. SiO<sub>2</sub> plot along with the compositions of Mt. Shasta primitive lavas and stratocone andesites and dacites. There are striking similarities between the Mt. Shasta region andesites and these other high-SiO<sub>2</sub>, high-MgO lavas. The primitive end members of the Setouchi lavas are quite similar to both the primitive basaltic andesite and magnesian andesites from the Mt. Shasta region, and high-pressure experiments by Tatsumi (1982) and Tatsumi et al. (1983) indicate a mantle origin and high H<sub>2</sub>O contents under conditions similar to those inferred for the Shasta lavas. Petrologic studies (Tatsumi and Ishizaka 1982) show mineral assemblages that are consistent with equilibration with a refractory mantle residue assemblage. High Mg# oliv (Stern and Kilian 1996) and cpx (Yogodzinski and Kelemen 1998) are also found in the Adak high-Mg lavas. Petrologic and experimental studies of the Mexican Volcanic Belt lavas (Blatter and Carmichael 1998; Moore and Carmichael 1998) indicate pre-eruptive H<sub>2</sub>O contents of at least 6 wt% and Luhr and Carmichael (1985) identify primitive parent magmas.

In these lavas there are strong trace element and isotopic contributions inherited from subducted lithosphere (basalt and sediment). Adak andesite lavas show strong MORB trace element and isotopic signatures and Setouchi andesite lavas contain a strong sediment signature. The Shasta region primitive lavas preserve both sediment-derived and MORB-derived trace element and isotopic signatures (Grove et al. 2002). However, the additional inference is often made that the major element characteristics of these lavas are produced by direct melting of the subducted slab (Defant and Drummond 1990). This reasoning has been extended to Archean times where SiO<sub>2</sub>-rich plutonic rocks, the trondhjemite-tonalite-granodiorite suite (TTG suite) are found to possess similar major and trace element characteristics (Martin 1987; Smithies 2000).

We propose an alternative to a single-stage slabmelting model that attempts to explain both major and trace element characteristics of these lavas. The major element characteristics of these magmas are inherited from a flux melting process that began at the base of the mantle wedge, and extended through the wedge. The result is a high-SiO<sub>2</sub>, high-MgO, high-H<sub>2</sub>O melt saturated with harzburgite at shallow mantle depth. Subsequent fractional crystallization under hydrous conditions produced SiO<sub>2</sub>-rich derivative liquids that follow calc-alkaline differentiation trends. This melting was fluxed by an H<sub>2</sub>O-rich and trace element-rich fluid component released by dehydration of the subducting ocean lithosphere (Grove et al. 2002). The trace element inventory is contributed by the H<sub>2</sub>O-rich component. The major element composition is determined by the temperature—depth structure in the wedge. The mantle melting process is the incongruent melting of opx to oliv + melt. This process can only occur at shallow depths. Thus, the high- $SiO_2$  signature of these lavas indicates shallow mantle melting under hydrous conditions.

#### **Conclusions**

This paper has discussed the influence of H<sub>2</sub>O in controlling the mantle melting and fractional crystallization of primitive, high-SiO<sub>2</sub>, subduction-related lavas. Preeruptive H<sub>2</sub>O contents of 4.5 to 6 wt% are present in these lavas, but some lavas may contain higher magmatic H<sub>2</sub>O as high as 10 wt% H<sub>2</sub>O. These high H<sub>2</sub>O contents, combined with high extents of shallow mantle melt reaction, may be the critical ingredients that lead to the production of the rare, but important low-FeO\*/MgO, high-SiO<sub>2</sub> primitive arc magmas. Thus, primitive high-SiO<sub>2</sub> lavas from young hot-slab subduction zones are characterized by high extents of mantle melting at high magmatic H<sub>2</sub>O contents.

Acknowledgements The authors thank J. Gill and J. Longhi for thoughtful reviews. This research was supported by National Science Foundation Grants EAR-9706214, EAR-0073766, and OCE-00001821. T.L.G. thanks A.N. Halliday and ETH, Zurich, for providing support as a Gastdozent during the preparation of this manuscript.

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