

Fermi National Accelerator Laboratory

FERMILAB-Conf-95/187-E

D0

Search for Light Top Squarks with the D0 Detector

S. Abachi et al.

The D0 Collaboration

*Fermi National Accelerator Laboratory
P.O. Box 500, Batavia, Illinois 60510*

July 1995

Submitted to the *International Europhysics Conference on High Energy Physics (HEP95)*,
Brussels, Belgium, July 27-August 2, 1995

Disclaimer

This report was prepared as an account of work sponsored by an agency of the United States Government. Neither the United States Government nor any agency thereof, nor any of their employees, makes any warranty, expressed or implied, or assumes any legal liability or responsibility for the accuracy, completeness, or usefulness of any information, apparatus, product, or process disclosed, or represents that its use would not infringe privately owned rights. Reference herein to any specific commercial product, process, or service by trade name, trademark, manufacturer, or otherwise, does not necessarily constitute or imply its endorsement, recommendation, or favoring by the United States Government or any agency thereof. The views and opinions of authors expressed herein do not necessarily state or reflect those of the United States Government or any agency thereof.

Search for Light Top Squarks with the DØ Detector

The DØ Collaboration¹
(July 1995)

ABSTRACT

We present preliminary results of a search for pair produced top squarks, the supersymmetric partners to the top quark, under the assumption that the two-body decay $\tilde{t}_1 \rightarrow b\tilde{W}_1$ and three-body decay $\tilde{t}_1 \rightarrow bW\tilde{Z}_1$ are kinematically forbidden. This scenario is realized if \tilde{t}_1 is considerably lighter than all other squarks as well as \tilde{W}_1 , all $\tilde{\nu}_s$, and the top quark. Such light top squarks will dominantly decay via $\tilde{t}_1 \rightarrow c\tilde{Z}_1$ yielding a signal of two acollinear jets with \cancel{E}_T . The data reported here were taken during the 1992-1993 run of the DØ detector at the Fermilab Tevatron $p\bar{p}$ collider operating at $\sqrt{s} = 1.8$ TeV.

S. Abachi,¹² B. Abbott,³⁴ M. Abolins,²³ B.S. Acharya,⁴¹ I. Adam,¹⁰ D.L. Adams,³⁵ M. Adams,¹⁵
S. Ahn,¹² H. Aihara,²⁰ J. Alitti,³⁷ G. Álvarez,¹⁶ G.A. Alves,⁸ E. Amidi,²⁷ N. Amos,²²
E.W. Anderson,¹⁷ S.H. Aronson,³ R. Astur,³⁹ R.E. Avery,²⁹ A. Baden,²¹ V. Balamurali,³⁰
J. Balderston,¹⁴ B. Baldin,¹² J. Bantly,⁴ J.F. Bartlett,¹² K. Bazizi,⁷ J. Bendich,²⁰ S.B. Beri,³²
I. Bertram,³⁵ V.A. Bezzubov,³³ P.C. Bhat,¹² V. Bhatnagar,³² M. Bhattacharjee,¹¹ A. Bischoff,⁷
N. Biswas,³⁰ G. Blazey,¹² S. Blessing,¹³ P. Bloom,⁵ A. Boehnlein,¹² N.I. Bojko,³³
F. Borchering,¹² J. Borders,³⁶ C. Boswell,⁷ A. Brandt,¹² R. Brock,²³ A. Bross,¹² D. Buchholz,²⁹
V.S. Burtovoi,³³ J.M. Butler,¹² D. Casey,³⁶ H. Castilla-Valdez,⁹ D. Chakraborty,³⁹
S.-M. Chang,²⁷ S.V. Chekulaev,³³ L.-P. Chen,²⁰ W. Chen,³⁹ L. Chevalier,³⁷ S. Chopra,³²
B.C. Choudhary,⁷ J.H. Christenson,¹² M. Chung,¹⁵ D. Claes,³⁹ A.R. Clark,²⁰ W.G. Cobau,²¹
J. Cochran,⁷ W.E. Cooper,¹² C. Cretsinger,³⁶ D. Cullen-Vidal,⁴ M.A.C. Cummings,¹⁴ D. Cutts,⁴
O.I. Dahl,²⁰ K. De,⁴² M. Demarteau,¹² R. Demina,²⁷ K. Denisenko,¹² N. Denisenko,¹²
D. Denisov,¹² S.P. Denisov,³³ W. Dharmaratna,¹³ H.T. Diehl,¹² M. Diesburg,¹² G. Di Loreto,²³
R. Dixon,¹² P. Draper,⁴² J. Drinkard,⁶ Y. Ducros,³⁷ S.R. Dugad,⁴¹ S. Durston-Johnson,³⁶
D. Edmunds,²³ J. Ellison,⁷ V.D. Elvira,^{12,†} R. Engelmann,³⁹ S. Eno,²¹ G. Eppley,³⁵
P. Ermolov,²⁴ O.V. Eroshin,³³ V.N. Evdokimov,³³ S. Fahey,²³ T. Fahland,⁴ M. Fatyga,³
M.K. Fatyga,³⁶ J. Featherly,³ S. Feher,³⁹ D. Fein,² T. Ferbel,³⁶ G. Finocchiaro,³⁹ H.E. Fisk,¹²
Yu. Fisyak,²⁴ E. Flattum,²³ G.E. Forden,² M. Fortner,²⁸ K.C. Frame,²³ P. Franzini,¹⁰ S. Fuess,¹²
A.N. Galjaev,³³ E. Gallas,⁴² C.S. Gao,^{12,*} S. Gao,^{12,*} T.L. Geld,²³ R.J. Genik II,²³ K. Genser,¹²
C.E. Gerber,^{12,§} B. Gibbard,³ V. Glebov,³⁶ S. Glenn,⁵ B. Gobbi,²⁹ M. Goforth,¹³
A. Goldschmidt,²⁰ B. Gómez,¹ P.I. Goncharov,³³ H. Gordon,³ L.T. Goss,⁴³ N. Graf,³
P.D. Grannis,³⁹ D.R. Green,¹² J. Green,²⁸ H. Greenlee,¹² G. Griffin,⁶ N. Grossman,¹²
P. Grudberg,²⁰ S. Grünendahl,³⁶ W. Gu,^{12,*} G. Guglielmo,³¹ J.A. Guida,³⁹ J.M. Guida,³
W. Gurny,³ S.N. Gurzhiev,³³ P. Gutierrez,³¹ Y.E. Gutnikov,³³ N.J. Hadley,²¹ H. Haggerty,¹²
S. Hagopian,¹³ V. Hagopian,¹³ K.S. Hahn,³⁶ R.E. Hall,⁶ S. Hansen,¹² R. Hatcher,²³
J.M. Hauptman,¹⁷ D. Hedin,²⁸ A.P. Heinson,⁷ U. Heintz,¹² R. Hernández-Montoya,⁹
T. Heuring,¹³ R. Hirosky,¹³ J.D. Hobbs,¹² B. Hoeneisen,^{1,¶} J.S. Hoftun,⁴ F. Hsieh,²² Ting Hu,³⁹
Tong Hu,¹⁶ T. Huehn,⁷ S. Igarashi,¹² A.S. Ito,¹² E. James,² J. Jaques,³⁰ S.A. Jerger,²³
J.Z.-Y. Jiang,³⁹ T. Joffe-Minor,²⁹ H. Johari,²⁷ K. Johns,² M. Johnson,¹² H. Johnstad,⁴⁰
A. Jonckheere,¹² M. Jones,¹⁴ H. Jöstlein,¹² S.Y. Jun,²⁹ C.K. Jung,³⁹ S. Kahn,³ G. Kalbfleisch,³¹

¹Submitted to the XVII International Symposium on Lepton-Photon Interactions (LP 95), Beijing, China, August 10-15, 1995.

J.S. Kang,¹⁸ R. Kehoe,³⁰ M.L. Kelly,³⁰ A. Kernan,⁷ L. Kerth,²⁰ C.L. Kim,¹⁸ S.K. Kim,³⁸
 A. Klatchko,¹³ B. Klima,¹² B.I. Klochov,³³ C. Klopfenstein,³⁹ V.I. Klyukhin,³³
 V.I. Kochetkov,³³ J.M. Kohli,³² D. Koltick,³⁴ A.V. Kostritskiy,³³ J. Kotcher,³ J. Kourlas,²⁶
 A.V. Kozelov,³³ E.A. Kozlovski,³³ M.R. Krishnaswamy,⁴¹ S. Krzywdzinski,¹² S. Kunori,²¹
 S. Lami,³⁹ G. Landsberg,¹² R.E. Lanou,⁴ J-F. Lebrat,³⁷ A. Leflat,²⁴ H. Li,³⁹ J. Li,⁴² Y.K. Li,²⁹
 Q.Z. Li-Demartean,¹² J.G.R. Lima,⁸ D. Lincoln,²² S.L. Linn,¹³ J. Linnemann,²³ R. Lipton,¹²
 Y.C. Liu,²⁹ F. Lobkowicz,³⁶ S.C. Loken,²⁰ S. Lökös,³⁹ L. Lueking,¹² A.L. Lyon,²¹
 A.K.A. Maciel,⁸ R.J. Madaras,²⁰ R. Madden,¹³ I.V. Mandrichenko,³³ Ph. Mangeot,³⁷ S. Mani,⁵
 B. Mansoulié,³⁷ H.S. Mao,^{12,*} S. Margulies,¹⁵ R. Markeloff,²⁸ L. Markosky,² T. Marshall,¹⁶
 M.I. Martin,¹² M. Marx,³⁹ B. May,²⁹ A.A. Mayorov,³³ R. McCarthy,³⁹ T. McKibben,¹⁵
 J. McKinley,²³ T. McMahon,³¹ H.L. Melanson,¹² J.R.T. de Mello Neto,⁸ K.W. Merritt,¹²
 H. Miettinen,³⁵ A. Milder,² A. Mincer,²⁶ J.M. de Miranda,⁸ C.S. Mishra,¹²
 M. Mohammadi-Baarmand,³⁹ N. Mokhov,¹² N.K. Mondal,⁴¹ H.E. Montgomery,¹² P. Mooney,¹
 M. Mudan,²⁶ C. Murphy,¹⁶ C.T. Murphy,¹² F. Nang,⁴ M. Narain,¹² V.S. Narasimham,⁴¹
 A. Narayanan,² H.A. Neal,²² J.P. Negret,¹ E. Neis,²² P. Nemethy,²⁶ D. Nešić,⁴ D. Norman,⁴³
 L. Oesch,²² V. Oguri,⁸ E. Oltman,²⁰ N. Oshima,¹² D. Owen,²³ P. Padley,³⁵ M. Pang,¹⁷ A. Para,¹²
 C.H. Park,¹² Y.M. Park,¹⁹ R. Partridge,⁴ N. Parua,⁴¹ M. Paterno,³⁶ J. Perkins,⁴² A. Peryshkin,¹²
 M. Peters,¹⁴ H. Piekarz,¹³ Y. Pischalnikov,³⁴ A. Pluquet,³⁷ V.M. Podstavkov,³³ B.G. Pope,²³
 H.B. Prosper,¹³ S. Protopopescu,³ D. Pušeljčić,²⁰ J. Qian,²² P.Z. Quintas,¹² R. Raja,¹²
 S. Rajagopalan,³⁹ O. Ramirez,¹⁵ M.V.S. Rao,⁴¹ P.A. Rapidis,¹² L. Rasmussen,³⁹ A.L. Read,¹²
 S. Reucroft,²⁷ M. Rijssenbeek,³⁹ T. Rockwell,²³ N.A. Roe,²⁰ P. Rubinov,³⁹ R. Ruchti,³⁰
 S. Rusin,²⁴ J. Rutherford,² A. Santoro,⁸ L. Sawyer,⁴² R.D. Schamberger,³⁹ H. Schellman,²⁹
 J. Sculli,²⁶ E. Shabalina,²⁴ C. Shaffer,¹³ H.C. Shankar,⁴¹ R.K. Shivpuri,¹¹ M. Shupe,²
 J.B. Singh,³² V. Sirotenko,²⁸ W. Smart,¹² A. Smith,² R.P. Smith,¹² R. Snihur,²⁹ G.R. Snow,²⁵
 S. Snyder,³⁹ J. Solomon,¹⁵ P.M. Sood,³² M. Sosebee,⁴² M. Souza,⁸ A.L. Spadafora,²⁰
 R.W. Stephens,⁴² M.L. Stevenson,²⁰ D. Stewart,²² D.A. Stoianova,³³ D. Stoker,⁶ K. Streets,²⁶
 M. Strovink,²⁰ A. Taketani,¹² P. Tamburello,²¹ J. Tarazi,⁶ M. Tartaglia,¹² T.L. Taylor,²⁹
 J. Teiger,³⁷ J. Thompson,²¹ T.G. Trippe,²⁰ P.M. Tuts,¹⁰ N. Varelas,²³ E.W. Varnes,²⁰
 P.R.G. Virador,²⁰ D. Vititoe,² A.A. Volkov,³³ A.P. Vorobiev,³³ H.D. Wahl,¹³ G. Wang,¹³
 J. Wang,^{12,*} L.Z. Wang,^{12,*} J. Warchol,³⁰ M. Wayne,³⁰ H. Weerts,²³ F. Wen,¹³ W.A. Wenzel,²⁰
 A. White,⁴² J.T. White,⁴³ J.A. Wightman,¹⁷ J. Wilcox,²⁷ S. Willis,²⁸ S.J. Wimpenny,⁷
 J.V.D. Wirjawan,⁴³ J. Womersley,¹² E. Won,³⁶ D.R. Wood,¹² H. Xu,⁴ R. Yamada,¹² P. Yamin,³
 C. Yanagisawa,³⁹ J. Yang,²⁶ T. Yasuda,²⁷ C. Yoshikawa,¹⁴ S. Youssef,¹³ J. Yu,³⁶ Y. Yu,³⁸
 Y. Zhang,^{12,*} Y.H. Zhou,^{12,*} Q. Zhu,²⁶ Y.S. Zhu,^{12,*} Z.H. Zhu,³⁶ D. Zieminska,¹⁶ A. Zieminski,¹⁶
 and A. Zylberstejn³⁷

¹Universidad de los Andes, Bogotá, Colombia

²University of Arizona, Tucson, Arizona 85721

³Brookhaven National Laboratory, Upton, New York 11973

⁴Brown University, Providence, Rhode Island 02912

⁵University of California, Davis, California 95616

⁶University of California, Irvine, California 92717

⁷University of California, Riverside, California 92521

⁸LAFEX, Centro Brasileiro de Pesquisas Físicas, Rio de Janeiro, Brazil

⁹CINVESTAV, Mexico City, Mexico

¹⁰Columbia University, New York, New York 10027

¹¹Delhi University, Delhi, India 110007

¹²Fermi National Accelerator Laboratory, Batavia, Illinois 60510

¹³Florida State University, Tallahassee, Florida 32306

¹⁴University of Hawaii, Honolulu, Hawaii 96822

¹⁵University of Illinois at Chicago, Chicago, Illinois 60607

- ¹⁶Indiana University, Bloomington, Indiana 47405
¹⁷Iowa State University, Ames, Iowa 50011
¹⁸Korea University, Seoul, Korea
¹⁹Kyungshung University, Pusan, Korea
²⁰Lawrence Berkeley Laboratory and University of California, Berkeley, California 94720
²¹University of Maryland, College Park, Maryland 20742
²²University of Michigan, Ann Arbor, Michigan 48109
²³Michigan State University, East Lansing, Michigan 48824
²⁴Moscow State University, Moscow, Russia
²⁵University of Nebraska, Lincoln, Nebraska 68588
²⁶New York University, New York, New York 10003
²⁷Northeastern University, Boston, Massachusetts 02115
²⁸Northern Illinois University, DeKalb, Illinois 60115
²⁹Northwestern University, Evanston, Illinois 60208
³⁰University of Notre Dame, Notre Dame, Indiana 46556
³¹University of Oklahoma, Norman, Oklahoma 73019
³²University of Panjab, Chandigarh 16-00-14, India
³³Institute for High Energy Physics, 142-284 Protvino, Russia
³⁴Purdue University, West Lafayette, Indiana 47907
³⁵Rice University, Houston, Texas 77251
³⁶University of Rochester, Rochester, New York 14627
³⁷CEA, DAPNIA/Service de Physique des Particules, CE-SACLAY, France
³⁸Seoul National University, Seoul, Korea
³⁹State University of New York, Stony Brook, New York 11794
⁴⁰SSC Laboratory, Dallas, Texas 75237
⁴¹Tata Institute of Fundamental Research, Colaba, Bombay 400005, India
⁴²University of Texas, Arlington, Texas 76019
⁴³Texas A&M University, College Station, Texas 77843

INTRODUCTION

A spacetime symmetry, supersymmetry (SUSY) links bosons to fermions, introducing supersymmetric partners (sparticles) to all the SM particles. In combination with Grand Unification Theories, SUSY results in models which successfully unify the U(1), SU(2) and SU(3) couplings at 10^{16} GeV yet remain consistent with the experimental proton lifetime limits. Together with the natural solution offered to the fine-tuning problem, SUSY is an attractive extension to the Standard Model (SM).

Early limits on squark (the SUSY partner of the quark) masses have been set under the assumption that all squarks have the same mass. This was justified by a model that argued all scalar particles share a common mass at the energy scale where SUSY is broken. The degeneracy among squarks is broken only slightly by the small differences in electroweak interactions between left and right states and the different Yukawa interactions of the various families. However, a heavy top quark (1) means *its* Yukawa interactions become substantial and can drive top squark masses lower than all other squarks. Mass-splitting by left/right mixing may split the mass eigenstates even further, making one state, \tilde{t}_1 , the lightest of all (2).

If kinematically accessible, the top squark is expected to decay through $\tilde{t}_1 \rightarrow b\tilde{W}_1$. If, however, $m_{\tilde{W}_1} > m_{\tilde{t}_1} + m_b$, the chargino becomes virtual and three-body decays $\tilde{t}_1 \rightarrow b\tilde{l}\tilde{\nu}$ or $\tilde{t}_1 \rightarrow b\tilde{\nu}\tilde{l}$ become accessible *unless* sleptons and sneutrinos are *also* heavier than \tilde{t}_1 . Under this additional assumption, top squarks will dominantly decay via $\tilde{t}_1 \rightarrow c\tilde{Z}_1$ yielding an

event signature of two acollinear jets with \cancel{E}_T (3). The major backgrounds expected for this signal are vector boson production associated with jets and Standard Model multijet events with faked \cancel{E}_T signals.

The top squark production occurs via gluon fusion and $q\bar{q}$ annihilation (4) and is thus fixed by QCD in terms of $m_{\tilde{t}_1}$. Its decay topology is determined by $m_{\tilde{Z}_1}$, the mass of the lightest neutralino. Thus the search is through a two-parameter phase space in $m_{\tilde{Z}_1}$ vs $m_{\tilde{t}_1}$. The region to be explored is the lower half-plane defined by the near-diagonal $m_{\tilde{t}_1} < m_{\tilde{Z}_1} + m_c$, hemmed in by $m_{\tilde{t}_1} > m_{\tilde{Z}_1} + m_b + m_W$ (to restrict competing 3-body decays).

A search was conducted for the SUSY partner of top quark, \tilde{t} , under the framework of a supergravity-GUT inspired Minimal Supersymmetric Standard Model (MSSM) (5). The model enforces conservation of R -parity which implies sparticles be produced in pairs, and that the Lightest Supersymmetric Particle (LSP) must be absolutely stable (6).

The results reported here on the \tilde{t}_1 search are preliminary.

THE DATA SET

Data corresponding to a total integrated luminosity of $13.5 \pm 0.7 \text{ pb}^{-1}$ were collected by the DØ detector during its 1992-1993 run. DØ is a general purpose detector consisting of a central tracking system and nearly hermetic liquid argon calorimeter surrounded by a toroidal muon spectrometer. A detailed description can be found elsewhere (7). Events for this analysis were collected under a missing E_T (\cancel{E}_T) trigger requiring 30 GeV \cancel{E}_T in hardware and 35 GeV in software.

The Single Interaction Cut

To ensure unambiguous \cancel{E}_T assignments, we demanded events be identified as having only one primary vertex. An algorithm that combined timing information from a set of trigger counters together with reconstructed scalar E_T and the number of vertices found by tracking was used to select single interaction events (8). Its use effectively reduced our data set to a *single interaction equivalent luminosity* of $7.4 \pm 0.4 \text{ pb}^{-1}$.

Angular Correlation Cuts

Badly mismeasured jets can produce false \cancel{E}_T , but such events usually show a correlation between the jet and \cancel{E}_T directions. If a jet is identified as the leading object in an event by an overestimate of its energy, a false \cancel{E}_T signal will be induced in a direction opposite to that of the jet. Jets with underestimated energy will tend to be aligned with the fake \cancel{E}_T . This is cited as the source of the clumping of events visible in the upper left-hand corner of Fig. 1, a feature not reproduced by either the SUSY signal or any of the expected physics backgrounds. This effect is observed not only in events from our \cancel{E}_T triggers, but even those from a low E_T threshold single jet trigger (used to study detector induced background) when mild \cancel{E}_T cuts are applied offline. A one-dimensional cut against high $\Delta\phi(\cancel{E}_T, jet1)$ values was applied to suppress these events. Additional cuts were made against jets aligned (within $\delta\phi = 0.1$) with the direction of the \cancel{E}_T .

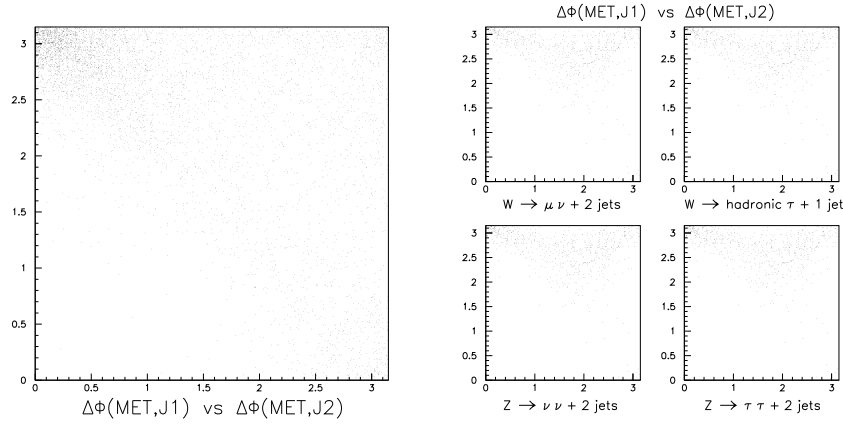


FIG. 1. The opening angle between \cancel{E}_T and the leading jet runs up the ordinate, its angle with the next leading jet along the abscissa. At left are events from a low E_T threshold single jet trigger, after an offline $\cancel{E}_T > 15$ GeV cut has been applied. At right are some Monte Carlo of vector boson backgrounds.

SIGNAL

Signal events were generated using ISAJET 7.13 (9), which incorporates the latest implementation of ISASUSY (10). These files were processed through a GEANT simulation of the DØ detector (11) and reconstructed. Kinematic distributions sampled from the middle of the parameter space to be probed (Figure 2, left) show that the leading jets are fairly high in E_T , and the \cancel{E}_T is large. The latter claim, however, does not hold over the full range of the region to be explored (Figure 2, right). Although it means sacrificing the rejection power of a high \cancel{E}_T cut, a relatively low cut will allow coverage of more of the parameter space. Trigger thresholds fixed this cut at 40 GeV.

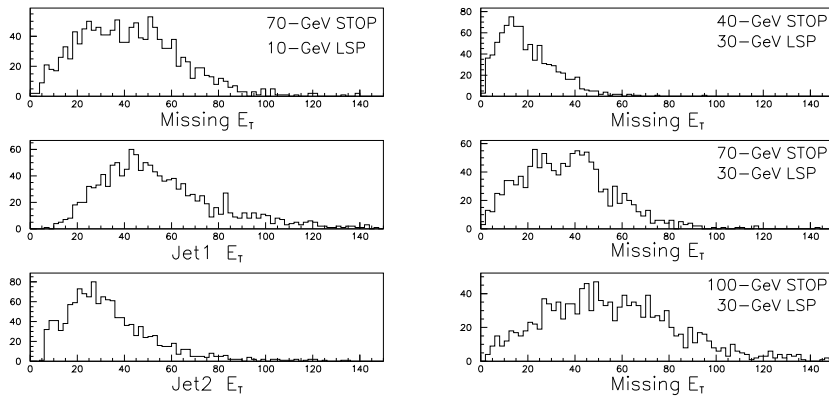


FIG. 2. The jet E_T 's are seen to be healthy (left) in this sample of $m_{\tilde{l}_1} = 70$ GeV/ c^2 , $m_{Z_1} = 10$ GeV/ c^2 ISAJET events. The \cancel{E}_T is respectable, although scanning the region to be searched (right) shows \cancel{E}_T to be rather feeble for low $m_{\tilde{l}_1}$.

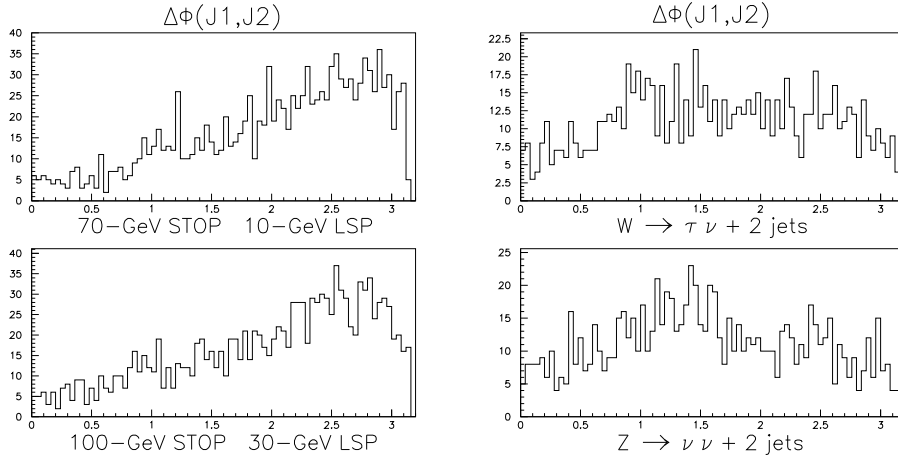


FIG. 3. The opening angle between the two leading jets tends toward higher values in the signal, though most of our background sources show relatively flat distributions.

FINAL SELECTION CUTS

Leptons are not primarily part of our signal, but appear only incidentally insofar as charm jets are in the final state. These tend to be low E_T objects. We gain some rejection against the vector boson background by discriminating against events with high- E_T electrons or muons.

The presence of two LSP's in the event suggests that the two leading jets in our signal be acollinear. But distributions of the opening angle between them (Fig.3) show that this angle tends to large values. A cut of $\Delta\phi(j_1, j_2) > 90^\circ$ preserves 70-75% of the signal. Background distributions tend to be flat.

Standard Model multijet events tend to be mostly back-to-back jet pairs. Thus we must also cut against two leading jets with an opening angle close to π . Monte Carlo distributions suggest an effective cut can be made at $\Delta\phi(j_1, j_2) < 165^\circ$.

Our final selection cuts of

$$\begin{aligned}
 \cancel{E}_T &> 40 \text{ GeV} \\
 E_T^{jet2} &> 30 \text{ GeV} \\
 90^\circ &< \Delta\phi(j_1, j_2) < 165^\circ \\
 10^\circ &< \Delta\phi(j_1, \cancel{E}_T) < 125^\circ \\
 10^\circ &< \Delta\phi(j_{3,4}, \cancel{E}_T)
 \end{aligned}$$

with a *VETO* against events with:

$$\begin{aligned}
 E_T^\mu &> 10 \text{ GeV} \\
 E_T^{el} &> 10 \text{ GeV}
 \end{aligned}$$

leave a total of two candidates in the single interaction Missing E_T triggers.

TABLE 1. Expected Vector Boson backgrounds to the \tilde{t} signal.

Background Process	Expected # of events
$W \rightarrow e\bar{\nu}$	0.52 ± 0.30
$W \rightarrow \mu\bar{\nu}$	0.84 ± 0.42
$W \rightarrow \tau\bar{\nu}$	0.99 ± 0.64
$Z \rightarrow \nu\bar{\nu}$	0.37 ± 0.32
$Z \rightarrow \mu\bar{\mu}$	0.06 ± 0.05
$Z \rightarrow \tau\bar{\nu}$	0.08 ± 0.05
TOTAL	2.86 ± 0.93

BACKGROUND

To estimate the vector boson associated background, we generated W/Z plus n jet samples using the Monte Carlo (MC) generator VECBOS (12), interfaced with ISAJET (9) to dress up the final parton states. VECBOS allowed us to specify n , the number of primary jets associated with the vector boson production. Since ISAJET also allows control over the decay of the tau, we were careful to count its hadronic decays as contributing to the jet total. Events were then passed through a GEANT simulation of the $D\bar{O}$ detector (11), reconstructed and subjected to our selection criteria. Table 1 lists the expected background from this source.

To estimate the contribution from Standard Model multijet production, we fit the \cancel{E}_T spectrum of a set of single low jet E_T triggers, and then determined the fraction of such events that passed our selection cuts, as a function of \cancel{E}_T . For our final selection cuts that contribution was predicted to be negligible.

The combined background is consistent with the number of observed candidates.

CONCLUSIONS

No excess of events, unexplained by the Standard Model, were observed. With two candidates, our preliminary background subtracted 95% Confidence Level exclusion contour is shown in Fig. 4. This contour intersects the $m_{\tilde{t}_1} = m_{\tilde{z}_1} + m_b + m_W$ line at $106 \text{ GeV}/c^2$. The gap between the LEP limit and our own exclusion region is due to the limitation of the \cancel{E}_T trigger threshold. The 1994-95 data run now carries a customized filter which employs some of the offline cuts against low- E_T jets aligned with $\phi(\cancel{E}_T)$ permitting the \cancel{E}_T cut to be lowered to 25 GeV, which may allow us to cover this area more completely. The region between $m_{\tilde{t}_1} = m_{\tilde{z}_1} + m_c$ and our excluded area requires additional statistics.

ACKNOWLEDGMENTS

We thank the Fermilab Accelerator, Computing, and Research Divisions, and the support staffs at the collaborating institutions for their contributions to the success of this work. We also acknowledge the support of the U.S. Department of Energy, the U.S. National Science Foundation, the Commissariat à l'Énergie Atomique in France, the Ministry for Atomic Energy and the Ministry of Science and Technology Policy in Russia, CNPq in Brazil, the Departments of Atomic Energy and Science and Education in India, Colciencias

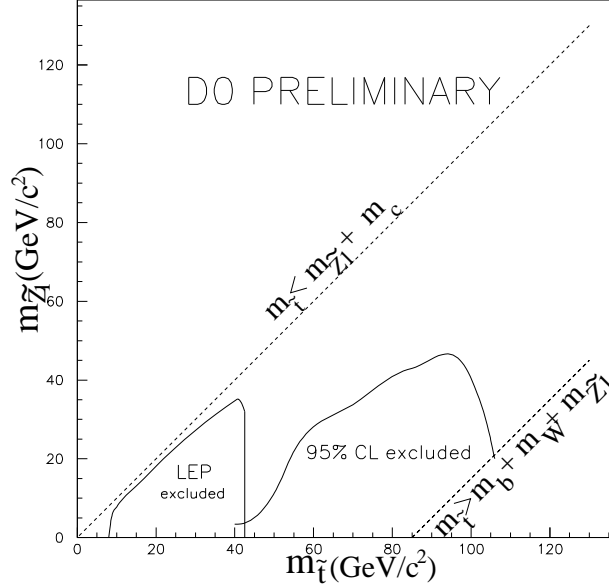


FIG. 4. The 95% Confidence Level contour. For comparison, the latest exclusion contour from OPAL(13) is shown.

in Colombia, CONACyT in Mexico, the Ministry of Education, Research Foundation and KOSEF in Korea and the A.P. Sloan Foundation.

REFERENCES

- * Visitor from IHEP, Beijing, China.
 - † Visitor from CONICET, Argentina.
 - § Visitor from Universidad de Buenos Aires, Argentina.
 - ¶ Visitor from Univ. San Francisco de Quito, Ecuador.
1. S. Abachi, et al. (D0 Collaboration), *Phys. Rev. Lett.* **74** 2632 (1995); F. Abe, et al. (CDF Collaboration), *Phys. Rev. Lett.* **74** 2626 (1995).
 2. J. Ellis and S. Rudaz, *Phys. Lett.* **128B**, 248 (1983); A. Bouquet, J. Kaplan and C. Savoy, *Nucl. Phys.* **B262**, 299 (1985).
 3. H. Baer, M. Drees, R. Godbole, J. F. Gunion, X. Tata, *Phys. Rev.* **D44**, 725 (1991).
 4. P. Harrison and C. Llewellyn Smith, *Nucl. Phys.* **B213**, 223 (1983); **B223**, 542E (1983); G. Kane and J. Leveille, *Phys. Lett.* **112B**, 227 (1982).
 5. X. Tata, in *The Standard Model and Beyond*, p. 304, ed. J. Kim, World Scientific, Singapore (1991); H. Nilles, *Phys. Rep.* **110**, 1 (1984); P. Nath, R. Arnowitt, and A. Chamseddine, *Applied N=1 Supergravity* (ICTP Series in Theoretical Physics, **Vol. 1**), World Scientific, Singapore (1984); H. Haber and G. Kane, *Phys. Rep.* **117**, 75 (1985).
 6. In these models the lightest neutralino \tilde{Z}_1 is the LSP.
 7. S. Abachi *et al.* (D0 Collaboration), *Nucl. Instr. and Meth.* **A338**, 185 (1993) and references therein.
 8. J. Bantly and Q. Li-Demarteau, D0 Internal Note 1691 (1993) (unpublished).

9. F. Paige and S. Protopopescu, Brookhaven National Laboratory Report BNL-37066 (1985) (unpublished).
10. ISASUSY was written by H. Baer and X. Tata and is an extension of ISAJET by F. Paige and S. Protopopescu. The physics included in ISASUSY was first incorporated into ISAJET with ISAJET 7.0.
11. W. Dharmaratna, R. Raja, and C. Stewart, DO Internal Note 1730 (1993) (unpublished). R. Brun and C. Carminati, "GEANT Detector Description and Simulation Tool," CERN, CERN Program Library Writeup W5013 (1993).
12. F. Berends, W. Geile, H. Kuijf, and B. Tausk, FERMILAB-Pub-90/213-T (1990).
13. R. Akers *et al.* (OPAL Collaboration), Phys. Lett. **B337**, 207 (1994).