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Observations and modelling of microphysical variability, aggregation and sedimentation in tropical storm cirrus outflow regions

M. W. Gallagher¹, P. J. Connolly¹, A. Heymsfield², K. N. Bower¹,
T. W. Choullarton¹, G. Allen¹, M. J. Flynn¹, G. Vaughan¹, and J. Hacker³

¹Centre for Atmospheric Science, University of Manchester, Manchester, UK

²NCAR, P.O. Box 3000, Boulder, CO, 80307-3000, USA

³School of the Environment, Flinders University, GPO Box 2100, Adelaide 5001, Australia

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Correspondence to: M. W. Gallagher (martin.gallagher@manchester.ac.uk)

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23761

Abstract

Aircraft measurements of the microphysics of a tropical convective anvil (at temperatures $\sim -60^\circ\text{C}$) forming above the HECTOR storm have been performed. The observed microphysics has been compared to a bulk and explicit microphysical model of the anvil region including crystal aggregation and sedimentation.

It has been found that in flights made using straight and level runs perpendicular to the storm that the number of ice crystals initially decreased with distance from the storm as aggregation took place resulting in larger crystals followed by their loss due to sedimentation. At still greater distances from the storm the number of very small crystals increased. This is attributed to the formation of new ice crystals on aerosol particles as the ice super saturation rose following the depletion of the larger ice particles following aggregation and sedimentation. Comparison with the explicit microphysics model showed that the changes in the shapes of the ice crystal spectra as a function of distance from the storm could be explained by the explicit microphysical model if the aggregation efficiency was set to $E \sim 0.02$.

It is noteworthy that this aggregation efficiency is much larger than values normally used in cloud resolving models at these temperatures (typically $E \sim 0.0016$). Furthermore if the bulk model is used then optimum agreement was reached with a collection efficiency for aggregation of $E \sim 0.05$. These results are important for the treatment of the evolution and lifetime of tropical cirrus clouds.

1 Introduction

The response of the tropical atmosphere to changes in climate forcing has received much attention lately (IPCC, 2007) since a positive feedback, arising from increased supply of water vapour by convection in the tropical tropo-pause layer (TTL), may be countered by negative feedbacks due to increasing amounts of thin cirrus clouds also being formed in this region. These cloud types have been identified as a significant

23762

source of uncertainty in the Earth's radiation budget (Lynch et al., 2002) with several recent studies showing that their contribution to the earth-atmosphere radiation and water budgets in the TTL can be very significant, (Edwards et al., 2007; Stephens et al., 2002; Liou et al., 1994; Donner et al., 1997). Many studies show that optically thin cirrus are prevalent in the region of the tropical tropopause: (McFarquhar et al., 2000; Massie et al., 2002; and Dessler et al., 2006) and recently it has been suggested that this frequency, and hence radiative contribution may be larger than hitherto assumed, (Lee et al., 2009). Furthermore, analysis of several large cloud microphysical data sets, Baum et al., (2005), when compared with MODIS simulated retrievals, have revealed differences between in situ and retrieved single scattering albedo and asymmetry parameters. These differences were found to be largest for ice clouds formed in areas of deep convection (e.g. the TRMM-KWAJEX project) compared to those formed in regions of low updraft velocities. Compared to other cloud types there are relatively few measurements of ice crystal particle size distributions (PSD) in those formed through tropical deep convection. It is known, Fusina et al. (2007) that depending on conditions, the transition between net warming and cooling due to cirrus depends on ice PSD. Hence, there is a general need to determine ice PSD within cirrus to allow better analysis of their mean radiative impacts and potential feedbacks. To reduce uncertainties in cloud model radiative predictions in these regions more and improved data sets of PSD are needed, particularly for high level cirrus. In general previous studies have focused on two different types of cirrus cloud; thin semi-laminar cirrus associated with in situ formation by synoptic and mesoscale processes, and those originating from deeper convection in mid-latitudes. Here we present a typical case study of the microphysical variation observed within cirrus generated in tropical storm anvil outflow.

2 Background

Tropical anvil cirrus form at the top of deep cumulonimbus clouds, consisting initially of ice "debris", that spreads out laterally from the convective storm cells, with the larger

23763

ice particles being removed rapidly by sedimentation. Aggregation and sedimentation then dominate ice removal in the anvil downwind of the convective centre. In the tropics these anvils give way to new in situ generated cirrus at the tropical tropopause and form extremely widespread cloud layers of several hundreds to thousands of km in extent, thus contributing to large-scale radiative effects (Collins et al., 1996). The prevalent conceptual model of in situ generated cirrus has been well described by Heymsfield and McFarquhar (2002) in which three distinct layers are identified. This conceptual model was based originally on in situ cirrus ice particle size distribution measurements made during the FIRE-I experiment, e.g. Gultepe and Starr (1994), using aircraft Lagrangian spiral descents and in the FIRE-II experiment using balloon-borne ice crystal replicators (Miloshevich et al., 1992). It was observed that average ice crystal sizes increased steadily from near cloud top towards cloud base. From the observed vertical distribution of ice crystal PSD and detailed photographs of their habits it was inferred that these cirrus can be represented, conceptually, by the following three layers with broadly distinct characteristics:

Layer 1 – The Nucleation layer in the uppermost part of the cloud is composed of small ice crystals, or proto-crystals, typically proto-bullet rosettes. In this layer ice production occurs when the relative humidity (RH) exceeds some critical value necessary for ice activation. Ice crystals in this layer are often up to a few tens of micrometers in size;

Layer 2 – The Growth layer which is much thicker and comprised of ice crystals which present pristine habits due to sufficient ice super-saturation (S_{ice}) that allows sustained ice crystal growth to large sizes, typically several hundreds of microns in length, but sometimes exceeding 1–2 mm;

Layer 3 – The Sublimation layer where conditions of ice sub-saturation ($S_{ice} < 1$) cause sublimation and eventually complete evaporation of the ice crystals falling from the layer above. In this layer ice crystals often present as rounded forms and are more difficult to classify than those in layer 2. The thickness of layer 3 depends on the ice water content (IWC) and the PSD near the base of layer 2 as well as the RH and temperature

23764

vs size relation suggested by Heymsfield (2004):

$$\text{IWC} = \sum_1^N N_i \alpha D_i^\beta \quad (2)$$

where N_i is the concentration in bin i , $\alpha = 3.4709$ is the multiplier of the mass-size relation, D_i is the mid-point diameter of bin i , and $\beta = 2.6$ is the index in the mass size relation.

Recent studies combining ensemble cirrus models with ice PSD schemes have been able to broadly predict IWC levels, ice water profile (IWP), total solar optical depth, volume extinction coefficients and radiative properties that compare well with those derived from in situ observations, in some cases without the need for information on ice crystal effective dimension, Baran et al. (2009). However, the uncertainties associated with in situ observations from 2-D imaging spectrometers can be large. The magnitudes of these uncertainties, whilst theoretically the same for most instrument types, can vary depending on the type, configuration and incept date of the instruments which are constantly being improved.

Uncertainties associated with ice PSD measurements include limited sample volumes (e.g. the Cloud Particle Imager, Lawson et al., 2006a), insufficient grey scale information to correct for depth of focus mis-sizing (e.g. CAPS-Cloud Imaging Probe, Baumgardner et al., 2006) probe dead time, probe triggering efficiency (particularly for small particles) and random collection within the instruments depth of field (Korolev et al., 2006; Connolly et al., 2007). More seriously they can be subject to artifacts, particularly in the small, sub $50 \mu\text{m}$ size range, due to ice particle shattering on aircraft as well instrument surfaces, which has been demonstrated to be an issue for all such instruments, Korolev (2009). These artifacts appear to manifest particularly in the presence of significant concentrations of large ice crystals (typically $> 4 \text{ mm}$). These measurement limitations are still being addressed and steps to minimize them are being generally adopted and depend on both operational, instrument and environmental conditions (Korolev, 2009; Field et al., 2006; Heymsfield, 2007) but they will continue to

23771

add to the general uncertainty associated with all such in situ measurements of small ice particles for the foreseeable future. The work by Korolev (2010) suggests the CIP, and similar instruments despite being open path instruments, also suffer from ice shattering artifacts. There has been little quantitative information published on the effect particle shattering has on the CPI.

For the analysis of the ACTIVE experimental data presented here the main assumption is that particle shattering, at least for the ARA aircraft platform, and the instruments used (CPI and CAPS-CIP), will be limited to particles $< 50 \mu\text{m}$ which will have a small influence on the total IWC. This assumption is justified having been demonstrated by a detailed analysis of data from an similar CAPS-CIP and CPI instruments flown on the DoE Proteus aircraft as part of the TWP-ICE experiment, Mc Farquhar et al. (2008). The TWP-ICE experiment was conducted simultaneously with the ACTIVE experiment and flights were often conducted jointly by the two aircraft, and which flew near identical cloud instruments (comparisons between the aircraft are being reported elsewhere).

Figures 3 and 4 demonstrates a typical comparison between the open-path CAPS-CIP instrument on the Egrett (which is assumed to be less prone, but not immune, to shattering artifacts) and the inlet based CPI (assumed to be more prone). The latter instrument, however, appears to be less sensitive (at least for the conditions reported here) to this artifact, likely due to limited triggering efficiency of its detection system for small particles. It should be noted that the data in this figure have been integrated across a full cloud penetration (10 min) to improve sampling statistics and so cover a wide range of PSD conditions. Figure 3 also shows the PSD measured by an open path Mie scattering spectrometer (DMT Cloud Droplet Probe, CDP-100 Version 1–2004, designed for sampling water PSD) mounted directly below the CAPS-CIP instrument, which again, considering the very different operating principles and resolutions, is encouraging, although there is a suggestion that both CIP and CPI overestimate concentrations below $50 \mu\text{m}$. It is known that version 1 of the CDP suffers from undercounting issues for particle concentrations above approximately 200 cm^{-3} , in water clouds, Lance et al. (2010). These concentrations are much larger than reported in ice clouds.

23772

Differences between the CPI and CIP reported concentrations can, however, become more pronounced on shorter sampling timescales depending on the PSD mode size. It is clear from these comparisons that for the conditions typically encountered during ACTIVE the CPI measured PSD also become statistically limited for particle sizes $>800\mu\text{m}$ whereas the coarser resolution afforded by the CAPS-CIP ($15\mu\text{m}$) will add to its uncertainties for small particles. We will expect the CPI PSD to be statistically less reliable for particle dimensions $D > 100\mu\text{m}$ if using <1 min integration periods compared to the CAPS-CIP and this limitation was mitigated in the subsequent analysis by optimizing the integration time for the CPI. The effect of these different probe issues on uncertainty in deriving values of λ_o is highlighted below.

5 Results – microphysical horizontal structure variability

The spatial variation in observed PSD characteristics within the outflow region was investigated by examining the time histories of IWC (g m^{-3}) and λ_o (cm^{-1}) across each track along with the ice crystal habits presented by the CPI images as a function of increasing downwind distance, and transverse to the outflow. Figures 5a–d and 6a–d show IWC and λ_o respectively, for each flight leg (1–4) as a function of latitude, and increasing distance), plates (a)–(d), from the storm centre. Figure 5, shows a general decrease in IWC with increasing distance due to ice loss by aggregation and sedimentation. Figure 6 shows there is a general decrease in λ_o with distance from the storm, again due to the effects of aggregation on the evolving PSD although there was significant variability.

Figures 5 and 6 suggest that there is a significant lateral shift in the core outflow region due to large-scale horizontal perturbation. Indeed there may have been two distinct outflow regions generated. To account for this when comparing with model predictions later we have computed average values of IWC and λ_o within $\pm 0.2^\circ$ of latitude of the central outflow region to the south, and take the core region as being represented by the peak in the IWC. This peak was also confirmed by the CPI measurements. As

23773

shown later this region was dominated by large aggregated ice crystals and the averaged values should be relatively unaffected by other processes such as re-nucleation which appears to occur and contribute to a greater extent in the wings of the observed outflow, as discussed later.

Flight leg 1, 19 km from the storm centre, displayed the highest IWC's, with median values $>1\text{ g m}^{-3}$ across the central outflow region although with low values of n_o (not shown). Thus small ice crystal concentration contributions (at least within the detection limits of the instruments) appear to contribute very little to the overall PSD here as shown by λ_o . The PSD in this region were clearly dominated by large complex aggregates comprised mainly of hexagonal plates. These large aggregates eventually sediment out downwind of the storm in the core region as can be seen by the changes in λ_o with distance. In the lateral edges of the outflow there is some evidence that λ_o (and n_o) increases as nucleation occurs. Most of the core region ice appears to have been precipitated within a distance of 30 km. Within track 2, IWC's peaked at $>1\text{ g m}^{-3}$ but showed greater variability. n_o (not shown) remained reasonably consistent across the outflow (albeit with small scale regions where significant enhancements occurred) as did values of λ_o downwind of the outflow. λ_o did show enhancement in several regions within track 4. In these regions the PSD were dominated by higher concentrations of small, recently formed, pristine ice crystals, which presented as small poly-crystals, or proto-bullet rosettes in some regions.

5.1 Habit variation

Images of the ice crystals recorded by the CPI corresponding to each flight leg are now considered in more detail. These are plotted in Figure 7a–d for each successive flight leg. The images are presented in a matrix plotted as a function of latitude ($^\circ$) along the horizontal axis, and which has been split into 5 latitudinal bands of approximately 0.2° width, while the vertical axis shows the ice crystals sorted according to size. The distance from the storm outflow increases from Fig. 7a to d. If we inspect the wings of the outflow there is evidence in most, but not all cases, of a gradual transition to a

23774

region where significant new ice particle production or re-nucleation is occurring likely due to an increase in super-saturation and subsequent haze freezing. The habit of the ice crystals in these transition regions is very different to those in the centre of the outflow, presenting as small polycrystalline or proto-bullet rosettes which are clearly imaged by the CPI, most of which are $<100\ \mu\text{m}$, although some have grown to significant sizes having fallen from higher in the cloud. However these crystals appear to have undergone very little, if any, aggregation. The patterns and extents of these zones are particularly evident in the later legs, Fig. 7b to d. It is interesting that although the overall concentration of large aggregates decreases they still persist even as far downstream as track 4, some 134 km, although overall λ_o decreases.

Figure 8, shows the ice crystal size distributions, again displayed in matrix form, showing the variation within the along-axis outflow region ($\pm 0.2^\circ$ latitude) for three of the penetrations at 19, 31 and 67 km downwind of the storm. The reduction in large aggregate contributions after 20 km becomes clear, leading to a gradual decrease in IWC.

If we consider the main processes governing new nucleation of ice particles in the anvil environment then key to this will be the precipitation of the larger aggregates that results in an increase in ambient ice super-saturation in the depleted region, subsequently allowing ice nucleation by haze freezing. This would lead to both an increase in n_o in such regions coupled with eventual recovery in λ_o through further aggregation. These particles, again due to the increased super-saturation, eventually grow into larger bullet-rosettes that are more characteristic of mature cirrus e.g. as seen in frontal cirrus. Figure 9 shows detailed images of the complex chain aggregates of hexagonal plates observed in this anvil outflow which should be compared with the pristine ice crystal habits (mainly large bullet rosettes) recorded in typical frontal cirrus, Fig. 10. (These were observed over Adelaide, Australia, during the EMERALD-1 experiment, Whiteway et al., 2004; Gallagher et al., 2005).

Observations within tropical anvil clouds have noted many aggregates of ice crystals, often formed in chains, Connolly et al. (2005). It is thought that a large majority of

23775

these aggregates originate within the storm, where temperatures are high enough for the cohesive properties of ice to be favourable. It is also thought that aggregation of the ice crystals may be enhanced by the electric field that is present due to the charge separation mechanism in thunderstorms. Davis (1964) calculated that gravitational settling would be the dominant collision mechanism for two conducting spheres, charged or uncharged, of mass similar to that of a typical cloud particle. Saunders (1975) hypothesized that the electric field may enable two ice crystals to come together for long enough for them to sinter.

Observations from the ACTIVE project also show clear evidence of larger aggregates occurring downwind of the main area of convection of the Hector thunderstorm, thus suggesting that significant aggregation may occur within the cirrus itself. The temperature of the cirrus investigated was approximately -60°C but it was previously thought that aggregation would not be very efficient at these temperatures. If significant aggregation does occur at these temperatures it is important to quantify as it may be an important contributory mechanism in determining cirrus lifetime. Also worthy of mention, but not wholly relevant to this study was the fact that, on occasion in localised regions, bullet rosette type crystals were also observed in the outflow cirrus. It is thought that the bullet rosette crystals form in situ, as in high level frontal cirrus, either by heterogeneous nucleation of ammonium sulphate aerosol or homogeneous freezing, Abbatt et al. (2009), and that two main criteria are necessary for their formation: (i) the cirrus has to become relatively thin, due to the fall out of larger aggregates and (ii) some reasonable uplift must be present, e.g. due to the presence of gravity waves.

6 Model Studies

The evolution of the ice crystal PSD's within the axial region of the anvil outflow was simulated using an explicit bin-aerosol-ice-microphysical model, based on the ACPIM (Aerosol Cloud Interaction Model). A description of the original model can be found

23776

in Connolly et al. (2009, 2010) so only the key features relevant to this study will be repeated briefly here. The model incorporates an aerosol and cloud particle moving centre multi-bin structure developed by Jacobson et al. (1994) with mass conservative, explicit liquid-ice, ice-ice collision, coalescence, riming and aggregation processes. Aggregation is described with the stochastic collection equation (see Pruppacher and Klett, 1997), and was solved using the algorithm described by Bott (2000). The kernel applied was the gravitational kernel described by Hall (1980)

$$k(x, y) = E \frac{\pi}{4} (x + y)^2 |u(x) - u(y)| \quad (3)$$

where x and y are the diameters of two particles, $u(\dots)$, is the terminal fall speed versus diameter relationship and E is the aggregation efficiency - assumed to be a constant in this study -, which is the product of the collision efficiency and the sticking efficiency. The ice mass-size relation used in the model follows Heymsfield et al. (2004) for consistency with the way the CPI and CIP observational data were treated as described above, viz $m(D) = 3.4709 D^{2.6}$ with m in kg and D in metres. Coalescence efficiencies in the model rely on experimental data from Low and List (1982) and Lew and Pruppacher (1983), and CFD calculations based on the work of Wang and Ji (1992).

Condensation is treated explicitly in the model whilst ice crystal growth and habits are treated using the model of Chen and Lamb (1994). The bin-microphysics model uses a 4th order monotonic advection and flux method for solving stochastic collision efficiencies for all bins. Ice nucleation can be represented in several different ways, either using the singular freezing hypothesis approach, using the equations described by Connolly et al. (2009), or a traditional stochastic nucleation approach. However, here we will focus on the process of aggregation and its sensitivity to the aggregation coefficient within the axial region of the anvil outflow. We assume contributions to the evolution of the PSD due to other processes in this core region, such as new particle nucleation, will be a minimum. The model was run in both bin and bulk modes and compared with observations. Finally we investigate the contribution of sedimentation to the results.

23777

6.1 Model initialisation

The model was initialised with the observed thermodynamic profiles and ice water contents from the CIP. Anvil cloud thickness was based on a combination of aircraft data, cloud top heights derived from MTSAT-1R (retrieved hourly resolution data provided by P. Minnis, personal communication, 2006, 2010) and radar observations, described by May et al. (2009). The simple conceptual model described earlier was adopted by assuming uniform ice saturation (within layer 2) with a below cloud sub-saturated layer. Ice crystal PSD predictions were selected from the model levels corresponding to the aircraft cloud penetrations (13 km) and exponential gamma functions were fitted to them conserving all three moments. The ice crystal aggregation efficiency was varied between $E = 0.005$ and 0.02 and the results compared with the in situ aircraft measurements.

We assume that the case study can be considered as a Lagrangian one. The sampling times and wind speeds for each leg with a corresponding distance downwind from the first leg were determined to be: Leg 1: ~15:10; Leg 2: ~15:31, 12 km down wind, ~10 m s⁻¹; Leg 3: ~15:57, 50 km down wind, ~24 m s⁻¹; and Leg 4: ~16:15, 75 km down wind, ~24 m s⁻¹. The actual horizontal wind speed at the altitude of observation was ~20 m s⁻¹ in the direction the anvil outflow, whereas the assumption that the flight is Lagrangian would necessitate velocities between the legs of 10, 24 and 24 m s⁻¹. Thus it can be seen that the flight is not strictly Lagrangian; however, another simplifying assumption is that the cloud was approximately steady-state, at least on the time scale of the observations, so that the important variable is, therefore, distance downwind of the first flight leg.

6.2 Model results

The size distributions measured by the CIP instrument and a mass vs size relation, Heymsfield (2004), were used to produce estimates of IWC using: $IWC = \sum_1^N N_i \alpha D_i^\beta$

23778

Table 1. Track Longitude, Latitude and Track lengths.

Track No.	Longitude ° (core distance km)	Latitude °	Track (km)
1	129.75 (19)	-10.820 to -12.115	144
2	129.64 (31)	-12.110 to -10.970	127
3	129.31 (67)	-11.145 to -12.252	123
4	129.07 (93)	-12.239 to -11.055	132

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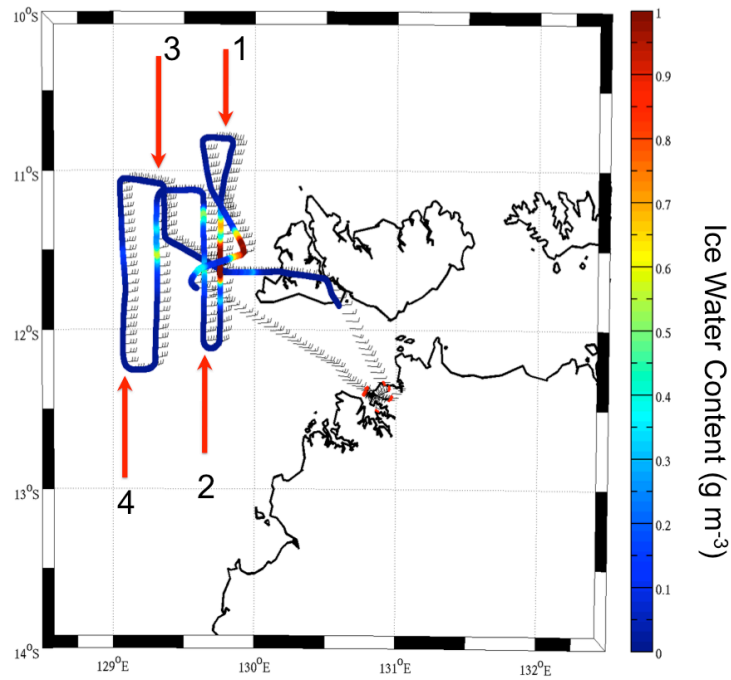


Fig. 1. Cirrus penetration flight tracks 1 to 4 during case study AE13 (9/11/2005) shown on a map of the Darwin region and the Melville and Bathurst islands in the Beagle Gulf. Darwin city is marked. The aircraft flight tracks are colour coded according to ice water content, IWC, (g m^{-3}) (see text).

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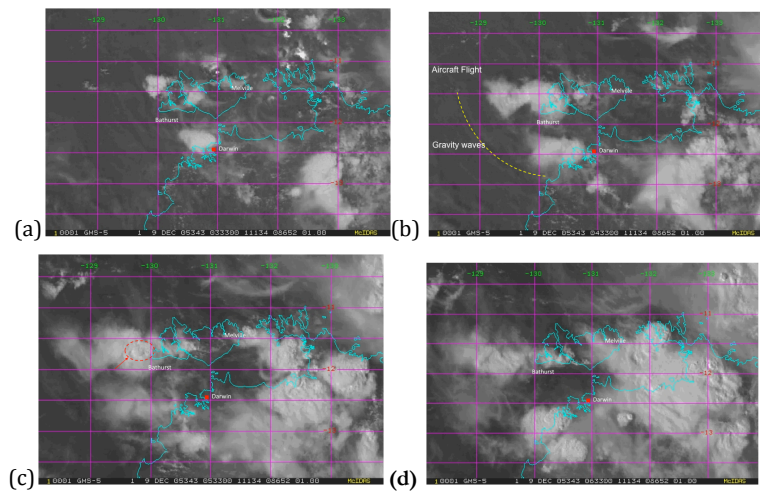


Fig. 2. MTSAT Visible images showing a Hector storm developing, north of Darwin and the cirrus outflow (see text). ACTIVE case AE13 on 9/12/2005 from 03:33–06:16 UTC. Melville and Bathurst islands are highlighted.

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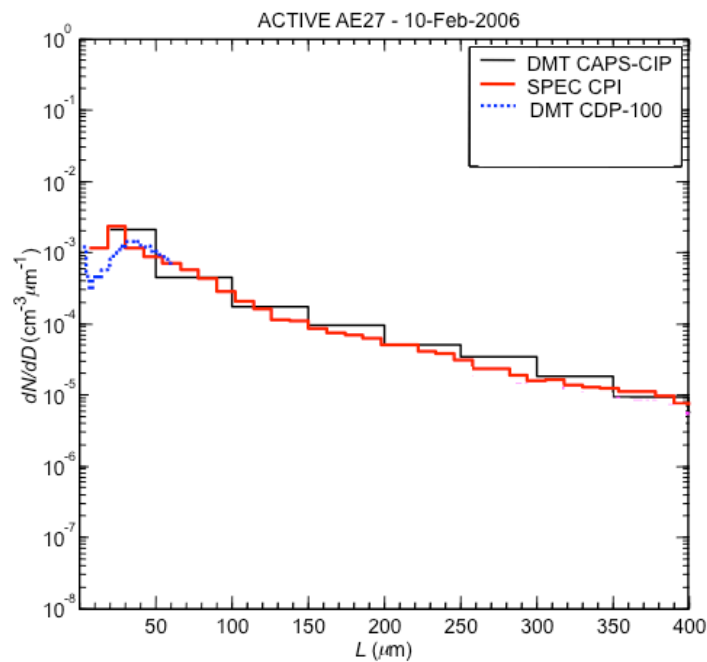


Fig. 3. Particle size distribution measurement comparison observed between the SPEC CPI and DMT CAPS-CIP instruments for (a) Particle size range $10 < L < 400 \mu\text{m}$ and (b) $10 < L < 1800 \mu\text{m}$. Also shown is the PSD in the size range $2 < D_p < 62 \mu\text{m}$ measured by a Mie scattering spectrometer (DMT CDP-100).

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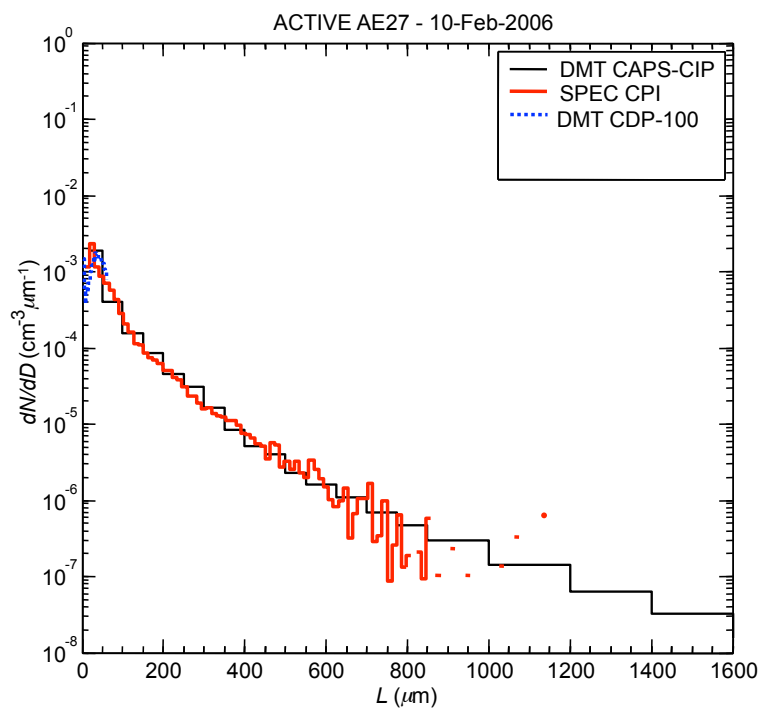


Fig. 4. As Fig. 3, but for particle size range, $10 < L < 1800 \mu\text{m}$. Also shown is the PSD in the size range $2 < D_p < 62 \mu\text{m}$ measured by a Mie scattering spectrometer (DMT CDP-100).

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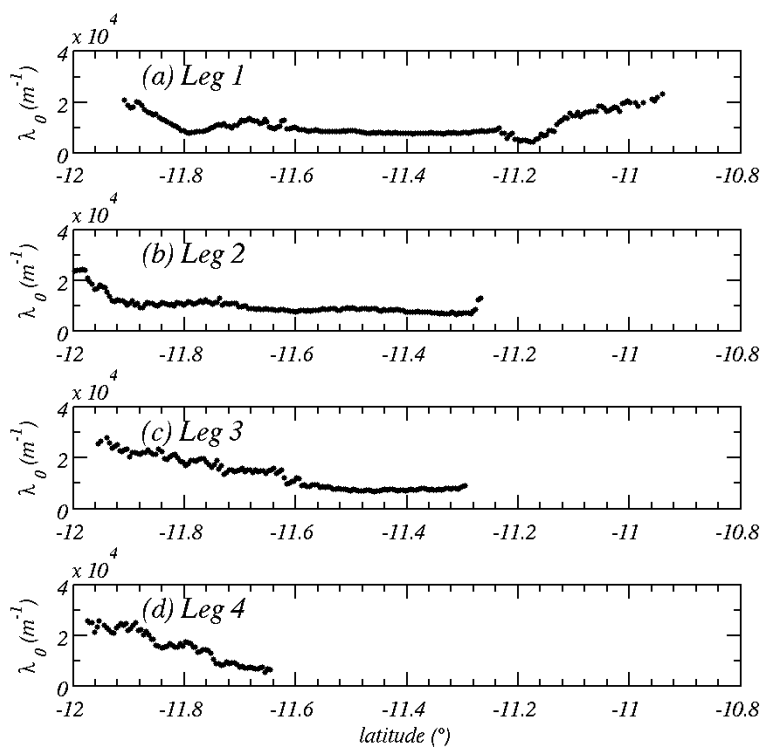


Fig. 5. IWC (g m^{-3}) determined by CIP across each flight leg 1–4, **(a)–(d)** as a function of increasing distance from the storm centre. There is a general decrease in IWC with distance as well significant lateral shift in core outflow.

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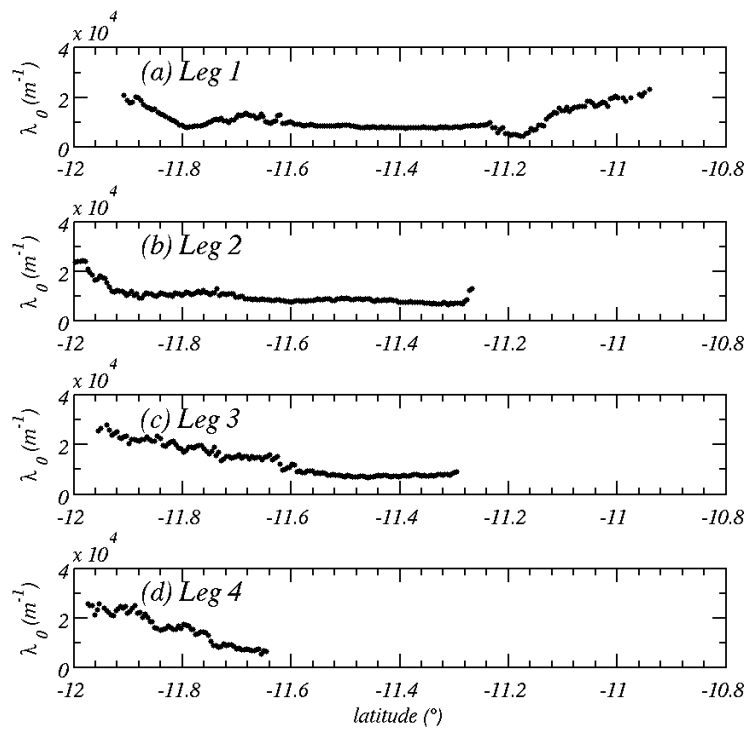


Fig. 6. The derived λ_0 (cm^{-1}) for each leg as a function of latitude with each flight as a function of increasing distance from the storm (a)–(d). There is a general decrease in λ_0 with distance from the storm.

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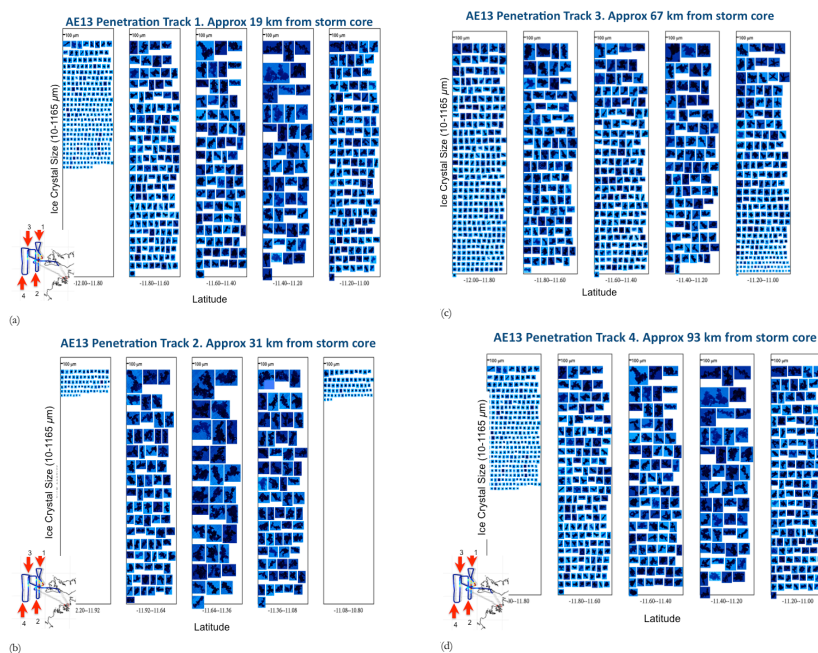


Fig. 7. CPI unprocessed images recorded during cloud penetrations 1 to 4 perpendicular to the storm outflow with increasing distances, 19, 31, 67 and 93 km, figures (a) to (d) respectively, from the storm. The crystal images observed are represented in a matrix with the horizontal axis split into 5 latitude bands of width: 0.2° , perpendicular to the outflow direction. The vertical axis is sorted as a function of the maximum ice crystal size (from 5 to $\sim 1120 \mu\text{m}$) observed in each latitude bin. The size scale at the top of each column represents $100 \mu\text{m}$.

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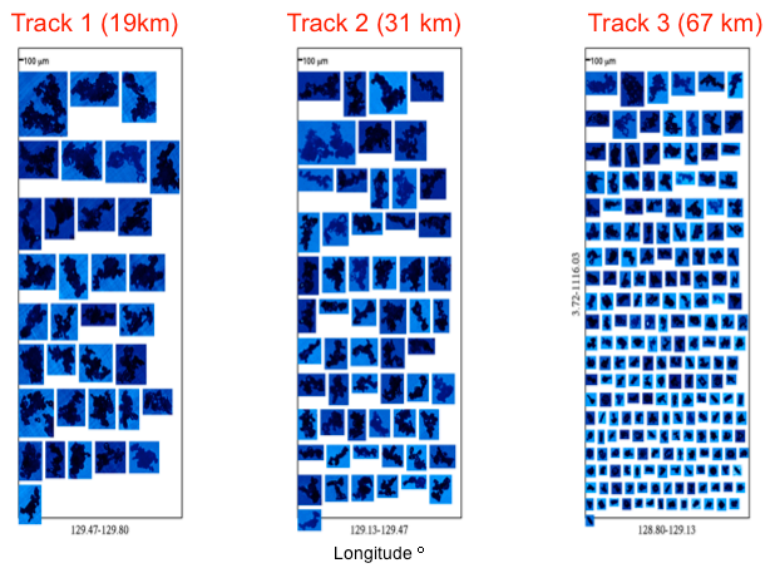


Fig. 8. Ice crystal habits and sizes along the centre-line of the storm outflow during AE13 at 19, 31 and 67 km respectively showing recovery in n_o and reduction in aggregate contributions.

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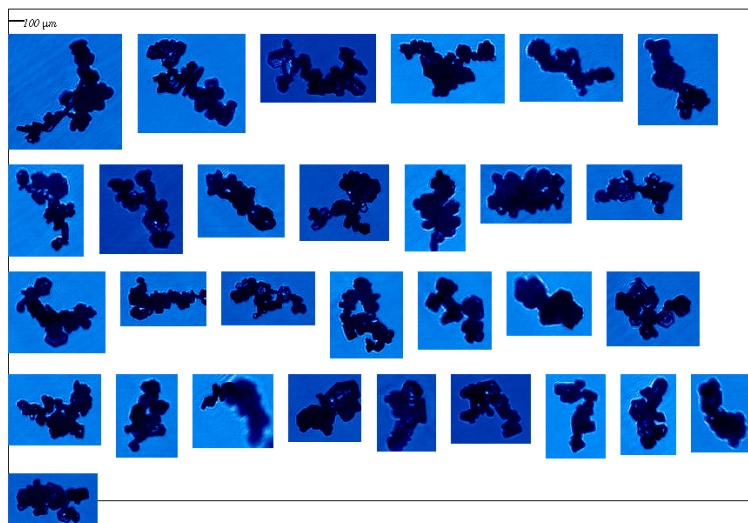


Fig. 9. Close up view of typical complex chain aggregates of hexagonal plates observed within a Hector storm outflow (temperatures $T < -50^\circ\text{C}$) during the ACTIVE-2 experiment, Connolly et al. (2007).

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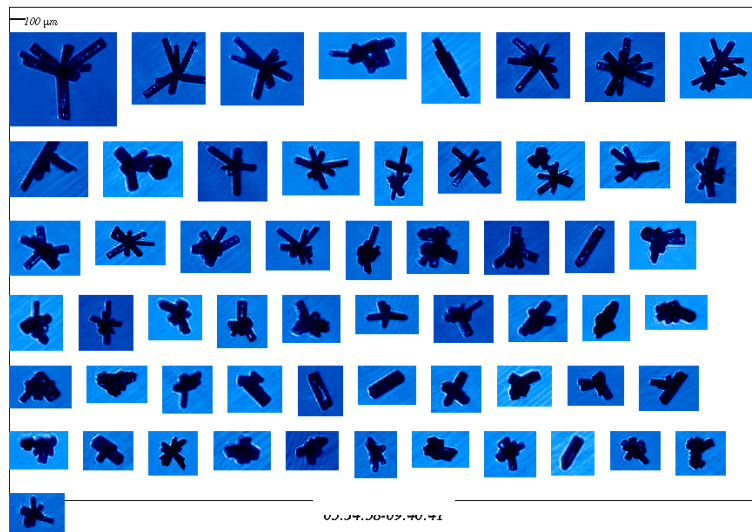


Fig. 10. Ice crystal habits observed in high-level frontal cirrus, at temperatures $T < -40^\circ\text{C}$, over Adelaide, Australia, during the EMERALD-1 experiment, Gallagher et al. (2005). Habits presented as mainly pristine bullet rosettes.

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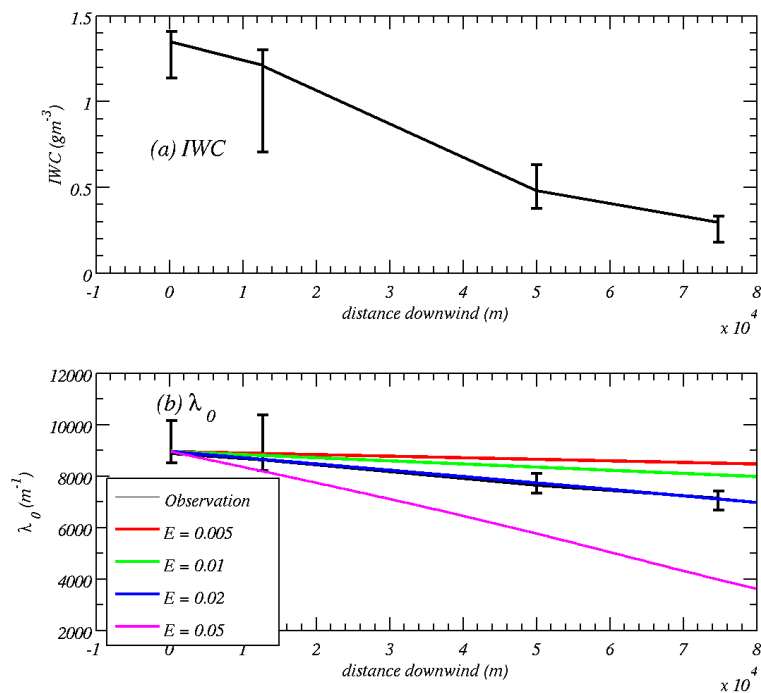


Fig. 11. (a); The IWC inferred from the CIP instrument as a function of distance from the main storm; **(b);** the λ_0 parameter as a function of distance from the main storm and as a function of aggregation efficiency, E , predicted by the model and compared to the λ_0 derived from CIP measurements.

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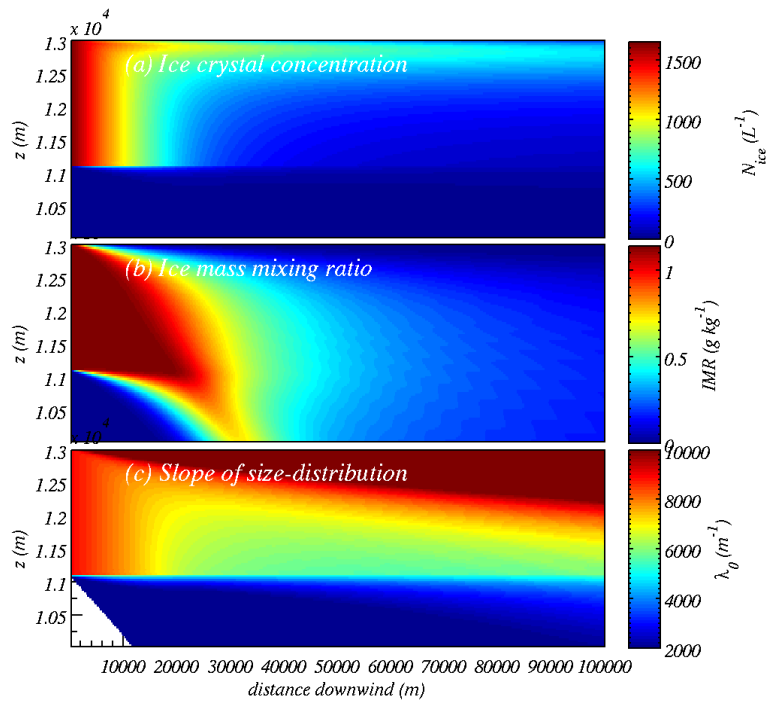


Fig. 12. Contour plots of a typical output from the sedimentation/aggregation model. **(a)** shows the time-height evolution of ice crystal number concentration, $N(L^{-1})$; **(b)** shows the time-height evolution of ice mixing ratio (IMR, $g\ kg^{-1}$); and **(c)** shows the time-height evolution of the slope of the distribution $\lambda_o\ (m^{-1})$.

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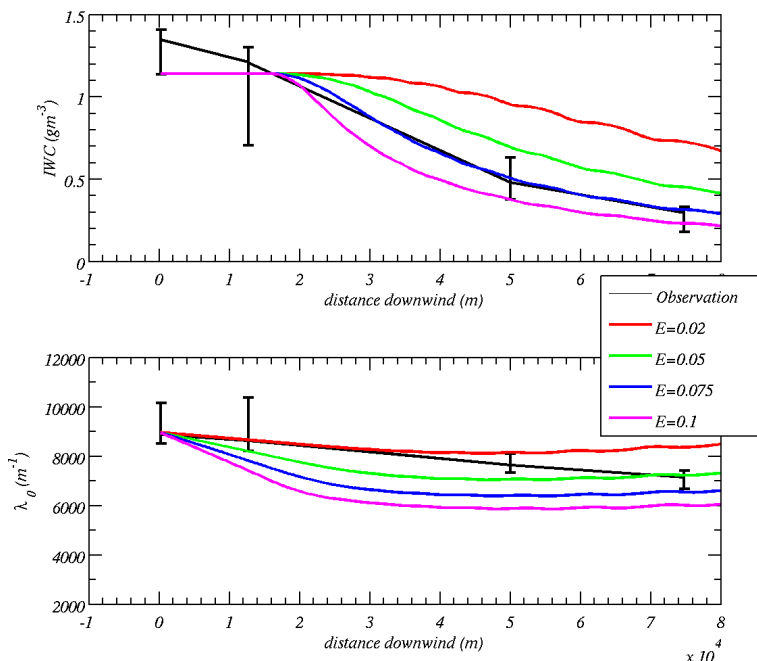


Fig. 13. (a) The IWC ($g\ m^{-3}$) inferred from the CIP instrument as a function of distance from the main storm and that predicted with the sedimentation model for differing collection efficiencies $E = 0.02, 0.05, 0.075,$ and 0.1 ; **(b)** Observed and predicted $\lambda_o\ (m^{-1})$ as a function of distance from the main storm.

23800