

## §8. Observation of Global Alfvén Eigenmodes Destabilized in the CHS Plasmas with Finite Magnetic Shear

Takechi, M., Matsunaga, G. (Dep. Energy Eng. Science, Nagoya Univ.), Toi, K., CHS Group

We are studying toroidal Alfvén eigenmodes (TAEs) in the CHS heliotron/torsatron, focusing on the destabilization condition and internal structure of TAEs with low toroidal mode numbers,  $n=1$  and 2. In CHS plasma, TAE gap can be formed due to its moderate magnetic shear. On the other hand global Alfvén eigenmodes (GAEs) are easily destabilized in the W7-AS stellarator with very low magnetic shear. Recently, two kinds of GAEs have been observed even in the CHS plasma. One is  $n=1$  GAE generated by the transition from the  $n=1$  TAE in a co-NBI heated plasma and the other the  $n=0$  GAE in a counter-NBI heated plasma.

Figure 1 (a) shows the time evolution of the magnetic fluctuations in the CHS plasma heated with co-NBI where the magnetic axis position of the vacuum field is  $R_{ax} = 0.949$  m and the toroidal field  $B_t = 0.9$  T. The mode of the frequency  $f \sim 180$  kHz is observed from  $t \sim 30$  ms. This mode is the  $n=1$  TAE related to the  $m=2$  and  $m=3$  coupling and disappears at  $t \sim 65$  ms. Then,  $n=2$  TAE split into multiple peaks is destabilized. The fluctuations of  $f > 200$  kHz are higher harmonics of the TAE. The  $n=1$  modes appear again from  $t \sim 90$  ms in the frequency range of  $f < 90$  kHz just below the  $n=2$  TAE with multiple peaks and in the range of  $f \sim 150$  kHz. Figure 1 (c) shows the TAE gap structure calculated by a simple dispersion relation for a large-aspect-ratio tokamak equilibrium at  $t=50$  ms. As seen from Fig. 1 (c) the observed frequency of  $n=1$  mode (indicated by the horizontal line) lies near the lower bound of the  $n=1, m=2/3$  TAE gap. Figure 1 (d) shows the TAE gap structure at  $t=120$  ms. The  $n=1, m=2/3$  TAE gap does not exist at the time. As seen from this figure, the higher frequency of  $n=1$  mode observed at  $t=120$  ms lies near the minimum of the  $n=1/m=3$  Alfvén spectrum, and the lower frequency of  $n=1$  mode lies near the maximum of  $n=1/m=2$ . These  $n=1$  modes are thought to be  $n=1/m=3$  and  $n=1/m=2$  GAEs. The rotational transform increases with the increase in the plasma current. Therefore, the  $1/q=0.4$  surface of  $n=1, m=2/3$  disappears in the plasma, and TAE cannot exist there. After that, GAE gap produced by the  $n=1/m=2$  and  $n=1/m=3$  continua is expanded due to the increase in the plasma current. The GAEs would be destabilized within the gap.

Figure 2 (a) shows the time evolution of the magnetic fluctuations in the CHS plasma heated with counter-NBI, where  $R_{ax} = 0.974$  m and  $B_t = 0.9$  T. Two  $n=0$  modes are observed. The figure 2 (c) shows the TAE gap structure at  $t=100$  ms. Because  $n=0$ , no TAE gap is formed. The Figure

shows the observed higher frequency of the mode lies near the minimum value of the  $n=0/m=1$  Alfvén spectrum which is indicated as  $f_{MIN}$  in Fig. 2 (a). The value of  $f_{MIN}$  is calculated using  $n_e(0)$ , where the impurity effect is taken into account. The calculated frequency  $f_{MIN}$  agrees well with the observed frequency of the  $n=0$  mode. Therefore the mod

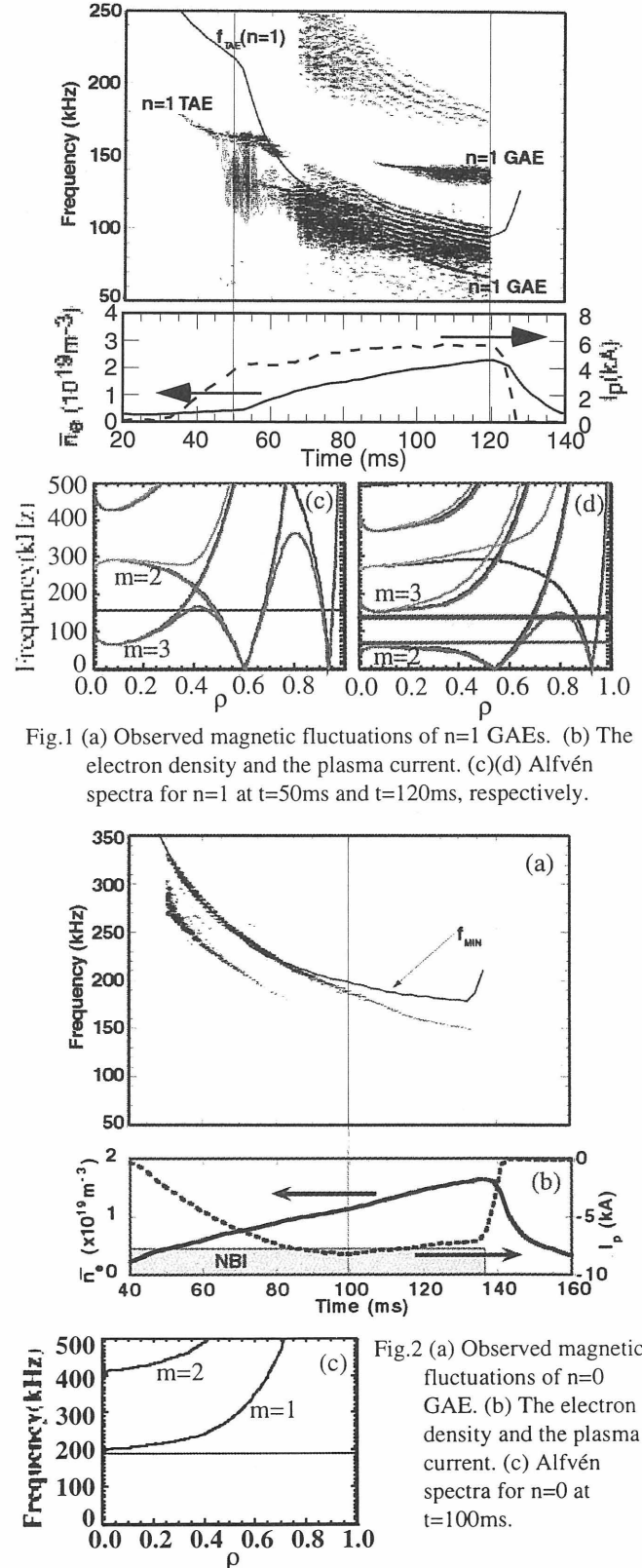


Fig.1 (a) Observed magnetic fluctuations of  $n=1$  GAEs. (b) The electron density and the plasma current. (c)(d) Alfvén spectra for  $n=1$  at  $t=50$  ms and  $t=120$  ms, respectively.

Fig.2 (a) Observed magnetic fluctuations of  $n=0$  GAE. (b) The electron density and the plasma current. (c) Alfvén spectra for  $n=0$  at  $t=100$  ms.