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Draft Manuscript for Review

**Rhyolite generation prior to a Yellowstone supereruption:
Insights from the Island Park-Mount Jackson rhyolite series**

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Keyword:	Yellowstone, rhyolite, crystal fractionation, recycling, petrogenesis

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7 1 **Rhyolite generation prior to a Yellowstone supereruption:**
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9 2 **Insights from the Island Park-Mount Jackson rhyolite series**
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30 **ABSTRACT**

31 *The Yellowstone volcanic field is one of the largest and best-studied centres of rhyolitic*
32 *volcanism on Earth, yet it still contains little-studied periods of activity. Such an example is*
33 *the Island Park-Mount Jackson series, which erupted between the Mesa Falls and Lava*
34 *Creek caldera-forming events as a series of rhyolitic domes and lavas. Here we present the*
35 *first detailed characterisation of these lavas and use our findings to provide a framework for*
36 *rhyolite generation in Yellowstone between 1.3 and 0.6 Ma, as well as to assess whether*
37 *magmatic evolution hints at the forthcoming super eruption.*

38 *These porphyritic (15-40% crystals) lavas contain mostly sanidine and quartz with lesser*
39 *amounts of plagioclase (consistent with equilibrium magmatic modelling via rhyolite-*
40 *MELTS) and a complex assemblage of mafic minerals. Mineral compositions vary*
41 *significantly between crystals in each unit, with larger ranges than expected from a single*
42 *homogeneous population in equilibrium with its host melt. Oxygen isotopes in quartz and*
43 *sanidine indicate slight depletions ($\delta^{18}O_{\text{magma}}$ of 5.0-6.1‰), suggesting some contribution of*
44 *localised remelting of hydrothermally altered material in the area of the previous Mesa Falls*
45 *Tuff-related caldera collapse. The preservation of variable O isotopic compositions in quartz*
46 *requires crystal entrainment less than a few thousand years prior to eruption. Late*
47 *entrainment of rhyolitic material is supported by the occurrence of subtly older sanidines*
48 *observable in single-grain $^{40}\text{Ar}/^{39}\text{Ar}$ geochronology. Eruption ages of the lavas show discrete*
49 *clusters illustrating that extended quiescence (>100 ka) in magmatic activity may be a re-*
50 *occurring feature in Yellowstone volcanism.*

51 *Ubiquitous crystal aggregates dominated by plagioclase, pyroxene and Fe-Ti oxides are*
52 *interpreted as cumulates co-erupted with their extracted liquid. Identical crystal aggregates*
53 *are found in both normal- $\delta^{18}O$ and low- $\delta^{18}O$ rocks from Yellowstone, indicating that common*

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7 54 *petrogenetic processes characterise both volcanic suites, including the late-stage extraction*
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9 55 *of melt from an incrementally-built upper crustal mush zone.*

10 56 **Keywords:** *Yellowstone, rhyolite, ~~super-eruption~~petrogenesis, crystal fractionation, recycling*
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16 59 INTRODUCTION

17
18 60 The largest silicic eruptions on Earth are produced from caldera-forming explosive
19
20 61 volcanism. While events on the scale of the 5 km³ eruption of Pinatubo in 1991 can result in
21
22 62 measurable global climatic effect (Minnis *et al.*, 1993), the geological record illustrates that
23
24 63 the largest silicic eruptions can be as much as three orders of magnitude larger (Self, 2006,
25
26 64 Hildreth & Wilson, 2007). The catastrophic consequences emphasise the importance of a
27
28 65 detailed scientific investigation of the mechanisms controlling the generation of these large
29
30 66 volumes of rhyolitic magma that can culminate in so-called super eruptions.

31
32 67 At large, long-lived caldera centres such as Yellowstone, the catastrophic explosive events
33
34 68 represent only a tiny fraction of the lifespan of the volcanic centre; the majority of ~~time~~
35
36 69 activity is taken up by relatively quiescent effusion of lava and intermittent minor explosive
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38 70 activity. Studying this effusive activity preceding and post-dating large-volume eruptions is
39
40 71 key to understanding the magmatic system as a whole, and can be addressed over shorter
41
42 72 timescales via geophysical methods (Chang *et al.*, 2007) and over longer timescales via
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44 73 assessment of the compositional evolution of the reservoir (Vazquez *et al.*, 2009). Such an
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46 74 assessment includes questions of how such huge volumes of magma are generated, stored in
47
48 75 the crust and erupted, with a particular focus on the processes and temporal relations in the
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50 76 generation of these magmas.

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52 77 The formation of rhyolites in the Yellowstone province remains a topic of active debate with
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54 78 two proposed end-member models: (1) a fractionation-dominated evolution with subordinate
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7 79 assimilation in an upper crustal mush zone (Vazquez & Reid, 2002, Vazquez *et al.*, 2009,
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9 80 Girard & Stix, 2010, Stelten *et al.*, 2015) and (2) bulk or partial crustal melting initiated by
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11 81 intrusion of basaltic and/or silicic magma as a heat source with limited mass contribution
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13 82 (Bindeman & Valley, 2001, Bindeman *et al.*, 2008, Simakin & Bindeman, 2012). While the
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15 83 first model is mainly based on mineralogical and geochronological studies of young
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17 84 Yellowstone lavas, the second scenario has received significant support from oxygen isotope
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19 85 studies. These revealed rhyolitic lavas with unusual $\delta^{18}\text{O}$ characteristics and isotopically
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21 86 diverse phenocrysts, suggesting melting of hydrothermally altered lithologies in the shallow
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23 87 crust. In this study, we evaluate the applicability of these models to the record of lavas
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25 88 erupted prior to Yellowstone's latest super eruption at 0.631 Ma (Matthews *et al.*, 2015). The
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27 89 Island Park-Mount Jackson rhyolite series represents a prime example of a series of effusive
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29 90 events that preceded an explosive eruption, but has yet received little attention. By coupling
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31 91 major and trace elements in whole rock, glasses and minerals to oxygen and Pb isotopic data
32
33 92 and $^{40}\text{Ar}/^{39}\text{Ar}$ -ages, we obtain a detailed petrological and geochemical characterisation of
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35 93 these lavas to provide a framework for rhyolite generation in Yellowstone and to test whether
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37 94 the mineralogical and geochemical records were hinting at the forthcoming super eruption.
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96 **GEOLOGICAL BACKGROUND**

41 97 The Yellowstone volcanic field represents the current focus of volcanism, which has been
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43 98 ongoing for at least 16.5 Ma covering vast areas of the Northwest of the USA. The Columbia
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45 99 River-Snake River Plain-Yellowstone province is compositionally bimodal (basalt-rhyolite)
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47 100 and most commonly interpreted as the result of a hotspot (Geist & Richards, 1993, Hooper *et*
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49 101 *al.*, 2007, Wolff *et al.*, 2008). Silicic volcanism in the province was initially widely dispersed
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51 102 (e.g. Coble & Mahood, 2012), prior to becoming focused along the track of the Snake River
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53 103 Plain from around 14 Ma. Here, numerous voluminous ignimbrites and lavas were produced,
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7 104 with more frequent explosive eruptions and higher magmatic temperatures compared to the
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9 105 younger part of the province (Cathey & Nash, 2009, Ellis *et al.*, 2013).
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11 106 Silicic volcanism at Yellowstone has been ongoing for the past 2 Ma and is characterised by
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13 107 large-volume explosive eruptions separated by periods of relative quiescence during which
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15 108 effusion of rhyolitic lava dominates (Christiansen, 1984, Christiansen, 2001). The explosive
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17 109 eruptions consist of the Huckleberry Ridge Tuff (HRT) at 2.1 Ma (Rivera *et al.*, 2014b), the
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19 110 Mesa Falls Tuff (MFT) at 1.3 Ma (Lanphere *et al.*, 2002), and the Lava Creek Tuff (LCT) at
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21 111 0.6 Ma (Lanphere *et al.*, 2002, Matthews *et al.*, 2015, Wotzlaw *et al.*, 2015). These
22
23 112 voluminous deposits provide regionally significant stratigraphic markers and separate
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25 113 Yellowstone volcanism into three volcanic cycles (Christiansen, 2001).
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27 114 Following the early studies which considered Yellowstone volcanism as a whole (e.g.
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29 115 Christiansen & Blank, 1972, Doe *et al.*, 1982, Hildreth *et al.*, 1984, Hildreth *et al.*, 1991,
30
31 116 Christiansen, 2001), more recent work has focused separately on the large explosive events
32
33 117 and the youngest volcanism after the LCT, namely the Upper Basin Member and Central
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35 118 Plateau Member (e.g. Obradovich, 1992, Gansecki *et al.*, 1996, Bindeman & Valley, 2001,
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37 119 Lanphere *et al.*, 2002, Vazquez & Reid, 2002, Bindeman *et al.*, 2008, Girard & Stix, 2009,
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39 120 Vazquez *et al.*, 2009, Girard & Stix, 2010, Ellis *et al.*, 2012, Watts *et al.*, 2012). To improve
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41 121 the understanding of the early history of Yellowstone, we investigate the ~~poorly known/less-~~
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43 122 documented Island Park domes and Mount Jackson lavas, which were erupted between the
44
45 123 MFT and the LCT, thus spanning the second and third volcanic cycle in Yellowstone
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47 124 (Christiansen, 2001).
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49 125
50
51 126 *The Island Park – Mount Jackson (IPMJ) rhyolite series*
52
53 127 The Island Park (IP) domes include the undated units Silver Lake dome, Lookout Butte and
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55 128 Elk Butte, as well as Osbourne Butte (1.28±0.01 Ma) and Warm River Butte (1.24±0.02 Ma,
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7 129 all K/Ar ages by Obradovich, 1992). These domes are located in a NW-SE trending zone
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9 130 inferred to be tectonically controlled (Christiansen, 2001), with only the Lookout Butte dome
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11 131 erupted along what is thought to be the caldera related to the MFT (Fig. 1). Historically, the
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13 132 notably small-volume IP domes are thought to represent post-caldera units relating to the
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15 133 preceding MFT-related caldera collapse, whereas the volumetrically larger Mount Jackson
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17 134 (MJ) rhyolite lava flows are interpreted as pre-caldera units of the following LCT eruption
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19 135 (Christiansen, 2001). Units from the MJ series are widely dispersed across the Yellowstone
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21 136 volcanic field and include the Moose Creek flow (1.22 ± 0.01 Ma), Wapiti Lake flow
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23 137 (1.16 ± 0.01 Ma), Flat Mountain Rhyolite (0.929 ± 0.034 Ma), Lewis Canyon Rhyolite
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25 138 (0.853 ± 0.007 Ma), Harlequin Lake flow (0.839 ± 0.008 Ma), Big Bear Lake flow (undated)
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27 139 and Mount Haynes Rhyolite (0.609 ± 0.006 Ma, all K/Ar ages by Obradovich, 1992). These
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29 140 lava flows were proposed to be tracing an arcuate fault structure, which was later exploited
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31 141 during the LCT eruption (Christiansen, 2007). Although the origin of such an arcuate
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33 142 structure prior to caldera collapse remains unclear, it could have been the result of piecemeal
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35 143 caldera-related faulting during the eruption of the earlier HRT. Notably, the MJ rhyolites are
36
37 144 morphologically similar to the voluminous rhyolite lavas of the Plateau Member, whereas
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39 145 rhyolites from the IP series are much smaller and have dome-like structures (Fig. 1). Within
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41 146 the MJ lavas lies the Lewis Canyon Rhyolite which is inferred to represent an unknown
42
43 147 number of lavas which are petrologically dissimilar to the other MJ lavas in containing
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45 148 significantly more plagioclase (Christiansen & Blank, 1972, Christiansen, 2001). We
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47 149 therefore mention this unit separately in the results and discussion sections.
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151 **METHODS**

152 **Major and trace element analyses**

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7 153 Samples were analysed for bulk geochemistry via XRF and ICP at the GeoAnalytical
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9 154 Laboratory at Washington State University following the procedures described in Johnson *et*
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11 155 *al.* (1999). For mineral separates, rocks were fragmented using a SELFRAG at ETH Zurich.
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13 156 Following cleaning, samples were separated further if necessary via heavy liquid separation
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15 157 and minerals picked by hand under a binocular microscope.

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17 158 Electron microprobe analyses of mafic minerals were performed at the University of Kiel
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19 159 (Germany) with a JXA 8900 R Electron Microanalyser. 20-30 grains per unit were measured
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21 160 for rim and core composition. Fayalite was analysed with 20 keV and 15 nA, pyroxene and
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23 161 hornblende with 15 keV and 15 nA. Standard measurements for all EMP and LA ICP MS
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25 162 analyses can be found in the electronic supplementary material.

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27 163 Cathodoluminescence images of 25-30 epoxy-mounted quartz grains per sample were
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29 164 obtained at the Scientific Centre for Optical and Electron Microscopy (ScopeM) at ETH
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31 165 Zurich on a FEI Quanta 200 scanning electron microscope. Individual grains were evaluated
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33 166 for zonation grade (strongly zoned, gradual changes or unzoned) and direction (unidirectional
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35 167 bright core-dark rim or dark core-bright rim? Oscillatory zoning?). In the same fashion,
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37 168 sanidine grains from five representative units (IP units Osbourne Butte and Silver Lake dome,
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39 169 MJ units Moose Creek flow, Lewis Canyon Rhyolite and Mt. Haynes Rhyolite) were imaged
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41 170 for CL prior to trace element analysis.

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43 171 Analyses on feldspars were performed at ETH Zurich with a Jeol JXA 8200 electron
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45 172 microprobe at 15 keV and 15 nA. On all grains a minimum of two points were measured to
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47 173 obtain rim and core compositions. Beam diameter was set to 10 μm and counting times were
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49 174 shortened for Na and K to avoid element mobility. Glass shards were analysed with a beam
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51 175 of 20 μm and 12 nA at 15 kV.

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53 176 A 193nm Resonetics Resolution 155 excimer laser ablation system coupled to a Thermo
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55 177 Element XR sector field mass spectrometer was used to determine trace element in quartz,
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7 178 feldspar and glass at ETH Zurich. Spot sizes were 43 or 67 microns and samples were ablated
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9 179 using 5 Hz for 40 seconds after 30 s of gas blank acquisition. NIST 612 was used as a
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11 180 primary standard and GSD-1G as a secondary standard for quality control. Data was reduced
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13 181 using SILLS software (Guillong *et al.*, 2008) with SiO₂ contents from previous EMP analyses
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15 182 used as internal standard (or taken as 100% for quartz). Trace element determinations are
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17 183 considered to be precise and accurate within 5% of the reported values based on long-term
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19 184 reproducibility of a variety of glass standards.

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21 22 186 ⁴⁰Ar/³⁹Ar dating

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24 187 For ⁴⁰Ar/³⁹Ar geochronology, sanidine separates were handpicked under a binocular
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26 188 microscope from the 0.5 to 1 mm size fraction with care taken to avoid inclusions of glass or
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28 189 minerals and visible alteration. Samples were tested with the method of Hynek *et al.* (2011)
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30 190 to ensure a pure sanidine separate, before being cleaned and packed for irradiation at the
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32 191 CLICIT facility of the OSU reactor. 16-31 crystals per unit were placed into high-purity Al
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34 192 irradiation disks with samples of Alder Creek (ACs, age 1.2056 Ma) and Fish Canyon
35
36 193 sanidine (FCs-EK, age 28.294 Ma, Morgan *et al.*, 2014). Most samples were irradiated for 1
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38 194 hour in June 2013, but were not run due to laboratory issues. They were re-irradiated (along
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40 195 with the same monitors) for 1 hour in November 2014 and analysed in January 2015. Wapiti
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42 196 Lake was irradiated for 2 hours in October 2014 and analysed in January 2015 (note this
43
44 197 sample was not a re-irradiation). Following irradiation and cooling, sanidine samples were
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46 198 fused using a CO₂ laser. Extracted gases were subjected to 300 s of purification by exposure
47
48 199 to two SAES GP50 getters at room temperature and at 450°C (for full system technical details
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50 200 see Mark *et al.*, 2012). Isotope measurements were made using a MAP 215-50 noble gas
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52 201 spectrometer.

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7 202 All Ar isotope data were corrected for backgrounds (average \pm standard deviation from entire
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9 203 run sequence), mass discrimination (calculated from air calibration shots of 7.32×10^{-14}
10 204 moles ^{40}Ar), and reactor-produced nuclides, before being processed using standard data
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12 205 reduction protocols (Mark *et al.*, 2005) and reported according to the criteria of Renne *et al.*
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14 206 (2009) and relative to the optimisation model of Renne *et al.* (2010, 2011). We employed the
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16 207 atmospheric argon isotope ratios of Lee *et al.* (2006), which have been independently verified
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18 208 by Mark *et al.* (2011) and are consistent with inverse isochron plots for individual units. The
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20 209 eruption age of each unit is calculated from the weighted means of eruption-age total fusion
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22 210 ages of individual sanidines. Ages were calculated relative to the optimisation model of Renne
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24 211 *et al.* (2010, 2011) and isare reported as $X \pm Y/Z$ (2 sigma confidence level), where Y is the
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26 212 analytical uncertainty and Z is the full external precision, including uncertainty from the
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28 213 decay constant.

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31 215 **Isotopic analyses**

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33 216 Lead isotope compositions in sanidine crystals were measured using laser ablation
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35 217 multicollector inductively-couple plasma mass spectrometry (LA-MC-ICP-MS), with a
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37 218 Photon Machines G2 excimer laser ablation system and NuPlasma MC-ICP-MS at Oregon
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39 219 State University. Analytical techniques follow those given in Kent (2008). Analyses were
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41 220 made using a single 85 μm laser spot translated at 5 $\mu\text{m}/\text{sec}$ and a pulse frequency of 7 Hz on
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43 221 each grain. Measured ratios were corrected for mass bias based on measurement of NIST-612
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45 222 standard glass at similar ablation conditions throughout the analysis session. Repeated
46
47 223 analysis on standard reference BCR-2G glasses were typically within 0.01-0.22% of the
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49 224 accepted values.

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51 225 Quartz and sanidine separates for O isotopic determinations were cleaned using dilute HNO_3
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53 226 to remove any trace of adhering groundmass. Aliquots of 1-2 mg (typically 2-4 grains) were
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7 227 then analysed by CO₂-laser fluorination at the University of Oregon using BrF₅ as reagent, an
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9 228 Hg diffusion pump fluorine getter and a MAT253 mass spectrometer (Bindeman *et al.*, 2008).
10 229 The long term reproducibility of standards is $\pm 0.07\%$ (1 stdev) during runs where standards
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12 230 are run concurrently with the unknowns and of $\pm 0.2\%$ (1 stdev) during airlock runs where
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14 231 samples are run one by one.
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18 233 RESULTS

20 234 Sample description

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22 235 Samples were collected of the rhyolitic domes of the Island Park series (Osbourne Butte,
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24 236 Lookout Butte, Warm River Butte, Elk Butte and Silver Lake dome), as well as of the
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26 237 rhyolitic lava flows of the Mount Jackson Member (Moose Creek flow, Wapiti Lake flow,
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28 238 Harlequin Lake flow and Mt. Haynes flow), and the Lewis Canyon Rhyolite. The
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30 239 groundmass in the IPMJ rhyolites is microcrystalline, often displaying spherulitic textures
31
32 240 and flow banding. The IP units Silver Lake dome, Osbourne Butte, Elk Butte and Lookout
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34 241 Butte are only partially devitrified and still contain some glass. IPMJ rhyolites are porphyritic
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36 242 with the IP rhyolites being more crystal-rich (28-39%) than the MJ rhyolites (13-19%). The
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38 243 mineral assemblage consists of ~6-26% sanidine, ~4-17% quartz, ~2-7% plagioclase, and ~1-
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40 244 3% Fe-Ti oxides and mafic minerals, such as clinopyroxene, orthopyroxene, amphibole and
41
42 245 fayalite (table 1). Biotite, zircon and allanite occur as accessory phases. Magnetite and
43
44 246 ilmenite are usually exsolved. Some units (Warm River Butte and Lookout Butte) contain
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46 247 blebs with myrmekitic intergrowth of K-feldspar and quartz. The Lewis Canyon Rhyolite is
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48 248 petrographically distinct and shows abundant plagioclase (~7%) with a concurrent decrease in
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50 249 the amount of sanidine and quartz.

53 251 Bulk and glass compositions

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7 252 All IPMJ samples are rhyolitic in composition with the majority being high-silica (>75 wt%
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9 253 SiO₂ on anhydrous basis) rhyolites (table 2), except for Osbourne Butte and Lookout Butte
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11 254 (74.0 and 73.5 wt% SiO₂ respectively). Like most Yellowstone rhyolites, the IPMJ rhyolites
12
13 255 are relatively potassic with K₂O and Na₂O typically ranging around 5.2 wt% and 3.3 wt%
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15 256 respectively.

16 257 Only the IP domes and the Wapiti Lake flow contain glass ~~that can be analysed~~, in all other
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18 258 flows the groundmass is microcrystalline. Glass contains ~76-78 wt% SiO₂ (recalculated
19
20 259 anhydrous) in the IP domes and ~78 wt% SiO₂ in the Wapiti Lake flow with original totals of
21
22 260 ~96-99 wt% (Supplement). Glass shards contain high amounts of Rb (100-640 ppm) and Zr
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24 261 (120-160 ppm) and notably low contents of Sr (<10 ppm) and Eu. REE patterns of the glasses
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26 262 mimic those of the bulk samples with both MJ and IP rhyolites exhibiting 'seagull-shaped'
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28 263 REE patterns with a deep negative Eu anomaly and relatively straight high concentrations in
29
30 264 LREE and HREE (Fig. 2). These signatures are typical for rhyolites from the Snake River
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32 265 Plain-Yellowstone volcanic suite (Ellis *et al.*, 2013) and characteristic for hot and dry
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34 266 volcanic settings (Bachmann & Bergantz, 2008).

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37 268 ⁴⁰Ar/³⁹Ar geochronology

39 269 Single crystal ⁴⁰Ar/³⁹Ar ages for the IPMJ series range between 1.2943±0.0020/0.0026 and
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41 270 1.2190±0.0144/0.0146 Ma for the IP domes and between 1.2856±0.0064/0.0066 and
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43 271 0.7016±0.0014/0.0016 Ma for the MJ rhyolites, thus both series overlap by 0.066 Ma. All
44
45 272 ages with their respective uncertainties are presented in table 3. Contrary to conclusions from
46
47 273 Christiansen (2001) based on field observations and preliminary K-Ar dates on some of the
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49 274 units, there is no time gap between the IP domes and the MJ rhyolite series and both the
50
51 275 Moose Creek flow (1.2856±0.0064/0.0066 Ma) and the Wapiti Lake flow
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53 276 (1.2187±0.0158/0.0160 Ma) overlap with ages from the IP domes. However, there is a
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7 277 significant time gap between the eruption of Wapiti Lake flow at $1.2187 \pm 0.0156/0.0160$ Ma
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9 278 and the eruption of the subsequent Flat Mountain Rhyolite at 0.93 Ma (K-Ar age from
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11 279 Obradovich (1992), not sampled in this study). Among the MJ rhyolites, the Moose Creek
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13 280 flow is the oldest and erupted close to the MFT-related Henry Forks caldera (Fig. 3). The MJ-
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15 281 rhyolite Wapiti Lake flow erupted contemporaneously with the last of the IP domes, but at
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17 282 the ~~northeastern~~north-eastern end of the volcanic field. Coherent with observations from
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19 283 other Yellowstone rhyolites (Gansecki *et al.*, 1996, Gansecki *et al.*, 1998, Lanphere *et al.*,
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21 284 2002, Dallegge, 2008, Ellis *et al.*, 2012, Rivera *et al.*, 2014b, Stelten *et al.*, 2015), many of
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23 285 the IP and MJ lavas show subtly older sanidine grains, which are too old to be part of the
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25 286 eruption-age population (Fig. 3). It cannot be entirely excluded that apparent “too-old”
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27 287 $^{40}\text{Ar}/^{39}\text{Ar}$ ages were caused by excess Ar hosted in minute melt or fluid inclusions, however,
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29 288 we note that all dated grains were screened for inclusions under a binocular microscope.
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31 289

32 **Quartz**

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34 291 Quartz is a ubiquitous mineral phase in all IPMJ rhyolites. Crystals are commonly subhedral,
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36 292 rounded in shape and ~1.5-2 mm in diameter, but reach up to 3 mm in the Warm River Butte.
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38 293 Quartz crystals commonly host melt inclusions with sizes from ~10-200 μm . CL imaging
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40 294 reveals a lack of systematic zonation (i.e. same zonation pattern in majority of crystals) with
41
42 295 most of the grains displaying gradual transitions between several generations of dark and
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44 296 bright zones (Fig. 4). Within a unit, several modes of zonation are observed, from non-zoned
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46 297 grains to a few that are strongly zoned. Quartz grains in most of the units feature darker rims
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48 298 whereas a few grains from the Lewis Canyon (~15%) and the Silver Lake dome (~14%) show
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50 299 an abrupt late change to bright rims. Minor discordant growth zones indicate periods of
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52 300 resorption followed by new crystallisation. No correlations were observed with changes in
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54 301 crystal sizes.
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7 302 | Ti contents correlate with ~~trace element~~ contents of Al (73-130 ppm) and Li (15-28 ppm) and
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9 303 are in good agreement with relative CL brightness distributions (Fig. 4), consistent with them
10 304 being the main control of CL brightness (e.g. Leeman *et al.*, 2012 and references therein).
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12 305 Quartz grains feature a wide range of Ti contents (35-206 ppm), similar to those observed in
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14 306 post-LCT rocks (e.g. Vazquez *et al.*, 2009, Girard & Stix, 2010, Leeman *et al.*, 2012).
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16 307 Different crystals in the same unit can be zoned from lower to higher Ti values from core to
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18 308 rim, from higher to lower Ti values, or show no zonation at all, thus do not record a
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20 309 systematic compositional evolution with time. In most units, the Ti contents in rims of
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22 310 different grains are relatively similar even when different crystals display a variable intra-
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24 311 grain Ti-evolution. All trace elemental data for quartz and details regarding CL brightness
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26 312 and zonation of grains are provided in the supplementary materials.

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30 314 **Feldspars**

31 315 The IPMJ rhyolites contain both sanidine and plagioclase. However, the two phases occur in
32
33 316 markedly different contexts: Sanidine is the dominant feldspar within all IPMJ rhyolites
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35 317 except for the Lewis Canyon Rhyolite, occurring as euhedral, single crystals showing scant
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37 318 optical zonation and rarely containing melt inclusions. Sizes range from ~0.6 to 2.5 mm in
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39 319 diameter with the largest sizes occurring in the Warm River Butte, yet even the largest
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41 320 sanidines observed are smaller than the cm-sized sanidines in the MFT (Christiansen, 2001).
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43 321 By contrast, plagioclase is relatively rare in the IPMJ rhyolites with the exception of the
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45 322 Lewis Canyon Rhyolite. Crystal sizes range from 0.2 to 2 mm, commonly displaying small
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47 323 melt inclusions. A few crystals have sieve textures, a feature that has been previously
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49 324 described for the post-LCT lavas (Hildreth *et al.*, 1984, Girard & Stix, 2009, Watts *et al.*,
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51 325 2012) and can be overgrown by sanidine. Optical zonation is rare and seemingly limited to
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53 326 very large crystals (>0.8 mm). Plagioclase is most commonly found within polymineralic

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7 327 glomerocrysts (*see later*), which are dominated by pyroxene, plagioclase and Fe-Ti oxides
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9 328 and may contain accessory zircon and apatite with some interstitial groundmass.

10 329 Compositionally, all sanidines from the IPMJ series are similar and range from Or₄₃ to Or₆₃,
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12 330 with the end member compositions from the Lewis Canyon (Or₄₃₋₅₅) and the Moose Creek
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14 331 (Or₅₉₋₆₃) samples. Intra-grain major elemental compositional variability is limited, typically
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16 332 less than Or₁₋₂. Plagioclase features compositions between An₁₁₋₃₆, with end members being
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18 333 the Mt. Haynes Rhyolite (An₁₁₋₂₂) and the Warm River Butte (An₂₁₋₃₆). Generally, plagioclase
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20 334 from the IP rhyolites shows slightly higher An-contents compared to MJ lavas (Fig. 5). Both
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22 335 IPMJ sanidine and plagioclase resemble previously reported compositions in Yellowstone
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24 336 rhyolites (Hildreth *et al.*, 1984, Gansecki *et al.*, 1996, Girard & Stix, 2009, 2010, Stelten *et*
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26 337 *al.*, 2015).

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28 338 Trace elemental data from sanidines does not distinguish the different units with the
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30 339 exception of sanidine from the Lewis Canyon Rhyolite, which contains more Sr and Ba.
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32 340 Rather, sanidines within each of the IPMJ rhyolites are characterised by large spreads in trace
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34 341 elemental compositions between different grains, often equally as large as the range of trace
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36 342 element contents within sanidines from the entire IPMJ series. CL imaging of sanidines in a
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38 343 subset of rhyolites reveals diverse modes of zonation between different grains similar to
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40 344 observations in quartz. IP sanidines show both oscillatory and unidirectional zoning. Strong
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42 345 reverse, mostly unidirectional zoning with bright CL-rims occurs in MJ rhyolites with strong
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44 346 variations of trace elements such as Ba, Sr, Eu, and Ti between core and rim. Sanidines in the
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46 347 Lewis Canyon rhyolite record Ba contents of up to 10,000 ppm in rims (compared to ~7,500
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48 348 ppm in cores), to a lesser degree rim-enrichment can also be observed in Sr contents. IP
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50 349 sanidines exhibit larger inter-grain variability and show no preferred direction of zonation in
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52 350 trace element space. Inter-grain variability and possible causes for Ba-enriched sanidine rims
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54 351 are discussed later in more detail.

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353 **Mafic minerals**

354 Mafic minerals are generally rare in the IPMJ rhyolites. Previously, these lavas were
355 described as containing clinopyroxene and orthopyroxene in addition to accessory fayalite
356 and Fe-hornblende (Hildreth *et al.*, 1984, Christiansen, 2001). However, in some of the
357 studied lavas, particularly those of the IP series, fayalite and hornblende (Fig. 6) are abundant
358 and reach up to 60% of the mafic mineral assemblage (table 1). Where both pyroxenes are
359 present, clinopyroxene is more abundant. However, there are cases where clinopyroxene
360 (Warm River Butte, Wapiti Lake flow, Harlequin Lake flow) or orthopyroxene (Mt. Haynes
361 Rhyolite) is sufficiently rare to be absent in mineral separates.

362 Despite limited intra-grain variability, compositional ranges in clinopyroxenes are large
363 between different grains ($\text{En}_{11-38}\text{Wo}_{30-48}$, Fig. 7). Mg# averages for each unit range from 32 ± 7
364 to 44 ± 2 , whereas individual crystals can reach Mg# of up to 70. The degree of scatter varies
365 markedly between units, e.g. clinopyroxenes from the IP domes Lookout Butte, Osbourne
366 Butte and Silver Lake dome cluster around $\text{Mg\#} = 33 \pm 7$ with individual grains as high as 56,
367 whereas grains from Lewis Canyon Rhyolite have values of around 45 ± 2 . Orthopyroxene has
368 compositions of $\text{En}_{18-65}\text{Wo}_{02-15}$. Additionally, the Mt. Haynes Rhyolite contains pure
369 ferrosilite with >2 wt% MnO and <0.1 wt% MgO. Both ortho- and clinopyroxenes can often
370 be found in crystal aggregates with plagioclase and iron oxides (see separate discussion);
371 these aggregate-hosted pyroxenes are compositionally identical to single pyroxenes in the
372 groundmass.

373 The REE abundances in clinopyroxene are similar to patterns observed in whole rock trace
374 elements and show elevated contents in LREE and HREE and a marked Eu anomaly (Fig. 8).
375 Orthopyroxenes show similar behaviour with the majority of grains having LREE $>$ HREE
376 and elevated contents in Zr, Sr and Sc. A minority of both clinopyroxene and orthopyroxene

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7 377 have LREE < HREE, lower overall abundances of REE, lack a pronounced Eu anomaly and
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9 378 have Mg# > 60.

10 379 Fayalite is restricted to the IP domes and the Moose Creek flow. Forsterite contents range
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12 380 from 6 to 15, slightly higher than the 2-7 reported by Vazquez *et al.* (2009) for post-LCT
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14 381 lavas. Fayalites commonly show altered rims and oxidation features involving the
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16 382 transformation to iron oxides and hydroxides along cracks (Fig. 6). In trace elements, the
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18 383 majority of grains show an evolved character with LREE > HREE and marked Eu anomalies,
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20 384 while a minority of crystals have low LREE contents and a less pronounced Eu anomaly (Fig.
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22 385 8).

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24 386 Amphiboles are Fe-hornblende with subhedral shapes without zoning or alteration features
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26 387 (Fig. 6). They occur commonly in IP domes, Moose Creek flow and the Lewis Canyon
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28 388 Rhyolite. All hornblendes display pronounced negative Eu anomalies and high contents of
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30 389 LREE and HREE (Fig. 8).

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33 34 391 **Oxygen isotopes**

35 392 Oxygen isotopic compositions were determined on quartz and sanidine with ranges of 4.7 to
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37 393 6.9 ‰ for quartz and 4.5 to 6.0 ‰ for sanidine (all data in supplementary materials). Intra-
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39 394 unit variability is large, reaching 1.3‰ in both quartz and sanidine (Fig. 9). Units with lower
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41 395 oxygen isotopic compositions such as the IP domes display more variability in $\delta^{18}\text{O}$. MSWDs
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43 396 of quartz populations are mostly higher than the critical MSWD for the analysed number of
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45 397 samples (Mahon, 1996), suggesting significantly more variation than expected for a single
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47 398 population.

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49 399 Sanidine $\delta^{18}\text{O}$ values are 0.5-1.0 ‰ lower than $\delta^{18}\text{O}$ for quartz in the same unit. While such
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51 400 divergence can be expected with equilibrium fractionation of $\Delta^{18}\text{O}(\text{Qtz-Fsp}) = 0.75$ ‰ in
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53 401 silicic magmas at $850 \pm 60^\circ\text{C}$ (Bindeman & Valley, 2001), overlaps in $\delta^{18}\text{O}$ values between
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7 402 quartz and sanidine crystals and the large variation of $\delta^{18}\text{O}$ within a rhyolite preclude a closed
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9 403 system evolution for the IPMJ rhyolites. Taking average quartz values and assuming a
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11 404 fractionation factor of $\Delta^{18}\text{O}(\text{Qtz-melt}) = 0.35 \text{ ‰}$ (Bindeman & Valley, 2001), we estimate
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13 405 magmatic $\delta^{18}\text{O}$ values between 4.9 and 6.1 ‰ for different units. Only the youngest MJ lavas
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15 406 are normal- $\delta^{18}\text{O}$ rhyolites (defined as 5.8-6.3 ‰, Bindeman, 2008), whereas the majority
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17 407 shows a slight depletion. All units are significantly lower than the $\delta^{18}\text{O}$ value of ambient crust
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19 408 (6.8-7.2‰, Fig. 9) or units that show a significant crustal component in oxygen, Pb, Sr and
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21 409 Nd-isotopic space, such as HRT member C and extra-caldera post-LCT rhyolite (Hildreth *et*
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23 410 *al.*, 1991).

24 411 The mildly depleted- $\delta^{18}\text{O}$ signature of both the IP and MJ lavas is in good agreement with
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26 412 values previously reported for some of the MJ units (Hildreth *et al.*, 1984, Bindeman &
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28 413 Valley, 2001) with the MJ lavas being slightly higher than the IP domes (Fig. 9). Both the
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30 414 MFT and LCT explosive eruptions have higher $\delta^{18}\text{O}$ than the intervening lavas. The pattern
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32 415 of dropping and recovering $\delta^{18}\text{O}$ magmatic values following large caldera-forming eruptions
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34 416 is well-documented for Yellowstone (Hildreth *et al.*, 1984, Bindeman & Valley, 2001) with
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36 417 the IPMJ series exhibiting similar behaviour on a smaller isotopic scale. Although inter-
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38 418 crystal O isotopic variability in the IPMJ rhyolites is not as large as in the low- $\delta^{18}\text{O}$ post-LCT
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40 419 lavas (Bindeman & Valley, 2001, Watts *et al.*, 2012), it still records a lack of closed-system
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42 420 fractionation throughout the history of Yellowstone.

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44 422 **Pb isotopes**

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46 423 Pb isotopes in sanidine have been proven to be a useful tool for the distinction of units that
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48 424 are very similar in trace element composition (e.g. Watts *et al.*, 2012, Stelten *et al.*, 2015).

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50 425 Data for the IPMJ rhyolites are in the range of previously published data for Yellowstone and
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52 426 differences between units are slight compared to the overall range of Pb isotopes in
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7 427 Yellowstone (Fig. 10). Values seem to transition from MFT-like to LCT-like compositions
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9 428 with Pb isotopes in the IP domes being close to those of the preceding MFT, whereas MJ
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11 429 units are closer to those of the LCT (Fig. 10). We note that only a single translated spot was
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13 430 analysed per grain and that these analyses were obtained prior to CL imaging and therefore
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15 431 yield no information on whether crystals are zoned in terms of Pb isotopes.
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18 433 **Intensive parameters**

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20 434 Zr-saturation thermometry (Hanchar & Watson, 2003, Boehnke *et al.*, 2013) returns
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22 435 temperatures of 795-858°C for bulk rock compositions with the Lewis Canyon Rhyolite
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24 436 having the highest inferred temperature (table 2). Where glass is present, results for glasses
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26 437 are typically 10-20°C higher than temperatures derived from bulk rock compositions, but
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28 438 within uncertainty. Bulk rock Zr-saturation temperatures from Boehnke *et al.* (2013) are 10-
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30 439 40°C lower than those from Hanchar and Watson (2003) and within error of results from
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32 440 alkali feldspar-liquid thermometry (Putirka, 2008), except for Warm River Butte which
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34 441 returns slightly lower temperatures (771±23°C instead of 814±17°C). All details and input
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36 442 parameters for thermometry calculations can be found in the supplementary material.

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38 443 Ti-in-quartz thermometry (TitaniQ, Wark & Watson, 2006) was not performed because we
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40 444 were unable to sufficiently constrain a_{TiO_2} . Fe-Ti oxides in these lavas are exsolved, therefore
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42 445 prohibiting calculation of a_{TiO_2} from co-existing mineral pairs. Given the compositional range
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44 446 of quartz grains in these lavas, magmatic conditions with similar a_{TiO_2} in space and time seem
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46 447 unlikely and previously employed activities range from 0.33 to 0.56 between different post-
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48 448 LCT lava flows (Vazquez *et al.*, 2009). Ti activities may change as a function of Ti content in
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50 449 the melt and its major element composition, co-crystallising Fe-Ti oxide composition, oxygen
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52 450 fugacity, and temperature and pressure conditions (e.g. Thomas & Watson, 2012 and
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54 451 references therein). Thus, a_{TiO_2} may have changed significantly throughout the growth
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7 452 history of even a single quartz grain, suggesting that an averaged a_{TiO_2} calculated from co-
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9 453 existing Fe-Ti oxides or TiO_2 solubility models provides only estimates at best.

10 454 The previously unrecognised occurrence of abundant amphibole and accessory biotite in the
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12 455 IPMJ rhyolites suggests sufficient presence of water to create hydrous mineral phases (e.g.
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14 456 >4.5 wt% H_2O for formation of amphibole and >2.5 wt% H_2O for biotite in experiments on
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16 457 Yellowstone hotspot track rhyolites by Almeev *et al.* (2012)). Estimates of water contents
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18 458 from plagioclase-melt hygrometers (Putirka, 2008) indicate that about 3.6 ± 1.1 to 5.7 ± 1.1
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20 459 wt% $-\text{H}_2\text{O}$ were present during the formation of the IPMJ rhyolites (see supplementary
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22 460 material). Although strongly depending on inferred temperatures, these values are consistent
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24 461 with the generally increasing water contents from the central Snake River Plain to
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26 462 Yellowstone (Almeev *et al.*, 2012, Bolte *et al.*, 2015).

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29 30 464 **DISCUSSION**

31 32 465 **Time-space evolution of the Yellowstone volcanic field at 1.3-0.6 Ma**

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34 466 The new $^{40}\text{Ar}/^{39}\text{Ar}$ -geochronology provides insights into the development of the volcanic
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36 467 field during the period between MFT and LCT eruptions. Broadly speaking, volcanism
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38 468 migrated from west to east: Post-caldera activity first focused on the area of the foregone
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40 469 MFT-related caldera collapse with the eruption of the IP domes and the Moose Creek flow,
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42 470 before transitioning to areas surrounding the younger LCT-related caldera, where activity is
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44 471 characterised by eruptions from widely separated regions.

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46 472 The overlap between the two magmatic series questions the traditional inference of two
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48 473 different cycles with the IP domes marking post-caldera activity of the MFT-forming
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50 474 eruption (2nd cycle) while the MJ succession records pre-caldera activity of the third cycle
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52 475 encompassing the eruption of the LCT (Christiansen, 2001). Both in terms of age and
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54 476 geochemical behaviour the Moose Creek flow is similar to the IP-domes, e.g. in mafic

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7 477 mineral abundances (table 1), feldspar and pyroxene major and trace element compositions
8 478 (Figs. 4, 5) and oxygen and Pb isotopic compositions (Figs. 9, 10). Geochemistry may thus
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10 479 largely be controlled by geographical location and source region rather than affiliation with a
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12 480 magmatic series defined by morphology.

14 481 The temporal overlap between the MJ and IP series suggests that the different morphologies
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16 482 of IP domes compared to the lava flows of the MJ series are not controlled by a time-
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18 483 progressive evolution. IPMJ volcanism occurred in unevenly distributed clusters of 2-4 lava
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20 484 flows (Fig. 3), coherent with discrete periods of activity observed in the post-LCT record
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22 485 (Christiansen, 2001, Christiansen, 2007). Notably, in some cases lava flows erupted
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24 486 contemporaneously within uncertainty with other flows at the opposite edge of the volcanic
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26 487 field (e.g. Lookout Butte/Wapiti Lake flow), suggesting that batches of broadly similar, yet
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28 488 isotopically subtly distinct liquids were present simultaneously within the crust.

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30 489 The largest time gap in the IPMJ record is the 290 ka of quiescence starting ~120 ka after the
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32 490 MFT-eruption between the eruptions of Wapiti Lake flow and the Flat Mountain Rhyolite.

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34 491 The post-LCT record equally contains periods of prolonged quiescence, such as the ~180 ka
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36 492 quiescence between the Dunraven Road flow and the following South Biscuit Basin flow
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38 493 (Christiansen, 2007, Bindeman *et al.*, 2008). This similarity suggests that prolonged
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40 494 quiescence (>100 ka) could be a common feature in the effusive record between two large-
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42 495 volume eruptions.

43 496 However, temporal gaps may be a function of an incomplete record: Prior to the Mesa Falls
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45 497 eruption, only 6 small-volume lavas are known (5 post-HRT lavas and the earlier Snake
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47 498 River Butte lava, Christiansen, 2001). Between the MFT and LCT eruptions, around 12 lavas
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49 499 are known, with the uncertainty coming from the unknown number of lavas comprising the
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51 500 Lewis Canyon Rhyolite. Following the LCT, the youngest period of volcanism contains at
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53 501 least 40 separate lavas and small volume pyroclastic deposits. This suggests that either the
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7 502 frequency of lavas erupted from Yellowstone has increased by a factor of 4 after the eruption
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9 503 of the LCT, or, as we prefer, the younger lavas are much better preserved due to a lack of
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11 504 later caldera-forming eruptions. Such a preservation bias remains a challenge during
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13 505 investigations of ~~caldera~~ cycles at caldera volcanoes.
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16 507 **Petrogenesis of the IPMJ series**

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18 508 Our detailed mineral-scale characterisation of the IPMJ record provides important constraints
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20 509 for rhyolite generation in Yellowstone. Here we discuss our observations from geochemistry,
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22 510 petrology, and isotopic studies, which together provide a framework for rhyolite petrogenesis,
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24 511 within which the IPMJ series may be assessed.
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27 28 513 *Glomerocrysts as recorders of melt extraction*

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30 514 As ~~previously~~ noted ~~before~~, plagioclase and pyroxenes in the IPMJ rhyolites mostly occur
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32 515 together in glomerocrysts. Notably, quartz and sanidine are absent in these aggregates,
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34 516 despite being the most abundant phases within the overall mineral assemblage (compare table
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36 517 1). Being dominated by plagioclase and mafic phases with only limited groundmass
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38 518 (typically <20%), these crystal aggregates are necessarily less evolved in bulk composition
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40 519 than the lavas within which they are hosted (Fig. 11). The stark ~~incoherence-contrast~~ between
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42 520 the bulk compositions of crystal aggregates and the host magma has been interpreted as
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44 521 reflecting the extraction of rhyolitic liquid from these crystal aggregates, thus rendering them
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46 522 cumulates (e.g. Ellis *et al.*, 2014). The lack of sanidine and quartz in these aggregates implies
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48 523 that these minerals predominantly crystallised after melt was extracted, leaving behind a
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50 524 refractory cumulate of plagioclase, pyroxene and Fe-Ti oxides (Fig. 11, compare Ellis *et al.*,
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52 525 2014). The observation that sanidine and quartz occur as large, mostly euhedral to subhedral
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54 526 crystals, whereas smaller plagioclase and mafic minerals are often anhedral, supports this
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7 527 interpretation. We note that the majority of mafic minerals show trace element patterns
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9 528 consistent with growth from an evolved liquid (Fig. 8), with high contents in LREE and
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11 529 HREE and a pronounced Eu anomaly, similar to whole rock and glass REE patterns in IPMJ
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13 530 rhyolites (Fig. 2). The crystal aggregates therefore record a fractionation step between two

14 531 ~~rhyolitic melt~~ evolved magmas.

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16 532 In bulk, glass, and pyroxene compositions, all IPMJ rhyolites contain notable negative Eu
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18 533 anomalies (Figs. 2, 7). The rarity of plagioclase and the low Sr contents in the erupted liquids
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20 534 (tables 1, 2) suggest extensive removal of plagioclase at depth to decrease the bulk Sr content
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22 535 from the 190-680 ppm in the basalts (Doe *et al.*, 1982) to the 10-60 ppm observed in the
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24 536 IPMJ rhyolites. Similarly, pyroxene fractionation could account for the extremely low
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26 537 contents of MgO in IPMJ rhyolites in whole rock (0.01-0.12 wt% MgO, table 2) and glasses
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28 538 (0.03±0.02 wt% MgO).

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30 539 In the Yellowstone province, identical crystal aggregates have been reported for the young
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32 540 post-LCT rhyolites (Girard & Stix, 2009, Watts *et al.*, 2012), the Heise eruptive centre (Watts
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34 541 *et al.*, 2011), in the central Snake River Plain (Cathey & Nash, 2009, Ellis & Wolff, 2012,
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36 542 Ellis *et al.*, 2014) and in the Jarbidge rhyolite (Brueseke *et al.*, 2014). We note that in the
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38 543 Yellowstone system, both low- and normal- $\delta^{18}\text{O}$ rhyolites contain these crystal aggregates
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40 544 and both groups have identical major and trace elemental compositions (as noted already by
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42 545 Hildreth *et al.*, 1984), including low Sr contents. Thus, both rhyolite groups must have
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44 546 experienced large-scale plagioclase and pyroxene fractionation during petrogenesis and the
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46 547 observed glomerocrysts may reflect a common process of crystal-liquid separation in a mush
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48 548 zone (Vazquez & Reid, 2002, Stelten *et al.*, 2015).

49 549 If quartz grew primarily after extraction, the time between liquid extraction from the
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51 550 pyroxene-plagioclase-dominated mush zone and eruption of rhyolite magma is equal to the
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53 551 *minimum* time for quartz to grow without intermittent resorption stages. Based on quartz sizes
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7 552 and growth rates, Ellis *et al.* (2014) suggested that timescales of extraction were on the order
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9 553 of a few thousand years for rhyolites of the Snake River Plain. Timescales inferred in the
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11 554 same fashion for IPMJ rhyolites with maximum quartz sizes of 1.5-2.9 mm and growth rates
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13 555 of $\sim 10^{-13}$ to $\sim 10^{-14}$ m/s (Gualda *et al.*, 2012b) result in timescales of $\sim 1,000$ -9,000 years. Such
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15 556 estimates are in good agreement with longevities of crystal-poor magma of $\sim 10,000$ years
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17 557 based on zircon chronology (Rivera *et al.*, 2014a, Wotzlaw *et al.*, 2014) and up to 9,000 years
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19 558 from ^{238}U - ^{230}Th dating of zircon rims (Stelten *et al.*, 2015).
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21 559

22 560 *Compositional variability and time scales of mineral entrainment*

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24 561 The large inter-grain variability seen in mineral compositions and oxygen isotopes is a
25
26 562 pervasive feature in all IPMJ rhyolites. Previously, we noted the large variability in oxygen
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28 563 isotopic compositions in both quartz and sanidine, with MSWD for quartz often exceeding
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30 564 critical MSWD values expected for a single population (Fig. 9, see also quartz results
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32 565 section). Large compositional differences for trace elements in sanidines and quartz (Fig. 4)
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34 566 and their complex, un-systematic CL zoning patterns (Figs. 4, 5) suggest that different
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36 567 crystals in the same unit did not share common magmatic histories. Additionally, a
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38 568 comparison of trace element contents such as Ba in glass with those in rims of CL-imaged
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40 569 sanidine grains shows that a large number of grains did not grow in equilibrium with their
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42 570 host melt, even if a wide range of published $K_D(\text{San-melt})$ values is explored (Fig. 12).
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44 571 Mg-numbers in clinopyroxene are similar for a majority of grains in one unit. Deviating
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46 572 mineral compositions in a small number of grains identify entrainment products not in
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48 573 equilibrium with their host melt, such as rare more mafic pyroxenes (Fig. 7). These
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50 574 pyroxenes could represent remnant crystals from less evolved regions of the magmatic
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52 575 system as indicated by their lower content in LREE (Fig. 8). Their lack of a pronounced
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54 576 negative Eu anomaly suggests crystallisation prior to significant plagioclase fractionation
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7 577 | (compare Szymanowski *et al.*, 2015). A ~~xenocrystic~~-basaltic origin is unlikely for these high-
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9 578 Mg# pyroxenes as Yellowstone basalts do not contain pyroxene phenocrysts (Christiansen,
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11 579 2001).

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13 580 Orthopyroxenes and fayalites are commonly thought to replace one another (e.g. Davidson &
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15 581 Lindsley, 1989) and the abundant resorption features in fayalites suggest that conditions
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17 582 mostly favoured the formation of Fe-rich orthopyroxene, which is abundant in both the Heise
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19 583 and Yellowstone volcanic fields (e.g. Christiansen, 2001, Watts *et al.*, 2011). A mafic
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21 584 assemblage of clinopyroxene, orthopyroxene, fayalite and hornblende, as observed in some of
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23 585 these lavas (table 1), reflects a disequilibrium mineral assemblage that requires late
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25 586 entrainment of crystals from at least one of these mineral groups.

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27 587 Un-systematic zoning patterns, variable major and trace element contents and isotopic
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29 588 compositions, and a complex mafic mineral assemblage point out that crystals did not share a
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31 589 common magmatic history and that a large number of crystals probably did not grow in the
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33 590 melt in which they erupted. This late-stage entrainment seems to be particularly strong in
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35 591 units close to the MFT-related caldera, where chemically homogenous crystals of different
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37 592 compositions derived either from the mush or the overlying caldera roof are assembled in the
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39 593 same magma batch, whereas younger MJ-rhyolites contain less variable crystal populations
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41 594 (Fig. 12, Fig. 9 plus see comments on compositional variability between different grains of
42
43 595 the same unit in the respective results section on different mineral groups).

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45 596 The preservation of oxygen isotopic disequilibrium between quartz and sanidine crystals
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47 597 allows the timescales of material entrainment to be estimated. Diffusion rates for oxygen in
48
49 598 quartz are about four magnitudes faster than those for feldspars, and quartz crystals of 4 mm
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51 599 diameter would be expected to have fully equilibrated at magmatic temperature after
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53 600 thousands of years (Bindeman, 2008 and references therein). The internal complexities
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55 601 observed in CL images of the quartz grains suggest a potential for dissolution and re-

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7 602 precipitation processes, which can accelerate isotopic exchange. Thus, the O isotopic
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9 603 disequilibria indicate that quartz was added to the magma less than a few thousand years (at a
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11 604 maximum) prior to eruption.

12 605 The recognition of the subtly older sanidine crystals via $^{40}\text{Ar}/^{39}\text{Ar}$ single crystal
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14 606 geochronology also allows an estimation of the timescale of entrainment. Diffusion
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16 607 calculations suggest that ages of older crystals are indistinguishable from phenocrysts after
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18 608 residence times of 1-2 years in the magma (Gansecki *et al.*, 1996), with similar short
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20 609 preservation timescales reported from other studies (Renne *et al.*, 2012). The preservation of
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22 610 such antecrystic sanidine requires that addition occurs immediately prior to eruption. The
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24 611 same process causing tails of older ages in sanidine could well be responsible for
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26 612 compositional ranges in all mineral groups, as sanidine is unlikely to be the only antecrystic
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28 613 mineral sampled prior to, and during eruption.

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31 615 *Assimilation and crustal component*

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33 616 Recurring patterns of sudden drops in $\delta^{18}\text{O}$ following caldera collapse represent the major
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35 617 argument for bulk remelting of hydrothermally altered material (Bindeman & Valley, 2001).

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37 618 Extra-caldera normal- $\delta^{18}\text{O}$ rhyolites indicate that the processes responsible for the depletions
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39 619 are limited to the spatial extent of the caldera and directly relate to the caldera collapse. Even
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41 620 if magma is stored in a mush-dominated environment, the caldera collapse can be expected to

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43 621 bring hydrothermally altered lithologies to sufficient depths for partial or bulk remelting, thus
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45 622 providing a mechanism for additional entrainment of roof material that could account for the
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47 623 higher crystal contents (table 1) and larger compositional mineral-scale ranges observed in
48
49 624 the units spatially associated with the caldera.-

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51 625 In the same manner, assimilation and remelting of hydrothermally altered roof material
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53 626 following caldera collapse~~The same process~~ could be responsible for adding sufficient

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7 627 amounts of H₂O to the shallow parts of the magma reservoir to allow for crystallisation of
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9 628 hydrous amphibole and accessory biotite. While we do not attempt to speculate about the
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11 629 mechanisms of this process, it is noteworthy that hydrous minerals are seemingly limited to
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13 630 units close to the MFT-associated caldera. ~~Furthermore, the additional input of material could~~
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15 631 ~~account for the higher crystal contents (table 1) and larger compositional mineral scale ranges~~
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17 632 ~~observed in these units.~~

18 633 IPMJ-like depletions in $\delta^{18}\text{O}$ are permissive of remelting and incorporation of (1) high
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20 634 degrees of very slightly altered materials or (2) lower degrees of more strongly altered (lower
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22 635 $\delta^{18}\text{O}$) protoliths. For a simple estimate, we take the lowest value advocated by Watts *et al.*
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24 636 (2011) for a hydrothermally altered protolithic assimilate of -1‰ and a normal- $\delta^{18}\text{O}$ rhyolite
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26 637 of 6.3‰ (i.e. value expected from fractionation of mantle-derived basalt), inferring that the
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28 638 depletions in the IPMJ series (+5.0 to +6.1‰) can be explained by $\leq 25\%$ assimilation.
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30 639 However, the mechanistic processes of assimilating a hydrothermally altered and
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32 640 mineralogically complex protolith remain poorly understood. Given that hydrothermal
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34 641 processes produce a wide variety of alteration assemblages with bulk $\delta^{18}\text{O}$ spanning more
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36 642 than 10‰ and potentially being lower than the assumed value of -1 (Hildreth *et al.*, 1984),
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38 643 the degree of interaction with hydrothermally altered material might deviate strongly.

39 644 In terms of radiogenic isotopes, Pb isotopes from this study (Fig. 10) show little variability
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41 645 and are similar to those of the majority of Yellowstone rhyolites (Doe *et al.*, 1982). Even
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43 646 given the strong leverage provided by Wyoming craton and the fact that the low Sr contents
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45 647 of these rhyolites make them particularly susceptible to Sr contamination, bulk $^{87}\text{Sr}/^{86}\text{Sr}$
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47 648 values for IPMJ-rhyolites vary between 0.7090 and 0.7139, only slightly higher than the most
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49 649 radiogenic Yellowstone basalts of 0.7089 (Doe *et al.*, 1982, Hildreth *et al.*, 1991). Thus, the
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51 650 rhyolites of the IPMJ series do not require significant contamination from crustal lithologies
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7 651 to pass from basalt to rhyolite and small variations likely reflect minor regional differences in
8 652 the magmatic source region.

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11 12 654 **A model for the Yellowstone magmatic system between 1.3 and 0.6 Ma**

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14 655 In the context of the parameters discussed above, we focus here on a coherent model for the
15 656 evolution of the Yellowstone magmatic system during the IPMJ period of activity. We
16 657 propose that during IPMJ time, the silicic magmas of the Yellowstone magmatic system are
17 658 derived from small-scale magma batches residing in a long-lived crystal-rich, mushy
18 659 reservoir with the crystal aggregates representing direct samples of this mushy refractory
19 660 material. These aggregates (Fig. 11) contain mostly plagioclase, pyroxene and Fe-Ti oxides,
20 661 suggesting that these phases dominate the mush. Aggregates from different lavas have subtly
21 662 different mineral compositions being identical to the compositions of the single crystals in
22 663 that lava, suggesting that some single crystals may represent disaggregated cumulate
23 664 material. If silicic lavas are a true reflection of the magma compositions beneath their
24 665 eruption sites, the widely dispersed, synchronous eruption of compositionally almost
25 666 indistinguishable rhyolite lavas (table 2, Fig. 2) suggests the presence of a large-volume
26 667 mushy source region. Small-scale compositional differences as observed in isotopes and trace
27 668 elements indicate that the composition of the mush varied both laterally and with time,
28 669 suggesting that the reservoir was incrementally built and subject to changes in the focus of
29 670 main activity over time.

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45 671 To investigate the conditions at which the erupted magmas were stored prior to eruption, we
46 672 performed simulations using rhyolite-MELTS (Gualda *et al.*, 2012a). Even though the
47 673 complex mineral assemblage requires some degree of disequilibrium, we can use equilibrium
48 674 magmatic modelling to approximate the last stages of crystallisation before eruption because
49 675 (1) recycling mainly focused on compositionally similar material, (2) entrainment of
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7 676 ~~antecrystic mineral phases (disequilibrium on trace elemental or isotopic scale) does not~~
8 677 ~~preclude them from being stable in the melt, and (3) true disequilibrium assemblages such as~~
9 678 ~~the mafic minerals make up a small fraction of the magma (table 1). In terms of mineral~~
10 679 ~~abundances. We choose Elk Butte as a starting material for rhyolite-MELTS calculations~~
11 680 ~~since Elk Butte represents an average typical MJIP member in terms of mineral compositions~~
12 681 ~~and does not contain any non-eruption age sanidine, therefore was not significantly affected~~
13 682 ~~by crystal contamination, and is therefore chosen as a starting material for r-MELTS~~
14 683 ~~calculations. As input parameters, we use a variety of combinations of pressure (0.5-5.0 kbar)~~
15 684 ~~and water contents (0-5 wt% H₂O) at an oxygen fugacity of QFM (supported by the~~
16 685 ~~observation of fayalite, quartz and magnetite in these units and in agreement with~~
17 686 ~~experimental data from Bolte *et al.* (2015)). Temperature input ranges start at the calculated~~
18 687 ~~liquidus of the assemblage and stop at 670°C with steps of 1°C. We define the solidus as the~~
19 688 ~~highest temperature at which r-MELTS rhyolite-MELTS returns <1% melt.~~
20 689 ~~As described above, the mineral record in MJIP rhyolites may be complex, so we focussed on~~
21 690 ~~a few important parameters in order to identify likely storage conditions: (1) the crystallinity~~
22 691 ~~of the r-MELTS rhyolite-MELTS run should be within the range of the IPMJ lavas (60-87%~~
23 692 ~~melt, table 1); (2) the composition of the melt should roughly correlate with results~~
24 693 ~~from correspond to glass analyses compositions for the same unit; (3) the composition of~~
25 694 ~~sanidine predicted by r-MELTS rhyolite-MELTS should be similar to sanidine compositions~~
26 695 ~~in the same unit (Or_{57,61}). These parameters are approximated at conditions of ~870°C, ~0.8-1~~
27 696 ~~kbar and ~2.38-3.3 wt% H₂O (see figure and additional details in supplementary file).~~
28 697 ~~At these conditions, the mineralogy of the r-MELTS rhyolite-MELTS output is dominated by~~
29 698 ~~quartz and sanidine with minor plagioclase, consistent with these phases being dominant in~~
30 699 ~~the mineral assemblage (table 1). The only other crystallising phase is magnetite, and the~~
31 700 ~~identified conditions border to the stability field of orthopyroxene. The fact that these~~

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701 ~~composition and conditions~~ cannot reproduce the mafic mineral assemblage is consistent with our
 702 ~~interpretation of them originating mostly from cumulate the glomerocrysts as cumulates,~~
 703 ~~which originate from record~~ an earlier petrogenetic stage (see section “Glomerocrysts as
 704 ~~recorders of melt extraction”).~~
 705 A magmatic temperature of 870°C ~~correlates agrees extremely well~~ the temperature inferred
 706 ~~from sanidine-melt geothermometry (~870°C for Elk Butte, supplementary material),~~
 707 ~~suggesting that thermometry on late crystallizing sanidine and melt pairs may result in the~~
 708 ~~most reliable temperature estimates. In terms of water content, the results are slightly higher~~
 709 ~~than previous estimates for rhyolites of the province based on re-homogenised melt~~
 710 ~~inclusions (1.5-2.0 wt.%, Befus & Gardner, 2016) and experimental petrology (1.5-2.5 wt%,~~
 711 ~~Bolte *et al.*, 2015) and lower than our estimates based on plagioclase-melt hygrometry~~
 712 ~~(3.6±1.1 to 5.7±1.1 wt% H₂O, see supplementary material). The latter indicates that~~
 713 ~~plagioclase-melt pairs should be treated with caution due to the longer crystallization range of~~
 714 ~~plagioclase compared to sanidine. Neither significant biotite nor amphibole crystallisation is~~
 715 ~~observed at the identified storage conditions or any other explored combination of pressure,~~
 716 ~~temperature and water content, suggesting that crystallisation of hydrous mafic minerals is~~
 717 ~~driven by a combination of less evolved melt composition at an earlier petrogenetic stage~~
 718 ~~and/or higher water contents, or that rhyolite-MELTS is not well-calibrated for these phases at~~
 719 ~~an earlier petrogenetic stage.~~
 720 The pressure estimate of 1 kbar translates to a depth of approximately 4 km given the
 721 relatively low density of the crust comprised of silicic volcanics and hydrothermally altered
 722 materials. ~~Similar pressures ranging between 0.5 and 1.5 kbar have been reported in other~~
 723 ~~studies (e.g. Befus & Gardner, 2016), albeit for the younger Yellowstone volcanism. These~~
 724 ~~pressure estimates is are~~ slightly lower than the upper reaches of the geophysically imaged
 725 ~~body beneath Yellowstone~~ (8-18 km, Husen *et al.*, 2004, Lowenstern *et al.*, 2006, Chang *et*

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7 726 *al.*, 2007, DeNosaquo *et al.*, 2009, Chu *et al.*, 2010, Farrell *et al.*, 2014) ~~Husen *et al.*, 2004,~~
8 ~~Lowenstern *et al.*, 2006, Chang *et al.*, 2007, DeNosaquo *et al.*, 2009, Chu *et al.*, 2010, Farrell *et*
9 ~~al.~~, 2014), where the eruptible magmas would collect prior to eruption (e.g. Bachmann &
10 728 Bergantz, 2004). ~~Similar pressures ranging between 0.5 and 1.5 kbar have been reported in~~
11 ~~other studies (e.g. Befus & Gardner, 2016), albeit for the younger Yellowstone volcanism.~~~~

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16 731 Geophysical studies have long been able to detect a large low- v_p body beneath the
17 732 Yellowstone volcanic field. The proportion of crystal-poor, eruptible magma has been
18 733 estimated at ~15% volume via gravity measurements (Krukoski, 2002), $\leq 30\%$ volume via P-
19 734 wave arrival times (Chu *et al.*, 2010), and $\leq 35\%$ volume via strain measurements from seiche
20 735 waves on Yellowstone lake (Luttrell *et al.*, 2013). ~~Importantly, some of the estimates above~~
21 ~~are dependent upon the degree of water saturation of the magma. The occurrence of biotite~~
22 ~~and amphibole in the IPMJ mineral assemblage indicates higher water contents in these units~~
23 ~~than often assumed for typical, dry Yellowstone rhyolites.~~ Regardless of the method of study
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32 739 or the proportion of crystal-poor melt estimated, geophysical studies overwhelmingly support
33 740 the occurrence of a crystal mush currently underlying Yellowstone. While this does not
34 741 necessarily show that such a situation occurred during IPMJ times, we have no reason to
35 742 believe that the volcanic system then behaved in a fundamentally different way to today.
36 743 Based on the age clustering of IPMJ rhyolites, magmatic activity occurred in discrete pulses
37 744 as is clear in the post-LCT record. Reactivation of the magmatic system probably occurred
38 745 due to heat transfer following recharge at larger depths (termed 'defrosting' by Mahood
39 746 (1990), compare Wolff *et al.* (2015)) and is recorded in late rims in some sanidine and quartz
40 747 crystals. The reverse zonation in Ba in some sanidine crystals coupled with other trace
41 748 elements (Sr, Rb, Eu) suggests that these rims result from melting of sanidine and plagioclase
42 749 from the cumulate pile or crystallised from less evolved rhyolite (e.g. 960 ppm Ba in Biscuit
43 750 Basin flow (Christiansen, 2001) with $K_D(\text{San-melt}) \approx 22$ (Leeman & Phelps, 1981) would
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7 751 allow crystallisation of sanidine with up to 21,000 ppm Ba). Significant mass transfer from
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9 752 replenishing magma is not favoured due to the homogeneity in Pb isotopes (Fig. 10). Crystal
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11 753 aggregates generally do not contain sanidine, thus the crystal aggregates could record the
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13 754 refractory part of the mushy reservoir after remelting of sanidine and plagioclase (Fig. 12).

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15 755 As noted above, the IPMJ rhyolites show a complex mineralogical record with entrainment of
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17 756 various mineral phases. Within the constraints of the preserved lavas, there appears to be little
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19 757 obvious compositional evolution within the magmatic system during IPMJ times, with the
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21 758 variability occurring in a mineral group in a single unit as large as that observed through
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23 759 time. Although the higher variability in our dataset compared to previous studies could be
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25 760 due to a larger number of analyses, this variation should not be large enough to obscure
26
27 761 possible trends, if these had the same magnitude as the clear trend towards cooler
28
29 762 temperatures and more juvenile isotopic compositions observed in the post-LCT lavas
30
31 763 (compare Vazquez *et al.*, 2009). The fact that there is neither a continuous evolution through
32
33 764 time nor sudden compositional changes that could hint at the forthcoming super eruption
34
35 765 suggests that (1) reactivation of the magmatic system occurs on a faster timescale than can be
36
37 766 recorded in most mineral compositions or (2) a significant preservation bias hinders
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39 767 conclusions on possible eruption triggers, and changes might be preserved in lavas erupted
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41 768 between the Mt. Haynes Rhyolite and LCT, which are now hidden under younger volcanic
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43 769 deposits or were removed by the LCT eruption.

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45 770 The limited depletion in $\delta^{18}\text{O}$ observed in the IPMJ rhyolites does not require high degrees of
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47 771 remelting of hydrothermally altered precursors, as is observed for post-LCT low- $\delta^{18}\text{O}$
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49 772 rhyolites with diverse phenocrysts. However, a number of similarities between the IPMJ
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51 773 series and the later low- $\delta^{18}\text{O}$ rhyolites hint that the processes involved in petrogenesis may
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53 774 not be so distinct. For example, within the IPMJ series the lowest $\delta^{18}\text{O}$ magmas and those
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55 775 with the highest proportion of hydrous mineral phases are those spatially associated with the
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7 776 MFT-related caldera. This association supports, on a smaller scale, the cannibalisation
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9 777 processes proposed for the youngest Yellowstone lavas. Understanding these processes fully,
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11 778 however, still requires a physical model for the melting of a mineralogically complex
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13 779 protolith with heterogeneous isotopic compositions.

14 780 Our combined data supports a petrogenesis in an incrementally-built mushy magma reservoir
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16 781 with compositional variations in space and time, similar to recent models of Vazquez *et al.*
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18 782 (2009) and Stelten *et al.* (2015) for similarly less depleted post-LCT rhyolites. While we do
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20 783 not find the same thermochemical trends indicating a cooling and evolving reservoir through
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22 784 time as is observed for the short-lived post-LCT rhyolites, we note that the compositional
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24 785 evolution described in these studies was revealed from 7 lavas erupted within the space of
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26 786 130 ka. Only among the IP domes ~~do~~, we find a similar density of deposits, and only the Mt.
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28 787 Haynes Rhyolite erupted within 130 ka before the LCT. Thus, the potential for direct
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30 788 comparison of the IPMJ period with the post-LCT record remains limited and potential
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32 789 precursory geochemical signals remain elusive in the IPMJ-record.

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35 791 **CONCLUSIONS**

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37 792 Our detailed mineralogical and geochemical study of the IPMJ rhyolites provides new
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39 793 insights into rhyolite generation in Yellowstone, complementing previous studies that
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41 794 focused on the youngest Yellowstone rhyolites. Multiple lines of evidence support rhyolite
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43 795 genesis in an incrementally-built upper crustal mush zone with remobilisation of
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45 796 hydrothermally altered material occurring in areas affected by caldera collapse (Fig. 13). The
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47 797 striking geochemical similarity of all Yellowstone rhyolites suggests fundamentally similar
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49 798 processes control their generation in the shallow parts of a crystal mush.

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51 799 New ages via $^{40}\text{Ar}/^{39}\text{Ar}$ -dating of sanidine reveal that the magmatic series IP domes and MJ
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53 800 rhyolites overlap by at least 66 ka. The decreased eruption frequency in the IPMJ record
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7 801 suggests that large parts of the depositional record are covered by younger lava flows,
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9 802 prohibiting a high-enough time resolution to decipher potential precursory activity before
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11 803 eruption of the LCT, if time scales allow minerals to record such processes.

12 804 Detailed investigation of the mineralogy of the IPMJ rhyolites indicates a complex
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14 805 assemblage of minerals within the lavas, particularly in the mafic mineral assemblage.
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16 806 Disequilibrium in O isotopic compositions and the preservation of subtly older single crystal
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18 807 $^{40}\text{Ar}/^{39}\text{Ar}$ ages in sanidines in most Yellowstone rhyolites are consistent with late stage
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20 808 entrainment of compositionally similar material. We infer that the processes giving rise to the
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22 809 variability in sanidine also generate the disequilibrium seen in other minerals. Since there is
23
24 810 no correlation between the amount of entrained material and the eruption style, this late stage
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26 811 contamination likely occurs in shallow, cooler portions of the crystal mush and is enhanced in
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28 812 the area of the preceding caldera collapse. Here, mushy roof material could be brought to
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30 813 sufficient depths to be remobilised and recycled. Additional research should be undertaken to
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32 814 assess the preservation potential of antecrystic sanidine in mush-like conditions, with or
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34 815 without partial re-equilibration.

35 816 An important conclusion of this work is that trace elements and isotopes suggest that
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37 817 significant proportions of the crystals found within the IPMJ suite of lavas did not crystallise
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39 818 from the magma with which they erupted, even if the same mineral phase can be stabilised in
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41 819 the melt according to equilibrium magma modelling via rhyolite-MELTS. This appreciation
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43 820 is particularly relevant for the derivation of magmatic parameters and in-situ studies (e.g.
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45 821 melt inclusion or diffusion timescale work), which necessarily investigate a limited number
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47 822 of grains due to the time-consuming nature of the work. In such cases, conclusions derived
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49 823 from a small number of grains must be exported to whole magmatic systems with care.

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51 824 Plagioclase-pyroxene-dominated glomerocrysts represent direct petrological samples from an
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53 825 underlying crystal-rich mushy zone at depth (Fig. 11). Such crystal aggregates occur in both
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7 826 low- $\delta^{18}\text{O}$ rhyolites such as the Upper Basin Member rhyolites (Girard & Stix, 2009, Watts *et*
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9 827 *al.*, 2012) and the mildly depleted rhyolites of the IPMJ series, and thus represent a common
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11 828 part of rhyolite petrogenesis. Depletions in Sr and Eu in whole rock and glass data support
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13 829 significant plagioclase fractionation at depth, coherent with the scarcity of glomerocryst-
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15 830 forming minerals in hand specimen. While plagioclase and mafic minerals remain trapped at
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17 831 depth and are rarely observed in the erupted magmas, subhedral to euhedral sanidine and
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19 832 quartz grow after extraction of the melt. The timescales for liquid extraction from an upper
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21 833 crustal mush zone as inferred from quartz growth are on the order of 10 ky, consistent with
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23 834 timescales from zircon chronology for both effusive and explosive eruptions in Yellowstone
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25 835 (Rivera *et al.*, 2014a, Wotzlav *et al.*, 2014, Stelten *et al.*, 2015).

26 836 Mafic minerals containing lower amounts of LREE and lacking a prominent Eu anomaly
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28 837 could represent glimpses into deeper, less evolved parts of the magmatic system, which are
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30 838 usually not sampled due to the sheer size of the magmatic reservoir (estimated today to be as
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32 839 large as c. 4000 km³ with 5-15% melt, Farrell *et al.* (2014)). The fact that similar lavas are
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34 840 recurrently produced throughout the entire magmatic history of the volcanic field supports
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36 841 magma genesis in a coherent crystal mush via several melt-rich magma batches that may
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38 842 erupt independently of each other.

39 843

41 844 **ACKNOWLEDGEMENTS**

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44
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48
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50
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Figure captions

Fig. 1: Distribution of IPMJ rhyolites in the Yellowstone area after geological map of Christiansen (2001). IP domes are shown in purple, MJ lavas in green and the Lewis Canyon Rhyolite in dark green. Dashed red lines indicate approximate caldera outlines related to eruption of HRT (I), MFT (II) and LCT (III). Black dashed line shows borders of Yellowstone National Park. Upper left inset shows unit abbreviations and affiliation of units with either the Island Park or Mount Jackson series. Upper right inset shows location of the Yellowstone volcanic province within North America, grey blobs mark eruptive centres along the Yellowstone hotspot track.

Fig. 2: REE contents normalised to BSE composition (McDonough & Sun, 1995) for bulk rock and glasses (grey area), where units contain glass (compare table 1). Note the strong

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7 1087 depletions in Ba, Sr and Eu, suggestive of extensive plagioclase fractionation. Depletions for
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9 1088 glass are more pronounced due to sanidine and plagioclase in whole rock.

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11 1089
12 1090 **Fig. 3:** Spatial distribution and appearance of IPMJ rhyolites based on $^{40}\text{Ar}/^{39}\text{Ar}$ -eruption
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14 1091 ages determined on sanidine. Coloured regions connect units from the same syn-eruptive
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16 1092 cluster as marked with a frame in bottom panel. Bottom panel shows eruption age as stated in
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18 1093 table 3 with error bars marking the full external precision (2 sigma, see text), noted in grey
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20 1094 are the two rhyolites not covered in this study: FM = Flat Mountain Rhyolite at c. 0.93 Ma K-
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22 1095 Ar age from Obradovich (1992), BB = Big Bear Lake flow (undated, probable age based on
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24 1096 stratigraphic relationships as recommended by Christiansen *et al.* (2001)). Inset diagrams
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26 1097 shows frequency distributions for some typical units, with the black solid line marking the
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28 1098 distribution of eruption age grains note the preservation of non-eruptive ages and the black
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30 1099 dashed line the preservation of non-eruptive ages.

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34 1101 **Fig. 4:** Cathodoluminescence images of quartz crystals sorted by eruption age (top left to
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36 1102 bottom right young to old). Titanium concentrations (ppm) measured by LA-ICPMS are
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38 1103 listed next to analysis locations (white circles). Bottom panel shows relative probability
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40 1104 distributions of Ti concentrations in quartz grains from each unit and the number of analyses.
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42 1105 Note the higher Ti contents in the Lewis Canyon Rhyolite compared to other MJ lavas.

43 1106
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45 1107 **Fig. 5:** CL images of typical sanidine grains with locations of trace element analysis; An-Ab-
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47 1108 Or ternary plot of sanidines and plagioclase from the IPMJ rhyolites, grey areas mark
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49 1109 compositions of other Yellowstone rhyolites (Hildreth *et al.*, 1984, Gansecki *et al.*, 1996).

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7 1111 **Fig. 6:** Examples of mafic minerals in back-scatter images (from upper left to bottom right):
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9 1112 1) exsolved magnetite, 2) clinopyroxene with zircon and apatite inclusions, 3) resorbed
10 1113 fayalite, 4) subhedral Fe-hornblende, 5) subhedral orthopyroxene with Fe-Ti inclusions, 6)
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12 1114 “fresh” anhedral inclusion-rich fayalite, 7) clinopyroxene overgrowth of orthopyroxene, 8)
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14 1115 euhedral fresh clinopyroxene, 9) resorbed fayalite with cracks.
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18 1117 **Fig. 7:** Pyroxene compositions from the MJ rhyolites (left) and the IP domes (right)
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20 1118 illustrating the wide compositional range between pyroxenes within the same unit.
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24 1120 **Fig. 8:** REE contents in mafic minerals normalised to BSE composition (McDonough & Sun,
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26 1121 1995). Dark grey fields show REE patterns of the majority of crystals, likely to have grown in
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28 1122 a highly evolved rhyolitic melt. Black lines indicate REE contents of crystals deviating from
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30 1123 the main trend. Crystals depleted in LREE and lacking characteristic Eu anomaly crystallised
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32 1124 in less evolved magma prior to significant plagioclase fractionation. ~~Non-parallel patterns~~
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34 1125 ~~suggest that parts of the mineral fraction are not co-genetic emphasising the complexity in the~~
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36 1126 ~~crystal heritage.~~
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40 1128 **Fig. 9:** Oxygen isotopes $\delta^{18}\text{O}$ in ‰ VSMOW for explosive end members MFT and LCT and
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42 1129 effusive domes and lavas in-between (abbreviations and colours as in Fig. 1). Wide boxes
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44 1130 indicate probable $\delta^{18}\text{O}_{\text{melt}}$ values calculated from average quartz compositions for each unit.
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46 1131 Grey background bar illustrates range of “normal”- $\delta^{18}\text{O}$ rhyolites (expected values by
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48 1132 fractionation of MOR and arc basalts, Bindeman (2008) and references therein). Upper panel
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50 1133 defines the MSWD of the quartz analyses. $\text{MSWD} > \text{MSWD}_{\text{critical}}$ indicates that the results do
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52 1134 not represent a single population. Lower panel represents trends in $\delta^{18}\text{O}$ throughout the
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7 1135 history of the Yellowstone volcanic field (after Bindeman and Valley, 2001; Hildreth et al.
8 1136 1984).

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12 1138 **Fig. 10:** Pb isotopes in sanidine in IPMJ rhyolites. Inlet diagram shows the entire range of Pb
13 1139 isotopic compositions for the entire Yellowstone volcanic field with red-framed box showing
14 1140 the extent of main diagram. Data for IPMJ rhyolites from this study, MFT from Ellis et al.
15 1141 (unpublished data), other Yellowstone rhyolites and basalts from Doe et al. (1982); Watts et
16 1142 al. (2012).

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23 1144 **Fig. 11:** Example of glomerocrysts from different Island Park domes. Comparison of mineral
24 1145 assemblage within the aggregate with free single crystals yields no significant compositional
25 1146 differences. Bulk glomerocryst compositions are less evolved compared to typical whole rock
26 1147 modal compositions due to the lack of late-forming sanidine and quartz.

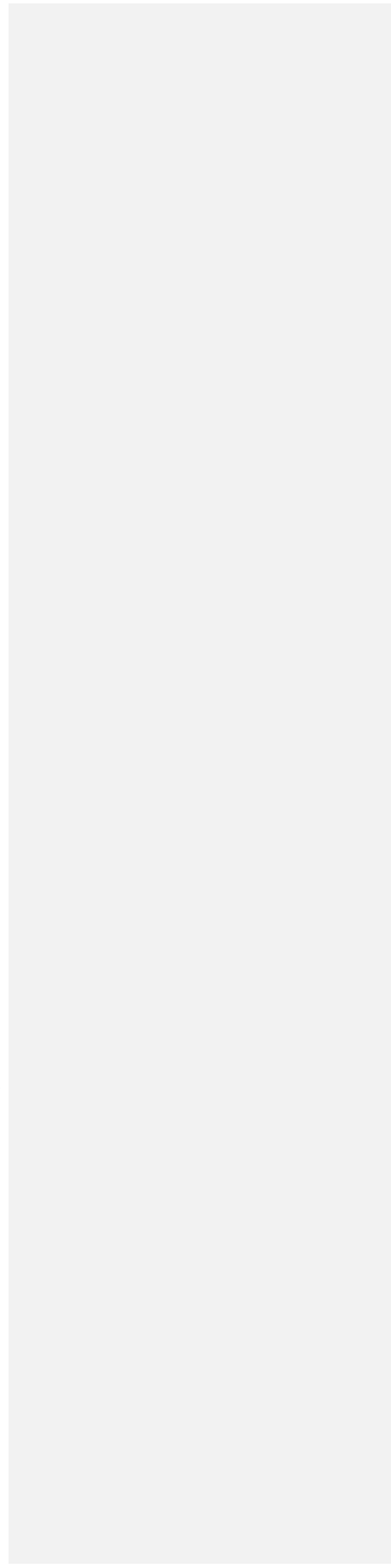
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33 1149 **Fig. 12:** (A) Ba contents in sanidine rims and cores do not fall into the calculated equilibrium
34 1150 range expected from Ba contents in glass from the same unit. Equilibrium range calculated
35 1151 with $K_D(\text{San-melt}) \approx 22$ in rhyolite from Leeman and Phelps (1981) and $K_D(\text{San-melt}) \approx 6.7$ in
36 1152 high-silica rhyolite (Mahood & Hildreth, 1983). Figure contains only data for those sanidine
37 1153 crystals that were imaged via CL. (B) Average Rb/Sr in San for each unit with error bars
38 1154 indicating one standard deviation. Numbers in italic correspond to number of analyses per
39 1155 unit. Note the large variability for IP domes and Moose Creek (MC) compared to MJ flows,
40 1156 indicating higher entrainment component for units close to MFT-related caldera. (C) Ba ratio
41 1157 between rim and core and Sr ratio respectively suggest that reverse zoning in San is due to
42 1158 cumulate remelting or crystallisation from less evolved rhyolite melt. Figure contains only
43 1159 data for those sanidine crystals that were imaged via CL.

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1161 **Fig. 13:** Schematic summary diagram of petrogenetic processes and characteristics of the
1162 IPMJ rhyolite series.

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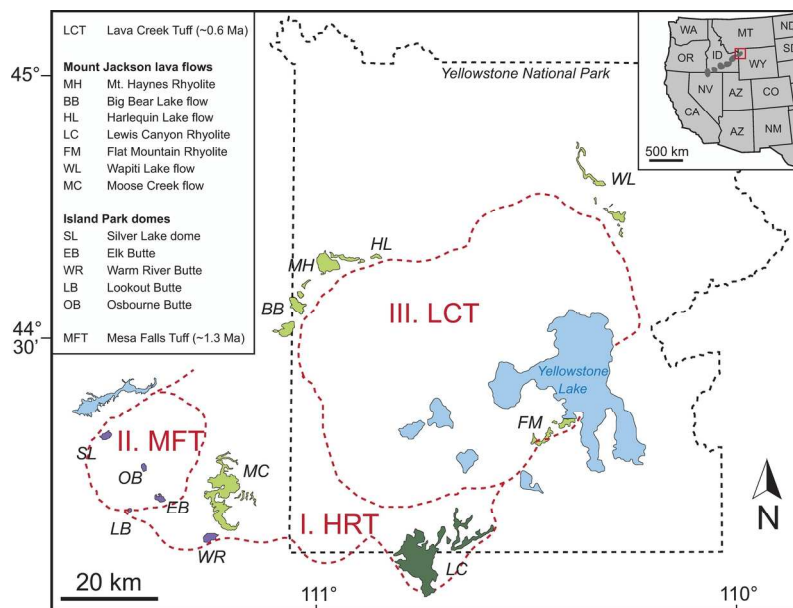


Fig. 1: Distribution of IPMJ rhyolites in the Yellowstone area after geological map of Christiansen (2001). IP domes are shown in purple, MJ lavas in green and the Lewis Canyon Rhyolite in dark green. Dashed red lines indicate approximate caldera outlines related to eruption of HRT (I), MFT (II) and LCT (III). Black dashed line shows borders of Yellowstone National Park. Upper left inset shows unit abbreviations and affiliation of units with either the Island Park or Mount Jackson series. Upper right inset shows location of the Yellowstone volcanic province within North America, grey blobs mark eruptive centres along the Yellowstone hotspot track.

Fig. 1
175x105mm (300 x 300 DPI)

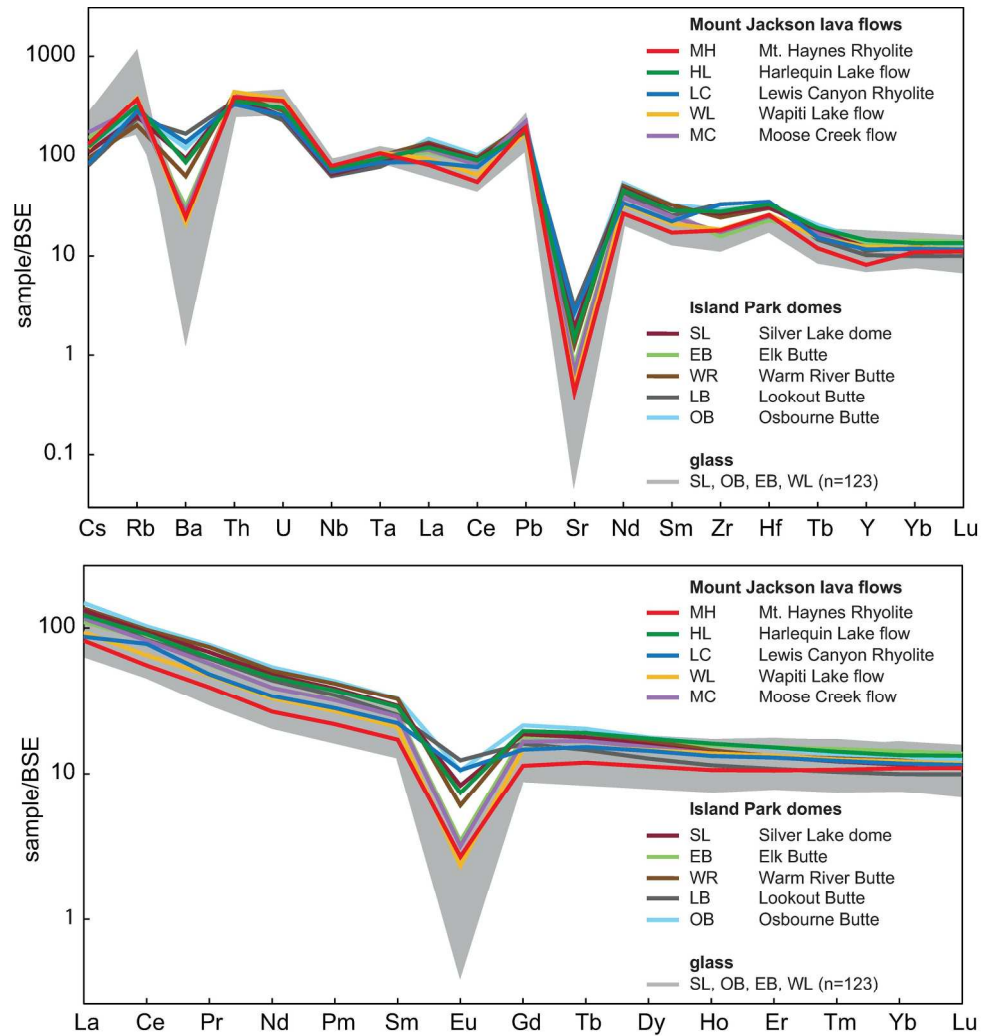


Fig. 2: REE contents normalised to BSE composition (McDonough & Sun, 1995) for bulk rock and glasses (grey area), where units contain glass (compare table 1). Note the strong depletions in Ba, Sr and Eu, suggestive of extensive plagioclase fractionation. Depletions for glass are more pronounced due to sanidine and plagioclase in whole rock.

Fig. 2

198x213mm (300 x 300 DPI)

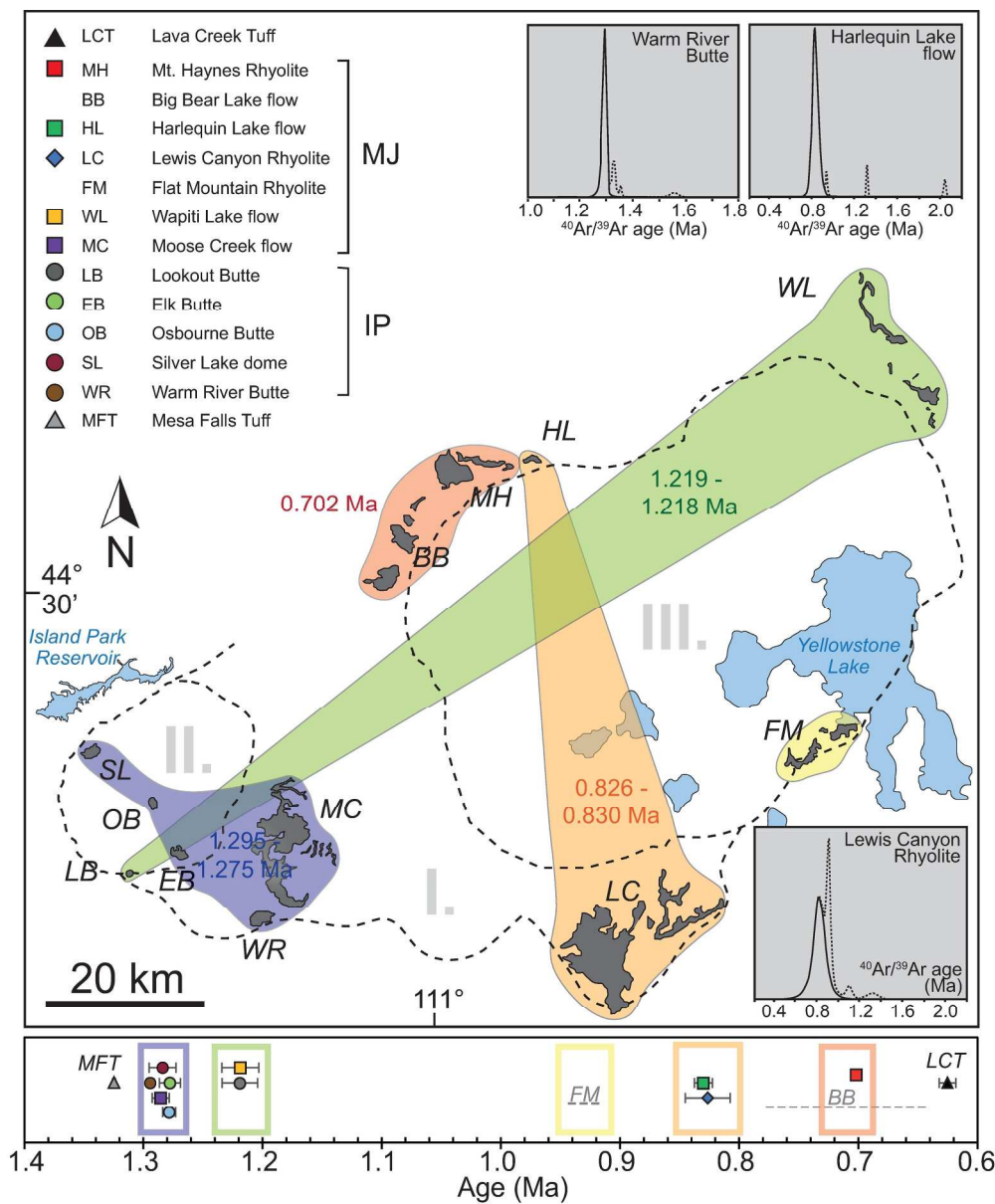


Fig. 3: Spatial distribution and appearance of IPMJ rhyolites based on $^{40}\text{Ar}/^{39}\text{Ar}$ -eruption ages determined on sanidine. Coloured regions connect units from the same syn-eruptive cluster as marked with a frame in bottom panel. Bottom panel shows eruption age as stated in table 3 with error bars marking the full external precision (2 sigma, see text), noted in grey are the two rhyolites not covered in this study: FM = Flat Mountain Rhyolite at c. 0.93 Ma K-Ar age from Obradovich (1992), BB = Big Bear Lake flow (undated, probable age based on stratigraphic relationships as recommended by Christiansen et al. (2001)). Inset diagrams show frequency distributions for some typical units, with the black solid line marking the distribution of eruption age grains and the black dashed line the preservation of non-eruptive ages.

Fig. 3

216x261mm (300 x 300 DPI)

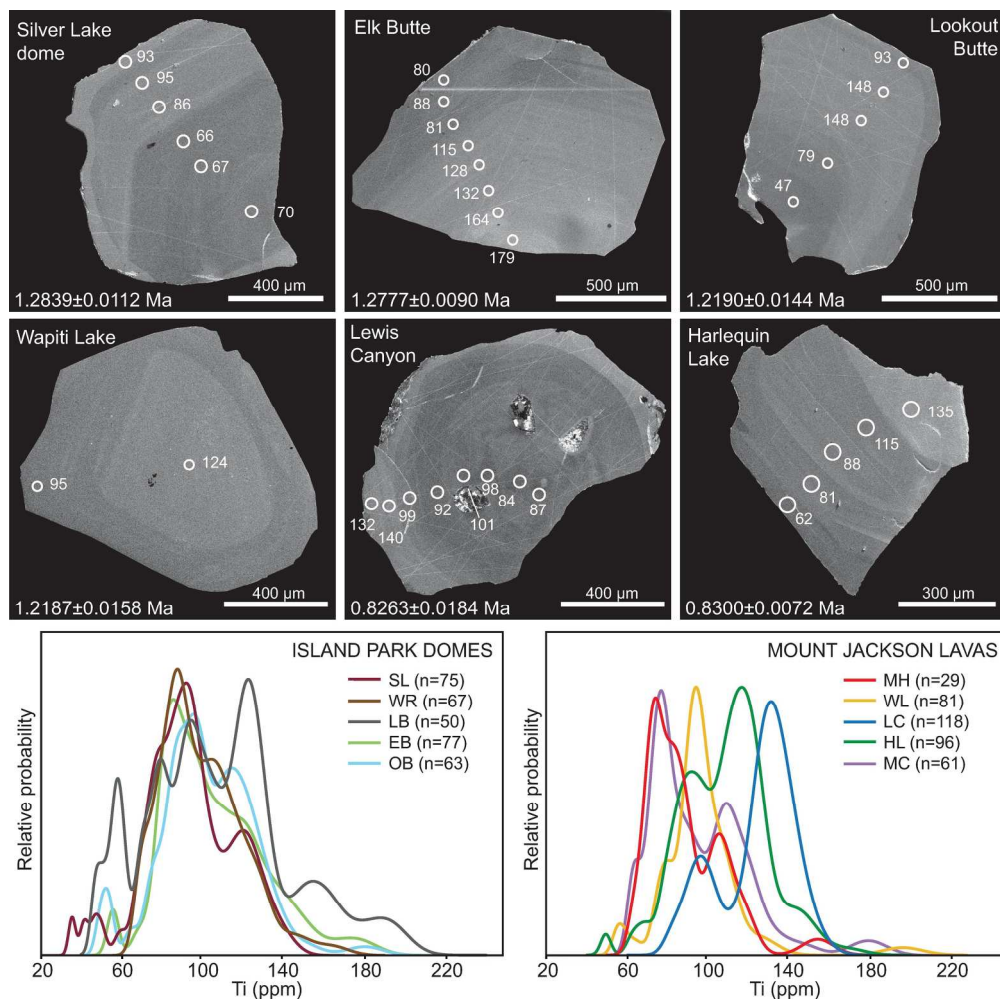


Fig. 4: Cathodoluminescence images of quartz crystals sorted by eruption age (top left to bottom right young to old). Titanium concentrations (ppm) measured by LA-ICPMS are listed next to analysis locations (white circles). Bottom panel shows relative probability distributions of Ti concentrations in quartz grains from each unit and the number of analyses. Note the higher Ti contents in the Lewis Canyon Rhyolite compared to other MJ lavas.

Fig. 4

199x197mm (300 x 300 DPI)

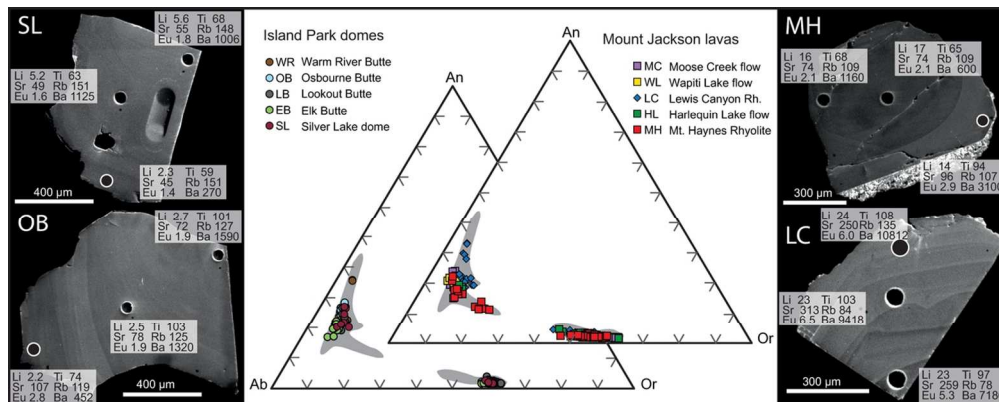


Fig. 5: CL images of typical sanidine grains with locations of trace element analysis; An-Ab-Or ternary plot of sanidines and plagioclase from the IPMJ rhyolites, grey areas mark compositions of other Yellowstone rhyolites (Hildreth et al., 1984, Gansecki et al., 1996).

Fig. 5

116x49mm (300 x 300 DPI)

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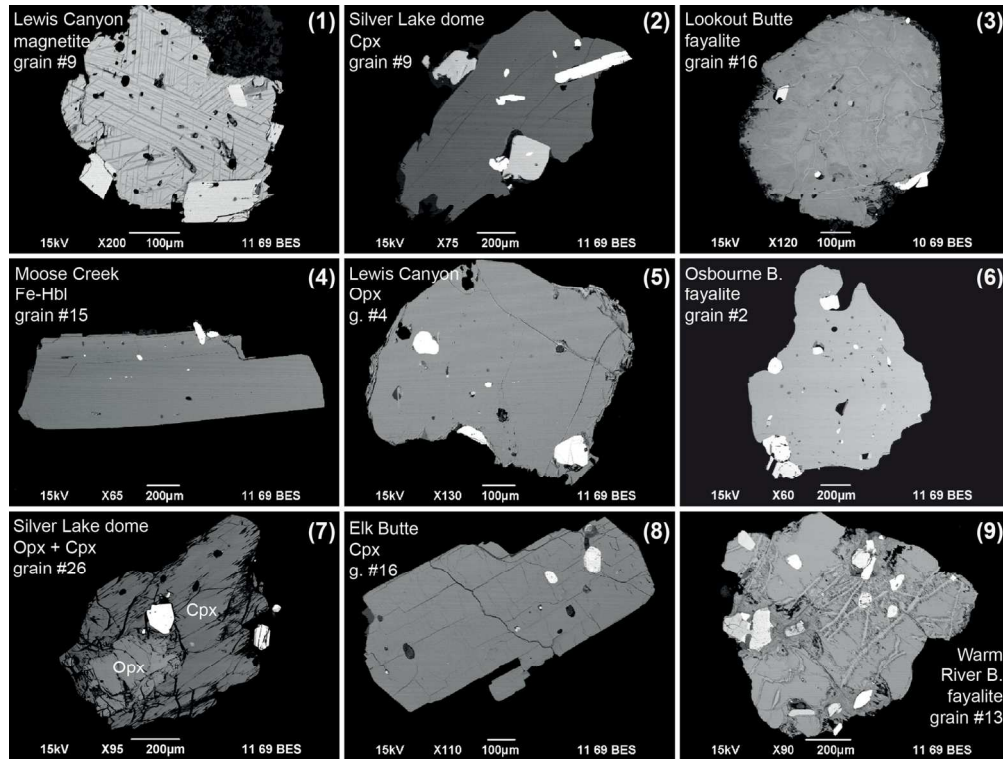


Fig. 6: Examples of mafic minerals in back-scatter images (from upper left to bottom right): 1) exsolved magnetite, 2) clinopyroxene with zircon and apatite inclusions, 3) resorbed fayalite, 4) subhedral Fe-hornblende, 5) subhedral orthopyroxene with Fe-Ti inclusions, 6) "fresh" anhedral inclusion-rich fayalite, 7) clinopyroxene overgrowth of orthopyroxene, 8) euhedral fresh clinopyroxene, 9) resorbed fayalite with cracks.

Fig. 6

152x114mm (300 x 300 DPI)

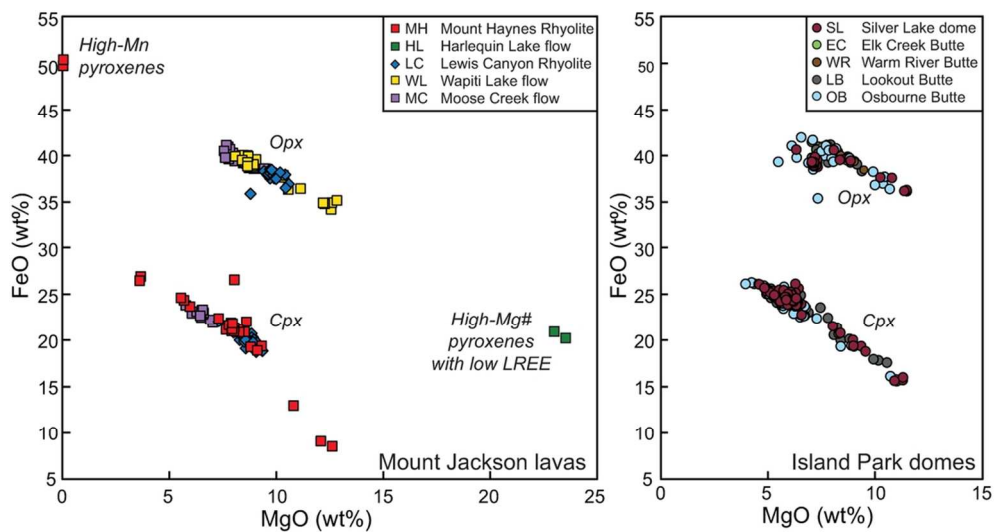


Fig. 7: Pyroxene compositions from the MJ rhyolites (left) and the IP domes (right) illustrating the wide compositional range between pyroxenes within the same unit.

Fig. 7
 99x51mm (300 x 300 DPI)

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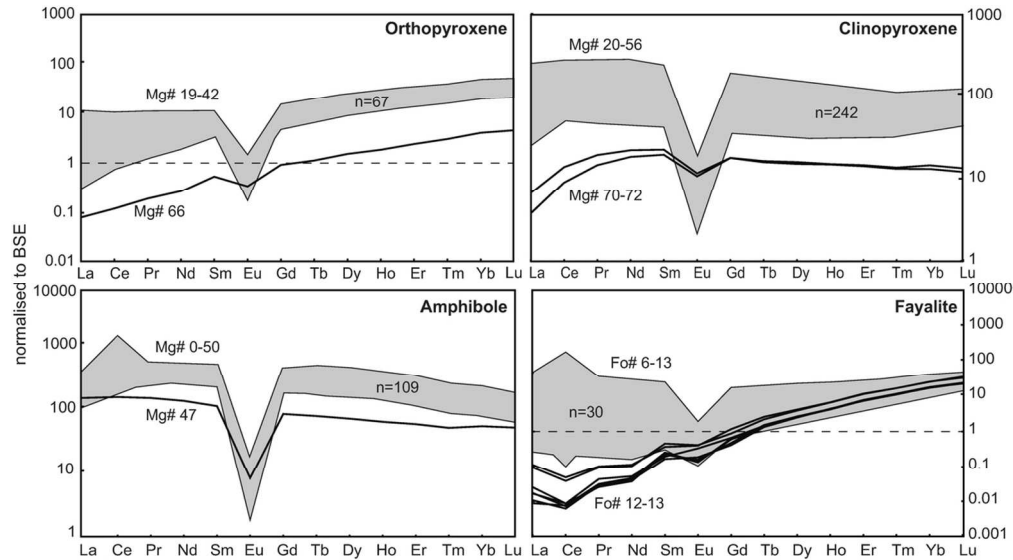


Fig. 8: REE contents in mafic minerals normalised to BSE composition (McDonough & Sun, 1995). Dark grey fields show REE patterns of the majority of crystals, likely to have grown in a highly evolved rhyolitic melt. Black lines indicate REE contents of crystals deviating from the main trend. Crystals depleted in LREE and lacking characteristic Eu anomaly crystallised in less evolved magma prior to significant plagioclase fractionation.

Fig. 8

114x63mm (300 x 300 DPI)

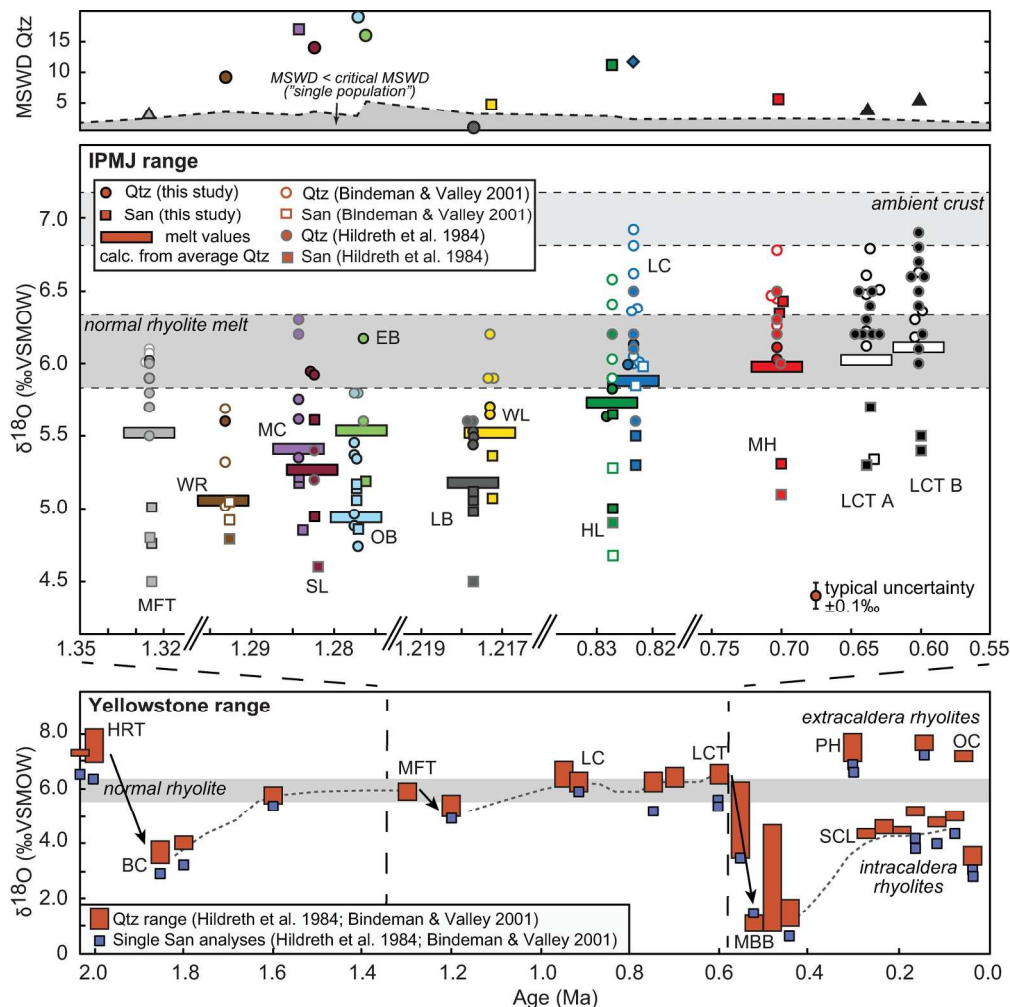


Fig. 9: Oxygen isotopes $\delta^{18}\text{O}$ in ‰ VSMOW for explosive end members MFT and LCT and effusive domes and lavas in-between (abbreviations and colours as in Fig. 1). Wide boxes indicate probable $\delta^{18}\text{O}_{\text{melt}}$ values calculated from average quartz compositions for each unit. Grey background bar illustrates range of "normal"- $\delta^{18}\text{O}$ rhyolites (expected values by fractionation of MOR and arc basalts, Bindeman (2008) and references therein). Upper panel defines the MSWD of the quartz analyses. $\text{MSWD} > \text{MSWD}_{\text{critical}}$ indicates that the results do not represent a single population. Lower panel represents trends in $\delta^{18}\text{O}$ throughout the history of the Yellowstone volcanic field (after Bindeman and Valley, 2001; Hildreth et al. 1984).

Fig. 9

188x188mm (300 x 300 DPI)

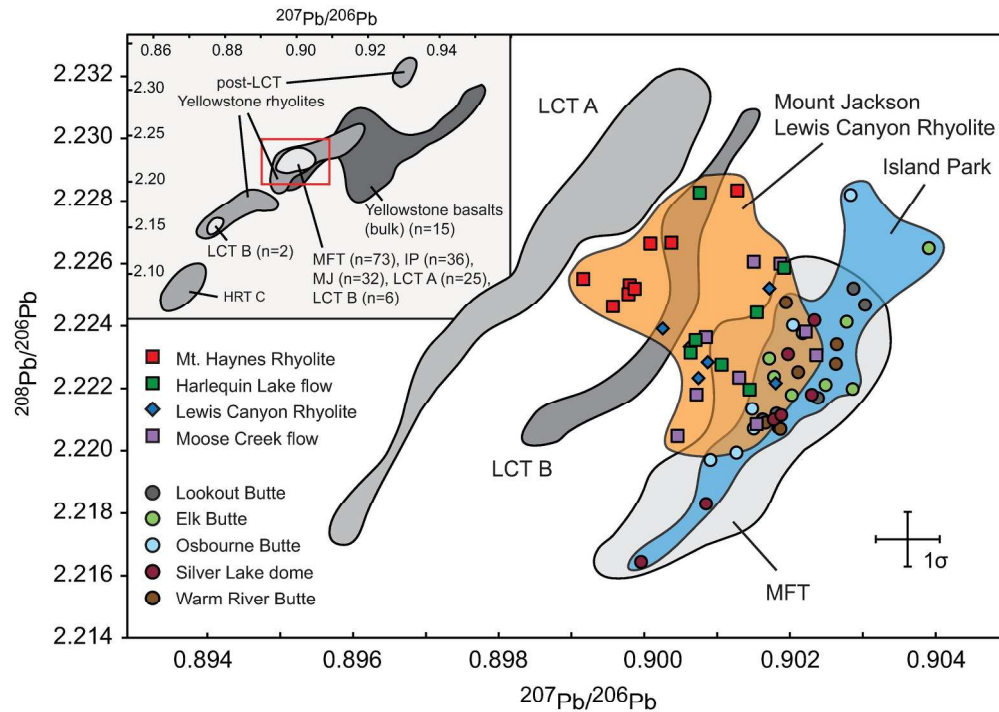


Fig. 10: Pb isotopes in sanidine in IPMJ rhyolites. Inlet diagram shows the entire range of Pb isotopic compositions for the entire Yellowstone volcanic field with red-framed box showing the extent of main diagram. Data for IPMJ rhyolites from this study, MFT from Ellis et al. (unpublished data), other Yellowstone rhyolites and basalts from Doe et al. (1982); Watts et al. (2012).

Fig. 10
212x253mm (300 x 300 DPI)

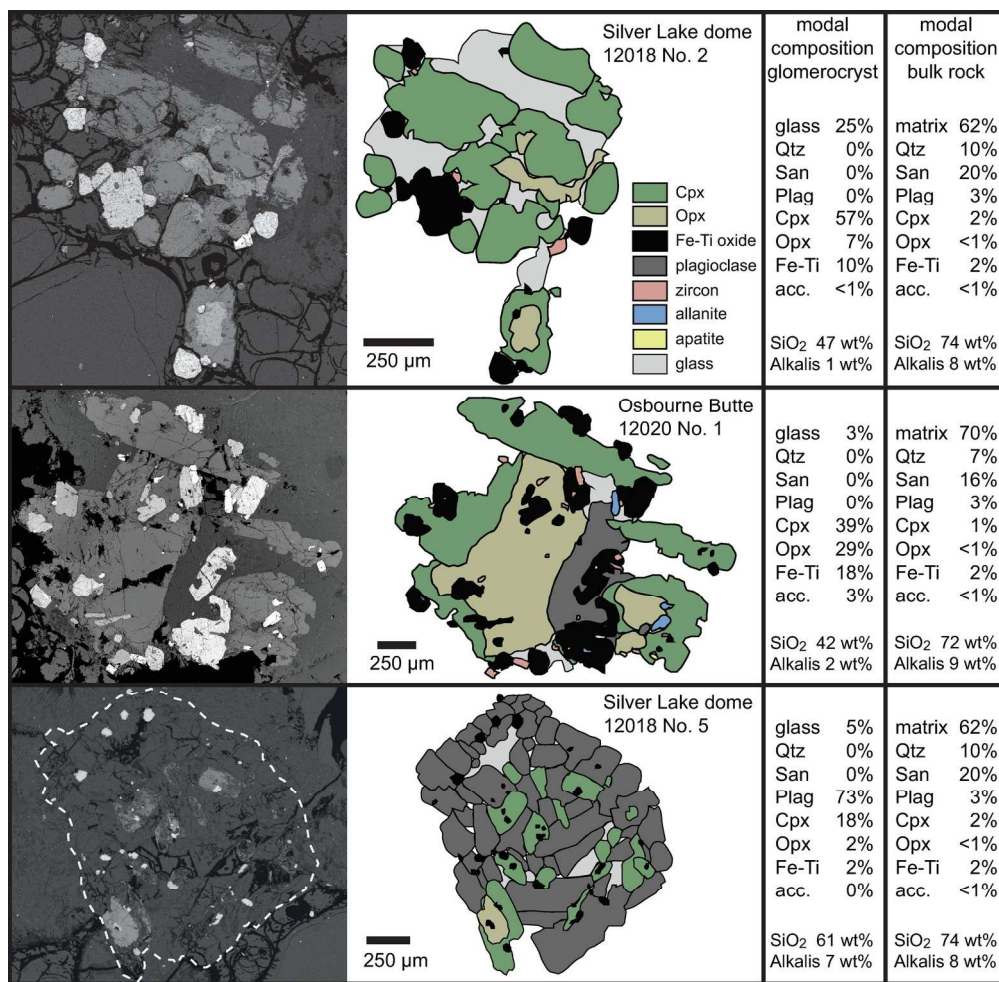


Fig. 11: Example of glomerocrysts from different Island Park domes. Comparison of mineral assemblage within the aggregate with free single crystals yields no significant compositional differences. Bulk glomerocryst compositions are less evolved compared to typical whole rock modal compositions due to the lack of late-forming sanidine and quartz.

Fig. 11
177x173mm (300 x 300 DPI)

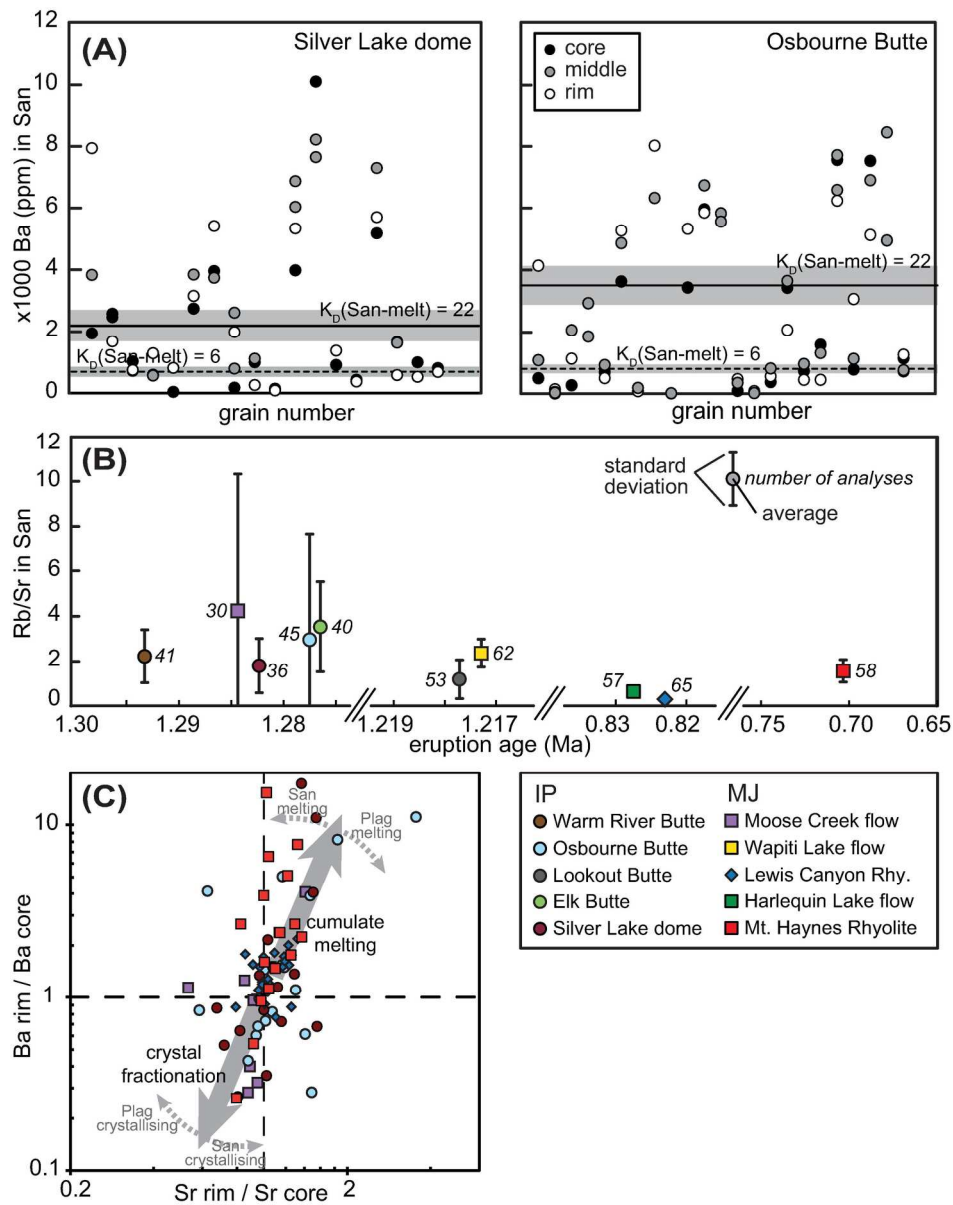


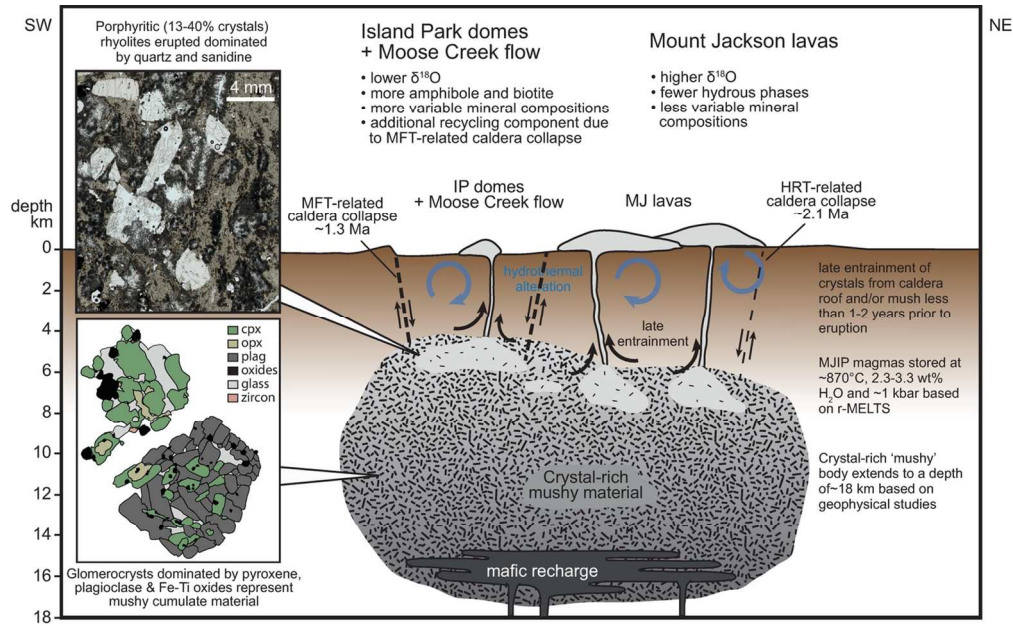
Fig. 12: (A) Ba contents in sanidine rims and cores do not fall into the calculated equilibrium range expected from Ba contents in glass from the same unit. Equilibrium range calculated with $K_D(\text{San-melt}) \approx 22$ in rhyolite from Leeman and Phelps (1981) and $K_D(\text{San-melt}) \approx 6.7$ in high-silica rhyolite (Mahood & Hildreth, 1983). Figure contains only data for those sanidine crystals that were imaged via CL. (B) Average Rb/Sr in San for each unit with error bars indicating one standard deviation. Numbers in italic correspond to number of analyses per unit. Note the large variability for units close to MFT-related caldera. (C) Ba ratio between rim and core and Sr ratio respectively suggest that reverse zoning in San is due to cumulate remelting or crystallisation from less evolved rhyolite melt. Figure contains only data for those sanidine crystals that were imaged via CL.

Fig. 12

174x222mm (300 x 300 DPI)

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28 Fig. 13: Schematic summary diagram of petrogenetic processes and characteristics of the IPMJ rhyolite series.

29 Fig. 13

30 124x77mm (300 x 300 DPI)

Table 1: Modal mineral abundances in percent determined via point counting on thin sections

Unit	WR	MC	SL	OB	EC	LB	WL	HL	LC	MH
devitrified	D	D	P	P	P	D	P	D	D	D
matrix	70.7	69.4	61.2	60.9	71.8	67.4	86.5	81.6	82.2	81.1
quartz	11.8	16.6	8.2	8.0	7.1	11.3	4.1	9.1	3.9	5.8
sanidine	11.4	8.8	25.5	19.8	16.3	14.8	7.0	5.5	5.9	10.1
plagioclase	3.9	3.2	2.5	7.2	4.3	4.5	1.8	3.1	7.1	2.5
Fe-Ti oxides	1.2	1.0	1.1	1.6	0.3	1.3	0.5	0.6	0.4	0.3
mafic minerals*	1.0	0.9	1.4	2.5	0.3	0.8	0.1	0.1	0.6	0.3
accessories	<0.1	<0.1	<0.1	<0.1	<0.1	<0.1	<0.1	<0.1	<0.1	<0.1

D = fully devitrified, P = partially devitrified, still contains glass shards

Mafic minerals*

clinopyroxene	0	20.3	61.8	73.6	22.5	34	0	0	67.1	93.3
orthopyroxene	28.6	28.4	15.4	24.2	10	18.4	100	100	25	6.7
fayalite	66.7	16.2	19.5	0	62.5	27.2	0	0	7.9	0
Fe-hornblende	4.8	35.1	3.3	2.2	5	20.4	0	0	0	0

* Due to the scarcity of mafic minerals in thin section mineral abundances of different mafic mineral groups were estimated by number of grains picked after density separation

Table 2: Bulk compositions of IPMJ lavas via XRF and ICP MS

Unit	Warm River Butte	Silver Lake dome	Os-bourne Butte	Elk Creek Butte	Look-out Butte	Moose Creek flow	Wapiti Lake flow	Lewis Canyon Rhy.	Harlequin Lake	Mt. Haynes flow
series	IP	IP	IP	IP	IP	MJ	MJ	MJ	MJ	MJ
sample #	BE SRP 12 027	BE SRP 12 018	BE SRP 12 022	BE SRP 12 024	BE SRP 12 025	BE SRP 12 001	JT 13 019	BE SRP 12 043	BE SRP 12 032	BE SRP 12 036
Major elements via XRF, normalised to 100% anhydrous (wt %)										
SiO ₂	76.11	75.54	73.97	78.43	73.54	77.73	77.77	75.74	76.28	77.65
TiO ₂	0.17	0.20	0.22	0.11	0.24	0.13	0.12	0.22	0.21	0.14
Al ₂ O ₃	13.29	13.08	13.70	11.69	13.71	12.09	12.04	12.66	12.40	12.02
FeO*	1.76	2.01	2.00	1.19	2.19	1.30	1.15	2.00	1.73	1.31
MnO	0.05	0.04	0.04	0.02	0.03	0.03	0.01	0.03	0.02	0.02
MgO	0.05	0.08	0.12	0.04	0.07	0.03	0.01	0.04	0.12	0.05
CaO	0.46	0.71	0.95	0.48	0.79	0.43	0.36	0.61	0.53	0.32
Na ₂ O	3.13	3.28	3.51	3.15	3.63	3.26	3.24	3.61	3.40	3.23
K ₂ O	4.95	5.04	5.45	4.85	5.65	4.97	5.30	5.06	5.28	5.24
P ₂ O ₅	0.02	0.03	0.03	0.04	0.14	0.02	0.01	0.03	0.02	0.01
Trace elements via LA-ICP MS (ppm)										
La	88	85	96	71	83	75	61	56	79	53
Ce	163	156	172	141	139	138	109	130	152	93
Pr	19	17	20	14	16	15	12	12	16	10
Nd	64	60	67	49	55	49	41	43	58	33
Sm	13	12	13	10	10	10	8	9	12	7
Eu	0.9	1.3	1.6	0.5	1.9	0.5	0.4	1.6	1.1	0.4
Gd	10.6	10.1	11.6	9.7	8.7	9.0	8.0	8.0	10.6	6.2
Tb	1.9	1.8	2.0	1.8	1.4	1.7	1.5	1.5	1.9	1.2
Dy	11.3	10.6	11.9	11.5	8.6	10.3	9.8	9.7	11.8	7.6
Ho	2.2	2.1	2.4	2.3	1.7	2.1	2.1	2.0	2.4	1.6
Er	5.6	5.6	6.4	6.5	4.6	5.6	5.7	5.5	6.5	4.5
Tm	0.9	0.8	0.9	1.0	0.7	0.8	0.9	0.8	1.0	0.7
Yb	5.4	5.1	5.7	6.3	4.4	5.3	5.4	5.2	5.9	4.8
Lu	0.8	0.8	0.9	0.9	0.7	0.8	0.8	0.8	0.9	0.7
Ba	420	622	789	207	1098	180	141	901	575	160
Th	34	30	29	31	28	33	35	26	29	31
Nb	49	44	42	43	42	48	50	46	50	52
Y	48	53	60	55	44	50	53	50	62	35
Hf	8.5	8.7	9.1	6.4	9.9	7.0	7.3	10.0	9.3	7.2
Ta	3.7	3.2	3.0	3.5	2.9	3.8	4.0	3.2	3.5	4.0
U	5.7	5.0	5.3	7.1	4.6	7.5	7.6	5.1	6.1	7.3
Pb	32	28	30	30	26	33	25	28	26	29
Rb	121	147	162	175	141	182	230	166	187	223
Cs	2.0	2.3	2.7	3.2	1.7	3.6	2.5	1.8	2.6	2.8
Sr	25	37	48	16	61	13	10	53	29	8
Sc	3.2	3.3	3.5	1.7	3.5	1.7	1.8	3.0	2.3	1.8
Zr	254	277	303	164	350	181	192	345	290	188
Zr sat. T(°C) ¹	843	843	843	795	856	804	807	858	842	806
Zr sat. T(°C) ²	814±17	810±19	806±21	753±17	820±22	763±18	766±18	824±21	805±20	766±18

¹Hanchar & Watson (2003); ²Boehnke et al. (2013)

Table 3: $^{40}\text{Ar}/^{39}\text{Ar}$ eruption ages

Unit	Abbr.	Series	$^{40}\text{Ar}/^{39}\text{Ar}$ age (Ma)	analytical uncertainty 2 sig	Full external precision 2 sig	MSWD	non-eruption age grains $n_{\text{non-eruptive}} / n_{\text{total}}$
Warm River Butte	WR	IP	1.2943	0.0020	0.0026	1.44	7 / 30
Moose Creek flow	MC	MJ	1.2856	0.0064	0.0066	0.57	0 / 16
Silver Lake dome	SL	IP	1.2839	0.0112	0.0114	0.88	0 / 28
Osbourne Butte	OB	IP	1.2784	0.0052	0.0054	1.12	1 / 27
Elk Creek Butte	EC	IP	1.2777	0.0090	0.0092	1.01	0 / 29
Lookout Butte	LB	IP	1.2190	0.0144	0.0146	0.9	9 / 19
Wapiti Lake flow	WL	MJ	1.2187	0.0158	0.0160	0.3	0 / 31
Harlequin Lake flow	HL	MJ	0.8300	0.0072	0.0074	1.09	3 / 29
Lewis Canyon Rhyolite	LC	MJ	0.8263	0.0184	0.0186	0.58	10 / 27
Mount Haynes Rhyolite	MH	MJ	0.7016	0.0014	0.0016	1.04	3 / 29