§19. Positive Ion and Secondary Electron Emission from Oxides by Ion Impact

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We have measured secondary electron emission (SEE) and secondary positive ion emission (SPIE) yields by ion impact on various oxides, for which there are few investigations. For uncleaned SrCeO3(5%Yb), where measurements were made in 10-6~ 10-7 Torr, we find incident ion beam current dependent SEE yields and anomalously high SPIE yields [1]. Measurements at 2.5 keV H+ and Ar+ were done in UHV. The SEE and SPIE yields appears to be insensitive to the surface conditions. The results are summarized in Table 1, where 2-3 means 2x10-3.

SEE yields  $\gamma_e$  taken at the incident ion beam current (I<sub>B</sub>) of 0.1 nA reach the constant values at the +90V bias of the electron collector. These values may not be the true SEE yields, because of the I<sub>B</sub> dependence of  $\gamma_e$ , but are regarded as the representative SEE yields. They are found to follow roughly the electronic stopping power Se (eV/nm). The ratios  $\gamma_e$ /Se of 0.012 -0.035 (nm/eV) are comparable with those for metals and other oxides such as HTC superconductors [2].

SPIE yields y+ do not much depend on the collector bias and  $I_B$ . Firstly,  $\gamma$ + is much larger than the reflection coefficient RN of TRIM calculation, confirming that SPIE is not due to the reflected ions. Secondary, y+ for H+ impact is much larger than the sputtering yields Ys of TRIM calculation with the surface binding energy of 6.1 eV (average of thermodynamically determined values for SrO and Ce<sub>2</sub>O<sub>3</sub>). Comparison of  $\gamma$ + for He+ and Ar+ impact with Ys implies high efficiency of charged fraction, if the SPIE originates from the sputtering. Thirdly, y+ does not follow the nuclear stopping power Sn. Anomalously high y+ is also observed for HTC superconductor oxides [2].

Rutherford backscattering analysis of graphite and Be which were used to collect the SPIE (neutral atoms are inevitably included) under 100 keV Ar<sup>+</sup> and 60 keV H2<sup>+</sup> show that Sr, Ce and Yb are deposited on the collector approximately in proportion to the composition. Larger amount of oxygen than expected from the composition is detected. A very small fraction of C is also detected, however, this amount of C can not explain SPIE.

IB dependence of  $\gamma_e$  and anomalously high  $\gamma$ + are under investigation.

References

[1] N. Matsunami, K. Hosaka and

H. Tawara, NIFS Report 1996.

[2] N. Matsunami, S. Majima and T. Kawamura(to be published).

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E(keV)	lon	۲e	γ+	RN	Ys	γ <sub>e</sub> /Se	γ+/Sn
						(nm/eV	()
2.5	H+	1.3	0.18	8.1-2	7.3-3	0.027	0.087
20	H+	1.5	0.13	1.6-2	1.8-3	0.012	0.15
100	H+	6	0.025	2.4-3	2.6-3	0.035	0.084
50	He+	3.5	0.05	3.4-2	1.3-2	0.018	8.8-3
100	He+	9	0.08	1.4-2	6.3-3	0.031	0.022
150	He+	10	0.11	7.0-3	2.9-3	0.030	0.039
2.5	Ar+	1	0.063	6.2-2	1.2	0.013	1.4-4
100	A r+	12	0.65	2.7-2	1.3	0.025	1.1-3
150	Ar+	12	0.5	1.9-2	1.1	0.020	8.9-4

Table	1
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