



## O<sub>2</sub>(a<sup>1</sup>Δ<sub>g</sub>) + Mg, Fe, and Ca: Experimental kinetics and formulation of a weak collision, multiwell master equation with spin-hopping

J. M. C. Plane, C. L. Whalley, L. Frances-Soriano, A. Goddard, J. N. Harvey et al.

Citation: *J. Chem. Phys.* **137**, 014310 (2012); doi: 10.1063/1.4730423

View online: <http://dx.doi.org/10.1063/1.4730423>

View Table of Contents: <http://jcp.aip.org/resource/1/JCPSA6/v137/i1>

Published by the [American Institute of Physics](http://www.aip.org).

---

### Additional information on *J. Chem. Phys.*

Journal Homepage: <http://jcp.aip.org/>

Journal Information: [http://jcp.aip.org/about/about\\_the\\_journal](http://jcp.aip.org/about/about_the_journal)

Top downloads: [http://jcp.aip.org/features/most\\_downloaded](http://jcp.aip.org/features/most_downloaded)

Information for Authors: <http://jcp.aip.org/authors>

## ADVERTISEMENT

A green rectangular banner for NVIDIA. On the left is the NVIDIA logo. To its right, the text reads 'ACCELERATE AMBER AND NAMD BY 5X. TRY IT ON A FREE, REMOTELY-HOSTED CLUSTER.' On the far right, there is a grey button with the text 'LEARN MORE' in white.

 **ACCELERATE AMBER AND NAMD BY 5X.**  
TRY IT ON A FREE, REMOTELY-HOSTED CLUSTER. [LEARN MORE](#)

# $O_2(a^1\Delta_g)$ + Mg, Fe, and Ca: Experimental kinetics and formulation of a weak collision, multiwell master equation with spin-hopping

J. M. C. Plane,<sup>1,a)</sup> C. L. Whalley,<sup>1</sup> L. Frances-Soriano,<sup>1,2</sup> A. Goddard,<sup>1</sup> J. N. Harvey,<sup>3</sup> D. R. Glowacki,<sup>3,b)</sup> and A. A. Viggiano<sup>4</sup>

<sup>1</sup>*School of Chemistry, University of Leeds, Leeds LS2 9JT, United Kingdom*

<sup>2</sup>*Department of Physical Chemistry, University of Valencia, Valencia 46100, Spain*

<sup>3</sup>*Centre for Computational Chemistry, University of Bristol, Bristol BS8 1TS, United Kingdom*

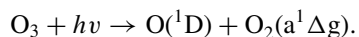
<sup>4</sup>*Air Force Research Laboratory, Kirtland AFB, New Mexico 87117, USA*

(Received 25 April 2012; accepted 5 June 2012; published online 5 July 2012)

The first excited electronic state of molecular oxygen,  $O_2(a^1\Delta_g)$ , is formed in the upper atmosphere by the photolysis of  $O_3$ . Its lifetime is over 70 min above 75 km, so that during the day its concentration is about 30 times greater than that of  $O_3$ . In order to explore its potential reactivity with atmospheric constituents produced by meteoric ablation, the reactions of Mg, Fe, and Ca with  $O_2(a)$  were studied in a fast flow tube, where the metal atoms were produced either by thermal evaporation (Ca and Mg) or by pulsed laser ablation of a metal target (Fe), and detected by laser induced fluorescence spectroscopy.  $O_2(a)$  was produced by bubbling a flow of  $Cl_2$  through chilled alkaline  $H_2O_2$ , and its absolute concentration determined from its optical emission at 1270 nm ( $O_2(a^1\Delta_g - X^3\Sigma_g^-)$ ). The following results were obtained at 296 K:  $k(Mg + O_2(a) + N_2 \rightarrow MgO_2 + N_2) = (1.8 \pm 0.2) \times 10^{-30} \text{ cm}^6 \text{ molecule}^{-2} \text{ s}^{-1}$ ;  $k(Fe + O_2(a) \rightarrow FeO + O) = (1.1 \pm 0.1) \times 10^{-13} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$ ;  $k(Ca + O_2(a) + N_2 \rightarrow CaO_2 + N_2) = (2.9 \pm 0.2) \times 10^{-28} \text{ cm}^6 \text{ molecule}^{-2} \text{ s}^{-1}$ ; and  $k(Ca + O_2(a) \rightarrow CaO + O) = (2.7 \pm 1.0) \times 10^{-12} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$ . The total uncertainty in these rate coefficients, which mostly arises from the systematic uncertainty in the  $O_2(a)$  concentration, is estimated to be  $\pm 40\%$ .  $Mg + O_2(a)$  occurs exclusively by association on the singlet surface, producing  $MgO_2(^1A_1)$ , with a pressure dependent rate coefficient.  $Fe + O_2(a)$ , on the other hand, shows pressure independent kinetics.  $FeO + O$  is produced with a probability of only  $\sim 0.1\%$ . There is no evidence for an association complex, suggesting that this reaction proceeds mostly by near-resonant electronic energy transfer to  $Fe(a^5F) + O_2(X)$ . The reaction of  $Ca + O_2(a)$  occurs in an intermediate regime with two competing pressure dependent channels: (1) a recombination to produce  $CaO_2(^1A_1)$ , and (2) a singlet/triplet non-adiabatic hopping channel leading to  $CaO + O(^3P)$ . In order to interpret the  $Ca + O_2(a)$  results, we utilized density functional theory along with multireference and explicitly correlated CCSD(T)-F12 electronic structure calculations to examine the lowest lying singlet and triplet surfaces. In addition to mapping stationary points, we used a genetic algorithm to locate minimum energy crossing points between the two surfaces. Simulations of the  $Ca + O_2(a)$  kinetics were then carried out using a combination of both standard and non-adiabatic Rice–Ramsperger–Kassel–Marcus (RRKM) theory implemented within a weak collision, multiwell master equation model. In terms of atmospheric significance, only in the case of Ca does reaction with  $O_2(a)$  compete with  $O_3$  during the daytime between 85 and 110 km. © 2012 American Institute of Physics. [<http://dx.doi.org/10.1063/1.4730423>]

## I. INTRODUCTION

The first excited state of molecular oxygen,  $O_2(a^1\Delta_g)$ , is produced in the mesosphere-lower thermosphere (MLT) region by photolysis of  $O_3$  at wavelengths shorter than 320 nm (Ref. 1)



$O_2(a)$  is comparatively long lived in the MLT. The quenching lifetime is more than 4 h,<sup>2</sup> much longer than the phosphorescence lifetime of 73.9 min for  $O_2(a^1\Delta_g - X^3\Sigma_g^-)$  emission at 1270 nm.<sup>3</sup> Since the rate of  $O_3$  photolysis in

the MLT is  $\sim 8 \times 10^{-3} \text{ s}^{-1}$ , the daytime steady-state ratio  $[O_2(a)]/[O_3]$  is about 30 (Ref. 4) and the daytime concentration of  $O_2(a)$  around 90–100 km is  $\sim 5 \times 10^9 \text{ cm}^{-3}$ .<sup>5</sup> After sunset,  $O_2(a)$  decays by an order of magnitude every 2.8 h.  $O_2(a)$  contains almost 1 eV of electronic excitation compared with ground-state  $O_2(X^3\Sigma_g^-)$ , and thus has the potential to be significantly more reactive, although it is not a radical species.

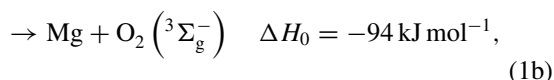
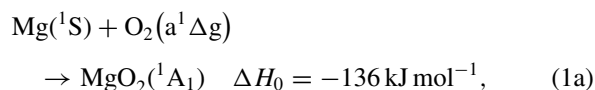
Here we describe a combined experimental and theoretical study of the reactions of  $O_2(a)$  with three metallic constituents of the MLT which are produced by the ablation of roughly 50 tonnes of interplanetary dust which enters the atmosphere each day from space.<sup>6</sup> Meteoric ablation gives rise to the permanent layers of metal atoms that occur globally between about 75 and 110 km. Two of these metals—Fe and Ca—have been studied intensively during the last two

<sup>a)</sup>J.M.C.Plane@leeds.ac.uk.

<sup>b)</sup>drglowacki@gmail.com.

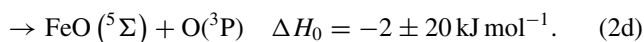
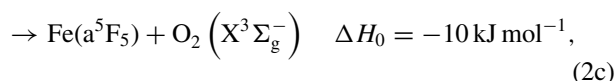
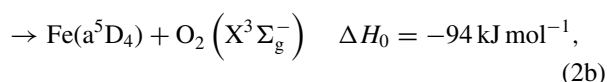
decades using the ground-based resonance lidar technique.<sup>7</sup> Most recently, satellite-borne spectrometric observations of the earth's dayglow have been used to retrieve the global layer of Mg.<sup>8</sup>

Ca, Mg, and Fe atoms all react rapidly with O<sub>3</sub>.<sup>9,10</sup> However, given the daytime [O<sub>2</sub>(a)]/[O<sub>3</sub>] ratio discussed above, the reactions of these metals with O<sub>2</sub>(a) could be atmospherically important if their reaction rate coefficients are larger than about  $1 \times 10^{-12} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$ . This requires the reactions to be exothermic (or very close to thermoneutral). For Mg + O<sub>2</sub>(a) there are only two possibilities



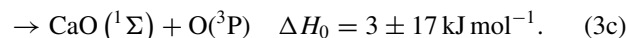
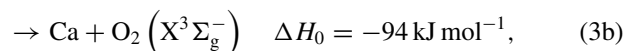
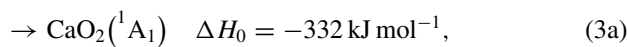
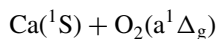
where the reaction enthalpies at 0 K are determined using electronic structure calculations (see below) for 1(a) and experimental energies for 1(b).<sup>11</sup> Reaction 1(b) would involve a spin change which might also allow the formation of triplet MgO<sub>2</sub>.<sup>12</sup> The Mg–O bond is relatively weak and so the formation of MgO + O is endothermic by 164 kJ mol<sup>-1</sup>.<sup>11</sup> The recombination reaction of Mg with O<sub>2</sub>(X) has a barrier of 24 kJ mol<sup>-1</sup>,<sup>12</sup> and thus is extremely slow at room temperature.

For the reaction Fe + O<sub>2</sub>(a) there are four possibilities



Recombination on the reactant quintet surface could yield FeO<sub>2</sub> (or the inserted OFeO isomer). The large exothermicity shown for 2(a) refers to the formation of the lowest quintet, which actually correlates with Fe + O<sub>2</sub>(X).<sup>13</sup> An excited state of quintet FeO<sub>2</sub> should therefore form initially, and a range of triplet and septet states<sup>13</sup> may also be available through spin crossings. In contrast, the reaction between Fe and ground-state O<sub>2</sub>(X<sup>3</sup>Σ<sub>g</sub><sup>-</sup>) can only produce FeO<sub>2</sub>, but this reaction has a large electronic barrier of about 17 kJ mol<sup>-1</sup>.<sup>14</sup> Another interesting possibility is near-resonant electronic energy transfer to yield Fe(a<sup>5</sup>F) + O<sub>2</sub>(X). Fe(a<sup>5</sup>F) is the first electronically excited state of Fe; depending on the a<sup>5</sup>F<sub>J</sub> multiplet produced this reaction channel ranges from being 10 kJ mol<sup>-1</sup> exothermic to 4 kJ mol<sup>-1</sup> endothermic. The final channel producing FeO + O is possibly slightly exothermic, using an Fe–O bond energy D<sub>0</sub> = 402 kJ mol<sup>-1</sup>.<sup>15</sup> However, the uncertainty in this bond energy is probably around ±20 kJ mol<sup>-1</sup>.<sup>16</sup>

Ca + O<sub>2</sub>(a) has three potential channels



In fact, there is considerable uncertainty in the bond strength of Ca–O. A fairly recent high level electronic structure study<sup>17</sup> concluded that the bond strength lay in the range 383 to 417 kJ mol<sup>-1</sup>, so that channel could be endothermic by up to 20 kJ mol<sup>-1</sup>. Recombination to form CaO<sub>2</sub>(<sup>1</sup>A<sub>1</sub>) (the most stable isomer<sup>18</sup>) is spin allowed, and may be considerably faster than the recombination of Ca and ground-state O<sub>2</sub>(X), which has a barrier of about 6 kJ mol<sup>-1</sup>.<sup>18</sup> Note that channels (3b) and (3c) involve a spin crossing from the reactant singlet onto a product triplet surface, which could also allow the formation of triplet CaO<sub>2</sub>.<sup>18</sup>

From a theoretical and computational perspective, the kinetics under investigation within this study (in particular, for the Ca + O<sub>2</sub>(a) system) presents an interesting challenge because they involve spin-hopping processes that occur within an intermediate pressure regime on a multiwell potential energy surface topology. Using a combination of both semiclassical molecular dynamics and statistical mechanics approaches, there has been significant prior work by a number of workers to formulate spin-hopping models which apply in two limits: (1) under zero-pressure, single collision conditions, and (2) under thermal conditions.<sup>19,20</sup> However, far less work has been done to formulate models capable of treating spin-hopping kinetics within intermediate pressure regimes like those characterizing combustion systems or planetary atmospheres. In such regimes, collisionless treatments are inadequate, and explicit semiclassical MD approaches are generally prohibitively expensive. In this work, we extend the multiwell master equation (ME) to include non-adiabatic microcanonical spin-hopping effects. The net result is a general computational kinetic model for polyatomic species which is capable of treating non-adiabatic hopping kinetics that simultaneously occurs alongside collisional relaxation processes.

The structure of the paper is as follows. First, we describe the experimental technique used to study the reactions between O<sub>2</sub>(<sup>1</sup>Δ<sub>g</sub>) and metal atoms. Second, we determine the experimental kinetics of the reactions (1)–(3). In the final section of the paper, we discuss the microscopic mechanisms that underly the phenomenological kinetics, and describe the theoretical/computational approach we have taken in order to simulate the kinetics, including electronic structure calculations and the development of a multiwell spin-hopping master equation.

## II. EXPERIMENTAL TECHNIQUE

Figure 1 is a schematic diagram of the fast flow tube apparatus used to study the kinetics of the Ca, Mg, and Fe reactions. The stainless steel flow tube has an internal diameter of 37.5 mm and consists of sections of tube, cross pieces,

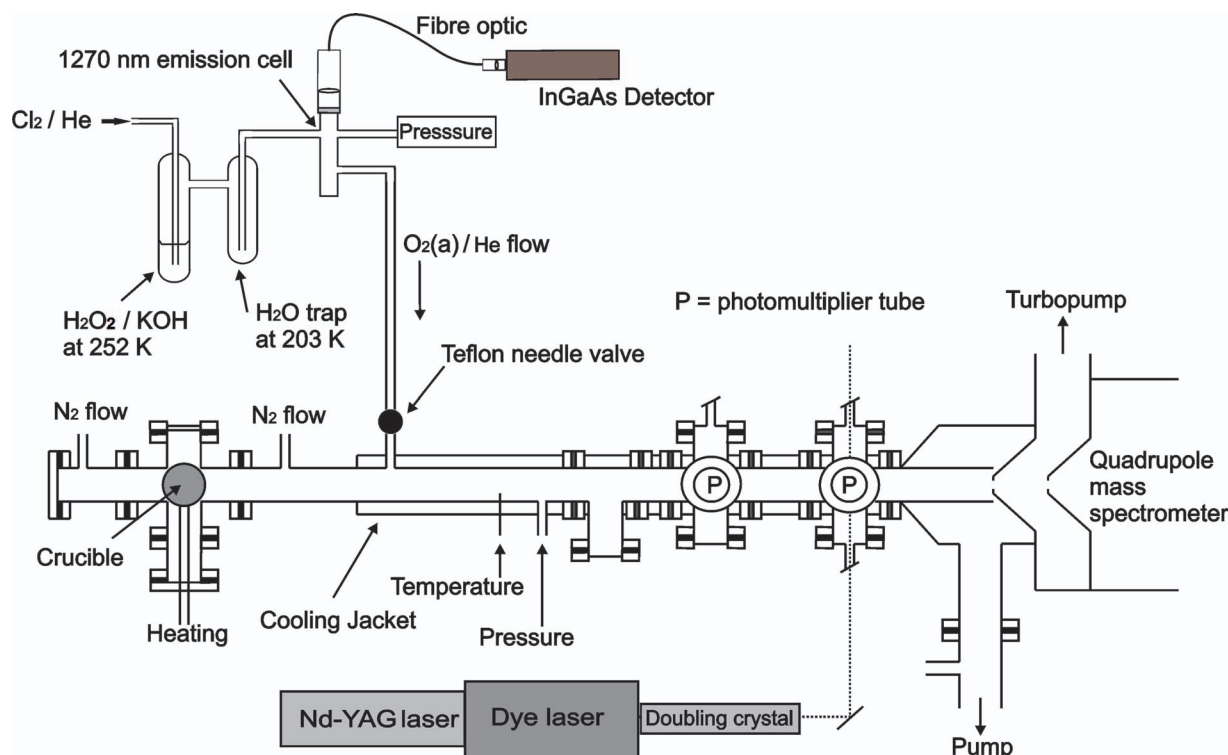


FIG. 1. Schematic diagram of the  $O_2(a)$  generator coupled to a fast flow with laser induced fluorescence detection for studying metal atom reactions (exemplified by Mg with  $O_2(a^1\Delta_g)$ ).

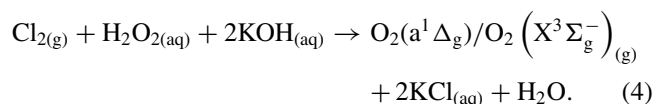
and nipple sections connected by conflat flanges sealed with copper gaskets. The tube has a total length of 1130 mm from the upstream entry point of the carrier gas to the downstream laser induced fluorescence (LIF) detection cell. Calcium atoms were produced continuously by heating calcium pellets (Aldrich, 99%) to 1070–1120 K. Magnesium atoms were produced by heating magnesium pellets (Aldrich, 99.5%) to a temperature between 700 and 800 K. The pellets were located in an aluminium oxide crucible placed inside a tungsten basket heater, positioned 1120 mm upstream of the LIF cell. The Ca or Mg atoms were entrained in the main carrier flow of  $N_2$ , which entered the tube upstream of the crucible. Ca was detected by resonant LIF at 422.7 nm ( $Ca(4^1P_1 - 4^1S)$ ) using a Nd-YAG-pumped dye laser (pulse rate 10 Hz; pulse energy 10 mJ). Mg was detected by resonant LIF at 285.2 nm ( $Mg(3^1P_1 - 3^1S_0)$ ), after frequency doubling the dye laser using a BBO crystal.

The source of Fe atoms in the flow tube was the pulsed ablation of a pure iron rod, using a Nd:YAG laser ( $\lambda = 532$  nm, pulse energy = 22–31 mJ, repetition rate = 8 Hz). The rod was coupled to a stepper motor (via a vacuum feedthrough in a sidearm of the flow tube), so that the rod could be rotated (2–4 Hz) and also translated slowly. This ensured that a fresh surface of the rod was presented to each successive laser shot, in order to keep the resulting pulses of ablated Fe as uniform as possible. The iron rod was long enough ( $\approx 5$  cm) to project across the central axis of the tube. The laser was loosely focused onto the rod through an orthogonal sidearm, so that the point of ablation was in the centre of the flow tube. The pulse of Fe atoms was then entrained in the  $N_2$  carrier gas and transported downstream to the LIF cell, where

the Fe was detected by resonant LIF at 248.3 nm  $Fe(x^5F_5 \leftarrow a^5D_4)$ .

The reactant flow of  $O_2(a)$  in He was injected via a side port downstream of the crucible/ablation cell (Figure 1). The gas flow exited the tube through a throttle valve to a booster pump backed by a rotary pump, providing a volume displacement rate of  $110 \text{ l s}^{-1}$ . Typically, a total gas flow rate of 3200 sccm was used with pressures ranging from 1 to 10 Torr. The Reynolds number was always below 80, ensuring laminar flow within the tube.

$O_2(a^1\Delta_g)$  was prepared using a new technique,<sup>21,22</sup> where  $Cl_2$  is bubbled through a chilled alkaline solution of  $H_2O_2$ ,



This produces  $O_2(a^1\Delta_g)$  at up to 30% yield.<sup>21</sup> The  $O_2(a)$  generator (Figure 1) consisted of two traps and the optical cell used to monitor 1270 nm emission from  $O_2(a)$ , all constructed from Pyrex glass. The first trap, containing 60 ml of 35% w/w  $H_2O_2$  held at  $-21^\circ\text{C}$ , was where reaction (4) took place. A total of 40 ml of chilled 4.0 M KOH was added slowly to this trap to create a slush, through which a 10%  $Cl_2/He$  flow was then bubbled at flow rates up to 100 sccm. The second trap, held at  $-70^\circ\text{C}$ , was used to dry the gas flow by freezing out  $H_2O$ . The  $O_2(a$  and X)/He flow then entered the cylindrical optical emission cell (length = 100 mm, radius = 10 mm), before flowing through a Teflon valve into the fast flow tube.

The weak emission at 1270 nm from O<sub>2</sub>(a-X), exiting through a window at one end of the optical cell, was focused by a lens (bi-convex BK-7 lens,  $f = 30$  mm) through an interference filter (centre wavelength = 1270 nm, FWHM = 4.2 nm, peak transmission = 33%) into a glass fibre optic bundle of length 610 mm. The light exiting the bundle was then focused by a second lens onto an InGaAs photodiode detector (Oriel, Model 71671). The photodiode current was read by a picoammeter (Keithley). The absolute sensitivity of this optical assembly—light collection, transmission through the fibre optic, and detection—was calibrated using a radiometric calibration standard (Ocean Optics LS-1-CAL NIST-traceable light source).

### III. EXPERIMENTAL RESULTS

#### A. Calibration of the O<sub>2</sub>(a) flow

Figure 2 shows that the InGaAs photodiode current resulting from 1270 nm emission in the optical cell was always proportional to the Cl<sub>2</sub> concentration entering the first trap of the O<sub>2</sub>(a) generator. This implies that a constant fraction of the O<sub>2</sub> produced by reaction (4) was in the a<sup>1</sup>Δ<sub>g</sub> state. The calibration of the absolute O<sub>2</sub>(a) concentration was achieved in two stages. First, a computer ray-tracing model was developed to determine the total collection efficiency of 1270 nm photons emitted in the gas cell. The model assumed that O<sub>2</sub>(a) has a uniform concentration in the emission cell. This should be the case given that the shortest quenching lifetime of O<sub>2</sub>(a) caused by the maximum O<sub>2</sub> in the flow in the generator was 42 s,<sup>2</sup> compared with a residence time of the flow in the emission cell of less than 2 s. The model then determined the probability of a photon emitted at each point in the cell being captured by the bi-convex lens and focused through the interference filter onto the entrance of the fibre optic bundle. The total number of photons entering the bundle was then computed by integrating over the cell volume. A correction was applied because the rovibrational line structure of

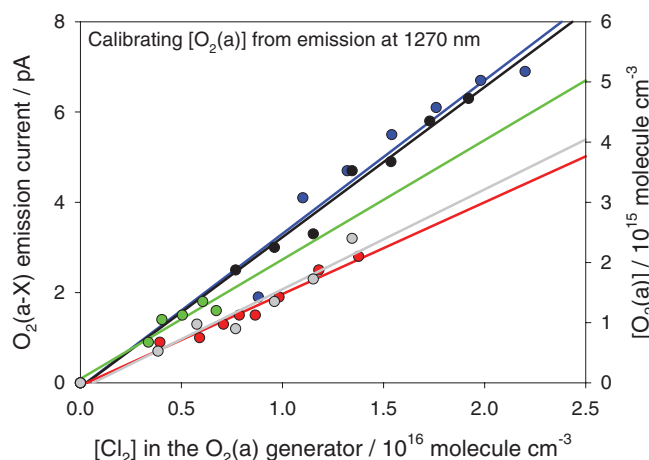


FIG. 2. O<sub>2</sub>(a) emission current measured with the In-Ga-As detector at 1270 nm, as a function of [Cl<sub>2</sub>] in the generator. The corresponding calibrated [O<sub>2</sub>(a)] is shown on the right-hand ordinate. Data from a selection of experimental runs over several months show that the efficiency for O<sub>2</sub>(a) production ranged from 16%–26% of the Cl<sub>2</sub>.

the O<sub>2</sub>(a-X) emission is broader than the bandpass of the interference filter. Convoluting the emission spectrum, calculated using PGOPHER,<sup>23</sup> with the interference bandpass indicates that only 20.1% of the total emission intensity was transmitted through the filter, compared with the signal which would have been measured if all the emission were at the transmission peak of the filter. The result is that if the concentration of O<sub>2</sub>(a) in the cell were  $1.0 \times 10^{16}$  cm<sup>-3</sup>, then the O<sub>2</sub>(a-X) emission power entering the fibre bundle would be 0.12 nW.

The second stage of the calibration procedure involved replacing the optical cell with the radiometric calibration source. The photodiode current was measured as a function of the distance between the radiometric calibration standard and the collection lens. This showed that the calibration factor was 103 pA nW<sup>-1</sup>. Hence, a concentration of  $1.0 \times 10^{16}$  cm<sup>-3</sup> O<sub>2</sub>(a) in the cell would produce a photodiode current of 12.4 pA. That is, the calibration factor was  $8.1 \times 10^{14}$  molecule cm<sup>-3</sup> pA<sup>-1</sup>.

The right-hand ordinate in Figure 2 shows the resulting O<sub>2</sub>(a) concentrations, calculated by applying this calibration factor to the detector current on the left-hand ordinate. The selection of experimental runs shown in Figure 2 covers the range of observed conversion efficiencies of Cl<sub>2</sub> into O<sub>2</sub>(a), which ranged from 16% to 26%. This variation in efficiency seemed most likely due to the cleanliness of the Pyrex glass traps.

#### B. Kinetics experiments

Reaction rate coefficients were determined using a procedure we have described in detail in Ref. 24. Taking reaction (3) as an example, the loss of Ca by diffusion to the flow tube walls and reaction with O<sub>2</sub>(a) can be described by a first-order decay coefficient,  $k'$ , since [O<sub>2</sub>(a)] ≫ [Ca],

$$k' = k_{\text{diff,Ca}} + k[\text{O}_2(\text{a})], \quad (5)$$

where  $k_{\text{diff,Ca}}$  describes the loss of Ca by diffusion and  $k$  is the rate coefficient for reaction (1) (which may depend on pressure). Experiments were carried out by varying [O<sub>2</sub>(a)] while keeping the total mass flow rate and pressure in the flow tube constant. This means that  $k_{\text{diff,Ca}}$  is constant, as well as the reaction time  $t$  between the point of injection of the O<sub>2</sub>(a) and the downstream LIF cell. Since the removal of Ca is pseudo first order

$$\frac{\ln\left(\frac{[\text{Ca}]_0^t}{[\text{Ca}]_{\text{O}_2(\text{a})}^t}\right)}{t} = \frac{\ln[\text{Ca}]_{\text{rel}}}{t} = k[\text{O}_2(\text{a})], \quad (6)$$

where [Ca]<sub>0</sub><sup>t</sup> is the concentration at the LIF detection cell in the absence of O<sub>2</sub>(a), [Ca]<sub>O<sub>2</sub>(a)</sub><sup>t</sup> is the Ca concentration at the LIF detection cell when O<sub>2</sub>(a) is added, and [Ca]<sub>rel</sub> is the ratio of these concentrations. Plots of ln[Ca]<sub>rel</sub>/t versus [O<sub>2</sub>(a)] are shown in Figure 3, for a range of pressures in the flow tube. The linear dependence expected from equation (6) is observed, and the slope of each plot gives the second-order rate coefficient  $k$ .

Figure 3 shows that reaction (3) is pressure dependent. This is confirmed in Figure 4, which is a plot of  $k$  against [N<sub>2</sub>]. The slope of this plot yields the third-order rate

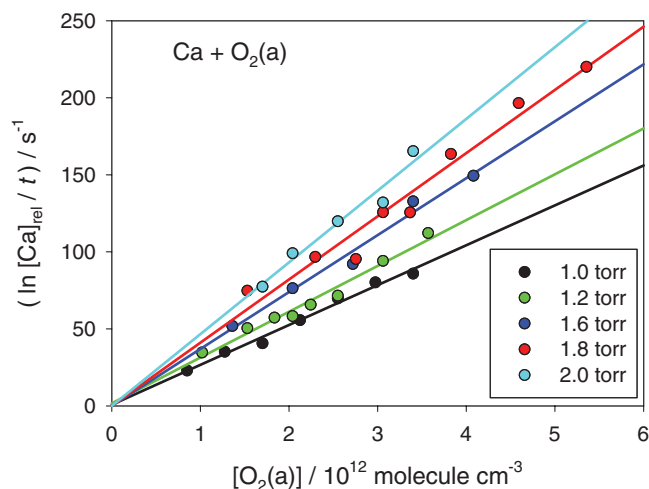


FIG. 3. Kinetic plots showing the first-order removal rate of Ca as a function of  $[O_2(a)]$ , at five different pressures of  $N_2$  in the flow tube.

coefficient listed in Table I. Note that there is also a significant intercept in Figure 4, which indicates that there is a second-order component to reaction (3). Since the only energetically accessible bimolecular products are  $CaO + O(^3P)$  (i.e., channel (3c)), this intercept is clear evidence for spin-hopping onto a triplet surface.

Figure 5 illustrates the first-order removal of Fe as a function of  $[O_2(a)]$ , at two pressures. Note that reaction (2) is essentially independent of pressure. This is confirmed in Figure 6, which also shows the second-order removal rate coefficient for reaction (1) between Mg and  $O_2(a)$ . The Mg reaction is clearly pressure dependent although, in contrast to reaction (3) (Figure 4), there is not a significant intercept on the ordinate. The rate coefficients for reactions (1) and (2) are also listed in Table I.

There are two final points to note here. The first is that Mg, Fe, and Ca react comparatively slowly with  $O_2(X)$ . At the highest pressures employed in the flow tube, the reactions of these atoms with  $O_2(a)$  are 4020, 116, and 82 times faster than their reactions with ground-state  $O_2(X)$  at 300 K, respectively.<sup>12,14,18,25</sup> Therefore, even though the ratio of  $O_2(a)/O_2(X)$  entering the flow tube was in some experimental runs as low as 16%, the reactions of these metal atoms with  $O_2(X)$  would have had a negligible impact on their removal rates and hence little effect on determination of the  $O_2(a)$  kinetics. The second point is that at the highest concentrations of  $O_2(a)$  employed in the generator ( $5 \times 10^{15} \text{ cm}^{-3}$ , Figure 2), the energy pooling reaction<sup>26</sup> between two  $O_2(a)$  molecules (which produces the so-called “dimol” emission at 635 nm)

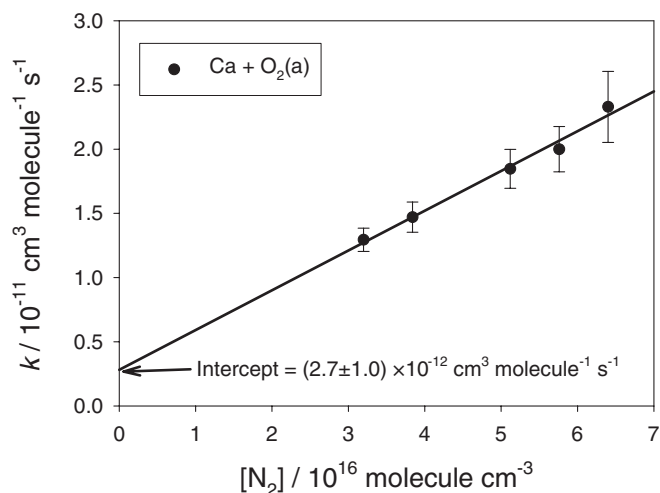


FIG. 4. Plot of the second-order rate coefficient for  $Ca + O_2(a)$  as a function of  $N_2$  concentration. This reaction exhibits third-order (pressure) dependence demonstrating the formation of  $CaO_2(^1A_1)$ ; the significant intercept indicates that the bimolecular channel to  $CaO + O$  is also active.

would have removed about 0.3% of the  $O_2(a)$  between the optical cell where the 1270 nm emission was monitored and injected into the flow tube. Once in the flow tube this reaction would have caused insignificant removal of  $O_2(a)$  because of the short flow time and dilution by the  $N_2$  carrier gas.

#### IV. DISCUSSION AND THEORETICAL ANALYSIS

The experimental results for these three systems show a striking range of kinetic behaviour. On the one hand, the pressure dependence observed for  $Mg + O_2$  is typical of a system in which an association complex undergoes collisional relaxation. On the other hand, for  $Fe + O_2$ , the kinetics are pressure independent, suggesting that all relevant kinetic channels involve prompt dissociation.  $Ca + O_2$  lies in between these extremes, with an association complex for which there is a competition between prompt dissociation and collisional stabilization, with non-adiabatic hopping dynamics clearly playing an important role. To provide microscopic insight into the origin of this behaviour, and to formulate a general kinetics model which allows the laboratory results to be extrapolated over a wider range of conditions, we carried out electronic structure calculations and master equation simulations. This involved a particularly detailed analysis of the  $Ca + O_2$  system, given that it incorporates aspects of both the pressure dependent Mg and pressure independent Fe kinetics.

TABLE I. Rate coefficients measured in the present study at 296 K. The quoted uncertainties are the standard errors from kinetics plots such as in Figures 3 and 5. The total uncertainty, which mostly arises from the systematic uncertainty in the  $O_2(a)$  concentration, is estimated to be  $\pm 40\%$ .

Reaction	Bimolecular rate coefficient ( $\text{cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$ )	Termolecular rate coefficient ( $\text{cm}^6 \text{ molecule}^{-2} \text{ s}^{-1}$ )
$Mg + O_2(a)$	$< 2.4 \times 10^{-14}$	$(1.8 \pm 0.2) \times 10^{-30}$
$Ca + O_2(a)$	$(2.7 \pm 1.0) \times 10^{-12}$	$(2.9 \pm 0.2) \times 10^{-28}$
$Fe + O_2(a)$	$(1.1 \pm 0.1) \times 10^{-13}$	$< 2.4 \times 10^{-31}$

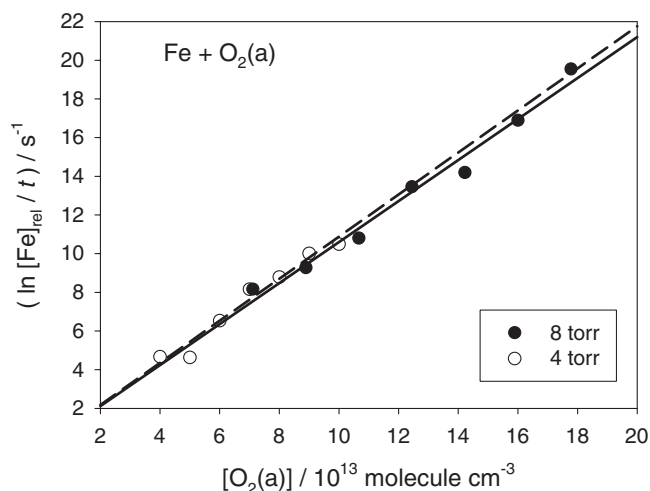


FIG. 5. Kinetic plots showing the first-order removal rate of Fe as a function of  $[O_2(a)]$ , at two different pressures of  $N_2$  in the flow tube.

The electronic structure calculations, which incorporated single and multireference approaches, allowed us to map stationary points and dynamically significant regions of the PESs for the experimental systems described above. The potential energy surface was initially mapped using hybrid density functional theory (DFT), which includes some exact Hartree-Fock exchange. In particular, we used the B3LYP method along with the 6-311+G(2d,p) triple zeta basis set. This is a large, flexible basis set which has both polarization and diffuse functions. At this level of theory, previous theoretical work estimates an expected uncertainty in the calculated reaction enthalpies on the order of  $\pm 20$  kJ mol $^{-1}$ .<sup>27</sup> All DFT calculations were performed using the GAUSSIAN 09 suite of programs.<sup>28</sup> The multireference calculations and explicitly correlated coupled cluster calculations with single, double, and perturbative triples excitations (CCSD(T)-F12) were carried out using MOLPRO,<sup>29</sup> and are described in further detail below.

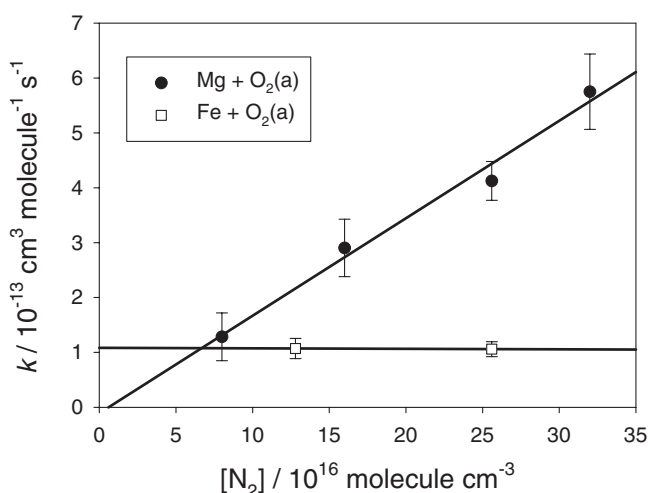


FIG. 6. Plot of the second-order rate coefficient for  $Mg + O_2(a)$  and  $Fe + O_2(a)$  as a function of  $[N_2]$ . The Mg reaction exhibits third-order kinetics forming  $MgO_2(^1A_1)$ . The Fe reaction shows no pressure dependence, indicating that the formation of  $FeO + O$  is the only reactive channel.

Our formulation of the multiwell energy-grained ME has been described in detail in Refs. 30 and 31, so only a brief description is given below. Beginning with a bimolecular  $A + B$  type reaction, the ME allows us to model subsequent adduct formation, isomerisation to form other intermediates, dissociation from the intermediates, and collisional relaxation of the intermediates. The aim of the ME is to provide a description of the reaction system at a macroscopic (or phenomenological) level which is formulated in terms of the behaviour of each of the isomers at an energy resolved (or microcanonical) level. The rovibrational state space of each intermediate is partitioned into “energy grains” with a width no larger than a few kJ mol $^{-1}$ . A differential rate equation is then constructed to describe the grain populations within each isomer and to model the rates of collisional energy transfer into and out of each grain, as well as the probability that population within each grain undergoes reactive processes. The whole set of coupled differential equations may be expressed in matrix form

$$\frac{d}{dt} \mathbf{p} = \mathbf{M} \mathbf{p}, \quad (7)$$

where  $\mathbf{p}$  is a vector containing the populations of each energy grain,  $n_{iE}$ , where  $i$  refers to the  $i$ th isomer and  $E$  to the energy of the grain belonging to a particular isomer.  $\mathbf{M}$  is the matrix that determines the evolution of grain populations due to collisional energy transfer and reaction. Solution of the matrix equation in (7) provides the time dependence of  $\mathbf{p}$ , which is of the form

$$\mathbf{p} = \mathbf{U} e^{\lambda t} \mathbf{U}^{-1} \mathbf{p}(0), \quad (8)$$

where  $\mathbf{p}(0)$  contains the initial ( $t = 0$ ) conditions for each grain (i.e.,  $n_{iE}(0)$ ),  $\mathbf{U}$  is a matrix of eigenvectors obtained from diagonalization of  $\mathbf{M}$ , and  $\lambda$  is a vector of the corresponding eigenvalues. The total number of eigenvalues is equal to the number of grains.

All Rice–Ramsperger–Kassel–Marcus (RRKM) and ME calculations reported in this work were carried out with the open source master equation program, MESMER (master equation solver for multi-well energy reactions).<sup>32</sup> Microcanonical rate coefficients for adiabatic isomerisation and dissociation reactions,  $k(E)$ , were calculated using RRKM theory as

$$k(E) = \frac{W(E)}{h\rho(E)}, \quad (9)$$

where  $W(E)$  is the sum of states at the transition state,  $h$  is Planck’s constant, and  $\rho(E)$  is the reactant density of states. Non-adiabatic modifications to (9) for treating singlet to triplet hopping are described below.

Within our formulation of the master equation, collisional energy transfer in the downward direction (i.e., from energy  $E'$  to  $E$ ) was treated using the so-called “exponential down” model<sup>33</sup> in which

$$P(E \leftarrow E') = C(E') \exp\left(-\frac{E' - E}{\langle \Delta E_d \rangle}\right), \quad (10)$$

where  $P(E \leftarrow E')$  is the probability of undergoing the downward transition,  $C(E')$  is a normalization constant, and  $\langle \Delta E_d \rangle$

is the average energy transferred per collision in a downward direction. The transition probabilities describing energy transfer in the upward direction are obtained from those calculated with (10) and invoking detailed balance. The exponential down model is an isolated binary collision model that typically calculates collision frequencies from Lennard-Jones parameters describing the intermolecular potentials.

Solution of (7) yields a full microcanonical description of the system time evolution; however, in order to link the master equation solution to experimental measurements of temperature and pressure dependent rate coefficients, it is generally necessary to transform the microcanonical information included in (8) to give a phenomenological rate coefficient. In order to accomplish this, we perform an eigenvector/eigenvalue analysis which is similar to methods described by Bartis and Widom.<sup>30,34</sup> The method implemented within MESMER relies on the fact that the eigenvalue spectrum obtained from solution of (7) generally shows separation between those eigenvalues which describe chemical change, and those which describe relaxation processes.

Finally, we note that the numerical algorithms we use to diagonalize  $\mathbf{M}$  in (7) are not immune from numerical instabilities—particularly at low temperatures and with very large wells.<sup>31,35</sup> At low temperatures, numerical instabilities arise from the fact that there is often a very large separation in timescales between those eigenvalues describing chemical change and those describing relaxation. With very large wells, numerical instabilities arise from the fact that the probabilities for energy transfer in the upward direction are often very small. The systems examined in this work are numerically problematic because they involve both relatively low temperatures, and also have very large wells. To limit numerical instabilities, all ME simulations within this work were carried out using high precision arithmetic libraries,<sup>36</sup> which can be called within MESMER using a simple keyword. Additionally, all the results reported in this paper were checked carefully to ensure that they are numerically reliable.

Using the methodology described above, we first discuss our analysis of reaction (1). Figure 7 (top panel) illustrates the stationary points on the singlet potential energy surface for  $\text{Mg} + \text{O}_2(\text{a})$  (red lines), as well as the triplet surface (black lines). If reaction (1) remains on the singlet surface then the outcome should be recombination to form  $\text{MgO}_2(^1\text{A}_1)$  ( $r_{\text{Mg-O}} = 1.81 \text{ \AA}$ ,  $\angle_{\text{O-Ca-O}} = 53.8^\circ$ ), a well which is  $136 \text{ kJ mol}^{-1}$  below the reactants. Although there is an  $\text{Mg-O}_2$  (Ref. 1)  $\text{A}'$  complex which forms initially on this surface ( $r_{\text{Mg-O}} = 1.96 \text{ \AA}$ ,  $r_{\text{O-O}} = 1.31 \text{ \AA}$ ,  $\angle_{\text{Mg-O-O}} = 126.7^\circ$ ), this is

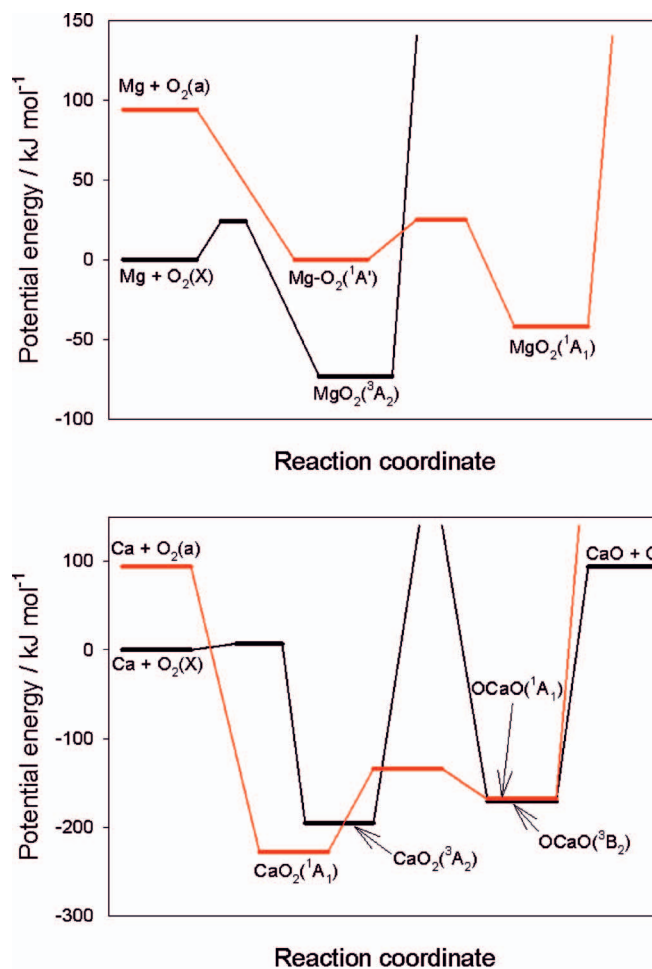


FIG. 7. Potential energy curves (calculated at the B3LYP/6-311+g(2d,p) level of theory) for:  $\text{Mg} + \text{O}_2(\text{a})$  (top panel); and  $\text{Ca} + \text{O}_2(\text{a})$  (bottom panel). Singlet surfaces are shown by red lines and triplet surfaces by black lines. For  $\text{Mg} + \text{O}_2(\text{a})$ , the only product is  $\text{MgO}_2(^1\text{A}_1)$ . Recombination of  $\text{Ca} + \text{O}_2(\text{a})$  produces mostly  $\text{CaO}_2(^1\text{A}_1)$ . However, there is a non-adiabatic crossing seam between  $\text{OCaO}(^1\text{A}_1)$  and  $\text{OCaO}(^3\text{B}_2)$ , where there is a small probability of switching onto the triplet surface and generating the bimolecular products  $\text{CaO} + \text{O}(^3\text{P})$ .

$42 \text{ kJ mol}^{-1}$  less stable than  $\text{MgO}_2(^1\text{A}_1)$ , and there is a barrier between these isomers which is well below the energy of the reactants (i.e.,  $67 \text{ kJ mol}^{-1}$  higher in energy than  $\text{MgO}_2(^1\text{A}_1)$ ). The vibrational frequencies and rotational constants of these stationary points are listed in Table II.

Reaction (1) is more than 3 orders of magnitude faster than  $\text{Mg} + \text{O}_2(\text{X})$ , which has a significant activation energy of  $23 \text{ kJ mol}^{-1}$ .<sup>12</sup> Figure 8 illustrates the singlet and triplet sur-

TABLE II. RRKM parameters for  $\text{Mg} + \text{O}_2(\text{a})$ .

Species	Vibrational frequencies <sup>a</sup>	Rotational constants <sup>a</sup>	Relative energy <sup>b</sup>
$\text{MgO}_2(^1\text{A}_1)^c$	575, 667, 831	0.783, 0.471, 0.294	0
TS from $\text{Mg-O}_2(^1\text{A}')$ to $\text{MgO}_2(^1\text{A}_1)^c$	320i, 592, 960	1.286, 0.297, 0.241	67
$\text{Mg-O}_2(^1\text{A}')^c$	240, 430, 1148	3.194, 0.192, 0.181	42
$\text{Mg} + \text{O}_2(\text{a})^d$	1484	1.426	136

<sup>a</sup>In  $\text{cm}^{-1}$ .

<sup>b</sup>In  $\text{kJ mol}^{-1}$ .

<sup>c</sup>Calculated at the B3LYP/6-311+g(2d,p) level of theory.

<sup>d</sup>NIST Webbook.<sup>52</sup>



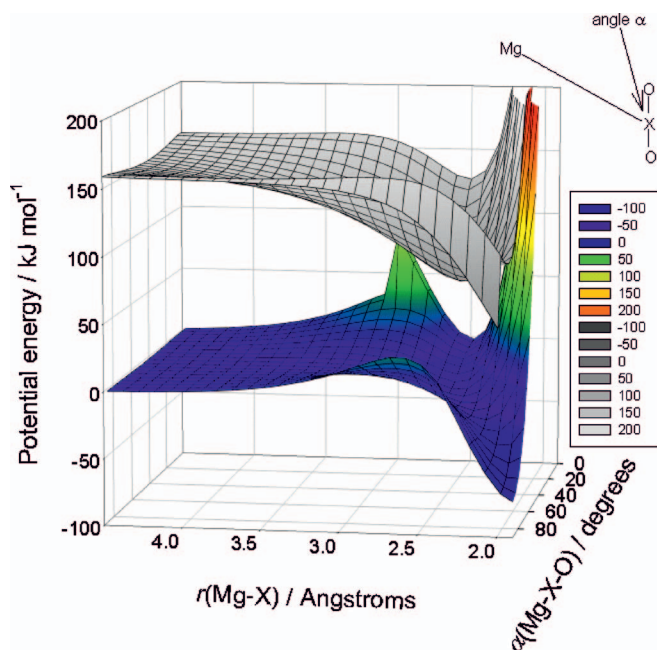


FIG. 8. Potential energy surfaces for  $\text{Mg} + \text{O}_2(\text{a})$  (monochrome shading) and  $\text{Mg} + \text{O}_2(\text{X})$  (coloured shading), calculated at the B3LYP/6-311+g(2d,p) level of theory using a relaxed scan where the O–O bond distance was optimized at each point on the surface. The diagram illustrates that there are no intersections between the surfaces. Thus, the only possible reaction of Mg with  $\text{O}_2(\text{a})$  is recombination to  $\text{MgO}_2(^1\text{A}_1)$ .

faces as a function of the distance between the Mg atom and X (the midpoint between the O atoms), and the Mg–X–O angle  $\alpha$ . These surfaces were calculated using a relaxed scan where the O–O bond distance was optimized at each point on the surface. The barrier on the triplet surface (coloured mesh plot) is particularly pronounced as  $\alpha$  approaches  $0^\circ$  and  $90^\circ$ , which accounts for the experimental activation energy.<sup>12</sup> In contrast, there is no barrier on the singlet surface over most angles of approach (monochrome mesh plot). There are no crossing points between the surfaces at any point below the reactant energy on the singlet surface. Thus, formation of the more stable  $\text{MgO}_2(^3\text{A}_2)$ , or dissociation to  $\text{Mg} + \text{O}_2(\text{X})$ , should be unlikely. On this assumption of negligible singlet-triplet crossing, we modelled reaction (1) using MESMER and the molecular parameters in Table II. The internal energy of each stationary point on the PES was divided into a set of energy grains, each with a width of  $100 \text{ cm}^{-1}$ . The grains associated with Mg– $\text{O}_2$  were then assigned a set of microcanonical rate coefficients for dissociation to  $\text{Mg} + \text{O}_2(\text{a})$ , which were determined using an inverse Laplace transformation to link them directly to  $k_{\text{rec},\infty}$ , the high pressure limiting recombination coefficient.<sup>37</sup> For these neutral reactions,  $k_{\text{rec},\infty}$  was set to a typical capture rate coefficient of  $3 \times 10^{-10} (T/300 \text{ K})^{1/6} \text{ cm}^3 \text{ molecule}^{-1} \text{ s}^{-1}$ ,<sup>38</sup> where the small positive temperature dependence is characteristic of a long-range potential governed by dispersion forces. Setting the collisional energy transfer parameters to typical values for an  $\text{N}_2$  bath gas ( $\langle \Delta E \rangle_{\text{down}} = 300 \text{ cm}^{-1}$ <sup>39</sup> and using  $\text{MgO}_2/\text{N}_2$  Lennard-Jones parameters of  $\sigma = 3.0 \text{ \AA}$  and  $\varepsilon/k = 300 \text{ K}$ ) yields a phenomenological rate coefficient,  $k_1(296 \text{ K})$ , of  $1.5 \times 10^{-30} \text{ cm}^6 \text{ molecule}^{-2} \text{ s}^{-1}$  (cf.  $k_1$  in Table I), in excellent agreement with the experimental rate coefficient. As shown in Figure 9 (top panel), the

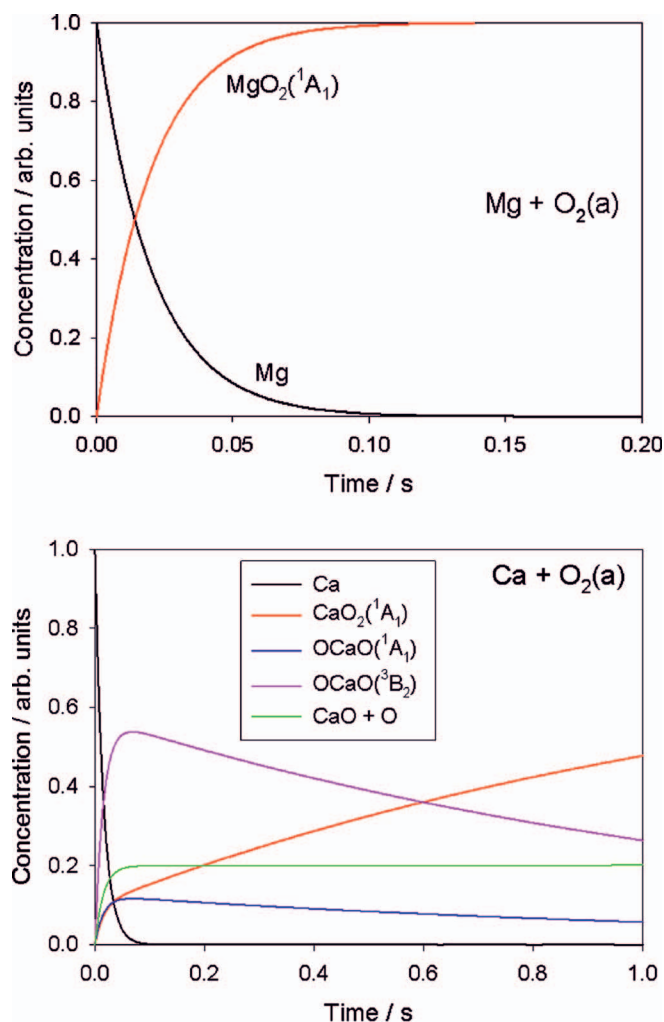


FIG. 9. Time-resolved concentration profiles predicted by MESMER. Top panel:  $\text{Mg} + \text{O}_2(\text{a})$ ,  $[\text{O}_2(\text{a})] = 1.0 \times 10^{14} \text{ cm}^{-3}$ ;  $[\text{N}_2] = 3.2 \times 10^{17} \text{ cm}^{-3}$ . Bottom panel:  $\text{Ca} + \text{O}_2(\text{a})$ ,  $[\text{O}_2(\text{a})] = 5.0 \times 10^{12} \text{ cm}^{-3}$ ;  $[\text{N}_2] = 3.9 \times 10^{16} \text{ cm}^{-3}$ .

only significant product is  $\text{MgO}_2(^1\text{A}_1)$ , which forms at essentially the same rate as Mg decays. When 1% of the Mg has reacted with  $\text{O}_2(\text{a})$ , only 0.3% of the product is in the intermediate Mg– $\text{O}_2$  well and this decreases to 0.004% by the time 90% of the Mg has been oxidised.

There are two surprising aspects to the reaction between Fe and  $\text{O}_2(\text{a})$ . The first is that there is no observable pressure dependence to the second-order rate coefficient (Figures 5 and 6), which implies that recombination is negligible (Table I). The second is that the bimolecular channel is comparatively slow. This reaction starts on a quintet surface with the formation of quintet  $\text{FeO}_2$ . Although the most stable state<sup>13</sup> of quintet  $\text{FeO}_2$  is  $376 \text{ kJ mol}^{-1}$  below the reactants, this state correlates with  $\text{Fe} + \text{O}_2(\text{X})$ .<sup>40</sup> Hence, an excited state of quintet  $\text{FeO}_2$  should form initially. Time-dependent B3LYP calculations indicate that the lowest lying quintet  $\text{FeO}_2$  which correlates with  $\text{Fe}(^5\text{D}) + \text{O}_2(\text{a})$  is only about  $70 \text{ kJ mol}^{-1}$  below these reactants, which may account in part for the absence of observable recombination. This state must then rearrange to  $\text{OFeO}$ , before dissociating to  $\text{FeO} + \text{O}$ ,<sup>40</sup> giving rise to the bimolecular reaction observed. Because of the high quintet spin multiplicity of Fe and FeO, reaction channel (3d)

TABLE III. Computed relative energies and molecular properties for Ca + O<sub>2</sub>(a).

Species	Vibrational frequencies <sup>a</sup>	Rotational constants <sup>a</sup>	Relative energy <sup>b,c</sup>
CaO <sub>2</sub> ( <sup>1</sup> A <sub>1</sub> ) <sup>c</sup>	501, 622, 807	0.914, 0.284, 0.217	0 (0)
TS from CaO <sub>2</sub> ( <sup>1</sup> A <sub>1</sub> ) to OCaO( <sup>1</sup> A <sub>1</sub> ) <sup>c</sup>	260i, 506, 618	0.400, 0.338, 0.183	94
OCaO( <sup>1</sup> A <sub>1</sub> ) <sup>c</sup>	106, 468, 532	1.355, 0.141, 0.128	46
OCaO( <sup>3</sup> B <sub>2</sub> ) <sup>c</sup>	103, 466, 521	1.460, 0.139, 0.127	45
MECP from OCaO( <sup>1</sup> A <sub>1</sub> ) to OCaO( <sup>3</sup> B <sub>2</sub> ) <sup>c</sup>	545, 458	0.834, 0.161, 0.135	59
Ca + O <sub>2</sub> (X)	1580 <sup>d</sup>	1.438 <sup>d</sup>	229 (240)
Ca + O <sub>2</sub> (a) <sup>e</sup>	1484 <sup>d</sup>	1.426 <sup>d</sup>	323 (335)
CaO + O <sup>d</sup>	732	0.445	315 (330)

<sup>a</sup>In cm<sup>-1</sup>.<sup>b</sup>In kJ mol<sup>-1</sup>.<sup>c</sup>Relative energies calculated at the B3LYP/6-311+g(2d,p) level of theory unless mentioned otherwise, with CCSD(T)-F12 results in brackets. All results were corrected for zpe using the indicated vibrational frequencies.<sup>d</sup>NIST Webbook.<sup>52</sup><sup>e</sup>The relative energy of O<sub>2</sub>(a) was not calculated, but assigned based on the experimental X—a splitting of 7918.1 cm<sup>-1</sup> taken from the NIST Webbook.<sup>52</sup>

may also involve crossing onto surfaces of triplet and septet multiplicities.

Nevertheless,  $k_2$  is relatively small (Table I), about 0.1% of the collision frequency between Fe and O<sub>2</sub>(a). This may be because reaction (2d) is more endothermic than the  $\Delta H_0 = -2$  kJ mol<sup>-1</sup> which results if  $D_0(\text{FeO}) = 402$  kJ mol<sup>-1</sup>. A value around  $D_0(\text{FeO}) = 388$  kJ mol<sup>-1</sup>, within the range of measured values for this bond energy,<sup>16</sup> would be required. However, another explanation for the small value of  $k_2$  is that quintet FeO<sub>2</sub>, which forms initially from Fe + O<sub>2</sub>(a), dissociates rapidly to Fe(a<sup>5</sup>F) + O<sub>2</sub>(X). This channel represents near-resonant energy transfer ( $\Delta E = -10$  to +4 kJ mol<sup>-1</sup>, depending on the Fe spin-orbit multiplet), and is thus likely to be the major channel of reaction (2). The excited Fe(a<sup>5</sup>F) atoms produced would be quenched<sup>41</sup> to ground-state Fe(a<sup>5</sup>D) in only 6  $\mu\text{s}$  at even the lowest pressure of N<sub>2</sub> used in the flow tube, so this energy transfer reaction would not have been observable as a loss of Fe atoms in the experiment.

In the case of Ca + O<sub>2</sub>(a), Figure 7 (bottom panel) illustrates the stationary points on the singlet potential energy surface (red lines), as well as the triplet surface (black lines) which links Ca + O<sub>2</sub>(X) with the bimolecular products CaO + O(<sup>3</sup>P). If reaction (3) remains on the singlet surface, then the outcome should be recombination to form CaO<sub>2</sub>(<sup>1</sup>A<sub>1</sub>) ( $r_{\text{Ca-O}} = 1.98$  Å,  $\angle_{\text{O-Ca-O}} = 45.1^\circ$ ), a deep well which is 322 kJ mol<sup>-1</sup> below the reactants. The inserted OCaO(<sup>1</sup>A<sub>1</sub>) isomer ( $r_{\text{Ca-O}} = 2.11$  Å,  $\angle_{\text{O-Ca-O}} = 133.2^\circ$ ) is 46 kJ mol<sup>-1</sup> less stable, and there is a barrier 94 kJ mol<sup>-1</sup> higher in energy than CaO<sub>2</sub>(<sup>1</sup>A<sub>1</sub>) between these singlet forms. The vibrational frequencies and rotational constants of these stationary points are listed in Table III. The energies mentioned in the text and Table III are based on the B3LYP calculations. For some species, single-point CCSD(T)-F12 calculations were carried out using the aug-cc-pVQZ basis on O and the cc-pCVQZ basis on Ca, correlating the O 2s,2p and Ca 3s,3p,4s electrons, and including a perturbative scalar relativistic correction. It can be seen that the agreement with B3LYP is good. The CCSD(T) calculations also agree with previous computational and experimental work concerning the Ca–O bond energy.<sup>17</sup>

Perhaps the most interesting feature of reaction (3) is the experimental indication of a dissociation channel, which must

correspond to the formation of CaO + O(<sup>3</sup>P), and thereby requires hopping from the singlet to triplet surface. In an effort to characterize the regions of most likely spin-hopping within the Ca+O<sub>2</sub> addition complex, we carried out a number of relaxed multireference complete active space self-consistent field (CASSCF)/cc-pVDZ scans along the O–Ca–O angle. All calculations were performed with a 10 electron, 7 orbital active space consisting of the: (1) in-plane  $\pi$  and  $\pi^*$  orbitals on O<sub>2</sub>; (2) out-of-plane  $\pi$  and  $\pi^*$  orbitals on O<sub>2</sub>; (3) in plane  $\sigma$  and  $\sigma^*$  orbitals on O<sub>2</sub>; and (4) the Ca 4s orbital (which transforms as A<sub>1</sub> in C<sub>2v</sub> symmetry). Figure 10 shows the  $\sigma$  and  $\pi$  orbitals used in the CAS calculations along with their corresponding irreducible representations in C<sub>2v</sub> symmetry. Figure 11 shows relaxed singlet and triplet scans carried out over the OCaO bond angle in each of the C<sub>2v</sub> symmetries (a<sub>1</sub>, a<sub>2</sub>, b<sub>1</sub>, b<sub>2</sub>). For the <sup>1</sup>B<sub>1</sub> state, results are not shown because we were unable to achieve convergence. For the <sup>1</sup>B<sub>2</sub> state, the energies are too high to appear on the plot.

An interesting question that arises from the results shown in Figure 11 is why the energy of <sup>3</sup>B<sub>1</sub> OCaO increases

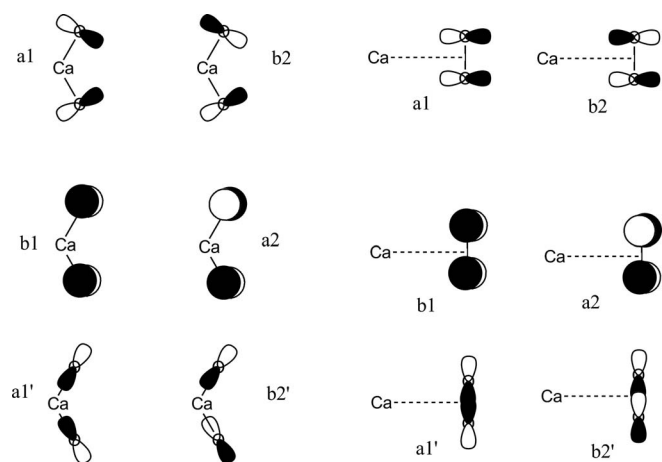


FIG. 10.  $\pi$  and  $\sigma$  orbitals used to carry out the CASSCF calculations described in the text. The left- and right-hand sides of the figure show the orbitals as they appear in OCaO and CaO<sub>2</sub>, respectively. Each orbital configuration is also identified by the corresponding symmetry label of its irreducible representation within the C<sub>2v</sub> point group. The last set of orbitals, a<sub>1</sub>' and b<sub>2</sub>', are labelled using a prime (') only to distinguish them from the top set of orbitals, which have the same symmetries.

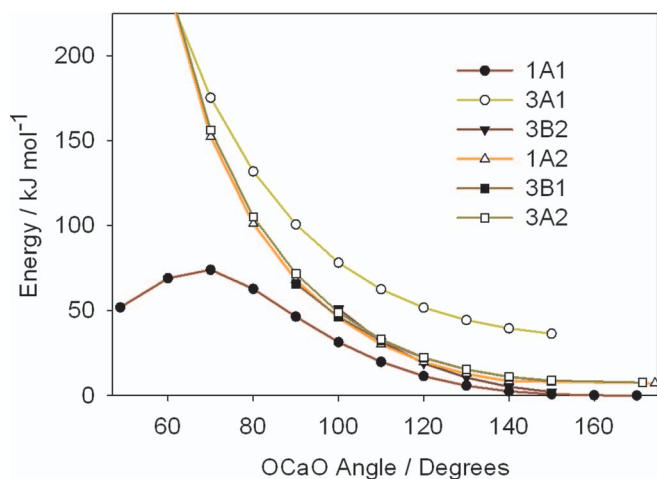


FIG. 11. Relaxed CASSCF potential energy scans along the OCaO angle. As discussed in the text,  $^1B_1$  and  $^1B_2$  are not shown. There is a crossing between  $^1A_1$  and  $^3B_2$ , and the triplet surfaces all increase rapidly to very high energies instead of evolving smoothly to a  $^3CaO_2$  superoxide type structure.

markedly with decreasing OCaO angle, despite the fact that its wavefunction symmetry is identical to that of  $^3CaO_2$ —i.e., why do the calculations not show a smooth transition to a triplet that then leads to  $Ca + O_2(X^3\Sigma_g^-)$ ? This appears to derive from the fact that triplet isomerisation from OCaO to  $CaO_2$ , despite the fact that it appears overall symmetry allowed, is forbidden based on the orbital occupation patterns within the different geometries. For  $Ca + O_2(X)$  at large separations, inspection of the CI vector obtained in CASSCF calculations shows a lowest energy  $^3B_1$  orbital occupation pattern of  $a_1^2(Ca)-a_1^2a_1^2b_1^2b_2^2a_2^1b_2^0$  (using the orbital labels shown in Figure 10). For OCaO, the  $^3B_1$  wavefunction has two dominant electronic configurations, and may be written as  $0.52 \times a_1^0(Ca)-a_1^2a_1^1b_1^1b_2^2b_2^2a_2^2 + 0.48 \times a_1^0(Ca)-a_1^2a_1^2b_1^2b_2^2b_2^1a_2^1$ , both of which differ substantially from the occupation of  $^3B_1$   $Ca + O_2(X)$ . This is reminiscent of what is observed in  $O_3$ ,<sup>42</sup> where both the open and cyclic forms have the same overall symmetry, but interconversion is symmetry forbidden owing to the orbital occupation pattern. In  $C_s$  symmetry,  $a_1$  and  $b_2$  orbitals (in  $C_{2v}$ ) become  $a'$  orbitals, and what were  $b_1$  and  $a_2$  orbitals become  $a''$  orbitals. Hence, the orbital occupation of  $Ca + O_2(X)$  becomes  $a'^2a'^2a'^2a''^2a''^1a''^1a''^0$ , giving an overall wavefunction symmetry of  $^3A''$ . For  $^3A''$  OCaO, the occupation pattern is  $0.52 \times a'^0(Ca)-a'^2a''^1a''^1a'^2a'^2a''^2 + 0.48 \times a'^0(Ca)-a'^2a''^2a''^2a'^2a''^1a''^1$ . Thus, the conversion of OCaO to  $Ca + O_2(X)$  is no longer symmetry forbidden, but there is likely to remain a significant barrier on account of the substantial electron reorganization that takes place on moving from one geometry to the other.

This orbital occupation analysis suggests that any OCaO triplet formed via non-adiabatic hopping from the OCaO singlet will be unlikely to dissociate to  $Ca + O_2(X)$ . This conclusion is compatible with the experimental observations, which see no evidence for a  $Ca + O_2(X)$  channel (which would appear as a slowing down of the Ca removal rate). It is also worth noting that the reaction of  $Ca + O_2(X)$  is much slower than  $Ca + O_2(a)$  because there is a small barrier ( $\sim 6$  kJ mol $^{-1}$ ) in the entrance channel and the well depth of  $CaO_2(^3A_2)$  ( $r_{Ca-O} = 2.21$  Å,  $\angle_{O-Ca-O} = 35.2^\circ$ ) is shallower,<sup>18</sup>

as shown in Figure 7. Hence, the role of  $Ca + O_2(X)$  should be limited in these experiments.

The results in Figure 11 show near degeneracy of the singlet and triplet surfaces near OCaO, with an actual singlet-triplet crossing between the  $^1A_1$  and  $^3B_2$  surfaces. This observation led us to initiate DFT and CASSCF searches for minimum energy crossing points between the  $^1A_1$  and  $^3B_2$  surfaces in the vicinity of OCaO. The minimum energy crossing point (MECP) optimization strategy that we utilized in this work builds on a method which relies on defining two orthogonal vectors  $f$  and  $g$  (Refs. 43 and 44)

$$f = (E_1 - E_3) \left[ \left( \frac{dE_1}{dq} \right) - \left( \frac{dE_3}{dq} \right) \right] = (E_1 - E_3)x_1, \quad (11)$$

$$g = \left( \frac{dE_1}{dq} \right) - \frac{x_1}{|x_1|} \left[ \left( \frac{dE_1}{dq} \right) \cdot \frac{x_1}{|x_1|} \right], \quad (12)$$

where  $q$  denotes the molecular coordinates,  $E_1$  denotes the energy on the singlet surface, and  $E_3$  the energy on the triplet surface. In the neighbourhood of the MECP,  $f$  is orthogonal to the singlet/triplet crossing seam, and  $g$ , which is parallel to the crossing seam, points toward the minimum energy along the seam. At the minimum energy crossing point, both  $f$  and  $g$  vanish.

Typical MECP optimization schemes generally exploit some sort of gradient following in an attempt to minimize  $f$  and  $g$ .<sup>43</sup> Instead of taking this approach, we implemented a genetic algorithm to minimize  $f$  and  $g$ . This was done for two reasons: (1) gradient following methods were plagued by slow convergence because the OCaO region of the PES is an extremely broad diradical basin with little structure and small energy gradients, and (2) numerical instabilities in our gradient following algorithm (which arose in part from the extremely small OCaO energy gradients) repeatedly destroyed the molecular  $C_{2v}$  symmetry. Utilizing the python PyGene library (available at <https://github.com/blaa/PyGene>), we implemented code which interfaced with MOLPRO and GAUSSIAN 09 to undertake an MECP optimization based on simple Mendelian genetics. The “alleles” that we chose corresponded to the OCaO angle and Ca–O bond distance. The algorithm terminated when  $f$  and  $g$  met our convergence criteria.

Genetic MECP optimizations were carried out at both the CASSCF and DFT levels of theory, with the DFT description of the  $^1A_1$  singlet diradical accomplished using broken spin symmetry. The CASSCF and DFT approaches yielded slightly different MECP geometries. CASSCF gave an MECP with an OCaO angle of  $134^\circ$  and CaO bond distance of 2.17 Å, while DFT searches gave an MECP with an angle of  $119^\circ$  and a bond distance of 2.11 Å. Using the CASSCF methodology described above, we calculated the spin-orbit coupling matrix elements<sup>45</sup> between the  $^1A_1$  and  $^3B_2$  states at both the DFT and CASSCF MECP geometries. The root mean square of the spin-orbit coupling matrix element is 0.07 cm $^{-1}$  at the CASSCF MECP, whereas it has a value of 1.07 cm $^{-1}$  at the DFT MECP. In the master equation calculations reported below, we examined the sensitivity of our results within this range of spin-orbit coupling values.

Having located the MECPs, we used code<sup>46</sup> based on a recently developed method<sup>43,44</sup> to calculate effective vibrational frequencies at the MECP, wherein Hessians for the non-adiabatic states are combined to give an overall effective Hessian within the non-adiabatic crossing seam. After projecting out overall rotations, overall translations, and the gradient for motion orthogonal to the MECP seam,<sup>47</sup> the effective Hessian was mass weighted and diagonalized to provide vibrational frequencies and eigenvectors. Because analytic vibrational frequencies are unavailable for  $C_{2v}$  CASSCF wavefunctions in MOLPRO, we were only able to carry out this vibrational analysis on the DFT geometry. The vibrational analysis clearly shows that the mode corresponding to passage through the MECP seam is the OCaO bending motion. The MECP vibrational frequencies and rotational constants are given in Table III.

With the above information in hand, we carried out non-adiabatic transition state theory calculations in MESMER using a modified RRKM expression<sup>43,48</sup>

$$k(E) = \frac{N_{MECP}(E)}{h\rho(E)}, \quad (13)$$

where  $N_{MECP}(E)$  is a convolution of  $\rho_{MECP}(E)$ , the rigid-rotor/harmonic oscillator density of states at the MECP geometry, and  $\rho_{SH}(E)$  the spin-forbidden hopping probabilities, calculated from Landau Zener (LZ) theory

$$N_{MECP}(E) = \int_0^E \rho_{MECP}(E - E_H) p_{SH}(E) dE_H, \quad (14)$$

where

$$p_{SH}(E_H) = (1 + P)(1 - P) \\ P = \exp\left(\frac{-2\pi V_{ST}^2}{h\Delta F} \sqrt{\frac{\mu}{2(E - E_{MECP})}}\right), \quad (15)$$

In Eq. (15),  $\rho_{SH}(E)$  corresponds to a double passage hopping probability, with non-adiabatic transit allowed on both forward and reverse passage through the MECP.  $V_{ST}$  is the matrix element for coupling between the two surfaces,  $\mu$  is the reduced mass for movement along  $f$  (31.4 amu), and  $\Delta F$  is identical to  $|x_1|$  ( $1.06 \times 10^{-3}$  au Bohr<sup>-1</sup>) in (12). The LZ surface hopping model is best suited to non-adiabatic systems with localized coupling regions and narrowly avoided crossings, such as that which occurs between the  $^1A_1$  and  $^3B_2$  surfaces in Figure 11. In addition to this crossing, Figure 11 also shows several weakly coupled states with nearly parallel energy surfaces, for which a Rosen-Zener-Demkov (RZD) type model is best suited to describe the hopping probabilities.<sup>20,49</sup> However, as discussed below, we obtain good agreement with experiment using a model that includes only LZ hopping. Thus, we did not additionally include RZD hopping probabilities, although we note that this would be possible.

The MESMER simulations of reaction (3) also required microcanonical rate coefficients for reaction of Ca with  $O_2(a)$  to form  $CaO_2(^1A_1)$ , and for reaction of CaO with O to form  $^3OCaO$ . Because these do not have a well defined energy barrier, they were calculated using inverse Laplace transformation<sup>31,37</sup> of the same high pressure limiting rate coefficient as for reaction (2) (see above). To estimate energy transfer efficiencies with  $N_2$ , we used values for  $\langle \Delta E \rangle_{down}$  and

Lennard-Jones parameters which were identical to those discussed above. The supplementary material<sup>50</sup> contains the resulting input file for MESMER.

The phenomenological rate coefficients calculated by MESMER are as follows:  $k(Ca + O_2(a) \rightarrow CaO_2) = 3.0 \times 10^{-27}$  cm<sup>6</sup> molecule<sup>-2</sup> s<sup>-1</sup> and  $k(Ca + O_2(a) \rightarrow CaO + O) = 2.7 \times 10^{-12}$  cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup>, in excellent agreement with experiment (Table I), if the bond energy of CaO is set to  $D_0 = 385$  kJ mol<sup>-1</sup>. Using the smaller CASSCF spin-orbit coupling constant of 0.07 cm<sup>-1</sup>, an equally good fit to the experimental data is obtained with  $D_0 = 391$  kJ mol<sup>-1</sup>. The MESMER estimate of  $k(Ca + O_2(a) \rightarrow CaO + O)$  is therefore not very sensitive to the coupling constant, which suggests that  $D_0$  is around 390 kJ mol<sup>-1</sup>. We note that the range of  $D_0$  values suggested by our MESMER calculations (385–391 kJ mol<sup>-1</sup>) is slightly smaller than a previous experimental estimate of 397 kJ mol<sup>-1</sup>,<sup>11</sup> and the 0 K CaO bond energy of 400 kJ mol<sup>-1</sup> calculated using CCSD(T)-F12 methods and the experimental CaO vibrational frequency (see Table III).

However, this range of values is within the estimated uncertainty of  $\pm 17$  kJ/mol determined by Radom and co-workers in a detailed study of the CaO enthalpy of formation. Using DFT, MP2, multireference, G3, G3X, W1, and W2 methods,<sup>17</sup> they noted that low level model chemistries provided accurate geometries and vibrational frequencies. In contrast, for heats of formation they noted significant discrepancies between the aforementioned methods and the available experimental data, with CaO being particularly pathological because a very large Ca basis set is required in order to obtain good results. For the calculation of relative energies (e.g., in Table III) we note that our explicitly correlated CCSD(T)-F12 results appear to converge faster with increasing basis set size than those reported previously.

Figure 9 (lower panel) shows that the time evolution of the Ca +  $O_2(a)$  products is more complex than for Mg +  $O_2(a)$ , and continues over a much longer timescale than the initial consumption of Ca and production of CaO + O. In addition to singlet OCaO, a major recombination product is

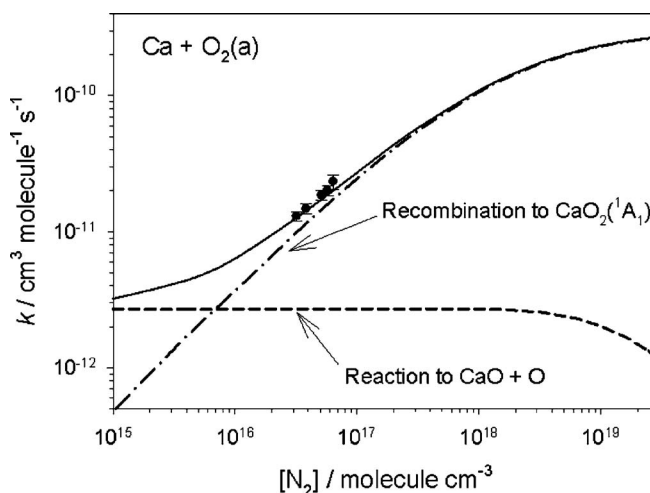


FIG. 12. Modelled dependence of the rate coefficients for the reaction of Ca with  $O_2(a)$  to form  $CaO_2(^1A_1)$  (dash-dot line) or  $CaO + O$  (dash line) as a function of  $N_2$  concentration. The total rate coefficient (solid line) is compared with the rate coefficients measured in the present study.

triplet OCaO (facilitated by spin-hopping), both of which are eventually converted to  $\text{CaO}_2(^1A_1)$ , the most stable of all the possible isomers. Figure 12 illustrates the predicted change with pressure of the overall rate coefficient and the rate coefficients for channels (3a) and (3c), and shows good agreement with the measurements over the narrow pressure range achievable in our experimental flow tube. Because  $k_{3a}$  is already within 5% of  $k_{\text{rec},\infty}$  at a pressure of 10 Torr, this reaction exhibits fall-off behaviour at an unusually low pressure for an atom + diatom recombination reaction. At  $\text{N}_2$  pressures below 0.2 Torr, the  $\text{Ca} + \text{O}(^3P)$  dissociation channel is “formally direct” (Ref. 51)—i.e., significantly faster than recombination. In contrast, above 100 Torr, quenching of singlet and triplet OCaO becomes rapid enough that the triplet dissociation channel starts to turn off.

## V. CONCLUSIONS

This is the first reported study (to our knowledge) of reactions between metallic atoms and  $\text{O}_2(\text{a})$ . The reactions of Ca, Mg, and Fe studied here show strikingly different kinetic behaviour. These metals were chosen for study because of their presence in the earth’s upper atmosphere, and the possibility of  $\text{O}_2(\text{a})$  competing as an oxidant with  $\text{O}_3$ . It turns out that only reaction with Ca is of atmospheric significance: when  $[\text{O}_2(\text{a})]$  is about 30 times larger than  $[\text{O}_3]$  during daytime, the rate of CaO production via  $\text{Ca} + \text{O}_3$ <sup>10</sup> is roughly four times faster than reaction (3). Reactions with Mg and Fe are much too slow to compete with  $\text{O}_3$  in the oxidation of Mg and Fe. Nevertheless, these reactions have proven to be of fundamental interest as a test for electronic structure calculations and rate theory, insofar as they have motivated us to formulate a non-adiabatic weak collision Master Equation model for reaction (3), which provides good agreement with the experimental observations. To our knowledge, this study represents one of the first successful fusions of non-adiabatic RRKM theory with a weak collision master equation. Alongside advances in electronic structure theory and our understanding of collisional energy transfer, we believe that this sort of approach will prove useful in future studies where polyatomic non-adiabatic hopping kinetics occur in intermediate pressure regimes. For this reason, we have included within the supplementary material<sup>50</sup> a sample input file for running spin-hopping simulations of the  $\text{CaO}_2$  system using MESMER.

## ACKNOWLEDGMENTS

This work was supported through award FA8655-10-1-3045 from the European Office of Aerospace Research and Development. Funding for D.R.G. was provided by the Engineering and Physical Sciences Research Council (United Kingdom) (EPSRC(GB)) Programme (Grant No. EP/G00224X).

<sup>1</sup>S. M. Ball, G. Hancock, and F. Winterbottom, *Faraday Discuss.* **100**, 215 (1995).

<sup>2</sup>K. H. Becker, K. H. Groth, and U. Schurath, *Chem. Phys. Lett.* **8**, 259 (1971).

<sup>3</sup>S. M. Newman, A. J. Orr-Ewing, D. A. Newnham, and J. Ballard, *J. Phys. Chem. A* **104**(42), 9467 (2000).

<sup>4</sup>G. P. Brasseur and S. Solomon, *Aeronomy of the Middle Atmosphere* (Springer, Dordrecht, 2005).

<sup>5</sup>P. P. Batista, H. Takahashi, B. R. Clemesha, and E. J. Llewellyn, *J. Geophys. Res.* **101**(A4), 7917, doi:10.1029/95JA03569 (1996).

<sup>6</sup>J. M. C. Plane, *Chem. Rev.* **103**(12), 4963 (2003).

<sup>7</sup>M. Gerding, M. Alpers, U. von Zahn, R. J. Rollason, and J. M. C. Plane, *J. Geophys. Res.* **105**(A12), 27131, doi:10.1029/2000JA900088 (2000); T. J. Kane and C. S. Gardner, *ibid.* **98**(D9), 16875, doi:10.1029/93JD01225 (1993).

<sup>8</sup>M. Scharringhausen, A. C. Aikin, J. P. Burrows, and M. Sinnhuber, *J. Geophys. Res.* **113**(D13), D13303, doi:10.1029/2007JD009043 (2008).

<sup>9</sup>M. Helmer and J. M. C. Plane, *J. Chem. Soc., Faraday Trans.* **90**(1), 31 (1994); J. M. C. Plane and M. Helmer, *Faraday Discuss.* **100**, 411 (1995).

<sup>10</sup>M. Helmer, J. M. C. Plane, and M. R. Allen, *J. Chem. Soc., Faraday Trans.* **89**(5), 763 (1993).

<sup>11</sup>D. R. Lide, *Handbook of Physics and Chemistry* (CRC, Boca Raton, FL, 2006).

<sup>12</sup>C. F. Nien, B. Rajasekhar, and J. M. C. Plane, *J. Phys. Chem.* **97**(47), 12422 (1993).

<sup>13</sup>J. M. C. Plane and R. J. Rollason, *Phys. Chem. Chem. Phys.* **1**(8), 1843 (1999).

<sup>14</sup>M. Helmer and J. M. C. Plane, *J. Chem. Soc., Faraday Trans.* **90**(3), 395 (1994).

<sup>15</sup>S. Smoes and J. Drowart, *High Temp. Sci.* **17**, 31 (1984).

<sup>16</sup>R. J. Rollason and J. M. C. Plane, *Phys. Chem. Chem. Phys.* **2**(10), 2335 (2000).

<sup>17</sup>M. B. Sullivan, M. A. Iron, P. C. Redfern, J. M. L. Martin, L. A. Curtiss, and L. Radom, *J. Phys. Chem. A* **107**(29), 5617 (2003).

<sup>18</sup>M. L. Campbell and J. M. C. Plane, *J. Phys. Chem. A* **105**(14), 3515 (2001).

<sup>19</sup>Q. Cui, K. Morokuma, J. M. Bowman, and S. J. Klippenstein, *J. Chem. Phys.* **110**(19), 9469 (1999); A. W. Jasper, C. Y. Zhu, S. Nangia, and D. G. Truhlar, *Faraday Discuss.* **127**, 1 (2004); J. N. Harvey and M. Aschi, *ibid.* **124**, 129 (2003).

<sup>20</sup>E. E. Nikitin, *Annu. Rev. Phys. Chem.* **50**, 1 (1999).

<sup>21</sup>A. Midey, I. Dotan, and A. A. Viggiano, *J. Phys. Chem. A* **112**, 3040 (2008).

<sup>22</sup>A. J. Midey, I. Dotan, S. Lee, W. T. Rawlins, M. A. Johnson, and A. A. Viggiano, *J. Phys. Chem. A* **111**, 5218 (2007); A. J. Midey, I. Dotan, J. V. Seeley, and A. A. Viggiano, *Int. J. Mass Spectrom.* **280**, 6 (2009).

<sup>23</sup>C. M. Western, PGOPHER—A program for simulating rotational structure, University of Bristol (2011).

<sup>24</sup>S. L. Broadley, T. Vondrak, and J. M. C. Plane, *Phys. Chem. Chem. Phys.* **9**(31), 4357 (2007).

<sup>25</sup>C. Vinciekier and P. Christiaens, *Bull. Soc. Chim. Belg.* **101**, 879 (1992).

<sup>26</sup>R. F. Heidner III, C. E. Gardner, T. M. El-Sayed, G. I. Segal, and J. V. V. Kasper, *J. Chem. Phys.* **74**, 5618 (1981).

<sup>27</sup>J. B. Foresman and A. Frisch, *Exploring Chemistry with Electronic Structure Methods* (Gaussian, Inc., Pittsburgh, PA, 1996).

<sup>28</sup>M. J. Frisch, G. W. Trucks, H. B. Schlegel, G. E. Scuseria, M. A. Robb, J. R. Cheeseman, G. Scalmani, V. Barone, B. Mennucci, G. A. Petersson, H. Nakatsuji, M. Caricato, X. Li, H. P. Hratchian, A. F. Izmaylov, J. Bloino, G. Zheng, J. L. Sonnenberg, M. Hada, M. Ehara, K. Toyota, R. Fukuda, J. Hasegawa, M. Ishida, T. Nakajima, Y. Honda, O. Kitao, H. Nakai, T. Vreven, J. J. A. Montgomery, J. E. Peralta, F. Ogliaro, M. Bearpark, J. J. Heyd, E. Brothers, K. N. Kudin, V. N. Staroverov, R. Kobayashi, J. Normand, K. Raghavachari, A. Rendell, J. C. Burant, S. S. Iyengar, J. Tomasi, M. Cossi, N. Rega, J. M. Millam, M. Klene, J. E. Knox, J. B. Cross, V. Bakken, C. Adamo, J. Jaramillo, R. Gomperts, R. E. Stratmann, O. Yazyev, A. J. Austin, R. Cammi, C. Pomelli, J. W. Ochterski, R. L. Martin, K. Morokuma, V. G. Zakrzewski, G. A. Voth, P. Salvador, J. J. Dannenberg, S. Dapprich, A. D. Daniels, O. Farkas, J. B. Foresman, J. V. Ortiz, J. Cioslowski, and D. J. Fox, GAUSSIAN 09, Revision A.1, Gaussian, Inc., Wallingford, CT, 2009.

<sup>29</sup>H. J. Werner, P. J. Knowles, G. Knizia, F. R. Manby, M. Schütz, P. Celani, T. Korona, R. Lindh, A. Mitrushenkov, G. Rauhut, K. R. Shamasundar, T. B. Adler, R. D. Amos, A. Bernhardsson, A. Berning, D. L. Cooper, M. J. O. Deegan, A. J. Dobbyn, F. Eckert, E. Goll, C. Hampel, A. Hesselmann, G. Hetzer, T. Hrenar, G. Jansen, G. Köppl, Y. Liu, A. W. Lloyd, R. A. Mata, A. J. May, S. J. McNicholas, W. Meyer, M. E. Mura, A. Nicklass, D. P. O’Neill, P. Palmieri, K. Pflüger, R. Pitzer, M. Reiher, T. Shiozaki, H. Stoll, A. J. Stone, R. Tarroni, T. Thorsteinsson, M. Wang, and

- A. Wolf, MOLPRO, version 2010.1, a package of *ab initio* programs, 2010, see [www.molpro.net](http://www.molpro.net).
- <sup>30</sup>J. A. Miller and S. J. Klippenstein, *J. Phys. Chem. A* **110**(36), 10528 (2006); S. H. Robertson, M. J. Pilling, L. C. Jitariu, and I. H. Hillier, *Phys. Chem. Chem. Phys.* **9**(31), 4085 (2007).
- <sup>31</sup>K. L. Gannon, M. A. Blitz, C. H. Liang, M. J. Pilling, P. W. Seakins, and D. R. Glowacki, *J. Phys. Chem. A* **114**(35), 9413 (2010).
- <sup>32</sup>S. H. Robertson, D. R. Glowacki, C.-H. Liang, C. Morley, R. Shannon, M. Blitz, and M. J. Pilling, MESMER (master equation solver for multi-energy well reactions), an object oriented C<sup>b</sup>b<sup>+</sup> program for carrying out ME calculations and eigenvalue-eigenvector analysis on arbitrary multiple well systems, 2008–2012, see <http://sourceforge.net/projects/mesmer>.
- <sup>33</sup>A. W. Jasper and J. A. Miller, *J. Phys. Chem. A* **115**(24), 6438 (2011); J. R. Barker, *Int. J. Chem. Kinet.* **41**(12), 748 (2009).
- <sup>34</sup>J. T. Bartis and B. Widom, *J. Chem. Phys.* **60**(9), 3474 (1974).
- <sup>35</sup>T. J. Frankcombe and S. C. Smith, *Theor. Chem. Acc.* **124**(5–6), 303 (2009); K. L. Gannon, D. R. Glowacki, M. A. Blitz, K. J. Hughes, M. J. Pilling, and P. W. Seakins, *J. Phys. Chem. A* **111**(29), 6679 (2007).
- <sup>36</sup>Y. Hida, X. S. Li, and D. H. Bailey, in *Arith-15 2001: 15th Symposium on Computer Arithmetic*, Proceedings (IEEE Computer Soc., Los Alamitos, 2001), pp. 155.
- <sup>37</sup>J. W. Davies, N. J. B. Green, and M. J. Pilling, *Chem. Phys. Lett.* **126**(3–4), 373 (1986).
- <sup>38</sup>Y. Georgievskii and S. J. Klippenstein, *J. Chem. Phys.* **122**, 194103 (2005).
- <sup>39</sup>R. G. Gilbert and S. C. Smith, *Theory of Unimolecular and Recombination Reactions* (Blackwell, Oxford, 1990).
- <sup>40</sup>D. E. Self and J. M. C. Plane, *Phys. Chem. Chem. Phys.* **5**, 1407 (2003).
- <sup>41</sup>S. A. Mitchell and P. A. Hackett, *J. Chem. Phys.* **93**, 7813 (1990).
- <sup>42</sup>S. S. Xantheas, G. J. Atchity, S. T. Elbert, and K. Ruedenberg, *J. Chem. Phys.* **94**(12), 8054 (1991).
- <sup>43</sup>J. N. Harvey and M. Aschi, *Phys. Chem. Chem. Phys.* **1**(24), 5555 (1999).
- <sup>44</sup>N. Koga and K. Morokuma, *Chem. Phys. Lett.* **119**(5), 371 (1985).
- <sup>45</sup>A. Berning, M. Schweizer, H. J. Werner, P. J. Knowles, and P. Palmieri, *Mol. Phys.* **98**(21), 1823 (2000).
- <sup>46</sup>D. R. Glowacki, The code for performing vibrational analysis at the MECP is available for download at <http://glow-wacky.com/code/>.
- <sup>47</sup>W. H. Miller, N. C. Handy, and J. E. Adams, *J. Chem. Phys.* **72**(1), 99 (1980).
- <sup>48</sup>K. L. Gannon, M. A. Blitz, C. H. Liang, M. J. Pilling, P. W. Seakins, D. R. Glowacki, and J. N. Harvey, *Faraday Discuss.* **147**, 173 (2010).
- <sup>49</sup>A. W. Jasper, M. D. Hack, and D. G. Truhlar, *J. Chem. Phys.* **115**(4), 1804 (2001).
- <sup>50</sup>See supplementary material at <http://dx.doi.org/10.1063/1.4730423> for a sample input file for running spin-hopping simulations using mesmer.
- <sup>51</sup>D. R. Glowacki and M. J. Pilling, *Chem. Phys. Chem.* **11**(18), 3836 (2010).
- <sup>52</sup>K. P. Huber and G. Herzberg, in *NIST Chemistry WebBook, NIST Standard Reference Database Number 69*, edited by P. J. Linstrom and W. G. Mallard (National Institute of Standards and Technology, Gaithersburg, MD, 2012), see <http://webbook.nist.gov>.