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3		The hydrothermal alteration of cooling lava domes
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23 ABSTRACT

Hydrothermal alteration is a recognized cause of volcanic instability and edifice collapse, 24 including that of lava domes or dome complexes. Alteration by percolating fluids transforms 25 26 primary minerals in dome lavas to weaker secondary products such as clay minerals; moreover, 27 secondary mineral precipitation can affect the porosity and permeability of dome lithologies. The location and intensity of alteration in a dome depend heavily on fluid pathways and availability 28 in conjunction with heat supply. Here we investigate post-emplacement lava dome weakening by 29 30 hydrothermal alteration using a finite element numerical model of water migration in simplified dome geometries. This is combined with the Rock Alteration Index (RAI) to predict zones of 31 32 alteration and secondary mineral precipitation. Our results show that alteration potential is 33 highest at the interface between the hot core of a lava dome and its clastic talus carapace. The 34 longest-lived alteration-potential fields occur in domes with persistent heat sources and permeabilities that allow sufficient infiltration of water for alteration processes, but not so much 35 that domes cool quickly. This leads us to conclude that alteration-induced collapses are most 36 37 likely to be shallow-seated and originate in the talus or talus/core interface in domes which have 38 a sustained supply of magmatic heat. Mineral precipitation at these zones of permeability 39 contrast could create barriers to fluid flow, potentially causing gas pressurization which might 40 promote deeper-seated and larger-volume collapses. This study contributes to our knowledge of 41 how hydrothermal alteration can affect lava domes and provides constraints on potential sites for 42 alteration-related collapses, which can be used to target hazard monitoring.

44 INTRODUCTION

45 Hydrothermal alteration of volcanic edifices promotes weakening and instability, increases the propensity for collapse and can lead to significant volcanic hazards (Voight et al. 46 47 2002; Reid et al. 2002b; McGuire 2003; Carrasco-Núñez et al. 2006; John et al. 2008; del Potro and Hürlimann 2009). The extent of alteration, and magnitude of any resultant collapses ranges 48 49 from those that involve parts of individual lava domes through to major volcanic sector collapses 50 (Siebert 2002). Here we focus on modeling hydrothermal alteration at the scale of an individual lava dome. Our motivation is that we know that hydrothermal systems in and around lava domes 51 52 are pervasive, hazardous and poorly understood. Lava dome internal structure can also be simplified for modeling in a reasonably valid way, which is significantly more difficult for 53 larger, heterogeneous edifices. 54

55 Direct evidence of important hydrothermal systems associated with lava domes include (i) soufrière systems, which are commonplace around the bases of domes (Boudon et al. 1998; 56 57 Walker et al. 2006; Bedrosian et al. 2007; Aizawa et al. 2009) and (ii) mass-flow deposits 58 (including debris-avalanches and debris-flow deposits) sourced from old domes, or dome 59 complexes which contain a high proportion of altered material and clay-rich matrix (Opfergelt et 60 al. 2006; Devoli et al. 2009). In fact much of what we know about alteration in domes comes 61 from the study of these mass-flow deposits, where the hydrothermally altered components are no 62 longer in situ, and their original position in the edifice can only be inferred indirectly from 63 stratigraphy and flow reconstructions. Lava domes sit as variably permeable caps often directly above the volcanic conduits from which they were extruded. Such conduits may be the source of 64 thermal and/or volatile fluxes for extended periods of time after an eruption has ceased 65 66 (Bedrosian et al. 2007; Salaün et al. 2011). This configuration may render lava domes more

67 susceptible to post-emplacement alteration than other more widely dispersed erupted units.
68 Indeed, a conduit-capping dome can host a hydrothermal system that would not otherwise exist
69 had an eruptive phase resulted in a more exposed upper conduit.

70 Lavas within domes can be altered to secondary mineral assemblages (e.g. smectite clays, 71 kaolinite and alunite), weakening the edifice, reducing slope stability, and ultimately resulting in 72 slope failure (Boudon et al. 1998; Opfergelt et al. 2006). Clay-rich alteration materials are not 73 only low-strength but have the potential to absorb and channel groundwater, locally increasing 74 fluid pore pressure and promoting the expansion and/or formation of low-strength zones and 75 exacerbating the risk of slope failure (Voight and Elsworth 1997). Secondary mineral formation 76 (including hydrous silica) may also contribute to internal pressurization of lava domes by 77 reducing gas permeability, thus provoking collapse through explosive decompression (Voight 78 and Elsworth 2000). Upon collapse, clay-hosted pore water can lubricate mass-flows, resulting in 79 the generation of more-mobile and cohesive debris-flows than would be generated by collapse of dry material (Boudon et al. 1998; Reid et al. 2002a; Opfergelt et al. 2006; John et al. 2008). 80

81 Two classic examples of edifice collapse involving hydrothermally altered lava domes/dome complexes are the 1998 event at Casita in Nicaragua (Scott et al. 2005; Opfergelt et 82 83 al. 2006; Devoli et al. 2009) and the 1997 debris-avalanche at Soufriere Hills, Montserrat 84 (Sparks et al. 2002; Voight et al. 2002). In both cases low-strength, low-permeability alteration products are thought to have hosted water, which reduced the effective stresses in, and shear 85 86 strength of, the rocks and which ultimately led to catastrophic destabilization of the edifices. At Casita, a *ca*. 8 ka dacite lava dome complex, a 1.6 million m³ collapse on 30 October 1998 was 87 triggered by intense rainfall associated with Hurricane Mitch. The collapse generated a debris-88 89 flow and lahar that resulted in more than 2500 fatalities (Sheridan et al. 1999; Kerle 2002). The

90 collapsed material was rich in smectite clays formed by intense hydrothermal alteration of the 91 original dacites (Opfergelt et al. 2006). A recent re-analysis of the collapse source area and 92 deposits suggests that all failure surfaces formed at or near an interface between units of 93 overlying volcanic breccia and underlying units of altered, clay-rich pyroclastic deposits and 94 lavas. Clay contents in the altered units were estimated to be 38-50 wt. % of the whole mass and 95 more than 90% of the fine fraction, with water contents in the remaining undisturbed clay-rich 96 material ranging from 56-81% (Devoli et al. 2009). For the 26 December 1997 debris-avalanche at Soufrière Hills Volcano, Montserrat, hydrothermal alteration of the dome-retaining crater wall, 97 98 itself made up of an older dome and pyroclastic deposits, (Galway's Mountain/Soufriere, ~113 99 ka; Harford et al., 2002) was implicated as a major contributor to the destabilization and 100 subsequent collapse and depressurization of the active lava dome (Sparks et al. 2002; Voight et 101 al. 2002). The resultant debris-avalanche deposits contained portions of the new dome but also 102 between 6-15 wt. % kaolinite and smectite group clays, alteration products typical of unsealed 103 acid-sulfate hydrothermal systems. Intact avalanche blocks showed repeating layers of these 104 alteration suites and suggested that the collapse slip surface intersected a layered hydrothermal 105 system. The authors suggested that collapse mechanisms could have included an increase in 106 pore-fluid pressure in the older dome materials due to the presence of low-permeability clay 107 layers (Voight et al. 2002).

While collapses from a given, actively extruding (fresh), lava dome are common during an eruptive phase, those that result from post-emplacement dome weakening by hydrothermal alteration are relatively infrequent, yet may pose very significant hazards that are harder to anticipate. The work reported here aims to increase our understanding of the collapse potential of young but inactive domes. Only sparse information exists on the collapse frequency of inactive,

113 altered lava domes, either young or old. But the increasing number of mapped debris-avalanche 114 deposits (many of which source from lava dome complexes) and the similar post-eruptive 115 processes at work in both lava domes and larger volcanic edifices suggests that it is important to 116 understand alteration in these systems. Salaün et al. (2011) and Friant et al. (2006) have mapped 117 debris-avalanche deposits and potential source areas on the Grande Découverte-Soufrière 118 volcano and lava dome in Guadeloupe, which indicate that the recurrence interval of such 119 collapses may be as high as one per 1000 years over the last 8 ky. These collapses have resulted 120 in debris-avalanche deposits rich in hydrothermally-altered material and Salaün et al. (2011) 121 suggest that hydrothermal alteration in the domes and flows that erupted after each collapse was 122 rapid and widespread..

123 Geophysical and geochemical investigations have been used to characterize the 124 hydrothermal system of specific lava domes at given points in time (Bedrosian et al. 2007; Finn 125 et al. 2007; Finn and Deszcz-Pan 2011; Brothelande et al. 2014). Such studies are logistically 126 challenging and cannot provide information about temporal variations of a system unless they 127 are repeated. At the Mount St. Helens lava dome (Bedrosian et al. 2007), electrical resistivity 128 surveys revealed that meteoric water circulated in the young dome due to heat input from a near-129 surface magmatic source, but did not capture longer-term changes occurring in the hydrothermal 130 system as the dome subsequently cooled. Aeromagnetic and electromagnetic surveys of Cascade 131 volcanoes (Finn et al. 2007; Finn and Deszcz-Pan 2011) and the La Soufriere volcano 132 (Brothelande et al. 2014) have indicated the presence of water and altered material in specific 133 locations in the edifices at the times of the surveys, but provide limited information about flow 134 pathways within the hydrothermal systems and how they might be expected to have evolved or 135 evolve in the future..

136 In this paper we present a physics-based numerical model of heat and fluid flow in a 137 generic lava dome combined with assessments of alteration potential, to determine where 138 alteration is most likely to occur within a lava dome. The dynamics of hydrothermal flow and 139 alteration processes in domes can be highly complex, and in detail each dome is unique; we do 140 not address all possible complexities but make a first step in quantitative modeling of major 141 aspects of the systems. While we do not seek to address the specific type of alteration in this 142 study, it is possible to distinguish likely regions of alteration based on knowledge of temperature 143 gradients and fluid flux. We use this understanding to make some inferences about the different 144 collapse styles and source areas that may occur in different hydrothermal settings. This work lays 145 the groundwork for future investigations to identify likely alteration minerals, in order to 146 distinguish whether individual collapses are related to weak alteration minerals (such as clays) or 147 precipitation that reduce the porosity/permeability of the upper dome (such as silica).

148

149 FINITE ELEMENT HEAT AND MASS TRANSFER CODE (FEHM)

150 The effect of meteorically-derived water on the hydrothermal system of cooling lava 151 domes is modeled using The Finite Element Heat and Mass transfer (FEHM) code. FEHM is an 152 extensively validated (Zyvoloski et al. 1999; Dash et al. 2003; Dash 2003) porous flow simulator 153 capable of modeling the flow of heat, water, air, and water vapor in a variably saturated porous 154 and/or fractured medium at temperatures up to 1500°C and fluid pressures of up to 1000 MPa. 155 The code employs a Newton-Raphson scheme to iteratively solve discretized conservation 156 equations for mass, energy and momentum for fluid and vapor on a Voronoi-conforming finite-157 volume computational mesh (Miller et al. 2007; Zyvoloski 2007). Basic governing equations of 158 state for conservation of mass and energy and modified Darcy flux are shown in Figure 1; for a

detailed derivation, see the supplemental material. The temperature and pressure ranges of FEHM are ideal for modeling volcanic systems and although it has previously been used to model cooling pyroclastic deposits (Keating 2005), volcanic seamount discharge and recharge (Hutnak et al. 2006), and mineral alteration in hydrothermal fault systems (Chaudhuri et al. 2009), it has never been applied to volcanic edifices or lava dome systems.

164 Active emplacement of a dome, whether endogenous or exogenous, is a principle forcing 165 mechanism for collapse (Calder et al. 2002; Calder et al. 2005);. However in systems where 166 emplacement has paused or ceased, other forcing mechanisms, including environmental ones, 167 come into play (Calder et al. 2005; Barclay et al. 2006). Our focus here is on young but 168 inactive/cooling lava dome systems. The rationale for this choice includes a number of 169 considerations. First, only in inactive domes is hydrothermal alteration likely to play a significant 170 role in collapses, and the relevant alteration minerals are most stable at temperatures of 200°C or 171 less (Ball et al., 2013; ;Giggenbach 1992). Second, in this initial application of FEHM, we 172 required a relatively simple system where lava extrusion, and its associated heterogeneous mass 173 and thermal fluxes can be negated. Furthermore, restricting the simulations to lower temperatures 174 (< 200°C) allows us to neglect high-temperature heat sources, which cause sudden short-term 175 phase changes that increase model instability and prevent convergence on a solution for the 176 governing equations (Ingebritsen et al. 2010).

177

178 Modeling methods

Our simulations are run for 100 years on a simplified lava dome geometry consisting of five material regions: A narrow conduit, a solid dome core, a clastic dome carapace or talus (covering and also tapering into a wedge away from the dome), a solid substrate, and colluvium 182 overlying the substrate (Figure 2). The conduit radius was selected based on inferences made for 183 actual domes as well as rheological models of dome eruptions (Fink and Pollard 1983; Costa et 184 al. 2007). Previous models of dome growth and evolution (Fink and Griffiths 1998a; Hale 2008; 185 Hale et al. 2009a; Hale et al. 2009b) and structural relationships of existing lava domes (Wadge 186 et al. 2009) were used to design the configuration of the material regions used in our models. 187 Two dome geometries are represented: a "crater-confined" dome similar to the domes at the 188 Santiaguito lava dome complex in Guatemala (Ball et al. 2013), which provided some of the 189 original motivation for this study, and a "perched" dome, essentially erupting on a sloping 190 substrate whose core and talus are unconfined by a crater wall, similar to that of the Unzen or 191 Merapi lava domes (Smithsonian Institution 1991; Anderson et al. 1999; Nakada et al. 1999; 192 Walter et al. 2013). These dome geometries were investigated to determine if there was a 193 difference in fluid migration when the dome was confined by a crater wall vs. unconfined.

194 Voronoi-conforming finite-volume computational meshes of these geometries were 195 generated with the LaGriT Grid Generation Toolset, which was developed at Los Alamos 196 National Laboratory (Fields et al. 1996; Miller et al. 2007). LaGriT accepts input files consisting 197 of coordinate data defining material regions, as well as commands choosing the level of 198 discretization in those regions (or in other subregions specified by the user). The dome meshes 199 for this study use a two-dimensional axisymmetric coordinate system; each mesh is a slice from 200 the center of a dome, including a thick substrate that is extended down to several thousand 201 meters in order to avoid boundary effects. The crater-confined and perched dome models are 202 triangulated to (i.e. to have a spatial resolution of) 20 meters within the top 1500 meters of the dome and substrate; the remaining (lower) 1000 meters of the mesh are triangulated to 40 meters 203

to save computational time and allow FEHM to calculate processes in the dome on a moredetailed scale.

206 Material properties (density, porosity, initial permeability, specific heat, and thermal 207 conductivity) of each dome region were taken from ranges given in the literature, summarized in 208 Table 1. All values were chosen from lithologies commonly associated with lava domes (dome 209 rocks, lava flows, and block-and-ash-flow deposits) of andesitic and dacitic composition. In most 210 of these studies the values were determined from hand sample and drill core analyses. Where 211 literature values had a wide range, a restricted range was chosen for modeling based on the most 212 commonly found values. A complete record of the values from which these ranges were defined 213 may be found in the supplementary material.

214 Although the boundaries of large lava dome structures such as shear lobes are well-215 described (Fink and Griffiths 1998b; Watts et al. 2002), there is a limited amount of structural 216 information available on the geometry of smaller, denser fracture networks in domes or their 217 carapaces (i.e., the orientation, depth of penetration, concentration of fractures in a given area, 218 aperture width of the fractures). Therefore, in this study we have chosen to treat material regions 219 as continua where the properties of the porous and/or fractured medium are averaged to account 220 for variations that are not captured in our mesh. The appropriateness of a continuum approach 221 depends on the ratio of the fracture density scale to the flow region scale. Khaleel (1989) 222 modeled two-dimensional planar laminar flow through the columnar-fractured Columbia River 223 Basalt Group lava flows and determined that for interconnected networks of filled/unfilled 224 fractures of uniform aperture and column diameters of 1 m, continuum models were appropriate for length scales of at least 6 times the column diameter. That author also indicated that a 225 226 continuum approach could be appropriate for other fractured rock masses if the entire rock mass

is at least 6 times the smallest spacing of the fractures, and preferably ~20-30 times the spacing in the case of variably sized and distributed fracture networks. The scale of fracturing and structural features on a lava dome is much smaller than the scale of dome and immediate substrate (even shear lobes of 10's to a few 100's m can be considered close to the 1/6th cutoff point), and so we feel that the continuum approach is resonable for our simplified domes.

232

233 Precipitation/recharge

234 Because the actual elevation of water tables within volcanic edifices is not well 235 constrained (Hurwitz 2003), we initiate model runs for these domes by assigning complete 236 saturation.. Variable saturation is then allowed to develop as the simulation runs. This 237 assumption is considered reasonable for volcanic systems/domes located in a tropical region that 238 receive significant (i.e. > 1000 mm/yr) amounts of precipitation, such as the lava dome 239 volcanoes detailed in Table 3. Previous models of groundwater in volcanoes using similar 240 recharge and permeability values (Join et al., 2005; Hurwitz et al., 2003) show that water tables 241 may rise to high levels within an edifice. As such, it is a reasonable assumption that there may be 242 conditions under which a lower-temperature (cooling) lava dome could become almost entirely 243 saturated (for example during an intense precipitation event). Even if there is an unsaturated zone 244 in the immediate rubbly surface of the dome, the models would be unable to represent it if it was 245 smaller than the 20 m mesh spacing. Given the paucity of information on water tables associated 246 with lava domes, water-saturation is taken as a reasonable first-order approach.

Ground surface recharge in the models (1300 mm/yr, or ~4.16 x 10^{-7} m³/[m² s]) is approximated using yearly precipitation rates in tropical regions. Actual infiltration rates into a dome would be expected to be less than precipitation due to evaporation, vegetation, localized

impermeable surfaces, and runoff, but such variations would be site-specific, localized and impractical to represent in these simulations. Additionally, Rad et al. (2007) state that infiltration in volcanic settings with exposed fresh lavas and pyroclastic flows may be as high as 80%. Thus we consider that using a recharge value at the low end of measured precipitation rates for five existing domes in tropical to temperate climate regions is a reasonable proxy for infiltration (Table 3).

256

257 Boundary and initial conditions

For simplicity of the simulations, a number of material properties and boundary 258 259 conditions were held constant throughout the simulations (Table 3). These include porosity, 260 density, thermal conductivity, and specific heat of the material region, and pressure, temperature 261 and recharge along the atmospheric surface (Zone 6, Figure 2). FEHM automatically treats the 262 boundaries of the modeling mesh as impermeable unless otherwise specified, so left side (Zone 263 7, Figure 2) and basal (Zone 8, Figure 2) no-flow boundary conditions are assigned in these 264 simulations. This reflects a radial dome geometry and a dome-topped volcanic edifice resting on 265 an impermeable base (equivalent to plutonic or metamorphic rock, which can have permeabilities as low as 1×10^{-16} to 1×10^{-18} m²; Brace, 1984). 266

Initial conditions for the material regions in all simulations are detailed in Table 4. Initial rock permeabilities are divided into 'low' permeability systems $(1.0 \times 10^{-15} \text{ m}^2 \text{ core} \text{ and}$ substrate, $1.0 \times 10^{-13} \text{ m}^2$ talus), 'intermediate' permeability systems $(1.0 \times 10^{-14} \text{ m}^2 \text{ core} \text{ and}$ substrate, $1.0 \times 10^{-12} \text{ m}^2$ talus) and 'high' permeability systems $(1.0 \times 10^{-13} \text{ m}^2 \text{ core} \text{ and} \text{ substrate}, 1.0 \times 10^{-11} \text{ m}^2$ talus) and 'high' permeability systems $(1.0 \times 10^{-13} \text{ m}^2 \text{ core} \text{ and} \text{ substrate}, 1.0 \times 10^{-11} \text{ m}^2$ talus) values. Two thermal conditions are used for simulations. One condition allows the dome and conduit to cool over time from their initial temperature conditions 273 (specified in Table 4 for each material region). The other maintains a constant 200°C heat source 274 in the conduit, simulating a recharging magmatic heat source such as a dike or other intrusion 275 providing heat to the system, slow solidification with release of latent heat, or heat fluxing from 276 outgassing. Relative permeabilities vary with time and saturation according to a linear 277 formulation (Zyvoloski et al. 1999; Zyvoloski 2007) which uses a residual liquid saturation of 278 0.2 and a residual vapor saturations at 0.1 (see supplementary material for linear function 279 equations). Again, we focus here on low-temperature domes because adding fluids to the model 280 under high temperature conditions results in the model attempting to simulate abrupt phase 281 changes and substantially slowing the modeling process. High temperatures may also result in 282 extremely low saturations in the modeling domain, which again slow calculations. We also note 283 that the temperature of formation of many of the minerals of concern with respect to edifice 284 weakening and instability (particularly smectite clay) fall within the 100-300°C temperature 285 range, which makes this a logical point in the temperature evolution of a cooling dome to 286 investigate. At higher temperatures, the alteration mineral suite changes significantly and 287 becomes dominated by minerals like biotite, actinolite, chlorite, and silica polymorphs (Henley 288 and Ellis 1983; Reyes 1990), which are not usually implicated in alteration-related edifice collapse. 289

290

291 Alteration potential determined with the rock alteration index (RAI)

Volumetric liquid and vapor fluxes are calculated directly within FEHM, while temperature gradients are post-processed based on the FEHM temperature fields. The temperature field and liquid volumetric flux vectors are then used to calculate alteration potential with the Rock Alteration Index (RAI) for all nodes in the models at regular time intervals for each simulation. The RAI was developed to describe the likelihood of alteration processes
occurring based on temperature gradients and patterns of fluid migration in subsurface aqueous
reservoirs. Steep temperature gradients along flow paths promote alteration by inducing
thermodynamic instability in the system and encouraging the formation of secondary minerals in
equilibrium with thermal conditions (Raffensperger and Vlassopoulos 1999; Wetzel et al. 2001).
The RAI is calculated from the temperature gradient and the fluid volumetric flux:

302

$$\mathbf{RAI} = \mathbf{q} \cdot \nabla \mathbf{T} \tag{1}$$

where **q** is the fluid volumetric flux (volume per cross-sectional area per time $(m^3/(m^2 s))$, and T 303 304 is the temperature (°C). High positive values of the RAI coincide with areas of higher fluid flux 305 and/or flow paths of rapidly increasing temperature and would promote mineral dissolution. 306 High negative values are found in areas of higher fluid flux and/or flow paths of rapidly 307 decreasing temperature and would promote mineral precipitation. An RAI value of zero does not 308 necessarily reflect zero fluid flow, but can also indicate flow along isotherms. It should be noted 309 that the RAI as calculated is an indicator of alteration potential only; it can be combined with 310 geochemical species models to determine the likelihood of specific mineral formation, but for 311 this initial study we focus on alteration potential only.

312

313 **RESULTS**

Twelve simulations were undertaken, varying the thermal and permeability conditions for each of the two dome geometries as described previously (Table 4). Temperature results highlight the differing evolution of temperature profiles in domes with a heat source and without, as in dome eruptions where magma either remains in some reservoir beneath the dome or has withdrawn. Variable permeability conditions were investigated because permeability is a primary 319 control on fluid flux (both direction and intensity); we are also trying to represent differences in 320 permeability between materials as well as cover the overall range of reasonable permeabilities 321 based on rock property measurements (see previous section). As temperature distribution and 322 fluid flux are the dependent variables by which the alteration indices are calculated, knowing 323 how they vary in space and time allows us to comment on why certain patterns of RAIs develop.

- 324
- 325

Temperature gradients and liquid/vapor flux patterns

326 The highest temperatures in the simulations persist in areas that are initially warm (dome 327 cores and conduits) and decrease over time; higher-than-background temperatures progressively 328 migrate downward into the substrate (Figure 3). In both dome geometries, domes with lower 329 initial permeabilities tend to cool slower and do not achieve background temperatures by the end 330 of the simulations, while initially higher permeability domes cool more rapidly, reaching 331 background temperatures by or before the 100-year time limit. Likewise, in domes without a 332 maintained heat source, both the onset of cooling as well as reaching background temperatures 333 occurs earlier than for domes with a heat source. A persistent 200°C heat source in the conduit 334 generally allowed the lower portions of the domes to remain hotter for longer periods of time, 335 provided permeabilities were low. The inner cores of the domes remain hot longer in simulations 336 with conduit heat, but there is little effect on outlying areas of the dome and talus. Crater-337 confined domes cool more slowly than perched domes, with perched domes cooling completely 338 by the 50-year mark whereas crater-confined domes still retain zones of elevated temperatures in 339 their cores (Figures 3e and 3f.) One persistent feature in every simulation is the development of a 340 zone of warmer temperatures (usually about 40-70°C) at the base of the dome (Figures 3c and d, 341 initial development), which migrates toward the break in slope created by the intersection of the

342 dome talus and slope talus. These zones develop faster in domes with medium or high343 permeabilities, but persist longer in lower permeability domes.

344 Liquid flux patterns are dominated by gravitational flow and show no evidence of 345 convection (Figure 4). The highest liquid fluxes (up to an order of magnitude higher than the 346 recharge flux) are found in both the talus of both the dome and slopes throughout the 100-year 347 simulations; liquid fluxes decrease where liquid saturations are low. In most simulations, this 348 decrease occurs at a front which migrates from the top of the dome core (Figures 5a and b) to its 349 center, and around the head of the conduit (Figure 5g). Over time, high fluid fluxes decrease to 350 background levels. Overall, fluxes vary between much less than the meteoric influx (min. ~ 1.0 $x10^{-8}$ m³/m² s) to an order of magnitude higher. Maximum fluxes occur in simulations where the 351 initial permeabilities are high (1.0 x 10⁻¹² m² for talus, 1.0 x 10⁻¹³ m² for dome core and 352 353 substrate). Perched domes contain larger areas of higher fluxes (Figure 4b) than crater-confined 354 domes (Figure 4a), which coincide with the location of talus layers beneath the domes. These 355 talus layers divert flow noticeably under the perched domes (Figures 4d-h), while in crater-356 confined domes this effect is much less pronounced and not as long-lived (Figure 4e). (Adding a 357 200°C heat source to the models does not appear to have a significant effect on fluid fluxes (i.e. 358 through increasing buoyancy of fluids, etc.), but indirect factors like low saturation zones related 359 to a heat source (which reduce relative permeability and thus fluid flux), may be masking lesser 360 effects.

Vapor fluxes (Figure 5) appear at the beginning of a simulation and a boiling point front (Figures 5a and b) migrates from the core/talus interface at the summit of the dome into the core of the dome as time progresses (Figures 5c - f). This front follows the high temperature gradient between the hot core of the dome and the recharge-cooled carapace. Once the dome has cooled,

365 vapor fluxes are confined to the area around the conduit head (Figures 5e, f and g), although they 366 only persist in simulations where the conduit is maintained at high temperature. The addition of a 367 persistent heat source does not increase the strength of the vapor fluxes but does affect the fluxes 368 indirectly (by increasing the areas that are not fully saturated and thus increasing the area over which vapor fluxes are produced). Vapor volume fluxes range from 1.0×10^{-6} to 9.0×10^{-6} m³ m⁻ 369 2 s⁻¹ and are higher when overall permeabilities are higher. Vapor fluxes are initially higher in 370 371 crater-confined domes than perched domes (Figures 5a and 5b), but more widespread in perched 372 domes. By the 50-year mark, vapor fluxes in both domes are approximately the same magnitude 373 (Figures 5e and 5f).

374

375 **Resulting alteration potential**

376 Alteration potential (Figure 6) in both domes is controlled primarily by the magnitude 377 and direction of fluid flux and the temperature gradients of the cooling domes, and is thus 378 intimately related to permeability contrasts in materials and the availability of heat. Because the 379 highest liquid fluxes occur in the higher-permeability talus layers (Figures 4a and b), and the 380 greatest temperature gradients are generated between the dome cores and the cooler substrates 381 (Figures 3a and b), the highest positive and negative RAI values occur at the interface between 382 the core and talus of the dome (Figures 6a and b). High positive values at the dome summits 383 suggest potential for mineral dissolution (alteration). Conversely, high negative values at the base of the dome and around the slope break in the talus indicate potential for mineral 384 385 precipitation. Crater-confined domes show more intense positive and negative RAI values 386 initially (Figures 6a and 6b), and are the only geometry that still shows non-zero RAI values in 387 the core of the dome by the 10-year simulation time (Figures 6c and 6d). Neither geometry has 388 non-zero RAI values in the dome after the 50-year simulation time (Figures 6e - h).

389 Permeability and thermal conditions also control the strength and longevity of alteration 390 zones. Potential alteration is more intense with low to moderately high permeabilities, but is not 391 sustained for long unless a conduit heat source is present. Alteration potential decreases 392 dramatically in these models within even the first 10 years (Figures 6c and d). However, very 393 high permeabilities preclude strong alteration at the talus/core interface and instead, high RAI 394 values are only generated around the conduit. Lower permeabilities combined with a heat source 395 sustain both positive and negative RAI values longer, but at lower intensities than those 396 developed at higher permeabilities. In both dome geometries, negative RAI values at the base of 397 the dome tend to persist longer than the positive ones in the upper dome (Figures 6c and d). The 398 ideal combination for forming sustained, localized areas with high RAI values (positive or 399 negative) appears to be a maintained conduit heat source combined with low to intermediate permeabilities (1 x 10^{-14} m² for core and substrate and 1 x 10^{-13} m² for talus). This enables water 400 401 flux across strong temperature gradients without allowing the dome to cool too quickly.

402

403 **DISCUSSION**

Domes cool from 150 to ~ 30°C within 100 years in these simulations, suggesting that the lifetime of a low-temperature hydrothermal system in a tropical lava dome is only years to decades if the dome lacks a sustained high-temperature magmatic heat source. A sustained 200°C heat source prolongs cooling, but 200°C appears to be insufficiently hot to effectively drive hydrothermal circulation within the domes. Perched dome geometries cool faster than crater-confined domes, likely due to the presence of high-permeability talus layers beneath portions of the dome that allow more water transport around the dome core. 411 Overall, higher permeabilities and fluid fluxes promote more-intense early-onset RAIs; 412 however, these values are not as persistent as those in domes of lower permeabilities. Crater-413 confined geometries retain intense RAIs somewhat longer than perched domes because of the 414 presence of high-permeability talus layers beneath perched domes (which helps cool the domes 415 faster), but geometry exerts a much weaker control than permeability. Therefore early-onset 416 RAIs might be expected in pervasively fractured domes, but longer-lasting RAIs would be found 417 in more-coherent domes. For both modeled dome geometries, high positive RAIs are formed 418 near the summit while negative RAIs occur at the base of the domes. This indicates that 419 dissolution processes are more likely during infiltration and precipitation processes are more 420 likely as water percolates out the base of a dome. In these models, the latter location is occupied 421 by talus, but it could also include features such as brecciated zones at the base of shear 422 lobes/flows (John et al. 2008); for example, the suites of hydrothermal alteration minerals in 423 clay-rich lahars on Mount Rainier are often sourced from brecciated lavas and pyroclastic 424 deposits with high primary permeabilities (Crandell 1971; Scott and Vallance 1995; John et al. 425 2008). However, because water permeates through domes – which are composed of both 426 fractured lavas and brecciated talus – relatively quickly, it is also possible that the water would 427 not have the opportunity to form advanced alteration assemblages. This is consistent with 428 observations at Santiaguito, where hot springs represent immature waters that have dissolved, but 429 not equilibrated with, dome rock or formed minerals such as clays, and alteration on the dome 430 surfaces is limited to the deposition of hydrous silica(Ball et al. 2013).

These results now provide a framework for investigations of specific mineral formation,
for example, by combining temperature and flux data with aqueous geochemical data using a
species model such as EQ3/6 (Wetzel et al. 2001). The simulation results suggest that alteration

434 minerals forming from dissolution processes (clays and kaolinite) are more likely to be found at 435 the dome summit, while minerals resulting from precipitation processes (alunite) are more likely 436 found at the base of the dome and talus around the slope break. Clays and alunite could 437 contribute to shallow-seated collapses of the carapace and in the talus. Collapses triggered by 438 talus erosion and dome undermining have been associated with intense rainfall (Calder et al. 439 2005). It might be that increased fluid pressure in clay-bearing horizons contributes to increased 440 instability associated with precipitation. In sufficient quantity, clays, alunite and silica minerals 441 could also form impermeable layers, facilitating gas sealing (Elsworth et al. 2004), a mechanism 442 that has been suggested for some deep-seated dome failures in active lava domes. Precipitation 443 of vapor-phase cristobalite in domes has been shown to decrease porosity (and by inference, also 444 permeability) of dome rock (Horwell et al. 2013). A similar effect could occur in low-445 temperature systems, such as modeled here, with other mineral precipitates (Figure 7); however, 446 more-complex models accounting for factors such as fracture networks and high-permeability 447 zones such as shear lobe boundaries would be necessary for a complete evaluation of this 448 scenario.

449 The presence of hot, magmatic acidic gases rising from a magma source will accelerate 450 alteration of the dome rock (Reyes 1990). However, Cox and Browne (1998) note that large-451 scale alteration of rock to smectite/montmorillonite is still possible even in neutral-to-alkaline pH 452 systems. Additionally, while advanced argillic alteration contains a suite of weak sulfate 453 minerals (such as alunite and jarosite) that require the presence of sulfur gases to form, clay 454 minerals such as smectite also depend on the presence of water. Water vapor is the primary gas released in any volcanic eruption and meteoric water composes a significant percentage of the 455 456 fluid available in any near-surface hydrothermal system (Goff & Janik, 2000) and it is reasonable 457 to assume that the degassing pathways followed by water vapor would be, at the shallow levels 458 depicted in these models, essentially the same as those traveled by acidic gases released from a 459 magmatic body. The behavior of water and water vapor are therefore useful both as first-order 460 information and as proxies for interpretations about additional gases in lava dome hydrothermal 461 systems.

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3 Comparison to existing domes and volcanoes

464 These simulations represent a first order approach into estimating the behavior of liquid 465 water and water vapor in the interior of cooling lava domes. Physically validating this 466 assessment of alteration is somewhat difficult, since there is currently little in the way of direct 467 or remotely sensed field data about the hydrothermal systems of lava domes. Physical mapping of dissected domes (Duffield et al. 1995; Riggs and Carrasco-Nunez 2004) is generally limited to 468 469 structural features or eruptive facies and neglects information about alteration mineral 470 assemblages or the location of hydrothermal flow paths or fumaroles. Remote sensing is 471 necessarily limited to surface materials, and while debris-avalanches may sample the interior of 472 domes, the percentages and types of altered material involved in volcanic collapses have not 473 been recorded in a systematic or detailed manner in most studies (Dufresne 2009 and personal 474 communication). Voight et al. (2002) were able to trace hydrothermally altered material in the 475 1996 collapse of the Soufriere Hills lava dome to the margins of the dome and buttressing crater 476 wall, suggesting that permeability contrasts did indeed come into play (the altered material 477 included pyroclastic deposits of brecciated lava). However in large collapses the deposits may 478 consist of entire lava domes, making it difficult, if not impossible, to reconstruct internal 479 structures and zones of alteration. As a result, the best available data about undisturbed dome 480 interiors comes from geophysical investigations.

481 Nicollin et al. (2006) completed an electrical tomographic study of La Soufriere of 482 Guadeloupe volcano and lava dome and created cross-sections of electrical resistivity indicating 483 likely areas of hydrothermal alteration (altered materials tend to be less resistive/more 484 conductive). They determined that there was a large area of low resistivity located at the base of 485 the dome, which they interpreted as a hydrothermally altered zone; in addition, a medium 486 resistivity layer on the western flank of the lava dome was found at depth below a layer of lower 487 resistivity. Other zones of low resistivity were interpreted as cross-cutting faults which provided 488 pathways for liquids and gases that promoted alteration. The authors interpreted the basal layer 489 as related to the collapse of a highly-altered summit of the volcano (producing the crater in 490 which the dome formed), but mentioned that the western low-resistivity zone could represent 491 massive unaltered areas of the lava dome overlaid by a layer of thick better-drained scoriaceous 492 or altered material from the dome carapace or brecciated units formed during the formation of 493 flow lobes. This would be consistent with the RAI interpretations reached in this model, where 494 alteration is likely to be concentrated at the interface between higher and lower permeability 495 regions (higher permeabilities allow the passage of more liquid and vapor, which are essential to 496 alteration processes). A recently published study by Brothelande et al. (2014) expands on this, 497 indicating that hydrothermal ascending flows in the volcano are limited to the dome and its 498 immediate proximity, confined in a collapse structure surrounding the dome (last modified in 1530 by a collapse followed by an eruption). 499

500 Finn et al. (2007) and Finn and Deszcz-Pan (2007) conducted helicopter magnetic and 501 electromagnetic surveys of Mounts Adams and Baker in order to determine the three-502 dimensional geometry of altered and saturated regions within those Cascade stratovolcanoes. 503 Hydrothermal alteration significantly reduces the magnetization in volcanic rocks, while the

504 presence of alteration minerals, or of water in the rock, reduces electrical resistivity; by 505 combining these methods, the authors were able to discriminate areas of dry, fresh rock from 506 saturated fresh or weakly altered rock and variably saturated intensely altered rock (Finn et al. 507 2007; Finn and Deszcz-Pan 2011). On Mount Adams, intensely altered and saturated rock is 508 found in the core of the volcano, but layers of fresh or weakly altered saturated rock also underlie 509 portion of the volcano's slopes. The authors interpreted these as 'fresh porous breccias', similar 510 to the kind of material found in talus units at lava domes (Finn et al. 2007). Alteration at Mount 511 Baker follows a different pattern, being restricted to thinner layers beneath the summit crater and 512 fumarole field on the volcano's northeast flank (Finn and Deszcz-pan 2011), but similar un- or 513 weakly-altered saturated layers underlie its slopes. While these methods were applied on an 514 entire stratovolcano and did not discern features on the scale of lava dome structures, ground-515 based application of magnetic and electromagnetic surveys could provide that information at the 516 scale of a lava dome. Muon radiography (Tanaka et al. 2007; Lesparre et al. 2012) is an 517 emerging geophysical method which shows promise in determining density contrasts in lava 518 domes, but it would be necessary to determine if altered material showed significant density 519 contrasts with unaltered material, and if it was distinguishable from density differences in 520 dome/conduit/talus material in the first place. Currently, muon radiography studies have been 521 able to locate conduits within and beneath domes, but are limited by the fact that the method 522 must encompass the whole thickness of the dome and cannot take a "slice" from it as in 523 resistivity studies.

In order to corroborate the results of this study, future field and geophysical investigations of both old and young domes would be useful, with attention paid to the location, degree and character of alteration.

527

528 CONCLUSIONS

This study applies a multi-phase porous flow model to determine the flow of water and heat in low-temperature cooling lava domes over 100-year timescales. A number of conclusions can be drawn from these first-order simulations of lava dome hydrothermal systems.

1. The alteration potential in these domes is controlled by the contrasts in material permeability and the heat sources driving hydrothermal flow, and is highest where permeability contrasts are greatest, particularly at the interface between the less permeable dome core and more permeable talus. This suggests that alteration mineral formation is most likely to occur at the boundaries of lava dome structures.

537 2. Areas of increased alteration likelihood are sustained longer in low-permeability 538 domes, but are more intense in domes with higher permeabilities and persistent heat sources. A 539 dome without a sustained heat input will cool on geologically short time periods, and even faster 540 if its overall permeability is high, denying the opportunity to develop alteration. However, at the 541 low temperatures and high infiltration rates in these simulations there is no evidence for 542 convection of water in the domes and flow is dominated by gravity, precluding the possibility of 543 long-lived hydrothermal circulation.

3. Potential for dissolution (clay mineral formation) is highest near the summit at the core/talus interface of the simulated domes, while the potential for mineral precipitation (alunite, silica formation) is highest at the base of the domes. If alteration forms weak minerals at the core/talus interface the area could source shallow-seated collapses of the carapace.

548 In combination with geophysical and field studies, numerical modeling can provide an
549 important first step in elucidating the behavior of post-eruptive volcanic systems. Incorporating

550 the results of numerical models with limited ground-based data and remote sensing can 551 strengthen the interpretations drawn from both, and provide valuable insight into dome evolution and hazards. Coupling these first-order flow models with mineral species models and 552 553 representing more complex dome structures and different fluid chemistries could allow 554 simulations such as these to be used to evaluate potential collapse mechanisms at specific domes. 555 The results of the models presented here provide a framework for future investigations, including 556 field, geochemical and geophysical, into the way post-eruptive lava domes are altered by 557 hydrothermal activity.

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765 FIGURE CAPTIONS

Fig. 1 Basic equations that combine to form the full governing partial differential equations in
FEHM, as reported in Zyvoloski et al. (1999). For a full derivation and comments, see the
supplemental material.

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Fig. 2 Schematic of the two modeling grids for (A) crater-confined lava domes (e.g. Santiaguito)
and (B) perched lava domes (not confined by a crater, e.g. Unzen).

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Fig. 3 Temperature outputs for the crater-confined (left) and perched (right) dome geometries
after 1 year for intermediate permeabilities and a conduit heat source. (Figures illustrate a subset
of the full modeling domain above 1500 m elevation).

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Fig. 4 Fluid flux outputs for crater-confined and perched dome geometries at 1, 10, 50, and 100
year intervals for intermediate permeabilities and a conduit heat source. Blues indicate higher
fluxes and greens lower fluxes. Streamlines indicate flow direction only and are arbitrarily
spaced.

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Fig. 5 Vapor flux outputs for crater-confined and perched dome geometries at 1, 10, 50, and 100
year intervals for intermediate permeabilities and a conduit heat source. Higher fluxes are
indicated by reds and areas of zero vapor flux by white.

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Fig. 6 RAI patterns for crater-confined and perched domes at 1, 10, 50, and 100 year intervals.
RAI magnitude scales differ between the two dome geometries, but warm colors indicate
positive RAIs (dissolution) and cool colors indicate negative RAIs (precipitation) in both.

Fig. 7 Summary of dome alteration and potential collapse loci based on RAI patterns. Alteration mineral formation is most likely to occur at the talus/core interface early in the lifetime of the dome, and depending on the mineral species involved could either strengthen or weaken the dome and/or promote internal gas pressurization.

Independent variable	Dome range	Talus range	Notes	References
Porosity (φ)	0.1-0.5	0.1-0.6	Talus layers are likely to be more porous than core (porosity increases with increasing fragmentation, void space)	(García et al. 1989; Ingebritsen and Hayba 1994; Alt-Epping et al. 2001; Barmin et al. 2002; Hurwitz 2003; Keating 2005; Bartetzko et al. 2006; Flint et al. 2006; Scheu et al. 2006; Bernard et al. 2007; Mueller et al. 2008; Watanabe et al. 2008; Ikeda et al. 2008; Aizawa et al. 2009; Hicks et al. 2009; Wicks et al. 2011)
Permeability (m ²)	1.0 x 10 ⁻¹⁹ to 1.0 x 10 ⁻¹²	1.0 x 10 ⁻¹⁷ to 1.0 x 10 ⁻ ¹⁰	Talus layers are likely to be much more permeable (connected to porosity; see previous) due to higher porosity, fractured/rubbly state	(Sammel et al. 1988; Sekioka 1988; Ingebritsen and Hayba 1994; Barmin et al. 2002; Keating 2005; Bartetzko et al. 2006; Bernard et al. 2007; Mueller et al. 2008; Watanabe et al. 2008; Aizawa et al. 2009; Hicks et al. 2009; Platz et al. 2012)
Density (kg/m ³)	593-2890		Dome and talus are assumed to be composed of the same material	(García et al. 1989; Smith et al. 2001; Hurwitz 2003; Keating 2005; Scheu et al. 2006; Bernard et al. 2007; Watanabe et al. 2008; Ikeda et al. 2008; Hicks et al. 2009)
Thermal conductivity (W/m K)	0.537- 3.430	0.35	Value for most volcanic rocks is ~2	(Sekioka 1988; García et al. 1989; Ingebritsen and Hayba 1994; Hurwitz 2003; Keating 2005; Hicks et al. 2009)
Specific heat (J/kg K)	730-1557	2036	Dome and talus are assumed to be composed of the same material	(Keating 2005; Hicks et al. 2009; Platz et al. 2012)

Table 1. Summary of ranges for material properties derived from literature review.

Parameter	Value
Porosity of Zones 2 and 3	0.3
Porosity of Zones 1 and 5	0.2
Porosity of Zone 4	0.1
Initial temperature of Zone 1	150 °C
Initial temperature of Zone 2	30 °C
Initial temperature of Zone 3	30 °C
Initial temperature of Zone 4	200 °C
Initial temperature of Zone 5	70 °C
Permeability of Zone 4	1 x 10 ⁻¹⁶ m ²
Thermal conductivity of all units	2.0 W m ⁻¹ K ⁻¹
Rock density of all units	2600 kg m ⁻³
Specific heat	1000
Atmospheric temperature (Zone 6)	25 °C
Atmospheric pressure (Zone 6)	0.1 MPa
Precipitation rate (Zone 6)	1700 mm/yr

Table 2. Constant simulation parameters and boundary conditions

Dome/ Location	Min rate (mm/yr)	Max rate (mm/yr)	Average (mm/yr)	References
Soufrière Hills Montserrat, W.I.	1250	2000	1625	(Barclay et al. 2006; Hemmings et al. 2015)
Soufrière of Guadeloupe Guadeloupe, WI	??	10000	10000	(Le Friant et al. 2004)
Merapi Central Java, Indonesia	2000	4500	3250	(Lavigne et al. 2000)
Casita Nicaragua			1250	(Velázquez and Gómez-Sal 2007)
Unzen Japan	2000	2600	2300	(Ogawa et al. 2007)
Galeras Columbia			1200	Meteorologia Aeronautica (Instituto de Hidrologia, Meteorologia y Estudios Ambientales) di Columbia (<u>http://www.meteoaeronautica.gov.co/</u>)
Santiaguito Guatemala	1800	4000	2900	(Lopez 2004)

Table 3. Yearly precipitation rates for different currently active, or recently active, lava domes.

Run	Geometry	Permeability Dome Talus (Zone 2)	Permeability Dome Core (Zone 1)	Permeability Slope Talus (Zone 3)	Permeability Substrate (Zone 5)	Conduit Heat
1	Crater-	1 x 10 ⁻¹² m ²	1 x 10 ⁻¹³ m ²	1 x 10 ⁻¹² m ²	1 x 10 ⁻¹³ m ²	
2	confined	$1 \ge 10^{-13} \text{ m}^2$	1 x 10 ⁻¹⁴ m ²	$1 \ge 10^{-13} \text{ m}^2$	1 x 10 ⁻¹⁴ m ²	Initial heat only
3		1 x 10 ⁻¹⁴ m ²	1 x 10 ⁻¹⁵ m ²	1 x 10 ⁻¹⁴ m ²	1 x 10 ⁻¹⁵ m ²	
4	Crater-	1 x 10 ⁻¹² m ²	1 x 10 ⁻¹³ m ²	1 x 10 ⁻¹² m ²	1 x 10 ⁻¹³ m ²	
5	confined	$1 \ge 10^{-13} \text{ m}^2$	1 x 10 ⁻¹⁴ m ²	$1 \ge 10^{-13} \text{ m}^2$	1 x 10 ⁻¹⁴ m ²	200°C sustained
6		1 x 10 ⁻¹⁴ m ²	1 x 10 ⁻¹⁵ m ²	1 x 10 ⁻¹⁴ m ²	1 x 10 ⁻¹⁵ m ²	
7	Daughad	1 x 10 ⁻¹² m ²	1 x 10 ⁻¹³ m ²	1 x 10 ⁻¹² m ²	1 x 10 ⁻¹³ m ²	
8	Perched	$1 \ge 10^{-13} \text{ m}^2$	1 x 10 ⁻¹⁴ m ²	$1 \ge 10^{-13} \text{ m}^2$	1 x 10 ⁻¹⁴ m ²	Initial heat only
9		1 x 10 ⁻¹⁴ m ²	1 x 10 ⁻¹⁵ m ²	1 x 10 ⁻¹⁴ m ²	1 x 10 ⁻¹⁵ m ²	
10		1 x 10 ⁻¹² m ²	1 x 10 ⁻¹³ m ²	1 x 10 ⁻¹² m ²	1 x 10 ⁻¹³ m ²	
11	Perched	$1 \ge 10^{-13} \text{ m}^2$	1 x 10 ⁻¹⁴ m ²	$1 \ge 10^{-13} \text{ m}^2$	1 x 10 ⁻¹⁴ m ²	200°C sustained
12		1 x 10 ⁻¹⁴ m ²	1 x 10 ⁻¹⁵ m ²	1 x 10 ⁻¹⁴ m ²	1 x 10 ⁻¹⁵ m ²	

Table 4. Individual simulation parameters

Figure 1









Elevation (meters)





Figure 7



1 Governing equations

2 In FEHM, the governing partial differential equations for mass and heat transfer are 3 discretized into a system of nonlinear algebraic equations, which are then solved using the 4 Newton-Raphson iteration method (a way to find successively better approximations to the roots 5 or zeroes of a real-valued function from an initial guess). In FEHM (as opposed to some other 6 multi-physics computer codes like TOUGH2 (Pruess et al. 2012), the Newton-Raphson 7 derivatives of the thermodynamic functions with respect to pressure and temperature are formed 8 analytically rather than numerically in order to achieve faster convergence of the nonlinear 9 system of equations (Zyvoloski 2007). FEHM equations of state are nonlinear because the 10 porosity, permeability, density, enthalpy and viscosity are strong functions of pressure and 11 temperature; in addition, relative permeabilities and capillary pressure can also be strong 12 functions of saturation, which varies significantly with temperature and pressure. Pressure and 13 temperature dependent behavior of density, enthalpy and viscosity are represented by rational 14 polynomials derived from National Bureau of Standards (NBS) steam table data (Zyvoloski et al. 15 1991).

All variables referenced in these equations are defined in Table 5. A more detailed breakdown of the equation derivation may be found in Zyvoloski et al. (1999) and Keating ((2000).

- 20
- 21
- 22

$A_{e,m}$	Energy/mass/noncondensible gas accumulation	R	Relative			
gas	terms (kg m ⁻¹ s ⁻²)/ (kg m ⁻³) / (kg m ⁻³)		permeability			
C_p	Specific heat $(m^2 s^{-2} \circ C^{-1})$	S	Saturation			
fe,m,gas	flux vectors for energy/mass/noncondensible gas equation (kg s ⁻³)/ (kg m ⁻² s ⁻¹)/ (kg m ⁻² s ⁻¹)	Τ	Temperature			
g	gravitational acceleration (m s ⁻²)	t	time			
h	enthalpy $(m^2 s^{-2})$	v	Velocity (m s ⁻¹)			
k	intrinsic rock permeability (m ²)	ϕ	Porosity			
K	Thermal conductivity (kg m T ⁻¹ s ⁻³)	$\eta_{vap,liq}$	Mass fraction of vapor/liquid			
Р	Pressure (kg $m^{-1} s^{-2}$)	μ	Viscosity (kg m ⁻¹ s ⁻¹)			
q_e	Energy/mass/noncondensible gas source term (kg m ⁻¹ s ⁻³)/ (kg m ⁻³ s ⁻¹)/ (kg m ⁻³ s ⁻¹)	ρ	Density			
Subscripts: gas = noncondensible gas, vap = vapor, liq = liquid, f = fracture, r = rock						

23 Table 5. Variables used in FEHM governing equations

24

25 Conservation of mass for water is

$$\frac{\partial A_m}{\partial t} + \overline{\nabla} \cdot \overline{f}_m + q_m = 0 \tag{1}$$

27 where the mass per unit volume A_m (a mass accumulation term) is a function of the porosity, the

28 saturation and the mass fraction of each phase:

29
$$A_m = \phi(S_{vap}\rho_{vap}(1-\eta_{vap}) + S_{liq}\rho_{liq}(1-\eta_{liq}))$$
(2)

30 the mass flux is

31
$$\bar{f}_m = (1 - \eta_{vap})\rho_{vap}\bar{v}_{vap} + (1 - \eta_{liq})\rho_{liq}\bar{v}_{liq}$$
(3)

and sources/sinks of mass are contained in the term q_m . The total of all water flux through the system, water accumulated in the system, and sources/sinks of water are assumed to equal zero.

34 Conservation of fluid-rock energy is

35
$$\frac{\partial A_e}{\partial t} + \overline{\nabla} \cdot \overline{f_e} + q_e = 0 \tag{4}$$

36 where the energy per unit volume (energy accumulation term) A_e is a combination of energy 37 transferred by the rock itself, energy transferred by water vapor and energy transferred by liquid 38 water:

39
$$A_e = (1 - \phi)\rho_f C_{pr}T + \phi(S_{vap}\rho_{vap}C_{pvap}T + S_{liq}\rho_{liq}C_{pliq}T)$$
(5)

40 The energy flux, \bar{f}_e , can be stated either as a sum of the products of density, specific enthalpy 41 and velocity of the vapor and liquid, or the product of the thermal conductivity and temperature 42 gradients:

43
$$\bar{f_e} = \rho_{vap} h_{vap} v_{vap} + \rho_{liq} h_{liq} v_{liq} = K \bar{\nabla} T$$
(6)

44 The conservation of noncondensible gas is described by

45
$$\frac{\partial A_{gas}}{\partial t} + \overline{\nabla} \cdot \overline{f}_{gas} + q_{gas} = 0$$
(7)

46 where the accumulation term A_{gas} is

47
$$A_{gas} = \phi(\eta_{vap}S_{vap}\rho_{vap} + \eta_{liq}S_{liq}\rho_{liq})$$
(8)

48 the gas flux is

49
$$\bar{f}_{gas} = \eta_{vap} \rho_{vap} \bar{v}_{vap} + \eta_{liq} \rho_{liq} \bar{v}_{liq}$$
(9)

50 and the source and sink term q_{gas} is

$$q_{gas} = \eta_{vap} q_{vap} + \eta_{liq} q_{liq} \tag{10}$$

It is assumed that Darcy's law applies to the movement of each phase; in this formulation, the hydraulic conductivity term is replaced by the quotient of the intrinsic and relative permeabilities divided by the viscosity of the phase, and the hydraulic head term is replaced with the difference between the pressure gradient and lithostatic gradient:

56
$$\bar{v}_{vap} = -\frac{kR_{vap}}{\mu_{vap}} (\bar{\nabla}P_{vap} - \rho_{vap}\bar{g})$$
(11)

57
$$\bar{\nu}_{liq} = -\frac{kR_{liq}}{\mu_{liq}} (\bar{\nabla}P_{liq} - \rho_{liq}\bar{g})$$
(12)

58 By combining equations 1 - 10 with Darcy's law (11 & 12), the full governing equations are 59 derived for mass,

$$60 \qquad -\overline{\nabla} \cdot \left((1 - \eta_{vap}) \Theta_{mvap} \overline{\nabla} P_{vap} \right) - \overline{\nabla} \cdot \left((1 - \eta_{liq}) \Theta_{mliq} \overline{\nabla} P_{liq} \right) + q_m +$$

$$61 \qquad \frac{\partial}{\partial z}g\left(\left(1-\eta_{vap}\right)\Theta_{mvap}\rho_{vap}+\left(1-\eta_{liq}\right)\Theta_{mliq}\rho_{liq}\right)+\frac{\partial A_m}{\partial t}=0 \tag{13}$$

62 Energy,

63
$$-\overline{\nabla} \cdot \left(\Theta_{evap}\overline{\nabla}P_{vap}\right) - \overline{\nabla} \cdot \left(\Theta_{eliq}\overline{\nabla}P_{liq}\right) + q_e + \frac{\partial}{\partial z}g\left(\Theta_{evap}\rho_{vap} + \Theta_{eliq}\rho_{liq}\right) + \frac{\partial A_{\eta gas}}{\partial t} = 0$$

(14)

64

65 and noncondensible gas:

$$66 \qquad -\overline{\nabla}\cdot\left(\eta_{vap}\Theta_{mvap}\overline{\nabla}P_{vap}\right) - \overline{\nabla}\cdot\left(\eta_{liq}\Theta_{mliq}\overline{\nabla}P_{liq}\right) + \overline{\nabla}\cdot\left(D_{va}\overline{\nabla}\eta_{vap}\right) + q_{\eta gas} +$$

67
$$\frac{\partial}{\partial z}g(\eta_{vap}\Theta_{mvap}\rho_{vap} + \eta_{liq}\Theta_{mliq}\rho_{liq}) + \frac{\partial A_{\eta gas}}{\partial t} = 0$$
(15)

68 Transmissibilities, a measure of how easily a given layer transmits water, are represented by Θ in 69 the equations and are given by

70
$$\Theta_{mvap} = \frac{kR_{vap}\rho_{vap}}{\mu_{vap}}$$
(16)

71
$$\Theta_{mliq} = \frac{kR_{liq}\rho_{liq}}{\mu_{liq}}$$
(17)

72
$$\Theta_{\text{evap}} = h_{\text{vap}}\Theta_{\text{mvap}}$$
(18)

73
$$\Theta_{eliq} = h_{liq} \Theta_{mliq} \tag{19}$$

74 The air/water diffusivity term D_{va} is given by

75
$$D_{va} = \tau \phi S_{vap} D_{va}^0 \rho_{vap} \frac{0.101325}{P} \left[\frac{T + 273.15}{273.15} \right]^m$$
(20)

where τ is the tortuosity factor (for diffusion in porous media, a measure of how circuitous a typical flow path is through a medium), *D* (standard conditions) = 2.4e-5 m²/s and *m* = 2.334 (Zyvoloski et al. 1997). Transmissibilities are given by

79
$$\Theta_{mvap} = \frac{kR_{vap}\rho_{vap}}{\mu_{vap}}$$
(21)

80
$$\Theta_{mliq} = \frac{kR_{liq}\rho_{liq}}{\mu_{liq}}$$
(22)

81
$$\Theta_{evap} = h_{vap}\Theta_{mvap}$$
(23)

82
$$\Theta_{eliq} = h_{liq} \Theta_{mliq}$$
(24)

The nonisothermal multiphase transport model in FEHM is formed by equation 13, 14 and 15, setting n = 0 in Equation 13 and leaving out Equation 15 in the pure-water model.

85

86 Relative permeability

In cases where the system being modeled in FEHM does not remain fully saturated, FEHM is capable of calculating relative permeability using either a linear, Corey-Brooks (Brooks and Corey 1964), or van Genuchten formulation (van Genuchten 1980). The linear formulation used in this study depends only on the residual liquid and vapor saturations and was chosen for simplicity. Linear functions for relative permeabilities of liquid (the wetting fluid, in this case water) and vapor (the 'nonwetting fluid') are defined by the equations

93
$$R_{l} = \begin{cases} 0, & S_{l} \leq S_{lr} \\ \frac{S_{l} - S_{lr}}{S_{lmax} - S_{lr}} & S_{lr} < S_{l} < S_{lmax} \\ 1, & S_{l} \geq S_{lmax} \end{cases}$$
(27)

94
$$R_{v} = \begin{cases} 0, & S_{v} \leq S_{vr} \\ \frac{S_{v} - S_{vr}}{S_{vmax} - S_{vr}} & S_{vr} < S_{v} < S_{vmax} \\ 1, & S_{v} \geq S_{vmax} \end{cases}$$
(28)

where S_l and S_v are liquid and vapor saturations, S_{lr} and S_{vr} are the residual liquid and vapor saturations, and S_{lmax} and S_{vmax} are the maximum liquid and vapor saturations.

- 97
- 98

Reference	Location	Sample size/type	Rock type	Density (kg/m ³)	Porosity (φ)	Permeability (m ²)	Thermal conductivity (W/mK)	Specific heat (J/kgK)
(Aizawa et	Conceptual	Unknown	Sealing zone		0.01	3.00E-16	(,)	(0,8/
al. 2009)		Unknown	Basement		0.05	3.00E-16		
		Unknown	Hydrothermal zone		0.1	3.00E-15		
		Unknown	Fractured rock surrounding		0.3	1.00E-13		
		Unknown	hydrothermal zone Conduit		0.1	1.00E-12		
		Unknown	Fresh volcanic rock		0.3	1E-13 - 2E-14		
(Bartetzko et al. 2006)	Undersea basin, PACMANUS field	Drill core	Dacite		0.22	1.25E-17		
		Drill core	Dacite		0.32	2.23E-17		
		Drill core	Dacite		0.24	4.48E-17		
		Drill core	Dacite		0.3	1.04E-16		
		Drill core	Dacite		0.24	4.46E-16		
		Drill core	Dacite		0.38	7.59E-16		
		Drill core	Dacite		0.21	1.50E-15		
		Drill core	Dacite		0.43	2.00E-15		
		Drill core	Dacite		0.01	1.17E-14		
		Drill core	Dacite		0.16			
		Drill core	Dacite		0.17			
		Drill core	Dacite		0.17			
		Drill core	Dacite		0.2			
		Drill core	Dacite		0.21			
		Drill core	Dacite		0.22			
(Bernard et	Mount Pelée	Hand	MB502	2700	0.038	1.00E-16		
al. 2007)		sample Hand sample	LPP Calebasse	2850	0.031	1.00E-15		
		Hand	LPP Plume	2890	0.035	1.00E-15		
		sample Hand sample	andesite, B&A flows	2780	0.099	1.00E-15		
		Hand	MB501	2740	0.102	1.00E-15		
		sample Hand sample	LPP Macouba	2760	0.134	1.00E-15		
		Hand	andesite dome	2720	0.098	2.00E-15		
		Hand sample	andesite, nuee ardente flows	2690	0.119	2.90E-15		
		Hand	andesite, indurated block and ash flows	2720	0.103	3.00E-15		
		Hand	andesite, indurated	2690	0.146	3.00E-15		
		Hand sample	andesite, indurated block and ash flows	2720	0.111	3.90E-15		
		Hand	andesite dome	2740	0.121	3.90E-15		
		sample Hand sample	andesite dome	2740	0.154	3.90E-15		

Table 6. Material properties of andesite and dacite lavas and pyroclastic deposits

Hand	andesite, nuee ardente	2670	0.157	3.90E-15
sample Hand	flows andesite, indurated	2670	0.353	3.90E-15
sample Hand	block and ash flows andesite, indurated	2770	0.145	4.90E-15
sample Hand	block and ash flows andesite, nuee ardente	2670	0.169	5.90E-15
sample Hand	nows andesite, indurated	2710	0.132	8.90E-15
Hand	andesite, nuee ardente	2670	0.19	1.09E-14
Hand	andesite, nuee ardente	2670	0.099	1.88E-14
Hand	andesite, nuee ardente	2690	0.158	3.85E-14
Hand	andesite, indurated	2720	0.149	5.53E-14
Hand	andesite, nuee ardente		0.312	5.92E-14
Hand	andesite, nuee ardente	2670	0.174	1.09E-13
sample Hand	andesite, ash-and-	2650	0.574	1.44E-13
sample Hand	andesite, nuee ardente	2700	0.197	1.67E-13
sample	andesite, nuee ardente	2690	0.146	2.04E-13
sample Hand	flows andesite, ash-and-	2630	0.547	2.07E-13
sample Hand	pumice flows andesite, indurated	2620	0.213	2.27E-13
sample Hand	block and ash flows andesite, indurated	2760	0.24	2.54E-13
sample Hand	block and ash flows andesite, ash-and-	2650	0.536	4.25E-13
sample Hand	pumice flows andesite, indurated	2620	0.232	5.54E-13
sample Hand	block and ash flows andesite, indurated	2630	0.321	5.70E-13
sample Hand	block and ash flows andesite, nuee ardente	2670	0.415	6.70E-13
sample Hand	flows andesite, scoria flows	2810	0.28	1.03E-12
sample Hand	andesite, ash-and-	2670	0.574	1.18E-12
sample Hand	pumice flows andesite, nuee ardente	2700	0.304	1.18E-12
sample Hand	flows andesite, nuee ardente	2670	0.428	1.45E-12
sample Hand	flows andesite, scoria flows	2860	0.363	1.54E-12
sample Hand	andesite, nuee ardente	2690	0.251	1.58E-12
sample Hand	flows andesite indurated	2680	0.305	1 77E-12
sample	block and ash flows	2670	0.578	2.05E-12
sample	pumice flows	2680	0.274	2.03E 12
sample	flows	2000	0.274	2.91E-12
sample	block and ash flows	2710	0.289	5.00E-12
sample	andesite, scoria flows	2840	0.352	5.69E-12
Hand sample	andesite, ash-and- pumice flows	2650	0.617	7.66E-12
Hand sample	andesite, block and ash flows	2710	0.328	1.02E-11
Hand sample	andesite, block and ash flows	2670	0.408	3.44E-11

		Hand sample	andesite, ash-and-	2670	0.585			
(Flint et al.	Yucca	Drill core	Pyroclastic unit	1490	0.341			
2006)	Mountain	Drill core	Pyroclastic unit	1600	0.322			
(García et al.	Los Azufres	Drill core	Andesite	2053	0.02		1.05	
1989)		Drill core	Andesite	2737	0.24		2.34	
(Hicks et al.	Soufriere	Hand	Andesite	2600	0.2	1.00E-10		
2009) (Hurwitz et	Hills	sample Conceptual	Conduit (numerical	2500	0.15		2	
(Hulwitz et al. 2003)	Cascades	Conceptual	model)	2500	0.15		2	
		Conceptual	model)	2500	0.01		2	
		Conceptual	Basal unit	2500	0.01		2.5	
(Ikeda et al. 2008)	Unzen	Drill core	Volcanic breccia	2400	0.15			
		Drill core	Lava dike	2500	0.15			
		Drill core	Lava dike	2600	0.15			
		Drill core	Volcanic breccia	2500	0.3			
(Ingebritsen	Cascades	Conceptual	Lava flows and domes		0.15	1.00E-14	1.55	
et al. 1994)		Conceptual	younger than 2.3 Ma Lava flows and minor pyroclastic rocks from		0.1	5.00E-16	1.55	
		Conceptual	Lava flows from 8 to		0.05	1.00E-16	1.65	
		Conceptual	Chiefly volcanic and volcaniclastic strata		0.05	1.00E-17	2	
		Conceptual	Quartz-bearing ash		0.02	2.50E-14	2	
(Keating	Mount St.	Hand	Dome rock	2200	0.6	2.10E-14	0.9	1557
2003)	Ticlens	Hand	Pyroclastic flow	2200	0.25	1.00E-15	0.35	2036
		sample Hand	Pyroclastic flow	2200	0.62	4.66E-14	0.35	2036
(Mueller et	Unzen	Hand	Dacite, breadcrust		0.367	8.90E-14		
al. 2008)		sample Hand	bomb Dacite, breadcrust		0.475	1.47E-13		
		sample Hand	bomb Dacite, dome rock		0.343	9.99E-13		
		sample Hand	Dacite, dome rock		0.349	3.41E-12		
		Hand	Dacite, dome rock		0.412	4.50E-12		
(Platz et al.	Mt. Taranaki	Cores from	andesite dome	2555		6.80E-13	14.47	918
(Reid 2004)	Cascades volcanoes	"typical of volcanic	Unknown	2650			2	1000
(Sammel et	Newberry	Conceptual	Fill			1E-14 - 5E-12		
al. 1988)		Conceptual	Dikes and pipe			5.00E-15		
		Conceptual	Flow 1			1E-15 - 1E-13		
		Conceptual	Flow 2			1E-16 - 1E-14		
		Conceptual	Flow 3			1E-17 - 1E-15		
		Conceptual	Flow 4			5E-16 - 1E-13		
		Conceptual	Flow 5			1E-18 - 1E-16		
		Conceptual	Magma			1.00E-18		

(Scheu et al.	Unzen	Hand	Dacite	2490	0.041		
2006)		sample Hand	Dacite	2420	0.073		
		sample					
		Hand	Dacite	2280	0.12		
		Hand	Dacite	2180	0.163		
		Hand	Dacite	2100	0.193		
		sample Hand sample	Dacite	1930	0.259		
(Sekioka 1988)	Japanese geothermal fields	Calculated	Unknown			1.10E-12	2.04
	Japanese geothermal	Calculated	Unknown			3.50E-13	2.33
	Japanese geothermal	Calculated	Unknown			1.90E-12	3.43
	Japanese geothermal	Calculated	Unknown			2.10E-12	2.87
	Japanese geothermal	Calculated	Unknown			2.00E-12	1.49
	Japanese geothermal	Calculated	Unknown			5.00E-14	0.93
	Japanese geothermal fields	Calculated	Unknown			1.70E-12	1.65
(Smith et al.	Mount St.	Cores from	dacite	2460	0.08		
2001)	Helens	spine					
		Cores from spine	dacite	2390	0.095		
		samples Cores from spine	dacite	2350	0.103		
		Cores from spine	dacite	2260	0.121		
		samples Cores from spine samples	dacite	2040	0.197		
(Watanabe et	Unzen	Drill core	C1-5-7, Brecciated	2410	0.08	1.00E-19	
al. 2008)		Dwill	dacite	2500	0.16	1.00E 10	
		Drill core		2500	0.16	1.00E-19	
		Drill core	C12, Volcanic breccia	2400	0.02	1.00E-17	
		Drill core	C14-2, Dacite dike	2560	0.04	1.00E-17	
		Drill core	C13, Dacite dike	2570	0.08	1.00E-17	

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