

PUBLISHED VERSION

G. Aad ... P. Jackson ... L. Lee ... A. Petridis ... N. Soni ... M.J. White ... et al. (ATLAS Collaboration)
Search for new phenomena in events with a photon and missing transverse momentum in pp collisions at $\sqrt{s} = 8\text{TeV}$ with the ATLAS detector
Physical Review D - Particles, Fields, Gravitation, and Cosmology, 2015; 91(1):012008-1-012008-25

© 2015 CERN, for the ATLAS Collaboration. Published by the American Physical Society under the terms of the Creative Commons Attribution 3.0 License. Further distribution of this work must maintain attribution to the author(s) and the published articles title, journal citation, and DOI.

Originally published at:

<http://doi.org/10.1103/PhysRevD.91.012008>

PERMISSIONS

<http://creativecommons.org/licenses/by/3.0/>



Attribution 3.0 Unported (CC BY 3.0)

This is a human-readable summary of (and not a substitute for) the [license](#).

[Disclaimer](#)



You are free to:

Share — copy and redistribute the material in any medium or format

Adapt — remix, transform, and build upon the material

for any purpose, even commercially.

The licensor cannot revoke these freedoms as long as you follow the license terms.

Under the following terms:



Attribution — You must give **appropriate credit**, provide a link to the license, and **indicate if changes were made**. You may do so in any reasonable manner, but not in any way that suggests the licensor endorses you or your use.

No additional restrictions — You may not apply legal terms or **technological measures** that legally restrict others from doing anything the license permits.

6 December 2016

<http://hdl.handle.net/2440/102043>

Search for new phenomena in events with a photon and missing transverse momentum in pp collisions at $\sqrt{s} = 8$ TeV with the ATLAS detector

G. Aad *et al.**

(ATLAS Collaboration)

(Received 7 November 2014; published 27 January 2015)

Results of a search for new phenomena in events with an energetic photon and large missing transverse momentum with the ATLAS experiment at the LHC are reported. Data were collected in proton-proton collisions at a center-of-mass energy of 8 TeV and correspond to an integrated luminosity of 20.3 fb⁻¹. The observed data are well described by the expected Standard Model backgrounds. The expected (observed) upper limit on the fiducial cross section for the production of events with a photon and large missing transverse momentum is 6.1 (5.3) fb at 95% confidence level. Exclusion limits are presented on models of new phenomena with large extra spatial dimensions, supersymmetric quarks, and direct pair production of dark-matter candidates.

DOI: 10.1103/PhysRevD.91.012008

PACS numbers: 13.85.Rm, 13.85.Qk, 14.70.Kv, 14.80.Rt

I. INTRODUCTION

Events that contain a high-momentum photon and large missing transverse momentum (referred to as $\gamma + E_T^{\text{miss}}$) constitute a low-background sample that provides powerful sensitivity to some models of new phenomena [1–7]. Theories with large extra spatial dimensions (LED), presence of dark matter (DM), or supersymmetric partners of the quarks (squarks) in a compressed mass spectrum scenario predict the production of $\gamma + E_T^{\text{miss}}$ events in pp collisions beyond Standard Model (SM) expectations.

The model of LED proposed by Arkani-Hamed, Dimopoulos, and Dvali [8] (ADD) aims to solve the hierarchy problem by hypothesizing the existence of n additional spatial dimensions of size R , leading to a new fundamental scale M_D related to the Planck mass, M_{Planck} , through $M_{\text{Planck}}^2 \approx M_D^{2+n} R^n$. If these dimensions are compactified, a series of massive graviton modes results. These gravitons may be invisible to the ATLAS detector, but if the graviton is produced in association with a photon, the detector signature is a $\gamma + E_T^{\text{miss}}$ event, as illustrated in Fig. 1.

Although the presence of DM is well established [9], its possible particle nature remains a mystery. A popular candidate is a weakly interacting massive particle (WIMP) denoted χ , which has an interaction strength with SM particles at the level of the weak interaction. If the WIMPs interact with quarks via a heavy mediator, they could be pair produced in collider events. The $\chi\bar{\chi}$ pair would be invisible, but $\gamma + E_T^{\text{miss}}$ events can be produced via radiation of an initial-state photon in $q\bar{q}\chi\bar{\chi}$ interactions [10].

As observations so far do not provide strong constraints on the nature of the WIMPs and the theoretical framework to which they belong, it is particularly interesting to study model-independent effective field theories (EFT) with various forms of interaction between the WIMPs and the Standard Model particles [10]. In this framework, the mediator is effectively integrated out from the propagator and the production mechanism at the LHC energy scale is considered as a contact interaction, as illustrated in Fig. 2. Several EFT operators for which the WIMP is a Dirac fermion are used as a representative set following the nomenclature of Ref. [10]: D5 (vector), D8 (axial vector), and D9 (tensor). The interactions of SM and DM particles are described by two parameters: the DM particle mass m_χ and the suppression scale (M_*) of the heavy mediator. In an ultraviolet complete theory, the contact interaction would be replaced by an interaction via an explicit mediator V ; the suppression scale is linked to the mediator mass m_V by the relation $M_* = m_V / \sqrt{g_f g_\chi}$, where g_f and g_χ represent the coupling factors of the mediator to SM particles and WIMPs, respectively. However, as the typical momentum transfer in LHC collisions can reach the scale of the microscopic interaction, it is also crucial to probe specific models that involve the explicit production of the intermediate state, as shown in Fig. 3. In this case, the interaction is effectively described by four parameters: m_χ , m_V , the width of the mediator Γ , and the overall

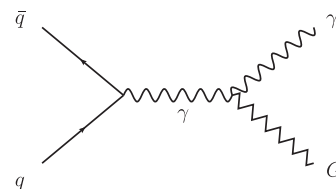


FIG. 1. Graviton (G) production in models of large extra dimensions.

* Full author list given at the end of the article.

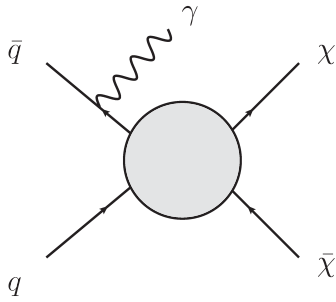


FIG. 2. Production of pairs of dark-matter particles ($\chi\bar{\chi}$) via an effective four-fermion $q\bar{q}\chi\bar{\chi}$ vertex.

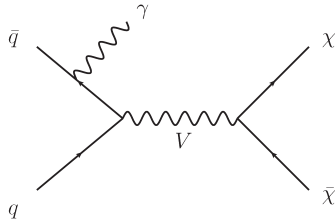


FIG. 3. Production of pairs of dark-matter particles ($\chi\bar{\chi}$) via an explicit s -channel mediator, V .

coupling $\sqrt{g_f g_\chi}$. In this paper, both the EFT approach presented in Ref. [10] and a specific model with a Z' -like mediator [11] are considered.

An alternative DM model hypothesizes interactions between the WIMPs and SM gauge bosons [12]. The effective coupling to different bosons is parametrized by the coupling strengths k_1 and k_2 , which control the strength of the coupling to the U(1) and SU(2) gauge sectors of the SM, respectively. In this model, dark-matter production proceeds via $pp \rightarrow \gamma + X \rightarrow \gamma\chi\bar{\chi} + X'$, requiring no initial-state radiation, as shown in Fig. 4. This model can also be used to describe the peak observed in the Fermi-LAT data [13], allowing a direct comparison of Fermi and ATLAS data in the same parameter space.

Supersymmetry [14–22] postulates the existence of a new supersymmetric partner for each SM particle, differing by half a unit of spin from, but with gauge coupling identical to those of their SM counterparts. Collisions of protons could result in pair production of squarks, \tilde{q} , which could decay to a SM quark and a neutralino $\tilde{\chi}_1^0$; the neutralino is assumed to be stable in R -parity-conserving

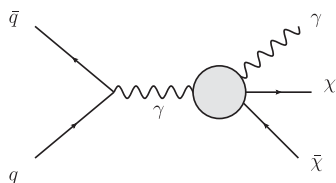


FIG. 4. Production of pairs of dark-matter particles ($\chi\bar{\chi}$) via an effective $\gamma\gamma\chi\bar{\chi}$ vertex.

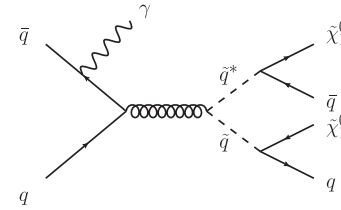


FIG. 5. Pair production of squarks (\tilde{q}), followed by decay into quarks and neutralinos ($\tilde{\chi}_1^0$). The photon may also be radiated from the squarks or final-state quarks.

models [23]. If the mass difference $m_{\tilde{q}} - m_{\tilde{\chi}_1^0}$ is small, the SM quarks would have very low momentum and would therefore not be reconstructed as jets. Again, the radiation of a photon either from an initial-state quark or an intermediate squark would result in $\gamma + E_T^{\text{miss}}$ events, as shown in Fig. 5.

The ATLAS [6] and CMS [7] collaborations have reported limits on various models of new physics based on searches for an excess in $\gamma + E_T^{\text{miss}}$ events using pp collisions at a center-of-mass energy $\sqrt{s} = 7$ TeV. This paper reports the result of a search for new phenomena in $\gamma + E_T^{\text{miss}}$ events in pp collisions at $\sqrt{s} = 8$ TeV.

The paper is organized as follows. Section II gives a brief description of the ATLAS detector. Section III explains the reconstruction of physics objects and Sec. IV describes the event selection applied. Section V describes the signal and background Monte Carlo simulation samples used. Section VI outlines how the SM backgrounds are estimated and discusses the systematic uncertainties on the background estimation. Section VII describes the results and their interpretation, and a summary is finally given in Sec. VIII.

II. THE ATLAS DETECTOR

The ATLAS detector [24] is a multipurpose particle physics apparatus with a forward-backward symmetric cylindrical geometry and near 4π coverage in solid angle [25]. The inner tracking detector (ID) covers the pseudorapidity range $|\eta| < 2.5$, and consists of a silicon pixel detector, a silicon microstrip detector, and, for $|\eta| < 2.0$, a transition radiation tracker (TRT). The ID is surrounded by a thin superconducting solenoid providing a 2 T magnetic field. A high-granularity lead/liquid-argon sampling electromagnetic calorimeter covers the region $|\eta| < 3.2$. An iron/scintillator-tile calorimeter provides hadronic coverage in the range $|\eta| < 1.7$. The liquid-argon technology is also used for the hadronic calorimeters in the end-cap region $1.5 < |\eta| < 3.2$ and for electromagnetic and hadronic measurements in the forward region up to $|\eta| = 4.9$. The muon spectrometer (MS) surrounds the calorimeters. It consists of three large air-core superconducting toroid systems, precision tracking chambers providing accurate muon tracking out to $|\eta| = 2.7$, and additional detectors for triggering in the region $|\eta| < 2.4$.

III. EVENT RECONSTRUCTION

Photons are reconstructed from clusters of energy deposits in the electromagnetic calorimeter measured in projective towers. Clusters without matching tracks are classified as unconverted photon candidates. A photon is considered as a converted photon candidate if it is matched to a pair of tracks that pass a TRT-hits requirement and form a vertex in the ID which is consistent with coming from a massless particle, or if it is matched to a single track passing a TRT-hits requirement and having a first hit after the innermost layer of the pixel detector [26]. The photon energy is corrected by applying the energy scales measured with $Z \rightarrow e^+e^-$ decays and cross-checked with $J/\psi \rightarrow e^+e^-$ and $Z \rightarrow \ell\ell\gamma$ decays [27]. Identification requirements are applied in order to reduce the contamination of the photon sample from π^0 or other neutral hadrons decaying to two photons. The photon identification is based on the profile of the energy deposit in the first and second layers of the electromagnetic calorimeter. Photons have to satisfy the tight identification criteria of Ref. [28]. They are also required to be isolated, i.e., the energy in the calorimeters in a cone of size $\Delta R = \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2} = 0.4$ around the cluster barycenter, excluding the energy associated with the photon cluster, is required to be less than 5 GeV. This cone energy is corrected for the leakage of the photon energy from the central core and for the effects of multiple pp interactions in the same or neighboring bunch crossings superimposed on the hard physics process (referred to as pileup interactions) [29].

Electrons are reconstructed from clusters in the electromagnetic calorimeter matched to a track in the ID and criteria for their identification and calibration procedure are similar to those used for photons. Electron candidates must satisfy the medium $++$ identification requirement of Ref. [27]. Muons are identified either as a combined track in the MS and ID systems, or as an ID track that, once extrapolated to the MS, is associated with at least one track segment in the MS [30].

Jets are reconstructed using the anti- k_t algorithm [31,32] with a radius parameter $R = 0.4$ from calibrated clusters of energy deposits in the calorimeters. These clusters are seeded by calorimeter cells with energies significantly above the measured noise. The differences in calorimeter response between electrons, photons, and hadrons are taken into account by classifying each cluster on the basis of its shape [33], prior to the jet reconstruction, as coming from an electromagnetic or hadronic shower. The jet energy thus accounts for electromagnetic and hadronic energy deposits at the cluster level with correction factors derived from Monte Carlo (MC) simulation. A further correction used to calibrate the jet energy to the scale of its constituent particles [33,34] is then applied. Jets are required to have transverse momentum $p_T > 30$ GeV, $|\eta| < 4.5$, and a distance to the closest preselected electron or photon of $\Delta R > 0.2$.

The vector momentum imbalance in the transverse plane is obtained from the negative vector sum of the reconstructed and calibrated physics objects and is referred to as missing transverse momentum, $\mathbf{E}_T^{\text{miss}}$. The symbol E_T^{miss} is used for its magnitude. Calorimeter energy deposits are associated with a reconstructed and identified high- p_T object in a specific order: electrons with $p_T > 10$ GeV, photons with $p_T > 10$ GeV, and jets with $p_T > 20$ GeV. Deposits not associated with any such objects are also taken into account in the $\mathbf{E}_T^{\text{miss}}$ calculation [35] using an energy-flow algorithm that considers calorimeter energy deposits as well as ID tracks [36].

IV. EVENT SELECTION

The data were collected in pp collisions at $\sqrt{s} = 8$ TeV. Events were selected using an E_T^{miss} trigger that requires a missing transverse momentum greater than 80 GeV [37]. Events selected using an e/γ trigger with a threshold of $p_T > 120$ GeV are also used in some control regions as described below [38]. Only data taken during periods when the calorimeters ID and MS were well functioning are considered. The data used correspond to an integrated luminosity of 20.3 fb $^{-1}$. Quality requirements are applied to photon candidates in order to reject those arising from instrumental problems. In addition, quality requirements are applied in order to remove jets arising from detector noise and out-of-time energy deposits in the calorimeter from cosmic rays or other noncollision sources [39]. Events are required to have a reconstructed primary vertex with at least five associated tracks; the primary vertex is defined as the vertex with the highest sum of the squared transverse momenta of its associated tracks.

The criteria for selecting events in the signal region (SR) are optimized to have good acceptance for the squark model and the dark-matter model with a Z' -like mediator described in Sec. I, as well as to suppress the background from SM processes. This signal region also provides good sensitivity to the other models described in Sec. I. Events in the SR are required to have $E_T^{\text{miss}} > 150$ GeV and a photon with $p_T^{\gamma} > 125$ GeV and $|\eta| < 1.37$. It is also required that the photon and $\mathbf{E}_T^{\text{miss}}$ are not overlapped in azimuth: $\Delta\phi(\gamma, \mathbf{E}_T^{\text{miss}}) > 0.4$. Events with more than one jet or with a jet with $\Delta\phi(\text{jet}, \mathbf{E}_T^{\text{miss}}) < 0.4$ are rejected. Events with one jet are retained to increase the signal acceptance and reduce systematic uncertainties related to the modeling of initial-state radiation. Events are required to have no electron ($p_T > 7$ GeV, $|\eta| < 2.47$) and no muon ($p_T > 6$ GeV, $|\eta| < 2.5$). The lepton veto mainly rejects W/Z events with charged leptons in the final state. For events satisfying these criteria, the E_T^{miss} trigger efficiency is 0.99 ± 0.01 , as determined using events selected with the e/γ trigger. The final data sample contains 521 events, where 319 and 202 events have zero and one jet, respectively.

V. MONTE CARLO SIMULATION SAMPLES

Monte Carlo simulated samples are used to estimate the signal acceptance, the detector efficiency, and to help in the estimation of the SM background contributions.

Simulated signal samples for ADD models are generated with PYTHIA 8 [40] version 1.7.5 using the MSTW2008LO [41] parton distribution function (PDF) set. Simulations were run for two values (2.0 and 2.5 TeV) of the scale parameter M_D and with the number of extra dimensions n varied from two to six.

Simulated samples of dark-matter production $pp \rightarrow \gamma + \chi\bar{\chi} + X$ via the $qq\chi\bar{\chi}$ interaction are generated using MADGRAPH5 [42] version 1.4.8.4, with showering and hadronization modeled by PYTHIA 8 version 1.6.5 using the set of parameters optimized to describe the properties of the events referred to as AU2 tune [43]; the MSTW2008LO PDFs are used. Values of m_χ from 1 to 1300 GeV are considered. In addition, simulated samples of $pp \rightarrow \gamma + \chi\bar{\chi}$ are produced using the simplified model with a Z' -like mediator [11] using the same simulation programs as for the EFT samples. Vector and axial-vector couplings are both considered. For each value of the mediator mass m_V , two different values of the mediator width are simulated: $\Gamma = m_V/8\pi$ and $\Gamma = m_V/3$. The smaller value corresponds to a mediator that can annihilate into only one quark flavor and helicity and has unit couplings; it can be regarded as an approximate lower limit on the mediator width. A value of $\Gamma = m_V/3$ is a reasonable upper bound for a narrow resonance approximation.

Samples of $pp \rightarrow \gamma + \chi\bar{\chi} + X$ are also produced via the $\gamma\gamma\chi\bar{\chi}$ interaction model [12] with a fermionic WIMP. These samples are generated with MADGRAPH5 version 1.4.2 for a WIMP mass of 130 GeV and over a grid of values of k_1 and k_2 .

Simulated samples of $pp \rightarrow \tilde{q}\tilde{q}^*\gamma + X \rightarrow q\bar{q}\gamma + \tilde{\chi}_1^0\tilde{\chi}_1^0 + X$ are generated with MADGRAPH5 version 1.5.11 with showering and hadronization modeled by PYTHIA 6 [44] version 4.2.7 and CTEQ6L1 PDFs [45], with the requirement of having one photon at parton level with $p_T^\gamma > 80$ GeV and $|\eta| < 2.5$. Only the first two generations of squarks are considered, and they are assumed to be degenerate in mass. Signal cross sections are calculated to next-to-leading order in the strong coupling constant including the resummation of soft gluon emission at next-to-leading-logarithm accuracy when available [46–50]. The nominal cross section and its uncertainty are taken from an envelope of cross-section predictions using different PDF sets and factorization and renormalization scales, as described in Ref. [51].

Simulated samples of $Z\gamma$ and $W\gamma$ events are generated with SHERPA version 1.4.1 [52], with parton-level requirements of $p_T^\gamma > 70$ GeV and $p_T^\gamma > 80$ GeV, respectively, and dilepton invariant mass $m_{\ell\ell} > 40$ GeV. A sample of simulated γ + jet events is generated with PYTHIA 8 version 1.6.5. The W/Z + jet processes are also simulated

using SHERPA version 1.4.1 with massive b/c quarks. Diboson samples are generated with HERWIG [53,54] version 6.520, the single-top samples with MC@NLO [55,56] version 4.06 for s -channel and Wt production, and ACERMC [57] version 3.8 for t -channel production. Simulated samples of top-quark pair production are generated with POWHEG [58] version r2129.

HERWIG version 6.520 is used for simulating the parton shower and fragmentation processes in combination with JIMMY [59] for underlying-event simulation for the MC@NLO samples, while PYTHIA 6 version 4.2.6 is used for the POWHEG and ACERMC samples. The proton PDFs used are CTEQ6L1 [45] for the PYTHIA 8 and ACERMC samples, and CT10 [60] for the MC@NLO, SHERPA, and POWHEG samples. The ATLAS underlying-event tune AUET2 [43] is used, except for the $t\bar{t}$ sample, which uses the new Perugia 2011C tune [61]. SHERPA uses its own parton shower, fragmentation, and underlying-event model.

Differing pileup conditions as a function of the instantaneous luminosity are taken into account by overlaying simulated minimum-bias events generated with PYTHIA 8 onto the hard-scattering process and reweighting their number according to the observed distribution of the average number of interactions per beam crossing.

The simulated samples are processed either with a full ATLAS detector simulation [62] based on GEANT4 [63] or a fast simulation based on the parametrization of the response to the electromagnetic and hadronic showers in the ATLAS calorimeters [64] and a simulation of the trigger system. The results based on fast simulation are validated against fully simulated samples. The simulated events are reconstructed and analyzed with the same analysis chain as for the data, using the same trigger and event selection criteria discussed in Sec. IV.

VI. BACKGROUND ESTIMATION

The SM background to the $\gamma + E_T^{\text{miss}}$ final state is dominated by the $Z(\rightarrow \nu\nu) + \gamma$ process, where the photon is due to initial-state radiation. Secondary contributions come from $W\gamma$ and $Z\gamma$ production with unidentified electrons, muons, or hadronically decaying τ leptons, or W/Z production where a lepton or an associated radiated jet is misidentified as a photon. In addition, there are smaller contributions from top-quark pair, diboson, γ + jet, and multijet production.

A. $Z\gamma$ and $W\gamma$ backgrounds

The E_T^{miss} distribution of events due to $Z\gamma$ and $W\gamma$ backgrounds is described using simulated samples, while the normalization is obtained via a likelihood fit to observed yields in several control regions (CRs), constructed to be enriched in specific backgrounds. Poisson likelihood functions are used for event counts in all regions; the systematic uncertainties described in Sec. VI E are

treated as Gaussian-distributed nuisance parameters in the likelihood function. Key ingredients of the fit are the normalization scale factors for the $W\gamma$ and $Z\gamma$ processes, which enable observations in the CRs to constrain background estimates in the SR. The same normalization factor is used for $Z(\nu\nu) + \gamma$, $Z(\mu\mu) + \gamma$, and $Z(ee) + \gamma$ events.

Three control regions are defined by inverting lepton vetoes. In the first control region, the $W\gamma$ contribution is enhanced by requiring the presence of a muon. The second (third) control region enhances the $Z\gamma$ background by requiring the presence of a pair of muons (electrons). In the muon control region, in order to ensure that the E_T^{miss} spectrum is similar to the one in the signal region, muons are treated as invisible particles in the E_T^{miss} calculation. The same procedure is followed for electrons in the electron control region. In each case, the CR lepton selection follows the same requirements as the SR lepton veto with the additional requirements that the lepton must be associated with an ID isolated track and that $\Delta R(\ell, \gamma) > 0.5$. In addition, the photon pseudorapidity requirement is relaxed with respect to the SR selection: $|\eta| < 2.37$, excluding the calorimeter barrel/end-cap transition region $1.37 < |\eta| < 1.52$, to increase the number of events in the CR. In both the $Z\gamma$ -enriched control regions, the dilepton mass $m_{\ell\ell}$ is required to be greater than 50 GeV. The normalization of the dominant $Z\gamma$ background process is largely constrained by the event yields in the two $Z(\ell\ell) + \gamma$ control regions. The results are cross-checked using the transfer-factor technique employed in the previous ATLAS analysis of the $\gamma + E_T^{\text{miss}}$ final state [6]; the two methods give consistent results.

B. Fake photons from misidentified electrons

Contributions from processes in which an electron is misidentified as a photon are estimated by scaling yields from a sample of $e + E_T^{\text{miss}}$ events by an electron-to-photon misidentification factor. This factor is measured in mutually exclusive samples of e^+e^- and $\gamma + e$ events. To establish a pure sample of electrons, m_{ee} and $m_{e\gamma}$ are both required to be consistent with the Z boson mass, and the multijet background estimated from sidebands is subtracted. The misidentification factor is parametrized as a function of p_T in three pseudorapidity bins. Similar estimates are made for the three control regions with leptons, scaling event yields from samples matching the control region requirements, but requiring an electron rather than a photon.

C. Fake photons from misidentified jets

Background contributions from events in which a jet is misidentified as a photon are estimated from samples of $\gamma + E_T^{\text{miss}}$ events where the photon does not fulfill the isolation requirement. The yield in this sample is scaled by a jet-to-photon misidentification factor, after subtraction of the contribution from real photons. The jet-to-photon misidentification factor is measured in samples enriched in

jets, selected by inverting some photon identification criteria, and is determined from the ratio of isolated jets to nonisolated jets. This estimate also accounts for the contribution from multijets, which can mimic the mono-photon signature if one jet is misreconstructed as a photon and one or more of the other jets are poorly reconstructed, resulting in large fake E_T^{miss} . The multijet background is found to be negligible in the SR.

D. $\gamma + \text{jet}$ background

The $\gamma + \text{jet}$ background in the signal region consists of events where the jet is poorly reconstructed and partially lost, creating fake E_T^{miss} . Despite the large production rate, this process is only a minor source of background as it is suppressed by the large E_T^{miss} and the large jet- E_T^{miss} separation requirements in the SR. This background is estimated from MC simulation and is cross-checked with a data-driven estimate, which gives a result in agreement with the MC simulation, but is limited by a large statistical uncertainty. The data-driven estimate is derived from a control region defined by requiring all the selection criteria of the SR but reversing the $\Delta\phi(\text{jet}, E_T^{\text{miss}})$ requirement, thereby selecting poorly reconstructed events in which the jet is aligned with the E_T^{miss} . Simulated samples are used to estimate and subtract electroweak backgrounds coming from $W/Z + \text{jet}$ and $Z/W + \gamma$ processes. As events with a jet with $p_T > 30$ GeV that is not well separated from E_T^{miss} are vetoed in the SR selection, the $\gamma + \text{jet}$ and multijet contribution in the SR is then estimated with a linear extrapolation of the jet p_T spectrum in this CR to the $p_T < 30$ GeV region.

E. Final estimation and systematic uncertainties

Background estimates in the SR are first derived from a fit using only data from the lepton CRs, in order to assess whether the observed SR yield is consistent with the background model. The values of the normalization factors for the $W\gamma$ and $Z\gamma$ backgrounds obtained from the fit to the CRs are $k_{W\gamma} = 0.81 \pm 0.05(\text{stat}) \pm 0.06(\text{syst})$ and $k_{Z\gamma} = 0.89 \pm 0.08(\text{stat}) \pm 0.08(\text{syst})$, where the systematic error takes into account the various sources of systematic uncertainties described below. Distributions of the missing transverse momentum in the three control regions are shown in Figs. 6–8.

The techniques used for the background estimation are checked in a validation region, where events are selected with the same criteria as used for the signal region, except for a lower E_T^{miss} (110–150 GeV) and a larger photon pseudorapidity range ($|\eta| < 2.37$, excluding the calorimeter barrel/end-cap transition region $1.37 < |\eta| < 1.52$) to increase the statistical power. To suppress the background from $\gamma + \text{jet}$ events and from fake photons to a level similar to that in the SR, a requirement on the azimuthal separation between the photon and the jet—when there is a jet in the event—is applied: $\Delta\phi(\gamma, \text{jet}) < 2.7$. To minimize the

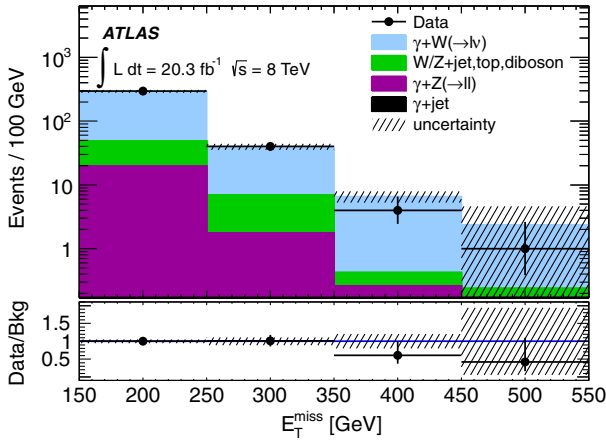


FIG. 6 (color online). Distribution of E_T^{miss} in the data and for the expected background in the single-muon control region. The total background expectation is normalized to the observed number of events in this control region. The dashed band includes statistical and systematic uncertainties. Overflows are included in the final bin. The lower part of the figure shows the ratios of data to expected-background event yields.

contamination of this region by signal events, a requirement on the azimuthal separation between the photon and E_T^{miss} is added: $\Delta\phi(\gamma, E_T^{\text{miss}}) < 3.0$. The number of events in the data in this region is 307 and the estimated total background obtained from the background-only fit to the control regions is $272 \pm 17 \pm 14$, resulting in agreement between the data and expectation within 2σ . Detailed results are shown in Table I; systematic uncertainties are computed as described below for the SR.

Systematic uncertainties on the background predictions in the signal region are presented here as percentages of the

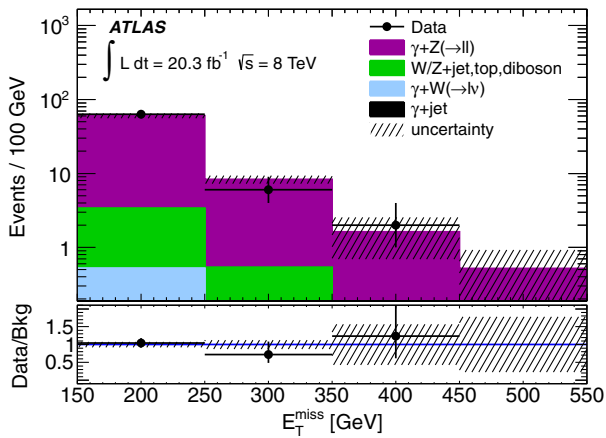


FIG. 7 (color online). Distribution of E_T^{miss} in the data and for the expected background in the two-muon control region. The total background expectation is normalized to the observed number of events in this control region. The dashed band includes statistical and systematic uncertainties. Overflows are included in the final bin. The lower part of the figure shows the ratios of data to expected-background event yields.

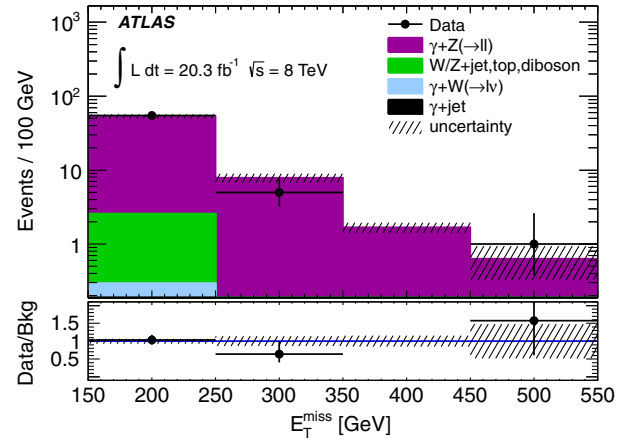


FIG. 8 (color online). Distribution of E_T^{miss} in the data and for the expected background in the two-electron control region. The total background expectation is normalized to the observed number of events in this control region. The dashed band includes statistical and systematic uncertainties. Overflows are included in the final bin. The lower part of the figure shows the ratios of data to expected-background event yields.

total background prediction. This prediction is obtained from the CR fit, which provides constraints on many of the sources of systematic uncertainty. The dominant contribution is due to the uncertainty on the electron fake rate, which contributes a 4.6% relative uncertainty, and to the reconstruction and identification efficiency corrections applied to electrons and muons in MC simulation, which contribute 1.3% and 0.7% relative uncertainty, respectively. The uncertainty on the absolute electron/photon energy scale translates into a 0.6% relative uncertainty on the total background prediction. Uncertainties in the simulation of the electron/photon energy resolution, isolation, and identification efficiency contribute a relative uncertainty of 0.1% on the total predicted background. The uncertainty on the absolute jet energy scale [34] and the jet energy resolution [65] contribute 0.1% and 0.5% relative

TABLE I. Observed event yield compared to predicted event yield from SM backgrounds in the SR and the validation region (VR), using estimates and uncertainties obtained from a fit in the control regions. Uncertainties are statistical followed by systematic. In the case of the γ + jet process a global uncertainty is quoted.

| Process | Event yield (SR) | Event yield (VR) |
|--|------------------------|-----------------------|
| $Z(\rightarrow \nu\nu) + \gamma$ | $389 \pm 36 \pm 10$ | $153 \pm 16 \pm 10$ |
| $W(\rightarrow \ell\nu) + \gamma$ | $82.5 \pm 5.3 \pm 3.4$ | $67 \pm 5 \pm 5$ |
| $W/Z + \text{jet}, \tau\tau, \text{diboson}$ | $83 \pm 2 \pm 28$ | $47 \pm 2 \pm 14$ |
| $Z(\rightarrow \ell\ell) + \gamma$ | $2.0 \pm 0.2 \pm 0.6$ | $2.9 \pm 0.3 \pm 0.6$ |
| $\gamma + \text{jet}$ | $0.4^{+0.3}_{-0.4}$ | $2.5^{+4.0}_{-2.5}$ |
| Total background | $557 \pm 36 \pm 27$ | $272 \pm 17 \pm 14$ |
| Data | 521 | 307 |

uncertainties, respectively. Uncertainties on the scale and resolution of the calorimeter energy deposits not associated with high- p_T physics objects affect the calculation of the E_T^{miss} and generate an uncertainty of 0.3% on the background prediction. Uncertainties on the PDF are evaluated by following, for the CT10 and MSTW2008LO PDF sets, the PDF4LHC recommendations [66]. The Hessian method is used to obtain asymmetric uncertainties at 68% confidence level (C.L.). In addition, to obtain inter-PDF uncertainties, the results are then compared with those obtained with the NNPDF set. Renormalization and factorization scale uncertainties are also taken into account by increasing and decreasing the scales used in the MC generators by a factor of 2. PDF and scale uncertainties contribute 0.7% to the background prediction uncertainty. After the fit, the uncertainty on the jet energy scale due to corrections for pileup, and the uncertainties on the trigger efficiency and luminosity [67], are found to have a negligible impact on the background estimation. The final total background prediction systematic uncertainty is about 5%, while the statistical uncertainty is about 6%.

VII. RESULTS

Table I presents the observed number of events and the SM background predictions obtained from a fit to the CRs. The E_T^{miss} distribution in the SR is shown in Fig. 9.

As the 521 events observed in data are well described by the SM background prediction of $557 \pm 36 \pm 27$, the results are interpreted in terms of exclusions on models that would produce an excess of $\gamma + E_T^{\text{miss}}$ events. Upper bounds are calculated using a one-sided profile likelihood ratio and the CL_s technique [68,69], evaluated using the asymptotic

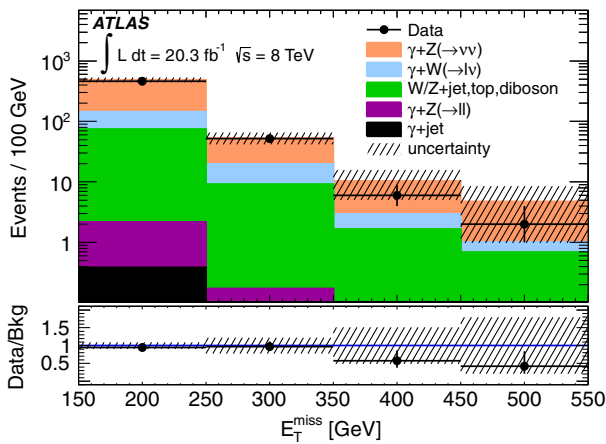


FIG. 9 (color online). Distribution of E_T^{miss} in the signal region for data and for the background predicted from a fit in the CRs. The dashed band includes statistical and systematic uncertainties. Overflows are included in the final bin. The lower part of the figure shows the ratios of data to expected-background event yields.

approximation [70], making use of data in the CRs as well as in the SR.

The most model-independent limits provided are those on the fiducial cross section of a potential new physics process, $\sigma \times A$, where σ is the cross section and A is the fiducial acceptance. The latter is defined using a selection identical to that defining the signal region but applied at particle level, where the particle-level E_T^{miss} is the vector sum of invisible particle momenta. The limit on $\sigma \times A$ is derived from a limit on the visible cross section $\sigma \times A \times \epsilon$, where ϵ is the fiducial reconstruction efficiency. A conservative estimate $\epsilon = 69\%$ is computed using ADD and WIMP samples with no quark/gluon produced from the main interaction vertex. The expected (observed) upper limit on the fiducial cross section is 6.1 (5.3) fb at 95% C.L. and 5.1 (4.4) fb at 90% C.L. These limits are applicable to any model that produces $\gamma + E_T^{\text{miss}}$ events in the fiducial region and has similar reconstruction efficiency ϵ .

For limits on specific models, the impact of systematic uncertainties on signal samples is evaluated separately for $A \times \epsilon$ [PDF, scale, initial-state radiation (ISR), and final-state radiation (FSR) uncertainties] and the cross section σ (PDF and scale uncertainties). Only uncertainties affecting $A \times \epsilon$ are included in the statistical analysis; uncertainties affecting the cross section are indicated as bands on observed limits and written as σ_{theo} . For the EFT and simplified-model DM samples, scale uncertainties are evaluated by varying the renormalization, factorization, and matching scales in MADGRAPH by a factor of 2. For the ADD samples, the PYTHIA 8 renormalization and factorization scale parameters are varied independently to 0.5 and 2.0. For these samples, the ISR and FSR signal uncertainties are assessed by varying the PYTHIA 8 parameters, as done in Ref. [71]. For the squark model described in Sec. I, systematic uncertainties arising from the treatment of ISR/FSR are studied with MC event samples by varying the value of α_s ; the renormalization and factorization scales and the MADGRAPH/PYTHIA matching parameter are also varied to estimate the related uncertainties. Radiation uncertainties are typically less than 10%, PDF uncertainties less than 30%, and scale uncertainties less than 20%.

Limits on dark-matter production are derived from the cross-section limits at a given WIMP mass m_χ , and expressed as 90% C.L. limits on the suppression scale M_* , for the D5 (Fig. 10), D8 (Fig. 11), and D9 (Fig. 12) operators. Values of M_* up to 760, 760, and 1010 GeV are excluded for the D5, D8, and D9 operators, respectively. As already mentioned, the effective field theory model becomes a poor approximation when the momentum transferred in the interaction Q_{tr} is comparable to the mass of the intermediate state $m_V = M_* \sqrt{g_f g_\chi}$ [10,72]. In order to illustrate the sensitivity to the unknown ultraviolet completion of the theory, limits computed retaining only simulated events with $Q_{\text{tr}} < m_V$ are also shown, for a value

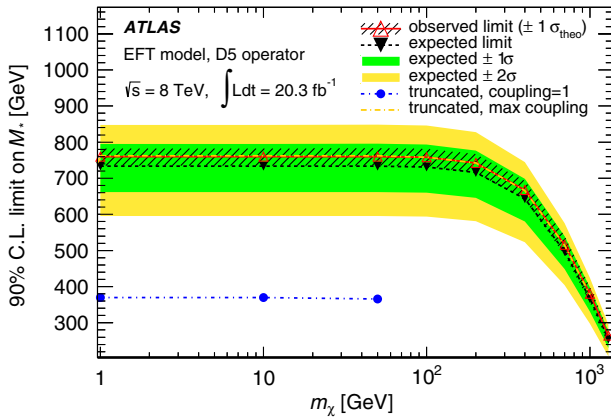


FIG. 10 (color online). Limits at 90% C.L. on the EFT suppression scale M_* as a function of the WIMP mass m_χ , for the vector operator D5. Results where EFT truncation is applied (see text) are also shown, assuming coupling values $\sqrt{g_f g_\chi} = 1, 4\pi$.

of the coupling $\sqrt{g_f g_\chi}$ equal to either unity or the maximum value (4π) that allows the perturbative approach to be valid. This procedure is referred to as truncation. As can be seen in Figs. 10–12, the truncated limits nearly overlap with the nontruncated limits for a 4π coupling. For unit coupling, the truncated limits are less stringent than the nontruncated limits at low m_χ , and the analysis loses sensitivity for $m_\chi > 100$ (200) GeV for the D5 and D8 (D9) operators. In this case, for the D5 and D8 operators, as no sample was generated between $m_\chi = 50$ GeV and $m_\chi = 100$ GeV, the limit is only shown up to $m_\chi = 50$ GeV; for the D9 operator, as no sample was generated between $m_\chi = 100$ GeV and $m_\chi = 200$ GeV, the limit is only shown up to $m_\chi = 100$ GeV. These lower limits on M_* can be translated into upper limits on the WIMP-nucleon interaction cross section as a function of m_χ using Eqs. (4) and (5) of

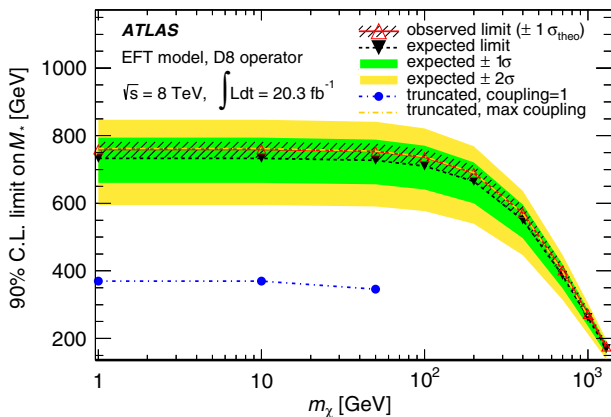


FIG. 11 (color online). Limits at 90% C.L. on the EFT suppression scale M_* as a function of the WIMP mass m_χ , for the axial-vector operator D8. Results where EFT truncation has been applied (see text) are also shown, assuming coupling values $\sqrt{g_f g_\chi} = 1, 4\pi$.

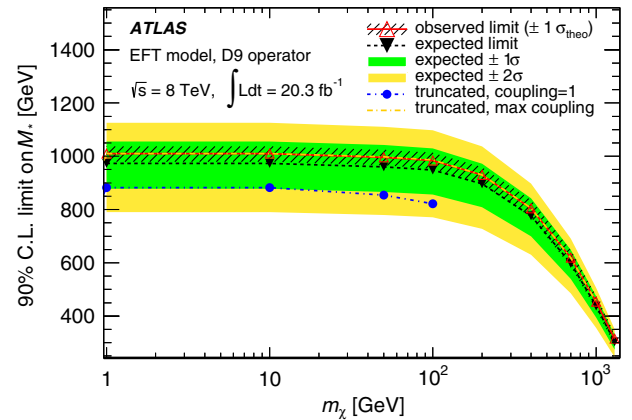


FIG. 12 (color online). Limits at 90% C.L. on the EFT suppression scale M_* as a function of the WIMP mass m_χ , for the tensor operator D9. Results where EFT truncation is applied (see text) are also shown, assuming coupling values $\sqrt{g_f g_\chi} = 1, 4\pi$.

Ref. [10]. Results are shown in Fig. 13 for spin-independent (D5) and spin-dependent (D8, D9) χ -nucleon interactions and are compared to measurements from various DM search experiments [73–85]. The search for dark-matter pair production in association with a γ at the LHC extends the limits on the χ -nucleon scattering cross section into the low-mass region $m_\chi < 10$ GeV where the astroparticle experiments have less sensitivity due to the very low-energy recoils such low-mass DM particles would induce.

Simplified models with explicit mediators are ultraviolet complete and therefore robust for all values of Q_{tr} . For the simplified Z' -like model with vector interactions and mediator width $\Gamma = m_V/3$, Fig. 14 shows the 95% C.L. limits on the coupling parameter $\sqrt{g_f g_\chi}$ calculated for various values of the WIMP and mediator particle masses, and compared to the lower limit resulting from the relic DM abundance [86]. In the region above the dashed line, the lower limits on the coupling resulting from the relic abundance of DM are higher than the upper limits found in this analysis. Figures 15 and 16 show, for vector and axial-vector interactions and different values of the WIMP mass, the corresponding 95% C.L. limits on the suppression scale M_* as a function of m_V . One can note how, when the mediator mass is greater than the LHC reach, the EFT model provides a good approximation of the simplified model with $M_* = m_V/\sqrt{g_f g_\chi}$. The truncation procedure is applied when computing the EFT limits; these limits are always more conservative than those from the simplified model as long as m_V is greater than or equal to the value used for EFT truncation. This can be seen by comparing the M_* limits derived from the EFT approach using truncation (Figs. 10 and 11) to those of the simplified model, recalling $m_V = M_*\sqrt{g_f g_\chi}$.

In the case of the model of $\gamma\gamma\chi\bar{\chi}$ interactions with an s -channel SM gauge boson inspired by the line near

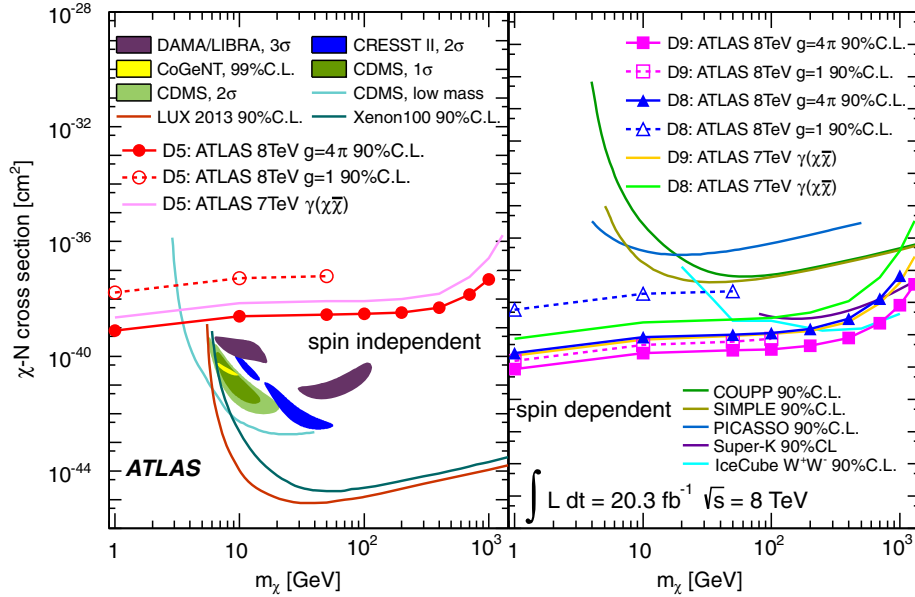


FIG. 13 (color online). Upper limits at 90% C.L. on the WIMP-nucleon (χ -N) scattering cross section as a function of m_χ for spin-independent (left) and spin-dependent (right) interactions, for a coupling strength $g = \sqrt{g_f g_\chi}$ of unity or the maximum value (4π) that keeps the model within its perturbative regime. The truncation procedure is applied for both cases. The results obtained from ATLAS with 7 TeV data for the same channel are shown for comparison. Also shown are results from various dark-matter search experiments [73–85].

130 GeV in the Fermi-LAT γ -ray spectrum, limits are placed on the effective mass scale M_* in the (k_2, k_1) parameter plane, as shown in Fig. 17. The exclusion line is drawn by considering the value of M_* needed to generate the $\chi\bar{\chi} \rightarrow \gamma\gamma$ annihilation rate consistent with the observed Fermi-LAT γ -ray line near 130 GeV. This model is able to

provide an effective constraint on the portion of the parameter space of the theory compatible with the Fermi-LAT peak.

In the ADD model of LED, limits on M_D for various values of n are provided in Fig. 18. Results incorporating truncation are also shown, for which the graviton

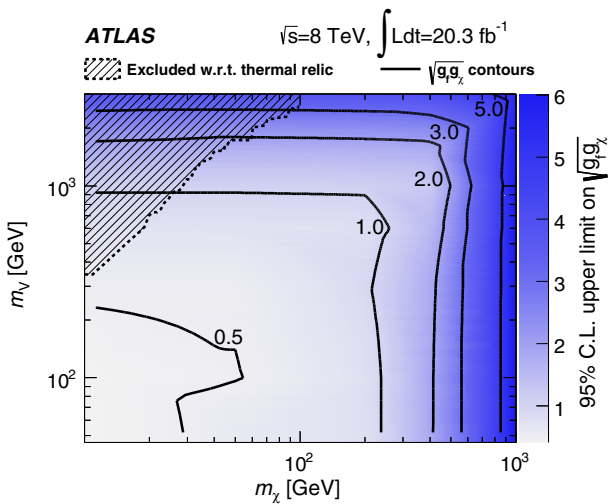


FIG. 14 (color online). Upper limits at 95% C.L. on the WIMP simplified-model coupling parameter, $\sqrt{g_f g_\chi}$, with vector coupling and mediator width $\Gamma = m_V/3$, as a function of the WIMP (m_χ) and the mediator particle (m_V) masses. Solid lines indicate contours in the coupling parameter. The lower limit on the coupling resulting from the relic abundance of DM is also shown.

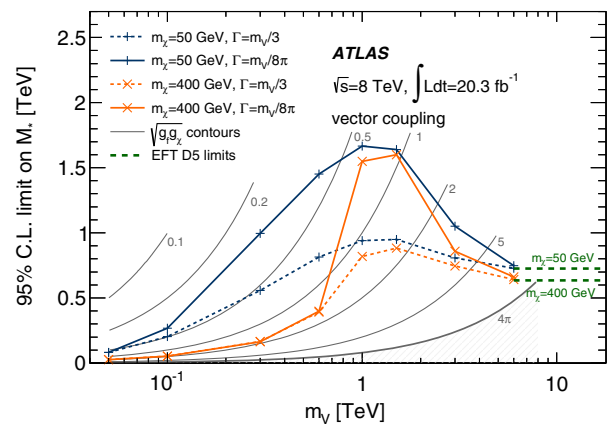


FIG. 15 (color online). Observed lower limits at 95% C.L. on the EFT suppression scale M_* as a function of the mediator mass m_V , for a Z' -like mediator with vector interactions. For a dark-matter mass m_χ of 50 or 400 GeV, results are shown for different values of the mediator total decay width Γ and compared to the EFT observed limit results for a D5 (vector) interaction. M_* vs m_V contours for an overall coupling $\sqrt{g_f g_\chi} = 0.1, 0.2, 0.5, 1, 2, 5, 4\pi$ are also shown. The corresponding limits from the D5 operator are shown as a dashed line.

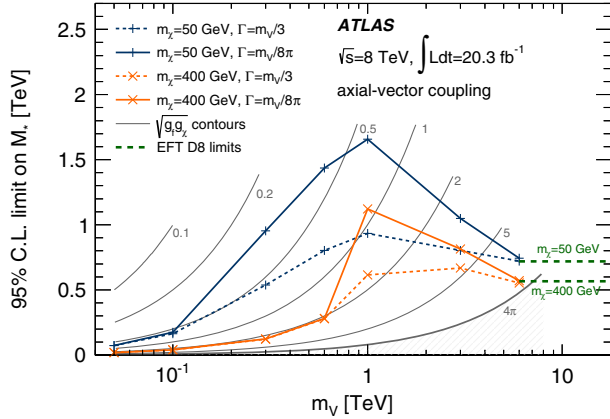


FIG. 16 (color online). Observed limits at 95% C.L. on the EFT suppression scale M_* as a function of the mediator mass m_V , for a Z' -like mediator with axial-vector interactions. For a dark-matter mass m_χ of 50 or 400 GeV, results are shown for different values of the mediator total decay width Γ and compared to the EFT observed limit results for a D8 (axial-vector) interaction. M_* vs m_V contours for an overall coupling $\sqrt{g_f g_\chi} = 0.1, 0.2, 0.5, 1, 2, 5, 4\pi$ are also shown. The corresponding limits from the D8 operator are shown as a dashed line.

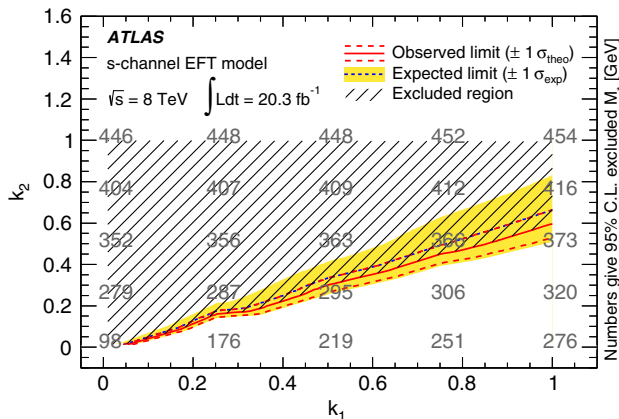


FIG. 17 (color online). Limits at 95% C.L. on the effective mass scale M_* in the (k_2, k_1) parameter plane for the s -channel EFT model inspired by Fermi-LAT γ -ray line, for $m_\chi = 130$ GeV. The upper part of the plane is excluded.

production cross section is suppressed by a factor M_D^4/\hat{s}^2 , where $\sqrt{\hat{s}}$ is the parton-parton center-of-mass energy. The analysis is able to exclude M_D up to 2.17 TeV, depending on the number of extra dimensions. The effect of truncation is larger for higher n as the graviton mass distribution is pushed to higher values.

In the case of squark pair production, limits on $\sigma(pp \rightarrow \tilde{q}\tilde{q}^* \gamma + X)$ as a function of $m_{\tilde{q}}$ and $m_{\tilde{q}} - m_{\tilde{\chi}_1^0}$ are presented in Fig. 19. The limit is presented down to $m_{\tilde{q}} - m_{\tilde{\chi}_1^0} = m_c$, below which the decay of the $\tilde{c} \rightarrow c\tilde{\chi}_1^0$ is off shell and not considered here. For very compressed spectra, the analysis is able to exclude squark masses up to 250 GeV. Some models of first- and second-generation squark pair

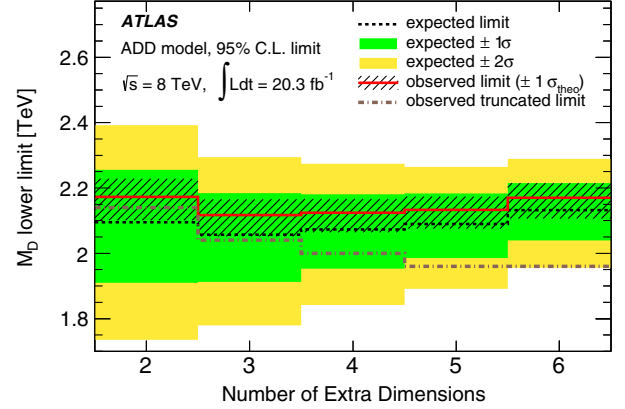


FIG. 18 (color online). Lower limits at 95% C.L. on the mass scale M_D in the ADD models of large extra dimensions, for several values of the number of extra dimensions. The expected and observed limits are shown, along with the limit obtained after applying truncation.

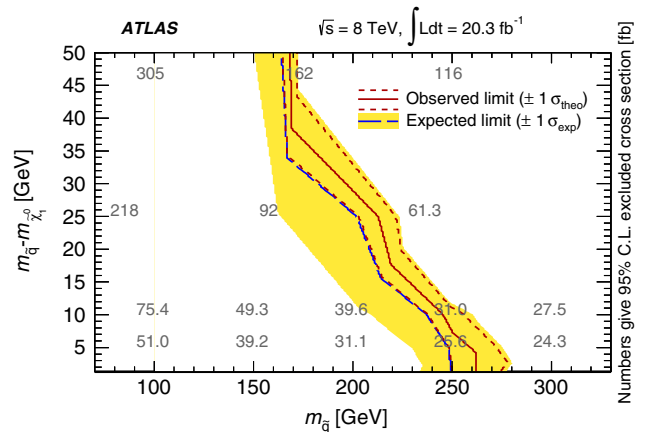


FIG. 19 (color online). Upper limits at 95% C.L. on the cross section for the compressed squark model, as a function of the squark mass, $m_{\tilde{q}}$, and of the difference between the squark mass and the mass of the neutralino, $m_{\tilde{q}} - m_{\tilde{\chi}_1^0}$, in the compressed region of $m_{\tilde{q}} - m_{\tilde{\chi}_1^0} < 50$ GeV. The observed (solid line) and expected (dashed line) upper limits from this analysis are shown; the upper limit on the cross section (in fb) is indicated for each model point.

production are also explored in Ref. [87]; the result presented here is complementary in that it probes very compressed spectra. Due to the reduced hadronic activity, the acceptance of the $\gamma + E_T^{\text{miss}}$ selection indeed increases as the mass difference between the squarks and the neutralino decreases, leading to an increased sensitivity to squark mass with decreasing mass difference.

VIII. SUMMARY

Results are reported from a search for new phenomena in events with a high- p_T photon and large missing transverse

momentum in pp collisions at $\sqrt{s} = 8$ TeV at the LHC, using ATLAS data corresponding to an integrated luminosity of 20.3 fb^{-1} . The observed data are in agreement with the SM background prediction. The expected (observed) upper limits on the fiducial cross section $\sigma \times A$ are 6.1 (5.3) fb at 95% C.L. and 5.1 (4.4) fb at 90% C.L. In addition, limits are placed on parameters of theories of large extra dimensions, WIMP dark matter, and supersymmetric quarks.

ACKNOWLEDGMENTS

We thank CERN for the very successful operation of the LHC, as well as the support staff from our institutions without whom ATLAS could not be operated efficiently. We acknowledge the support of ANPCyT, Argentina; YerPhI, Armenia; ARC, Australia; BMWFW and FWF, Austria; ANAS, Azerbaijan; SSTC, Belarus; CNPq and FAPESP, Brazil; NSERC, NRC, and CFI, Canada; CERN; CONICYT, Chile; CAS, MOST, and NSFC, China; COLCIENCIAS, Colombia; MSMT CR, MPO CR, and VSC CR, Czech Republic; DNRF, DNSRC, and Lundbeck Foundation, Denmark; EPLANET, ERC,

and NSRF, European Union; IN2P3-CNRS, CEA-DSM/IRFU, France; GNSF, Georgia; BMBF, DFG, HGF, MPG, and AvH Foundation, Germany; GSRT and NSRF, Greece; ISF, MINERVA, GIF, I-CORE, and Benoziyo Center, Israel; INFN, Italy; MEXT and JSPS, Japan; CNRST, Morocco; FOM and NWO, Netherlands; BRF and RCN, Norway; MNiSW and NCN, Poland; GRICES and FCT, Portugal; MNE/IFA, Romania; MES of Russia and ROSATOM, Russian Federation; JINR; MSTD, Serbia; MSSR, Slovakia; ARRS and MIZŠ, Slovenia; DST/NRF, South Africa; MINECO, Spain; SRC and Wallenberg Foundation, Sweden; SER, SNSF, and Cantons of Bern and Geneva, Switzerland; NSC, Taiwan; TAEK, Turkey; STFC, the Royal Society, and Leverhulme Trust, United Kingdom; DOE and NSF, United States of America. The crucial computing support from all WLCG partners is acknowledged gratefully, in particular from CERN and the ATLAS Tier-1 facilities at TRIUMF (Canada), NDGF (Denmark, Norway, Sweden), CC-IN2P3 (France), KIT/GridKA (Germany), INFN-CNAF (Italy), NL-T1 (Netherlands), PIC (Spain), ASGC (Taiwan), RAL (UK), and BNL (USA) and in the Tier-2 facilities worldwide.

-
- [1] T. Aaltonen *et al.* (CDF Collaboration), *Phys. Rev. Lett.* **101**, 181602 (2008).
- [2] V. Abazov *et al.* (D0 Collaboration), *Phys. Rev. Lett.* **101**, 011601 (2008).
- [3] G. Abbiendi *et al.* (OPAL Collaboration), *Eur. Phys. J. C* **18**, 253 (2000).
- [4] P. Achard *et al.* (L3 Collaboration), *Phys. Lett. B* **587**, 16 (2004).
- [5] J. Abdallah *et al.* (DELPHI Collaboration), *Eur. Phys. J. C* **38**, 395 (2005).
- [6] ATLAS Collaboration, *Phys. Rev. Lett.* **110**, 011802 (2013).
- [7] CMS Collaboration, *Phys. Rev. Lett.* **108**, 261803 (2012).
- [8] N. Arkani-Hamed, S. Dimopoulos, and G. Dvali, *Phys. Lett. B* **429**, 263 (1998).
- [9] G. Bertone, D. Hooper, and J. Silk, *Phys. Rep.* **405**, 279 (2005).
- [10] J. Goodman, M. Ibe, A. Rajaraman, W. Shepherd, T. M. P. Tait, and H.-B. Yu, *Phys. Rev. D* **82**, 116010 (2010).
- [11] P. J. Fox, R. Harnik, J. Kopp, and Y. Tsai, *Phys. Rev. D* **85**, 056011 (2012).
- [12] A. Nelson, L. M. Carpenter, R. Cotta, A. Johnstone, and D. Whiteson, *Phys. Rev. D* **89**, 056011 (2014).
- [13] C. Weniger, *J. Cosmol. Astropart. Phys.* **08** (2012) 007.
- [14] H. Miyazawa, *Prog. Theor. Phys.* **36**, 1266 (1966).
- [15] P. Ramond, *Phys. Rev. D* **3**, 2415 (1971).
- [16] Y. A. Golfand and E. P. Likhtman, *Pisma Zh. Eksp. Teor. Fiz.* **13**, 452 (1971) [*JETP Lett.* **13**, 323 (1971)].
- [17] A. Neveu and J. H. Schwarz, *Nucl. Phys.* **B31**, 86 (1971).
- [18] A. Neveu and J. H. Schwarz, *Phys. Rev. D* **4**, 1109 (1971).
- [19] J. Gervais and B. Sakita, *Nucl. Phys.* **B34**, 632 (1971).
- [20] D. V. Volkov and V. P. Akulov, *Phys. Lett.* **46B**, 109 (1973).
- [21] J. Wess and B. Zumino, *Phys. Lett.* **49B**, 52 (1974).
- [22] J. Wess and B. Zumino, *Nucl. Phys.* **B70**, 39 (1974).
- [23] P. Fayet, *Phys. Lett.* **64B**, 159 (1976).
- [24] ATLAS Collaboration, *J. Instrum.* **3**, S08003 (2008).
- [25] ATLAS uses a right-handed coordinate system with its origin at the nominal interaction point (IP) in the center of the detector and the z axis along the beam pipe. The x axis points from the IP to the center of the LHC ring, and the y axis points upward. Cylindrical coordinates (r, ϕ) are used in the transverse plane, ϕ being the azimuthal angle around the beam pipe. The pseudorapidity is defined in terms of the polar θ angle as $\eta = -\ln[\tan(\theta/2)]$.
- [26] ATLAS Collaboration, Report No. ATL-PHYS-PUB-2011-007, 2011 [<http://cdsweb.cern.ch/record/1345329>].
- [27] ATLAS Collaboration, *Eur. Phys. J. C* **74**, 3071 (2014).
- [28] ATLAS Collaboration, Report No. ATLAS-CONF-2012-123, 2012 [<http://cdsweb.cern.ch/record/1473426>].
- [29] ATLAS Collaboration, *Phys. Rev. D* **83**, 052005 (2011).
- [30] ATLAS Collaboration, *Eur. Phys. J. C* **74**, 3130 (2014).
- [31] M. Cacciari, G. P. Salam, and G. Soyez, *J. High Energy Phys.* **04** (2008) 063.
- [32] M. Cacciari and G. P. Salam, *Phys. Lett. B* **641**, 57 (2006).
- [33] ATLAS Collaboration, *Eur. Phys. J. C* **73**, 2304 (2013).
- [34] ATLAS Collaboration, *Eur. Phys. J. C* **73**, 2305 (2013).
- [35] ATLAS Collaboration, Report No. ATLAS-CONF-2013-082, 2013 [<http://cdsweb.cern.ch/record/1570993>].
- [36] ATLAS Collaboration, *Eur. Phys. J. C* **72**, 1844 (2012).

- [37] ATLAS Collaboration, Report No. ATLAS-CONF-2014-002, 2014 [<http://cdsweb.cern.ch/record/1647616>].
- [38] ATLAS Collaboration, Report No. ATLAS-CONF-2012-048, 2012 [<http://cdsweb.cern.ch/record/1450089>].
- [39] ATLAS Collaboration, Report No. ATLAS-CONF-2012-020, 2012 [<http://cdsweb.cern.ch/record/1430034>].
- [40] T. Sjostrand, S. Mrenna, and P. Z. Skands, *Comput. Phys. Commun.* **178**, 852 (2008).
- [41] A. Martin, W. Stirling, R. Thorne, and G. Watt, *Eur. Phys. J. C* **63**, 189 (2009).
- [42] J. Alwall, M. Herquet, F. Maltoni, O. Mattelaer, and T. Stelzer, *J. High Energy Phys.* **06** (2011) 128.
- [43] ATLAS Collaboration, Report No. ATL-PHYS-PUB-2011-008, 2011 [<http://cdsweb.cern.ch/record/1345343>].
- [44] T. Sjostrand, S. Mrenna, and P. Skands, *J. High Energy Phys.* **05** (2006) 026.
- [45] J. Pumplin, D. R. Stump, J. Huston, H.-L. Lai, P. Nadolsky, and W.-K. Tung, *J. High Energy Phys.* **07** (2002) 012.
- [46] W. Beenakker, R. Hopker, M. Spira, and P. Zerwas, *Nucl. Phys.* **B492**, 51 (1997).
- [47] A. Kulesza and L. Motyka, *Phys. Rev. Lett.* **102**, 111802 (2009).
- [48] A. Kulesza and L. Motyka, *Phys. Rev. D* **80**, 095004 (2009).
- [49] W. Beenakker *et al.*, *J. High Energy Phys.* **12** (2009) 041.
- [50] W. Beenakker, S. Brensing, M. Krämer, A. Kulesza, E. Laenen, L. Motyka, and I. Niessen, *Int. J. Mod. Phys. A* **26**, 2637 (2011).
- [51] M. Kramer *et al.*, [arXiv:1206.2892](https://arxiv.org/abs/1206.2892).
- [52] T. Gleisberg, S. Höche, F. Krauss, M. Schönherr, S. Schumann, F. Siegert, and J. Winter, *J. High Energy Phys.* **02** (2009) 007.
- [53] G. Corcella, I. G. Knowles, G. Marchesini, S. Moretti, K. Odagiri, P. Richardson, M. H. Seymour, and B. R. Webber, *J. High Energy Phys.* **01** (2001) 010.
- [54] G. Corcella *et al.*, [arXiv:hep-ph/0210213](https://arxiv.org/abs/hep-ph/0210213).
- [55] S. Frixione and B. R. Webber, *J. High Energy Phys.* **06** (2002) 029.
- [56] S. Frixione, P. Nason, and B. R. Webber, *J. High Energy Phys.* **08** (2003) 007.
- [57] B. P. Kersevan and E. Richter-Was, [arXiv:hep-ph/0405247](https://arxiv.org/abs/hep-ph/0405247).
- [58] S. Frixione, P. Nason, and G. Ridolfi, *J. High Energy Phys.* **09** (2007) 126.
- [59] J. Butterworth, J. R. Forshaw, and M. Seymour, *Z. Phys. C* **72**, 637 (1996).
- [60] H.-L. Lai, M. Guzzi, J. Huston, Z. Li, P. M. Nadolsky, J. Pumplin, and C.-P. Yuan, *Phys. Rev. D* **82**, 074024 (2010).
- [61] B. Cooper *et al.*, [arXiv:1109.5295](https://arxiv.org/abs/1109.5295)
- [62] ATLAS Collaboration, *Eur. Phys. J. C* **70**, 823 (2010).
- [63] S. Agostinelli *et al.* (GEANT4 Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **506**, 250 (2003).
- [64] ATLAS Collaboration, Report No. ATL-PHYS-PUB-2010-013, 2010 [<http://cdsweb.cern.ch/record/1300517>].
- [65] ATLAS Collaboration, Report No. ATLAS-CONF-2010-054, 2010 [<http://cdsweb.cern.ch/record/1281311>].
- [66] M. Botje *et al.*, [arXiv:1101.0538](https://arxiv.org/abs/1101.0538).
- [67] ATLAS Collaboration, *Eur. Phys. J. C* **73**, 2518 (2013).
- [68] A. L. Read, *J. Phys. G* **28**, 2693 (2002).
- [69] T. Junk, *Nucl. Instrum. Methods Phys. Res., Sect. A* **434**, 435 (1999).
- [70] G. Cowan, K. Cranmer, E. Gross, and O. Vitells, *Eur. Phys. J. C* **71**, 1 (2011).
- [71] ATLAS Collaboration, *Phys. Rev. D* **90**, 012004 (2014).
- [72] G. Busoni, A. De Simone, E. Morgante, and A. Riotto, *Phys. Lett. B* **728**, 412 (2014).
- [73] E. Aprile *et al.* (XENON100 Collaboration), *Phys. Rev. Lett.* **109**, 181301 (2012).
- [74] Z. Ahmed *et al.* (CDMS Collaboration), *Phys. Rev. Lett.* **106**, 131302 (2011).
- [75] C. E. Aalseth *et al.* (CoGeNT Collaboration), *Phys. Rev. Lett.* **106**, 131301 (2011).
- [76] S. Archambault *et al.* (PICASSO Collaboration), *Phys. Lett. B* **711**, 153 (2012).
- [77] M. Felizardo *et al.* (SIMPLE Collaboration), *Phys. Rev. Lett.* **108**, 201302 (2012).
- [78] D. S. Akerib *et al.* (LUX Collaboration), *Phys. Rev. Lett.* **112**, 091303 (2014).
- [79] S. Desai *et al.* (Super-Kamiokande Collaboration), *Phys. Rev. D* **70**, 083523 (2004).
- [80] M. G. Aartsen *et al.* (IceCube Collaboration), *Phys. Rev. Lett.* **110**, 131302 (2013).
- [81] E. Behnke *et al.* (COUPP Collaboration), *Phys. Rev. D* **86**, 052001 (2012).
- [82] G. Angloher *et al.* (CRESST-II Collaboration), *Eur. Phys. J. C* **72**, 1971 (2012).
- [83] R. Agnese *et al.* (CDMS Collaboration), *Phys. Rev. Lett.* **111**, 251301 (2013).
- [84] R. Bernabei *et al.* (DAMA Collaboration), *Eur. Phys. J. C* **56**, 333 (2008).
- [85] A. Abramowski *et al.* (HESS Collaboration), *Phys. Rev. Lett.* **110**, 041301 (2013).
- [86] E. Komatsu *et al.* (WMAP Collaboration), *Astrophys. J. Suppl. Ser.* **192**, 18, (2011).
- [87] ATLAS Collaboration, *J. High Energy Phys.* **09** (2014) 176.

G. Aad,⁸⁵ B. Abbott,¹¹³ J. Abdallah,¹⁵² S. Abdel Khalek,¹¹⁷ O. Abdinov,¹¹ R. Aben,¹⁰⁷ B. Abi,¹¹⁴ M. Abolins,⁹⁰ O. S. AbouZeid,¹⁵⁹ H. Abramowicz,¹⁵⁴ H. Abreu,¹⁵³ R. Abreu,³⁰ Y. Abulaiti,^{147a,147b} B. S. Acharya,^{165a,165b,b} L. Adamczyk,^{38a} D. L. Adams,²⁵ J. Adelman,¹⁰⁸ S. Adomeit,¹⁰⁰ T. Adye,¹³¹ T. Agatonovic-Jovin,^{13a} J. A. Aguilar-Saavedra,^{126a,126f} M. Agustoni,¹⁷ S. P. Ahlen,²² F. Ahmadov,^{65,c} G. Aielli,^{134a,134b} H. Akerstedt,^{147a,147b} T. P. A. Åkesson,⁸¹ G. Akimoto,¹⁵⁶ A. V. Akimov,⁹⁶ G. L. Alberghi,^{20a,20b} J. Albert,¹⁷⁰ S. Albrand,⁵⁵ M. J. Alconada Verzini,⁷¹ M. Aleksa,³⁰ I. N. Aleksandrov,⁶⁵ C. Alexa,^{26a} G. Alexander,¹⁵⁴ G. Alexandre,⁴⁹ T. Alexopoulos,¹⁰ M. Alhroob,¹¹³ G. Alimonti,^{91a} L. Alio,⁸⁵ J. Alison,³¹ B. M. M. Allbrooke,¹⁸ L. J. Allison,⁷² P. P. Allport,⁷⁴

A. Aloisio,^{104a,104b} A. Alonso,³⁶ F. Alonso,⁷¹ C. Alpigiani,⁷⁶ A. Altheimer,³⁵ B. Alvarez Gonzalez,⁹⁰ M. G. Alviggi,^{104a,104b}
 K. Amako,⁶⁶ Y. Amaral Coutinho,^{24a} C. Amelung,²³ D. Amidei,⁸⁹ S. P. Amor Dos Santos,^{126a,126c} A. Amorim,^{126a,126b}
 S. Amoroso,⁴⁸ N. Amram,¹⁵⁴ G. Amundsen,²³ C. Anastopoulos,¹⁴⁰ L. S. Ancu,⁴⁹ N. Andari,³⁰ T. Andeen,³⁵ C. F. Anders,^{58b}
 G. Anders,³⁰ K. J. Anderson,³¹ A. Andreazza,^{91a,91b} V. Andrei,^{58a} X. S. Anduaga,⁷¹ S. Angelidakis,⁹ I. Angelozzi,¹⁰⁷
 P. Anger,⁴⁴ A. Angerami,³⁵ F. Anghinolfi,³⁰ A. V. Anisenkov,^{109,d} N. Anjos,¹² A. Annovi,⁴⁷ M. Antonelli,⁴⁷ A. Antonov,⁹⁸
 J. Antos,^{145b} F. Anulli,^{133a} M. Aoki,⁶⁶ L. Aperio Bella,¹⁸ G. Arabidze,⁹⁰ Y. Arai,⁶⁶ J. P. Araque,^{126a} A. T. H. Arce,⁴⁵
 F. A. Arduh,⁷¹ J-F. Arguin,⁹⁵ S. Argyropoulos,⁴² M. Arik,^{19a} A. J. Armbruster,³⁰ O. Arnaez,³⁰ V. Arnal,⁸² H. Arnold,⁴⁸
 M. Arratia,²⁸ O. Arslan,²¹ A. Artamonov,⁹⁷ G. Artoni,²³ S. Asai,¹⁵⁶ N. Asbah,⁴² A. Ashkenazi,¹⁵⁴ B. Åsman,^{147a,147b}
 L. Asquith,¹⁵⁰ K. Assamagan,²⁵ R. Aсталos,^{145a} M. Atkinson,¹⁶⁶ N. B. Atlay,¹⁴² B. Auerbach,⁶ K. Augsten,¹²⁸
 M. Auresseau,^{146b} G. Avolio,³⁰ B. Axen,¹⁵ G. Azelos,^{95,e} Y. Azuma,¹⁵⁶ M. A. Baak,³⁰ A. E. Baas,^{58a} C. Bacci,^{135a,135b}
 H. Bachacou,¹³⁷ K. Bachas,¹⁵⁵ M. Backes,³⁰ M. Backhaus,³⁰ E. Badescu,^{26a} P. Bagiacchi,^{133a,133b} P. Bagnaia,^{133a,133b}
 Y. Bai,^{33a} T. Bain,³⁵ J. T. Baines,¹³¹ O. K. Baker,¹⁷⁷ P. Balek,¹²⁹ F. Balli,⁸⁴ E. Banas,³⁹ Sw. Banerjee,¹⁷⁴ A. A. E. Bannoura,¹⁷⁶
 H. S. Bansil,¹⁸ L. Barak,¹⁷³ S. P. Baranov,⁹⁶ E. L. Barberio,⁸⁸ D. Barberis,^{50a,50b} M. Barbero,⁸⁵ T. Barillari,¹⁰¹
 M. Barisonzi,¹⁷⁶ T. Barklow,¹⁴⁴ N. Barlow,²⁸ S. L. Barnes,⁸⁴ B. M. Barnett,¹³¹ R. M. Barnett,¹⁵ Z. Barnovska,⁵
 A. Baroncelli,^{135a} G. Barone,⁴⁹ A. J. Barr,¹²⁰ F. Barreiro,⁸² J. Barreiro Guimarães da Costa,⁵⁷ R. Bartoldus,¹⁴⁴ A. E. Barton,⁷²
 P. Bartos,^{145a} V. Bartsch,¹⁵⁰ A. Bassalat,¹¹⁷ A. Basye,¹⁶⁶ R. L. Bates,⁵³ S. J. Batista,¹⁵⁹ J. R. Batley,²⁸ M. Battaglia,¹³⁸
 M. Battistin,³⁰ F. Bauer,¹³⁷ H. S. Bawa,^{144,f} J. B. Beacham,¹¹⁰ M. D. Beattie,⁷² T. Beau,⁸⁰ P. H. Beauchemin,¹⁶²
 R. Beccherle,^{124a,124b} P. Bechtel,²¹ H. P. Beck,^{17,g} K. Becker,¹²⁰ S. Becker,¹⁰⁰ M. Beckingham,¹⁷¹ C. Becot,¹¹⁷
 A. J. Beddall,^{19c} A. Beddall,^{19c} S. Bedikian,¹⁷⁷ V. A. Bednyakov,⁶⁵ C. P. Bee,¹⁴⁹ L. J. Beemster,¹⁰⁷ T. A. Beermann,¹⁷⁶
 M. Begel,²⁵ K. Behr,¹²⁰ C. Belanger-Champagne,⁸⁷ P. J. Bell,⁴⁹ W. H. Bell,⁴⁹ G. Bella,¹⁵⁴ L. Bellagamba,^{20a} A. Bellerive,²⁹
 M. Bellomo,⁸⁶ K. Belotskiy,⁹⁸ O. Beltramello,³⁰ O. Benary,¹⁵⁴ D. Bencheekroun,^{136a} K. Bendtz,^{147a,147b} N. Benekos,¹⁶⁶
 Y. Benhammou,¹⁵⁴ E. Benhar Noccioli,⁴⁹ J. A. Benitez Garcia,^{160b} D. P. Benjamin,⁴⁵ J. R. Bensinger,²³ S. Bentvelsen,¹⁰⁷
 D. Berge,¹⁰⁷ E. Bergeaas Kuutmann,¹⁶⁷ N. Berger,⁵ F. Berghaus,¹⁷⁰ J. Beringer,¹⁵ C. Bernard,²² N. R. Bernard,⁸⁶
 C. Bernius,¹¹⁰ F. U. Bernlochner,²¹ T. Berry,⁷⁷ P. Berta,¹²⁹ C. Bertella,⁸³ G. Bertoli,^{147a,147b} F. Bertolucci,^{124a,124b}
 C. Bertsche,¹¹³ D. Bertsche,¹¹³ M. I. Besana,^{91a} G. J. Besjes,¹⁰⁶ O. Bessidskaia,^{147a,147b} M. Bessner,⁴² N. Besson,¹³⁷
 C. Betancourt,⁴⁸ S. Bethke,¹⁰¹ A. J. Bevan,⁷⁶ W. Bhimji,⁴⁶ R. M. Bianchi,¹²⁵ L. Bianchini,²³ M. Bianco,³⁰ O. Biebel,¹⁰⁰
 S. P. Bieniek,⁷⁸ K. Bierwagen,⁵⁴ M. Biglietti,^{135a} J. Bilbao De Mendizabal,⁴⁹ H. Bilokon,⁴⁷ M. Bindi,⁵⁴ S. Binet,¹¹⁷
 A. Bingul,^{19c} C. Bini,^{133a,133b} C. W. Black,¹⁵¹ J. E. Black,¹⁴⁴ K. M. Black,²² D. Blackburn,¹³⁹ R. E. Blair,⁶ J.-B. Blanchard,¹³⁷
 T. Blazek,^{145a} I. Bloch,⁴² C. Blocker,²³ W. Blum,^{83,a} U. Blumenschein,⁵⁴ G. J. Bobbink,¹⁰⁷ V. S. Bobrovnikov,^{109,d}
 S. S. Bocchetta,⁸¹ A. Bocci,⁴⁵ C. Bock,¹⁰⁰ C. R. Boddy,¹²⁰ M. Boehler,⁴⁸ T. T. Boek,¹⁷⁶ J. A. Bogaerts,³⁰
 A. G. Bogdanchikov,¹⁰⁹ A. Bogouch,^{92,a} C. Bohm,^{147a} V. Boisvert,⁷⁷ T. Bold,^{38a} V. Boldea,^{26a} A. S. Boldyrev,⁹⁹
 M. Bomben,⁸⁰ M. Bona,⁷⁶ M. Boonekamp,¹³⁷ A. Borisov,¹³⁰ G. Borissoy,⁷² S. Borroni,⁴² J. Bortfeldt,¹⁰⁰ V. Bortolotto,^{60a}
 K. Bos,¹⁰⁷ D. Boscherini,^{20a} M. Bosman,¹² H. Boterenbrood,¹⁰⁷ J. Boudreau,¹²⁵ J. Bouffard,² E. V. Bouhova-Thacker,⁷²
 D. Boumediene,³⁴ C. Bourdarios,¹¹⁷ N. Bousson,¹¹⁴ S. Boutouil,^{136d} A. Boveia,³¹ J. Boyd,³⁰ I. R. Boyko,⁶⁵ I. Bozic,^{13a}
 J. Bracinik,¹⁸ A. Brandt,⁸ G. Brandt,¹⁵ O. Brandt,^{58a} U. Bratzler,¹⁵⁷ B. Brau,⁸⁶ J. E. Brau,¹¹⁶ H. M. Braun,^{176,a}
 S. F. Brazzale,^{165a,165c} B. Brelier,¹⁵⁹ K. Brendlinger,¹²² A. J. Brennan,⁸⁸ R. Brenner,¹⁶⁷ S. Bressler,¹⁷³ K. Bristow,^{146c}
 T. M. Bristow,⁴⁶ D. Britton,⁵³ F. M. Brochu,²⁸ I. Brock,²¹ R. Brock,⁹⁰ J. Bronner,¹⁰¹ G. Brooijmans,³⁵ T. Brooks,⁷⁷
 W. K. Brooks,^{32b} J. Brosamer,¹⁵ E. Brost,¹¹⁶ J. Brown,⁵⁵ P. A. Bruckman de Renstrom,³⁹ D. Bruncko,^{145b} R. Bruneliere,⁴⁸
 S. Brunet,⁶¹ A. Bruni,^{20a} G. Bruni,^{20a} M. Bruschi,^{20a} L. Bryngemark,⁸¹ T. Buanes,¹⁴ Q. Buat,¹⁴³ F. Bucci,⁴⁹ P. Buchholz,¹⁴²
 A. G. Buckley,⁵³ S. I. Buda,^{26a} I. A. Budagov,⁶⁵ F. Buehrer,⁴⁸ L. Bugge,¹¹⁹ M. K. Bugge,¹¹⁹ O. Bulekov,⁹⁸ A. C. Bundock,⁷⁴
 H. Burkhardt,³⁰ S. Burdın,⁷⁴ B. Burghgrave,¹⁰⁸ S. Burke,¹³¹ I. Burmeister,⁴³ E. Busato,³⁴ D. Büscher,⁴⁸ V. Büscher,⁸³
 P. Bussey,⁵³ C. P. Buszello,¹⁶⁷ B. Butler,⁵⁷ J. M. Butler,²² A. I. Butt,³ C. M. Buttar,⁵³ J. M. Butterworth,⁷⁸ P. Butti,¹⁰⁷
 W. Buttinger,²⁸ A. Buzatu,⁵³ M. Byszewski,¹⁰ S. Cabrera Urbán,¹⁶⁸ D. Caforio,^{20a,20b} O. Cakir,^{4a} P. Calafiura,¹⁵
 A. Calandri,¹³⁷ G. Calderini,⁸⁰ P. Calfayan,¹⁰⁰ L. P. Caloba,^{24a} D. Calvet,³⁴ S. Calvet,³⁴ R. Camacho Toro,⁴⁹ S. Camarda,⁴²
 D. Cameron,¹¹⁹ L. M. Caminada,¹⁵ R. Caminal Armadans,¹² S. Campana,³⁰ M. Campanelli,⁷⁸ A. Campoverde,¹⁴⁹
 V. Canale,^{104a,104b} A. Canepa,^{160a} M. Cano Bret,⁷⁶ J. Cantero,⁸² R. Cantrill,^{126a} T. Cao,⁴⁰ M. D. M. Capeans Garrido,³⁰
 I. Caprini,^{26a} M. Caprini,^{26a} M. Capua,^{37a,37b} R. Caputo,⁸³ R. Cardarelli,^{134a} T. Carli,³⁰ G. Carlino,^{104a} L. Carminati,^{91a,91b}
 S. Caron,¹⁰⁶ E. Carquin,^{32a} G. D. Carrillo-Montoya,^{146c} J. R. Carter,²⁸ J. Carvalho,^{126a,126c} D. Casadei,⁷⁸ M. P. Casado,¹²
 M. Casolino,¹² E. Castaneda-Miranda,^{146b} A. Castelli,¹⁰⁷ V. Castillo Gimenez,¹⁶⁸ N. F. Castro,^{126a} P. Catastini,⁵⁷

A. Catinaccio,³⁰ J. R. Catmore,¹¹⁹ A. Cattai,³⁰ G. Cattani,^{134a,134b} J. Caudron,⁸³ V. Cavaliere,¹⁶⁶ D. Cavalli,^{91a} M. Cavalli-Sforza,¹² V. Cavalinni,^{124a,124b} F. Ceradini,^{135a,135b} B. C. Cerio,⁴⁵ K. Cerny,¹²⁹ A. S. Cerqueira,^{24b} A. Cerri,¹⁵⁰ L. Cerrito,⁷⁶ F. Cerutti,¹⁵ M. Cerv,³⁰ A. Cervelli,¹⁷ S. A. Cetin,^{19b} A. Chafaq,^{136a} D. Chakraborty,¹⁰⁸ I. Chalupkova,¹²⁹ P. Chang,¹⁶⁶ B. Chapleau,⁸⁷ J. D. Chapman,²⁸ D. Charfeddine,¹¹⁷ D. G. Charlton,¹⁸ C. C. Chau,¹⁵⁹ C. A. Chavez Barajas,¹⁵⁰ S. Cheatham,¹⁵³ A. Chegwidan,⁹⁰ S. Chekanov,⁶ S. V. Chekulaev,^{160a} G. A. Chelkov,^{65,h} M. A. Chelstowska,⁸⁹ C. Chen,⁶⁴ H. Chen,²⁵ K. Chen,¹⁴⁹ L. Chen,^{33d,i} S. Chen,^{33c} X. Chen,^{33f} Y. Chen,⁶⁷ H. C. Cheng,⁸⁹ Y. Cheng,³¹ A. Cheplakov,⁶⁵ E. Cheremushkina,¹³⁰ R. Cherkaoui El Moursli,^{136e} V. Chernyatin,^{25,a} E. Cheu,⁷ L. Chevalier,¹³⁷ V. Chiarella,⁴⁷ G. Chiefari,^{104a,104b} J. T. Childers,⁶ A. Chilingarov,⁷² G. Chiodini,^{73a} A. S. Chisholm,¹⁸ R. T. Chislett,⁷⁸ A. Chitan,^{26a} M. V. Chizhov,⁶⁵ S. Chouridou,⁹ B. K. B. Chow,¹⁰⁰ D. Chromek-Burckhart,³⁰ M. L. Chu,¹⁵² J. Chudoba,¹²⁷ J. J. Chwastowski,³⁹ L. Chytka,¹¹⁵ G. Ciapetti,^{133a,133b} A. K. Ciftci,^{4a} R. Ciftci,^{4a} D. Cinca,⁵³ V. Cindro,⁷⁵ A. Ciocio,¹⁵ Z. H. Citron,¹⁷³ M. Citterio,^{91a} M. Ciubancan,^{26a} A. Clark,⁴⁹ P. J. Clark,⁴⁶ R. N. Clarke,¹⁵ W. Cleland,¹²⁵ J. C. Clemens,⁸⁵ C. Clement,^{147a,147b} Y. Coadou,⁸⁵ M. Cobal,^{165a,165c} A. Coccaro,¹³⁹ J. Cochran,⁶⁴ L. Coffey,²³ J. G. Cogan,¹⁴⁴ B. Cole,³⁵ S. Cole,¹⁰⁸ A. P. Colijn,¹⁰⁷ J. Collot,⁵⁵ T. Colombo,^{58c} G. Compostella,¹⁰¹ P. Conde Muiño,^{126a,126b} E. Coniavitis,⁴⁸ S. H. Connell,^{146b} I. A. Connelly,⁷⁷ S. M. Consonni,^{91a,91b} V. Consorti,⁴⁸ S. Constantinescu,^{26a} C. Conta,^{121a,121b} G. Conti,⁵⁷ F. Conventi,^{104a,j} M. Cooke,¹⁵ B. D. Cooper,⁷⁸ A. M. Cooper-Sarkar,¹²⁰ N. J. Cooper-Smith,⁷⁷ K. Copic,¹⁵ T. Cornelissen,¹⁷⁶ M. Corradi,^{20a} F. Corriveau,^{87,k} A. Corso-Radu,¹⁶⁴ A. Cortes-Gonzalez,¹² G. Cortiana,¹⁰¹ G. Costa,^{91a} M. J. Costa,¹⁶⁸ D. Costanzo,¹⁴⁰ D. Côté,⁸ G. Cottin,²⁸ G. Cowan,⁷⁷ B. E. Cox,⁸⁴ K. Cranmer,¹¹⁰ G. Cree,²⁹ S. Crépe-Renaudin,⁵⁵ F. Crescioli,⁸⁰ W. A. Cribbs,^{147a,147b} M. Crispin Ortuzar,¹²⁰ M. Cristinziani,²¹ V. Croft,¹⁰⁶ G. Crosetti,^{37a,37b} T. Cuhadar Donszelmann,¹⁴⁰ J. Cummings,¹⁷⁷ M. Curatolo,⁴⁷ C. Cuthbert,¹⁵¹ H. Czirr,¹⁴² P. Czodrowski,³ S. D'Auria,⁵³ M. D'Onofrio,⁷⁴ M. J. Da Cunha Sargedas De Sousa,^{126a,126b} C. Da Via,⁸⁴ W. Dabrowski,^{38a} A. Dafinca,¹²⁰ T. Dai,⁸⁹ O. Dale,¹⁴ F. Dallaire,⁹⁵ C. Dallapiccola,⁸⁶ M. Dam,³⁶ A. C. Daniells,¹⁸ M. Danninger,¹⁶⁹ M. Dano Hoffmann,¹³⁷ V. Dao,⁴⁸ G. Darbo,^{50a} S. Darmora,⁸ J. Dassoulas,⁷⁴ A. Dattagupta,⁶¹ W. Davey,²¹ C. David,¹⁷⁰ T. Davidek,¹²⁹ E. Davies,^{120,l} M. Davies,¹⁵⁴ O. Davignon,⁸⁰ A. R. Davison,⁷⁸ P. Davison,⁷⁸ Y. Davygora,^{58a} E. Dawe,¹⁴³ I. Dawson,¹⁴⁰ R. K. Daya-Ishmukhametova,⁸⁶ K. De,⁸ R. de Asmundis,^{104a} S. De Castro,^{20a,20b} S. De Cecco,⁸⁰ N. De Groot,¹⁰⁶ P. de Jong,¹⁰⁷ H. De la Torre,⁸² F. De Lorenzi,⁶⁴ L. De Nooij,¹⁰⁷ D. De Pedis,^{133a} A. De Salvo,^{133a} U. De Sanctis,¹⁵⁰ A. De Santo,¹⁵⁰ J. B. De Vivie De Regie,¹¹⁷ W. J. Dearnaley,⁷² R. Debbé,²⁵ C. Debenedetti,¹³⁸ B. Dechenaux,⁵⁵ D. V. Dedovich,⁶⁵ I. Deigaard,¹⁰⁷ J. Del Peso,⁸² T. Del Prete,^{124a,124b} F. Deliot,¹³⁷ C. M. Delitzsch,⁴⁹ M. Deliyergiyev,⁷⁵ A. Dell'Acqua,³⁰ L. Dell'Asta,²² M. Dell'Orso,^{124a,124b} M. Della Pietra,^{104a,j} D. della Volpe,⁴⁹ M. Delmastro,⁵ P. A. Delsart,⁵⁵ C. Deluca,¹⁰⁷ D. A. DeMarco,¹⁵⁹ S. Demers,¹⁷⁷ M. Demichev,⁶⁵ A. Demilly,⁸⁰ S. P. Denisov,¹³⁰ D. Derendarz,³⁹ J. E. Derkaoui,^{136d} F. Derue,⁸⁰ P. Dervan,⁷⁴ K. Desch,²¹ C. Deterre,⁴² P. O. Deviveiros,³⁰ A. Dewhurst,¹³¹ S. Dhaliwal,¹⁰⁷ A. Di Ciaccio,^{134a,134b} L. Di Ciaccio,⁵ A. Di Domenico,^{133a,133b} C. Di Donato,^{104a,104b} A. Di Girolamo,³⁰ B. Di Girolamo,³⁰ A. Di Mattia,¹⁵³ B. Di Micco,^{135a,135b} R. Di Nardo,⁴⁷ A. Di Simone,⁴⁸ R. Di Sipio,^{20a,20b} D. Di Valentino,²⁹ F. A. Dias,⁴⁶ M. A. Diaz,^{32a} E. B. Diehl,⁸⁹ J. Dietrich,¹⁶ T. A. Dietzsch,^{58a} S. Diglio,⁸⁵ A. Dimitrievska,^{13a} J. Dingfelder,²¹ P. Dita,^{26a} S. Dita,^{26a} F. Dittus,³⁰ F. Djama,⁸⁵ T. Djobava,^{51b} J. I. Djuvsland,^{58a} M. A. B. do Vale,^{24c} D. Dobos,³⁰ C. Doglioni,⁴⁹ T. Doherty,⁵³ T. Dohmae,¹⁵⁶ J. Dolejsi,¹²⁹ Z. Dolezal,¹²⁹ B. A. Dolgoshein,^{98,a} M. Donadelli,^{24d} S. Donati,^{124a,124b} P. Dondero,^{121a,121b} J. Donini,³⁴ J. Dopke,¹³¹ A. Doria,^{104a} M. T. Dova,⁷¹ A. T. Doyle,⁵³ M. Dris,¹⁰ J. Dubbert,⁸⁹ S. Dube,¹⁵ E. Dubreuil,³⁴ E. Duchovni,¹⁷³ G. Duckeck,¹⁰⁰ O. A. Ducu,^{26a} D. Duda,¹⁷⁶ A. Dudarev,³⁰ F. Dudziak,⁶⁴ L. Duflost,¹¹⁷ L. Duguid,⁷⁷ M. Dührssen,³⁰ M. Dunford,^{58a} H. Duran Yildiz,^{4a} M. Düren,⁵² A. Durglishvili,^{51b} D. Duschinger,⁴⁴ M. Dwuznik,^{38a} M. Dyndal,^{38a} W. Edson,² N. C. Edwards,⁴⁶ W. Ehrenfeld,²¹ T. Eifert,³⁰ G. Eigen,¹⁴ K. Einsweiler,¹⁵ T. Ekelof,¹⁶⁷ M. El Kacimi,^{136c} M. Ellert,¹⁶⁷ S. Elles,⁵ F. Ellinghaus,⁸³ A. A. Elliot,¹⁷⁰ N. Ellis,³⁰ J. Elmsheuser,¹⁰⁰ M. Elsing,³⁰ D. Emelianov,¹³¹ Y. Enari,¹⁵⁶ O. C. Endner,⁸³ M. Endo,¹¹⁸ R. Engelmann,¹⁴⁹ J. Erdmann,⁴³ A. Ereditato,¹⁷ D. Eriksson,^{147a} G. Ernis,¹⁷⁶ J. Ernst,² M. Ernst,²⁵ J. Ernwein,¹³⁷ S. Errede,¹⁶⁶ E. Ertel,⁸³ M. Escalier,¹¹⁷ H. Esch,⁴³ C. Escobar,¹²⁵ B. Esposito,⁴⁷ A. I. Etiennevire,¹³⁷ E. Etzion,¹⁵⁴ H. Evans,⁶¹ A. Ezhilov,¹²³ L. Fabbri,^{20a,20b} G. Facini,³¹ R. M. Fakhruudinov,¹³⁰ S. Falciano,^{133a} R. J. Falla,⁷⁸ J. Faltova,¹²⁹ Y. Fang,^{33a} M. Fanti,^{91a,91b} A. Farbin,⁸ A. Farilla,^{135a} T. Farooque,¹² S. Farrell,¹⁵ S. M. Farrington,¹⁷¹ P. Farthouat,³⁰ F. Fassi,^{136e} P. Fassnacht,³⁰ D. Fassouliotis,⁹ A. Favareto,^{50a,50b} L. Fayard,¹¹⁷ P. Federic,^{145a} O. L. Fedin,^{123,m} W. Fedorko,¹⁶⁹ S. Feigl,³⁰ L. Felgioni,⁸⁵ C. Feng,^{33d} E. J. Feng,⁶ H. Feng,⁸⁹ A. B. Fenyuk,¹³⁰ P. Fernandez Martinez,¹⁶⁸ S. Fernandez Perez,³⁰ S. Ferrag,⁵³ J. Ferrando,⁵³ A. Ferrari,¹⁶⁷ P. Ferrari,¹⁰⁷ R. Ferrari,^{121a} D. E. Ferreira de Lima,⁵³ A. Ferrer,¹⁶⁸ D. Ferrere,⁴⁹ C. Ferretti,⁸⁹ A. Ferretto Parodi,^{50a,50b} M. Fiascaris,³¹ F. Fiedler,⁸³ A. Filipčić,⁷⁵ M. Filipuzzi,⁴² F. Filthaut,¹⁰⁶ M. Fincke-Keeler,¹⁷⁰

K. D. Finelli,¹⁵¹ M. C. N. Fiolhais,^{126a,126c} L. Fiorini,¹⁶⁸ A. Firan,⁴⁰ A. Fischer,² J. Fischer,¹⁷⁶ W. C. Fisher,⁹⁰
 E. A. Fitzgerald,²³ M. Flechl,⁴⁸ I. Fleck,¹⁴² P. Fleischmann,⁸⁹ S. Fleischmann,¹⁷⁶ G. T. Fletcher,¹⁴⁰ G. Fletcher,⁷⁶ T. Flick,¹⁷⁶
 A. Floderus,⁸¹ L. R. Flores Castillo,^{60a} M. J. Flowerdew,¹⁰¹ A. Formica,¹³⁷ A. Forti,⁸⁴ D. Fournier,¹¹⁷ H. Fox,⁷² S. Fracchia,¹²
 P. Francavilla,⁸⁰ M. Franchini,^{20a,20b} S. Franchino,³⁰ D. Francis,³⁰ L. Franconi,¹¹⁹ M. Franklin,⁵⁷ M. Fraternali,^{121a,121b}
 S. T. French,²⁸ C. Friedrich,⁴² F. Friedrich,⁴⁴ D. Froidevaux,³⁰ J. A. Frost,¹²⁰ C. Fukunaga,¹⁵⁷ E. Fullana Torregrosa,⁸³
 B. G. Fulsom,¹⁴⁴ J. Fuster,¹⁶⁸ C. Gabaldon,⁵⁵ O. Gabizon,¹⁷⁶ A. Gabrielli,^{20a,20b} A. Gabrielli,^{133a,133b} S. Gadatsch,¹⁰⁷
 S. Gadowski,⁴⁹ G. Gagliardi,^{50a,50b} P. Gagnon,⁶¹ C. Galea,¹⁰⁶ B. Galhardo,^{126a,126c} E. J. Gallas,¹²⁰ B. J. Gallop,¹³¹
 P. Gallus,¹²⁸ G. Galster,³⁶ K. K. Gan,¹¹¹ J. Gao,^{33b,i} Y. S. Gao,^{144,f} F. M. Garay Walls,⁴⁶ F. Garbersson,¹⁷⁷ C. García,¹⁶⁸
 J. E. García Navarro,¹⁶⁸ M. Garcia-Sciveres,¹⁵ R. W. Gardner,³¹ N. Garelli,¹⁴⁴ V. Garonne,³⁰ C. Gatti,⁴⁷ G. Gaudio,^{121a}
 B. Gaur,¹⁴² L. Gauthier,⁹⁵ P. Gauzzi,^{133a,133b} I. L. Gavrilenko,⁹⁶ C. Gay,¹⁶⁹ G. Gaycken,²¹ E. N. Gazis,¹⁰ P. Ge,^{33d} Z. Gece,¹⁶⁹
 C. N. P. Gee,¹³¹ D. A. A. Geerts,¹⁰⁷ Ch. Geich-Gimbel,²¹ K. Gellerstedt,^{147a,147b} C. Gemme,^{50a} A. Gemmell,⁵³
 M. H. Genest,⁵⁵ S. Gentile,^{133a,133b} M. George,⁵⁴ S. George,⁷⁷ D. Gerbaudo,¹⁶⁴ A. Gershon,¹⁵⁴ H. Ghazlane,^{136b}
 N. Ghodbane,³⁴ B. Giacobbe,^{20a} S. Giagu,^{133a,133b} V. Giangiobbe,¹² P. Giannetti,^{124a,124b} F. Gianotti,³⁰ B. Gibbard,²⁵
 S. M. Gibson,⁷⁷ M. Gilchriese,¹⁵ T. P. S. Gillam,²⁸ D. Gillberg,³⁰ G. Gilles,³⁴ D. M. Gingrich,^{3,e} N. Giokaris,⁹
 M. P. Giordani,^{165a,165c} R. Giordano,^{104a,104b} F. M. Giorgi,^{20a} F. M. Giorgi,¹⁶ P. F. Giraud,¹³⁷ D. Giugni,^{91a} C. Giuliani,⁴⁸
 M. Giulini,^{58b} B. K. Gjelsten,¹¹⁹ S. Gkaitatzis,¹⁵⁵ I. Gkialas,¹⁵⁵ E. L. Gkoukousis,¹¹⁷ L. K. Gladilin,⁹⁹ C. Glasman,⁸²
 J. Glatzer,³⁰ P. C. F. Glaysher,⁴⁶ A. Glazov,⁴² G. L. Glonti,⁶² M. Goblirsch-Kolb,¹⁰¹ J. R. Goddard,⁷⁶ J. Godlewski,³⁰
 S. Goldfarb,⁸⁹ T. Golling,⁴⁹ D. Golubkov,¹³⁰ A. Gomes,^{126a,126b,126d} L. S. Gomez Fajardo,⁴² R. Gonçalo,^{126a}
 J. Goncalves Pinto Firmino Da Costa,¹³⁷ L. Gonella,²¹ S. González de la Hoz,¹⁶⁸ G. Gonzalez Parra,¹² S. Gonzalez-Sevilla,⁴⁹
 L. Goossens,³⁰ P. A. Gorbounov,⁹⁷ H. A. Gordon,²⁵ I. Gorelov,¹⁰⁵ B. Gorini,³⁰ E. Gorini,^{73a,73b} A. Gorišek,⁷⁵ E. Gornicki,³⁹
 A. T. Goshaw,⁴⁵ C. Gössling,⁴³ M. I. Gostkin,⁶⁵ M. Gouighri,^{136a} D. Goujdami,^{136c} M. P. Goulette,⁴⁹ A. G. Goussiou,¹³⁹
 C. Goy,⁵ H. M. X. Grabas,¹³⁸ L. Graber,⁵⁴ I. Grabowska-Bold,^{38a} P. Grafström,^{20a,20b} K.-J. Grahn,⁴² J. Gramling,⁴⁹
 E. Gramstad,¹¹⁹ S. Grancagnolo,¹⁶ V. Grassi,¹⁴⁹ V. Gratchev,¹²³ H. M. Gray,³⁰ E. Graziani,^{135a} O. G. Grebenyuk,¹²³
 Z. D. Greenwood,^{79,n} K. Gregersen,⁷⁸ I. M. Gregor,⁴² P. Grenier,¹⁴⁴ J. Griffiths,⁸ A. A. Grillo,¹³⁸ K. Grimm,⁷²
 S. Grinstein,^{12,o} Ph. Gris,³⁴ Y. V. Grishkevich,⁹⁹ J.-F. Grivaz,¹¹⁷ J. P. Grohs,⁴⁴ A. Grohsjean,⁴² E. Gross,¹⁷³
 J. Grosse-Knetter,⁵⁴ G. C. Grossi,^{134a,134b} Z. J. Grout,¹⁵⁰ L. Guan,^{33b} J. Guenther,¹²⁸ F. Guescini,⁴⁹ D. Guest,¹⁷⁷ O. Gueta,¹⁵⁴
 C. Guicheney,³⁴ E. Guido,^{50a,50b} T. Guillemin,¹¹⁷ S. Guindon,² U. Gul,⁵³ C. Gumpert,⁴⁴ J. Guo,³⁵ S. Gupta,¹²⁰ P. Gutierrez,¹¹³
 N. G. Gutierrez Ortiz,⁵³ C. Gutsche,⁷⁸ N. Guttman,¹⁵⁴ C. Guyot,¹³⁷ C. Gwenlan,¹²⁰ C. B. Gwilliam,⁷⁴ A. Haas,¹¹⁰
 C. Haber,¹⁵ H. K. Hadavand,⁸ N. Haddad,^{136e} P. Haefner,²¹ S. Hageböck,²¹ Z. Hajduk,³⁹ H. Hakobyan,¹⁷⁸ M. Haleem,⁴²
 D. Hall,¹²⁰ G. Halladjian,⁹⁰ G. D. Hallewell,⁸⁵ K. Hamacher,¹⁷⁶ P. Hamal,¹¹⁵ K. Hamano,¹⁷⁰ M. Hamer,⁵⁴ A. Hamilton,^{146a}
 S. Hamilton,¹⁶² G. N. Hamity,^{146c} P. G. Hamnett,⁴² L. Han,^{33b} K. Hanagaki,¹¹⁸ K. Hanawa,¹⁵⁶ M. Hance,¹⁵ P. Hanke,^{58a}
 R. Hanna,¹³⁷ J. B. Hansen,³⁶ J. D. Hansen,³⁶ P. H. Hansen,³⁶ K. Hara,¹⁶¹ A. S. Hard,¹⁷⁴ T. Harenberg,¹⁷⁶ F. Hariri,¹¹⁷
 S. Harkusha,⁹² R. D. Harrington,⁴⁶ P. F. Harrison,¹⁷¹ F. Hartjes,¹⁰⁷ M. Hasegawa,⁶⁷ S. Hasegawa,¹⁰³ Y. Hasegawa,¹⁴¹
 A. Hasib,¹¹³ S. Hassani,¹³⁷ S. Haug,¹⁷ M. Hauschild,³⁰ R. Hauser,⁹⁰ M. Havranek,¹²⁷ C. M. Hawkes,¹⁸ R. J. Hawkings,³⁰
 A. D. Hawkins,⁸¹ T. Hayashi,¹⁶¹ D. Hayden,⁹⁰ C. P. Hays,¹²⁰ J. M. Hays,⁷⁶ H. S. Hayward,⁷⁴ S. J. Haywood,¹³¹ S. J. Head,¹⁸
 T. Heck,⁸³ V. Hedberg,⁸¹ L. Heelan,⁸ S. Heim,¹²² T. Heim,¹⁷⁶ B. Heinemann,¹⁵ L. Heinrich,¹¹⁰ J. Hejbal,¹²⁷ L. Helary,²²
 M. Heller,³⁰ S. Hellman,^{147a,147b} D. Hellmich,²¹ C. Helsens,³⁰ J. Henderson,¹²⁰ R. C. W. Henderson,⁷² Y. Heng,¹⁷⁴
 C. Hengler,⁴² A. Henrichs,¹⁷⁷ A. M. Henriques Correia,³⁰ S. Henrot-Versille,¹¹⁷ G. H. Herbert,¹⁶ Y. Hernández Jiménez,¹⁶⁸
 R. Herrberg-Schubert,¹⁶ G. Herten,⁴⁸ R. Hertenberger,¹⁰⁰ L. Hervas,³⁰ G. G. Hesketh,⁷⁸ N. P. Hessey,¹⁰⁷ R. Hickling,⁷⁶
 E. Higón-Rodríguez,¹⁶⁸ E. Hill,¹⁷⁰ J. C. Hill,²⁸ K. H. Hiller,⁴² S. J. Hillier,¹⁸ I. Hinchliffe,¹⁵ E. Hines,¹²² M. Hirose,¹⁵⁸
 D. Hirschbuehl,¹⁷⁶ J. Hobbs,¹⁴⁹ N. Hod,¹⁰⁷ M. C. Hodgkinson,¹⁴⁰ P. Hodgson,¹⁴⁰ A. Hoecker,³⁰ M. R. Hoferkamp,¹⁰⁵
 F. Hoenig,¹⁰⁰ D. Hoffmann,⁸⁵ M. Hohlfeld,⁸³ T. R. Holmes,¹⁵ T. M. Hong,¹²² L. Hooft van Huysduynen,¹¹⁰
 W. H. Hopkins,¹¹⁶ Y. Horii,¹⁰³ A. J. Horton,¹⁴³ J.-Y. Hostachy,⁵⁵ S. Hou,¹⁵² A. Hoummada,^{136a} J. Howard,¹²⁰ J. Howarth,⁴²
 M. Hrabovsky,¹¹⁵ I. Hristova,¹⁶ J. Hrivnac,¹¹⁷ T. Hryn'ova,⁵ A. Hrynevich,⁹³ C. Hsu,^{146c} P. J. Hsu,¹⁵² S.-C. Hsu,¹³⁹ D. Hu,³⁵
 X. Hu,⁸⁹ Y. Huang,⁴² Z. Hubacek,³⁰ F. Hubaut,⁸⁵ F. Huegging,²¹ T. B. Huffman,¹²⁰ E. W. Hughes,³⁵ G. Hughes,⁷²
 M. Huhtinen,³⁰ T. A. Hülsing,⁸³ M. Hurwitz,¹⁵ N. Huseynov,^{65,c} J. Huston,⁹⁰ J. Huth,⁵⁷ G. Iacobucci,⁴⁹ G. Iakovidis,¹⁰
 I. Ibragimov,¹⁴² L. Iconomidou-Fayard,¹¹⁷ E. Ideal,¹⁷⁷ Z. Idrissi,^{136e} P. Ingo,^{104a} O. Igonkina,¹⁰⁷ T. Iizawa,¹⁷² Y. Ikegami,⁶⁶
 K. Ikematsu,¹⁴² M. Ikeno,⁶⁶ Y. Ilchenko,^{31,p} D. Iliadis,¹⁵⁵ N. Ilic,¹⁵⁹ Y. Inamaru,⁶⁷ T. Ince,¹⁰¹ P. Ioannou,⁹ M. Iodice,^{135a}
 K. Iordanidou,⁹ V. Ippolito,⁵⁷ A. Irls Quiles,¹⁶⁸ C. Isaksson,¹⁶⁷ M. Ishino,⁶⁸ M. Ishitsuka,¹⁵⁸ R. Ishmukhametov,¹¹¹

C. Issever,¹²⁰ S. Istin,^{19a} J. M. Iturbe Ponce,⁸⁴ R. Iuppa,^{134a,134b} J. Ivarsson,⁸¹ W. Iwanski,³⁹ H. Iwasaki,⁶⁶ J. M. Izen,⁴¹ V. Izzo,^{104a} B. Jackson,¹²² M. Jackson,⁷⁴ P. Jackson,¹ M. R. Jaekel,³⁰ V. Jain,² K. Jakobs,⁴⁸ S. Jakobsen,³⁰ T. Jakoubek,¹²⁷ J. Jakubek,¹²⁸ D. O. Jamin,¹⁵² D. K. Jana,⁷⁹ E. Jansen,⁷⁸ J. Janssen,²¹ M. Janus,¹⁷¹ G. Jarlskog,⁸¹ N. Javadov,^{65,c} T. Javůrek,⁴⁸ L. Jeanty,¹⁵ J. Jejelava,^{51a,q} G.-Y. Jeng,¹⁵¹ D. Jennens,⁸⁸ P. Jenni,^{48,r} J. Jentzsch,⁴³ C. Jeske,¹⁷¹ S. Jézéquel,⁵ H. Ji,¹⁷⁴ J. Jia,¹⁴⁹ Y. Jiang,^{33b} M. Jimenez Belenguer,⁴² S. Jin,^{33a} A. Jinaru,^{26a} O. Jinnouchi,¹⁵⁸ M. D. Joergensen,³⁶ P. Johansson,¹⁴⁰ K. A. Johns,⁷ K. Jon-And,^{147a,147b} G. Jones,¹⁷¹ R. W. L. Jones,⁷² T. J. Jones,⁷⁴ J. Jongmanns,^{58a} P. M. Jorge,^{126a,126b} K. D. Joshi,⁸⁴ J. Jovicevic,¹⁴⁸ X. Ju,¹⁷⁴ C. A. Jung,⁴³ P. Jussel,⁶² A. Juste Rozas,^{12,o} M. Kaci,¹⁶⁸ A. Kaczmarska,³⁹ M. Kado,¹¹⁷ H. Kagan,¹¹¹ M. Kagan,¹⁴⁴ E. Kajomovitz,⁴⁵ C. W. Kalderon,¹²⁰ S. Kama,⁴⁰ A. Kamenshchikov,¹³⁰ N. Kanaya,¹⁵⁶ M. Kaneda,³⁰ S. Kaneti,²⁸ V. A. Kantserov,⁹⁸ J. Kanzaki,⁶⁶ B. Kaplan,¹¹⁰ A. Kapliy,³¹ D. Kar,⁵³ K. Karakostas,¹⁰ A. Karamaoun,³ N. Karastathis,¹⁰ M. J. Kareem,⁵⁴ M. Karnevskiy,⁸³ S. N. Karpov,⁶⁵ Z. M. Karpova,⁶⁵ K. Karthik,¹¹⁰ V. Kartvelishvili,⁷² A. N. Karyukhin,¹³⁰ L. Kashif,¹⁷⁴ G. Kasieczka,^{58b} R. D. Kass,¹¹¹ A. Kastanas,¹⁴ Y. Kataoka,¹⁵⁶ A. Katre,⁴⁹ J. Katzy,⁴² V. Kaushik,⁷ K. Kawagoe,⁷⁰ T. Kawamoto,¹⁵⁶ G. Kawamura,⁵⁴ S. Kazama,¹⁵⁶ V. F. Kazanin,¹⁰⁹ M. Y. Kazarinov,⁶⁵ R. Keeler,¹⁷⁰ R. Kehoe,⁴⁰ M. Keil,⁵⁴ J. S. Keller,⁴² J. J. Kempster,⁷⁷ H. Keoshkerian,⁵ O. Kepka,¹²⁷ B. P. Kerševan,⁷⁵ S. Kersten,¹⁷⁶ K. Kessoku,¹⁵⁶ J. Keung,¹⁵⁹ R. A. Keyes,⁸⁷ F. Khalil-zada,¹¹ H. Khandanyan,^{147a,147b} A. Khanov,¹¹⁴ A. Kharlamov,¹⁰⁹ A. Khodinov,⁹⁸ A. Khomich,^{58a} T. J. Khoo,²⁸ G. Khoraiuli,²¹ V. Khovanskii,⁹⁷ E. Khramov,⁶⁵ J. Khubua,^{51b} H. Y. Kim,⁸ H. Kim,^{147a,147b} S. H. Kim,¹⁶¹ N. Kimura,¹⁵⁵ O. Kind,¹⁶ B. T. King,⁷⁴ M. King,¹⁶⁸ R. S. B. King,¹²⁰ S. B. King,¹⁶⁹ J. Kirk,¹³¹ A. E. Kiryunin,¹⁰¹ T. Kishimoto,⁶⁷ D. Kisielewska,^{38a} F. Kiss,⁴⁸ K. Kiuchi,¹⁶¹ E. Kladiva,^{145b} M. Klein,⁷⁴ U. Klein,⁷⁴ K. Kleinknecht,⁸³ P. Klimek,^{147a,147b} A. Klimentov,²⁵ R. Klingenberg,⁴³ J. A. Klinger,⁸⁴ T. Klioutchnikova,³⁰ P. F. Klok,¹⁰⁶ E.-E. Kluge,^{58a} P. Kluit,¹⁰⁷ S. Kluth,¹⁰¹ E. Kneringer,⁶² E. B. F. G. Knoops,⁸⁵ A. Knue,⁵³ D. Kobayashi,¹⁵⁸ T. Kobayashi,¹⁵⁶ M. Kobel,⁴⁴ M. Kocian,¹⁴⁴ P. Kodys,¹²⁹ T. Koffas,²⁹ E. Koffeman,¹⁰⁷ L. A. Kogan,¹²⁰ S. Kohlmann,¹⁷⁶ Z. Kohout,¹²⁸ T. Kohriki,⁶⁶ T. Koi,¹⁴⁴ H. Kolanoski,¹⁶ I. Koletsou,⁵ J. Koll,⁹⁰ A. A. Komar,^{96,a} Y. Komori,¹⁵⁶ T. Kondo,⁶⁶ N. Kondrashova,⁴² K. Köneke,⁴⁸ A. C. König,¹⁰⁶ S. König,⁸³ T. Kono,^{66,s} R. Konoplich,^{110,t} N. Konstantinidis,⁷⁸ R. Kopeliansky,¹⁵³ S. Koperny,^{38a} L. Köpke,⁸³ A. K. Kopp,⁴⁸ K. Korcyl,³⁹ K. Kordas,¹⁵⁵ A. Korn,⁷⁸ A. A. Korol,^{109,d} I. Korolkov,¹² E. V. Korolkova,¹⁴⁰ V. A. Korotkov,¹³⁰ O. Kortner,¹⁰¹ S. Kortner,¹⁰¹ V. V. Kostyukhin,²¹ V. M. Kotov,⁶⁵ A. Kotwal,⁴⁵ A. Kourkoumeli-Charalampidi,¹⁵⁵ C. Kourkoumelis,⁹ V. Kouskoura,²⁵ A. Koutsman,^{160a} R. Kowalewski,¹⁷⁰ T. Z. Kowalski,^{38a} W. Kozanecki,¹³⁷ A. S. Kozhin,¹³⁰ V. A. Kramarenko,⁹⁹ G. Kramberger,⁷⁵ D. Krasnopevtsev,⁹⁸ M. W. Krasny,⁸⁰ A. Krasznahorkay,³⁰ J. K. Kraus,²¹ A. Kravchenko,²⁵ S. Kreiss,¹¹⁰ M. Kretz,^{58c} J. Kretzschmar,⁷⁴ K. Kreutzfeldt,⁵² P. Krieger,¹⁵⁹ K. Krizka,³¹ K. Kroeninger,⁴³ H. Kroha,¹⁰¹ J. Kroll,¹²² J. Kroseberg,²¹ J. Krstic,^{13a} U. Kruchonak,⁶⁵ H. Krüger,²¹ N. Krumnack,⁶⁴ Z. V. Krumshteyn,⁶⁵ A. Kruse,¹⁷⁴ M. C. Kruse,⁴⁵ M. Kruskal,²² T. Kubota,⁸⁸ H. Kucuk,⁷⁸ S. Kудay,^{4c} S. Kuehn,⁴⁸ A. Kugel,^{58c} F. Kuger,¹⁷⁵ A. Kuhl,¹³⁸ T. Kuhl,⁴² V. Kukhtin,⁶⁵ Y. Kulchitsky,⁹² S. Kuleshov,^{32b} M. Kuna,^{133a,133b} T. Kunigo,⁶⁸ A. Kupco,¹²⁷ H. Kurashige,⁶⁷ Y. A. Kurochkin,⁹² R. Kurumida,⁶⁷ V. Kus,¹²⁷ E. S. Kuwertz,¹⁴⁸ M. Kuze,¹⁵⁸ J. Kvita,¹¹⁵ D. Kyriazopoulos,¹⁴⁰ A. La Rosa,⁴⁹ L. La Rotonda,^{37a,37b} C. Lacasta,¹⁶⁸ F. Lacava,^{133a,133b} J. Lacey,²⁹ H. Lacker,¹⁶ D. Lacour,⁸⁰ V. R. Lacuesta,¹⁶⁸ E. Ladygin,⁶⁵ R. Lafaye,⁵ B. Laforge,⁸⁰ T. Lagouri,¹⁷⁷ S. Lai,⁴⁸ H. Laier,^{58a} L. Lambourne,⁷⁸ S. Lammers,⁶¹ C. L. Lampen,⁷ W. Lampl,⁷ E. Lançon,¹³⁷ U. Landgraf,⁴⁸ M. P. J. Landon,⁷⁶ V. S. Lang,^{58a} A. J. Lankford,¹⁶⁴ F. Lanni,²⁵ K. Lantzsck,³⁰ S. Laplace,⁸⁰ C. Lapoire,²¹ J. F. Laporte,¹³⁷ T. Lari,^{91a} F. Lasagni Manghi,^{20a,20b} M. Lassnig,³⁰ P. Laurelli,⁴⁷ W. Lavrijsen,¹⁵ A. T. Law,¹³⁸ P. Laycock,⁷⁴ O. Le Dortz,⁸⁰ E. Le Guirriec,⁸⁵ E. Le Menedeu,¹² T. LeCompte,⁶ F. Ledroit-Guillon,⁵⁵ C. A. Lee,^{146b} H. Lee,¹⁰⁷ S. C. Lee,¹⁵² L. Lee,¹ G. Lefebvre,⁸⁰ M. Lefebvre,¹⁷⁰ F. Legger,¹⁰⁰ C. Leggett,¹⁵ A. Lehan,⁷⁴ G. Lehmann Miotto,³⁰ X. Lei,⁷ W. A. Leight,²⁹ A. Leisos,¹⁵⁵ A. G. Leister,¹⁷⁷ M. A. L. Leite,^{24d} R. Leitner,¹²⁹ D. Lellouch,¹⁷³ B. Lemmer,⁵⁴ K. J. C. Leney,⁷⁸ T. Lenz,²¹ G. Lenzen,¹⁷⁶ B. Lenzi,³⁰ R. Leone,⁷ S. Leone,^{124a,124b} C. Leonidopoulos,⁴⁶ S. Leontsinis,¹⁰ C. Leroy,⁹⁵ C. G. Lester,²⁸ C. M. Lester,¹²² M. Levchenko,¹²³ J. Levêque,⁵ D. Levin,⁸⁹ L. J. Levinson,¹⁷³ M. Levy,¹⁸ A. Lewis,¹²⁰ A. M. Leyko,²¹ M. Leyton,⁴¹ B. Li,^{33b,u} B. Li,⁸⁵ H. Li,¹⁴⁹ H. L. Li,³¹ L. Li,⁴⁵ L. Li,^{33e} S. Li,⁴⁵ Y. Li,^{33c,v} Z. Liang,¹³⁸ H. Liao,³⁴ B. Liberti,^{134a} P. Lichard,³⁰ K. Lie,¹⁶⁶ J. Liebal,²¹ W. Liebig,¹⁴ C. Limbach,²¹ A. Limosani,¹⁵¹ S. C. Lin,^{152,w} T. H. Lin,⁸³ F. Linde,¹⁰⁷ B. E. Lindquist,¹⁴⁹ J. T. Linnemann,⁹⁰ E. Lipeles,¹²² A. Lipniacka,¹⁴ M. Lisovyi,⁴² T. M. Liss,¹⁶⁶ D. Lissauer,²⁵ A. Lister,¹⁶⁹ A. M. Litke,¹³⁸ B. Liu,¹⁵² D. Liu,¹⁵² J. B. Liu,^{33b} K. Liu,^{33b,x} L. Liu,⁸⁹ M. Liu,⁴⁵ M. Liu,^{33b} Y. Liu,^{33b} M. Livan,^{121a,121b} A. Lleres,⁵⁵ J. Llorente Merino,⁸² S. L. Lloyd,⁷⁶ F. Lo Sterzo,¹⁵² E. Lobodzinska,⁴² P. Loch,⁷ W. S. Lockman,¹³⁸ F. K. Loebinger,⁸⁴ A. E. Loevschall-Jensen,³⁶ A. Loginov,¹⁷⁷ T. Lohse,¹⁶ K. Lohwasser,⁴² M. Lokajicek,¹²⁷ B. A. Long,²² J. D. Long,⁸⁹ R. E. Long,⁷² K. A. Looper,¹¹¹ L. Lopes,^{126a} D. Lopez Mateos,⁵⁷ B. Lopez Paredes,¹⁴⁰ I. Lopez Paz,¹² J. Lorenz,¹⁰⁰ N. Lorenzo Martinez,⁶¹ M. Losada,¹⁶³

- P. Loscutoff,¹⁵ X. Lou,^{33a} A. Lounis,¹¹⁷ J. Love,⁶ P. A. Love,⁷² A. J. Lowe,^{144,f} F. Lu,^{33a} N. Lu,⁸⁹ H. J. Lubatti,¹³⁹
 C. Luci,^{133a,133b} A. Lucotte,⁵⁵ F. Luehring,⁶¹ W. Lukas,⁶² L. Luminari,^{133a} O. Lundberg,^{147a,147b} B. Lund-Jensen,¹⁴⁸
 M. Lungwitz,⁸³ D. Lynn,²⁵ R. Lysak,¹²⁷ E. Lytken,⁸¹ H. Ma,²⁵ L. L. Ma,^{33d} G. Maccarrone,⁴⁷ A. Macchiolo,¹⁰¹
 J. Machado Miguens,^{126a,126b} D. Macina,³⁰ D. Madaffari,⁸⁵ R. Madar,⁴⁸ H. J. Maddocks,⁷² W. F. Mader,⁴⁴ A. Madsen,¹⁶⁷
 M. Maeno,⁸ T. Maeno,²⁵ A. Maevskiy,⁹⁹ E. Magradze,⁵⁴ K. Mahboubi,⁴⁸ J. Mahlstedt,¹⁰⁷ S. Mahmoud,⁷⁴ C. Maiani,¹³⁷
 C. Maidantchik,^{24a} A. A. Maier,¹⁰¹ A. Maio,^{126a,126b,126d} S. Majewski,¹¹⁶ Y. Makida,⁶⁶ N. Makovec,¹¹⁷ P. Mal,^{137,y}
 B. Malaescu,⁸⁰ Pa. Malecki,³⁹ V. P. Maleev,¹²³ F. Malek,⁵⁵ U. Mallik,⁶³ D. Malon,⁶ C. Malone,¹⁴⁴ S. Maltezos,¹⁰
 V. M. Malyshev,¹⁰⁹ S. Malyukov,³⁰ J. Mamuzic,^{13b} B. Mandelli,³⁰ L. Mandelli,^{91a} I. Mandić,⁷⁵ R. Mandrysch,⁶³
 J. Maneira,^{126a,126b} A. Manfredini,¹⁰¹ L. Manhaes de Andrade Filho,^{24b} J. A. Manjarres Ramos,^{160b} A. Mann,¹⁰⁰
 P. M. Manning,¹³⁸ A. Manousakis-Katsikakis,⁹ B. Mansoulie,¹³⁷ R. Mantifel,⁸⁷ M. Mantoani,⁵⁴ L. Mapelli,³⁰ L. March,^{146c}
 J. F. Marchand,²⁹ G. Marchiori,⁸⁰ M. Marcisovsky,¹²⁷ C. P. Marino,¹⁷⁰ M. Marjanovic,^{13a} F. Marroquim,^{24a} S. P. Marsden,⁸⁴
 Z. Marshall,¹⁵ L. F. Marti,¹⁷ S. Marti-Garcia,¹⁶⁸ B. Martin,³⁰ B. Martin,⁹⁰ T. A. Martin,¹⁷¹ V. J. Martin,⁴⁶
 B. Martin dit Latour,¹⁴ H. Martinez,¹³⁷ M. Martinez,^{12,o} S. Martin-Haugh,¹³¹ A. C. Martyniuk,⁷⁸ M. Marx,¹³⁹ F. Marzano,^{133a}
 A. Marzin,³⁰ L. Masetti,⁸³ T. Mashimo,¹⁵⁶ R. Mashinistov,⁹⁶ J. Masik,⁸⁴ A. L. Maslennikov,^{109,d} I. Massa,^{20a,20b}
 L. Massa,^{20a,20b} N. Massol,⁵ P. Mastrandrea,¹⁴⁹ A. Mastroberardino,^{37a,37b} T. Masubuchi,¹⁵⁶ P. Mättig,¹⁷⁶ J. Mattmann,⁸³
 J. Maurer,^{26a} S. J. Maxfield,⁷⁴ D. A. Maximov,^{109,d} R. Mazini,¹⁵² L. Mazzaferro,^{134a,134b} G. Mc Goldrick,¹⁵⁹ S. P. Mc Kee,⁸⁹
 A. McCann,⁸⁹ R. L. McCarthy,¹⁴⁹ T. G. McCarthy,²⁹ N. A. McCubbin,¹³¹ K. W. McFarlane,^{56a} J. A. MCFayden,⁷⁸
 G. Mchedlidze,⁵⁴ S. J. McMahon,¹³¹ R. A. McPherson,^{170,k} J. Mechnich,¹⁰⁷ M. Medinnis,⁴² S. Meehan,³¹ S. Mehlhase,¹⁰⁰
 A. Mehta,⁷⁴ K. Meier,^{58a} C. Meineck,¹⁰⁰ B. Meirose,⁴¹ C. Melachrinou,³¹ B. R. Mellado Garcia,^{146c} F. Meloni,¹⁷
 A. Mengarelli,^{20a,20b} S. Menke,¹⁰¹ E. Meoni,¹⁶² K. M. Mercurio,⁵⁷ S. Mergelmeyer,²¹ N. Meric,¹³⁷ P. Mermod,⁴⁹
 L. Merola,^{104a,104b} C. Meroni,^{91a} F. S. Merritt,³¹ H. Merritt,¹¹¹ A. Messina,^{30,z} J. Metcalfe,²⁵ A. S. Mete,¹⁶⁴ C. Meyer,⁸³
 C. Meyer,¹²² J.-P. Meyer,¹³⁷ J. Meyer,³⁰ R. P. Middleton,¹³¹ S. Migas,⁷⁴ S. Miglioranza,^{165a,165c} L. Mijović,²¹ G. Mikenberg,¹⁷³
 M. Mikestikova,¹²⁷ M. Mikuž,⁷⁵ A. Milic,³⁰ D. W. Miller,³¹ C. Mills,⁴⁶ A. Milov,¹⁷³ D. A. Milstead,^{147a,147b}
 A. A. Minaenko,¹³⁰ Y. Minami,¹⁵⁶ I. A. Minashvili,⁶⁵ A. I. Mincer,¹¹⁰ B. Mindur,^{38a} M. Mineev,⁶⁵ Y. Ming,¹⁷⁴ L. M. Mir,¹²
 G. Mirabelli,^{133a} T. Mitani,¹⁷² J. Mitrevski,¹⁰⁰ V. A. Mitsou,¹⁶⁸ A. Miucci,⁴⁹ P. S. Miyagawa,¹⁴⁰ J. U. Mjörnmark,⁸¹
 T. Moa,^{147a,147b} K. Mochizuki,⁸⁵ S. Mohapatra,³⁵ W. Mohr,⁴⁸ S. Molander,^{147a,147b} R. Moles-Valls,¹⁶⁸ K. Mönig,⁴²
 C. Monini,⁵⁵ J. Monk,³⁶ E. Monnier,⁸⁵ J. Montejo Berlingen,¹² F. Monticelli,⁷¹ S. Monzani,^{133a,133b} R. W. Moore,³
 N. Morange,⁶³ D. Moreno,¹⁶³ M. Moreno Llácer,⁵⁴ P. Morettini,^{50a} M. Morgenstern,⁴⁴ M. Morii,⁵⁷ V. Morisbak,¹¹⁹
 S. Moritz,⁸³ A. K. Morley,¹⁴⁸ G. Mornacchi,³⁰ J. D. Morris,⁷⁶ A. Morton,⁴² L. Morvaj,¹⁰³ H. G. Moser,¹⁰¹ M. Mosidze,^{51b}
 J. Moss,¹¹¹ K. Motohashi,¹⁵⁸ R. Mount,¹⁴⁴ E. Mountricha,²⁵ S. V. Mouraviev,^{96,a} E. J. W. Moyses,⁸⁶ S. Muanza,⁸⁵
 R. D. Mudd,¹⁸ F. Mueller,^{58a} J. Mueller,¹²⁵ K. Mueller,²¹ T. Mueller,²⁸ D. Muenstermann,⁴⁹ Y. Munwes,¹⁵⁴
 J. A. Murillo Quijada,¹⁸ W. J. Murray,^{171,131} H. Musheghyan,⁵⁴ E. Musto,¹⁵³ A. G. Myagkov,^{130,aa} M. Myska,¹²⁸
 O. Nackenhorst,⁵⁴ J. Nadal,⁵⁴ K. Nagai,¹²⁰ R. Nagai,¹⁵⁸ Y. Nagai,⁸⁵ K. Nagano,⁶⁶ A. Nagarkar,¹¹¹ Y. Nagasaka,⁵⁹
 K. Nagata,¹⁶¹ M. Nagel,¹⁰¹ A. M. Nairz,³⁰ Y. Nakahama,³⁰ K. Nakamura,⁶⁶ T. Nakamura,¹⁵⁶ I. Nakano,¹¹²
 H. Namasivayam,⁴¹ G. Nanava,²¹ R. F. Naranjo Garcia,⁴² R. Narayan,^{58b} T. Nattermann,²¹ T. Naumann,⁴² G. Navarro,¹⁶³
 R. Nayyar,⁷ H. A. Neal,⁸⁹ P. Yu. Nechaeva,⁹⁶ T. J. Neep,⁸⁴ P. D. Nef,¹⁴⁴ A. Negri,^{121a,121b} G. Negri,³⁰ M. Negrini,^{20a}
 S. Nektarijevic,⁴⁹ C. Nellist,¹¹⁷ A. Nelson,¹⁶⁴ T. K. Nelson,¹⁴⁴ S. Nemecek,¹²⁷ P. Nemethy,¹¹⁰ A. A. Nepomuceno,^{24a}
 M. Nessi,^{30,bb} M. S. Neubauer,¹⁶⁶ M. Neumann,¹⁷⁶ R. M. Neves,¹¹⁰ P. Nevski,²⁵ P. R. Newman,¹⁸ D. H. Nguyen,⁶
 R. B. Nickerson,¹²⁰ R. Nicolaidou,¹³⁷ B. Nicquevert,³⁰ J. Nielsen,¹³⁸ N. Nikiforou,³⁵ A. Nikiforov,¹⁶ V. Nikolaenko,^{130,aa}
 I. Nikolic-Audit,⁸⁰ K. Nikolics,⁴⁹ K. Nikolopoulos,¹⁸ P. Nilsson,²⁵ Y. Ninomiya,¹⁵⁶ A. Nisati,^{133a} R. Nisius,¹⁰¹ T. Nobe,¹⁵⁸
 M. Nomachi,¹¹⁸ I. Nomidis,²⁹ S. Norberg,¹¹³ M. Nordberg,³⁰ O. Novgorodova,⁴⁴ S. Nowak,¹⁰¹ M. Nozaki,⁶⁶ L. Nozka,¹¹⁵
 K. Ntekas,¹⁰ G. Nunes Hanninger,⁸⁸ T. Nunnemann,¹⁰⁰ E. Nurse,⁷⁸ F. Nuti,⁸⁸ B. J. O'Brien,⁴⁶ F. O'grady,⁷ D. C. O'Neil,¹⁴³
 V. O'Shea,⁵³ F. G. Oakham,^{29,e} H. Oberlack,¹⁰¹ T. Obermann,²¹ J. Ocariz,⁸⁰ A. Ochi,⁶⁷ I. Ochoa,⁷⁸ S. Oda,⁷⁰ S. Odaka,⁶⁶
 H. Ogren,⁶¹ A. Oh,⁸⁴ S. H. Oh,⁴⁵ C. C. Ohm,¹⁵ H. Ohman,¹⁶⁷ H. Oide,³⁰ W. Okamura,¹¹⁸ H. Okawa,¹⁶¹ Y. Okumura,³¹
 T. Okuyama,¹⁵⁶ A. Olariu,^{26a} A. G. Olchevski,⁶⁵ S. A. Olivares Pino,⁴⁶ D. Oliveira Damazio,²⁵ E. Oliver Garcia,¹⁶⁸
 A. Olszewski,³⁹ J. Olszowska,³⁹ A. Onofre,^{126a,126e} P. U. E. Onyisi,^{31,p} C. J. Oram,^{160a} M. J. Oreglia,³¹ Y. Oren,¹⁵⁴
 D. Orestano,^{135a,135b} N. Orlando,^{73a,73b} C. Oropeza Barrera,⁵³ R. S. Orr,¹⁵⁹ B. Osculati,^{50a,50b} R. Ospanov,¹²²
 G. Otero y Garzon,²⁷ H. Otono,⁷⁰ M. Ouchrif,^{136d} E. A. Ouellette,¹⁷⁰ F. Ould-Saada,¹¹⁹ A. Ouraou,¹³⁷ K. P. Oussoren,¹⁰⁷
 Q. Ouyang,^{33a} A. Ovcharova,¹⁵ M. Owen,⁸⁴ V. E. Ozcan,^{19a} N. Ozturk,⁸ K. Pachal,¹²⁰ A. Pacheco Pages,¹²

C. Padilla Aranda,¹² M. Pagáčová,⁴⁸ S. Pagan Griso,¹⁵ E. Paganis,¹⁴⁰ C. Pahl,¹⁰¹ F. Paige,²⁵ P. Pais,⁸⁶ K. Pajchel,¹¹⁹
 G. Palacino,^{160b} S. Palestini,³⁰ M. Palka,^{38b} D. Pallin,³⁴ A. Palma,^{126a,126b} J. D. Palmer,¹⁸ Y. B. Pan,¹⁷⁴ E. Panagiotopoulou,¹⁰
 J. G. Panduro Vazquez,⁷⁷ P. Pani,¹⁰⁷ N. Panikashvili,⁸⁹ S. Panitkin,²⁵ D. Pantea,^{26a} L. Paolozzi,^{134a,134b}
 Th. D. Papadopoulou,¹⁰ K. Papageorgiou,¹⁵⁵ A. Paramonov,⁶ D. Paredes Hernandez,¹⁵⁵ M. A. Parker,²⁸ F. Parodi,^{50a,50b}
 J. A. Parsons,³⁵ U. Parzefall,⁴⁸ E. Pasqualucci,^{133a} S. Passaggio,^{50a} A. Passeri,^{135a} F. Pastore,^{135a,135b,a} Fr. Pastore,⁷⁷
 G. Pásztor,²⁹ S. Patarraia,¹⁷⁶ N. D. Patel,¹⁵¹ J. R. Pater,⁸⁴ S. Patricelli,^{104a,104b} T. Pauly,³⁰ J. Pearce,¹⁷⁰ L. E. Pedersen,³⁶
 M. Pedersen,¹¹⁹ S. Pedraza Lopez,¹⁶⁸ R. Pedro,^{126a,126b} S. V. Peleganchuk,¹⁰⁹ D. Pelikan,¹⁶⁷ H. Peng,^{33b} B. Penning,³¹
 J. Penwell,⁶¹ M. M. Perego,^{91a,91b} D. V. Perepelitsa,²⁵ E. Perez Codina,^{160a} M. T. Pérez García-Estañ,¹⁶⁸ L. Perini,^{91a,91b}
 H. Pernegger,³⁰ S. Perrella,^{104a,104b} R. Peschke,⁴² V. D. Peshekhonov,⁶⁵ K. Peters,³⁰ R. F. Y. Peters,⁸⁴ B. A. Petersen,³⁰
 T. C. Petersen,³⁶ E. Petit,⁴² A. Petridis,^{147a,147b} C. Petridou,¹⁵⁵ E. Petrolo,^{133a} F. Petrucci,^{135a,135b} N. E. Pettersson,¹⁵⁸
 R. Pezoa,^{32b} P. W. Phillips,¹³¹ G. Piacquadio,¹⁴⁴ E. Pianori,¹⁷¹ A. Picazio,⁴⁹ E. Piccaro,⁷⁶ M. Piccinini,^{20a,20b}
 M. A. Pickering,¹²⁰ R. Piegai,²⁷ D. T. Pignotti,¹¹¹ J. E. Pilcher,³¹ A. D. Pilkington,⁷⁸ J. Pina,^{126a,126b,126d}
 M. Pinamonti,^{165a,165c,cc} A. Pinder,¹²⁰ J. L. Pinfold,³ A. Pingel,³⁶ B. Pinto,^{126a} S. Pires,⁸⁰ M. Pitt,¹⁷³ C. Pizio,^{91a,91b}
 L. Plazak,^{145a} M.-A. Pleier,²⁵ V. Pleskot,¹²⁹ E. Plotnikova,⁶⁵ P. Plucinski,^{147a,147b} D. Pluth,⁶⁴ S. Poddar,^{58a} F. Podlyski,³⁴
 R. Poettgen,⁸³ L. Poggioli,¹¹⁷ D. Pohl,²¹ M. Pohl,⁴⁹ G. Polesello,^{121a} A. Policicchio,^{37a,37b} R. Polifka,¹⁵⁹ A. Polini,^{20a}
 C. S. Pollard,⁵³ V. Polychronakos,²⁵ K. Pommès,³⁰ L. Pontecorvo,^{133a} B. G. Pope,⁹⁰ G. A. Popeneciu,^{26b} D. S. Popovic,^{13a}
 A. Poppleton,³⁰ S. Pospisil,¹²⁸ K. Potamianos,¹⁵ I. N. Potrap,⁶⁵ C. J. Potter,¹⁵⁰ C. T. Potter,¹¹⁶ G. Poulard,³⁰ J. Poveda,³⁰
 V. Pozdnyakov,⁶⁵ P. Pralavorio,⁸⁵ A. Pranko,¹⁵ S. Prasad,³⁰ S. Prell,⁶⁴ D. Price,⁸⁴ J. Price,⁷⁴ L. E. Price,⁶ D. Prieur,¹²⁵
 M. Primavera,^{73a} S. Prince,⁸⁷ M. Proissl,⁴⁶ K. Prokofiev,^{60c} F. Prokoshin,^{32b} E. Protopapadaki,¹³⁷ S. Protopopescu,²⁵
 J. Proudfoot,⁶ M. Przybycien,^{38a} H. Przysieznik,⁵ E. Ptacek,¹¹⁶ D. Puddu,^{135a,135b} E. Pueschel,⁸⁶ D. Puldon,¹⁴⁹
 M. Purohit,^{25,dd} P. Puzo,¹¹⁷ J. Qian,⁸⁹ G. Qin,⁵³ Y. Qin,⁸⁴ A. Quadt,⁵⁴ D. R. Quarrie,¹⁵ W. B. Quayle,^{165a,165b}
 M. Queitsch-Maitland,⁸⁴ D. Quilty,⁵³ A. Qureshi,^{160b} V. Radeka,²⁵ V. Radescu,⁴² S. K. Radhakrishnan,¹⁴⁹ P. Radloff,¹¹⁶
 P. Rados,⁸⁸ F. Ragusa,^{91a,91b} G. Rahal,¹⁷⁹ S. Rajagopalan,²⁵ M. Rammensee,³⁰ C. Rangel-Smith,¹⁶⁷ K. Rao,¹⁶⁴
 M. G. Ratti,^{91a,91b} F. Rauscher,¹⁰⁰ S. Rave,⁸³ T. C. Rave,⁴⁸ T. Ravenscroft,⁵³ M. Raymond,³⁰ A. L. Read,¹¹⁹ N. P. Readioff,⁷⁴
 D. M. Rebuzzi,^{121a,121b} A. Redelbach,¹⁷⁵ G. Redlinger,²⁵ R. Reece,¹³⁸ K. Reeves,⁴¹ L. Rehnisch,¹⁶ H. Reisin,²⁷ M. Relich,¹⁶⁴
 C. Rembser,³⁰ H. Ren,^{33a} Z. L. Ren,¹⁵² A. Renaud,¹¹⁷ M. Rescigno,^{133a} S. Resconi,^{91a} O. L. Rezanova,^{109,d} P. Reznicek,¹²⁹
 R. Rezvani,⁹⁵ R. Richter,¹⁰¹ M. Ridel,⁸⁰ P. Rieck,¹⁶ J. Rieger,⁵⁴ M. Rijssenbeek,¹⁴⁹ A. Rimoldi,^{121a,121b} L. Rinaldi,^{20a}
 E. Ritsch,⁶² I. Riu,¹² F. Rizatdinova,¹¹⁴ E. Rizvi,⁷⁶ S. H. Robertson,^{87,k} A. Robichaud-Veronneau,⁸⁷ D. Robinson,²⁸
 J. E. M. Robinson,⁸⁴ A. Robson,⁵³ C. Roda,^{124a,124b} L. Rodrigues,³⁰ S. Roe,³⁰ O. Røhne,¹¹⁹ S. Rolli,¹⁶² A. Romaniouk,⁹⁸
 M. Romano,^{20a,20b} E. Romero Adam,¹⁶⁸ N. Rompotis,¹³⁹ M. Ronzani,⁴⁸ L. Roos,⁸⁰ E. Ros,¹⁶⁸ S. Rosati,^{133a} K. Rosbach,⁴⁹
 M. Rose,⁷⁷ P. Rose,¹³⁸ P. L. Rosendahl,¹⁴ O. Rosenthal,¹⁴² V. Rossetti,^{147a,147b} E. Rossi,^{104a,104b} L. P. Rossi,^{50a} R. Rosten,¹³⁹
 M. Rotaru,^{26a} I. Roth,¹⁷³ J. Rothberg,¹³⁹ D. Rousseau,¹¹⁷ C. R. Royon,¹³⁷ A. Rozanov,⁸⁵ Y. Rozen,¹⁵³ X. Ruan,^{146c}
 F. Rubbo,¹² I. Rubinskiy,⁴² V. I. Rud,⁹⁹ C. Rudolph,⁴⁴ M. S. Rudolph,¹⁵⁹ F. Rühr,⁴⁸ A. Ruiz-Martinez,³⁰ Z. Rurikova,⁴⁸
 N. A. Rusakovich,⁶⁵ A. Ruschke,¹⁰⁰ H. L. Russell,¹³⁹ J. P. Rutherford,⁷ N. Ruthmann,⁴⁸ Y. F. Ryabov,¹²³ M. Rybar,¹²⁹
 G. Rybkin,¹¹⁷ N. C. Ryder,¹²⁰ A. F. Saavedra,¹⁵¹ G. Sabato,¹⁰⁷ S. Sacerdoti,²⁷ A. Saddique,³ H. F. W. Sadrozinski,¹³⁸
 R. Sadykov,⁶⁵ F. Safai Tehrani,^{133a} H. Sakamoto,¹⁵⁶ Y. Sakurai,¹⁷² G. Salamanna,^{135a,135b} A. Salamon,^{134a} M. Saleem,¹¹³
 D. Salek,¹⁰⁷ P. H. Sales De Bruin,¹³⁹ D. Salihagic,¹⁰¹ A. Salnikov,¹⁴⁴ J. Salt,¹⁶⁸ D. Salvatore,^{37a,37b} F. Salvatore,¹⁵⁰
 A. Salvucci,¹⁰⁶ A. Salzburger,³⁰ D. Sampsonidis,¹⁵⁵ A. Sanchez,^{104a,104b} J. Sánchez,¹⁶⁸ V. Sanchez Martinez,¹⁶⁸
 H. Sandaker,¹⁴ R. L. Sandbach,⁷⁶ H. G. Sander,⁸³ M. P. Sanders,¹⁰⁰ M. Sandhoff,¹⁷⁶ T. Sandoval,²⁸ C. Sandoval,¹⁶³
 R. Sandstroem,¹⁰¹ D. P. C. Sankey,¹³¹ A. Sansoni,⁴⁷ C. Santoni,³⁴ R. Santonico,^{134a,134b} H. Santos,^{126a} I. Santoyo Castillo,¹⁵⁰
 K. Sapp,¹²⁵ A. Saponov,⁶⁵ J. G. Saraiva,^{126a,126d} B. Sarrazin,²¹ G. Sartisohn,¹⁷⁶ O. Sasaki,⁶⁶ Y. Sasaki,¹⁵⁶ K. Sato,¹⁶¹
 G. Sauvage,^{5,a} E. Sauvan,⁵ G. Savage,⁷⁷ P. Savard,^{159,e} C. Sawyer,¹²⁰ L. Sawyer,^{79,n} D. H. Saxon,⁵³ J. Saxon,³¹ C. Sbarra,^{20a}
 A. Sbrizzi,^{20a,20b} T. Scanlon,⁷⁸ D. A. Scannicchio,¹⁶⁴ M. Scarcella,¹⁵¹ V. Scarfone,^{37a,37b} J. Schaarschmidt,¹⁷³ P. Schacht,¹⁰¹
 D. Schaefer,³⁰ R. Schaefer,⁴² S. Schaepe,²¹ S. Schaezel,^{58b} U. Schäfer,⁸³ A. C. Schaffer,¹¹⁷ D. Schaile,¹⁰⁰
 R. D. Schamberger,¹⁴⁹ V. Scharf,^{58a} V. A. Schegelsky,¹²³ D. Scheirich,¹²⁹ M. Schernau,¹⁶⁴ C. Schiavi,^{50a,50b} J. Schieck,¹⁰⁰
 C. Schillo,⁴⁸ M. Schioppa,^{37a,37b} S. Schlenker,³⁰ E. Schmidt,⁴⁸ K. Schmieden,³⁰ C. Schmitt,⁸³ S. Schmitt,^{58b} B. Schneider,¹⁷
 Y. J. Schnellbach,⁷⁴ U. Schnoor,⁴⁴ L. Schoeffel,¹³⁷ A. Schoening,^{58b} B. D. Schoenrock,⁹⁰ A. L. S. Schorlemmer,⁵⁴
 M. Schott,⁸³ D. Schouten,^{160a} J. Schovancova,²⁵ S. Schramm,¹⁵⁹ M. Schreyer,¹⁷⁵ C. Schroeder,⁸³ N. Schuh,⁸³
 M. J. Schultens,²¹ H.-C. Schultz-Coulon,^{58a} H. Schulz,¹⁶ M. Schumacher,⁴⁸ B. A. Schumm,¹³⁸ Ph. Schune,¹³⁷

C. Schwanenberger,⁸⁴ A. Schwartzman,¹⁴⁴ T. A. Schwarz,⁸⁹ Ph. Schwegler,¹⁰¹ Ph. Schwemling,¹³⁷ R. Schwienhorst,⁹⁰ J. Schwindling,¹³⁷ T. Schwindt,²¹ M. Schwoerer,⁵ F. G. Sciacca,¹⁷ E. Scifo,¹¹⁷ G. Sciolla,²³ F. Scuri,^{124a,124b} F. Scutti,²¹ J. Searcy,⁸⁹ G. Sedov,⁴² E. Sedykh,¹²³ P. Seema,²¹ S. C. Seidel,¹⁰⁵ A. Seiden,¹³⁸ F. Seifert,¹²⁸ J. M. Seixas,^{24a} G. Sekhniaidze,^{104a} S. J. Sekula,⁴⁰ K. E. Selbach,⁴⁶ D. M. Seliverstov,^{123,a} G. Sellers,⁷⁴ N. Semprini-Cesari,^{20a,20b} C. Serfon,³⁰ L. Serin,¹¹⁷ L. Serkin,⁵⁴ T. Serre,⁸⁵ R. Seuster,^{160a} H. Severini,¹¹³ T. Sfiligoi,⁷⁵ F. Sforza,¹⁰¹ A. Sfyrla,³⁰ E. Shabalina,⁵⁴ M. Shamim,¹¹⁶ L. Y. Shan,^{33a} R. Shang,¹⁶⁶ J. T. Shank,²² M. Shapiro,¹⁵ P. B. Shatalov,⁹⁷ K. Shaw,^{165a,165b} A. Shcherbakova,^{147a,147b} C. Y. Shehu,¹⁵⁰ P. Sherwood,⁷⁸ L. Shi,^{152,ee} S. Shimizu,⁶⁷ C. O. Shimmin,¹⁶⁴ M. Shimojima,¹⁰² M. Shiyakova,⁶⁵ A. Shmeleva,⁹⁶ D. Shoaleh Saadi,⁹⁵ M. J. Shochet,³¹ S. Shojaii,^{91a,91b} D. Short,¹²⁰ S. Shrestha,¹¹¹ E. Shulga,⁹⁸ M. A. Shupe,⁷ S. Shushkevich,⁴² P. Sicho,¹²⁷ O. Sidiropoulou,¹⁵⁵ D. Sidorov,¹¹⁴ A. Sidoti,^{133a} F. Siegert,⁴⁴ Dj. Sijacki,^{13a} J. Silva,^{126a,126d} Y. Silver,¹⁵⁴ D. Silverstein,¹⁴⁴ S. B. Silverstein,^{147a} V. Simak,¹²⁸ O. Simard,⁵ Lj. Simic,^{13a} S. Simion,¹¹⁷ E. Simioni,⁸³ B. Simmons,⁷⁸ D. Simon,³⁴ R. Simoniello,^{91a,91b} P. Sinervo,¹⁵⁹ N. B. Sinev,¹¹⁶ G. Siragusa,¹⁷⁵ A. Sircar,⁷⁹ A. N. Sisakyan,^{65,a} S. Yu. Sivoklokov,⁹⁹ J. Sjölin,^{147a,147b} T. B. Sjursen,¹⁴ H. P. Skottowe,⁵⁷ P. Skubic,¹¹³ M. Slater,¹⁸ T. Slavicek,¹²⁸ M. Slawinska,¹⁰⁷ K. Sliwa,¹⁶² V. Smakhtin,¹⁷³ B. H. Smart,⁴⁶ L. Smestad,¹⁴ S. Yu. Smirnov,⁹⁸ Y. Smirnov,⁹⁸ L. N. Smirnova,^{99,ff} O. Smirnova,⁸¹ K. M. Smith,⁵³ M. Smizanska,⁷² K. Smolek,¹²⁸ A. A. Snesarev,⁹⁶ G. Snidero,⁷⁶ S. Snyder,²⁵ R. Sobie,^{170,k} F. Socher,⁴⁴ A. Soffer,¹⁵⁴ D. A. Soh,^{152,ee} C. A. Solans,³⁰ M. Solar,¹²⁸ J. Solc,¹²⁸ E. Yu. Soldatov,⁹⁸ U. Soldevila,¹⁶⁸ A. A. Solodkov,¹³⁰ A. Soloshenko,⁶⁵ O. V. Solovyanov,¹³⁰ V. Solovyev,¹²³ P. Sommer,⁴⁸ H. Y. Song,^{33b} N. Soni,¹ A. Sood,¹⁵ A. Sopczak,¹²⁸ B. Sopko,¹²⁸ V. Sopko,¹²⁸ V. Sorin,¹² M. Sosebee,⁸ R. Soualah,^{165a,165c} P. Soueid,⁹⁵ A. M. Soukharev,^{109,d} D. South,⁴² S. Spagnolo,^{73a,73b} F. Spanò,⁷⁷ W. R. Spearman,⁵⁷ F. Spettel,¹⁰¹ R. Spighi,^{20a} G. Spigo,³⁰ L. A. Spiller,⁸⁸ M. Spousta,¹²⁹ T. Spreitzer,¹⁵⁹ R. D. St. Denis,^{53,a} S. Staerz,⁴⁴ J. Stahlman,¹²² R. Stamen,^{58a} S. Stamm,¹⁶ E. Stanecka,³⁹ C. Stanescu,^{135a} M. Stanescu-Bellu,⁴² M. M. Stanitzki,⁴² S. Stapnes,¹¹⁹ E. A. Starchenko,¹³⁰ J. Stark,⁵⁵ P. Staroba,¹²⁷ P. Starovoitov,⁴² R. Staszewski,³⁹ P. Stavina,^{145a,a} P. Steinberg,²⁵ B. Stelzer,¹⁴³ H. J. Stelzer,³⁰ O. Stelzer-Chilton,^{160a} H. Stenzel,⁵² S. Stern,¹⁰¹ G. A. Stewart,⁵³ J. A. Stillings,²¹ M. C. Stockton,⁸⁷ M. Stoebe,⁸⁷ G. Stoicea,^{26a} P. Stolte,⁵⁴ S. Stonjek,¹⁰¹ A. R. Stradling,⁸ A. Straessner,⁴⁴ M. E. Stramaglia,¹⁷ J. Strandberg,¹⁴⁸ S. Strandberg,^{147a,147b} A. Strandlie,¹¹⁹ E. Strauss,¹⁴⁴ M. Strauss,¹¹³ P. Strizenec,^{145b} R. Ströhmer,¹⁷⁵ D. M. Strom,¹¹⁶ R. Stroynowski,⁴⁰ A. Strubig,¹⁰⁶ S. A. Stucci,¹⁷ B. Stugu,¹⁴ N. A. Styles,⁴² D. Su,¹⁴⁴ J. Su,¹²⁵ R. Subramaniam,⁷⁹ A. Succurro,¹² Y. Sugaya,¹¹⁸ C. Suhr,¹⁰⁸ M. Suk,¹²⁸ V. V. Sulin,⁹⁶ S. Sultansoy,^{4d} T. Sumida,⁶⁸ S. Sun,⁵⁷ X. Sun,^{33a} J. E. Sundermann,⁴⁸ K. Suruliz,¹⁵⁰ G. Susinno,^{37a,37b} M. R. Sutton,¹⁵⁰ Y. Suzuki,⁶⁶ M. Svatos,¹²⁷ S. Swedish,¹⁶⁹ M. Swiatlowski,¹⁴⁴ I. Sykora,^{145a} T. Sykora,¹²⁹ D. Ta,⁹⁰ C. Taccini,^{135a,135b} K. Tackmann,⁴² J. Taenzer,¹⁵⁹ A. Taffard,¹⁶⁴ R. Tafirout,^{160a} N. Taiblum,¹⁵⁴ H. Takai,²⁵ R. Takashima,⁶⁹ H. Takeda,⁶⁷ T. Takeshita,¹⁴¹ Y. Takubo,⁶⁶ M. Talby,⁸⁵ A. A. Talyshev,^{109,d} J. Y. C. Tam,¹⁷⁵ K. G. Tan,⁸⁸ J. Tanaka,¹⁵⁶ R. Tanaka,¹¹⁷ S. Tanaka,¹³² S. Tanaka,⁶⁶ A. J. Tanasijczuk,¹⁴³ B. B. Tannenwald,¹¹¹ N. Tannoury,²¹ S. Tapprogge,⁸³ S. Tarem,¹⁵³ F. Tarrade,²⁹ G. F. Tartarelli,^{91a} P. Tas,¹²⁹ M. Tasevsky,¹²⁷ T. Tashiro,⁶⁸ E. Tassi,^{37a,37b} A. Tavares Delgado,^{126a,126b} Y. Tayalati,^{136d} F. E. Taylor,⁹⁴ G. N. Taylor,⁸⁸ W. Taylor,^{160b} F. A. Teischinger,³⁰ M. Teixeira Dias Castanheira,⁷⁶ P. Teixeira-Dias,⁷⁷ K. K. Temming,⁴⁸ H. Ten Kate,³⁰ P. K. Teng,¹⁵² J. J. Teoh,¹¹⁸ S. Terada,⁶⁶ K. Terashi,¹⁵⁶ J. Terron,⁸² S. Terzo,¹⁰¹ M. Testa,⁴⁷ R. J. Teuscher,^{159,k} J. Therhaag,²¹ T. Theveneaux-Pelzer,³⁴ J. P. Thomas,¹⁸ J. Thomas-Wilsker,⁷⁷ E. N. Thompson,³⁵ P. D. Thompson,¹⁸ R. J. Thompson,⁸⁴ A. S. Thompson,⁵³ L. A. Thomsen,³⁶ E. Thomson,¹²² M. Thomson,²⁸ W. M. Thong,⁸⁸ R. P. Thun,^{89,a} F. Tian,³⁵ M. J. Tibbetts,¹⁵ V. O. Tikhomirov,^{96,gg} Yu. A. Tikhonov,^{109,d} S. Timoshenko,⁹⁸ E. Tiouchichine,⁸⁵ P. Tipton,¹⁷⁷ S. Tisserant,⁸⁵ T. Todorov,⁵ S. Todorova-Nova,¹²⁹ J. Tojo,⁷⁰ S. Tokár,^{145a} K. Tokushuku,⁶⁶ K. Tollefson,⁹⁰ E. Tolley,⁵⁷ L. Tomlinson,⁸⁴ M. Tomoto,¹⁰³ L. Tompkins,³¹ K. Toms,¹⁰⁵ N. D. Topilin,⁶⁵ E. Torrence,¹¹⁶ H. Torres,¹⁴³ E. Torró Pastor,¹⁶⁸ J. Toth,^{85,hh} F. Touchard,⁸⁵ D. R. Tovey,¹⁴⁰ H. L. Tran,¹¹⁷ T. Trefzger,¹⁷⁵ L. Tremblet,³⁰ A. Tricoli,³⁰ I. M. Trigger,^{160a} S. Trincaz-Duvoid,⁸⁰ M. F. Tripania,¹² W. Trischuk,¹⁵⁹ B. Trocmé,⁵⁵ C. Troncon,^{91a} M. Trottier-McDonald,¹⁵ M. Trovatelli,^{135a,135b} P. True,⁹⁰ M. Trzebinski,³⁹ A. Trzupek,³⁹ C. Tsarouchas,³⁰ J. C.-L. Tseng,¹²⁰ P. V. Tsiarehshka,⁹² D. Tsionou,¹³⁷ G. Tsipolitis,¹⁰ N. Tsirintanis,⁹ S. Tsiskaridze,¹² V. Tsiskaridze,⁴⁸ E. G. Tskhadadze,^{51a} I. I. Tsukerman,⁹⁷ V. Tsulaia,¹⁵ S. Tsuno,⁶⁶ D. Tsybychev,¹⁴⁹ A. Tudorache,^{26a} V. Tudorache,^{26a} A. N. Tuna,¹²² S. A. Tuppiti,^{20a,20b} S. Turchikhin,^{99,ff} D. Turecek,¹²⁸ I. Turk Cakir,^{4c} R. Turra,^{91a,91b} A. J. Turvey,⁴⁰ P. M. Tuts,³⁵ A. Tykhonov,⁴⁹ M. Tylmad,^{147a,147b} M. Tyndel,¹³¹ I. Ueda,¹⁵⁶ R. Ueno,²⁹ M. Ughetto,⁸⁵ M. Ugland,¹⁴ M. Uhlenbrock,²¹ F. Ukegawa,¹⁶¹ G. Unal,³⁰ A. Undrus,²⁵ G. Unel,¹⁶⁴ F. C. Ungaro,⁴⁸ Y. Unno,⁶⁶ C. Unverdorben,¹⁰⁰ J. Urban,^{145b} D. Urbaniec,³⁵ P. Urquijo,⁸⁸ G. Usai,⁸ A. Usanova,⁶² L. Vacavant,⁸⁵ V. Vacek,¹²⁸ B. Vachon,⁸⁷ N. Valencic,¹⁰⁷ S. Valentineti,^{20a,20b} A. Valero,¹⁶⁸ L. Valery,³⁴ S. Valkar,¹²⁹ E. Valladolid Gallego,¹⁶⁸ S. Vallecorsa,⁴⁹ J. A. Valls Ferrer,¹⁶⁸ W. Van Den Wollenberg,¹⁰⁷

P. C. Van Der Deijl,¹⁰⁷ R. van der Geer,¹⁰⁷ H. van der Graaf,¹⁰⁷ R. Van Der Leeuw,¹⁰⁷ D. van der Ster,³⁰ N. van Eldik,³⁰ P. van Gemeren,⁶ J. Van Nieuwkoop,¹⁴³ I. van Vulpen,¹⁰⁷ M. C. van Woerden,³⁰ M. Vanadia,^{133a,133b} W. Vandelli,³⁰ R. Vanguri,¹²² A. Vaniachine,⁶ P. Vankov,⁴² F. Vannucci,⁸⁰ G. Vardanyan,¹⁷⁸ R. Vari,^{133a} E. W. Varnes,⁷ T. Varol,⁸⁶ D. Varouchas,⁸⁰ A. Vartapetian,⁸ K. E. Varvell,¹⁵¹ F. Vazeille,³⁴ T. Vazquez Schroeder,⁵⁴ J. Veatch,⁷ F. Veloso,^{126a,126c} T. Velz,²¹ S. Veneziano,^{133a} A. Ventura,^{73a,73b} D. Ventura,⁸⁶ M. Venturi,¹⁷⁰ N. Venturi,¹⁵⁹ A. Venturini,²³ V. Vercesi,^{121a} M. Verducci,^{133a,133b} W. Verkerke,¹⁰⁷ J. C. Vermeulen,¹⁰⁷ A. Vest,⁴⁴ M. C. Vetterli,^{143,e} O. Viazlo,⁸¹ I. Vichou,¹⁶⁶ T. Vickey,^{146c,ii} O. E. Vickey Boeriu,^{146c} G. H. A. Viehhauser,¹²⁰ S. Viel,¹⁶⁹ R. Vigne,³⁰ M. Villa,^{20a,20b} M. Villaplana Perez,^{91a,91b} E. Vilucchi,⁴⁷ M. G. Vincter,²⁹ V. B. Vinogradov,⁶⁵ J. Virzi,¹⁵ I. Vivarelli,¹⁵⁰ F. Vives Vaque,³ S. Vlachos,¹⁰ D. Vladoiu,¹⁰⁰ M. Vlasak,¹²⁸ A. Vogel,²¹ M. Vogel,^{32a} P. Vokac,¹²⁸ G. Volpi,^{124a,124b} M. Volpi,⁸⁸ H. von der Schmitt,¹⁰¹ H. von Radziewski,⁴⁸ E. von Toerne,²¹ V. Vorobel,¹²⁹ K. Vorobev,⁹⁸ M. Vos,¹⁶⁸ R. Voss,³⁰ J. H. Vosseveld,⁷⁴ N. Vranjes,¹³⁷ M. Vranjes Milosavljevic,^{13a} V. Vrba,¹²⁷ M. Vreeswijk,¹⁰⁷ T. Vu Anh,⁴⁸ R. Vuillermet,³⁰ I. Vukotic,³¹ Z. Vykydal,¹²⁸ P. Wagner,²¹ W. Wagner,¹⁷⁶ H. Wahlberg,⁷¹ S. Wahrmund,⁴⁴ J. Wakabayashi,¹⁰³ J. Walder,⁷² R. Walker,¹⁰⁰ W. Walkowiak,¹⁴² R. Wall,¹⁷⁷ P. Waller,⁷⁴ B. Walsh,¹⁷⁷ C. Wang,^{33c} C. Wang,⁴⁵ F. Wang,¹⁷⁴ H. Wang,¹⁵ H. Wang,⁴⁰ J. Wang,⁴² J. Wang,^{33a} K. Wang,⁸⁷ R. Wang,¹⁰⁵ S. M. Wang,¹⁵² T. Wang,²¹ X. Wang,¹⁷⁷ C. Wanotayaroj,¹¹⁶ A. Warburton,⁸⁷ C. P. Ward,²⁸ D. R. Wardrope,⁷⁸ M. Warsinsky,⁴⁸ A. Washbrook,⁴⁶ C. Wasicki,⁴² P. M. Watkins,¹⁸ A. T. Watson,¹⁸ I. J. Watson,¹⁵¹ M. F. Watson,¹⁸ G. Watts,¹³⁹ S. Watts,⁸⁴ B. M. Waugh,⁷⁸ S. Webb,⁸⁴ M. S. Weber,¹⁷ S. W. Weber,¹⁷⁵ J. S. Webster,³¹ A. R. Weidberg,¹²⁰ B. Weinert,⁶¹ J. Weingarten,⁵⁴ C. Weiser,⁴⁸ H. Weits,¹⁰⁷ P. S. Wells,³⁰ T. Wenaus,²⁵ D. Wendland,¹⁶ Z. Weng,^{152,ee} T. Wengler,³⁰ S. Wenig,³⁰ N. Vermes,²¹ M. Werner,⁴⁸ P. Werner,³⁰ M. Wessels,^{58a} J. Wetter,¹⁶² K. Whalen,²⁹ A. White,⁸ M. J. White,¹ R. White,^{32b} S. White,^{124a,124b} D. Whiteson,¹⁶⁴ D. Wicke,¹⁷⁶ F. J. Wickens,¹³¹ W. Wiedenmann,¹⁷⁴ M. Wielers,¹³¹ P. Wienemann,²¹ C. Wiglesworth,³⁶ L. A. M. Wiik-Fuchs,²¹ P. A. Wijeratne,⁷⁸ A. Wildauer,¹⁰¹ M. A. Wildt,^{42,ij} H. G. Wilkens,³⁰ H. H. Williams,¹²² S. Williams,²⁸ C. Willis,⁹⁰ S. Willocq,⁸⁶ A. Wilson,⁸⁹ J. A. Wilson,¹⁸ I. Wingerter-Seez,⁵ F. Winklmeier,¹¹⁶ B. T. Winter,²¹ M. Wittgen,¹⁴⁴ J. Wittkowski,¹⁰⁰ S. J. Wollstadt,⁸³ M. W. Wolter,³⁹ H. Wolters,^{126a,126c} B. K. Wosiek,³⁹ J. Wotschack,³⁰ M. J. Woudstra,⁸⁴ K. W. Wozniak,³⁹ M. Wright,⁵³ M. Wu,⁵⁵ S. L. Wu,¹⁷⁴ X. Wu,⁴⁹ Y. Wu,⁸⁹ T. R. Wyatt,⁸⁴ B. M. Wynne,⁴⁶ S. Xella,³⁶ M. Xiao,¹³⁷ D. Xu,^{33a} L. Xu,^{33b,kk} B. Yabsley,¹⁵¹ S. Yacoob,^{146b,ll} R. Yakabe,⁶⁷ M. Yamada,⁶⁶ H. Yamaguchi,¹⁵⁶ Y. Yamaguchi,¹¹⁸ A. Yamamoto,⁶⁶ S. Yamamoto,¹⁵⁶ T. Yamamura,¹⁵⁶ T. Yamanaka,¹⁵⁶ K. Yamauchi,¹⁰³ Y. Yamazaki,⁶⁷ Z. Yan,²² H. Yang,^{33e} H. Yang,¹⁷⁴ Y. Yang,¹¹¹ S. Yanush,⁹³ L. Yao,^{33a} W-M. Yao,¹⁵ Y. Yasu,⁶⁶ E. Yatsenko,⁴² K. H. Yau Wong,²¹ J. Ye,⁴⁰ S. Ye,²⁵ I. Yeletsikh,⁶⁵ A. L. Yen,⁵⁷ E. Yildirim,⁴² M. Yilmaz,^{4b} K. Yorita,¹⁷² R. Yoshida,⁶ K. Yoshihara,¹⁵⁶ C. Young,¹⁴⁴ C. J. S. Young,³⁰ S. Youssef,²² D. R. Yu,¹⁵ J. Yu,⁸ J. M. Yu,⁸⁹ J. Yu,¹¹⁴ L. Yuan,⁶⁷ A. Yurkewicz,¹⁰⁸ I. Yusuff,^{28,mmm} B. Zabinski,³⁹ R. Zaidan,⁶³ A. M. Zaitsev,^{130,aa} A. Zaman,¹⁴⁹ S. Zambito,²³ L. Zanello,^{133a,133b} D. Zanzi,⁸⁸ C. Zeitnitz,¹⁷⁶ M. Zeman,¹²⁸ A. Zemla,^{38a} K. Zengel,²³ O. Zenin,¹³⁰ T. Ženiš,^{145a} D. Zerwas,¹¹⁷ G. Zevi della Porta,⁵⁷ D. Zhang,⁸⁹ F. Zhang,¹⁷⁴ H. Zhang,⁹⁰ J. Zhang,⁶ L. Zhang,¹⁵² R. Zhang,^{33b} X. Zhang,^{33d} Z. Zhang,¹¹⁷ Y. Zhao,^{33d} Z. Zhao,^{33b} A. Zhemchugov,⁶⁵ J. Zhong,¹²⁰ B. Zhou,⁸⁹ C. Zhou,⁴⁵ L. Zhou,³⁵ L. Zhou,⁴⁰ N. Zhou,¹⁶⁴ C. G. Zhu,^{33d} H. Zhu,^{33a} J. Zhu,⁸⁹ Y. Zhu,^{33b} X. Zhuang,^{33a} K. Zhukov,⁹⁶ A. Zibell,¹⁷⁵ D. Zieminska,⁶¹ N. I. Zimine,⁶⁵ C. Zimmermann,⁸³ R. Zimmermann,²¹ S. Zimmermann,²¹ S. Zimmermann,⁴⁸ Z. Zinonos,⁵⁴ M. Ziolkowski,¹⁴² G. Zobernig,¹⁷⁴ A. Zoccoli,^{20a,20b} M. zur Nedden,¹⁶ G. Zurzolo,^{104a,104b} and L. Zwalinski³⁰

(ATLAS Collaboration)

¹Department of Physics, University of Adelaide, Adelaide, Australia²Physics Department, SUNY Albany, Albany New York, USA³Department of Physics, University of Alberta, Edmonton, AB, Canada^{4a}Department of Physics, Ankara University, Ankara, Turkey^{4b}Department of Physics, Gazi University, Ankara, Turkey^{4c}Istanbul Aydin University, Istanbul, Turkey^{4d}Division of Physics, TOBB University of Economics and Technology, Ankara, Turkey⁵LAPP, CNRS/IN2P3 and Université de Savoie, Annecy-le-Vieux, France⁶High Energy Physics Division, Argonne National Laboratory, Argonne, Illinois, USA⁷Department of Physics, University of Arizona, Tucson, Arizona, USA⁸Department of Physics, The University of Texas at Arlington, Arlington, Texas, USA⁹Physics Department, University of Athens, Athens, Greece¹⁰Physics Department, National Technical University of Athens, Zografou, Greece

- ¹¹*Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan*
- ¹²*Institut de Física d'Altes Energies and Departament de Física de la Universitat Autònoma de Barcelona, Barcelona, Spain*
- ^{13a}*Institute of Physics, University of Belgrade, Belgrade, Serbia*
- ^{13b}*Vinca Institute of Nuclear Sciences, University of Belgrade, Belgrade, Serbia*
- ¹⁴*Department for Physics and Technology, University of Bergen, Bergen, Norway*
- ¹⁵*Physics Division, Lawrence Berkeley National Laboratory and University of California, Berkeley, California, USA*
- ¹⁶*Department of Physics, Humboldt University, Berlin, Germany*
- ¹⁷*Albert Einstein Center for Fundamental Physics and Laboratory for High Energy Physics, University of Bern, Bern, Switzerland*
- ¹⁸*School of Physics and Astronomy, University of Birmingham, Birmingham, United Kingdom*
- ^{19a}*Department of Physics, Bogazici University, Istanbul, Turkey*
- ^{19b}*Department of Physics, Dogus University, Istanbul, Turkey*
- ^{19c}*Department of Physics Engineering, Gaziantep University, Gaziantep, Turkey*
- ^{20a}*INFN Sezione di Bologna, Italy*
- ^{20b}*Dipartimento di Fisica e Astronomia, Università di Bologna, Bologna, Italy*
- ²¹*Physikalisches Institut, University of Bonn, Bonn, Germany*
- ²²*Department of Physics, Boston University, Boston, Massachusetts, USA*
- ²³*Department of Physics, Brandeis University, Waltham, Massachusetts, USA*
- ^{24a}*Universidade Federal do Rio De Janeiro COPPE/EE/IF, Rio de Janeiro, Brazil*
- ^{24b}*Electrical Circuits Department, Federal University of Juiz de Fora (UFJF), Juiz de Fora, Brazil*
- ^{24c}*Federal University of Sao Joao del Rei (UFSJ), Sao Joao del Rei, Brazil*
- ^{24d}*Instituto de Fisica, Universidade de Sao Paulo, Sao Paulo, Brazil*
- ²⁵*Physics Department, Brookhaven National Laboratory, Upton, New York, USA*
- ^{26a}*National Institute of Physics and Nuclear Engineering, Bucharest, Romania*
- ^{26b}*National Institute for Research and Development of Isotopic and Molecular Technologies, Physics Department, Cluj Napoca, Romania*
- ^{26c}*University Politehnica Bucharest, Bucharest, Romania*
- ^{26d}*West University in Timisoara, Timisoara, Romania*
- ²⁷*Departamento de Física, Universidad de Buenos Aires, Buenos Aires, Argentina*
- ²⁸*Cavendish Laboratory, University of Cambridge, Cambridge, United Kingdom*
- ²⁹*Department of Physics, Carleton University, Ottawa, ON, Canada*
- ³⁰*CERN, Geneva, Switzerland*
- ³¹*Enrico Fermi Institute, University of Chicago, Chicago, Illinois, USA*
- ^{32a}*Departamento de Física, Pontificia Universidad Católica de Chile, Santiago, Chile*
- ^{32b}*Departamento de Física, Universidad Técnica Federico Santa María, Valparaíso, Chile*
- ^{33a}*Institute of High Energy Physics, Chinese Academy of Sciences, Beijing, China*
- ^{33b}*Department of Modern Physics, University of Science and Technology of China, Anhui, China*
- ^{33c}*Department of Physics, Nanjing University, Jiangsu, China*
- ^{33d}*School of Physics, Shandong University, Shandong, China*
- ^{33e}*Physics Department, Shanghai Jiao Tong University, Shanghai, China*
- ^{33f}*Physics Department, Tsinghua University, Beijing 100084, China*
- ³⁴*Laboratoire de Physique Corpusculaire, Clermont Université and Université Blaise Pascal and CNRS/IN2P3, Clermont-Ferrand, France*
- ³⁵*Nevis Laboratory, Columbia University, Irvington, New York, USA*
- ³⁶*Niels Bohr Institute, University of Copenhagen, Kobenhavn, Denmark*
- ^{37a}*INFN Gruppo Collegato di Cosenza, Laboratori Nazionali di Frascati, Italy*
- ^{37b}*Dipartimento di Fisica, Università della Calabria, Rende, Italy*
- ^{38a}*AGH University of Science and Technology, Faculty of Physics and Applied Computer Science, Krakow, Poland*
- ^{38b}*Marian Smoluchowski Institute of Physics, Jagiellonian University, Krakow, Poland*
- ³⁹*The Henryk Niewodniczanski Institute of Nuclear Physics, Polish Academy of Sciences, Krakow, Poland*
- ⁴⁰*Physics Department, Southern Methodist University, Dallas, Texas, USA*
- ⁴¹*Physics Department, University of Texas at Dallas, Richardson, Texas, USA*
- ⁴²*DESY, Hamburg and Zeuthen, Germany*
- ⁴³*Institut für Experimentelle Physik IV, Technische Universität Dortmund, Dortmund, Germany*
- ⁴⁴*Institut für Kern- und Teilchenphysik, Technische Universität Dresden, Dresden, Germany*
- ⁴⁵*Department of Physics, Duke University, Durham, North Carolina, USA*
- ⁴⁶*SUPA - School of Physics and Astronomy, University of Edinburgh, Edinburgh, United Kingdom*

- ⁴⁷*INFN Laboratori Nazionali di Frascati, Frascati, Italy*
- ⁴⁸*Fakultät für Mathematik und Physik, Albert-Ludwigs-Universität, Freiburg, Germany*
- ⁴⁹*Section de Physique, Université de Genève, Geneva, Switzerland*
- ^{50a}*INFN Sezione di Genova, Italy*
- ^{50b}*Dipartimento di Fisica, Università di Genova, Genova, Italy*
- ^{51a}*E. Andronikashvili Institute of Physics, Iv. Javakhishvili Tbilisi State University, Tbilisi, Georgia*
- ^{51b}*High Energy Physics Institute, Tbilisi State University, Tbilisi, Georgia*
- ⁵²*II Physikalisches Institut, Justus-Liebig-Universität Giessen, Giessen, Germany*
- ⁵³*SUPA - School of Physics and Astronomy, University of Glasgow, Glasgow, United Kingdom*
- ⁵⁴*II Physikalisches Institut, Georg-August-Universität, Göttingen, Germany*
- ⁵⁵*Laboratoire de Physique Subatomique et de Cosmologie, Université Grenoble-Alpes, CNRS/IN2P3, Grenoble, France*
- ⁵⁶*Department of Physics, Hampton University, Hampton, Virginia, USA*
- ⁵⁷*Laboratory for Particle Physics and Cosmology, Harvard University, Cambridge, Massachusetts, USA*
- ^{58a}*Kirchhoff-Institut für Physik, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany*
- ^{58b}*Physikalisches Institut, Ruprecht-Karls-Universität Heidelberg, Heidelberg, Germany*
- ^{58c}*ZITI Institut für technische Informatik, Ruprecht-Karls-Universität Heidelberg, Mannheim, Germany*
- ⁵⁹*Faculty of Applied Information Science, Hiroshima Institute of Technology, Hiroshima, Japan*
- ^{60a}*Department of Physics, The Chinese University of Hong Kong, Shatin, N.T., Hong Kong, China*
- ^{60b}*Department of Physics, The University of Hong Kong, Hong Kong, China*
- ^{60c}*Department of Physics, The Hong Kong University of Science and Technology, Clear Water Bay, Kowloon, Hong Kong, China*
- ⁶¹*Department of Physics, Indiana University, Bloomington, Indiana, USA*
- ⁶²*Institut für Astro- und Teilchenphysik, Leopold-Franzens-Universität, Innsbruck, Austria*
- ⁶³*University of Iowa, Iowa City, Iowa, USA*
- ⁶⁴*Department of Physics and Astronomy, Iowa State University, Ames, Iowa, USA*
- ⁶⁵*Joint Institute for Nuclear Research, JINR Dubna, Dubna, Russia*
- ⁶⁶*KEK, High Energy Accelerator Research Organization, Tsukuba, Japan*
- ⁶⁷*Graduate School of Science, Kobe University, Kobe, Japan*
- ⁶⁸*Faculty of Science, Kyoto University, Kyoto, Japan*
- ⁶⁹*Kyoto University of Education, Kyoto, Japan*
- ⁷⁰*Department of Physics, Kyushu University, Fukuoka, Japan*
- ⁷¹*Instituto de Física La Plata, Universidad Nacional de La Plata and CONICET, La Plata, Argentina*
- ⁷²*Physics Department, Lancaster University, Lancaster, United Kingdom*
- ^{73a}*INFN Sezione di Lecce, Italy*
- ^{73b}*Dipartimento di Matematica e Fisica, Università del Salento, Lecce, Italy*
- ⁷⁴*Oliver Lodge Laboratory, University of Liverpool, Liverpool, United Kingdom*
- ⁷⁵*Department of Physics, Jožef Stefan Institute and University of Ljubljana, Ljubljana, Slovenia*
- ⁷⁶*School of Physics and Astronomy, Queen Mary University of London, London, United Kingdom*
- ⁷⁷*Department of Physics, Royal Holloway University of London, Surrey, United Kingdom*
- ⁷⁸*Department of Physics and Astronomy, University College London, London, United Kingdom*
- ⁷⁹*Louisiana Tech University, Ruston, Louisiana, USA*
- ⁸⁰*Laboratoire de Physique Nucléaire et de Hautes Energies, UPMC and Université Paris-Diderot and CNRS/IN2P3, Paris, France*
- ⁸¹*Fysiska institutionen, Lunds universitet, Lund, Sweden*
- ⁸²*Departamento de Física Teórica C-15, Universidad Autónoma de Madrid, Madrid, Spain*
- ⁸³*Institut für Physik, Universität Mainz, Mainz, Germany*
- ⁸⁴*School of Physics and Astronomy, University of Manchester, Manchester, United Kingdom*
- ⁸⁵*CPPM, Aix-Marseille Université and CNRS/IN2P3, Marseille, France*
- ⁸⁶*Department of Physics, University of Massachusetts, Amherst, Massachusetts, USA*
- ⁸⁷*Department of Physics, McGill University, Montreal, QC, Canada*
- ⁸⁸*School of Physics, University of Melbourne, Victoria, Australia*
- ⁸⁹*Department of Physics, The University of Michigan, Ann Arbor, Michigan, USA*
- ⁹⁰*Department of Physics and Astronomy, Michigan State University, East Lansing, Michigan, USA*
- ^{91a}*INFN Sezione di Milano, Italy*
- ^{91b}*Dipartimento di Fisica, Università di Milano, Milano, Italy*
- ⁹²*B.I. Stepanov Institute of Physics, National Academy of Sciences of Belarus, Minsk, Republic of Belarus*
- ⁹³*National Scientific and Educational Centre for Particle and High Energy Physics, Minsk, Republic of Belarus*
- ⁹⁴*Department of Physics, Massachusetts Institute of Technology, Cambridge, Massachusetts, USA*

- ⁹⁵*Group of Particle Physics, University of Montreal, Montreal, QC, Canada*
- ⁹⁶*P.N. Lebedev Institute of Physics, Academy of Sciences, Moscow, Russia*
- ⁹⁷*Institute for Theoretical and Experimental Physics (ITEP), Moscow, Russia*
- ⁹⁸*National Research Nuclear University MEPhI, Moscow, Russia*
- ⁹⁹*D.V.Skobel'syn Institute of Nuclear Physics, M.V. Lomonosov Moscow State University, Moscow, Russia*
- ¹⁰⁰*Fakultät für Physik, Ludwig-Maximilians-Universität München, München, Germany*
- ¹⁰¹*Max-Planck-Institut für Physik (Werner-Heisenberg-Institut), München, Germany*
- ¹⁰²*Nagasaki Institute of Applied Science, Nagasaki, Japan*
- ¹⁰³*Graduate School of Science and Kobayashi-Maskawa Institute, Nagoya University, Nagoya, Japan*
- ^{104a}*INFN Sezione di Napoli, Italy*
- ^{104b}*Dipartimento di Fisica, Università di Napoli, Napoli, Italy*
- ¹⁰⁵*Department of Physics and Astronomy, University of New Mexico, Albuquerque, New Mexico, USA*
- ¹⁰⁶*Institute for Mathematics, Astrophysics and Particle Physics, Radboud University Nijmegen/Nikhef, Nijmegen, Netherlands*
- ¹⁰⁷*Nikhef National Institute for Subatomic Physics and University of Amsterdam, Amsterdam, Netherlands*
- ¹⁰⁸*Department of Physics, Northern Illinois University, DeKalb, Illinois, USA*
- ¹⁰⁹*Budker Institute of Nuclear Physics, SB RAS, Novosibirsk, Russia*
- ¹¹⁰*Department of Physics, New York University, New York, New York, USA*
- ¹¹¹*Ohio State University, Columbus, Ohio, USA*
- ¹¹²*Faculty of Science, Okayama University, Okayama, Japan*
- ¹¹³*Homer L. Dodge Department of Physics and Astronomy, University of Oklahoma, Norman, Oklahoma, USA*
- ¹¹⁴*Department of Physics, Oklahoma State University, Stillwater, Oklahoma, USA*
- ¹¹⁵*Palacký University, RCPTM, Olomouc, Czech Republic*
- ¹¹⁶*Center for High Energy Physics, University of Oregon, Eugene, Oregon, USA*
- ¹¹⁷*LAL, Université Paris-Sud and CNRS/IN2P3, Orsay, France*
- ¹¹⁸*Graduate School of Science, Osaka University, Osaka, Japan*
- ¹¹⁹*Department of Physics, University of Oslo, Oslo, Norway*
- ¹²⁰*Department of Physics, Oxford University, Oxford, United Kingdom*
- ^{121a}*INFN Sezione di Pavia, Italy*
- ^{121b}*Dipartimento di Fisica, Università di Pavia, Pavia, Italy*
- ¹²²*Department of Physics, University of Pennsylvania, Philadelphia, Pennsylvania, USA*
- ¹²³*Petersburg Nuclear Physics Institute, Gatchina, Russia*
- ^{124a}*INFN Sezione di Pisa, Italy*
- ^{124b}*Dipartimento di Fisica E. Fermi, Università di Pisa, Pisa, Italy*
- ¹²⁵*Department of Physics and Astronomy, University of Pittsburgh, Pittsburgh, Pennsylvania, USA*
- ^{126a}*Laboratorio de Instrumentacao e Fisica Experimental de Particulas - LIP, Lisboa, Portugal*
- ^{126b}*Faculdade de Ciências, Universidade de Lisboa, Lisboa, Portugal*
- ^{126c}*Department of Physics, University of Coimbra, Coimbra, Portugal*
- ^{126d}*Centro de Fisica Nuclear da Universidade de Lisboa, Lisboa, Portugal*
- ^{126e}*Departamento de Fisica, Universidade do Minho, Braga, Portugal*
- ^{126f}*Departamento de Fisica Teorica y del Cosmos and CAFPE, Universidad de Granada, Granada, Spain*
- ^{126g}*Dep Fisica and CEFITEC of Faculdade de Ciencias e Tecnologia, Universidade Nova de Lisboa, Caparica, Portugal*
- ¹²⁷*Institute of Physics, Academy of Sciences of the Czech Republic, Praha, Czech Republic*
- ¹²⁸*Czech Technical University in Prague, Praha, Czech Republic*
- ¹²⁹*Faculty of Mathematics and Physics, Charles University in Prague, Praha, Czech Republic*
- ¹³⁰*State Research Center Institute for High Energy Physics, Protvino, Russia*
- ¹³¹*Particle Physics Department, Rutherford Appleton Laboratory, Didcot, United Kingdom*
- ¹³²*Ritsumeikan University, Kusatsu, Shiga, Japan*
- ^{133a}*INFN Sezione di Roma, Italy*
- ^{133b}*Dipartimento di Fisica, Sapienza Università di Roma, Roma, Italy*
- ^{134a}*INFN Sezione di Roma Tor Vergata, Italy*
- ^{134b}*Dipartimento di Fisica, Università di Roma Tor Vergata, Roma, Italy*
- ^{135a}*INFN Sezione di Roma Tre, Italy*
- ^{135b}*Dipartimento di Matematica e Fisica, Università Roma Tre, Roma, Italy*
- ^{136a}*Faculté des Sciences Ain Chock, Réseau Universitaire de Physique des Hautes Energies - Université Hassan II, Casablanca, Morocco*
- ^{136b}*Centre National de l'Energie des Sciences Techniques Nucleaires, Rabat, Morocco*

- ^{136c}*Faculté des Sciences Semlalia, Université Cadi Ayyad, LPHEA-Marrakech, Morocco*
- ^{136d}*Faculté des Sciences, Université Mohamed Premier and LPTPM, Oujda, Morocco*
- ^{136e}*Faculté des sciences, Université Mohammed V-Agdal, Rabat, Morocco*
- ¹³⁷*DSM/IRFU (Institut de Recherches sur les Lois Fondamentales de l'Univers), CEA Saclay (Commissariat à l'Energie Atomique et aux Energies Alternatives), Gif-sur-Yvette, France*
- ¹³⁸*Santa Cruz Institute for Particle Physics, University of California Santa Cruz, Santa Cruz, California, USA*
- ¹³⁹*Department of Physics, University of Washington, Seattle, Washington, USA*
- ¹⁴⁰*Department of Physics and Astronomy, University of Sheffield, Sheffield, United Kingdom*
- ¹⁴¹*Department of Physics, Shinshu University, Nagano, Japan*
- ¹⁴²*Fachbereich Physik, Universität Siegen, Siegen, Germany*
- ¹⁴³*Department of Physics, Simon Fraser University, Burnaby, BC, Canada*
- ¹⁴⁴*SLAC National Accelerator Laboratory, Stanford, California, USA*
- ^{145a}*Faculty of Mathematics, Physics & Informatics, Comenius University, Bratislava, Slovak Republic*
- ^{145b}*Department of Subnuclear Physics, Institute of Experimental Physics of the Slovak Academy of Sciences, Kosice, Slovak Republic*
- ^{146a}*Department of Physics, University of Cape Town, Cape Town, South Africa*
- ^{146b}*Department of Physics, University of Johannesburg, Johannesburg, South Africa*
- ^{146c}*School of Physics, University of the Witwatersrand, Johannesburg, South Africa*
- ^{147a}*Department of Physics, Stockholm University, Sweden*
- ^{147b}*The Oskar Klein Centre, Stockholm, Sweden*
- ¹⁴⁸*Physics Department, Royal Institute of Technology, Stockholm, Sweden*
- ¹⁴⁹*Departments of Physics & Astronomy and Chemistry, Stony Brook University, Stony Brook, New York, USA*
- ¹⁵⁰*Department of Physics and Astronomy, University of Sussex, Brighton, United Kingdom*
- ¹⁵¹*School of Physics, University of Sydney, Sydney, Australia*
- ¹⁵²*Institute of Physics, Academia Sinica, Taipei, Taiwan*
- ¹⁵³*Department of Physics, Technion: Israel Institute of Technology, Haifa, Israel*
- ¹⁵⁴*Raymond and Beverly Sackler School of Physics and Astronomy, Tel Aviv University, Tel Aviv, Israel*
- ¹⁵⁵*Department of Physics, Aristotle University of Thessaloniki, Thessaloniki, Greece*
- ¹⁵⁶*International Center for Elementary Particle Physics and Department of Physics, The University of Tokyo, Tokyo, Japan*
- ¹⁵⁷*Graduate School of Science and Technology, Tokyo Metropolitan University, Tokyo, Japan*
- ¹⁵⁸*Department of Physics, Tokyo Institute of Technology, Tokyo, Japan*
- ¹⁵⁹*Department of Physics, University of Toronto, Toronto, ON, Canada*
- ^{160a}*TRIUMF, Vancouver, BC, Canada*
- ^{160b}*Department of Physics and Astronomy, York University, Toronto, ON, Canada*
- ¹⁶¹*Faculty of Pure and Applied Sciences, University of Tsukuba, Tsukuba, Japan*
- ¹⁶²*Department of Physics and Astronomy, Tufts University, Medford, Massachusetts, USA*
- ¹⁶³*Centro de Investigaciones, Universidad Antonio Narino, Bogota, Colombia*
- ¹⁶⁴*Department of Physics and Astronomy, University of California Irvine, Irvine, California, USA*
- ^{165a}*INFN Gruppo Collegato di Udine, Sezione di Trieste, Udine, Italy*
- ^{165b}*ICTP, Trieste, Italy*
- ^{165c}*Dipartimento di Chimica, Fisica e Ambiente, Università di Udine, Udine, Italy*
- ¹⁶⁶*Department of Physics, University of Illinois, Urbana, Illinois, USA*
- ¹⁶⁷*Department of Physics and Astronomy, University of Uppsala, Uppsala, Sweden*
- ¹⁶⁸*Instituto de Física Corpuscular (IFIC) and Departamento de Física Atómica, Molecular y Nuclear and Departamento de Ingeniería Electrónica and Instituto de Microelectrónica de Barcelona (IMB-CNM), University of Valencia and CSIC, Valencia, Spain*
- ¹⁶⁹*Department of Physics, University of British Columbia, Vancouver, BC, Canada*
- ¹⁷⁰*Department of Physics and Astronomy, University of Victoria, Victoria, BC, Canada*
- ¹⁷¹*Department of Physics, University of Warwick, Coventry, United Kingdom*
- ¹⁷²*Waseda University, Tokyo, Japan*
- ¹⁷³*Department of Particle Physics, The Weizmann Institute of Science, Rehovot, Israel*
- ¹⁷⁴*Department of Physics, University of Wisconsin, Madison, Wisconsin, USA*
- ¹⁷⁵*Fakultät für Physik und Astronomie, Julius-Maximilians-Universität, Würzburg, Germany*
- ¹⁷⁶*Fachbereich C Physik, Bergische Universität Wuppertal, Wuppertal, Germany*
- ¹⁷⁷*Department of Physics, Yale University, New Haven, Connecticut, USA*
- ¹⁷⁸*Yerevan Physics Institute, Yerevan, Armenia*

¹⁷⁹*Centre de Calcul de l'Institut National de Physique Nucléaire et de Physique des Particules (IN2P3),
Villeurbanne, France*

- ^aDeceased.
- ^bAlso at Department of Physics, King's College London, London, United Kingdom.
- ^cAlso at Institute of Physics, Azerbaijan Academy of Sciences, Baku, Azerbaijan.
- ^dAlso at Novosibirsk State University, Novosibirsk, Russia.
- ^eAlso at TRIUMF, Vancouver, BC, Canada.
- ^fAlso at Department of Physics, California State University, Fresno, California, USA.
- ^gAlso at Department of Physics, University of Fribourg, Fribourg, Switzerland.
- ^hAlso at Tomsk State University, Tomsk, Russia.
- ⁱAlso at CPPM, Aix-Marseille Université and CNRS/IN2P3, Marseille, France.
- ^jAlso at Università di Napoli Parthenope, Napoli, Italy.
- ^kAlso at Institute of Particle Physics (IPP), Canada.
- ^lAlso at Particle Physics Department, Rutherford Appleton Laboratory, Didcot, United Kingdom.
- ^mAlso at Department of Physics, St. Petersburg State Polytechnical University, St. Petersburg, Russia.
- ⁿAlso at Louisiana Tech University, Ruston, Louisiana, USA.
- ^oAlso at Institutio Catalana de Recerca i Estudis Avancats, ICREA, Barcelona, Spain.
- ^pAlso at Department of Physics, The University of Texas at Austin, Austin, Texas, USA.
- ^qAlso at Institute of Theoretical Physics, Ilia State University, Tbilisi, Georgia.
- ^rAlso at CERN, Geneva, Switzerland.
- ^sAlso at O Chadai Academic Production, Ochanomizu University, Tokyo, Japan.
- ^tAlso at Manhattan College, New York, New York, USA.
- ^uAlso at Institute of Physics, Academia Sinica, Taipei, Taiwan.
- ^vAlso at LAL, Université Paris-Sud and CNRS/IN2P3, Orsay, France.
- ^wAlso at Academia Sinica Grid Computing, Institute of Physics, Academia Sinica, Taipei, Taiwan.
- ^xAlso at Laboratoire de Physique Nucléaire et de Hautes Energies, UPMC and Université Paris-Diderot and CNRS/IN2P3, Paris, France.
- ^yAlso at School of Physical Sciences, National Institute of Science Education and Research, Bhubaneswar, India.
- ^zAlso at Dipartimento di Fisica, Sapienza Università di Roma, Roma, Italy.
- ^{aa}Also at Moscow Institute of Physics and Technology State University, Dolgoprudny, Russia.
- ^{bb}Also at Section de Physique, Université de Genève, Geneva, Switzerland.
- ^{cc}Also at International School for Advanced Studies (SISSA), Trieste, Italy.
- ^{dd}Also at Department of Physics and Astronomy, University of South Carolina, Columbia, South Carolina, USA.
- ^{ee}Also at School of Physics and Engineering, Sun Yat-sen University, Guangzhou, China.
- ^{ff}Also at Faculty of Physics, M.V.Lomonosov Moscow State University, Moscow, Russia.
- ^{gg}Also at National Research Nuclear University MEPhI, Moscow, Russia.
- ^{hh}Also at Institute for Particle and Nuclear Physics, Wigner Research Centre for Physics, Budapest, Hungary.
- ⁱⁱAlso at Department of Physics, Oxford University, Oxford, United Kingdom.
- ^{jj}Also at Institut für Experimentalphysik, Universität Hamburg, Hamburg, Germany.
- ^{kk}Also at Department of Physics, The University of Michigan, Ann Arbor, Michigan, USA.
- ^{ll}Also at Discipline of Physics, University of KwaZulu-Natal, Durban, South Africa.
- ^{mm}Also at University of Malaya, Department of Physics, Kuala Lumpur, Malaysia.