

1 Comment on “Low time resolution analysis of polar ice cores cannot detect impulsive nitrate events”  
2 by D.F. Smart et al.

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17 **Key points:**

- 18 • Sharp nitrate spikes in ice cores are not repeatably sampled in parallel cores
- 19 • Most nitrate spikes in previous work cannot be acidic
- 20 • Nitrate in ice cannot be used to document solar energetic particle event statistics

## 21 **Abstract**

22 Smart et al. [2014] suggested that the detection of nitrate spikes in polar ice cores from solar  
23 energetic particle (SEP) events could be achieved if an analytical system with sufficiently high  
24 resolution was used. Here we show that the spikes they associate with SEP events are not reliably  
25 recorded in cores from the same location, even when the resolution is clearly adequate. We explain  
26 the processes that limit the effective resolution of ice cores. Liquid conductivity data suggests that  
27 the observed spikes are associated with sodium or another non-acidic cation, making it likely that  
28 they result from deposition of sea salt or similar aerosol that has scavenged nitrate, rather than from  
29 a primary input of nitrate in the troposphere. We consider that there is no evidence at present to  
30 support the identification of any spikes in nitrate as representing SEP events. Although such events  
31 undoubtedly create nitrate in the atmosphere, we see no plausible route to using nitrate spikes to  
32 document the statistics of such events.

### 33 1. Introduction

34 Large solar energetic particle (SEP) events have the potential to severely disrupt satellite,  
35 communications and electronic systems. There is therefore strong motivation to establish a proxy  
36 that could document the statistics of occurrence of SEP events of different magnitudes, and in  
37 particular the recurrence frequency of the largest events. There has long been a controversy as to  
38 whether spikes in the concentration of nitrate in polar ice cores can be used as such a proxy  
39 [Legrand and Delmas, 1986; Zeller et al., 1986; 1989; McCracken et al., 2001; Palmer et al., 2001;  
40 Wolff et al., 2008]. The perspective of the proponents, coming from the space physics community,  
41 has been that SEP events will produce  $\text{NO}_x$  in the atmosphere, some of which will be deposited in  
42 ice. They have then attempted to align measured spikes with known SEP event dates to establish a  
43 link. The perspective of the opponents, coming from the atmospheric chemistry and ice core  
44 community, has been that any signal would be too small and broad to be detected as a spike, and  
45 that there are other causes of such spikes, unrelated to SEPs.

46 In an attempt to answer the questions posed by the controversy, and to reach both scientific  
47 communities, we presented a study [Wolff *et al.*, 2012] in which numerous high resolution ice core  
48 profiles of nitrate from Greenland and Antarctic ice cores were compiled and compared for a 40 year  
49 period surrounding the well-known Carrington space weather event of 1859. We showed that a  
50 peak corresponding to that event was not present in most cores, and that most nitrate peaks in  
51 Greenland during the 40 year period were due to the transit of biomass burning plumes over the ice  
52 core site (simultaneously depositing ammonium, formate, and to a lesser extent nitrate [Savarino  
53 and Legrand, 1998], along with specific fire tracers).

54 Subsequently, a new paper [Smart *et al.*, 2014] has been published, essentially as a critique of our  
55 paper [Wolff *et al.*, 2012]. This has three main conclusions:

- 56 a. The Carrington Event was a poor choice of test case because it is not clear whether it had  
57 the characteristics that the authors would expect to lead to a significant and sharp nitrate  
58 enhancement.
- 59 b. The resolution typically used to discern nitrate spikes in ice cores is insufficient to discover  
60 the kind of events the authors have in mind.
- 61 c. Analysis at higher resolution and measuring multiple chemical species may allow nitrate  
62 spikes caused by SEPs to be isolated, reinstating the possibility to assess their occurrence  
63 using the ice core record.

64 In this comment we would like to discuss each of these points in turn. Before doing so, we think it  
65 would be helpful to summarise two issues on which we think that the authors of the commented  
66 paper [Smart *et al.*, 2014] and ourselves can now agree, which may not have been so obvious in  
67 earlier discussions.

68 SEPs deposit most of their energy, and therefore produce most NO<sub>x</sub> in the middle stratosphere or  
69 above. A source of NO<sub>x</sub> at these altitudes would take months to years to reach ground level, would

70 be broad and diffuse over time, and would be diluted by other sources. The signature of such a  
71 source could never be a sharp spike deposited almost immediately after the event. It seems  
72 therefore to now be common ground that only hard spectrum events that can deposit energy into  
73 the troposphere could possibly produce the kind of spikes being described, and only nitrate  
74 produced at low altitudes should be seen as spikes.

75 It also is agreed that there are other causes of sharp nitrate spikes in polar ice, and in particular that  
76 some nitrate spikes, previously attributed to SEPs, are in fact caused by biomass burning plumes.  
77 This is critical because it instantly establishes that inventories of nitrate spikes without additional  
78 chemical information cannot be used to establish the statistics of SEP events, and this was actually  
79 the main message intended in our earlier paper [*Wolff et al.*, 2012].

80 We now discuss the three main issues we have identified from the paper on which we are  
81 commenting [*Smart et al.*, 2014].

## 82 2. The role of the Carrington Event

83 We acknowledge that this event may not have deposited energy at low altitudes, and that a nitrate  
84 spike of the kind the authors now propose would therefore not be expected. We emphasise that  
85 our earlier paper [*Wolff et al.*, 2012] used the 40 years around the Carrington Event as an example  
86 period, in which we demonstrated that most peaks previously claimed to be SEP-related actually  
87 have a different origin. Our conclusion about the use of nitrate to identify SEP events did not rely on  
88 the Event itself or alone. However, the apparent coincidence of timing of the largest integrated  
89 nitrate peak in a 400 year period, and the Carrington Event, had previously been used as a major  
90 statistical underpinning of the hypothesis that nitrate spikes were indeed caused by SEPs  
91 [*McCracken et al.*, 2001]. Once that coincidence is removed (because the spike has the signature of  
92 a biomass burning event), so that there is no evidence for any spike associated with the Carrington

93 Event, the idea that any nitrate spikes above background are due to SEPs reverts to being  
94 speculation.

95 We agree that it is worthwhile to assess whether other events, known to have a hard spectrum,  
96 might show a signature, but the idea cannot be considered to have any prior support. In the reply to  
97 this comment [*Smart et al.*, Submitted], it is claimed that there is a nitrate peak in the GISP2-H core  
98 associated with the February 1956 SEP event. However, given the rapidly changing snow  
99 accumulation rate deduced in this part of the core [*Smart et al.*, 2014, Fig. 1] (14 samples in 1955 but  
100 35 in 1956), the dating of the nitrate peak must be considered uncertain within several months, and  
101 the evidence that the nitrate peak is related to the SEP is very weak. As an additional point, our  
102 earlier work does establish that events such as the Carrington, which was associated with a huge  
103 geomagnetic storm, cannot be logged through nitrate.

### 104 3. Ice core resolution

105 Smart et al. [2014] carried out power spectral analysis to assess the resolution of different ice core  
106 records, and on this basis they suggest that many of them were inadequate to detect the kind of  
107 signals that SEPs might cause. Firstly, we comment that the resolution of most of the analytical  
108 systems in use has been characterised directly by applying a rapid change in concentration at the  
109 melter and observing the character of the signal at the detector. This is an empirical and direct way  
110 to discover what kind of smoothing is caused by mixing of water on the melthead, in the tubes  
111 leading to the detector, and in the detector itself. It therefore includes all the sources of dispersion  
112 that would occur for a real sample. The resolutions quoted in Table 1 of our earlier paper [*Wolff et*  
113 *al.*, 2012] were generally derived in this way. It is certainly likely that the system used on the BU  
114 core, analysing only one component, has a higher resolution than the multi-component systems  
115 used at Zoe and D4. However, this analysis ignores many issues concerning the way in which signals  
116 are recorded in ice cores, which inherently limit the resolution that is useful and reliable.

117 Chemicals can be deposited either by wet or dry deposition. If they are dry deposited, they may give  
118 a very thin layer between snowfalls. If they are deposited by wet deposition, then they may give a  
119 signal that is initially the width of a snowfall (typically mm to cm). While there are numerous  
120 snowfall events each year at Summit (84 events were recorded in 2001), they are far from uniformly  
121 distributed through the year [Dibb and Fahnestock, 2004]. This has two consequences: firstly it is  
122 impossible to accurately attach a calendar date to a layer in an ice core, because the snowfall  
123 quantity varies strongly from month to month. Based on Dibb and Fahnestock [2004], linear  
124 interpolation of the quantity of new snow with time within a year would lead to an error of up to  
125 about 2 months at some parts of the year. Secondly, there are months when there are only one or  
126 two snowfall events, so that the effective resolution, even in the deposited fresh snow, is 1-  
127 2/month. With no a priori way of knowing which months this applies to, the reliable resolution (if  
128 one is going to assess event frequencies) in the deposited snow in central Greenland is of order 1  
129 month.

130 However after snowfall, snow is redistributed through wind, leading to mixing of different snowfall  
131 layers, and to inhomogeneous deposition across sastrugi, where some parts of the surface may  
132 contain a thick layer, and the layer may be completely lost from other parts of the surface. Recent  
133 model studies suggest that drifting snow occurs on 50 or more days per year in central Greenland  
134 [Lenaerts *et al.*, 2012, Fig. 8].

135 Finally nitrate that is deposited as nitric acid is known to be mobile in the snow pack, so that even  
136 sharp peaks become smoothed through vapour redistribution. Nitrate that is deposited as aerosol  
137 will not be subject to this process, but suffers from the issue that the deposition rate may be  
138 controlled by atmospheric concentrations of the counter-cation [Wolff *et al.*, 2008; Duderstadt *et al.*,  
139 2014] rather than by concentrations of nitrate itself. For example, an influx of marine air containing  
140 high concentrations of sea salt will scavenge acidic nitrate from the atmosphere, leading to  
141 deposition of a nitrate spike even in the absence of any primary input of nitrate.

142 Of course, despite these issues, an event lasting only a few days can be detected if it is large enough.  
143 Indeed the numerous biomass burning events that have been detected on the basis of markers such  
144 as ammonium, formate [e.g. *Whitlow et al.*, 1994; *Legrand and de Angelis*, 1996], and vanillic acid ,  
145 clearly derive from events that would only have passed over the ice core site for days, but many give  
146 clear nitrate signals [*Wolff et al.*, 2012]. However, a method that would rely on very narrow signals  
147 being reliably present in an ice core at such high resolution cannot be successful.

148 This is illustrated in Smart et al's [2014] own paper. The BU core shows perhaps 5 peaks that could  
149 be identified as sharp nitrate spikes in the period 1937-51. These spikes appear to be 1-2 cm across,  
150 and therefore should clearly show up in the GISP2-H core (which has a resolution of 1.5 cm from  
151 discretely cut samples, which equates to about 0.03 years at this age). We have (Fig. 1) binned the  
152 data from the BU core into 0.03 year sections to mimic what the H core should show if the same  
153 peaks were recorded. We bin rather than smooth [*Smart et al.*, Submitted], because discrete  
154 samples are indeed bins of the ice section. Some years have more than 30 samples in the H core;  
155 however changing our bins to 0.04 or even 0.05 years does not affect the result. Assuming the dating  
156 of both cores is correct, it is immediately obvious that the largest peak in BU (in late 1946), which  
157 should show a peak of 250-340 ppb after binning (depending on the position of the bin boundaries),  
158 is not seen in the H core. Other significant sharp spikes in BU are also not seen at the expected  
159 depth in the H core; although one can find peaks in the H core at ages near to those in BU in some  
160 cases (eg at the end of 1949), they actually occur in the wrong part of the seasonal nitrate curve. The  
161 attempts to find plausible peak matches [*Smart et al.*, Submitted, section 3.3] requires such a flexible  
162 attitude to dating, seasonality and peak size and shape that it is hard to imagine a situation in which  
163 an apparent match would not be found. In fact the GISP2-H core also shows one very clear spike,  
164 but it is in a year when BU shows no spike. Even if these spikes could be attributed to SEPs, they are  
165 not reliably recorded in adjacent cores.

166 We note that the very high resolution Law Dome (Antarctica) ice core was sampled discretely at  
167 about 20-12 samples/year (decreasing with depth), and also showed no significant nitrate  
168 enhancement, let alone a sharp spike, after individual SEP events [Palmer *et al.*, 2001], although a  
169 small, broad, enhancement of nitrate concentrations was found between 3-15 months after the SEP  
170 date, when averaged over the event population. We note also that we see no clear nitrate signal in  
171 the Law Dome DSS core around the  $^{10}\text{Be}$  peak associated with an event in 775 AD [Miyake *et al.*,  
172 2012], recently identified as most likely a hard spectrum event 25-50 times as strong as the February  
173 1956 event [Usoskin *et al.*, 2013].

#### 174 4. A signature of SEP events

175 It has been suggested that replicate sampling of several cores, with multiple chemical species, and at  
176 very high resolution, will allow SEP signals to be isolated [Smart *et al.*, 2014]. In principle, deposition  
177 of acidic nitrate might, at least in the preindustrial era, be indicative of SEP or other events of  
178 atmospheric origin, distinguishable from biomass burning events (with associated ammonium), or  
179 deposition of nitrate as a consequence of scavenging by sea salt or dust. This criterion (of nitrate  
180 peaks with no measured counter-cation) was used to isolate candidate peaks in surface snow in a  
181 recent paper [Duderstadt *et al.*, 2014]. However, if an additional requirement is that the peak  
182 preserved in ice at depth must be very narrow, such as those discussed [Smart *et al.*, 2014], then we  
183 would predict that they must have been deposited as aerosol, and will therefore not have an acidic  
184 signature that might be a fingerprint of an SEP event. No cation information exists for the BU or  
185 GISP2-H core data. However, we can infer the nature of the cation from the liquid conductivity,  
186 which was also presented [Kepko *et al.*, 2009; Smart *et al.*, 2014]. The largest nitrate peak in the BU  
187 core, in 1946, rises about 400 ppb ( $6.5 \mu\text{eq L}^{-1}$ ) above the background. Using well documented ionic  
188 conductances, we would expect such a peak to be accompanied by a liquid conductivity increase of  
189 almost  $3 \mu\text{S cm}^{-1}$  if the nitrate was present as acid ( $\text{HNO}_3$ ). The actual measured increase was about  
190  $0.8 \mu\text{S cm}^{-1}$ , exactly what would be expected in the case where  $\text{Na}^+$  is the counter-cation (and similar



191 for ammonium or calcium). A similar calculation confirms that none of the sharp peaks shown in the  
192 BU core between 1937 and 1951 are acidic.

193 Finally, we need to consider whether it is plausible that the events discussed [Smart *et al.*, 2014]  
194 could arise from a hard SEP event acting in the lower atmosphere. Modelling studies are needed to  
195 assess this. Newly-accepted work [Duderstadt *et al.*, In Press] using the WACCM model found that  
196 very large hypothetical SEP events, 2-3 orders of magnitude larger than those observed in recent  
197 decades, produced, as expected, NO<sub>y</sub> enhancements above the troposphere, but insufficient NO<sub>x</sub> in  
198 the lowest 15 km of the atmosphere to produce spikes in Greenland ice cores. Earlier work [Calisto  
199 *et al.*, 2013] would also not support the kind of enhancements in nitrate deposition flux that would  
200 be needed to observe spikes for events such as those in the 1940s and 1950s. We understand that  
201 other modelling studies are in progress and we await their results with interest, emphasising that, to  
202 be relevant to the production of very sharp spikes, such models will need to focus on production in  
203 the troposphere or lowest stratosphere.

## 204 5. Conclusion

205 Smart *et al* [2014] showed that very sharp nitrate peaks may be detected in a system with very high  
206 resolution. However, there is no evidence that these arise from SEPs. Liquid conductivity data  
207 confirm that many sharp peaks are not acidic, making it likely that they are a by-product of influx of  
208 sea salt, biomass burning or dust aerosol, rather than a consequence of a significant influx of nitrate.  
209 There seems no a priori way to distinguish SEP signals from other causes of nitrate enhancement,  
210 and we again emphasise that the evidence that was used previously, regarding the coincidence in  
211 date of SEP events and nitrate spikes, is no longer valid. In any case, we have shown clearly that such  
212 spikes are not reliably deposited in adjacent cores, for reasons we have discussed, which makes the  
213 use of such sharp features to log the power and frequency of SEP events impractical. We do not  
214 doubt that the largest SEPs will cause an enhancement of nitrate in the middle atmosphere and in  
215 extreme cases the troposphere. As ice core scientists, we would very much like to be able to use a

216 proxy such as that envisaged, but we cannot see a plausible route to identifying and using nitrate  
217 deposition in snow to diagnose past SEPs. The use of  $^{10}\text{Be}$  [e.g. *Usoskin et al.*, 2013], although at  
218 much lower time resolution, seems a much more promising path.

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221

222 **Figure caption**

223 Figure 1. Nitrate in ice cores from the Greenland Summit region. Top: BU core data at its original  
224 resolution. Bottom: In black on the right axis, H core data. In blue offset by 100 ppb (left axis), BU  
225 core binned into sections with the same resolution (0.03 years) as the H core at this age. This was  
226 done with a range of starting points, so that the width of the blue line represents the variability  
227 caused by binning with different boundaries. Down arrows are at the position of prominent BU  
228 peaks, the up arrow marks the most prominent H core peak.

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- 231 Calisto, M., I. Usoskin, and E. Rozanov (2013), Influence of a Carrington-like event on the  
232 atmospheric chemistry, temperature and dynamics: revised, *Environmental Research Letters*,  
233 8(4), doi: 10.1088/1748-9326/8/4/045010.
- 234 Dibb, J. E., and M. Fahnstock (2004), Snow accumulation, surface height change, and firn  
235 densification at Summit, Greenland: Insights from 2 years of in situ observation, *J. Geophys.*  
236 *Res.-Atmos.*, 109(D24), doi: 10.1029/2003jd004300.
- 237 Duderstadt, K. A., J. E. Dibb, C. H. Jackman, C. E. Randall, S. C. Solomon, M. J. Mills, N. A. Schwadron,  
238 and H. E. Spence (2014), Nitrate deposition to surface snow at Summit, Greenland, following  
239 the 9 November 2000 solar proton event, *J. Geophys. Res.-Atmos.*, 119(11), 6938-6957, doi:  
240 10.1002/2013jd021389.
- 241 Duderstadt, K. A., J. E. Dibb, C. H. Jackman, C. E. Randall, N. A. Schwadron, S. C. Solomon, H. E.  
242 Spence, and V. A. Yudin (In Press), Nitrate ions spikes in ice cores are not suitable proxies for  
243 solar proton events, *J. Geophys. Res.*, <http://arxiv.org/abs/1511.03358>.
- 244 Kepko, L., H. Spence, D. F. Smart, and M. A. Shea (2009), Interhemispheric observations of impulsive  
245 nitrate enhancements associated with the four large ground-level solar cosmic ray events  
246 (1940-1950), *Journal of Atmospheric and Solar-Terrestrial Physics*, 71(17-18), 1840-1845, doi:  
247 10.1016/j.jastp.2009.07.002.
- 248 Legrand, M., and M. de Angelis (1996), Light carboxylic acids in Greenland ice: A record of past forest  
249 fires and vegetation emissions from the boreal zone, *J. Geophys. Res.*, 101(D2), 4129-4145.
- 250 Legrand, M. R., and R. J. Delmas (1986), Relative contributions of tropospheric and stratospheric  
251 sources to nitrate in Antarctic snow, *Tellus*, 38B, 236-249.
- 252 Lenaerts, J. T. M., M. R. van den Broeke, J. H. van Angelen, E. van Meijgaard, and S. J. Dery (2012),  
253 Drifting snow climate of the Greenland ice sheet: a study with a regional climate model,  
254 *Cryosphere*, 6(4), 891-899, doi: 10.5194/tc-6-891-2012.
- 255 McCracken, K. G., G. A. M. Dreschhoff, E. J. Zeller, D. F. Smart, and M. A. Shea (2001), Solar cosmic  
256 ray events for the period 1561-1994 1. Identification in polar ice, 1561-1950, *J. Geophys.*  
257 *Res.*, 106(A10), 21585-21598.
- 258 Miyake, F., K. Nagaya, K. Masuda, and T. Nakamura (2012), A signature of cosmic-ray increase in AD  
259 774-775 from tree rings in Japan, *Nature*, 486(7402), 240-242, doi: 10.1038/nature11123.
- 260 Palmer, A. S., T. D. van Ommen, M. A. J. Curran, and V. Morgan (2001), Ice-core evidence for a small  
261 solar-source of atmospheric nitrate, *Geophys. Res. Lett.*, 28(10), 1953-1956.
- 262 Savarino, J., and M. Legrand (1998), High northern latitude forest fires and vegetation emissions over  
263 the last millennium inferred from the chemistry of a central Greenland ice core, *J. Geophys.*  
264 *Res.*, 103(D7), 8267-8279, doi: 10.1029/97jd03748.
- 265 Smart, D. F., M. A. Shea, A. L. Melott, and C. M. Laird (2014), Low time resolution analysis of polar ice  
266 cores cannot detect impulsive nitrate events, *J. Geophys. Res-Space Phys.*, 119(12), doi:  
267 10.1002/2014ja020378.
- 268 Smart, D. F., M. A. Shea, A. L. Melott, and C. M. Laird (Submitted), Reply to "Comment on Low time  
269 resolution analysis of polar ice cores cannot detect impulsive nitrate events by D.F. Smart et  
270 al." by E.W. Wolff et al., *J. Geophys. Res.*
- 271 Usoskin, I. G., B. Kromer, F. Ludlow, J. Beer, M. Friedrich, G. A. Kovaltsov, S. K. Solanki, and L. Wacker  
272 (2013), The AD775 cosmic event revisited: the Sun is to blame, *A&A*, 552, L3.
- 273 Whitlow, S., P. Mayewski, J. Dibb, G. Holdsworth, and M. Twickler (1994), An ice-core-based record  
274 of biomass burning in the Arctic and Subarctic, 1750-1980, *Tellus*, 46B(3), 234-242.
- 275 Wolff, E. W., A. E. Jones, S. J.-B. Bauguitte, and R. A. Salmon (2008), The interpretation of spikes and  
276 trends in concentration of nitrate in polar ice cores, based on evidence from snow and  
277 atmospheric measurements, *Atmos. Chem. Phys.*, 8, 5627-5634.

278 Wolff, E. W., M. Bigler, M. A. J. Curran, J. E. Dibb, M. M. Frey, M. Legrand, and J. R. McConnell (2012),  
279 The Carrington Event not observed in most ice core nitrate records, *Geophys. Res. Lett.*, *39*,  
280 L08503, doi: 10.1029/2012GL051603.

281 Zeller, E. J., G. A. M. Dreschhoff, and C. M. Laird (1986), Nitrate flux on the Ross Ice Shelf, Antarctica  
282 and its relation to solar cosmic rays, *Geophys. Res. Lett.*, *13*(12), 1264-1267.

283 Zeller, E. J., G. A. M. Dreschhoff, and C. M. Laird (1989), A record of solar proton events in a firn core  
284 from Windless Bight, *Ant. J. U.S.*, *24*(5), 92-94.

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