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A Compton-Vetoed Germanium Detector with Increased Sensitivity at Low Energies

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The difficulty to directly detect plutonium in spent nuclear fuel due to the high Compton background of the fission products motivates the design of a Gamma detector with improved sensitivity at low energies. We have built such a detector by operating a thin high-purity Ge detector with a large scintillator Compton veto directly *behind* it. The Ge detector is thin to absorb just the low-energy Pu radiation of interest while minimizing Compton scattering of high-energy radiation from the fission products. The subsequent scintillator is large so that forward-scattered photons from the Ge detector interact in it at least once to provide an anti-coincidence veto for the Ge detector. For highest sensitivity, additional material in the line-of-sight is minimized, the radioactive sample is kept thin, and its radiation is collimated. We will discuss the instrument design, and demonstrate the feasibility of the approach with a prototype that employs two large CsI scintillator vetoes. Initial spectra of a thin Cs-137 calibration source show a background suppression of a factor of ~ 2.5 at ~ 100 keV, limited by an unexpectedly thick 4 mm dead layer in the Ge detector.

Key words: isotope analysis, germanium detectors, Compton veto, background suppression, spent fuel analysis, plutonium detection

INTRODUCTION

Gamma spectroscopy with high-purity germanium detectors can be limited in sensitivity at low energies by the Compton background from high-energy radiation. This prevents, for example, the direct detection of plutonium in spent nuclear fuel due to the high-intensity Compton background from the fission products. This problem is of significant concern in nuclear safeguards, since it limits the accuracy of the Pu assay and thus the accounting for tons of fissile material worldwide [1].

The sensitivity at low energies can be greatly increased with a large Compton veto located *behind* the primary Ge detector, which selectively rejects low-energy background events in the Ge by anti-coincidence vetoing [2, 3]. We have designed such a composite detector that is optimized for background suppression at low energies (≤ 200 keV) where the characteristic actinide lines of interest lie. Here we discuss the design considerations for increased sensitivity at low energies, and demonstrate the feasibility of the approach with a prototype instrument that employs a dual CsI scintillator veto.

INSTRUMENT DESIGN

The instrument design exploits the idea that the low-energy Compton background in Ge detectors is caused by *forward*-scattering of high-energy gamma-rays. A scintillator anti-coincidence veto should therefore be placed *behind* the primary Ge detector, rather than around it as is typical for general-purpose Compton shields [4, 5]. The Ge detector itself should be just thick enough to absorb the low-energy Pu radiation of interest without providing unnecessary extra mass for Compton scattering. Additional material in the line of sight between the source and the detectors should be minimized to reduce scattering to low energies in material other than

the Ge detector, which cannot be vetoed. This includes the windows of the cryostat, the support structure of the Ge crystal, the case of the scintillator and the radioactive source itself. Scattering can be further reduced when the collimator between the source and the Ge detector is conical so that its inside is not illuminated directly. Additional shielding can be placed around the source to suppress backscattering from the laboratory environment, as long as the shield is open in the back to avoid direct Compton scattering from the shield into the detector (Figure 1).

Geant4 Monte-Carlo simulations illustrate the background suppression that can be achieved with a 10 cm (diameter) \times 14 cm (length) scintillator behind an 8 mm thick Ge detector, assuming idealized conditions of zero additional mass in the line-of-sight between them and the source (Figure 2). The high-energy background, which is caused by large-angle scattering that bypasses the veto, is largely unaffected. The background starts to decrease at energies below \sim 350 keV, which corresponds to a scattering angle of \sim 90 degrees where the scattered radiation starts to pass through the veto. The low-energy background can be suppressed by several orders of magnitude, depending on the size and the absorption efficiency of the scintillator material. The suppression is largest for a scattering angle towards the corner of the cylindrical veto where the path through the scintillator is longest and the probability of interaction is therefore highest (figure 1, red arrow).

In this idealized case, the background suppression can be estimated analytically from the ratio of interacting to transmitted photons in the scintillator according to

$$\frac{B_{NoVeto}}{B_{Veto}} = \frac{1 - e^{-\mu(E)\rho L(E)}}{e^{-\mu(E)\rho L(E)}}. \quad (1)$$

Here $\mu(E)$ and ρ are absorption coefficient and density of the scintillator material, and $L(E)$ is the length of the scintillator in direction of the scattering angle for Compton photons of energy E .

The analytical estimate according to equation (1) is shown in grey in figure 2 for the case of a NaI veto at a distance of 5 and 7 mm from the Ge detector.

For more realistic cases with a finite amount of intervening material between the source and the detectors, the background suppression is reduced by the fraction of photons that interact in the intervening material. This adds a term $\Sigma(1-e^{-\mu(E)\rho T})$ to the denominator of equation (1), where $\mu_i(E)$, ρ_i and T_i are the absorption coefficient, density and thickness of the intervening materials, and the summation extends over all materials in the line of sight.

The degree to which the idealized case can be implemented in practice depends on how much the amount of intervening material in the line-of-sight can be reduced. We have built a prototype instrument to examine this question experimentally (Figure 3). It uses an 8 mm thick planar Ge detector with a diameter of 35 mm that is read out with a custom-designed preamplifier at room temperature. The Ge crystal is held and heat sunk to the LN_2 cold stage by three plastic clamps at its edges that are not illuminated directly by the source. Two 0.025 mm Al windows in front and behind the Ge detector reduce IR heating to the detector cold stage, and the vacuum barrier in the line-of-sight consists of two 0.75 mm Al windows with 40 mm diameter. The source is collimated onto the central part of the Ge detector by a 10 cm thick 10 cm diameter tungsten collimator (Hevimet® MT-185) that has a conical aperture with a 12.5 mm entrance and a 25 mm exit opening. Sources that are placed at least 10 cm away from the entrance opening within the cone of the collimator will therefore not illuminate its inside directly (cf. Figure 1). The primary veto consists of a 3.5" \times 6" CsI scintillator in a 0.8 mm Al housing. It is surrounded by a secondary annular CsI veto with an outer diameter of 5.5". All detectors are biased and read out with standard commercial electronics.

RESULTS AND DISCUSSION

We have tested the response of the instrument with a 5 μCi Cs-137 calibration source supported between two clear plastic sheets. The source is so dilute to be invisible to the naked eye and can be considered a point source with negligible in-source scattering. We have also measured the laboratory background radiation, which is not negligible in this setup, since the source is weak and held at a distance of ~ 25 cm from the Ge detector to avoid illuminating the inside of the W collimator. Figure 4 shows the background-corrected Cs-137 gamma spectra without the vetoes (dark green), and with both CsI vetoes installed and active (light green). It also shows Geant4 Monte-Carlo simulations of our geometry for the two cases, redistributed with a FWHM of 2.5 keV to account for the finite energy resolution of the Ge detector (dark and light grey). All spectra are normalized to 10^5 counts in the photopeak at 662 keV (which causes the noise in the simulations to be unrealistically low) to allow a comparison of the Compton background level.

As expected, the spectra show the broad background with the Compton edge at 470 eV and a small backscatter peak at ~ 190 eV. The background is suppressed preferentially at low energies caused by forward-scattered Compton photons. The absence of W X-rays at ~ 60 keV shows that the collimator works as expected without contributing any X-rays to the background. Still, the two vetoes suppress the background in the region of interest around 100 keV only by a factor of ~ 2.5 , significantly less than the expected factor of ~ 15 for this geometry based on the Monte-Carlo simulations. This limitation turns out to be due to an unexpectedly large dead layer at the back side electrode of the Ge detector. The thickness of such a dead layer can easily be measured in our geometry by comparing spectra for front and back-side illumination. Figure 5 shows two spectra from an Am-241 / Eu-152 calibration source and the measured line ratio at

low energies. The observed line ratios (squares) are explained well by the absorption of a 4 mm dead layer (grey line), implying that half of the 8 mm Ge detector does not contribute to any signal, while fully contributing to the Compton background that cannot be vetoed. Including a uniform 4 mm Ge dead layer in the Monte-Carlo simulations qualitatively accounts for much of the discrepancy with the data, although it quantitatively differs significantly in the fine structure that it predicts in the Compton background. This implies that the dead layer inside the Ge detector is either not uniform, or that there are additional yet unidentified scattering sources in our setup. In future work, the Ge detector will therefore be replaced by a crystal with an amorphous Ge electrode whose dead layer should be only a few μm thick. In addition, the CsI vetoes will be replaced by BGO vetoes to further improve the background suppression at low energies, and the liquid N_2 can be replaced by a mechanical cooler for improved deployability. If a suppression factor above ~ 20 at ~ 100 keV can be achieved, it would enable the direct detection of Pu lines for realistic spent nuclear fuel compositions.

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LIST OF FIGURES

Figure 1: Schematic cross section of the instrument

Figure 2: Monte-Carlo simulations of the Compton suppression that can be attained with different 10 cm (diameter) \times 14 cm (length) scintillators under idealized conditions. The grey lines are analytical approximations to the Compton suppression for a NaI scintillator veto 5 and 7 mm behind the Ge detector.

Figure 3: Picture of the prototype with the conical W collimator, the thin Ge detector, and the dual CsI veto directly behind it.

Figure 4: Background-corrected spectra from a Cs-137 source (green), compared to Geant4 Monte Carlo simulations without (grey), both with and without the two CsI vetoes.

Figure 5: Eu-152 / Am-241 spectra for front (dark) and back side (light green) illumination of the Ge detector. The observed line ratios (squares) correspond to a 4 mm dead layer on the back side electrode of the Ge detector (black).

FIGURE 1

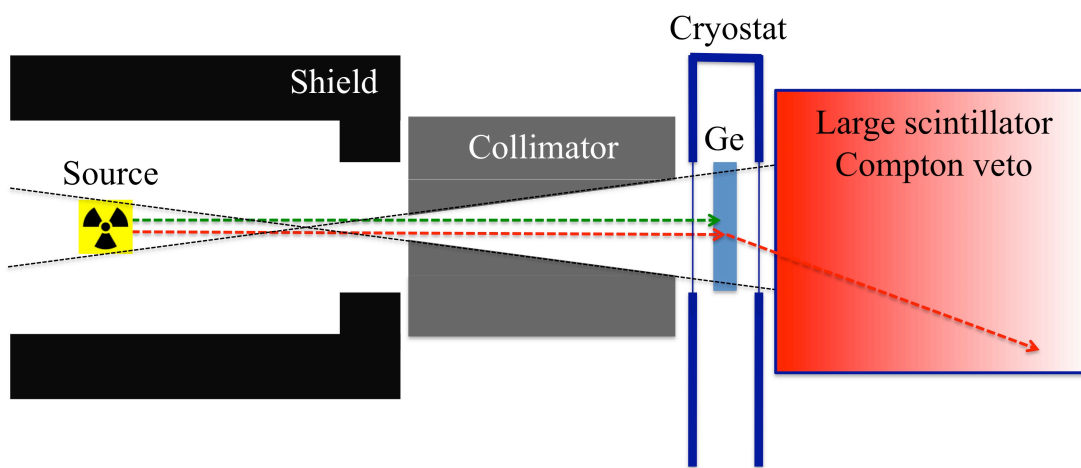


FIGURE 2

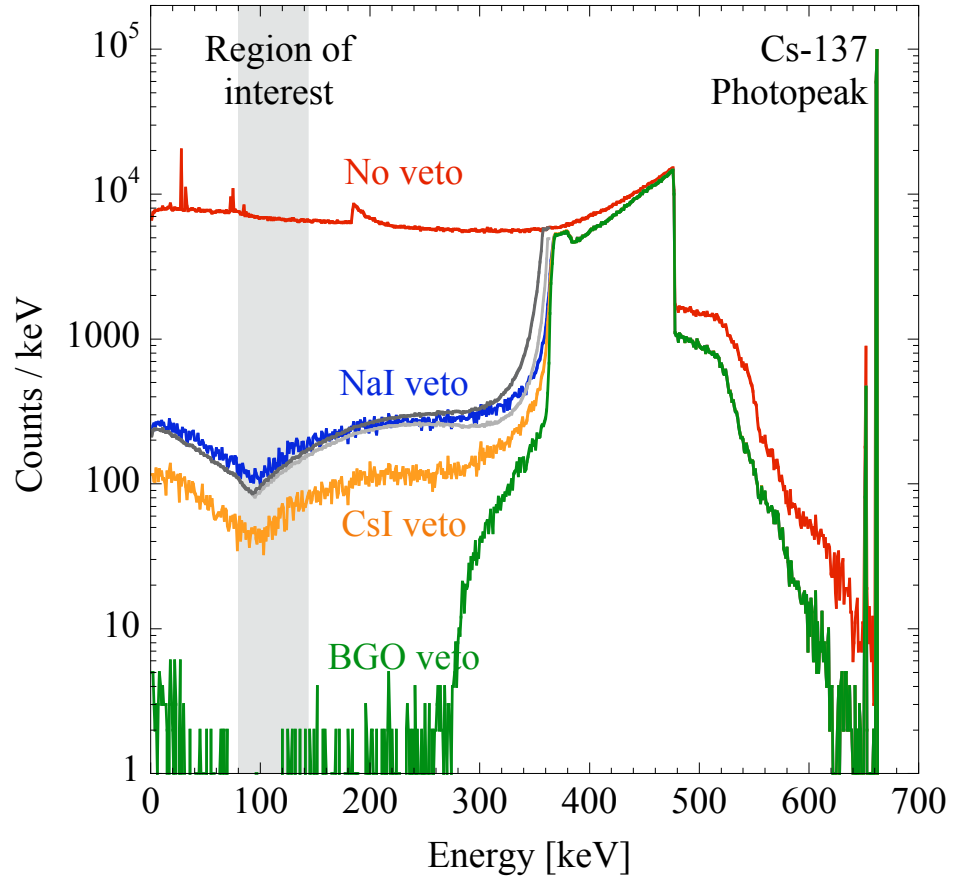


FIGURE 3

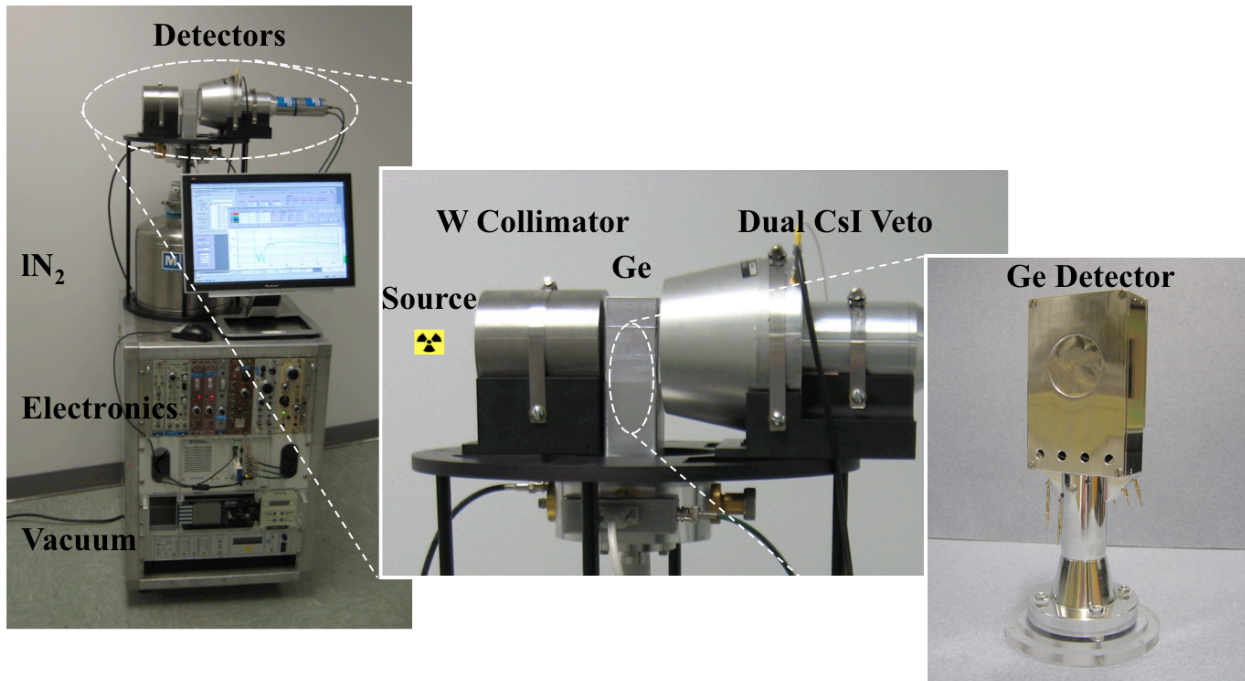


FIGURE 4

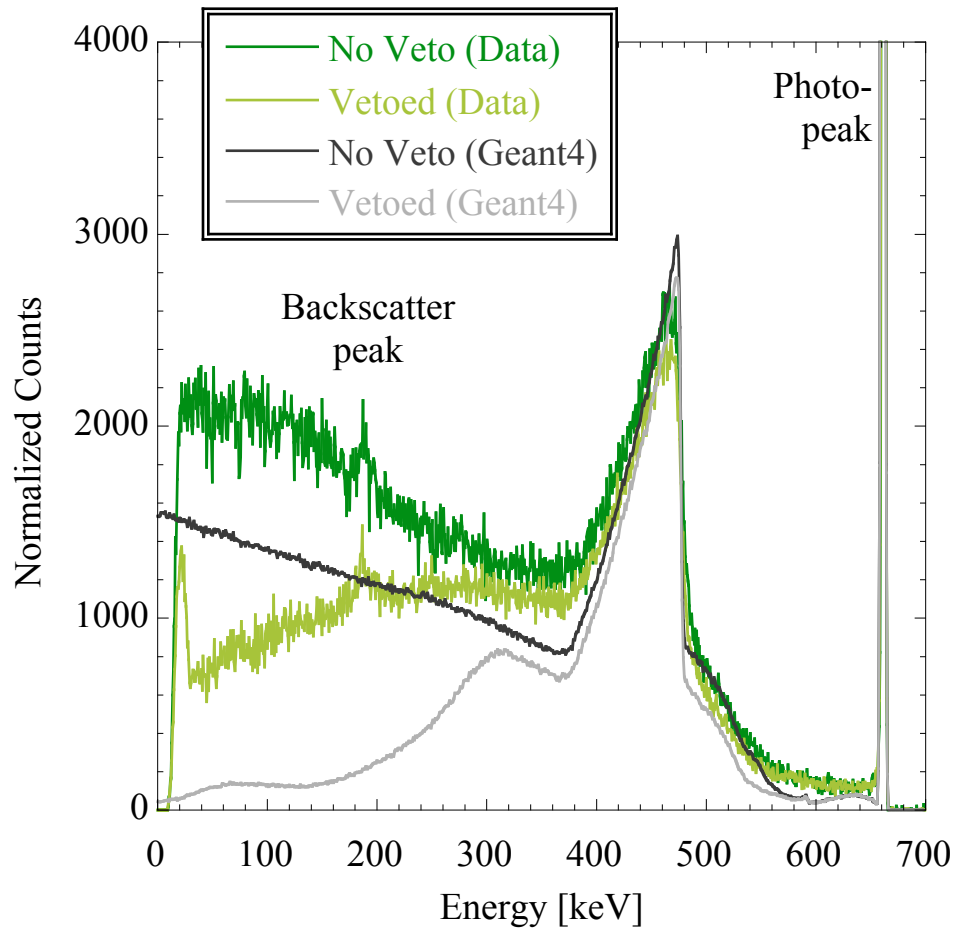


FIGURE 5

