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# Evaluating the construction and evolution of upper crustal magma reservoirs with coupled U/Pb zircon geochronology and thermal modeling: A case study from the Mt. Capanne pluton (Elba, Italy) 

Mélanie Barboni ${ }^{1}$, Catherine Annen ${ }^{2}$ and Blair Schoene ${ }^{3}$<br>1 Department of Earth, Planetary and Space Sciences, University of California Los Angeles<br>Phone number: +1 (609) 510 4782; Fax number: +1 (310) 825 2279; Email: mbarboni@epss.ucla.edu<br>2 School of Earth Sciences, University of Bristol (UK)<br>3 Department of Geosciences, Princeton University

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#### Abstract

Evaluating mechanisms and rates for magma transport and emplacement in the upper crust is important in order to predict the thermal and rheological state of the crust, and understand the relationship between plutonism and volcanism. U-Pb geochronology on zircon is commonly used to constrain magma emplacement and storage time in the crust, but interpreting complex zircon age populations in terms of in-situ crystallization versus crystallization at a deeper level is not trivial. This study focuses on the Mt. Capanne pluton in Elba (Italy), a well-documented example of arc-related laccolith emplaced in the upper continental crust. Previous studies proposed that the Mt. Capanne intrusion was accreted in less than 10,000 years by distinct and mappable magma pulses. Here, we couple high-precision ID-TIMS U-Pb zircon geochronology with numerical


thermal simulations to evaluate emplacement rates, test different emplacement models, inform zircon age interpretations and evaluate the potential for melt storage during construction of the Mt. Capanne pluton. Our results require that the Mt. Capanne intrusion was built in at least 300,000 years by multiple magma injections. A variety of emplacement scenarios show that melt was preserved for $<60^{\prime} 000$ years after each pulse and that the maximum eruptible volumes were approximately equal to the volume of each pulse. Our results also require that the majority of zircon crystallization occurred in zircon saturated reservoirs at deeper crustal levels prior to final magma emplacement and cooling, which has implications for using zircon U-Pb geochronology to infer upper crustal magma residence times.

## 1. Introduction

The flux of magma to the upper crust plays an important role in how continental crust is made and the relationship between volcanic and plutonic systems (e.g. Lipman, 2007; Bartley, 2008; Glazner et al., 2004; Deering et al., 2011; Miller and Miller, 2002; Miller et al., 2011; Bachmann et al., 2007). Low upper crustal temperatures and the associated brittle rheology of host rocks limit the rate at which large magma bodies can be emplaced (Clemens, 1992; Coleman et al., 2004), which has led to the hypothesis that many igneous bodies grow incrementally by accretion of smaller sills and dikes (e.g. Coleman et al., 2004; Lipman, 2007; Menand, 2011). Therefore, geochronology that constrains magma emplacement rates into the upper crust is important for building models linking magma reservoirs with volcanic and plutonic products (e.g. Barboni and Schoene, 2014).

Plutons are time-integrated records of magma reservoirs residing beneath volcanic systems (e.g. Bachmann et al., 2007; de Silva and Gosnold, 2007; Bartley, 2008; Menand et al., 2015), and
deciphering intrusion geometries and emplacement modes is critical to testing models for volcanic-plutonic systems (Farina et al., 2010). However, interpreting plutonic records in the field is rarely straightforward, as critical information such as contacts between different sills or pulses and original magmatic textures (Bartley, 2008) can be obscured. Quantifying models for upper crustal pluton construction thus involves a combination of thermal and mechanical models tested by field observation, geophysics, petrology and geochronology.

Heat transfer calculations using a range of "typical" magma emplacement rates show that in the cool shallow crust successive magma pulses separated by long time intervals solidify before the next one is emplaced (Annen, 2009; Schöpa and Annen, 2013; de Saint Blanquat et al., 2010). This has led some authors to consider that upper crustal magma chambers with volumes equivalent to plutons are rare and short-lived (Glazner et al., 2004; Coleman et al., 2004; de Saint Blanquat et al., 2010). However, further evidence from thermal models (Annen, 2009; Gelman et al., 2013) has shown that high magma emplacement rates can result in magma chambers equivalent in size to large-volume ignimbrites associated with caldera collapse (Lipman, 2007; Bachmann and Bergantz, 2008). Testing model results with examples from the geologic record requires estimates of magma flux in the crust derived from geochronology (e.g. Caricchi et al., 2014). U-Th-Pb and U-series geochronology of magmatic accessory minerals have been used to infer rates of emplacement and melt storage in the crust (e.g. Coleman et al., 2004; Schmitt et al., 2011; Schoene et al., 2012; Barboni and Schoene., 2014). However, an increasing number of studies from both plutonic and volcanic rocks reveal complexities in zircon age populations that are difficult to interpret in terms of emplacement time (e.g. Lissenberg et al., 2009; Schoene et al., 2012). In essence, the ability of zircon to retain age information at magmatic temperatures which itself makes zircon such an attractive geochrometer in magmatic systems - also increases
its propensity to record protracted crystallization and recycling that obscure the connection between zircon date and magma emplacement rates.

In this study, we couple high-precision U-Pb zircon geochronology with thermal modeling to test models of pluton construction. We use differing interpretations of zircon dates to model magma emplacement rates, and use the model results to in turn place constraints on zircon age interpretations. Our case study is the Miocene Mt. Capanne intrusion located on Elba Island (Italy), which is a well-documented example of an upper-crustal laccolith (e.g. Westerman et al., 2004; Rocchi et al., 2010). Previous work has proposed that different mappable magma batches contributed to pluton construction (Farina et al., 2010), therefore providing an excellent opportunity to test and refine the incremental emplacement paradigm for upper crustal plutons and to measure the potential for such systems to generate and store eruptible liquid.

## 2. Geological background

Elba Island formed behind the eastwards-progressing compressive front of the Apennine orogeny (Malinverno and Ryan, 1986) during the stacking of five tectonic complexes in the early Miocene. The arc-related Elba intrusive complex was emplaced as a nested "Christmas-tree" laccolith at ca. 6 km depth within the tectonic complexes in what is now western and central Elba (Fig. 1; Rosenbaum and Lister, 2004; Gasparon et al., 2009; Bussy, 1991; Westerman et al., 2004; Rocchi et al., 2010). Crosscutting relationships show the following sequence of intrusion (Westerman et al., 2004; Farina et al., 2010; Rocchi et al., 2010; Fig.1): The Capo Bianco aplite, the Portoferraio porphyry, the San Martino porphyry, the ca. $\sim 200 \mathrm{~km}^{3}$ Mt. Capanne granitic pluton, and lastly late dykes of intermediate compositions (Orano dyke swarm; Fig.1; Dini et al., 2008).

Early intrusions such as the Capo Bianco aplite and the Portoferraio porphyry have been interpreted as anatectic crustal melts. Large chemical variability within the later intermediate to felsic pluton facies and within individual megacrystic K-feldspars have been used to generate models whereby the magma mixing and mingling occurred at lower crustal depths synchronous with upper crustal emplacement of the Mt. Capanne pluton (Dini et al., 2002; Dini et al., 2004; Farina et al., 2010; Gagnevin et al., 2004; Gagnevin et al., 2005; Gagnevin et al., 2008; Westerman et al., 2004). Several mixing models have been proposed to explain variation observed in the Elba magmatic system. Liquid-liquid mixing of mantle and crustal melts has been used to explain textures and chemical/isotopic zoning in plagioclase and K-feldspar within the Mt. Capanne granite (e.g. Bussy, 1991; Gagnevin et al., 2004; Dini et al., 2002). Previously hybridized products, represented in the field by abundant mafic microgranular enclaves in the Mt . Capanne facies, were also proposed as potential mixing sources (e.g. Bussy, 1991; Dini et al., 2002). Melt-solid interaction such as assimilation and fragmentation of cumulates has also been used to explain isotopic variability and observed feldspar xenocrysts (Bussy, 1991). Farina et al. (2012), however, proposed that much of the chemical and isotopic variability observed is related to progressive lower- to mid-crustal anatexis of isotopically heterogeneous sources, and that hybridization was most important in generating the intermediate Orano dykes and mafic enclaves of the Mt. Capanne granite.

## 3. Model for the Mt. Capanne pluton construction

The calc-alkaline Mt. Capanne pluton has monzogranitic composition $\left(\mathrm{SiO}_{2}\right.$ between 66 and 70 $\mathrm{wt} \%$ ) and slightly peraluminous character (ASI=1.11) (Dini et al., 2002; Farina et al., 2012). Farina et al. (2010) distinguish three geochemically- and texturally-distinct facies within the Mt.

Capanne pluton that correlate with the abundance of K-feldspar megacrysts. The Sant'Andrea facies (SA) contains the highest content of megacrysts and mafic enclaves and is mostly preserved along the pluton margin and at high elevation (Fig.1), suggesting it forms the roof of the pluton (Farina et al., 2010; Westerman et al., 2015). The San Piero (SP) is the structurally lowest facies and contains a low abundance of K-feldspar megacrysts and mafic enclaves. The two facies display subtle but systematic geochemical differences such as higher $\mathrm{SiO}_{2}$ contents and biotite Mg\# in the SA compared to the SP (Farina et al., 2010). The San Francesco facies (SF) is intermediate between the SA and SP in geochemical signature, K-feldspar megacryst content, and structural level. Westerman et al. (2015) interpreted the three facies as three distinct magma batches that were emplaced from the top down (i.e., under accretion) in the order SA, SF, SP. Injection of the SF and SP successively deformed the preexisting sheets into the geometry observed today (Farina et al., 2010; Westerman et al., 2015; Fig.1). Construction by underaccretion is supported by the absence of feeder dikes and contractional strain in the roof (Farina et al., 2010). The original diameters and thicknesses of the SA, SF and SP sills are estimated at $9.5 \times 0.25,9.0 \times 0.65$ and $8.0 \times 1.5 \mathrm{~km}$, respectively (Farina et al., 2010), yielding to volumes of about 17, 41 and $75 \mathrm{~km}^{3}$ (assuming a cylinder geometry; Table 2). Although the dimensions of SA and SF are derived from field observations, the thickness of the SP facies has been estimated through a magnetic data model (Dini et al., 2008). Only the top of the SP is outcropping and the buried portion may be composed of another facies, and the possibility that the Mt. Capanne intrusion is thicker than proposed by Dini et al. (2008) cannot be excluded at present.

A striking feature of the Mt. Capanne pluton is the absence of clear contacts between the different pulses of Farina et al., (2010). This led these authors to propose that melt was preserved between injections, and using a simple numerical model, concluded that the whole Mt. Capanne
intrusion was built in less than 10,000 years. Barboni and Schoene, (2014) showed that one facies of the Cappane (SA) intruded and cooled to below its solidus in $<40 \mathrm{ka}$, but could not further test the Farina et al. (2010) model for the whole pluton. A recent study by Gagnevin et al. (2011) presented U-Pb dates on zircon by Secondary Ion Mass Spectrometry (SIMS) from the Mt. Capanne pluton, and found no age differences within the pluton. Those data are consistent with rapid intrusion of the Mt. Capanne pluton but large uncertainties on single zircons $(\geq \pm 0.15$ Ma) prevent a rigorous testing of the intrusion model of Farina et al. (2010).

## 4. ID-TIMS U-Pb geochronology

## 4.1. sample description

In this study, we carry out $\mathrm{U}-\mathrm{Pb}$ geochronology on zircon using Isotope Dilution Thermal Ioniazation Mass Spectrometry (ID-TIMS), which provides the temporal resolution necessary to test incremental emplacement models for the Mt. Capanne pluton. Sample locations are shown in Fig. 1 and representative CL images of zircon are presented in Fig. 2. Additional CL images are given in Supplementary Materials. Detailed descriptions of hand samples, zircon populations, and field photos are given in the Supplementary Materials. One sample each of the Portoferraio (MB11-11) and San Martino (MB11-14) porphyries, as well as six samples from the Mt. Capanne intrusion and one sample from the Orano dykes were collected for U-Pb ID-TIMS analysis. Following the nomenclature of Farina et al., (2010), we collected two samples from the San Piero facies (MB11-1 and MB11-2, elevation of 150 and 350 m respectively), two from the San Francesco facies (MB12-4 and MB12-8, elevation of 608 and 1000 m respectively) and two from the Sant'Andrea facies (MB11-6 and MB12-9, elevation of 0 and 50 m respectively). Unfolding
the sills to estimate the geometry at the time of emplacement following the model of Farina et al. (2010), the SP samples would have been close to the same structural level. On the other hand, the SA and SF samples represent a range of depths within the sheets. MB11-6 would have been ca. 150 m higher than MB12-9, which is located close from the contact with the SF facies (Fig.1). MB12-4 would have been ca. 100 m higher than MB12-8 in the SF sheet. The Orano sample was collected from a dyke crosscutting the Sant'Andrea facies (Fig.1; Supplement Fig.1).

### 4.2 ID-TIMS U-Pb geochronology

A total of 164 zircons were measured for U-Pb ID-TIMS geochronology. CL imaging was performed on each grain before dissolution to identify old metamorphic cores and characterize growth zoning. When visible cores were present, zircons were broken so that only the tips were dissolved and analyzed (Fig. 2). Many zircons showed complex magmatic textures, but our dating of $\sim 200$ zircons from Elba (including the data presented in Barboni and Schoene, 2014) illustrates qualitatively that intragrain complexity does not correlate with date (Fig. 2; Supplementary Material). Fig. 1 presents the ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ dates for each sample. A detailed methodology section and data tables can be found in the Supplementary Material. All uncertainties in the text, figures, tables and supplements are reported to the 2 -sigma level and include internal uncertainties only.

Fourteen zircons were analyzed from the Portoferraio porphyry and the results show a spread of ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ dates ranging from $7.942 \pm 0.008$ to $8.009 \pm 0.012 \mathrm{Ma}(\Delta \mathrm{t}=67 \pm 14 \mathrm{ka})$. A cluster of 9 zircons yield a weighted mean of $8.001 \pm 0.002 \mathrm{Ma}$ (MSWD of 2.3 ), which does not overlap with the youngest zircon measured in the sample ( $7.942 \pm 0.008 \mathrm{Ma}$; Fig.1). Sixteen zircons were analyzed for the San Martino porphyry (data published in Barboni and Schoene, 2014). ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ ages spread between $7.437 \pm 0.011$ and $7.947 \pm 0.005 \mathrm{Ma}$, with most of the ages ranging between
$7.437 \pm 0.011$ to $7.541 \pm 0.006 \mathrm{Ma}(\Delta \mathrm{t}=105 \pm 12 \mathrm{ka})$. Weighted means yield unreasonably high MSWDs of $>100$ (Wendt and Carl, 1991). Three grains are distinctively older (7.783, 7.947 and 9.149 Ma).

Thirty-two zircons were analyzed for the Sant'Andrea facies of the Mt. Capanne pluton (twenty for MB11-6; previously published in Barboni and Schoene, 2014; and twelve for MB129; Fig.1). Both samples show a large spread in ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ dates with data spanning over 0.3 Ma $(\Delta \mathrm{t}=331 \pm 8 \mathrm{ka}$; Fig.1), from $7.236 \pm 0.005$ to $7.567 \pm 0.006 \mathrm{Ma}$ (ignoring older outliers at 7.65 and 7.87 Ma; Fig.1). MB12-9 yields younger dates compared to MB11-6 (youngest zircons of $7.236 \pm 0.005$ and $7.323 \pm 0.019 \mathrm{Ma}$, respectively).

Forty-three zircons were analyzed for the San Francesco Facies (twenty-seven for MB12-4 and sixteen for MB12-8; Fig.1). Most ${ }^{206} \mathrm{~Pb} / /^{238} \mathrm{U}$ ages span between $7.166 \pm 0.007$ and $7.404 \pm 0.005$ $\mathrm{Ma}(\Delta t=238 \pm 9 \mathrm{ka}$ ), with four older outliers in sample MB12-4 (from 7.563 $\pm 0.005$ to $9.316 \pm 0.006$ $\mathrm{Ma})$ and one in MB12-8 (7.715 $\pm 0.028 \mathrm{Ma})$. The youngest zircons from both samples overlap within uncertainties at $7.166 \pm 0.007 \mathrm{Ma}(\mathrm{MB} 12-4)$ and $7.171 \pm 0.005 \mathrm{Ma}$ (MB12-8).

Thirty-seven zircons were analyzed from the San Piero facies (fifteen for MB11-1 and twentytwo for MB11-2). The SP samples display the largest spread in ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ ages within the Mt. Capanne pluton ( $\Delta \mathrm{t}=531 \pm 9 \mathrm{ka}$ ), between $7.007 \pm 0.007$ and $7.538 \pm 0.005 \mathrm{Ma}$ (excluding older outlier at $7.72,7.81,12.06$ and 12.74 Ma ; Fig.1; Supplement Table 1). The youngest zircon from MB11-1 overlaps within uncertainties with the youngest zircon from MB11-2 (7.007 $\pm 0.007$ and $7.009 \pm 0.004 \mathrm{Ma}$, respectively).

Twenty-six zircons were measured for the Orano dyke sample (Fig.1). Similar to the Mt. Capanne intrusion samples, a large spread in ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ ages is observed ( $\Delta \mathrm{t}=423 \pm 9 \mathrm{ka}$ ), ranging from $7.080 \pm 0.005$ to $7.503 \pm 0.007 \mathrm{Ma}$ (with two older out layers at 7.73 and 7.90 Ma ).

All the samples measured in this study display a large spread in ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ dates of ca. 200400 ka (with average uncertainties of $\pm 8 \mathrm{ka}$ ), raising the question of the significance of any single date or series of dates in terms of magma emplacement and solidification. We suggest two endmember interpretations of the zircon record given the constraint that all analyzed zircons show magmatic zoning (Fig. 2) and therefore crystallized in the presence of melt: (1) the youngest date represents the best estimate for intrusion and the older dates correspond to zircon recycling from a deeper level of the system, and (2) the oldest grain represents post-emplacement zircon saturation and all other dates represent in situ growth. Below we use numerical thermal modeling to test these possibilities.

## 5. Thermal model setup

We numerically simulated the assembly of the Mt. Capanne pluton by addition of sills and calculated the evolution of temperature and melt fraction in the igneous body and country rock system. The equation of heat:

$$
\begin{equation*}
\rho c \frac{\partial T}{\partial t}+\rho L \frac{\partial f}{\partial t}=k \nabla^{2} T \tag{1}
\end{equation*}
$$

where $\rho$ is density, $c$ is specific heat, $T$ is temperature, $L$ is latent heat of crystallization or fusion, $f$ is melt fraction, and $k$ is thermal conductivity, was solved with an explicit finite difference scheme on a 2D numerical grid using cylindrical coordinates. The numerical domain extends from the surface to a depth of 13 km and radially to a distance of 10 km from the central axis of
symmetry. We used $\rho=2700 \mathrm{~kg} / \mathrm{m}^{3}$ (Rocchi et al., 2010), $c=1200 \mathrm{~J} / \mathrm{kg} \mathrm{K}$ (Robertson, 1988) and $L=3.5 \times 10^{5} \mathrm{~J} / \mathrm{kg}$ (Hale et al., 2007). The conductivity $k$ varies with temperature as in Whittington et al. (2009). More information on the numerical method can be found in Annen et al. (2008).

The top of the pluton was emplaced at 6 km depth (Rocchi et al., 2010 Farina et al., 2010). The initial system temperature as a function of depth is determined by linear geothermal gradients of either $25^{\circ} \mathrm{C} / \mathrm{km}\left(150^{\circ} \mathrm{C}\right.$ at 6 km depth $)$ or $40^{\circ} \mathrm{C} / \mathrm{km}\left(240^{\circ} \mathrm{C} / \mathrm{km}\right.$ at 6 km depth). Early and distant intrusions such as the Portoferraio and San Martino porphyries do not significantly affect crustal temperature at the time of Mt. Capanne emplacement, supported by the lack of contact metamorphism associated with those earlier sills.

Solving eq. 1 and calculating melt fraction with time requires knowledge of the temperaturemelt fraction relationship. We used the Rhyolite-MELTS phase equilibria package (Gualda et al., 2012) with the chemical composition of the SA, SF and SP facies as inputs (see Supplementary material for details on the MELTS model). Inaccuracies in Rhyolite-MELTS results have been reported for felsic, hydrated melts (Gardner et al., 2014), and so we compared a series of Rhyolite-MELTS outputs (by varying, e.g., pressure and water content) to petrological observations, modal proportions and microprobe mineral measurements made on Mt. Capanne granite thin sections (e.g., Barboni and Schoene, 2014). The runs that did not closely match observed records were discarded. The remaining runs that were used in our numerical simulations (Fig. 3) are in good agreement with experimental data acquired on similar granitic compositions (Naney, 1983; Whitney, 1988).

According to Rocchi et al., (2010) and Barboni and Schoene, (2014), the magma was rich in feldspar and quartz at the time of emplacement. The zircon saturation temperature was estimated at about $805{ }^{\circ} \mathrm{C}$ for all three Mt. Capanne facies at the emplacement level (see Supplementary

Material for details on the Zr saturation model). In a first series of simulations, we test the hypothesis that the magma is emplaced above zircon saturation temperature and that zircons are crystallised after emplacement. Accordingly, in those simulations the magma is emplaced with $20 \%$ crystals and emplacement temperatures of 878,906 , and $856{ }^{\circ} \mathrm{C}$ for $\mathrm{SA}, \mathrm{SF}$ and SP respectively (Fig.3). We determined the period $\tau$ during which zircon can crystallize by calculating the time spent by the magma between the saturation temperature and the solidus (688, 676 and $666{ }^{\circ} \mathrm{C}$ for SA, SF and SP facies respectively). We infer from those simulations (c.f. section 5.1 below) that zircon saturation preceded emplacement because the range of observed zircon dates is far longer than the time that liquid is present as permitted by the numerical models. In a second series of simulations, the magma is emplaced at $800^{\circ} \mathrm{C}$ with a crystal fraction of $40 \%$. In this case $\tau$ is the time spent by the magma above solidus.

The results are reported in Table 2 and 3 as $\tau_{\text {sample }}$, which is $\tau$ at an estimated sample paleodepth (Fig. 4 b ) and $\tau_{\text {batch }}$, which is the maximum $\tau$ within a given batch, so that:

$$
\begin{align*}
& \tau_{\text {sample }}=\tau\left(z=z_{\text {sample }}\right)  \tag{2}\\
& \tau_{\text {batch }}=\operatorname{Max}\left[\tau\left(z_{u}<z<z_{l}\right)\right] \tag{3}
\end{align*}
$$

with $z$, the depth, $z_{\text {sample }}$ the depth of a sample, $z_{u}$ the depth of a batch upper boundary and $z_{l}$ the depth of a batch lower boundary.

Since cooling times are the longest close to the batch center, $\tau_{\text {batch }}$ is significantly longer than $\tau_{\text {sample }}$ if a sample has been collected far away from the batch center. This is the case for SP (Fig. 4b).

We tested two scenarios for the construction of the Mt. Capanne pluton. In the first scenario, (3-batch scenario) we followed Farina et al., (2010) in assuming that SA, SF, and SP facies each correspond to a single magma batch that are respectively 250,650 and 1500 m thick (Fig.4a,

Table 2) with emplacement times determined by samples MB11-6 (SA), MB12-4 (SF) and MB11-2 (SP). In the second scenario, the magma emplacement sequence is determined by the zircon dates from samples MB11-6, MB12-9, MB12-8, MB12-4, and MB11-2, such that each sample is treated as an individual batch of magma (Fig.4b, Table 3). In other words, the second scenario assumes multiple pulses per facies, each one of them represented by one of our samples. The thicknesses of those pulses were approximated on the basis of the cross-section after removing the deformation of successive pulses (unfolding the cross-section on Fig.1). We did not include MB11-1 because in the unfolded intrusion it is located at the same vertical level than MB11-2 and hence should belong to the same sill. In all cases, we used a diameter for the intrusions of 9 km (Farina et al., 2010).

## 6. Numerical simulation results

### 6.1 Testing the significance of zircon date spectra

The numerical simulations were used to test the hypothesis that magma emplacement is recorded by the oldest zircon ages from each pulse, outliers excluded (Table 1), and that the range of zircon ages reflects in-situ crystallization. We decided to consider as outliers the grains that were older than the main population by a large gap in dates (Fig.1; Table 1).

As a preliminary test, we emplaced the entire Mt. Cappane pluton instantly and observed whether the cooling time between zircon saturation and solidus within the pluton matched the zircon age spectra. In this scenario (using an intrusion temperature of $906{ }^{\circ} \mathrm{C}$ and composition equivalent to the SF ), the maximum predicted range in zircon dates is 78 ka and 52 ka for a geotherm of 40 and $25^{\circ} \mathrm{C} / \mathrm{km}$, respectively (Fig.5a). This is much shorter than the observed
zircon age range ( $\Delta \mathrm{t}=560 \pm 10 \mathrm{ka}$ for the all Mt. Capanne intrusion; Fig. 1 and Table 1), and is thus insufficient to explain the zircon dataset.

In the 3-batch scenario, the oldest zircon date of two samples within a facies was initially taken as the emplacement age for the facies and emplacement is by underaccretation (Table 2). By contrast, in the 5-batch scenario, successive batches are injected randomly within the intrusive body rather than below, as dictated by the zircon dates (Fig.4a; Table 3), and thus there is no relationship between facies and timing. In this set of models, magmas were emplaced at low crystallinity $(20 \%)$ to test the duration of liquid residence from above zircon saturation to the solidus. Table 2 and 3 reports $\tau_{\text {batch }}$ the longest magma residence time between zircon saturation and solidus temperature for each of the 3 and 5 batches, respectively. For both scenarios, the maximum zircon crystallization time predicted by our simulation is less than 60,000 years in the SP facies (Table 3; Fig.5) and is much shorter than the range of observed zircon dates (minimum 0.3 Ma ; Fig. 1) even when a steep $40^{\circ} \mathrm{C} / \mathrm{km}$ geothermal gradient is considered. The modeled maximum zircon crystallization time in Sant'Andrea facies is 750 years for a 3-batch scenario and 1100 years in a 5-batch scenario and incommensurate with the observed spread in SA zircon dates ( $\Delta \mathrm{t}=331 \pm 8 \mathrm{ka} ;$ Fig.1). This discrepancy between the model results and zircon dates suggests that our working hypothesis - that the oldest zircon dates record emplacement and that the range of dates records continuous in situ crystallization - is not correct. The alternate hypothesis, that most of the zircons crystallized at depth and were recycled (antecrystic, using terminology of Miller et al., 2007), is more accurate.

### 6.2. Testing for the presence of melt between batches

Farina et al. (2010) argue that retaining melts between the three injections is necessary in order to erase contacts between the sheets. We tested their hypothesis in the light of our new temporal constraints. Based on the result described above, we now test the other end-member assumption that magma is emplaced at or below zircon saturation temperature (Figs. 6 and 7) and that the youngest zircon date from each facies or sample corresponds to the pulse emplacement time (Table 2 and 3). This results in emplacement by underaccretion for both construction scenarios. In these sets of models, magmas were emplaced at high crystallinity (40\%) to be in line with the Zr-saturation temperature (ca. $805^{\circ} \mathrm{C}$; Supplement Material) The simulations unequivocally show that for both the 3- and 5-batch scenarios, no melt remains in the system between the successive batches even with a high initial geotherm of $40^{\circ} \mathrm{C} / \mathrm{km}$ (Table 2 and 3; Fig. 6 and 7). The maximum volumes of melts present in the system do not exceed the volume of the last pulse (Fig.6A and 7), which indicates that remelting of former pulse is absent or very limited. For the 3-batch model, the geotherm is only slightly perturbed at the time of emplacement of San Francesco and San Piero (Fig.6B and 6B). Our simulations also show that increasing the number of batches (while still respecting the final volume and duration of intrusion of the Mt. Capanne pluton) would only give shorter liquid residence time, and would not satisfy the hypothesis of residual melt persistence. We concur with Farina et al. (2010) that fast emplacement of Mt. Capanne over less than 10,000 year is necessary for maintaining melt between pulses but our zircon ages indicate much longer emplacement timescale. Therefore, the integration of thermal models and zircon ages does not support the hypothesis that the absence of mappable contacts between facies in the pluton is due to the presence of melt between pulses.

### 6.3. The size of magma reservoirs

Both for the 3-pulse and 5-pulse scenarios each pulse completely solidifies before the next one so that the maximum volume of melt corresponds to the volume of individual pulses. According to our geochronologic data, and assuming that the youngest zircon has been sampled and dates emplacement, the minimum number of pulses is 4 . The youngest dates of MB12-4 and MB12-8 overlap and those two samples may belong to the same pulse. The largest possible pulse has the inferred thickness ( 1.5 km ) and diameter ( 8 km ) of the San Piero facies, which corresponds to a volume of about $75 \mathrm{~km}^{3}$. However, pulses may have been much more numerous, in which case any magma reservoir would have been of smaller volume.

## 7. Discussion

7.1 Constraints on pluton assembly provided by combined zircon geochronology and thermal modeling

### 7.1.1 Zircon sources and crystallization histories

The ID-TIMS U-Pb geochronology presented here illustrates that temporal resolution of tens of thousands of years is necessary to resolve the intrusion history of the Mt. Capanne pluton. Previous U-Pb data obtained by SIMS (Gagnevin et al., 2011) was consistent with a rapid emplacement of the pluton, but due to inherently larger uncertainties associated with that method ( $\geq \pm 150 \mathrm{ka}$ in Gagnevin et al., 2011) were unable to differentiate between the ages of different magma pulses and therefore could not test the top-down laccolith emplacement model (Farina et al., 2010). The drawback of ID-TIMS geochronology is the larger volume of zircon analyzed, which can integrate protracted zircon growth histories into a single precise crystallization date.

We show in this study that careful CL-imaging followed by microsampling can avoid analysis of complexities such as inherited cores (Fig.2; Supplementary Figure 2) in an effort to date the youngest zone of a zircon. However, we observe that removing and selectively analyzing zircon tips does not necessarily yield the youngest dates within a sample (Fig. 2; Supplementary Figure 2), suggesting that complexities in zircon saturation and nucleation at small spatial scales and/or grain armoring affects the zircon growth record.

Experimental data for zircon saturation show that granodioritic to granitic melts reach zircon saturation at temperatures well above the solidus and therefore could carry significant amounts of early-crystallized zircon during remobilization and transport of magma (Watson and Harrison, 1983; Boehnke et al., 2013; Harrison et al., 2007; Miller et al., 2007). As a consequence, the physical integrity and also age information in zircon can survive transport, reheating, and reincorporation in subsequent batches of magma, and interpreting zircon dates in terms of magmatic processes is difficult (Lissenberg et al., 2009; Miller et al., 2007; Schoene et al., 2012). One approach to this problem is to target zircon populations included within specific phases versus the bulk rock, and combine those data with petrologic observations and zircon saturation and phase equilibria modeling to estimate the solidification age of certain pulses (see Barboni and Schoene, 2014, for an example using the SA). Building on that approach for the entire pluton, we estimated zircon saturation temperatures for our samples by modeling the Zr and bulk composition evolution of the melt given by our Rhyolite-MELTS results (see Supplement for detailed methodology and figures) and using the zircon saturation models of Boehnke et al., (2013). Our results suggest that saturation was reached in the Mt. Capanne magmas at temperatures of ca. $805^{\circ} \mathrm{C}$, well above the modeled solidi of $688-666{ }^{\circ} \mathrm{C}$. Textures revealed by CL imaging (Fig. 2) also record resorption events within the Mt. Capanne zircons (see also Gagnevin et al., 2011), suggesting that zircon saturation was not constant over the time span of

200 to 400 ka recorded by zircon in our samples. This conclusion is supported by the results of Barboni and Schoene, (2014), who show that at least 100 ka of zircon crystallization is recorded in the SA prior to its emplacement, in that zircon dates from that pulse predate the intrusion of the demonstrably cross-cut San Martino porphyry by ca. 100 ka .

Our thermal modeling and geochonology combine these data to test whether the observed range of zircon dates ( $\sim 300 \mathrm{ka}$ ) from each sample could represent post-emplacement cooling. Because the maximum melt residence time determined in our models is $\sim 58 \mathrm{ka}$ for any batch of magma, it is unavoidable that at least ca. 200 ka of pre-intrusion zircon crystallization (i.e. zircons carried from depth) occurred and is recorded within the Mt Capanne pluton. Our results therefore limit the duration of a magma reservoir in the shallow Elba magmatic system and suggest that most of the zircons were recycled from a deeper crustal reservoir.

### 7.1.2 Construction of Mt. Capanne pluton

Farina et al., (2010) hypothesized that the SA, SF and SP facies correspond to distinct magma batches injected in the upper crust and contacts between them are absent because melt was preserved between magma pulses, which required pluton emplacement in less than 10 ka . However, the difference between the youngest zircons dated in this study from the SA (MB11-6) and the SP (MB11-2) exceeds 300 ka . A possibility is that the pluton was emplaced rapidly ca. 7.3 Ma (the youngest SA zircon) and that older SA zircons were inherited from deeper in the crust whereas SP zircons $<7.3 \mathrm{Ma}$ crystallized in situ post emplacement. We tested this hypothesis but found that the crystallization time of the SP in such a model is limited to 58 ka , requiring the SP to intrude at least 240 ka after the SA , containing a substantial amount of inherited zircon (Fig. 5 and 6). For this reason, our zircon U-Pb data and thermal modeling show
that the conclusion of Farina et al. (2010), that the Mt. Capanne pluton must have been emplaced in $<10 \mathrm{ka}$, is incorrect; this incorrect assertion stems from the assumption that melt was present between pulses, explored below.

Our thermal modeling results therefore argue that the timing of magma emplacement for a given pulse (as represented by a handsample) is closer to the youngest zircon date, and that the intrusion time is constrained by the youngest zircon date and the limits of the thermal model. Though the exact number of pulses is not determined by our data, we can reach several important conclusions and highlight remaining uncertainties. The SA was constructed by under-accretion of at least two different pulses 150 vertical meters apart represented by samples MB11-6 and MB12-9 (Fig.4b). Our two samples from the SF, MB12-4 and MB12-8, have youngest zircons that are nearly indistinguishable though MB12-4 was structurally 100 m higher than MB12-8. The large gap in dates between the youngest and the second youngest zircon in MB12-4 imposes significant uncertainty on whether the youngest zircon is in fact representative of the magma solidus. Therefore any conclusion as to whether the SF intruded as one or more than one pulse based on zircon dates is speculative. Similarly, zircon dates from the SP samples are very similar and the youngest dates overlap within uncertainty. However, the SP outcrops sampled were emplaced at a similar structural level, and may represent only one of potentially many sills. Therefore, while collecting more samples within the Mt. Capanne cross-section would allow better estimation of the number of pulses, our main conclusions would not likely change. These are that our zircon data coupled with thermal modeling require that the Mt. Capanne pluton intruded in at least 4 pulses (potentially many more) over $\sim 250$ ka and that the maximum liquid residence time is $\sim 60$ ka. Our data support previous models for pluton assembly by underaccretion (Farina et al., 2010), similar to some other recently studied intrusions (de Saint-Blanquat et al., 2006; Michel et al., 2008; Barboni et al., 2013).

### 7.2 Effect of unknown pluton geometry on the numerical simulation results

Increasing the volume of magma underplated beneath the exposed pluton (i.e. increasing the thickness of the San Piero facies) could effect the melt residence time in the deepest part of the intrusion. In this case, the outcropping part of the Mt. Capanne would represent only the roof of a larger magmatic system extending at depth. This possibility is in contradiction with Dini et al. (2008), who used a detailed magnetic susceptibility survey of the Mt. Capanne pluton to resolve the laccolith pluton shape with a maximum thickness of ca. 2.5 km (thickness used in our simulation), but we consider the possibility nonetheless. Our geochronologic data for the SA facies requires solidification prior to the intrusion of the SF and SP facies, consistent with the results from Barboni and Schoene (2014), and without subsequent remobilization. An early SA crystallization is also suggested by our Orano Dyke zircon dates, which are younger than those from the SA facies it crosscuts (Supplement Fig.1), but much older that the youngest San Piero zircon (7.080 $\pm 0.005 \mathrm{Ma}$ for Orano versus $7.007 \pm 0.007 \mathrm{Ma}$ for SP; Fig.1). If the San Piero facies was much thicker than predicted by Dini et al., (2008), its thermal effect did not significantly reheat the upper part of the intrusion.

Another unknown that could affect our numerical results is the possibility that part of the intrusion roof (represented in our case by the SA facies) was eroded away. Our models show that the maximum melt residence time for the SA facies is ca. 1 ka , which is much shorter than the prediction from Barboni and Schoene (2014) based on the information recorded by the SA Kfeldspar megacrysts (10-40 ka). This suggests a SA facies thicker than observed today, which potentially biases our thermal model outputs by underestimating the original volume of the pulse.

### 7.3 Absence of internal contacts within the Mt. Capanne pluton

Integrating age data with numerical modeling shows that for the simulations we ran, melt was not preserved between magma batches and therefore the absence of internal contacts within the Mt. Capanne intrusion is not related to the presence of residual melt. Alternative hypotheses that could explain the lack of internal contacts within an incrementally built granitic pluton include (1) remelting of previously emplaced pulses by a new injection that obscures contacts (Bartley, 2008), and (2) sustained amphibolite facies conditions triggering subsolidus textural change and erasing contacts (e.g. Hanson and Glazner, 1995). Farina et al. (2010) discarded the latter hypothesis based on lack of contact metamorphism in the Mt. Capanne host rock and absence of macro- or microscopic recrystallization evidence in the Mt. Capanne granite. The results of Barboni and Schoene (2014), which show core-to-rim younging in zircon included within megacrystic K-feldspar, require rapid cooling and solidification of early pulses without substantial textural modification. Numerical models also predict only slightly perturbed geotherms by successive injections at the depth and of the size of the Mt. Capanne pluton. Another hypothesis for obscuring internal contacts is that (3) contacts between magma of similar composition and texture, as is the case in the Mt. Capanne intrusion, could be difficult to identify in the field if there is no major changes in mineralogy, mineral modal proportions or mineral size. Chilled margins (i.e. reduction of the mineral size approaching the contact with the cold facies) are usually the best way to identify contacts between various injections. However, if the temperature gradient between the new injection and the already solidified pulse is not large enough (as is usually the case for intermediate to felsic melts), then chilled margins may not be expected. While our thermal models preclude large scale reheating or remelting of previous sheets, heating and or physical abrasion (plucking, remobilization) of contacts at a centimeter to
meter scale are possible and would act to obscure pulse contacts. Though our geochronological data were able to identify at least 4 pulses, it is possible that the Mt. Capanne pluton is composed of many more sheets with gradational composition whose contacts could be obscured by these processes. Other field examples where 100 m to km scale 3D cross-sections are visible show that it is possible to identify subtle contacts within similar composition sheets (e.g. Torres de Paines laccoliths; Michel et al., 2008; Himalayan leucogranites; Searle et al., 2010), but such relationships are elusive at the outcrop level (Bartley, 2008). We suspect the Mt. Capanne pluton is one of these cases where the compositions and textures of the different increments are too similar to generate obvious contacts in the field.
7.4 Eruptible volumes for the upper-crustal Mt. Capanne reservoir and implication for modern arc volcanism.

A difficulty in understanding active volcanic systems is that very little information is available about the longevity and size of subvolcanic reservoirs. "Fossilized" reservoirs such as the Mt. Capanne pluton can be used to constrain emplacement rates and reservoir lifespan, and contrast volcanic and plutonic records. Calc-alkaline magmatism in Elba occurred in a similar tectonic context as modern arc systems (Rosenbaum and Lister, 2004; Gasparon et al., 2009). Our thermal models show that small upper-crustal systems such as the Mt. Capanne pluton do not contain magma chambers on timescales of hundreds of ka, regardless of the intrusion mechanism. For the scenarios modeled, each pulse injected in the Mt. Capanne system solidifies before the next one intrudes, and therefore the maximum eruptible volume for the system is the volume of each injection. Those volumes range from 2 to $60 \mathrm{~km}^{3}$, depending upon the crystallinity of the magma at time of melt extraction and eruption (eruption between $40 \%$ and $60 \%$ crystals; e.g. Bachmann,
2004) and are similar to those observed in recent arc eruptions (Mt. Pinatubo 1991, 8-10 $\mathrm{km}^{3}$, Wolf et Hoblitt, 1996; Mt. Saint Helens 1980, $1.25 \mathrm{~km}^{3}$, Tilling, 1984). If magma was indeed emplaced at about $40 \%$ crystals, then our results show that melt remained eruptible for only a very short period of time ( $<50 \mathrm{ka}$, but for most pulses in less than several ka ), which is in line with the conclusions of Barboni and Schoene (2014) and Cooper and Kent (2014). Volatiles would also have a strong effect on the eruptibility of the system but their nature and role are unfortunately currently poorly understood in Elba.

## 8. Conclusions

Our study on the Mt. Capanne intrusion from Elba island shows that coupling a large dataset of high-precision ages with thermal modeling can help assess the emplacement history and thermal evolution of incrementally built upper-crustal reservoirs. Our results suggest that the Mt. Capanne intrusion was built in minimum 250,000 years by multiple magma increments in the crust. Numerical models of sill accretion constrained by U-Pb ID-TIMS dates indicate that no melt was preserved between individual injections. The thermal contribution of each pulse is not large enough to generate long-lasting magma reservoirs and the maximum volumes that can be erupted correspond to the volume of each pulse injected. These results provide interesting insight into understanding active volcanism in modern arcs. Elba magmatism was very similar, both in tectonic context and magma volumes, to some modern arc volcanoes (e.g. Mt. St. Helen; Mt. Pinatubo) and would have produced similar eruptible volumes $\left(<10 \mathrm{~km}^{3}\right)$. While active magma chambers can only be assessed indirectly, information recorded in a "fossilized" reservoir such as the Mt. Capanne can give insight into time-integrated rates of magma recharge and duration of potential eruption windows in active system.

We also show that absence of contacts within intermediate to granitic intrusions does not require melt preservation in the contact zones between magma sheets, as was previously proposed by Farina et al. (2010) for the Mt. Capanne intrusion. Our thermal models unambiguously indicate that every new magma increment will have a very short liquid residence time in the upper crust, with no melt remaining in-between pulses. Contacts between sheets of similar texture and composition might be unidentifiable at the outcrop scale, as suggested by Bartley et al., (2008). In such cases, the use of high-precision dating on zircon can discriminate between different magma injections (Schoene et al., 2012).

Complexities in $\mathrm{U}-\mathrm{Pb}$ age populations resulting from zircon recycling introduce complexities in resolving incremental emplacement in evolved melts. However, coupling zircon ages with thermal modeling aids interpretation of complex age spectrums and identifying zircon inherited from deeper crustal levels. Our results suggest that for the Mt. Capanne pluton, the majority of the zircon record is inherited from a deeper level of the system and that only a small portion of the ages record in-situ crystallization.

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## Figure captions

Figure 1. Study area and ID-TIMS U-Pb geochronology. A) Geological map of Elba with sample locations and cross-section of the Mt. Capanne pluton (modified after Westerman et al., 2004 and Farina et al., 2010). Numbers I-V refer to tectonic complexes (as described in Westerman et al., 2004). B) Rank-order plot of ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ zircon dates from the Elba intrusives (with youngest zircon age indicated for each samples with 2-sigma uncertainties). Individual bars correspond to single zircons or zircon fragments with height of bars representing 2-sigma uncertainties. See Table S 1 for full $\mathrm{U}-\mathrm{Pb}$ data table.

Figure 2. Cathodoluminescence images of selected zircons from the Elba intrusive rocks, with ID-TIMS ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ dates and 2-sigma uncertainties. Sample name and facies labeled in each panel. White dashed line indicates where grains were fractured and solid line points to which fragment was analyzed. See Table S 1 for full $\mathrm{U}-\mathrm{Pb}$ data table.

Figure 3. Melt fraction and proportion of crystallizing phases versus temperature used in the numerical simulations for the three Mt. Capanne facies, as predicted by Rhyolite-MELTS. The modal mineralogy observed in thin section is plotted on the left (colour boxes); colour curves represent modal proportions predicted by MELTS. The zircon saturation temperatures were determined for the evolving melt composition using MELTS and mineral-melt Zr partitioning (Supplementary Methods).

Figure 4. Setup for the numerical simulation for the 3-pulse (A) and 5-pulse (B) scenario. Successive magma pulses have the geometry of sills. The colors represent the different facies (Fig. 1), whose thicknesses are determined by field relations; see text. The stars show the approximate position of the $\mathrm{U}-\mathrm{Pb}$ geochronology samples.

Figure 5. Numerical thermal model results for simulations where the oldest zircon dates approximate intrusion of magma pulse for the 1-pulse (A), 3-pulse (B) and 5-pulse (C) scenarios. Here magma emplacement temperature is above zircon saturation temperature (see text). The colours and contour lines show the $\log 10$ of the time (i.e., $4=10,000$ years) spent by the magma between zircon saturation $\left(\sim 805^{\circ} \mathrm{C}\right)$ and the solidus on a cross section through the intrusion. Initial geothermal gradient is $40^{\circ} \mathrm{C} / \mathrm{km}$.

Figure 6. Numerical thermal model results for simulations where the youngest zircon dates approximate intrusion of magma pulse for the 3-pulses scenario. Magma emplacement temperature is below zircon saturation temperature at $\sim 40 \%$ crystals. (A) Volumes of melt over time. The spikes correspond to intrusion of SF and SP. SA is intruded at time 0 and the magma is too short-lived to be visible on the diagram. (B) and (C) Snapshot of temperatures on a cross section of the system at time 70,300 years (B) and 229,500 years (C) just before the emplacement of SF (B) and SP (C). Curves are labelled in ${ }^{\circ} \mathrm{C}$. The dashed line show the contour of already emplaced pulses and the arrows show the level of the next pulse. In both cases, temperatures at the time and depth of the new pulse are several hundred degrees below the wet granite solidus (Withney, 1988) and no melt remains in the system between pulses. Initial geothermal gradient is $40^{\circ} \mathrm{C} / \mathrm{km}$.

Figure 7. Numerical thermal model results for simulations where the youngest zircon dates approximate intrusion of magma pulse for the 5-pulses scenario. Curves showing volume of magma with $>50 \%$ and $>0 \%$ melt in the entire modelled magmatic systems are shown as a function of time. Each spike corresponds to a pulse but the first pulse emplaced at time 0 is not visible because it is too short-lived. No melt is retained between pulses. Initial geothermal gradient is $40^{\circ} \mathrm{C} / \mathrm{km}$.


B






A B


Distance from axis (km)



A



Table 1 Zircon dates used in the thermal model

| Facies | sample | Youngest zircon <br> $(\mathbf{M a})$ | 2-sigma <br> $\mathbf{( M a )}$ | Oldest zircon <br> $\mathbf{( M a )}$ | 2-sigma <br> $\mathbf{( M a )}$ | Total age dispersion <br> $\mathbf{( k a )}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  |
| SA | MB11-6 | 7.323 | 0.019 | 7.567 | 0.006 | $244 \pm 20$ |
| SA | MB12-9 | 7.236 | 0.005 | 7.406 | 0.005 | $170 \pm 7$ |
| SF | MB12-4 | 7.166 | 0.007 | 7.471 | 0.009 | $305 \pm 11$ |
| SF | MB12-8 | 7.172 | 0.005 | 7.404 | 0.003 | $232 \pm 5$ |
| SP | MB11-2 | 7.007 | 0.007 | 7.411 | 0.005 | $404 \pm 9$ |

## Table 2 Three batches emplacement

| Pulse | Pulse thickness <br> $\mathbf{( m )}$ | Facies | Volume <br> $\left(\mathbf{k m}^{\mathbf{3}} \mathbf{)}\right.$ |
| :---: | :---: | :---: | :---: |
|  |  |  |  |
| 1 | 250 | SA | 17 |
| 2 | 650 | SF | 41 |
| 3 | 1500 | SP | 75 |

Case 1: Magma emplaced above ziron saturation temperature (20\% crystals); the oldest zircon dates emplacement.

| Pulse | Ages <br> (Ma) | Sample | Emplacement time year | $\begin{gathered} \tau_{\text {sample }} \text { (year) } \\ \text { Geotherm }=25^{\circ} \mathrm{C} / \mathrm{km}(1) \\ \hline \end{gathered}$ | $\begin{gathered} \tau_{\text {sample }} \text { (year) } \\ \text { Geotherm }=40^{\circ} \mathrm{C} / \mathrm{km}(1) \\ \hline \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 7.567 | MB11-6 | 0 | 523 | 732 |
| 2 | 7.471 | MB12-4 | 96317 | 1759 | 3576 |
| 3 | 7.411 | MB11-2 | 156817 | 5673 | 20975 |
| Pulse | Ages <br> (Ma) | Sample | $\begin{gathered} \tau_{\text {batch }} \text { (year) } \\ \text { Geotherm }=25^{\circ} \mathrm{C} / \mathrm{km}(2) \end{gathered}$ | $\begin{gathered} \tau_{\text {batch }}(\text { year }) \\ \text { Geotherm }=40^{\circ} \mathrm{C} / \mathrm{km}(2) \end{gathered}$ | Volumetric emplacement rate ( $\mathrm{km}^{3} \mathbf{y r}^{-1}$ ) |
| 1 | 7.567 | MB11-6 | 567 | 775 | $1.25 \mathrm{E}-04$ |
| 2 | 7.471 | MB12-4 | 4217 | 6150 | 4.26E-04 |
| 3 | 7.411 | MB11-2 | 32895 | 52060 | $4.78 \mathrm{E}-04$ |

Case 2: Magma emplaced below zircon saturation temperature (40\% crystals); the youngest zircon dates emplacement

| Pulse | Ages <br> (Ma) | Sample | Emplacement time year | $\begin{gathered} \tau_{\text {sample }} \text { (year) } \\ \text { Geotherm }=25^{\circ} \mathrm{C} / \mathrm{km}(1) \\ \hline \end{gathered}$ | $\begin{gathered} \tau_{\text {sample }}(\text { year }) \\ \text { Geotherm }=40^{\circ} \mathrm{C} / \mathrm{km}(1) \\ \hline \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 7.236 | MB12-9 | 0 | 141 | 165 |
| 2 | 7.166 | MB12-4 | 70300 | 440 | 1098 |
| 3 | 7.007 | MB11-2 | 229500 | 555 | 1593 |
| Pulse | Ages <br> (Ma) | Sample | $\begin{gathered} \tau_{\text {tatch }} \text { (year) } \\ \text { Geotherm }=25^{\circ} \mathrm{C} / \mathrm{km}(2) \\ \hline \end{gathered}$ | $\begin{gathered} \tau_{\text {batch }} \text { (year) } \\ \text { Geotherm }=40^{\circ} \mathrm{C} / \mathrm{km}(2) \end{gathered}$ | Volumetric emplacement rate ( $\mathbf{k m}^{3} \mathbf{y r}^{-1}$ ) |
| 1 | 7.236 | MB12-9 | 574 | 763 | $8.47 \mathrm{E}-05$ |
| 2 | 7.166 | MB12-4 | 4185 | 5433 | 5.83E-04 |
| 3 | 7.007 | MB11-2 | 24490 | 32144 | $3.27 \mathrm{E}-04$ |

(1) time spent by the magma between zircon saturation temperature and solidus at the sample paleodepth
(2) maximum time spent by the magma between zircon saturation temperature and solidus anywhere within a batch (see text)

Since cooling times are the longest close to the batch center, $\boldsymbol{T}_{\text {batch }}$ is significantly longer than $\boldsymbol{T}_{\text {sample }}$ if a sample has been collected far away from the batch center.

## Table 3 Five batches emplacement

Case 1: Magma emplaced above zircon saturation temperature; the oldest zircon dates emplacement

| Pulse | Facies | sample | Ages <br> (Ma) | Pulse thickness (m) | Volume $\left(\mathrm{km}^{3}\right)$ | Sample depth (m) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | SA | MB11-6 | 7.567 | 150 | 9 | 6050 |
| 2 | SF | MB12-4 | 7.471 | 200 | 6 | 6300 |
| 3 | SP | MB11-2 | 7.411 | 1500 | 13 | 6950 |
| 4 | SA | MB12-9 | 7.406 | 100 | 27 | 6200 |
| 5 | SF | MB12-8 | 7.404 | 450 | 75 | 6600 |
| Pulse | Facies | sample | Ages <br> (Ma) | Emplacement time year | $\begin{gathered} \tau_{\text {sample }} \text { (year) } \\ \text { Geotherm }=25^{\circ} \mathrm{C} / \mathrm{km}(1) \\ \hline \end{gathered}$ | $\begin{gathered} \tau_{\text {sample }} \text { (year) } \\ \text { Geotherm }=40^{\circ} \mathrm{C} / \mathrm{km}(1) \end{gathered}$ |
| 1 | SA | MB11-6 | 7.567 | 0 | 236 | 310 |
| 2 | SF | MB12-4 | 7.471 | 96300 | 271 | 2132 |
| 3 | SP | MB11-2 | 7.411 | 156800 | 44396 | 58548 |
| 4 | SA | MB12-9 | 7.406 | 161900 | 803 | 1133 |
| 5 | SF | MB12-8 | 7.404 | 163700 | 21050 | 30012 |
| Pulse | Facies | sample | Ages <br> (Ma) | $\begin{gathered} \tau_{\text {batch }}(\text { year }) \\ \text { Geotherm }=25^{\circ} \mathrm{C} / \mathrm{km}(2) \end{gathered}$ | $\begin{gathered} \tau_{\text {batch }} \text { (year) } \\ \text { Geotherm }=40^{\circ} \mathrm{C} / \mathrm{km}(2) \end{gathered}$ | Volumetric emplacement rate ( $\mathrm{km}^{3} \mathrm{yr}^{-1}$ ) |
| 1 | SA | MB11-6 | 7.567 | 236 | 314 | 6.63E-05 |
| 2 | SF | MB12-4 | 7.471 | 8486 | 15112 | 1.35E-04 |
| 3 | SP | MB11-2 | 7.411 | 52167 | 65686 | 4.78E-04 |
| 4 | SA | MB12-9 | 7.406 | 803 | 1133 | 3.71E-05 |
| 5 | SF | MB12-8 | 7.404 | 37200 | 48670 | $1.65 \mathrm{E}-04$ |

(1) time spent by the magma between zircon saturation temperature and solidus at the sample paleodepth
(2) maximum time spent by the magma between zircon saturation temperature and solidus anywhere within a batch (see text)

Since cooling times are the longest close to the batch center, $\boldsymbol{T}_{\text {batch }}$ is significantly longer than $\boldsymbol{T}_{\text {sample }}$ if a sample has been collected far away from the batch center.

## Table 3 (cont) Five batches emplacement

Case 2: Magma emplaced below zircon saturation temperature; the youngest zircon dates emplacement

| Pulse | Facies | sample | Ages <br> (Ma) | Pulse thickness (m) | Volume (km ${ }^{3}$ ) | Sample depth (m) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | SA | MB11-6 | 7.323 | 150 | 9 | 6050 |
| 2 | SA | MB12-9 | 7.236 | 100 | 6 | 6200 |
| 3 | SF | MB12-4 | 7.166 | 200 | 13 | 6300 |
| 4 | SF | MB12-8 | 7.172 | 450 | 27 | 6600 |
| 5 | SP | MB11-2 | 7.007 | 1500 | 75 | 6950 |
| Pulse | Facies | sample | Ages <br> (Ma) | Emplacement time year | $\begin{gathered} \tau_{\text {sample }} \text { (year) } \\ \text { Geotherm }=25^{\circ} \mathrm{C} / \mathrm{km}(1) \\ \hline \end{gathered}$ | $\begin{gathered} \tau_{\text {sample }} \text { (year) } \\ \text { Geotherm }=40^{\circ} \mathrm{C} / \mathrm{km}(1) \\ \hline \end{gathered}$ |
| 1 | SA | MB11-6 | 7.323 | 0 | 248 | 299 |
| 2 | SA | MB12-9 | 7.236 | 89600 | 98 | 122 |
| 3 | SF | MB12-4 | 7.166 | 159900 | 405 | 500 |
| 4 | SF | MB12-8 | 7.172 | 165500 | 2518 | 3297 |
| 5 | SP | MB11-2 | 7.007 | 333200 | 814 | 1424 |
| Pulse | Facies | sample | Ages <br> (Ma) | $\begin{gathered} \tau_{\text {batch }} \text { (year) } \\ \text { Geotherm }=25^{\circ} \mathrm{C} / \mathrm{km}(2) \\ \hline \end{gathered}$ | $\begin{gathered} \tau_{\text {batch }} \text { (year) } \\ \text { Geotherm }=40^{\circ} \mathrm{C} / \mathrm{km}(2) \\ \hline \end{gathered}$ | Volumetric emplacement rate ( $\mathrm{km}^{3} \mathrm{yr}^{-1}$ ) |
| 1 | SA | MB11-6 | 7.323 | 248 | 299 | 4.49E-05 |
| 2 | SA | MB12-9 | 7.236 | 98 | 122 | $6.70 \mathrm{E}-05$ |
| 3 | SF | MB12-4 | 7.166 | 409 | 503 | 8.13E-05 |
| 4 | SF | MB12-8 | 7.172 | 2636 | 3438 | 1.63E-04 |
| 5 | SP | MB11-2 | 7.007 | 24314 | 30783 | 2.25E-04 |
| (1) <br> (2) | time spent by the magma above solidus at the sample paleodepth |  |  |  |  |  |

Since cooling times are the longest close to the batch center, $\boldsymbol{\tau}_{\text {batch }}$ is significantly longer than $\boldsymbol{T}_{\text {sample }}$ if a sample has been collected far away from the batch center.

## S1. Hand sample and zircon description:

## Portoferraio and San Martino Porphyries:

The Portoferraio sample was collected near the "Aquavita" locality (Fig. 1 of the main text). Latitude and Longitude (WGS84) are $\mathrm{N} 42^{\circ} 49^{\prime} 21.6^{\prime \prime} / \mathrm{E} 10^{\circ} 17^{\prime} 13.6^{\prime \prime}$. It is a biotite-bearing monzogranite containing quartz phenocrysts set in a very fine matrix (Supplement Fig.1A). Zircons are very abundant, included both in the matrix and as inclusions in biotite. The present a euhedral and prismatic shape and range in size from $50-300 \mu \mathrm{~m}$ in length, but mostly 70-200 $\mu \mathrm{m}$. Cathodoluminescence (CL) imaging of zircon shows oscillatory zoning typical of igneous zircon, with some grains displaying rounded cores and truncated oscillatory zoning or sector zoning (Fig. 2 of the main text).
Barboni and Schoene, (2014), described the San Martino (SM) porphyry sample (MB11-14). It was collected near the "La Focce" locality (Fig. 1 of the main text). Latitude and longitude (WGS84) are: N $42^{\circ} 74$ ' 65.4 " /E $10^{\circ} 25^{\prime} 02.8^{\prime \prime}$. It is a biotite-bearing monzogranite containing prominent sanidine megacrysts set in in a finegrained groundmass (Supplement Fig.1B). The zircon grains are euhedral and prismatic ranging in size from $50-400 \mu \mathrm{~m}$ in length, but mostly 100-250 $\mu \mathrm{m}$. Cathodoluminescence (CL) imaging of zircon shows oscillatory zoning typical of igneous zircon, though some grains contain rounded cores with truncated oscillatory zoning or sector zoning (Fig. 2 of the main text).

## Orano Dyke:

We collected the Orano sample from a dyke that was crosscutting the Sant'Andrea facies of the Mt. Capanne intrusion, West of the town of "Chiessi (Fig. 1 of the main text; supplement Fig.3C). Latitude and longitude (WGS84) are $\mathrm{N} 42^{\circ} 75^{\prime} 89.8^{\prime \prime} / \mathrm{E} 10^{\circ} 29^{\prime} 68.2$ ". It is a quartz monzodiorite containing resorbed xenocrysts of quartz and K-feldspar, as well as mafic microgranular enclaves (MME) and xenoliths (insert in Supplement Fig.3C). Zircon is located in the groundmass and in inclusion in biotite. Although the largest enclaves and xenocrysts from the sample prior to crushing and zircon extraction, some smaller ones were unavoidable during crushing. The zircons therefore represent a mixed population including grains from the host granite, the MME and smaller megacrysts. The zircon grains are euhedral and prismatic ranging in size from 50-200 $\mu \mathrm{m}$ in length, but mostly $50-150 \mu \mathrm{~m}$. Cathodoluminescence (CL) imaging of zircon shows oscillatory zoning typical of igneous zircon, though some grains contain rounded cores with truncated oscillatory zoning or sector zoning (Fig. 2 of the main text).

## Mt. Capanne pluton:

The Sant'Andrea (SA) facies rock samples were collected on the Sant'Andrea beach for MB11-6 (Fig. 1 of the main text; WGS84 latitude and longitude of N $42^{\circ} 80^{\prime} 80.1^{\prime \prime} / \mathrm{E} 10^{\circ} 14^{\prime} 09.1^{\prime \prime}$; described in Barboni and Schoene, 2014), and in the Pomonte quarry for MB12-9 (Fig. 1 of the main text; WGS84 latitude and longitude of N $42^{\circ} 75^{\prime} 08.4^{\prime \prime} / \mathrm{E} 10^{\circ} 12^{\prime} 93.9^{\prime \prime}$ ). Both are biotite-bearing monzogranites that contains numerous mafic microgranular enclaves (MME) and K-feldspar megacrysts (Supplementary Fig.1D). Modal proportions are the following: $38 \% \mathrm{An}_{35-12}$ plagioclase, $27 \%$ quartz, $22 \%$ orthoclase $\left(\mathrm{Or}_{65-81}\right), 13 \%$ biotite and accessories (apatite, zircon, tour-
maline, allanite, titanite and oxides). Zircon is located both in the matrix and included in plagioclase, orthoclase and biotite. Although the largest MME and K-feldspar were removed from the sample prior to crushing and zircon extraction, some smaller ones were unavoidable during crushing. The zircons therefore represent a mixed population including grains from the host granite, the MME and smaller megacrysts. Zircons are euhedral and prismatic, mostly elongate; grain size ranges from 50-500 $\mu \mathrm{m}$ in length, but are mostly 100-300 $\mu \mathrm{m}$. Cathodoluminescence (CL) imaging of zircon shows oscillatory zoning typical of an igneous origin. Most of the Sant'Andrea zircon grains have rounded cores with truncated oscillatory zoning or sector zoning, sometimes showing a patchy texture (Fig. 2 of the main text).
The San Francesco (SF) facies rock samples were collected near the contact with the San Piero facies on the road above the "Torre Giovanni" for sample MB12-8 (Fig. 1 of the main text; WGS84 latitude and longitude of N $42^{\circ} 76^{\prime} 93.2^{\prime \prime} / \mathrm{E} 10^{\circ} 18^{\prime} 77.9^{\prime \prime}$ ), and 50 m bellow the summit of the Mt.Capanne for sample MB12-4 (Fig. 1 of the main text; WGS84 latitude and longitude of N $42^{\circ} 77^{\prime} 09.5^{\prime \prime} / \mathrm{E} 10^{\circ} 16^{\prime} 89.5^{\prime \prime}$ ). Both samples are biotite-bearing monzogranites that are very similar in compositions and modal proportion as the Sant'Andrea samples described above, at the exception of a lower content of MME and K-feldspar megacrysts and a slightly lower $\mathrm{SiO}_{2} \mathrm{wt} \%$ composition (Supplement Fig.1E). Zircons from both samples present similar characteristics that the one from the two SA samples.
The San Piero (SP) facies rocks were collected in the San Piero quarry for MB11-1 (Fig. 1 of the main text; WGS84 latitude and longitude of N $42^{\circ} 74^{\prime} 67.2 " / E 10^{\circ} 20^{\prime} 88.5 "$ "), and on the road bellow the "Torre Giovanni" (close to the contact with the San Francesco facies) for MB11-2 (Fig. 1 of the main text; WGS84 latitude and longitude of $\mathrm{N} 42^{\circ} 76^{\prime} 23.4^{\prime \prime} / \mathrm{E} 10^{\circ} 20^{\prime} 07.6^{\prime \prime}$ ). Both samples are monzogranites and did not contain any MME or K-feldspar megacrysts (Supplement Fig.1F). Zircons from both samples present similar characteristics that the one from the SA and SF samples.

## S2. Methodology

### 2.1 U-Pb methodology

All samples were processed and analyzed at Princeton University. Zircon separates were prepared by standard density and magnetic mineral separation methods (crushing and milling; sieving to $<500 \mu \mathrm{~m}$; concentration via hand-panning; magnetic separation; hand-picking). All the zircons were dated by removing grains that were imaged by CL from the epoxy mount, in order to document any correlations between internal textures and dates and also to target the simplest zircons (representative images are shown in fig. 2 of the main text). However, there was no obvious correlation between the dates of the grains and the amount of core material observed in the zircon, the type of zoning (sector vs. oscillatory), the number of growth episodes, nor the brightness of the zoning. When xenocrystic cores were identified, we cut the zircon following the CL imaging in order to insolate and date only the tips (representative images are shown in fig. 2 of the main text).
Analyses were performed following the same procedure as described in Barboni and Schoene, (2014) and is repeated here. Annealing was performed by loading the zircons of each sample in quartz crucibles, which were heated at $900^{\circ} \mathrm{C}$ for ca. 48 h . Zircons were removed from Epoxy grainmount following CL imaging, loaded into $200 \mu 1$ savillex capsules, leached in $\mathrm{HF}+$ trace $\mathrm{HNO}_{3}$ for ca. 12 hours at $190^{\circ} \mathrm{C}$ and rinsed with water, 6 N HCl


Supplementary Fig.1. A) Portoferraio Porphyry (sample MB11-11). B) San Martino Porphyry (sample MB1114). C) Orano dyke crosscutting the Sant'Andrea facies of the Mt. Capanne intrusions (sample MB11-5). D) Mafic enclaves and K-feldspar megacrysts in the Sant'Andrea facies of the Capanne intrusions. F) San Francesco facies of the Capanne intrusion. D) San Piero facies of the Capanne intrusion.
and HF. Each grain was spiked with ca. 0.006 g of the EARTHTIME ${ }^{20} 5 \mathrm{~Pb}-{ }^{233} \mathrm{U}-{ }^{235} \mathrm{U}$ tracer solution (Condon et al., in press; Mclean et al., in press). Zircons were subsequently dissolved in ca. $70 \mu 140 \% \mathrm{HF}$ and trace $\mathrm{HNO}_{3}$ at $210^{\circ} \mathrm{C}$ for $48+$ hours, dried down and redissolved in 6 N HCl overnight. Samples were then dried down and redissolved in 3 N HCl and put through a modified single $50 \mu 1$ column HCl-based anion exchange chemistry (Krogh, 1973). U and Pb were collected in single beakers, dried down with a drop of 0.02 M H 3 PO 4 , and analyzed on a single outgassed Re filament in Si-gel emitter (modified from Gerstenberg and Haase, 1997).
Measurements were performed on an IsotopX Phoenix62 thermal ionization mass spectrometer at Princeton University. Pb was measured in dynamic mode on an axial ion-counting Daly photomultiplier. Deadtime for the Daly was determined at 40.5 ns by repeated measurements of NBS-981 and NBS-982 for up to 2.5 Mcps. Lead mass fractionation was calibrated by repeated NBS-981 measurements (mean $\alpha{ }^{208} \mathrm{~Pb}-{ }^{206} \mathrm{~Pb}=0.18 \pm 0.04 \% / \mathrm{amu}$, 2-sigma standard deviation) on mixed $\mathrm{Pb}-\mathrm{U}$ aliquots of $<100 \mathrm{pg} \mathrm{Pb}$ to closely imitate sample running behavior. Baseline measurements were made at each half-mass and the average intensity bounding each measured peak was subtracted. Isobaric interferences on ${ }^{205} \mathrm{~Pb}$ were monitored by measuring mass 203 , but repeated analyses of unspiked zircon show that the intensity of non- 205 Pb ions under mass 205 is trivial for this study. As a result, no corrections were applied, and the decay of mass 203 over the duration of the analysis relative to Pb is used as an indicator of declining isobaric interferences under all Pb masses. Data culling was done using decreasing 203/205 and increasing 206/204 ratios over the course of an analysis. U was measured in static mode on Faraday cups on $10^{12}$ ohm resistors as $\mathrm{UO}_{2}+.{ }^{233} \mathrm{UO}_{2}$ and ${ }^{235} \mathrm{UO}_{2}$ were corrected for an oxygen isotopic composition of 0.002055 (see discussion in Condon et al., in press). Because ${ }^{18} \mathrm{O} /{ }^{16} \mathrm{O}$ typically grows at the beginning of an analysis before stabilizing, early blocks of data were deleted. Baselines were measured at $\pm 0.5$ mass units for 15 seconds every 10 ratios. Correction for mass-fractionation of U was done using the EARTHTIME ${ }^{205} \mathrm{~Pb}-{ }^{233} \mathrm{U}-{ }^{235} \mathrm{U}$ tracer solution assuming a sample ${ }^{238} \mathrm{U} /{ }^{235} \mathrm{U}$ ratio of $137.818 \pm 0.021$ (Hiess et al., 2012). All data reduction, error propagation and plotting of $\mathrm{U}-\mathrm{Pb}$ data was done using the $\mathrm{U}-\mathrm{Pb} \_$Redux software package (Bowring et al., 2011; McLean et al., 2011). All reported uncertainties are 2 -sigma and include internal sources of uncertainty only. Including systematic sources of uncertainty such as tracer composition and decay constants should be carried out for comparison with U-Pb data collected using a different tracer or with other isotopic systems, and can be done by accessing data from this study on the Geochron data storage website, given the uncertainties for tracer composition reported in Condon et al. (in press) and uncertainties for decay constants discussed in Schoene et al. (2006).
21 procedural blanks were measured over the course of this study, spiked with the same tracer. The amount of Pbc in the total procedural blanks ( $0.5-3.6 \mathrm{pg}$; avg. 1.1 pg ) agreed well with that found in zircon analyses, and therefore all common Pb is assumed to derive from procedural blanks. After 2-sigma outlier rejection, the composition of $19{ }^{205} \mathrm{~Pb}-{ }^{233} \mathrm{U}-{ }^{235} \mathrm{U}$-spiked blanks was: ${ }^{206} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}=18.50 \pm 0.10,{ }^{207} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}=15.56 \pm 0.21,{ }^{208} \mathrm{~Pb} /{ }^{204} \mathrm{~Pb}=$ $37.48 \pm 0.34$ (2-sigma standard deviation), and these uncertainties were propagated into each U-Pb analysis.
An important consideration in U-Pb geochronology of ca. 7 Ma zircons is the correction for initial secular disequilibrium in the U-Pb decay chain. During zircon crystallization, intermediate daughters products can be incorporated or excluded from the crystal depending on the zircon/magma distribution coefficient for each element. Our primary concern is the exclusion of ${ }^{230} \mathrm{Th}\left(\mathrm{t}_{1 / 2}=75,380\right.$ years $)$, a long-lived intermediate daughter product of ${ }^{238} \mathrm{U}$, as initial depletion leads to a deficiency of ${ }^{206} \mathrm{~Pb}$ and therefore apparent ${ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}$ dates that are too young. This effect is generally corrected by using a model $\mathrm{Th} / \mathrm{U}_{\text {zircon }}$ calculated from the blank-subtracted ${ }^{208} \mathrm{~Pb} /{ }^{206} \mathrm{~Pb}$ zircon measured by ID-TIMS and an estimate of the $\mathrm{Th} / \mathrm{U}_{\text {magma }}$ at the time of the zircon crystallization.

We calculated the $\mathrm{Th} / \mathrm{U}_{\text {magma }}$ for each dated zircon by using the model $\mathrm{Th} / \mathrm{U}_{\text {zircon }}$ and $\mathrm{Th} / \mathrm{U}$ zircon/melt distribution coefficients (D) experimentally determined by Rubatto and Hermann (2007) for hydrous granitic melt at $800^{\circ} \mathrm{C}$ ( $\mathrm{DTh}=41 \pm 4 ; ~ \mathrm{DU}=167 \pm 17$ ). Uncertainties on the distribution coefficients were not propagated into our age uncertainties because this uncertainty is regarded as systematic for each grain, assuming a restricted temperature and compositional range of the magma. Additionally, the effect of changing intensive variables on the ratio of partition coefficients ( $\mathrm{DTh} / \mathrm{U}$ ) is far less than the absolute values of each, further supporting the systematic nature of this uncertainty. As such, the differences between dated grains are insensitive to the disequilbrium correction. Our calculated values for $\mathrm{Th} / \mathrm{U}_{\text {magma }}$ are reported in Supplementary Table 1, and while these are not meant to be robust estimates of magma composition, they do illustrate that they yield $\mathrm{Th} / \mathrm{U}_{\text {magma }}$ ratios that are reasonable for the Elba samples.

### 2.2 Ryholite-MELTS model

The crystallization sequence of the Sant'Andrea, San Francesco and San Piero magmas was simulated using the modeling package Rhyolite-MELTS optimized for silica-rich, fluid-bearing magmatic systems (Gualda et al., 2012). We assumed closed system crystallization at isobaric conditions, regulated by the QFM oxygen fugacity buffer. Starting liquid equivalent to the whole-rock compositions of the Sant'Andrea facies sample PP-334 (Dini et al., 2002), San Francesco facies sample PP-364 (Dini et al., 2002) and San Piero facies sample MB11-2 (this study) were used. All these samples are close to the average composition of all samples reported for each facies in the literature (Dini et al., 2002; Farina et al., 2010). Other compositions do not significantly change the results. We performed multiple runs for pressures ranging from 2 to 5 kbar and water contents between 1 and $6 \mathrm{wt} \%$, with temperatures decreasing from 1200 to $500^{\circ} \mathrm{C}$ for each of the Capanne facies. Each run with different conditions was discarded if not closely matching observed petrological observations or modal proportions and microprobe mineral measurements (Bussy, 1991) made on Sant'Andrea granite thin sections. We noticed that MELTS could not produce results bellow a pressure of 2.3 kbar . As our study tries to model the cooling evolution of the Capanne granite at emplacement level (ca. 2 Kbar; Bussy, 1991), we narrowed our selection to runs computed for 2.3 kbar. Our best-fit models for the three samples were performed with initial water content of $2 \mathrm{wt} \%$ (Supplement Fig.2A-2C). Though this water content best fit the observed mineral assemblages, we note that the absolute temperatures calculated (Supplementary Table 2) are very sensitive to the assumed water content. Uncertainties in these temperatures are therefore on the order of $\pm 20^{\circ} \mathrm{C}$ for water contents of $1.5-2.5 \%$, with the added constraint that zircon saturation ( $807 \pm 11^{\circ} \mathrm{C}$, see Supplementary Methods section 2.3) must have occurred at $<40 \%$ crystal content as indicated by pre-emplacement zircon crystallization and the megacryst inclusion history reported by Barboni and Schoene, (2014). Orthopyroxene was intentionally excluded as they were not observed, and MELTS produced them at the expense of biotite. The MELTS raw data are presented in Supplementary Table 2.

### 2.3 Zr-saturation temperature

Zircon saturation temperature was estimated using the same technique as described in Barboni and Schoene, (2014). Published saturation experiments (Watson and Harrison, 1983; Boehnke et al., 2013) were integrated with our MELTS model. We assumed an initial melt Zr concentration equal to bulk rock and calculated liquid Zr con-
tent during crystallization using bulk partition coefficients (from the GERM database: http://earthref.org/GERM/) for the saturated phases predicted by MELTS. The major element composition of the coexisting liquid was used to determine the M parameter used in the zircon saturation calculation (Supplementary Table 3 and Supplementary Fig. 2). We then calculated the Zr concentration required for saturation in the evolving liquid using Watson and Harrison (1983) and Boehnke et al. (2013), yielding temperatures of ca. $805^{\circ} \mathrm{C}$ for all three Capanne facies, $\sim 10-$ $50{ }^{\circ} \mathrm{C}$ hotter than that predicted using only bulk rock chemistry alone (Watson and Harrison, 1983; Supplement Fig.2). Uncertainties of $\pm 11^{\circ} \mathrm{C}$ in this temperature were calculated by calculating the maximum and minimum temperature permitted by the calibration uncertainties reported in Boehnke et al. (2013).

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## MB11-1 San Piero



## MB1 1-2 San Piero



## MB11-5 Orano Dyke



Supplementary Fig.2. CL images of the zircon measured in this study with $u$ - Pb ages (2-sigma uncertainties)

MB11-6 Sant Andrea


MB11-11 Portoferraio Porphyry


MB11-14 San Martino Porphyry


Supplementary Fig.2. CL images of the zircon measured in this study with $u$ - Pb ages ( 2 -sigma uncertainties)

## MB12-4 San Francesco



## MB12-8 San Francesco



## MB12-9 Sant'Andrea



Supplementary Fig.2. CL images of the zircon measured in this study with u-Pb ages (2-sigma uncertainties)
Supplementary Table 2: U-Pb isotopic data

|  | Compositional parameters |  |  |  |  | Radiogenic Isotope Ratios |  |  |  |  |  |  |  | Dates (Ma) |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sample (a) | $\begin{gathered} \frac{\mathrm{Th}}{\mathrm{U}} \\ \text { zircon } \end{gathered}$ <br> (b) | $\frac{\mathrm{Th}}{\mathrm{U}}$ melt (c) | $\mathrm{Pb}^{*}$ <br> Pbc <br> (d) | Pbc <br> (pg) <br> (d) | ${ }^{2006} \mathrm{~Pb}$ <br> (e) | ${ }^{2008} \mathrm{~Pb}$ (f) | ${ }^{2006} \mathrm{~Pb}$ (f) | \% err <br> (g) | ${ }^{207 \mathrm{~Pb}}$ <br> (f) | \% err <br> (g) | ${ }^{2068 \mathrm{~Pb}}$ <br> (f) | \% err <br> (g) | corr. coef. | ${ }^{2006} \mathrm{~Pb}$ <br> (h) | (g) | ${ }^{2007} \mathrm{~Pb}$ <br> (h) | (g) | ${ }^{206} \mathrm{~Pb}$ | (g) | ${ }_{206}^{208 \mathrm{~Pb}}$ <br> (h) | (g) |
| MB11-6 | Sant'Andrea granite (Capanne pluton) - ( $\mathrm{N} 42^{\circ} 80^{\prime} 80.1^{\prime \prime} / \mathrm{E} 10^{\circ} 14^{\prime} 09.1^{\prime \prime}$ ) $n=20$; as published in Barboni and Schoene (2014) |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 211 | 0.21 | 0.84 | 1.5 | 5.90 | 115 | 0.067 | 0.048482 | 4.560 | 0.007598 | 4.735 | 0.001137 | 0.261 | 0.739 | 122.9 | 107.41 | 7.685 | 0.362 | 7.240 | 0.020 | 7.323 | 0.019 |
| z9 | 0.17 | 0.69 | 1.8 | 1.43 | 136 | 0.056 | 0.045340 | 4.109 | 0.007142 | 4.254 | 0.001143 | 0.190 | 0.840 | -37.4 | 99.74 | 7.227 | 0.306 | 7.279 | 0.015 | 7.361 | 0.014 |
| z2 | 0.16 | 0.65 | 1.8 | 2.04 | 139 | 0.053 | 0.046022 | 3.863 | 0.007265 | 3.994 | 0.001145 | 0.170 | 0.856 | -1.2 | 93.13 | 7.350 | 0.292 | 7.295 | 0.013 | 7.376 | 0.013 |
| z7 | 0.13 | 0.53 | 6.4 | 0.99 | 449 | 0.043 | 0.046038 | 1.099 | 0.007270 | 1.143 | 0.001145 | 0.072 | 0.708 | -0.4 | 26.50 | 7.355 | 0.084 | 7.297 | 0.006 | 7.378 | 0.005 |
| z15 | 0.16 | 0.65 | 1.4 | 3.49 | 113 | 0.051 | 0.047116 | 4.793 | 0.007453 | 4.961 | 0.001147 | 0.216 | 0.860 | 55.1 | 114.34 | 7.540 | 0.373 | 7.309 | 0.017 | 7.392 | 0.016 |
| z13 | 0.14 | 0.57 | 1.9 | 3.49 | 145 | 0.046 | 0.046807 | 3.584 | 0.007414 | 3.706 | 0.001149 | 0.163 | 0.850 | 39.4 | 85.74 | 7.500 | 0.277 | 7.320 | 0.013 | 7.401 | 0.012 |
| z7_2 | 0.18 | 0.73 | 7.9 | 1.16 | 538 | 0.058 | 0.046618 | 0.901 | 0.007385 | 0.938 | 0.001149 | 0.080 | 0.543 | 29.7 | 21.60 | 7.471 | 0.070 | 7.320 | 0.006 | 7.402 | 0.006 |
| z23 | 0.28 | 1.14 | 2.7 | 2.97 | 189 | 0.090 | 0.047055 | 2.642 | 0.007489 | 2.754 | 0.001154 | 0.141 | 0.835 | 52.0 | 63.07 | 7.576 | 0.208 | 7.354 | 0.011 | 7.437 | 0.010 |
| z3 | 0.16 | 0.65 | 3.2 | 3.64 | 233 | 0.052 | 0.046744 | 2.124 | 0.007441 | 2.198 | 0.001155 | 0.121 | 0.703 | 36.2 | 50.85 | 7.527 | 0.165 | 7.356 | 0.009 | 7.438 | 0.009 |
| z11_2 | 0.16 | 0.65 | 2.4 | 3.65 | 175 | 0.052 | 0.046846 | 2.888 | 0.007482 | 2.991 | 0.001158 | 0.139 | 0.825 | 41.4 | 69.08 | 7.569 | 0.226 | 7.381 | 0.011 | 7.463 | 0.010 |
| z22 | 0.25 | 1.02 | 5.7 | 2.09 | 388 | 0.081 | 0.046349 | 1.178 | 0.007406 | 1.244 | 0.001159 | 0.100 | 0.710 | 15.8 | 28.31 | 7.492 | 0.093 | 7.384 | 0.008 | 7.466 | 0.007 |
| z6 | 0.12 | 0.49 | 3.8 | 1.21 | 275 | 0.039 | 0.046040 | 1.854 | 0.007363 | 1.914 | 0.001160 | 0.092 | 0.764 | -0.3 | 44.70 | 7.449 | 0.142 | 7.391 | 0.007 | 7.473 | 0.007 |
| z18 | 0.16 | 0.67 | 7.9 | 2.41 | 542 | 0.053 | 0.046263 | 0.884 | 0.007410 | 0.922 | 0.001162 | 0.066 | 0.647 | 11.3 | 21.27 | 7.496 | 0.069 | 7.402 | 0.005 | 7.484 | 0.005 |
| $z 19$ | 0.07 | 0.29 | 5.6 | 2.62 | 404 | 0.023 | 0.043999 | 1.465 | 0.007058 | 1.471 | 0.001163 | 0.137 | 0.202 | -110.7 | 36.08 | 7.141 | 0.105 | 7.413 | 0.011 | 7.495 | 0.010 |
| z4 | 0.14 | 0.57 | 10.7 | 1.75 | 730 | 0.045 | 0.046928 | 0.661 | 0.007539 | 0.685 | 0.001165 | 0.058 | 0.517 | 45.6 | 15.80 | 7.626 | 0.052 | 7.424 | 0.004 | 7.507 | 0.004 |
| z21 | 0.24 | 0.98 | 8.8 | 1.75 | 587 | 0.077 | 0.046743 | 0.802 | 0.007524 | 0.836 | 0.001167 | 0.066 | 0.590 | 36.1 | 19.20 | 7.611 | 0.063 | 7.439 | 0.005 | 7.521 | 0.005 |
| z16 | 0.23 | 0.94 | 1.3 | 3.72 | 101 | 0.075 | 0.046618 | 5.557 | 0.007533 | 5.763 | 0.001172 | 0.254 | 0.865 | 29.7 | 133.20 | 7.620 | 0.438 | 7.468 | 0.020 | 7.551 | 0.019 |
| z16_2 | 0.56 | 2.28 | 7.2 | 2.47 | 446 | 0.181 | 0.046765 | 1.063 | 0.007574 | 1.109 | 0.001175 | 0.074 | 0.666 | 37.3 | 25.44 | 7.661 | 0.085 | 7.485 | 0.006 | 7.568 | 0.006 |
| z20 | 0.17 | 0.70 | 17.9 | 1.19 | 1202 | 0.056 | 0.046207 | 0.407 | 0.007563 | 0.439 | 0.001187 | 0.063 | 0.580 | 8.4 | 9.79 | 7.650 | 0.033 | 7.566 | 0.005 | 7.648 | 0.005 |
| $z 17$ | 0.24 | 0.96 | 4.6 | 2.31 | 319 | 0.076 | 0.046835 | 1.526 | 0.007892 | 1.588 | 0.001222 | 0.102 | 0.668 | 40.8 | 36.50 | 7.982 | 0.126 | 7.791 | 0.008 | 7.873 | 0.008 |


Supplementary Table 2: U-Pb isotopic data (cont.)

|  | Compositional parameters |  |  |  |  | Radiogenic Isotope Ratios |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sample | Th | Th | $\underline{\mathrm{Pb}}{ }^{*}$ | Pbc | ${ }^{206} \mathrm{~Pb}$ | ${ }^{208} \mathrm{~Pb}$ | ${ }^{207} \mathrm{~Pb}$ |  |
|  | U | U | Pbc | (pg) | ${ }^{204} \mathrm{~Pb}$ | ${ }^{206} \mathrm{~Pb}$ | ${ }^{206} \mathrm{~Pb}$ | \% err |
|  | zircon | melt |  |  |  |  |  |  |
| (a) | (b) | (c) | (d) | (d) | (e) | (f) | (f) | (g) |





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 San Piero granite (Capanne pluton) - (N $\left.42^{\circ} 76^{\prime} 23.4^{\prime \prime} / E 10^{\circ} 20^{\prime} 07.6^{\prime \prime}\right) n=22$
Supplementary Table 2：U－Pb isotopic data（cont．）

| Sample | Compositional parameters |  |  |  |  | Radiogenic Isotope Ratios |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Th | Th | $\mathrm{Pb}^{*}$ | Pbc | ${ }^{206} \mathrm{~Pb}$ | ${ }^{209} \mathrm{~Pb}$ | ${ }^{207} \mathrm{~Pb}$ |  |
|  | U | U | Pbc | （pg） | ${ }^{204} \mathrm{~Pb}$ | ${ }^{206} \mathrm{~Pb}$ | ${ }^{206} \mathrm{~Pb}$ | \％err |
|  | zircon | melt |  |  |  |  |  |  |
| （a） | （b） | （c） | （d） | （d） | （e） | （f） | （f） | （g） |



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Supplementary Table 2: U-Pb isotopic data (cont.)
Radiogenic Isotope Ratios
Dates (Ma)

| Sample (a) | $\begin{aligned} & \frac{\mathrm{Th}}{U} \\ & \text { zircon } \\ & \text { (b) } \end{aligned}$ | $\begin{gathered} \frac{\mathrm{Th}}{\mathrm{U}} \\ \text { melt } \\ \text { (c) } \end{gathered}$ | $\frac{\mathrm{Pb}^{*}}{\mathrm{Pbc}}$ <br> (d) | Pbc <br> (pg) <br> (d) | $\begin{aligned} & 200 \mathrm{~Pb} \\ & 2004 \mathrm{pb} \end{aligned}$ | ${ }_{200}^{209} \mathrm{~Pb}$ <br> (f) | $\begin{aligned} & { }^{27} \mathrm{~Pb} \\ & { }_{26 \mathrm{~Pb}} \\ & (\mathrm{f}) \end{aligned}$ | \% err <br> (g) | ${ }^{2027}{ }^{235} \mathrm{Ub}$ <br> (f) | \% err (g) | $\begin{aligned} & { }^{206} \mathrm{~Pb} \\ & { }_{238} \mathrm{U} \\ & \text { (f) } \end{aligned}$ | \% err <br> (g) | corr. coef. | ${ }^{207} \mathrm{~Pb}$ <br> ${ }^{206} \mathrm{~Pb}$ <br> (h) | (g) | ${ }_{2023}^{203 \mathrm{~Pb}}$ <br> (h) | $\pm$ (g) | $\begin{aligned} & { }^{206} \mathrm{~Pb} \\ & { }^{238} \mathrm{U} \\ & \text { (i) } \end{aligned}$ | $\pm$ (g) | ${ }^{2208} \mathrm{~Pb}$ <br> (h) | (g) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| MB11-11 | Portoferraio Porphyry ( $\mathrm{N} 42^{\circ} 49^{\prime} 21.6^{\prime \prime} / \mathrm{E} 10^{\circ} 17^{\prime} 13.6^{\prime \prime}$ ) (cont.) |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 221 | 0.17 | 0.69 | 5.3 | 1.23 | 370 | 0.055 | 0.0461816 | 1.386 | 0.007907 | 1.444 | 0.001243 | 0.083 | 0.705 | 30.8 | 33.04 | 8.002 | 0.115 | 7.926 | 0.007 | 8.008 | 0.007 |
| 21 | 0.19 | 0.78 | 3.4 | 3.34 | 243 | 0.062 | 0.0464509 | 2.024 | 0.007957 | 2.104 | 0.001243 | 0.146 | 0.596 | 45.8 | 48.12 | 8.052 | 0.168 | 7.927 | 0.012 | 8.009 | 0.012 |
| z3 | 0.37 | 1.51 | 8.4 | 2.77 | 543 | 0.119 | 0.0494844 | 0.814 | 0.017911 | 0.851 | 0.002627 | 0.054 | 0.657 | 181.2 | 18.96 | 18.030 | 0.152 | 16.830 | 0.009 | 16.913 | 0.009 |


Supplementary Table 2: U-Pb isotopic data (cont.)
Radiogenic Isotope Ratios

| Sample (a) | $\underset{\text { zircon }}{\frac{\mathrm{Th}}{\mathrm{U}}}$ <br> (b) | $\frac{T h}{U}$ melt <br> (c) | $\frac{\mathrm{Pb}^{*}}{\mathrm{Pbc}}$ <br> (d) | Pbc (pg) <br> (d) | ${ }^{206} \mathrm{~Pb}$ <br> ${ }^{204} \mathrm{~Pb}$ <br> (e) | ${ }^{208} \mathrm{~Pb}$ <br> ${ }^{206} \mathrm{~Pb}$ <br> (f) | ${ }^{207} \mathrm{~Pb}$ <br> (f) | \% err <br> (g) | ${ }^{207 \mathrm{~Pb}}$ <br> (f) | \% err <br> (g) | ${ }^{2068 \mathrm{~Pb}}$ <br> (f) | \% err <br> (g) | coef. | ${ }^{207} \mathrm{~Pb}$ <br> (h) | (g) | ${ }^{207 \mathrm{~Pb}}$ <br> (h) | (g) | ${ }^{206} \mathrm{~Pb}$ <br> ${ }^{238} \mathrm{U}$ <br> (i) | (g) | ${ }^{206 \mathrm{~Pb}}{ }^{238} \mathrm{U}$ <br> (h) | (g) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| MB12-8 | San Francesco granite (Cappane pluton) - ( $\mathrm{N} 42^{\circ} 76^{\prime} 93.2^{\prime \prime} / \mathrm{E} 10^{\circ} 18^{\prime} 77.9^{\prime \prime}$ ) $n=16$ |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| z2 | 0.17 | 0.67 | 9.4 | 1.65 | 640 | 0.054 | 0.045977 | 0.798 | 0.007052 | 0.844 | 0.001113 | 0.064 | 0.633 | 24.1 | 19.0 | 7.140 | 0.059 | 7.089 | 0.005 | 7.172 | 0.005 |
| z1 | 0.19 | 0.77 | 6.7 | 2.69 | 459 | 0.062 | 0.046177 | 1.064 | 0.007090 | 1.111 | 0.001114 | 0.064 | 0.682 | 34.5 | 25.3 | 7.179 | 0.079 | 7.097 | 0.005 | 7.180 | 0.005 |
| z21 | 0.26 | 1.06 | 2.4 | 0.63 | 176 | 0.083 | 0.046673 | 3.077 | 0.007204 | 3.198 | 0.001121 | 0.169 | 0.739 | 59.0 | 73.0 | 7.293 | 0.232 | 7.138 | 0.013 | 7.220 | 0.012 |
| z20 | 0.22 | 0.90 | 5.0 | 0.65 | 347 | 0.072 | 0.045618 | 1.589 | 0.007045 | 1.673 | 0.001121 | 0.106 | 0.774 | 4.0 | 38.1 | 7.133 | 0.118 | 7.142 | 0.008 | 7.224 | 0.008 |
| z13 | 0.26 | 1.06 | 5.6 | 1.33 | 377 | 0.085 | 0.046175 | 1.526 | 0.007149 | 1.598 | 0.001124 | 0.112 | 0.642 | 34.2 | 36.4 | 7.238 | 0.115 | 7.157 | 0.008 | 7.239 | 0.008 |
| $z 10$ | 0.11 | 0.45 | 11.7 | 0.93 | 807 | 0.036 | 0.046767 | 0.656 | 0.007258 | 0.693 | 0.001126 | 0.072 | 0.449 | 64.7 | 15.5 | 7.348 | 0.050 | 7.174 | 0.005 | 7.257 | 0.005 |
| z16 | 0.14 | 0.57 | 6.9 | 0.86 | 476 | 0.045 | 0.046607 | 1.029 | 0.007237 | 1.071 | 0.001127 | 0.059 | 0.683 | 56.5 | 24.4 | 7.327 | 0.078 | 7.178 | 0.004 | 7.260 | 0.004 |
| z22 | 0.13 | 0.53 | 25.1 | 0.63 | 1699 | 0.044 | 0.046632 | 0.302 | 0.007251 | 0.341 | 0.001129 | 0.046 | 0.395 | 56.2 | 7.2 | 7.341 | 0.023 | 7.193 | 0.003 | 7.274 | 0.003 |
| z14 | 0.13 | 0.53 | 15.4 | 0.40 | 1052 | 0.041 | 0.046609 | 0.512 | 0.007254 | 0.539 | 0.001129 | 0.047 | 0.358 | 56.7 | 12.1 | 7.344 | 0.038 | 7.194 | 0.003 | 7.277 | 0.003 |
| 211 | 0.14 | 0.57 | 7.9 | 0.82 | 546 | 0.046 | 0.046558 | 0.954 | 0.007269 | 1.001 | 0.001133 | 0.079 | 0.584 | 53.7 | 22.6 | 7.359 | 0.073 | 7.218 | 0.006 | 7.300 | 0.006 |
| z4 | 0.17 | 0.68 | 4.9 | 1.28 | 343.2 | 0.0541 | 0.0465217 | 1.444 | 0.007270 | 1.501 | 0.001134 | 0.086 | 0.678 | 51.9 | 34.3 | 7.360 | 0.110 | 7.224 | 0.007 | 7.307 | 0.006 |
| $z 15$ | 0.22 | 0.90 | 20.4 | 0.53 | 1348 | 0.0711 | 0.0466782 | 0.389 | 0.007297 | 0.428 | 0.001134 | 0.060 | 0.420 | 59.9 | 9.2 | 7.387 | 0.030 | 7.226 | 0.004 | 7.309 | 0.004 |
| z5 | 0.22 | 0.90 | 9.3 | 0.99 | 621.8 | 0.0717 | 0.0464455 | 0.772 | 0.007298 | 0.811 | 0.001140 | 0.053 | 0.613 | 47.8 | 18.4 | 7.388 | 0.059 | 7.264 | 0.004 | 7.347 | 0.004 |
| 217 | 0.17 | 0.65 | 7.0 | 0.63 | 482.1 | 0.0535 | 0.0462571 | 1.095 | 0.007284 | 1.142 | 0.001143 | 0.065 | 0.685 | 37.7 | 26.0 | 7.374 | 0.083 | 7.281 | 0.005 | 7.363 | 0.005 |
| $z 18$ | 0.10 | 0.41 | 19.6 | 0.73 | 1345 | 0.0327 | 0.0465314 | 0.375 | 0.007368 | 0.407 | 0.001149 | 0.038 | 0.429 | 52.0 | 8.9 | 7.459 | 0.029 | 7.322 | 0.003 | 7.404 | 0.003 |
| $z 19$ | 0.22 | 0.90 | 0.9 | 1.43 | 76 | 0.0719 | 0.0461928 | 8.111 | 0.007619 | 8.413 | 0.001197 | 0.358 | 0.879 | 32.3 | 193.4 | 7.711 | 0.646 | 7.633 | 0.029 | 7.715 | 0.028 |
| MB12-9 | San Francesco granite (Cappane pluton) - (N 42 $\left.{ }^{\circ} 76^{\prime} 93.2^{\prime \prime} / \mathrm{E} 10^{\circ} 18^{\prime} 77.9^{\prime \prime}\right) n=12$ |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| z12 | 0.21 | 0.82 | 18.7 | 0.61 | 1240 | 0.0684 | 0.046551 | 0.419 | 0.007204 | 0.465 | 0.001123 | 0.070 | 0.478 | 53.2 | 10.0 | 7.294 | 0.032 | 7.155 | 0.005 | 7.236 | 0.005 |
| z13 | 0.23 | 0.94 | 2.5 | 2.72 | 184 | 0.0751 | 0.046815 | 2.735 | 0.007257 | 2.843 | 0.001125 | 0.133 | 0.835 | 67.0 | 64.8 | 7.346 | 0.208 | 7.166 | 0.010 | 7.248 | 0.010 |
| z16 | 0.30 | 1.22 | 3.0 | 0.49 | 208.9 | 0.0983 | 0.045715 | 2.663 | 0.007114 | 2.775 | 0.001130 | 0.140 | 0.801 | 8.8 | 63.9 | 7.202 | 0.199 | 7.197 | 0.011 | 7.279 | 0.010 |
| z18 | 0.21 | 0.86 | 4.8 | 1.19 | 334.2 | 0.0673 | 0.046972 | 1.464 | 0.007313 | 1.528 | 0.001130 | 0.086 | 0.734 | 74.0 | 34.6 | 7.403 | 0.112 | 7.200 | 0.006 | 7.282 | 0.006 |
| z5 | 0.16 | 0.67 | 18.0 | 1.37 | 1209 | 0.053 | 0.046247 | 0.407 | 0.007203 | 0.466 | 0.001130 | 0.099 | 0.520 | 37.9 | 9.7 | 7.293 | 0.033 | 7.200 | 0.007 | 7.283 | 0.007 |
| z6 | 0.11 | 0.44 | 3.0 | 3.72 | 218.3 | 0.0347 | 0.045296 | 2.699 | 0.007060 | 2.798 | 0.001131 | 0.151 | 0.739 | -12.1 | 64.6 | 7.148 | 0.199 | 7.205 | 0.012 | 7.288 | 0.011 |
| z3 | 0.09 | 0.38 | 19.3 | 1.65 | 1324 | 0.0305 | 0.046560 | 0.386 | 0.007298 | 0.442 | 0.001137 | 0.072 | 0.609 | 53.7 | 9.1 | 7.388 | 0.031 | 7.247 | 0.005 | 7.329 | 0.005 |
| z4 | 0.17 | 0.70 | 11.2 | 1.54 | 756.3 | 0.0554 | 0.046360 | 0.689 | 0.007293 | 0.738 | 0.001142 | 0.073 | 0.585 | 43.4 | 16.4 | 7.383 | 0.053 | 7.273 | 0.005 | 7.355 | 0.005 |
| z19 | 0.16 | 0.65 | 10.4 | 1.07 | 706.3 | 0.0526 | 0.046598 | 0.707 | 0.007332 | 0.741 | 0.001142 | 0.058 | 0.484 | 54.3 | 16.8 | 7.422 | 0.054 | 7.278 | 0.004 | 7.360 | 0.004 |
| z1 | 0.14 | 0.55 | 8.4 | 2.55 | 579.7 | 0.0441 | 0.046106 | 0.859 | 0.007292 | 0.908 | 0.001148 | 0.098 | 0.503 | 30.0 | 20.4 | 7.382 | 0.066 | 7.313 | 0.007 | 7.395 | 0.007 |
| z14 | 0.30 | 1.22 | 9.3 | 0.63 | 607.4 | 0.0976 | 0.046740 | 0.855 | 0.007397 | 0.901 | 0.001149 | 0.079 | 0.525 | 61.6 | 20.3 | 7.488 | 0.067 | 7.321 | 0.006 | 7.403 | 0.006 |
| $z 15$ | 0.13 | 0.53 | 19.1 | 0.69 | 1298 | 0.0422 | 0.046593 | 0.387 | 0.007376 | 0.433 | 0.001149 | 0.061 | 0.518 | 54.1 | 9.2 | 7.467 | 0.031 | 7.323 | 0.005 | 7.406 | 0.005 |

(a) z1, z2 etc. are labels for single zircon grains or fragment of grain; all zircons annealed and chemically abraded after Mattinson (2005).
(b) Model Th/U ratio calculated from radiogenic $208 \mathrm{~Pb} / 206 \mathrm{~Pb}$ ratio and $207 \mathrm{~Pb} / 235 \mathrm{U}$ age

San Francesco ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content $=2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Temperature ( $\left.{ }^{\circ} \mathrm{C}\right)$ | wt\% SiO2 <br> melt | wt\% TiO2 <br> melt | wt $\%$ Al2O3 <br> melt | wt\% FeO <br> melt | wt $\%$ MgO <br> melt | wt\% CaO <br> melt | wt $\%$ Na2O <br> melt | wt\% K2O <br> melt | wt $\mathbf{~ H 2 O ~}$ <br> melt |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1006 | 67.44 | 0.58 | 16.03 | 1.94 | 1.21 | 2.51 | 3.50 | 4.08 | 1.99 |
| 996 | 67.50 | 0.58 | 16.09 | 1.88 | 1.13 | 2.51 | 3.51 | 4.10 | 2.00 |
| 986 | 67.56 | 0.58 | 16.14 | 1.82 | 1.05 | 2.51 | 3.53 | 4.11 | 2.01 |
| 976 | 67.69 | 0.59 | 16.12 | 1.76 | 0.97 | 2.48 | 3.52 | 4.16 | 2.03 |
| 966 | 67.99 | 0.61 | 15.96 | 1.70 | 0.89 | 2.36 | 3.45 | 4.26 | 2.10 |
| 956 | 68.28 | 0.62 | 15.80 | 1.65 | 0.82 | 2.25 | 3.37 | 4.37 | 2.16 |
| 946 | 68.56 | 0.64 | 15.64 | 1.58 | 0.75 | 2.15 | 3.29 | 4.48 | 2.22 |
| 936 | 68.82 | 0.66 | 15.49 | 1.52 | 0.69 | 2.06 | 3.20 | 4.59 | 2.29 |
| 926 | 69.09 | 0.68 | 15.35 | 1.45 | 0.64 | 1.98 | 3.10 | 4.69 | 2.35 |
| 916 | 69.34 | 0.70 | 15.21 | 1.38 | 0.59 | 1.91 | 3.00 | 4.80 | 2.42 |
| 906 | 69.58 | 0.71 | 15.07 | 1.31 | 0.54 | 1.84 | 2.89 | 4.90 | 2.48 |
| 896 | 69.89 | 0.68 | 14.93 | 1.22 | 0.50 | 1.78 | 2.79 | 5.00 | 2.55 |
| 886 | 70.20 | 0.64 | 14.80 | 1.13 | 0.47 | 1.72 | 2.67 | 5.10 | 2.62 |
| 876 | 70.49 | 0.61 | 14.67 | 1.04 | 0.44 | 1.67 | 2.56 | 5.20 | 2.69 |
| 866 | 70.77 | 0.57 | 14.54 | 0.96 | 0.41 | 1.62 | 2.45 | 5.30 | 2.76 |
| 856 | 71.05 | 0.54 | 14.42 | 0.88 | 0.38 | 1.58 | 2.33 | 5.39 | 2.82 |
| 846 | 71.31 | 0.51 | 14.30 | 0.80 | 0.36 | 1.54 | 2.22 | 5.47 | 2.89 |
| 836 | 71.56 | 0.48 | 14.18 | 0.74 | 0.34 | 1.50 | 2.11 | 5.55 | 2.95 |
| 826 | 71.80 | 0.45 | 14.06 | 0.67 | 0.32 | 1.47 | 2.00 | 5.63 | 3.02 |
| 816 | 72.03 | 0.43 | 13.95 | 0.61 | 0.30 | 1.44 | 1.89 | 5.70 | 3.08 |
| 806 | 72.05 | 0.40 | 13.93 | 0.56 | 0.29 | 1.41 | 1.79 | 5.82 | 3.18 |
| 796 | 71.84 | 0.37 | 14.02 | 0.51 | 0.27 | 1.39 | 1.71 | 5.97 | 3.33 |
| 786 | 71.60 | 0.34 | 14.12 | 0.49 | 0.27 | 1.40 | 1.69 | 5.91 | 3.58 |
| 776 | 71.30 | 0.30 | 14.22 | 0.49 | 0.25 | 1.41 | 1.67 | 5.85 | 3.85 |
| 766 | 70.93 | 0.26 | 14.35 | 0.53 | 0.21 | 1.43 | 1.66 | 5.77 | 4.14 |
| 756 | 70.53 | 0.22 | 14.48 | 0.56 | 0.18 | 1.45 | 1.65 | 5.68 | 4.46 |
| 746 | 70.02 | 0.19 | 14.70 | 0.57 | 0.16 | 1.49 | 1.63 | 5.59 | 4.78 |
| 736 | 69.54 | 0.17 | 14.89 | 0.55 | 0.14 | 1.51 | 1.63 | 5.49 | 5.16 |
| 726 | 68.08 | 0.16 | 16.28 | 0.50 | 0.13 | 1.68 | 1.29 | 5.44 | 5.18 |
| 716 | 65.94 | 0.15 | 16.85 | 0.42 | 0.12 | 2.58 | 1.12 | 5.25 | 5.28 |
| 706 | 64.78 | 0.16 | 17.93 | 0.33 | 0.11 | 2.73 | 0.94 | 5.13 | 5.34 |
| 69 | 63.63 | 0.15 | 18.86 | 0.27 | 0.10 | 2.94 | 0.81 | 4.99 | 5.41 |
| 686 | 62.44 | 0.15 | 19.69 | 0.22 | 0.09 | 3.21 | 0.71 | 4.84 | 5.49 |
| 676 | 58.24 | 0.13 | 19.39 | 0.23 | 0.09 | 4.67 | 0.70 | 4.18 | 6.33 |

Supplementary Table 2 Rhyolite-MELTS modeling results: melt composition
Sant'Andrea ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content $=2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Temperature ( ${ }^{\circ} \mathrm{C}$ ) | wt\% SiO2 <br> melt | wt\% TiO2 <br> melt | wt $\%$ Al2O3 <br> melt | wt\% FeO <br> melt | wt $\%$ MgO <br> melt | wt\% CaO <br> melt | wt $\%$ Na2O <br> melt | wt\% K2O <br> melt | wt\% H2O <br> melt |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  |  |  |  |
| 968 | 69.08 | 0.39 | 15.89 | 1.68 | 0.85 | 2.20 | 3.09 | 4.25 | 1.99 |
| 958 | 69.25 | 0.39 | 15.85 | 1.62 | 0.79 | 2.15 | 3.06 | 4.31 | 2.02 |
| 948 | 69.55 | 0.40 | 15.69 | 1.56 | 0.72 | 2.04 | 2.98 | 4.41 | 2.08 |
| 938 | 69.85 | 0.41 | 15.55 | 1.50 | 0.66 | 1.94 | 2.90 | 4.51 | 2.13 |
| 928 | 70.13 | 0.42 | 15.40 | 1.43 | 0.61 | 1.85 | 2.80 | 4.61 | 2.19 |
| 918 | 70.41 | 0.44 | 15.26 | 1.37 | 0.56 | 1.77 | 2.71 | 4.70 | 2.24 |
| 908 | 70.67 | 0.45 | 15.13 | 1.30 | 0.51 | 1.69 | 2.61 | 4.80 | 2.30 |
| 898 | 70.93 | 0.46 | 15.00 | 1.22 | 0.47 | 1.63 | 2.51 | 4.89 | 2.36 |
| 888 | 71.19 | 0.47 | 14.88 | 1.15 | 0.44 | 1.56 | 2.40 | 4.98 | 2.41 |
| 878 | 71.43 | 0.48 | 14.75 | 1.08 | 0.40 | 1.51 | 2.29 | 5.07 | 2.47 |
| 868 | 71.67 | 0.49 | 14.64 | 1.01 | 0.37 | 1.45 | 2.19 | 5.16 | 2.53 |
| 858 | 71.90 | 0.50 | 14.52 | 0.94 | 0.35 | 1.40 | 2.08 | 5.24 | 2.58 |
| 848 | 72.13 | 0.51 | 14.41 | 0.87 | 0.32 | 1.36 | 1.97 | 5.32 | 2.64 |
| 838 | 72.39 | 0.49 | 14.30 | 0.79 | 0.30 | 1.32 | 1.86 | 5.40 | 2.69 |
| 828 | 72.23 | 0.45 | 14.40 | 0.73 | 0.29 | 1.27 | 1.77 | 5.58 | 2.81 |
| 818 | 72.05 | 0.42 | 14.51 | 0.67 | 0.27 | 1.24 | 1.68 | 5.77 | 2.93 |
| 808 | 71.86 | 0.40 | 14.62 | 0.62 | 0.26 | 1.20 | 1.59 | 5.94 | 3.05 |
| 798 | 71.65 | 0.37 | 14.75 | 0.57 | 0.25 | 1.18 | 1.52 | 6.05 | 3.21 |
| 788 | 71.40 | 0.34 | 14.89 | 0.54 | 0.24 | 1.17 | 1.49 | 6.01 | 3.44 |
| 778 | 71.14 | 0.31 | 15.05 | 0.51 | 0.23 | 1.17 | 1.46 | 5.95 | 3.69 |
| 768 | 70.82 | 0.27 | 15.22 | 0.51 | 0.22 | 1.17 | 1.43 | 5.88 | 3.96 |
| 758 | 70.43 | 0.23 | 15.43 | 0.54 | 0.19 | 1.16 | 1.41 | 5.80 | 4.24 |
| 748 | 70.01 | 0.20 | 15.66 | 0.56 | 0.16 | 1.16 | 1.39 | 5.71 | 4.55 |
| 738 | 69.48 | 0.18 | 16.05 | 0.53 | 0.14 | 1.16 | 1.34 | 5.62 | 4.84 |
| 728 | 68.86 | 0.17 | 16.56 | 0.50 | 0.13 | 1.16 | 1.27 | 5.53 | 5.12 |
| 718 | 67.44 | 0.16 | 17.61 | 0.41 | 0.12 | 1.49 | 1.07 | 5.41 | 5.16 |
| 708 | 66.37 | 0.17 | 18.69 | 0.32 | 0.11 | 1.62 | 0.90 | 5.30 | 5.19 |
| 698 | 65.37 | 0.16 | 19.66 | 0.26 | 0.10 | 1.76 | 0.78 | 5.18 | 5.24 |
| 688 | 64.04 | 0.16 | 20.33 | 0.21 | 0.09 | 2.19 | 0.68 | 5.02 | 5.31 |

## Supplementary Table 2 Rhyolite-MELTS modeling results: melt composition

San Piero ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content = $2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Temperature ( ${ }^{\circ} \mathrm{C}$ ) | wt\% SiO2 melt | wt\% TiO2 melt | wt\% Al2O3 melt | wt\% FeO melt | wt\% MgO melt | wt\% CaO melt | wt\% Na2O melt | wt\% K2O melt | wt\% H2O melt |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1076 | 68.21 | 0.58 | 14.87 | 2.81 | 1.72 | 1.95 | 2.23 | 4.84 | 2.01 |
| 1066 | 68.29 | 0.58 | 14.94 | 2.73 | 1.60 | 1.96 | 2.24 | 4.86 | 2.02 |
| 1056 | 68.37 | 0.59 | 15.01 | 2.65 | 1.49 | 1.96 | 2.25 | 4.88 | 2.03 |
| 1046 | 68.45 | 0.59 | 15.07 | 2.57 | 1.39 | 1.97 | 2.26 | 4.91 | 2.04 |
| 1036 | 68.53 | 0.59 | 15.13 | 2.49 | 1.29 | 1.97 | 2.27 | 4.93 | 2.05 |
| 1026 | 68.61 | 0.59 | 15.20 | 2.40 | 1.20 | 1.98 | 2.28 | 4.96 | 2.06 |
| 1016 | 68.69 | 0.60 | 15.25 | 2.31 | 1.11 | 1.98 | 2.29 | 4.98 | 2.07 |
| 1006 | 68.77 | 0.60 | 15.31 | 2.22 | 1.03 | 1.98 | 2.30 | 5.00 | 2.08 |
| 996 | 68.85 | 0.60 | 15.37 | 2.12 | 0.96 | 1.98 | 2.31 | 5.02 | 2.08 |
| 986 | 68.93 | 0.60 | 15.42 | 2.03 | 0.89 | 1.98 | 2.32 | 5.05 | 2.09 |
| 976 | 69.01 | 0.61 | 15.48 | 1.93 | 0.83 | 1.98 | 2.33 | 5.07 | 2.10 |
| 966 | 69.09 | 0.61 | 15.53 | 1.83 | 0.77 | 1.98 | 2.34 | 5.09 | 2.11 |
| 956 | 69.17 | 0.61 | 15.58 | 1.74 | 0.71 | 1.98 | 2.35 | 5.11 | 2.12 |
| 946 | 69.25 | 0.61 | 15.63 | 1.64 | 0.66 | 1.97 | 2.36 | 5.13 | 2.13 |
| 936 | 69.33 | 0.61 | 15.67 | 1.54 | 0.61 | 1.97 | 2.37 | 5.15 | 2.14 |
| 926 | 69.41 | 0.62 | 15.72 | 1.45 | 0.57 | 1.96 | 2.38 | 5.17 | 2.14 |
| 916 | 69.62 | 0.63 | 15.65 | 1.36 | 0.53 | 1.90 | 2.32 | 5.24 | 2.18 |
| 906 | 69.89 | 0.64 | 15.52 | 1.27 | 0.49 | 1.81 | 2.24 | 5.35 | 2.23 |
| 896 | 70.15 | 0.65 | 15.40 | 1.18 | 0.45 | 1.73 | 2.15 | 5.45 | 2.28 |
| 886 | 70.40 | 0.67 | 15.28 | 1.10 | 0.42 | 1.65 | 2.06 | 5.55 | 2.33 |
| 876 | 70.68 | 0.65 | 15.16 | 1.01 | 0.39 | 1.58 | 1.96 | 5.65 | 2.39 |
| 866 | 70.97 | 0.61 | 15.05 | 0.92 | 0.37 | 1.52 | 1.86 | 5.75 | 2.44 |
| 856 | 71.25 | 0.58 | 14.94 | 0.84 | 0.35 | 1.46 | 1.76 | 5.84 | 2.49 |
| 846 | 71.51 | 0.55 | 14.84 | 0.77 | 0.32 | 1.41 | 1.67 | 5.93 | 2.54 |
| 836 | 71.76 | 0.52 | 14.74 | 0.70 | 0.31 | 1.36 | 1.57 | 6.01 | 2.59 |
| 826 | 71.99 | 0.49 | 14.64 | 0.63 | 0.29 | 1.31 | 1.47 | 6.09 | 2.63 |
| 816 | 71.84 | 0.46 | 14.74 | 0.58 | 0.27 | 1.27 | 1.38 | 6.27 | 2.74 |
| 806 | 71.63 | 0.42 | 14.88 | 0.54 | 0.27 | 1.27 | 1.33 | 6.25 | 2.94 |
| 796 | 71.41 | 0.39 | 15.02 | 0.51 | 0.26 | 1.27 | 1.29 | 6.21 | 3.16 |
| 786 | 71.16 | 0.36 | 15.18 | 0.48 | 0.25 | 1.27 | 1.25 | 6.17 | 3.38 |
| 776 | 70.89 | 0.32 | 15.35 | 0.46 | 0.25 | 1.27 | 1.22 | 6.11 | 3.63 |
| 766 | 70.52 | 0.28 | 15.56 | 0.49 | 0.21 | 1.27 | 1.18 | 6.04 | 3.88 |
| 756 | 70.13 | 0.24 | 15.79 | 0.52 | 0.18 | 1.26 | 1.15 | 5.96 | 4.15 |
| 746 | 69.70 | 0.21 | 16.05 | 0.54 | 0.16 | 1.26 | 1.12 | 5.87 | 4.44 |
| 736 | 69.10 | 0.19 | 16.56 | 0.51 | 0.14 | 1.26 | 1.05 | 5.78 | 4.68 |
| 726 | 68.36 | 0.17 | 16.47 | 0.49 | 0.12 | 1.52 | 1.12 | 5.60 | 5.10 |
| 716 | 67.03 | 0.17 | 17.57 | 0.38 | 0.11 | 1.77 | 0.92 | 5.50 | 5.14 |
| 706 | 65.87 | 0.17 | 18.56 | 0.30 | 0.11 | 1.98 | 0.77 | 5.37 | 5.19 |
| 696 | 64.75 | 0.17 | 19.46 | 0.25 | 0.10 | 2.18 | 0.66 | 5.23 | 5.25 |
| 686 | 63.54 | 0.16 | 20.22 | 0.20 | 0.09 | 2.50 | 0.58 | 5.06 | 5.33 |
| 676 | 58.87 | 0.13 | 19.55 | 0.22 | 0.09 | 4.36 | 0.58 | 4.34 | 6.20 |

Supplementary Table 2 (cont) Rhyolite-MELTS modeling results: modal proportions
San Francesco ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content = $2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Temperature ( ${ }^{\circ} \mathrm{C}$ ) | \% liquid | \% crystal | \% KId | \% qz | \% plag | \% biotite | \% illmnite | \% cpx |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1006 | 1.00 | 0.00 | - | - | - | - | - | 0.23 |
| 996 | 1.00 | 0.00 | - | - | - | - | - | 0.62 |
| 986 | 0.99 | 0.01 | - | - | - | - | - | 1.01 |
| 976 | 0.98 | 0.02 | - | - | 0.74 | - | - | 1.44 |
| 966 | 0.95 | 0.05 | - | - | 3.12 | - | - | 1.94 |
| 956 | 0.93 | 0.07 | - | - | 5.42 | - | - | 2.42 |
| 946 | 0.90 | 0.10 | - | - | 7.62 | - | - | 2.85 |
| 936 | 0.87 | 0.13 | - | - | 9.74 | - | - | 3.26 |
| 926 | 0.85 | 0.15 | - | - | 11.77 | - | - | 3.64 |
| 916 | 0.83 | 0.17 | - | - | 13.71 | - | - | 4.00 |
| 906 | 0.81 | 0.19 | - | - | 15.56 | - | - | 4.33 |
| 896 | 0.78 | 0.22 | - | - | 17.39 | - | 0.08 | 4.59 |
| 886 | 0.76 | 0.24 | - | - | 19.13 | - | 0.17 | 4.81 |
| 876 | 0.74 | 0.26 | - | - | 20.79 | - | 0.26 | 5.02 |
| 866 | 0.73 | 0.27 | - | - | 22.36 | - | 0.33 | 5.20 |
| 856 | 0.71 | 0.29 | - | - | 23.84 | - | 0.40 | 5.37 |
| 846 | 0.69 | 0.31 | - | - | 25.26 | - | 0.46 | 5.52 |
| 836 | 0.68 | 0.32 | - | - | 26.59 | - | 0.52 | 5.66 |
| 826 | 0.66 | 0.34 | - | - | 27.86 | - | 0.57 | 5.78 |
| 816 | 0.65 | 0.35 | - | - | 29.07 | - | 0.62 | 5.89 |
| 806 | 0.63 | 0.37 | - | 0.55 | 30.36 | - | 0.67 | 5.99 |
| 796 | 0.60 | 0.40 | 0.34 | 1.70 | 31.56 | - | 0.72 | 6.08 |
| 786 | 0.56 | 0.44 | 3.10 | 3.16 | 31.44 | - | 0.79 | 6.15 |
| 776 | 0.51 | 0.49 | 5.39 | 4.90 | 31.61 | 0.43 | 0.87 | 5.84 |
| 766 | 0.47 | 0.53 | 7.25 | 6.83 | 31.99 | 1.23 | 0.95 | 5.24 |
| 756 | 0.43 | 0.57 | 8.92 | 8.53 | 32.33 | 1.87 | 1.01 | 4.75 |
| 746 | 0.38 | 0.62 | 9.22 | 12.65 | 33.72 | 4.44 | 0.63 | - |
| 736 | 0.35 | 0.65 | 10.74 | 13.73 | 33.85 | 4.57 | 0.68 | - |
| 726 | 0.23 | 0.77 | 15.79 | 17.83 | 35.53 | 4.68 | 0.74 | - |
| 716 | 0.10 | 0.90 | 22.79 | 21.31 | 35.03 | 3.56 | 1.18 | - |
| 706 | 0.08 | 0.92 | 23.84 | 22.06 | 35.33 | 3.39 | 1.18 | - |
| 696 | 0.06 | 0.94 | 24.43 | 22.51 | 35.56 | 3.28 | 1.18 | - |
| 686 | 0.05 | 0.95 | 24.79 | 22.81 | 35.75 | 3.20 | 1.18 | - |
| 676 | 0.00 | 1.00 | 26.44 | 24.36 | 35.74 | 3.27 | 1.20 | - |

Sant'Andrea ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content = $2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Temperature ( ${ }^{\circ} \mathrm{C}$ ) | \% liquid | \% crystal | \% KId | \% qz | \% plag | \% biotite | \% illmnite | \% cpx |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 968 | 1.00 | 0.00 | - | - | - | - | - | 0.94 |
| 958 | 0.99 | 0.01 | - | - | 0.98 | - | - | 1.34 |
| 948 | 0.96 | 0.04 | - | - | 3.16 | - | - | 1.78 |
| 938 | 0.94 | 0.06 | - | - | 5.26 | - | - | 2.20 |
| 928 | 0.91 | 0.09 | - | - | 7.28 | - | - | 2.59 |
| 918 | 0.89 | 0.11 | - | - | 9.21 | - | - | 2.95 |
| 908 | 0.87 | 0.13 | - | - | 11.06 | - | - | 3.29 |
| 898 | 0.85 | 0.15 | - | - | 12.83 | - | - | 3.61 |
| 888 | 0.83 | 0.17 | - | - | 14.52 | - | - | 3.90 |
| 878 | 0.81 | 0.19 | - | - | 16.13 | - | - | 4.18 |
| 868 | 0.79 | 0.21 | - | - | 17.66 | - | - | 4.43 |
| 858 | 0.78 | 0.23 | - | - | 19.12 | - | - | 4.66 |
| 848 | 0.76 | 0.24 | - | - | 20.50 | - | - | 4.88 |
| 838 | 0.74 | 0.26 | - | - | 21.84 | - | 0.05 | 5.03 |
| 828 | 0.71 | 0.29 | - | 1.29 | 23.45 | - | 0.13 | 5.18 |
| 818 | 0.68 | 0.32 | - | 2.53 | 24.96 | - | 0.20 | 5.30 |
| 808 | 0.66 | 0.35 | - | 3.70 | 26.39 | - | 0.26 | 5.41 |
| 798 | 0.62 | 0.38 | 0.89 | 4.94 | 27.29 | - | 0.33 | 5.51 |
| 788 | 0.58 | 0.42 | 3.60 | 6.41 | 27.22 | - | 0.40 | 5.58 |
| 778 | 0.54 | 0.46 | 6.03 | 7.77 | 27.23 | - | 0.46 | 5.64 |
| 768 | 0.50 | 0.50 | 8.11 | 9.28 | 27.40 | 0.27 | 0.52 | 5.48 |
| 758 | 0.46 | 0.54 | 9.79 | 10.99 | 27.77 | 0.91 | 0.58 | 5.00 |
| 748 | 0.43 | 0.57 | 11.32 | 12.52 | 28.10 | 1.45 | 0.63 | 4.61 |
| 738 | 0.38 | 0.62 | 11.53 | 16.34 | 29.34 | 3.73 | 0.24 | - |
| 728 | 0.34 | 0.66 | 13.63 | 17.95 | 29.76 | 3.84 | 0.29 | - |
| 718 | 0.18 | 0.83 | 21.99 | 22.39 | 29.53 | 2.65 | 0.75 | - |
| 708 | 0.14 | 0.86 | 23.47 | 23.52 | 29.98 | 2.52 | 0.76 | - |
| 698 | 0.12 | 0.88 | 24.38 | 24.24 | 30.30 | 2.43 | 0.77 | - |
| 688 | 0.00 | 1.00 | 25.64 | 25.32 | 30.52 | 2.38 | 0.78 | - |

San Piero ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content $=2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Temperature ( ${ }^{\circ} \mathrm{C}$ ) | \% liquid | \% crystal | \% KId | \% qz | \% plag | \% biotite | \% illmnite | \% cpx |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1076 | 1.00 | 0.00 | - | - | - | - | - | 0.26 |
| 1066 | 0.99 | 0.01 | - | - | - | - | - | 0.78 |
| 1056 | 0.99 | 0.01 | - | - | - | - | - | 1.27 |
| 1046 | 0.98 | 0.02 | - | - | - | - | - | 1.76 |
| 1036 | 0.98 | 0.02 | - | - | - | - | - | 2.23 |
| 1026 | 0.97 | 0.03 | - | - | - | - | - | 2.69 |
| 1016 | 0.97 | 0.03 | - | - | - | - | - | 3.13 |
| 1006 | 0.96 | 0.04 | - | - | - | - | - | 3.57 |
| 996 | 0.96 | 0.04 | - | - | - | - | - | 3.99 |
| 986 | 0.96 | 0.05 | - | - | - | - | - | 4.40 |
| 976 | 0.95 | 0.05 | - | - | - | - | - | 4.80 |
| 966 | 0.95 | 0.05 | - | - | - | - | - | 5.19 |
| 956 | 0.94 | 0.06 | - | - | - | - | - | 5.57 |
| 946 | 0.94 | 0.06 | - | - | - | - | - | 5.93 |
| 936 | 0.94 | 0.06 | - | - | - | - | - | 6.29 |
| 926 | 0.93 | 0.07 | - | - | - | - | - | 6.63 |
| 916 | 0.92 | 0.08 | - | - | - | - | - | 7.00 |
| 906 | 0.90 | 0.10 | - | - | - | - | - | 7.36 |
| 896 | 0.88 | 0.12 | - | - | - | - | - | 7.70 |
| 886 | 0.86 | 0.14 | - | - | - | - | - | 8.00 |
| 876 | 0.84 | 0.16 | - | - | - | - | 0.06 | 8.24 |
| 866 | 0.82 | 0.18 | - | - | - | - | 0.14 | 8.43 |
| 856 | 0.80 | 0.20 | - | - | - | - | 0.22 | 8.60 |
| 846 | 0.79 | 0.21 | - | - | 1.20 | - | 0.29 | 8.76 |
| 836 | 0.77 | 0.23 | - | - | 4.32 | - | 0.36 | 8.89 |
| 826 | 0.76 | 0.24 | - | - | 7.05 | - | 0.42 | 9.02 |
| 816 | 0.73 | 0.27 | 0.03 | 1.20 | 9.45 | - | 0.49 | 9.13 |
| 806 | 0.68 | 0.32 | 3.08 | 2.90 | 11.58 | - | 0.59 | 9.21 |
| 796 | 0.63 | 0.37 | 5.80 | 4.48 | 13.49 | - | 0.68 | 9.29 |
| 786 | 0.59 | 0.41 | 8.24 | 5.95 | 15.22 | - | 0.75 | 9.35 |
| 776 | 0.55 | 0.45 | 10.42 | 7.39 | 16.58 | 0.07 | 0.82 | 9.35 |
| 766 | 0.50 | 0.50 | 11.96 | 9.63 | 17.55 | 1.23 | 0.91 | 8.47 |
| 756 | 0.46 | 0.54 | 13.35 | 11.56 | 18.41 | 2.17 | 0.97 | 7.75 |
| 746 | 0.42 | 0.58 | 14.60 | 13.27 | 19.06 | 2.97 | 1.03 | 7.15 |
| 736 | 0.37 | 0.63 | 14.08 | 18.53 | 20.28 | 6.45 | 0.42 | - |
| 726 | 0.23 | 0.77 | 21.84 | 21.97 | 24.53 | 5.24 | 1.06 | - |
| 716 | 0.15 | 0.85 | 25.46 | 24.33 | 26.30 | 4.76 | 1.16 | - |
| 706 | 0.12 | 0.88 | 26.82 | 25.38 | 27.32 | 4.57 | 1.16 | - |
| 696 | 0.10 | 0.90 | 27.64 | 26.05 | 27.96 | 4.43 | 1.17 | - |
| 686 | 0.08 | 0.92 | 28.33 | 26.67 | 30.74 | 4.34 | 1.17 | - |
| 676 | 0.00 | 1.00 | 31.08 | 29.32 | 31.08 | 4.51 | 1.20 | - |

Supplementary Table 2 (cont) Rhyolite-MELTS modeling results: feldspar compositions
San Francesco ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content $=2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

## Plagioclase

| Temperature ( ${ }^{\circ} \mathrm{C}$ ) | wt\% SiO2 | wt\% Al2O3 | wt\% CaO | wt\% Na2O | wt\% K2O | \% ab | \% an | \% or |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 976 | 58.79 | 26.01 | 7.74 | 6.85 | 0.62 | 0.59 | 0.37 | 0.04 |
| 966 | 59.12 | 25.77 | 7.47 | 6.98 | 0.66 | 0.61 | 0.36 | 0.04 |
| 956 | 59.44 | 25.55 | 7.21 | 7.10 | 0.70 | 0.62 | 0.35 | 0.04 |
| 946 | 59.73 | 25.35 | 6.97 | 7.21 | 0.74 | 0.62 | 0.33 | 0.04 |
| 936 | 60.00 | 25.16 | 6.75 | 7.31 | 0.78 | 0.63 | 0.32 | 0.04 |
| 926 | 60.24 | 24.98 | 6.55 | 7.40 | 0.83 | 0.64 | 0.31 | 0.05 |
| 916 | 60.47 | 24.82 | 6.37 | 7.48 | 0.87 | 0.65 | 0.30 | 0.05 |
| 906 | 60.68 | 24.67 | 6.19 | 7.55 | 0.91 | 0.65 | 0.30 | 0.05 |
| 896 | 60.86 | 24.53 | 6.04 | 7.61 | 0.96 | 0.66 | 0.29 | 0.05 |
| 886 | 61.03 | 24.41 | 5.90 | 7.66 | 1.01 | 0.66 | 0.28 | 0.06 |
| 876 | 61.19 | 24.29 | 5.76 | 7.70 | 1.05 | 0.67 | 0.28 | 0.06 |
| 866 | 61.33 | 24.19 | 5.64 | 7.74 | 1.10 | 0.67 | 0.27 | 0.06 |
| 856 | 61.46 | 24.09 | 5.53 | 7.77 | 1.15 | 0.67 | 0.26 | 0.07 |
| 846 | 61.58 | 24.00 | 5.43 | 7.80 | 1.20 | 0.67 | 0.26 | 0.07 |
| 836 | 61.69 | 23.91 | 5.33 | 7.82 | 1.25 | 0.67 | 0.25 | 0.07 |
| 826 | 61.79 | 23.83 | 5.25 | 7.84 | 1.29 | 0.68 | 0.25 | 0.07 |
| 816 | 61.88 | 23.76 | 5.17 | 7.85 | 1.34 | 0.68 | 0.25 | 0.08 |
| 806 | 61.95 | 23.70 | 5.10 | 7.84 | 1.41 | 0.68 | 0.24 | 0.08 |
| 796 | 61.98 | 23.66 | 5.06 | 7.82 | 1.48 | 0.67 | 0.24 | 0.08 |
| 786 | 61.81 | 23.80 | 5.21 | 7.79 | 1.39 | 0.67 | 0.25 | 0.08 |
| 776 | 61.65 | 23.92 | 5.35 | 7.77 | 1.30 | 0.67 | 0.26 | 0.07 |
| 766 | 61.51 | 24.04 | 5.48 | 7.75 | 1.23 | 0.67 | 0.26 | 0.07 |
| 756 | 61.39 | 24.13 | 5.59 | 7.73 | 1.16 | 0.67 | 0.27 | 0.07 |
| 746 | 61.25 | 24.24 | 5.71 | 7.71 | 1.09 | 0.67 | 0.27 | 0.06 |
| 736 | 61.16 | 24.31 | 5.79 | 7.71 | 1.03 | 0.67 | 0.28 | 0.06 |
| 726 | 61.07 | 24.39 | 5.87 | 7.70 | 0.98 | 0.66 | 0.28 | 0.06 |
| 716 | 61.21 | 24.30 | 5.76 | 7.78 | 0.95 | 0.67 | 0.27 | 0.05 |
| 706 | 61.18 | 24.33 | 5.80 | 7.79 | 0.90 | 0.67 | 0.28 | 0.05 |
| 696 | 61.16 | 24.35 | 5.82 | 7.80 | 0.86 | 0.67 | 0.28 | 0.05 |
| 686 | 61.15 | 24.37 | 5.84 | 7.82 | 0.82 | 0.68 | 0.28 | 0.05 |
| 676 | 61.15 | 24.38 | 5.84 | 7.85 | 0.78 | 0.68 | 0.28 | 0.04 |

K-feldspar

| Temperature ( ${ }^{\circ} \mathbf{C}$ ) | wt\% SiO2 | wt\% Al2O3 | $\mathbf{w t} \% \mathbf{C a O}$ | $\mathbf{w t} \% \mathbf{N a 2 O}$ | $\mathbf{w t \%} \mathbf{K 2 O}$ | $\% \mathbf{a b}$ | $\%$ an | $\%$ or |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 796 |  |  |  |  |  |  |  |  |
| 786 | 65.56 | 19.08 | 0.44 | 3.79 | 11.13 | 0.33 | 0.02 | 0.64 |
| 776 | 65.55 | 19.04 | 0.41 | 3.66 | 11.34 | 0.32 | 0.02 | 0.66 |
| 766 | 65.54 | 19.00 | 0.39 | 3.53 | 11.54 | 0.31 | 0.02 | 0.67 |
| 756 | 65.52 | 18.97 | 0.36 | 3.41 | 11.73 | 0.30 | 0.02 | 0.68 |
| 746 | 65.51 | 18.94 | 0.34 | 3.30 | 11.91 | 0.29 | 0.02 | 0.69 |
| 736 | 65.50 | 18.91 | 0.31 | 3.19 | 12.09 | 0.28 | 0.02 | 0.70 |
| 726 | 65.49 | 18.88 | 0.29 | 3.10 | 12.24 | 0.27 | 0.01 | 0.71 |
| 716 | 65.48 | 18.85 | 0.27 | 3.00 | 12.40 | 0.27 | 0.01 | 0.72 |
| 706 | 65.49 | 18.83 | 0.25 | 2.96 | 12.47 | 0.26 | 0.01 | 0.73 |
| 696 | 65.48 | 18.81 | 0.24 | 2.88 | 12.60 | 0.25 | 0.01 | 0.73 |
| 686 | 65.47 | 18.78 | 0.22 | 2.80 | 12.72 | 0.25 | 0.01 | 0.74 |
| 676 | 65.46 | 18.76 | 0.21 | 2.73 | 12.85 | 0.24 | 0.01 | 0.75 |
|  | 65.45 | 18.74 | 0.19 | 2.66 | 12.96 | 0.24 | 0.01 | 0.76 |

Supplementary Table 2 (cont) Rhyolite-MELTS modeling results: feldspar compositions
Sant'Andrea ( $\mathrm{P}=2.3$ Kbar; initial water content $=2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Plagioclase |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Temperature ( ${ }^{\circ} \mathrm{C}$ ) | wt\% SiO2 | wt\% Al2O3 | wt\% CaO | wt\% Na2O | wt\% K2O | \% ab | \% an | \% or |
| 958 | 59.21 | 25.71 | 7.39 | 7.01 | 0.67 | 0.61 | 0.35 | 0.04 |
| 948 | 59.52 | 25.49 | 7.14 | 7.13 | 0.71 | 0.62 | 0.34 | 0.04 |
| 938 | 59.81 | 25.29 | 6.91 | 7.24 | 0.75 | 0.63 | 0.33 | 0.04 |
| 928 | 60.07 | 25.10 | 6.69 | 7.34 | 0.80 | 0.63 | 0.32 | 0.05 |
| 918 | 60.32 | 24.93 | 6.49 | 7.43 | 0.84 | 0.64 | 0.31 | 0.05 |
| 908 | 60.54 | 24.77 | 6.31 | 7.50 | 0.88 | 0.65 | 0.30 | 0.05 |
| 898 | 60.74 | 24.62 | 6.14 | 7.57 | 0.92 | 0.65 | 0.29 | 0.05 |
| 888 | 60.93 | 24.49 | 5.98 | 7.63 | 0.97 | 0.66 | 0.29 | 0.05 |
| 878 | 61.10 | 24.36 | 5.84 | 7.69 | 1.01 | 0.66 | 0.28 | 0.06 |
| 868 | 61.25 | 24.25 | 5.71 | 7.74 | 1.05 | 0.67 | 0.27 | 0.06 |
| 858 | 61.40 | 24.14 | 5.59 | 7.78 | 1.09 | 0.67 | 0.27 | 0.06 |
| 848 | 61.52 | 24.05 | 5.48 | 7.81 | 1.14 | 0.67 | 0.26 | 0.06 |
| 838 | 61.64 | 23.96 | 5.39 | 7.84 | 1.18 | 0.68 | 0.26 | 0.07 |
| 828 | 61.71 | 23.90 | 5.32 | 7.82 | 1.25 | 0.68 | 0.25 | 0.07 |
| 818 | 61.77 | 23.83 | 5.25 | 7.81 | 1.34 | 0.67 | 0.25 | 0.08 |
| 808 | 61.84 | 23.77 | 5.18 | 7.79 | 1.42 | 0.67 | 0.25 | 0.08 |
| 798 | 61.82 | 23.77 | 5.19 | 7.76 | 1.46 | 0.67 | 0.25 | 0.08 |
| 788 | 61.64 | 23.92 | 5.35 | 7.73 | 1.37 | 0.67 | 0.26 | 0.08 |
| 778 | 61.47 | 24.05 | 5.50 | 7.70 | 1.28 | 0.66 | 0.26 | 0.07 |
| 768 | 61.33 | 24.17 | 5.63 | 7.68 | 1.21 | 0.66 | 0.27 | 0.07 |
| 758 | 61.19 | 24.27 | 5.75 | 7.66 | 1.14 | 0.66 | 0.27 | 0.06 |
| 748 | 61.08 | 24.36 | 5.85 | 7.64 | 1.07 | 0.66 | 0.28 | 0.06 |
| 738 | 60.97 | 24.45 | 5.94 | 7.63 | 1.01 | 0.66 | 0.28 | 0.06 |
| 728 | 60.89 | 24.52 | 6.02 | 7.62 | 0.96 | 0.66 | 0.29 | 0.05 |
| 718 | 60.99 | 24.45 | 5.94 | 7.69 | 0.93 | 0.66 | 0.28 | 0.05 |
| 708 | 60.94 | 24.50 | 5.99 | 7.69 | 0.88 | 0.66 | 0.29 | 0.05 |
| 698 | 60.91 | 24.53 | 6.02 | 7.70 | 0.84 | 0.66 | 0.29 | 0.05 |
| 688 | 60.87 | 24.56 | 6.06 | 7.71 | 0.80 | 0.67 | 0.29 | 0.05 |

K-feldspar

| Temperature $\left({ }^{\circ} \mathbf{C}\right)$ | wt\% SiO2 | wt\% Al2O3 | wt $\% \mathbf{C a O}$ | $\mathbf{w t} \% \mathbf{N a 2 O}$ | wt\% K2O | \% ab | \% an | \% or |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 798 |  |  |  |  |  |  |  |  |
| 788 | 65.55 | 19.08 | 0.45 | 3.75 | 11.18 | 0.33 | 0.02 | 0.65 |
| 778 | 65.53 | 19.04 | 0.42 | 3.61 | 11.40 | 0.32 | 0.02 | 0.66 |
| 768 | 65.52 | 19.00 | 0.39 | 3.49 | 11.60 | 0.31 | 0.02 | 0.67 |
| 758 | 65.51 | 18.97 | 0.36 | 3.37 | 11.79 | 0.30 | 0.02 | 0.68 |
| 748 | 65.49 | 18.94 | 0.34 | 3.26 | 11.97 | 0.29 | 0.02 | 0.70 |
| 738 | 65.48 | 18.91 | 0.32 | 3.16 | 12.13 | 0.28 | 0.02 | 0.71 |
| 728 | 65.47 | 18.88 | 0.30 | 3.06 | 12.29 | 0.27 | 0.01 | 0.72 |
| 718 | 65.46 | 18.85 | 0.28 | 2.97 | 12.44 | 0.26 | 0.01 | 0.72 |
| 708 | 65.47 | 18.83 | 0.26 | 2.92 | 12.52 | 0.26 | 0.01 | 0.73 |
| 698 | 65.46 | 18.81 | 0.24 | 2.84 | 12.66 | 0.25 | 0.01 | 0.74 |
| 688 | 65.45 | 18.78 | 0.23 | 2.76 | 12.78 | 0.24 | 0.01 | 0.74 |
|  | 65.44 | 18.76 | 0.21 | 2.68 | 12.91 | 0.24 | 0.01 | 0.75 |

Supplementary Table 2 (cont) Rhyolite-MELTS modeling results: feldspar compositions
San Piero ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content $=2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Plagioclase |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Temperature ( ${ }^{\circ}$ C) | wt\% SiO2 | wt\% Al2O3 | wt\% CaO | wt $\%$ Na2O | wt\% K2O | \% ab | \% an | \% or |
|  |  |  |  |  |  |  |  |  |
| 916 | 59.47 | 25.50 | 7.16 | 7.02 | 0.86 | 0.61 | 0.34 | 0.05 |
| 906 | 59.75 | 25.29 | 6.93 | 7.12 | 0.91 | 0.62 | 0.33 | 0.05 |
| 896 | 60.01 | 25.11 | 6.71 | 7.21 | 0.96 | 0.62 | 0.32 | 0.05 |
| 886 | 60.24 | 24.94 | 6.52 | 7.29 | 1.01 | 0.63 | 0.31 | 0.06 |
| 876 | 60.45 | 24.78 | 6.34 | 7.35 | 1.07 | 0.64 | 0.30 | 0.06 |
| 866 | 60.64 | 24.65 | 6.18 | 7.41 | 1.12 | 0.64 | 0.30 | 0.06 |
| 856 | 60.81 | 24.52 | 6.04 | 7.45 | 1.18 | 0.64 | 0.29 | 0.07 |
| 846 | 60.97 | 24.40 | 5.90 | 7.49 | 1.24 | 0.65 | 0.28 | 0.07 |
| 836 | 61.11 | 24.29 | 5.78 | 7.52 | 1.30 | 0.65 | 0.28 | 0.07 |
| 826 | 61.24 | 24.19 | 5.67 | 7.54 | 1.37 | 0.65 | 0.27 | 0.08 |
| 816 | 61.32 | 24.11 | 5.59 | 7.52 | 1.46 | 0.65 | 0.27 | 0.08 |
| 806 | 61.02 | 24.34 | 5.84 | 7.45 | 1.35 | 0.64 | 0.28 | 0.08 |
| 796 | 60.76 | 24.54 | 6.07 | 7.38 | 1.26 | 0.64 | 0.29 | 0.07 |
| 786 | 60.51 | 24.72 | 6.27 | 7.32 | 1.17 | 0.63 | 0.30 | 0.07 |
| 776 | 60.29 | 24.88 | 6.46 | 7.27 | 1.10 | 0.63 | 0.31 | 0.06 |
| 766 | 60.08 | 25.04 | 6.64 | 7.21 | 1.03 | 0.62 | 0.32 | 0.06 |
| 756 | 59.91 | 25.18 | 6.79 | 7.16 | 0.96 | 0.62 | 0.32 | 0.05 |
| 746 | 59.76 | 25.29 | 6.92 | 7.13 | 0.91 | 0.62 | 0.33 | 0.05 |
| 736 | 59.60 | 25.41 | 7.06 | 7.09 | 0.85 | 0.61 | 0.34 | 0.05 |
| 726 | 59.88 | 25.22 | 6.83 | 7.22 | 0.84 | 0.63 | 0.33 | 0.05 |
| 716 | 59.79 | 25.29 | 6.92 | 7.21 | 0.80 | 0.62 | 0.33 | 0.05 |
| 706 | 59.69 | 25.37 | 7.00 | 7.19 | 0.75 | 0.62 | 0.33 | 0.04 |
| 696 | 59.63 | 25.41 | 7.05 | 7.18 | 0.72 | 0.62 | 0.34 | 0.04 |
| 686 | 59.59 | 25.45 | 7.09 | 7.18 | 0.68 | 0.62 | 0.34 | 0.04 |
| 676 | 59.58 | 25.47 | 7.11 | 7.20 | 0.65 | 0.62 | 0.34 | 0.04 |

K-feldspar

| Temperature ( $\left.{ }^{\circ} \mathbf{C}\right)$ | $\mathbf{w t} \% \mathbf{S i O 2}$ | $\mathbf{w t} \% \mathbf{A l 2 O 3}$ | $\mathbf{w t} \% \mathbf{C a O}$ | $\mathbf{w t} \% \mathbf{N a 2 O}$ | $\mathbf{w t} \% \mathbf{K 2 O}$ | $\% \mathbf{a b}$ | $\%$ an | $\%$ or |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 816 |  |  |  |  |  |  |  |  |
| 806 | 65.47 | 19.13 | 0.50 | 3.71 | 11.19 | 0.33 | 0.02 | 0.65 |
| 796 | 65.45 | 19.08 | 0.47 | 3.55 | 11.45 | 0.31 | 0.02 | 0.66 |
| 786 | 65.44 | 19.04 | 0.44 | 3.40 | 11.68 | 0.30 | 0.02 | 0.68 |
| 776 | 65.42 | 19.00 | 0.41 | 3.27 | 11.90 | 0.29 | 0.02 | 0.69 |
| 766 | 65.41 | 18.97 | 0.39 | 3.14 | 12.10 | 0.28 | 0.02 | 0.70 |
| 756 | 65.39 | 18.93 | 0.36 | 3.03 | 12.28 | 0.27 | 0.02 | 0.71 |
| 746 | 65.38 | 18.90 | 0.34 | 2.92 | 12.46 | 0.26 | 0.02 | 0.72 |
| 736 | 65.37 | 18.88 | 0.32 | 2.82 | 12.61 | 0.25 | 0.02 | 0.73 |
| 726 | 65.36 | 18.85 | 0.30 | 2.73 | 12.77 | 0.24 | 0.01 | 0.74 |
| 716 | 65.38 | 18.83 | 0.28 | 2.72 | 12.79 | 0.24 | 0.01 | 0.75 |
| 706 | 65.37 | 18.81 | 0.26 | 2.64 | 12.92 | 0.23 | 0.01 | 0.75 |
| 696 | 65.36 | 18.78 | 0.24 | 2.56 | 13.06 | 0.23 | 0.01 | 0.76 |
| 686 | 65.35 | 18.76 | 0.23 | 2.48 | 13.18 | 0.22 | 0.01 | 0.77 |
| 676 | 65.34 | 18.74 | 0.21 | 2.41 | 13.29 | 0.21 | 0.01 | 0.78 |

Supplementary Table 2 (cont) Rhyolite-MELTS modeling results: biotite compositions
San Francesco ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content $=2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Temperature ( ${ }^{\circ} \mathrm{C}$ ) | wt\% SiO2 | wt\% Al2O3 | wt\% FeO | wt\% MgO | wt\% K2O | wt\% H2O |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 776 | 42.93 | 12.14 | 1.40 | 28.02 | 11.22 | 4.29 |
| 766 | 42.82 | 12.11 | 1.98 | 27.62 | 11.19 | 4.28 |
| 756 | 42.68 | 12.07 | 2.74 | 27.09 | 11.15 | 4.27 |
| 746 | 42.51 | 12.02 | 3.61 | 26.49 | 11.11 | 4.25 |
| 736 | 42.40 | 11.99 | 4.23 | 26.07 | 11.08 | 4.24 |
| 726 | 42.34 | 11.97 | 4.55 | 25.85 | 11.06 | 4.23 |
| 716 | 42.41 | 12.00 | 4.15 | 26.12 | 11.08 | 4.24 |
| 706 | 42.57 | 12.04 | 3.33 | 26.68 | 11.12 | 4.25 |
| 696 | 42.67 | 12.07 | 2.79 | 27.06 | 11.15 | 4.26 |
| 686 | 42.75 | 12.09 | 2.39 | 27.33 | 11.17 | 4.27 |
| 676 | 42.73 | 12.08 | 2.49 | 27.26 | 11.16 | 4.27 |

Sant'Andrea (P = 2.3 Kbar; initial water content = $2 \mathrm{wt} \mathrm{\%}$; OPX out; Qz-Fa-Mag buffer)

| Temperature $\left({ }^{\circ} \mathbf{C}\right)$ | wt\% SiO2 | wt\% Al2O3 | wt\% FeO | $\mathbf{w t} \% \mathbf{M g O}$ | $\mathbf{w t} \% \mathbf{K 2 O}$ | $\mathbf{w t \%} \mathbf{~ H 2 O}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | ---: |
| 768 | 42.86 | 12.12 | 1.79 | 27.75 | 11.20 | 4.28 |
| 758 | 42.73 | 12.09 | 2.47 | 27.28 | 11.17 | 4.27 |
| 748 | 42.56 | 12.04 | 3.36 | 26.67 | 11.12 | 4.25 |
| 738 | 42.46 | 12.01 | 3.90 | 26.29 | 11.09 | 4.24 |
| 728 | 42.35 | 11.98 | 4.48 | 25.90 | 11.07 | 4.23 |
| 718 | 42.49 | 12.02 | 3.74 | 26.40 | 11.10 | 4.25 |
| 708 | 42.62 | 12.06 | 3.03 | 26.89 | 11.14 | 4.26 |
| 698 | 42.72 | 12.08 | 2.55 | 27.23 | 11.16 | 4.27 |
| 688 | 42.77 | 12.10 | 2.24 | 27.44 | 11.18 | 4.28 |

San Piero ( $\mathrm{P}=2.3 \mathrm{Kbar}$; initial water content $=2 \mathrm{wt} \%$; OPX out; Qz-Fa-Mag buffer)

| Temperature ( $\left.{ }^{\circ} \mathbf{C}\right)$ | wt\% SiO2 | wt\% Al2O3 | wt\% FeO | wt\% MgO | wt\% K2O | wt\% H2O |
| :---: | :---: | :---: | :---: | :---: | :---: | ---: |
| 776 | 42.95 | 12.15 | 1.32 | 28.07 | 11.22 | 4.29 |
| 766 | 42.85 | 12.12 | 1.86 | 27.70 | 11.20 | 4.28 |
| 756 | 42.71 | 12.08 | 2.57 | 27.21 | 11.16 | 4.27 |
| 746 | 42.54 | 12.03 | 3.48 | 26.58 | 11.12 | 4.25 |
| 736 | 42.44 | 12.00 | 3.98 | 26.24 | 11.09 | 4.24 |
| 726 | 42.31 | 11.97 | 4.68 | 25.76 | 11.06 | 4.23 |
| 716 | 42.50 | 12.02 | 3.69 | 26.44 | 11.11 | 4.25 |
| 706 | 42.62 | 12.06 | 3.03 | 26.89 | 11.14 | 4.26 |
| 696 | 42.71 | 12.08 | 2.57 | 27.21 | 11.16 | 4.27 |
| 686 | 42.77 | 12.10 | 2.24 | 27.44 | 11.18 | 4.28 |
| 676 | 42.74 | 12.09 | 2.44 | 27.30 | 11.17 | 4.27 |

Supplementary Table 3a Zircon saturation temperature modeling (Sant'Andrea granite)

| $\mathbf{T}^{\circ} \mathbf{C}$ | M values | Zr for saturation (ppm) | Zr for saturation (ppm) <br> MELTS outputs | Calculated from MELTS compositions |
| :---: | :---: | :---: | :---: | :---: | | Watson and Harrison (1983) |
| :---: |
| Boehnke et al., (2013) | using kd's and MELTS outputs

Supplementary Table 3b Zircon saturation temperature modeling (San Francesco granite)

| $\mathrm{T}^{\circ} \mathrm{C}$ | $M$ values | Zr for saturation (ppm) | Zr for saturation (ppm) | Zr melt (ppm) |
| :---: | :---: | :---: | :---: | :---: |
| MELTS outputs | Calculated from MELTS compositions | Watson and Harrison (1983) | Boehnke et al., (2013) | sing kd's and MELTS outputs |
| 1006 | 1.53 | 1454 | 1493 | - |
| 996 | 1.53 | 1342 | 1402 | 160 |
| 986 | 1.53 | 1237 | 1315 | 161 |
| 976 | 1.52 | 1132 | 1222 | 163 |
| 966 | 1.50 | 1021 | 1114 | 167 |
| 956 | 1.47 | 920 | 1016 | 171 |
| 946 | 1.45 | 830 | 927 | 176 |
| 936 | 1.43 | 748 | 847 | 181 |
| 926 | 1.42 | 674 | 773 | 185 |
| 916 | 1.40 | 607 | 706 | 189 |
| 906 | 1.38 | 546 | 645 | 193 |
| 896 | 1.37 | 490 | 588 | 200 |
| 886 | 1.35 | 440 | 536 | 205 |
| 876 | 1.34 | 395 | 489 | 210 |
| 866 | 1.32 | 353 | 445 | 212 |
| 856 | 1.31 | 316 | 405 | 218 |
| 846 | 1.30 | 282 | 369 | 224 |
| 836 | 1.28 | 252 | 335 | 227 |
| 826 | 1.27 | 224 | 304 | 233 |
| 816 | 1.26 | 199 | 276 | 236 |
| 806 | 1.25 | 177 | 250 | 243 |
| 796 | 1.24 | 157 | 226 | 255 |
| 786 | 1.23 | 138 | 204 | 272 |
| 776 | 1.22 | 122 | 184 | 297 |
| 766 | 1.20 | 107 | 165 | 321 |
| 756 | 1.19 | 94 | 148 | 349 |
| 746 | 1.18 | 82 | 132 | 396 |
| 736 | 1.16 | 72 | 118 | 428 |
| 726 | 1.07 | 58 | 96 | 639 |
| 716 | 1.27 | 60 | 108 | 1421 |
| 706 | 1.21 | 50 | 92 | 1761 |
| 696 | 1.19 | 43 | 81 | 2320 |
| 686 | 1.20 | 38 | 73 | 2763 |
| Kd K-fld | 0.03 | Initial melt $\mathrm{Zr}=$ bulk rock Zr content (sample PP-364 from Dini et al., 2002) |  |  |
| Kd plag | 0.05 |  |  |  |
| Kd quartz | 0 |  |  |  |
| Kd bt | 0.432 |  |  |  |
| Kd cpx | 0.278 |  |  |  |
| Kd illmenite | 0.568 |  |  |  |

Supplementary Table 3c Zircon saturation temperature modeling (San Piero granite)

| $\mathrm{T}^{\circ} \mathrm{C}$ | M values | Zr for saturation (ppm) | Zr for saturation (ppm) | Zr melt (ppm) |
| :---: | :---: | :---: | :---: | :---: |
| MELTS outputs | Calculated from MELTS compositions | Watson and Harrison (1983) | Boehnke et al., (2013) | sing kd's and MELTS outputs |
| 1076 | 1.33 | 2070 | 1787 | - |
| 1066 | 1.33 | 1928 | 1690 | 136 |
| 1056 | 1.33 | 1793 | 1597 | 136 |
| 1046 | 1.33 | 1666 | 1508 | 137 |
| 1036 | 1.33 | 1546 | 1422 | 137 |
| 1026 | 1.33 | 1433 | 1339 | 138 |
| 1016 | 1.33 | 1326 | 1260 | 138 |
| 1006 | 1.33 | 1225 | 1184 | 139 |
| 996 | 1.33 | 1131 | 1112 | 139 |
| 986 | 1.33 | 1042 | 1042 | 140 |
| 976 | 1.33 | 959 | 976 | 140 |
| 966 | 1.33 | 882 | 913 | 140 |
| 956 | 1.33 | 809 | 853 | 141 |
| 946 | 1.32 | 741 | 796 | 141 |
| 936 | 1.32 | 678 | 741 | 142 |
| 926 | 1.32 | 619 | 689 | 142 |
| 916 | 1.30 | 558 | 631 | 144 |
| 906 | 1.28 | 500 | 573 | 146 |
| 896 | 1.26 | 447 | 520 | 149 |
| 886 | 1.24 | 400 | 472 | 151 |
| 876 | 1.23 | 358 | 428 | 153 |
| 866 | 1.21 | 319 | 388 | 156 |
| 856 | 1.19 | 285 | 352 | 158 |
| 846 | 1.17 | 253 | 319 | 161 |
| 836 | 1.16 | 225 | 289 | 165 |
| 826 | 1.14 | 200 | 261 | 169 |
| 816 | 1.13 | 177 | 236 | 176 |
| 806 | 1.11 | 157 | 212 | 189 |
| 796 | 1.09 | 138 | 190 | 202 |
| 786 | 1.07 | 121 | 170 | 216 |
| 776 | 1.05 | 106 | 151 | 232 |
| 766 | 1.03 | 92 | 134 | 253 |
| 756 | 1.00 | 80 | 119 | 274 |
| 746 | 0.98 | 69 | 105 | 297 |
| 736 | 0.93 | 59 | 90 | 343 |
| 726 | 1.01 | 55 | 89 | 540 |
| 716 | 0.99 | 48 | 79 | 794 |
| 706 | 0.97 | 41 | 69 | 985 |
| 696 | 0.96 | 36 | 62 | 1169 |
| 686 | 0.99 | 32 | 57 | 1421 |
| Kd K-fld | 0.03 | Kd from EarthRef.org |  |  |
| Kd plag | 0.05 | Initial melt $\mathrm{Zr}=$ bulk rock Zr co | ntent (sample MB11-2) |  |
| Kd quartz | 0 |  |  |  |
| Kd bt | 0.432 |  |  |  |
| Kd cpx | 0.278 |  |  |  |
| Kd illmenite | 0.568 |  |  |  |



Supplementary Fig. 3a. Zr saturation model (Sant'Andrea facies)


Supplementary Fig. 3b. Zr saturation model (San Francesco facies)


Supplementary Fig. 3c. Zr saturation model (San Piero facies)

