

Bilarge Neutrino Mixing and Abelian Flavor Symmetry

Gui-Jun Ding^a, S. Morisi^{b,c}, and J. W. F. Valle^b

^a*Department of Modern Physics,*

University of Science and Technology of China, Hefei, Anhui 230026, China

^b *AHEP Group, Instituto de Física Corpuscular – C.S.I.C./Universitat de València
Edificio de Institutos de Paterna, Apartado 22085, E-46071 València, Spain*

^c *Institut für Theoretische Physik und Astrophysik,
Universität Würzburg, 97074 Würzburg, Germany*

Abstract

We explore two bilarge Neutrino Mixing ansätze within the context of Abelian flavor symmetry theories: (BL₁) $\sin \theta_{12} \sim \lambda$, $\sin \theta_{13} \sim \lambda$, $\sin \theta_{23} \sim \lambda$, and (BL₂) $\sin \theta_{12} \sim \lambda$, $\sin \theta_{13} \sim \lambda$, $\sin \theta_{23} \sim 1 - \lambda$. The first pattern is proposed by two of us and is favored if the atmospheric mixing angle θ_{23} lies in the first octant, while the second one is preferred for the second octant of θ_{23} . In order to reproduce the second texture, we find that the flavor symmetry should be $U(1) \times Z_m$, while for the first pattern the flavor symmetry should be extended to $U(1) \times Z_m \times Z_n$ with m and n of different parity. Explicit models for both mixing patterns are constructed based on the flavor symmetries $U(1) \times Z_3 \times Z_4$ and $U(1) \times Z_2$. The models are extended to the quark sector within the framework of $SU(5)$ grand unified theory in order to give a successful description of quark and lepton masses and mixing simultaneously. Phenomenological implications are discussed.

1 Introduction

Our knowledge of the neutrino oscillation parameters has enormously improved in recent years. In particular the Daya Bay [1], RENO [2] and Double Chooz [3] Collaborations have established that the reactor mixing angle $\theta_{13} > 0$ at about 5σ confidence level, confirming the early hints for a nonzero θ_{13} [4, 5]. Recent global analyses [6, 7] of neutrino oscillation parameters, including the data released at the Neutrino-2012 conference, find that θ_{13} is nonzero at about 10σ , and non-maximal atmospheric mixing angle θ_{23} is preferred. However, it still isn't clear which octant θ_{23} lies in. The global fit of Ref. [6], prefers θ_{23} in the second octant with the best fit value $\sin^2 \theta_{23} = 0.613$ (0.600) for normal (inverted) neutrino mass hierarchy, although this hint is quite marginal and first octant values of θ_{23} are well inside the 1σ range for normal hierarchy and at 1.2σ for inverted spectrum. While the independent phenomenological analyses of atmospheric neutrino data in Ref. [7] obtain a preference for θ_{23} in the first octant for both mass hierarchies and exclude maximal mixing at the 2σ level, the best fit value is found to be $\sin^2 \theta_{23} = 0.386$ (0.392) for normal (inverted) neutrino spectrum. Alternative recent global fits claim both the first and second θ_{23} octants are possible [8]. As for the mass-squared difference, the best fit values of Δm_{sol}^2 and Δm_{atm}^2 are $7.62 \times 10^{-5} \text{eV}^2$ and $2.55(2.43) \times 10^{-3} \text{eV}^2$ respectively, which lead to $\Delta m_{\text{sol}}^2 / \Delta m_{\text{atm}}^2 \simeq 0.030(0.031)$. Here the values shown in parentheses correspond to the inverted neutrino mass hierarchy. Note that the three groups give almost the same 3σ ranges for the lepton mixing parameters.

From the theoretical or model-building point of view, one implication of this significant experimental progress is that it excludes the tri-bimaximal mixing ansatz for neutrino mixing [9], unless the underlying theory is capable of providing sufficiently large corrections. So far many suggestions have been advanced to explain the new data, in particular the largish θ_{13} [10–15]. Instead of seeking for new mass-independent lepton mixing matrices to replace the tri-bimaximal pattern [11–13], which may be derived from certain discrete flavor symmetries, Ref. [14] proposed a novel Wolfenstein-like *ansatz* for the neutrino mixing matrix. In this scheme, all three lepton mixing angles are assumed to be of the same order to first approximation

$$\sin \theta_{12} \sim \lambda, \quad \sin \theta_{13} \sim \lambda, \quad \sin \theta_{23} \sim \lambda, \quad (1)$$

where $\lambda \simeq 0.23$ is the Cabibbo angle, and the symbol “ \sim ” implies that the above relations contain unknown factors of order one, the freedom in these factors can be used to obtain an adequate description of the neutrino mixing. Inspecting the global data fitting [6–8], we see that $\sin \theta_{12} \simeq 2.5\lambda$ and $\sin \theta_{13} \simeq \lambda/\sqrt{2}$, which is proposed in the so-called Tri-bimaximal-Cabibbo mixing [15] and also appeared in the context of quark-lepton complementarity [16]. Such bilarge mixing pattern [14] would clearly provide a good leading order approximation for the current neutrino mixing pattern, if the atmospheric neutrino mixing angle θ_{23} turns out to lie in the first octant. However, the second octant of θ_{23} can not be ruled out and is supported by the analyses in Refs. [6, 8]. In this case the texture

$$\sin \theta_{12} \sim \lambda, \quad \sin \theta_{13} \sim \lambda, \quad \sin \theta_{23} \sim 1 - \lambda \quad (2)$$

could be taken as a viable model-building standard. We shall refer to two mixing patterns as BL_1 and BL_2 textures respectively. The difference between BL_1 and BL_2 mixing lies in the order of magnitude of the atmospheric mixing angle θ_{23} ; the BL_1 mixing pattern would be favored if future experiments establish that θ_{23} belongs to the first octant and the deviation from maximal mixing is somewhat large; otherwise, BL_2 mixing is preferred. It is well-known that the observed hierarchies of masses and flavor mixing in the quarks and charged leptons sectors can be conveniently characterized by the Cabibbo angle. As a result the BL_1 and BL_2 parametrization may have deep implications for the theoretical formulation of the ultimate unified theory of flavor. A lot of work in the literature has demonstrated that the smallness and hierarchy of the quark masses and mixing angles can be naturally generated in theories which, at low energy, are described effectively by an Abelian horizontal symmetry, which is explicitly broken by a small parameter [17–19]. It certainly follows a natural path to try and apply these ideas on Abelian family symmetries developed for the quarks to the lepton sector. In this work, we shall investigate whether and how the BL_1 and BL_2 textures can be reproduced naturally from the Abelian horizontal flavor symmetry. For generality we assume that the light neutrino masses arise from lepton-number-violating effective Weinberg-like operators.

The paper is organized as follows. In section 2, we present the effective low energy theory for the Abelian $U(1)$ flavor symmetry and its extension to $U(1) \times Z_m \times Z_n$. We find that, in order to produce the BL_1 texture without fine-tuning, the family symmetry should be $U(1) \times Z_m \times Z_n$ with m and n of opposite parity. Models for the BL_1 and BL_2 schemes are constructed in section 3 and section 4 respectively. These models are extended to include quarks within the $SU(5)$ grand unified theory (GUT), the observed patterns of both quark and lepton masses and flavor mixings are reproduced, and the general phenomenological predictions of the models are discussed. Finally, our conclusions are summarized in section 5.

2 Theoretical Framework

Our theoretical framework is defined as follows. For definiteness we consider a low energy effective theory with the same particle content as the supersymmetric Standard Model (SM). In addition to supersymmetry and the SM gauge symmetry, we introduce a horizontal $U(1)$ symmetry and a SM singlet chiral superfield Θ which is charged under the $U(1)$ family symmetry; without loss of generality, we normalize its charge to -1 . The effective Yukawa couplings of the quarks and leptons are generated from nonrenormalizable superpotential terms of the form

$$\begin{aligned}
W = & (y_u)_{ij} Q_i U_j^c H_u \left(\frac{\Theta}{\Lambda}\right)^{F(Q_i)+F(U_j^c)} + (y_d)_{ij} Q_i D_j^c H_d \left(\frac{\Theta}{\Lambda}\right)^{F(Q_i)+F(D_j^c)} \\
& + (y_e)_{ij} L_i E_j^c H_d \left(\frac{\Theta}{\Lambda}\right)^{F(L_i)+F(E_j^c)} + (y_\nu)_{ij} \frac{1}{\Lambda} L_i L_j H_u H_u \left(\frac{\Theta}{\Lambda}\right)^{F(L_i)+F(L_j)} \quad (3)
\end{aligned}$$

where $H_{u,d}$ are Higgs doublets, Q_i and L_i are the left-handed quark and lepton doublets respectively, U_j^c , D_j^c and E_j^c are the right-handed up-type quark, down-type quark and charged lepton superfields respectively, and i, j are generation indices. The parameter Λ is the cutoff scale of the $U(1)$ symmetry, and $F(\psi)$ denotes the $U(1)$ charge of the field ψ . Note that $F(H_u)$ and $F(H_d)$ do not appear in the exponents since one can always set the horizontal charges of the Higgs doublet H_u and H_d to zero by redefinition of the $U(1)$ charges. The last term of Eq.(3) is the high-dimensional version of the effective lepton-number-violating Weinberg operator.

For the Froggatt-Nielsen flavon field Θ , the supersymmetric action contains a Fayet-Iliopoulos term and the associated D-term in the scalar potential provides a large vacuum expectation value (VEV) for the scalar component of Θ . The D-term in the potential is given by

$$V_D = \frac{1}{2}(M_{FI}^2 - g_\Theta |\Theta|^2)^2 \quad (4)$$

where M_{FI}^2 is the Fayet-Iliopoulos term. The vanishing of V_D requires

$$|\langle \Theta \rangle| = M_{FI}/\sqrt{g_\Theta} \quad (5)$$

We note that this flavor symmetry breaking mechanism is also frequently exploited in discrete flavor symmetry model building [20]. Once the horizontal symmetry is broken by the VEV $\langle \Theta \rangle$, one obtains the quark and lepton mass matrices whose elements are suppressed by powers of the small parameter $\langle \Theta \rangle/\Lambda$, which for simplicity is usually assumed to be characterized by the Cabibbo angle, i.e., $\lambda = \langle \Theta \rangle/\Lambda$, then we have

$$\begin{aligned} (M_u)_{ij} &= (y_u)_{ij} \lambda^{F(Q_i)+F(U_j^c)} v_u, & (M_d)_{ij} &= (y_d)_{ij} \lambda^{F(Q_i)+F(D_j^c)} v_d, \\ (M_e)_{ij} &= (y_e)_{ij} \lambda^{F(L_i)+F(E_j^c)} v_u, & (M_\nu)_{ij} &= (y_\nu)_{ij} \lambda^{F(L_i)+F(L_j)} \frac{v_u^2}{\Lambda} \end{aligned} \quad (6)$$

where $v_{u,d} = \langle H_{u,d} \rangle$ is the electroweak scale VEV of the Higgs doublet $H_{u,d}$. The factors $(y_u)_{ij}$, $(y_d)_{ij}$, $(y_e)_{ij}$ and $(y_\nu)_{ij}$ are not constrained by the flavor symmetry and are usually assumed to be of order one, the freedom in these factors is used in order to obtain a quantitative description of the fermion masses and flavor mixings. Since the holomorphicity of the superpotential forbids nonrenormalizable terms with a negative power of the superfield Θ , one has $(M_u)_{ij} = 0$ if $F(Q_i) + F(U_j^c) < 0$. Similarly $(M_d)_{ij} = 0$ if $F(Q_i) + F(D_j^c) < 0$, $(M_e)_{ij} = 0$ if $F(L_i) + F(E_j^c) < 0$, and $(M_\nu)_{ij} = 0$ if $F(L_i) + F(L_j) < 0$.

In our framework, the light neutrino masses are generated by the high-dimensional effective Weinberg operators shown in the last term of Eq.(3), consequently, the light neutrinos are Majorana particles and its mass matrix M_ν is symmetric with $(M_\nu)_{ij} = (M_\nu)_{ji}$ ¹. Furthermore, if all the horizontal charges are positive, the hierarchial structure of the mass

¹ Note that if we introduce three right-handed neutrino superfields N_i^c to generate light neutrino mass via type I seesaw mechanism, the structure of the light neutrino mass matrix M_ν is independent of the N_i^c charge assignments [24, 25], unless there are holomorphic zeros in neutrino Dirac mass matrix M_D or in Majorana mass matrix M_N for the heavy fields N^c .

matrices shown in Eq.(6) allows a simple order of magnitude estimate for the various mass ratios and mixing angles:

$$\begin{aligned} \frac{m_{u_i}}{m_{u_j}} &\sim \lambda^{F(Q_i)-F(Q_j)+F(U_i^c)-F(U_j^c)}, & \frac{m_{d_i}}{m_{d_j}} &\sim \lambda^{F(Q_i)-F(Q_j)+F(D_i^c)-F(D_j^c)}, & V_{ij} &\sim \lambda^{F(Q_i)-F(Q_j)}, \\ \frac{m_i}{m_j} &\sim \lambda^{2[F(L_i)-F(L_j)]}, & \frac{m_{\ell_i}}{m_{\ell_j}} &\sim \lambda^{F(L_i)-F(L_j)+F(E_i^c)-F(E_j^c)}, & \sin \theta_{ij} &\sim \lambda^{F(L_i)-F(L_j)} \end{aligned} \quad (7)$$

where m_i is the light neutrino mass, and V_{ij} denotes the element of the quark Cabibbo-Kobayashi-Maskawa quark-mixing matrix (CKM) mixing matrix. We note that the sign “ \sim ” implies that there is an unknown order one coefficient in each relation, so that the actual value of the mass ratios and mixing angles may slightly depart from the naive “power counting” estimate. Moreover, if some fields carry negative F charges, then holomorphy plays an important role and the estimates (7) could be violated as well. For the BL_1 mixing pattern, both $\sin \theta_{12}$ and $\sin \theta_{23}$ are of order λ , then we should require

$$F(L_1) = F(L_2) + 1, \quad F(L_2) = F(L_3) + 1 \quad (8)$$

This implies $F(L_1) = F(L_3) + 2$, as a result, we have $\sin \theta_{13} \sim \lambda^2$. Therefore we conclude that the BL_1 mixing pattern can not be naturally produced from a pure $U(1)$ flavor symmetry. Turning to the BL_2 mixing pattern given by $\sin \theta_{23} \sim 1$, $\sin \theta_{12} \sim \lambda$ and $\sin \theta_{13} \sim \lambda$, one should choose

$$F(L_2) = F(L_3) = F(L_1) - 1 \quad (9)$$

Then we have the $(2i)$ and $(3i)$ ($i = 1, 2, 3$) entries of the charged lepton mass matrix are of the same order, hence the diagonalization of the charged lepton mass matrix leads to large 2-3 mixing. In addition, we obtain

$$M_\nu \sim \lambda^{2F(L_3)} \begin{pmatrix} \lambda^2 & \lambda & \lambda \\ \lambda & 1 & 1 \\ \lambda & 1 & 1 \end{pmatrix} \frac{v_u^2}{\Lambda} \quad (10)$$

Clearly the (2-3) sector of the light neutrino mass matrix has a democratic structure, thus large mixing in this (2-3) sector is naturally obtained. However, barring the presence of special cancellations, the masses of the second and the third light neutrinos are typically expected to be of the same order in this case. As a result, the three light neutrinos are quasi-degenerate and strong parameter fine-tuning is required in order to account for the hierarchy between the measured mass squared differences Δm_{sol}^2 and Δm_{atm}^2 .

In order to avoid this kind of fine-tuning in obtaining an acceptable pattern of neutrino oscillation parameters, we must go beyond the pure $U(1)$ flavor symmetry case considered above. Let us now move to the extended flavor symmetry $U(1) \times Z_m \times Z_n \subset U(1) \times$

$U(1)' \times U(1)''$. This kind of Abelian symmetry is somewhat complex and not yet fully discussed, as far as we know, since most of the previous work concentrated on $U(1)$ or $U(1) \times Z_m \subset U(1) \times U(1)'$ flavor symmetry. We now consider [18, 25] three SM singlet superfields Θ_1 , Θ_2 and Θ_3 with the horizontal charges

$$\Theta_1 : (-1, 0, 0), \quad \Theta_2 : (0, -1, 0), \quad \Theta_3 : (0, 0, -1) \quad (11)$$

In exactly the same way as the single $U(1)$ case, the three flavons Θ_1 , Θ_2 and Θ_3 could get non-vanishing VEVs determined by corresponding the D-terms. In general the VEVs $\langle \Theta_1 \rangle$, $\langle \Theta_2 \rangle$ and $\langle \Theta_3 \rangle$ are different [18, 25]. For simplicity, we take in what follows: $\langle \Theta_1 \rangle / \Lambda \sim \lambda$, $\langle \Theta_2 \rangle / \Lambda \sim \lambda$ and $\langle \Theta_3 \rangle / \Lambda \sim \lambda$. The effective Yukawa couplings are given by extending Eq.(3) with new flavons Θ_1 , Θ_2 and Θ_3 as follows:

$$\begin{aligned} W = & (y_u)_{ij} Q_i U_j^c H_u \left(\frac{\Theta_1}{\Lambda} \right)^{F(Q_i)+F(U_j^c)} \left(\frac{\Theta_2}{\Lambda} \right)^{[Z_m(Q_i)+Z_m(U_j^c)]} \left(\frac{\Theta_3}{\Lambda} \right)^{[Z_n(Q_i)+Z_n(U_j^c)]} \\ & + (y_d)_{ij} Q_i D_j^c H_d \left(\frac{\Theta_1}{\Lambda} \right)^{F(Q_i)+F(D_j^c)} \left(\frac{\Theta_2}{\Lambda} \right)^{[Z_m(Q_i)+Z_m(D_j^c)]} \left(\frac{\Theta_3}{\Lambda} \right)^{[Z_n(Q_i)+Z_n(D_j^c)]} \\ & + (y_e)_{ij} L_i E_j^c H_d \left(\frac{\Theta_1}{\Lambda} \right)^{F(L_i)+F(E_j^c)} \left(\frac{\Theta_2}{\Lambda} \right)^{[Z_m(L_i)+Z_m(E_j^c)]} \left(\frac{\Theta_3}{\Lambda} \right)^{[Z_n(L_i)+Z_n(E_j^c)]} \\ & + (y_\nu)_{ij} \frac{1}{\Lambda} L_i L_j H_u H_u \left(\frac{\Theta_1}{\Lambda} \right)^{F(L_i)+F(L_j)} \left(\frac{\Theta_2}{\Lambda} \right)^{[Z_m(L_i)+Z_m(L_j)]} \left(\frac{\Theta_3}{\Lambda} \right)^{[Z_n(L_i)+Z_n(L_j)]} \end{aligned} \quad (12)$$

where $Z_{m,n}(\psi)$ is the $Z_{m,n}$ charge of the field ψ , and the brackets $[\dots]$ around the exponents denote that we are modding out by m (n) according to the Z_m (Z_n) addition rule, namely,

$$[Z_m(Q_i) + Z_m(U_j^c)] = \begin{cases} r & \text{if } r < m \\ r - m & \text{if } r \geq m \end{cases} \quad (13)$$

where $r = Z_m(Q_i) + Z_m(U_j^c)$. We note that the charge assignments of the Higgs doublets H_u and H_d have been set to $(0, 0, 0)$ by redefining the flavor symmetry charges of the fields. Thus, the fermion mass matrix can be expressed in term of the horizontal charges as

$$\begin{aligned} (M_u)_{ij} &= (y_u)_{ij} \lambda^{F(Q_i)+F(U_j^c)+[Z_m(Q_i)+Z_m(U_j^c)]+[Z_n(Q_i)+Z_n(U_j^c)]} v_u, \\ (M_d)_{ij} &= (y_d)_{ij} \lambda^{F(Q_i)+F(D_j^c)+[Z_m(Q_i)+Z_m(D_j^c)]+[Z_n(Q_i)+Z_n(D_j^c)]} v_d, \\ (M_e)_{ij} &= (y_e)_{ij} \lambda^{F(L_i)+F(E_j^c)+[Z_m(L_i)+Z_m(E_j^c)]+[Z_n(L_i)+Z_n(E_j^c)]} v_d, \\ (M_\nu)_{ij} &= (y_\nu)_{ij} \lambda^{F(L_i)+F(L_j)+[Z_m(L_i)+Z_m(L_j)]+[Z_n(L_i)+Z_n(L_j)]} \frac{v_u^2}{\Lambda}. \end{aligned} \quad (14)$$

Consider the quark sector, the flavor mixing angles there are given by

$$V_{ij}^u \sim \lambda^{(F(Q_i)+F(U_j^c))-(F(Q_j)+F(U_j^c))+[Z_m(n)(Q_i)+Z_m(n)(U_j^c)]-[Z_m(n)(Q_j)+Z_m(n)(U_j^c)]}, \quad (15)$$

$$V_{ij}^d \sim \lambda^{(F(Q_i)+F(D_j^c))-(F(Q_j)+F(D_j^c))+[Z_m(n)(Q_i)+Z_m(n)(D_j^c)]-[Z_m(n)(Q_j)+Z_m(n)(D_j^c)]}. \quad (16)$$

For $m = n = 0$ the CKM matrix elements describing the charged current weak interaction of quarks behave approximatively as $V_{ij}^{u,d} \sim \lambda^{F(Q_i)-F(Q_j)}$ and therefore the CKM mixing

$V_{CKM} = V^{u\dagger} \cdot V^d$ is expected to scale as $V_{CKM_{ij}} \sim \lambda^{F(Q_i)-F(Q_j)}$. In order to compare with the pure $U(1)$ horizontal symmetry case, we can define an effective flavor charge in the general case $m \neq n \neq 0$ as

$$F_{eff}(\psi) = F(\psi) + Z_m(\psi) + Z_n(\psi). \quad (17)$$

Then it is clear that

$$V_{CKM_{ij}} \sim \lambda^{F_{eff}(Q_i)-F_{eff}(Q_j)\pm\alpha m\pm\beta n}, \quad (18)$$

where $\alpha, \beta = 0, 1$ and we have used Eq. (13) and the fact that

$$[Z_m(Q_i) + Z_m(U_j^c)] - [Z_m(Q_j) + Z_m(U_j^c)] = Z_m(Q_i) - Z_m(Q_j) \pm \alpha m \quad (19)$$

$$[Z_m(Q_i) + Z_m(D_j^c)] - [Z_m(Q_j) + Z_m(D_j^c)] = Z_m(Q_i) - Z_m(Q_j) \pm \alpha m \quad (20)$$

where $\alpha = 0, 1$. The condition for the value $\pm\beta n$ follows similarly. Likewise for the lepton sector, one obtains

$$V_{ij}^l \sim \lambda^{F_{eff}(L_i)-F_{eff}(L_j)\pm\alpha m\pm\beta n}. \quad (21)$$

Therefore, the masses and mixing angles can be enhanced or suppressed by $\lambda^{\pm m \pm n}$ relative to the scaling predictions obtained when the family symmetry is the continuous flavor symmetry $U(1) \times U(1)' \times U(1)''$ because of the discrete nature of $Z_m \times Z_n$. Note that in the case where the light neutrino masses are generated by the type I seesaw mechanism and all fermion charges are positive, the neutrino masses and mixing angles still do not depend on the details of the right-handed neutrino sector, except for the possible enhancement or suppression associated to the $Z_m \times Z_n$ flavor symmetry.

Furthermore, when the flavor symmetry is reduced to $U(1) \times Z_m$ by taking $n = 0$, all the above results remain valid. It is remarkable that we can employ the $U(1) \times Z_m$ flavor symmetry to maintain the BL_2 mixing while achieving very different neutrino masses without fine-tuning. We shall restrict our attention to the case of a Z_2 symmetry which is the minimal nontrivial Z_m group (see, for example, the explicit model construction given in sec. 4 below). In this case just the Z_m symmetry can reproduce a hierarchy in neutrino masses of order λ^2 consistent with the observed ratio of solar-to-atmospheric splittings.

In contrast, note that since the reactor neutrino mixing is necessarily of order $\sin \theta_{13} \sim \lambda^{2 \pm \alpha m}$ the $U(1) \times Z_m$ flavor symmetry can not produce the BL_1 mixing pattern. Indeed for such BL_1 texture one has $\sin \theta_{12} \sim \lambda$ and $\sin \theta_{23} \sim \lambda$, which is in conflict with the required linear behavior of the reactor mixing angle $\sin \theta_{13} \sim \lambda$. Note parenthetically that the Z_1 group consists of only the identity element, so the group $U(1) \times Z_1$ is isomorphic to $U(1)$, and the Z_1 charge of field is 0, hence the flavor symmetry $U(1) \times Z_1$ produces a wrong scaling behavior $\sin \theta_{13} \sim \lambda^2$.

We now turn to the realistic case of the $U(1) \times Z_m \times Z_n$ flavor symmetry. If both solar and atmospheric neutrino mixing angles are of order λ then the reactor angle would be

constrained to be of order $\sin\theta_{13} \sim \lambda^{2\pm\alpha m\pm\beta n}$. As a result, one can have $\sin\theta_{13} \sim \lambda$ if the parity of m and n is opposite. This is an interesting observation of the present work. In section 3, a concrete model for the BL_1 mixing pattern is presented based on the flavor symmetry $U(1) \times Z_3 \times Z_4$.

Since an Abelian flavor symmetry can not predict the exact value of the $\mathcal{O}(1)$ coefficients in front of each invariant operator, we must content ourselves with explaining the orders of magnitude of fermion masses and flavor mixing parameters. To identify the phenomenologically acceptable mass matrices, we will estimate the various mass ratios and mixing angles as approximate powers of the small parameter λ . The hierarchies in the quark mixing angles are clearly displayed in Wolfenstein's truncated form [21] of the parametrization of the CKM matrix [22]:

$$V_{CKM} = \begin{pmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix} \quad (22)$$

where the quantities A , ρ and η are experimentally determined to be of order one. Therefore the order of magnitude of the three mixing angles is given in terms of the λ as

$$|V_{us}| \sim \lambda, \quad |V_{cb}| \sim \lambda^2, \quad |V_{ub}| \sim \lambda^3 - \lambda^4 \quad (23)$$

The charged fermion mass ratios at the grand unified theory (GUT) scale should satisfy [23]

$$\begin{aligned} \frac{m_u}{m_c} &\sim \lambda^4, & \frac{m_c}{m_t} &\sim \lambda^3 - \lambda^4, \\ \frac{m_d}{m_s} &\sim \lambda^2, & \frac{m_s}{m_b} &\sim \lambda^2, \\ \frac{m_e}{m_\mu} &\sim \lambda^2 - \lambda^3, & \frac{m_\mu}{m_\tau} &\sim \lambda^2 \end{aligned} \quad (24)$$

as well as

$$\frac{m_b}{m_\tau} \sim 1, \quad \frac{m_b}{m_t} \sim \lambda^3 \quad (25)$$

for the intrafamily hierarchy. The first identity is the well-known $b - \tau$ unification relation. For the neutrinos, we required that the lepton mixing is of BL_1 or BL_2 type depending on the octant of θ_{23} . For the quark sector, all the explicit models are properly constructed to meet the requirement $m_t/v_u \sim 1$ and $m_b/v_d \sim \lambda^3$

3 Model for BL_1 mixing

As has been shown in the previous section, one can reproduce the BL_1 texture within the framework of $U(1) \times Z_m \times Z_n$ family symmetry, where m and n should have different parity. For concreteness, we shall use $m = 3$ and $n = 4$ for our model. For such symmetry choice the possible model realization of the BL_1 texture is not unique. As a concrete example, here the horizontal charges of the lepton fields are taken to be

$$\begin{aligned} L_1 &: (4, 1, 3), & L_2 &: (3, 2, 2), & L_3 &: (1, 1, 1), \\ E_1^c &: (3, 2, 2), & E_2^c &: (1, 2, 2), & E_3^c &: (0, 0, 0). \end{aligned} \quad (26)$$

One immediately obtains the charged lepton mass matrix

$$M_e \sim \begin{pmatrix} \lambda^8 & \lambda^6 & \lambda^8 \\ \lambda^7 & \lambda^5 & \lambda^7 \\ \lambda^7 & \lambda^5 & \lambda^3 \end{pmatrix} v_d \quad (27)$$

which yields the mass ratios

$$\frac{m_e}{m_\mu} \sim \lambda^3, \quad \frac{m_\mu}{m_\tau} \sim \lambda^2, \quad (28)$$

that are consistent with the experimental requirements. For the charged assignments in Eq.(26), the light neutrino mass matrix is given by

$$M_\nu \sim \begin{pmatrix} \lambda^{12} & \lambda^8 & \lambda^7 \\ \lambda^8 & \lambda^7 & \lambda^7 \\ \lambda^7 & \lambda^7 & \lambda^6 \end{pmatrix} \frac{v_u^2}{\Lambda}. \quad (29)$$

It predicts the light neutrino mass eigenvalues as follows:

$$m_1 \sim \lambda^8 \frac{v_u^2}{\Lambda}, \quad m_2 \sim \lambda^7 \frac{v_u^2}{\Lambda}, \quad m_3 \sim \lambda^6 \frac{v_u^2}{\Lambda} \quad (30)$$

The neutrino mass spectrum is normal hierarchy, this is confirmed by subsequent numerical analysis. It is remarkable that this model gives rise to $m_2/m_3 \sim \lambda$ and $\Delta m_{\text{sol}}^2/\Delta m_{\text{atm}}^2 \sim \lambda^2$, which is in excellent agreement with the experimental data. In conventional $U(1)$ or $U(1) \times Z_m$ flavor symmetries, if any ratio between neutrino masses is an odd power of the small breaking parameter, generally the mixing angle between the two neutrinos will vanish [25]. The crucial point is that the element $(M_\nu)_{22}$, which would have been $\mathcal{O}(\lambda^{14})$ under the continuous $U(1) \times U(1)' \times U(1)''$ symmetry, is enhanced to $\mathcal{O}(\lambda^7)$ due to the discrete symmetry $Z_3 \times Z_4$. Diagonalizing the mass matrices in Eq.(27) and Eq.(29) by the standard perturbative techniques described in Refs. [18,25,26], we get the three lepton flavor mixing angles

$$\sin \theta_{12} \sim \lambda, \quad \sin \theta_{13} \sim \lambda, \quad \sin \theta_{23} \sim \lambda. \quad (31)$$

Hence the BL_1 pattern is produced automatically. Note that the solar neutrino mixing $\sin \theta_{12}$ arises from order λ contributions from the diagonalization of both M_e and M_ν , while at leading order the reactor and the atmospheric neutrino mixing angles receive contribution only from the neutrino mass matrix M_ν . The off-diagonal elements $(M_\nu)_{13}$ and $(M_\nu)_{23}$ are enhanced by Z_4 and Z_3 respectively, hence we have $\sin \theta_{13} \sim \lambda$ and $\sin \theta_{23} \sim \lambda$ instead of the naive expectations $\sin \theta_{13} \sim \lambda^5$ and $\sin \theta_{23} \sim \lambda^4$ characteristic of the continuous flavor symmetry case.

In the following, we shall extend the model to encompass also quark sector. Since GUT relates quarks and leptons, the transformation properties of quark fields can be determined from those of leptons. In order to give a successful description of the observed fermion mass hierarchies and mixings simultaneously under the same flavor symmetry acting on quarks and leptons we work in the framework of $SU(5)$, for definiteness. Another motivation of

considering $SU(5)$ unification is the anomaly cancellation. If the $U(1)$ flavor symmetry is gauged then a general assignment of flavor charges to the fields will be anomalous. One can imagine the anomaly to be canceled via the Green-Schwarz mechanism [27], however, one must check whether the correct relations are satisfied [28]. A convenient way to ensure that the flavor charges are amenable to cancellation is to have the flavor symmetry to commute with the $SU(5)$ group [29].

Here we propose a model with the quark and lepton matter assignments manifestly compatible with potential unification within $SU(5)$. A complete study of a realistic grand unified model addressing the well-known problems such as the doublet-triplet splitting, the proton lifetime and gauge coupling unification, is beyond the scope of the present paper and will be studied elsewhere.

In the conventional $SU(5)$ grand unified theory, the fields D_i^c and L_i of the same generation are assigned to a $\bar{\mathbf{5}}$ multiplet, the fields Q_i , U_i^c and E_i^c are unified in the $\mathbf{10}$ representation. Since the flavor symmetry is required to commute with the gauge symmetry, this means that the fields in each gauge multiplet transform in the same way under the flavor symmetry. Consequently, the quantum numbers of the quark fields under the flavor symmetry $U(1) \times Z_3 \times Z_4$ are as follows:

$$\begin{aligned}
Q_{L1} &: (3, 2, 2), & Q_{L2} &: (1, 2, 2), & Q_{L3} &: (0, 0, 0), \\
U_1^c &: (3, 2, 2), & U_2^c &: (1, 2, 2), & U_3^c &: (0, 0, 0), \\
D_1^c &: (4, 1, 3), & D_2^c &: (3, 2, 2), & D_3^c &: (1, 1, 1).
\end{aligned} \tag{32}$$

We note that although there are many possible assignments to produce the BL_1 texture in the neutrino sector, only a few of them can satisfy the quark sector phenomenological constraints within $SU(5)$. It is well-known that the minimal $SU(5)$ grand unified theory predicts that the down-type quark mass matrix is the transpose of the charged lepton mass matrix, therefore the down-type quarks and charged lepton masses are closely related: $m_e = m_d$, $m_\mu = m_s$ and $m_\tau = m_b$, which are in gross disagreement with the measured fermion masses and must be corrected [30]. This can be done through the contribution of renormalizable [30] or non-renormalizable [31] operators to the Yukawa matrices. Following Ref. [32], we introduce an additional $U(1) \times Z_3 \times Z_4$ singlet superfield Σ transforming as a $\mathbf{75}$ of $SU(5)$, which has non-renormalizable couplings to fermions of the form $\bar{\mathbf{5}} \mathbf{10} H_{\bar{\mathbf{5}}} \Sigma / \Lambda$. The Yukawa couplings of the down-type quark and charged leptons then arise from the two $SU(5) \times U(1) \times Z_3 \times Z_4$ invariant superpotential terms ².

$$\begin{aligned}
W_d &= \left(\mathbf{10}_i (C_1)_{ij} \bar{\mathbf{5}}_j H_{\bar{\mathbf{5}}} + \frac{\Sigma}{\Lambda} \mathbf{10}_i (C_2)_{ij} \bar{\mathbf{5}}_j H_{\bar{\mathbf{5}}} \right) \left(\frac{\Theta_1}{\Lambda} \right)^{F(\mathbf{10}_i) + F(\bar{\mathbf{5}}_j)} \left(\frac{\Theta_2}{\Lambda} \right)^{[Z_3(\mathbf{10}_i) + Z_3(\bar{\mathbf{5}}_j)]} \\
&\quad \times \left(\frac{\Theta_3}{\Lambda} \right)^{[Z_4(\mathbf{10}_i) + Z_4(\bar{\mathbf{5}}_j)]},
\end{aligned} \tag{33}$$

²The $\mathbf{75}$ could in principle also give a contribution in the up sector. However, following Ref. [32] we neglect such a term since it is not needed to reproduce the up-type quark masses.

which, after the scalar components of Σ acquires a VEV, lead to:

$$\begin{aligned} (\mathbf{Y}_d)_{ij} &= ((C_1)_{ij} + \kappa (C_2)_{ij}) \lambda^{F(Q_i)+F(D_j^c)+[Z_3(Q_i)+Z_3(D_j^c)]+[Z_4(Q_i)+Z_4(D_j^c)]}, \\ (\mathbf{Y}_e)_{ij} &= ((C_1)_{ij} - 3\kappa (C_2)_{ij}) \lambda^{F(Q_j)+F(D_i^c)+[Z_3(Q_j)+Z_3(D_i^c)]+[Z_4(Q_j)+Z_4(D_i^c)]}, \end{aligned} \quad (34)$$

where $\kappa = \langle \Sigma \rangle / \Lambda$, which breaks the transposition relation between \mathbf{Y}_d and \mathbf{Y}_e and can explain the difference between down-type quarks and charged lepton masses. In our numerical fits, we take $\kappa = 0.3$ for illustration and find that realistic values for down-type quarks and charged lepton masses can be reproduced. The superpotential for the up-type quark mass is

$$W_u = \mathbf{10}_i (C_3)_{ij} \mathbf{10}_j H_5 \left(\frac{\Theta_1}{\Lambda} \right)^{F(\mathbf{10}_i)+F(\mathbf{10}_j)} \left(\frac{\Theta_2}{\Lambda} \right)^{[Z_3(\mathbf{10}_i)+Z_4(\mathbf{10}_j)]} \left(\frac{\Theta_3}{\Lambda} \right)^{[Z_4(\mathbf{10}_i)+Z_4(\mathbf{10}_j)]}, \quad (35)$$

where one has $(C_3)_{ij} = (C_3)_{ji}$ due to the constraint of the $SU(5)$ gauge symmetry. Then one can express the effective Yukawa couplings for the up-type quark in terms of the flavor symmetry charges as

$$(\mathbf{Y}_u)_{ij} = (C_3)_{ij} \lambda^{F(Q_i)+F(Q_j)+[Z_3(Q_i)+Z_3(Q_j)]+[Z_4(Q_i)+Z_4(Q_j)]} \quad (36)$$

With the assignments dictated by Eq.(32), one has the following patterns for the up- and down-type quark mass matrices,

$$M_u \sim \begin{pmatrix} \lambda^7 & \lambda^5 & \lambda^7 \\ \lambda^5 & \lambda^3 & \lambda^5 \\ \lambda^7 & \lambda^5 & 1 \end{pmatrix} v_u, \quad M_d \sim \begin{pmatrix} \lambda^8 & \lambda^7 & \lambda^7 \\ \lambda^6 & \lambda^5 & \lambda^5 \\ \lambda^8 & \lambda^7 & \lambda^3 \end{pmatrix} v_d, \quad (37)$$

which yield

$$|V_{us}| \sim \lambda^2, \quad |V_{cb}| \sim \lambda^2, \quad |V_{ub}| \sim \lambda^4. \quad (38)$$

We note that both the up and down quark sector contribute λ^2 to the mixing element $|V_{us}|$, therefore an accidental enhancement of $\mathcal{O}(\lambda^{-1})$ among the undetermined order one coefficients $(C_1)_{ij}$, $(C_2)_{ij}$ and $(C_3)_{ij}$ is required in order to describe the correct Cabibbo angle. The remaining CKM mixing angles $|V_{cb}|$ and $|V_{ub}|$ arise solely from the diagonalization of the down-type quark mass matrix M_d . In addition, the pattern given by Eq.(37) leads to the following quark mass scalings:

$$\begin{aligned} m_u &\sim \lambda^7 v_u, & m_c &\sim \lambda^3 v_u, & m_t &\sim v_u, \\ m_d &\sim \lambda^8 v_d, & m_s &\sim \lambda^5 v_d, & m_b &\sim \lambda^3 v_d, \end{aligned} \quad (39)$$

which describe the experimental data satisfactorily. Note that the second term in Eq.(33) accounts for the mass difference between the down-type quarks and charged leptons, allowing for an acceptable charged fermion mass pattern.

In order to see in a quantitative way how well the model describes the observed values of the fermion masses and mixings, we perform a numerical analysis, within three independent

different seeding methods: namely flat, Gaussian and exponential distributions. The modulus of the undetermined order one coefficients are taken to be random numbers with flat, Gaussian and exponential distributions in turn, the corresponding phases are varied between 0 and 2π . The probability density function $f(x)$ of the three distributions is well known

$$f(x) = \begin{cases} \frac{1}{b-a} & a \leq x \leq b \\ 0 & x < a \text{ or } x > b \end{cases} . \quad (40)$$

For flat distribution, we take $a = 1/3$ and $b = 3$ for illustration in the present work. In the case of gaussian distribution,

$$f(x) = \frac{a}{\sqrt{2\pi}\sigma} e^{-\frac{(x-\mu)^2}{2\sigma^2}} . \quad (41)$$

We set the mean $\mu = 1$ and the standard deviation $\sigma = 1.5$ in our numerical calculation. The probability density function for the exponential distribution is

$$f(x) = \begin{cases} \lambda e^{-\lambda x} & x \geq 0 \\ 0 & x < 0 \end{cases} . \quad (42)$$

Its statistic mean is $1/\lambda$, and λ is taken to be 1 as a typical value for numerical simulation. To the extent that our results are independent of the choice of seeding method, they are robust and not simply an artifact of the choice of the seed function.

The coefficients $(C_1)_{ij}$, $(C_2)_{ij}$, $(C_3)_{ij}$ and $(y_\nu)_{ij}$ are treated as random complex numbers with arbitrary phases and absolute value in the interval of $[1/3, 3]$. Then we calculate the quark and lepton masses as well as the CKM and lepton mixing matrix entries which are required to lie in the experimentally allowed ranges. The numerical results are found to be nicely consistent with the above theoretical estimates and qualitative discussions. Since the flavor parameters of the quark sector are precisely measured, here we focus on the neutrino sector. As an example the predicted distributions for the light neutrino masses and atmospheric mixing parameter are shown in Fig. 1. The light neutrino masses follow the normal hierarchy pattern and, for all the points produced, though all non-vanishing, they are rather tiny, with most of the expected m_1 values below 0.015 eV. As to the mixing angles, no specific values of θ_{12} and θ_{13} are favored within 3σ , and hence they are not shown in the figure³. In contrast, however, the atmospheric neutrino mixing angle θ_{23} obeys $\sin^2 \theta_{23} < 1/2$, which means that non-maximal θ_{23} values are preferred, as indicated by current neutrino oscillation global analyses post-Neutrino 2012 [6–8], with a preference for the first octant. This has been one of our motivations for introducing BL_1 mixing pattern, which leads to $\sin \theta_{23}$ of order λ at leading order.

The rare process, neutrinoless double beta decay ($0\nu 2\beta$), constitutes an important probe for the Majorana nature of neutrino and lepton number violation [33], a sizable number of new experiments are currently running, under construction, or in the planing phase. The

³Similarly, we can hardly see any specific preferred pattern for the CP violating phases δ , φ_1 and φ_2 , hence, as before, these are not shown.

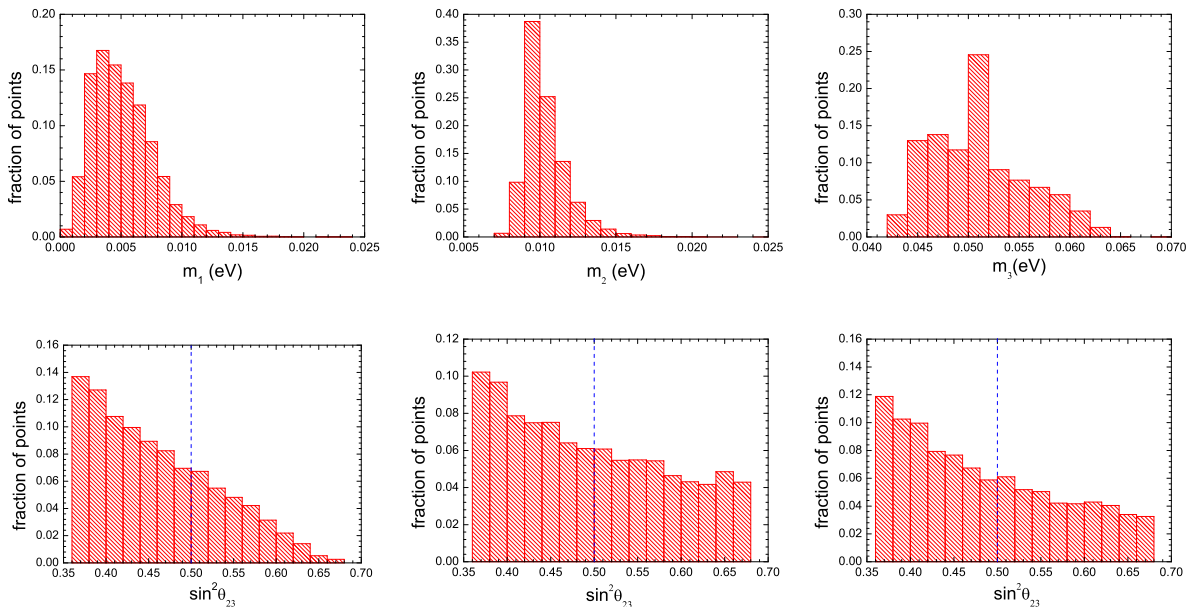


Figure 1: Histograms for the distribution of light neutrino masses and atmospheric neutrino mixing parameter in the BL_1 model. In the second row, the left, middle and right panels are obtained with using different seed procedures for the order one Yukawa coefficients, namely flat, exponential and Gaussian, respectively, from left to right.

histogram for the distribution of the effective $0\nu 2\beta$ -decay mass $|m_{ee}|$ and its correlation with the lightest neutrino mass m_1 are given in Fig. 2. We also show the future sensitivity on the lightest neutrino mass of 0.2 eV from the KATRIN experiment [34]. The horizontal lines represent the sensitivities of the future $0\nu 2\beta$ -decay experiments CUORE [35] and MAJORANA [36]/GERDA III [37], which are approximately 18 meV and 12 meV respectively. Clearly the expected effective mass $|m_{ee}|$ is predicted to be far below the sensitivities of the planned $0\nu 2\beta$ experiments. The reason for this is the strong destructive interference amongst the three light neutrinos, as seen in the right panel. As a result, if $0\nu 2\beta$ decay will be detected in the near future, our construction would be ruled out.

To keep our discussion as generic as possible, we describe the light neutrino masses by the effective higher-dimensional Weinberg operators as shown in Eq.(3) and Eq.(12), which could come from the so-called type I seesaw mechanism, by integrating out the right-handed neutrinos. It is interesting to note that $U(1)$ flavor symmetry models have particularly simple factorization properties [24, 25]: our various predictions for the light neutrino parameters given above, are independent of the $U(1)$ charge assignments of the right-handed neutrinos. For example, suppose we introduce three right-handed neutrinos transforming under the flavor symmetry $U(1) \times Z_3 \times Z_4$ as follows:

$$N_1^c : (n_1, 0, 1), \quad N_2^c : (n_2, 0, 3), \quad N_3^c : (n_3, 2, 2), \quad (43)$$

where n_i , which are positive integers denoting the $U(1)$ charges of the heavy Majorana neutrinos. Then one can straightforwardly read out the Dirac neutrino mass matrix M_D and

the Majorana mass matrix M_N of the right-handed neutrinos,

$$M_D \sim \begin{pmatrix} \lambda^{5+n_1} & \lambda^{7+n_2} & \lambda^{5+n_3} \\ \lambda^{8+n_1} & \lambda^{6+n_2} & \lambda^{4+n_3} \\ \lambda^{4+n_1} & \lambda^{2+n_2} & \lambda^{4+n_3} \end{pmatrix} v_u, \quad M_N \sim \begin{pmatrix} \lambda^{2+2n_1} & \lambda^{n_1+n_2} & \lambda^{5+n_1+n_3} \\ \lambda^{n_1+n_2} & \lambda^{2+2n_2} & \lambda^{3+n_2+n_3} \\ \lambda^{5+n_1+n_3} & \lambda^{3+n_2+n_3} & \lambda^{1+2n_3} \end{pmatrix} \Lambda. \quad (44)$$

The resulting effective light neutrino mass matrix is given by the seesaw formula

$$M_\nu = -M_D M_N^{-1} M_D^T \sim \begin{pmatrix} \lambda^9 & \lambda^8 & \lambda^7 \\ \lambda^8 & \lambda^7 & \lambda^7 \\ \lambda^7 & \lambda^7 & \lambda^6 \end{pmatrix} \frac{v_u^2}{\Lambda}. \quad (45)$$

This is the same as obtained in the above effective approach given in Eq.(29) except that the smallest element $(M_\nu)_{11}$ is of order λ^9 instead of λ^{12} , both of them are too small to affect the predictions for the neutrino oscillation parameters. We get the same light neutrino masses in Eq.(30) and the same neutrino mixing angles in Eq.(31) as in the above effective Weinberg operator neutrino mass generation. We would like to emphasize again that the predictions for the neutrino masses and mixing parameters are independent of the charges n_i , which drop out in the seesaw formula for the light neutrino mass matrix. However, different values of the charges n_i obviously give rise to different Dirac neutrino Yukawa coupling $Y_\nu \equiv M_D/v_u$. As a result, the predictions for charged lepton flavor violation (LFV) processes such as $\mu \rightarrow e\gamma$, $\tau \rightarrow \mu\gamma$ and $\mu \rightarrow 3e$ are quite different [38]. Recalling that the branching ratio of the LFV process is generally proportional to Y_ν^4 , the stringent bound on LFV, in particular from $\mu \rightarrow e\gamma$, can be easily satisfied for only slightly large n_i [38] while keeping the predictions for neutrino parameters intact.

4 Model for BL_2 mixing

As explained in section 2, the order one atmospheric neutrino mixing $\sin\theta_{23} \sim 1$ generically implies that the corresponding masses of ν_2 and ν_3 are of the same order of magnitude within pure $U(1)$ family symmetry schemes. As a result, the neutrino mass spectrum is quasi-degenerate and strong fine-tuning is required in order to account for the measured mass-squared differences Δm_{sol}^2 and Δm_{atm}^2 . Furthermore, the renormalization group evolution effects could drastically enhance the neutrino mixing angles due to the degeneracy, so that the BL_2 texture would be spoiled at the electroweak scale. This can be avoided by extending the flavor symmetry to $U(1) \times Z_m$. Now the whole flavor symmetry is chosen to be $U(1) \times Z_2$, the lepton fields carry the following $U(1) \times Z_2$ charges:

$$\begin{aligned} L_1 &: (3, 0), & L_2 &: (3, 1), & L_3 &: (2, 0), \\ E_1^c &: (4, 0), & E_2^c &: (2, 1), & E_3^c &: (0, 1), \end{aligned} \quad (46)$$

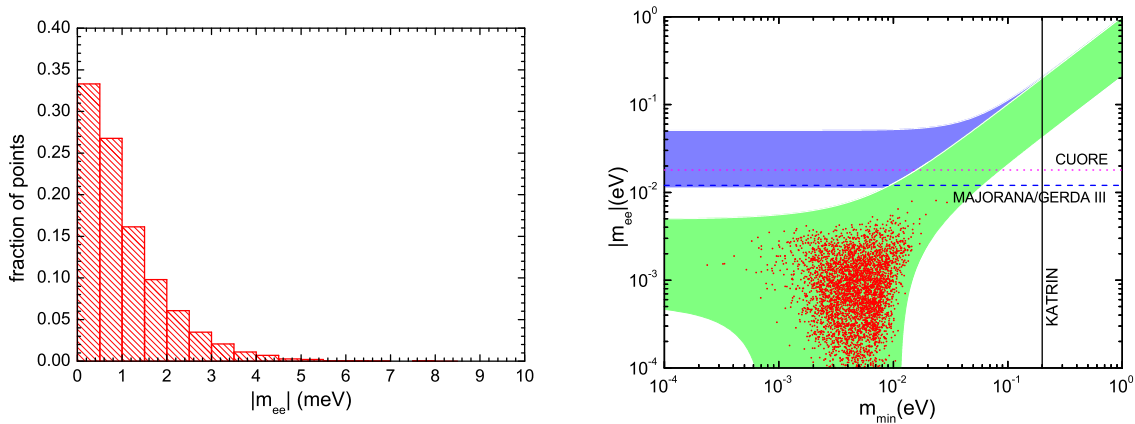


Figure 2: Histogram of the effective mass $|m_{ee}|$ (left panel) and the scatter plot of $|m_{ee}|$ versus the lightest neutrino mass m_1 (right panel) for the BL_1 model. The colored bands represent the regions for the 3σ ranges of the oscillation parameters in the normal and inverted neutrino mass spectrum respectively. The future sensitivity of 0.2 eV of the KATRIN experiment is shown by the vertical solid line, while the future expected bounds on $|m_{ee}|$ from the CUORE and MAJORANA/GERDA III experiments are represented by horizontal lines.

Then the light neutrino mass matrix is given, apart from the order one coefficients, as

$$M_\nu \sim \begin{pmatrix} \lambda^6 & \lambda^7 & \lambda^5 \\ \lambda^7 & \lambda^6 & \lambda^6 \\ \lambda^5 & \lambda^6 & \lambda^4 \end{pmatrix} \frac{v_u^2}{\Lambda}, \quad (47)$$

which yields

$$m_1 \sim \lambda^6 v_u^2/\Lambda, \quad m_2 \sim \lambda^6 v_u^2/\Lambda, \quad m_3 \sim \lambda^4 v_u^2/\Lambda \quad (48)$$

One sees that the first two light neutrinos are quasi-degenerate in this model, and their masses are suppressed by $\mathcal{O}(\lambda^2)$ with respect to the third one. This prediction is consistent with the observation that the solar neutrino mass difference Δm_{sol}^2 is much smaller than the atmospheric neutrino mass difference Δm_{atm}^2 . Moreover, the neutrino mass spectrum is predicted to be of the normal hierarchy type here, the same as in the previous BL_1 model (this is also confirmed our numerical analysis). The next generation of higher precision neutrino oscillation experiments is designed to be able to measure neutrino mass hierarchy and the CP phase [39]. Should the latter be determined to be of the inverted type by future experiments, both of our models would be ruled out. On the other hand, the charged lepton mass matrix takes the following form:

$$M_e \sim \begin{pmatrix} \lambda^7 & \lambda^6 & \lambda^4 \\ \lambda^8 & \lambda^5 & \lambda^3 \\ \lambda^6 & \lambda^5 & \lambda^3 \end{pmatrix} v_d, \quad (49)$$

which has a “lopsided” structure, a large 2-3 mixing arises from the diagonalization of M_e . Obviously it also gives the correct order of magnitude for the charged lepton mass ratios.

Combining the contribution from both the neutrino and the charged lepton mass matrices diagonalization, the leptonic mixing angles are given by

$$\sin \theta_{12} \sim \lambda, \quad \sin \theta_{13} \sim \lambda \quad \sin \theta_{23} \sim 1. \quad (50)$$

This is exactly the desired BL_2 mixing pattern, Eq. (2). Here we would like to point out that since the Super-Kamiokande data indicated large atmospheric neutrino mixing, perhaps even maximal [41], there have been several attempts to account for the large atmospheric neutrino mixing $\sin \theta_{23} \sim 1$ in terms of Abelian flavor symmetries [40]. However, it was usually assumed that the reactor angle θ_{13} was rather small, at most of order λ^2 at that time [42]. In contrast, in our construction the consistency between large $\sin \theta_{23}$ and sizeable $\sin \theta_{13}$ mixing angles emerges naturally.

In what follows, we extend the model to include quarks within the $SU(5)$ unified framework. The fields Q_i and U_i^c together with E_i^c within the same generation fill out the $\mathbf{10}$ representation, while D_i^c and the left-handed lepton doublet L_i make up the $\bar{\mathbf{5}}$ representation. As a result, we can determine the transformation properties of the quark fields under the $U(1) \times Z_2$ flavor symmetry as follows:

$$\begin{aligned} Q_1 &: (4, 0), & Q_2 &: (2, 1), & Q_3 &: (0, 1), \\ U_1^c &: (4, 0), & U_2^c &: (2, 1), & U_3^c &: (0, 1), \\ D_1^c &: (3, 0), & D_2^c &: (3, 1), & D_3^c &: (2, 0). \end{aligned} \quad (51)$$

The up and down quark mass matrices can be determined in a straightforward way as follows:

$$M_u \sim \begin{pmatrix} \lambda^8 & \lambda^7 & \lambda^5 \\ \lambda^7 & \lambda^4 & \lambda^2 \\ \lambda^5 & \lambda^2 & 1 \end{pmatrix} v_u, \quad M_d \sim \begin{pmatrix} \lambda^7 & \lambda^8 & \lambda^6 \\ \lambda^6 & \lambda^5 & \lambda^5 \\ \lambda^4 & \lambda^3 & \lambda^3 \end{pmatrix} v_d. \quad (52)$$

which lead to

$$\begin{aligned} |V_{us}| &\sim \lambda, & |V_{cb}| &\sim \lambda^2, & |V_{ub}| &\sim \lambda^3, \\ \frac{m_u}{m_c} &\sim \lambda^4, & \frac{m_c}{m_t} &\sim \lambda^4, & \frac{m_d}{m_s} &\sim \lambda^2, & \frac{m_s}{m_b} &\sim \lambda^2, & \frac{m_b}{m_t} &\sim \lambda^3, \end{aligned} \quad (53)$$

which are in excellent agreement with observed quark mass hierarchies and CKM mixing angles. As in section 3, we perform a numerical simulation of the expected neutrino oscillation parameters. In Fig. 3 we display the resulting histograms for the neutrino mass eigenvalues⁴. As expected on the basis of the qualitative estimate in Eq. (48), the light neutrino mass spectrum is normal hierarchy, the degenerate spectrum being strongly disfavored, and almost all the generated points lie in the region of the lightest neutrino mass m_1 smaller than 0.015 eV. The neutrinoless double beta decay predictions are shown in Fig. 4. One sees that, in contrast with the BL_1 case, although the effective mass $|m_{ee}|$ is also quite small, with

⁴Insofar as the neutrino mixing angles θ_{ij} and CP phases δ , φ_1 and φ_2 are concerned, we do not obtain any special predicted pattern, hence the results are not displayed.

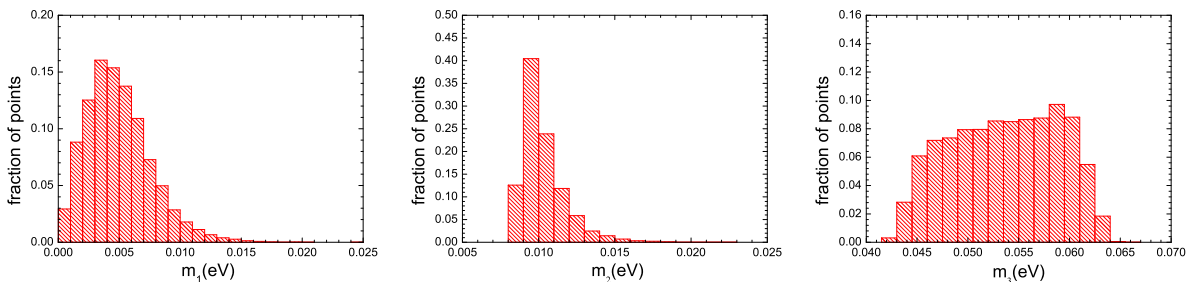


Figure 3: Light neutrino masses in the BL_2 model.

$|m_{ee}|$ around 5 meV preferred, there is a small portion of the parameter space of the model where the predictions for $|m_{ee}|$ approach the future experimental sensitivities. However, the points above the sensitivity limits on next generation experiments are statistically rather low. Therefore, if the signal of $0\nu 2\beta$ decay would be observed by upcoming experiments, the present BL_2 model would also be ruled out, although not completely. We expect that the future $0\nu 2\beta$ -decay experiments with sensitivity much higher than MAJORANA/GERDA III should be able to provide a better test of the model.

Now we turn to the seesaw realization of this model, the assignments for the right-handed neutrinos are not unique. As an example, we can introduce three right-handed neutrinos transforming as

$$N_1^c : (n_1, 0), \quad N_2^c : (n_2, 1), \quad N_3^c : (n_3, 0). \quad (54)$$

Then we obtain the Dirac neutrino mass matrix M_D as well as the right-handed neutrino mass matrix M_N ,

$$M_D \sim \begin{pmatrix} \lambda^{3+n_1} & \lambda^{4+n_2} & \lambda^{3+n_3} \\ \lambda^{4+n_1} & \lambda^{3+n_2} & \lambda^{4+n_3} \\ \lambda^{2+n_1} & \lambda^{3+n_2} & \lambda^{2+n_3} \end{pmatrix} v_u, \quad M_N \sim \begin{pmatrix} \lambda^{2n_1} & \lambda^{1+n_1+n_2} & \lambda^{n_1+n_3} \\ \lambda^{1+n_1+n_2} & \lambda^{2n_2} & \lambda^{1+n_2+n_3} \\ \lambda^{n_1+n_3} & \lambda^{1+n_2+n_3} & \lambda^{2n_3} \end{pmatrix} \Lambda. \quad (55)$$

The effective light neutrino mass matrix is given by the seesaw relation

$$M_\nu = -M_D M_M^{-1} M_D^T \sim \begin{pmatrix} \lambda^6 & \lambda^7 & \lambda^5 \\ \lambda^7 & \lambda^6 & \lambda^6 \\ \lambda^5 & \lambda^6 & \lambda^4 \end{pmatrix} \frac{v_u^2}{\Lambda} \quad (56)$$

This is exactly Eq.(47), consequently the predictions for neutrino parameters in Eq.(48) and Eq.(50) remain, note that dependence on the right-handed neutrino charges n_i drops out. However, different values of the charges n_i result in different LFV predictions, and the model would be less constrained for slightly large n_i assignments [38].

5 Conclusions

The recent neutrino oscillation experimental highlights: (i) rather large value of reactor mixing angle θ_{13} and (ii) indication of significant deviation of the atmospheric neutrino

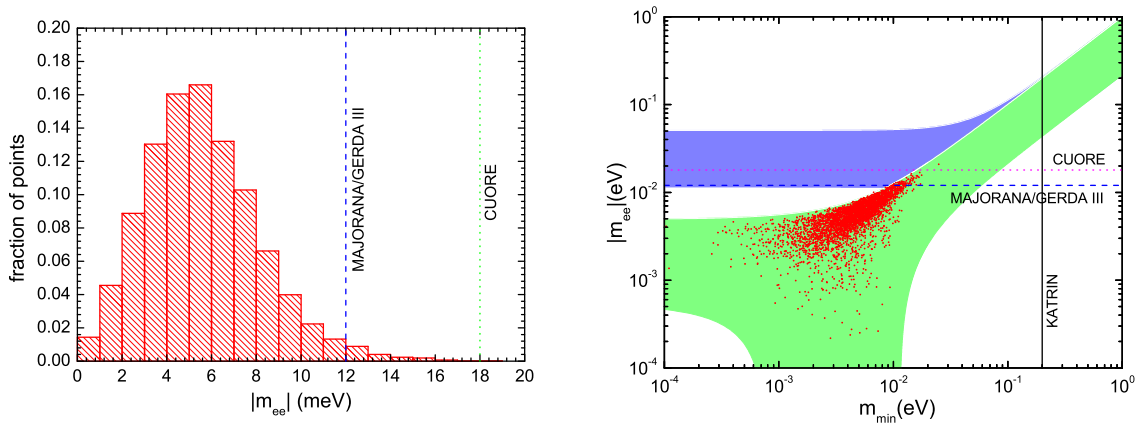


Figure 4: Histogram of the effective mass $|m_{ee}|$ (left panel) and the $|m_{ee}|$ versus the lightest neutrino mass m_1 correlation (right panel) predicted in the BL_2 model.

mixing angle θ_{23} from maximality may change our theoretical approach for constructing neutrino mass models. In this paper, we study the Wolfenstein-like mixing schemes: BL_1 mixing in which $\sin \theta_{12} \sim \lambda$, $\sin \theta_{13} \sim \lambda$, $\sin \theta_{23} \sim \lambda$, and BL_2 mixing, in which $\sin \theta_{12} \sim \lambda$, $\sin \theta_{13} \sim \lambda$, $\sin \theta_{23} \sim 1$. The largish θ_{13} can be naturally accommodated in both of them, the two mixing patterns differ in the order of magnitude of $\sin \theta_{23}$, the BL_1 texture is favored for θ_{23} in the first octant, while BL_2 is preferred for the second octant θ_{23} . In order to produce the BL_1 mixing without invoking unnatural cancellation, the Abelian flavor symmetry should be $U(1) \times Z_m \times Z_n$ with the parity of m and n being opposite. A concrete model based on $U(1) \times Z_3 \times Z_4$ family symmetry is constructed, where the light neutrino mass hierarchy $m_2/m_3 \sim \lambda$ is realized due to the discrete nature of $Z_3 \times Z_4$. The ratio $\Delta m_{\text{sol}}^2/\Delta m_{\text{atm}}^2$ is expected to be of order λ^2 in this model, which is in good agreement with experimental data in contrast with conventional $U(1)$ or $U(1) \times Z_m$ flavor symmetry constructions. Furthermore, the model is embedded into the $SU(5)$ grand unified theory to describe the quark masses and mixing simultaneously. As for the BL_2 mixing, it can be reproduced within the framework of pure $U(1)$ flavor symmetry. However, the light neutrino mass spectrum is expected to be quasi-degenerate, hence fine-tuning of the neutrino mass parameters is needed in order to achieve the observed mass-squared differences. To improve upon this situation, the family symmetry is enlarged to $U(1) \times Z_2$, which gives rise to both large atmospheric neutrino mixing $\sin \theta_{23} \sim 1$ and hierarchical neutrino masses. The model is extended to $SU(5)$ grand unified theory as well.

We show that both models can give a successful description of the observed quark and lepton masses and mixing angles, and the numerical results are nicely in agreement with the theoretical estimates and the qualitative discussions. The light neutrinos are normal mass hierarchy in both models, quasi-degenerate spectrum is strongly disfavored. If the next generation high precision neutrino oscillation experiments determine that the neutrino mass spectrum is inverted hierarchy, both our constructions will be ruled out. The present

framework can not predict the CP violating phases δ , φ_1 and φ_2 . The $0\nu 2\beta$ -decay effective mass $|m_{ee}|$ is predicted to be rather small in both constructions, substantial part of the data are below the sensitivity of future experiments except for a region of the BL_2 model indicated in Fig. 4. Therefore future $0\nu 2\beta$ -decay experiments such as CUORE, MAJORANA and GERDA III will provide another important test of the present models.

Acknowledgements

This work was supported by the National Natural Science Foundation of China under Grant No 10905053, Chinese Academy KJCX2-YW-N29, DFG grant WI 2639/4-1 and the 973 project with Grant No. 2009CB825200; by the Spanish MINECO under grants FPA2011-22975 and MULTIDARK CSD2009-00064 (Consolider-Ingenio 2010 Programme), by Prometeo/2009/091 (Generalitat Valenciana), and by the EU ITN UNILHC PITN-GA-2009-237920. S.M. is supported by a Juan de la Cierva grant.

References

- [1] F. P. An *et al.* [DAYA-BAY Collaboration], Phys. Rev. Lett. **108**, 171803 (2012) [arXiv:1203.1669 [hep-ex]].
- [2] J. K. Ahn *et al.* [RENO Collaboration], Phys. Rev. Lett. **108**, 191802 (2012) [arXiv:1204.0626 [hep-ex]].
- [3] Y. Abe *et al.* [DOUBLE-CHOOZ Collaboration], Phys. Rev. Lett. **108**, 131801 (2012) [arXiv:1112.6353 [hep-ex]].
- [4] K. Abe *et al.* [T2K Collaboration], Phys. Rev. Lett. **107**, 041801 (2011) [arXiv:1106.2822 [hep-ex]].
- [5] P. Adamson *et al.* [MINOS Collaboration], Phys. Rev. Lett. **107**, 181802 (2011) [arXiv:1108.0015 [hep-ex]].
- [6] D. V. Forero, M. Tortola and J. W. F. Valle, Phys. Rev. D **86**, 073012 (2012) [arXiv:1205.4018 [hep-ph]].
- [7] G. L. Fogli, E. Lisi, A. Marrone, D. Montanino, A. Palazzo and A. M. Rotunno, Phys. Rev. D **86**, 013012 (2012) [arXiv:1205.5254 [hep-ph]].
- [8] M. C. Gonzalez-Garcia, M. Maltoni, J. Salvado and T. Schwetz, arXiv:1209.3023 [hep-ph].
- [9] P. F. Harrison, D. H. Perkins and W. G. Scott, Phys. Lett. B **530** (2002) 167 [hep-ph/0202074]; P. F. Harrison and W. G. Scott, Phys. Lett. B **535** (2002) 163

[hep-ph/0203209]; Z. -z. Xing, Phys. Lett. B **533** (2002) 85 [hep-ph/0204049]; X. G. He and A. Zee, Phys. Lett. B **560**, 87 (2003) [hep-ph/0301092].

- [10] H. -J. He and F. -R. Yin, Phys. Rev. D **84**, 033009 (2011) [arXiv:1104.2654 [hep-ph]]; Y. Shimizu, M. Tanimoto and A. Watanabe, Prog. Theor. Phys. **126**, 81 (2011) [arXiv:1105.2929 [hep-ph]]; J. -M. Chen, B. Wang and X. -Q. Li, Phys. Rev. D **84**, 073002 (2011) [arXiv:1106.3133 [hep-ph]]; Z. -z. Xing, [arXiv:1106.3244 [hep-ph]]; E. Ma and D. Wegman, Phys. Rev. Lett. **107**, 061803 (2011) [arXiv:1106.4269 [hep-ph]]; Y. -j. Zheng and B. -Q. Ma, [arXiv:1106.4040 [hep-ph]]; X. -G. He and A. Zee, Phys. Rev. D **84**, 053004 (2011) [arXiv:1106.4359 [hep-ph]]; S. Zhou, Phys. Lett. B **704**, 291 (2011) [arXiv:1106.4808 [hep-ph]]; N. Haba and R. Takahashi, Phys. Lett. B **702**, 388 (2011) [arXiv:1106.5926 [hep-ph]]; D. Meloni, JHEP **1110**, 010 (2011) [arXiv:1107.0221 [hep-ph]]; S. Morisi, K. M. Patel and E. Peinado, Phys. Rev. D **84**, 053002 (2011) [arXiv:1107.0696 [hep-ph]]; W. Chao and Y. -j. Zheng, arXiv:1107.0738 [hep-ph]; H. Zhang and S. Zhou, Phys. Lett. B **704**, 296 (2011) [arXiv:1107.1097 [hep-ph]]; X. Chu, M. Dhen and T. Hambye, JHEP **1111**, 106 (2011) [arXiv:1107.1589 [hep-ph]]; P. S. Bhupal Dev, R. N. Mohapatra and M. Severson, Phys. Rev. D **84**, 053005 (2011) [arXiv:1107.2378 [hep-ph]]; S. Antusch and V. Maurer, Phys. Rev. D **84**, 117301 (2011) [arXiv:1107.3728 [hep-ph]]; W. Rodejohann, H. Zhang and S. Zhou, Nucl. Phys. B **855**, 592 (2012) [arXiv:1107.3970 [hep-ph]]; Y. H. Ahn, H. -Y. Cheng and S. Oh, [arXiv:1107.4549 [hep-ph]]; S. F. King and C. Luhn, JHEP **1109**, 042 (2011) [arXiv:1107.5332 [hep-ph]]; Q. -H. Cao, S. Khalil, E. Ma and H. Okada, Phys. Rev. D **84**, 071302 (2011) [arXiv:1108.0570 [hep-ph]]; D. Marzocca, S. T. Petcov, A. Romanino and M. Spinrath, JHEP **1111**, 009 (2011) [arXiv:1108.0614 [hep-ph]]; S. -F. Ge, D. A. Dicus and W. W. Repko, [arXiv:1108.0964 [hep-ph]]; Riazuddin, arXiv:1108.1469 [hep-ph]; F. Bazzocchi, [arXiv:1108.2497 [hep-ph]]; T. Araki and C. -Q. Geng, JHEP **1109**, 139 (2011) [arXiv:1108.3175 [hep-ph]]; S. Antusch, S. F. King, C. Luhn and M. Spinrath, Nucl. Phys. B **856**, 328 (2012) [arXiv:1108.4278 [hep-ph]]; H. Fritzsch, Z. -z. Xing and S. Zhou, JHEP **1109**, 083 (2011) [arXiv:1108.4534 [hep-ph]]; R. N. Mohapatra and M. K. Parida, Phys. Rev. D **84**, 095021 (2011) [arXiv:1109.2188 [hep-ph]]; A. Rashed and A. Datta, [arXiv:1109.2320 [hep-ph]]; P. O. Ludl, S. Morisi and E. Peinado, [arXiv:1109.3393 [hep-ph]]; N. Okada and Q. Shafi, arXiv:1109.4963 [hep-ph]; A. Aranda, C. Bonilla and A. D. Rojas, [arXiv:1110.1182 [hep-ph]]; G. -J. Ding, L. L. Everett and A. J. Stuart, Nucl. Phys. B **857**, 219 (2012) [arXiv:1110.1688 [hep-ph]]; D. Meloni, arXiv:1110.5210 [hep-ph]; S. Dev, S. Gupta, R. R. Gautam and L. Singh, Phys. Lett. B **706**, 168 (2011) [arXiv:1111.1300 [hep-ph]]; K. N. Deepthi, S. Gollu and R. Mohanta, arXiv:1111.2781 [hep-ph]; A. Rashed, arXiv:1111.3072 [hep-ph]; W. Buchmuller, V. Domcke and K. Schmitz, arXiv:1111.3872 [hep-ph]; I. d. M. Varzielas, arXiv:1111.3952 [hep-ph]; R. de Adelhart Toorop, F. Feruglio and C. Hagedorn, [arXiv:1112.1340 [hep-ph]]; S. F. King and C. Luhn, arXiv:1112.1959 [hep-ph]; S. Gupta, A. S. Joshipura and K. M. Patel, arXiv:1112.6113 [hep-ph]. H. Ishimori and T. Kobayashi, Phys. Rev.

- D **85**, 125004 (2012) [arXiv:1201.3429 [hep-ph]]; S. Dev, R. R. Gautam and L. Singh, Phys. Lett. B **708**, 284 (2012) [arXiv:1201.3755 [hep-ph]]. P. S. Bhupal Dev, B. Dutta, R. N. Mohapatra and M. Severson, Phys. Rev. D **86**, 035002 (2012) [arXiv:1202.4012 [hep-ph]]; X. Zhang and B. -Q. Ma, Phys. Lett. B **710**, 630 (2012) [arXiv:1202.4258 [hep-ph]]; I. K. Cooper, S. F. King and C. Luhn, JHEP **1206**, 130 (2012) [arXiv:1203.1324 [hep-ph]]; Z. -Z. Xing, Chin. Phys. C **36**, 281 (2012) [arXiv:1203.1672 [hep-ph]]; Y. -L. Wu, Phys. Lett. B **714**, 286 (2012) [arXiv:1203.2382 [hep-ph]]; Y. BenTov and A. Zee, Phys. Lett. B **714**, 80 (2012) [arXiv:1203.2671 [hep-ph]]; H. -J. He and X. -J. Xu, arXiv:1203.2908 [hep-ph]; D. Meloni, JHEP **1205**, 124 (2012) [arXiv:1203.3126 [hep-ph]]. Y. H. Ahn and S. K. Kang, arXiv:1203.4185 [hep-ph]; I. d. M. Varzielas and G. G. Ross, arXiv:1203.6636 [hep-ph]; D. Hernandez and A. Y. .Smirnov, arXiv:1204.0445 [hep-ph]; C. Hagedorn and D. Meloni, Nucl. Phys. B **862**, 691 (2012) [arXiv:1204.0715 [hep-ph]]; M. Fukugita, Y. Shimizu, M. Tanimoto and T. T. Yanagida, Phys. Lett. B **716**, 294 (2012) [arXiv:1204.2389 [hep-ph]]; S. Zhou, arXiv:1205.0761 [hep-ph]; C. Hagedorn, S. F. King and C. Luhn, arXiv:1205.3114 [hep-ph]; G. Altarelli, F. Feruglio, L. Merlo and E. Stamou, JHEP **1208**, 021 (2012) [arXiv:1205.4670 [hep-ph]]; G. Altarelli, F. Feruglio and L. Merlo, arXiv:1205.5133 [hep-ph]; A. Meroni, S. T. Petcov and M. Spinrath, arXiv:1205.5241 [hep-ph]; M. J. Baker, J. Bordes, H. M. Chan and S. T. Tsou, arXiv:1206.0199 [hep-ph]; X. Zhang and B. -Q. Ma, arXiv:1206.0519 [hep-ph]; E. Ma, A. Natale and A. Rashed, Int. J. Mod. Phys. A **27**, 1250134 (2012) [arXiv:1206.1570 [hep-ph]]; N. Haba and R. Takahashi, arXiv:1206.2793 [hep-ph]; A. G. Dias, A. C. B. Machado and C. C. Nishi, arXiv:1206.6362 [hep-ph]; P. M. Ferreira, W. Grimus, L. Lavoura and P. O. Ludl, arXiv:1206.7072 [hep-ph]; G. Altarelli, F. Feruglio, I. Masina and L. Merlo, arXiv:1207.0587 [hep-ph]; Y. BenTov, X. -G. He and A. Zee, arXiv:1208.1062 [hep-ph]; W. Rodejohann and H. Zhang, arXiv:1207.1225 [hep-ph]; R. N. Mohapatra and C. C. Nishi, arXiv:1208.2875 [hep-ph].
- [11] R. d. A. Toorop, F. Feruglio and C. Hagedorn, Phys. Lett. B **703**, 447 (2011) [arXiv:1107.3486 [hep-ph]].
- [12] G. -J. Ding, Nucl. Phys. B **862**, 1 (2012) [arXiv:1201.3279[hep-ph]].
- [13] S. F. King, C. Luhn and A. J. Stuart, arXiv:1207.5741 [hep-ph].
- [14] S. M. Boucenna, S. Morisi, M. Tortola and J. W. F. Valle, Phys. Rev. D **86**, 051301 (2012) [arXiv:1206.2555 [hep-ph]].
- [15] S. F. King, arXiv:1205.0506 [hep-ph]; S. Antusch, C. Gross, V. Maurer and C. Sluka, arXiv:1205.1051 [hep-ph];
- [16] H. Minakata and A. Y. .Smirnov, Phys. Rev. D **70**, 073009 (2004) [hep-ph/0405088]; N. Li and B. -Q. Ma, Eur. Phys. J. C **42**, 17 (2005) [hep-ph/0504161]; N. Qin and B. Q. Ma,

- Phys. Lett. B **702**, 143 (2011) [arXiv:1106.3284 [hep-ph]]; Y. H. Ahn, H. -Y. Cheng and S. Oh, Phys. Rev. D **83**, 076012 (2011) [arXiv:1102.0879 [hep-ph]].
- [17] C. D. Froggatt and H. B. Nielsen, Nucl. Phys. B **147** (1979) 277.
- [18] M. Leurer, Y. Nir and N. Seiberg, Nucl. Phys. B **398**, 319 (1993) [hep-ph/9212278]; Nucl. Phys. B **420**, 468 (1994) [hep-ph/9310320].
- [19] E. Dudas, C. Grojean, S. Pokorski and C. A. Savoy, Nucl. Phys. B **481**, 85 (1996) [hep-ph/9606383]; K. Choi, E. J. Chun and H. D. Kim, Phys. Lett. B **394**, 89 (1997) [hep-ph/9611293]; N. Irges, S. Lavignac and P. Ramond, Phys. Rev. D **58**, 035003 (1998) [hep-ph/9802334]; J. K. Elwood, N. Irges and P. Ramond, Phys. Rev. Lett. **81**, 5064 (1998) [hep-ph/9807228]; J. R. Ellis, G. K. Leontaris and J. Rizos, JHEP **0005**, 001 (2000) [hep-ph/0002263]; H. K. Dreiner and M. Thormeier, Phys. Rev. D **69**, 053002 (2004) [hep-ph/0305270]; H. K. Dreiner, H. Murayama and M. Thormeier, Nucl. Phys. B **729**, 278 (2005) [hep-ph/0312012]; M. -C. Chen, D. R. T. Jones, A. Rajaraman and H. -B. Yu, Phys. Rev. D **78**, 015019 (2008) [arXiv:0801.0248 [hep-ph]]; L. F. Duque, D. A. Gutierrez, E. Nardi and J. Norena, Phys. Rev. D **78**, 035003 (2008) [arXiv:0804.2865 [hep-ph]].
- [20] G. Altarelli, F. Feruglio and L. Merlo, JHEP **0905**, 020 (2009) [arXiv:0903.1940 [hep-ph]].
- [21] L. Wolfenstein, Phys. Rev. Lett. **51**, 1945 (1983).
- [22] J. Schechter and J. W. F. Valle, Phys. Rev. D **22** (1980) 2227; W. Rodejohann and J. W. F. Valle, Phys. Rev. D **84**, 073011 (2011) [arXiv:1108.3484 [hep-ph]].
- [23] H. Fusaoka and Y. Koide, Phys. Rev. D **57**, 3986 (1998) [hep-ph/9712201]; G. Ross and M. Serna, Phys. Lett. B **664**, 97 (2008) [arXiv:0704.1248 [hep-ph]]; Z. -z. Xing, H. Zhang and S. Zhou, Phys. Rev. D **77**, 113016 (2008) [arXiv:0712.1419 [hep-ph]].
- [24] A. Rasin and J. P. Silva, Phys. Rev. D **49**, 20 (1994) [hep-ph/9309240].
- [25] Y. Grossman and Y. Nir, Nucl. Phys. B **448**, 30 (1995) [hep-ph/9502418]; Y. Grossman, Y. Nir and Y. Shadmi, JHEP **9810**, 007 (1998) [hep-ph/9808355]; Y. Nir and Y. Shadmi, JHEP **9905**, 023 (1999) [hep-ph/9902293].
- [26] L. J. Hall and A. Rasin, Phys. Lett. B **315**, 164 (1993), hep-ph/9303303.
- [27] M. B. Green and J. H. Schwarz, Phys. Lett. B **149**, 117 (1984).
- [28] L. E. Ibanez and G. G. Ross, Phys. Lett. B **332**, 100 (1994) [hep-ph/9403338]; P. Binetruy and P. Ramond, Phys. Lett. B **350**, 49 (1995) [hep-ph/9412385]; V. Jain and R. Shrock, Phys. Lett. B **352**, 83 (1995) [hep-ph/9412367]; E. Dudas, S. Pokorski and C. A. Savoy, Phys. Lett. B **356**, 45 (1995) [hep-ph/9504292]; P. Binetruy, S. Lavignac and

- P. Ramond, Nucl. Phys. B **477**, 353 (1996) [hep-ph/9601243]; P. H. Chankowski, K. Kowalska, S. Lavignac and S. Pokorski, Phys. Rev. D **71**, 055004 (2005) [hep-ph/0501071].
- [29] A. E. Nelson and D. Wright, Phys. Rev. D **56**, 1598 (1997) [hep-ph/9702359].
- [30] H. Georgi and C. Jarlskog, Phys. Lett. B **86**, 297 (1979).
- [31] J. R. Ellis and M. K. Gaillard, Phys. Lett. B **88**, 315 (1979).
- [32] G. Altarelli, F. Feruglio and I. Masina, JHEP **0011**, 040 (2000) [hep-ph/0007254].
- [33] J. Schechter and J. W. F. Valle, Phys. Rev. D **25**, 2951 (1982).
- [34] A. Osipowicz *et al.* [KATRIN Collaboration], arXiv:hep-ex/0109033; see also: <http://www-ik.fzk.de/~katrin/index.html>
- [35] C. Arnaboldi *et al.* [CUORE Collaboration], Nucl. Instrum. Meth. A **518**, 775 (2004) [hep-ex/0212053]; R. Ardito, C. Arnaboldi, D. R. Artusa, F. T. Avignone, III, M. Balata, I. Bandac, M. Barucci and J. W. Beeman *et al.*, hep-ex/0501010.
- [36] S. R. Elliott [MAJORANA Collaboration], J. Phys. Conf. Ser. **173**, 012007 (2009) [arXiv:0807.1741 [nucl-ex]]; V. E. Guiseppe *et al.* [Majorana Collaboration], IEEE Nucl. Sci. Symp. Conf. Rec. **2008**, 1793 (2008) [arXiv:0811.2446 [nucl-ex]].
- [37] I. Abt, M. F. Altmann, A. Bakalyarov, I. Barabanov, C. Bauer, E. Bellotti, S. T. Belyaev and L. B. Bezrukov *et al.*, hep-ex/0404039; S. Schonert (GERDA Collaboration), J. Phys. Conf. Ser. **203**, 012014 (2010).
- [38] M. Cannoni, J. Ellis, M. E. Gomez and S. Lola, arXiv:1301.6002 [hep-ph].
- [39] L. Zhan, Y. Wang, J. Cao and L. Wen, Phys. Rev. D **79**, 073007 (2009) [arXiv:0901.2976 [hep-ex]]; X. Qian, D. A. Dwyer, R. D. McKeown, P. Vogel, W. Wang, and C. Zhang, arXiv:1208.1551 [hep-ex].
- [40] P. Binetruy, S. Lavignac, S. T. Petcov and P. Ramond, Nucl. Phys. B **496**, 3 (1997) [hep-ph/9610481]; S. Lola and G. G. Ross, Nucl. Phys. B **553**, 81 (1999) [hep-ph/9902283]; M. Tanimoto, Phys. Lett. B **456**, 220 (1999) [hep-ph/9901210]; M. S. Berger and K. Siyeon, Phys. Rev. D **62**, 033004 (2000) [hep-ph/0003121]; M. S. Berger and K. Siyeon, Phys. Rev. D **64**, 053006 (2001) [hep-ph/0005249]; J. Sato and K. Tobe, Phys. Rev. D **63**, 116010 (2001) [hep-ph/0012333].
- [41] Y. Fukuda *et al.* [Super-Kamiokande Collaboration], Phys. Rev. Lett. **81**, 1562 (1998) [hep-ex/9807003].
- [42] M. Maltoni, T. Schwetz, M. A. Tortola and J. W. F. Valle, New J. Phys. **6**, 122 (2004) [hep-ph/0405172].