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# 1 Sulphide mineral evolution and metal 2 mobility during alteration of the oceanic 3 crust: Insights from ODP Hole 1256D

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## 10 **Abstract**

11 Fluxes of metals during the hydrothermal alteration of the oceanic crust have far reaching effects  
12 including buffering of the compositions of the ocean and lithosphere, supporting microbial life and the  
13 formation of sulphide ore deposits. The mechanisms responsible for metal mobilisation during the  
14 evolution of the oceanic crust are complex and are neither fully constrained nor quantified. Investigations  
15 into the mineral reactions that release metals, such as sulphide leaching, would generate better  
16 understanding of the controls on metal mobility in the oceanic crust. We investigate the sulphide and  
17 oxide mineral paragenesis and the extent to which these minerals control the metal budget in samples from  
18 Ocean Drilling Program (ODP) Hole 1256D. The ODP Hole 1256D drill core provides a unique sample  
19 suite representative of a complete section of a fast-spreading oceanic crust from the volcanic section down  
20 to the plutonic complex. The sulphide population at Hole 1256D is divided into five groups based on  
21 mineralogical assemblage, lithological location and texture: the magmatic, metasomatised, high  
22 temperature hydrothermal, low temperature and patchy sulphides. The initiation of hydrothermal alteration

23 by downward flow of moderate temperature (250-350 °C) hydrothermal fluids under oxidising conditions  
24 leads to metasomatism of the magmatic sulphides in the sheeted dyke and plutonic complexes. Subsequent  
25 increase in the degree of hydrothermal alteration at temperatures >350 °C under reducing conditions then  
26 leads to the leaching of the metasomatised sulphides by rising hydrothermal fluids. Mass balance  
27 calculations show that the mobility of Cu, Se and Au occurs through sulphide leaching during high  
28 temperature hydrothermal alteration and that the mobility of Zn, As, Sb and Pb is controlled by silicate  
29 rather than sulphide alteration. Sulphide leaching is not complete at Hole 1256D and more advanced  
30 alteration would mobilise greater masses of metals. Alteration of oxide minerals does not release  
31 significant quantities of metal into the hydrothermal fluid at Hole 1256D. Mixing of rising high  
32 temperature fluids with low temperature fluids, either in the upper sheeted dyke section or in the  
33 transitional zone, triggers local high temperature hydrothermal sulphide precipitation and trapping of Co,  
34 Ni, Cu, Zn, As, Ag, Sb, Se, Te, Au, Hg and Pb. In the volcanic section, low temperature fluid circulation  
35 (<150 °C) leads to low temperature sulphide precipitation in the form of pyrite fronts that have high As  
36 concentrations due to uptake from the circulating fluids. Deep late low temperature circulation in the  
37 sheeted dyke and the plutonic complexes results in local precipitation of patchy sulphides and local metal  
38 remobilisation. Control of sulphides over Au, Se and Cu throughout fast-spreading mid-oceanic crust  
39 history implies that the generation of hydrothermal fluids enriched in these metals, which can eventually  
40 form VMS deposits, is strongly controlled by sulphide leaching.

## 41 **1. Introduction**

42 The cycling of metals in the oceanic crust is an important process which buffers the composition of the  
43 lithosphere and oceans, contributes to development of microbial life, and leads to the formation of  
44 hydrothermal ore deposits such as volcanogenic massive sulphide (VMS) deposits. Investigations from  
45 modern-day oceanic crust (Nesbitt et al., 1987; Alt et al., 1989; Alt, 1995; Alt et al., 1996; Gillis et al.,  
46 2001; Bach et al., 2003; Heft et al., 2008; Teagle et al., 2010; Alt et al., 2010; Coogan and Dosso, 2012;

47 Patten et al., 2015) and ophiolites (e.g. Richardson et al., 1987; Schiffman et al., 1987; Jowitt et al., 2009)  
48 suggest that mobilisation of metals from the lower sheeted dyke complex during on-axis high temperature  
49 hydrothermal alteration of the oceanic crust is a systematic phenomenon. The high temperature (>350 °C)  
50 hydrothermal fluids that leach metals from the lower sheeted dyke section rise buoyantly through upflow  
51 zones towards the seafloor where venting occurs (Alt et al., 1989; Alt et al., 2010; Hannington, 2014).  
52 During this ascent metals can be lost to mineralised horizons within the upper oceanic crust (Honnorez et  
53 al., 1985; Alt et al., 1998; Bach et al., 2003; Alt et al., 2010; Hannington 2013), trapped within VMS  
54 deposits (e.g. Fouquet et al., 1988; Herzig and Hannington, 1995; Hannington et al., 1998; Wohlgemuth-  
55 Ueberwasser et al., 2015) or vented in hydrothermal plumes and associated sediments (e.g. Feely et al.,  
56 1994a; 1994b; Hannington 2013). Mobilisation of metals in the lower sheeted dykes occurs at high  
57 temperature and under reduced conditions (>350 °C; e.g. Alt et al., 2010) when solubility of metals is  
58 optimal (Seewald and Seyfried; 1990). It is generally assumed that this mobilisation is partly achieved  
59 through leaching of magmatic sulphides by hydrothermal fluids (e.g. Keays, 1987; Alt, 1995; Jowitt et al.,  
60 2012). The extent, however, to which this mineral reaction controls the release of base metals and other  
61 trace metals such as Au, As, Sb, Se and Te into the hydrothermal fluids is neither constrained nor  
62 quantified. Reactions involving oxide and silicate minerals have also been shown to mobilise metals  
63 including Zn (Doe, 1994; Jowitt et al., 2012). The mineral reactions that produce metal-rich fluids in the  
64 modern-day oceanic crust are not well constrained and systematic investigation of the sulphide population  
65 evolution during hydrothermal alteration would greatly improve our understanding of metal mobilisation  
66 in the oceanic crust.

67 Located in the Cocos Plate, Pacific Ocean, ODP (Ocean Drilling Program) Hole 1256D is an ideal  
68 location to investigate the mobility of metals in a fast-spreading mid-oceanic crust. The 1256D drill core  
69 recovers a complete section of the oceanic crust down to the plutonic complex (Wilson et al., 2003; Teagle  
70 et al., 2006). The hydrothermal system in the crust at Hole 1256D and the resulting alteration are well  
71 defined (Teagle et al., 2006; Alt et al., 2010; Alt and Shanks, 2011; Violay et al., 2013), providing a solid

72 framework for investigation of mineral reactions leading to metal mobility. Recent investigations on Hole  
73 1256D have shown depletions in whole rock of Cu, Zn and Pb and also of Au, As, and Se concentrations  
74 in the deeper parts of the crust (Alt et al., 2010; Patten et al., 2015). These metals are depleted at different  
75 crustal levels ranging from the upper sheeted dykes down to the plutonic section implying that multiple  
76 mineral reactions may control metal mobility.

77 This investigation aims to determine the paragenesis and metal contents of the sulphide and oxide  
78 minerals in the ODP Hole 1256D crust in order to identify the mineral reactions that release metals into  
79 the hydrothermal fluids. We employ in-situ analytical methods including electron-probe microanalysis  
80 (EPMA) and laser ablation-inductively coupled plasma mass spectrometer (LA-ICP-MS) analysis to  
81 enable quantification of major, minor and trace elements in sulphide and oxide minerals present in the  
82 Hole 1256D drill core. We compare our data with published whole-rock concentrations for the same  
83 samples from Hole 1256D (Patten et al., 2015) in order to identify the host minerals for this suite of metals  
84 and the mineral reactions that liberate metals into hydrothermal fluids during alteration.

## 85 **2. Geological setting**

### 86 2.1 ODP Hole 1256D lithologic units

87 The oceanic crust at drilling site Hole 1256D is a 15 Myr old crust generated from a superfast spreading  
88 ridge (~200 mm/yr) and is located in the Cocos Plate (6.736° N, 91.934° W; Wilson et al., 2003). The  
89 oceanic crust can be divided into four main lithological units: the volcanic section, the transitional zone,  
90 the sheeted dyke complex and the plutonic complex (Fig. 1). The volcanic section is capped by a lava  
91 pond on the first 100 m that was formed during off-axis volcanism events (Wilson et al., 2003) underlined  
92 by phyrlic to aphyric sheeted flows, aphyric massive units, pillow lavas and hyaloclastites (Wilson et al.,  
93 2003; Teagle et al., 2006). The transitional zone is mainly composed of aphyric sheeted flows with  
94 presence of breccias characterised by angular aphyric cryptocrystalline basaltic clasts cemented by

95 chalcedony, saponite, carbonate, albite, anhydrite and sulphides (Teagle et al., 2006). The sheeted dyke  
96 complex comprises massive aphyric basalts with common sub-vertical intrusive contacts (Teagle et al.,  
97 2006). Due to intrusion of underlying gabbros, high-temperature recrystallisation occurs in the lower part  
98 of the sheeted dyke complex giving a granoblastic texture (Teagle et al., 2006). The plutonic complex is  
99 characterised by two gabbro intrusive bodies (Teagle et al., 2006; Koepke et al., 2008). The contacts  
100 between the gabbro bodies and the dykes are intrusive, with dyke fragments occurring in the margin of the  
101 lower gabbro body (Alt et al., 2010).

## 102 2.2 Hydrothermal system and metal mobility

103 The oceanic hydrothermal system at Hole 1256D can be divided into two simplified domains: an upper  
104 low temperature domain and a deeper high temperature domain (Teagle et al., 2006; Alt et al., 2010). The  
105 low temperature domain occurs mostly in the volcanic section (250-1004 mbsf) and is characterised by  
106 alteration formed from circulation of low-temperature (50°C-185°C) oxidised fluids (Fig. 1, Alt et al.,  
107 2010). The associated alteration style is a relatively low-intensity background alteration (2-20 %  
108 recrystallisation) where the primary mineralogy is partially replaced with saponite, celadonite, iron  
109 oxyhydroxides and chalcedony (Teagle et al., 2006; Alt et al., 2010). Minor pyrite and trace chalcopryrite  
110 are associated with this alteration and occur as disseminated grains in the groundmass or in veins.

111 Substantial As enrichment of the volcanic section is associated with this alteration (Patten et al., 2015).  
112 Pyrite fronts occur on the borders of oxidised alteration halos (Alt and Shanks, 2011). Significant As and  
113 Sb (~20 %) enrichments are associated with these alteration halos (Patten et al., 2015).

114 The high temperature domain is characterised by alteration formed from circulation of high-temperature  
115 (300°C to >650°C; Fig. 1) reduced fluids and occurs in the sheeted dyke and plutonic complexes (1061-  
116 1507 mbsf; Alt et al., 2010). The pervasive background alteration is characterised by chlorite, actinolite,  
117 albite, titanite and pyrite which correspond to sub-greenschist and greenschist facies conditions (Teagle et  
118 al., 2006; Alt et al., 2010). Mobilisation of Au, As and Se associated with the background alteration occurs

119 in the upper part of the sheeted dyke complex (Patten et al., 2015). Veins of quartz, chlorite, epidote,  
120 pyrite, chalcopyrite, magnetite and sphalerite overprint the background alteration and have been  
121 interpreted to represent the pathways of rising high temperature fluids from the lower sheeted dyke section  
122 and the plutonic complex (Alt et al., 2010). Below 1300 mbsf actinolite is more common than chlorite,  
123 indicating an increase in the temperature of alteration (Alt et al., 2010), which also corresponds to an  
124 increase in the extent of metal mobilisation in the crust (Cu, Zn, Pb, Au, As and Se; Alt et al., 2010; Patten  
125 et al., 2015). The transitional zone and the upper sheeted dyke section correspond to the zone of  
126 interaction of the two fluid domains (Fig. 1). The large difference in temperature between the upper and  
127 lower fluid domains triggers sulphide mineralisation characterised by a 2.8 m mineralised breccia at 1028  
128 mbsf enriched in Zn, Au, As, Sb, Se and Te (Teagle et al., 2006; Alt et al., 2010; Patten et al., 2015). The  
129 sulphide assemblage associated with this mineralisation is pyrite and sphalerite with trace chalcopyrite. A  
130 number of similar, but smaller, mineralised veins also occur in the upper sheeted dykes (Teagle et al.,  
131 2006; Alt et al., 2010).

### 132 **3. Sampling and analytical methods**

#### 133 3.1 EPMA analyses

134 Rock chips from the Hole 1256D drill core were obtained from the collection at the National  
135 Oceanographic Centre (NOC) at Southampton, U.K. Optical microscopy by transmitted and reflective  
136 light was carried out on 45 polished thin sections representative of the whole crust section. Of the 45 thin  
137 polished sections selected, 14 had suitable sulphide and oxide minerals for EPMA and LA-ICP-MS  
138 analyses. Electron microprobe analyses were carried out at Uppsala University, Sweden, with a JXA-  
139 8530F JEOL SUPERPROBE. For sulphide analyses an accelerating voltage of 20 kV and a beam current  
140 of 20 nA were used for a 5µm beam, whereas for oxide analyses an accelerating voltage of 15 kV and a  
141 beam current of 10 nA were used. Acquisition time of 20 s and 10 s background was used for sulphide

142 analyses and acquisition time of 10 s and 5 s background was used for oxide analyses. A total of 314 spot  
143 analyses were carried out for sulphide minerals on pyrite, chalcopyrite, pyrrhotite, sphalerite, pentlandite,  
144 millerite and marcasite (Table 1-5), whilst 180 spot analyses were carried out on oxide minerals on  
145 magnetite, titanomagnetite and ilmenite (Table 6). One sigma detection limits for sulphide analyses are 35  
146 ppm for S, 43 ppm for Fe, 76 ppm for Cu, 77 ppm for Zn, 65 ppm for Ni and 58 ppm for Co. For oxide  
147 minerals detection limits are 104 ppm for Si, 64 ppm for Al, 64 ppm for Mg, 329 ppm for Mn, 161 ppm  
148 for Ti, 83 ppm for Ca and 119 ppm for Fe. The Fe<sub>2</sub>O<sub>3</sub> content of oxide minerals are calculated using the  
149 methods described in Carmichael (1966) and Lepage (2003).

### 150 3.2 LA-ICP-MS analyses

151 Analyses by LA-ICP-MS were carried out at the PetroTectonics facility at Stockholm University using a  
152 NWR-213 nm solid state laser coupled to a ThermoFisher XSeries II quadrupole mass spectrometer. Spot  
153 sizes of 30 and 60 µm were used with a laser pulse frequency of 7 Hz and a laser energy density of 8.4  
154 J.cm<sup>-2</sup> for sulphide analyses. A total of 337 spot analyses were carried out on pyrite, chalcopyrite,  
155 pyrrhotite, sphalerite and marcasite (Table 1-5). Sulphide grains smaller than 30 µm including pentlandite  
156 and millerite could not be efficiently analysed. Each analysis consisted of 60 s background, 40 s ablation  
157 and 30 s wash-out. After wash-out, Se isotopes showed long-lived residual signals and additional wash-  
158 outs were carried out until the residual signals faded away. Data from EPMA analyses of <sup>33</sup>S were used as  
159 an internal standard. The following isotopes were monitored: <sup>33</sup>S, <sup>57</sup>Fe, <sup>60</sup>Ni, <sup>65</sup>Cu, <sup>66</sup>Zn, <sup>75</sup>As, <sup>76</sup>Se, <sup>77</sup>Se,  
160 <sup>78</sup>Se, <sup>82</sup>Se, <sup>105</sup>Pd, <sup>106</sup>Pd, <sup>107</sup>Ag, <sup>108</sup>Pd, <sup>109</sup>Ag, <sup>121</sup>Sb, <sup>123</sup>Sb, <sup>125</sup>Te, <sup>193</sup>Ir, <sup>195</sup>Pt, <sup>197</sup>Au, <sup>200</sup>Hg, <sup>202</sup>Hg, <sup>208</sup>Pb and the  
161 preferred isotopes used for calculations are listed in Appendix A.1. Dwell times were 10 ms/peak for <sup>33</sup>S,  
162 <sup>57</sup>Fe, <sup>200</sup>Hg and <sup>202</sup>Hg; 20 ms/peak for <sup>60</sup>Ni, <sup>65</sup>Cu, <sup>66</sup>Zn, <sup>76</sup>Se, <sup>106</sup>Pd, <sup>108</sup>Pd and <sup>208</sup>Pb; 30 ms/peak for <sup>82</sup>Se,  
163 <sup>105</sup>Pd, <sup>107</sup>Ag, <sup>109</sup>Ag, <sup>125</sup>Te, <sup>193</sup>Ir and <sup>195</sup>Pt; and 40 ms/peak for <sup>121</sup>Sb, <sup>123</sup>Sb and <sup>197</sup>Au. Calibration was carried  
164 out using three reference materials: Trans-1, a synthetic Fe-Ni sulphide doped with 17-25 ppm As, Sb, Se  
165 and Te (Wohlgemuth-Ueberwasser et al., 2014); NiS<sub>3</sub>, a fused nickel sulphide doped with 20-24 ppm Au,  
166 Pt, Pd, Ir, Os, Ru and Rh (Gilbert et al., 2013); and MASS-1, a pressed pellet (Fe-Zn-Cu-S) doped with



167 50–70 ppm As, Ag, Pb, Sb, Se and Te (USGS, Wilson et al., 2002; Appendix A.1). Additionally, MASS-  
168 1, NiS<sub>3</sub> and PGE\_8a were also used as monitors (Appendix A.2). Accuracy and precision of calibrants  
169 and monitors are shown in Appendix A.1 and A.2. Isotopes monitored are falling within 10 % of the  
170 published values except for <sup>121</sup>Sb and <sup>197</sup>Au, which are 32 % and 25 % higher respectively than MASS-1.  
171 This discrepancy could be due to oxide interferences due to the water present in MASS-1 (Wohlgemuth-  
172 Ueberwasser et al., 2014). No appropriate reference materials for monitoring <sup>65</sup>Cu, <sup>66</sup>Zn, <sup>107</sup>Ag, <sup>125</sup>Te,  
173 <sup>200</sup>Hg and <sup>208</sup>Pb were available and the implications on the study are detailed in Appendix B. Data  
174 reduction and calculation of trace metal concentrations were performed off-line through IOLITE software  
175 (Hellstrom et al., 2008, Paton et al., 2011). Detection limits were calculated by IOLITE software at three  
176 times the square root of the blanks. Trace metal signals in sulphide minerals are inclusion free which  
177 suggests no nugget effect.

178 Analyses by LA-ICP-MS of oxide minerals were carried out with the same equipment as sulphide  
179 analyses. A spot size of 35 µm was used with a laser pulse frequency of 10 Hz. Each analysis consisted of  
180 50 s background, 40 s ablation and 30 s wash-out. Data from EPMA analyses of <sup>56</sup>Fe were used for an  
181 internal standard. Coarse oxide grains in the plutonic section show fine exsolution textures and their  
182 ablation by LA-ICP-MS results in a mixed signal of magnetite and ilmenite. The true Fe concentration is  
183 calculated following the Dare et al. (2014) method and using EPMA data for both ilmenite and magnetite.  
184 The analyses are thus likely to represent the composition of the oxide before sub-solidus exsolution (Dare  
185 et al. 2012). The following isotopes were monitored: <sup>23</sup>Na, <sup>29</sup>Si, <sup>31</sup>P, <sup>33</sup>S, <sup>44</sup>Ca, <sup>47</sup>Ti, <sup>49</sup>Ti, <sup>51</sup>V, <sup>52</sup>Cr, <sup>56</sup>Fe, <sup>57</sup>Fe,  
186 <sup>60</sup>Ni, <sup>65</sup>Cu, <sup>66</sup>Zn, <sup>75</sup>As, <sup>82</sup>Se, <sup>121</sup>Sb, <sup>125</sup>Te, <sup>195</sup>Pt, <sup>197</sup>Au, <sup>208</sup>Pb. Dwell times were 2 ms/peak for <sup>23</sup>Na, <sup>29</sup>Si and  
187 <sup>44</sup>Ca; 10 ms/peak for <sup>51</sup>V, <sup>52</sup>Cr, <sup>60</sup>Ni, <sup>65</sup>Cu, <sup>66</sup>Zn and <sup>208</sup>Pb; 15 ms/peak for <sup>75</sup>As; 20 ms/peak for <sup>82</sup>Se and  
188 <sup>121</sup>Sb; 25 ms/peak for <sup>47</sup>Ti, <sup>49</sup>Ti and <sup>125</sup>Te; 40 ms/peak for <sup>33</sup>S; 45 ms/peak for <sup>195</sup>Pt; and 60 ms/peak for  
189 <sup>197</sup>Au. Calibration was performed using NIST 610, a fused glass doped with sixty-one trace elements  
190 (Jochum et al., 2011, Appendix A.3). Accuracy and precision of calibration was checked by monitoring  
191 the isotopes using BCR-2G a natural basaltic glass (Jochum et al., 2005; Jochum and Nohl, 2008; Jochum

192 et al., 2014; Appendix A.3). Matrix match standards were not available for oxide analyses and NIST 610  
193 and BCR-2G were used following the methods of (Dare et al., 2012; Nadoll et al., 2014). Good accuracy  
194 and precision are reached for  $^{47}\text{Ti}$ ,  $^{49}\text{Ti}$ ,  $^{51}\text{V}$ ,  $^{52}\text{Cr}$ ,  $^{65}\text{Cu}$ ,  $^{66}\text{Zn}$ ,  $^{121}\text{Sb}$  and  $^{208}\text{Pb}$  (Appendix A.3). Sample  
195 analyses for  $^{82}\text{Se}$ ,  $^{125}\text{Te}$ ,  $^{195}\text{Pt}$  and  $^{197}\text{Au}$  are below detection limits. Sulphur was monitored in order to  
196 identify possible sulphide inclusions, and Si, Ca and P were monitored to control matrix silicate  
197 contamination or titanite alteration of the oxides. Arsenic analyses give signals close to the background  
198 with high standard deviation, and furthermore, count per second (CPS) values correlate with Ca  
199 suggesting a possible interference of  $^{43}\text{Ca}^{16}\text{O}_2$  on  $^{75}\text{As}$ . The  $^{75}\text{As}$  data are thus considered to be below the  
200 detection limit. The  $^{65}\text{Cu}$  values are likely to be a real Cu signal rather than due to  $^{49}\text{Ti}^{16}\text{O}^+$  interference, as  
201 the oxides with high  $^{65}\text{Cu}$  CPS values show no correlation with Ti oxide content. Interference of  $^{50}\text{Ti}^{16}\text{O}^+$   
202 on  $^{66}\text{Zn}$  is interpreted to be negligible due to the low natural abundance of  $^{50}\text{Ti}$  (5.25 %).

## 203 **4. Results**

### 204 4.1 Sulphide classification

205 The sulphide population in Hole 1256D is classified using assemblage, texture, grain size and distribution.  
206 Three main processes can account for the majority of sulphide mineral formation: 1) magmatic  
207 crystallisation, 2) high temperature fluid circulation, and 3) low temperature fluid circulation.

#### 208 *4.1.1 Magmatic sulphides*

209 Magmatic sulphides consist of pyrrhotite, chalcopyrite and pentlandite that are characteristic of high  
210 temperature magmatic sulphides (Fig. 3A and B; Czamanske and Moore, 1977; Mathez, 1979; Peach et  
211 al., 1990; Barnes et al. 2006; Patten et al., 2012). They have been observed in two samples from the  
212 volcanic section and the plutonic complex at 445 mbsf and 1493 mbsf respectively (Fig. 2) and represent  
213 only a minor proportion of the Hole 1256D sulphide population. In the volcanic section, magmatic  
214 sulphides occur as 10-50 micron-wide blebs composed of pyrrhotite and chalcopyrite intergrowths

215 representative of quenched monosulphide solid solution (MSS) and intermediate solid solution (ISS; Fig  
216 3A; e.g. MacLean, 1977; Czamanske and Moore, 1977; Patten et al., 2012). Similar sulphide droplets are  
217 observed in fresh MORB glass (Czamanske and Moore, 1977; Mathez, 1979; Francis et al., 1990; Patten  
218 et al., 2012) and from other oceanic crust drill core material (MacLean, 1977; Alt et al., 1989). In the  
219 plutonic complex magmatic sulphides do not have systematic shape but are larger than in the volcanic  
220 section (up to hundreds of microns in width), and are composed of pyrrhotite, chalcopyrite and pentlandite  
221 (Fig. 3B). The latter occurs as a lamellae exsolution within the pyrrhotite, which is characteristic of slowly  
222 cooled magmatic sulphides that have exsolved from MSS and ISS (Fig. 3B; Kelly and Vaughan, 1983; Alt  
223 and Anderson, 1991; Barnes et al., 2006; Dare et al., 2010; Djon et al., 2012). Mineral proportions by  
224 visual estimation are 90 % pyrrhotite with trace exsolution pentlandite and 10 % chalcopyrite.

225 Quenched sulphide blebs from the volcanic section are a mixture of MSS and ISS and have different  
226 composition from the magmatic sulphides in the plutonic complex (Table 1). The sulphide blebs from the  
227 volcanic section have the Fe and S composition of pyrrhotite but show high and variable Cu, Zn, Pb and  
228 trace metals concentrations due to contamination from small chalcopyrite inclusions (Table 1). Large  
229 pyrrhotite grains from the plutonic complex contain trace amounts of Co, Ni, Cu, Zn, As, Se, Ag, Te, Au  
230 and Pb. Concentrations of Sb, platinum group elements (PGE; Pt, Pd, Ir, Os, Ru and Rh) and Hg are below  
231 detection limits. Chalcopyrite from the plutonic complex has higher concentrations of Zn (1169 ppm, 2-  
232 5000 ppm), Ag ( $5.38 \pm 3.38$ ), Te ( $2.82 \pm 1.36$  ppm), and Pb (3.29 ppm, 0.28-10.6 ppm) than pyrrhotite  
233 and pentlandite, and contains trace amounts of Co, Ni, Se and Au (Table 1). Concentrations of As, Sb,  
234 PGEs and Hg are below detection limit in chalcopyrite. Selenium and Au concentrations in pyrrhotite and  
235 chalcopyrite are similar ( $45.2 \pm 8.2$  ppm and  $49.6 \pm 18.7$  ppm for Se respectively, and  $0.023 \pm 0.019$  ppm  
236 and  $0.027 \pm 0.017$  for Au respectively, Table 1). Pentlandite, unlike pyrrhotite and chalcopyrite, does not  
237 show stoichiometric composition due to high concentrations of Co ( $16.5 \pm 3.6$  wt.%; Table 1) that can  
238 substitute for Fe ( $19.8 \pm 6.4$  wt.%) and Ni ( $21.5 \pm 4.0$  wt.%; Riley, 1977). The pentlandite Cu and Zn

239 contents are below the detection limit for EPMA and analyses by LA-ICP-MS were impossible due to too  
240 small grain size.

#### 241 4.1.2 *Metasomatised sulphides*

242 Metasomatised sulphides consist of pyrite, chalcopyrite, millerite and magnetite, an assemblage that is  
243 characteristic of recrystallised magmatic sulphides (Fig. 3C, D, E and F; Table 2; Alt., 1989). This group  
244 occurs mostly in the sheeted dyke and the plutonic complexes with a few occurrences in the volcanic  
245 section (Fig. 2). Pyrite is the most abundant mineral of the group and occurs as grains up to hundreds of  
246 microns in size that are commonly anhedral in shape and can be porous in places (Fig. 3D). Chalcopyrite  
247 is homogeneous and occurs either as grains associated with the pyrite or as isolated grains. It ranges in  
248 size from a few microns up to tens of microns. Millerite generally occurs at the contact between pyrite and  
249 chalcopyrite and is the smallest phase of the assemblage with a size that does not exceed a few microns  
250 (Fig. 3F). Magnetite occurs as small micron-sized grains associated with pyrite, often forming rims  
251 surrounding the sulphide assemblage (Fig. 3C, D and E).

252 Metasomatised pyrites contain traces of Co, Ni, Cu, Zn, As, Sb, Te, Hg, Au and Pb (Table 2). Few grains  
253 show signals above the detection limit for Pd and Pt, and other PGEs are below detection limits. Trace  
254 metal concentrations in metasomatised pyrites are highly variable and positively skewed as highlighted by  
255 the discrepancy between the average and median values (Table 2). This discrepancy is due to localised  
256 enrichment of trace metals during metasomatism and leaching (see discussion). Selenium and As can  
257 substitute for S in pyrite (e.g. Genna and Gaboury, 2015) and the Se/As ratio is a useful parameter to  
258 determine pyrite genesis in magmatic or hydrothermal systems (Duran et al., 2015). Metasomatised  
259 pyrites are characterised by high Se/As, suggesting a magmatic origin (Duran et al., 2015), and  
260 intermediate to high Ni content (Fig. 4). Chalcopyrite contains traces of Co, Ni, As, Se, Ag, Sb, Te, Hg,  
261 Au and Pb (Table 2) whereas PGEs are below detection limit. Trace metal concentrations in  
262 metasomatised chalcopyrite are also variable but not as pronounced as in the metasomatised pyrites (Table

263 2). Millerite has low Ni ( $51.3 \pm 12.3$  wt.%) relative to stoichiometric value, which can be accounted for by  
264 the high concentration of Co and Fe ( $4 \pm 1.9$  wt.% and  $6.4 \pm 7.6$  wt.%, respectively).

#### 265 4.1.3 High temperature hydrothermal sulphides

266 High temperature hydrothermal sulphides represent the largest sulphide occurrence of the Hole 1256D.  
267 They occur mostly in the upper sheeted dyke section and the transitional zone (Fig. 2) within mineralised  
268 breccias, quartz-chlorite dominated veins or as disseminated grains (Fig. 1). They are characterised by  
269 euhedral to subhedral grains of pyrite, chalcopyrite and sphalerite. Pyrite grains vary in size from small  
270 (tens of microns) grains that form porous aggregates, to large (hundreds of microns) grains (Fig. 5A, B  
271 and C). Chalcopyrite is common in the upper sheeted dyke section but scarce in the transitional zone (Fig.  
272 2). Conversely, sphalerite occurs in trace amounts in veins from the upper sheeted dyke section but is the  
273 second most common sulphide mineral in the transitional zone mineralised breccias (Fig. 2). In the upper  
274 sheeted dyke section sphalerite commonly shows the chalcopyrite disease texture (Fig. 5A) whereas in the  
275 transitional zone it is euhedral to subhedral (Fig. 5C). This distribution of sulphide phases is similar to that  
276 reported in the mineralised zone in Hole 504B (Honnorez et al., 1985), although the sulphide  
277 mineralisation in Hole 1256D is less well developed than in Hole 504B.

278 High temperature pyrites contain trace amounts of Co, Ni, Cu, Zn, As, Sb, Te, Hg, Au and Pb (Table 3).  
279 Few grains show signals above the detection limit for Pd and other PGEs are below detection limit.  
280 Relative to metasomatised pyrites, high temperature pyrites have higher median values of As (23.7 ppm),  
281 Ag (2.39 ppm) and Au (0.12 ppm) but lower Ni (30.0 ppm), Cu (25.9 ppm) and Pb (14.6 ppm); other trace  
282 metals are within the same order of magnitude (Table 2 and 3). The variability in trace metal  
283 concentrations in high temperature pyrites is most likely due to different trapping efficiency during  
284 sulphide precipitation. Differences in composition between high temperatures pyrite from the transitional  
285 zone and the upper sheeted dyke section can be observed (Table 3). The former shows higher As ( $31.1 \pm$   
286  $18.7$  ppm), Sb (0.67 ppm, 0.04-4.07 ppm), Au ( $0.12 \pm 0.10$  ppm) and Pb (104.6 ppm, 0.5-408 ppm) where  
287 the latter shows higher Ni (236 ppm, 11.8-1225 ppm). Other trace metals are within the same order of

288 magnitude in the two different groups. High temperature pyrites are overall characterised by low Ni  
289 concentration and large Se/As (Fig. 4). Chalcopyrites contain trace amounts of Ni, Zn, As, Pd, Sb, Te, Au  
290 and Pb whereas Co, PGE and Hg are below detection limit (Table 3). Zinc concentrations in sphalerite are  
291 low ( $55.2 \pm 2.8$  wt.%) due to variable concentration of Fe ( $8.5 \pm 3.1$  wt.%). The Fe/Zn ratio of inclusion-  
292 free sphalerite can be used to calculate the minimum fluid temperature from which sphalerite precipitates  
293 (Keith et al., 2014). An average temperature of  $378 \pm 24$  °C is calculated for sphalerite from the  
294 transitional zone from EPMA data. Calculation for sphalerite from the upper sheeted dyke section cannot  
295 be carried out because of systematic chalcopyrite disease texture (Keith et al., 2014). This texture also  
296 explains the high Cu variability in sphalerite (up to 3.8 wt.%; Table 3) especially in the upper sheeted  
297 dyke section. Sphalerite contain trace amounts of Co, Ni, Cu, Zn, As, Se, Ag, Sb, Hg and Pb. Within the  
298 high temperature hydrothermal sulphides, pyrites have the highest Co, As, Te and Au concentrations,  
299 whereas Ni, Se and Ag concentrations are higher in chalcopyrite, and Sb, Hg and Pb in sphalerite (Table  
300 3; e.g. Monecke et al., 2016).

#### 301 *4.1.4 Low temperature sulphides*

302 Sulphides from the low temperature sulphide group are common throughout the volcanic section and are  
303 characterised by pyrite, millerite and trace chalcopyrite (Fig. 2). Micron sized pyrite and chalcopyrite  
304 grains are commonly disseminated in the ground mass associated with the background alteration, but the  
305 bulk of the low temperature sulphide occurs as “pyrite fronts” composed of pyrite and marcasite  
306 (Andrews, 1979). They occur as narrow zones hundreds of microns wide and centimeters long in which  
307 they replace the silicate matrix and fill pores and cracks (Figs. 5D, E and F). These pyrite fronts occur on  
308 the border of alteration halos (Wilson et al., 2003; Alt. and Shanks, 2011) or oxidation zones (Andrews,  
309 1979).

310 Distinction between pyrite and marcasite grains during in-situ analyses was not possible and the data in  
311 Table 4 represents average content of both minerals. They contain trace amounts of Co, Ni, Cu, Zn, As,  
312 Se, Ag, Sb and Pb (Table 4). Tellurium, PGEs, Au and Hg are below detection limits. Low temperature

313 pyrites show the highest As concentrations of the different sulphide groups ( $99 \pm 66$  ppm) but have low  
314 Cu (9.6 ppm, 1.01-57 ppm), Zn ( $3.3 \pm 2.8$  ppm), Ag ( $0.20 \pm 0.10$  ppm) and Pb ( $1.37 \pm 1.26$  ppm) relative  
315 to metasomatised and high temperature pyrites (Table 4). Because of high As concentrations, low  
316 temperature pyrites show low As/Se and are also characterized by high Ni content (Figure 4). Rare  
317 chalcopyrite grains were too small for in-situ analysis.

#### 318 *4.1.5 Patchy sulphides*

319 Patchy sulphides occur in the lower volcanic section, transitional zone, sheeted dyke and plutonic  
320 complexes (Fig. 2). They are characterised by a spherical web-like network up to one centimetre wide that  
321 replaces the silicate matrix (Fig 5H and I). Silicate phenocrysts, oxides and other sulphides are also  
322 affected. The patches do not seem to be associated with any specific type of silicate alteration. Pyrite is the  
323 most common mineral and is generally homogeneous and anhedral, but can also be porous. Minor  
324 chalcopyrite also occurs and one sample from the plutonic complex contains patches of marcasite.

325 Metal concentrations in the patchy sulphides (Table 5) are similar to the low temperature sulphides and  
326 are low relative to metasomatised and high temperature sulphides (Table 1-5). Pyrites contain trace  
327 amounts of Co, Ni, Cu, Zn, As, Se, Ag, Sb, Te, Au and Pb (Table 5). Platinum group elements and Hg are  
328 below detection limits. Similarly pyrite fronts, patchy pyrites are characterised by low Se/As and high Ni  
329 content (Fig. 4) which suggest similar paragenesis. A few in-situ analyses of chalcopyrite show trace  
330 amounts of Co, Ni, Zn, Ag and Pb (Table 5).

#### 331 4.2 Oxide classification

332 Oxide minerals are common in Hole 1256D and are classified into three main groups based on their  
333 texture, size, lithological occurrence and mineralogical assemblage: skeletal magmatic titanomagnetite,  
334 coarse magmatic magnetite-ilmenite and hydrothermal-related magnetite. Skeletal titanomagnetites are the  
335 most common oxide mineral and occur in the volcanic section and the sheeted dyke complex. They are  
336 present as micrometric disseminated skeletal grains or as dendritic chains (tens of microns, Fig. 6A;

337 Wilson et al., 2003). Their presence in the ground mass is ubiquitous (Wilson et al., 2003) and is generally  
338 confined to interstitial areas (Fig. 6A; Kempton et al., 1985). Disseminated micrometric ilmenite and  
339 magnetite grains are also locally observed. Titanomagnetites have  $\text{TiO}_2 > 2$  wt.%, average  $17.8 \pm 6.5$  wt.%  
340  $\text{TiO}_2$  and have the highest V, Cu, Zn, Sb and Pb concentrations of the three oxide groups (Table 6).

341 The second group of oxide minerals in Hole 1256D are magmatic related coarse grained magnetite and  
342 ilmenite (few mm) occurring in the plutonic complex (e.g. oxide gabbro; Teagle et al., 2006). They are  
343 characterised by coarse magnetite with trellis patterns of ilmenite exsolution or by ilmenite with magnetite  
344 exsolution (Fig. 6B; e.g. Kent et al., 1978). Homogeneous grains of magnetite or ilmenite are also present.  
345 Ilmenites average  $49.7 \pm 6.9$  wt.%  $\text{TiO}_2$  and magnetites average of  $3.4 \pm 1.4$  wt.%  $\text{TiO}_2$  (Table 6). Trace  
346 element concentrations analysed by LA-ICP-MS are a mixed signal of both magnetite and ilmenite and are  
347 listed in Table 6. Coarse magnetite-ilmenites have the highest Cr and the lowest Cu content of the three  
348 oxide group (Table 6).

349 The third group corresponds to secondary magnetite grains that are associated with hydrothermal fluid  
350 circulation. They either occur as precipitated euhedral grains within veins, as disseminated subhedral  
351 grains or as replacement of primary silicate and sulphide phases (Fig. 6C and D). They are present mostly  
352 in the sheeted dyke complex. Hydrothermal magnetites have low  $\text{TiO}_2$  concentrations ( $1.7 \pm 1.5$  wt.%) and  
353 are characterised by the highest Ni content and the lowest V, Zn and Pb content (Table 6). The magnetite  
354 rims associated with the metasomatised sulphide are part of the hydrothermal oxide group but, given their  
355 small size, only two grains were large enough to be analysed by LA-ICP-MS, with one of the two showing  
356 considerable sulphide contamination.

## 357 **5. Discussion**

358 The sulphide classification described above indicates that sulphides are sensitive to the different styles of  
359 hydrothermal alteration in the oceanic crust. In-situ analyses reveal significant variations in trace metal



360 content between the different sulphide mineral groups. The comparison between trace element  
361 concentrations in the different sulphide groups with whole rock metal concentrations allows some  
362 assessment of the degree to which the sulphide minerals control the trace metal budget in these rocks. This  
363 mass balance indicates which mineral reactions are responsible for metal mobilisation during the  
364 hydrothermal alteration of the oceanic crust.

## 365 5.1 The magmatic assemblage

366 Some primary sulphides and oxides have been sufficiently shielded from later hydrothermal fluid  
367 circulation that their magmatic characteristics have been preserved. Magmatic sulphides are, however,  
368 scarce in the Hole 1256D due to the extensive hydrothermal alteration. Magmatic S concentrations in Hole  
369 1256D average  $1250 \pm 200$  ppm (Alt and Shanks, 2011), which is high enough for a MOR basaltic melt to  
370 be sulphide-saturated before eruption (e.g. Mathez, 1979; Li and Ripley, 2005). Magmatic sulphides  
371 observed in Hole 1256D were most likely present as sulphide liquid within the magma prior to  
372 crystallisation, similar to the sulphide droplets observed in fresh MORB glass (Mathez, 1976; Peach et al.,  
373 1990; Patten et al., 2013). These sulphide droplets are quenched from  $\sim 1200$  °C to seawater temperature in  
374 a few minutes (Moore, 1975), resulting in their crystallisation as MSS and ISS intergrowths characteristic  
375 of magmatic temperatures (1200-1000 °C, Fig. 3A and 7; e.g. Czamanske and Moore, 1977; Mathez,  
376 1979; Patten et al., 2012). Magmatic sulphide droplets in the volcanic section are considered to represent  
377 these high-temperature sulphide assemblages. In the plutonic complex, however, where sulphide liquid is  
378 preserved from quenching, the observation of pyrrhotite with pentlandite exsolution as lamellae implies  
379 that the sulphides cooled below 650 °C, enabling the MSS and the ISS to exsolve to the observed phases  
380 (Fig. 3A and 7; Kelly and Vaughan, 1983; Patten et al., 2012).

381 The composition of magmatic sulphides from Hole 1256D (Table 1) differs slightly from previous reports  
382 of fresh MORB sulphide droplets through having lower chalcophile element concentrations (Co, Ni, Cu,  
383 Se, Ag, Te, Au; e.g. Patten et al., 2013). Differences in concentrations are attributed to different primary

384 magmatic composition rather than effect of alteration. Arsenic can be used as an indicator for fluid  
385 interaction with sulphides as it is a highly mobile element in fluids (e.g. Smedley and Kinniburgh, 2002).  
386 Pyrrhotite grains from the plutonic section have low As concentration ( $0.95 \pm 0.11$  ppm, Table 1)  
387 suggesting limited interaction with circulating fluids whereas sulphide blebs in the volcanic section have  
388 relatively high As concentration ( $5.12 \pm 1.34$  ppm) suggesting possible contamination by low temperature  
389 fluids. The magmatic sulphides from the plutonic complex can be interpreted to have conserved their true  
390 primary magmatic composition most likely due to limited interaction with hydrothermal fluids as  
391 implicated by the low water-rock ratio ( $<1$ ) below 1300 mbsf (Gao et al., 2012).

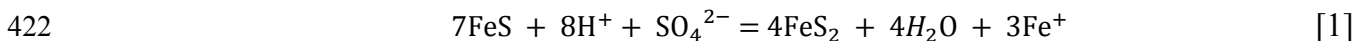
392 In the volcanic section, disseminated titanomagnetites have skeletal texture characteristic of a magmatic  
393 origin and occur in interstitial areas suggesting late-stage crystallisation (Kempton et al., 1985). Trellis  
394 pattern of ilmenite-magnetite exsolutions in coarse grain oxides in the plutonic complex represent sub-  
395 solidus oxy-exsolution during magma cooling (Dare et al., 2012), also indicating its magmatic origin. The  
396 trace element composition of the skeletal titanomagnetite composition indicates temperatures around  
397  $\sim 500^\circ\text{C}$  (Fig. 8D). Magmatic oxide minerals host minor Cu and Zn, traces of Pb and Sb and no As, Se, Te  
398 or Au. This elemental distribution can be partly explained by the Goldschmidt rule for cation substitution  
399 within the oxide structure, with  $\text{Cu}^{2+}$  and  $\text{Zn}^{2+}$  being substituted with  $\text{Fe}^{2+}$  in the octahedral sites, whereas  
400 other metallic cations cannot be substituted with Fe cations (e.g. Dare et al., 2012, Nadoll et al., 2014).  
401 Skeletal titanomagnetites are richer in Cu, Zn, Sb and Pb than the coarse magnetite-ilmenites (Table 6).  
402 Samples hosting skeletal titanomagnetites from the volcanic section are slightly more evolved than those  
403 hosting coarse magnetite-ilmenites from the plutonic section (Fig. 1). The primary metal content of the  
404 magma, however, is unlikely to account for the different metal contents of the two oxide groups as Cu, Zn,  
405 Pb and Sb show variable behaviour during magmatic differentiation with Cu for example, being enriched  
406 in more primitive melts and Pb, Zn and Sb being enriched in more evolved melts (Jenner et al., 2012). A  
407 more likely alternative is that the metal content of the magmatic oxides is related to the cooling rate of the  
408 magma and the presence of sulphide with which competition for metal trapping occurs (Dare et al., 2014).

409 In the volcanic rocks hosting skeletal titanomagnetite, the quick cooling rate of the magma at eruption  
410 may have inhibited effective diffusion of metals to the sulphide blebs (low R-factor; Barnes and Lightfoot,  
411 2005) allowing metals to be partly incorporated into oxides, with the converse occurring in the plutonic  
412 rocks that host coarse magnetite-ilmenite and coarse magmatic sulphides.

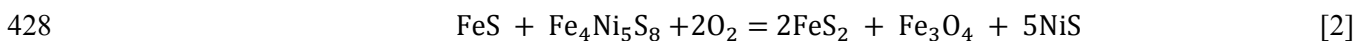
## 413 5.2 Effects of hydrothermal fluid circulation in the lower oceanic crust

### 414 5.2.1 *Sulphide metasomatism*

415 Hydrothermal fluid circulation in the lower portion of the oceanic crust leads to sub-solidus  
416 recrystallisation of the magmatic sulphides into metasomatised sulphides (Alt et al., 1989). The pyrrhotite,  
417 chalcopyrite and pentlandite are recrystallised to secondary pyrite, chalcopyrite, millerite and magnetite.  
418 Two processes can trigger the recrystallisation of pyrrhotite to secondary metasomatised pyrite: reduction  
419 of seawater sulphate at high temperature (>250 °C) or oxidation of the magmatic sulphides. Seawater  
420 sulphate ( $\text{SO}_4^{2-}$ ) reduction in the oceanic crust provides S as sulphide to the rock, leading to pyrrhotite  
421 (FeS) recrystallisation to secondary pyrite ( $\text{FeS}_2$ ; Shanks and Seyfried; 1987):



423  
424 The isotopic sulphur signature of metasomatised pyrite suggests that this reaction occurs in the sheeted  
425 dyke complex at Hole 504B (Alt et al., 1989). Alternatively, oxidation of pyrrhotite and pentlandite  
426 ( $\text{Fe}_4\text{Ni}_5\text{S}_8$ ) can lead to the formation of pyrite, magnetite ( $\text{Fe}_3\text{O}_4$ ) and millerite (NiS) at moderate  
427 temperature (up to 350 °C, Craig and Kullerud, 1969; Craig, 1973; Barnes et al., 2009; Djon et al., 2012):



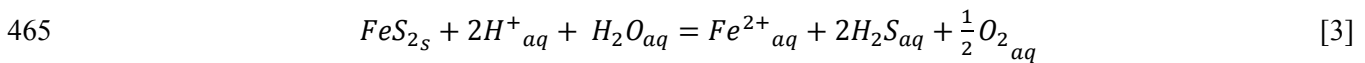
429  
430 The textures of the metasomatised sulphides in Hole 1256D support the latter process as they are  
431 characterised by metasomatised pyrite surrounded by a magnetite rim which represents an oxidation front  
432 where fluids have reacted with the pyrrhotite (Fig. 3C, D and E). Formation of metasomatised pyrite by

433 sulphate reduction is also likely to occur in Hole 1256D but the almost ubiquitous presence of magnetite  
434 associated with metasomatised sulphides suggests that the oxidation is the dominant reaction. Trace  
435 element concentrations of the magnetite rims suggest oxidation reaction at temperatures  $\sim 300$  °C (Fig. 8B;  
436 Nadoll et al., 2014). These temperatures and the mineralogical assemblage of metasomatised sulphides  
437 suggest that metasomatism is most likely driven by deep influx of recharging hydrothermal fluids at  
438 moderate temperature (200-350 °C; Alt et al., 1986; Alt and Chaussidon, 1989; Coumou et al. 2008; Harris  
439 2015) into the lower part of Hole 1256D during on-axis hydrothermal alteration (Tolstoy et al., 2008;  
440 Coumou et al., 2008). These fluids penetrate into the sheeted dyke complex along dyke margins (Harris et  
441 al., 2015) most probably during the early stages of hydrothermal alteration (Alt et al., 2010; Harris et al.,  
442 2015).

#### 443 5.2.2 *Sulphide leaching*

444 During on-axis hydrothermal fluid circulation, equilibration of the “recharging” moderate temperature  
445 hydrothermal fluids with the surrounding rocks in the sheeted dyke and plutonic complexes increases their  
446 temperature ( $>350$ °C) and acidity, and causes reduction of the fluids (e.g. Alt et al., 1986; Alt et al., 2010).  
447 These buoyant, acidic and reduced high-temperature hydrothermal fluids eventually rise within the  
448 sheeted dyke and plutonic complexes towards the seafloor following the similar pathways to those taken  
449 by the recharging fluids (Alt et al., 2010; Harris et al., 2015). Under these conditions, metal solubility in  
450 hydrothermal fluids significantly increases (Seewald and Seyfried, 1990) and sulphide minerals undergo  
451 leaching and dissolution (Fig. 7). These conditions correlate with the onset of depletion of metals from the  
452 Hole 1256D whole rock samples (Alt et al., 2010; Patten et al., 2015). The local preservation of magmatic  
453 and metasomatised sulphides in Hole 1256D indicates, however, that sulphide leaching is not complete  
454 (Fig. 2 and 9). Leaching of sulphides is likely to be dependent upon the degree of interaction with the high  
455 temperature fluids on a grain-scale as highlighted by the preservation of variably leached sulphide grains  
456 within the same sample. Degrees of leaching range from full preservation of pristine magmatic and  
457 metasomatised sulphides (Fig. 3E), to extensively leached sulphides (Fig. 9). In the extensively leached

458 sulphides, chalcopyrite grains are more abundant than pyrite (Fig. 9A, C and D) which suggests that  
459 chalcopyrite is less reactive to high temperature fluid circulation than pyrite (Fig. 10). This is supported by  
460 experimental studies that show Cu is less soluble than other base metals in hydrothermal fluids at  
461 temperature >300 °C in reduced conditions (e.g. Hemley et al., 1992). Pyrite occurs either as porous  
462 patches or isolated grains within the magnetite rim, and these textures indicate a significant sulphide  
463 volume loss (Fig. 9B, C and D). This volume loss is due to progressive release of S and Fe to the high  
464 temperature fluids (Crerar et al., 1978):



466 Behaviour of trace metals during sulphide leaching is dependent on their solubilities in the high  
467 temperature fluids, which are controlled by temperature, pressure, pH,  $fO_2$ ,  $fS_2$  and fluid composition (e.g.  
468 Cl<sup>-</sup>; Crerar et al., 1978; Hannington and Scott, 1989; Seewald and Seyfried, 1990; Hemley et al., 1992;  
469 Gibert et al., 1998).

### 470 5.2.3 Trace metal behaviour during metasomatism and leaching

471 Major and trace metal concentrations vary considerably during metasomatism and leaching of the  
472 magmatic sulphides (Table 1 and 2). Using median values, metasomatised pyrite shows higher  
473 concentrations of Co (+7 %), Cu (+2528%), Zn (+334 %), As (+265 %), Se (+91 %), Ag (+469 %), Te  
474 (+89 %), Au (+115 %) and Pb (+1466 %) relative to pyrrhotite (Table 1 and 2). Only Ni shows lower  
475 concentrations (-21 %; Table 1 and 2). Although the moderate temperature hydrothermal fluids that drive  
476 sulphide metasomatism probably contain some trace metals, it is highly unlikely that they systematically  
477 enrich the metasomatised sulphides in trace metals to such extents. Increases in trace metal concentrations  
478 in metasomatised sulphides are most likely due to residual enrichment during successive sulphide volume  
479 loss and local metal diffusion between mineral grains. During metasomatism, diffusion of Fe from  
480 pyrrhotite forms the magnetite rim (Equation 2). This process leads to sulphide volume loss and  
481 recrystallisation of pyrrhotite to pyrite with a corresponding S increase of +34.3 % and Fe decrease of -  
482 21.4 % of the sulphide assemblage (Table 1 and 2, Fig. 10). The residual metal enrichment caused by the

483 conversion of pyrrhotite to pyrite, however, only partially explains the large enrichments of trace metals  
484 observed in metasomatised pyrite. Additional concentration of trace metals occurs during leaching where  
485 preferential mobilisation of Fe and S into the hydrothermal fluids leads to further sulphide loss (equation  
486 [3]). The behaviour of Se clearly highlights this two-step process as demonstrated by the variation in the  
487 S/Se ratio during metasomatism and leaching (Fig. 11A). During metasomatism the S/Se is constant as S  
488 and Se are equally concentrated during Fe diffusion to magnetite, but as leaching occurs the S/Se  
489 decreases as S is preferentially mobilised by the hydrothermal fluids along with Fe (Fig. 11A). The high  
490 concentrations of Co, Zn, Ag, Se, Te and Au in pyrite are most likely explained by the combined effects of  
491 metasomatism and leaching (Fig. 11). Although Sb is below detection limit in pyrrhotite the range of  
492 concentrations in metasomatised pyrites suggest similar behaviour (Table 2). The extreme concentration  
493 increases of Cu (+ 2528%) and Pb (+1466 %), and the decrease in Ni concentrations (-21 %) in  
494 metasomatised pyrites cannot be explained solely by sulphide volume loss and we suggest that diffusion  
495 of these elements between the different sulphide phases present in the assemblage occurs. Nickel is likely  
496 to diffuse into the millerite from both pyrrhotite and pentlandite (Table 1 and 2; Fig. 10) whereas Cu is  
497 likely to diffuse from the chalcopyrite to the pyrite during metasomatism, and Pb from surrounding  
498 silicates (Table 1 and 2; Fig. 10). Changes in the trace metal distribution between magmatic and  
499 metasomatised sulphides suggests that localised grain-scale metal mobility occurs during  
500 metasomatisation (Fig 10), but bulk mobilisation of metals from the rock does not occur until the onset of  
501 sulphide leaching. Ultimately, if sulphide leaching goes to completion then all trace elements hosted in  
502 metasomatised sulphides will be mobilised into the high temperature fluids (Fig. 10).

#### 503 *5.2.4 Oxide alteration by hydrothermal fluid circulation*

504 Although oxide minerals show some evidence of alteration they are less affected by high temperature fluid  
505 circulation than sulphide minerals. Laverne et al. (2006) report breakdown of titanomagnetite to  
506 hydroschlormite during low temperature alteration. Additionally Alt et al. (2010) reports increasing  
507 titanomagnetite alteration below 660 mbsf. The Si+Ca/Ti ratio increases with depth indicating greater

508 alteration (Fig. 8B). No systematic variation of metal concentrations with the Si+Ca/Ti ratio can be  
509 observed in oxides indicating that, the contrary to sulphides, alteration of these minerals in Hole 1256D  
510 does not release metals (Fig. 8E, F, G and H).

### 511 5.3 Mineral precipitation during high temperature fluid circulation

512 In the upper sheeted dyke and transitional zone, on-axis mixing of ascending high temperature fluids with  
513 the low temperature fluids prevailing in the volcanic section triggers significant sulphide precipitation  
514 forming a 2.8 m mineralised breccia at 1028 mbsf (Teagle et al., 2006; Alt et al., 2010; Fig. 7). Fluid  
515 inclusion microthermometry suggests that fluid mixing and sulphide precipitation occurred at  
516 temperatures between 320 °C and 450 °C (Alt et al., 2010) and at high fluid-rock ratios (Harris et al.,  
517 2015). Such temperatures are consistent with the minimum hydrothermal fluid temperature of  $378 \pm 24$  °C  
518 estimated from Fe/Zn in sphalerite (Keith et al., 2014). The mineralogical assemblages and trace metal  
519 concentrations in pyrite vary between the transitional zone and the upper sheeted dykes suggesting  
520 different mineralisation in these two areas. Temperature controlled metal zonation is common in VMS  
521 deposits: precipitation of chalcopyrite and Ni-bearing pyrite occur at higher temperatures than sphalerite  
522 and As, Sb, Au and Pb-bearing pyrite (Hannington and Scott, 1989; Metz and Trefry, 2000; Genna and  
523 Gaboury, 2015; Keith et al., 2016; Monecke et al., 2016) and a similar zonation is observed in Hole  
524 1256D between the upper sheeted dyke section and the transitional zone (Table 3). The Se/As ratio tends  
525 to be lower in pyrites from the transitional zone relative to those in the upper sheeted dyke section (Fig. 4)  
526 which can be attributed to an increased input of low temperature fluids leading to temperature decrease  
527 (Genna and Gaboury, 2015), with possible As contamination from these fluids. The low Ni content and  
528 variable Se/As ratios of hydrothermal pyrites are distinctly different from metasomatised pyrites  
529 highlighting the different origin between the two sulphide groups. Although high temperature  
530 hydrothermal sulphides have significant trace metal concentrations (Table 3) their impact on the bulk  
531 fluxes of metals from the lower sheeted dykes is minimal (Nesbitt et al., 1987; Patten et al., 2015).  
532 Hydrothermal magnetites are also interpreted to be generated from the circulation of high temperature

533 fluids. Their composition, characterised by low Ti content (Fig. 8C), and texture confirm their  
534 hydrothermal origin and suggests that they formed at similar temperatures (~300-500 °C) to the high  
535 temperature sulphides (Fig. 8D).

#### 536 5.4 Sulphide formation during low temperature fluid circulation

537 In the volcanic section pyrite fronts locally occur on the margins of significantly oxidised zones referred  
538 to as alteration halos (e.g. Andrews, 1979; Alt and Shanks, 2001; Fig. 7). Andrews (1979) proposes that  
539 the pyrite fronts form from leaching of magmatic sulfides from the alteration halos by low temperature  
540 fluids. The presence of marcasite indicates acidic solution (<5 pH) and partly oxidised soluble species  
541 such as thiosulphate (Murowchick and Barnes, 1986). Alt and Shanks (2011), however, argue that there is  
542 insufficient sulfide in the alteration halos to account for that in the pyrite fronts, and additional S is  
543 required from microbial seawater sulfate reduction. The formation of alteration halos and pyrite fronts  
544 occurs at the latest stages of fluid circulation during off-axis low temperature alteration (Alt et al., 2010;  
545 Harris et al., 2015) at higher water-rock ratio than the background alteration (Teagle et al., 2006; Gao et  
546 al., 2012). Significant As enrichment occurs in the alteration halos and the pyrite fronts occurs (Patten et  
547 al., 2015) causing the low Se/As ratio due to As input from the low temperature fluids (Fig. 4). In sulphide-  
548 free samples, which also show As enrichment (Patten et al., 2015), As is likely to be hosted in Fe-  
549 oxyhydroxide minerals (Wolthers et al., 1998; Smedley and Kinniburgh, 2002; Han et al., 2012). High Ni  
550 concentrations in pyrite fronts (Table 4; Fig. 4) can be explained by Ni incorporation in pyrite at low  
551 temperature and under anoxic conditions (Morse and Luther, 1999). Along with As, significant whole-  
552 rock enrichment of Sb occurs in the volcanic section (Patten et al., 2015), but this enrichment is associated  
553 with secondary silicates rather than the low temperature sulphides (Fig. 12; Jochum and Verna, 1996).

554 In the sheeted dyke and plutonic complexes the patchy sulphides share numerous characteristics with low  
555 temperature sulphides although they are not associated with any specific alteration. They have similar  
556 mineralogical assemblages of pyrite, minor marcasite and trace chalcopyrite (Fig. 2), and replacement



557 textures (Fig. 5). Pyrites from the patchy sulphide group are characterised by low Se/As ratio and high Ni  
558 and falls within the field of the low temperature pyrite (Fig. 4). These similarities between the two  
559 sulphide groups suggest similar paragenesis, implying that the patchy sulphide group formed during late-  
560 stage off axis low temperature fluid circulation in the sheeted dyke and plutonic complexes. Patchy  
561 sulphides have moderate trace metal concentrations (Table 5), suggesting metal remobilisation from the  
562 surrounding rocks during off axis fluid circulation.

### 563 5.5 Mass balance

564 A mass balance calculation has been carried out to investigate the degree to which sulphide minerals  
565 control the whole-rock metal concentrations in Hole 1256D samples. The mass balance has been carried  
566 out only on samples containing one dominant sulphide group. Hypothetical whole rock compositions were  
567 calculated for each sample assuming that all the metals present in the sample are solely hosted by  
568 sulphides. The total mass of sulfide is calculated using whole rock S concentration (Patten et al., 2015).  
569 The total mass of sulfide and the compositions of the different phases in the sulphide group are then used  
570 to calculate the hypothetical whole rock compositions, which are then compared to the real whole rock  
571 compositions to give the % of each metal that is hosted as sulphide (Appendix B). Several significant  
572 errors are associated with the mass balance calculation, such as the problem of comparing observations  
573 from a 2D thin section surface with analyses of a 3D sample volume, and error propagations are discussed  
574 in Appendix B. Because of these errors, the mass balance is considered to be semi-quantitative but the  
575 results still enable evaluation of the extent to which the sulphides control the whole rock metals budget in  
576 Hole 1256D. The results are classified into >50 %, 10-50 % and <10 % sulphide hosting.

577 Although based on a low number of samples (n=2), the mass balance for samples hosting magmatic  
578 sulphides suggests that Cu, Se, Te and Au are strongly controlled (>50 %) by sulphides (Fig. 12) whereas  
579 Zn, As, Sb and Pb are mostly controlled by silicate and oxide minerals (sulphide <10 %, Fig. 12; Gurney  
580 and Ahrens, 1973; Doe et al., 1994; Alt, 1995). Similarly to magmatic sulphides, metasomatised sulphides

581 strongly control Cu, Se, Te and Au (>50 %), partly control As and Pb (10-50 %) and weakly control Zn  
582 and Sb (<10 %; Fig. 12). The similarity in mass balance between the two sulphide groups suggests that  
583 during metasomatism, metals undergo local grain-scale diffusion but are not extensively mobilised by the  
584 moderate temperature hydrothermal fluids (250-350 °C; Fig. 10). Samples hosting high temperature  
585 sulphides, located in the upper sheeted dykes and the transitional zone, are the only ones where sulphides  
586 control most of the whole rock metal budget (Fig. 12). In these samples pyrite hosts most of the Se, Te,  
587 Au and Pb whereas Cu is controlled by chalcopyrite and Zn by sphalerite (Fig. 12). Arsenic and Sb are  
588 only partly controlled by sulphides in these samples. In the volcanic section, mass balance calculations of  
589 samples hosting low temperature sulphides suggest that Cu and As are strongly controlled and Se partly  
590 controlled by sulphides, whereas Zn, Sb and Pb are poorly controlled by sulphides (Fig. 12). Finally, the  
591 mass balance calculation of the patchy sulphide group suggests that only Cu is strongly hosted by these  
592 sulphides, whereas As, Se, Sb, Te and Au are partly controlled and Zn and Pb are poorly controlled by  
593 sulphides (Fig. 12).

## 594 5.6 Mineralogical controls on metal mobilisation

595 In the oceanic crust, significant mobilisation of metals occurs from the lower sheeted dyke section during  
596 high temperature hydrothermal fluid circulation (e.g. Hole 504B: Alt et al., 1996; Nesbitt et al., 1987;  
597 Bach et al., 2003; Hole 1256D: Teagle et al., 2006; Alt et al., 2010; Patten et al., 2015; Pito deep: Heft et  
598 al., 2008). In Hole 1256D, Au, As, Se, Cu, Zn and Pb are variably depleted relative to primary crustal  
599 compositions (-46 %, -27 %, -27 %, -10 %, -8 % and -44 % respectively; Patten et al., 2015). As  
600 magmatic and metasomatised sulphides have strong control over Cu, Se and Au budgets (Fig. 12) at the  
601 time of high temperature hydrothermal alteration, sulphide leaching is clearly a key mineral reaction for  
602 mobilisation of these metals. This is strongly supported by the large whole-rock metal depletions in  
603 samples hosting metasomatised and leached sulphide minerals (i.e in samples below 1061 mbsf; Fig. 13).  
604 Additionally, variability in whole rock metal depletions can be partly explained by the variable behaviour  
605 of the metals during sulphide leaching. Copper, Se and Au are strongly hosted by magmatic and

606 metasomatised sulphides but Cu shows lower whole rock depletions (-10 %) than Se and Au (-27 % and -  
607 46 % respectively) in the sheeted dyke and plutonic complexes. Selenium and Au are principally hosted in  
608 pyrite (Fig. 12) which is reactive to sulphide leaching (Fig. 9) explaining the higher whole rock depletions,  
609 whereas Cu is hosted in chalcopyrite (Fig. 12) which is more resistant to leaching (Fig. 9A, B and 10). The  
610 depletions of As, Zn and Pb in Hole 1256D (-27 %, -8% and -44 % respectively) are neither associated  
611 with sulphide leaching (Fig. 12) or oxide alteration (Fig. 8F) but most likely with alteration of silicate  
612 minerals (Doe et al., 1994). Oxide minerals host significant Zn and some Sb (Table 6; Doe et al., 1994;  
613 Gurney and Ahrens, 1973) and increased alteration of these minerals could lead to greater depletion of  
614 these metals. Antimony and Te show no significant whole rock variation in Hole 1256D during high  
615 temperature hydrothermal alteration (Patten et al., 2015) and more intensive alteration would be required  
616 to mobilise these elements (Fig. 12).

617 The different metal behaviour during sulphide leaching, as summarised in Figure 10, and the control of  
618 oxide and silicate mineral alteration on mobility of some elements indicates that peak mobilisation for  
619 each metal occurs at different stages, and under different conditions. Trace metals hosted in pyrite are  
620 likely to be released during initiation of leaching, whereas Cu and trace metals in chalcopyrite would be  
621 released at higher temperature during sulphide leaching leading to significant fractionation of metals in the  
622 hydrothermal fluid. This process may affect the relative metal contents of the high temperature  
623 hydrothermal fluids through time.

## 624 5.7 Metal-rich hydrothermal fluids and formation of VMS deposits

625 At ODP Hole 1256D, the masses of metals mobilised during thih temperature hydrothermal alteration of  
626 the sheeted dyke and plutonic complexes is sufficient to form modern-day low tonnage mafic type VMS  
627 deposits (Patten et al., 2015) such as those found at the EPR 13 °N (Fouquet et al., 1988) or the TAG  
628 VMS deposit (e.g. Hannington et al., 1998). Different reactions provide different metals to the fluids with  
629 sulphide leaching providing the Cu, Se and Au, and silicate reactions providing the As, Sb, Zn and Pb

630 (Figs. 12 and 13). Lack of known VMS deposits in the vicinity of Hole 1256D, however, highlights the  
631 importance of metal trapping mechanisms for formation of VMS deposits. The preservation of both  
632 magmatic and metasomatised (i.e. not fully leached) sulphides in Hole 1256D crust implies that sulphide  
633 leaching has not gone to completion. The preservation of these sulphides in Hole 1256D is most likely due  
634 to the relative low water-rock ratio in the lower part of Hole 1256D ( $w/r < 1$ ; Gao et al., 2012) which may  
635 be a feature of fast-spreading ridge oceanic crust compared to other tectonic settings such as ophiolites  
636 which are more extensively altered (Alt and Teagle; 2000). More extensive hydrothermal alteration would  
637 lead to total sulphide leaching and an increase in the quantity of metal mobilised from the oceanic crust.  
638 Epidosite zones observed in ophiolitic systems and interpreted as the source areas for metals enriched in  
639 associated VMS deposits (Jowitt et al., 2012) are likely to represent this advanced stage of sulphide  
640 leaching.

## 641 **6. Conclusions**

642 In-situ analyses of the sulphide and oxide minerals present in the ODP Hole 1256D drill core enables  
643 determination of their paragenesis during the evolution of the oceanic crust and evaluation of the extent to  
644 which extend they control bulk metal behaviour in the crust (Fig. 7). The main outcomes of this study are:

- 645 • Five main groups of sulphides occur in ODP Hole 1256D: magmatic, metasomatised, high  
646 temperature, low temperature and patchy sulphides. Three groups of oxides occur: magmatic skeletal  
647 titanomagnetite, magmatic coarse magnetite-ilmenite and hydrothermal related magnetite.
- 648 • Initiation of hydrothermal fluid circulation in the lower sheeted dyke section and the plutonic  
649 complex at moderate temperature ( $\sim 300$  °C) and relatively oxidising conditions leads to  
650 metasomatism of the magmatic sulphides. Recrystallisation of pyrrhotite, chalcopyrite and  
651 pentlandite to form secondary pyrite, chalcopyrite, millerite and magnetite drives significant diffusion  
652 of metals between the sulphide phases and the surrounding silicates and oxides, although little bulk  
653 metal mobilisation to the circulating hydrothermal fluids occurs at this stage.

- 654 • Increase in hydrothermal alteration temperature (>350 °C) under reduced conditions leads to  
655 progressive leaching and dissolution of the metasomatised sulphides. Semi-quantitative mass balance  
656 calculations suggest that sulphide leaching is responsible for the release of most Cu, Se and Au  
657 during the hydrothermal alteration of the lower sheeted dyke section and the plutonic complex. The  
658 different behaviour of trace metals during sulphide leaching suggests that they have different timing  
659 for peak mobilisation. Chalcopyrite is less reactive to leaching than pyrite implying that Cu and the  
660 metals hosted in chalcopyrite are released during the highest temperature alteration conditions. The  
661 mobilisation of As, Zn and Pb is controlled by silicate reactions rather than sulfide leaching.
- 662 • A fraction of the metal mobilised by the high temperature hydrothermal fluids are trapped within  
663 mineralised veins and breccias as high temperature sulphides in the upper sheeted dyke section and  
664 the transitional zone during mixing with low temperature fluids. The high temperature pyrite,  
665 sphalerite and chalcopyrite have high and variable trace metal concentrations which highlight  
666 different behaviour during sulphide precipitation and metal trapping.
- 667 • Late off-axis circulation of low temperature fluids (<150 °C) in the oceanic crust leads to the local  
668 formation of pyrite fronts in the volcanic section and to patchy sulphides in the sheeted dyke and  
669 plutonic complexes. High As enrichments in the pyrite fronts suggests significant uptake from the  
670 circulating low temperature fluids in the volcanic section whereas the moderate trace metal  
671 concentrations in the patchy sulphides suggests late metal remobilisation within the sheeted dyke and  
672 plutonic complexes.
- 673 • The hydrothermal fluids that form during alteration of the lower sheeted dyke units in Hole 1256D  
674 are sufficiently metal-rich to form low tonnage VMS deposits on the seafloor. The metal enrichments  
675 in the fluids are driven by different mineral reactions with Cu, Se and Au mobilised by sulfide  
676 leaching, and As, Zn and Pb mobilised by silicate reactions. An increase in alteration would likely  
677 lead to total sulphide leaching and metal release leading to higher metal budget of the hydrothermal  
678 fluids increasing the likeliness of VMS deposits formation.

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## 946 **9. Figure captions**

947 Figure 1: Lithostratigraphy at IODP 1256D with Mg# and sulphide proportions in veins. Major  
948 secondary minerals and temperature are also shown. Modified from Teagle et al. (2006) and Alt et al.  
949 (2010).

950 Figure 2. Sulphide distribution in IODP Hole1256D. The scale corresponds to sulphide proportions with  
951 1=trace, 2=discrete, 3=common and 4=major. Yellow boxes are samples with sulphides analysed by  
952 EPMA and LA-ICP-MS. Meta. stands for metasomatised sulphides, high T. stands for high temperature  
953 sulphides and low T. stands for low temperature sulphides.

954 Figure 3. Reflected light photomicrographs of the magmatic and metasomatised sulphides. Magmatic  
955 sulphides: A) Sulphide blebs preserved in the volcanic section composed of pyrrhotite (Po) and  
956 chalcopyrite (Cpy). The Po represents the MSS and the Cpy the ISS, see discussion for details. B) Base  
957 metal sulphides pyrrhotite-chalcopyrite-pentlandite (Pn) from the plutonic section. Pentlandite is present  
958 as lamellae exsolution within the Po which is specific of MSS. Metasomatised sulphides: C) Spherical  
959 blebs composed of Py and magnetite (Mag). D) and E) Assemblage from the plutonic section composed of  
960 Py-Cpy-Mag; note that the Mag preferentially form a rim. F) Assemblage of Py-Cpy and millerite (Mi)  
961 from the plutonic section.

962 Figure 4. Variation of the Ni content versus the Se/As of the pyrites from the different sulphide groups.  
963 Metasomatised pyrites are characterised by high Se/As and high Ni, high temperature pyrites by variable  
964 Se/As and low Ni and low temperature and patchy pyrites by low Se/As and high Ni. See text for details.

965 Figure 5. Reflected light photomicrographs of the high temperature, low temperature and patchy  
966 sulphides. High temperature sulphides: A) pyrite (Py) aggregate in quartz-chlorite vein from the upper  
967 sheeted dyke section. The core is formed of small Py grains whereas the rim is formed by large euhedral  
968 grains. Presence of chalcopyrite (Cpy) and sphalerite (Sp) with the Cpy disease. B) Large euhedral Py  
969 grains in the transitional zone. C) Py and Sp from the mineralised breccia in the transitional zone. Low  
970 temperature sulphides: D) Pyrite front in border of an alteration halo. E) Details of the pyrite front with Py  
971 and marcasite (Mrc) filling a crack. F) Euhedral Py filling a vesicle. G) Py and trace Cpy in saponite  
972 dominated vein. Patchy sulphides: H) and I) Pyrite replacing the silicate matrix and partly the silicate  
973 minerals resulting in a patchy aspect.

974 Figure 6. Reflected light photomicrographs of the oxide population in Hole 1256D. A) Skeletal  
975 titanomagnetite (T-Mag) in the volcanic section (group 1). M.S.= metasomatised sulphides B) Coarse  
976 grain magmatic oxide composed of magnetite (Mag) with ilmenite (Ilm) exsolution (group 2); note the  
977 trellis pattern and grain texture. Po=pyrrhotite; Cpy=chalcopyrite. C) Hydrothermal related oxide  
978 corresponding to magnetite replacing a silicate phase (group 3). D) Hydrothermal related oxide  
979 corresponding to magnetite replacing a sulphide phase (group 3); see discussion. Py=pyrite.

980 Figure 7. Schematic diagram illustrating the paragenesis of the sulphide population at IODP Hole  
981 1256D. 1) Magmatic sulphides forming spherical sulphide droplets in the volcanic section and large  
982 sulphide aggregates in the plutonic complex. 2) Metasomatised sulphides showing different degrees of  
983 sulphide leaching from a) pristine to b) extensively leached. 3) High temperature hydrothermal sulphides  
984 precipitated from rising hydrothermal fluids in a) the upper sheeted dykes and b) the mineralised breccia  
985 at 1028 mbsf. 4) Low temperature sulphides occurring a) in background altered samples due to low

986 temperature fluid circulation and b) on borders of alteration halos forming pyrite fronts. 5) Patchy  
987 sulphides formed from deep low temperature fluid circulation in the lower part of Hole 1256D.  
988 Po=pyrrhotite, Pn=pentlandite, Cpy= chalcopyrite, Py=pyrite, Mag= magnetite, Mi=millerite.

989 Figure 8. Oxide minerals trace elements. A) Ternary FeO<sub>2</sub>, Fe<sub>2</sub>O<sub>3</sub> and TiO<sub>2</sub> diagram. Mag=magnetite,  
990 Ilm=ilmenite. B) Content of Si and Ca in oxide minerals versus depth. Silica and Ca contents are used as a  
991 proxy for the degree of alteration. C) Discrimination diagram for magmatic or hydrothermal origin of  
992 magnetite and titanomagnetite with TiO<sub>2</sub><20 wt.%; modified from Dare et al. (2014). Note that skeletal  
993 titanomagnetite and coarse magnetite plot in the magmatic field whereas hydrothermal magnetite plots in  
994 the hydrothermal field. D) Discrimination diagram showing and titanomagnetite with TiO<sub>2</sub><20 wt.%  
995 estimated formation temperature; modified from Nadoll et al. (2014). E) to H) Trace metal content in  
996 oxide minerals versus the degree of alteration. See text for details.

997 Figure 9. Reflected light photomicrographs of metasomatised sulphides showing different degree of  
998 sulphide leaching. A) Shadow of magmatic sulphide blebs in the volcanic section with a magnetite (Mag)  
999 rim and residual chalcopyrite (Cpy) inside. B) Porous pyrite (Py) surrounded by Mag and trace Cpy in the  
1000 surroundings. C) and D) Trace leached sulphides with the Mag remaining. B), C) and D) are from the  
1001 plutonic complex.

1002 Figure 10. Sulphide evolution and metal content variation during metasomatism and leaching. A)  
1003 Magmatic assemblage characterised by monosulphide solution (MSS) and intermediate solid solution  
1004 (ISS). During subsolidus reaction MSS exsolve to pyrrhotite (Po) and pentlandite (Pn), and ISS exsolve to  
1005 chalcopyrite (Cpy). M\*=trace metals hosted in Cpy, their behaviour during sulphide evolution is not  
1006 determined in this study. B) Metal diffusion during metasomatism. Py=pyrite, Mag=magnetite,  
1007 Mi=millerite. C) and D) Metal mobilisation during partial and extensive sulphide leaching respectively.  
1008 See text for details.

1009 Figure 11. Trace metal variations in pyrrhotite and metasomatised pyrite during metasomatism and  
1010 leaching. Selenium is used as proxy to highlight trace metal behaviours both during metasomatism and  
1011 leaching. Increase concentrations of Se, Te, As, Ag, Au and Zn in metasomatised pyrites are due to  
1012 residual enrichment by progressive sulphide volume loss first during metasomatism and second by  
1013 sulphide leaching. See text for details.

1014 Figure 12. Pie charts illustrating the semi-quantitative results of the mass balance calculations. Charts  
1015 show the proportion to which metals are hosted by different sulfides in %. Magmatic sulphide assemblage  
1016 is shown as a combination of monosulphide solid solution (MSS) and intermediate solid solution (ISS)  
1017 because of the too few data available for each sulphide phases in this group. Mass balance calculation  
1018 cannot be determined for Te and Au in the low temperature sulphides and for Sb in the magmatic  
1019 sulphides. See Appendix A.4 and B for mass balance calculation results and uncertainties.

1020 Figure 13. Mass variations in the sheeted dyke and plutonic complexes and the associated sulphide  
1021 population. Mass variations are from whole rock data (Patten et al., 2015). The main sulphide group  
1022 observed in thin section is shown. No thin sections are available for several samples. Samples hosting  
1023 metasomatised sulphides show significant metal depletion. See text for details. SD = sheeted dyke  
1024 complex; PC = plutonic complex.

Table 1. Major and trace metals composition of magmatic sulphides

		S	Fe	Co	Ni	Cu	Zn	As	Se	Ag	Pd	Sb	Te	Ir	Pt	Au	Hg	Pb
		wt.%	wt.%	wt.%	wt.%	wt.%	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm
		EPMA	EPMA	EPMA	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS
Po-Cpy intergrowth EPMA n=11 LA-ICP-MS n=8	mean	39.24	58.0	0.07	0.21	0.58	1054	5.12	56.00	3.44	<D.L.	0.24	2.23	<D.L.	<D.L.	0.066	<D.L.	409.5
	s.d.	0.7	1.9	0.02	0.02	0.47	987	1.34	15.49	0.93		0.12	0.26			0.049		680.6
	median	39.55	59.0	0.08	0.21	0.69	855	4.90	54.50	3.51		0.24	2.20			0.066		35.1
	range min	37.6	55.3	0.03	0.18	7.9E-03	47	3.90	36.00	2.12		0.09	2.00			0.031		8.1
	range max	39.9	59.6	0.10	0.24	1.28	2900	8.00	89.00	4.80		0.41	2.50			0.100		2100.0
Pyrrhotite EPMA n=17 LA-ICP-MS n=18	mean	39.57	58.4	0.14	2.18	0.02	9.9	0.95	45.52	0.21	<D.L.	<D.L.	1.25	<D.L.	<D.L.	0.023	<D.L.	0.49
	s.d.	0.3	0.8	0.27	1.49	0.04	18.1	0.11	8.38	0.11			0.34			0.019		0.22
	median	39.60	58.7	0.07	1.67	1.5E-03	3.8	0.95	44.60	0.17			1.20			0.019		0.57
	range min	38.5	56.0	0.03	1.15	1.9E-04	1.9	0.87	32.80	0.10			0.67			0.005		0.10
	range max	40.0	59.2	1.20	7.50	0.14	78.0	1.02	58.80	0.42			1.81			0.053		0.82
Chalcopyrite EPMA n=14 LA-ICP-MS n=9	mean	35.0	31.7	0.28	0.80	31.6	1169	<D.L.	49.64	5.98	<D.L.	<D.L.	2.82	<D.L.	<D.L.	0.027	<D.L.	3.29
	s.d.	0.2	2.0	0.04	0.99	3.8	1748		18.71	3.38			1.36			0.017		3.64
	median	35.1	30.9	0.30	0.68	33.4	160		46.80	6.76			2.80			0.025		1.30
	range min	34.6	29.8	0.24	0.01	23.0	2.4		31.50	0.37			0.85			0.012		0.28
	range max	35.3	36.8	0.30	3.03	33.7	5000		89.00	10.20			4.40			0.045		10.60
Pentlandite (n=6)	mean	42.0	19.8	16.5	21.5	<D.L.	<D.L.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.
	s.d.	0.4	6.4	3.9	4.0													
	range min	41.5	11.9	9.5	16.1													
	range max	42.3	27.6	21.0	28.5													

The range of composition is reported in the text if the standard deviation is higher than the average; po=pyrrhotite; cpy=chalcopyrite; s.d.=standard deviation; D.L.=detection limit; N.D.= not determined

Table 2. Major and trace metals composition of metasomatised sulphides

		S	Fe	Co	Ni	Cu	Zn	As	Se	Ag	Pd	Sb	Te	Ir	Pt	Au	Hg	Pb	
		wt.%	wt.%	wt.%	wt.%	wt.%	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	
		EPMA	EPMA	EPMA	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	
Pyrite EPMA n=58 LA-ICP-MS n=82	mean	53.0	45.8	0.60	0.51	0.16	76	6.58	108	2.34	4.39	0.77	4.87	<D.L.	1.65	0.07	3.60	16.99	
	s.d.	0.9	1.3	1.32	0.67	0.25	299	9.67	79	5.35	8.48	1.75	8.52		1.62	0.09	2.05	19.38	
	median	53.3	46.2	0.08	0.17	0.04	17	3.45	86	0.97	1.16	0.25	2.27		1.16	0.04	3.10	8.93	
	range	min	50.2	40.5	0.01	1.8E-03	3.5E-04	1.2	0.42	4	0.08	0.21	0.03	0.50		0.28	0.002	1.40	1.60
		max	54.4	46.9	6.00	2.50	0.91	1870	48.40	470	38.00	25.00	11.80	51.00		4.00	0.31	6.70	94.00
Chalcopyrite EPMA n=35 LA-ICP-MS n=20	mean	35.5	31.2	0.106	0.75	32.1	79.1	3.35	76	9.78	<D.L.	0.55	3.55	<D.L.	<D.L.	0.13	11.80	11.48	
	s.d.	2.4	1.9	0.084	1.25	3.6	138	3.64	36	7.37		1.00	3.72			0.18	5.84	8.27	
	median	35.1	30.7	0.063	0.14	33.2	27	1.86	72	7.40		0.15	2.03			0.07	12.20	9.85	
	range	min	34.5	29.6	0.019	3.5E-03	14.2	7.3	0.42	12	1.46		0.02	0.93			0.01	6.40	1.10
		max	48.9	40.6	0.280	4.10	34.3	500	14.40	150	29.50		3.71	14.00			0.66	18.00	37.80
Millerite (n=6)	mean	38.2	4.0	6.4	51.3	<D.L.	<D.L.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	N.D.	
	s.d.	3.5	1.9	7.6	12.3														
	range	min	35.2	2.5	0.02	33.4													
max		42.5	7.6	14.8	61.6														

s.d.=standard deviation; D.L.=detection limit; N.D.= not determined

Table 3. Major and trace metals composition of high temperature sulphides

		S	Fe	Co	Ni	Cu	Zn	As	Se	Ag	Pd	Sb	Te	Ir	Pt	Au	Hg	Pb	
		wt.%	wt.%	wt.%	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	
		EPMA	EPMA	EPMA	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	
All pyrites EPMA n=62 LA-ICP-MS n=105	mean	53.8	46.6	0.11	124	44.1	46.0	23.94	66.76	2.45	0.17	0.61	1.63	<D.L.	<D.L.	0.12	1.89	69.72	
	s.d.	0.5	0.5	0.09	205	60.8	187.6	19.93	170.07	2.40	0.10	0.78	2.22			0.10	1.76	97.73	
	median	53.9	46.8	0.09	30.0	25.9	5.9	23.70	31.50	2.39	0.13	0.33	1.09			0.12	1.24	14.60	
	range	min	52.3	44.5	0.01	0.95	0.7	0.6	0.67	1.27	0.09	0.10	0.04	0.46			0.02	0.64	0.50
		max	54.6	47.8	0.32	1225	390.0	1600.0	93.00	1243.00	12.10	0.29	4.07	14.20			0.58	7.60	408.00
Transitional zone pyrite EPMA n=28 LA-ICP-MS n=36	mean	53.7	46.8	0.10	50	34.7	52.0	31.11	88.50	2.49	0.11	0.67	2.45	<D.L.	<D.L.	0.12	1.71	104.63	
	s.d.	0.5	0.3	0.11	111	28.7	222.9	18.68	231.45	2.18	0.02	0.80	4.14			0.10	1.71	107.52	
	median	53.8	46.8	0.05	19.3	29.6	5.7	32.15	34.15	2.57	0.11	0.47	1.17			0.12	1.02	68.80	
	range	min	52.5	46.3	0.01	0.95	0.7	0.6	0.67	1.27	0.09	0.10	0.04	0.51			0.02	0.64	0.50
		max	54.4	47.8	0.32	578	136.0	1600.0	93.00	1243.00	9.59	0.13	4.07	14.20			0.58	7.60	408.00
Upper sheeted dyke pyrite EPMA n=34 LA-ICP-MS n=69	mean	53.9	46.5	0.12	236	60.4	35.0	6.15	42.41	2.38	<D.L.	0.07	1.33	<D.L.	<D.L.	<D.L.	3.45	8.13	
	s.d.	0.5	0.6	0.04	260	91.9	94.6	8.59	35.13	2.81		0.03	0.72				1.91	11.05	
	median	53.9	46.7	0.11	174.5	21.3	6.5	3.44	30.80	1.58		0.06	1.09				3.45	4.22	
	range	min	52.3	44.5	0.06	11.80	0.9	0.9	0.69	12.90	0.12		0.04	0.46				2.10	0.69
		max	54.6	47.1	0.18	1225	390.0	419.0	41.40	187.00	12.10		0.13	2.88				4.80	52.60
Chalcopyrite EPMA n=29 LA-ICP-MS n=12	mean	35.0	30.2	<D.L.	237	33.5	1701	1.73	160.25	5.95	<D.L.	0.31	1.04	<D.L.	<D.L.	<D.L.	<D.L.	11.49	
	s.d.	0.4	0.8		203	0.5	3755	0.89	197.92	4.56		0.28	0.44					8.99	
	median	35.1	30.4		197	33.7	714	1.90	67.90	6.20		0.17	1.10					8.97	
	range	min	34.1	28.5		49.0	32.5	2.2	0.37	11.00	0.37		0.12	0.49					1.53
		max	35.5	31.3		591	34.0	12900	2.50	566.00	14.50		0.80	1.50					28.70
Sphalerite EPMA n=19 LA-ICP-MS n=10	mean	33.7	8.5	0.03	39.1	6683	55.2	12.24	17.42	3.59	<D.L.	22.42	<D.L.	<D.L.	<D.L.	0.02	95.42	65.45	
	s.d.	0.3	3.1	0.01	35.6	12071	2.9	8.38	11.13	3.49		17.47				0.01	113.16	58.45	
	median	33.6	9.8	0.03	30.3	1650	54.2	11.50	14.40	1.73		20.90				0.02	55.50	52.55	
	range	min	33.2	2.1	0.02	6.0	126	51.9	0.57	4.30	1.46		0.06			0.01	10.50	8.49	
		max	34.2	12.3	0.06	110.0	37900	61.3	23.30	29.50	12.00		52.90			0.03	310.00	190.00	

s.d.=standard deviation; D.L.=detection limit; N.D.= not determined



Table 4. Major and trace metlas composition of low temperature sulphides

		S	Fe	Co	Ni	Cu	Zn	As	Se	Ag	Pd	Sb	Te	Ir	Pt	Au	Hg	Pb
		wt.%	wt.%	wt.%	wt.%	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm
		EPMA	EPMA	EPMA	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS	LA-ICP- MS
Pyrite/Marcasite	mean	52.7	46.3	0.06	0.20	9.6	3.3	102	37	0.20	<D.L.	0.38	<D.L.	<D.L.	<D.L.	<D.L.	<D.L.	1.37
EPMA n=28	s.d.	0.6	0.5	0.08	0.30	11.2	2.8	67	14	0.10		0.24						1.26
LA-ICP-MS n=28	median	52.9	46.5	0.04	0.06	6.0	3.3	98	38	0.17		0.33						0.9
	range																	
	min	51.1	45.1	0.01	0.01	1.0	0.7	9.1	18	0.10		0.09						0.12
	max	53.6	47.0	0.26	1.07	57.0	14	271	61	0.42		1.03						4

s.d.=standard deviation; D.L.=detection limit; N.D.= not determined

Table 5. Major and trace metals composition of patchy sulphides

		S	Fe	Co	Ni	Cu	Zn	As	Se	Ag	Pd	Sb	Te	Ir	Pt	Au	Hg	Pb
		wt.%	wt.%	wt.%	wt.%	wt.%	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm
		EPMA	EPMA	EPMA	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS	LA-ICP-MS
Pyrite	mean	53.2	46.2	0.28	0.34	0.022	9.34	43.60	34.44	0.60	<D.L.	2.67	3.12	<D.L.	<D.L.	0.028	<D.L.	9.58
	s.d.	0.9	1.0	0.34	0.61	0.088	18.37	35.32	26.87	0.72		1.62	1.68			0.027		14.80
	median	53.4	46.5	0.15	0.11	0.001	2.22	43.00	28.00	0.28		2.82	3.00			0.015		4.01
	range																	
	min	50.4	42.2	0.01	4E-03	8E-05	0.65	0.65	3.10	0.09		0.09	1.50			0.003		0.30
	max	53.9	47.2	1.26	3.11	0.530	89.00	118.80	104.40	2.54		6.64	4.86			0.081		75.70
Chalcopyrite	mean	34.8	30.4	0.55	0.14	32.72	1170	<D.L.	<D.L.	2.01	<D.L.	<D.L.	<D.L.	<D.L.	<D.L.	<D.L.	<D.L.	20.75
	s.d.	0.3	0.9	0.76	0.17	0.98	1556			0.01								16.19
	min	34.2	29.4	0.01	0.02	31.25	69			2.00								9.30
	max	34.8	31.6	1.08	0.26	33.28	2270			2.01								32.20

s.d.=standard deviation; D.L.=detection limit; N.D.= not determined

Table 6. Composition of oxide minerals at IODP Hole 1256D;  $\sigma$ = standard deviation

		Magmatic Skeletal Titanomagnetite	Magmatic Ilmenite	Magmatic Magnetite	Hydrothermal Magnetite
<b>EPMA (wt.%)</b>					
SiO <sub>2</sub>	average	<b>0.49</b>	<b>0.12</b>	<b>0.17</b>	<b>0.59</b>
	$\sigma$	0.41	0.07	0.08	0.65
Al <sub>2</sub> O <sub>3</sub>	average	<b>1.62</b>	<b>0.07</b>	<b>1.13</b>	<b>0.63</b>
	$\sigma$	0.47	0.04	0.47	0.36
MgO	average	<b>0.24</b>	<b>0.22</b>	<b>0.07</b>	<b>0.15</b>
	$\sigma$	0.10	0.16	0.04	0.23
MnO	average	<b>0.68</b>	<b>0.92</b>	<b>0.11</b>	<b>0.11</b>
	$\sigma$	0.43	0.31	0.06	0.09
TiO <sub>2</sub>	average	<b>17.83</b>	<b>48.75</b>	<b>3.44</b>	<b>1.73</b>
	$\sigma$	6.49	1.45	1.41	1.45
CaO	average	<b>0.17</b>	<b>0.06</b>	<b>0.05</b>	<b>0.25</b>
	$\sigma$	0.17	0.06	0.04	0.28
FeO	average	<b>46.5</b>	<b>42.6</b>	<b>34.3</b>	<b>32.3</b>
	$\sigma$	5.70	1.26	0.83	1.18
Fe <sub>2</sub> O <sub>3</sub> *	average	<b>29.2</b>	<b>6.2</b>	<b>57.9</b>	<b>61.9</b>
	$\sigma$	12.51	2.53	2.13	3.24
Total	average	<b>97.4</b>	<b>99.4</b>	<b>98.9</b>	<b>98.1</b>
	$\sigma$	1.87	0.67	0.44	0.88
<b>LA-ICP-MS (ppm)</b>					
Magmatic Mixed signal					
51V	average	<b>4204</b>	<b>3959</b>		<b>2666</b>
	$\sigma$	1905	2925		610
	median	4505	2710		2580
	min	88	1142		2110
	max	8850	10100		3700
52Cr	average	<b>16</b>	<b>174</b>		<b>107</b>
	$\sigma$	27	238		119
	median	8	58		70
	min	2	6		16
	max	139	1008		314
60Ni	average	<b>101</b>	<b>229</b>		<b>676</b>
	$\sigma$	73	482		499
	median	84	112		477
	min	24	51		224
	max	310	3390		1420
65Cu	average	<b>55</b>	<b>31</b>		<b>44</b>
	$\sigma$	76	64		41
	median	21	9		33
	min	5	2		8
	max	316	277		102
66Zn	average	<b>1582</b>	<b>251</b>		<b>118</b>
	$\sigma$	764	195		82
	median	1517	158		86
	min	37	72		68
	max	3120	732		262
121Sb	average	<b>0.26</b>	<b>0.17</b>		<b>&gt;d.l.</b>
	$\sigma$	0.23	0.01		
	median	0.14	0.17		
	min	0.12	0.16		
	max	0.72	0.18		
208Pb	average	<b>2.62</b>	<b>2.30</b>		<b>1.76</b>
	$\sigma$	4.00	4.65		0.70
	median	1.31	0.43		2.00
	min	0.08	0.11		0.73
	max	19.60	23.50		2.30

$\sigma$ = standard deviation; Fe<sub>2</sub>O<sub>3</sub>\* is calculated using Carmichael (1966) and Lepage (2003) method.

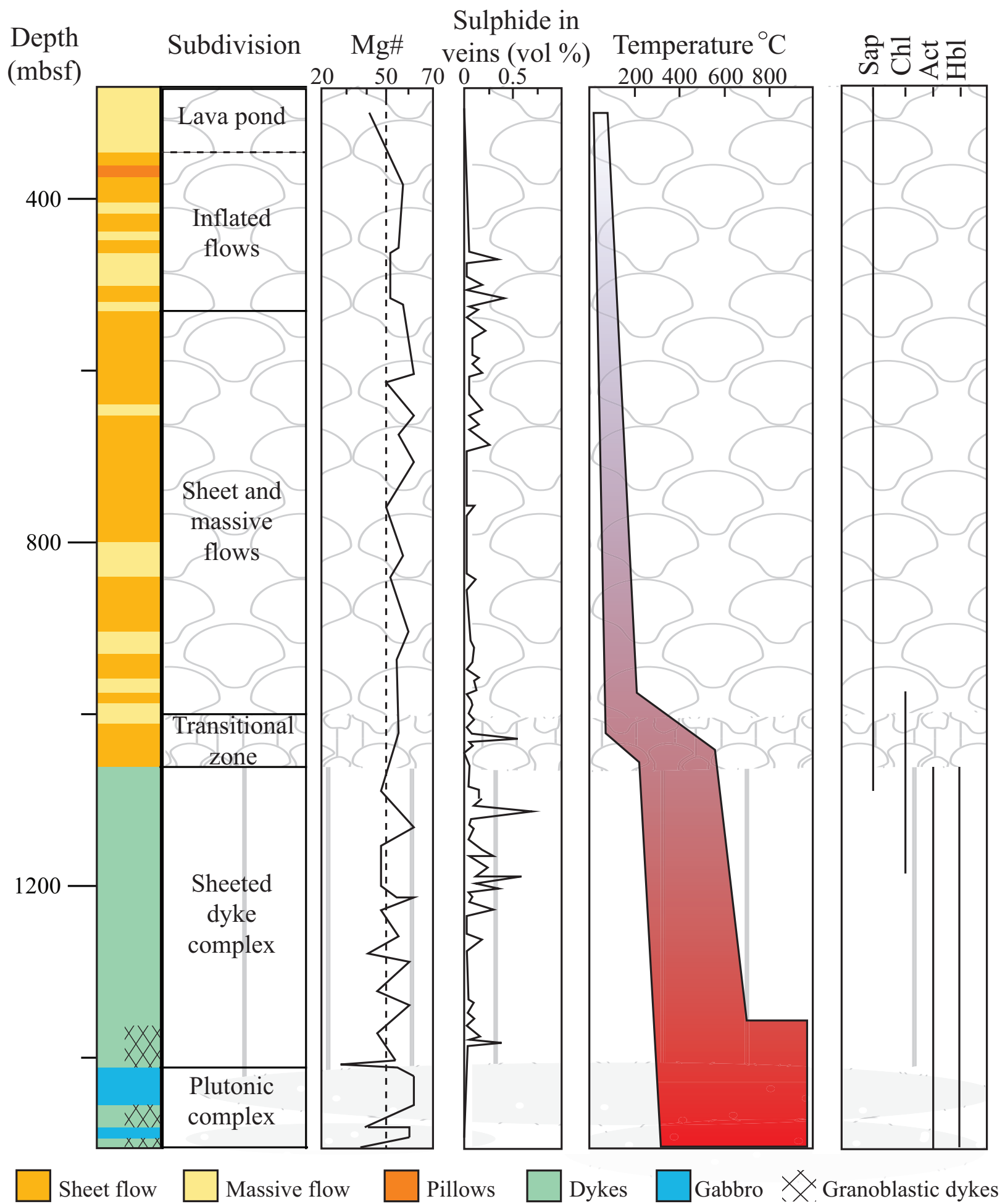


Figure 1.

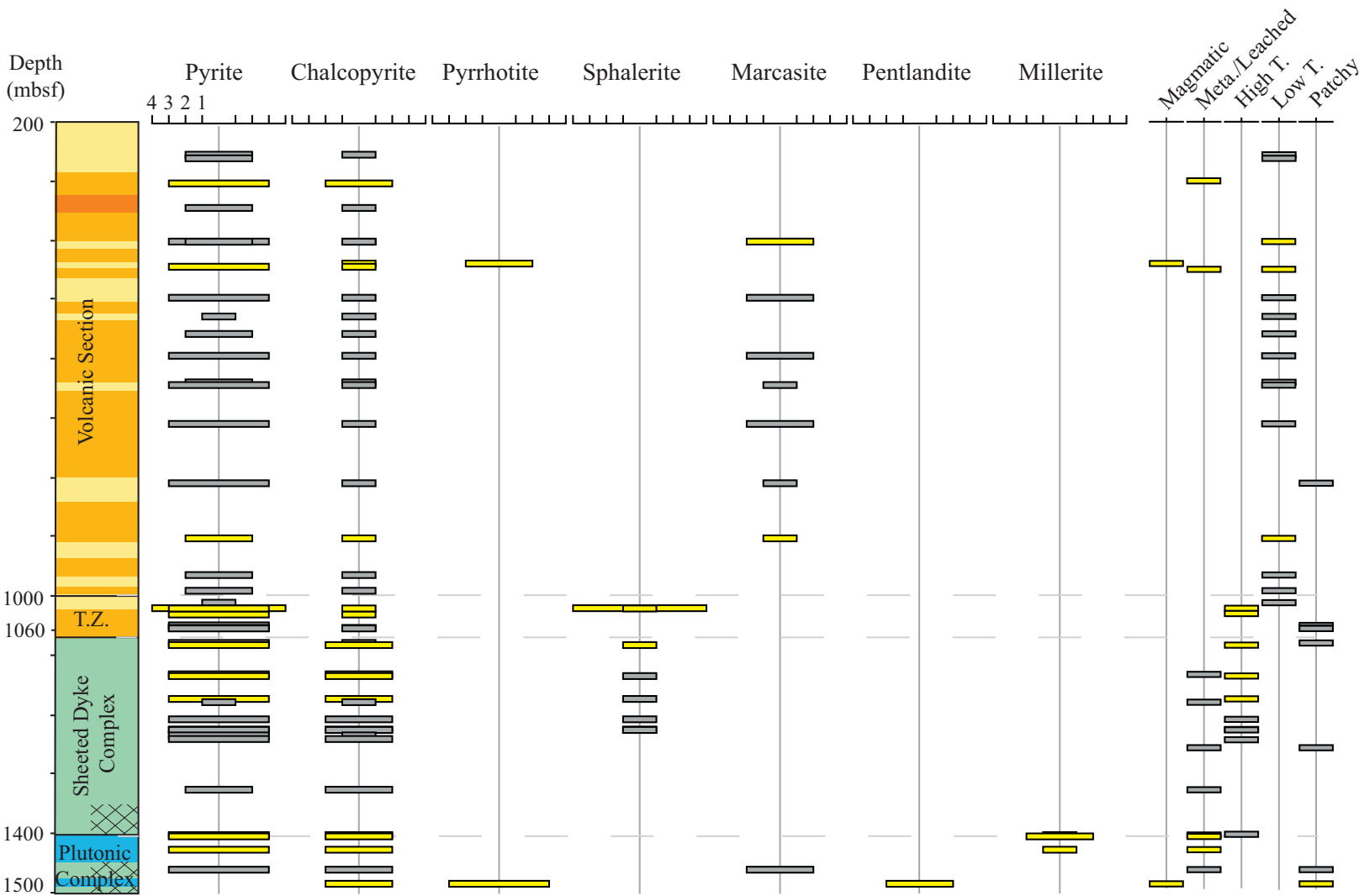


Figure 2.

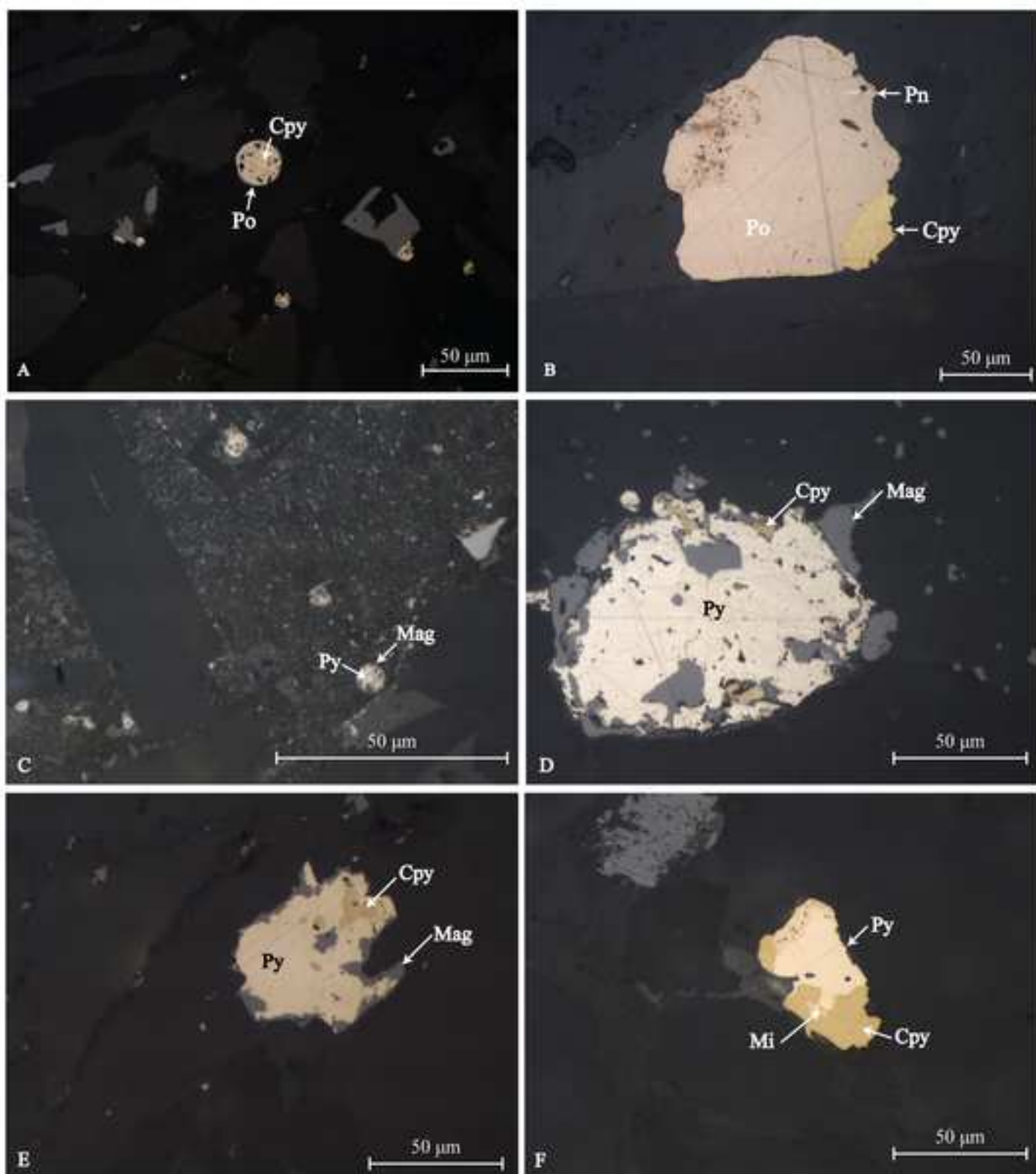


Figure 3.

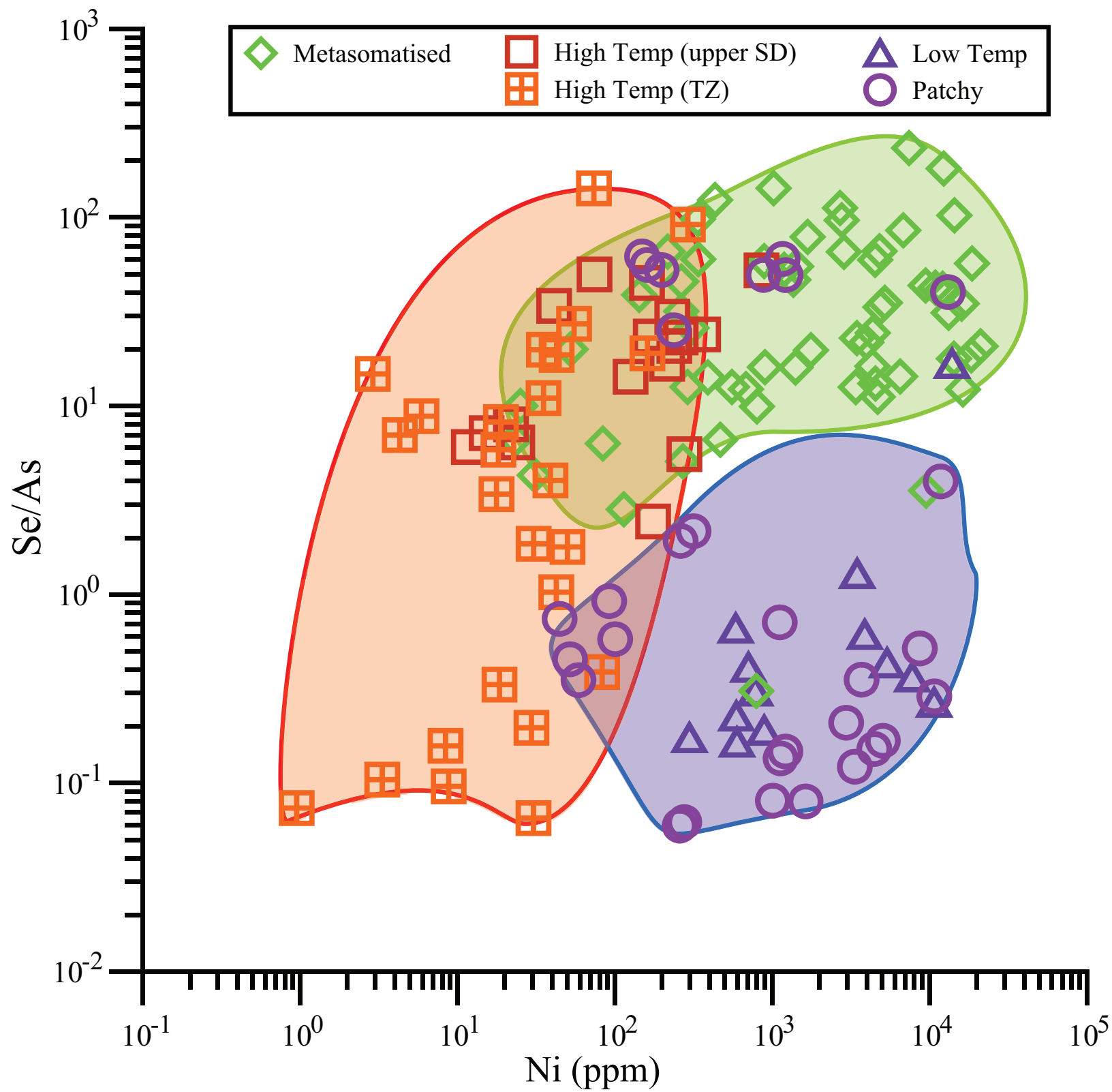


Figure 4.

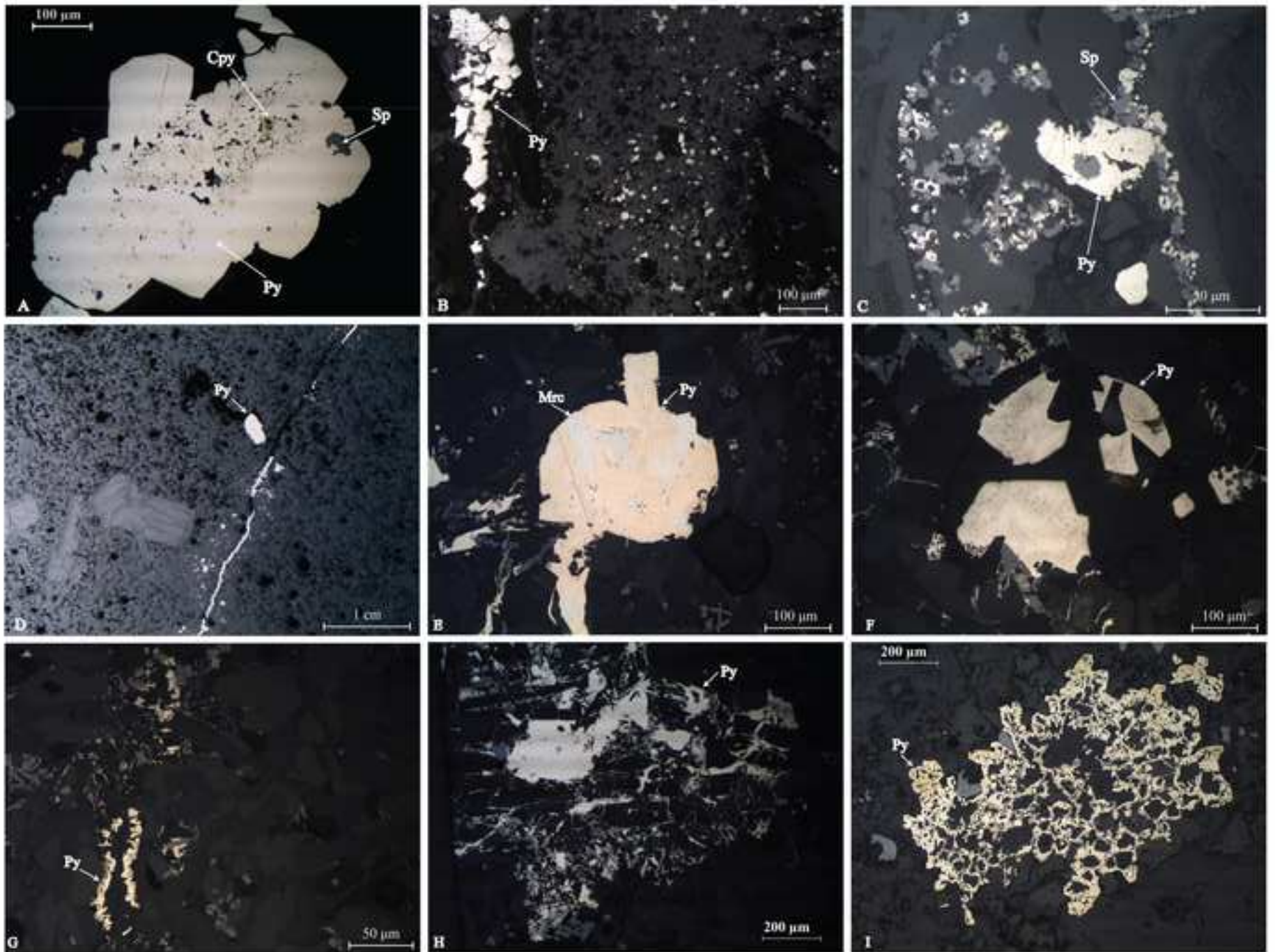


Figure 5.



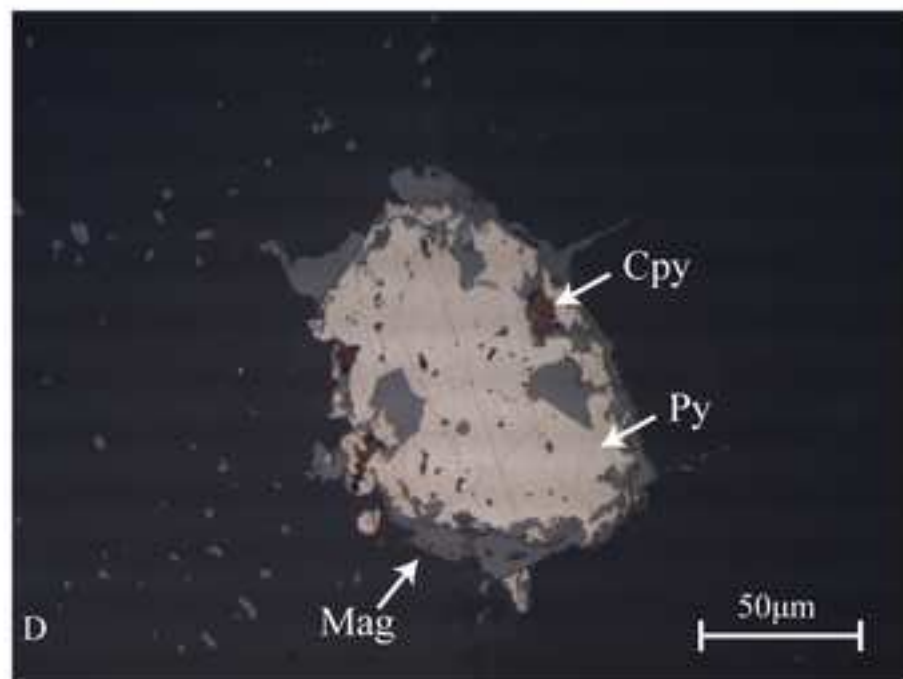
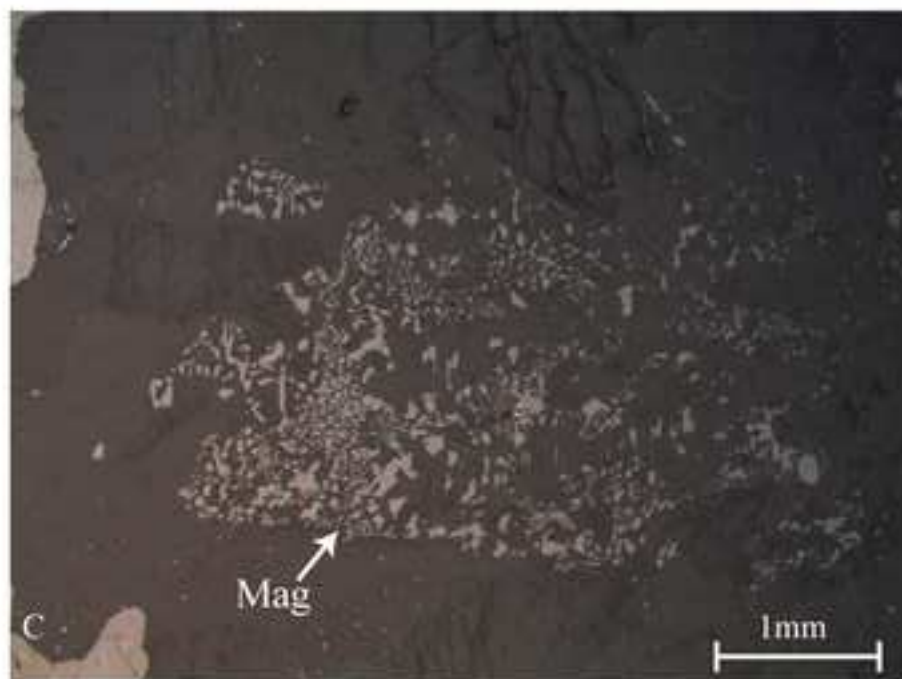
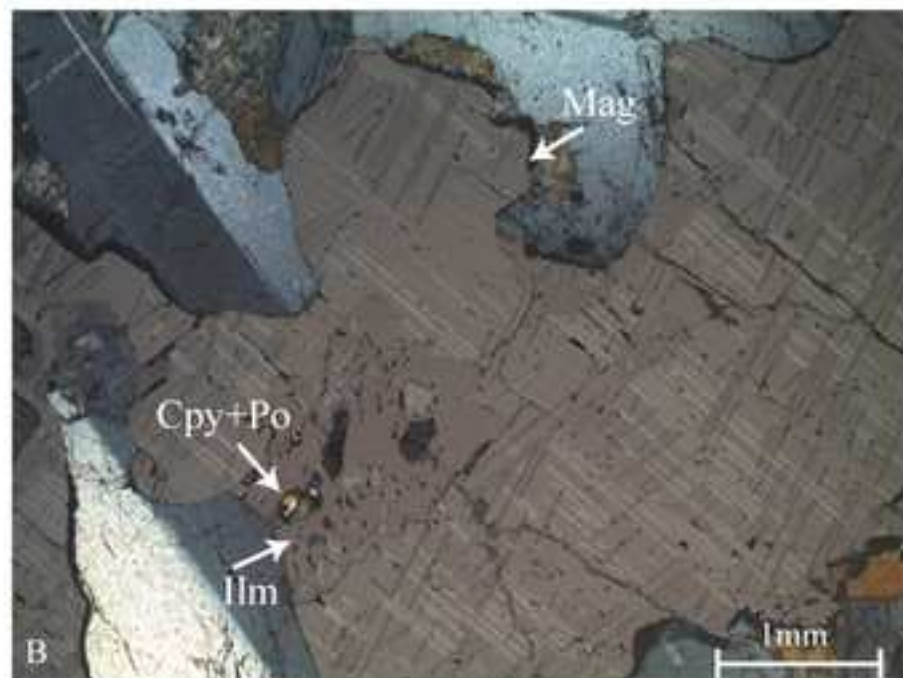
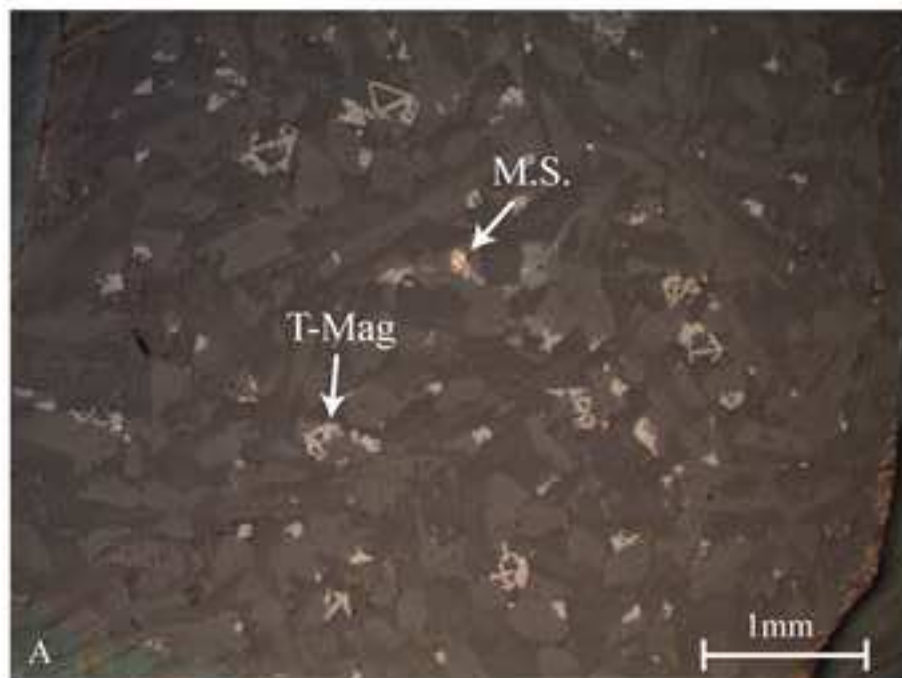


Figure 6.

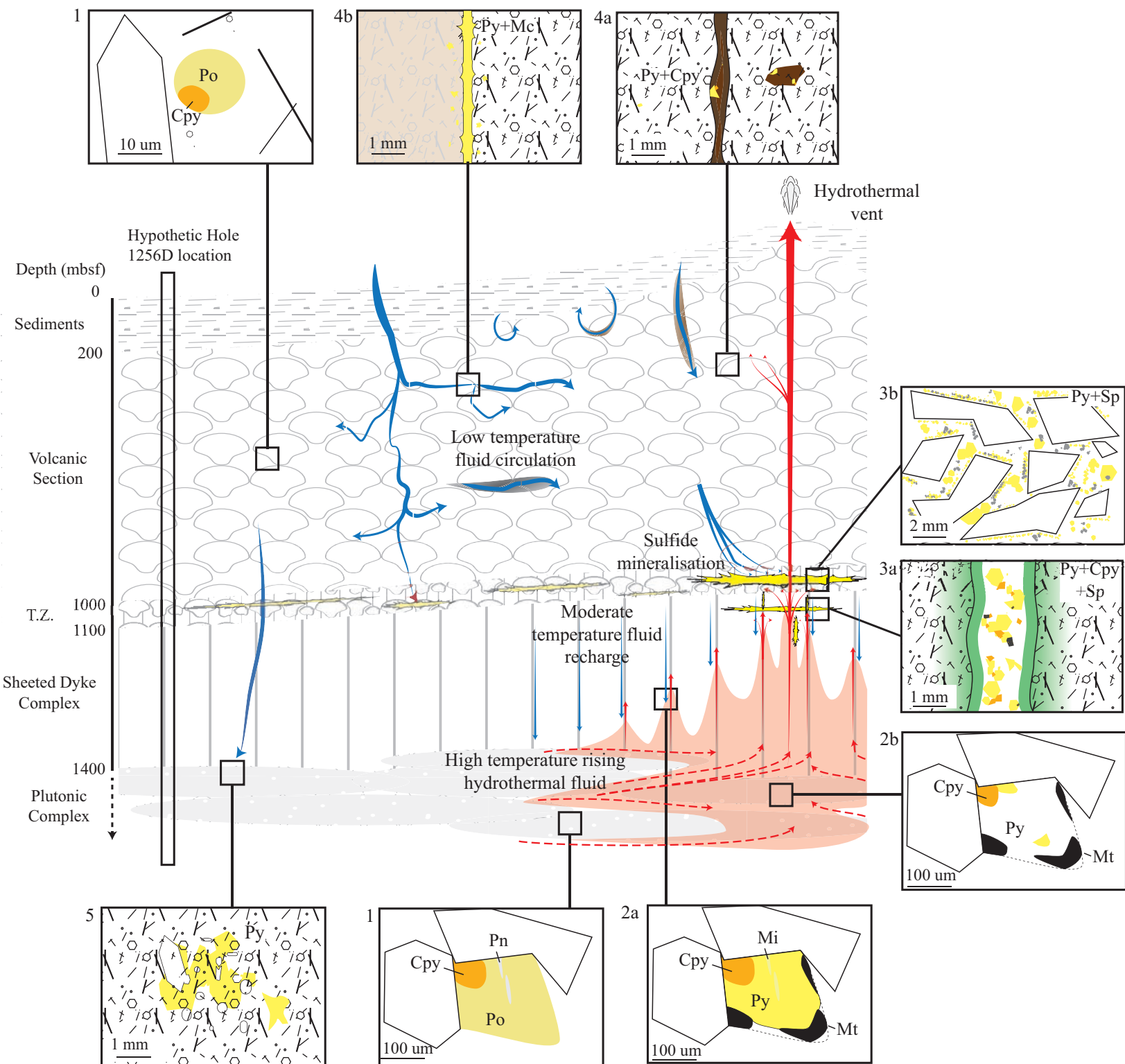


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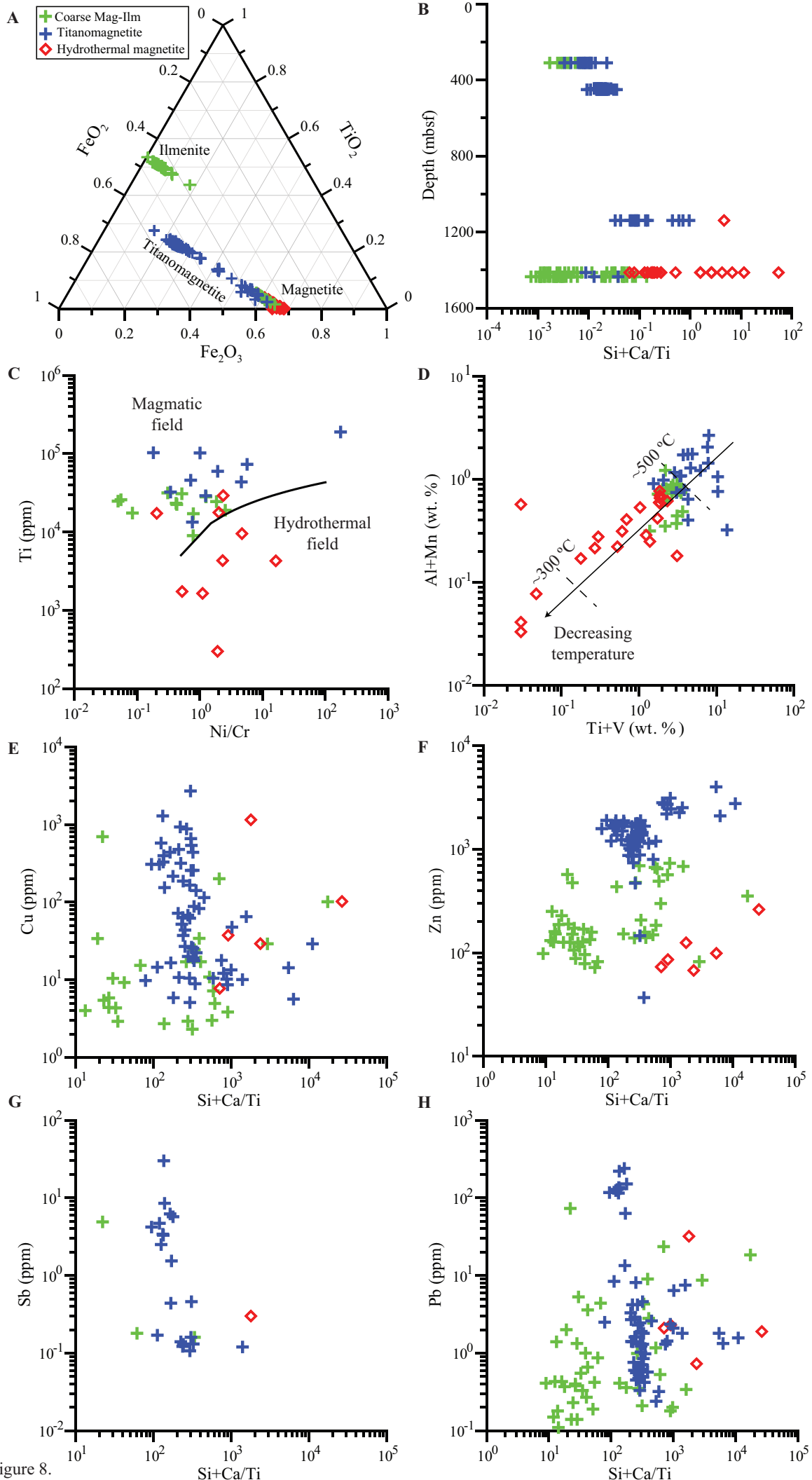


Figure 8.

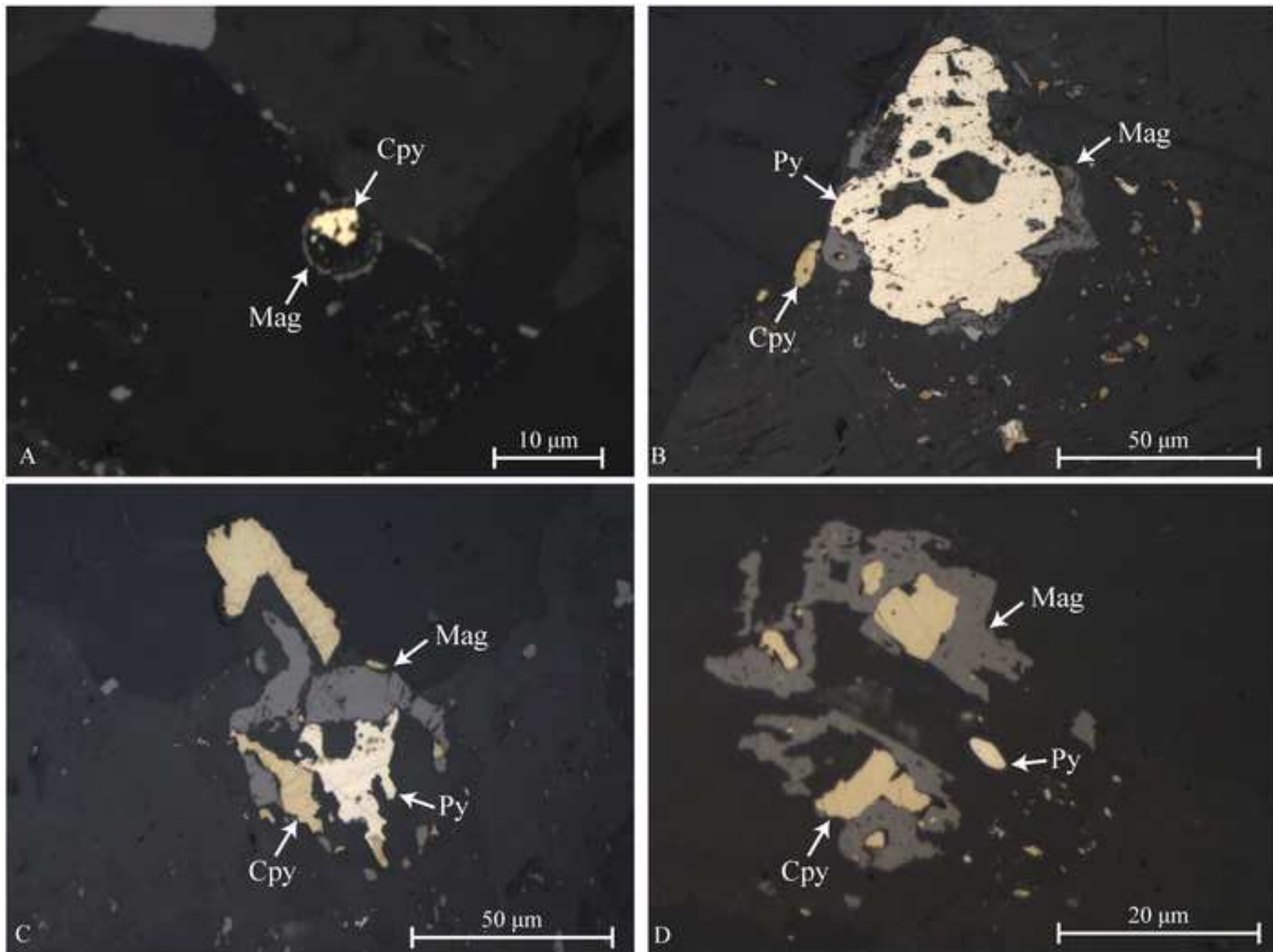
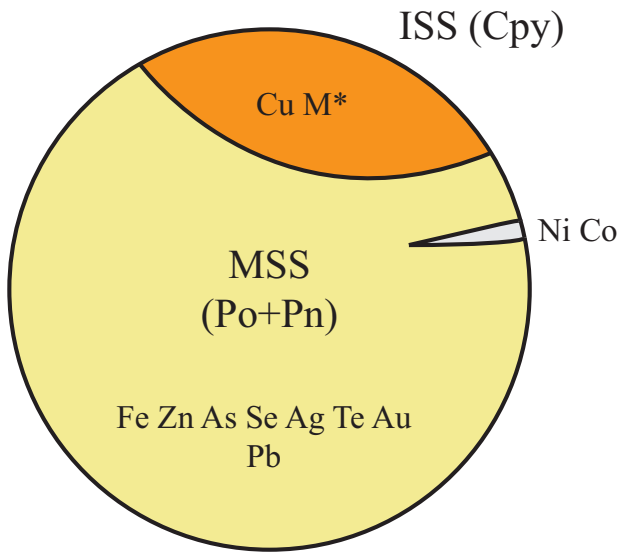
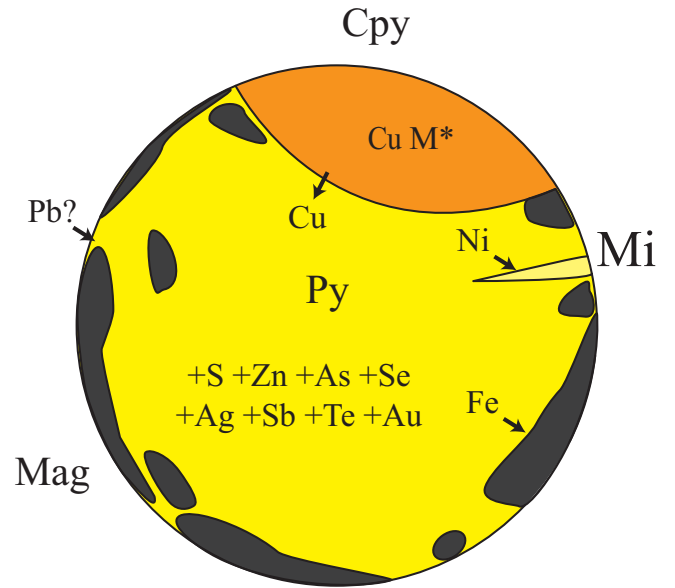


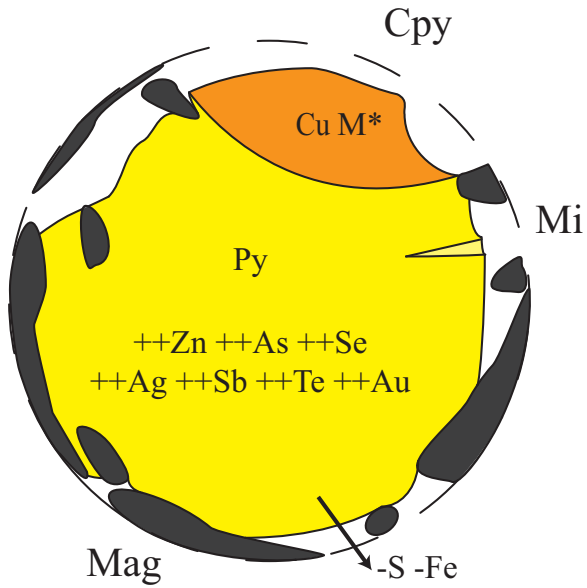
Figure 9.



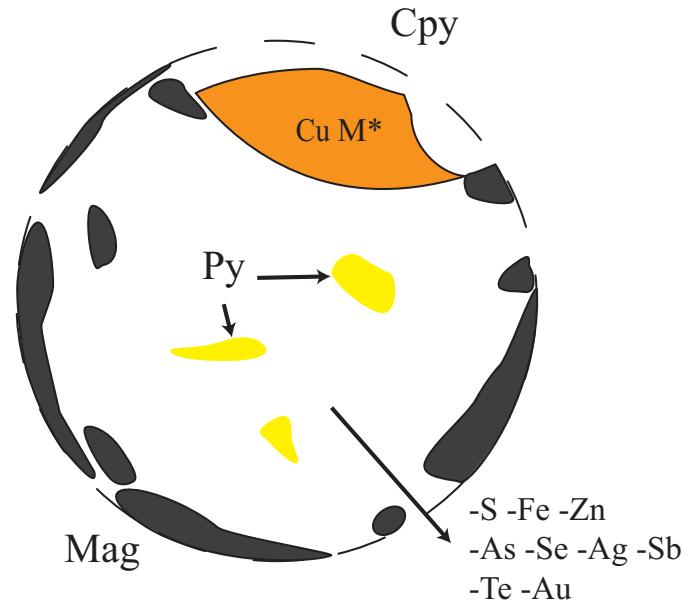
A. Magmatic assemblage  
High temperature 1200-650°C



B. Metasomatism  
Oxidation by moderate temperature hydrothermal fluids (250-350°C)



C. Partial sulphide leaching  
Leaching by high temperature hydrothermal fluids (>350°C)



D. Extensive sulphide leaching  
Leaching by high temperature hydrothermal fluids (>350°C)

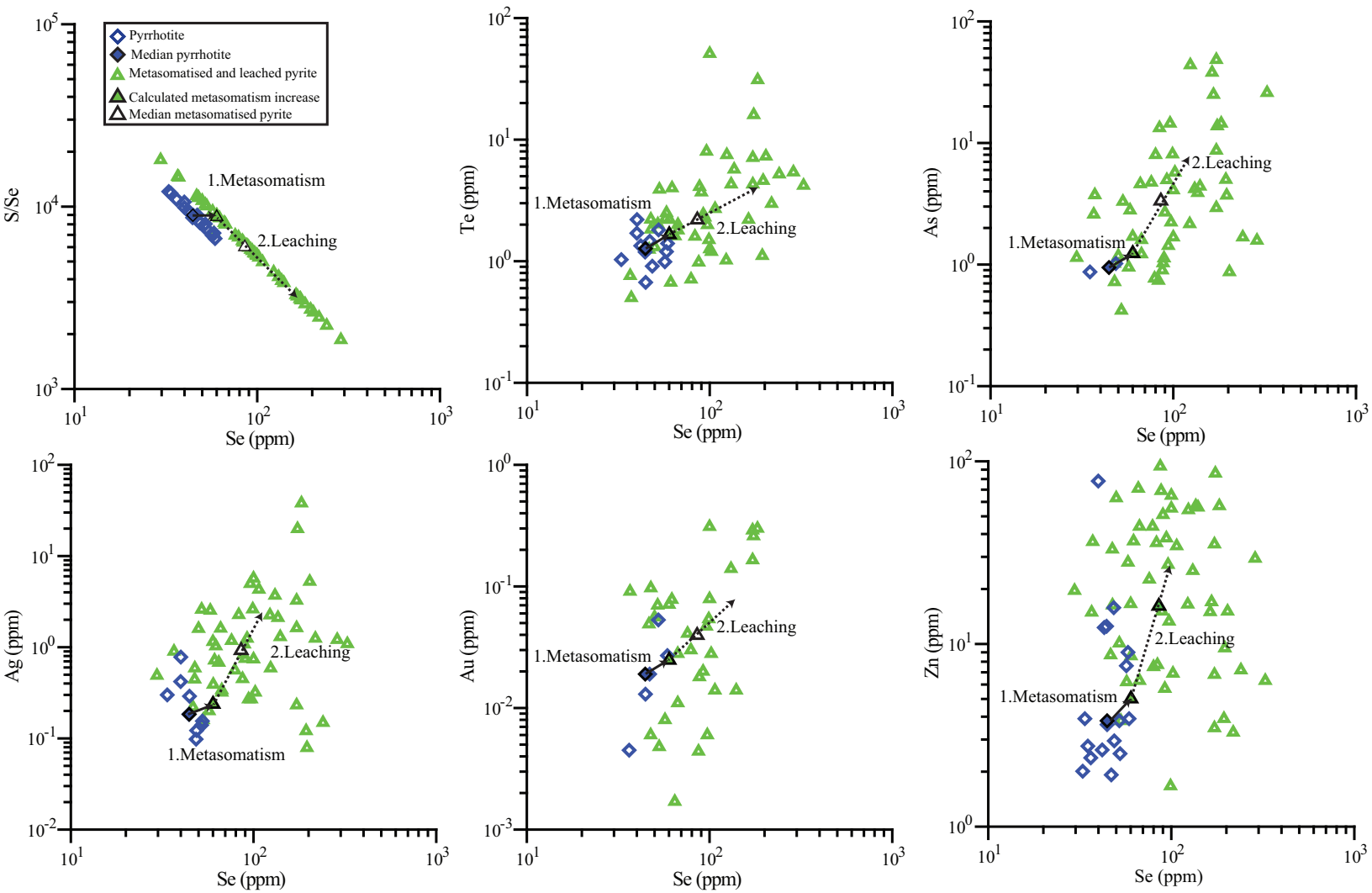
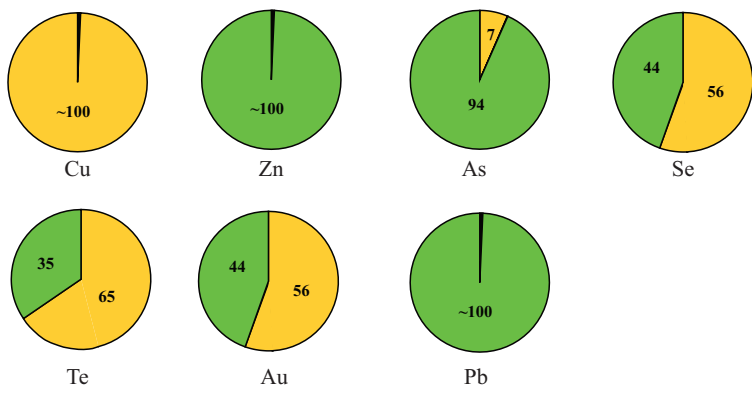
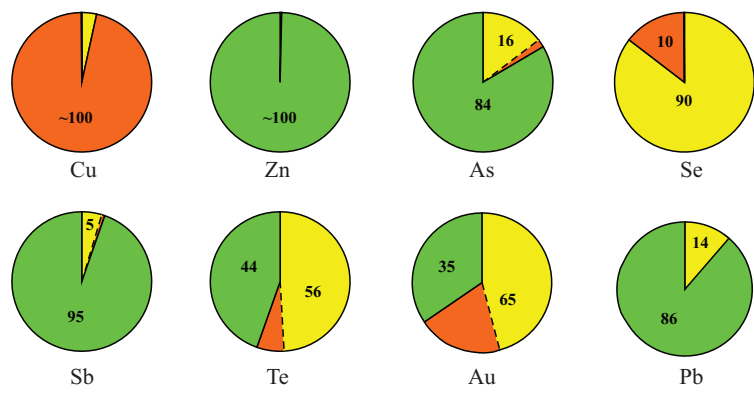


Figure. 11

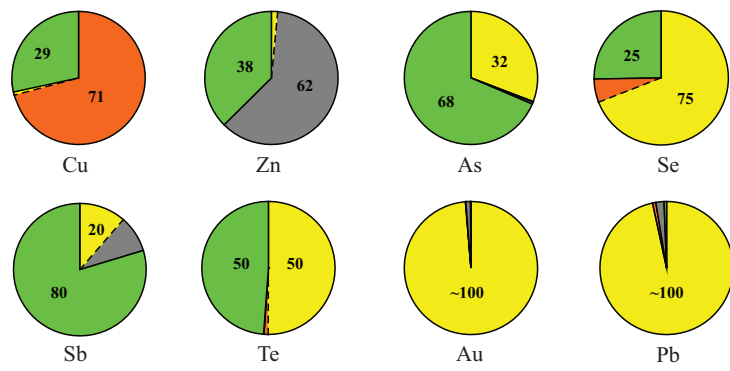
**Magmatic sulphide group**



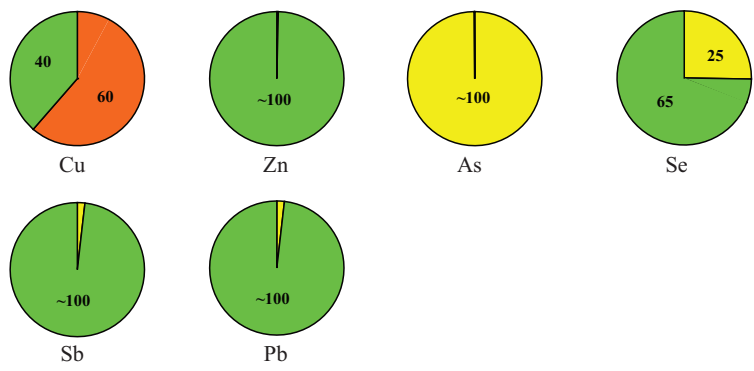
**Metasomatised sulphide group**



**High temperature sulphide group**



**Low temperature sulphide group**



**Patchy sulphide group**

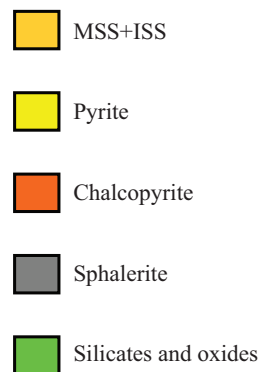
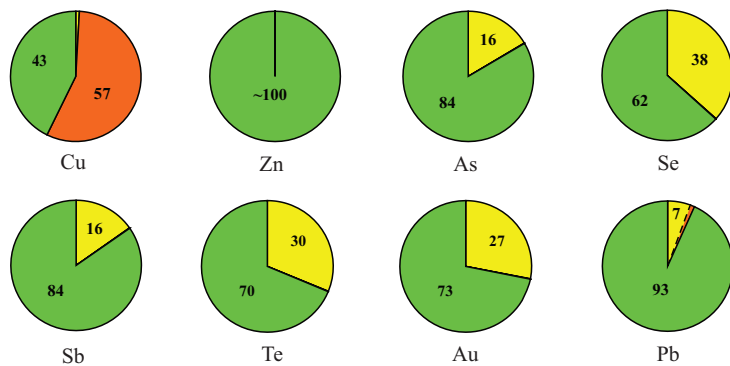


Figure 12.

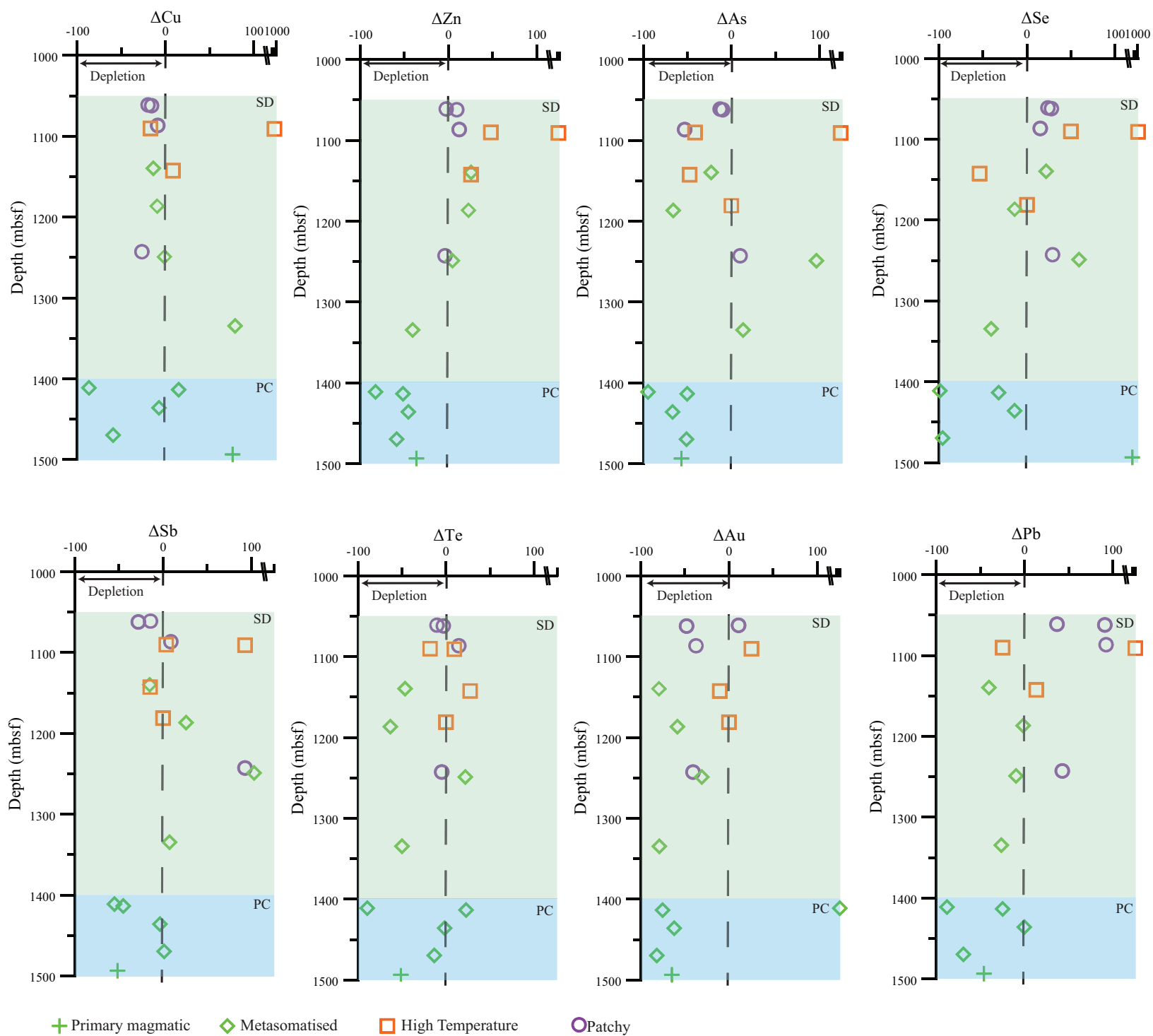


Figure 13.