

Open Archive TOULOUSE Archive Ouverte (OATAO)

OATAO is an open access repository that collects the work of Toulouse researchers and makes it freely available over the web where possible.

This is an author-deposited version published in: http://oatao.univ-toulouse.fr/Eprints ID: 17369

To cite this version: Belloir, Jean-Marc and Goiffon, Vincent and Magnan, Pierre and Virmontois, Cédric and Gilard, Olivier and Raine, Mélanie and Paillet, Philippe *Dark Current Spectroscopy in neutron, proton and ion irradiated CMOS Image Sensors.* (2016) In: IEEE Nuclear and Space Radiation Effects Conference (NSREC 2016), 11 July 2016 - 15 July 2016 (Portland, Oregon, United States).

Any correspondence concerning this service should be sent to the repository administrator: staff-oatao@listes-diff.inp-toulouse.fr







SUPAERO

Dark current spectroscopy in proton, neutron and ion irradiated CMOS Image Sensors

J.-M. Belloir, V. Goiffon, P.Magnan, ISAE-SUPAERO, Toulouse, France,

C. Virmontois, O. Gilard, CNES, Toulouse, France,

M. Raine, P. Paillet, CEA DAM DIF, Arpajon, France.





Space and nuclear particles can degrade the performance of CMOS image sensors (CIS)





- Space and nuclear particles can degrade the performance of CMOS image sensors (CIS)
- They produce displacement damage by coulomb or nuclear interactions and create silicon bulk defects which can generate dark current





- Space and nuclear particles can degrade the performance of CMOS image sensors (CIS)
- They produce displacement damage by coulomb or nuclear interactions and create silicon bulk defects which can generate dark current
- The dark current distribution depends on the type and distribution of the radiation-induced defects





- Space and nuclear particles can degrade the performance of CMOS image sensors (CIS)
- They produce displacement damage by coulomb or nuclear interactions and create silicon bulk defects which can generate dark current
- The dark current distribution depends on the type and distribution of the radiation-induced defects
- Identifying the defects generated by each type of interaction could help predicting the dark current increase for various particles





- Space and nuclear particles can degrade the performance of CMOS image sensors (CIS)
- They produce displacement damage by coulomb or nuclear interactions and create silicon bulk defects which can generate dark current
- The dark current distribution depends on the type and distribution of the radiation-induced defects

In this work, the dark current spectroscopy is tested on neutron, proton and ion irradiated CIS to identify the defects generated by coulomb and nuclear interactions





$$U = \frac{\sigma v_{th} n_i}{2 \cosh \left(\frac{|\boldsymbol{E_t} - E_i|}{kT} \right)}$$





$$U = \frac{\sigma v_{th} n_i}{2 \cosh\left(\frac{|\boldsymbol{E_t} - E_i|}{kT}\right)}$$

- If each pixel contains a maximum of one radiation-induced bulk defect:
 - The generation rate of individual defects can be measured





$$U = \frac{\sigma v_{th} n_i}{2 \cosh\left(\frac{|\boldsymbol{E_t} - E_i|}{kT}\right)}$$

- If each pixel contains a maximum of one radiation-induced bulk defect:
 - The generation rate of individual defects can be measured
 - Pixels which contain similar defects have a similar dark current increase after irradiation



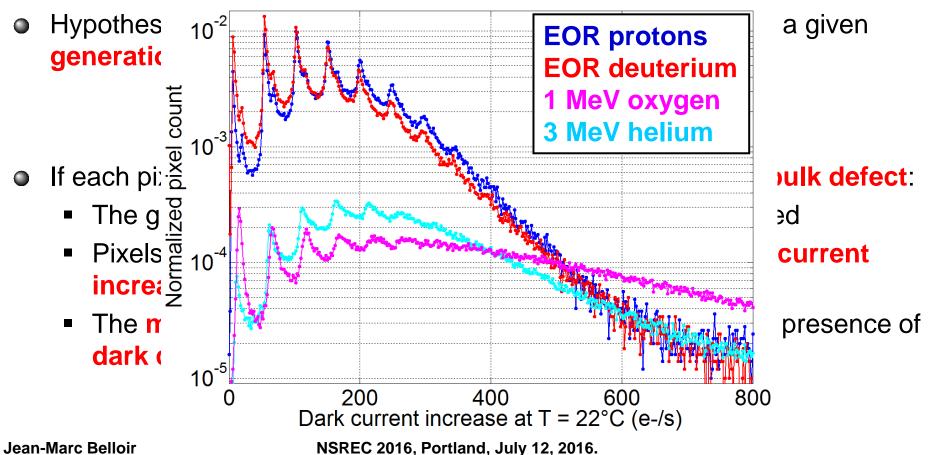


$$U = \frac{\sigma v_{th} n_i}{2 \cosh\left(\frac{|\boldsymbol{E_t} - E_i|}{kT}\right)}$$

- If each pixel contains a maximum of one radiation-induced bulk defect:
 - The generation rate of individual defects can be measured
 - Pixels which contain similar defects have a similar dark current increase after irradiation
 - The main radiation-induced defects are detected by the presence of dark current peaks in the dark current distribution





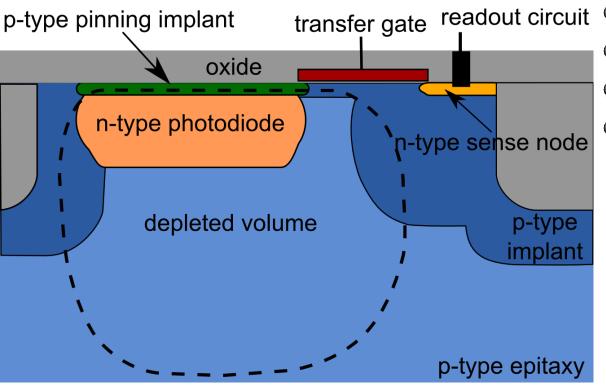




Tested device:

IS a e R O

Pinned photodiode CMOS image sensor

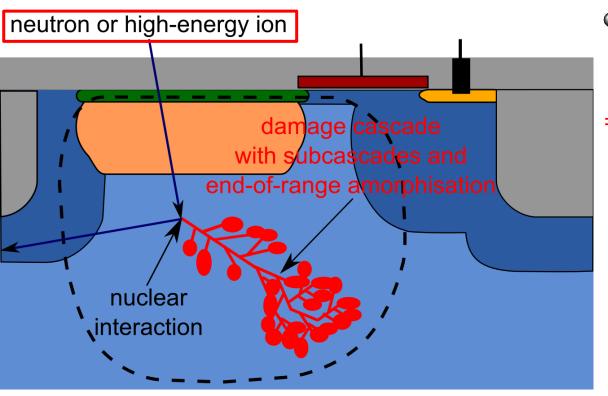


- 512 x 512 pixels
- 7 µm pixel pitch
- Pinned photodiode
- Low intrinsic dark current (6 e-/s @ 22°C)



Neutrons and high-energy ions: nuclear interactions





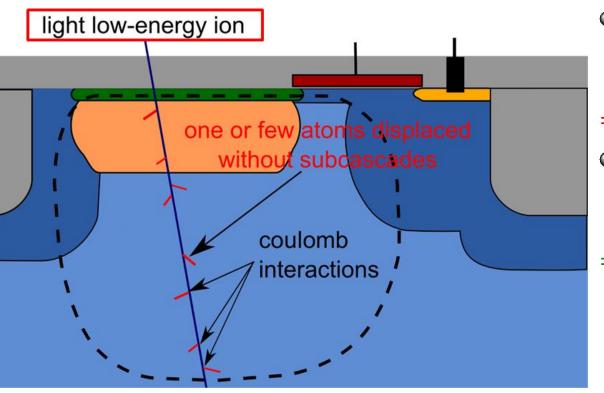
- Nuclear (elastic, inelastic):
 high energy silicon PKA
 (~ 100 keV)
- ⇒ dense damage

PKA = Primary Knock-on Atom (primary recoil)



Low energy light ions: low NIEL coulomb interactions





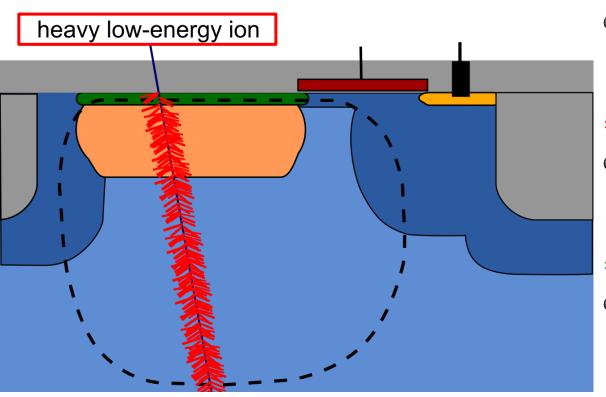
- Nuclear (elastic, inelastic):
 high energy silicon PKA
 (~ 100 keV)
- ⇒ dense damage
- Low NIEL coulomb: low energy PKA (~ 100 eV) which are well separated
- ⇒ sparse damage

PKA = Primary Knock-on Atom (primary recoil)



Low energy heavy ions: high NIEL coulomb interactions



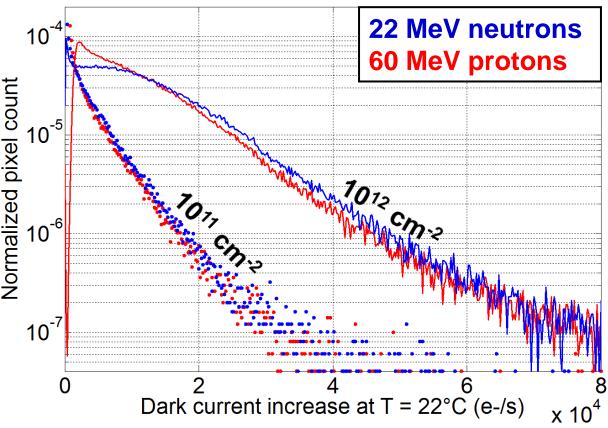


- Nuclear (elastic, inelastic):
 high energy silicon PKA
 (~ 100 keV)
- ⇒ dense damage
 - Low NIEL coulomb: low energy PKA (~ 100 eV) which are well separated
- ⇒ sparse damage
- High NIEL coulomb: low energy PKA but very close to each other
- ⇒ dense damage

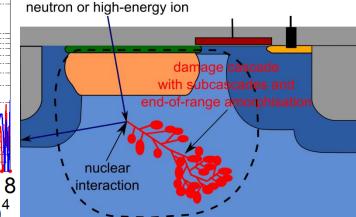


Nuclear interactions: Exponential hot pixel tail





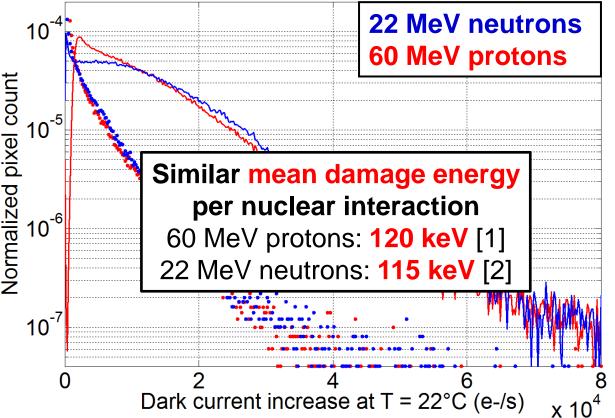
tail (several 10,000 e⁻/s) with many hot pixels





Nuclear interactions: Exponential hot pixel tail





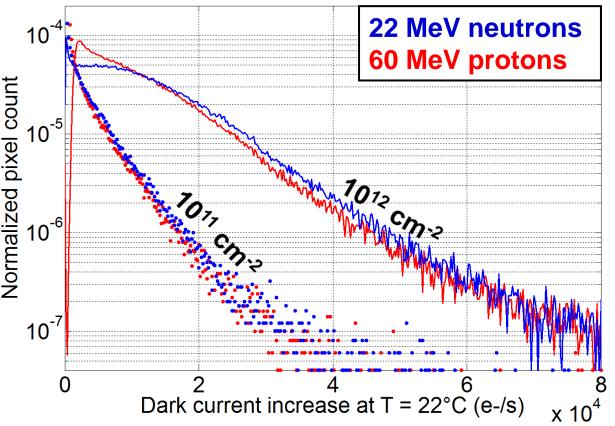
- Exponential hot pixel tail (several 10,000 e⁻/s) with many hot pixels
- Similar exponential mean for neutrons and protons (4.3.10³ e⁻/s)

- [1] Dale et al., IEEE TNS, 1994.
- 1 [1] Dale et al., IEEE 1NS, 1994.



Nuclear interactions: Exponential hot pixel tail





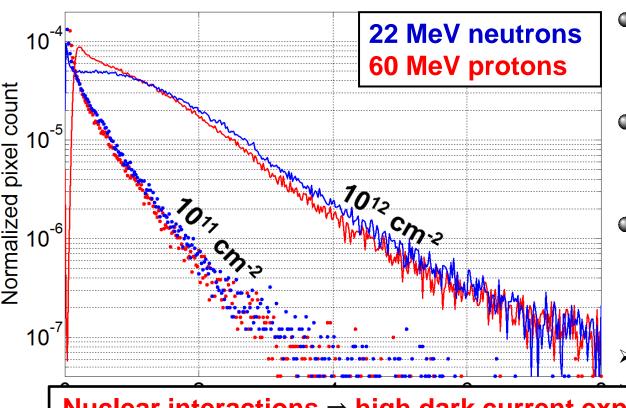
- **Exponential hot pixel** tail (several 10,000 e⁻/s) with many hot pixels
- Similar exponential mean for neutrons and protons (4.3.10³ e⁻/s)
- Similar number of hot pixels at similar fluence due to similar nuclear **NIEL**
 - Neutrons: 4.0 keVcm²/g

Protons: 2.7 keVcm²/g



Nuclear interactions: Exponential hot pixel tail





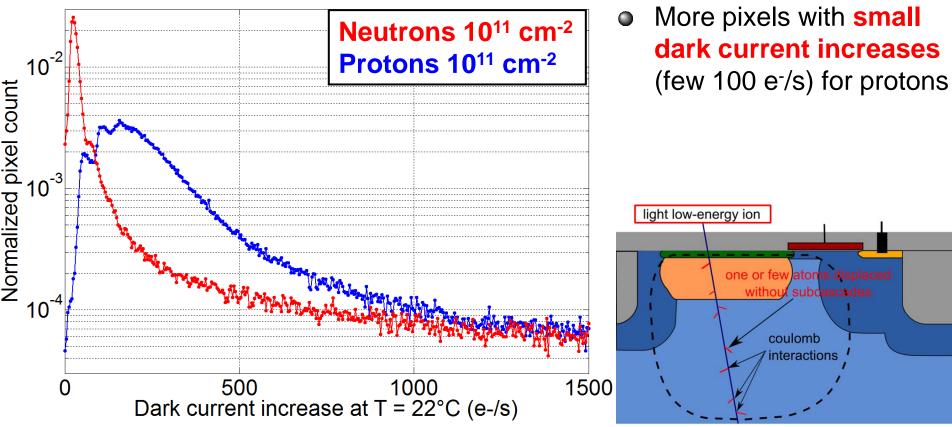
- tail (several 10,000 e⁻/s) with many hot pixels
- Similar exponential mean for neutrons and protons (4.3.10³ e⁻/s)
- Similar number of hot pixels at similar fluence due to similar nuclear NIEL
- > Neutrons: 4.0 keVcm²/g

Nuclear interactions ⇒ high dark current exponential hot pixel tail



Low NIEL coulomb interactions: dark current peaks

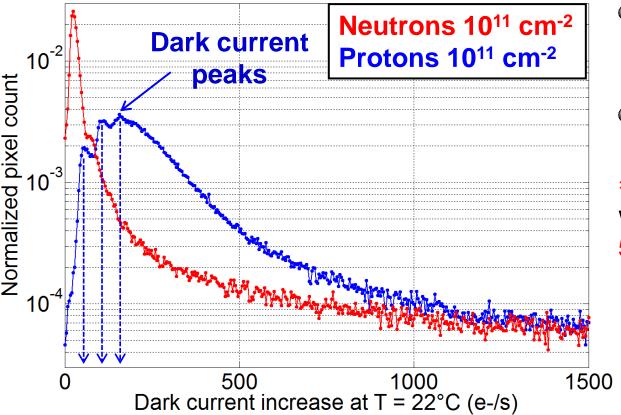






Low NIEL coulomb interactions: dark current peaks



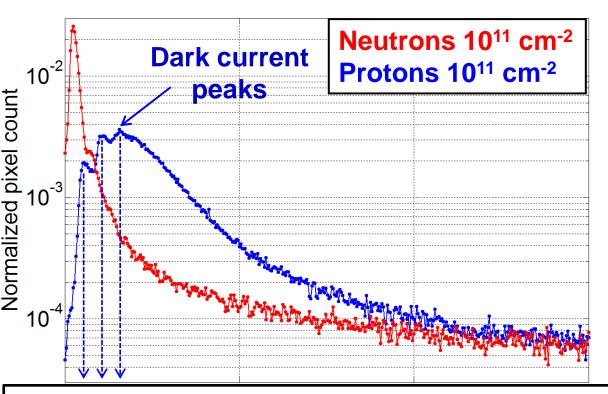


- More pixels with small dark current increases (few 100 e⁻/s) for protons
- Dark current peaks separated by 50 e⁻/s
- ⇒ specific point defectswith generation rate of50 e-/s



Low NIEL coulomb interactions: dark current peaks





- More pixels with small dark current increases (few 100 e⁻/s) for protons
- Dark current peaks separated by 50 e⁻/s
- ⇒ specific point defectswith generation rate of

50 e-/s

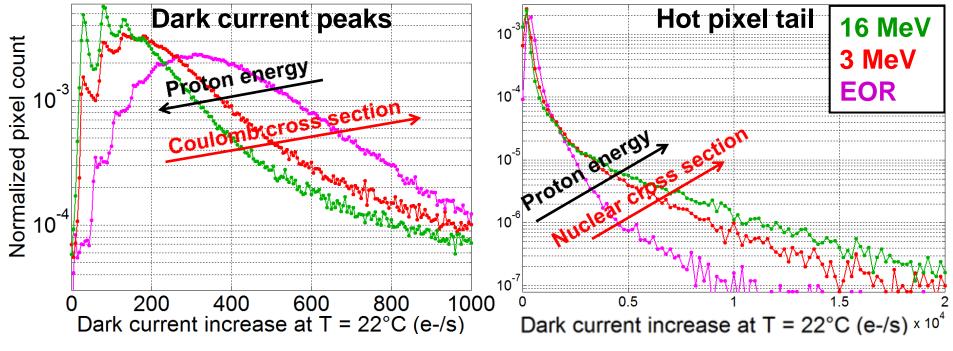
⇒ generated by coulomb interactions (only for protons)

Low NIEL coulomb interactions ⇒ dark current peaks (point defects)



Effect of the coulomb and nuclear cross section



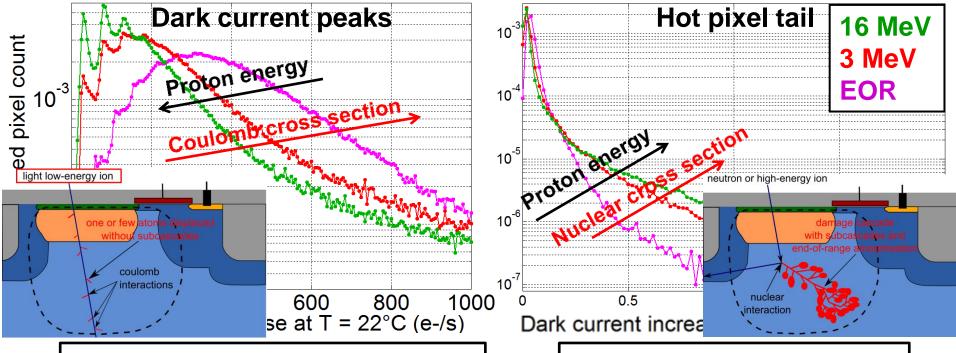




Jean-l

Effect of the coulomb and nuclear cross section





ե, July 12

Low NIEL coulomb interactions

W

Dark current peaks

Nuclear interactions

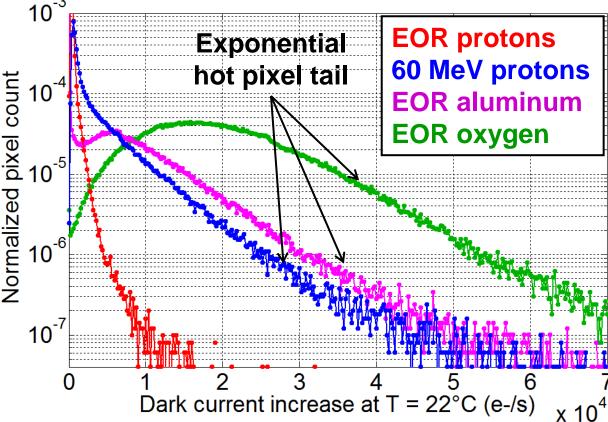
W

Exponential hot pixel tail

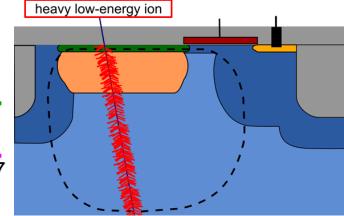


High NIEL coulomb interactions: exponential hot pixel tail





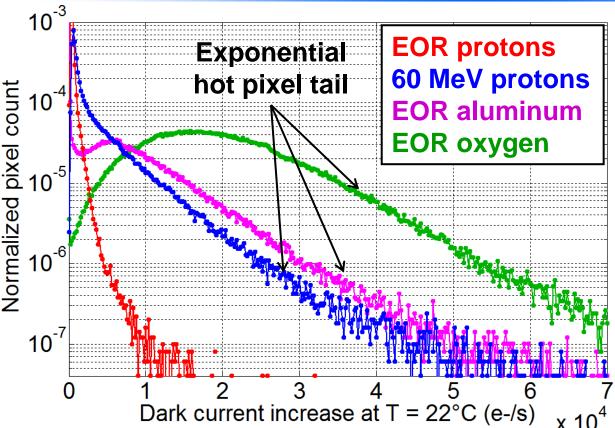
- EOR heavy ions (oxygen and aluminum):
- exponential hot pixel tail





High NIEL coulomb interactions: exponential hot pixel tail



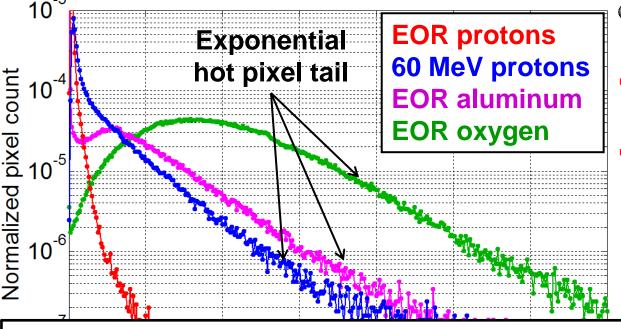


- EOR heavy ions (oxygen and aluminum):
- exponential hot pixel tail
- similar slope than 60 MeV protons (nuclear interactions)



High NIEL coulomb interactions: exponential hot pixel tail





- EOR heavy ions (oxygen and aluminum):
- exponential hot pixel tail
- similar slope than 60 MeV protons (nuclear interactions)

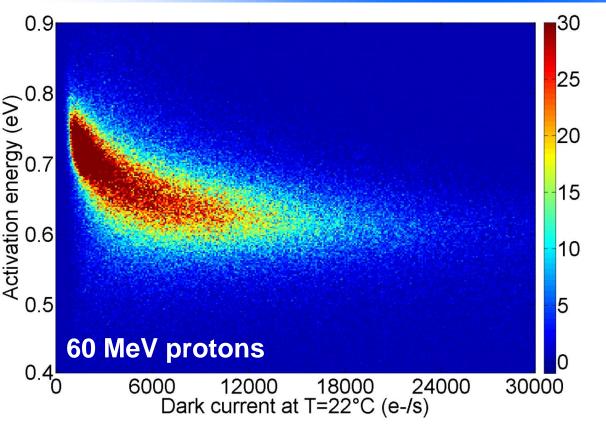
Dense damage (nuclear interactions or high NIEL coulomb interactions)

Exponential hot pixel tail



Dark current activation energy of the exponential hot pixel tail



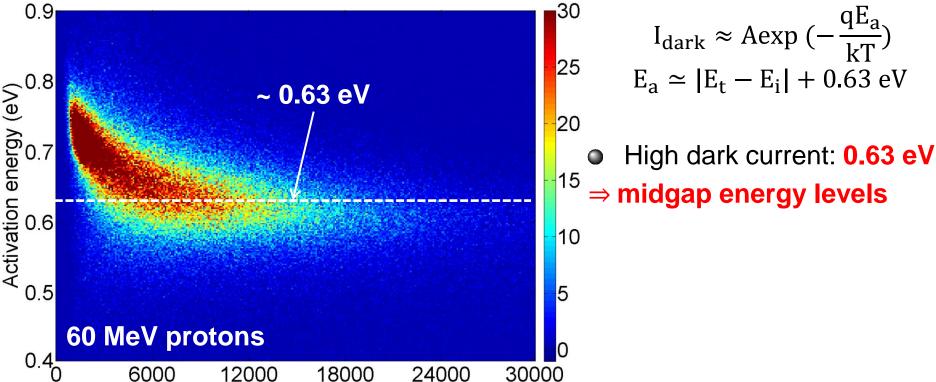


 $I_{dark} \approx Aexp \left(-\frac{qE_a}{kT}\right)$ $E_a \simeq |E_t - E_i| + 0.63 \text{ eV}$



Dark current activation energy of the exponential hot pixel tail



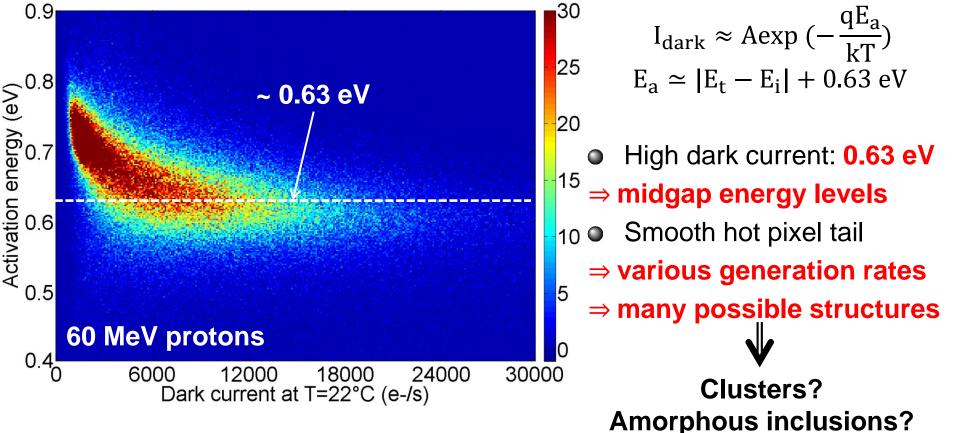


Dark current at T=22°C (e-/s)



Dark current activation energy of the exponential hot pixel tail

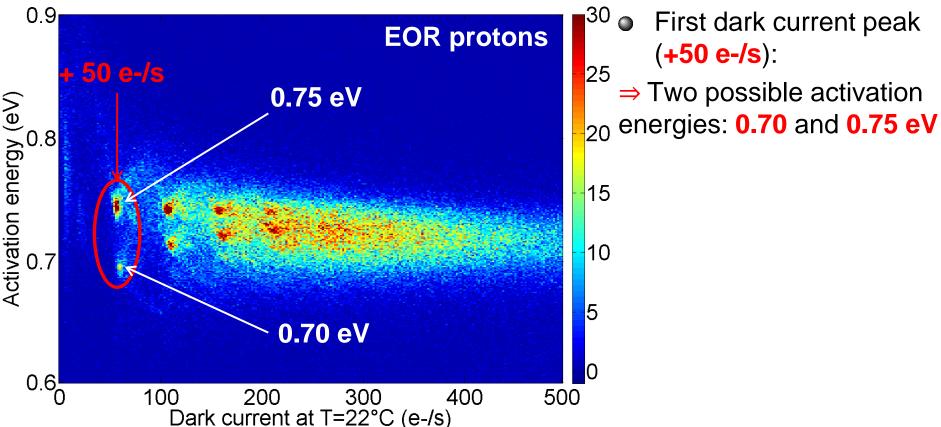






Dark current activation energy of the dark current peaks



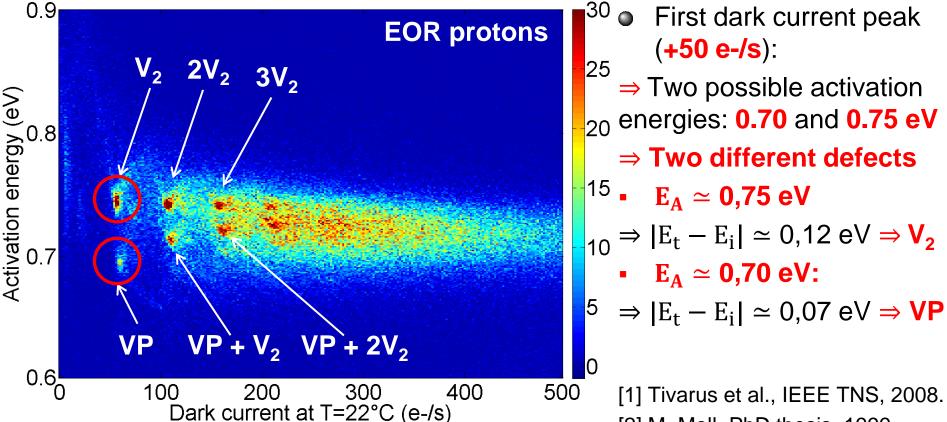




Jean-Marc Belloir

Dark current activation energy of the dark current peaks





(+50 e-/s): ⇒ Two possible activation ₂₀ energies: **0.70** and **0.75 eV** ⇒ Two different defects

$$E_{A} \simeq 0.75 \text{ eV}$$
 $E_{A} \simeq 0.75 \text{ eV}$
 $E_{t} - E_{t} \simeq 0.12 \text{ eV} \Rightarrow V_{2}$

• $E_A \simeq 0.70 \text{ eV}$:

[1] Tivarus et al., IEEE TNS, 2008.

[2] M. Moll, PhD thesis, 1999.

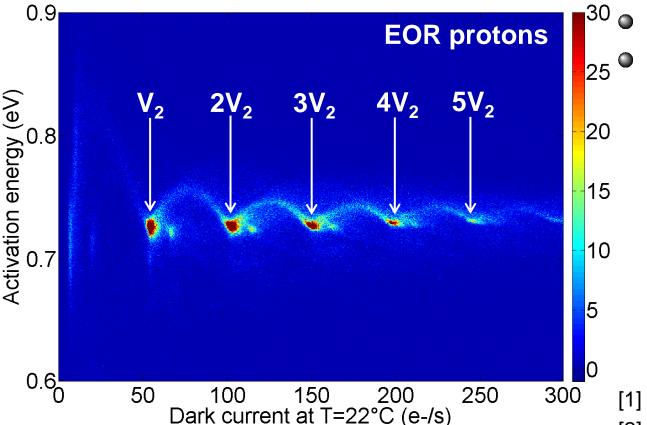


Jean-Marc Belloir

Dark current activation energy after 200°C annealing

NSREC 2016, Portland, July 12, 2016.







VP anneals below 200°C

V₂ is stable at 200°C

[1] Tivarus et al., IEEE TNS, 2008.

[2] M. Moll, PhD thesis, 1999.





The effect of coulomb and nuclear interactions on the dark current distribution have been dissociated:





- The effect of coulomb and nuclear interactions on the dark current distribution have been dissociated:
 - Low NIEL coulomb interactions:
 - Sparse displacement damage
 - Mainly dark current peaks (few 100 e-/s) and few hot pixels





- The effect of coulomb and nuclear interactions on the dark current distribution have been dissociated:
 - Low NIEL coulomb interactions:
 - Sparse displacement damage
 - Mainly dark current peaks (few 100 e⁻/s) and few hot pixels
 - Nuclear interactions or <u>high NIEL</u> coulomb interactions:
 - Dense displacement damage
 - Exponential hot pixel tail with many hot pixels (few 10,000 e⁻/s)





- The effect of coulomb and nuclear interactions on the dark current distribution have been dissociated:
 - Low NIEL coulomb interactions:
 - > Sparse displacement damage
 - Mainly dark current peaks (few 100 e-/s) and few hot pixels
 - Nuclear interactions or high NIEL coulomb interactions:
 - Dense displacement damage
 - Exponential hot pixel tail with many hot pixels (few 10,000 e⁻/s)
- The main radiation-induced defects depend on the damage density:
 - Sparse displacement damage: point defects such as V2 and VP
 - Clustered displacement damage: midgap defects such as clusters







SUPAERO

Thank you for your attention!

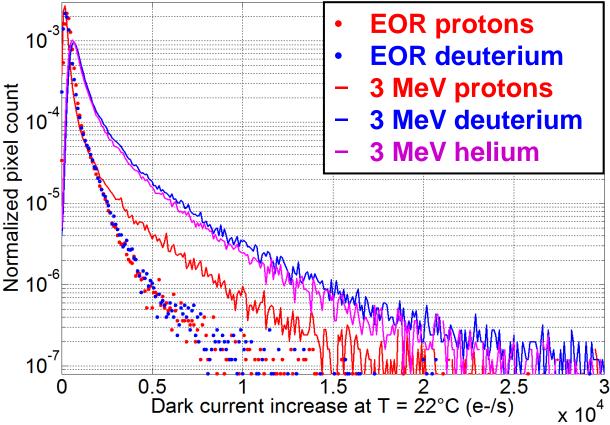
J.-M. Belloir, V. Goiffon, P.Magnan, ISAE-SUPAERO, Toulouse, France,

C. Virmontois, O. Gilard, CNES, Toulouse, France,

M. Raine, P. Paillet, CEA DAM DIF, Arpajon, France.



End-of-range light ions (low NIEL coulomb) | Sale From Sight ions (some nuclear interactions) | Sale From Si



- End-of-range:
 - non-exponential hot pixel tail
 - very few hot pixels
- ⇒ mainly point defects
- 3 MeV:
- exponential hot pixel tail
- less hot pixels than highenergy ions
- [™]3 ⇒ only few clusters

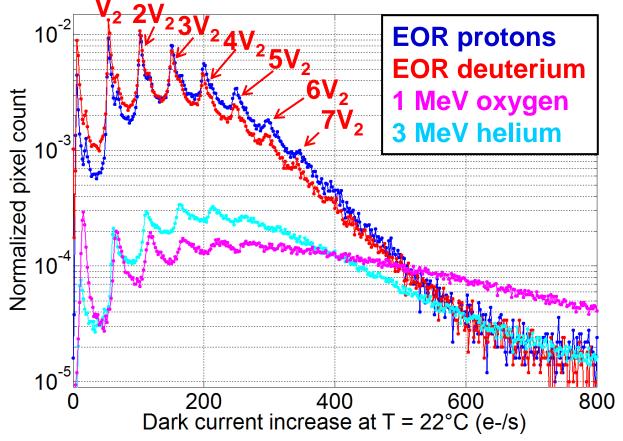
Jean-Marc Belloir

selloir NSREC 2016, Portland, July 12, 2016.



Dark current peaks after 200°C annealing for low energy light ions



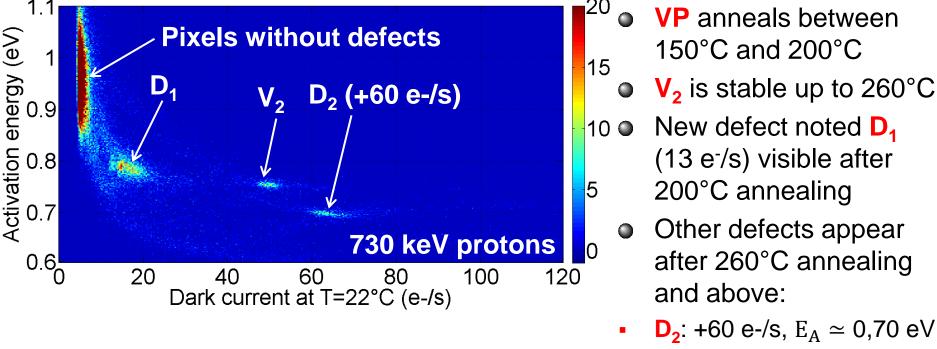




Dark current activation energy after 300°C annealing



 D_3 : +700 e-/s, $E_A \simeq 0.60 \text{ eV}$





Displacement damage density effect



Particle	Maximum coulomb nuclear stopping power (eV/Å)	
End-of-range protons	0.1	Non-exponential hot pixel
End-of-range deuterium	0.2	tail with low probability
End-of-range carbon	11	
End-of-range oxygen	17	Exponential hot pixel tail
End-of-range aluminum	36	Exponential not pixel tall
End-of-range silicon	40	
(nuclear interaction)	40	

- 9 eV/Å ~ 1 displaced atom per lattice plan (E_D ~ 21 eV)
- ⇒ Exponential hot pixel tail (clusters) for clustered displacement damage