

# Thermal Radiation from Vapour Cloud Explosions

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## Abstract

The current study estimates the radiation flux emitted from hot extended gas clouds characteristic of vapour cloud explosions along with the corresponding level of irradiance posed on particles suspended in the unburnt part of the cloud ahead of an advancing flame front. The data presented permits an assessment of the plausibility of combustion initiation by such particles due to forward thermal radiation. The thermal radiation will depend on the emissivity of the burned volume, which relates to the concentration of gaseous and particulate combustion products. A sensitivity analysis has been carried out to account for variations in the equivalence ratio, mixture pressure and radiative heat losses. The spatial distribution of irradiance ahead of the flame front has been computed by introducing appropriate geometrical factors to explore the impact of cloud size. Using fuel rich ethylene-air mixtures it has been shown that high flame emissivities can be achieved at path lengths of order 1 m even in the presence of very low soot volume fractions. The emissivity of gas-soot mixtures will hence be mainly determined by the soot concentration and to a lesser extent by the mixture temperature. Our analysis suggests that the role of forward thermal radiation as a contributing factor to flame propagation in large scale vapour cloud explosions can not currently be ruled out.

*Keywords:* Radiation Induced Ignition, Vapour Cloud Explosions, Soot

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## 1. Introduction

Radiation often dominates heat transfer process at high temperatures (Hottel, 1958). Consequently, thermal radiation makes a decisive contribution to the overall energy transport in many combustion systems (Nathan et al., 2012). However, the influence of radiative heat transfer in unconfined vapour cloud explosions (UVCE) and on the corresponding rate of flame propagation is not yet fully understood. Particles heated by high levels of radiation can induce ignition of an adjacent explosive charge. Moore and Weinberg (1981, 1983, 1987) have shown that this may become important in vapour cloud explosions (VCE). The emission of strong radiative heat loads, emanating from the hot product cloud, on particles situated in the reactants can be sufficient to ignite the surrounding fuel-air mixture. In order to have a notable effect, ignition centres have to be formed well ahead of the advancing flame, thus relatively long length scales and short time scales are essential. Beyrau et al. (2013) explored the potential of fine particles acting as initiators of combustion in flammable mixtures upon irradiation using a near infrared (NIR) laser source. The experimental investigation featured powders with widely different characteristics (type, size, morphology, etc.) and times to ignition were established. In particular, ignition time scales  $\simeq 100$  ms were obtained in a stoichiometric butane-air mixture at an irradiance  $< 600$  kW/m<sup>2</sup> using substrates coated with a commercially available carbon black powder (acetylene black). In a recent study, Beyrau et al. (2014) quantified the heating process of such irradiated powders using time-resolved emission spectroscopy. The particle surface temperatures necessary to cause ignition of a surrounding charge were also obtained revealing two different ignition

26 regimes based on the reactivity of the powder.

27 Fine particles may be raised by an expanding gas cloud and become sus-  
28 pended in the unburnt gas mixture. The dispersion of dusts/particulates  
29 ahead of a propagating flame front is a well established phenomenon. Ac-  
30 cording to Klemens et al. (2006), fine dusts can be raised by expansion waves  
31 induced from a moderate local explosion. For example, in coal mines the pres-  
32 sure wave of a weak methane explosion can disperse dust deposits leading  
33 to the formation of an explosive dust-air cloud. The dust can be ignited by  
34 the hot methane-air products causing a (strong) secondary explosion. The  
35 phenomenon has been the subject of studies exploring the interaction of de-  
36 posited dust layers with shockwaves (e.g. Fedorov (2004); Gerrard (1963)).  
37 In addition, the dispersion of coal dust deposits by an advancing methane-air  
38 flame has been studied experimentally by Lu et al. (2002) in a laboratory  
39 scale flame tube. Hydrogen-air explosions can exhibit visible luminosity due  
40 to suspended inert particles while, in hydrogen jet flames, naturally occur-  
41 ring particulates present in the air entrained into the reaction region can also  
42 be a source of visible light emission (Shirvill et al., 2012). Finally, inert dust  
43 can suppress dust explosions and hence can be employed for the prevention  
44 and mitigation of dust explosions in coal mines (Amyotte, 2006).

45 The levels of flame surface flux reported in literature from various com-  
46 bustion systems can be seen in Table 1. There is a notable absence of data  
47 on the premixed systems considered in the current study. However, Holbrow  
48 et al. (2000) examined the radiative power densities from fireballs produced  
49 from vented dust explosions. Average surface emissive power (ASEP) of up  
50 to 275 kW/m<sup>2</sup> have been measured with coal dust and up to 2900 kW/m<sup>2</sup>

51 with aluminium. In heterogeneous combustion systems, reaction takes place  
52 at the surface of the condensed fuel, hence, dust explosions emit continuous  
53 Planck's radiation which is a function of the particle temperature. This can  
54 explain the discrepancy between results obtained with aluminium and coal  
55 dust. Thermal radiation from fireballs produced in Boiling Liquid Expand-  
56 ing Vapour Explosion (BLEVE) have also been examined. These turbulent  
57 flames emit non-luminous infrared radiation emanating from the emission  
58 bands of gaseous combustion products and luminous continuous radiation  
59 by soot particles in the visible and infrared (Tien and Lee, 1982; Viskanta  
60 and Mengüç, 1987). High emissivities can be achieved due to the high soot  
61 concentration and large burnt gas volume. Measurements by Roberts et al.  
62 (2000) indicate  $SEP_{max}$  up to  $550 \text{ kW/m}^2$  while extrapolated results from  
63 Roberts (1981) suggest that  $SEP_{max}$  up to  $450 \text{ kW/m}^2$  can be achieved. Aver-  
64 age SEPs from optically thick diffusion flames can typically be expected to be  
65 of the order  $200\text{-}300 \text{ kW/m}^2$  with maximum spot values of  $350\text{-}450 \text{ kW/m}^2$  as  
66 shown in Table 1. Similarly, radiation emanating from gaseous products and  
67 soot is a well known design consideration in gas turbine burners. Theoretical  
68 results, obtained from spray-stabilised flames in pressurised enclosures, sug-  
69 gest that flame surface flux around  $1500 \text{ kW/m}^2$  can be achieved (Lefebvre,  
70 1984; Mengüç et al., 1986; Najjar, 1985).

71 Experiments by Hardee et al. (1978) involving fireballs, produced by non-  
72 premixed as well as premixed stoichiometric methane-air mixtures (1.5 and 10  
73 kg of  $\text{CH}_4$ ), showed that premixed clouds, although appearing less luminous  
74 and relatively more transparent than the corresponding non-premixed case,  
75 emit higher flame surface fluxes due to the increased temperature of the gas.

Table 1: Reported radiation heat flux from selected combustion systems.

System	Regime	Remarks	Flux (kW/m <sup>2</sup> )	Reference
CH <sub>4</sub> -air cloud	Premixed	Scaled	450	Hardee et al. (1978)
C <sub>3</sub> H <sub>8</sub> -air cloud	Premixed		690	The Steel Construction Institute (2014)
LPG fireball	Non-Premixed	Extrapolated	450	Roberts (1981)
LPG fireball	Non-Premixed	ASEP	350	Roberts et al. (2000)
LPG fireball	Non-Premixed	SEP <sub>max</sub>	550	Roberts et al. (2000)
GT				Lefebvre (1984); Mengüç et al. (1986); Najjar (1985)
Combustor	Non-Premixed		1500	
Dust explosion	Heterogeneous	ASEP	275	Holbrow et al. (2000)

76 Scaled results suggest that maximum flame surface fluxes up to  $450 \text{ kW/m}^2$   
77 could be expected from a premixed cloud. The argument is corroborated  
78 by considering results from Dorofeev et al. (1996) who collected light from  
79 stoichiometric and fuel rich propane-air detonations. Measurements showed  
80 that significantly more light is emitted during the premixed burning phase  
81 than at any subsequent excess fuel burnout. In a detonation wave, both the  
82 temperature and pressure are much higher than in conventional deflagration,  
83 which will ultimately induce increased gas emissivities. Radiation measure-  
84 ment obtained from premixed propane-air clouds suggest spot values of ap-  
85 proximately  $700 \text{ kW/m}^2$  (The Steel Construction Institute, 2014). In these  
86 particular tests, carbon based dusts were laid on the floor of the explosion  
87 chamber to examine if they would cause secondary ignition due to forward  
88 thermal radiation. While no acceleration that could be attributed to radia-  
89 tive heating was observed, previous work has shown (Beyrau et al., 2013)  
90 that ignition timescales can vary by orders of magnitude for different carbon  
91 black powders. Explosions often deviate from stable deflagrations or detona-  
92 tions occurring under ideal conditions and in unimpeded geometries (Oran  
93 and Williams, 2012). Accordingly, in a real incident local gas pockets may  
94 achieve high pressure and temperature without subsequently developing into  
95 a detonation. Hence, radiative properties obtained from local events may  
96 still be important for an unsuccessful deflagration-to-detonation transition  
97 (DDT).

98 There is an obvious lack of radiation measurements in large scale pre-  
99 mixed systems related to explosions and the actual level of thermal radiation  
100 emitted from a VCE remains conjectural. Radiation emanates from both

101 gaseous and particulate combustion products, which are at higher temper-  
102 atures for premixed flames and hence higher radiation levels are expected.  
103 Moore and Weinberg (1981) reported theoretical values of blackbody radia-  
104 tion up to  $1 \text{ MW/m}^2$  assuming a burnt gas temperature ( $T_b$ ) at 2050 K as  
105 representative of a stoichiometric mixture and unit emissivity. Although lab-  
106 oratory premixed flames vary from the blackbody condition, in vapour cloud  
107 explosions, the shear size of combustion product cloud is believed to yield  
108 higher emissivities (Finkelburg, 1949). Additionally, soot can be generated  
109 as a result of local inhomogeneities in the equivalence ratio or in fuel rich  
110 regions. The presence of a large number of very small unburned carbon par-  
111 ticles, initially expected to be in thermal equilibrium with the surrounding  
112 combustion products, will induce higher flame emissivities. The true level of  
113 radiant heat emitted will hence be affected by the local mixture stoichiometry  
114 and vapour clouds resulting from accidental leaks are likely to be stratified.

115 Atkinson and Cusco (2011) have further proposed that the theory of ra-  
116 diatively ignited particulates may explain the unusual flame propagation rate  
117 observed in the 2005 Buncefield explosion. The objective of the current study  
118 is, hence, to estimate the flame surface flux expected from large premixed  
119 systems and examine the corresponding level of irradiance posed on particles  
120 suspended in the unburned gas mixture. A comparison of such theoretical es-  
121 timates with the experimentally measured ignition time data (Beyrau et al.,  
122 2013) is vital for evaluation purposes.

123 The current study extends previous efforts by consideration of parameters  
124 relevant to the mechanism proposed by Atkinson and Cusco (2011). Flame  
125 radiation emitted from the principal gaseous products  $\text{H}_2\text{O}$  and  $\text{CO}_2$  at large

126 path lengths is obtained along with the corresponding emissivity. More-  
127 over, a sensitivity analysis is carried out based on laminar flame calculations  
128 for fuel-air mixtures using detailed chemistry to account for variations in the  
129 equivalence ratio, pressure and heat losses. The resulting spatial distribution  
130 of irradiance on particles present in the unburned gas mixture is calculated  
131 using appropriate view factors. Finally, the influence of the flame tempera-  
132 ture, size and location relative to the irradiated particle is considered.

## 133 **2. Material and Methods**

### 134 *2.1. Estimation of Flame Radiation*

135 Flame radiation originates from gaseous combustion products like water  
136 vapour ( $\text{H}_2\text{O}$ ), carbon dioxide ( $\text{CO}_2$ ), carbon monoxide ( $\text{CO}$ ) and particles  
137 such as soot. Emissions from carbon monoxide and pollutants such as sul-  
138 phur dioxide and nitrous oxide, are minimal compared to the water vapour  
139 and carbon dioxide and can therefore be neglected. The product gas quan-  
140 tities and temperatures are intrinsic flame properties which depend on the  
141 type of fuel, initial temperature and pressure, and equivalence ratio. The  
142 absorption/emission spectrum of each species is banded even at the high  
143 temperatures encountered in flames. Consequently, spectral considerations  
144 have to be taken into account without, however, the need for detailed line-by-  
145 line calculations. Typically, computations can be performed by dividing the  
146 spectrum of interest in smaller (narrow or wide) bands and assume that the  
147 discrete absorption lines of each gas can be represented by a smooth profile.  
148 Theoretical narrow band models provide the mean spectral emissivity over  
149 these smaller spectral ranges by utilising statistical methods to characterise



150 the exact emission lines. Wide-band models provide the total absorption  
 151 over individual bands for each radiating gas using empirical relations fitted  
 152 to data obtained from experimental measurements. Detailed discussion on  
 153 narrow and wide band models can be found in Tien and Lee (1982); Viskanta  
 154 and Mengüç (1987). In this study, the emissivity of combustion products has  
 155 been computed using the spectrally resolved absorption coefficient data at  
 156 flame temperatures obtained by Ludwig et al. (1973).

157 For luminous flames, radiation originates from soot particles and gaseous  
 158 combustion products. Soot particles emit continuous radiation over the vis-  
 159 ible and infrared spectrum. The structure of soot consists of fused carbon  
 160 particles ranging from a few nanometers to a few hundred nanometers in  
 161 diameter. The interaction of soot with incident radiation follows the Mie  
 162 theory (Van de Hulst, 1957) and scattering is negligible compared to ab-  
 163 sorption since the radiation wavelengths are larger than the soot particle  
 164 diameter ( $\pi D/\lambda < 1$ ). It has been shown by Yuen and Tien (1977) that in  
 165 luminous flames the exact closed-form expressions for soot emissivity ( $\varepsilon_s$ ),  
 166 can be approximated based by,

$$\varepsilon_s = 1 - \exp(-k_s L) \quad (1)$$

167 where  $L$  is the path length of the flame and  $k_s$  is a soot-emission parameter  
 168 given by,

$$k_s = 3.6 \frac{cT_b}{c_2} \quad (2)$$

169 where  $c_2 = 1.44 \times 10^{-2}$  mK is the Planck's second constant and  $T_b$  is the

170 flame temperature. The constant  $c$  is given by,

$$c = 36\pi f_v \frac{n^2 k}{[n^2 - (nk)^2 + 2]^2 + 4n^2 k^2} \quad (3)$$

171 where  $n$  and  $k$  are the infrared-average optical constants of soot and  $f_v$  is  
172 the soot volume fraction.

173 The above non-grey analysis has been adopted in a number of studies  
174 (Mason et al., 2009; Wiedenhoefer and Reitz, 2003; Yoshikawa and Reitz,  
175 2009). Howell et al. (2011) suggested that  $c/(c_2 f_v) = 350 \text{ m}^{-1} \text{K}^{-1}$  irrespective  
176 of the type of soot. However, current evidence suggests that the constant  
177  $c$ , should be computed on the basis of refractive index  $\tilde{m} = n - ik$ . If,  
178 for example,  $\tilde{m} = 1.8 - 1.0i$  (Shaddix and Williams, 2007) is used higher  
179 absorptivities are obtained. Dalzell and Sarofim (1969) proposed a set of  
180 experimental values for the refractive index of soot ( $\tilde{m} = n - ik$ ) determined  
181 from reflectance measurements and a dispersion model for fitting to the data.  
182 Lee and Tien (1981) used a revised model in conjunction with transmission  
183 measurements and reported different values for the optical constants of soot.  
184 Furthermore, Habib and Vervisch (1988) suggested that the variation of the  
185 refractive index of soot with respect to the hydrogen content of the fuel can be  
186 calculated via two bound-one free dispersion equations. In the current study,  
187 the refractive index  $\tilde{m} = 1.56 - 0.56i$  is used which has been frequently cited  
188 by the combustion community (Smyth and Shaddix, 1996).

189 Assuming that soot behaves like a grey body, the emissivity of a luminous  
190 flame ( $\varepsilon_f$ ) emanating from soot and combustion products can be expressed

191 by Eq. (4).

$$\varepsilon_f = \varepsilon_g + \varepsilon_s - \varepsilon_g \varepsilon_s \quad (4)$$

192 The emissivity of the gaseous combustion products ( $\varepsilon_g$ ) can be obtained  
193 using non-luminous flame analysis alone. Therefore, the problem of calculat-  
194 ing the emissivity from luminous flames can be greatly simplified by using  
195 Eq. (4). Moreover, it can be deduced that radiation from a luminous flame  
196 is equal to the emissivity of gas and soot alone minus a correction factor.  
197 Mixtures of gas and soot have been considered in this article to quantitatively  
198 highlight the importance of the presence of soot particles in vapour cloud  
199 explosions.

200 Ideally, a hybrid model should be used to account for potential irregular-  
201 ities in luminosity expected in the event of cloud stratification. Similarly, in  
202 pool fires the flame is split into to a lower clear luminous burning zone and an  
203 upper sooty black smoke zone (Hailwood et al., 2009; Rew et al., 1997). The  
204 inhomogeneities in the concentration of combustion products (gas and soot)  
205 and fluctuations in temperature caused by turbulence will have an effect on  
206 the resulting thermal radiation. However, detailed accounting for variations  
207 induced by turbulence, fuel stratification and spatial luminosity variations  
208 are likely to be strongly scenario dependent and correspondingly complex.  
209 The difficulties associated with considering such effects outweigh the cur-  
210 rent objective of providing estimated radiation fluxes. Hence, a homogenous  
211 temperature and concentration model was implemented.

212 *2.2. Geometric View Factor*

213 The level of irradiance received by particles located ahead of an advanc-  
214 ing flame will depend on (i) the flame surface flux, (ii) the geometric view  
215 factor between the flame and particles and (iii) the absorption of radiation by  
216 the unburned fuel-air mixture. To estimate the irradiance received by such  
217 a particle, the flame front is represented by an appropriate physical model  
218 which entails knowledge of the flame heat release rate and shape. Moore and  
219 Weinberg (1983) represented the flame with a planar circular shape propa-  
220 gating along its central axis. This is considered a solid flame model since  
221 the flame is approximated as a solid body of equivalent shape (Davis and  
222 Bagster, 1989).

223 The total heat transfer by radiation ( $Q_{1-2}$ ) from the flame (Body 1) with  
224 emissive power  $E_f$  to the target (Body 2) is given by Eq. (5)

$$Q_{1-2} = E_f F_{1-2} A_1 \quad (5)$$

225 where,  $F_{1-2}$  is the geometric view factor between the flame and the target  
226 and  $A_1$  is emitting area.

227 The irradiance received by a target ( $q_2$ ) of area  $A_2$  is calculated via Eq. (6).

$$q_2 = Q_{1-2}/A_2 \quad (6)$$

228 The flame is assumed to emit radiation like a solid body thus the corre-  
229 sponding emissive power ( $E_f$ ) can be expressed by,

$$E_f = \varepsilon_f \sigma T_b^4 \quad (7)$$

230 where  $\varepsilon_f$  is the flame emissivity and  $\sigma = 5.6704 \times 10^{-8} \text{ W/m}^2\text{K}^4$  is the  
231 Stefan-Boltzmann constant.

232 Equations (5 - 7) allow the calculation of irradiance received by a particle  
233 located in the unburned region assuming the flame emits radiation like a  
234 solid body and there is no attenuation from the interleaving unburned fuel-  
235 air mixture. In fact, the level of absorption is essentially determined by  
236 the spectral overlap of the emitted radiation and absorption bands of the  
237 unburned gas mixture. In order to examine to what extent the assumption  
238 of no attenuation is correct, the case of collimated blackbody radiation into  
239 an ethylene-air mixture of unit stoichiometry at 1 atm, 298 K and 100%  
240 humidity was considered. In this case, there are no geometrical effects and,  
241 hence, the attenuation of radiation with distance ( $x$ ) can be described by  
242 Beer's law (Eq. 8), assuming a mean absorption coefficient ( $k$ ). Ethylene  
243 was specifically chosen due to its strong near and mid infrared absorption  
244 bands (Moore and Weinberg, 1983).

$$I = I_o \exp^{-kx} \quad (8)$$

245 Lastly, an appropriate geometrical factor between the flame and parti-  
246 cles  $F_{1-2}$  is required. This depends on the geometrical characteristics of the  
247 emitter and target. For the purpose of this study, three well known view fac-  
248 tors, that of straight cylinder (Mudan, 1987), and two parallel coaxial disks  
249 and coaxial squares have been selected (Howell et al., 2011). In summary,  
250 in the calculation of the irradiance received by particles present in the reac-  
251 tant gas, the flame is (i) modelled as a circular or rectangular planar shape  
252 propagating along its centre axis or as a squat cylinder propagating radially,

253 (ii) flame surface fluxes are obtained from Planck’s radiation law and (iii)  
254 attenuation from the fuel-air mixture is neglected unless stated. Hence, a  
255 symmetric cloud shape with a vertical axis of symmetry is assumed and up-  
256 wards radiative fluxes were not considered due to the reduced likelihood of  
257 particulate material being suspended above the cloud.

### 258 *2.3. Estimation of Flame Properties*

259 Flame properties required for this study have been computed using an in-  
260 house code developed by Jones and Lindstedt (1988). The chemistry is based  
261 on the work of Lindstedt and coworkers (Lindstedt and Meyer, 2002; Lindst-  
262 edt et al., 2011), the mechanism consists of 168 reactions and 33 species. A  
263 laminar flame, propagating freely through a premixed mixture was consid-  
264 ered based on the constant pressure assumption. The boundary conditions  
265 of pressure ( $P_0$ ) and temperature ( $T_0$ ) were set to 101325 Pa and 298 K  
266 respectively. Adiabatic combustion was assumed in all cases except when  
267 the effect of radiative heat losses was examined. The burnt gas tempera-  
268 ture and species concentration, necessary for the calculation of the flame  
269 emissivity, were extracted from the simulations. The computational domain  
270 for methane-air cases was resolved using 318 nodes featuring a mesh size of  
271  $\sim 3 \mu\text{m}$  in the reaction zone, while for ethylene and ethane-air 214 nodes  
272 were used corresponding to a mesh size of  $\sim 4 \mu\text{m}$  in the reaction zone.

273 Illustrative species profiles are shown in Fig. 1, these include  $\text{CH}_4$ ,  $\text{O}_2$  and  
274  $\text{CO}$  and final product species  $\text{H}_2\text{O}$  and  $\text{CO}_2$ . Carbon monoxide is formed dur-  
275 ing the combustion process followed by further oxidation to  $\text{CO}_2$ . Therefore,  
276 the  $\text{CO}$  profile features a maximum within the reaction zone. Principal com-  
277 bustion products, water and carbon monoxide, increase steadily throughout

278 the domain. The middle plot shown in Fig. 1 shows the rate of production  
 279 (+ve) and consumption (-ve) of fuel, oxygen and carbon monoxide. The tem-  
 280 perature profile rises steadily during the carbon monoxide to carbon dioxide  
 281 oxidisation phase as seen in the bottom plot.

282 Heat losses have been included in the computation by correcting the flame  
 283 temperature ( $T_b$ ) via Eq. (9), where  $T_{ad}$  is the adiabatic flame temperature  
 284 and  $\beta$  is the heat loss factor. The heat loss factor ( $\beta$ ) approach has been  
 285 used in a number of studies over a considerable period of time, e.g. Jones and  
 286 Lindstedt (1988), to include the  $T^4$  law impact on laminar flame structures  
 287 as part of radiation calculations. For example, the approach was used by  
 288 Fairweather et al. (1992) as part of a calculation procedure for the estimation  
 289 of radiative transfer from turbulent reacting jets.

$$T_b = T[1 - \beta(\frac{T}{T_{ad}})^4] \quad (9)$$

290 The objective of the laminar flame calculation is to determine the tem-  
 291 perature and concentrations of principal combustion products. The com-  
 292 putationally determined major species concentrations ( $x_i$ ) and temperatures  
 293 can seen in Table 2 along with the corresponding boundary conditions used  
 294 for each case. The thermal expansion ratio ( $\tau$ ) calculated via Eq. (10),  
 295 where  $\rho$  is the density and the subscripts '0' and 'b' indicate the values in  
 296 the reactants and burnt products, is also listed for each case.

$$\tau = \rho_0/\rho_b - 1 \quad (10)$$

297 The calculated flame temperatures obtained for all methane-air cases are

298 shown in Fig. 2. The top and middle rows illustrate the influence of the  
299 mixture stoichiometry and heat losses via radiation at an initial pressure  
300 of 1 atm. The flame temperature shows a peak at close to stoichiometric  
301 concentration. Furthermore, it is evident that a pressure rise will increase the  
302 flame temperature as shown in the bottom row. The corresponding influence  
303 of the flame temperature on the flame surface emissive flux is discussed in  
304 Section 3.1.

### 305 **3. Results and Discussion**

#### 306 *3.1. Flame Radiation*

307 The spectral radiance from gaseous combustion products from a stoichio-  
308 metric methane-air flame at  $T_b = 2212$  K and ambient pressure for different  
309 path lengths can be seen in Fig. 3 along with the blackbody distribution.  
310 Flame radiation will approximate that of a blackbody at large path lengths,  
311 Moore and Weinberg (1987) have also reported high emissivities of a stoi-  
312 chiometric propane-air flame at 2000 K using the same method at 25 m path  
313 length. The absorption bands of carbon dioxide and water vapour overlap at  
314 2.7 and 4.3  $\mu\text{m}$ , which explains the high emissivity reached in these spectral  
315 regions even at short path lengths. Furthermore, it can be seen why labora-  
316 tory scale flames (i.e.  $L = 1$  m) exhibit low emissivity,  $\varepsilon_g = 0.16$  and thus  
317 are optically thin. The flame emissivity reaches  $\varepsilon_g = 0.68$  at a path length  
318 of 50 m. The path length corresponds to the physical path through the hot  
319 combustion products. Due to the thermal expansion, a 50 m burnt cloud typ-  
320 ically corresponds to an unburnt cloud size of less than 10 m. Finkelburg  
321 (1949) has discussed the conditions for blackbody radiation from extensive



Table 2: Flame properties and boundary conditions for cases studied. Due to thermal expansion, a 10 m path length in combustion products corresponds to an initial could size of around 2 m.

Case	Fuel	$\phi$ (-)	$p_T$ (atm)	Mole Fraction of Major Radiating Species H <sub>2</sub> O	CO <sub>2</sub>	$\beta$ (-)	$T_{ad}$ (K)	$T_b$ (K)	$\tau$ (-)	$\varepsilon^*$ (-)	Flux* (kW/m <sup>2</sup> )
1	CH <sub>4</sub>	0.8	1	0.1524	0.0757	0	1984	1984	5.67	0.458	402
2	CH <sub>4</sub>	0.9	1	0.1685	0.0824	0	2121	2121	6.15	0.444	510
3	CH <sub>4</sub>	1.0	1	0.1821	0.0840	0	2212	2212	6.49	0.436	592
4	CH <sub>4</sub>	1.1	1	0.1865	0.0738	0	2178	2178	6.48	0.447	571
5	CH <sub>4</sub>	1.2	1	0.1834	0.0619	0	2086	2086	6.47	0.462	496
6	CH <sub>4</sub>	1.0	1	0.1841	0.0869	0.05	2212	2126	6.18	0.456	528
7	CH <sub>4</sub>	1.0	1	0.1854	0.0888	0.10	2212	2018	5.81	0.477	449
8	CH <sub>4</sub>	1.0	1	0.1865	0.0907	0.20	2212	1773	4.97	0.534	299
9	CH <sub>4</sub>	1.0	2	0.1835	0.0858	0	2232	2232	6.54	0.541	762
10	CH <sub>4</sub>	1.0	4	0.1846	0.0872	0	2248	2248	6.59	0.644	932
11	CH <sub>4</sub>	1.0	6	0.1852	0.0881	0	2257	2257	6.61	0.694	1021
12	CH <sub>4</sub>	1.0	8	0.1857	0.0888	0	2264	2264	6.63	0.724	1078
13	C <sub>2</sub> H <sub>4</sub>	1.0	1	0.1213	0.1076	0	2362	2362	7.05	0.348	615
14	C <sub>2</sub> H <sub>6</sub>	1.0	1	0.1574	0.0965	0	2251	2251	6.84	0.407	593
15	C <sub>2</sub> H <sub>4</sub>	2.0	1	0.0776	0.0259	0	1903	1903	6.56	0.398	286
16	C <sub>2</sub> H <sub>4</sub>	2.0	1	0.0953	0.0427	0.20	1903	1515	4.78	0.502	150

\*Based on 10 m path length.

322 gas masses of equal temperature, one example typically used to corroborate  
323 this argument is the continuous spectrum of the sun. In reality, any isother-  
324 mal gas can potentially emit radiation according to Planck’s law which is  
325 solely controlled by the absolute temperature given sufficient layer thickness  
326 and temperature equilibrium. Reasonable temperature homogeneity is in-  
327 deed expected in the burnt gas volume which is subject mainly to radiative  
328 cooling. This simple analysis shows that a vapour cloud explosion can theo-  
329 retically approximate a blackbody radiator due to the large gas volumes and  
330 elevated temperatures involved.

331 The effect of mixture stoichiometry, heat loss and pressure were consid-  
332 ered for a methane-air mixture from laminar flame calculations. The total  
333 mixture emissivity can be seen on the left hand side of Fig. 4 while the cor-  
334 responding flame surface flux, calculated via Eq. (7), is shown on the right  
335 hand side. The flame emissivity remains almost constant with changes in the  
336 fuel-air concentration. The resulting flame temperature is a function of the  
337 equivalence ratio, hence, the latter has a clear impact on the emitted sur-  
338 face flux which is proportional to the forth power of the temperature. For a  
339 methane-air mixture the emissive flux is expected to peak near stoichiomet-  
340 ric concentration. Heat losses occurring during combustion influence both  
341 the resulting temperature and product concentration. Whilst lower flame  
342 temperatures led to higher emissivities, the surface flux is reduced due to the  
343 strong temperature dependence.

344 The spectral radiance of a blackbody will shift to the infrared with de-  
345 creasing temperature (i.e.  $\lambda_p = 1.64 \mu\text{m}$  at 1773 K to  $\lambda_p = 1.31 \mu\text{m}$  at  
346 2210 K, where  $\lambda_p$  is the wavelength peak of the blackbody the spectrum)

347 providing a better overlap with the infrared absorption bands of the princi-  
348 pal combustion products leading to higher emissivity. In addition, the effect  
349 of total mixture pressure is examined in order to account for over-pressures  
350 that may occur during the flame propagation. High pressure increases the  
351 partial pressure of gaseous combustion products and to a lesser extent the  
352 temperature which raises significantly both the emissivity and surface flux.  
353 Furthermore, the same mechanism implies that detonation products will be-  
354 come highly emissive. Based on data obtained using GASEQ (Morley, 2013)  
355 for a stoichiometric methane-air CJ detonation, assuming unit path length,  
356 combustion products can reach an emissivity of 0.40 compared to 0.16 for a  
357 deflagration. In addition, the increased temperature observed in a detonation  
358 wave will induce a significant increase of the flame surface flux.

359 The emissivity from other fuels was calculated also for stoichiometric con-  
360 centration and assuming adiabatic combustion. These, include ethylene and  
361 ethane as shown in Fig. 5. Again, the shift of the blackbody distribution to  
362 visible wavelengths with increasing temperature is responsible for the lower  
363 emissivities observed with more reactive fuels. The flame surface flux is simi-  
364 lar for all mixtures because temperature and emissivity counter interact each  
365 other. For the case of a path length of 50 m the corresponding difference in  
366 the magnitude of flame surface flux obtained with different fuels is less than  
367 10 %. Carbon dioxide absorption bands (i.e.  $2.7 \mu\text{m}$ ,  $4.3 \mu\text{m}$  and  $15 \mu\text{m}$ )  
368 saturate at much shorter path lengths than the corresponding water bands.  
369 Additionally, at wavelengths shorter than  $2.7 \mu\text{m}$  radiation is emitted only  
370 by water molecules, this coincides with the peak blackbody distribution at  
371 such flame temperatures ( $\lambda_p = 1.27 \mu\text{m}$ ). This observation leads to the con-

372 clusion that hydrocarbon compounds with more hydrogen atoms will cause  
373 saturation at shorter path lengths given similar flame temperatures.

374 The emissivity from a rich ethylene-air flame ( $\phi = 2.0$ ) is shown in Fig. 6.  
375 This equivalence ratio is above the sooting limit and a modest soot volume  
376 fraction of  $1 \times 10^{-7}$  is included in the analysis. The gaseous emissivity is  
377 computed using the method described above using data from laminar flame  
378 calculation at adiabatic conditions. The emissivity of soot has been obtained  
379 from Eqs. (1) and (4) using the soot-emission parameter ( $k_s$ ) from Yuen and  
380 Tien (1977). The emissivity of a gas-soot mixture is of the same order as that  
381 of the gas alone for path lengths below 0.1 m, as the path length increases,  
382 the combined emissivity rises steeply reaching almost unity at 10 m.

383 Furthermore, the influence of the temperature and soot concentration on  
384 the corresponding emissivity of a gas-soot mixture can be seen in Fig. 7. The  
385 temperature sensitivity was carried out by comparing the adiabatic ethylene-  
386 air case ( $\phi = 2.0$ ) with one that included 20 % heat loss. In addition, three  
387 soot volume fractions have been included in each mixture that relate to a low,  
388 moderate and high soot concentrations (Geitlinger et al., 1998; DeIuliis et al.,  
389 1998; McEnally et al., 1997; Nathan et al., 2012; Wal and Weiland, 1994).  
390 Unit emissivities are reached around 0.1 m path length with  $f_v = 1.0 \times 10^{-5}$   
391 compared to 10 m with  $f_v = 1.0 \times 10^{-7}$  at the same temperature. The flame  
392 temperature also affects the mixture emissivity but to a lesser extent. The  
393 combined emissivity at a fixed path length depends on the soot-emission  
394 parameter ( $k_s$ ) which is directly proportional to both the temperature and  
395 soot volume fraction. Flame temperatures vary by a few hundreds degrees,  
396 however, the soot volume fraction can vary by orders of magnitude, hence, it

397 can greatly impact the total emissivity. The extreme case of 20 % heat loss  
398 which led to a flame temperature of 1515 K has been specifically selected to  
399 highlight the comparatively weak dependence on temperature. Therefore, in  
400 a real incident, soot will ultimate lead to higher emissivity at much shorter  
401 path lengths, irrespective of the overall temperature. Of course, in case  
402 of a stratified cloud featuring local temperature inhomogeneities the overall  
403 radiative heat emitted will be a superposition over the total flame surface.

### 404 *3.2. Forward Thermal Radiation*

405 In order to calculate the radiation received by particles ahead of the flame  
406 front, the flame surface flux has to be estimated first. Assuming that the  
407 flame front radiates like a solid radiator the emitted heat flux will strongly  
408 depend on the temperature of the combustion products as outlined above.  
409 The maximum surface radiation flux expected for different flame tempera-  
410 tures and emissivities obtained from Eq. (7) is shown in Fig. 8. Note that the  
411 flame surface flux corresponds to the emitted power density rather than the  
412 heat transferred to a potential target and is therefore a function of the radi-  
413 ator temperature only. For a stoichiometric methane-air mixture at ambient  
414 pressure the burned products would be at  $T_b = 2212$  K. It can be deduced  
415 that surface fluxes ranging from  $1350 \text{ kW/m}^2$  for  $\varepsilon_f = 1.0$  to  $810 \text{ kW/m}^2$  for  
416  $\varepsilon_f = 0.6$  can be expected.

417 The spatial variation of the ratio of the incident to source flux ( $I/I_0$ )  
418 due to the absorption from unburned gas mixture (including the presence of  
419 water vapour and carbon dioxide), obtained from Eq. (8), for three product  
420 temperatures is shown in Fig. 9. As the blackbody temperature decreases,  
421 the spectral radiance shifts towards the infrared providing a better overlap

422 with the infrared absorption bands of the fuel-air mixture and, hence, higher  
423 attenuation is observed. Nevertheless, even at the limiting case of 2000 K  
424 the attenuation remains at 10 and 25 % at 10 and 25 m from the flame front,  
425 respectively. This suggests that attenuation from the unburned gas mixture  
426 is not appreciable up to 10 m from the main flame front which is the order  
427 of length scale of interest.

428 The radial irradiance field assuming an average unburned cloud depth of  
429 2 m and an isobaric expansion coefficient of 6 is shown in Fig. 10a. The  
430 products were modelled as a squat cylinder and the receptor was assumed  
431 to be on the ground. Due to thermal expansion, a burnt cloud height of 12  
432 m corresponds to an unburnt cloud depth of approximately 2 m. A flame  
433 temperature of 2212 K and unit emissivity was selected as a representative of  
434 a stoichiometric methane-air mixture. As the flame propagates radially into  
435 the unburned gas mixture, the products approximate larger characteristic  
436 ratios (i.e.  $R/H \geq 5$ ) increasing the level of irradiance received by parti-  
437 cles. Therefore, the magnitude of flux will depend on the surface area of the  
438 source (i.e. the radius of the products) highlighting the specific application to  
439 vapour cloud explosions. Also, off the ground particles may receive different  
440 irradiance.

441 The distribution of irradiance in the unburned region as a function of  
442 the distance between the particle and flame front can be seen in Fig. 10b.  
443 As shown in Fig. 8, the flame surface flux expected from a stoichiometric  
444 product cloud of unit emissivity is around  $1350 \text{ kW/m}^2$ . The squat cylinder  
445 model used (Mudan, 1987) predicts lower irradiance close to the burnt cloud  
446 edge. This is a geometric effect due to the curvature of the cylinder and the

447 fact that the receptor is assumed to be located on the ground, which causes  
448 attenuation in the near field. Assuming a characteristic ratio of  $R/H = 1$ ,  
449 irradiance greater than  $250 \text{ kW/m}^2$  can be expected up to a distance of 10 m  
450 ahead the flame front. As the burned products grow radially, they approx-  
451 imate higher  $R/H$  ratios and the same level of irradiance ( $250 \text{ kW/m}^2$ ) can  
452 be received up to 25 m ahead of the flame. The first 10 to 25 m into the  
453 unburned gas mixture is critical for the short length and time scales required  
454 for radiation to have an impact on the overall flame propagation. Beyond  
455 this region, lower irradiance is expected reducing the possibility of a kernel  
456 forming within a reasonable ignition time. Nevertheless, it has been shown  
457 that the radiation levels obtained from the above analysis are sufficient to  
458 ignite a fuel-air mixture (Beyrau et al., 2013). In particular, ignition time  
459 scales  $\simeq 100 \text{ ms}$  at an irradiance  $< 600 \text{ kW/m}^2$  were obtained using sub-  
460 strates coated with a commercially available carbon black powder (acetylene  
461 black) in a stoichiometric butane-air mixture. For length scales larger than  
462 10 m, absorption by the fuel-air mixture may become important. However,  
463 as shown earlier the effect is small compared to that of the view factor.

464 Whilst the types of particles examined by Beyrau et al. (2013) may be  
465 present in industrial facilities, the ignition times alone do not prove that  
466 radiative induced ignition can occur. As shown above, the level of irradiance  
467 posed on such particles cannot be spatially uniform across the reactants  
468 since, the irradiance will attenuate with distance from the main flame front  
469 due to geometric effects and possibly due to absorption from the intervening  
470 medium. In practical vapour cloud incidents, the plausibility of formation  
471 of local exothermic centres ahead of the flame will depend on the time to

472 ignition which is determined by the level of irradiance and, hence, on the  
473 distance from the flame front. In order to have a notable effect, particles  
474 have to be ignited before the advancing flame consumes them. Therefore  
475 there are likely to be influences from (i) the average rate of flame propagation,  
476 (ii) the location (relative the advancing flame front), (iii) the density of the  
477 suspended particles in the unburnt region and (iv) the lower flame surface flux  
478 due to convective and radiative losses. The flame speed is a key parameter as  
479 it will influence the thermal dose received a particle as well as the duration of  
480 the irradiation before it is consumed by the propagating flame. Nevertheless,  
481 the current study shows that the estimated and required irradiances are  
482 comparatively close and hence, the flame propagation mechanism proposed  
483 by Atkinson and Cusco (2011) cannot currently be ruled out.

484 Moreover, Fig. 11 shows a comparison with other radiation heat trans-  
485 fer models, that of coaxial disks and squares, which are more appropriate  
486 for confined geometries. In these coaxial models, the vertical coordinate of  
487 the target is linked with the source size (i.e. burned gas height) which is  
488 not a realistic assumption for heavy vapour clouds. A squat cylinder is a  
489 more suitable for representing unconfined systems as considered in the cur-  
490 rent study. Nevertheless, these models can be used to highlight further the  
491 importance of the source size. From a heat transfer point of view, the radi-  
492 ation emitted from a modest (e.g. laboratory) scale solid radiator with unit  
493 area (i.e.  $L_{char} = 1$  m) will, even with unit emissivity, attenuate simply due  
494 to geometric factors at a distance of 2 - 3 m from the source. Distributions  
495 predicted using the model of coaxial squares are slightly higher than those  
496 obtained using the coaxial disks model. This, of course, is expected since



497 a larger emitting area is taken into account when a square is used. Also,  
 498 the coaxial models predict higher irradiance near the flame front and lower  
 499 irradiance in the far field compared to the squat cylinder. Again, this is  
 500 expected since the curvature of the cylinder causes a geometric attenuation  
 501 in the near field while in the far field, the surface area of the squat shape is  
 502 overall larger than the symmetric ones.

503 Another point that can be raised is the effect of flame area enhancement  
 504 due to turbulence. The total forward thermal radiation is proportional to  
 505 the emitting area ( $A_1$ ) as shown in Eq. (5). In the above analysis, the flame  
 506 has been modelled as smooth surface which is arguably not accurate since a  
 507 turbulent flame front will be wrinkled. Gouldin (1987) and Gouldin et. al.  
 508 (1989) have proposed that the area of flamelet surfaces in turbulent flames  
 509 can be estimated using fractals. Fractal surfaces are characterised by self-  
 510 similarity over wide range of scales (Mandelbrot, 1982) and hence, allow the  
 511 explicit consideration of multi-scale wrinkling. Turbulence will increase the  
 512 ensemble-average flame area by an enhancement factor,

$$\frac{A_1}{A_o} = \left(\frac{L_I}{L_K}\right)^{D-2} \quad (11)$$

513 where  $A_1/A_o$  is the ratio of the areas at the inner and outer cutoffs,  $L_I$  is  
 514 the integral length scale,  $L_K$  the Kolmogorov length scale and  $D$  the fractal  
 515 dimension. The ratio of the integral and Kolmogorov length scales follows,

$$\frac{L_I}{L_K} = Re_T^{3/4} \quad (12)$$

516 where  $Re_T$  is the turbulence Reynolds number. Assuming a mean fractal

517 dimension of 7/3 (Gülder, 1990; Kerstein, 1988), the flame surface enhance-  
 518 ment factor can be expressed as a function of  $Re_T$ .

$$\frac{A_1}{A_o} = Re_T^{1/4} \quad (13)$$

519 The influence of the turbulence Reynolds number on the flame surface en-  
 520 hancement factor is shown in Table 3, The results suggest a significant in-  
 521 crease at high turbulence Reynolds numbers. Hence, turbulence can be ex-  
 522 pected to enhance the overall heat transfer, though the local flame surface  
 523 characteristics may affect the corresponding view factor. Moreover any spa-  
 524 tial inhomogeneities in the temperature or the concentration of the combus-  
 525 tion products will influence the corresponding flame surface flux.

Table 3: The influence of the turbulence Reynolds number on the flame surface enhancement factor.

$Re_T$	$A_1/A_o$
1	1
10	1.79
100	3.16
1000	5.62

526 The fuel dependency on the distribution of irradiance from a square source  
 527 ( $L_{char} = 10$  m) as function of the distance from the flame front is shown in  
 528 Fig. 12. Assuming unit emissivity and adiabatic flame temperatures obtained  
 529 earlier using laminar flame calculations, the forward thermal radiation from  
 530 methane, ethane and ethylene flames can be estimated. The flame tem-  
 531 perature will influence both the maximum intensity and the corresponding  
 532 distribution of irradiance over a given distance. Hence, in practical incidents  
 533 the level of irradiance expected on particles likely to be present ahead of the

534 propagating front will strongly depend on the fuel type.

535 Finally, while not of direct relevance to the current study, radiation from a  
536 detonation event may be estimated for comparison purposes. For example, a  
537 stoichiometric methane-air CJ detonation yields a temperature of 2767 K and  
538 pressure of 16.6 atm (Morley, 2013) and the peak irradiance is estimated  $\sim$   
539 3350 kW/m<sup>2</sup>. A detonation event is indeed expected to be highly emissive as  
540 discussed earlier. Although not presented here, it is estimated that radiation  
541 levels higher than 820 kW/m<sup>2</sup> can be readily obtained 10 m ahead of a local  
542 detonation event ( $R/H = 1$ ) for an unburnt cloud height of 2 m.

#### 543 4. Conclusions

544 A study has been performed in order to estimate the radiation levels ex-  
545 pected from flame fronts appearing as part of vapour cloud explosions and  
546 to examine the corresponding forward thermal radiation posed on particles  
547 in the unburnt part of the cloud. The radiation emitted from the principal  
548 gaseous products H<sub>2</sub>O and CO<sub>2</sub> at large optical path lengths has been esti-  
549 mated and a sensitivity analysis performed to assess the impact of variations  
550 in the equivalence ratio, mixture pressure and radiative heat losses. It has  
551 ben shown that the flux from the gas phase is expected to peak near a stoi-  
552 chiometric concentration due to the strong temperature dependence. It has  
553 also been shown that in the presence of soot, high flame emissivities can be  
554 achieved at path lengths of the order 1 m and that the emissivity of gas-soot  
555 mixture will be mainly determined by the soot concentration and to a lesser  
556 extent by the mixture temperature. Hence, in a real incident, fuel-air mix-  
557 tures prone to soot formation will ultimately yield higher thermal radiation at

558 short path lengths. In summary, the distribution of irradiance on particles  
559 suspended ahead of a flame front will strongly depend on the presence of any  
560 soot, the flame temperature and, to a lesser extent, the burnt gas composi-  
561 tion. Assuming a stoichiometric, squat methane-air cloud of unit emissivity  
562 and a characteristic ratio  $R/H = 10$ , irradiance greater than  $250 \text{ kW/m}^2$   
563 can be expected up to a distance of 25 m from the flame front without ac-  
564 counting for any area enhancement caused by turbulence. Radiation levels  
565  $\sim 600 \text{ kW/m}^2$ , necessary for sufficiently short ignition time scales (Beyrau  
566 et al., 2013), can be readily obtained for distances up to 8 m from the flame  
567 front. The estimated and required fluxes are comparatively close and, hence,  
568 the episodal flame propagation mechanism proposed by Atkinson and Cusco  
569 (2011) cannot currently be ruled out.

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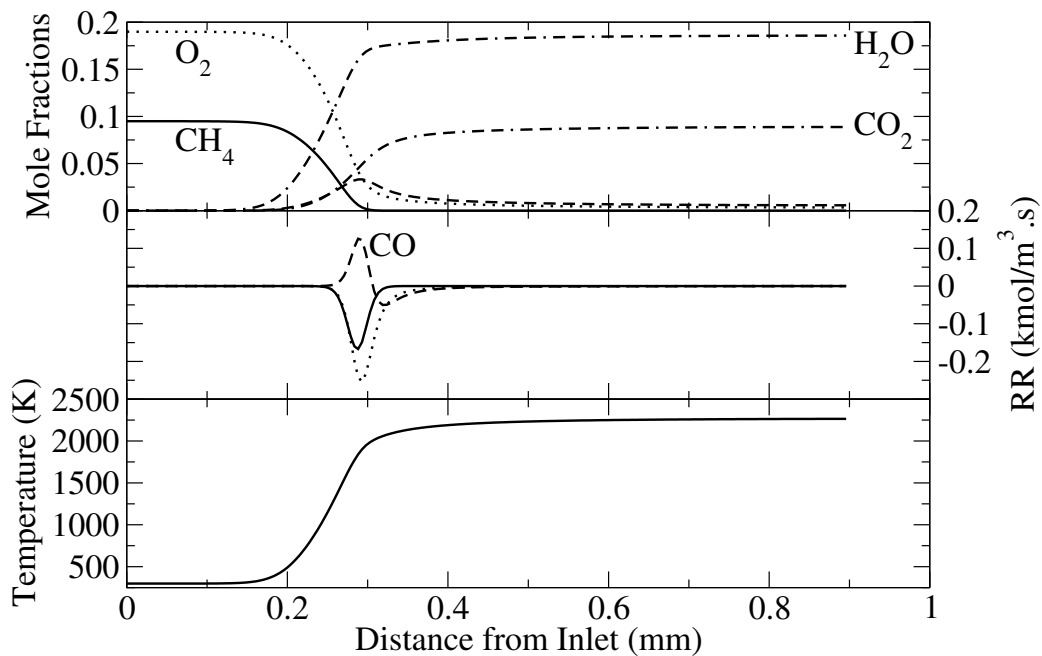


Fig. 1: Sample laminar flame calculation for a stoichiometric methane-air mixture,  $T_i = 298$  K,  $p_T = 8$  atm. Top: Species concentration ( $CH_4$ ,  $O_2$ ,  $H_2O$ ,  $CO_2$ ,  $CO$ ); Middle: The corresponding reaction rates for  $CH_4$ ,  $O_2$  and  $CO$ ; Bottom: Temperature profile.

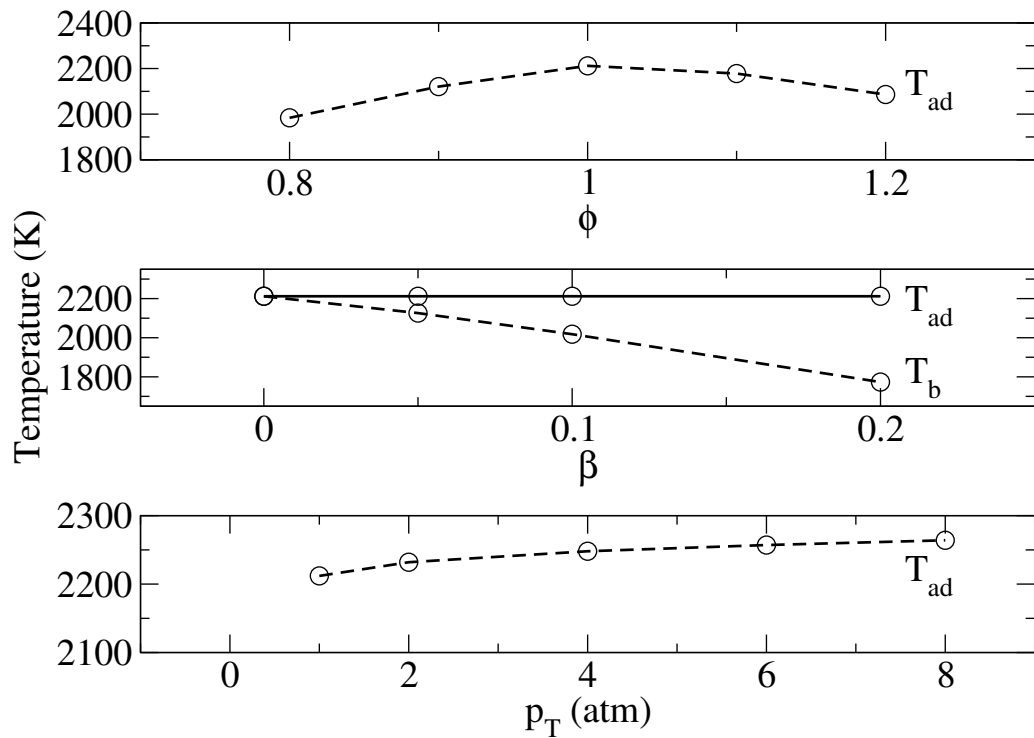


Fig. 2: Flame temperatures for all methane-air mixtures examined,  $T_i = 298$  K. The corresponding flame properties for selected cases can be found in Table 2. Top: Influence of mixture stoichiometry; Middle: Influence of radiation losses; Bottom: Influence of total pressure.

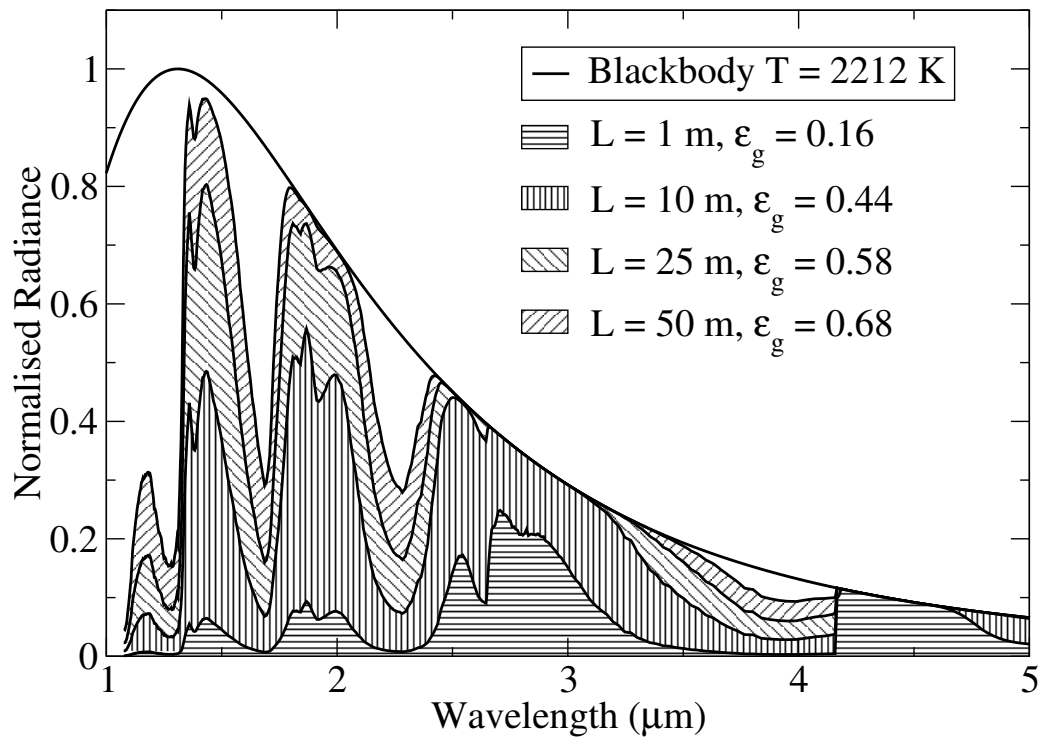


Fig. 3: Normalised spectral radiance of principal combustion products of a stoichiometric methane-air flame ( $T_b = 2212$  K) along with blackbody at 2212 K for comparison. Partial pressure of principal combustion products:  $p_{\text{H}_2\text{O}} = 0.182$  atm,  $p_{\text{CO}_2} = 0.084$  atm,  $f_v = 0$ .

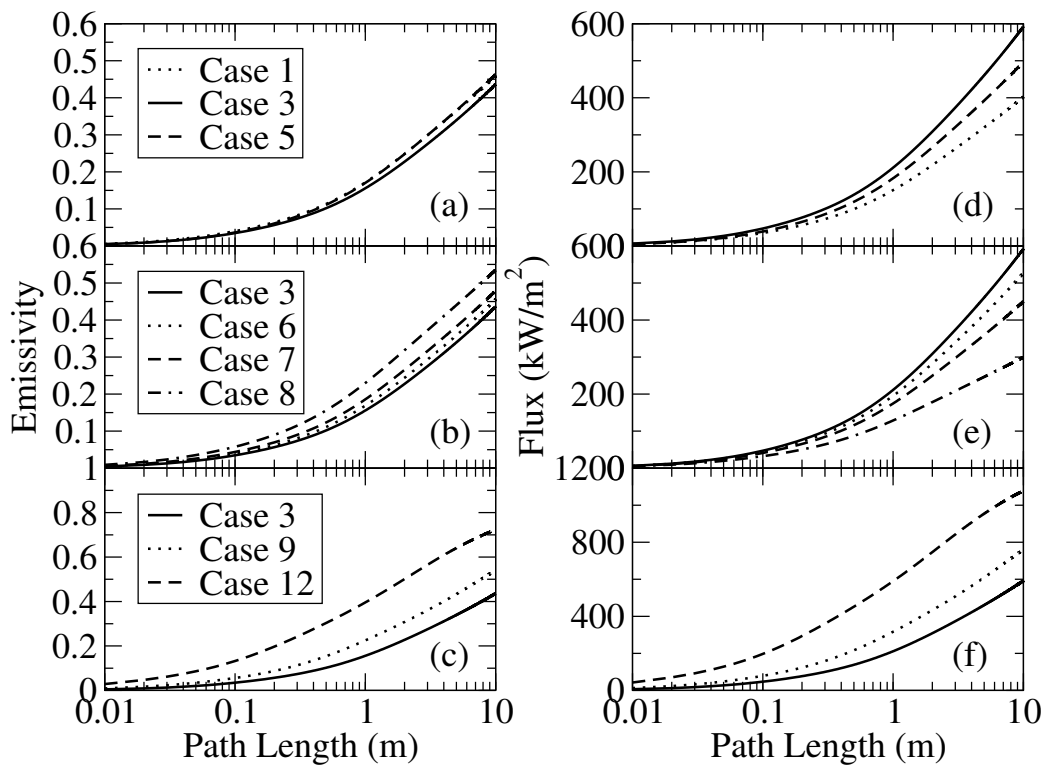


Fig. 4: Total emissivity and surface emissive flux from a methane-air flame as a function of the path length. Top: Influence of mixture stoichiometry; Middle: Influence of radiation losses; Bottom: Influence of total pressure. The corresponding flame properties for selected cases can be found in Table 2.



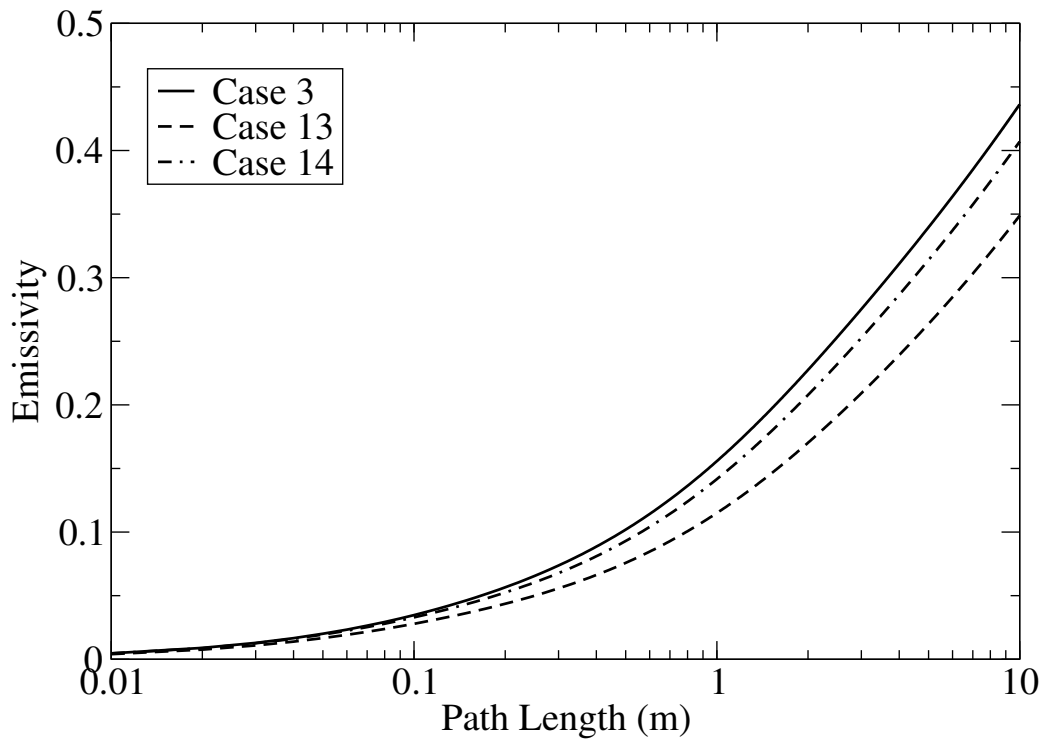


Fig. 5: The influence of the fuel type on the total emissivity for selected stoichiometric fuel-air mixtures as a function of the path length at total pressure,  $p_T = 1$  atm. The corresponding flame properties for selected cases can be found in Table 2.

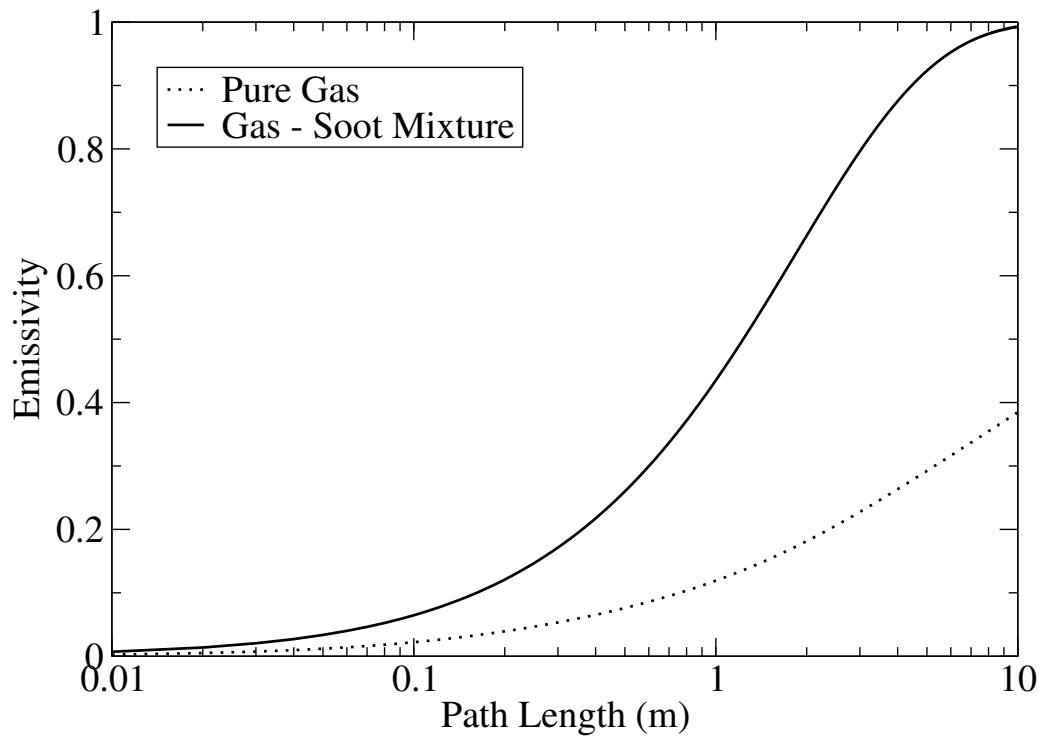


Fig. 6: Total emissivity of homogeneous H<sub>2</sub>O-CO<sub>2</sub>-soot mixture obtained from an ethylene-air flame ( $\phi = 2.0$ ,  $T_b = 1903$  K,  $f_v = 1.0 \times 10^{-7}$ ) as a function of the path length. The corresponding flame properties for Case 15 can be found in Table 2.

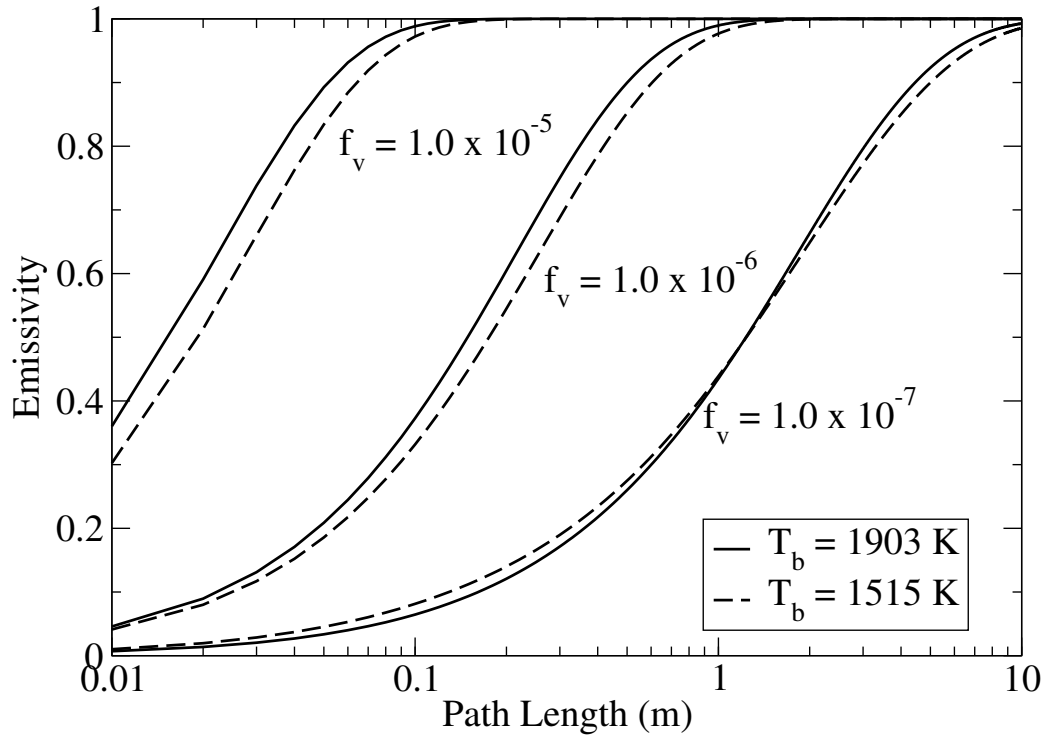


Fig. 7: The influence of combustion heat losses and soot volume fraction ( $f_v$ ) on the total emissivity of homogeneous  $\text{H}_2\text{O}$ - $\text{CO}_2$ -soot mixture as a function of the path length. Case 15 (—) and 16 (---) from Table 2 have been considered corresponding to a rich ethylene-air flame of 0 and 20 % heat loss respectively.

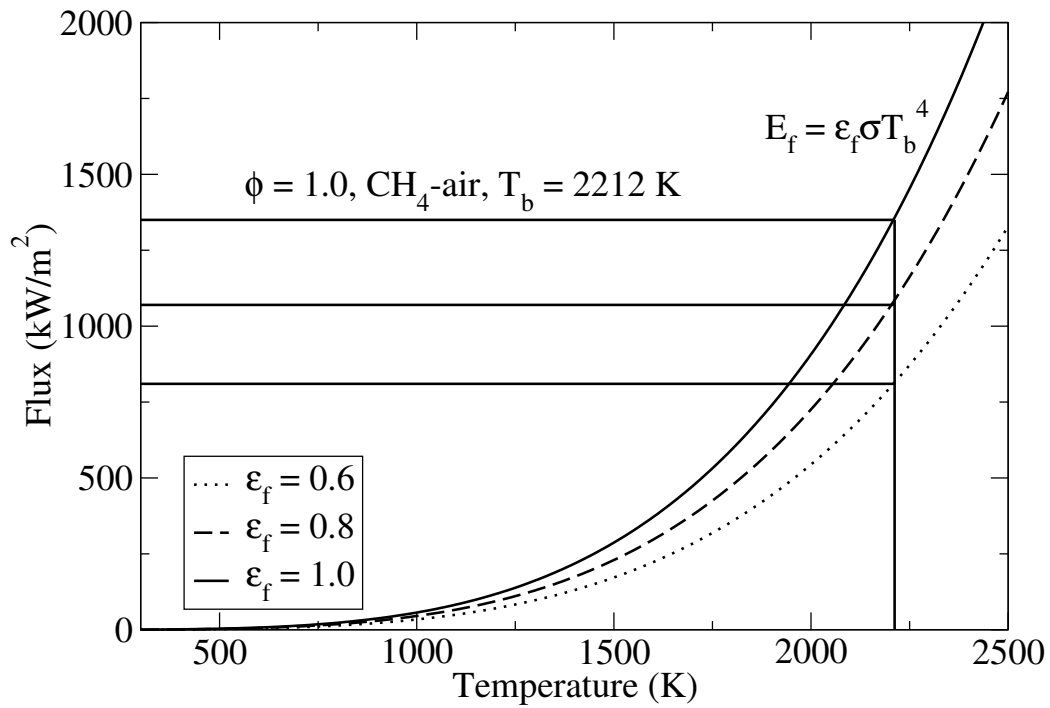


Fig. 8: Surface emissive flux as a function of the flame temperature for different emissivities. Note that the surface emissive flux corresponds to the emitted power density rather than the heat transferred to a potential target, thus is a function of the radiator temperature only.

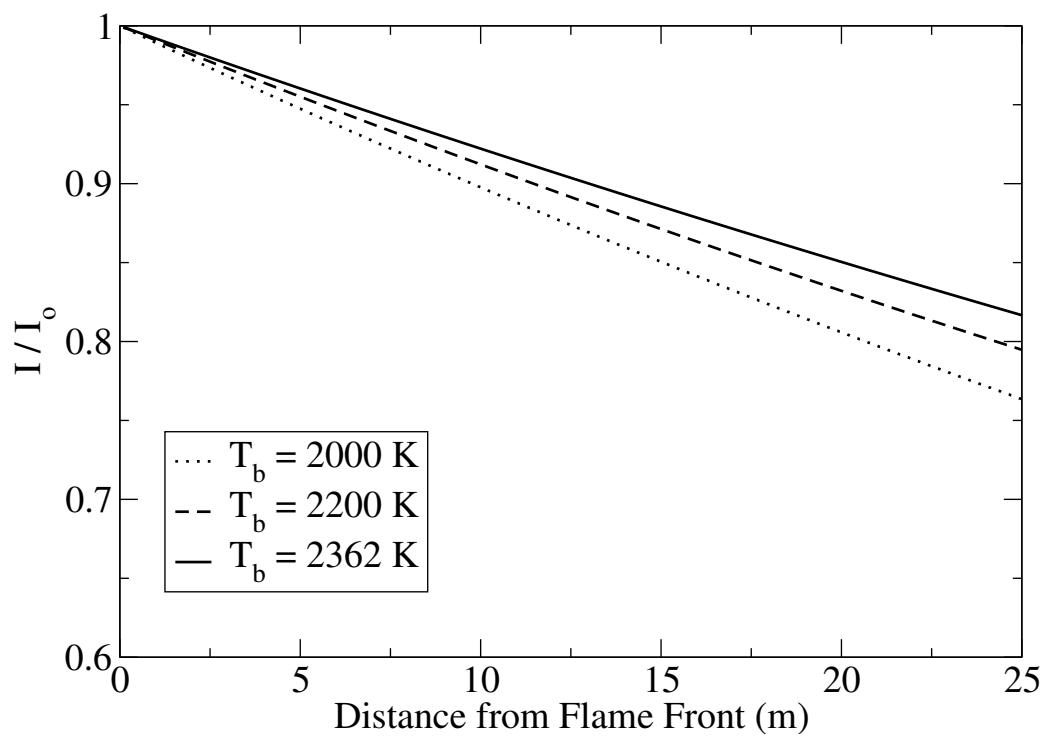
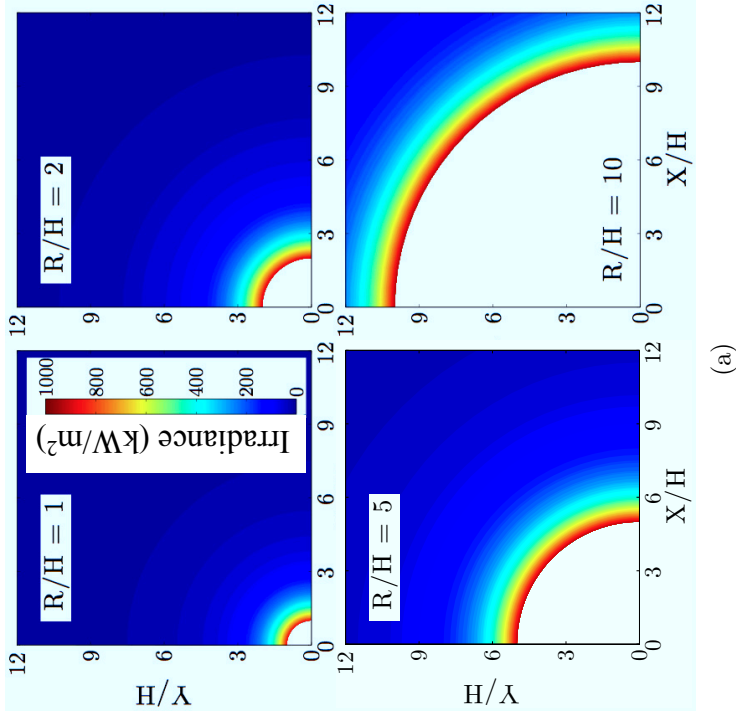
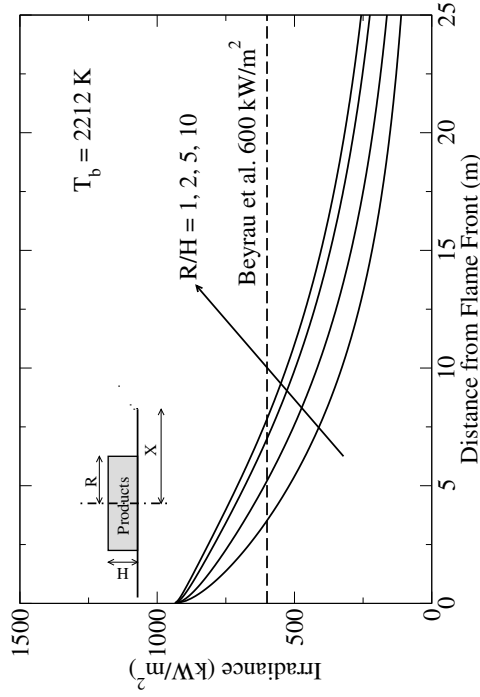


Fig. 9: The ratio of the incident to source flux ( $I/I_0$ ) transmitted through a fuel-air mixture for different product cloud temperatures. Conditions: collimated blackbody radiation into a stoichiometric ethylene/air mixture at 1 atm, 298 K and 100 % humidity.



(a)



(b)

Fig. 10: Spatial distribution of irradiance in the unburned region due to flame/particle geometry for different product cloud radius ( $R/H = 1, 2, 5, 10$ ). Products were modelled as a squat cylinder of constant unburnt cloud height of 2 m. A flame temperature of 2212 K and unit emissivity was selected as a representative of a stoichiometric methane-air mixture. (a) The radial irradiance field – combustion products are shown in white, (b) variation of irradiance ahead of the flame front. Horizontal line represents the minimum radiation flux necessary to induce ignitions times below 100 ms as reported by Beyrau et al. (2013).

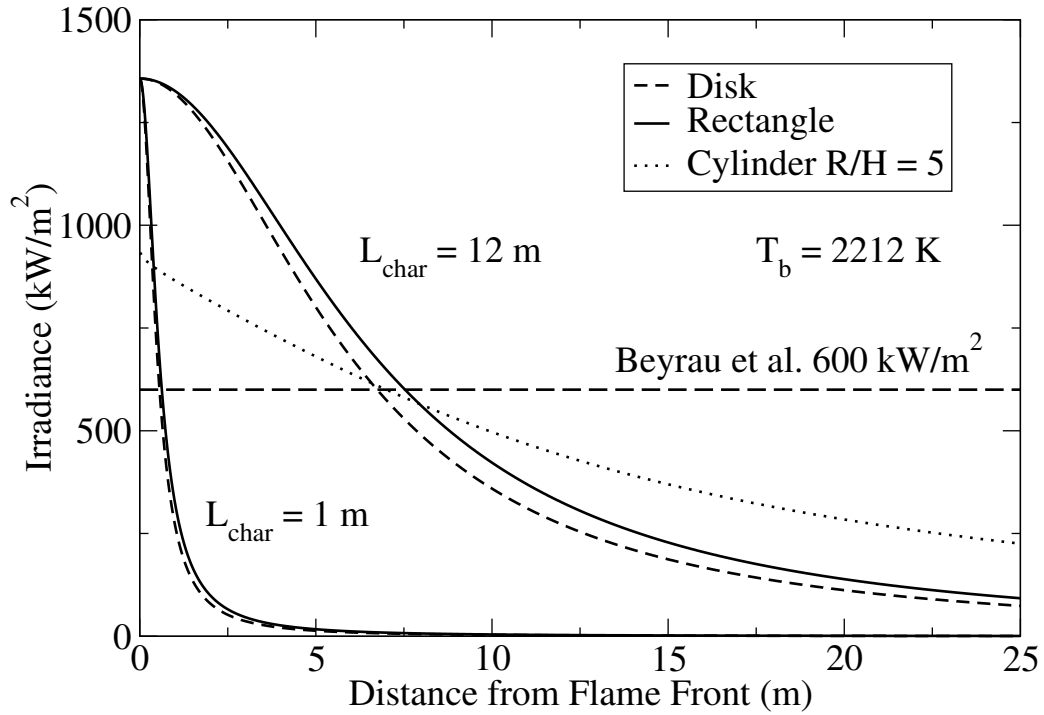


Fig. 11: Spatial distribution of irradiance in the unburned region due to flame/particle geometry. Flame modelled as a straight cylinder, and planar circular and square surface. A flame temperature of 2212 K and unit emissivity was selected as a representative of a stoichiometric methane-air mixture. Horizontal line represents the minimum radiation flux necessary to induce ignitions times below 100 ms as reported by Beyrau et al. (2013).

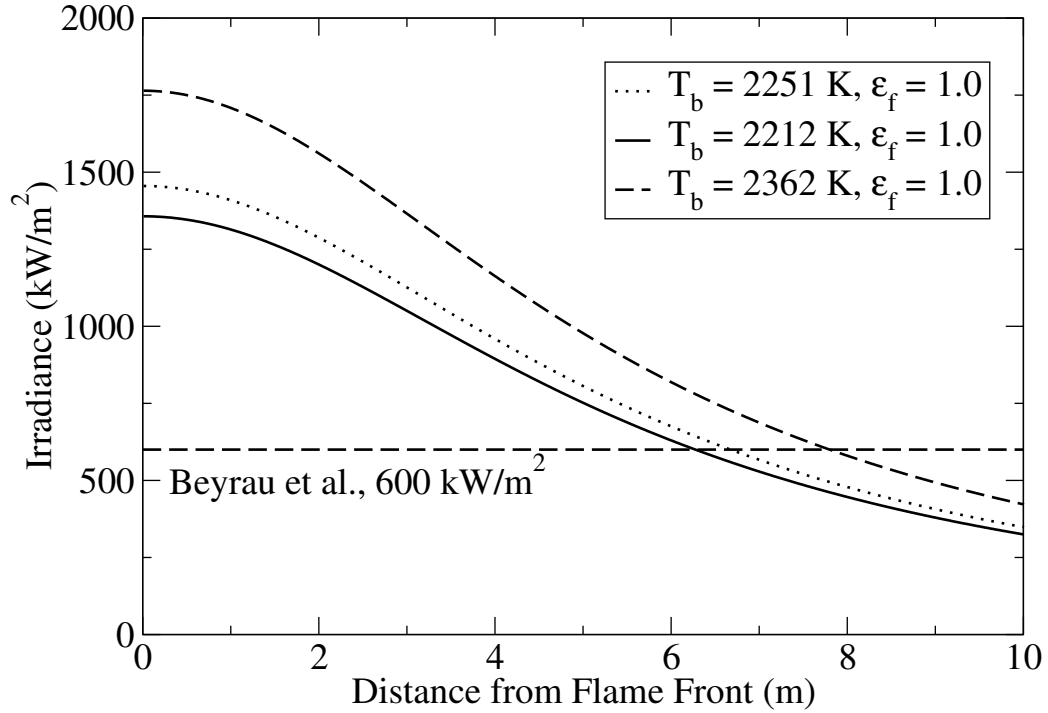


Fig. 12: The influence of the fuel type on the spatial distribution of irradiance in the unburned region due to flame/particle geometry. Case 3 (—), 13 (---) and 14 (···) from Table 2 have been considered, corresponding to a stoichiometric methane, ethane and ethylene-air mixture respectively. Flame modelled as a square source ( $L_{char} = 10$  m). Horizontal line represents the minimum radiation flux necessary to induce ignition times below 100 ms as reported by Beyrau et al. (2013).



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